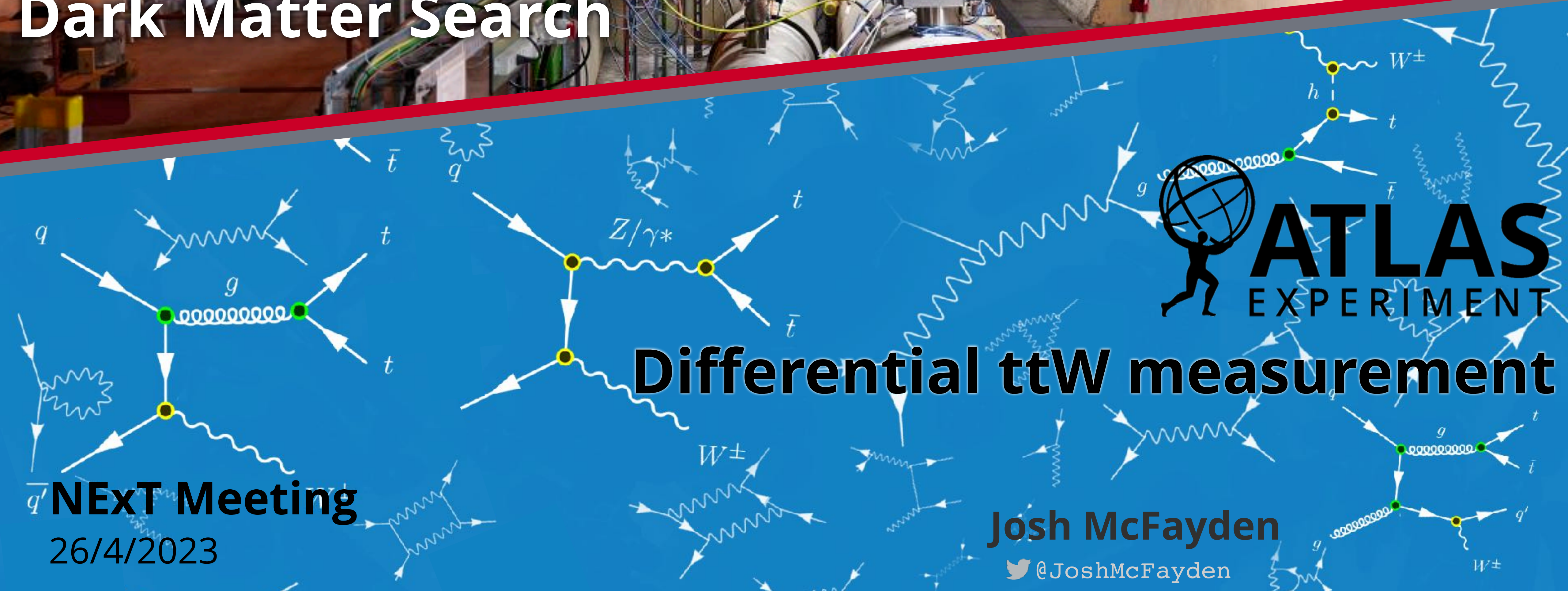


FASER

THE
ROYAL
SOCIETY

US
UNIVERSITY
OF SUSSEX


Observation of collider neutrinos Dark Matter Search



 **ATLAS**
EXPERIMENT

Differential ttW measurement

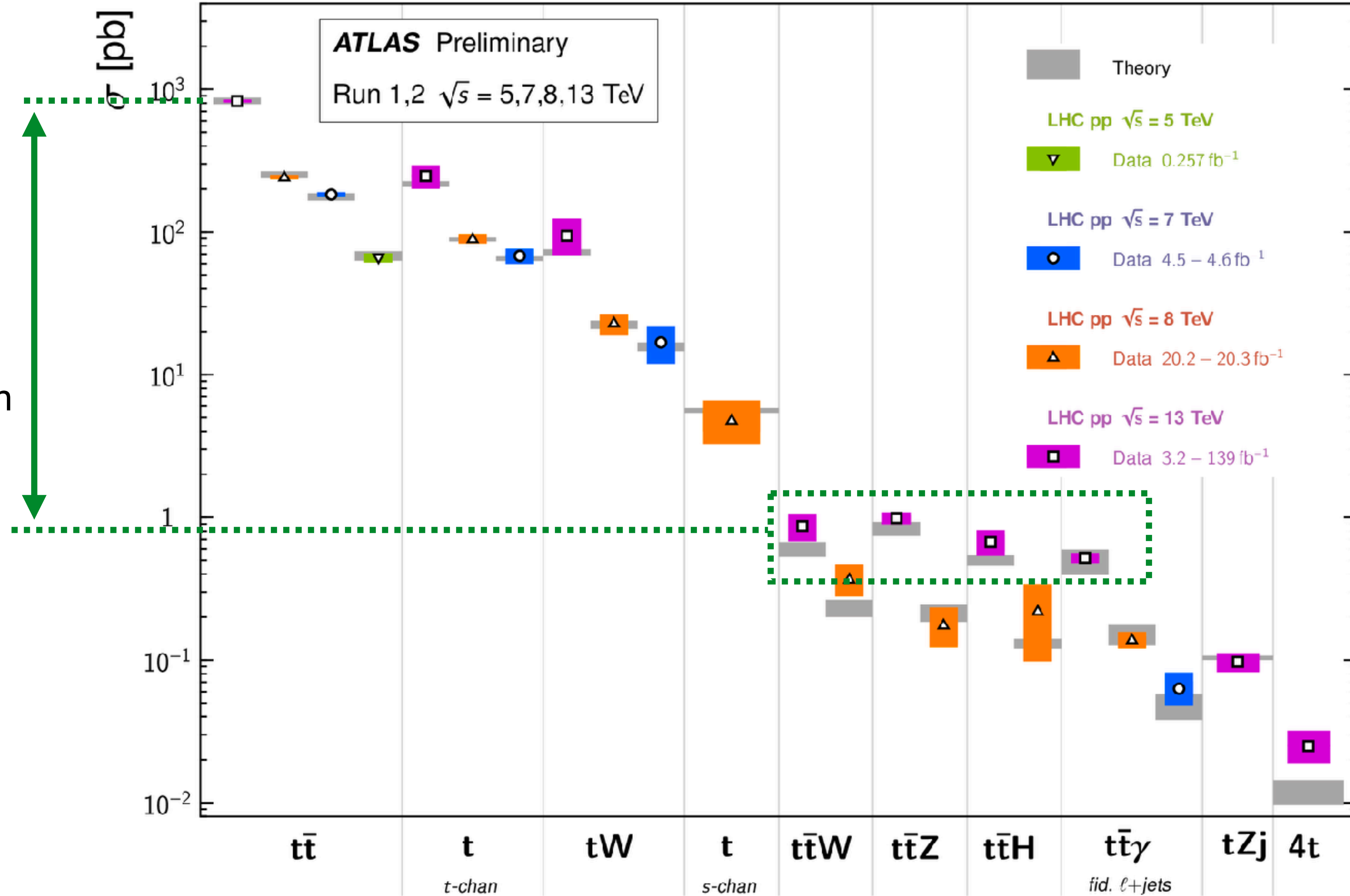
NExT Meeting
26/4/2023

Josh McFayden
 @JoshMcFayden

Top Quark Production Cross Section Measurements

Status: May 2021

3 orders of magnitude below the inclusive $t\bar{t}$ cross section



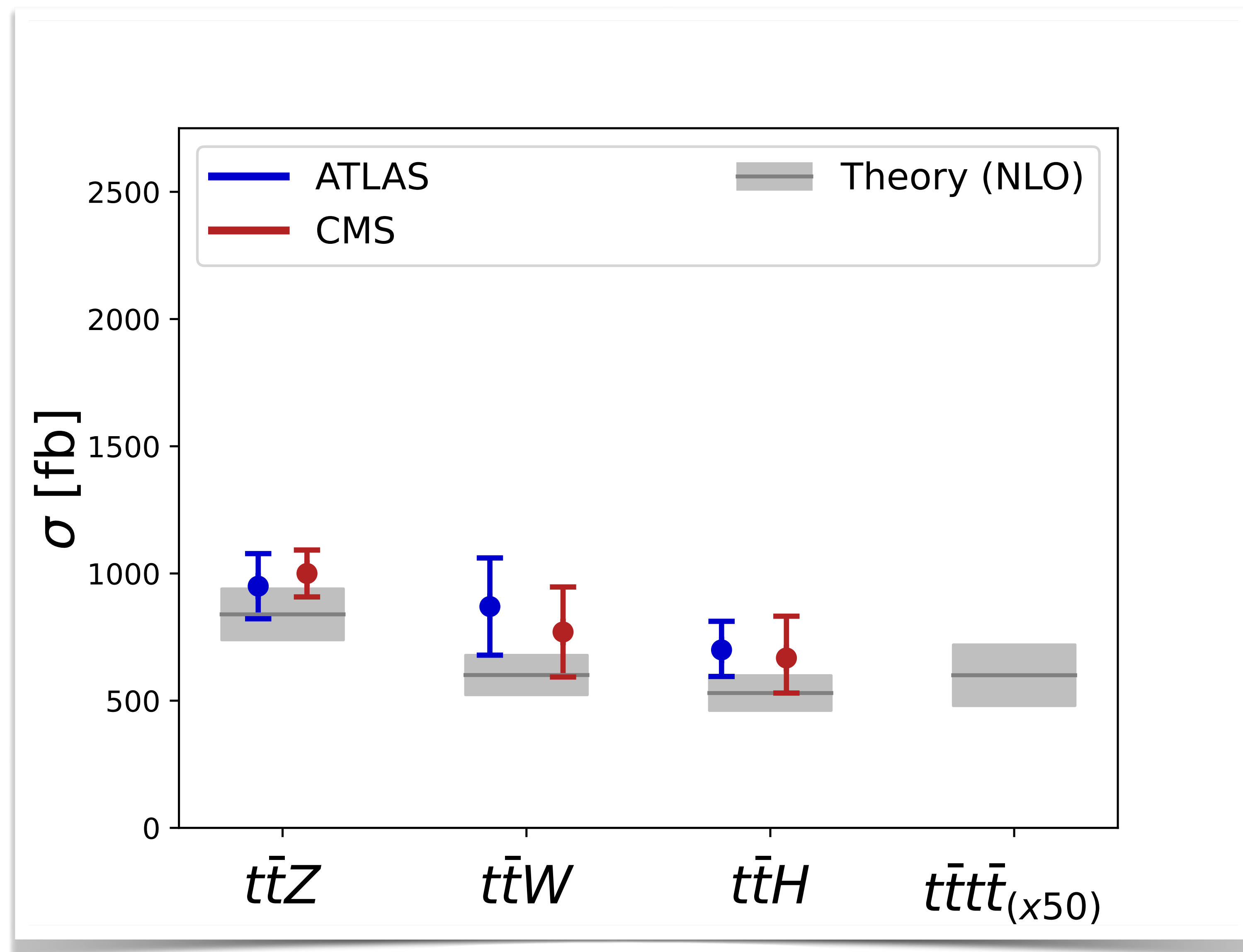


Rare top quark associated production

- ▶ Rare Standard Model processes accessible only at the LHC.
 - ▶ Just reaching sufficient sensitivity for differential measurements
 - ▶ Measurement sensitivity just starting to exceed that of theoretical predictions
- ▶ Important backgrounds to new physics searches
 - ▶ ttZ in top+MET SUSY searches
 - ▶ ttW is major background in any same-sign di-lepton final state
- ▶ Rich phenomenology
 - ▶ Many final states, EFT interpretations
- ▶ ttW deviation from SM...?

