

On anomaly-free dark matter models

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with Pavel Fileviez Perez, Elliot Golias, Rui-Hao Li and Clara Murgui

[arXiv: 1904.01017]



PHENO 2019 - University of Pittsburgh

WIMP paradigm

- Initially, dark matter particle in thermal equilibrium with plasma. Later on, due to weak interactions dark matter freezes-out
- Current experiments already impose strong constraints on the simplest WIMPs scenarios, where dark matter is coupled to Standard Model gauge bosons or the Higgs
- However, it is natural to expect new mediator between the SM and the Dark Sector
- In recent years, there has been a strong interest in studying simplified models for dark matter, where the new particles in the spectrum are only the DM and the mediator **Review: [1506.03116]**

GOAL: Study the phenomenology of consistent completions of simplified models.

Spin-1 mediator: New gauge boson that couples to SM fermions

One motivation, is to promote the global symmetries in the SM to gauge symmetries

$U(1)_B$

This talk

$U(1)_L$

$U(1)_{B-L}$



See Clara Murgui
talk Tuesday 15:30

Dark matter is predicted by anomaly cancellation and its stability is guaranteed by remnant $U(1) \rightarrow Z_2$ symmetry

Simplified Dark Matter

χ : Majorana DM

Z_B : Leptophobic mediator

Motivated by gauging **baryon number**, we choose

$$\mathcal{L} \supset \underbrace{\frac{3}{4} g_B \bar{\chi} \gamma^\mu \gamma^5 \chi Z_\mu^B}_{\text{Axial}} - \underbrace{\frac{1}{3} g_B \bar{q} \gamma^\mu q Z_\mu^B}_{\text{Vector}} + \frac{M_\chi}{2v_B} \sin \theta_B \bar{\chi} \chi h_1 - \frac{M_\chi}{2v_B} \cos \theta_B \bar{\chi} \chi h_2$$

Axial

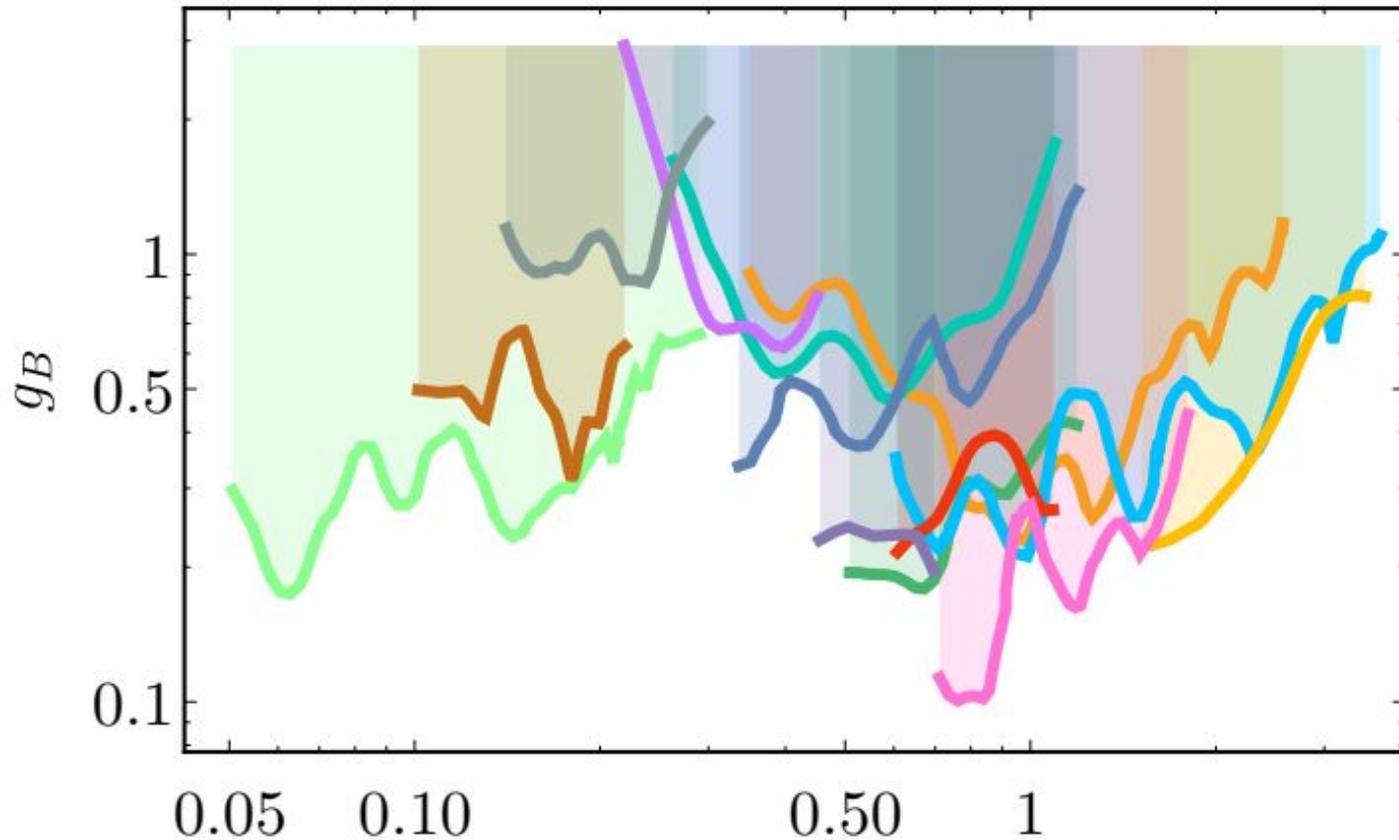
Vector

the free parameters in the model are:

