

# NNLL resummation for global observables

Pier Francesco Monni CERN  Perturbative accuracy now defined in terms of how many towers of logarithms one sums up

e.g.

LL ~ 100% uncertainty

NLL ~ 20% uncertainty

NNLL ~ 5% uncertainty

. . .

SCET Message approaches Merwig spanching algorithm agaism. Merwig Factorisation theorems

QCD at all

orders!??

$$\Sigma(v) = \int_0^v \frac{1}{\sigma_{\text{Born}}} \frac{d\sigma}{dv'} dv' \sim e^{\alpha_s^n L^{n+1} + \alpha_s^n L^n + \alpha_s^n L^{n-1} + \dots}$$

### Introduction

#### Approach I :

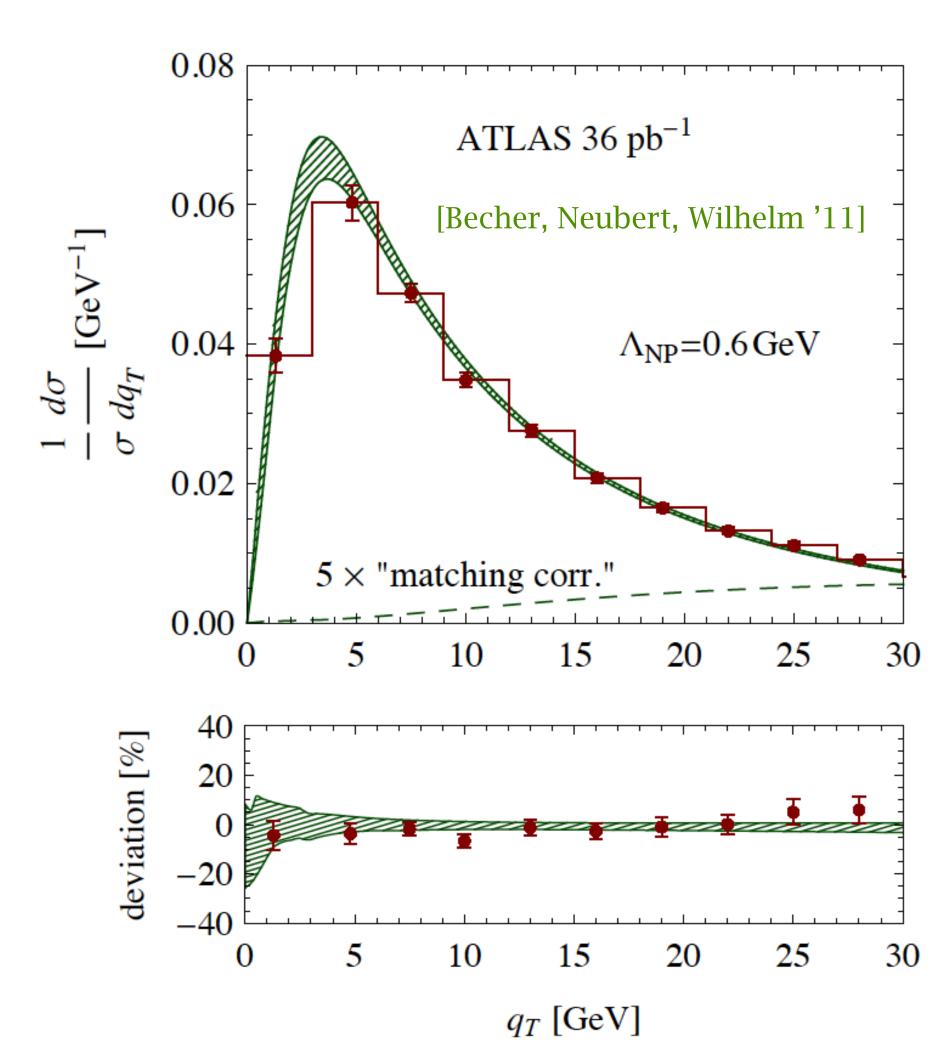
i) Build an effective field theory of QCD by integrating out the irrelevant d.o.f. (i.e. hard radiation)

$$\mathcal{L}_{\text{QCD}} \longrightarrow \mathcal{L}_{\text{SCET}}$$

$$d\sigma \sim \sum_{n} C_{n} \langle \mathcal{O}_{n} \mathcal{O}_{n}^{\dagger} \rangle$$