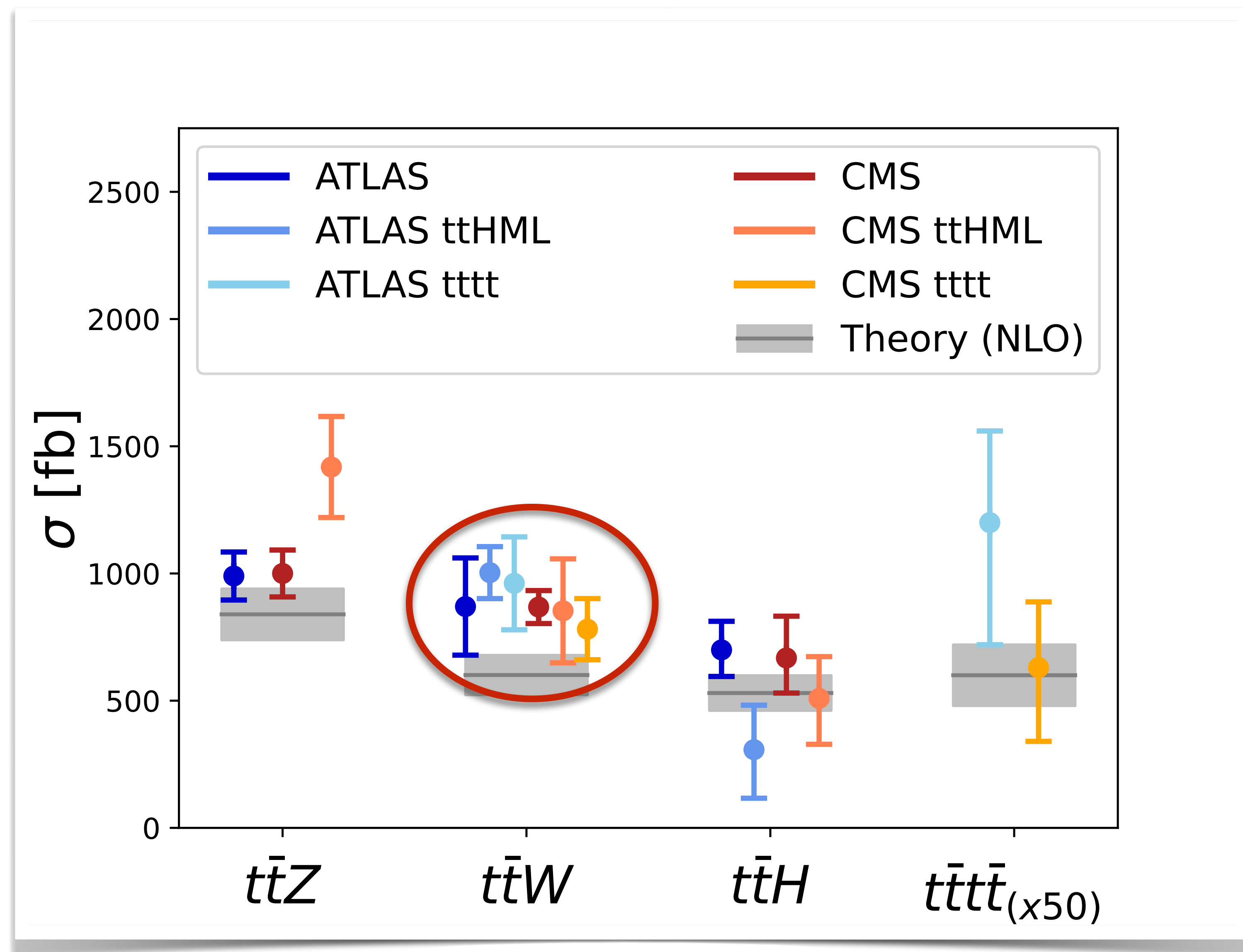
Why $t\bar{t}W$?

- ▶ Started off as barely a hint of deviation from the SM



Why ttW?

- ▶ Started off as barely a hint of deviation from the SM
- ▶ Then mounting evidence that something is going on with ttW (or ttW-like final states)



ttW as background to ttH

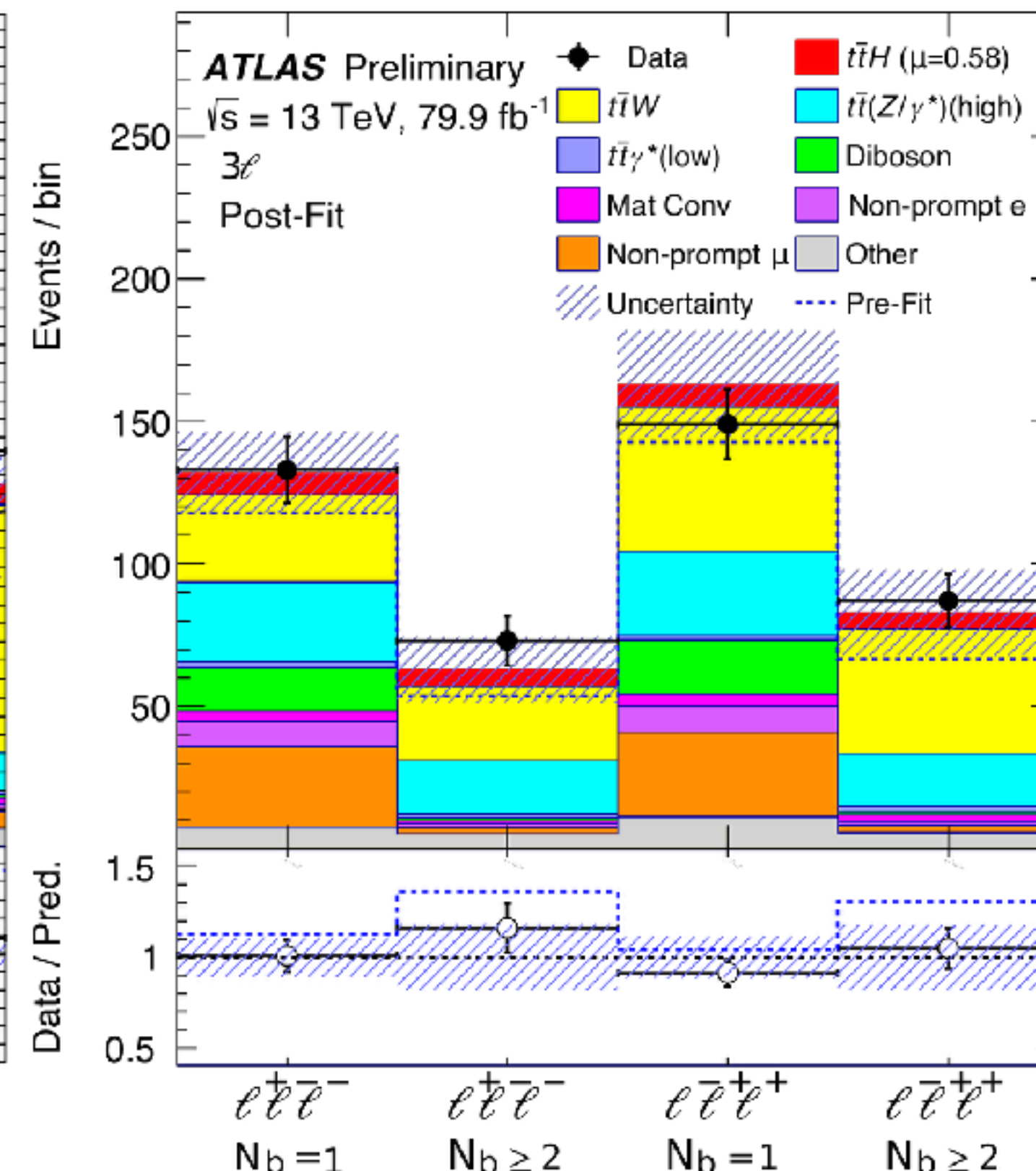
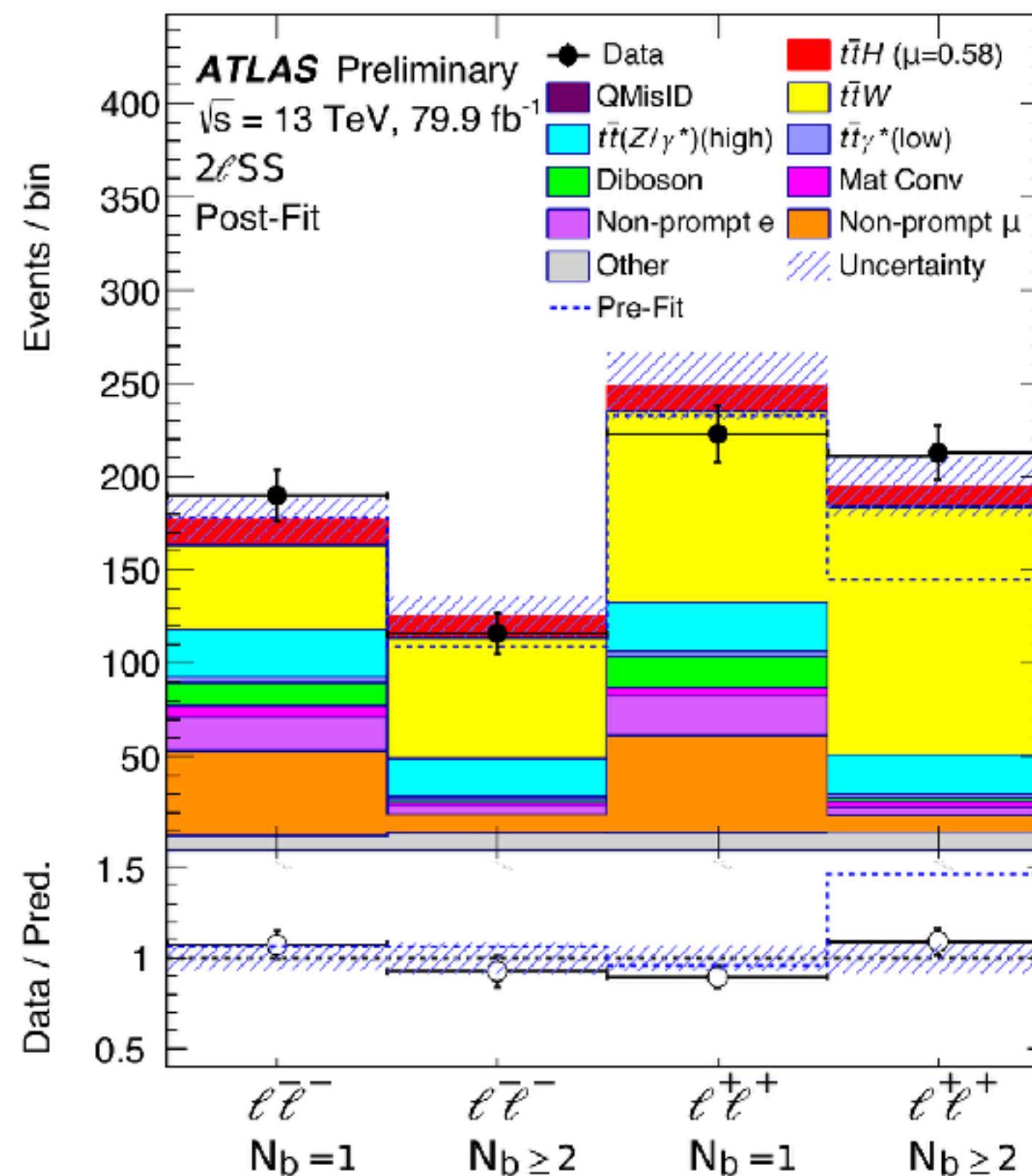
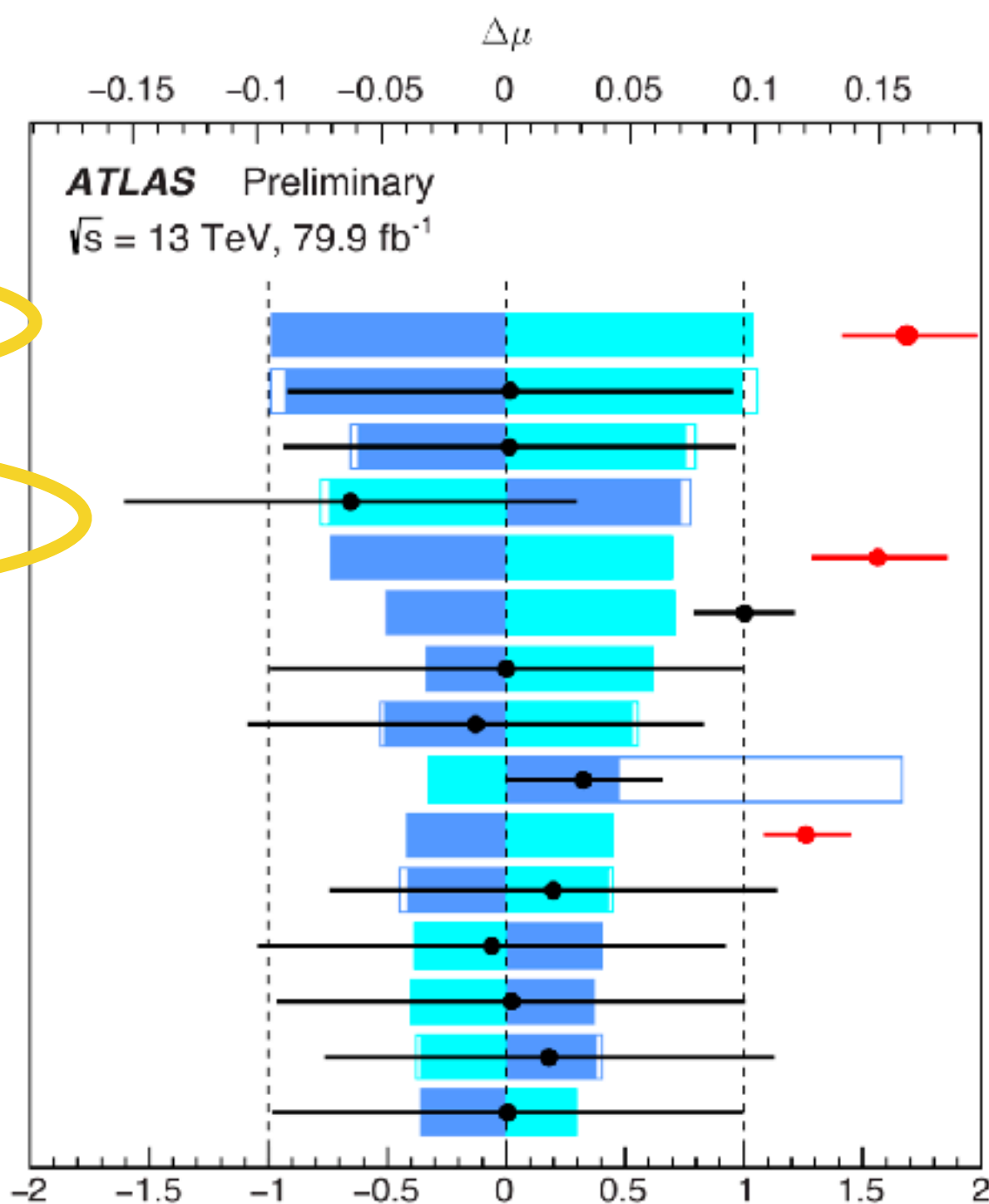
▶ In $ttH \rightarrow WW/ZZ/\tau\tau \rightarrow$ multi-leptons measurements ttW is a dominant irreducible background

