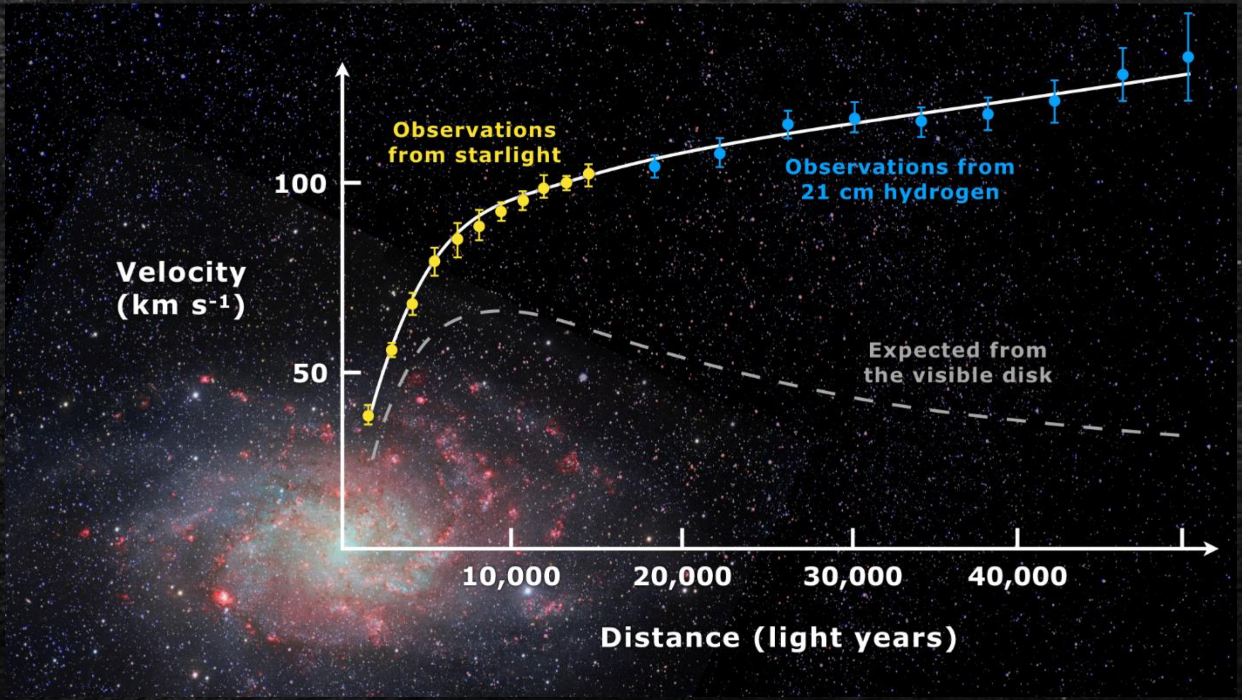
Higgsino Dark Matter in Electron Electric Dipole Moments

Pheno 2021, 5/25/2021

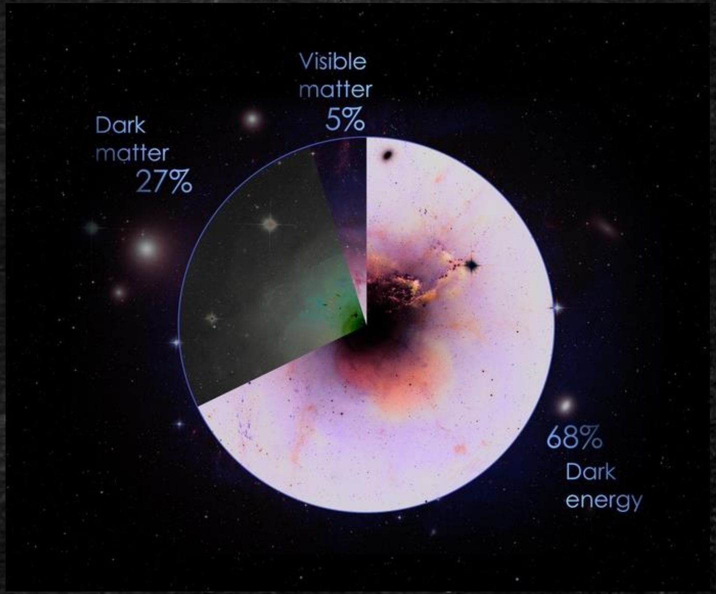
Ben Sheff – University of Michigan

Based on work in collaboration with
Raymond Co – University of Minnesota, James Wells – University of Michigan