ii) Evolution Equations arise from RG invariance

$$\frac{d\sigma}{d\ln\mu} = 0$$



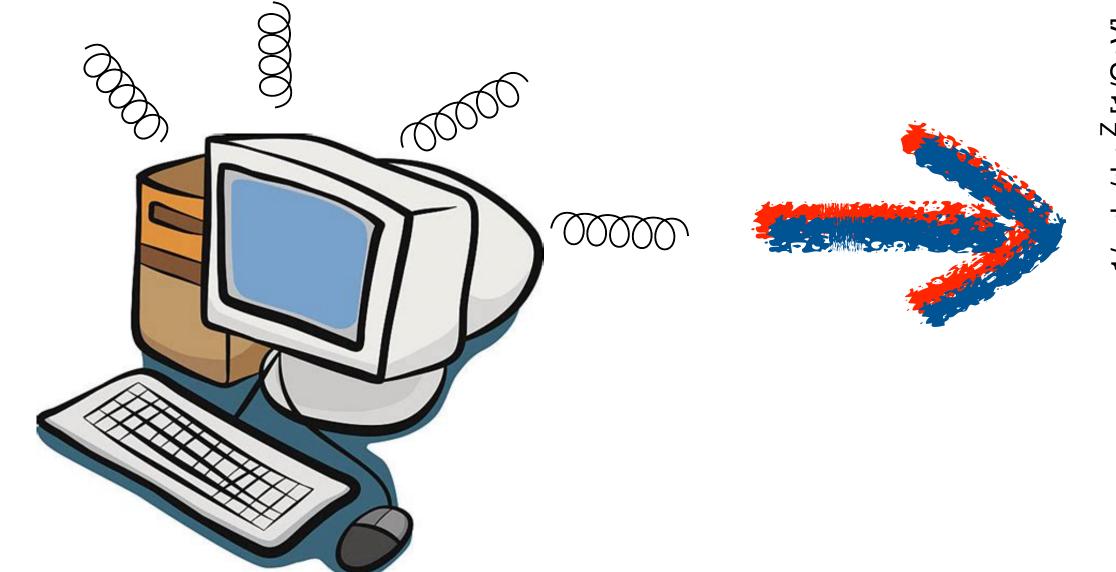
#### Introduction

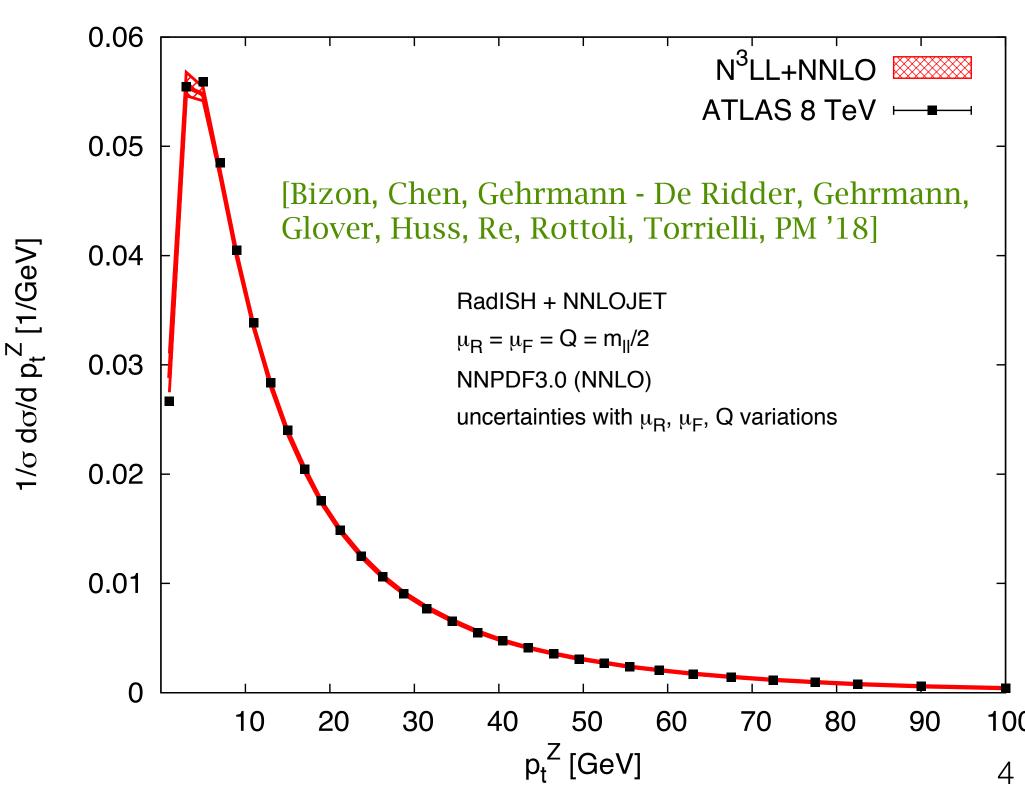
#### Approach II :

i) Build a simplified model of QCD amplitudes in soft and collinear limits, simulate radiation at all orders

$$d\sigma \sim \sum_{n=1}^{\infty} \prod_{i=1}^{n} \int dk_i |\mathcal{M}(k_i)|^2$$

ii) Evolution formulated in algorithmic form

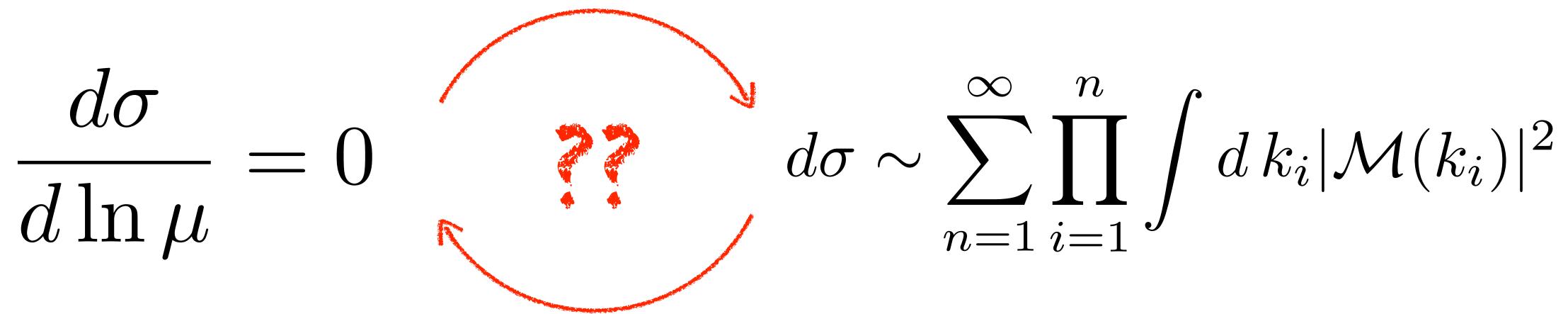




## Exploiting synergies in resummation

Formal link between the two formulations is tricky, some progress recently

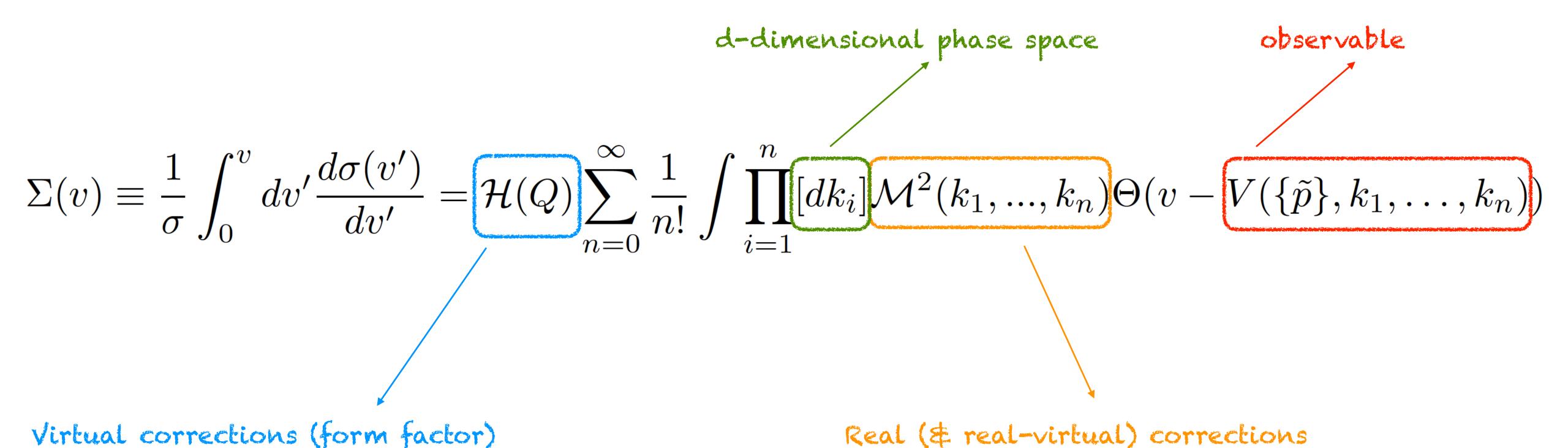
[Bauer, PM '18 + mainly ongoing work]



- A solid connection will shed more light on infrared structure of QCD, and allow us to address some complicated problems from a different angle
  - e.g.
    - multi-leg processes
    - multi-scale problems
    - non-global observables
    - •