▶ Also a dominant uncertainty:

Pre-fit impact on μ :
 $\theta = \hat{\theta} + \Delta\theta$ $\theta = \hat{\theta} - \Delta\theta$

Post-fit impact on μ :
 $\theta = \hat{\theta} + \Delta\hat{\theta}$ $\theta = \hat{\theta} - \Delta\hat{\theta}$

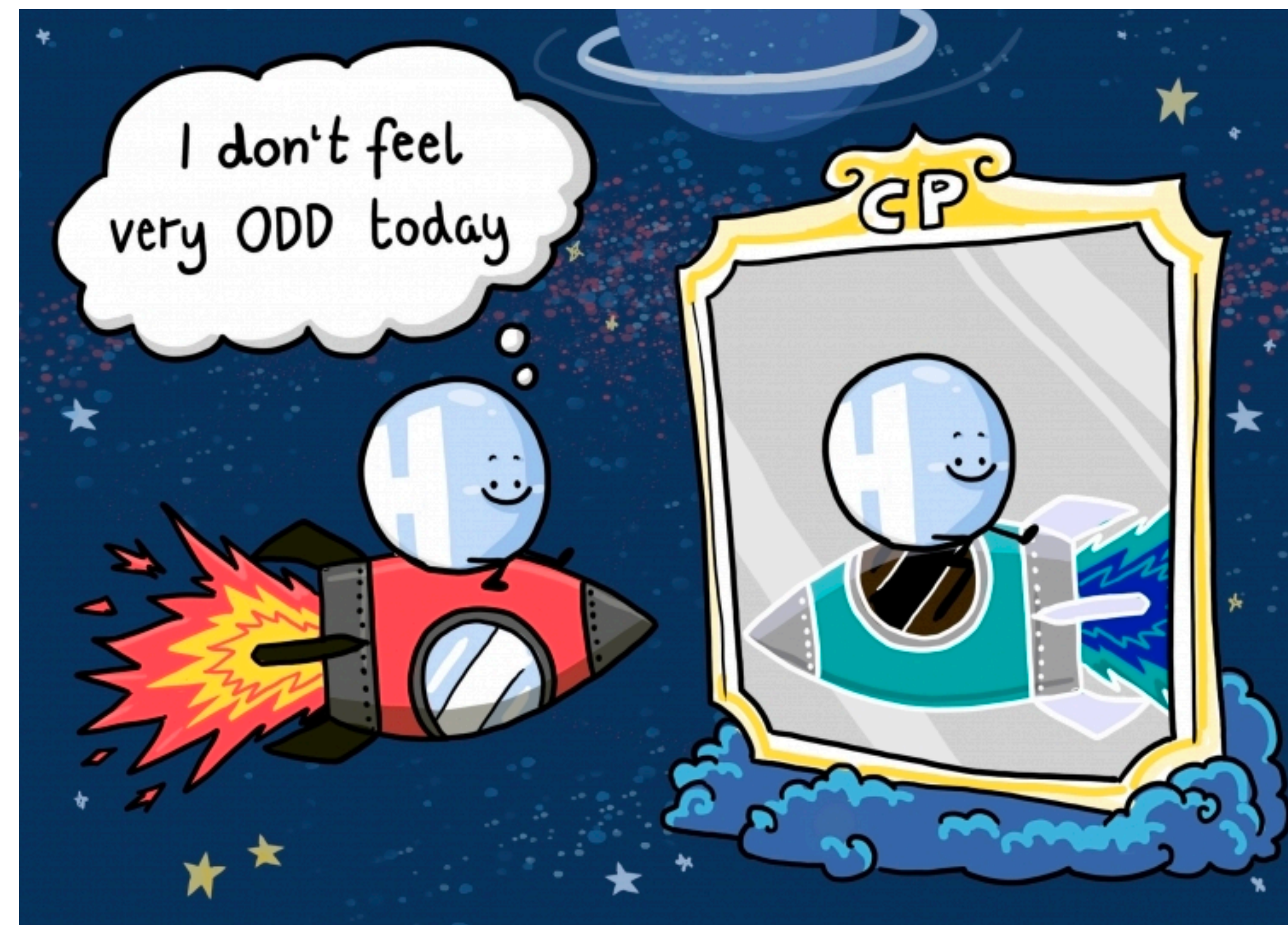
—●— Pull: $(\hat{\theta} - \theta_0) / \Delta\theta$
 —●— Norm. Factor



▶ Normalisation measured to be ~40% higher than best SM prediction at the time.

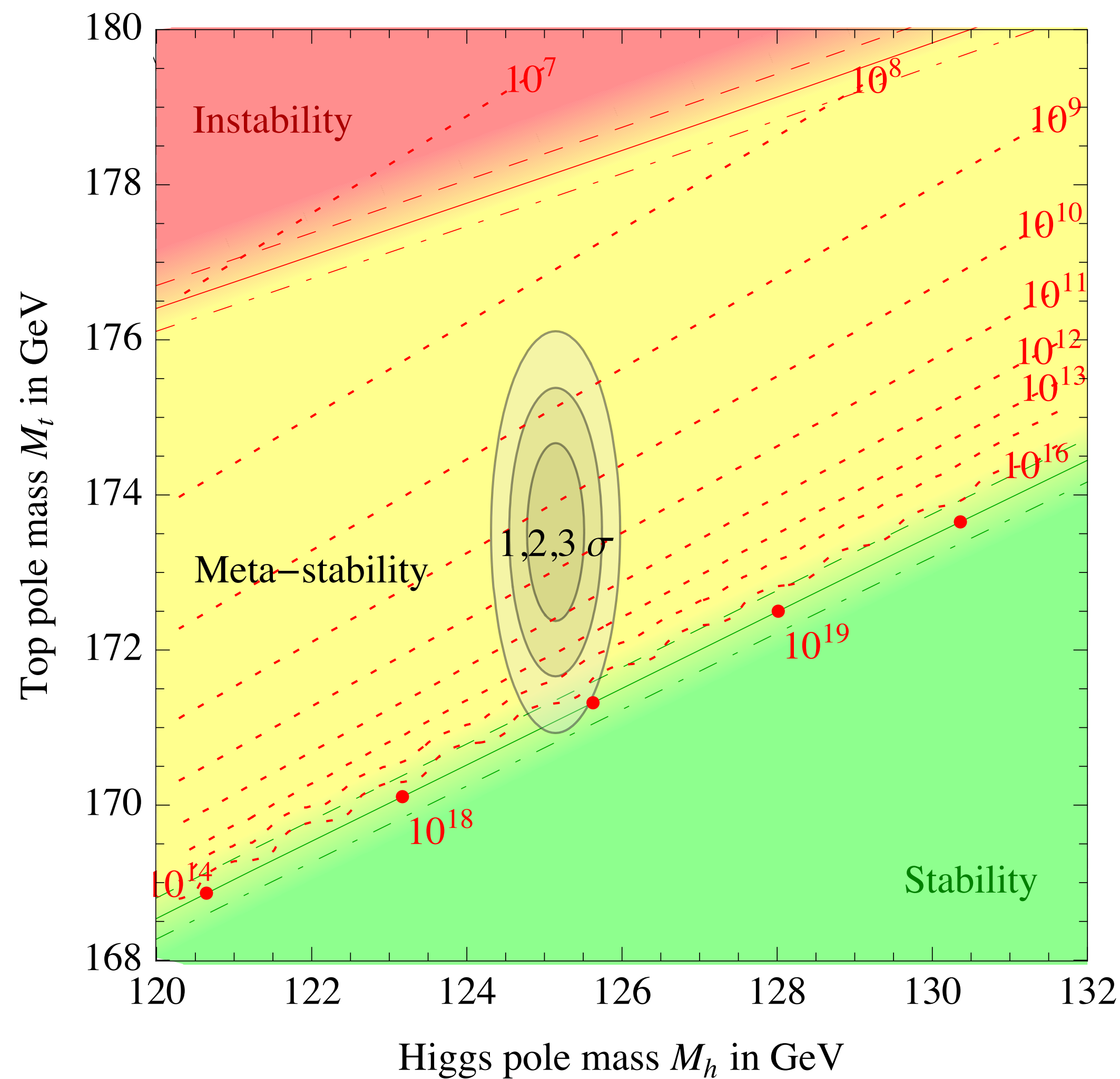
Why ttH?

