

# Is a Miracle-less WIMP Ruled out?

Jason Arakawa

University of California, Irvine

arXiv:2101.11031 with Tim M.P. Tait



# Miracle-less WIMP

WIMPs: popular DM candidate for many years.

- ▶ Predicted by many BSM theories (e.g. SUSY)
- ▶ Freeze out of WIMPs  $\rightarrow$  correct relic abundance (WIMP Miracle)

However,

- ▶ Direct, indirect, and collider searches  $\rightarrow$  constrain WIMP miracle regime
- ▶ No data before BBN  $\rightarrow$  Early universe is unknown

What are the constraints for WIMPs if we stay agnostic to the production mechanism?

# Miracle-less WIMP

We make only the following assumptions:

1. WIMP is a real scalar  $SU(2)_{EW}$  triplet,  $\phi$
2. Masses between  $0.1 \text{ GeV} \lesssim m_\phi \lesssim 100 \text{ TeV}$ .
3. DM abundance matches observed value

Outline:

1. General Model
2. Collider Constraints
3. Direct Detection
4. Indirect Detection

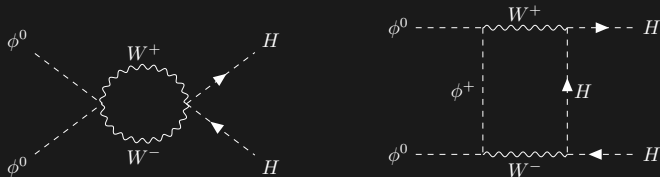
## $SU(2)_{EW}$ Triplet Scalar DM

For a real  $SU(2)_{EW}$  triplet scalar,  $\phi$ , the most general, renormalizable Lagrangian is:

$$\mathcal{L} = \frac{1}{2}(D_\mu\phi)_i(D^\mu\phi)_i - \frac{\mu_\phi^2}{2}\phi^2 - \frac{\lambda_\phi}{4!}\phi^4 - \lambda H^\dagger\phi^2 H$$

# $SU(2)_{EW}$ Triplet Scalar DM

Small Higgs coupling  $\lambda \rightarrow$  fine tuning.  $\lambda$  is renormalized by the diagrams



with one-loop correction,

$$\lambda(\mu) \simeq \lambda(\mu_0) + \frac{g_2^4}{\pi^2} \ln \left( \frac{\mu^2}{\mu_0^2} \right)$$

# Collider Searches

High energy colliders may be able to produce dark matter.

LHC  $\rightarrow$  few search methods:

- ▶ Monojet searches
- ▶ Disappearing Charged Tracks (DCTs)
- ▶ Invisible Higgs decays

$m_\phi \lesssim 287$  GeV excluded<sup>1</sup> if:  $\Delta m \equiv m_{\phi^\pm} - m_{\phi^0} \approx 166$  MeV  
(electroweak effects)

Can evade bounds by adding a dim. 6 operator  $\rightarrow$  increases  $\Delta m$ :

$$\mathcal{L}_{\text{Mass}} = -\frac{1}{\Lambda^2} |\phi^a H^\dagger T^a H|^2 = -\frac{1}{16\Lambda^2} (\phi^0)^2 (v+h)^4$$

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<sup>1</sup>C.-W. Chiang, G. Cottin, Y. Du, K. Fuyuto, and M. J. Ramsey-Musolf, arXiv:2003.07867

# Invisible Higgs Decays

If  $m_\phi \leq M_h/2 \rightarrow h \rightarrow \phi^0\phi^0$

$$\Gamma_{h \rightarrow \phi^0\phi^0} = \frac{\sqrt{M_h^2 - 4m_{\phi^0}^2}}{16\pi M_h^2} v^2 \lambda_{\text{eff}}^2,$$

Branching ratio is bounded<sup>2</sup>,

$$\mathcal{B}(h \rightarrow \text{inv}) = \frac{\Gamma_{\text{DM}}}{\Gamma_{\text{SM}} + \Gamma_{\text{DM}}} \leq 0.19$$

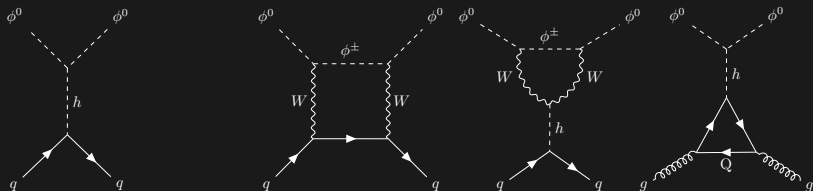
Requires  $\lambda_{\text{eff}} \lesssim 10^{-2}$  for  $m_{\phi^0} \ll M_h$ .