# Two scale problems: rIRC safe & global

• e.g. e+e- —> jets



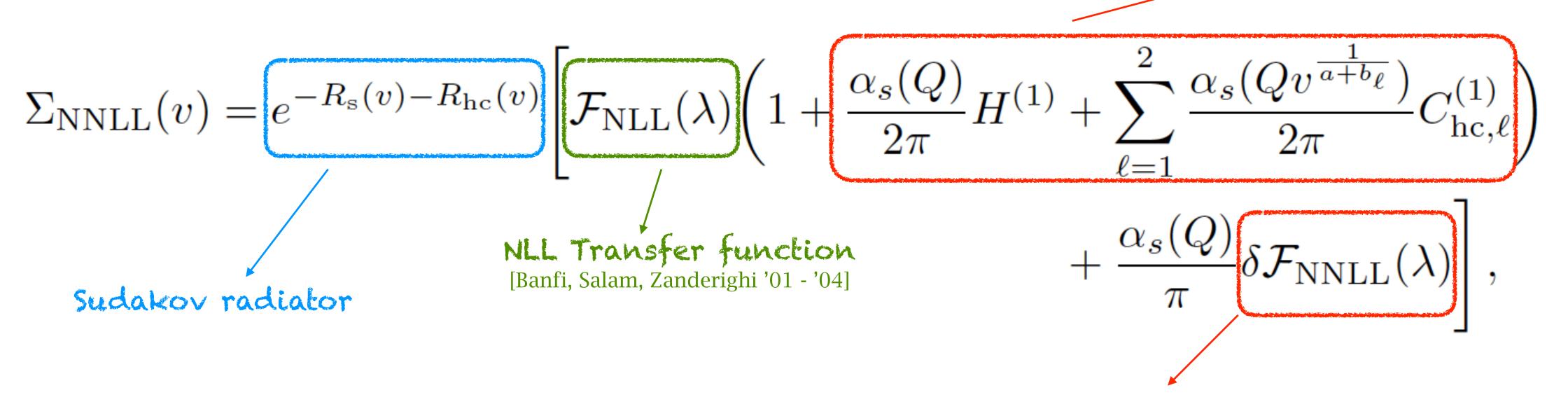
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# Two scale problems: rIRC safe & global

This problem has a general solution at NNLL ( $\lambda = \alpha_s(Q)\beta_0 \ln \frac{1}{v}$ )

[Banfi, El Menoufi, PM 1807.11487] [Banfi, McAslan, Zanderighi, PM 1412.2126, 1607.03111]

NNLL virtual and collinear constants

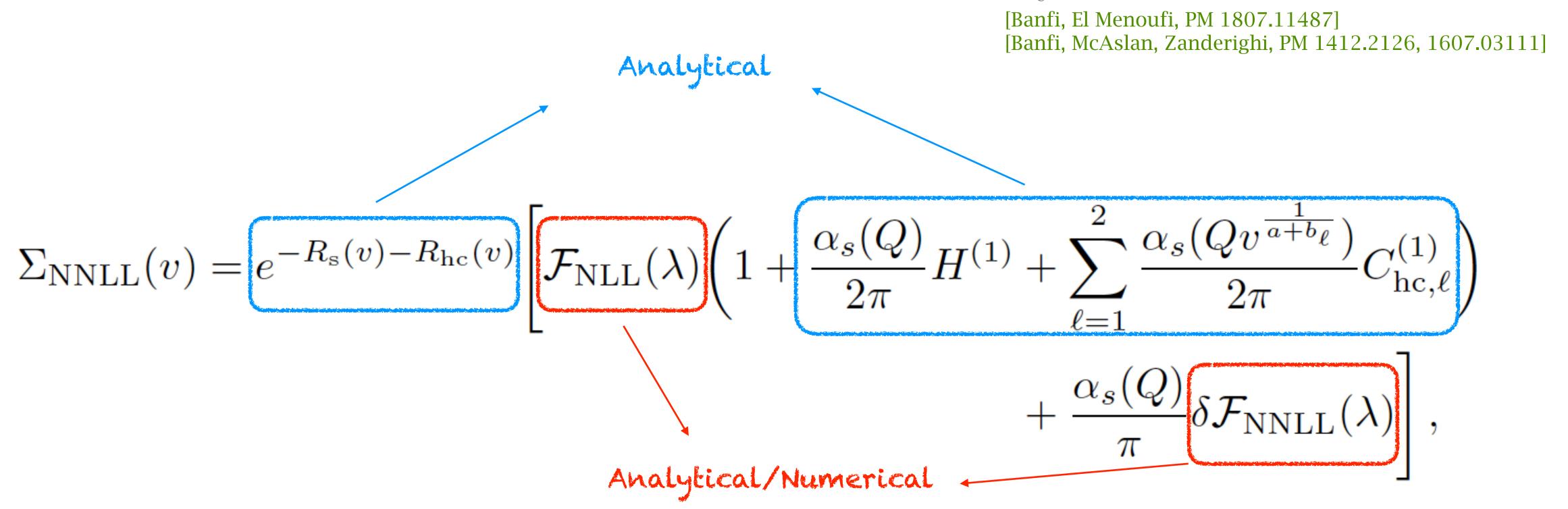


NNLL Transfer function

$$\delta \mathcal{F}_{\mathrm{NNLL}} = \delta \mathcal{F}_{\mathrm{sc}} + \delta \mathcal{F}_{\mathrm{hc}} + \delta \mathcal{F}_{\mathrm{rec}} + \delta \mathcal{F}_{\mathrm{wa}} + \delta \mathcal{F}_{\mathrm{correl}} + \delta \mathcal{F}_{\mathrm{clust}}$$

# Two scale problems: rIRC safe & global

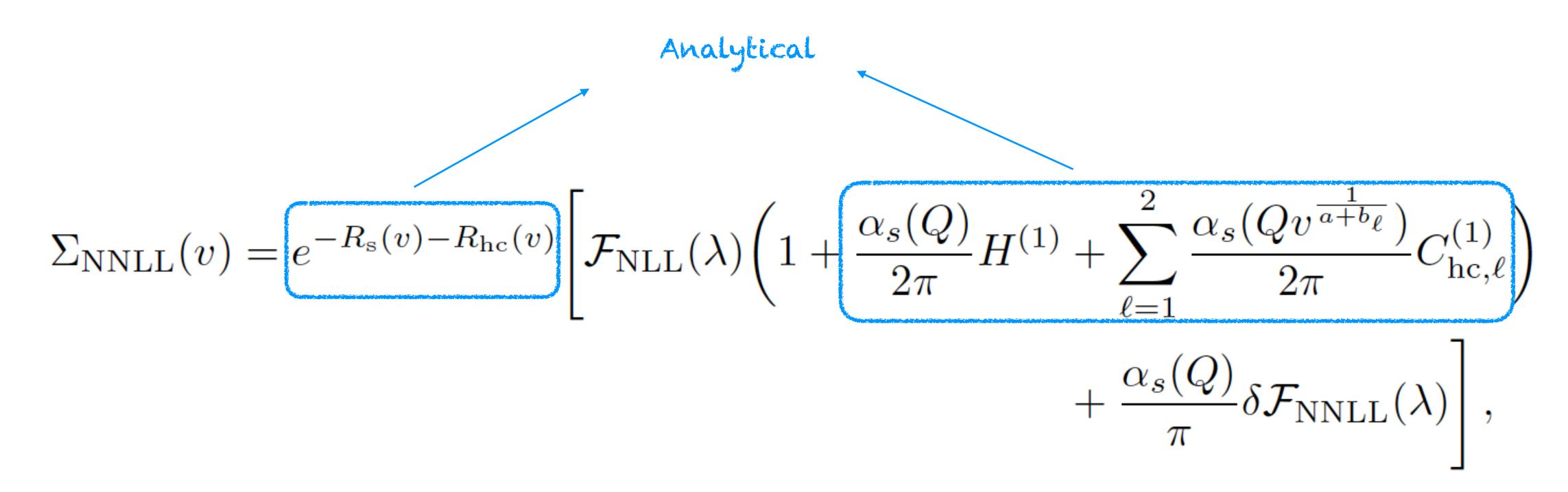
This problem has a general solution at NNLL ( $\lambda = \alpha_s(Q)\beta_0 \ln \frac{1}{v}$ )



