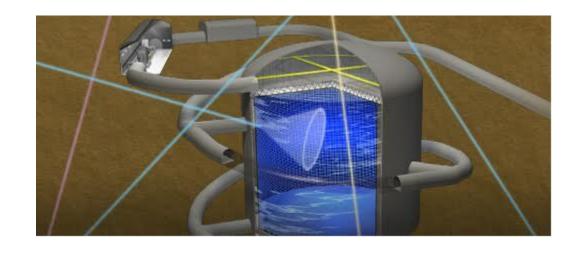


http://hyperk.org



Status of Hyper-Kamiokande T2HK

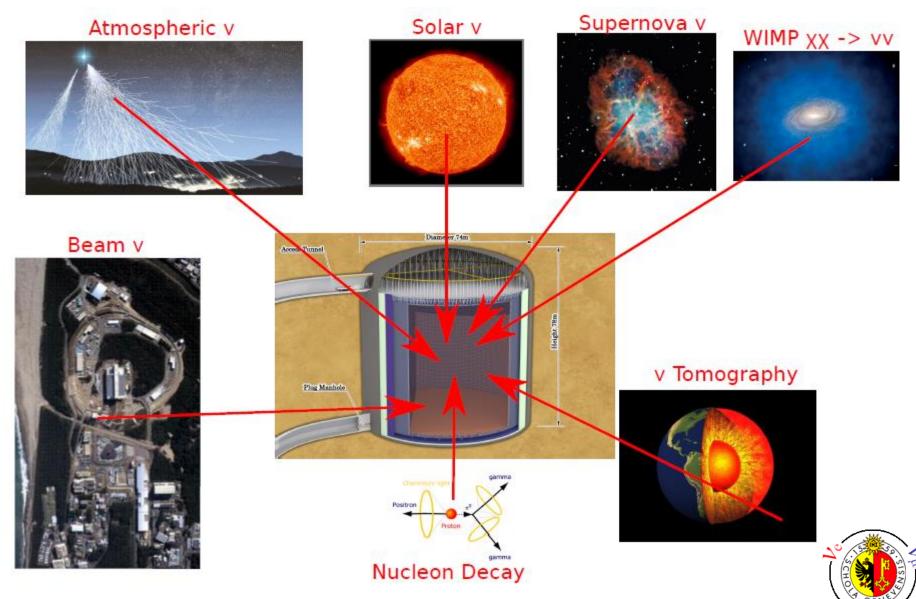
Alessandro Bravar on behalf of the HK Collaboration



TAU2021 Oct 01, 2021

Hyper-K Physics Overview





Broad Science Program with Hyper-K



Neutrino oscillation physics

comprehensive study with beam and atmospheric neutrinos determination of neutrino mass hierarchy determination of θ_{23} octant measurement of CP Violation in leptonic sector reveal exotic scenarios

Search for nucleon decay

possible discovery with ~10 × SK sensitivity all visible modes including p \rightarrow e⁺ π^0 and p $\rightarrow \overline{v}$ K⁺ reach 1035 years sensitivity

Solar neutrino physics

precision measurement of Δm_{21}^2 measurement of energy spectrum upturn discovery & measurement of hep neutrinos

Neutrino Astrophysics

high statistics measurement of SN burst neutrinos detection and study of relic SN neutrinos indirect Dark Matter search from Galactic Core, Sun, Earth

Geophysics ("neutrinography" of Earth's interior)