- ▶ Window into top-Higgs coupling
 - ▶ Very important test of EW symmetry breaking
- ▶ Higgs CP properties
- ▶ Higgs self-coupling

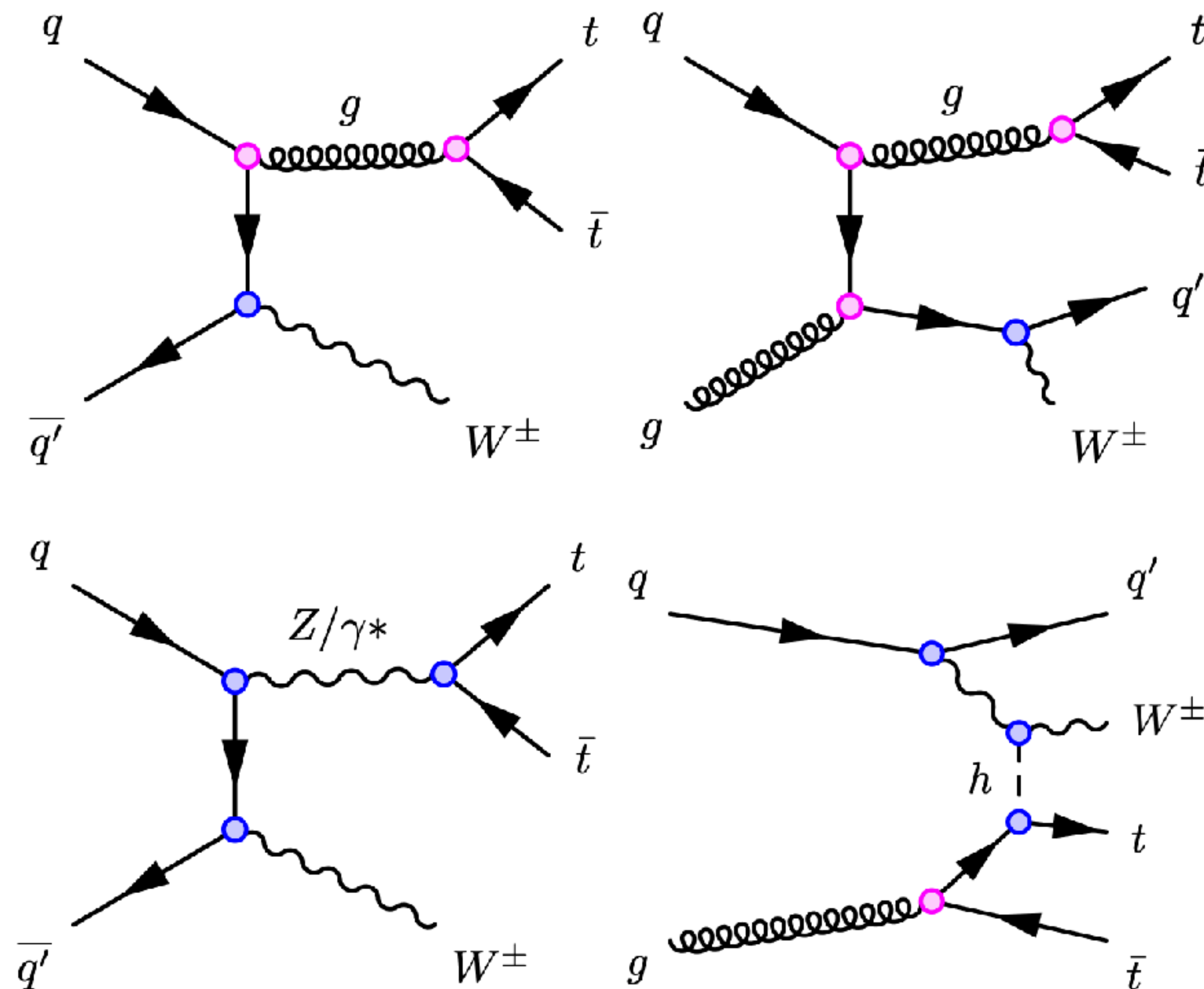


Why ttH?

- ▶ Window into top-Higgs coupling
 - ▶ Very important test of EW symmetry breaking
- ▶ Higgs CP properties
- ▶ Higgs self-coupling



- ▶ qq-initiated process
- ▶ Large NLO QCD corrections
 - ▶ NLO multi-leg merging is important
- ▶ Large NLO EWK corrections
- ▶ Effects of resummation are small
- ▶ No NNLO calculations!

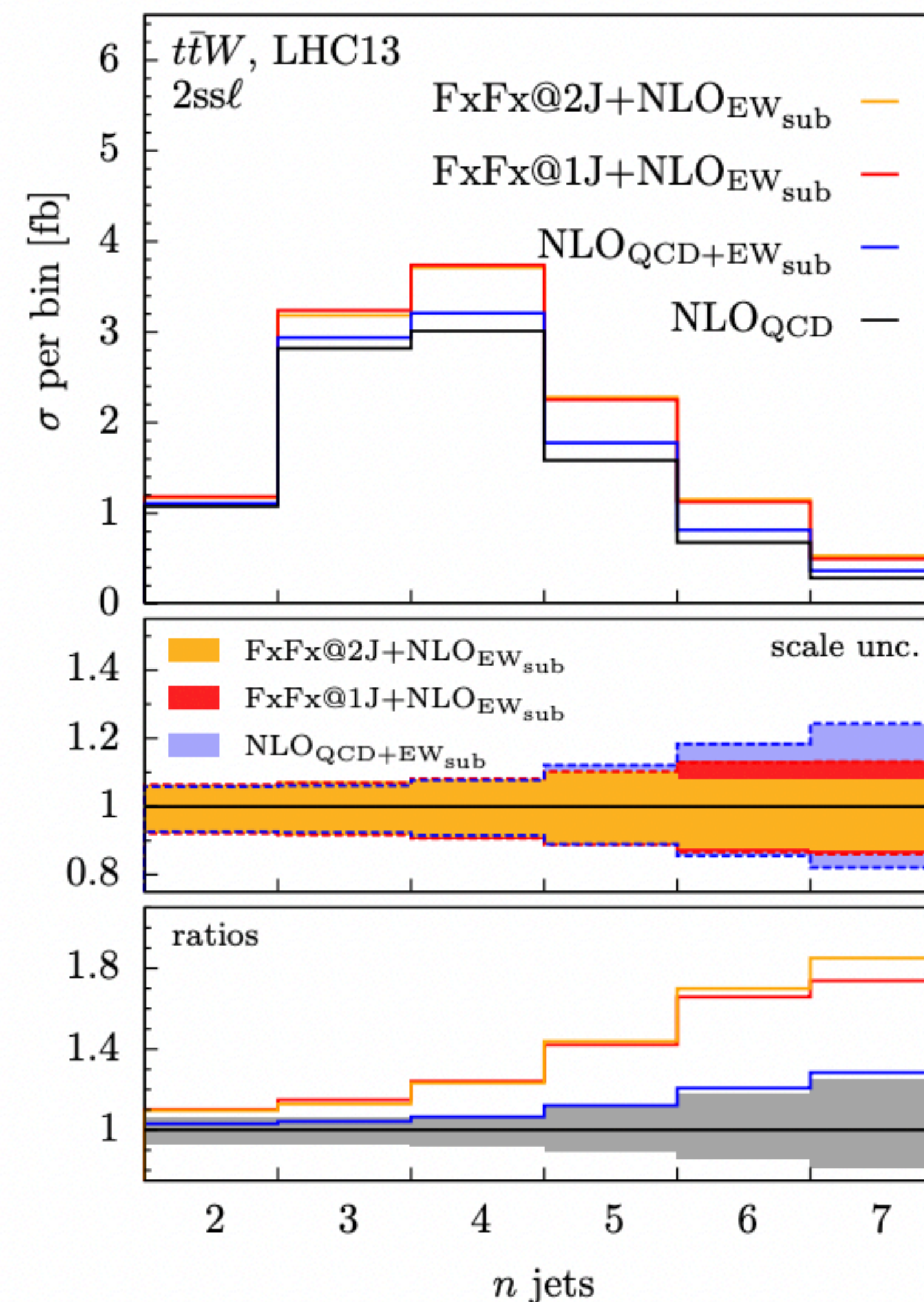


Theory predictions

► NLO + EW corrections from Powheg, MG5_aMC & Sherpa [[2101.11808](#)]

► Multi-leg merged 0-2j@NLO + EW corrections from Frederix et al. [[2108.07826](#)]

Order (default scale)	$\sigma \pm \text{scale} \pm \text{PDF}$ [fb]
FxFx@2J	$691.1(8)^{+65.7(+9.5\%)}_{-74.1(-10.7\%)} \quad +7.3(+1.1\%)_{-7.3(-1.1\%)}$
FxFx@2J+NLO _{EW} ^{sub}	$738.8(8)^{+75.0(+10.1\%)}_{-81.3(-11.0\%)} \quad +7.5(+1.0\%)_{-7.5(-1.0\%)}$
FxFx@2J+NLO _{EW} ^{lead} +NLO _{EW} ^{sub}	$722.4(8)^{+70.2(+9.7\%)}_{-77.7(-10.8\%)} \quad +7.2(+1.0\%)_{-7.2(-1.0\%)}$