$$M_\chi, M_{Z_B}, M_{h_2}, \theta_B, g_B.$$

LHC bounds on leptophobic gauge boson

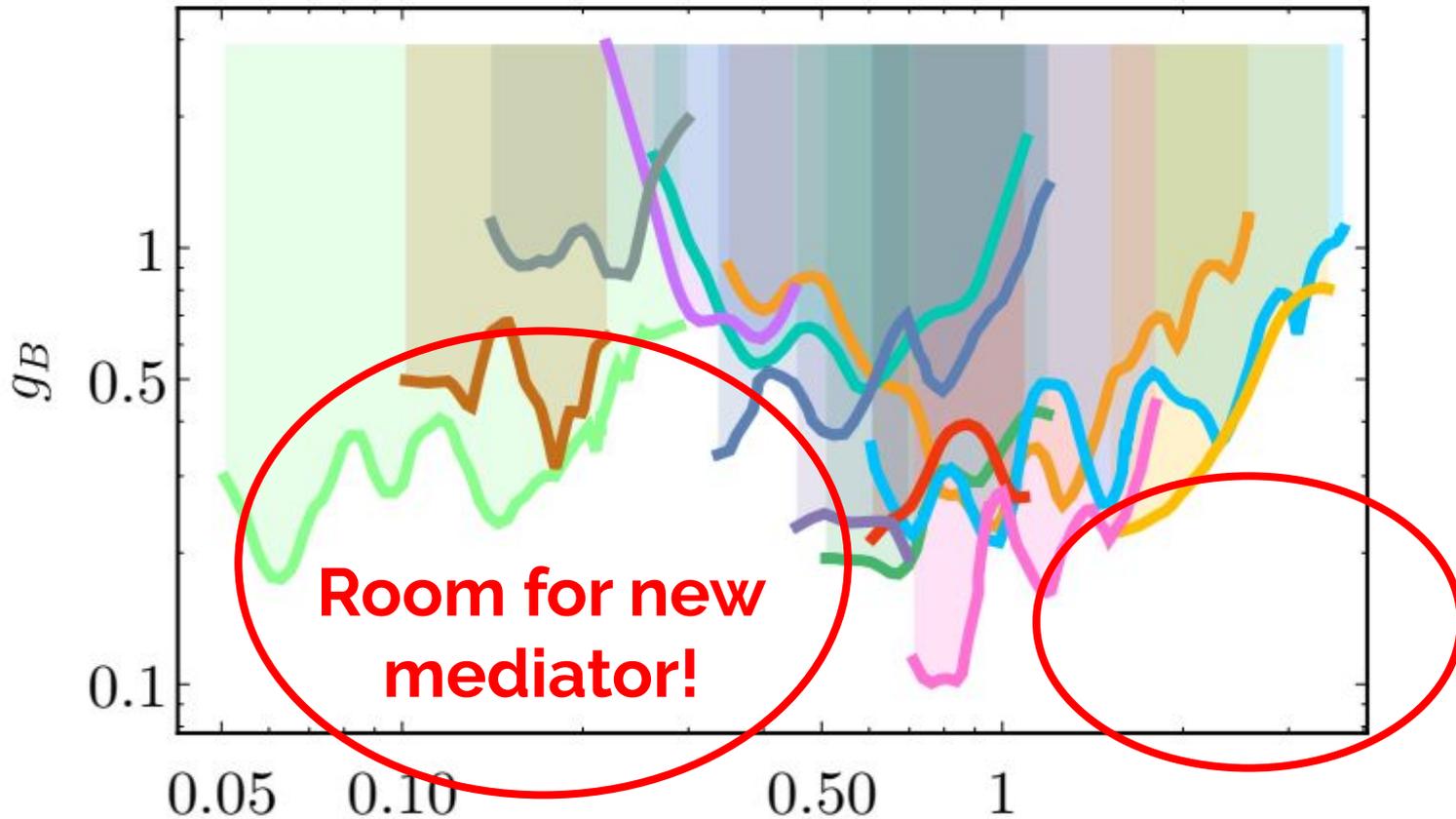
- No LEP bound for this scenario
- Di-jet searches at CMS and ATLAS - Run I & II



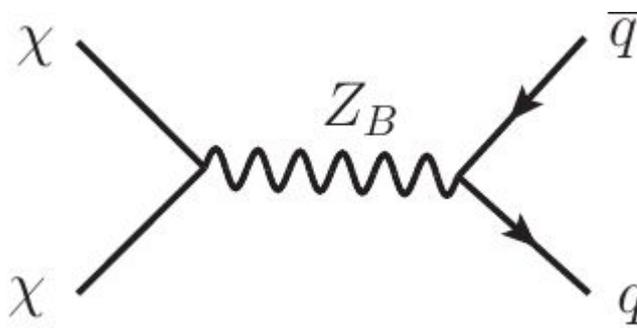
M_{Z_B} [TeV] [Fileviez Perez, Goliás, Li, Murgui 2018]

