

Phenomenology of the minimal inverse-seesaw model with Abelian flavour symmetries

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Inverse-seesaw (ISS) mechanism

ISS(n_R, n_s) Mohapatra; Mohapatra & Valle'86; Gonzalez-Garcia & Valle'89

➤ Sterile neutrino fields: ν_{Ri} ($i = 1, \dots, n_R$), s_i ($i = 1, \dots, n_s$)

$$-\mathcal{L}_{\text{mass}}^{\text{ISS}} = \bar{e}_L \mathbf{M}_\ell e_R + \bar{\nu}_L \mathbf{M}_D \nu_R + \bar{\nu}_R \mathbf{M}_R s + \frac{1}{2} \bar{s}^c \mathbf{M}_s s + \text{H.c.} \longrightarrow \mathcal{M} = \begin{pmatrix} 0 & \mathbf{M}_D^* & 0 \\ \mathbf{M}_D^\dagger & 0 & \mathbf{M}_R \\ 0 & \mathbf{M}_R^T & \mathbf{M}_s \end{pmatrix}$$

$(3 + n_R + n_s) \times (3 + n_R + n_s)$

➤ Effective neutrino mass matrix ($m_\nu, \mu_s \ll M$):

$$\mathbf{M}_{\text{eff}} = -\mathbf{M}_D^* (\mathbf{M}_R^T)^{-1} \mathbf{M}_s \mathbf{M}_R^{-1} \mathbf{M}_D^\dagger \longrightarrow m_\nu \sim \mu_s \frac{m_D^2}{M^2}$$

➤ Active-sterile mixing:

$$U_{\text{Hl}} \simeq \mathbf{V}_L^\dagger (0, \mathbf{M}_D (\mathbf{M}_R^\dagger)^{-1}) \mathbf{U}_s \longrightarrow U_{\text{Hl}} \sim \frac{m_D}{M} \sim \sqrt{\frac{m_\nu}{\mu_s}}$$

Type-I seesaw: $m_\nu \sim \frac{m_D^2}{M}, U_{\text{Hl}} \sim \frac{m_D}{M} \sim \sqrt{\frac{m_\nu}{M}}$

- The ISS is a **low-scale** neutrino mass generation mechanism
- ISS provides a **natural template** for (active) neutrino mass suppression with **sizeable** active-sterile neutrino mixing

Minimal Inverse Seesaw:

ISS(n_R, n_s) \longrightarrow ISS(2, 2)

Abada & Lucente'14

- One massless neutrino
- Neutrino data can be accommodated
- Still 17 parameters (in the \mathbf{M}_s diagonal basis)

Neutrino oscillation data

Minimal Inverse Seesaw ISS(2,2):

17 parameters vs 7 observables

$$\theta_{12}, \theta_{23}, \theta_{13}, \Delta m_{21,31}^2, \delta, \alpha$$

Parameter	Best Fit $\pm 1\sigma$	3σ range
$\theta_{12}(\circ)$	34.3 ± 1.0	$31.4 \rightarrow 37.4$
$\theta_{23}(\circ)$ [NO]	$48.79^{+0.93}_{-1.25}$	$41.63 \rightarrow 51.32$
$\theta_{23}(\circ)$ [IO]	$48.79^{+1.04}_{-1.30}$	$41.88 \rightarrow 51.30$
$\theta_{13}(\circ)$ [NO]	$8.58^{+0.11}_{-0.15}$	$8.16 \rightarrow 8.94$
$\theta_{13}(\circ)$ [IO]	$8.63^{+0.11}_{-0.15}$	$8.21 \rightarrow 8.99$
$\delta(\circ)$ [NO]	216^{+41}_{-25}	$144 \rightarrow 360$
$\delta(\circ)$ [IO]	277^{+23}_{-24}	$205 \rightarrow 342$
$\Delta m_{21}^2 (\times 10^{-5} \text{ eV}^2)$	$7.50^{+0.22}_{-0.20}$	$6.94 \rightarrow 8.14$
$ \Delta m_{31}^2 (\times 10^{-3} \text{ eV}^2)$ [NO]	$2.56^{+0.03}_{-0.04}$	$2.46 \rightarrow 2.65$
$ \Delta m_{31}^2 (\times 10^{-3} \text{ eV}^2)$ [IO]	2.46 ± 0.03	$2.37 \rightarrow 2.55$

de Salas et. al, 20; Capozzi et. al'20; Esteban et. al'20

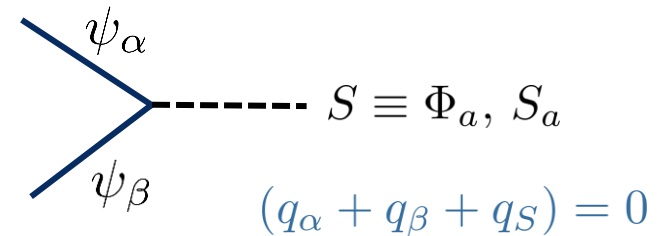
Abelian flavour symmetries:

- All mass terms generated dynamically

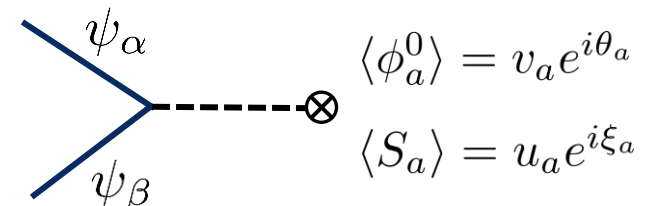
Mass matrices

$$\mathbf{M}_\ell, \mathbf{M}_D$$

$$\mathbf{M}_R, \mathbf{M}_s$$



- Impose texture zeros in the mass matrices **reducing the number of parameters**
- CPV from vacuum phases (SCPV)



Scalar content and Yukawa Lagrangian

- Need to add a **second Higgs doublet** to be able to realise the charged-lepton mass matrix textures.
- Add **two neutral complex scalar singlets** to dynamically generate \mathbf{M}_S and \mathbf{M}_R .

$$\Phi_{1,2} = \begin{pmatrix} \phi_{1,2}^+ \\ \phi_{1,2}^0 \end{pmatrix} = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2}\phi_{1,2}^+ \\ v_{1,2}e^{i\theta_{1,2}} + \rho_{1,2} + i\eta_{1,2} \end{pmatrix}, \quad S_{1,2} = \frac{1}{\sqrt{2}} (u_{1,2}e^{i\xi_{1,2}} + \rho_{3,4} + i\eta_{3,4})$$

Yukawa Lagrangian

$$-\mathcal{L}_{\text{Yuk.}} = \bar{\ell}_L (\mathbf{Y}_\ell^1 \Phi_1 + \mathbf{Y}_\ell^2 \Phi_2) e_R + \bar{\ell}_L (\mathbf{Y}_D^1 \tilde{\Phi}_1 + \mathbf{Y}_D^2 \tilde{\Phi}_2) \nu_R \\ + \frac{1}{2} \bar{s}^c (\mathbf{Y}_s^1 S_1 + \mathbf{Y}_s^2 S_1^*) s + \bar{\nu}_R (\mathbf{Y}_R^1 S_2 + \mathbf{Y}_R^2 S_2^*) s + \text{H.c.}$$

Scalar potential

$$V(\Phi_a, S_a) = V_{\text{sym.}} + V_{\text{soft}}(\Phi_a, S_a)$$

$$V_{\text{soft}}(\Phi_a, S_a) = \mu_{12}^2 \Phi_1^\dagger \Phi_2 + \mu_3^2 S_1^2 + \mu_4 |S_1|^2 S_1 \\ + \mu_5 |S_2|^2 S_2 + \text{H.c.}$$

- SCPV is achieved by: $\theta, \xi_2 = 0, \xi_1 = \arctan \left(\frac{\sqrt{32\mu_3^4 - \mu_4^2 u_1^2}}{\mu_4 u_1} \right)$

Abelian flavour symmetries

- Maximally-restrictive texture sets compatible with neutrino oscillation data that are realisable by Abelian symmetries:

$$\mathbf{G}_F = \mathbf{U}(1) \times \mathbb{Z}_n \times \mathbf{U}(1)_F, \quad n = 2, 4$$

Fields	U(1)	$(5_{1,I}^\ell, T_{45})$	$(4_3^\ell, T_{124})$	$(4_3^\ell, T_{456})$	$(4_3^\ell, T_{136,I})$	$(4_3^\ell, T_{146,I})$
		$\mathbb{Z}_2 \times \mathbf{U}(1)_F$	$\mathbb{Z}_2 \times \mathbf{U}(1)_F$	$\mathbb{Z}_2 \times \mathbf{U}(1)_F$	$\mathbb{Z}_4 \times \mathbf{U}(1)_F$	$\mathbb{Z}_4 \times \mathbf{U}(1)_F$
Φ_1	0	(1, 1)	(0, -5)	(1, 1)	(1, 2)	(0, 1)
Φ_2	0	(0, -1)	(1, -3)	(0, -1)	(0, 1)	(3, 0)
S_1	0	(0, 2)	(0, -2)	(0, -2)	(0, -2)	(0, -2)
S_2	1	(0, 0)	(0, 0)	(1, 0)	(0, 0)	(0, 0)
ℓ_{eL}	1	(1, 0)	(0, 0)	(0, 0)	(2, 0)	(2, 0)
$\ell_{\mu L}$	1	(0, 2)	(1, 2)	(1, -2)	(1, -1)	(1, -1)
$\ell_{\tau L}$	1	(0, -2)	(0, 4)	(0, -4)	(0, -2)	(0, -2)
e_R	1	(1, -3)	(0, 9)	(1, -5)	(3, -4)	(0, -3)
μ_R	1	(0, 3)	(1, 7)	(0, -3)	(0, -3)	(1, -2)
τ_R	1	(0, -1)	(0, 5)	(1, -1)	(1, -2)	(2, -1)
ν_{R1}	1	(0, 1)	(0, -1)	(0, -1)	(0, -1)	(0, -1)
ν_{R2}	1	(1, -1)	(1, 1)	(1, 1)	(2, 1)	(2, 1)
s_1	0	(1, -1)	(1, 1)	(0, 1)	(2, 1)	(2, 1)
s_2	0	(0, 1)	(0, -1)	(1, -1)	(0, -1)	(0, -1)