$$\delta \mathcal{F}_{\mathrm{NNLL}} = \delta \mathcal{F}_{\mathrm{sc}} + \delta \mathcal{F}_{\mathrm{hc}} + \delta \mathcal{F}_{\mathrm{rec}} + \delta \mathcal{F}_{\mathrm{wa}} + \delta \mathcal{F}_{\mathrm{correl}} + \delta \mathcal{F}_{\mathrm{clust}}$$

## Two scale problems: the Radiator

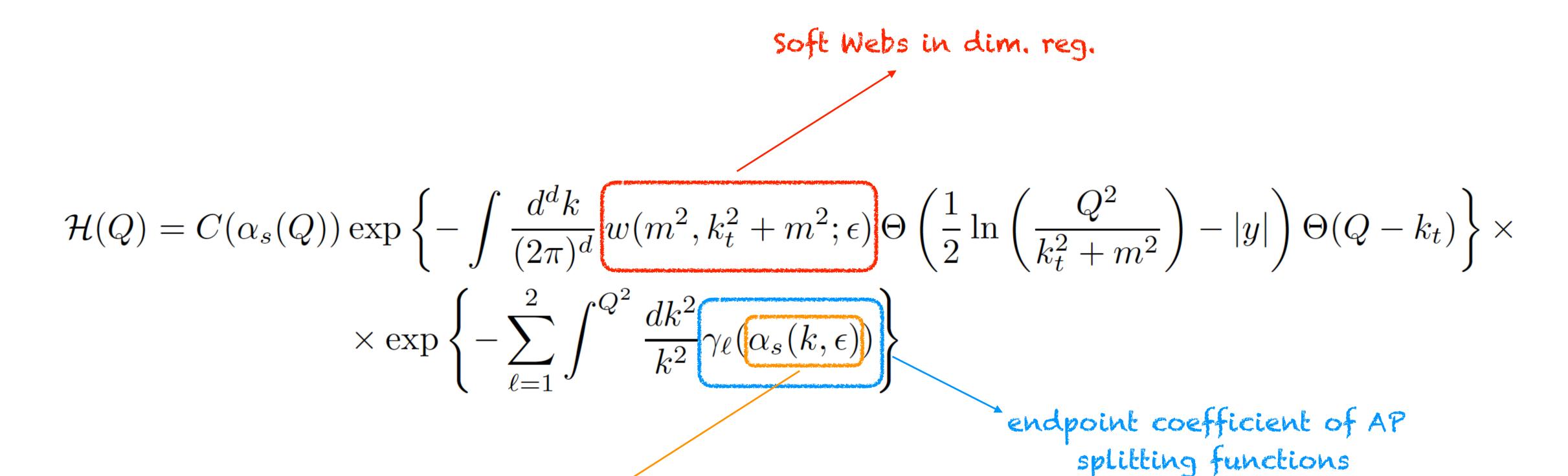
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$$\delta \mathcal{F}_{\text{NNLL}} = \delta \mathcal{F}_{\text{sc}} + \delta \mathcal{F}_{\text{hc}} + \delta \mathcal{F}_{\text{rec}} + \delta \mathcal{F}_{\text{wa}} + \delta \mathcal{F}_{\text{correl}} + \delta \mathcal{F}_{\text{clust}}$$

### Sudakov radiator: virtual corrections

Consider the resummed form factor



strong coupling in dim. reg.

$$\mu_R^2 \frac{d\alpha_s}{d\mu_R^2} = -\epsilon \,\alpha_s + \beta^{(d=4)}(\alpha_s)$$

## Sudakov radiator: real (unresolved) corrections

- Decompose squared amplitude in terms of correlated clusters of emissions
  - e.g. soft limit

$$M_{\rm s}^2(k_1) \equiv \tilde{M}_{\rm s}^2(k_1)$$

two clusters  $(k_1,k_2)$  cluster  $k_{\text{clust.}}=k_1+k_2$ 

Each cluster is dressed by virtual corrections

$$\tilde{M}_{\mathrm{s}}^{2}(k_{1},\ldots,k_{n}) = \tilde{M}_{\mathrm{s},0}^{2}(k_{1},\ldots,k_{n}) + \frac{\alpha_{s}(\mu_{R})}{2\pi}\tilde{M}_{\mathrm{s},1}^{2}(k_{1},\ldots,k_{n}) + \ldots$$

## Sudakov radiator: real (unresolved) corrections

Introduce a resolution scale (slicing parameter) such that unresolved clusters satisfy

$$V_{
m sc}(k_{
m clust.}) < \delta v$$
 
$$V_{
m sc}(k) \equiv \sum_{\ell=1}^{2} d_{\ell} \left(\frac{k_{t}^{(\ell)}}{Q}\right)^{d} e^{-b_{\ell} \eta^{(\ell)}} g_{\ell}(\phi^{(\ell)}) \Theta(\eta^{(\ell)})$$

Unresolved radiation is unconstrained by the observable —> logarithmic counting

	Unresolved ro	diation	Resolved radiation			
	$[nPC^{(j)}]_{sc}$	$[nPC^{(j)}]_{sc}$	$[nPC^{(j)}]_{sc}$	$[nPC^{(j)}]_{sc}$		
LL	$n+j \leq 1$					
NLL	$n+j \leq 2$	$n+j \leq 1$	$n+j \leq 1$			
NNLL	$n+j \leq 3$	$n+j \leq 2$	$n+j \leq 2$	$n+j \leq 1$		
$N^k LL$	$n+j \le k+1$	$n+j \leq k$	$n+j \leq k$	$n+j \le k-1$		

### Sudakov radiator at NNLL

• Combination of virtuals and unresolved clusters defines the radiator (cutoff dependence cancels against that of *resolved* radiation)

Starts at LL (double logs) 
$$2 \\ R_{\mathrm{S}}(v) = \sum_{\ell=1}^{2} \int \frac{d^4k}{(2\pi)^4} w(m^2, k_t^2 + m^2) \Theta\left(d_\ell \left(\frac{k_t}{Q}\right)^a e^{-b_\ell \eta^{(\ell)}} g_\ell(\phi^{(\ell)}) - v\right) \Theta(\eta^{(\ell)})$$

$$R_{\rm hc} = \sum_{\ell=1}^{2} \int_{Q^2 v^{\frac{2}{a+b_\ell}}}^{Q^2} \frac{dk^2}{k^2} \frac{\alpha_s(k)}{2\pi} \left[ \gamma_\ell^{(0)} + \left( \frac{\alpha_s}{2\pi} \right) \gamma_\ell^{(1)} \right]$$