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<sup>2</sup>A. M. Sirunyan et al. (CMS), Phys. Lett. B 793, 520 (2019), arXiv:1809.05937

# Direct Detection

DM can scatter off nucleons in terrestrial experiments:

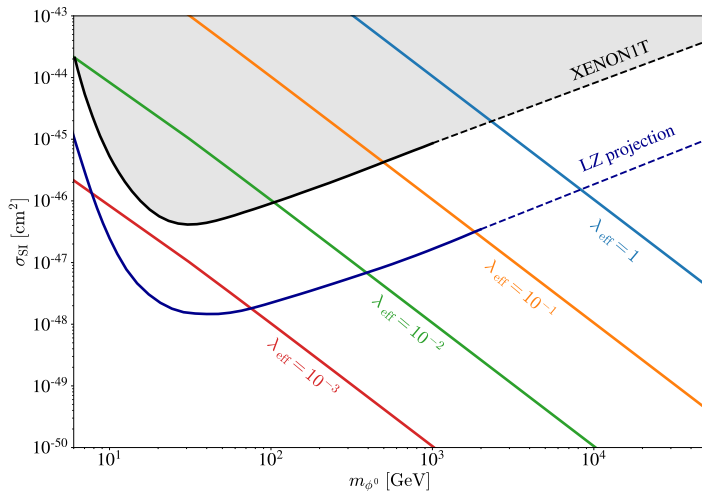


DM-nucleon cross section

$$\sigma_{\text{SI}}(\phi n \rightarrow \phi n) = \frac{\lambda^2}{8\pi m_\phi^2} \frac{\mu_{\phi n}^2 m_n^2}{m_H^4} (f_{T,u}^n + f_{T,d}^n + f_{T,s}^n + \frac{2}{9} f_{T,g}^n)^2$$



# Direct Detection

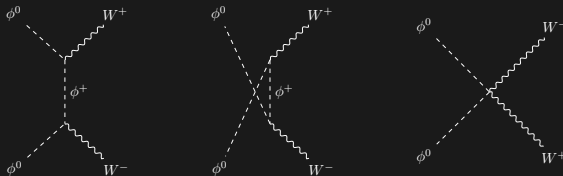


E. Aprile et al. (XENON), Phys. Rev. Lett. 121, 111302 (2018), arXiv:1805.12562

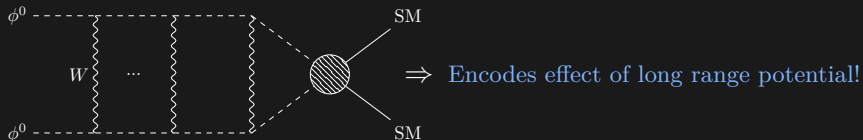
D. Akerib et al. (LUX-ZEPLIN), Phys. Rev. D 101, 052002 (2020), arXiv:1802.06039

# Indirect Detection

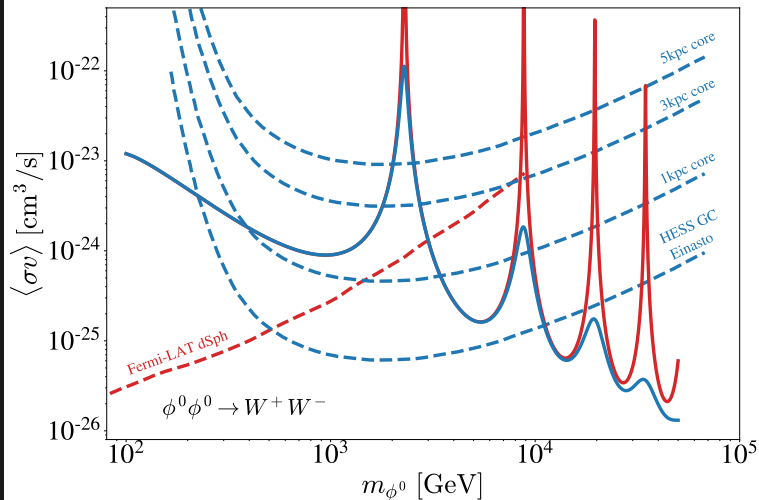
- ▶ Most constraining annihilation final state,  $WW$ .