The Hyper-Kamiokande Detector



Large Water Cherenkov Detector

Larger mass for more statistics

Better sensitivity by more photons with improved sensors

Cherenkov light

Neutrino

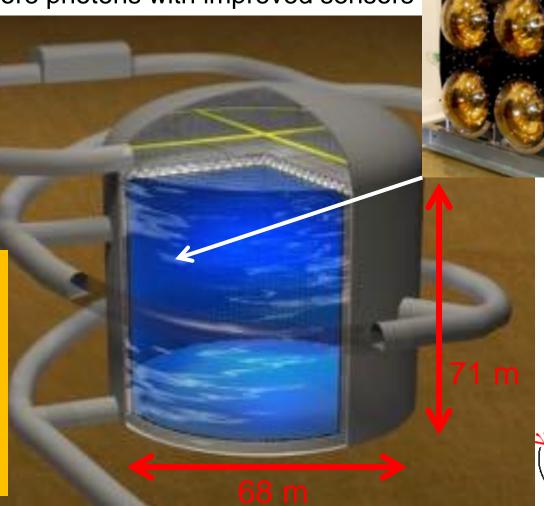
Charged particle in water

Photosensors

total volume 260 kton fiducial volume 190 kton

inner detector 20,000 50 cm PMTs few 1,000 M-PMTs

outer detector 6,700 8 cm PMTs



Upgraded Photo-Sensors

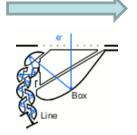




high QE photocathode



dynode improvement



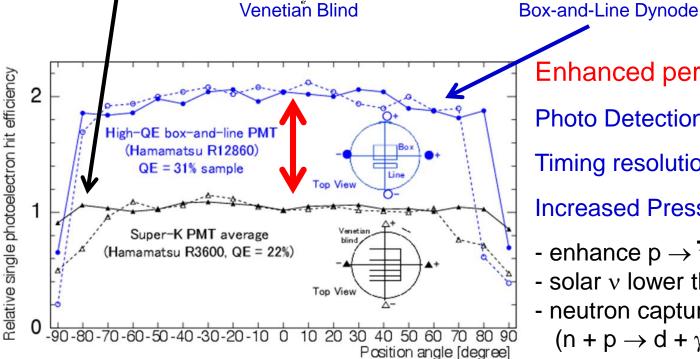
50 cm HQE **Box&Line PMT**

under validation





under validation



Enhanced performance

Photo Detection Efficiency 2 × bigger

Timing resolution 2 × as good

Increased Pressure tolerance × 2

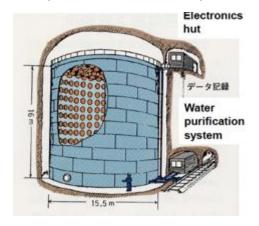
- enhance p $\rightarrow \overline{v}$ K⁺ signal
- solar v lower threshold
- neutron capture signature $(n + p \rightarrow d + \gamma - 2.2 MeV \gamma)$



3 Generations of Kamioka Detectors



Kamiokande (1983 – 1996)

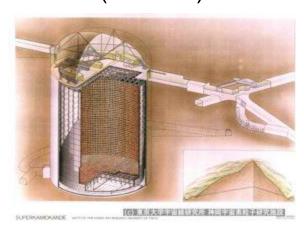


3 kton 20% coverage with 50 cm PMT



Observation of SN1987A

Super-Kamiokande (1996 –)

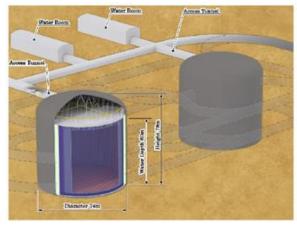


50 kton 40% coverage with 50 cm PMT



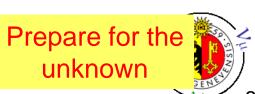
Discovery of v oscillations





260 kton 20% coverage with high-QE 50 cm PMT





The Hyper-K Collaboration





19 countries 93 institutes ~450 members (and growing)

January 2015

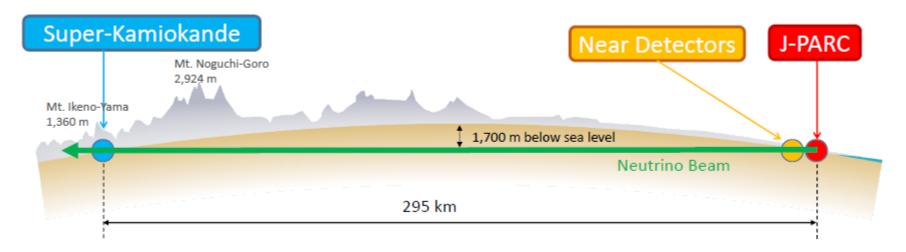




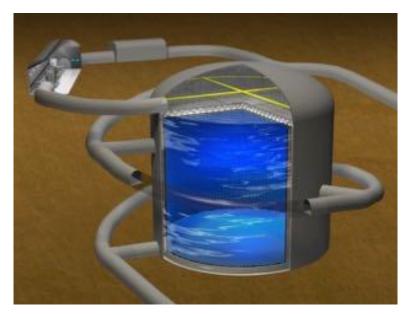
ICRR

From J-PARC to Kamioka (T2HK)