LHC bounds on leptophobic gauge boson

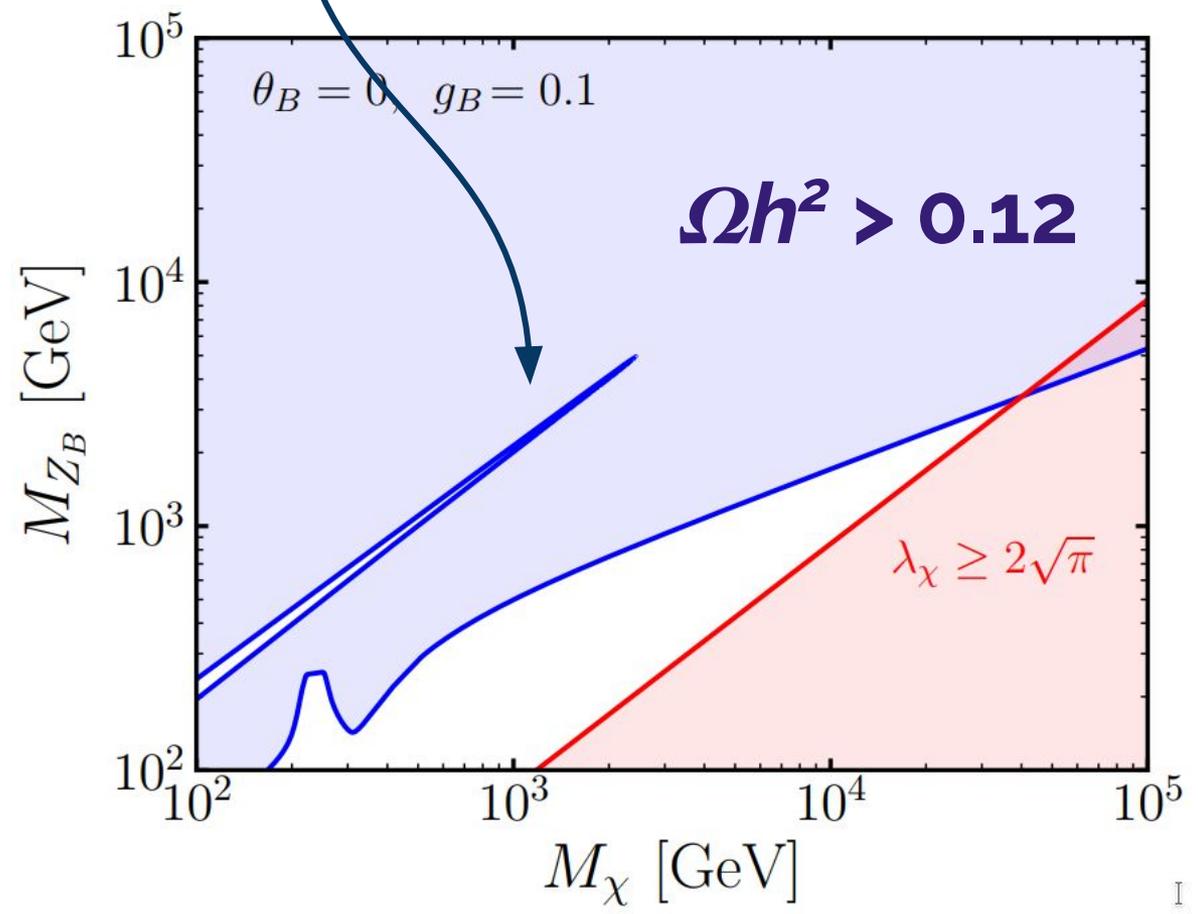
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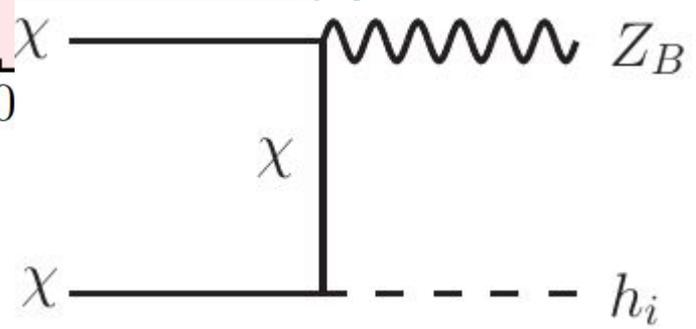
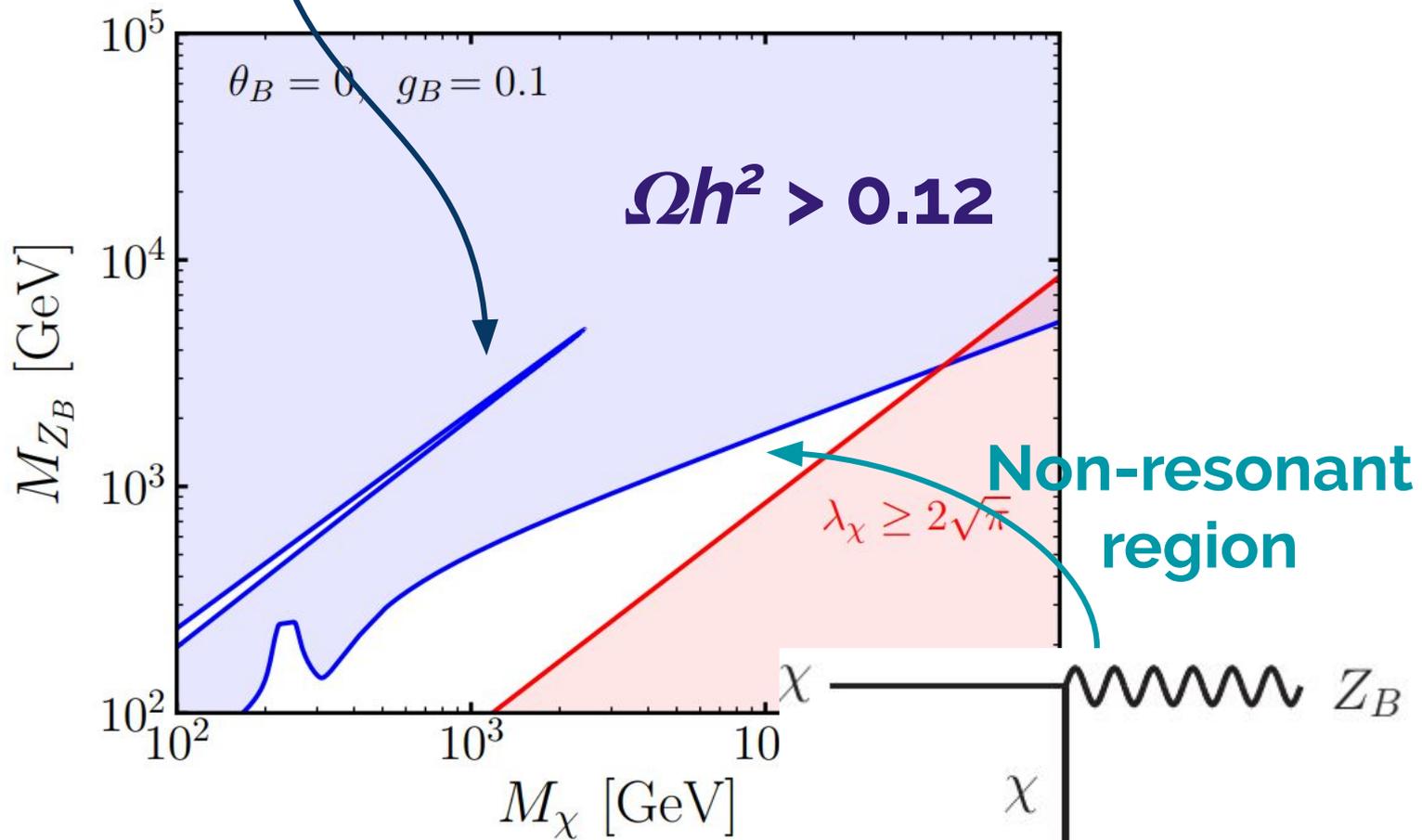
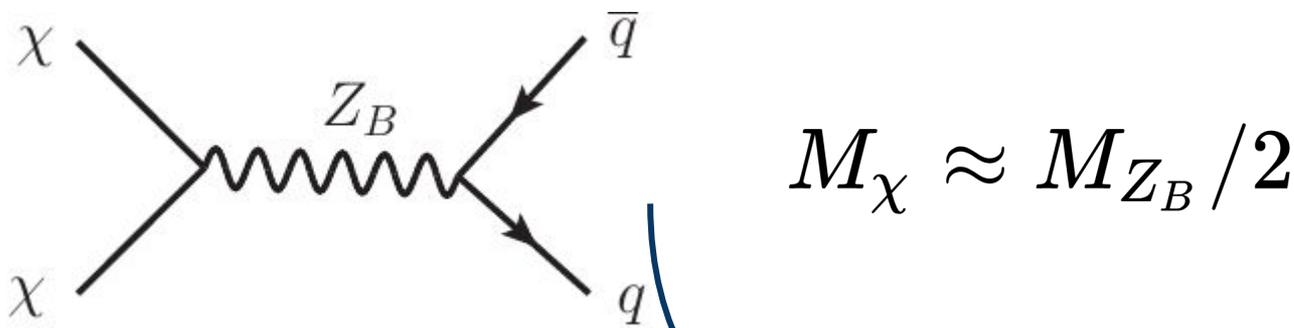


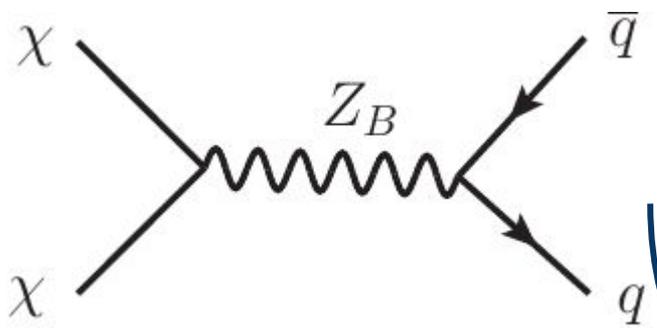
M_{Z_B} [TeV] [Fileviez Perez, Goliás, Li, Murgui 2018]