Abelian flavour symmetries

ONLY INTERESTING CASE

Fields	U(1)	$(5_{1,I}^\ell, T_{45})$ $\mathbb{Z}_2 \times U(1)_F$
Φ_1	0	(1, 1)
Φ_2	0	(0, -1)
S_1	0	(0, 2)
S_2	1	(0, 0)
ℓ_{eL}	1	(1, 0)
$\ell_{\mu L}$	1	(0, 2)
$\ell_{\tau L}$	1	(0, -2)
e_R	1	(1, -3)
μ_R	1	(0, 3)
τ_R	1	(0, -1)
ν_{R1}	1	(0, 1)
ν_{R2}	1	(1, -1)
s_1	0	(1, -1)
s_2	0	(0, 1)

$$-\mathcal{L}_{\text{Yuk.}} = \bar{\ell}_L (\mathbf{Y}_\ell^1 \Phi_1 + \mathbf{Y}_\ell^2 \Phi_2) e_R + \bar{\ell}_L (\mathbf{Y}_D^1 \tilde{\Phi}_1 + \mathbf{Y}_D^2 \tilde{\Phi}_2) \nu_R + \frac{1}{2} \bar{s}^c (\mathbf{Y}_s^1 S_1 + \mathbf{Y}_s^2 S_1^*) s + \bar{\nu}_R (\mathbf{Y}_R^1 S_2 + \mathbf{Y}_R^2 S_2^*) s + \text{H.c.}$$

Mass matrices Yukawa decompositions

$$\begin{array}{c} \hline \mathbf{M}_\ell \quad \mathbf{Y}_\ell^1 \quad \mathbf{Y}_\ell^2 \\ \hline 5_{1,I}^\ell \quad \begin{pmatrix} 0 & 0 & \times \\ 0 & 0 & 0 \\ \times & 0 & 0 \end{pmatrix} \quad \begin{pmatrix} 0 & 0 & 0 \\ 0 & \times & 0 \\ 0 & 0 & \times \end{pmatrix} \\ \hline \end{array}$$

$$\begin{array}{c} \hline \mathbf{M}_R \quad \mathbf{Y}_R \\ \hline T_{14} \quad \begin{pmatrix} 0 & \times \\ \times & 0 \end{pmatrix} \\ \hline \end{array}$$

$$\begin{array}{c} \hline \mathbf{M}_D \quad \mathbf{Y}_D^1 \quad \mathbf{Y}_D^2 \\ \hline T_{45} \quad \begin{pmatrix} \times & 0 \\ 0 & 0 \\ 0 & \times \end{pmatrix} \quad \begin{pmatrix} 0 & \times \\ \times & 0 \\ 0 & 0 \end{pmatrix} \\ \hline \end{array}$$

$$\begin{array}{c} \hline \mathbf{M}_s \quad \mathbf{Y}_s^1 \quad \mathbf{Y}_s^2 \\ \hline T_{23} \quad \begin{pmatrix} \times & 0 \\ 0 & 0 \end{pmatrix} \quad \begin{pmatrix} 0 & 0 \\ 0 & \times \end{pmatrix} \\ \hline \end{array}$$

A common origin for Leptonic CPV

- Parameterisation of the **charged lepton-mass matrix**:

$$5_1^\ell : \mathbf{M}_\ell = \begin{pmatrix} 0 & 0 & a_1 \\ 0 & m_{\ell_1}^2 & 0 \\ a_2 & 0 & a_4 \end{pmatrix}, \quad \mathbf{H}_\ell = \begin{pmatrix} a_1^2 & 0 & a_1 a_4 \\ 0 & a_3^2 & 0 \\ a_1 a_4 & 0 & a_2^2 + a_4^2 \end{pmatrix}, \quad \mathbf{V}'_L = \begin{pmatrix} c_L & 0 & s_L \\ 0 & 1 & 0 \\ -s_L & 0 & c_L \end{pmatrix} \theta_L$$

$$5_1^e : \mathbf{V}_{L,R} = \mathbf{V}'_{L,R} \mathbf{P}_{12}, \quad 5_1^\mu : \mathbf{V}_{L,R} = \mathbf{V}'_{L,R}, \quad 5_1^\tau : \mathbf{V}_{L,R} = \mathbf{V}'_{L,R} \mathbf{P}_{23},$$

NO_{e,μ,τ}, IO_{e,μ,τ} \longrightarrow 6 distinct cases to be analysed

Real Yukawa couplings (CP is conserved at the Lagrangian level)

$$\mathbf{Y}_D^1 = \begin{pmatrix} b_1 & 0 \\ 0 & 0 \\ 0 & b_2 \end{pmatrix}, \quad \mathbf{Y}_D^2 = \begin{pmatrix} 0 & b_3 \\ b_4 & 0 \\ 0 & 0 \end{pmatrix}, \quad \mathbf{Y}_R = \begin{pmatrix} 0 & d_2 \\ d_1 & 0 \end{pmatrix}, \quad \mathbf{Y}_s^1 = \begin{pmatrix} f_2 & 0 \\ 0 & 0 \end{pmatrix}, \quad \mathbf{Y}_s^2 = \begin{pmatrix} 0 & 0 \\ 0 & f_1 \end{pmatrix}$$

VEV configuration:

$$\langle \phi_1^0 \rangle = v \cos \beta$$

$$\langle \phi_2^0 \rangle = v \sin \beta$$

$$\langle S_1 \rangle = u_1 e^{i\xi}, \quad \langle S_2 \rangle = u_2$$



$$\mathbf{M}_D = \begin{pmatrix} m_{D_1} & m_{D_3} \\ m_{D_4} & 0 \\ 0 & m_{D_2} \end{pmatrix}, \quad \mathbf{M}_R = \begin{pmatrix} 0 & M \\ qM & 0 \end{pmatrix}, \quad \mathbf{M}_s = \begin{pmatrix} p \mu_s e^{i\xi} & 0 \\ 0 & \mu_s e^{-i\xi} \end{pmatrix}$$