Why Dark Matter



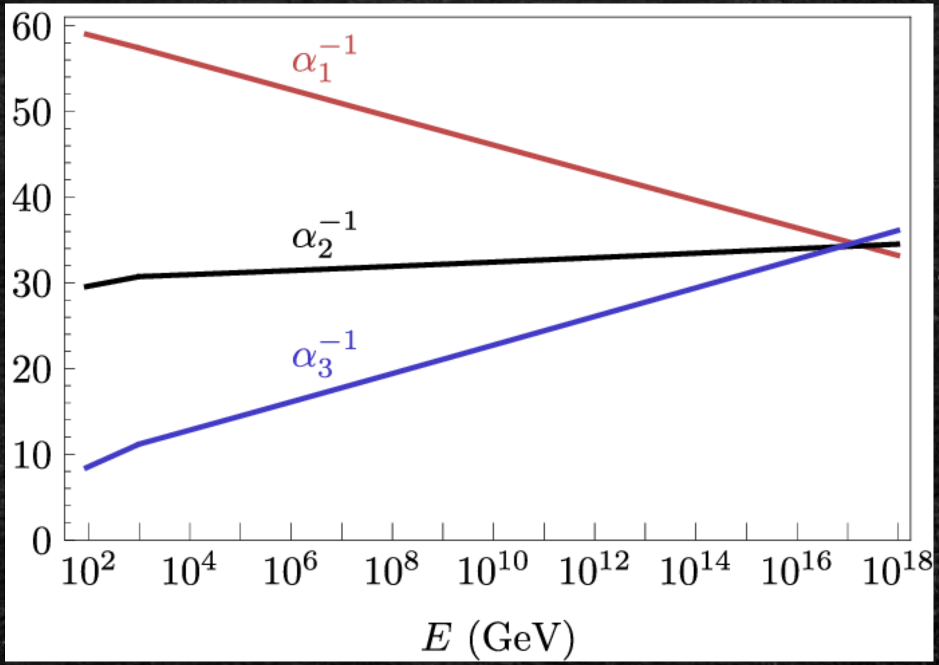
By Mario De Leo - Own work, CC BY-SA 4.0, https://commons.wikimedia.org/w/index.php?curid=74398525



credit to NASA's Goddard Space Flight Center https://svs.gsfc.nasa.gov/12307

Why *Higgsino* Dark Matter

- Most of the advantages of SUSY, with very few parameters
  - Gauge coupling unification
  - Electroweak scale stabilization
- Unify scalar masses and give SUSY breaking gaugino masses by anomaly to reduce to 3 degrees of freedom
  - M_2, μ, $m_{3/2}$



Illustrative example

The Model, Split SUSY

$$-\mathcal{L}_{\text{eff}} = \frac{M_2}{2}\tilde{W}^a\tilde{W}^a + \frac{M_1}{2}\tilde{B}\tilde{B} + \mu\tilde{H}_u\epsilon\tilde{H}_d + \frac{H^\dagger}{\sqrt{2}}\left(\tilde{g}_u\sigma^a\tilde{W}^a + \tilde{g}'_u\tilde{B}\right)\tilde{H}_u + \frac{H^T\epsilon}{\sqrt{2}}\left(\tilde{g}_d\sigma^a\tilde{W}^a + \tilde{g}'_d\tilde{B}\right)\tilde{H}_d + h.c.$$

- Decouple scalars
  - Masses set to $m_{3/2}$=O(PeV)

Wells hep-ph/0411041

- Anomaly generated gaugino masses
  - $M_3 \approx 10 M_2 \approx 3 M_1$
  - $m_{3/2} \sim 300\ M_2$