260 kton Water Cherenkov Detector

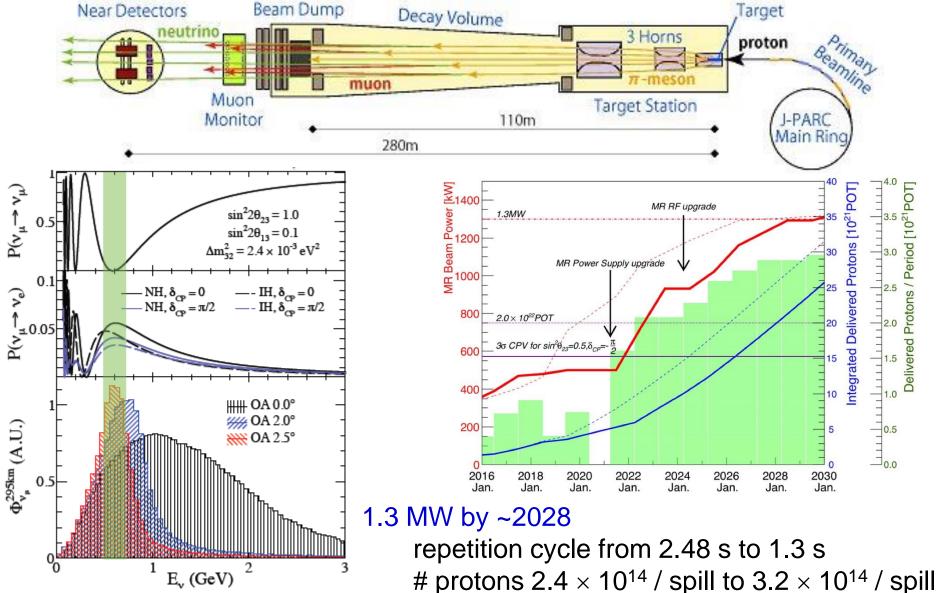


Upgraded J-PARC neutrino beam New / upgraded near detectors



Beamline Upgrade

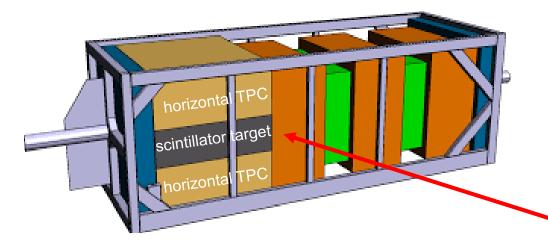




The Near Detectors @ J-PARC



upgraded ND280 Near Detector



designed to address v – Nucleus interactions and modeling enlarge phase space (4π coverage) efficiency for short hadron tracks with proton reconstruction

improve electron neutrino selection

new: horizontal TPCs scintillator target ToF

2.5°

Intermediate Water Cherenkov

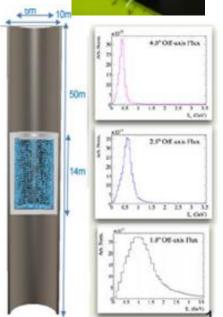
~600 ton water Cherenkov located at ~1 km from v source

off-axis angle spanning orientation vary ν peak energy

probe neutrino energy vs. reconstructed energy

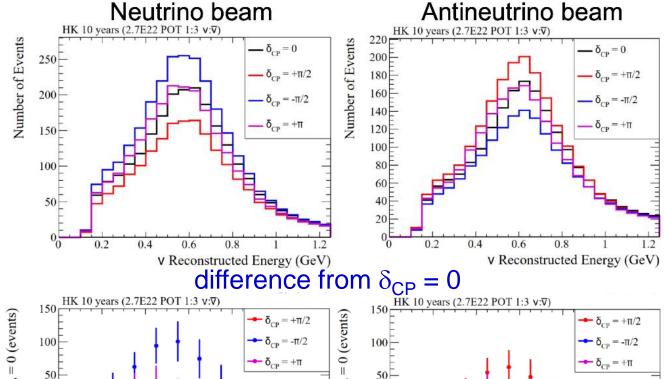
multi-PMT units (good reconstruction despite small det.)