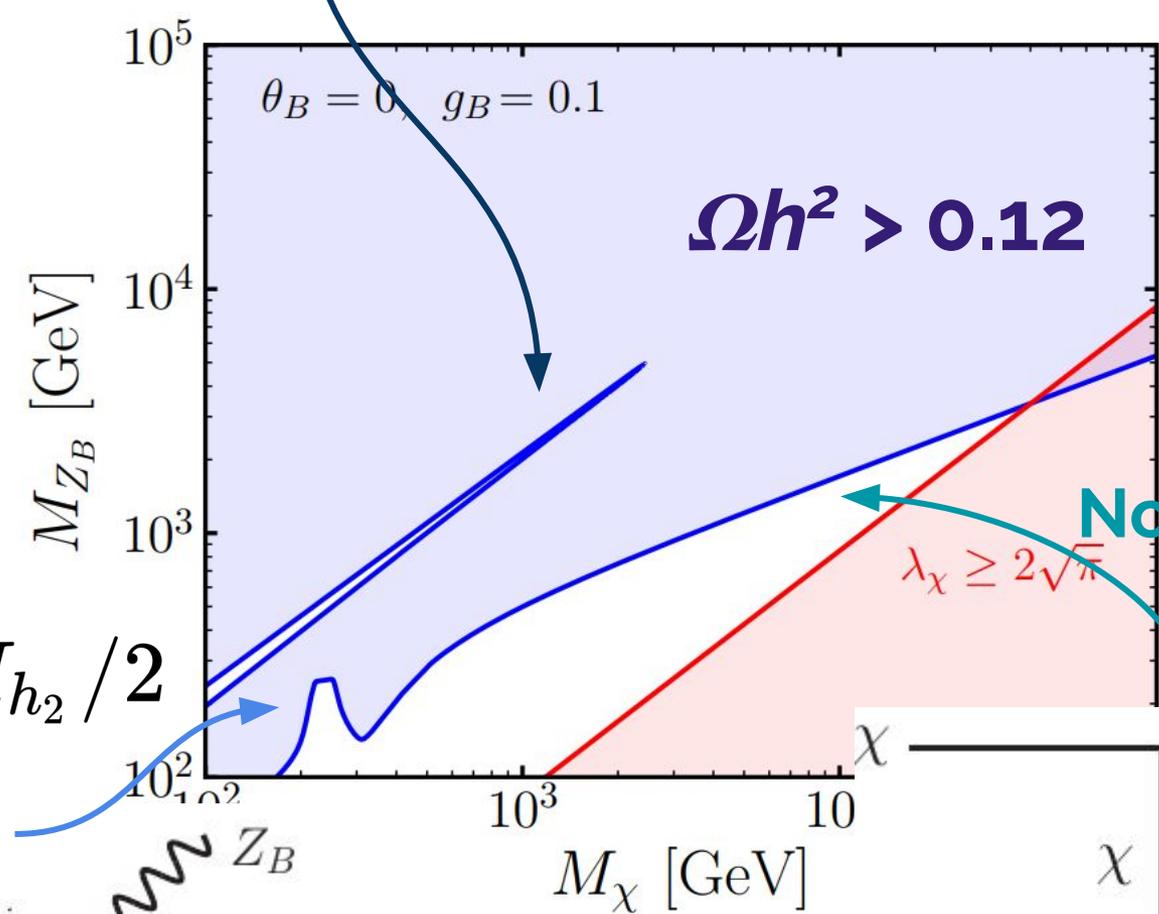
$$M_\chi \approx M_{Z_B} / 2$$



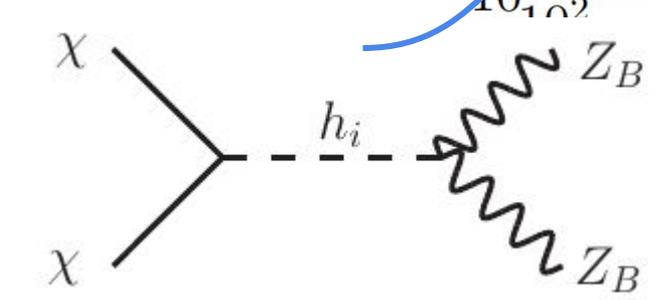




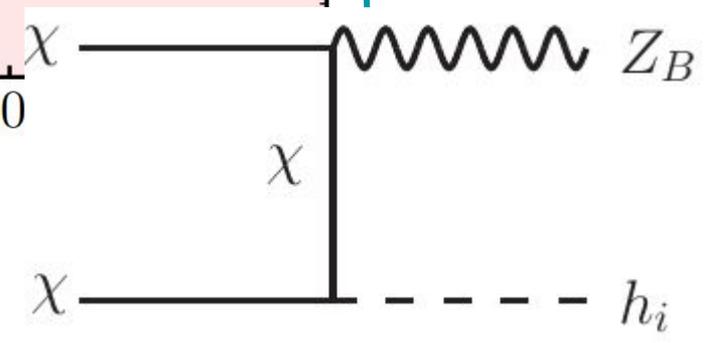
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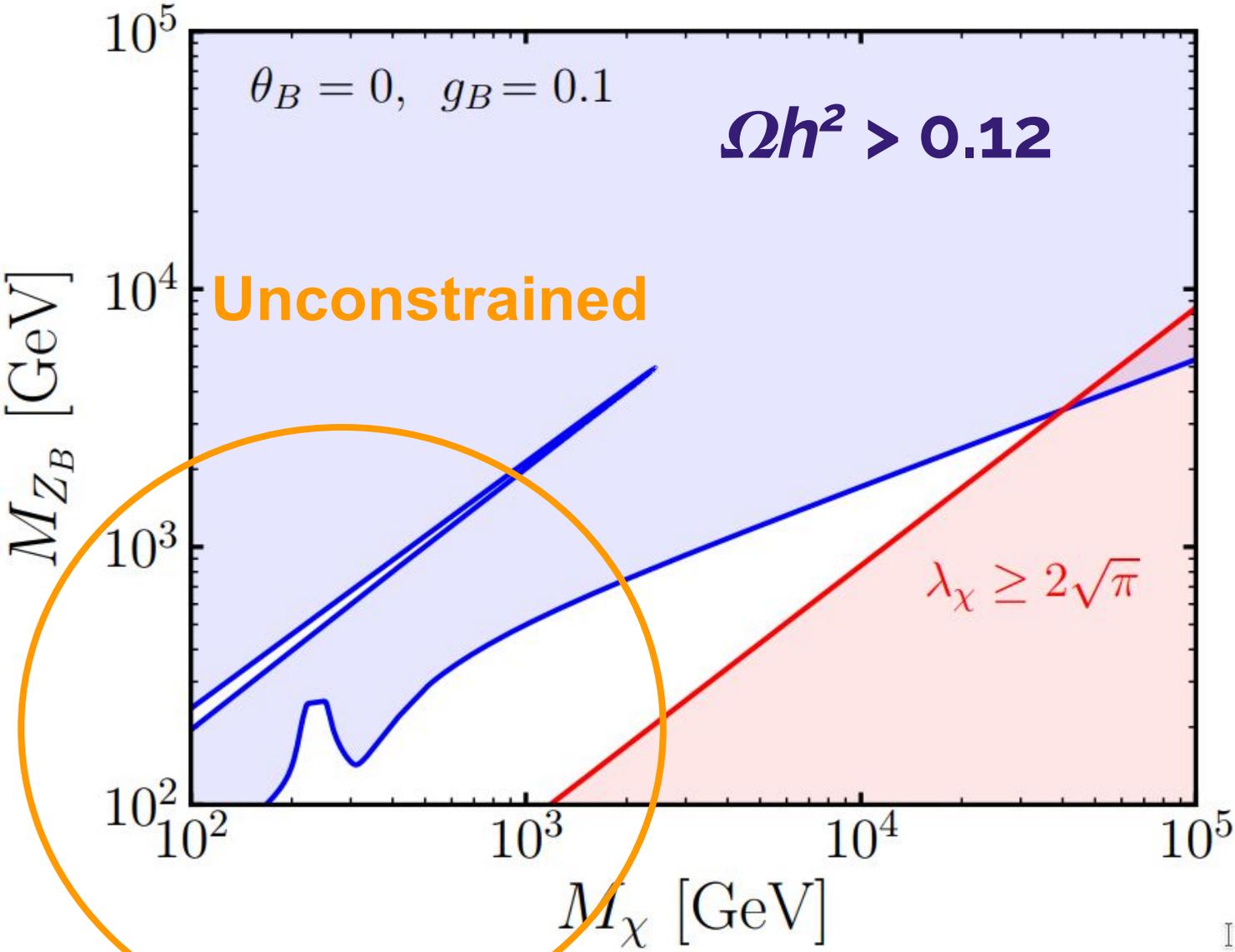
$$M_\chi \approx M_{h_2} / 2$$