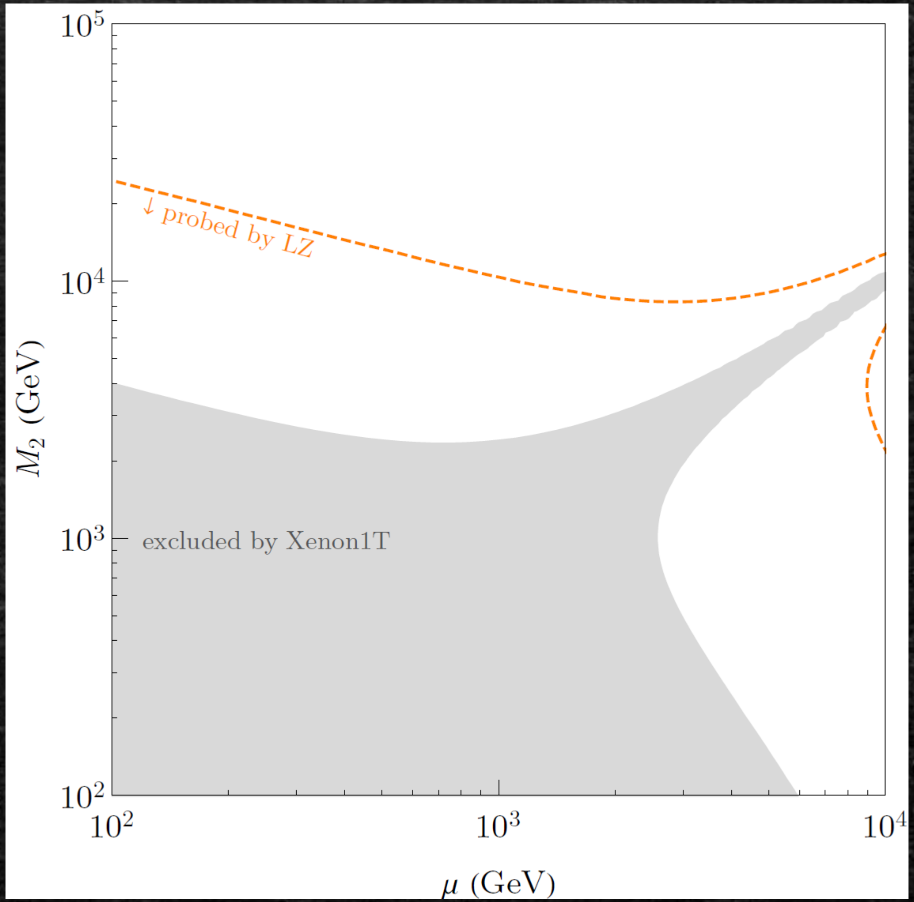
Randall, Sundrum hep-th/9810155

- Higgsino DM as WIMP DM
  - Standard freeze-out WIMP scenario
  - Mass set to $m_{DM} \sim \mu \approx 1.2$ TeV

Profuno, Yaguna hep-ph/0407036

Limited Accessibility to Usual Approaches

- Colliders are very limited for heavy, non-colored particles
- Direct detection cross section falls rapidly as the higgsino-gaugino mixing angle
- Indirect detection has limited reach on higgsino mass
  - CMB measurements limited to O(100 GeV)
  - CTA can reach near μ = 1 TeV



