

Neutrino mass: The origins

Miha Nemevšek

~ JSI Ljubljana ~

w. Maiezza, Nesti, Popara, Senjanović, Tello, Vasquez, Zhang

PITT PACC NuTheories workshop

November 2018

Fermion mass

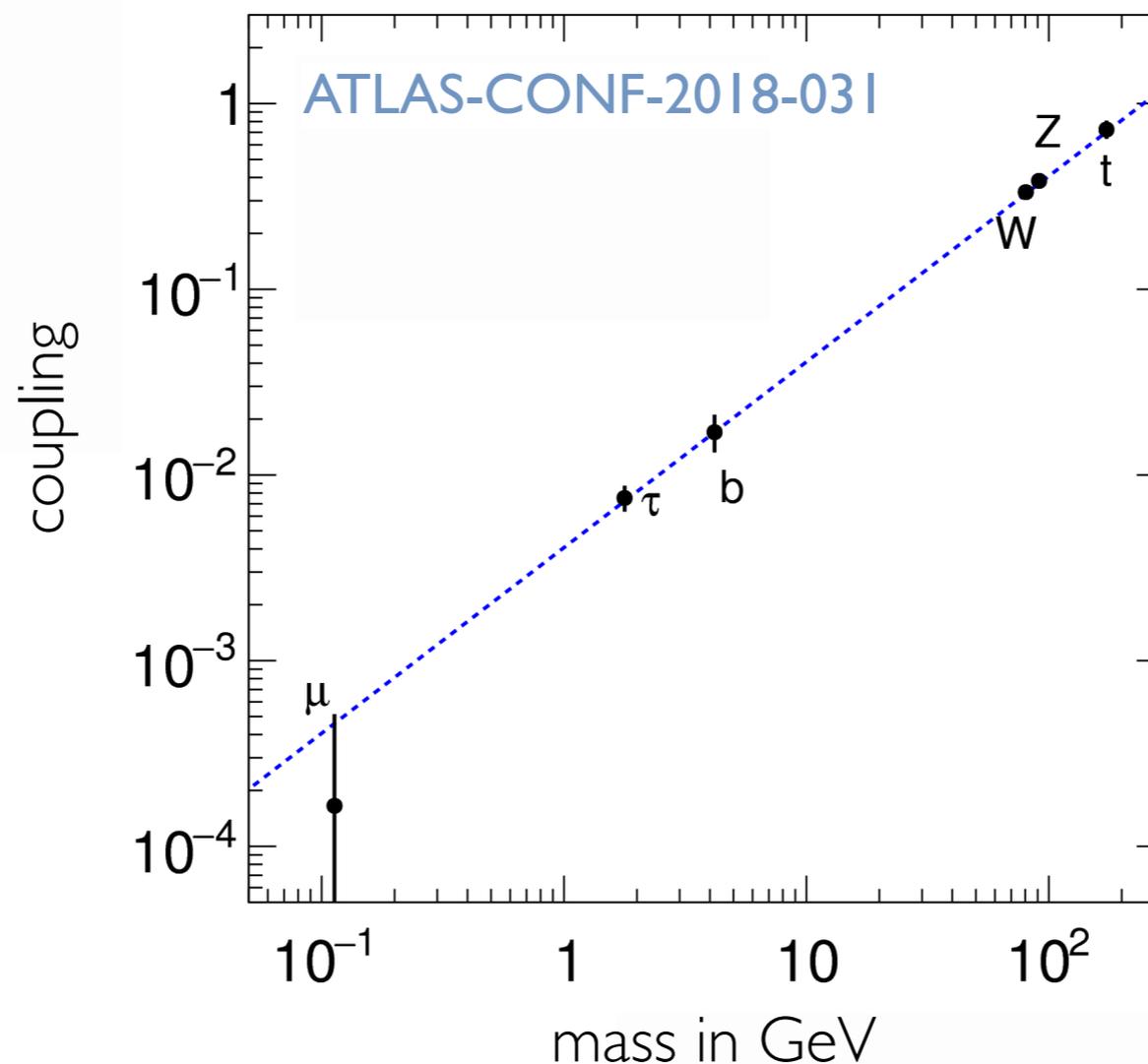
Origin of fermion mass from SSB, e.g. the $SU(2)_L \times U(1)_Y$

Weinberg '67

Dirac term only

$$\mathcal{L}_f = y_\ell \bar{L} H \ell_R$$

$$\Gamma_{h \rightarrow \ell \bar{\ell}} \propto y_\ell^2$$



EXP: all masses come from SSB

$$m_\ell = y_\ell v$$

Neutrinos

Massless by fiat in SM, does nature abhor gauge singlets?

Gauge extension motivations

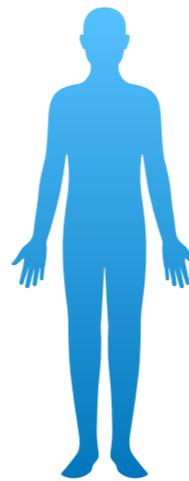


unification of forces

$$SU(5), SO(10), E_6, \dots$$

Glashow '79

Gell-Mann et al. '79,...

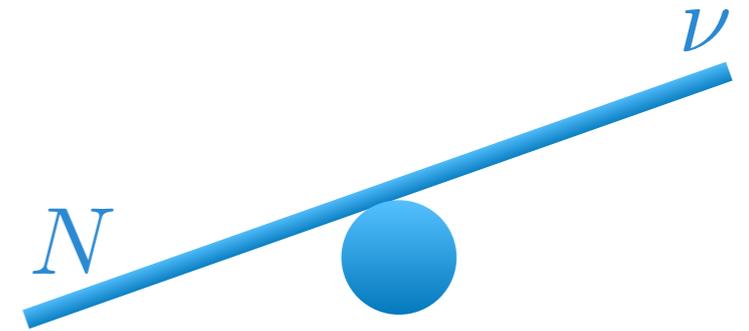


weak force asymmetry

$$SU(4)_c \times SU(2)_L \times SU(2)_R$$

$$SU(3)^3$$

$$SU(2)_L \times SU(2)_R \times U(1)_{B-L}$$



minimal gauged seesaw

see Pavel's talk

$$U(1)_{B-L}$$

Origin of neutrino mass

Not as simple as charged fermions - enter Majorana - two sources

$$\frac{M_D}{v} \bar{L} H N + M_N N^T C N$$

Majorana '37

$$M_\nu = M_D^T M_N^{-1} M_D$$

Origin of neutrino mass

Not as simple as charged fermions - enter Majorana - two sources

oscillations,
 $0\nu 2\beta$, endpoint,
cosmology,
 N_{eff} , CMB,
...

$$\frac{M_D}{v} \bar{L} H N + M_N N^T C N$$

$$M_\nu = M_D^T M_N^{-1} M_D$$

Origin of neutrino mass

Not as simple as charged fermions - enter Majorana - two sources

$$\frac{M_D}{v} \bar{L} H N + M_N N^T C N$$

$$M_\nu = M_D^T M_N^{-1} M_D$$

colliders,
meson decays,
warm DM,
leptogenesis,
...

Origin of neutrino mass

Not as simple as charged fermions - enter Majorana - two sources

$$\frac{M_D}{v} \bar{L} H N + M_N N^T C N$$

$$\begin{aligned} M_\nu &= M_D^T M_N^{-1} M_D \\ &= M_D^T m_N^{-1/2} \underbrace{m_N^{-1/2} M_D}_{OS} \end{aligned} \quad O^T O = 1,$$

Even so - connection to M_D is ambiguous

$$S^T = S, \quad S = \sqrt{M_\nu}$$

non-unitarity,
colliders, meson
decays, $0\nu 2\beta$,
eEDMs,
...

$$M_D = m_N^{-1/2} O \sqrt{M_\nu}$$

Finding the origins

Gauge extensions: additional W's, Z's and Higgses



unification

$$N \ni 16_F, 27_F, \dots$$

$$p\text{-decay: } M_{GUT} \gtrsim 10^{16} \text{ GeV}$$

Finding the origins

Gauge extensions: additional W's, Z's and Higgses



unification

$$N \ni 16_F, 27_F, \dots$$

$$p\text{-decay: } M_{GUT} \gtrsim 10^{16} \text{ GeV}$$

Pati, Salam '75



partial unification
Left-Right

$$L_R = \begin{pmatrix} N \\ \ell_R \end{pmatrix}$$

$$K\text{-decay: } M_{PS} \gtrsim 10^8 \text{ GeV}$$

Finding the origins

Gauge extensions: additional W's, Z's and Higgses



unification

$$N \ni 16_F, 27_F, \dots$$

p -decay : $M_{GUT} \gtrsim 10^{16} \text{ GeV}$

Pati, Salam '75



partial unification
Left-Right

$$L_R = \begin{pmatrix} N \\ \ell_R \end{pmatrix}$$

K -decay : $M_{PS} \gtrsim 10^8 \text{ GeV}$

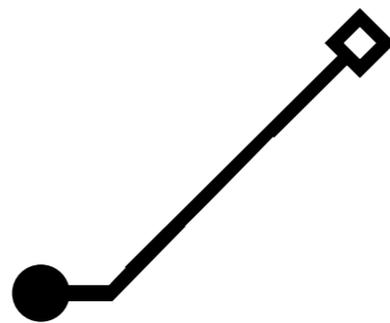
K & B oscillations :

$$M_{W_R} \gtrsim 3 - 4 \text{ TeV}$$

Senjanović, Mohapatra '79,
Beall, Bander, Soni '82, many ...
Bertolini, Nesti, Maiezza '14

mLRSM : flavor fixed

$$V_R^q \simeq V_L^q$$



$nEDM$, strong P breaking

Maiezza, MN '14

Finding the origins

Gauge extensions: additional W's, Z's and Higgses



unification

$$N \ni 16_F, 27_F, \dots$$

$$p\text{-decay} : M_{GUT} \gtrsim 10^{16} \text{ GeV}$$

Pati, Salam '75



partial unification
Left-Right

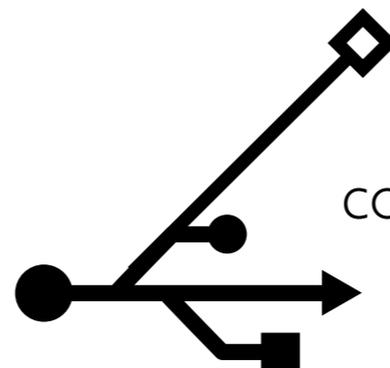
$$L_R = \begin{pmatrix} N \\ \ell_R \end{pmatrix}$$

$$K\text{-decay} : M_{PS} \gtrsim 10^8 \text{ GeV}$$

$$K \text{ \& } B \text{ oscillations} : M_{W_R} \gtrsim 3 - 4 \text{ TeV}$$

mLRSM : flavor fixed

$$V_R^q \simeq V_L^q$$



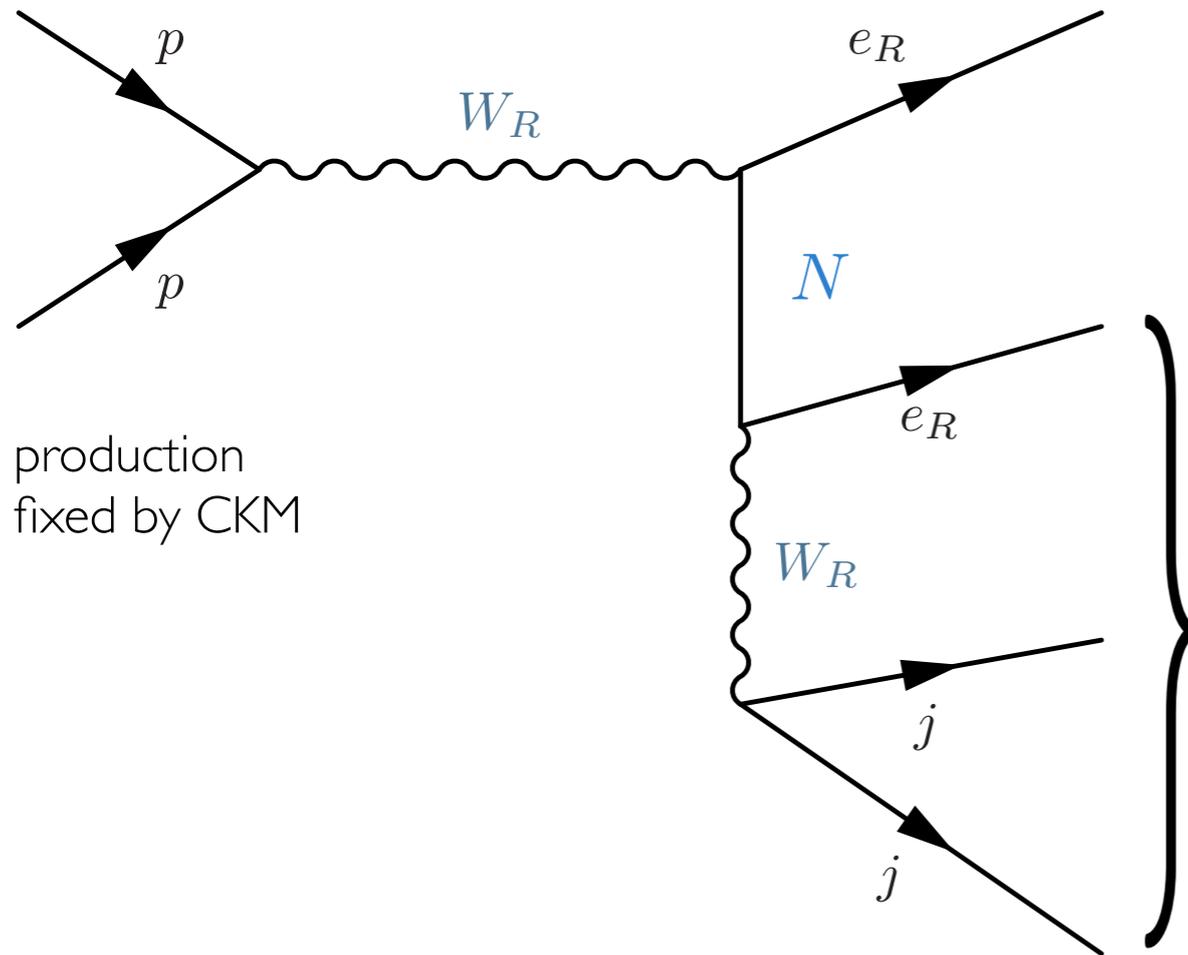
colliders

$$0\nu 2\beta$$

eEDM, wDM, ...

Colliders

Keung, Senjanović '83



Main feature: Lepton Number Violation

On-shell Majorana fermion

$N \rightarrow \ell^\pm jj$ 50-50% same-opposite sign

$$m_{\ell jj} = m_{N_i}$$

narrow mass peaks for $m_N < M_{W_R}$

~no missing energy

flavor states measure V_R^ℓ (free)

more on the Majorana nature

Gluza, Jelinski '15 '16
Das, Dev, Mohapatra '17

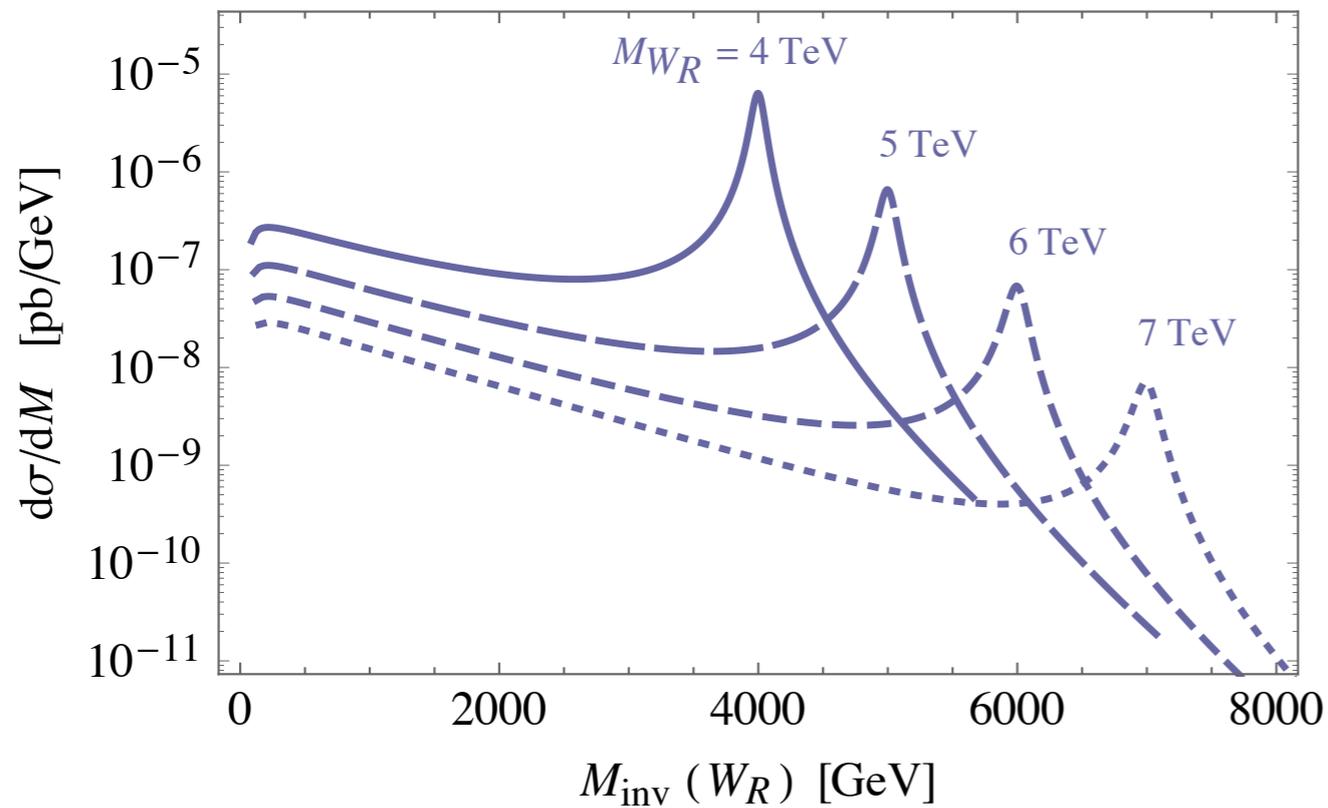
$$V_R^{\ell T} M_N V_R^\ell = m_N$$

Golden channel: $pp \rightarrow W_R \rightarrow \ell_R N$

Keung, Senjanović '83

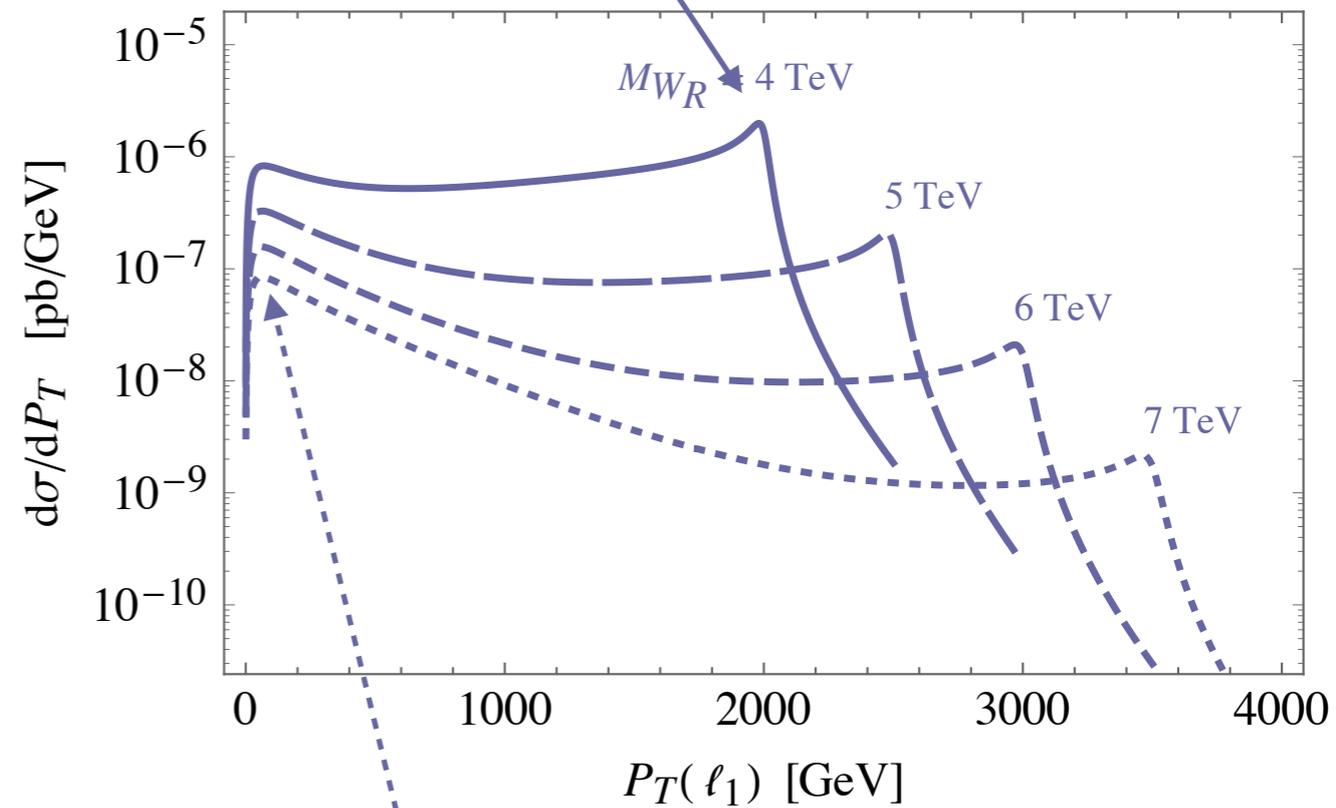
$$\hat{\sigma}_{ij}^{\ell N}(\hat{s}) = \frac{\alpha_2^2 \pi}{72 \hat{s}^2} |V_{ij}^{\text{CKM}}|^2 \frac{(\hat{s} - m_N^2)^2 (2\hat{s} + m_N^2)}{(\hat{s} - M_{W_R}^2)^2 + M_{W_R}^2 \Gamma_{W_R}^2}$$