Correlation between low-energy observables

➤ Effective neutrino mass matrix: $\mathbf{V}_L^\dagger \mathbf{M}_{\text{eff}} \mathbf{V}_L$

$$\mathbf{M}_{\text{eff}} = e^{-i\xi} \begin{pmatrix} \frac{y^2}{x} + \frac{z^2}{w} e^{2i\xi} & y & ze^{2i\xi} \\ x & y & 0 \\ ze^{2i\xi} & 0 & we^{2i\xi} \end{pmatrix}, \quad \mathbf{V}_L = \begin{pmatrix} \cos \theta_L & 0 & \sin \theta_L \\ 0 & 1 & 0 \\ -\sin \theta_L & 0 & \cos \theta_L \end{pmatrix}$$

$$\begin{aligned} 5_1^e &: \mathbf{V}_{L,R} = \mathbf{V}'_{L,R} \mathbf{P}_{12} \\ 5_1^\mu &: \mathbf{V}_{L,R} = \mathbf{V}'_{L,R} \\ 5_1^\tau &: \mathbf{V}_{L,R} = \mathbf{V}'_{L,R} \mathbf{P}_{23} \end{aligned}$$

$$z = \mu_s \frac{m_{D_2} m_{D_3}}{M^2} \frac{p}{q^2}, \quad w = \mu_s \frac{m_{D_2}^2}{M^2} \frac{p}{q^2}, \quad x = \mu_s \frac{m_{D_4}^2}{M^2}, \quad y = \mu_s \frac{m_{D_1} m_{D_4}}{M^2}$$

➤ The effective light neutrino mass matrix is written solely in terms of 6 effective parameters:

$$(x, y, z, w, \theta_L, \xi) \longrightarrow \mathcal{O}_i \equiv (\Delta m_{21}^2, \Delta m_{31}^2, \theta_{ij}, \delta, \alpha)$$

$$\text{NO} : M_{ij} = \left[\mathbf{U}'^* \text{diag} \left(0, \sqrt{\Delta m_{21}^2}, \sqrt{\Delta m_{31}^2} \right) \mathbf{U}'^\dagger \right]_{ij}$$

$$\text{IO} : M_{ij} = \left[\mathbf{U}'^* \text{diag} \left(\sqrt{\Delta m_{31}^2}, \sqrt{\Delta m_{21}^2 + \Delta m_{31}^2}, 0 \right) \mathbf{U}'^\dagger \right]_{ij}$$

$$D_{ij} = M_{ii} M_{jj} - M_{ij}^2$$

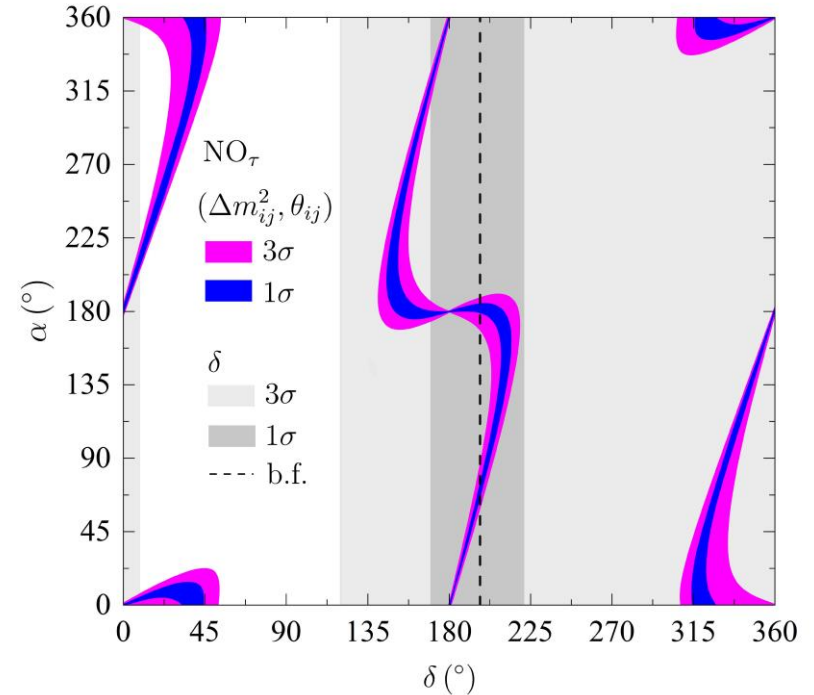
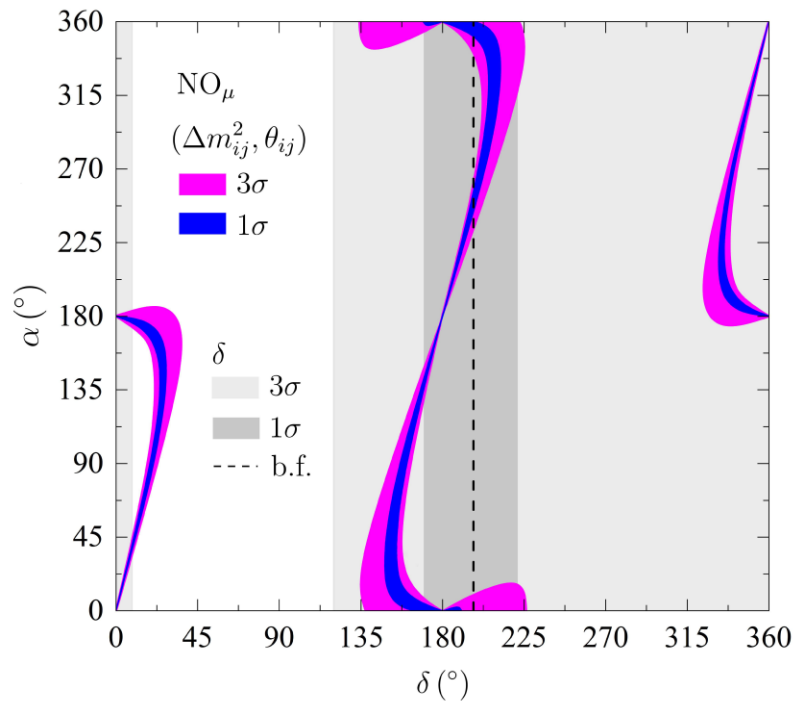
Low-energy relations:

$$5_1^e : \arg \left[M_{11}^{*2} M_{13}^2 \frac{D_{12}}{D_{23}} \right] = 0$$

$$5_1^\mu : \arg \left[M_{12}^{*2} M_{23}^2 \frac{D_{12}}{D_{23}} \right] = 0$$

$$5_1^\tau : \arg \left[M_{13}^{*2} M_{33}^2 \frac{D_{12}}{D_{23}} \right] = 0$$

Leptonic CP violation



➤ Strong correlation between α and δ ,

➤ No Dirac CPV implies no Majorana CPV, ➔

$$\langle S_1 \rangle = u_1 e^{i\xi}$$

$$\mathcal{J}_{\text{Dirac}}^{\text{CP}}, \mathcal{J}_{\text{Maj}}^{\text{CP}} \propto \sin(2\xi)$$

as in Branco, Felipe, FRJ, Serôdio (2012)

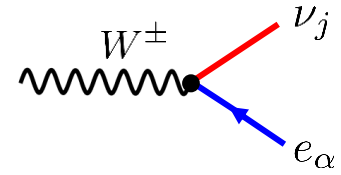
➤ A measurement of δ in the intervals $[45^\circ, 135^\circ]$ and $[225^\circ, 315^\circ]$ would exclude the NO_μ and NO_τ cases,

➤ $\beta\beta_{0\nu}$ analysis done in the paper (no time to discuss here).

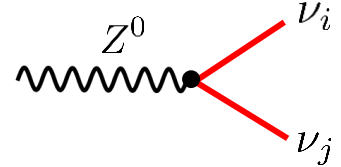
Heavy-light mixing relations

- Charged and neutral current interactions

$$\mathcal{L}_{W^\pm} = \frac{g}{\sqrt{2}} W_\mu^- \sum_{\alpha=1}^3 \sum_{j=1}^{n_f} \mathbf{B}_{\alpha j} \bar{e}_\alpha \gamma^\mu P_L \nu_j + \text{H.c.}$$



$$\mathcal{L}_Z = \frac{g}{4c_W} Z_\mu \sum_{i,j=1}^{n_f} \bar{\nu}_i \gamma^\mu (\mathbf{C}_{ij} P_L - \mathbf{C}_{ij}^* P_R) \nu_j, \quad \mathbf{C}_{ij} = \sum_{\alpha=1}^3 \mathbf{B}_{\alpha i}^* \mathbf{B}_{\alpha j}$$



$$\frac{\mathbf{B}_{e4}}{\mathbf{B}_{\mu4}} \simeq \frac{\mathbf{B}_{e5}}{\mathbf{B}_{\mu5}} \simeq \frac{x}{y c_L}, \quad \frac{\mathbf{B}_{\tau4}}{\mathbf{B}_{\mu4}} \simeq \frac{\mathbf{B}_{\tau5}}{\mathbf{B}_{\mu5}} \simeq \tan \theta_L, \quad \frac{\mathbf{B}_{\mu6}}{\mathbf{B}_{\tau6}} \simeq \frac{\mathbf{B}_{\mu7}}{\mathbf{B}_{\tau7}} \simeq \frac{z - w \tan \theta_L}{w + z \tan \theta_L}, \quad \mathbf{B}_{e6} \simeq \mathbf{B}_{e7} \simeq 0$$