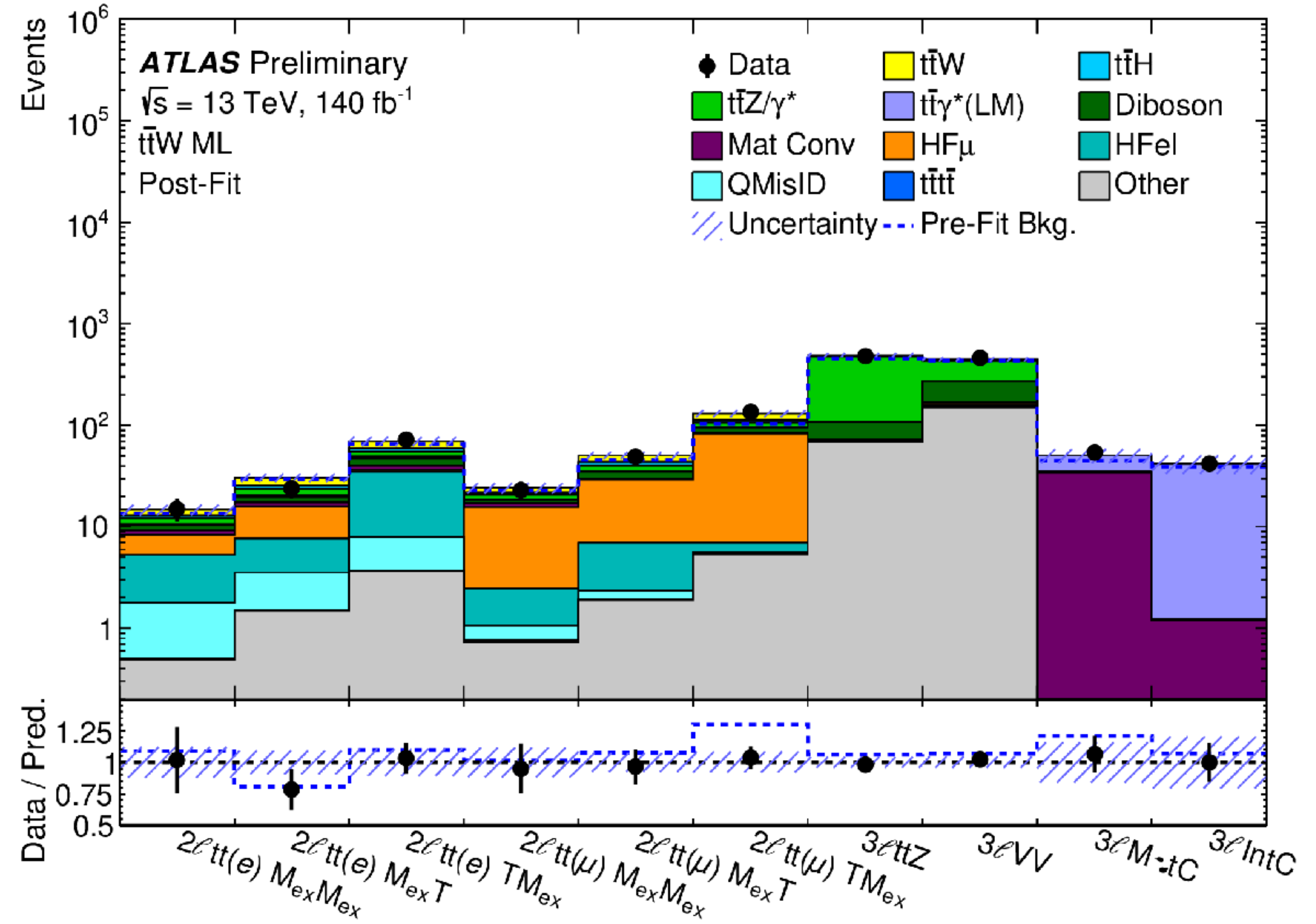
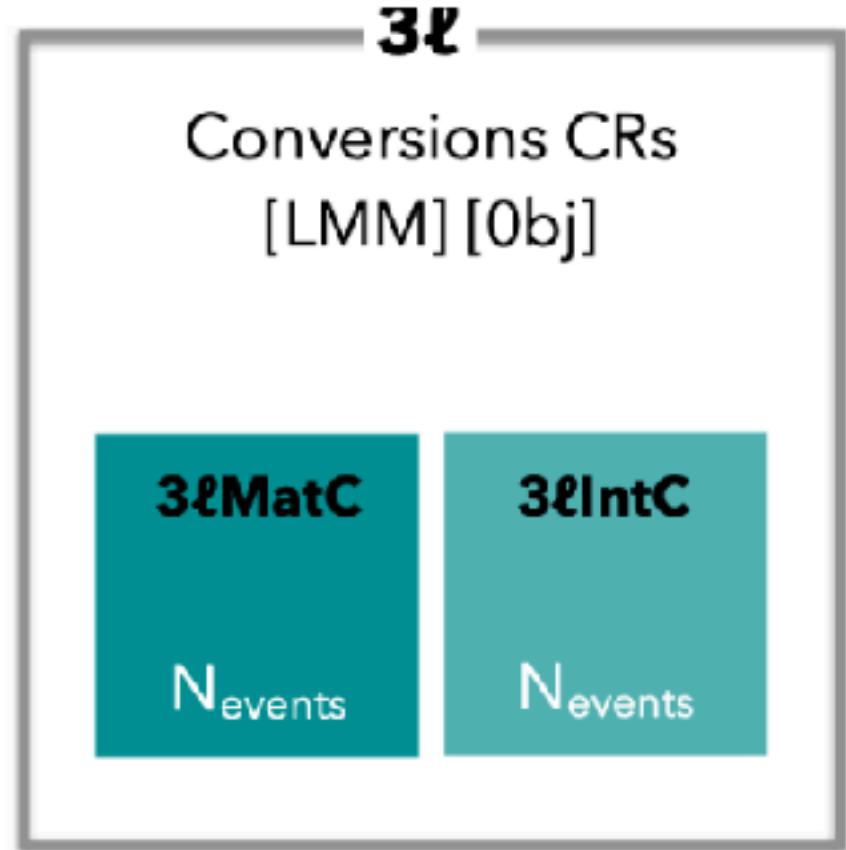
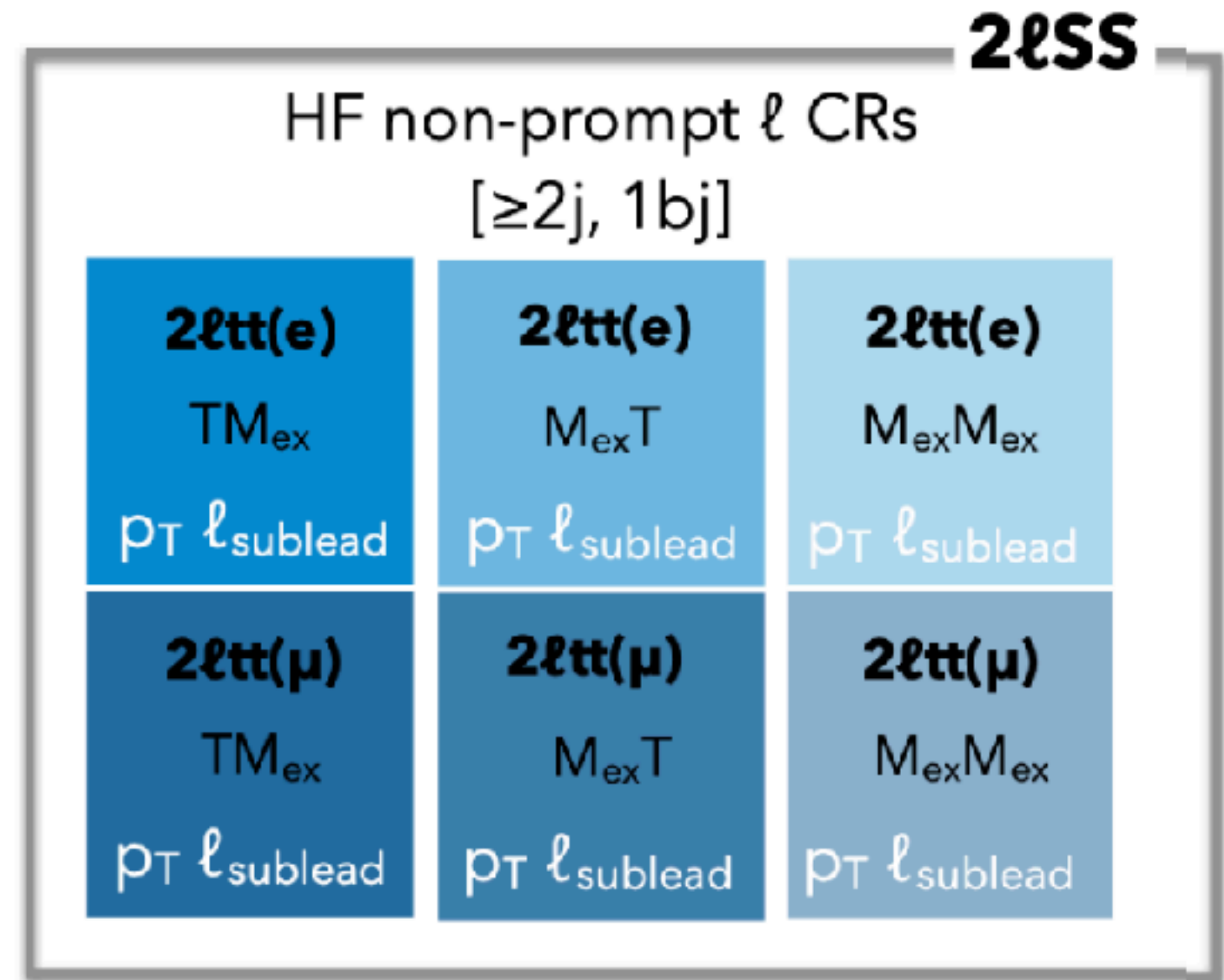


- ▶ Presentation of recent ATLAS ttW measurement:
 - ▶ <https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/CONFNOTES/ATLAS-CONF-2023-019/>
- ▶ Trying to understand issues observed in previous measurements (ttW, ttH, 4-tops).
- ▶ First ever differential measurement of ttW!
 - ▶ Using relatively new profile likelihood unfolding technique
 - ▶ Testing state-of-the-art theoretical predictions