clear peak



m_{inv} disappears

mostly on-shell, N boosted



off-shell = soft lepton and N

Ruiz '17

see Richard's talk

Sketch of a search for $pp \rightarrow W_R \rightarrow \ell_R N$

MN, Nesti, Senjanović, Zhang '11

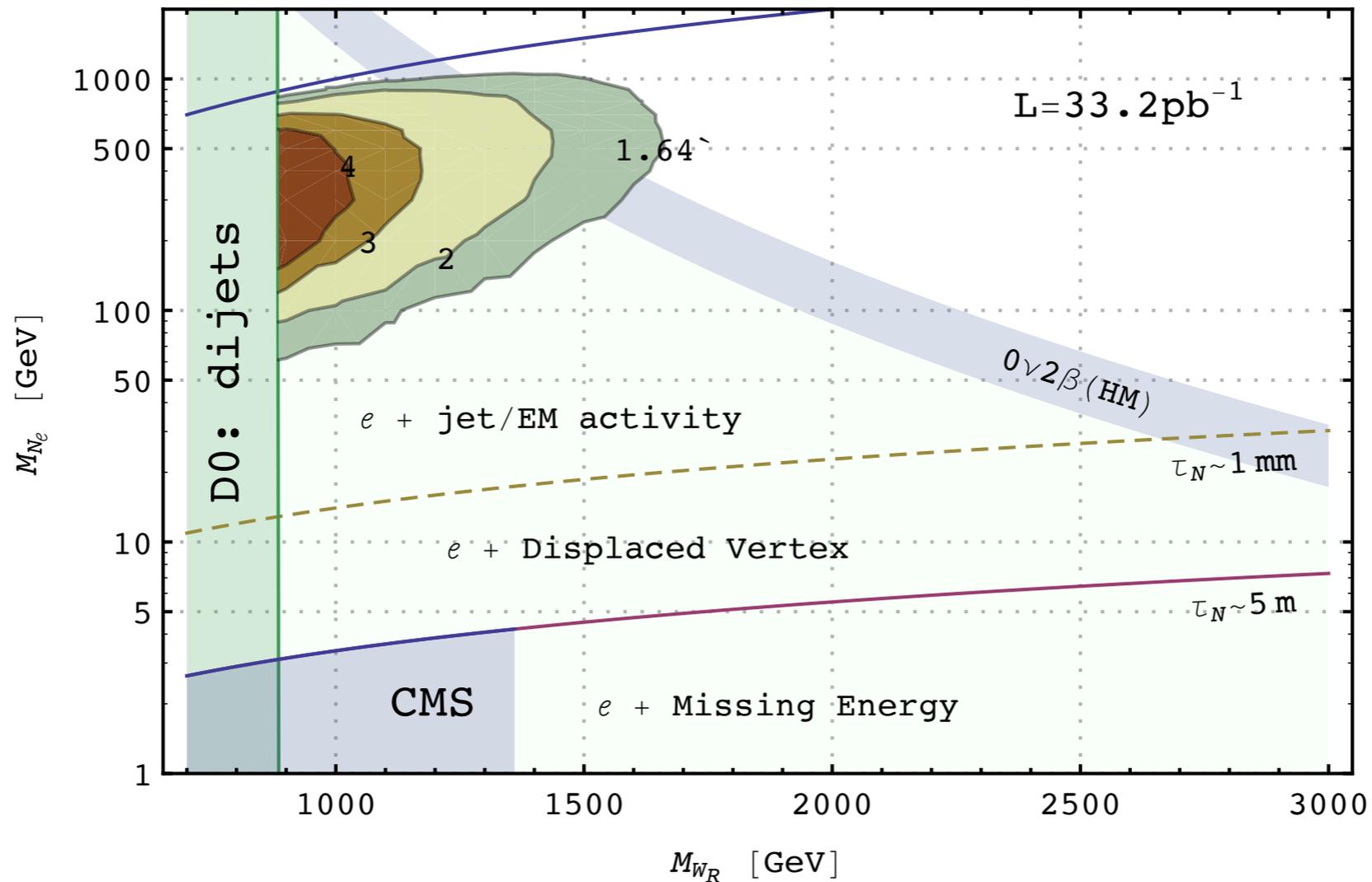
separated
eejj

merged
neutrino jet

Richard's talk

displaced jet

missing
energy



first LHC data,
low bound

LNv relation to
 $0\nu 2\beta$

Reach of 5-6 TeV at 14 TeV

ATLAS: Ferrari et al. '00
CMS: Gninenko et al. '07

Isolation and displacement $pp \rightarrow W_R \rightarrow \ell_R N$

MN, Nesti, Popara '18

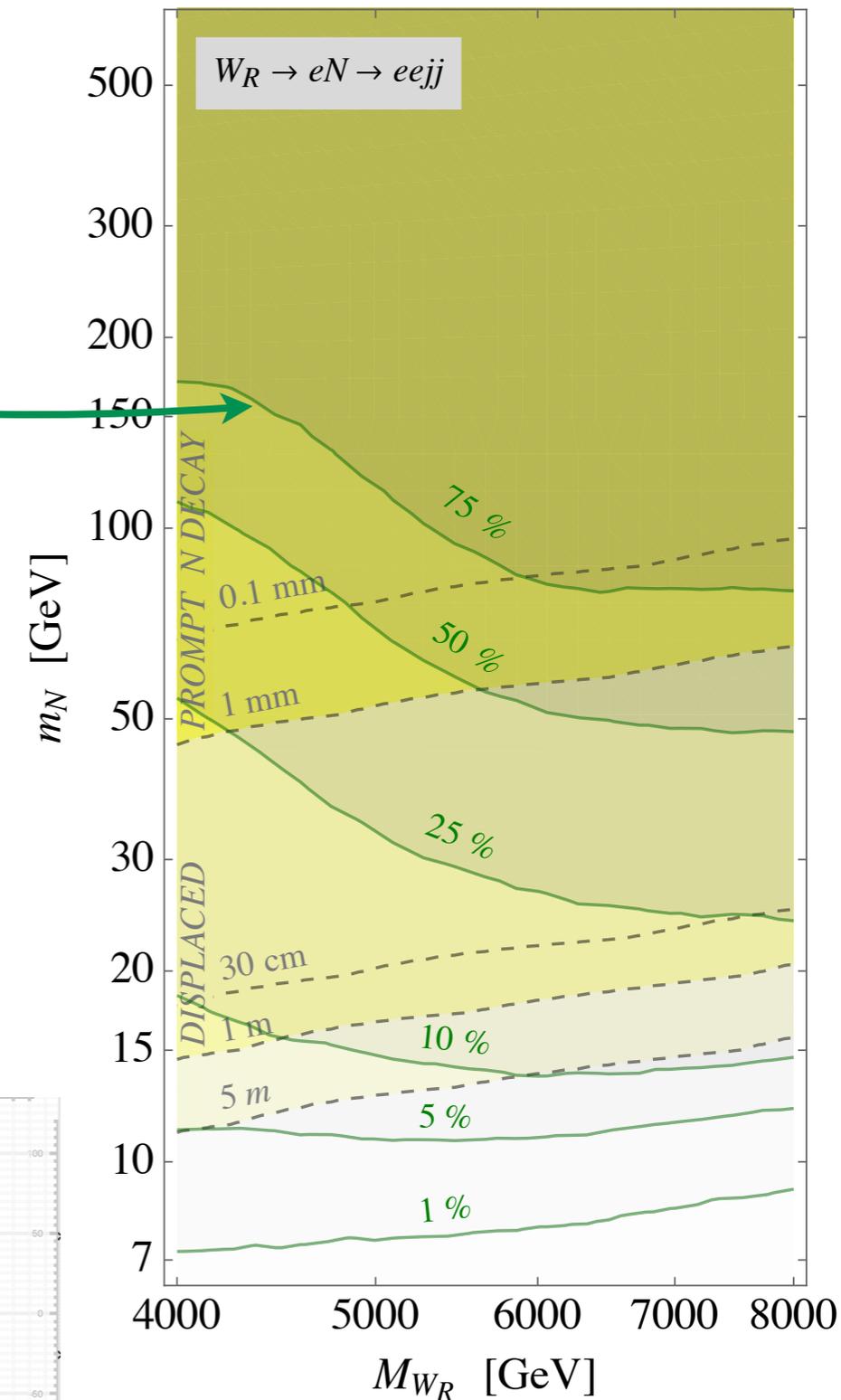
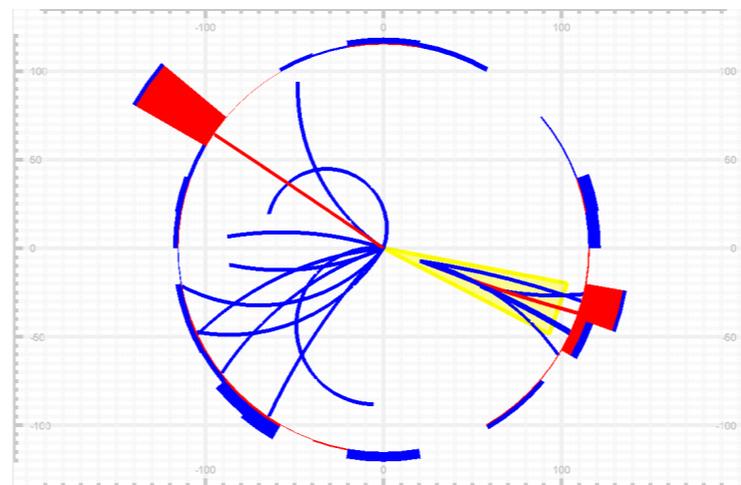
2nd lepton isolation depends on the boost of N

$$\gamma_N \simeq \begin{cases} \frac{M_{W_R}}{2m_N}, & W_R \rightarrow \text{on-shell}, \\ \frac{1 \text{ TeV}}{m_N}, & W_R \rightarrow \text{off-shell} \end{cases}$$

Lab decay length very sensitive to m_N

$$\Gamma_N^0 \simeq \frac{\alpha_2^2 m_N^5}{64\pi M_{W_R}^4} \simeq \frac{1}{2.5 \text{ mm}} \frac{(m_N/10 \text{ GeV})^5}{(M_{W_R}/3 \text{ TeV})^4}$$

Simultaneous transition from prompt isolated to displaced merged - look for displaced merged jets (tracks)



Displaced jet discrimination

MN, Nesti, Popara '18

Event generation: custom generator KSEG, small width issues with MG5

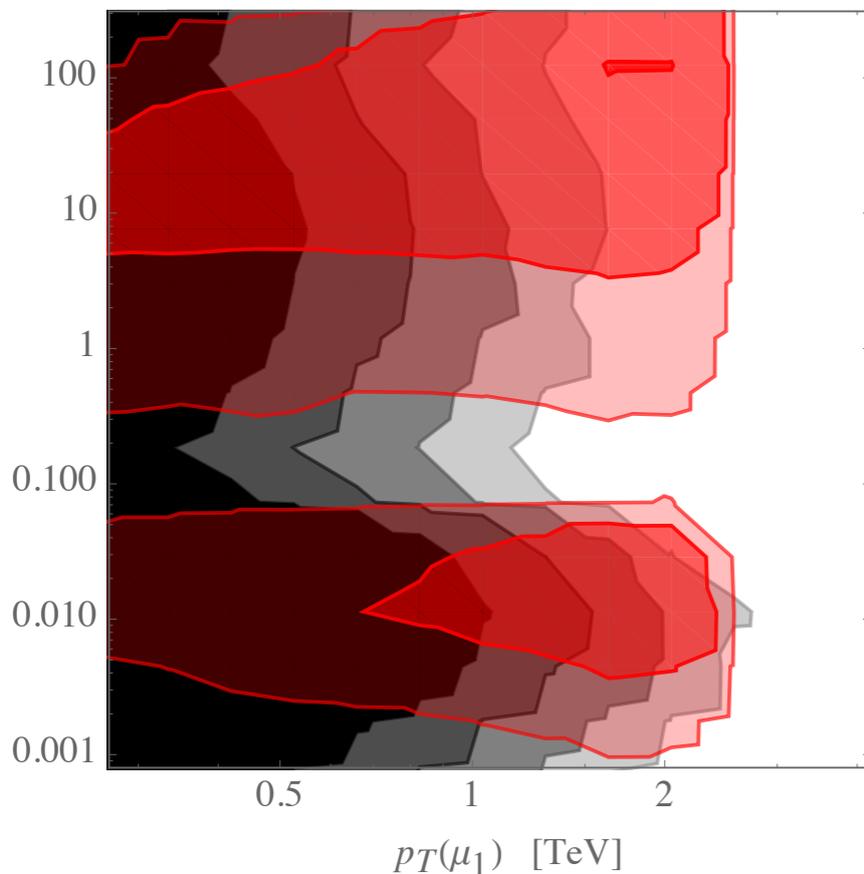
also Feynrules model file and Delphes,
Madanalysis displacement hack

sites.google.com/site/leftrighthep

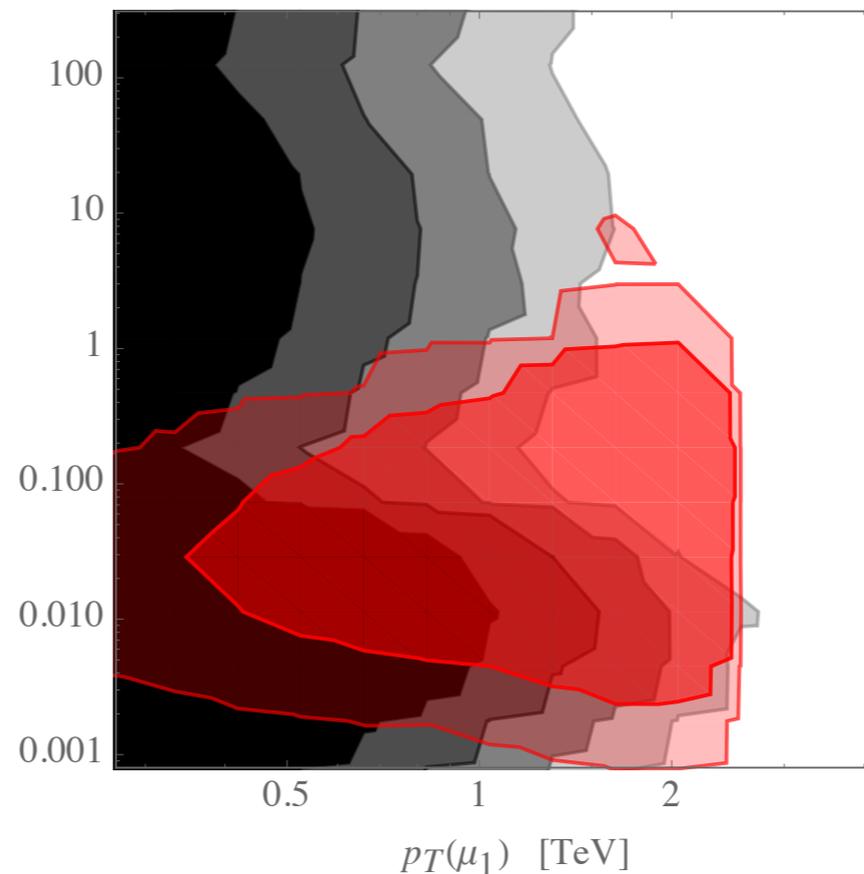
Main bckgs

background	# generator	weight	# detector
$V + 012j$	22.46 M	0.021	9.93M
$VV + 012j$	10.55 M	0.0028	4.61M
$t\bar{t} + 012j$	10.47 M	0.024	4.38M

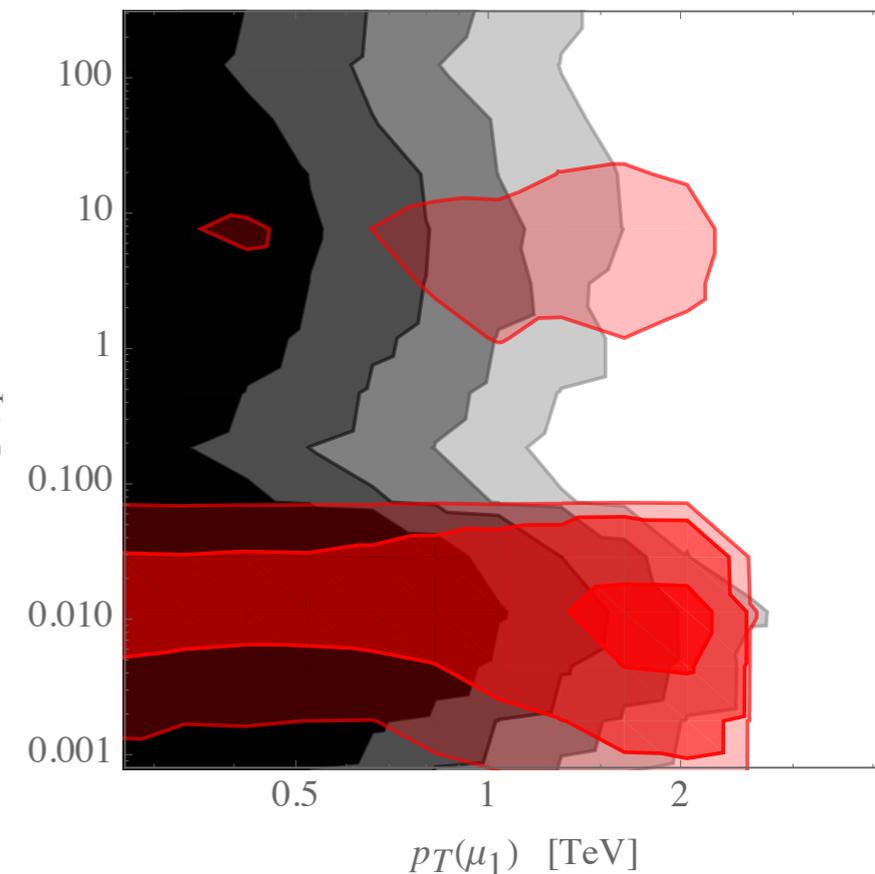
$M_{WR}=4$ TeV $m_N=20$ GeV



$M_{WR}=4$ TeV $m_N=60$ GeV



$M_{WR}=4$ TeV $m_N=150$ GeV



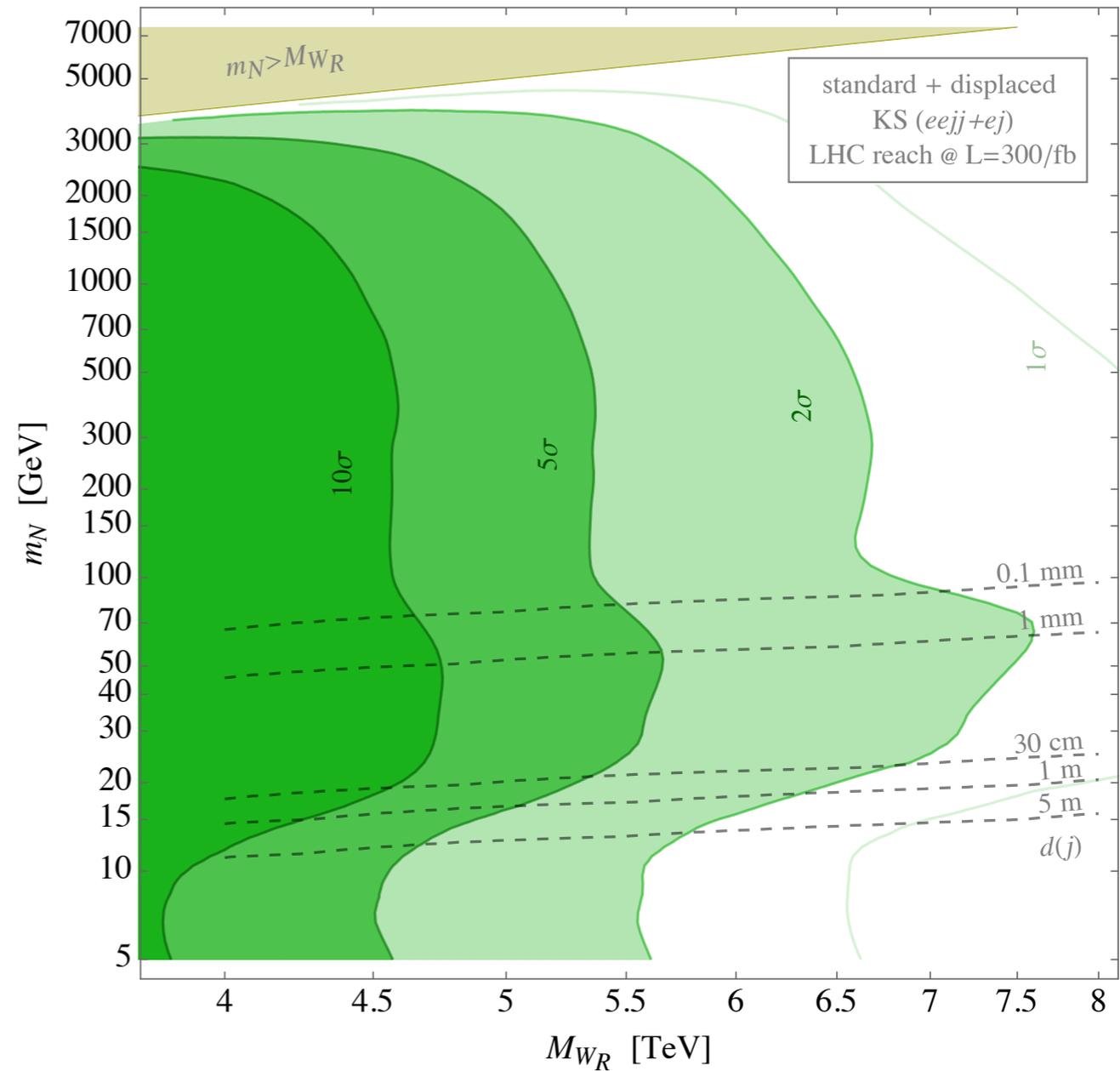
Sensitivity estimate

MN, Nesti, Popara '18

rough pre-selection

bin over 6 variables below

$$\text{sensitivity} = \sqrt{\sum_{i \in \text{bins}} \frac{s_i^2}{s_i + b_i}}$$

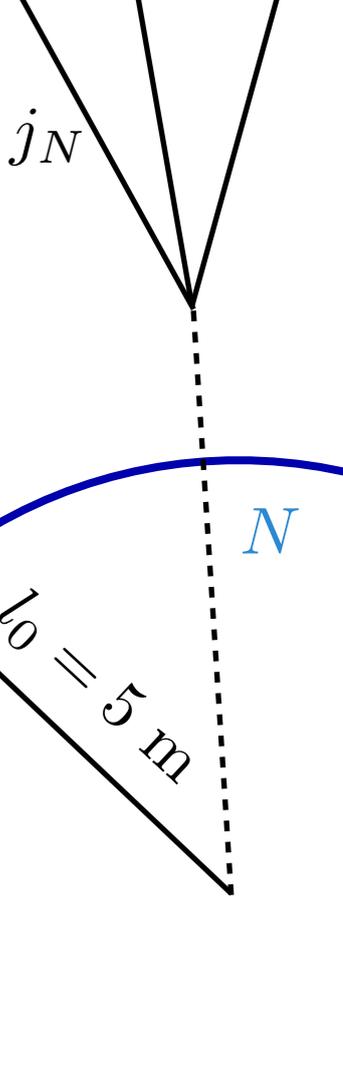


Sensitivities

Electron Channel $\mathcal{L} = 300 \text{ fb}^{-1}$			M_{WR} :	4 TeV	4 TeV	4 TeV	6 TeV	6 TeV
variable	range	# bins	m_N :	20 GeV	300 GeV	2 TeV	20 GeV	300 GeV
$p_T(\ell_1)$	{150, 4500} GeV	35		14.19	13.82	7.19	1.03	1.77
$d_T(j_1)$	{0.001, 300} mm	100		17.57	14.04	7.60	2.02	1.91
#(jets)	1, 2, 3, 4	4		17.88	14.20	7.94	2.24	2.04
#(leptons)	1, 2	2		17.97	14.90	9.08	2.30	2.23
#(same sign)	0, 1	2		18.00	15.71	9.85	2.32	2.61
$m_{\ell_1 j_1}^{\text{inv}}$	{200, 8500} GeV	20		18.82	17.24	10.91	2.81	3.03