Electron Electric Dipole Moments

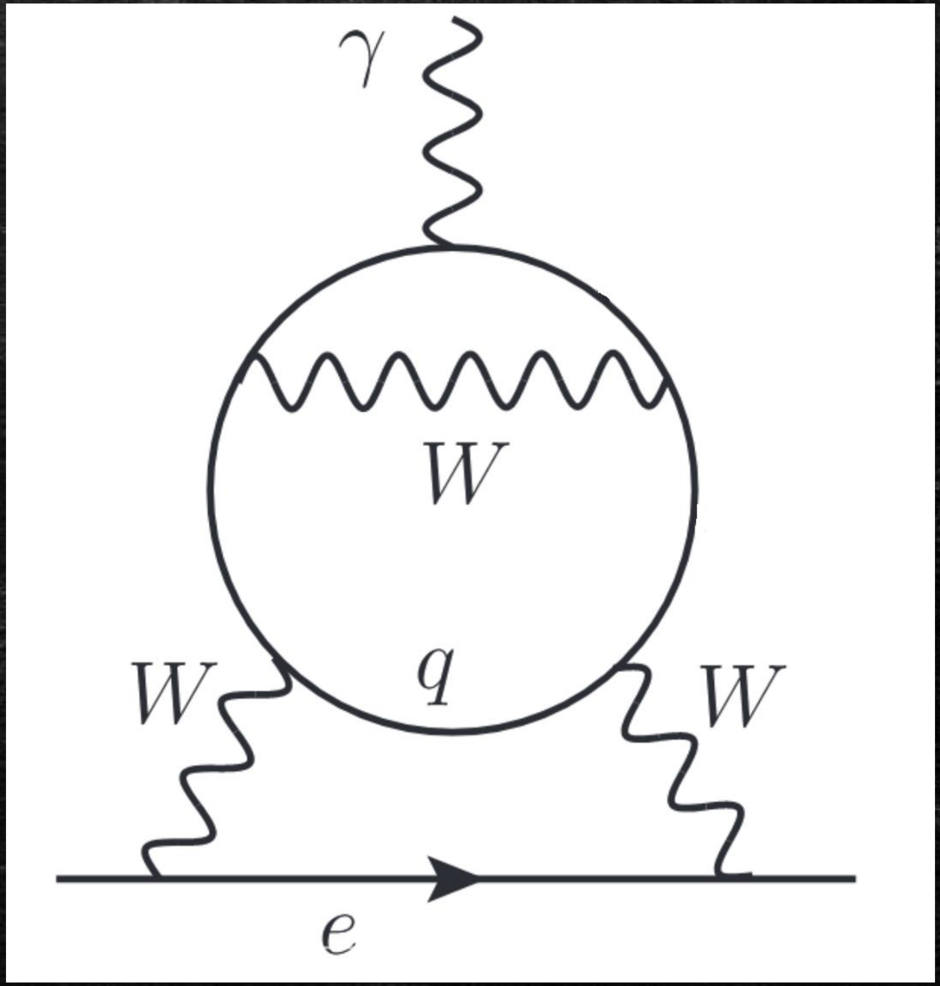
- SUSY is well understood
  - Generically has large complex phases
  - Lead to charge parity violations
- Very little background to worry about
  - Electron EDM can come from SUSY charge parity breaking
  - In SM it's at most ~10^{-38} e cm
  - Current limit is at 1.1×10^{-29} e cm

Scale reference for EDM:

Water: 3.9×10^{-9} e cm
Naïve neutron: 4×10^{-14} e cm
Neutron limit: 10^{-26} e cm

What's going on in the SM?

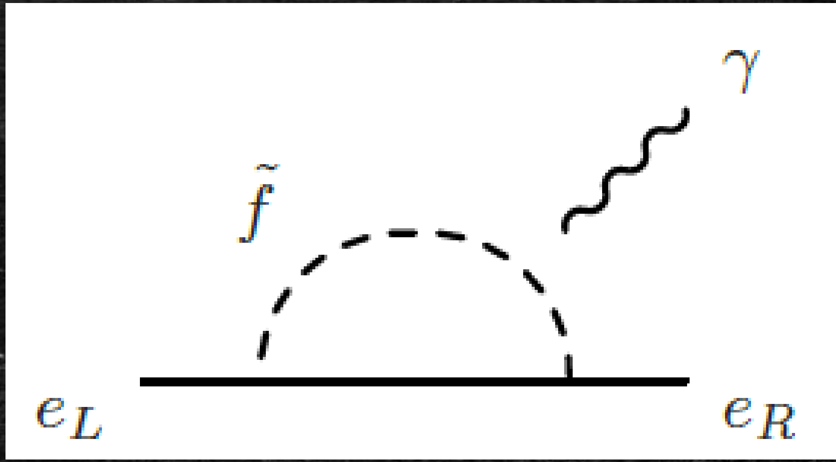
- Need to involve CKM matrix
  - Need all three doublets involved
- 3 loop diagrams are all that work



What about SUSY?

- CP phase can be large
- Limit can be (for O(1) phases)