Non-resonant region

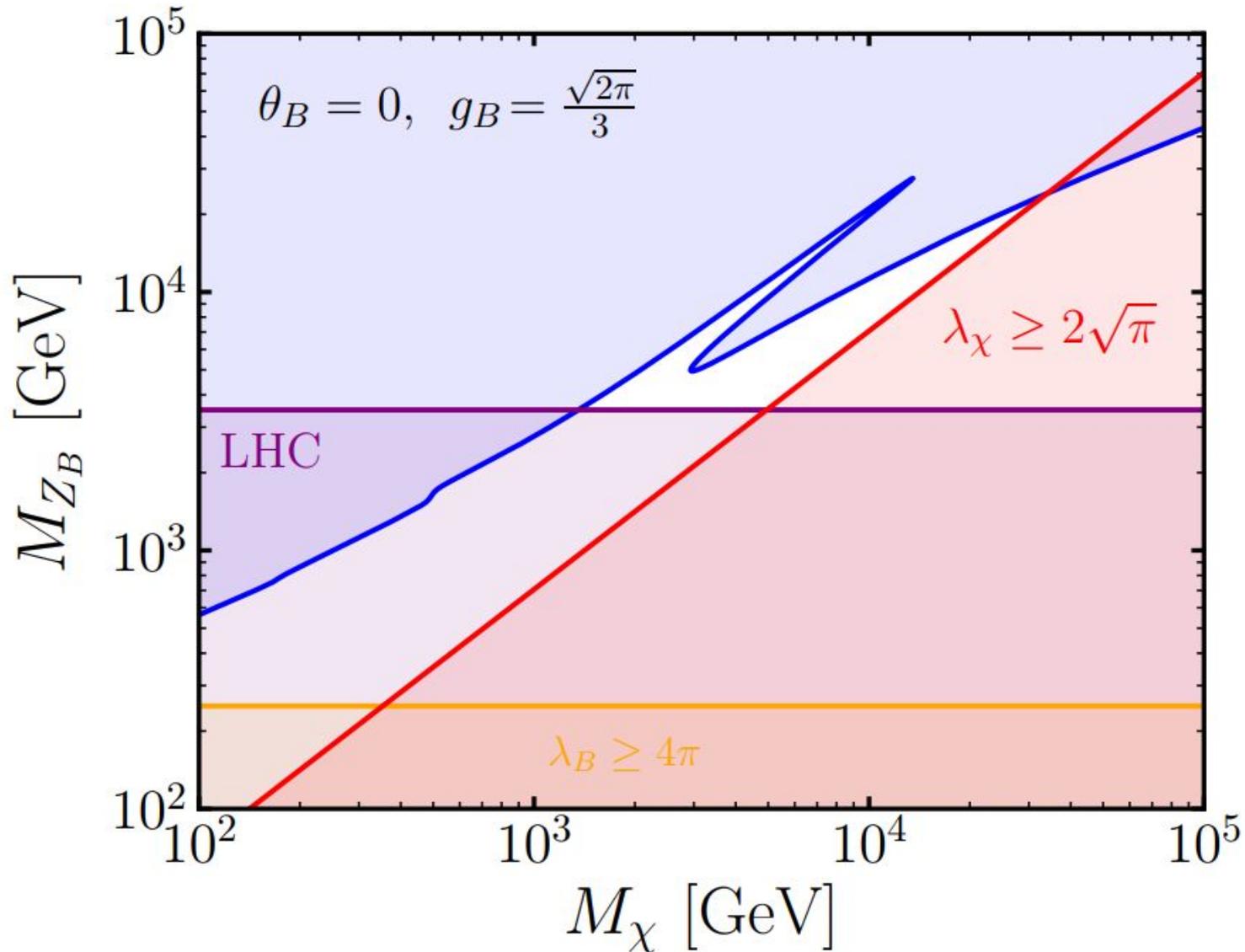


Results



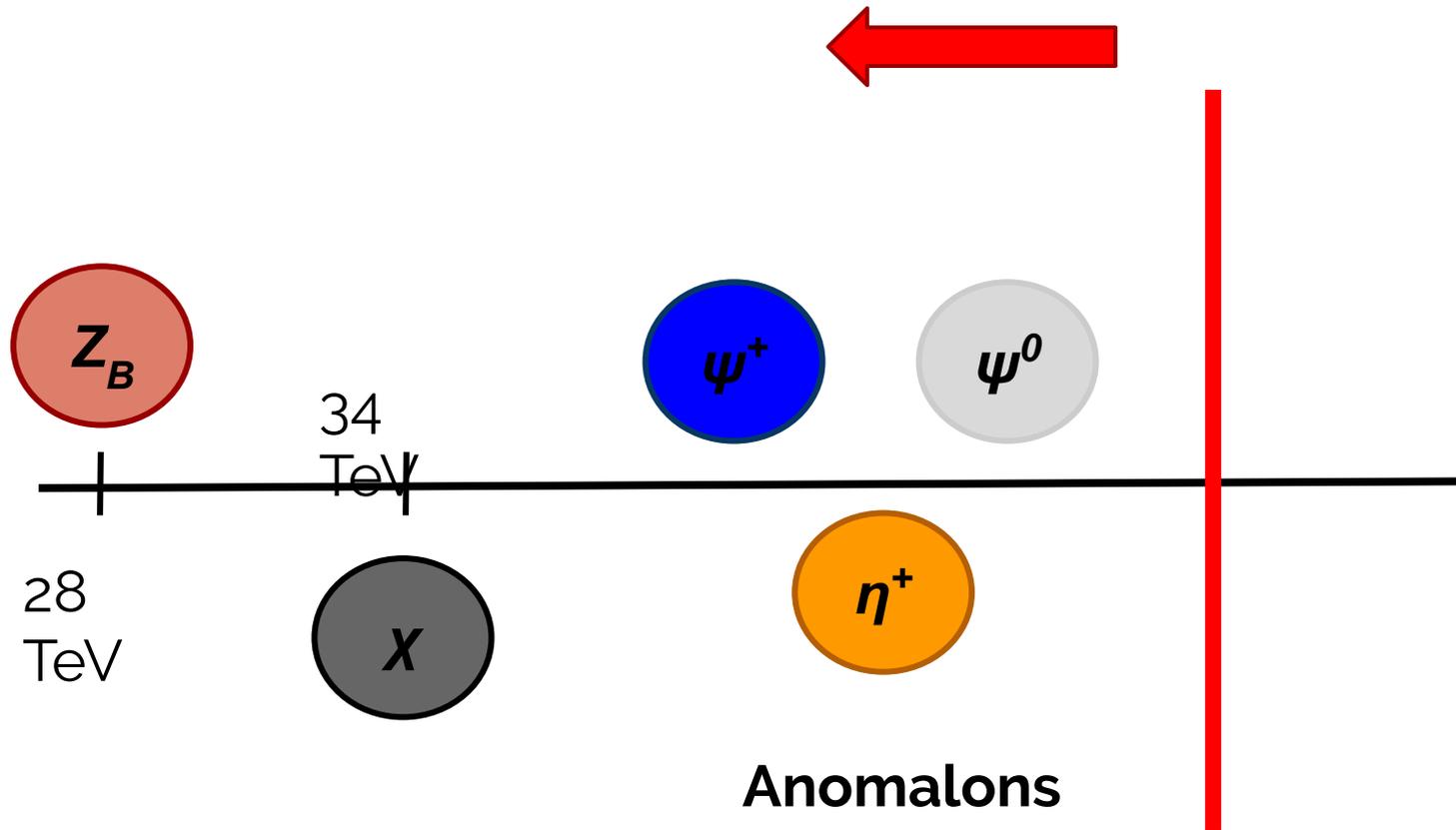
Perturbativity $g_B \leq \frac{\sqrt{2\pi}}{3} \approx 0.84$ and $\Omega h^2 \leq 0.12$

 Give an upper bound on the scale



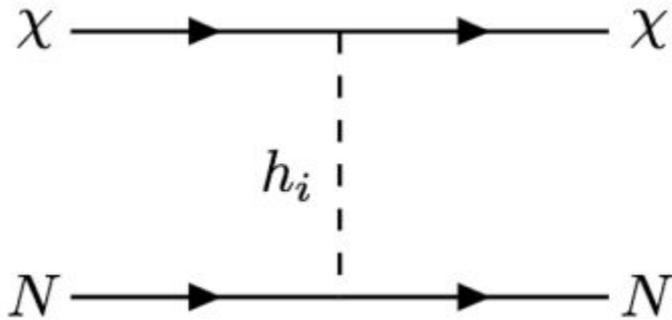
Upper bound on baryon number breaking scale