Recasting the $W' \rightarrow \ell\nu$

MN, Nesti, Popara '18

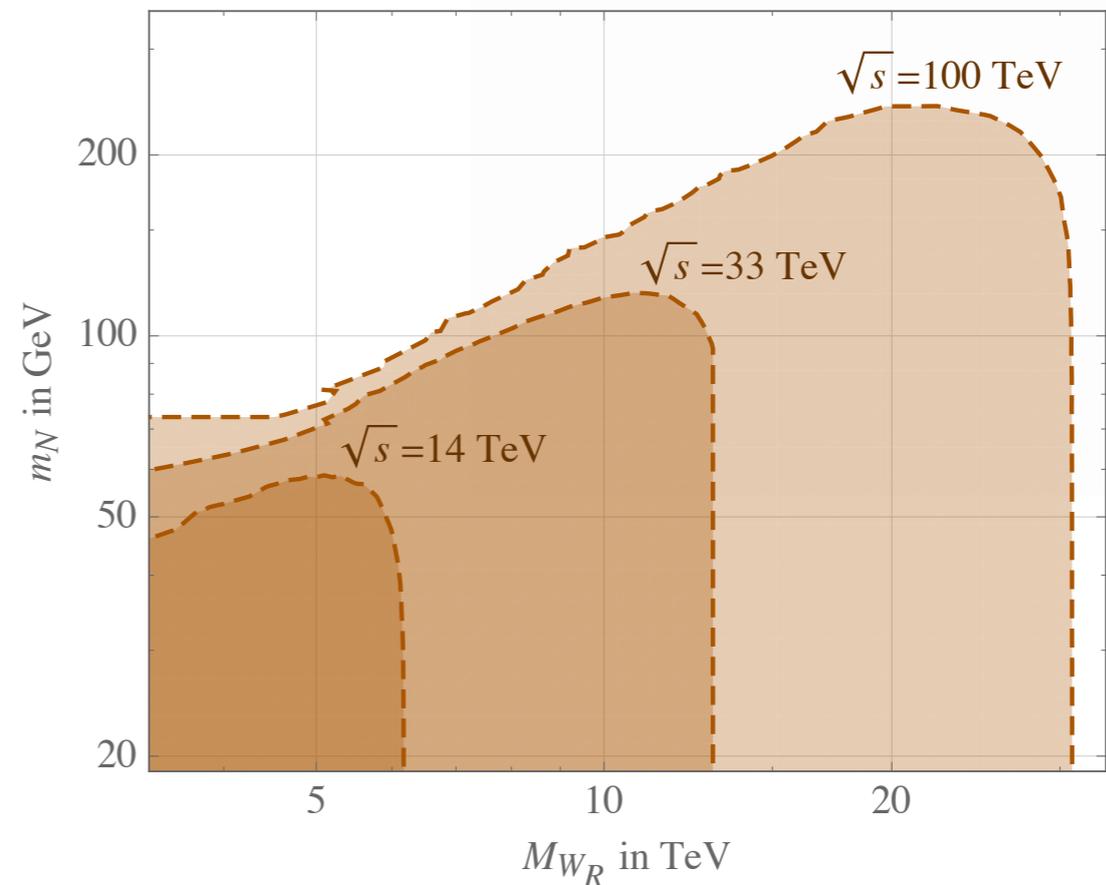
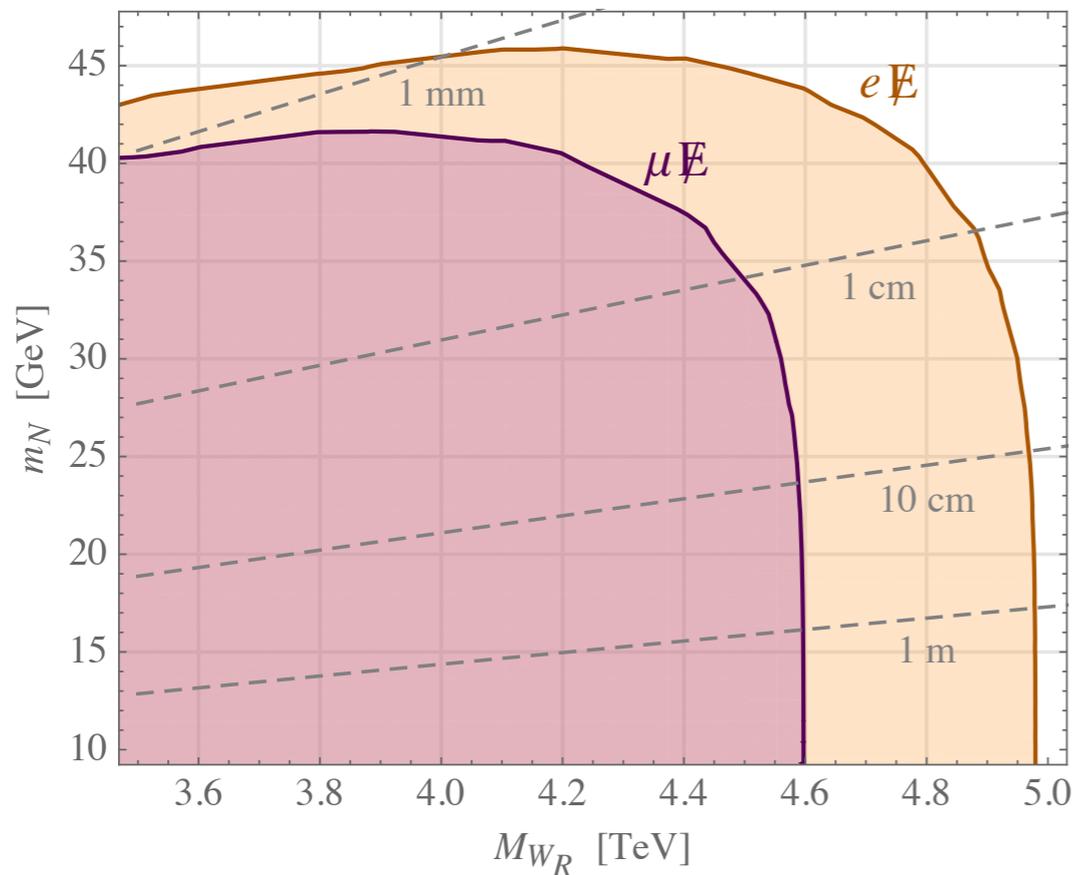


prompt hard leading lepton and significant missing energy

$$\frac{d\sigma}{dm_T} = \alpha_2^2 \frac{\pi}{24} p_T \int_{\tau_-}^1 \int_{\frac{\tau_-}{x_1}}^1 dx_{1,2} \frac{(\hat{s} - m_N^2 - 2p_T^2) \pm 1}{\sqrt{(\hat{s} - m_N^2)^2 - 4p_T^2 \hat{s}}}$$

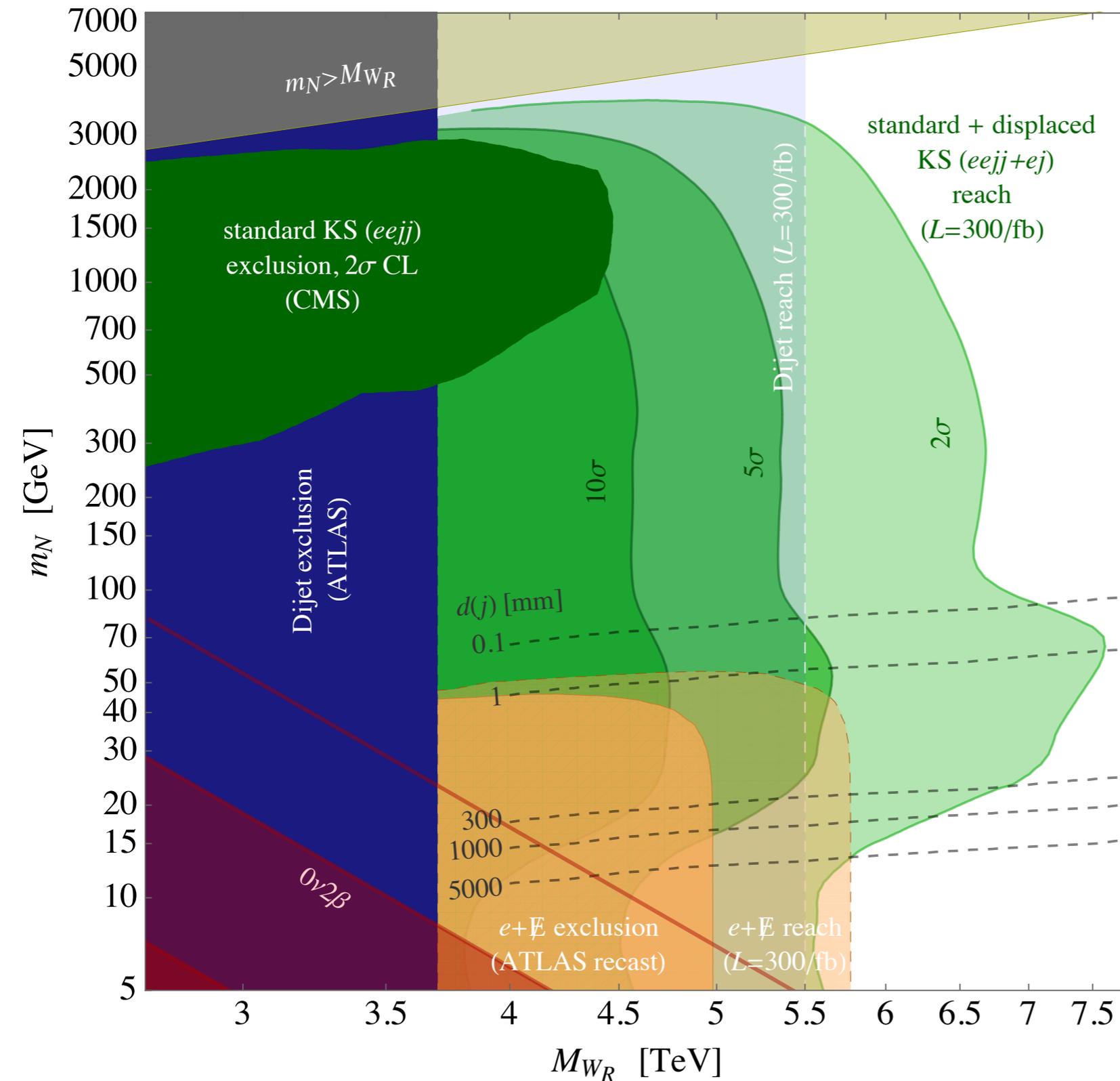
$$\frac{\varepsilon_\ell^\pm(p_T, \eta_\ell)}{(\hat{s} - M^2)^2 + (\Gamma M)^2} |V_{ud}V_{\ell N}|^2 f_u(x_{1,2}) f_{\bar{d}}(x_{2,1}) e^{-l_0/L_\pm}$$

exponential distributions have tails



Search overview $pp \rightarrow W_R \rightarrow \ell_R N$

MN, Nesti, Popara '18



standard prompt isolated mode

Ng et al. '15, Ruiz '17

merged neutrino jet ℓj_N

Mitra, Ruiz, Spannowsky '16

Richard's talk

displaced jet

ℓj_N^d

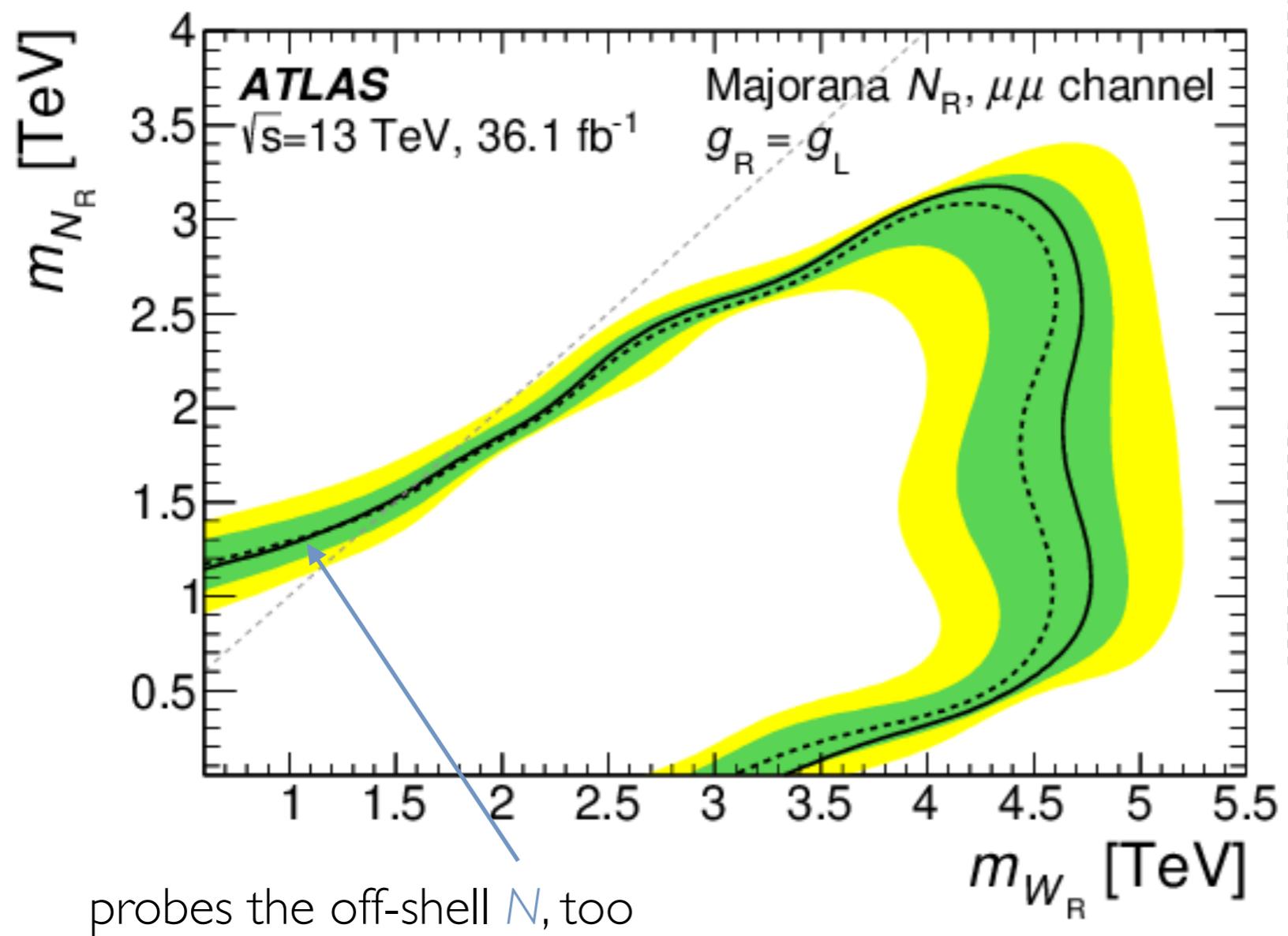
Cottin, Helo, Hirsch '18

invisible: prompt

$\ell + E_{miss}$

relevant for any light N search (SHIP, FASER, MATHUSLA, etc.)

Experimental limits review $pp \rightarrow W_R \rightarrow \ell_R N, \dots$



standard prompt isolated mode

*e, mu	4.7 TeV	ATLAS 1809.11105
tau	3.5 TeV	CMS 1811.00806

lepton + missing energy

*e, mu	$M_{W_R} \gtrsim 5$ TeV	ATLAS 1706.04786
tau _h	$M_{W_R} > 3.7$ TeV	ATLAS 1801.06992

*interplay with $0\nu 2\beta$

Mohapatra Senjanović '79
Tello, MN, Nesti, Senjanović, Vissani '10

dijets

$M_{W_R} > 3.6$ TeV	ATLAS 1703.09127
---------------------	---------------------

tb

$M_{W_R} > 3$ TeV	ATLAS 1801.07893
-------------------	---------------------

di-boson WZ mode $\propto \xi_{LR}$

up to ~ 5 TeV	ATLAS 1808.02380
--------------------	---------------------

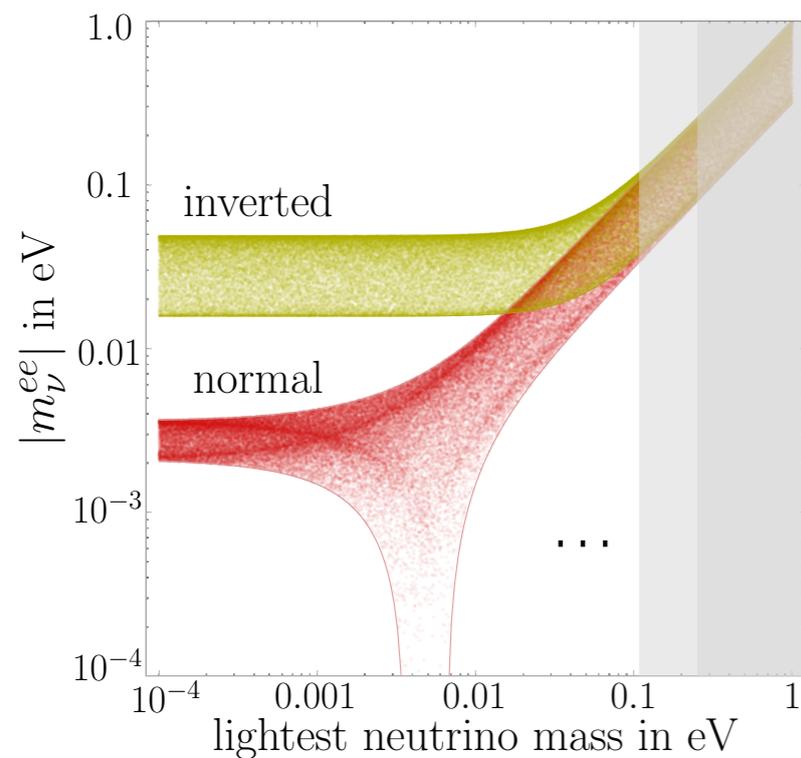
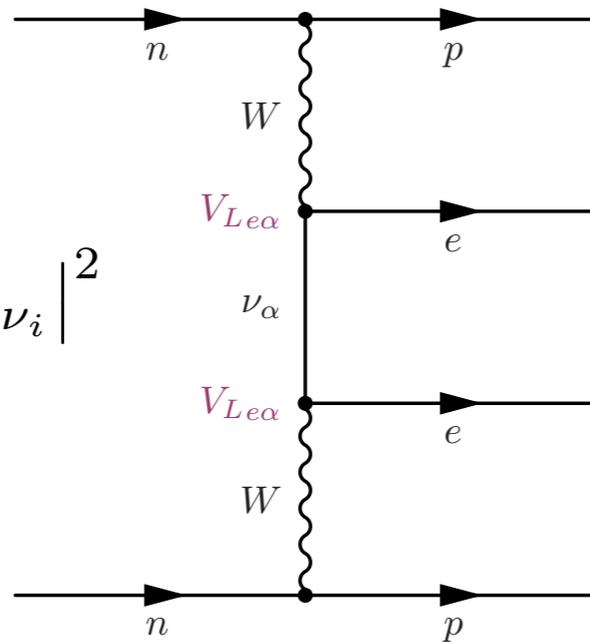
LNV interplay

Nuclear vs. collider physics

Immediate proposal after Majorana: $0\nu 2\beta$ via light neutrino exchange

Furry '39

$$\Gamma \propto |m_{\nu}^{ee} = V_{ei}^2 m_{\nu_i}|^2$$



Cosmology,
KATRIN, ...