#### Sudakov radiator at NNLL

Soft contribution can be further decomposed as follows

$$R_{\rm s}(v) \simeq \sum_{\ell} \left( R_{\ell}(v) + R'_{\ell}(v) \int_{0}^{2\pi} \frac{d\phi^{(\ell)}}{2\pi} \ln(d_{\ell}g_{\ell}(\phi^{(\ell)})) + R''_{\ell}(v) \int_{0}^{2\pi} \frac{d\phi^{(\ell)}}{2\pi} \frac{1}{2} \ln^{2}(d_{\ell}g_{\ell}(\phi^{(\ell)})) \right)$$

$$R_{\ell}(v) \simeq R_{\ell}^{0}(v) + \delta R_{\ell}(v)$$

massless term

$$R_{\ell}^{0}(v) = \int \frac{d^{4}k}{(2\pi)^{4}} w(m^{2}, k_{t}^{2} + m^{2}) \Theta\left(\ln\frac{Q}{k_{t}} - \eta^{(\ell)}\right) \Theta\left(\left(\frac{k_{t}}{Q}\right)^{a} e^{-b_{\ell}\eta^{(\ell)}} - v\right) \Theta(\eta^{(\ell)})$$

#### Sudakov radiator at NNLL

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$$R_{\ell}(v) \simeq R_{\ell}^{0}(v) + \delta R_{\ell}(v)$$

mass correction

$$\delta R_{\ell}(v) = C_{\ell} \int_{Qv^{\frac{1}{a+b_{\ell}}}}^{Q} \frac{dk_{t}}{k_{t}} \left(\frac{\alpha_{s}(k_{t})}{\pi}\right)^{2} \int_{0}^{\infty} \frac{d\mu^{2}}{\mu^{2}(1+\mu)} \left(C_{A} \ln \frac{1+\mu^{2}}{\mu^{4}} - 2\pi\beta_{0}\right) \ln \left(\sqrt{\frac{1}{1+\mu^{2}}}\right) \\
= \pi \beta_{0} \zeta_{2} C_{\ell} \int_{Qv^{\frac{1}{a+b_{\ell}}}}^{Q} \frac{dk_{t}}{k_{t}} \left(\frac{\alpha_{s}(k_{t})}{\pi}\right)^{2}.$$

## Physical coupling in the soft limit

The massless terms defines a physical coupling in the soft limit

$$R_{\ell}^{0}(v) = \int \frac{d^{4}k}{(2\pi)^{4}} w(m^{2}, k_{t}^{2} + m^{2}) \Theta\left(\ln\frac{Q}{k_{t}} - \eta^{(\ell)}\right) \Theta\left(\left(\frac{k_{t}}{Q}\right)^{a} e^{-b_{\ell}\eta^{(\ell)}} - v\right) \Theta(\eta^{(\ell)})$$

$$\int_0^\infty dm^2 w(m^2, k_t^2 + m^2) \equiv (4\pi)^2 \frac{2C_\ell}{k_t^2} \alpha_s^{\text{phys}}(k_t) \qquad \qquad \alpha_s^{\text{phys}} = \alpha_s \left(1 + \sum_{n=1}^\infty \left(\frac{\alpha_s}{2\pi}\right)^n K^{(n)}\right)$$

$$K^{(1)}=C_A\left(rac{67}{18}-rac{\pi^2}{6}
ight)-rac{5}{9}n_f$$
 , -> [Catani, Marchesini, Webber '91]

This is the only, universal, as³ ingredient at NNLL!

$$K^{(2)} = C_A^2 \left( \frac{245}{24} - \frac{67}{9} \zeta_2 + \frac{11}{6} \zeta_3 + \frac{11}{5} \zeta_2^2 \right) + C_F n_f \left( -\frac{55}{24} + 2\zeta_3 \right)$$

$$+ C_A n_f \left( -\frac{209}{108} + \frac{10}{9} \zeta_2 - \frac{7}{3} \zeta_3 \right) - \frac{1}{27} n_f^2 + \frac{\pi \beta_0}{2} \left( C_A \left( \frac{808}{27} - 28\zeta_3 \right) - \frac{224}{54} n_f \right)$$

## Two scale problems: the transfer functions

- The definition of the *unresolved* radiation ensures that all LL are in the Radiator
- One's left with the computation of the transfer functions

$$\Sigma_{\mathrm{NNLL}}(v) = e^{-R_{\mathrm{s}}(v) - R_{\mathrm{hc}}(v)} \left[ \mathcal{F}_{\mathrm{NLL}}(\lambda) \left( 1 + \frac{\alpha_{s}(Q)}{2\pi} H^{(1)} + \sum_{\ell=1}^{2} \frac{\alpha_{s}(Qv^{\frac{1}{a+b_{\ell}}})}{2\pi} C_{\mathrm{hc},\ell}^{(1)} \right) + \frac{\alpha_{s}(Q)}{\pi} \delta \mathcal{F}_{\mathrm{NNLL}}(\lambda) \right],$$
Analytical/Numerical

$$\delta \mathcal{F}_{\text{NNLL}} = \delta \mathcal{F}_{\text{sc}} + \delta \mathcal{F}_{\text{hc}} + \delta \mathcal{F}_{\text{rec}} + \delta \mathcal{F}_{\text{wa}} + \delta \mathcal{F}_{\text{correl}} + \delta \mathcal{F}_{\text{clust}}$$

## Two scale problems: the transfer functions

- Transfer functions describe radiation in well specified kinematical regimes
  - e.g. NLL (CAESAR): ensemble of soft-collinear gluons strongly ordered in rapidity

$$\mathcal{F}_{\mathrm{NLL}}(\lambda) = \int d\mathcal{Z}[\{R'_{\mathrm{NLL},\ell_i}, k_i\}] \Theta\left(1 - \lim_{v \to 0} \frac{V_{\mathrm{sc}}(\{\tilde{p}\}, \{k_i\})}{v}\right)$$

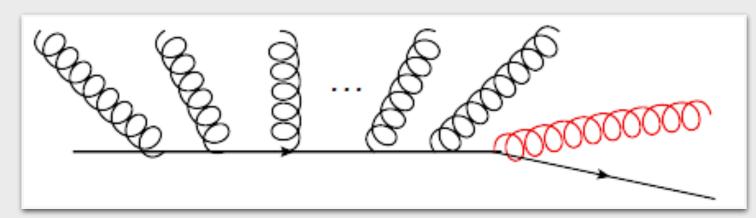
$$\int d\mathcal{Z}[\{R'_{\text{NLL},\ell_i}, k_i\}]G(\{\tilde{p}\}, \{k_i\}) = \epsilon^{R'_{\text{NLL}}} \sum_{n=0}^{\infty} \frac{1}{n!} \prod_{i=1}^{n} \int_{\epsilon}^{\infty} \frac{d\zeta_i}{\zeta_i} \int_{0}^{2\pi} \frac{d\phi_i}{2\pi} \sum_{\ell_i = 1, 2} R'_{\text{NLL},\ell_i}G(\{\tilde{p}\}, k_1, \dots, k_n)$$