All masses connected to v_B and hence there is an upper bound for the full model



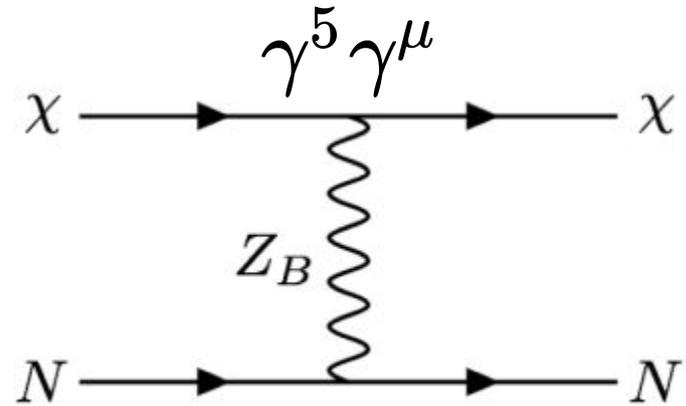
Direct Detection

$$\sigma_{\chi N}^{\text{TOT}} = \sigma_{\chi N}(h_i) + \sigma_{\chi N}^0(Z_B)v^2$$



Higgs mixing suppressed

$$\theta < 0.3 \quad \text{for } M_{H_2} > 200 \text{ GeV}$$



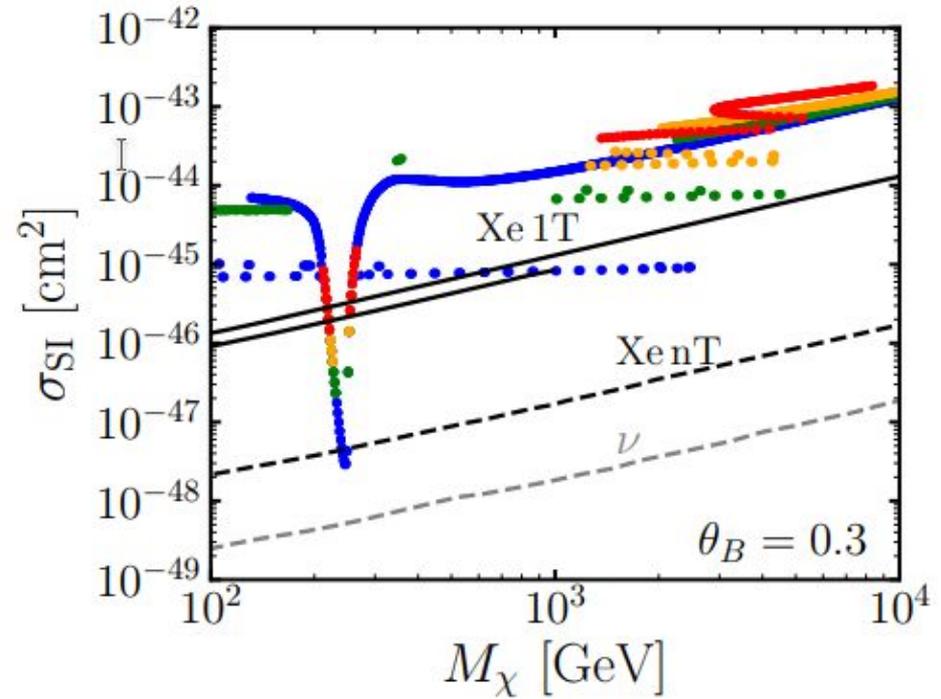
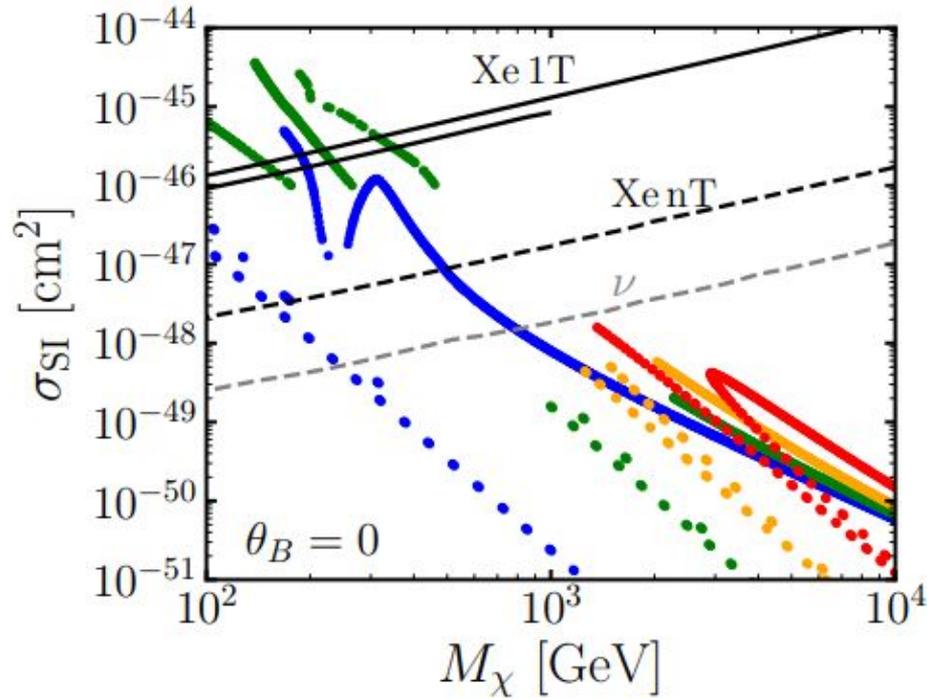
Due to axial coupling,

velocity suppressed $v \sim 10^{-3}$

[Inicka, Robens, Stefaniak 2018]

Direct detection constraints can be avoided

Direct Detection



Gauging baryon number

- Baryon number is an accidental global symmetry in the SM
- Only broken by non-perturbative effects - SU(2) instantons
- Spontaneous breaking

$$\underbrace{U(1)_B}$$

Local gauge symmetry

gauge boson: Z_B

$$\langle S_B \rangle \neq 0$$

[Pais 1973]

[Fileviez Perez & Wise 2011]

Anomaly cancellation

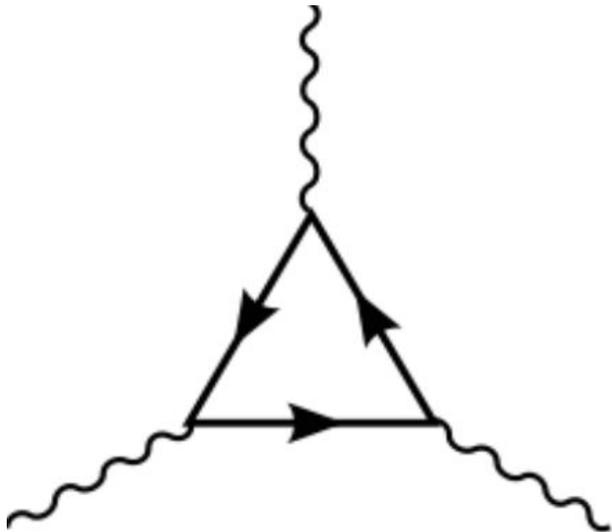
- Baryon number broken by 3 units: $\Delta B = \pm 3$ interactions