Possible
tension

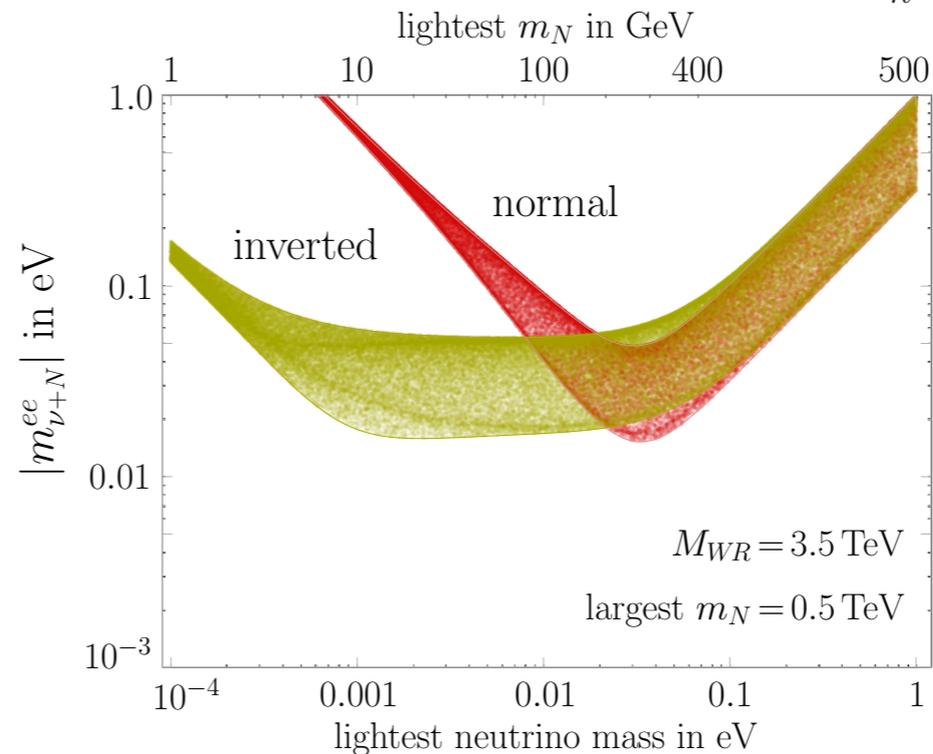
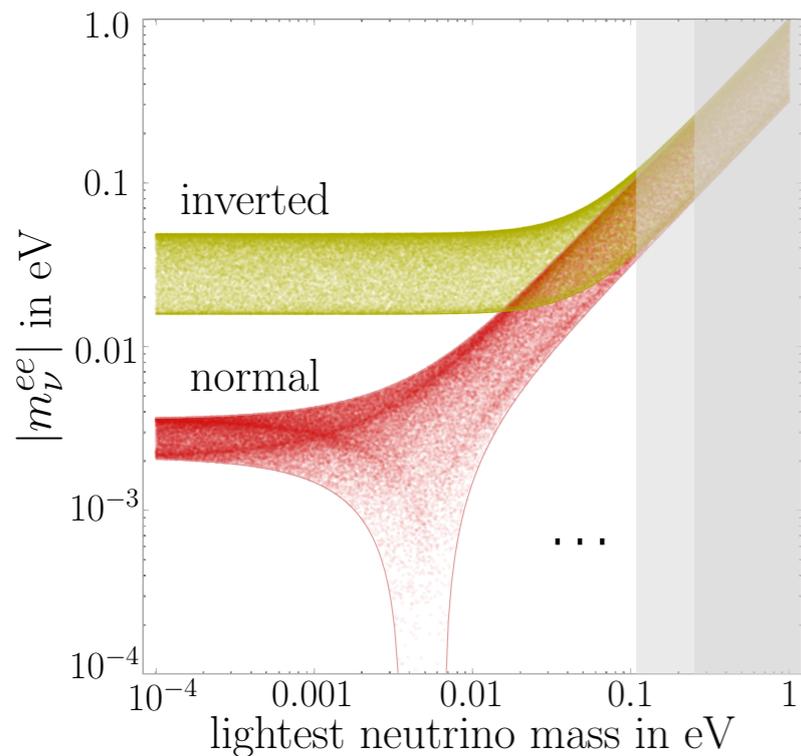
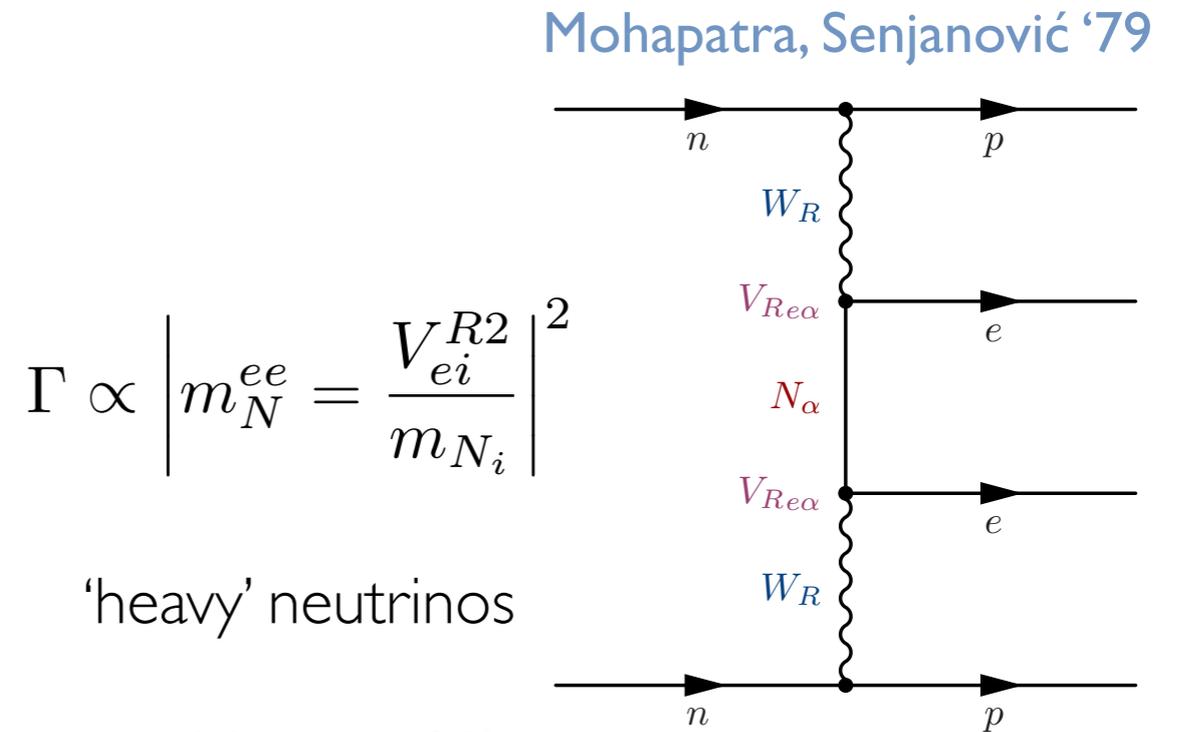
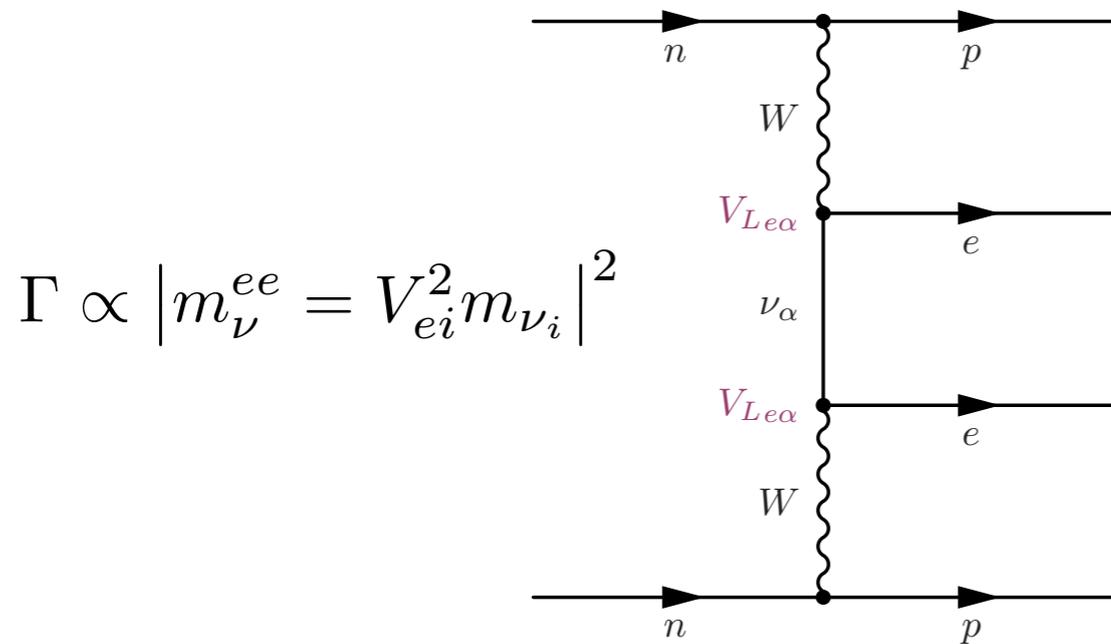
Vissani '02

LNV interplay

Nuclear vs. collider physics

Immediate proposal after Majorana: $0\nu 2\beta$ via light neutrino exchange

Furry '39



Tello, MN, Nesti, Senjanović, Vissani '11

no tension with LR

$$V_R = V_{PMNS}$$

connection to LFV

+other diagrams

Majorana - Dirac connection

$$M_\nu = M_D^T M_N^{-1} M_D + M_L,$$

MN, Senjanović, Tello, '12

$$\mathcal{C} : M_D^T = M_D,$$

$$M_L = \frac{v_L}{v_R} M_N$$

$$M_D = M_N \sqrt{M_N^{-1} M_\nu - \frac{v_L}{v_R}}$$

Majorana - Dirac connection

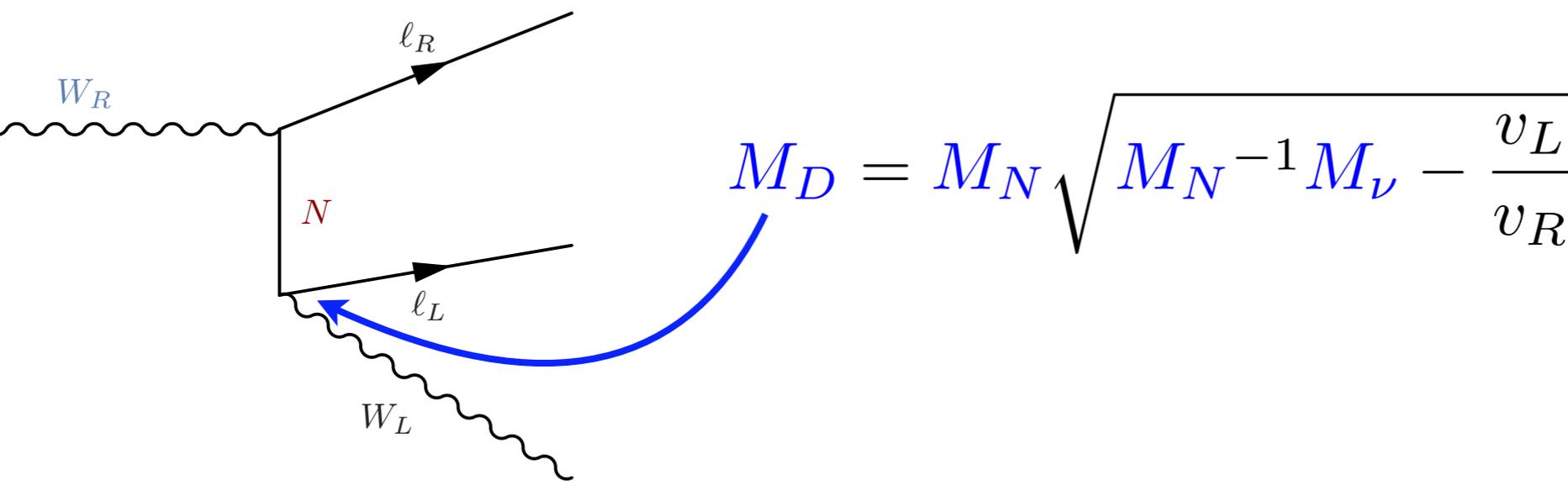
Colliders

$$M_\nu = M_D^T M_N^{-1} M_D + M_L,$$

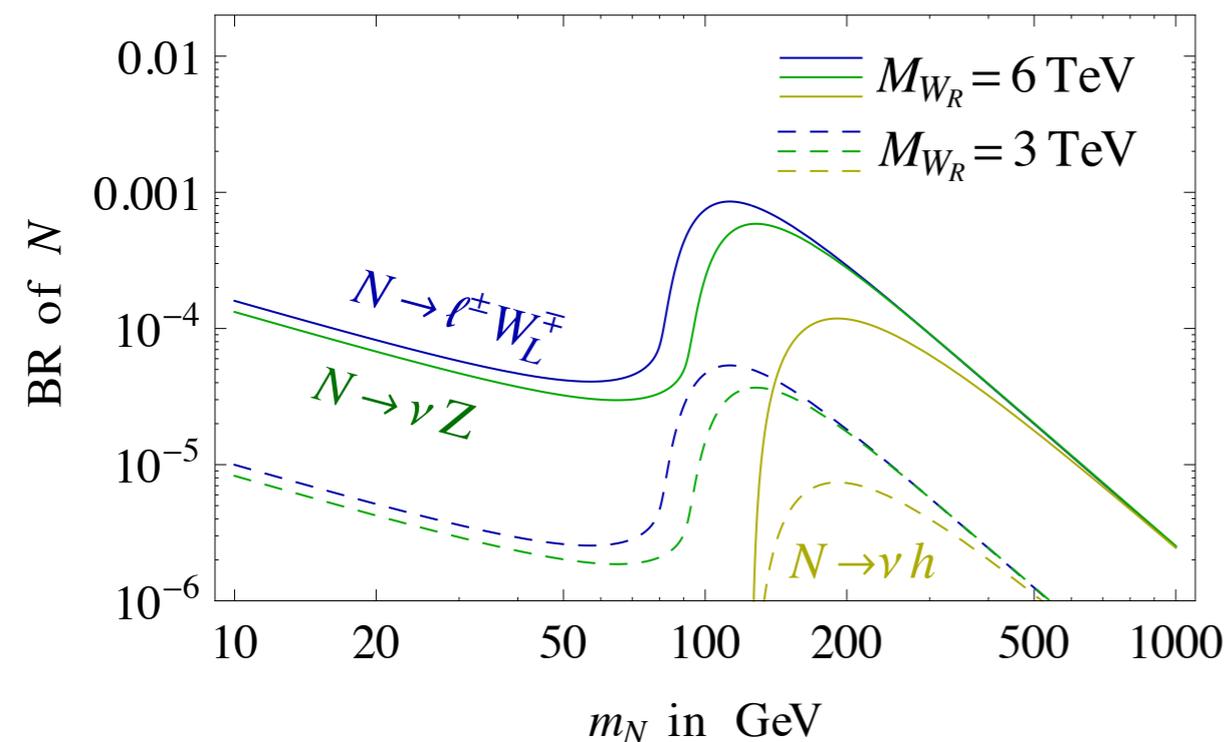
MN, Senjanović, Tello, '12

$$C : M_D^T = M_D,$$

$$M_L = \frac{v_L}{v_R} M_N$$



$$M_D = M_N \sqrt{M_N^{-1} M_\nu - \frac{v_L}{v_R}}$$



Sub-dominant decays, needs high lumi

Six flavor channels for M_D

Polarimetry

Han, Luiz, Ruiz, Si '12

Majorana - Dirac connection

eEDM

$$M_\nu = M_D^T M_N^{-1} M_D + M_L,$$

MN, Senjanović, Tello, '12

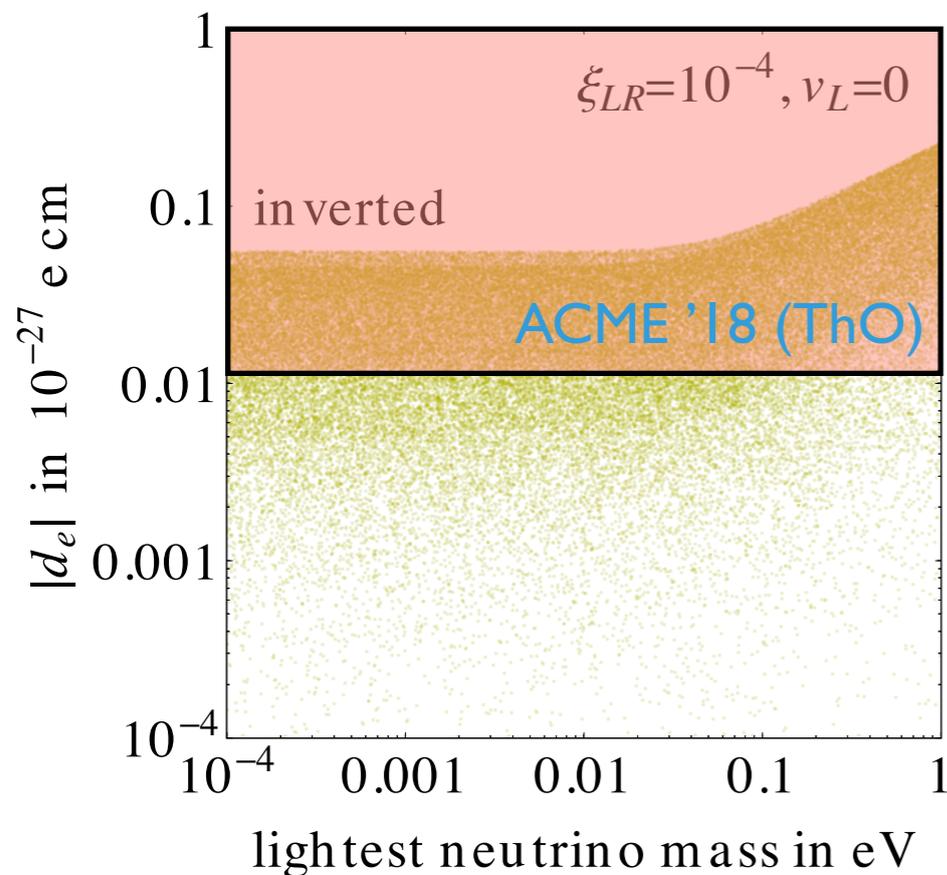
$$\mathcal{C} : M_D^T = M_D,$$

$$M_L = \frac{v_L}{v_R} M_N$$

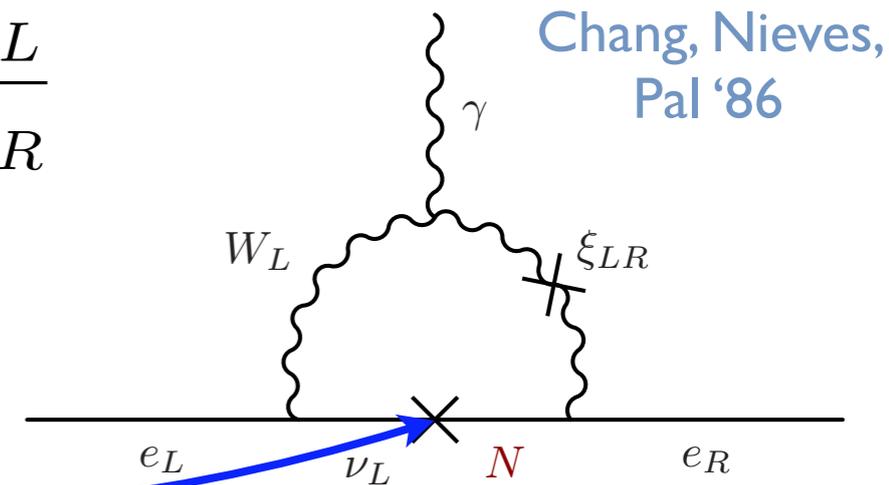
SM is \sim zero

$$d_e^{SM} \lesssim 10^{-38} \text{ e cm}$$

Pospelov, Ritz '05



$$M_D = M_N \sqrt{M_N^{-1} M_\nu - \frac{v_L}{v_R}}$$



$$d_e = \frac{eG_F}{4\sqrt{2}\pi^2} \text{Im} [\xi_{LR} V_R \mathcal{F}(t) V_R^\dagger M_D]_{ee}$$

T-odd, sensitive to Majorana phases

$$\mathcal{F}(t) = \frac{t^2 - 11t + 4 + 6t^2 \log t / (t - 1)}{2(t - 1)^2},$$

$$t = (m_N / M_W)^2$$

Higgs sector

$$\Delta_L(3, 1, 2), \Phi(2, 2, 0), \Delta_R(1, 3, 2)$$

Minkowski '77

Mohapatra, Senjanović '79

$$\Phi = \begin{pmatrix} \phi_1^0 & \phi_2^+ \\ \phi_1^- & \phi_2^0 \end{pmatrix}$$

$$\langle \Phi \rangle = \begin{pmatrix} v & 0 \\ 0 & 0 \end{pmatrix}$$

$$\Delta_R = \begin{pmatrix} \Delta^+/\sqrt{2} & \Delta^{++} \\ \Delta^0 & -\Delta^+/\sqrt{2} \end{pmatrix}_R \quad \langle \Delta_R \rangle = \begin{pmatrix} 0 & 0 \\ v_R & 0 \end{pmatrix}$$

SSB of parity

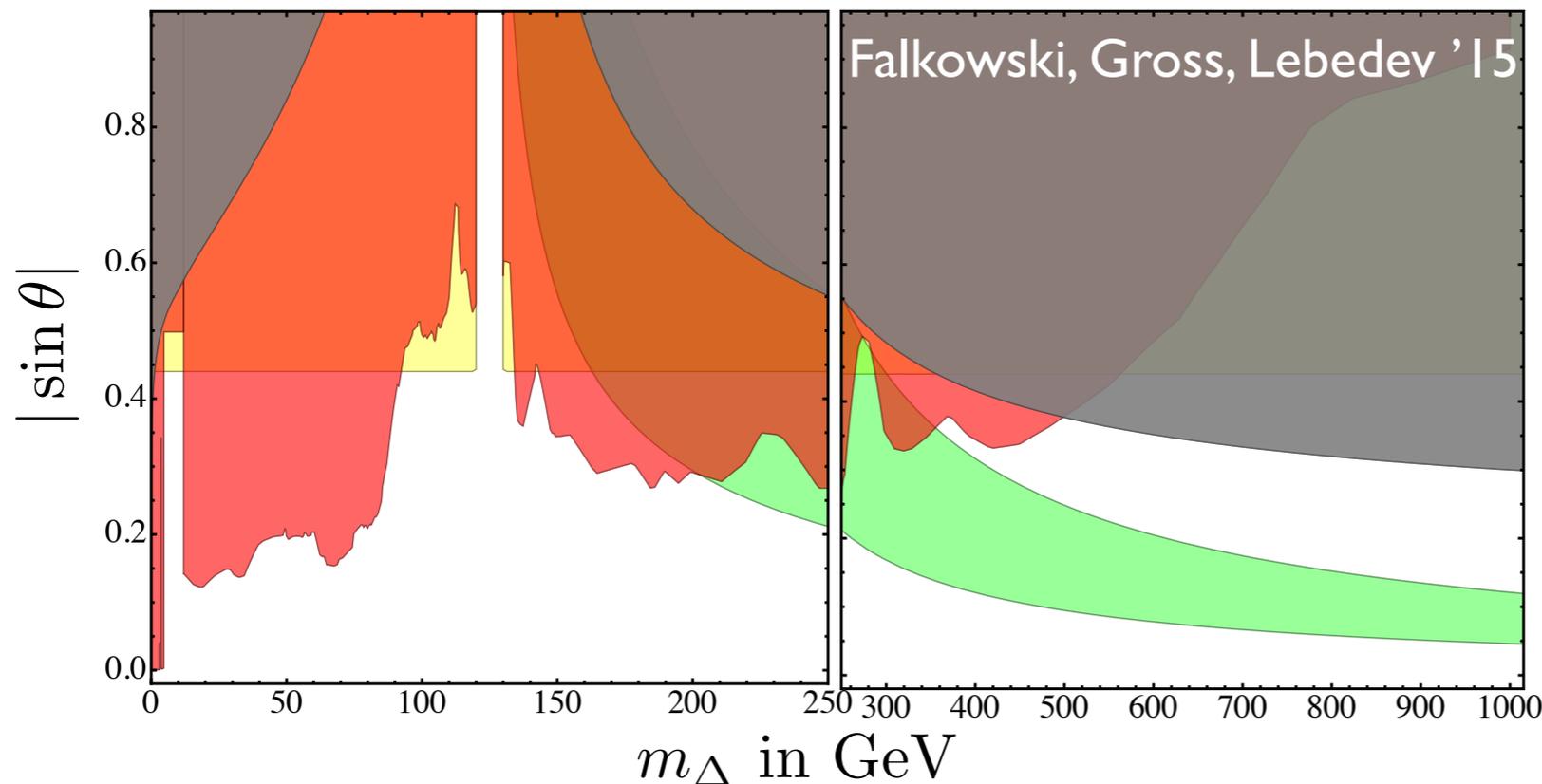
$$\mathcal{P} : \begin{cases} \Delta_L \leftrightarrow \Delta_R, \Phi \rightarrow \Phi^\dagger \\ Q_L \leftrightarrow Q_R, L_L \leftrightarrow L_R \end{cases}$$