$M_{SUSY} > 10$ TeV   Cesarotti, et al. 1810.07736



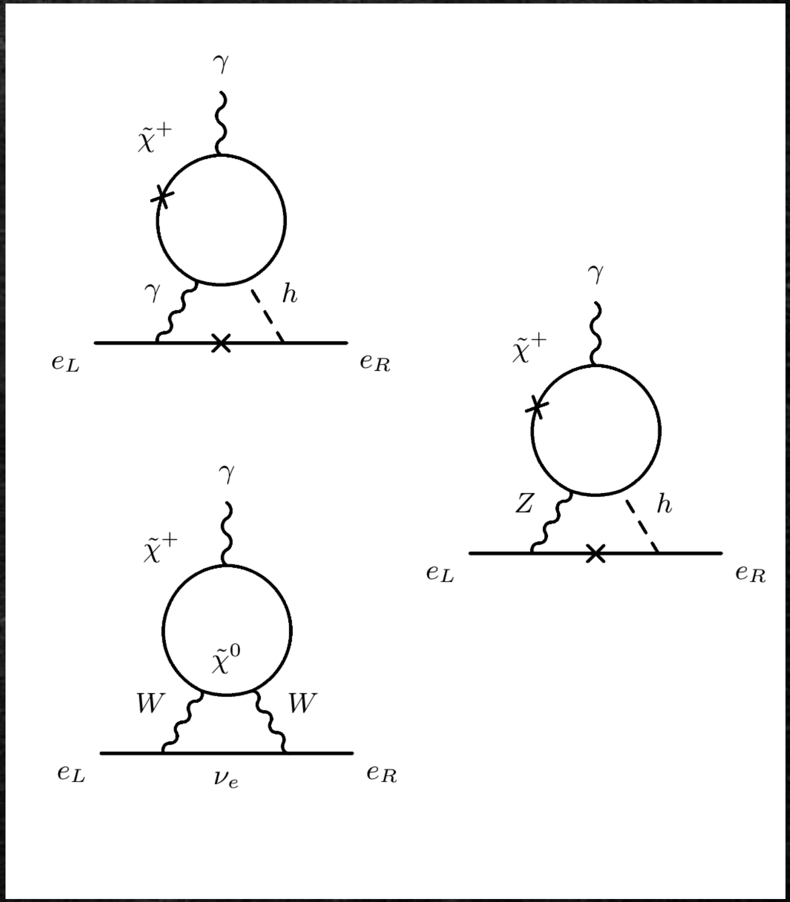
Electron EDM in Split SUSY

- Heavy scalars suppresses previous slide loops
- No 1 loop EDM
  - Non-trivial to find CP phase that can't be absorbed
  - Move to two loop

Barr-Zee Diagram

- Leading order diagrams for EDM

$M_2, \mu >> m_Z$ gives:



$$d_{\gamma h} \simeq \frac{-e\alpha m_e}{8\pi^3} \frac{\tilde{g}_u \tilde{g}_d}{M_2 \mu} \sin\phi_2 F_{\gamma h}\left(\frac{M_2^2}{\mu^2}, \frac{M_2\mu}{m_h^2}\right)$$

$$d_{Zh} \simeq \frac{e\left(4\sin^2\theta_W - 1\right)\alpha m_e}{32\pi^3 \cos^2\theta_W} \frac{\tilde{g}_u \tilde{g}_d}{M_2 \mu} \sin\phi_2 F_{Zh}\left(\frac{m_Z^2}{m_h^2}, \frac{M_2^2}{\mu^2}, \frac{M_2\mu}{m_h^2}\right)$$

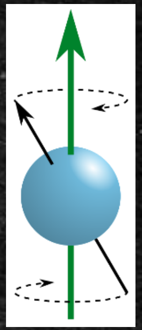
$$d_{WW} \simeq \frac{-e\alpha m_e}{32\pi^3 \sin^2\theta_W}\left(\frac{\tilde{g}_u \tilde{g}_d}{M_2 \mu} \sin\phi_2 F_{WW}^{(2)}\left(\frac{M_2^2}{\mu^2}, \frac{M_2\mu}{m_h^2}\right) + \frac{\tilde{g}'_u \tilde{g}'_d}{M_1 \mu} \sin\phi_1 F_{WW}^{(1)}\left(\frac{M_1^2}{\mu^2}, \frac{M_1\mu}{m_h^2}\right)\right)$$