## Two scale problems: the transfer functions

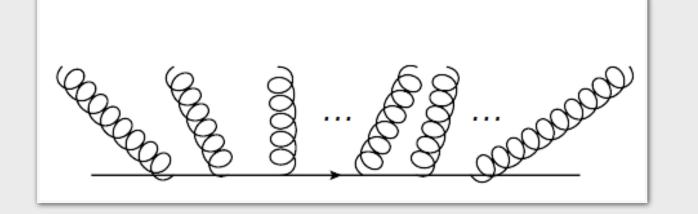
- Transfer functions describe radiation in well specified kinematical regimes
  - e.g. NNLL (ARES): at most one single emission probes less singular kinematics

$$\delta \mathcal{F}_{\text{NNLL}} = \delta \mathcal{F}_{\text{sc}} + \delta \mathcal{F}_{\text{hc}} + \delta \mathcal{F}_{\text{rec}} + \delta \mathcal{F}_{\text{wa}} + \delta \mathcal{F}_{\text{correl}} + \delta \mathcal{F}_{\text{clust}}$$

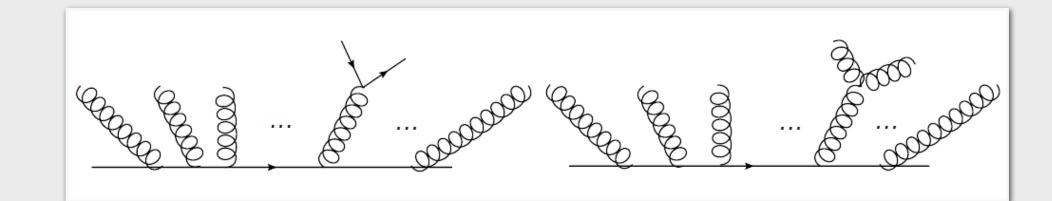
- collinear emission carries a significant energy fraction
  - correction to the amplitude: hard-collinear corrections
  - correction to the observable: recoil corrections



- soft-collinear emission gets close in rapidity to another
  - sensitive to the exact rapidity bounds: rapidity (SC) corrections
  - different clustering history for a jet algorithm: clustering corrections



- insertion of double-soft current and corresponding virtual corrections
  - correlated corrections

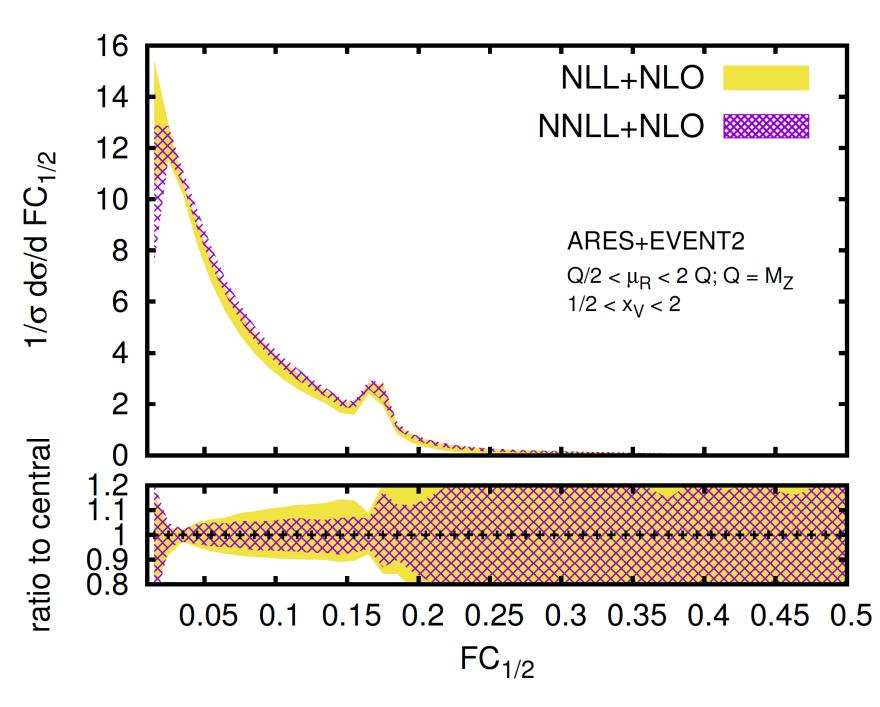


- soft emission is allowed to propagate at small rapidities
- soft-wide-angle corrections
- Restores correct rapidity dependence in observable
- Gets complicated for multi leg case (interference between hard emitters)

### Automation

- Suitable for automation in a computer program
  - observable (in various limits) is the only external input
- Agnostic to factorisation structure of the measurement function

correction type	$p_{ m t,veto}$	1-T	$B_{T}$	$B_{W}$	C	$ ho_H$	$T_{M}$	O	$y_3^{\text{Dur.}}$	$y_3^{\text{Cam}}$ .	$p_t$
$\mathcal{F}_{\mathrm{NLL}}$	<b>√</b>	✓	<b>√</b>	<b>√</b>	<b>√</b>						
$\delta \mathcal{F}_{ ext{SC}}$	X	✓	✓	✓	✓	✓	✓	✓	✓	X	✓
$\delta\mathcal{F}_{\mathbf{wa}}$	X	X	X	X	✓	X	X	X	✓	✓	X
$\delta \mathcal{F}_{ m hc}$	X	✓	✓	✓	✓	✓	✓	✓	✓	X	✓
$\delta \mathcal{F}_{ m rec}$	X	✓	✓	✓	✓	✓	✓	✓	✓	X	X
$\delta \mathcal{F}_{ ext{clust}}$	✓	X	X	X	X	X	X	X	✓	✓	X
$\delta \mathcal{F}_{ ext{correl}}$	✓	X	<b>√</b>	✓	X	X	✓	✓	✓	✓	X