Senjanović,
Mohapatra '75

$$V \in \lambda (\Phi^\dagger \Phi)^2 + \alpha (\Phi^\dagger \Phi) (\Delta_R^\dagger \Delta_R) + \rho (\Delta_R^\dagger \Delta_R)^2$$

same for \mathcal{C} -symmetry

$$h - \Delta \text{ mixing: } \theta \simeq \left(\frac{\alpha}{2\rho} \right) \left(\frac{v}{v_R} \right) \lesssim .44$$



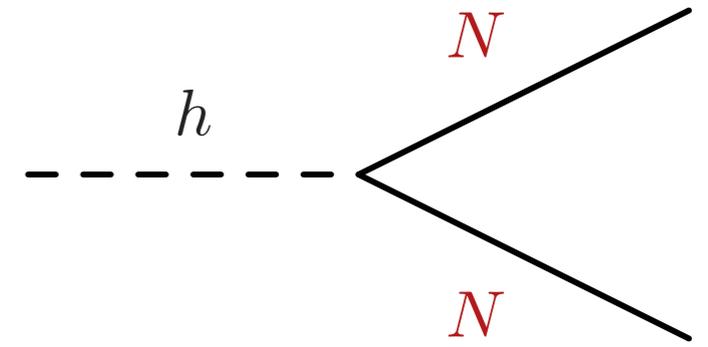
Future collider
outlook

$$|\sin \theta| < .34$$

Buttazzo, Sala, Tesi '15

'Majorana' SM Higgs

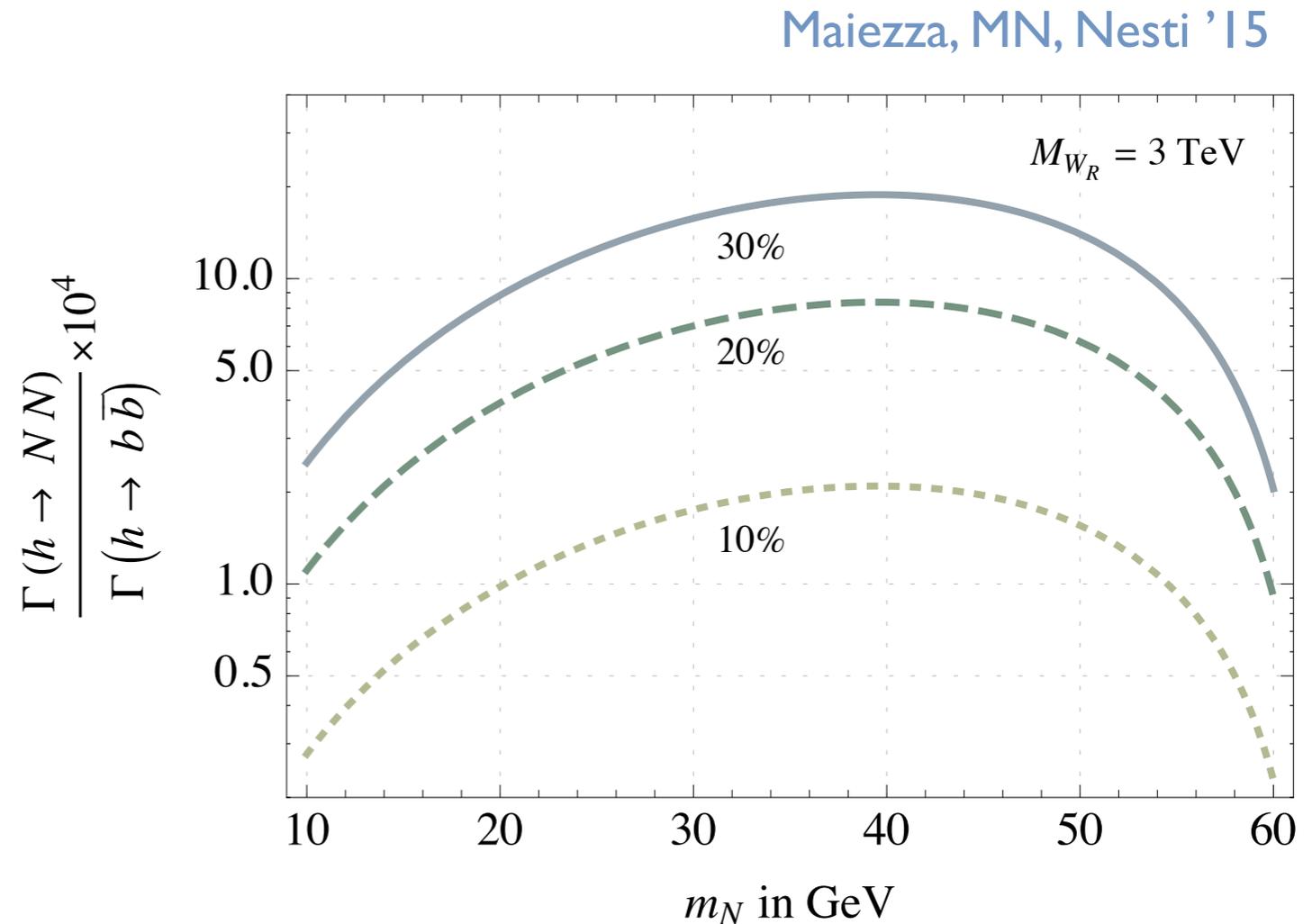
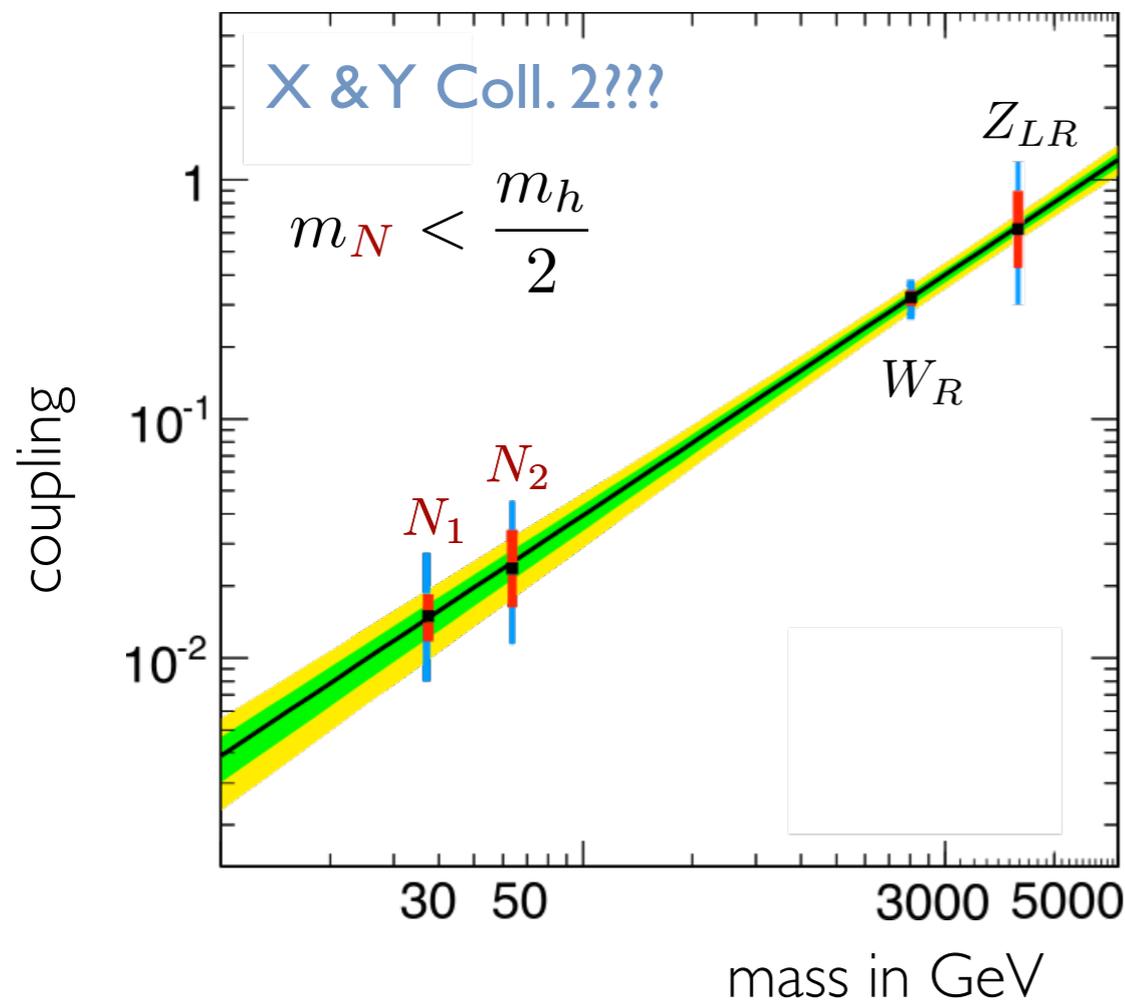
h decays



$$\Gamma_{h \rightarrow NN} \propto s_\theta^2 m_N^2 \quad \frac{\Gamma_{h \rightarrow NN}}{\Gamma_{h \rightarrow b\bar{b}}} \simeq \frac{\theta^2}{3} \left(\frac{m_N}{m_b} \right)^2 \left(\frac{M_W}{M_{W_R}} \right)^2$$

Gunion et al. Snowmass '86

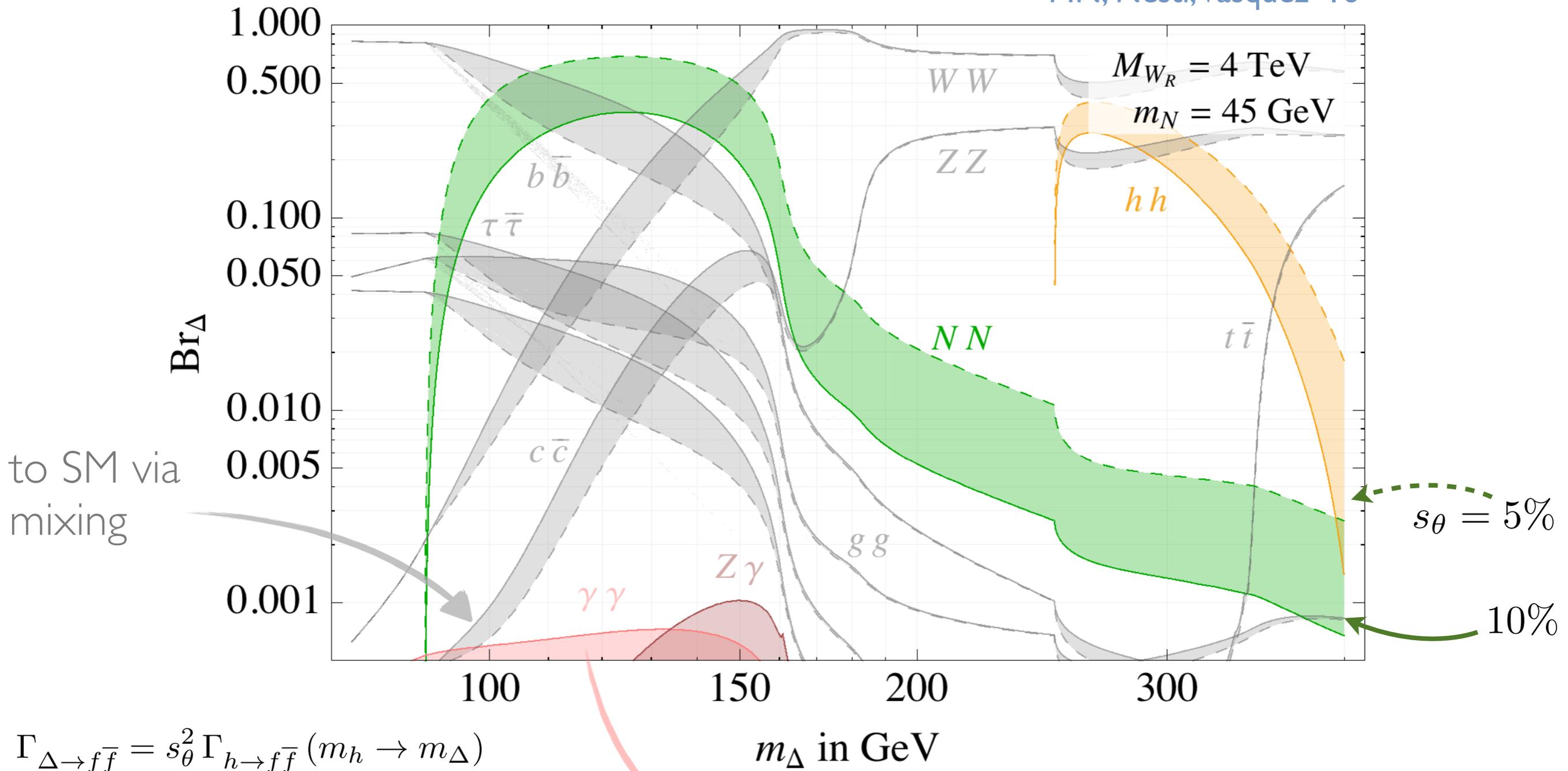
EFT SM+h+N Graesser '07



'Right-handed' Higgs

Δ_R^0 decays

MN, Nesti, Vasquez '16



radiative loops
(SM, W_R , $\Delta_{L,R}^{++}$)

Displaced photons Dev,
Mohapatra, Zhang '16

$$\Gamma_{\Delta \rightarrow \gamma \gamma} = \frac{m_\Delta^3}{64\pi} \left(\frac{\alpha}{4\pi} \right)^2 |F_\Delta|^2$$

'Right-handed' Higgs

Δ decays

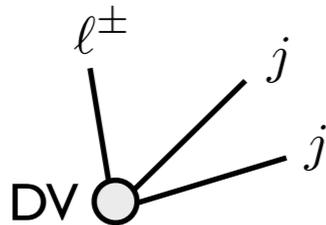
Region of interest for $\Delta \rightarrow NN$

$$20 \text{ GeV} \lesssim m_\Delta \lesssim 170 \text{ GeV}$$

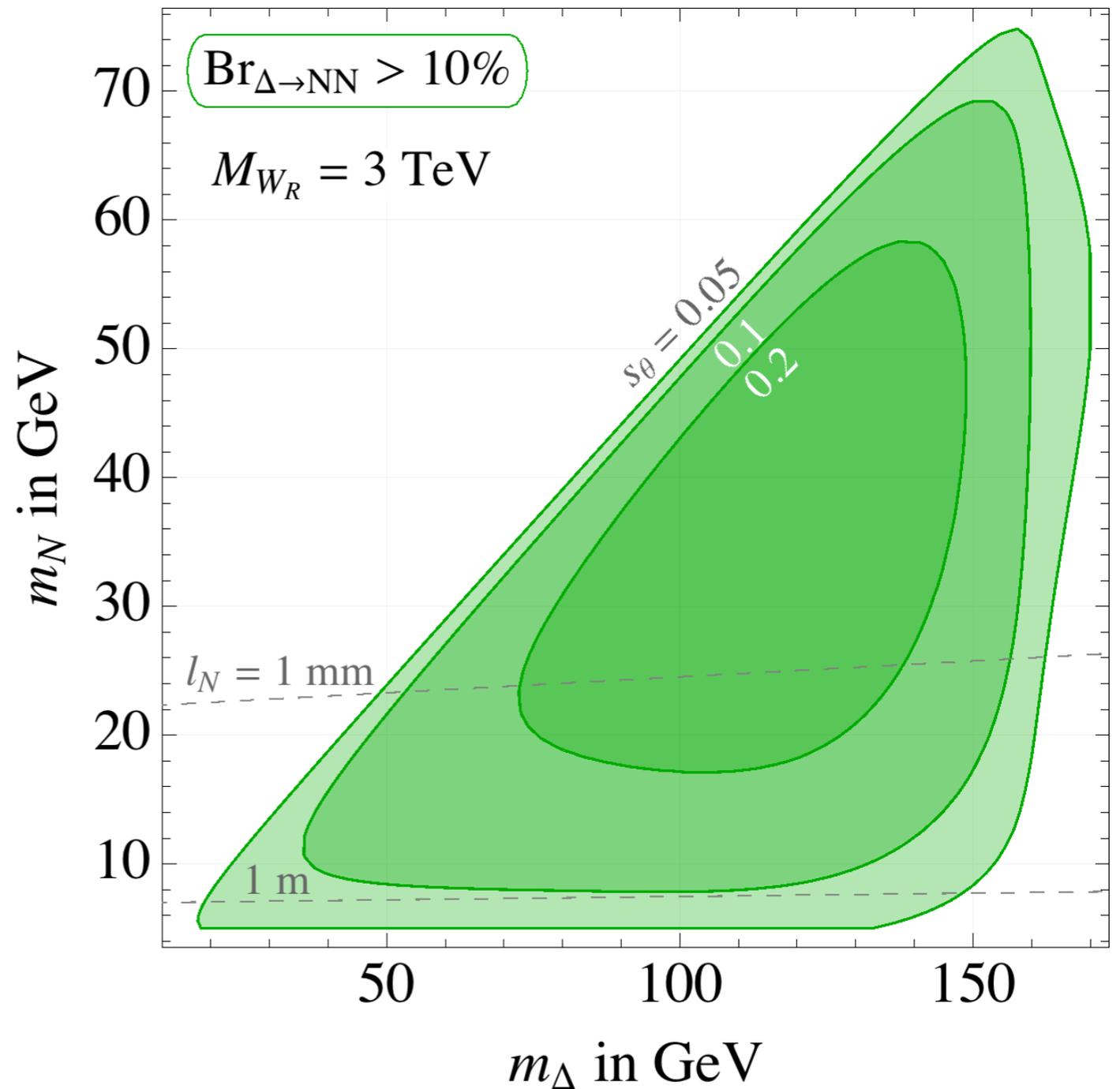
Decay length

$$c\tau_N^0 \simeq 0.1 \text{ mm} \left(\frac{40 \text{ GeV}}{m_N} \right)^5 \left(\frac{M_{W_R}}{5 \text{ TeV}} \right)^4$$

Leads to two DV with LNV



resol. $\mathcal{O}(10) \mu m$



'Right-handed' Higgs

Δ production

single

$$\sigma(gg \rightarrow \Delta) = s_\theta^2 \sigma(gg \rightarrow h)$$

$$\sigma(pp \rightarrow V\Delta) = s_\theta^2 \sigma(pp \rightarrow Vh)$$

N³LO

Anastasiou et al.'16

pair &
associated

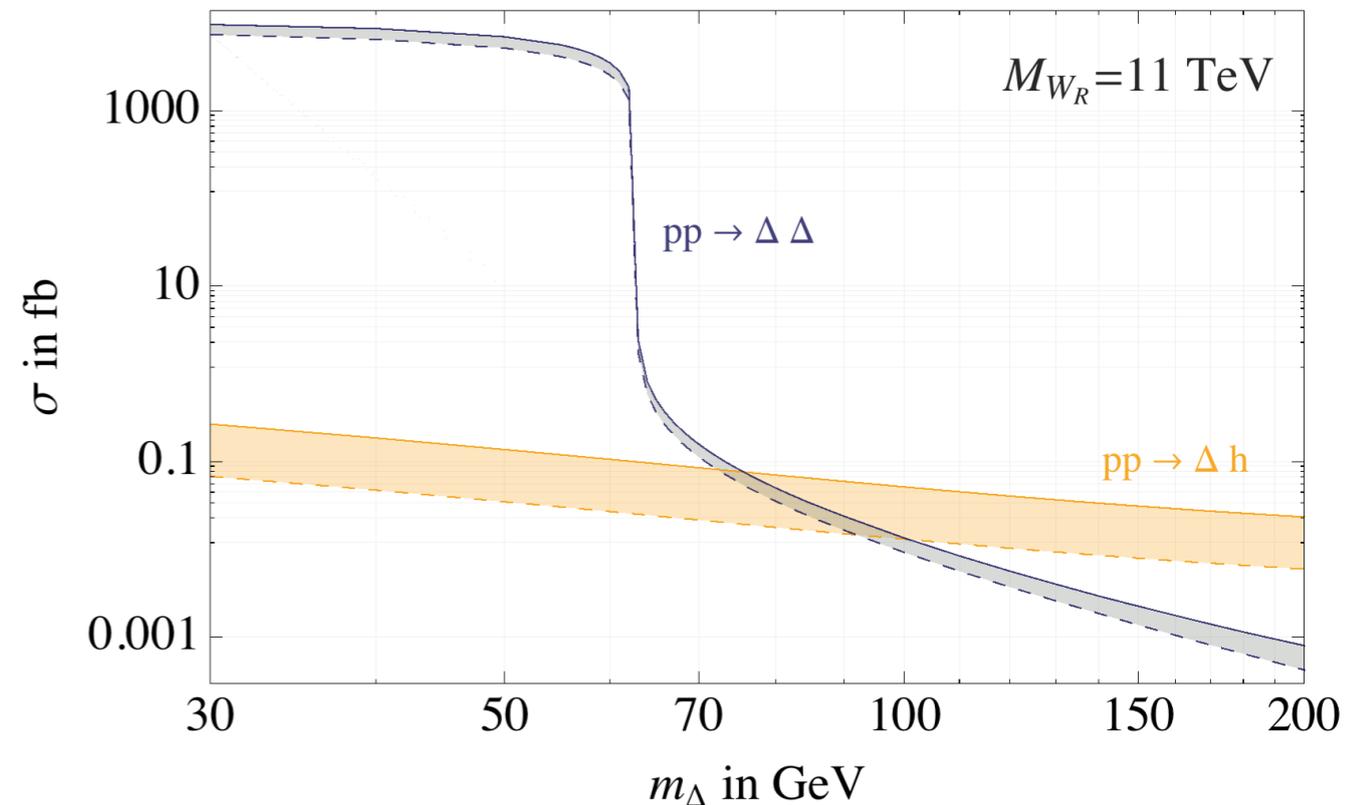
$$\hat{\sigma}_{gg \rightarrow \Delta S} \simeq \frac{c_\theta^2}{64\pi(1 + \delta_{\Delta S})} \hat{s} \left(\frac{\alpha_s}{4\pi}\right)^2 \frac{v_h^2 S_\Delta}{(\hat{s} - m_h^2)^2 + \hat{s}\Gamma_h^2} |F_b + F_t|^2 \sqrt{\beta_{\hat{s}\Delta S}}$$