- Numerical estimates

	NO _e	NO _μ	NO _τ	IO _e	IO _μ	IO _τ
$\mathbf{B}_{e4}/\mathbf{B}_{\mu4} \simeq \mathbf{B}_{e5}/\mathbf{B}_{\mu5}$	0.21	0.17	0.17	2.73	0.21	0.41
$\mathbf{B}_{\tau4}/\mathbf{B}_{\mu4} \simeq \mathbf{B}_{\tau5}/\mathbf{B}_{\mu5}$	0.27	0.88	0.87	0.51	1.09	1.24
$\mathbf{B}_{\tau4}/\mathbf{B}_{e4} \simeq \mathbf{B}_{\tau5}/\mathbf{B}_{e5}$	1.27	5.07	5.24	0.19	5.33	5.02
$\mathbf{B}_{e6}/\mathbf{B}_{\mu6} \simeq \mathbf{B}_{e7}/\mathbf{B}_{\mu7}$	0	–	0.36	0	–	4.96
$\mathbf{B}_{\tau6}/\mathbf{B}_{\mu6} \simeq \mathbf{B}_{\tau7}/\mathbf{B}_{\mu7}$	0.61	–	0	1.14	–	0
$\mathbf{B}_{\tau6}/\mathbf{B}_{e6} \simeq \mathbf{B}_{\tau7}/\mathbf{B}_{e7}$	–	1.64	0	–	0.23	0

- The $\mathbf{B}_{\alpha i}$ ($\alpha = e, \mu, \tau$) ($i = 4, \dots, 7$) are related to each other;
- The relations are expressed solely in terms of the low-energy neutrino observables;
- Due to the flavour symmetries the heavy-light mixing parameters are not independent.

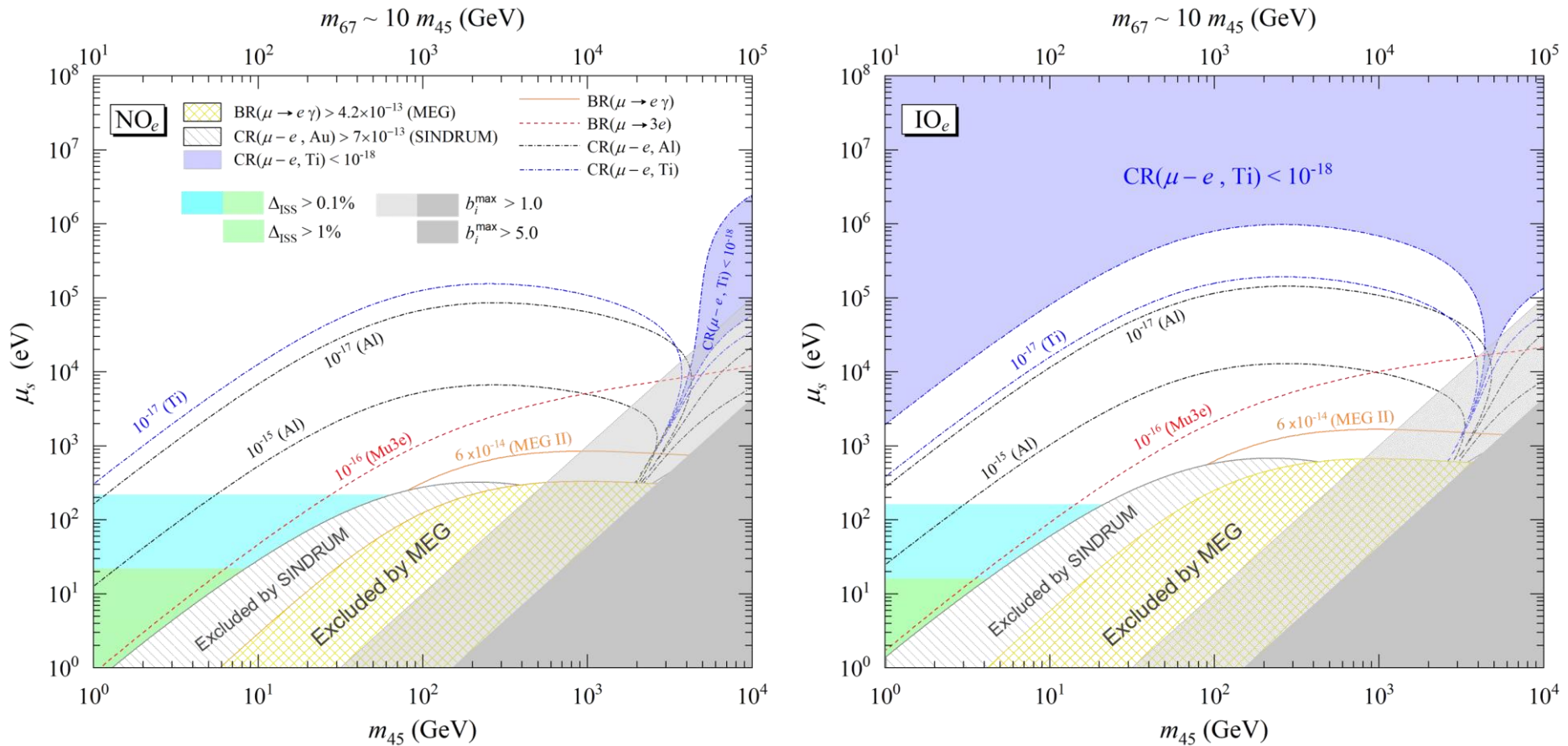
Relations among cLFV processes (no time to discuss here)