Analysis

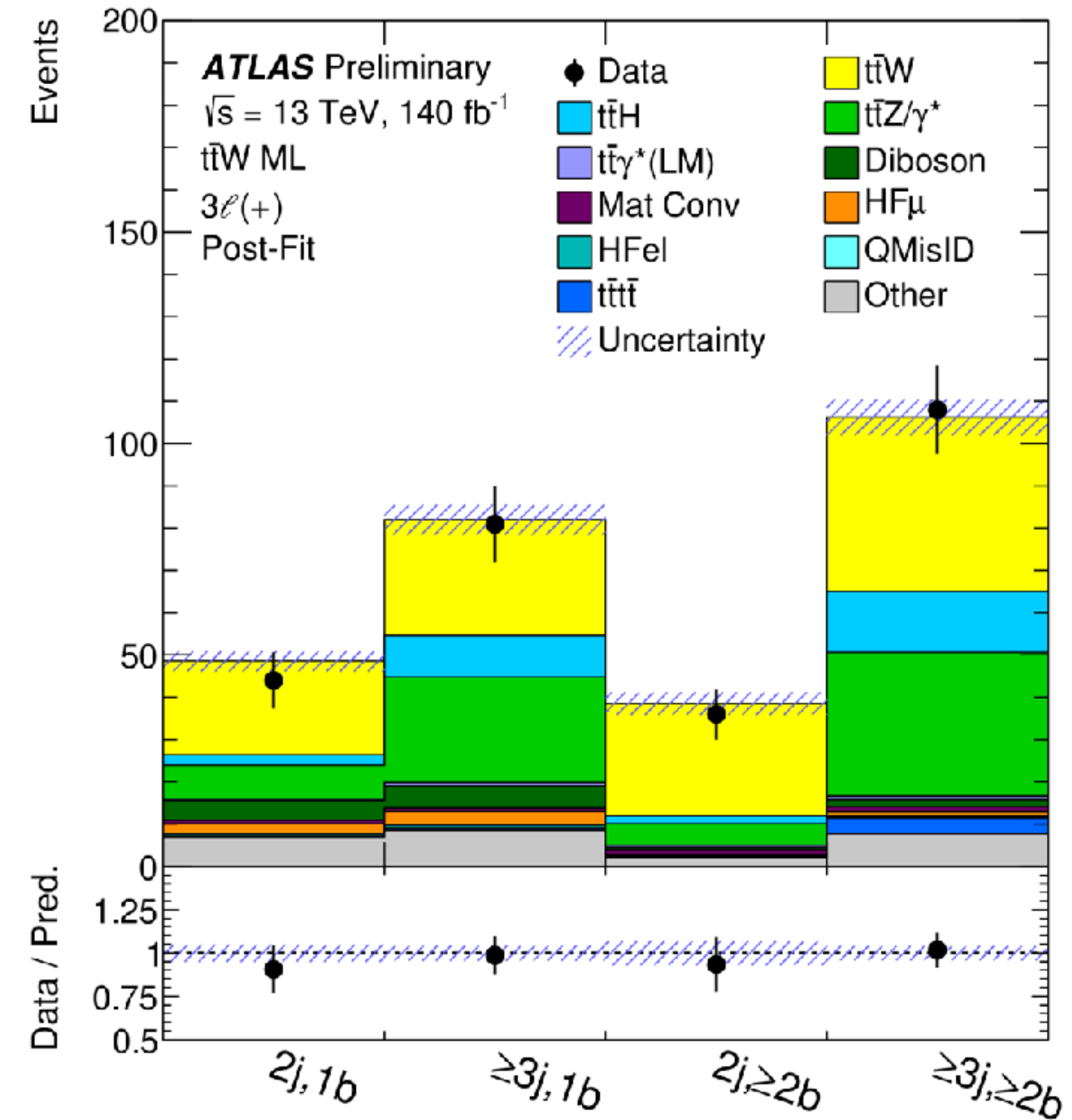
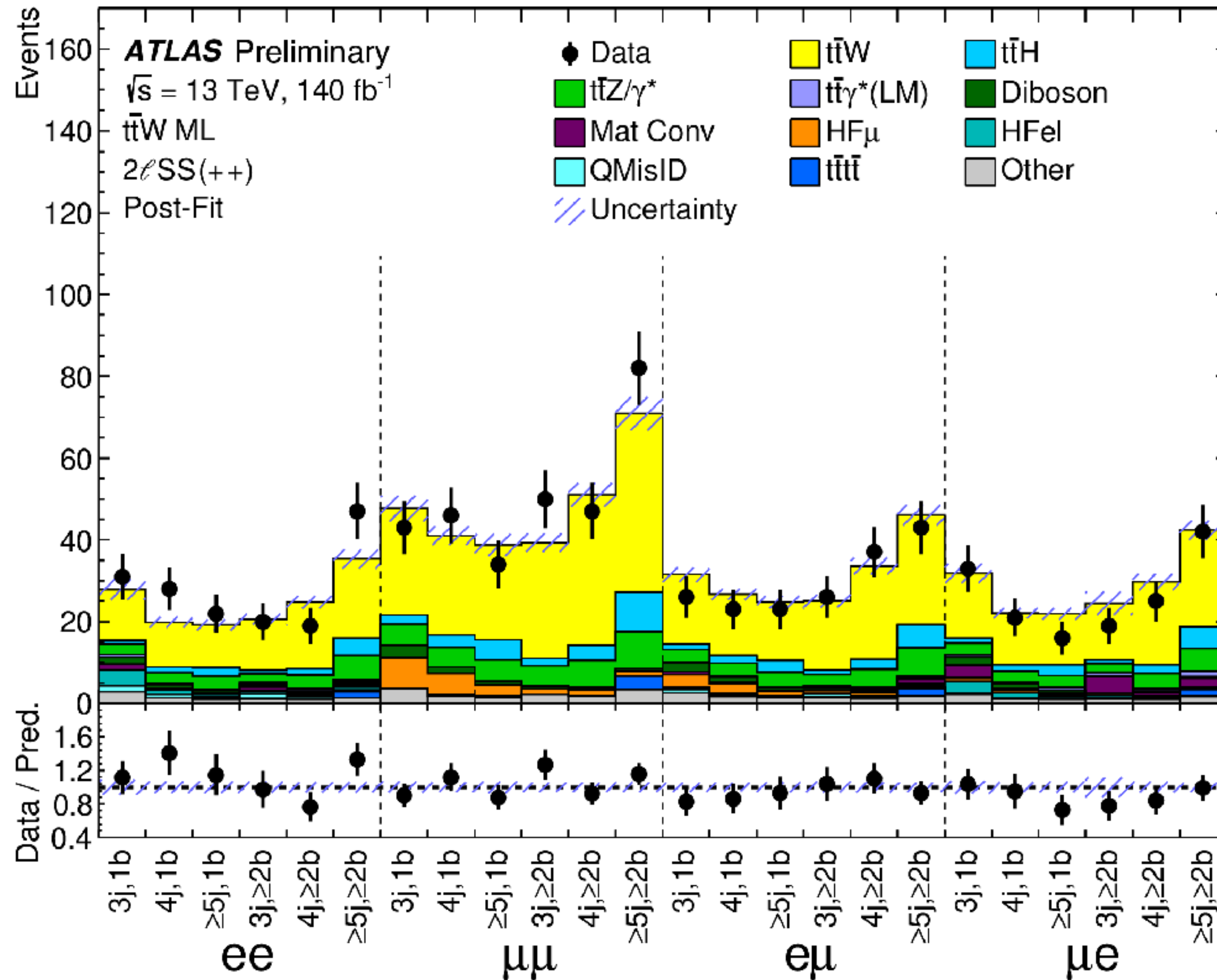
► Background estimate through "Template fit"





Inclusive fit

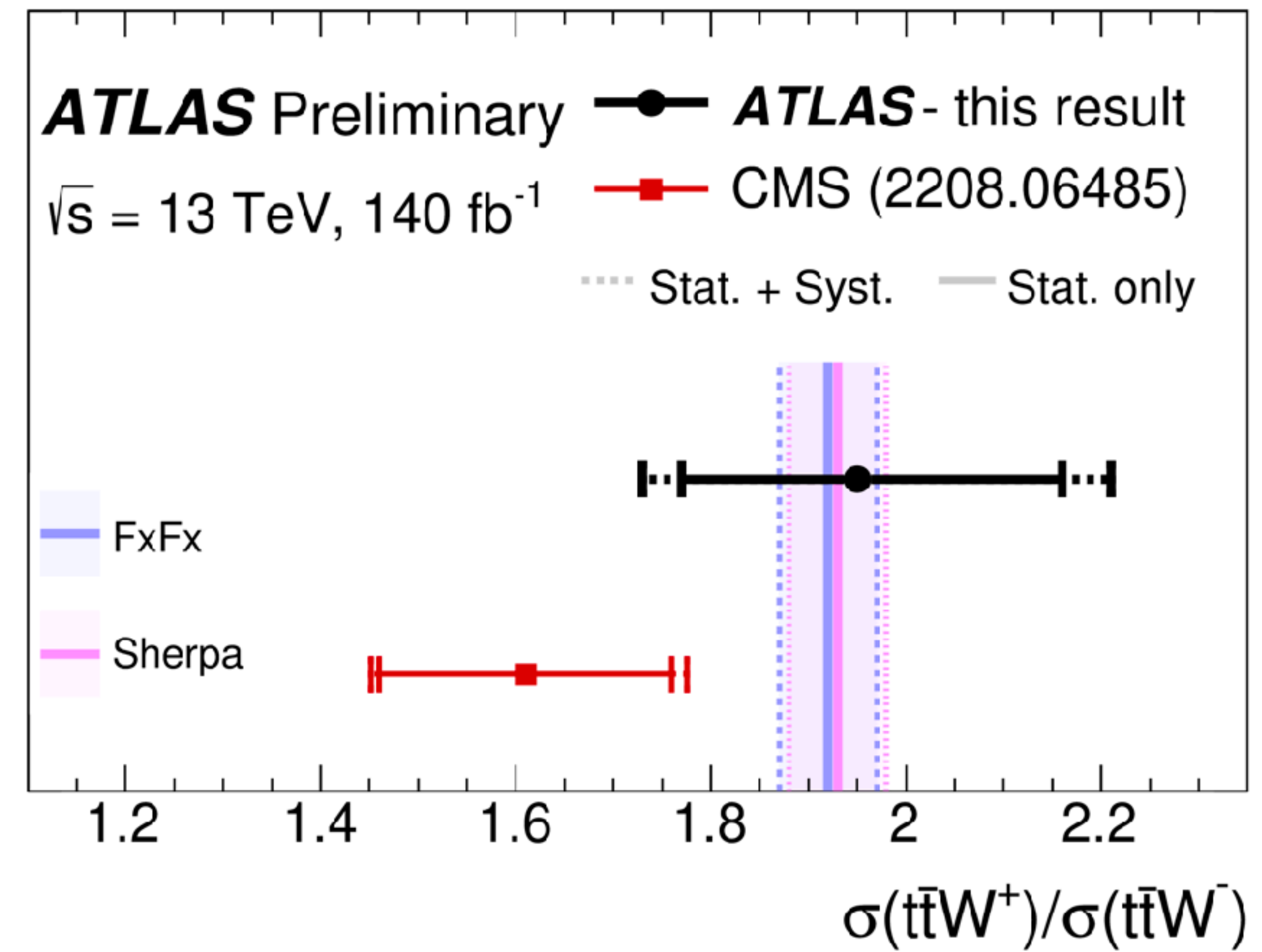
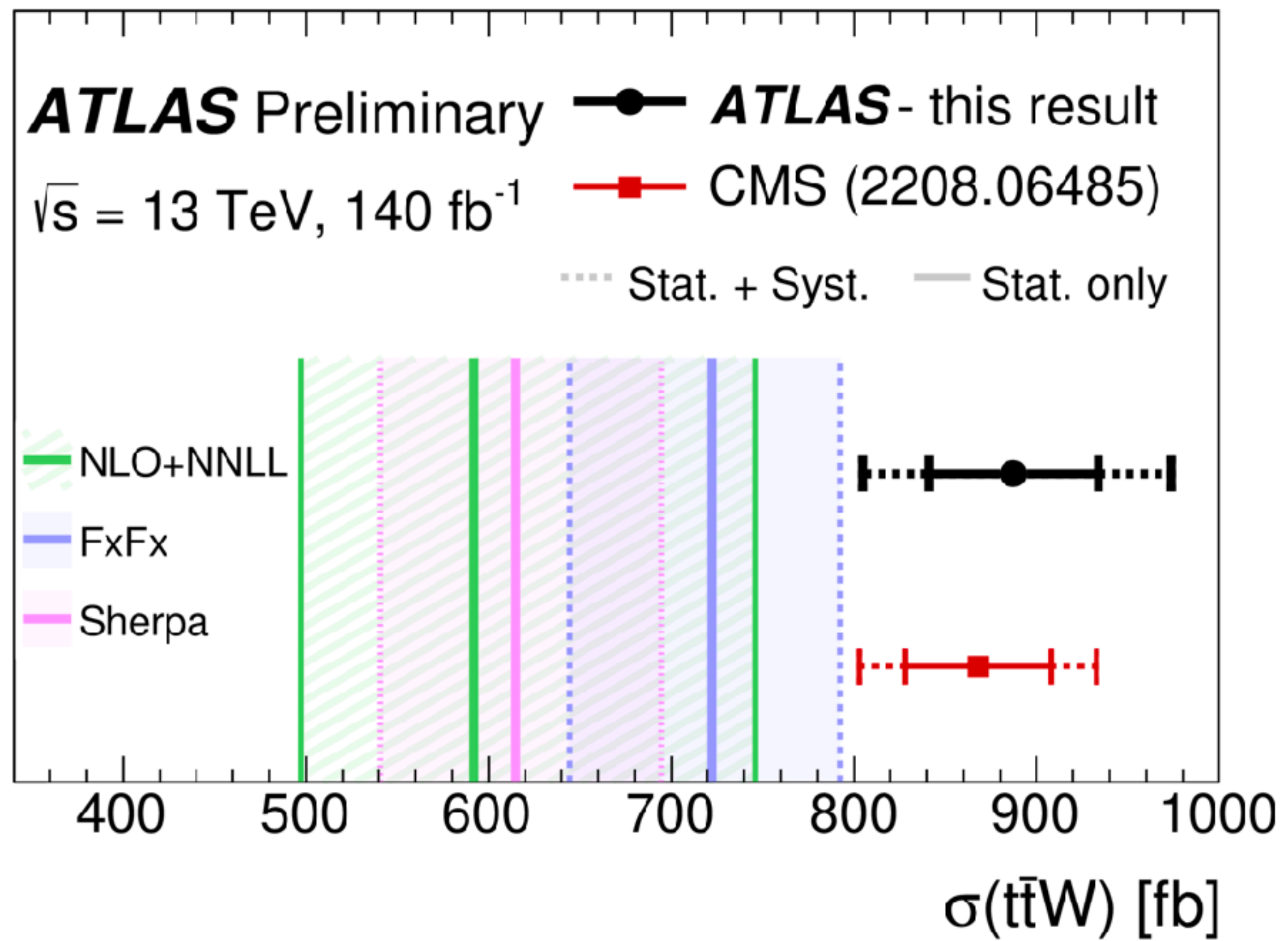
► Split by Njets, Nb and lepton charge and lepton flavour = 48 SS + 8 3L SRs