large rate for $m_\Delta < m_h/2$

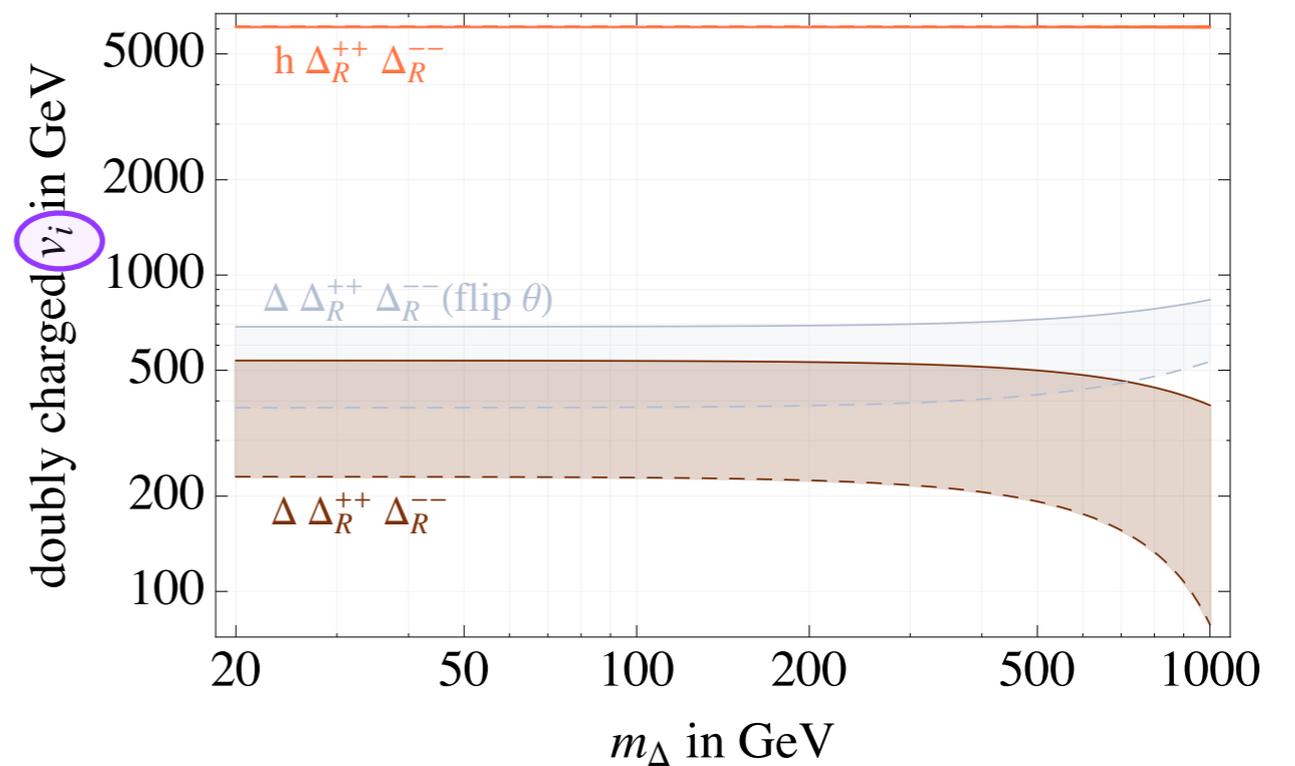
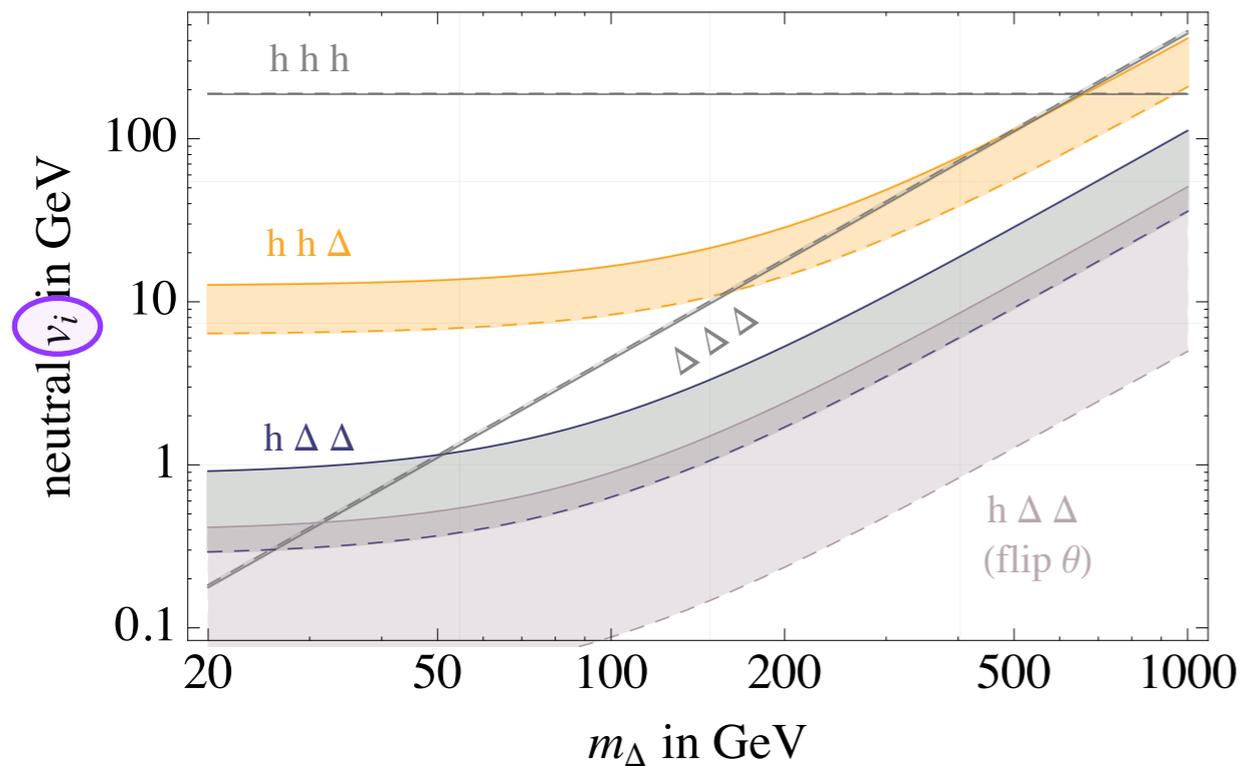
$$\sigma_{gg \rightarrow \Delta\Delta} \simeq \sigma_{gg \rightarrow h} \text{Br}_{h \rightarrow \Delta\Delta}$$

not very significant

(accidental cancellation)



Tri-linear Higgs @ LO



2 x 2 matrix, mixing suppressed by flavor and $\langle \Delta_L \rangle$

tree level

$$v_{hhhh} = \frac{3g}{2} m_h^2 \left[\frac{c_\theta^3}{M_W} - \sqrt{2} \frac{s_\theta^3}{M_{W_R}} \right]$$

$$v_{hh\Delta} = \frac{g}{4} s_{2\theta} (m_\Delta^2 + 2m_h^2) \left[\frac{c_\theta}{M_W} + \sqrt{2} \frac{s_\theta}{M_{W_R}} \right] \xrightarrow{\theta \rightarrow 0} 0$$

$$v_{h\Delta\Delta} = \frac{g}{4} s_{2\theta} (m_\Delta^2 + 2m_h^2) \left[\frac{s_\theta}{M_W} - \sqrt{2} \frac{c_\theta}{M_{W_R}} \right] \xrightarrow{\theta \rightarrow 0} 0$$

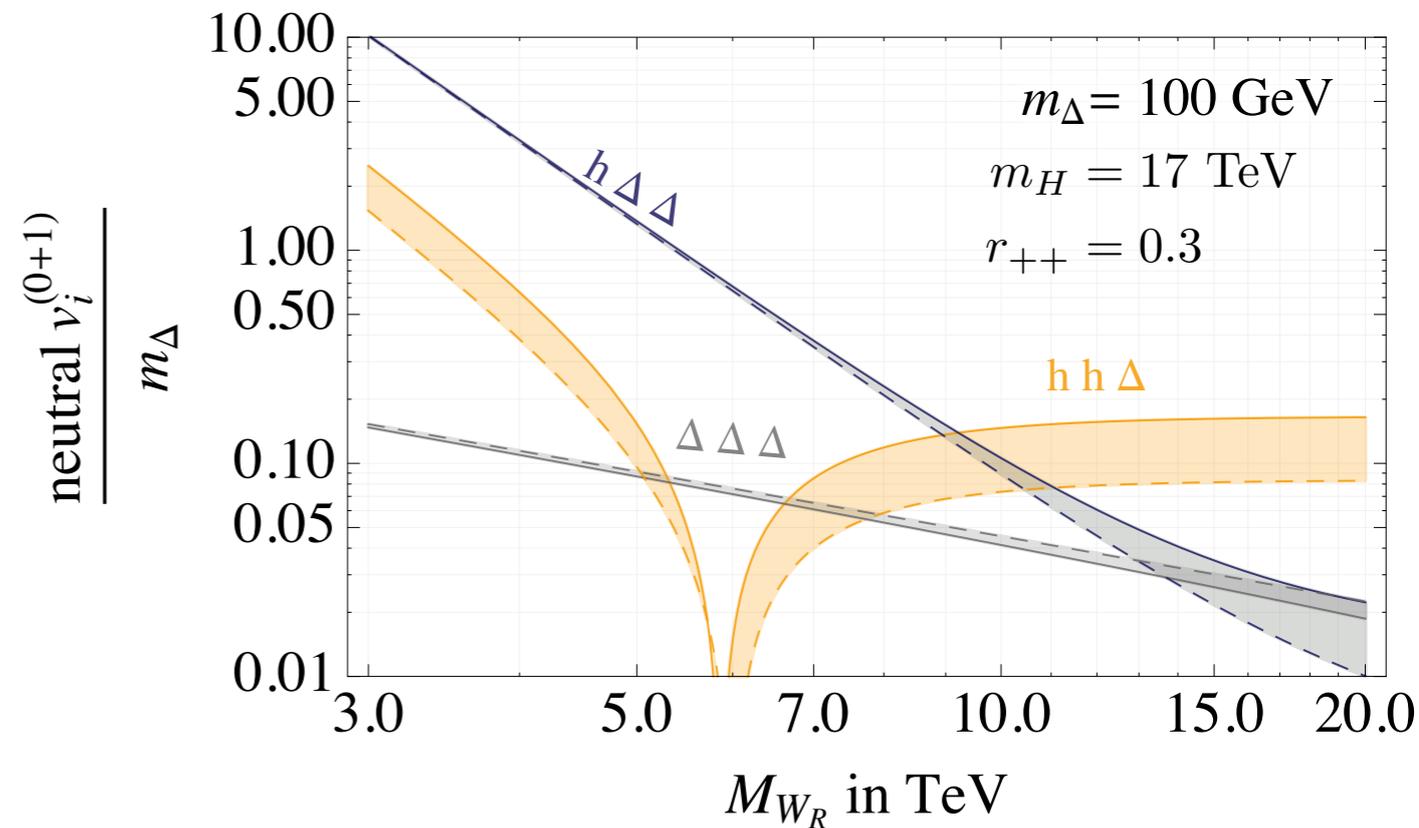
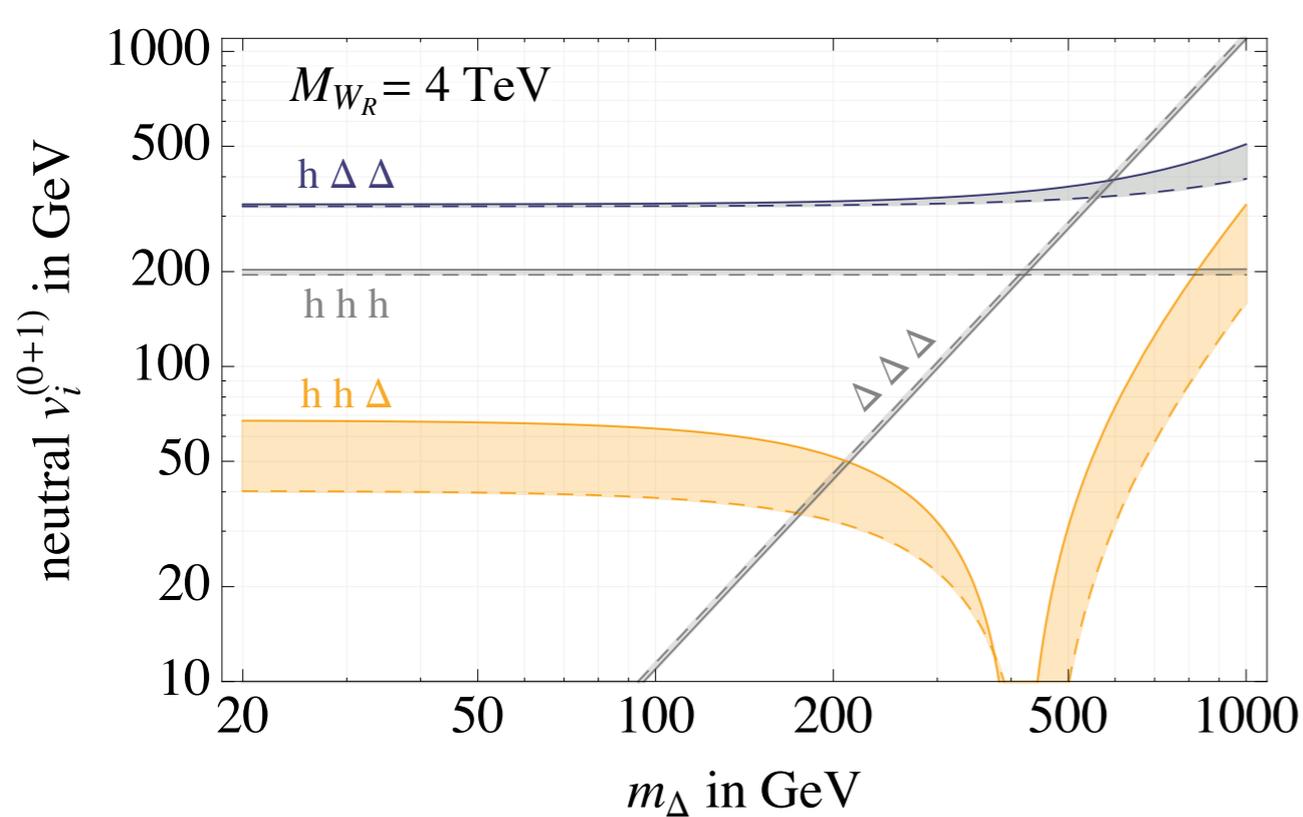
$$v_{\Delta\Delta\Delta} = \frac{3g}{2} m_\Delta^2 \left[\frac{s_\theta^3}{M_W} + \sqrt{2} \frac{c_\theta^3}{M_{W_R}} \right]$$

cancellation

+ corrections due to H mixing

Tri-linear Higgs @ NLO

loop corrections, \sim top in the hhh vertex of the SM



$$v_{hhh}^{(1)} \simeq c^{(1)} \left(1 + \frac{17}{3} \frac{1}{r_{++}} \right) \left(\frac{v}{v_R} \right)^2 v$$

$$v_{h\Delta\Delta}^{(1)} \simeq c^{(1)} (4 + 10 r_{++}) v$$

$$c^{(1)} = \frac{1}{\sqrt{2}(4\pi)^2} \left(\frac{m_H}{v_R} \right)^4,$$

$$r_{++} = \left(\frac{m_{\Delta^{++}, \Delta_L^{0,+,++}}}{m_H} \right)^2$$

$$v_{hh\Delta}^{(1)} \simeq c^{(1)} 11 \left(\frac{v}{v_R} \right) v$$

$$v_{\Delta\Delta\Delta}^{(1)} \simeq c^{(1)} (8 + 16 r_{++}^2) v_R$$

decouple with v_R

upper bound $v_{\Delta\Delta\Delta}^{(1)} \leq \left(\frac{7}{3} \right) v_{\Delta\Delta\Delta}^{\text{tree level}}$ from vacuum stability

Linde '76, Weinberg '76
Mohapatra '86
Basecq, Wyler '89

Δ production

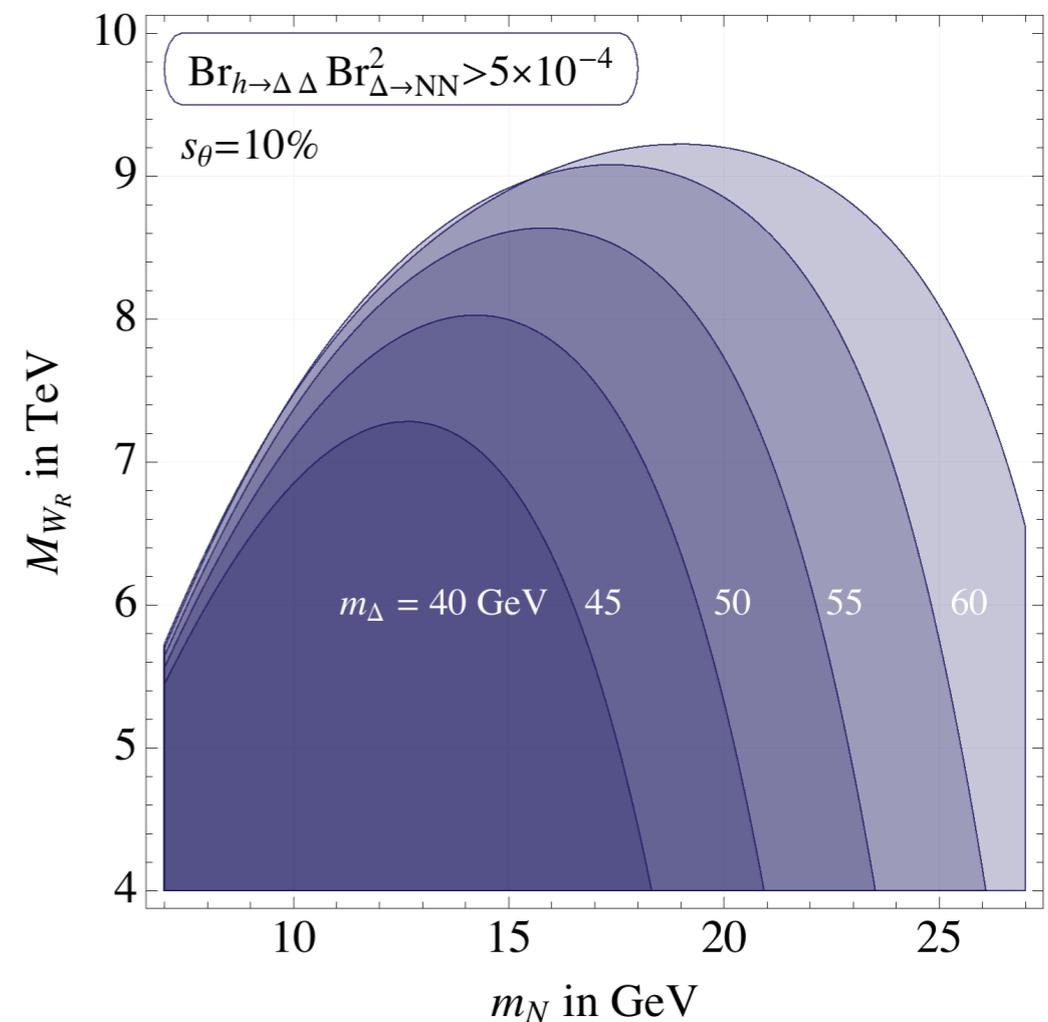
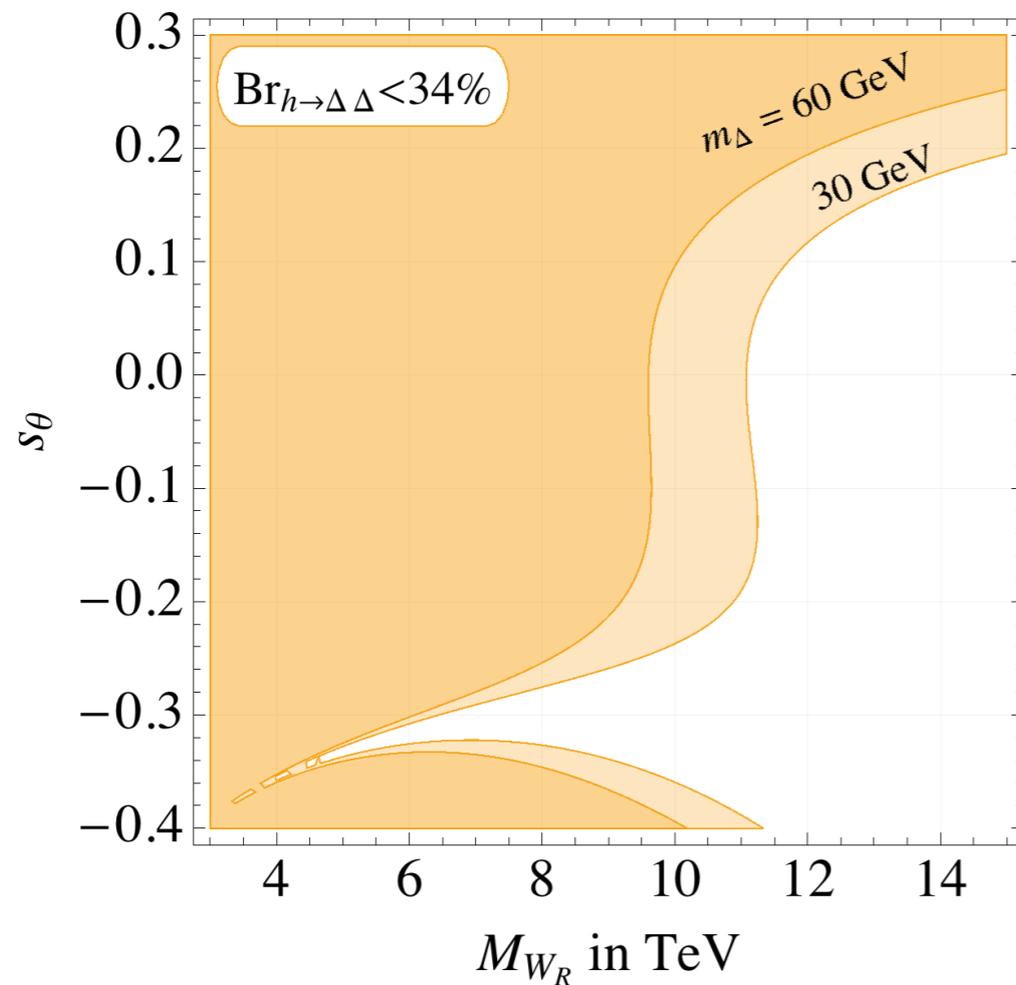
Δ^* suppressed

pair &
associated

$$\hat{\sigma}_{gg \rightarrow \Delta S} \simeq \frac{c_\theta^2}{64\pi(1 + \delta_{\Delta S})} \hat{s} \left(\frac{\alpha_s}{4\pi}\right)^2 \frac{v_{hS\Delta}^2}{(\hat{s} - m_h^2)^2 + \hat{s}\Gamma_h^2} |F_b + F_t|^2 \sqrt{\beta_{\hat{s}\Delta S}}$$