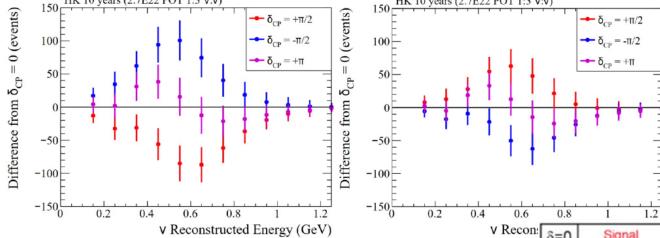
potential to loading with Gd to measure neutron production



Beam Events in T2HK: v_e appearance







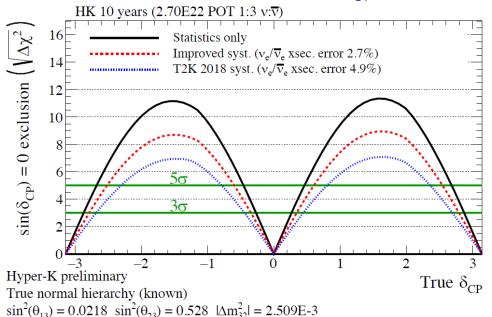
10 years data taking 190 kton fiducial × 1.3 MW

δ=0	Signal $(v_{\mu} \rightarrow v_e CC)$	Wrong sign appearance	$ \begin{array}{c} \nu_{\mu} , \overline{\nu}_{\mu} \\ \hline CC \end{array} $	Beam v_e , \bar{v}_e contamination	NC
V beam	2300	21	10	362	188
$\overline{\overline{V}}$ beam	1656	289	6	444	274

T2HK Sensitivity to δ_{CP}

assumes normal ordering known

significance to exclude $\delta_{CP} = 0$



$\sin \delta_{CP} = 0$ exclusion:

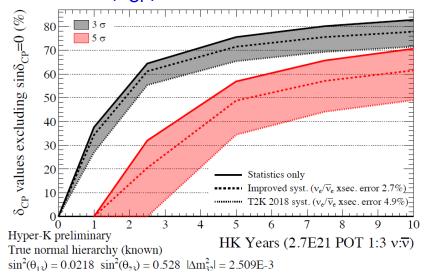
~8- σ significance if δ_{CP} = ±90°

 \sim 6-σ significance if δ_{CP} = ±45°

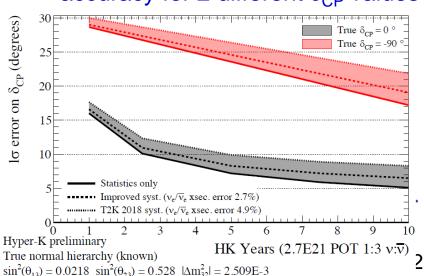
~80% coverage of δ_{CP} parameter space



fraction of δ_{CP} values for which $\sin(\delta_{CP}) = 0$ can be excluded



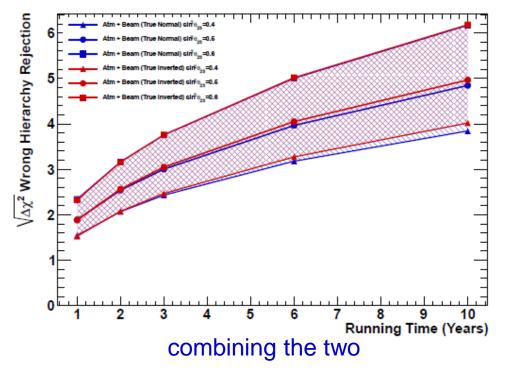
accuracy for 2 different δ_{CP} values



Mass Hierarchy

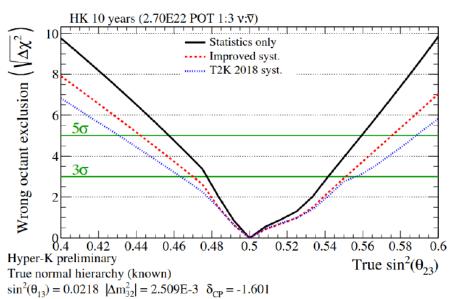


295 km baseline does not produce significant matter effects atmospheric neutrinos are sensitive to matter effects in Earth beam v + atmospheric v \rightarrow sensitivity to mass ordering atmospheric neutrinos allow to break possible degeneracies between MH and δ_{CP} when MH is unknown