Giudice, Romanino hep-ph/0510197

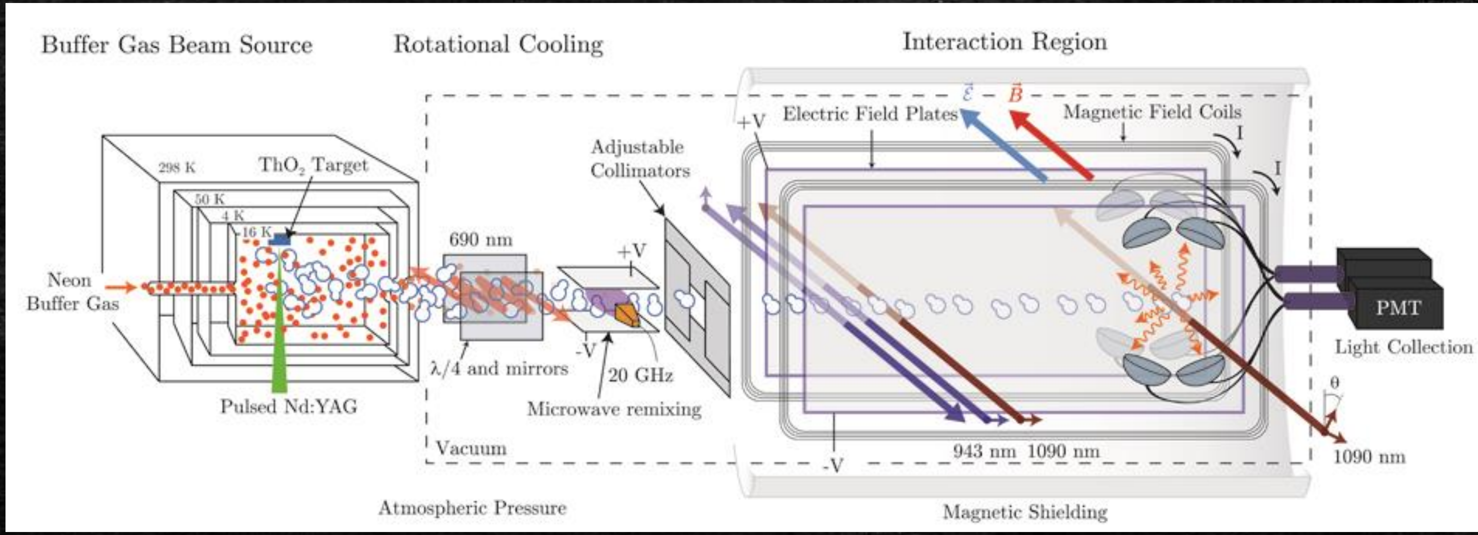
DOI: 10.1038/s41586-018-0599-8

How to measure EDM: ACME II

- Precession of EDM in a strong electric field
  - Field inside ThO molecule is one of the strongest known: 80 GV/cm
- Propagate molecules through shielded chamber
  - Known time-of-flight
  - excite electron to particular spin angle in xy-plane at start
  - measure final angle with fluorescence by linearly polarized laser
- Current measurements at 1.1×10^{-29} e-cm



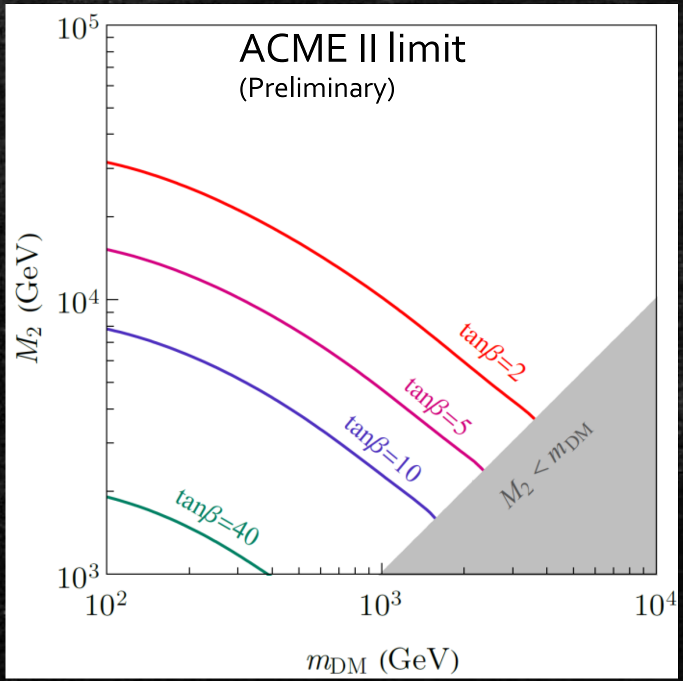
By Mario De Leo - https://commons.wikimedia.org/wiki/File:Precession_in_magnetic_field.svg