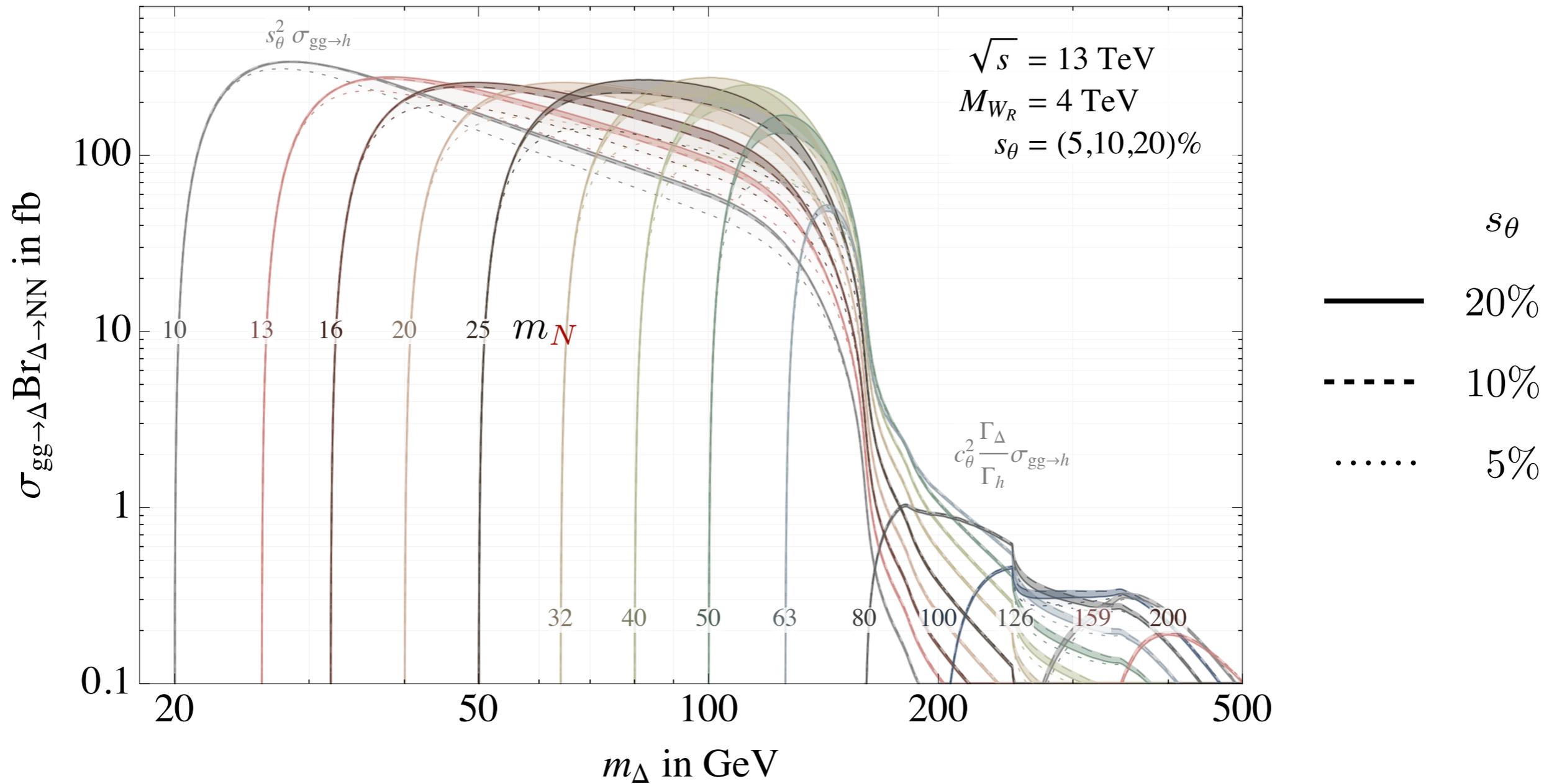
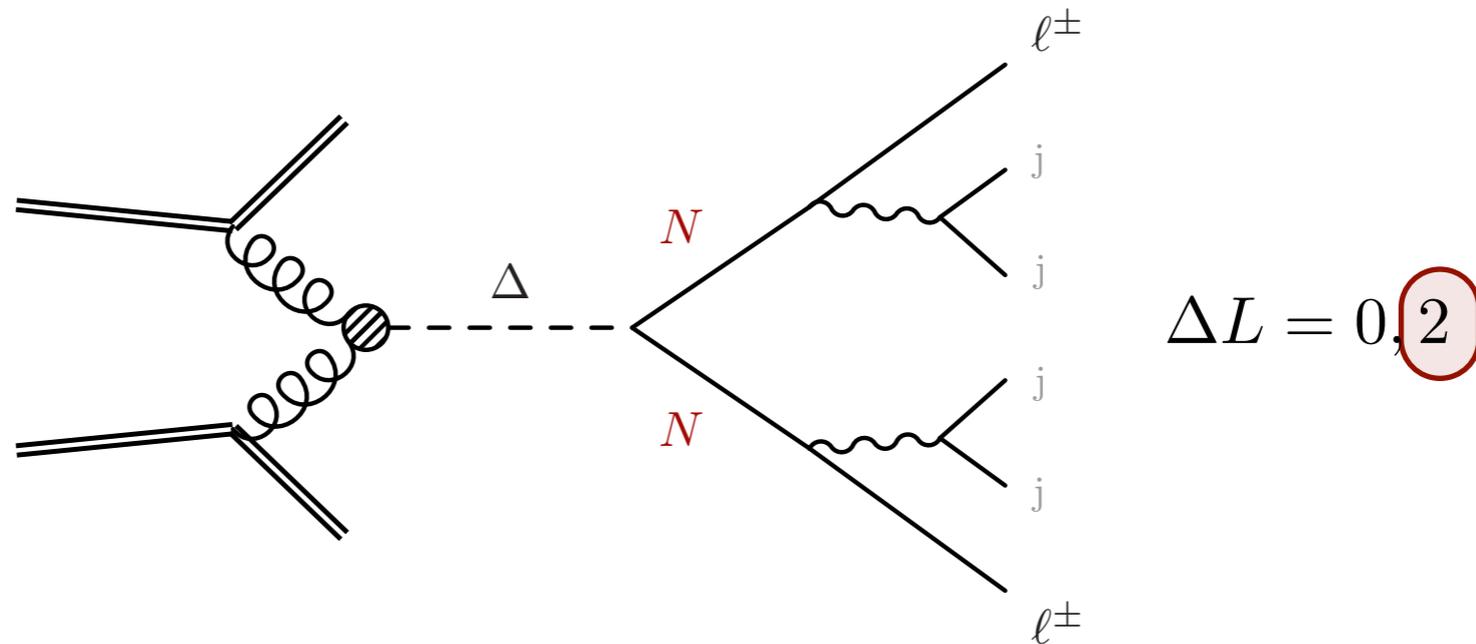
$\sigma_{gg \rightarrow \Delta\Delta} \simeq \sigma_{gg \rightarrow h} \text{Br}_{h \rightarrow \Delta\Delta}$ leads to $pp \rightarrow NNNN$

$\sigma_{gg \rightarrow h}$ N³LO Anastasiou et al. '16



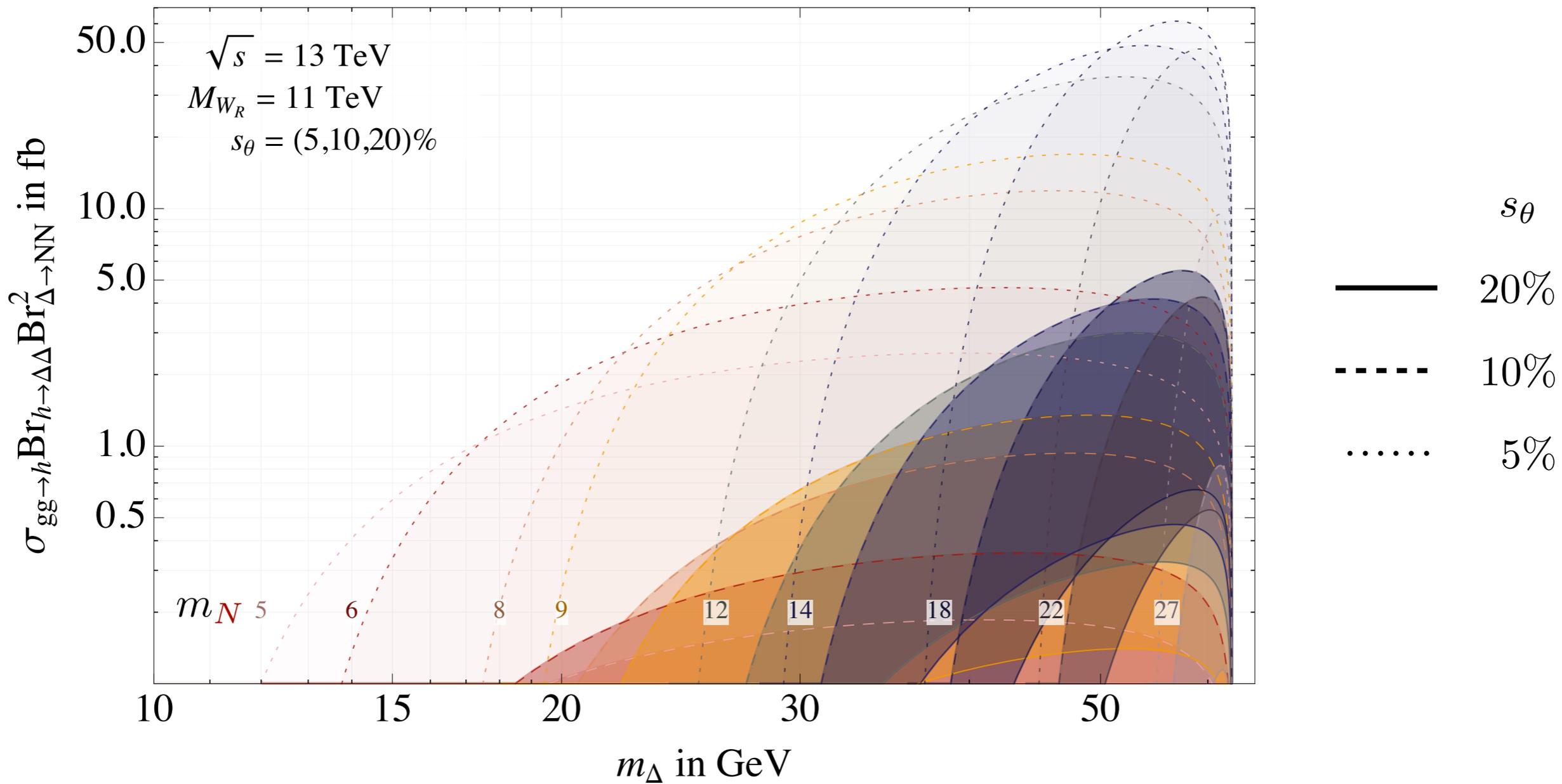
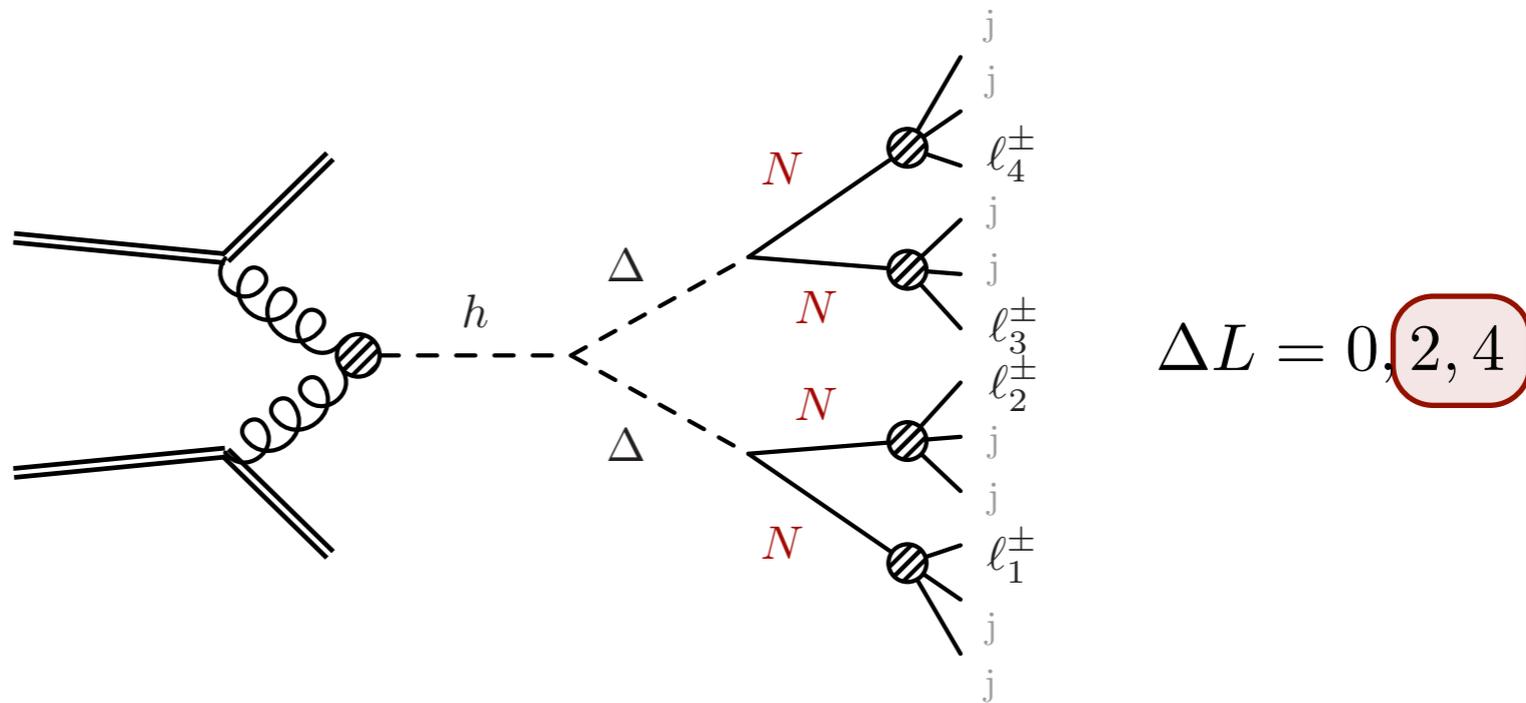
Δ signals

single



Δ signals

pair



'Majorana' Higgses at LHC

ggF production $\sigma_{gg \rightarrow h} \simeq 45 \text{ pb}$

$N^3\text{LO}$

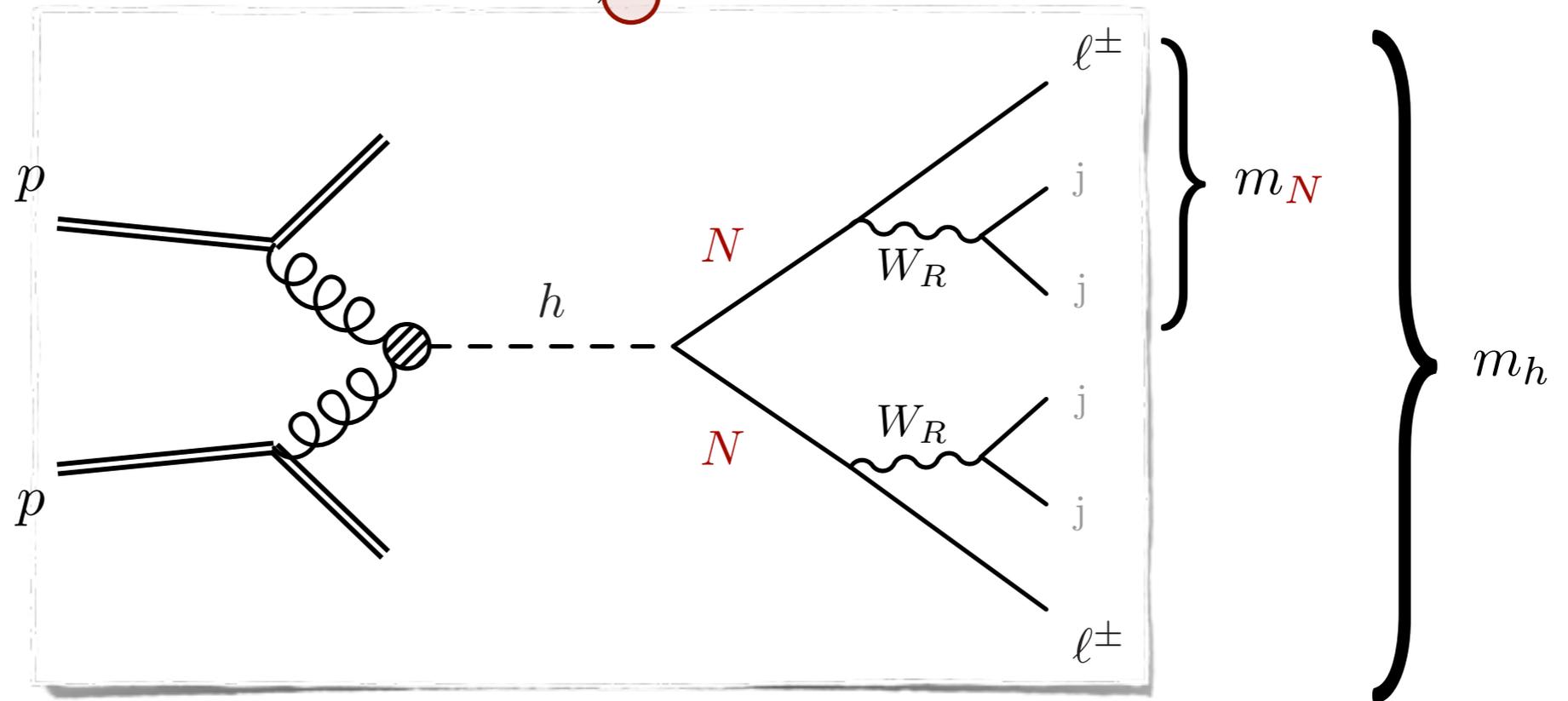
Anastasiou et al. '14

$\Delta L = 0$, 2

MN, Nesti, Vasquez '16

$$\Gamma_{h \rightarrow NN} \propto s_\theta^2 m_N^2$$

$$\text{Br}_{h \rightarrow NN} \simeq 10^{-3}$$



small couplings, no tuning

no missing energy

light jets only $V_L^q = V_R^q$

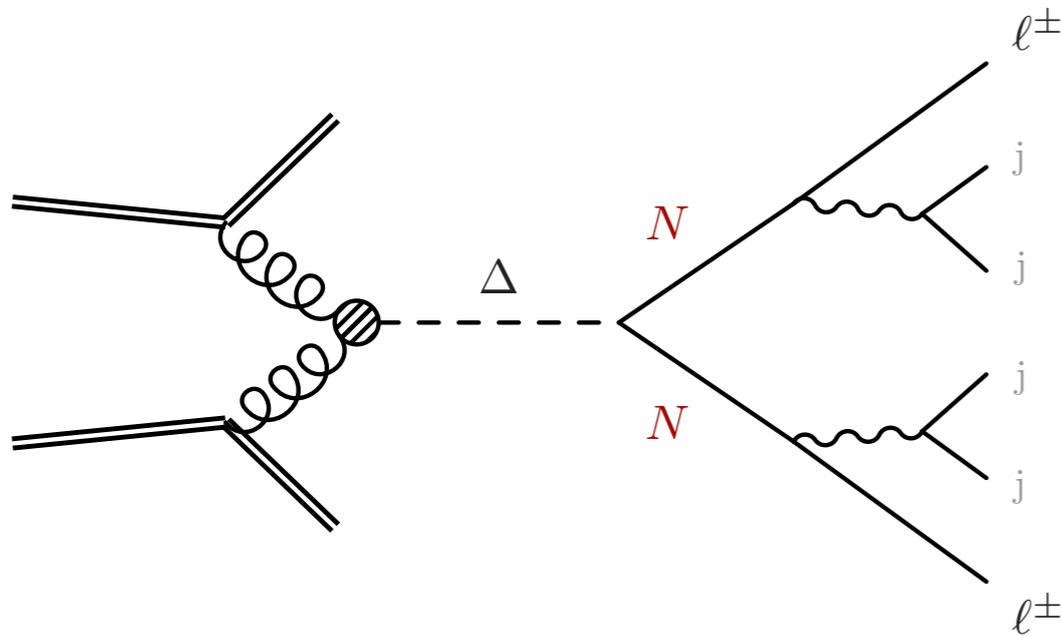
soft products $p_T \simeq m_h/6 \sim 20 \text{ GeV}$

Kiers et al. '02, Zhang et al. '07
Maiezza et al. '10, Senjanović, Tello '14

low background (LNV)

