# All $\nu_{\tau}$ 's, Great and Small

#### Tau Neutrino Physics with IceCube Spanning Six Orders of Magnitude in Energy

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# The IceCube+DeepCore Detector

- IceCube built in 2010 to map the  $\nu$  sky at  $E_{\nu} \sim 1~{\rm TeV}$ 
  - Find astrophysical v
  - Find astrophysical v sources
  - Help solve mystery of UHECR
- Enhanced with DeepCore
  - more densely instrumented region for DM and atm. ν osc.





# Neutrinos in IceCube: Sources

#### Atmospheric neutrinos

- cosmic rays (mainly protons) interact in the earth's atmosphere
- $\bullet$  resulting particle showers include  $\nu ' s$
- IceCube threshold  $E_{\nu} \sim 5$  GeV,  $E_{\nu}^{\text{atm.}} < \sim 10$  TeV;  $E_{\nu} \approx 10^{9-12}$  eV



- Astrophysical high energy neutrinos
  - created in cosmic accelerators, e.g., in particle jets created by black holes
  - Evident at  $E_{\nu} > \sim 50 \text{ TeV}$  in IceCube
    - IceCube has seen PeV-scale ( $10^{15}$  eV)  $\nu$ 's



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At higher energies, neutrino flavors can be distinguished:



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• IceCube has seen PeV-scale ( $10^{15}$  eV)  $\nu$ 's



# Atmospheric $\nu_{\tau}$

- Atmospheric  $\nu$ 's arise mainly from the decay of light mesons
  - At production, expect flux ratio of roughly

 $(\nu_e : \nu_\mu : \nu_\tau) :: (1 : 2 : 0)$ 

• No  $\nu_{\tau}$  there.

•  $\nu_{\tau}$  arise from  $\nu$  oscillations as they cross the earth:

$$P_{\nu_{\mu} \to \nu_{\tau}} \approx \sin^2(2\theta) \sin^2\left(1.27 \frac{\Delta m^2 L_{\nu}}{E_{\nu}}\right) \frac{[\text{eV}^2][\text{km}]}{\text{GeV}}$$

• E.g., 
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# Atmospheric $\nu_{\tau}$ Reconstruction

- Atm.  $\nu_{\tau}$  will produce particle showers in the detector
- We reconstruct the energy & direction of these showers, and can distinguish them from track-like events
- Look for excess of upward-going, shower-like interactions around 20 GeV

#### 21 GeV $\nu_{\tau}$ (sim.)





IceCube/Upgrade  $(N_{\gamma} much \text{ larger})$ 



#### IceCube/DeepCore

### First IceCube $\nu_{\tau}$ Appearance Measurement



### Future IceCube $\nu_{\tau}$ Appearance Measurements



# Astrophysical $\nu: E_{\nu} > 10^{13-14} \text{ eV}$

- $\nu$  mainly from  $\pi^{\pm}$  decay in astrophysical beamdumps
- Needle in a haystack!
  - $10^{11}$  atmospheric  $\mu$ /yr,
  - $10^5$  atmospheric  $\nu$ /yr, and
  - $10^1$  astrophysical  $\nu/{
    m yr}$

- Beat down atm. μ using part of detector as veto (see below)
- Separate atm.  $\nu$  from astrophys.  $\nu$  using  $E_{\nu}$ , spatiotemporal concidence, and/or event topology



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# Astrophysical $\nu_{\tau}$

- Measurements to date:
  - Search for clean "double pulse" waveforms arising from "twocascade" signature
  - Two  $u_{\tau}$  candidates found:
    - With
      - 1:1:1 flavor ratio at earth
      - $\Phi(\nu) \propto E_{\nu}^{-2.87}$

expect 1.5 signal + 0.8 background in 7.5 yrs

• (Estimate that 98% and 76% of events like the two seen are  $\nu_{\tau}$ -induced)



# Astrophysical $\nu_{\tau}$

- Joint flavor analysis:
  - First time best fit point with  $(\Phi_{\nu_e}, \Phi_{\nu_{\mu}}, \underline{\text{and}} \ \Phi_{\nu_{\tau}}) \neq 0$
  - $\bullet$  First probe of  $\nu$  flavor oscillations over cosmic baselines & at the TeV scale
  - Rules out no- $u_{ au}^{\mathrm{astro.}}$  hypothesis at  $2.8\sigma$



#### "Flavor Triangle"



## Future Astrophysical $\nu_{\tau}$ Measurement

- Waiting for a clean "double bang" would require much patience:  $E_{\nu_{\tau}} > \sim \text{PeV}$  are rare.
- Instead use more plentiful "double pulse"  $\nu_{\tau}$  events at lower threshold energies:  $E_{\nu_{\tau}} > \sim 50 \text{ TeV}$
- Follow in footsteps of previous analyses, but look for DP signature on 3 strings (180 vs. 1–2 modules)
  - Render each string into a 2-D image
  - Identify DP signal(s) using deep convolutional neural networks



## Future Astrophysical $\nu_{\tau}$ Measurement

- Preliminarily predict ~5.5  $\nu_{\tau}^{CC}$  on background of ~0.3 events
  - 10 years livetime
  - background dominated by other–flavor astrophysical  $\nu$
  - systematic effects appear to have minimal impact
- With ~5 events, can rule out  $no-\nu_{\tau}^{astro.}$  at high confidence
- May be able to better constrain astrophysical neutrino "flavor triangle"
- Also: Exploring supra-PeV  $\nu_{\tau}$  producing kms-long  $\tau$  tracks
  - Potentially distinguishable from  $\mu$  tracks (smoother:  $m_{\tau} \gg m_{\mu}$ )

After opening the box, here's what the triangle plot might look like for two selected values of events seen:



Blue lines from IceCube Collaboration, Phys. Rev. D 99, 032004 Orange lines lack full systematic treatment.

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# Conclusions

- $\bullet$  IceCube is unique in its broad sensitivity to  $\nu_{\tau}$  and  $\tau$ 
  - •~6 orders of magnitude in  $E_{\nu_{\tau}}$  and  $E_{\tau}$
  - ~20 orders of magnitude in  $L_{\nu_{\tau}}$
- $\bullet$  IceCube makes both inclusive and exclusive measurements of  $\nu_{\tau}$  and  $\tau$ 
  - Inclusive:  $\nu_{\mu}^{\text{atm}} \rightarrow \nu_{\tau}^{\text{atm}}$  appearance with world's largest sample (thousands of  $\nu_{\tau}$  and  $\tau$ )
    - Fundamental  $\nu$  oscillation measurement
    - Current measurement in agreement with standard  $\nu$  osc. picture; future measurements will have compelling sensitivity to non-standard physics
  - Exclusive:  $\nu_{\tau}^{\text{astro}} \rightarrow \tau \rightarrow \text{double pulse, not with world's largest sample (yet)}$ 
    - Powerful probe of ultra-long baseline, ultra-high energy  $\nu$  oscillations, and of astrophysical accelerator  $\nu$  production scenarios
    - Very early days for  $u_{ au}^{\mathrm{astro}}$ ...but maturing rapidly!