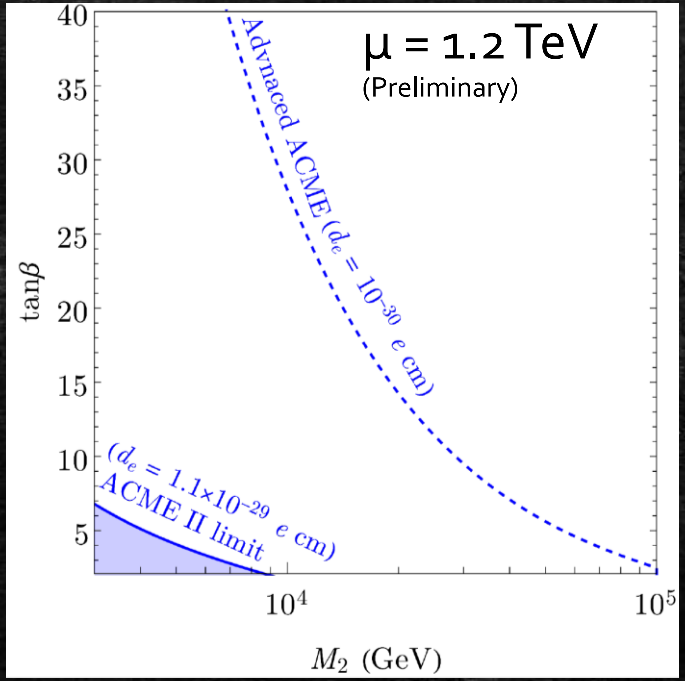


https://www.danielang.net/2016/10/16/guide-to-the-acme-edm-experiment-a-simple-overview/

EDM Results

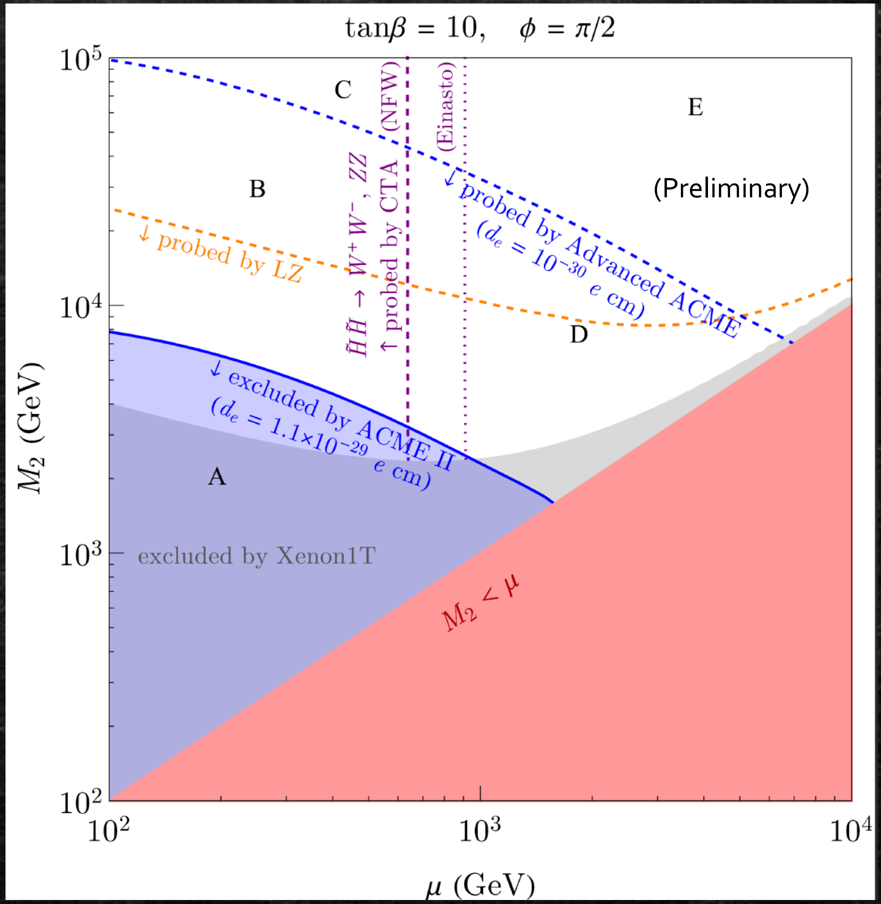
- Current limits confined to low tanβ or low (M_2μ)
- Advanced ACME expected to be an order of magnitude more sensitive
  - Would reach, for μ = 1.2 TeV, M2 ~ O(10 TeV), tanβ ~ O(10) covering our region of interest