➔ No proton decay

- Need to add new fermions to cancel anomalies

Neutral fermion required for anomaly cancellation

➔ DM Candidate 😊



$$\mathcal{A}_1 (SU(3)^2 \otimes U(1)_B), \mathcal{A}_2 (SU(2)^2 \otimes U(1)_B), \\ \mathcal{A}_3 (U(1)_Y^2 \otimes U(1)_B), \mathcal{A}_4 (U(1)_Y \otimes U(1)_B^2), \\ \mathcal{A}_5 (U(1)_B), \mathcal{A}_6 (U(1)_B^3).$$

In the SM the non-zero values are:

$$\mathcal{A}_2 = -\mathcal{A}_3 = 3/2$$

Anomaly cancellation

[Duerr, Fileviez Perez, Wise 2013]

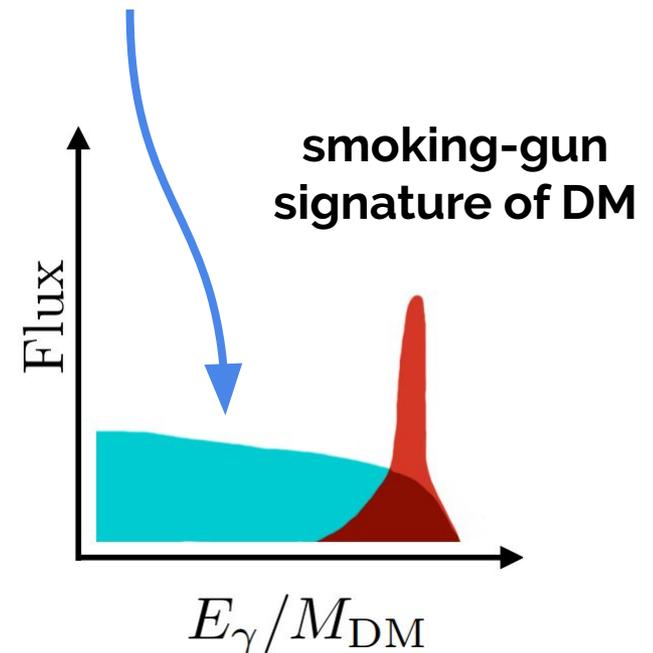
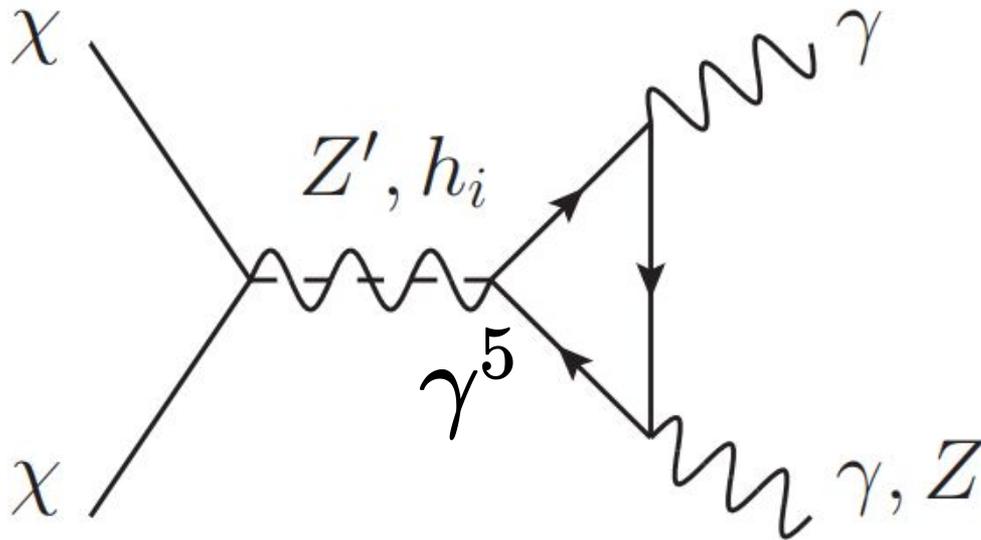
Fields	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_B$
$\Psi_L = \begin{pmatrix} \Psi_L^0 \\ \Psi_L^- \end{pmatrix}$	1	2	$-\frac{1}{2}$	$-\frac{3}{2}$
$\Psi_R = \begin{pmatrix} \Psi_R^0 \\ \Psi_R^- \end{pmatrix}$	1	2	$-\frac{1}{2}$	$\frac{3}{2}$
η_R^-	1	1	-1	$-\frac{3}{2}$
η_L^-	1	1	-1	$\frac{3}{2}$
χ_R^0	1	1	0	$-\frac{3}{2}$
χ_L^0	1	1	0	$\frac{3}{2}$

DM

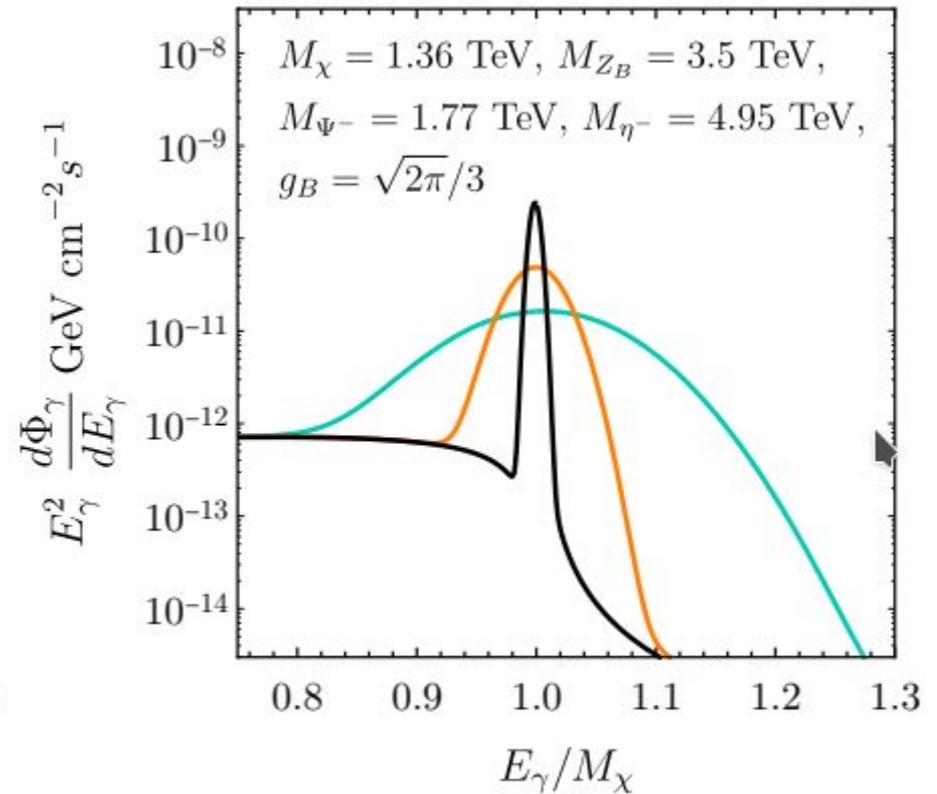
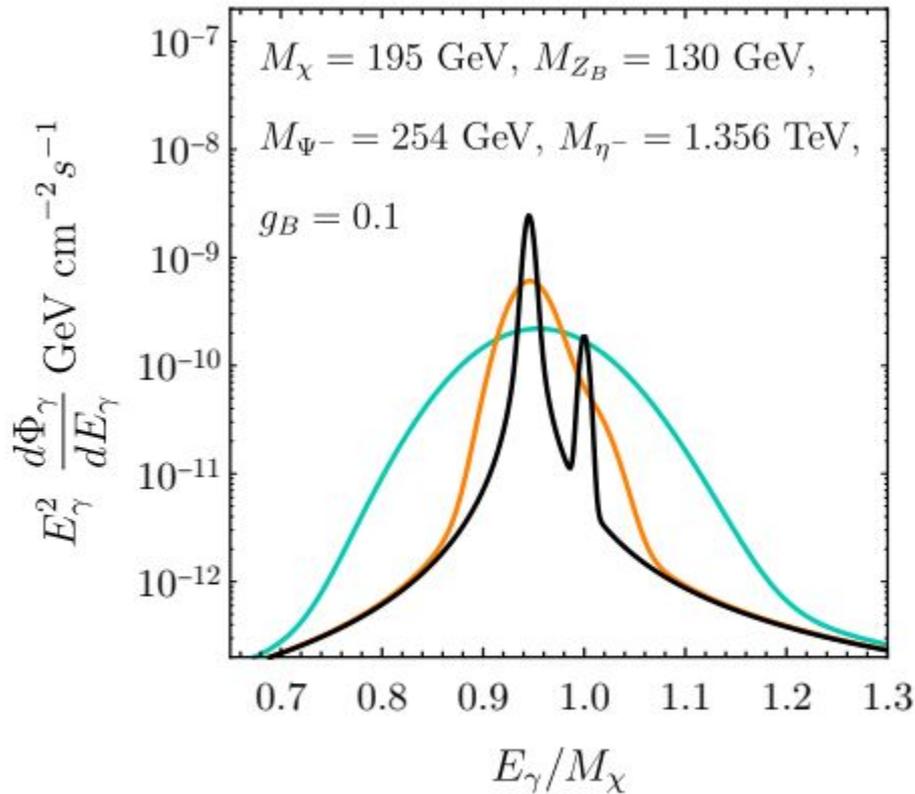
**For Model II see
[Ohmer, Fileviez Perez, Patel 2014]**