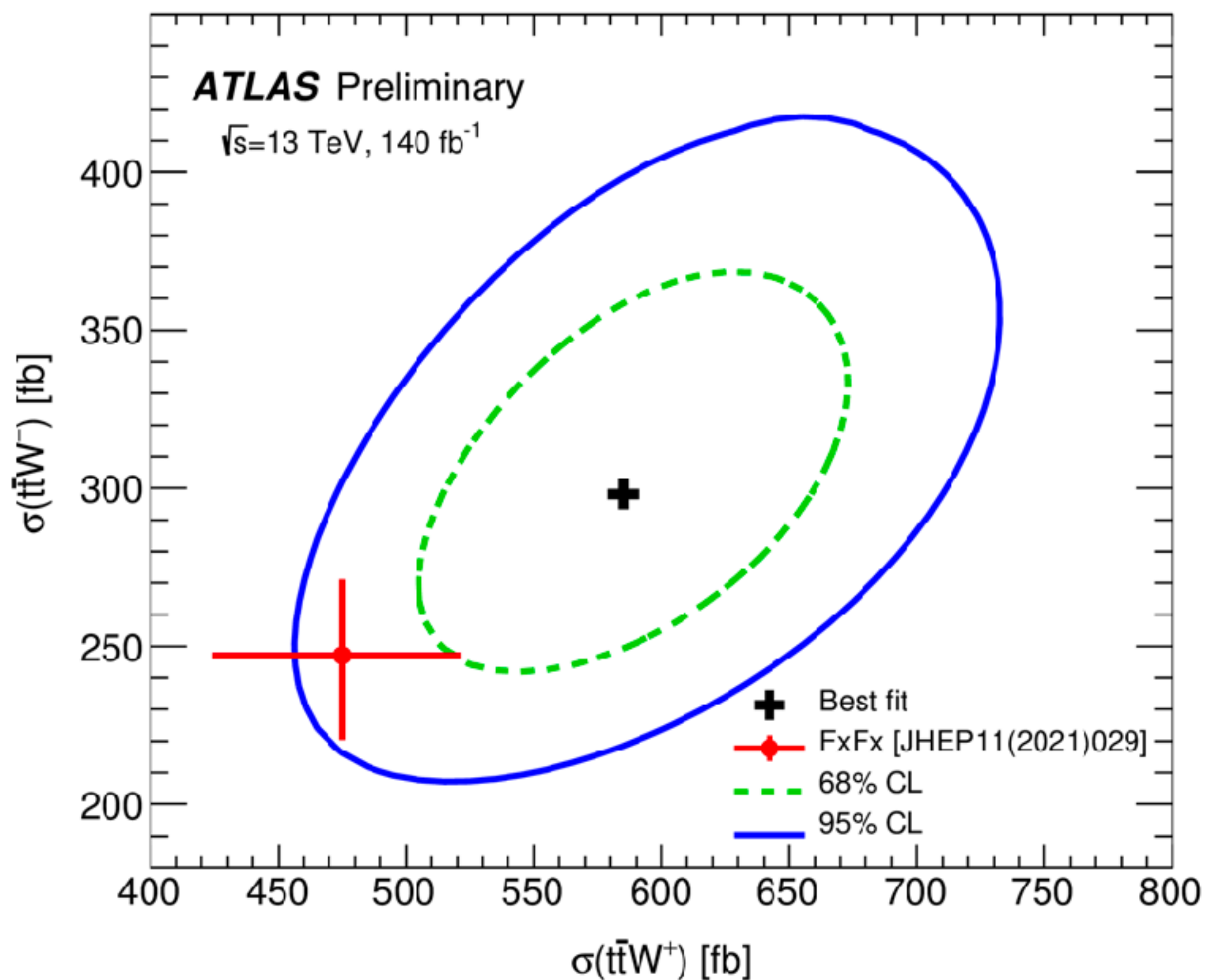
Inclusive results | XS and charge ratio

- ▶ $\sigma(ttW) = 890 \pm 50$ (stat) ± 70 (syst) fb (9% relative uncertainty)
- ▶ Consistent at 1.5σ with FxFx calculation: $\sigma(ttW) = 722^{+70}_{-78}$ (scale) ± 7 (PDF) fb
- ▶ $\sigma(ttW+)/\sigma(ttW-) = 1.95 \pm 0.21$ (stat) ± 0.16 (syst) \rightarrow consistent with SM