Results in Context

- A bit of a race between approaches
  - LZ could significantly increase DD limits, but limited by M_2^2 scaling
  - CTA can have reach in ID search, dependent on astrophysical unknowns
  - e-EDM reach is fairly robust
    - Up to complex phase or large M_2



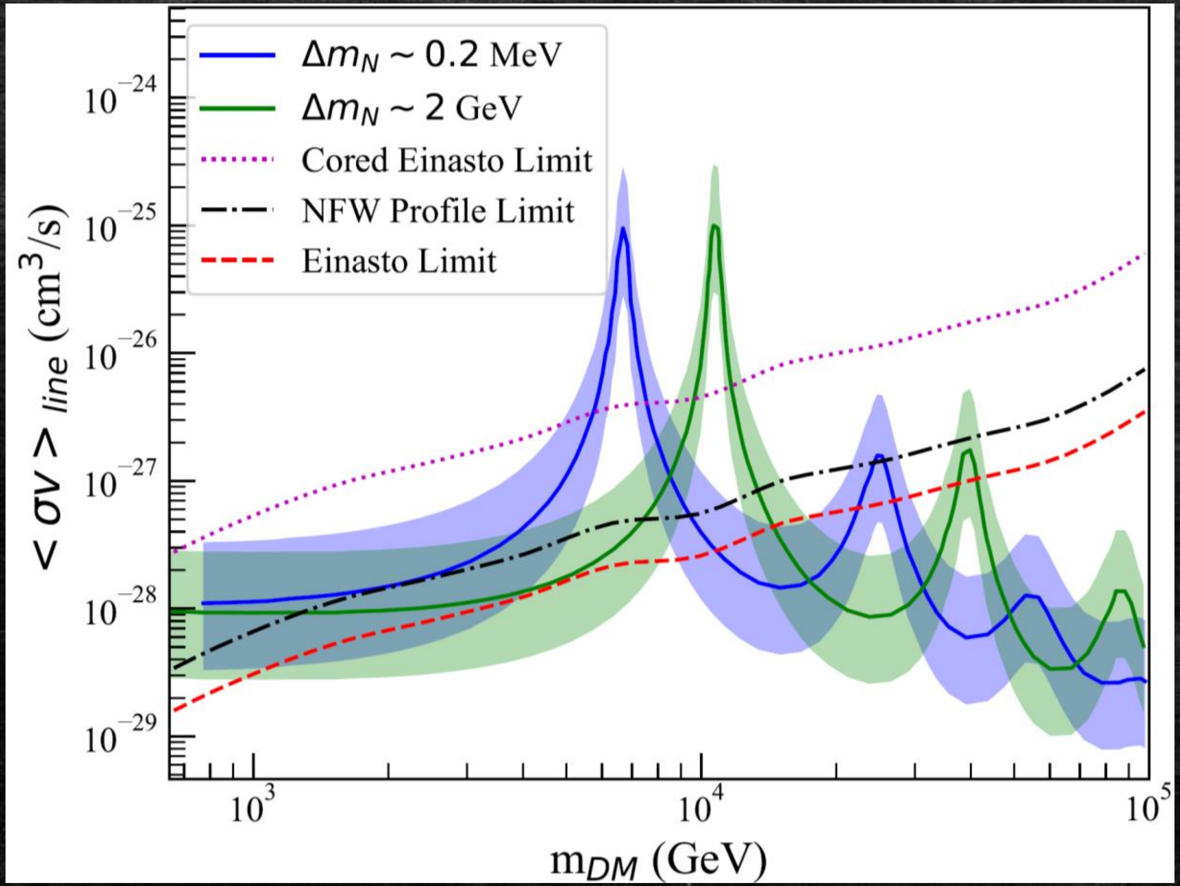
Thank you!

U.S. DEPARTMENT OF ENERGY

Office of Science

NSF

Line photon indirect detection searches



Data for curves courtesy of Rinchiuso, et al. 1905.00315, Hryczuk, et al 2008.00692