Charged lepton flavour violation (cLFV)

cLFV process	Present limit (90% CL)	Future sensitivity
$\text{BR}(\mu \rightarrow e\gamma)$	4.2×10^{-13} (MEG)	6×10^{-14} (MEG II)
$\text{BR}(\tau \rightarrow e\gamma)$	3.3×10^{-8} (BaBar)	3×10^{-9} (Belle II)
$\text{BR}(\tau \rightarrow \mu\gamma)$	4.4×10^{-8} (BaBar)	10^{-9} (Belle II)
$\text{BR}(\mu^- \rightarrow e^- e^+ e^-)$	1.0×10^{-12} (SINDRUM)	10^{-16} (Mu3e)
$\text{BR}(\tau^- \rightarrow e^- e^+ e^-)$	2.7×10^{-8} (Belle)	5×10^{-10} (Belle II)
$\text{BR}(\tau^- \rightarrow e^- \mu^+ \mu^-)$	2.7×10^{-8} (Belle)	5×10^{-10} (Belle II)
$\text{BR}(\tau^- \rightarrow e^+ \mu^- \mu^-)$	1.7×10^{-8} (Belle)	3×10^{-10} (Belle II)
$\text{BR}(\tau^- \rightarrow \mu^- e^+ e^-)$	1.8×10^{-8} (Belle)	3×10^{-10} (Belle II)
$\text{BR}(\tau^- \rightarrow \mu^+ e^- e^-)$	1.5×10^{-8} (Belle)	3×10^{-10} (Belle II)
$\text{BR}(\tau^- \rightarrow \mu^- \mu^+ \mu^-)$	2.1×10^{-8} (Belle)	4×10^{-10} (Belle II)
$\text{CR}(\mu - e, \text{Al})$	–	3×10^{-17} (Mu2e) $10^{-15} - 10^{-17}$ (COMET I-II)
$\text{CR}(\mu - e, \text{Ti})$	4.3×10^{-12} (SINDRUM II)	10^{-18} (PRISM/PRIME)
$\text{CR}(\mu - e, \text{Au})$	7×10^{-13} (SINDRUM II)	–
$\text{CR}(\mu - e, \text{Pb})$	4.6×10^{-11} (SINDRUM II)	–

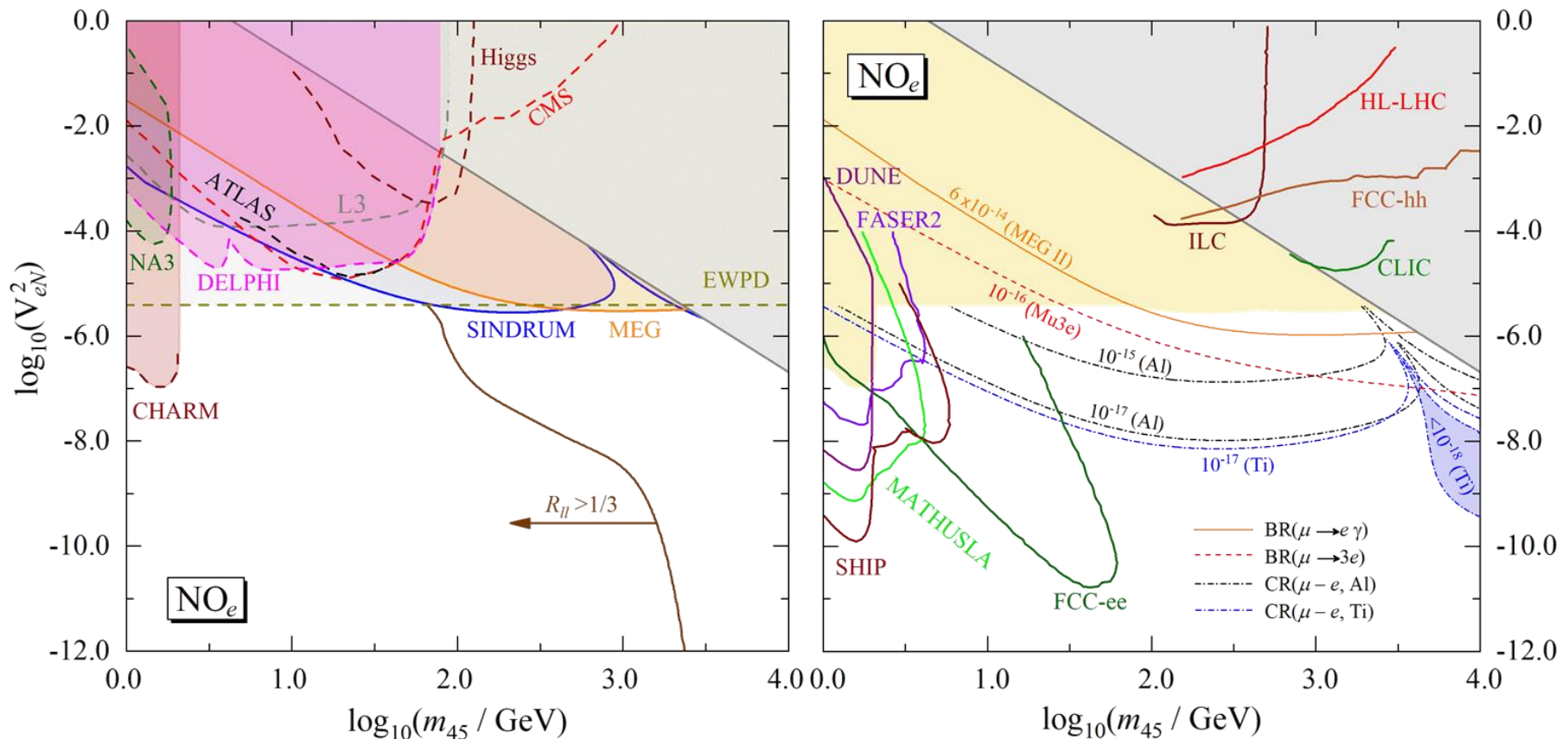
➤ **Muon cLFV:** strongest current constraints and future lowest sensitivities