'Majorana' Higgses at LHC

$$\Delta L = 0, 2$$

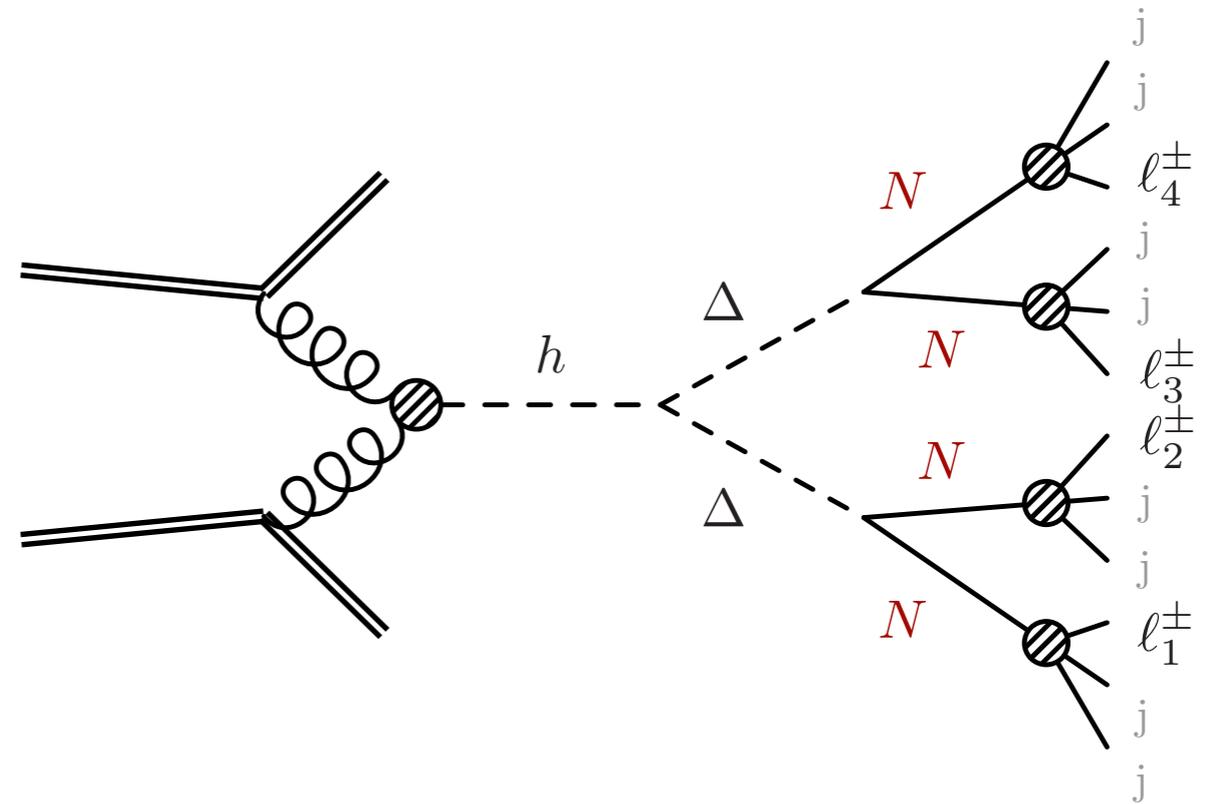


similar to $h \rightarrow NN$

ggF of CP even scalar

Anastasiou et al. '16

$$\Delta L = 0, 2, 4 \text{ MN, Nesti, Vasquez '16}$$



(same-sign) multi-leptons

$2^4 = 16$ possibilities

$$\Delta L_0 : \Delta L_2 : \Delta L_4 = 3 : 4 : 1$$

$$\mathcal{R}_{\Delta L}^{\#\ell} \Rightarrow \mathcal{R}_2^2, \mathcal{R}_3^3, \mathcal{R}_2^4, \mathcal{R}_4^4$$

Backgrounds

Selection criteria

	$t\bar{t}$	$t\bar{t}h$	$t\bar{t}Z$	$t\bar{t}W$	WZ	Wh	ZZ	Zh	$WWjj$	fakes
--	------------	-------------	-------------	-------------	------	------	------	------	--------	-------

Selection

$$\ell^\pm \ell^\pm + n_j$$

$$\cancel{E}_T$$

$$\cancel{E}_T < 30 \text{ GeV}$$

$$p_T$$

$$p_T(\ell_1) < 55 \text{ GeV}$$

$$m_T$$

$$m_{\ell p_T}^T < 30 \text{ GeV}$$

$$m_{\text{inv}}$$

$$m_{\ell\ell} < 80 \text{ GeV}$$

$$m_{\ell p_T} < 60 \text{ GeV}$$

$$l_{T\ell}$$

$$l_{T\ell} > 0.1 \text{ mm}$$

all contain missing energy

one prompt, one displaced lepton

Backgrounds

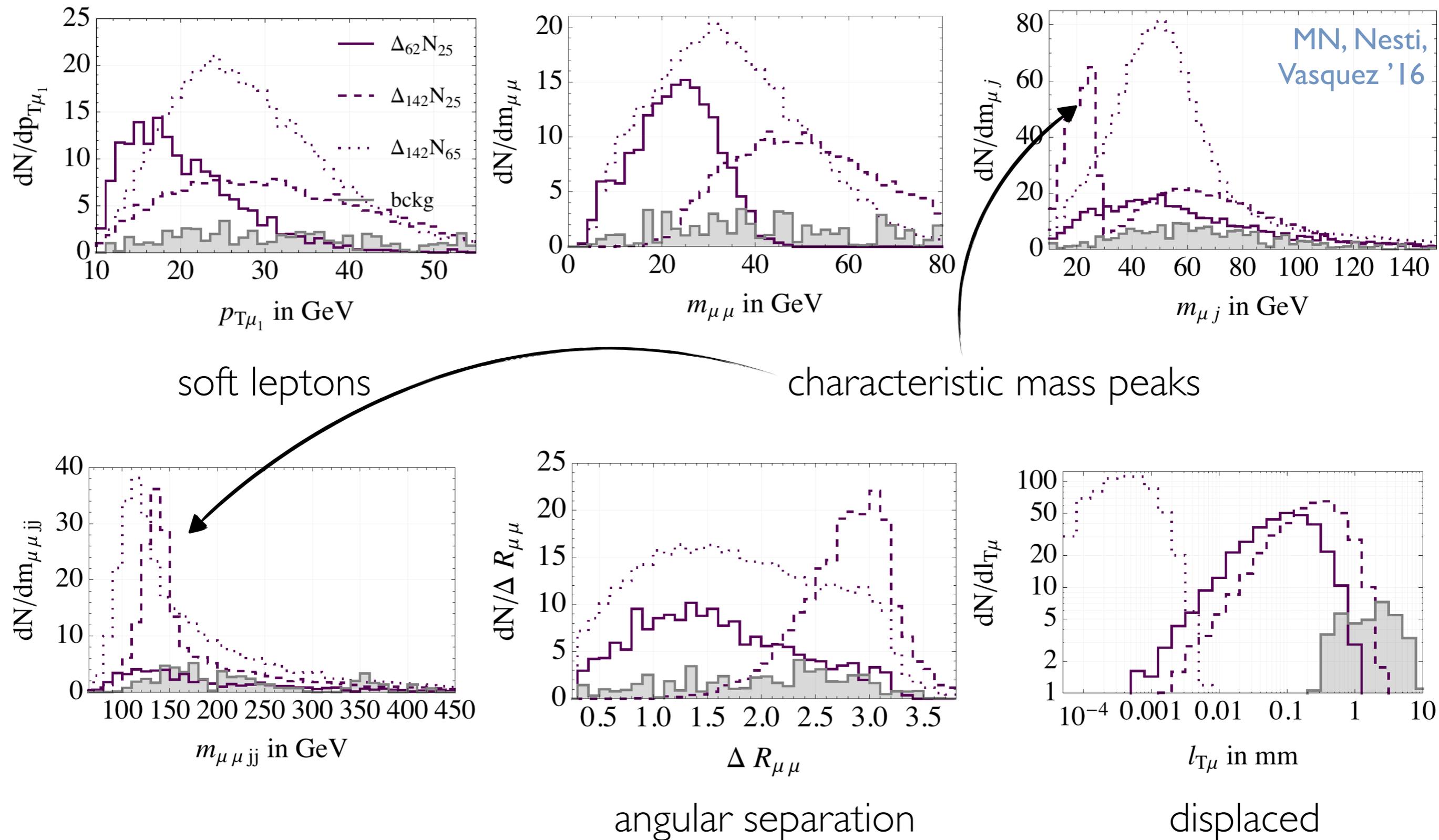
$$\ell^\pm \ell^\pm + n_j$$

	$t\bar{t}$	$t\bar{t}h$	$t\bar{t}Z$	$t\bar{t}W$	WZ	Wh	ZZ	Zh	$WWjj$	fakes
select	806	4	5	26	1241	87	147	16	1.5	2651
\cancel{E}_T	313	0.5	0.7	3	400	21	129	7	0.2	782
p_T	112	0.2	0.1	0.7	174	8.4	63	4	0.05	284
m_T	60	0.1	0.04	0.3	80	4	56	2	0.03	106
m^{inv}	35	0.03	0.03	0.2	25	2	36	2	0	80
l_{Te}	0	0	0	0	0.7	0.1	0.9	0.05	0.001	2
	$t\bar{t}$	$t\bar{t}h$	$t\bar{t}Z$	$t\bar{t}W$	WZ	Wh	ZZ	Zh	$WWjj$	fakes
select	670	4	6	32	750	133	68	16	2	1676
\cancel{E}_T	130	0.5	0.9	3.5	200	32	33	6	0.3	391
p_T	57	0.2	0.2	1	95	17	16	3	0.1	152
m_T	32	0.1	0.1	0.5	51	9	12	2	0.05	49
m^{inv}	17	0.04	0.04	0.2	23	5	8	1	0.01	40
$l_{T\mu}$	0	0	0	0	1.4	0.4	1	0.15	0.005	3

all contain missing energy

one prompt, one displaced lepton

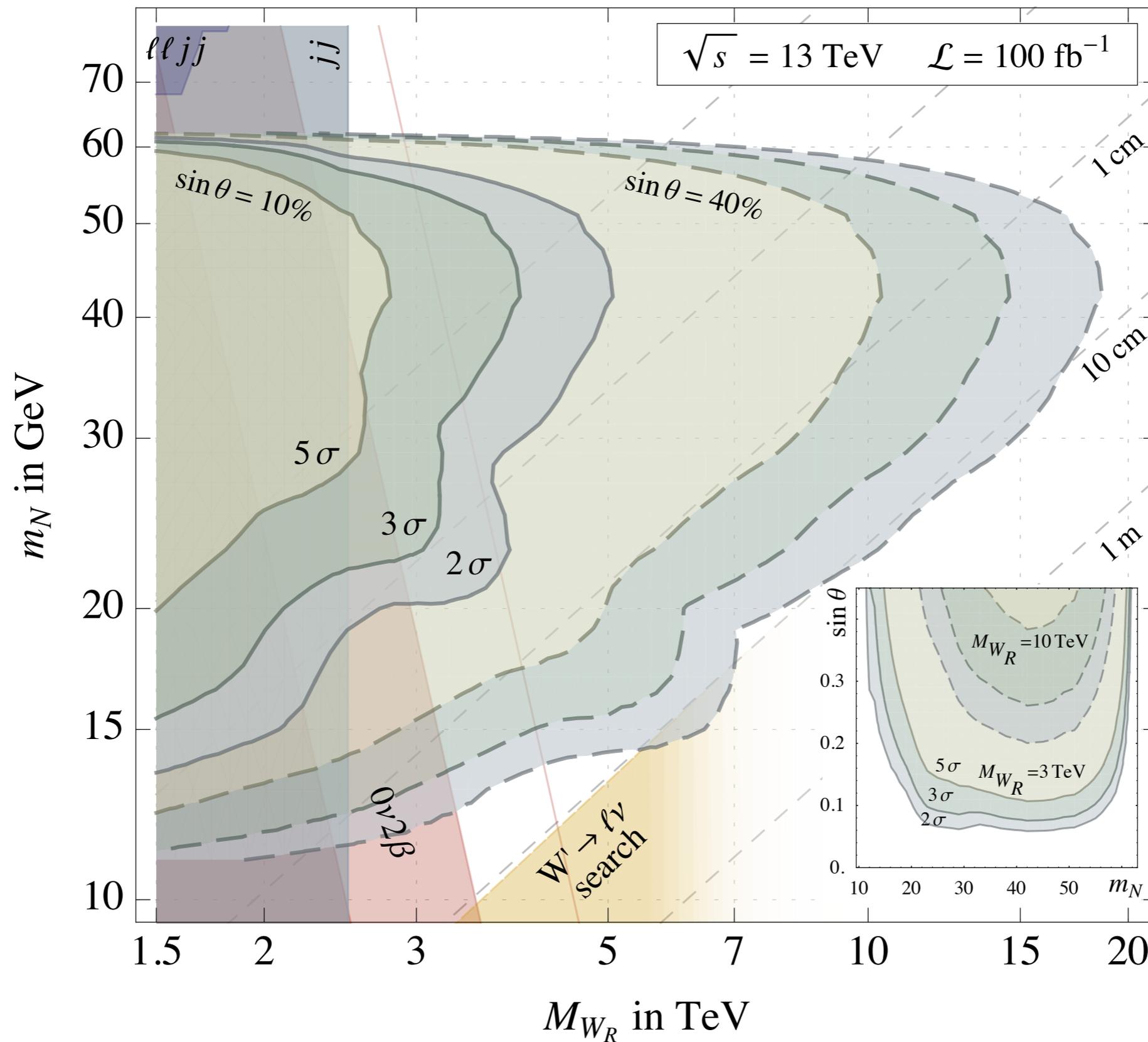
Signal features



Sensitivity

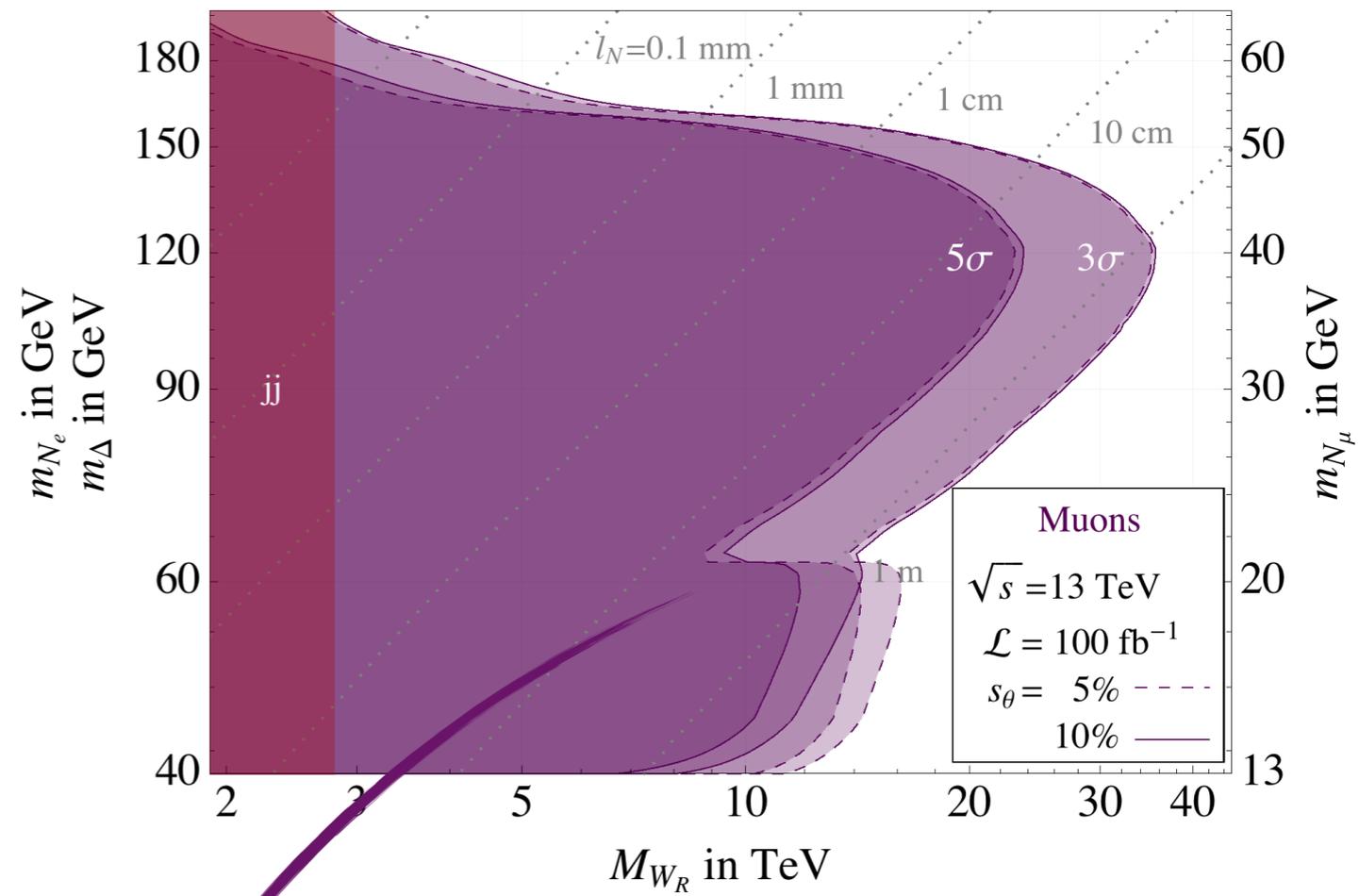
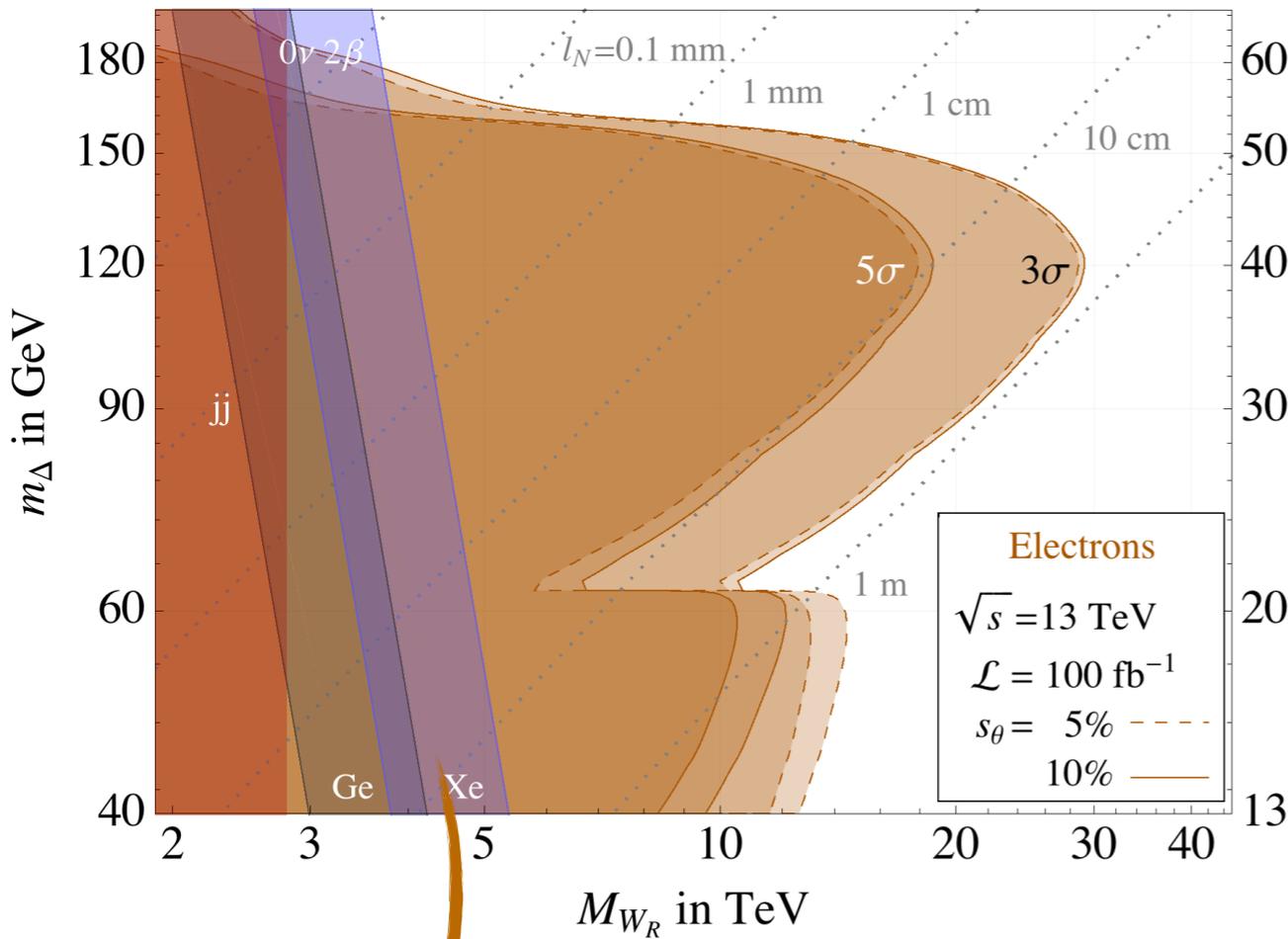
$h \rightarrow NN$

Maiezza, MN, Nesti '14



Sensitivity

Combined $h \rightarrow NN$ $\Delta \rightarrow NN$ $\Delta\Delta \rightarrow NNNN$



connection to $0\nu 2\beta$

GERDA, Neutrino '16

KamLAND-Zen '16

$h \rightarrow \Delta\Delta \rightarrow NNNN$

displaced $0.01 \text{ mm} - > 1 \text{ m}$

discovery reach beyond direct searches

more LNV Higgs candidates

No-go for vanilla see-saw(s)

Fourth generation $h \rightarrow \nu_4 \nu_4$

Pilaftsis '92
Carpenter '11

EFT from SM + h + N

Graesser '07
Caputo, Hernandez, Lopez-Pavon '17

SM + h + N + singlet scalar

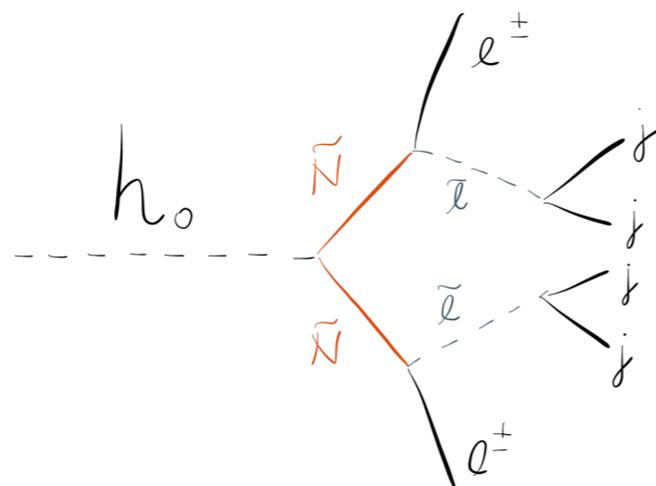
Shoemaker, Petraki, Kusenko '10

Spontaneous B-L

$SU(2)_L \times U(1)_R \times U(1)_{B-L}$

Deppisch, Mitra '18

RPV Susy



LNV disfavored

$$m_{\tilde{l}} \simeq m_{\tilde{\nu}}$$

needs post-LHC revision

Banks, Carpenter Fortin '08

Summary

Conclusions

Neutrino masses provide a clear path beyond the SM

Testing LNV is a fundamental issue, similar to baryon number

TeV Colliders can discover the origin of neutrino mass

Outlook

Perform the non-trivial searches, including other channels

Improve theoretical tools, vertexing, data driven backgrounds

More studies needed for future colliders, HE-HL LHC, FCC-ee, eh, hh

Thanks to

organizers for hospitality and
participants for talks and discussions