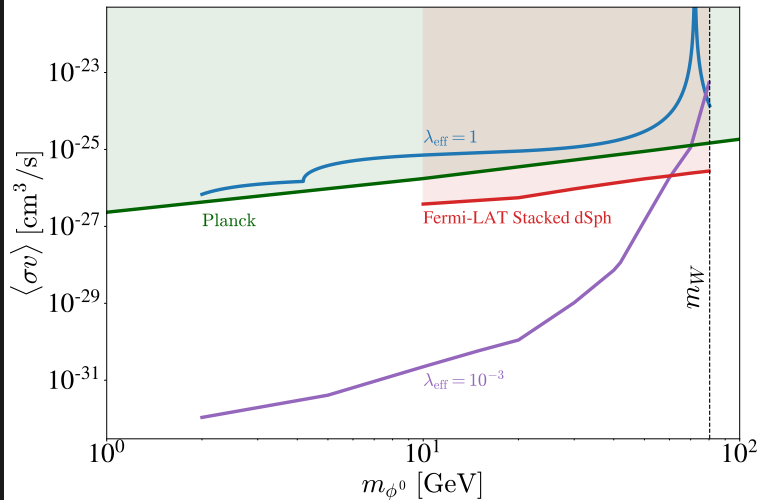
- ▶ Must include Sommerfeld, generated by ladder diagrams.



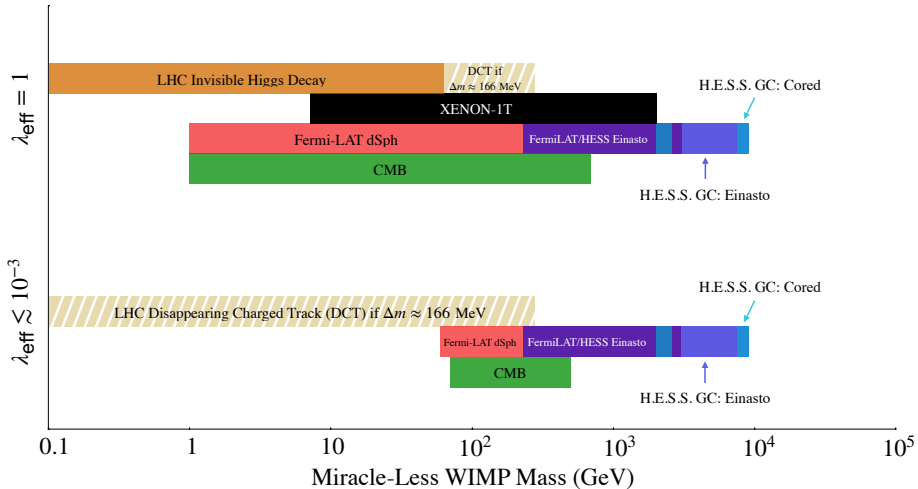
# Indirect Detection



# Indirect Detection: Below the $W$ mass



# Is a Miracle-Less WIMP Ruled Out?

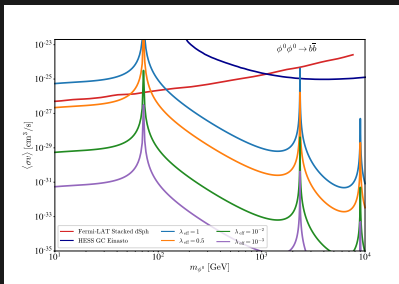
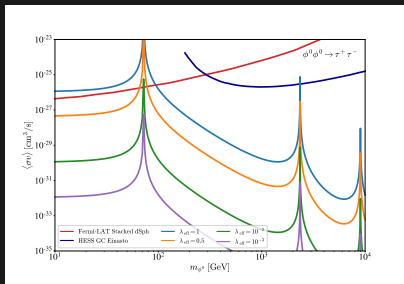


Questions?

# Back-up Slides

# Indirect Detection

We can have annihilation into SM fermions through the Higgs portal.





# Indirect Detection