cLFV in the ISS(2,2) with Abelian symmetries



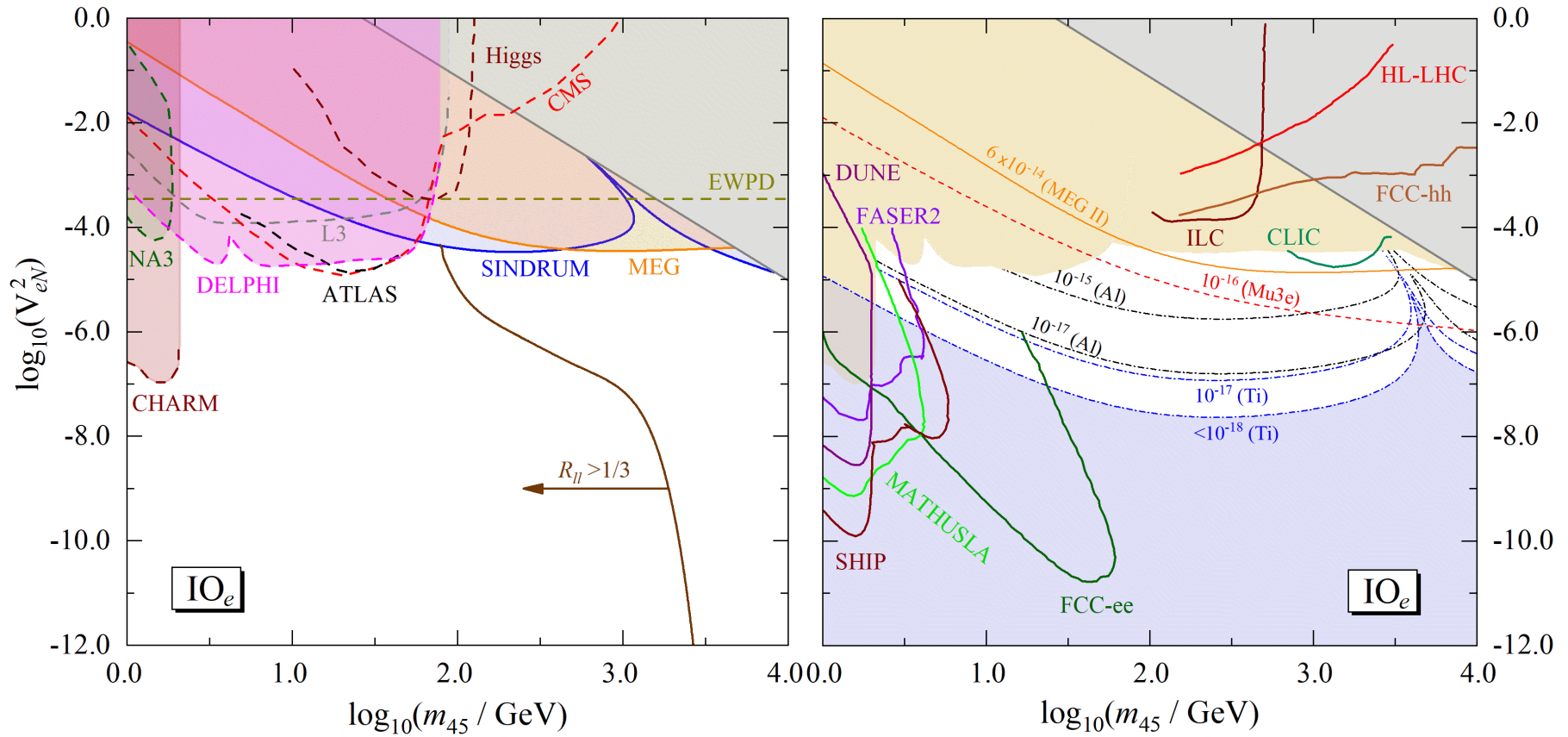
- For NO, almost the **whole parameter space will be scrutinized by future μ - e conversion** experiments (Mu2e, COMET, PRISM/PRIME);
- For IO, the prospects are less optimistic.

Constraints on heavy sterile neutrinos



- Current data implies an upper bound $V_{eN}^2 \sim 10^{-6} - 10^{-5}$;
 - Future probes will be sensitive to much smaller mixings.
- Indirect LFV experiments fully complementary to other direct searches.

Constraints on heavy sterile neutrinos



- EWPD is less constraining in the IO case;
- Future CLV probes will be sensitive to $V_{eN}^2 \sim 10^{-7}$.

Conclusion

- Comprehensive study of the **minimal inverse seesaw model** constrained by **Abelian flavour symmetries** with all mass terms generated via SSB;
- Majorana and Dirac-type CP violation **are related**;
- Relations among LFV parameters in our framework provide a **very constrained setup for phenomenological studies**;
- **Constraining power of cLFV processes** in the model's parameter space;
- **Alternative probes** such as beam-dump, hadron-collider, linear-collider, displaced-vertex experiments as well as EWPD.

Analysed in paper: Impact of radiative correction on neutrino masses, neutrinoless double beta decay, relations among tau and muon LFV decays, ...

Thank you!