>3 σ ability to reject wrong MH 5 σ for larger values of $\sin^2 \theta_{23}$

sensitivity to θ_{23} octant



More Physics with Atmospheric v



Atmospheric v:

neutrinos with various energy, flight length, and flavor

 v_{τ} cross section measurement

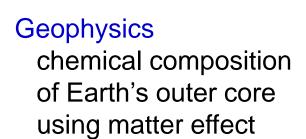
Sterile neutrinos

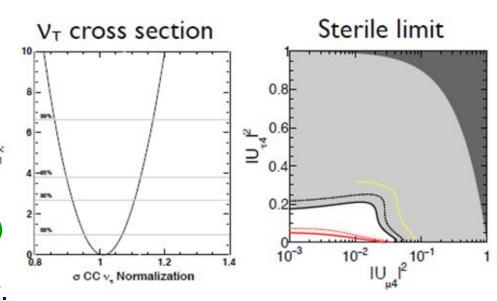
Lorentz violation studies

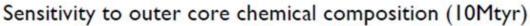
 $(3 - 4 \times \text{stronger than current SK limits})$

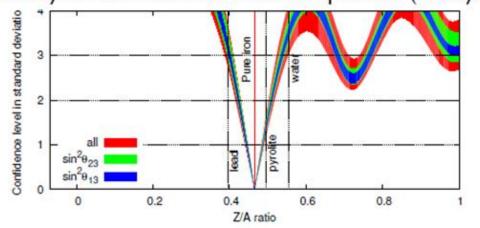
dark matter annihilation into SM part.

 $(3 - 5 \times \text{stronger than current SK limits})$



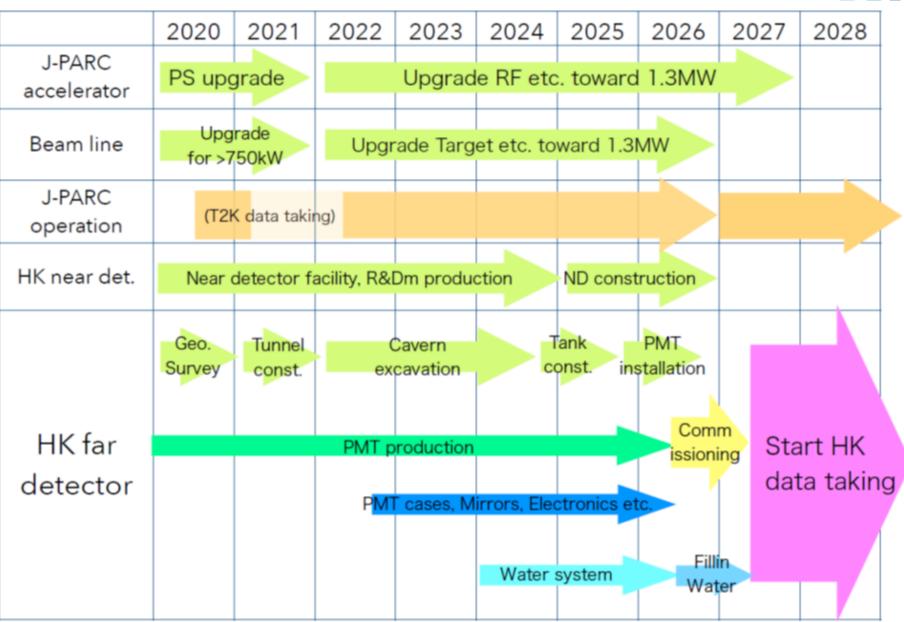






Hyper-K Schedule





5

Summary



A new adventure in v Physics to start

Hyper-K next generation neutrino experiment, 190 kton fiducial volume Hyper-K will address major open questions in science

Neutrino CP violation and mass ordering

Proton decay

Astrophysics (SN neutrinos)

.

Construction of Hyper-K water Cherenkov is ongoing (started in 2020)

J-PARC beam upgrade to 1.3 MW

Upgrade Near Detectors

Intermediate Water Cherenkov Detector

Start to take data in 2027