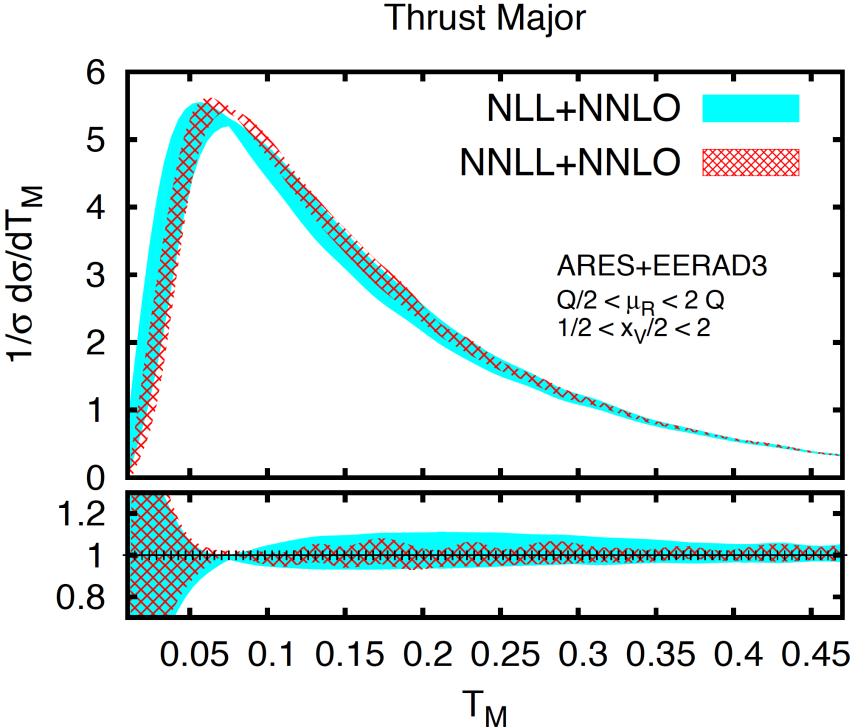
## Some examples

#### Moments of EEC & angularities

$$FC_x = \sum_{i \neq j} \frac{E_i E_j |\sin \theta_{ij}|^x (1 - |\cos \theta_{ij}|)^{1-x}}{(\sum_i E_i)^2} \Theta\left[ (\vec{q}_i \cdot n_T)(\vec{q}_j \cdot \vec{n}_T) \right]$$

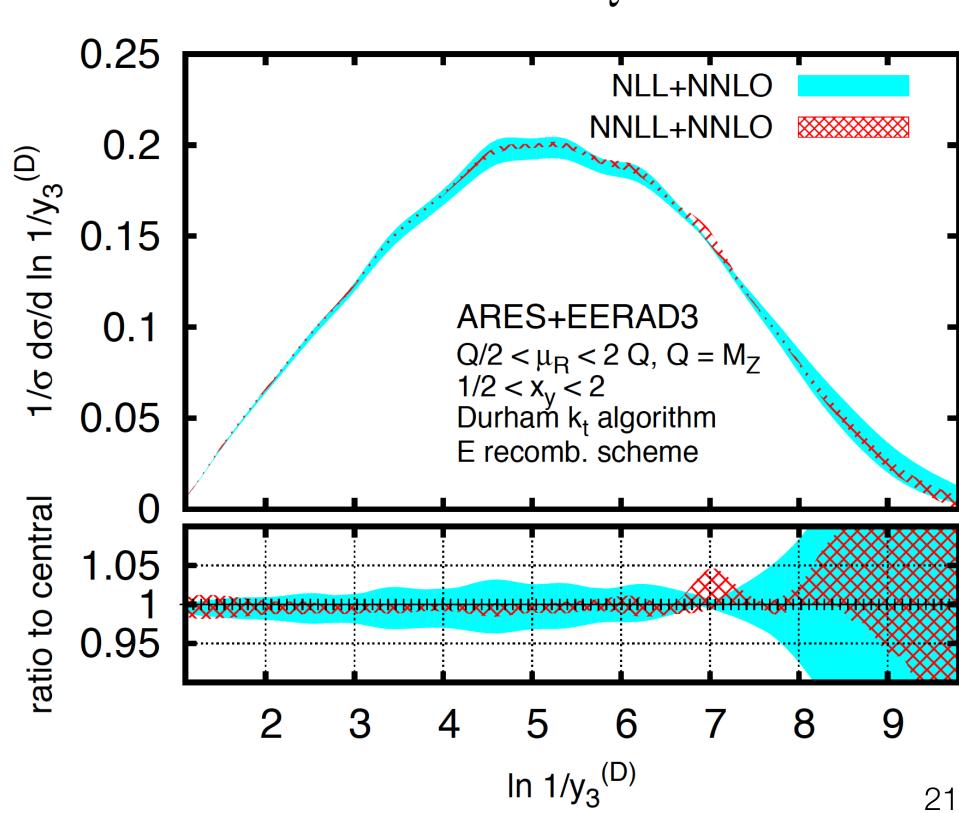
#### Durham ke algorithm

$$y_{ij}^{(D)} = v_{ij}^{(D)} = 2 \frac{\min\{E_i, E_j\}^2}{Q^2} (1 - \cos \theta_{ij})$$



#### Thrust Major

$$T_M = \frac{1}{\sum_{i} |\vec{p}_i|} \max_{\vec{n} \cdot \vec{n}_T = 0} \sum_{i} |\vec{p}_i \cdot \vec{n}|$$



#### Conclusions

- Global problems with two scales have a general solution at NNLL
  - no need for a factorisation theorem
  - suitable for automation
  - can be systematically extended to higher orders if higher precision is needed (e.g. p<sub>T</sub>)
- For more complicated observables, more efforts are required to understand the general structure (known at NLL)
  - Multi differential distributions
  - Only a few results available for multi-leg observables at NNLL
  - Full control of IR physics at NNLL requires solution of (next to leading) non-global logarithms
- Matching ambiguity to fixed order ~few-%: pheno impact of subleading power corrections to be established, might be relevant in high precision observables

# Backup material

# Sudakov radiator: web exponentiation

Soft unresolved clusters obey non-abelian (web) exponentiation theorem

$$\mathcal{H}(Q) \exp\left\{\int^Q \frac{d^dk}{(2\pi)^d} w(m^2, k_t^2 + m^2; \epsilon) \Theta(\delta v - V_{\rm sc}(k))\right\}$$
 Soft radiator 
$$= C(\alpha_s(Q)) e^{-R_{\rm s}(\delta v)} \exp\left\{-\sum_{\ell=1}^2 \int^{Q^2} \frac{dk^2}{k^2} \gamma_\ell(\alpha_s(k, \epsilon))\right\}$$
 the form factor

Cut virtual-collinear at collinear scale, and expand to cancel real divergences

$$\int^{Q^{2}} \frac{dk^{2}}{k^{2}} \gamma_{\ell}(\alpha_{s}(k,\epsilon)) = \int_{Q^{2}v^{\frac{2}{a+b_{\ell}}}}^{Q^{2}} \frac{dk^{2}}{k^{2}} \gamma_{\ell}(\alpha_{s}(k)) + \int_{0}^{Q^{2}v^{\frac{2}{a+b_{\ell}}}} \frac{dk^{2}}{k^{2}} \gamma_{\ell}(\alpha_{s}(k,\epsilon))$$

$$= \exp\left\{-\int_{0}^{Q^{2}v^{\frac{2}{a+b_{\ell}}}} \frac{dk^{2}}{k^{2}} \gamma_{\ell}(\alpha_{s}(k,\epsilon))\right\} = 1 - \int_{0}^{Q^{2}v^{\frac{2}{a+b_{\ell}}}} \frac{dk^{2}}{k^{2}} \frac{\alpha_{s}(k,\epsilon)}{2\pi} \gamma_{\ell}^{(0)} + \mathcal{O}(\alpha_{s}^{2}(Qv^{\frac{1}{a+b_{\ell}}}))$$