Gamma lines

- Annihilation into $\gamma\gamma$ possible. New fermions required for anomaly cancellation in the loop.
- Peak at $E = M_{DM}$ in the gamma spectrum
- Continuum is velocity suppressed, because of axial coupling

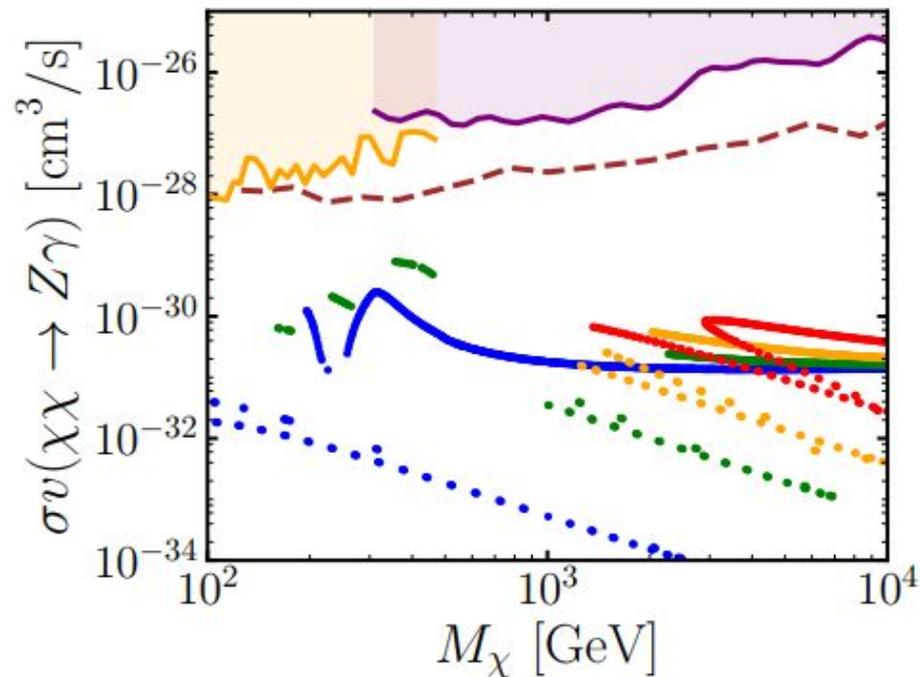
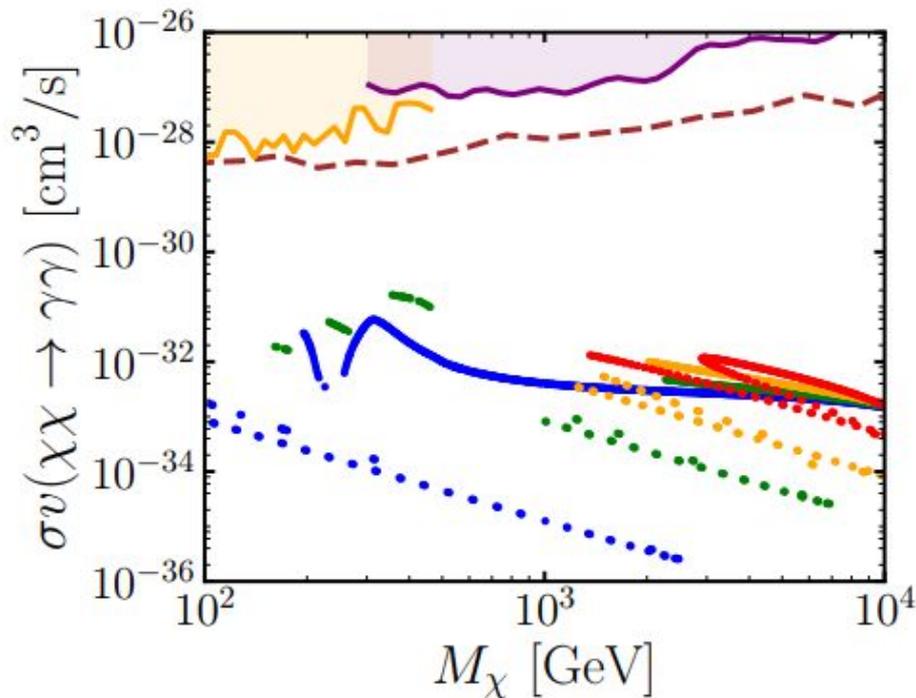


Gamma lines



— $\xi = 0.15$
 — $\xi = 0.05$
 — $\xi = 0.01$

Gamma lines



Conclusions

- ✓ Consistent completions of simplified models of dark matter give rise to interesting phenomenology
- ✓ DM predicted from anomaly cancellation and stable from remnant $U(1)_B \rightarrow Z_2$ symmetry
- ✓ Direct detection is velocity suppressed, Xenon nT will probe Higgs portal interaction
- ✓ Leptophobic gauge boson Z_B can appear as resonance in dijet searches at LHC
- ✓ Final state radiation velocity suppressed \rightarrow gamma-ray lines can be observed in future

Back-up

Model II

Fields	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_B$
$\Psi_L = \begin{pmatrix} \Psi_L^+ \\ \Psi_L^0 \end{pmatrix}$	1	2	$\frac{1}{2}$	$\frac{3}{2}$
$\Psi_R = \begin{pmatrix} \Psi_R^+ \\ \Psi_R^0 \end{pmatrix}$	1	2	$\frac{1}{2}$	$-\frac{3}{2}$
$\Sigma_L = \frac{1}{\sqrt{2}} \begin{pmatrix} \Sigma_L^0 & \sqrt{2}\Sigma_L^+ \\ \sqrt{2}\Sigma_L^- & -\Sigma_L^0 \end{pmatrix}$	1	3	0	$-\frac{3}{2}$
χ_L^0	1	1	0	$-\frac{3}{2}$