### Collinear corrections

- Extension to NNLL involves additional kinematic configurations:
  - (at most) one collinear emission can carry a significant fraction of the energy of the hard emitter (which recoils against it)



$$\delta \mathcal{F}_{hc}(\lambda) = \sum_{\ell=1,2} \frac{\alpha_s(v^{1/(a+b_\ell)}Q)}{\alpha_s(Q)(a+b_\ell)} \int_0^\infty \frac{d\zeta}{\zeta} \int_0^{2\pi} \frac{d\phi}{2\pi} \int d\mathcal{Z}[\{R'_{\text{NLL},\ell_i}, k_i\}] \times$$

$$\times \int_0^1 \frac{dz}{z} \left(zp_\ell(z) - 2C_\ell\right) \left[\Theta\left(1 - \lim_{v \to 0} \frac{V_{\text{sc}}(\{\tilde{p}\}, k, \{k_i\})}{v}\right) - \Theta\left(1 - \lim_{v \to 0} \frac{V_{\text{sc}}(\{\tilde{p}\}, \{k_i\})}{v}\right) \Theta(1 - \zeta)\right]$$

$$\delta \mathcal{F}_{rec}(\lambda) = \sum_{\ell=1,2} \frac{\alpha_s(v^{1/(a+b_\ell)}Q)}{\alpha_s(Q)(a+b_\ell)} \int_0^\infty \frac{d\zeta}{\zeta} \int_0^{2\pi} \frac{d\phi}{2\pi} \int d\mathcal{Z}[\{R'_{\text{NLL},\ell_i}, k_i\}] \times$$

$$\times \int_0^1 dz \, p_\ell(z) \left[\Theta\left(1 - \lim_{v \to 0} \frac{V_{\text{hc}}^{(k')}(\{\tilde{p}\}, k', \{k_i\})}{v}\right) - \Theta\left(1 - \lim_{v \to 0} \frac{V_{\text{sc}}(\{\tilde{p}\}, k, \{k_i\})}{v}\right)\right]$$

## Single-soft corrections

• (at most) one soft-collinear emission has the correct rapidity bounds (approximated in the NLL ensemble) - rapidity corrections

$$\delta \mathcal{F}_{\text{sc}}(\lambda) = \frac{\pi}{\alpha_s(Q)} \int_0^\infty \frac{d\zeta}{\zeta} \int_0^{2\pi} \frac{d\phi}{2\pi} \sum_{\ell=1,2} \left( \delta R'_{\text{NNLL},\ell} + R''_{\ell} \ln \frac{d_{\ell} g_{\ell}(\phi)}{\zeta} \right) \int d\mathcal{Z}[\{R'_{\text{NLL},\ell_i}, k_i\}] \times \left[ \Theta\left( 1 - \lim_{v \to 0} \frac{V_{\text{sc}}(\{\tilde{p}\}, k, \{k_i\})}{v} \right) - \Theta(1 - \zeta) \Theta\left( 1 - \lim_{v \to 0} \frac{V_{\text{sc}}(\{\tilde{p}\}, \{k_i\})}{v} \right) \right],$$

• (at most) one soft emission can propagate at very small rapidities (wide angle corrections)\*\*

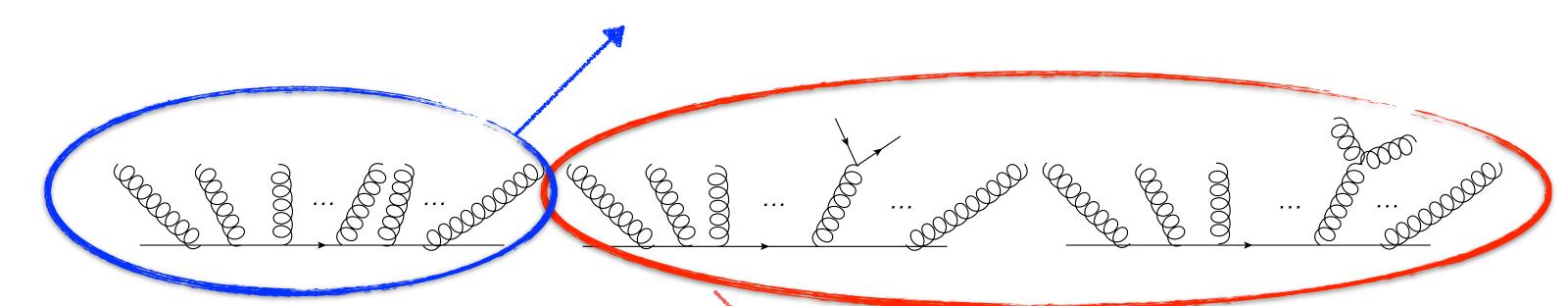
$$\delta \mathcal{F}_{wa}(\lambda) = \frac{2C_F}{a} \frac{\alpha_s(v^{1/a}Q)}{\alpha_s(Q)} \int_0^\infty \frac{d\zeta}{\zeta} \int_{-\infty}^\infty d\eta \int_0^{2\pi} \frac{d\phi}{2\pi} \int d\mathcal{Z}[\{R'_{NLL,\ell_i}, k_i\}]$$

$$\times \left[ \Theta\left(1 - \lim_{v \to 0} \frac{V_{wa}^{(k)}(\{\tilde{p}\}, k, \{k_i\})}{v}\right) - \Theta\left(1 - \lim_{v \to 0} \frac{V_{sc}(\{\tilde{p}\}, k, \{k_i\})}{v}\right) \right]$$

\*\*With n>2 there are additional contributions due to the quantum interference between hard emitters starting at NLL order, due to soft-wide-angle radiation

### Double-soft corrections

- (at most) two soft-collinear emissions get close in rapidity:
  - Relax strong angular ordering (clustering corrections, e.g. jet algorithms)



Treat (at most) one correlated branching exactly (correlated corrections)

$$\delta \mathcal{F}_{\text{correl}}(\lambda) = \int_{0}^{\infty} \frac{d\zeta_{a}}{\zeta_{a}} \int_{0}^{2\pi} \frac{d\phi_{a}}{2\pi} \sum_{\ell_{a}=1,2} \left( \frac{2C_{\ell_{a}}\lambda}{a\beta_{0}} \frac{R_{\ell_{a}}^{"}(v)}{\alpha_{s}(Q)} \right) \int_{0}^{\infty} \frac{d\kappa}{\kappa} \int_{-\infty}^{\infty} d\eta \int_{0}^{2\pi} \frac{d\phi}{2\pi} \frac{1}{2!} C_{ab}(\kappa, \eta, \phi) \times \int d\mathcal{Z}[\{R_{\text{NLL},\ell_{i}}^{"}, k_{i}\}] \left[ \Theta\left(v - V_{\text{sc}}(\{\tilde{p}\}, k_{a}, k_{b}, \{k_{i}\})\right) - \Theta\left(v - V_{\text{sc}}(\{\tilde{p}\}, k_{a} + k_{b}, \{k_{i}\})\right) \right]$$

$$C_{ab}(\kappa, \eta, \phi) = \frac{\tilde{M}^{2}(k_{a}, k_{b})}{M_{\text{sc}}^{2}(k_{a}) M_{\text{sc}}^{2}(k_{b})}$$