Inclusive results | 2D result and uncertainties

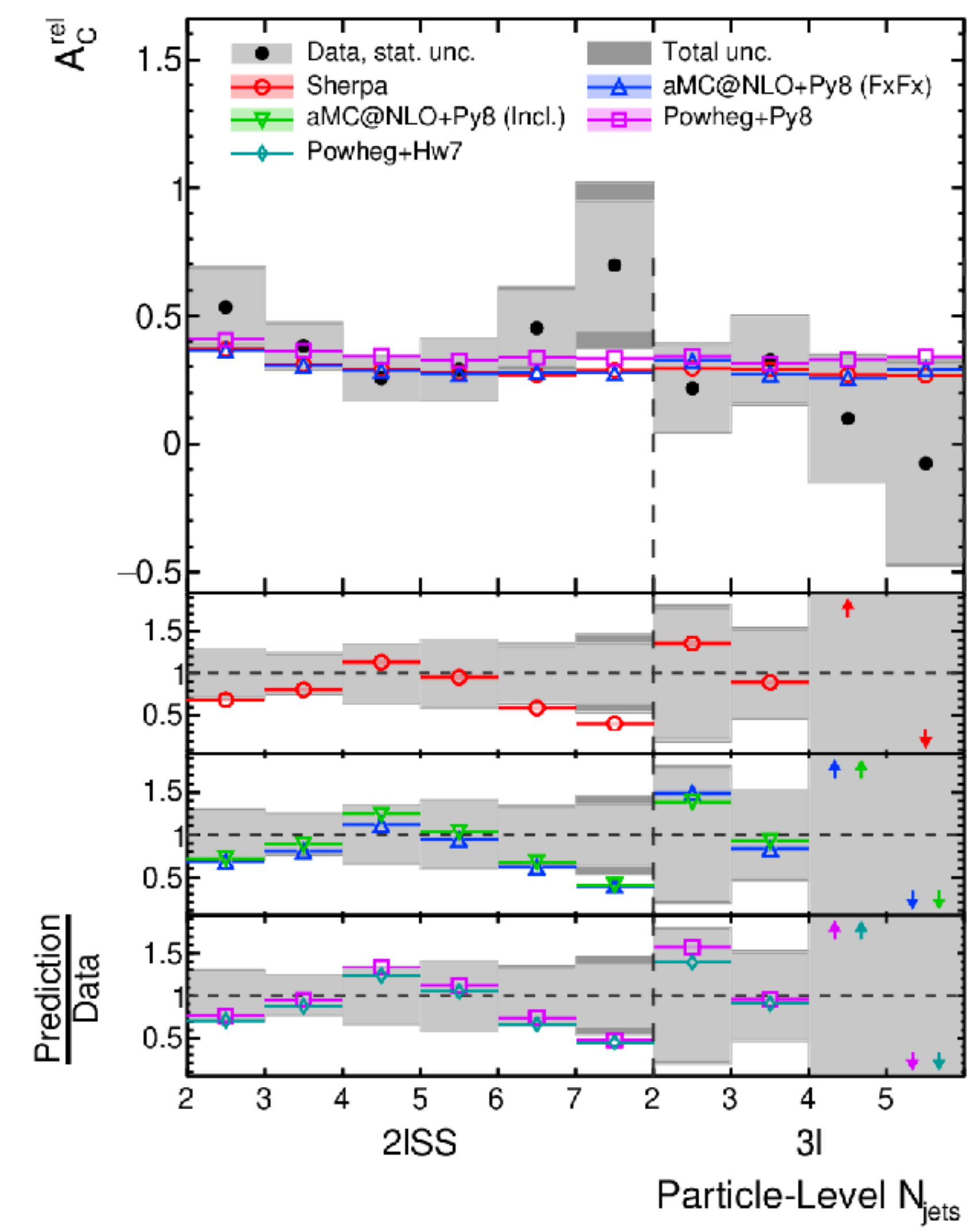
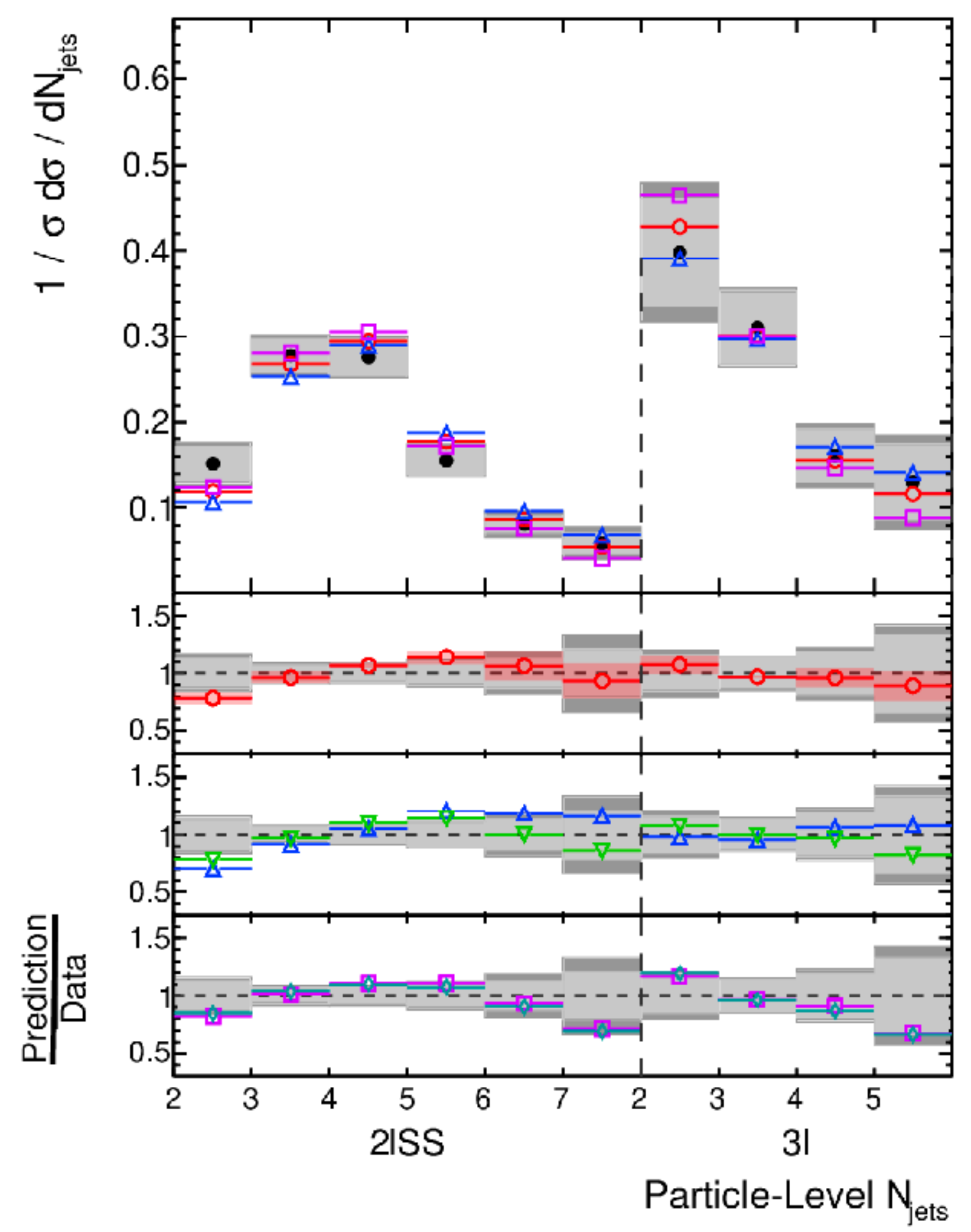
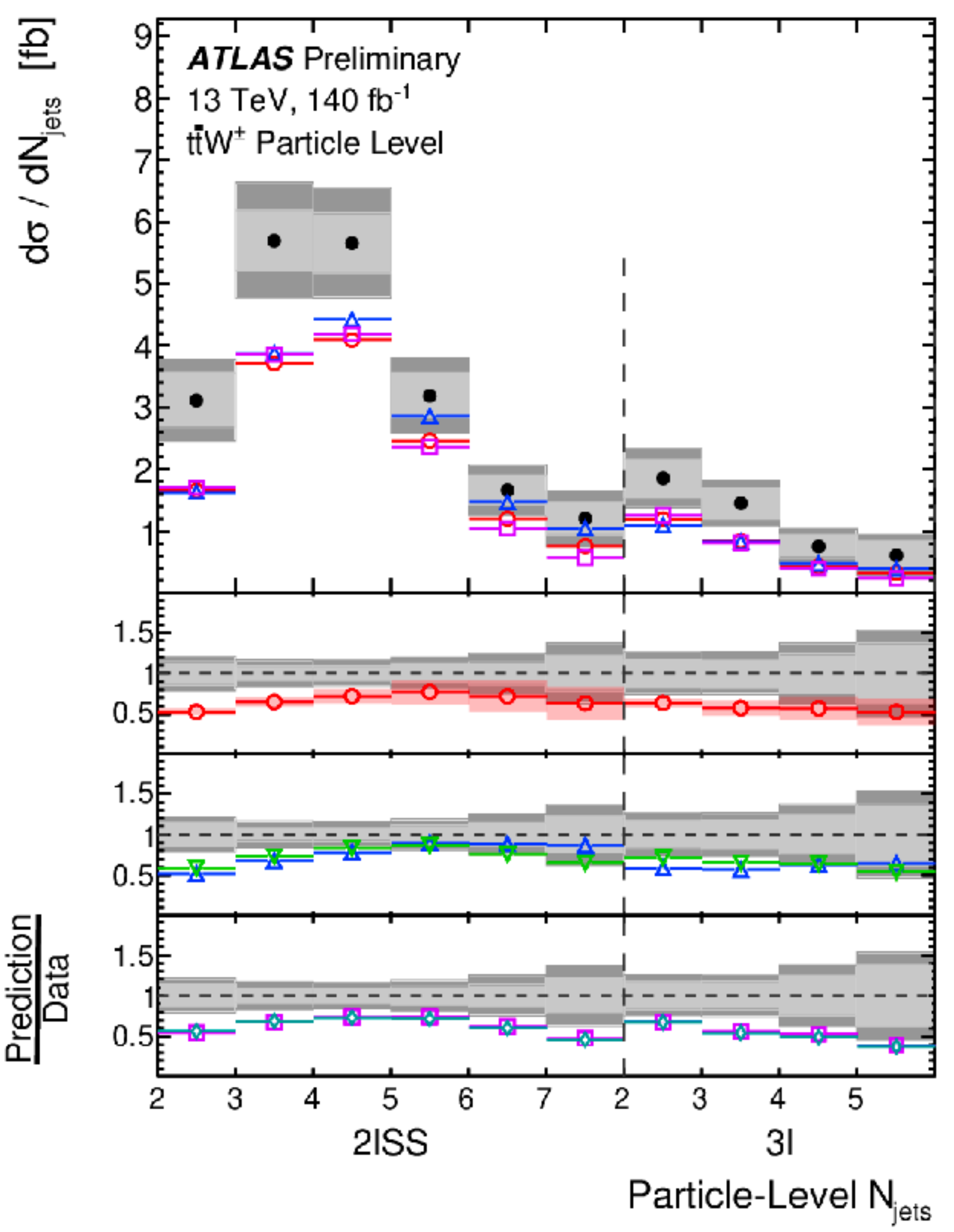


	$\frac{\Delta\sigma(ttW)}{\sigma(ttW)}$ [%]	$\frac{\Delta\sigma_{fid}(ttW)}{\sigma_{fid}}$ [%]	$\frac{\Delta R(ttW)}{R(ttW)}$ [%]	$\frac{\Delta A_C^{rel}}{A_C^{rel}}$ [%]
<i>ttW</i> ME and PS modelling	6.0	7.0	6.0	8.0
Prompt lepton bkg. norm.	2.6	2.5	1.6	2.2
Lepton isolation BDT	2.3	2.3	1.0	1.2
Fakes/ <i>VV</i> / <i>ttZ</i> norm. (free-floated)	2.3	2.7	1.8	2.5
Non-prompt lepton bkg. modelling	1.9	1.7	2.3	3.1
Trigger	1.9	1.8	0.5	0.7
MC statistics	1.5	1.6	1.9	2.5
<i>ttW</i> PDF	1.5	1.4	2.1	2.8
Jet energy scale	1.4	1.9	0.8	1.1
Prompt lepton bkg. modelling	1.3	1.3	1.3	1.9
Luminosity	1.0	1.0	0.08	0.13
Charge Mis-ID	0.7	0.7	0.4	0.5
Jet energy resolution	0.5	0.6	0.7	0.31
Flavour tagging	0.28	0.33	0.5	1.0
<i>ttW</i> Scale	0.21	0.9	1.4	1.9
Electron/photon reco.	0.15	0.2	0.12	0.3
MET	<0.10	<0.10	0.17	0.4
Muon	<0.10	<0.10	<0.10	0.4
Pile-up	<0.10	0.25	<0.10	0.3
Total syst.	8	10	8	10
Data statistics	5	5	10	16
Total	9	11	13	19



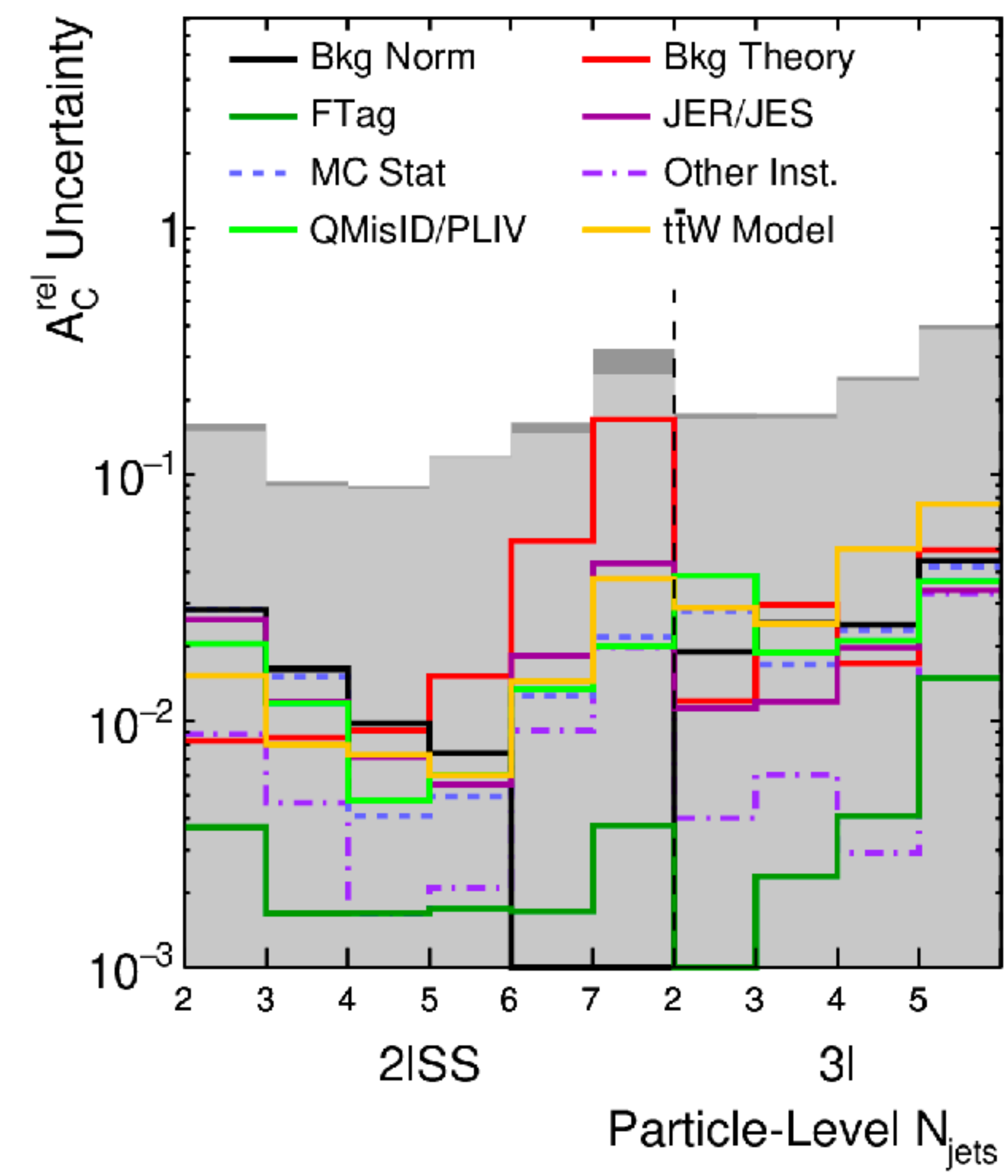
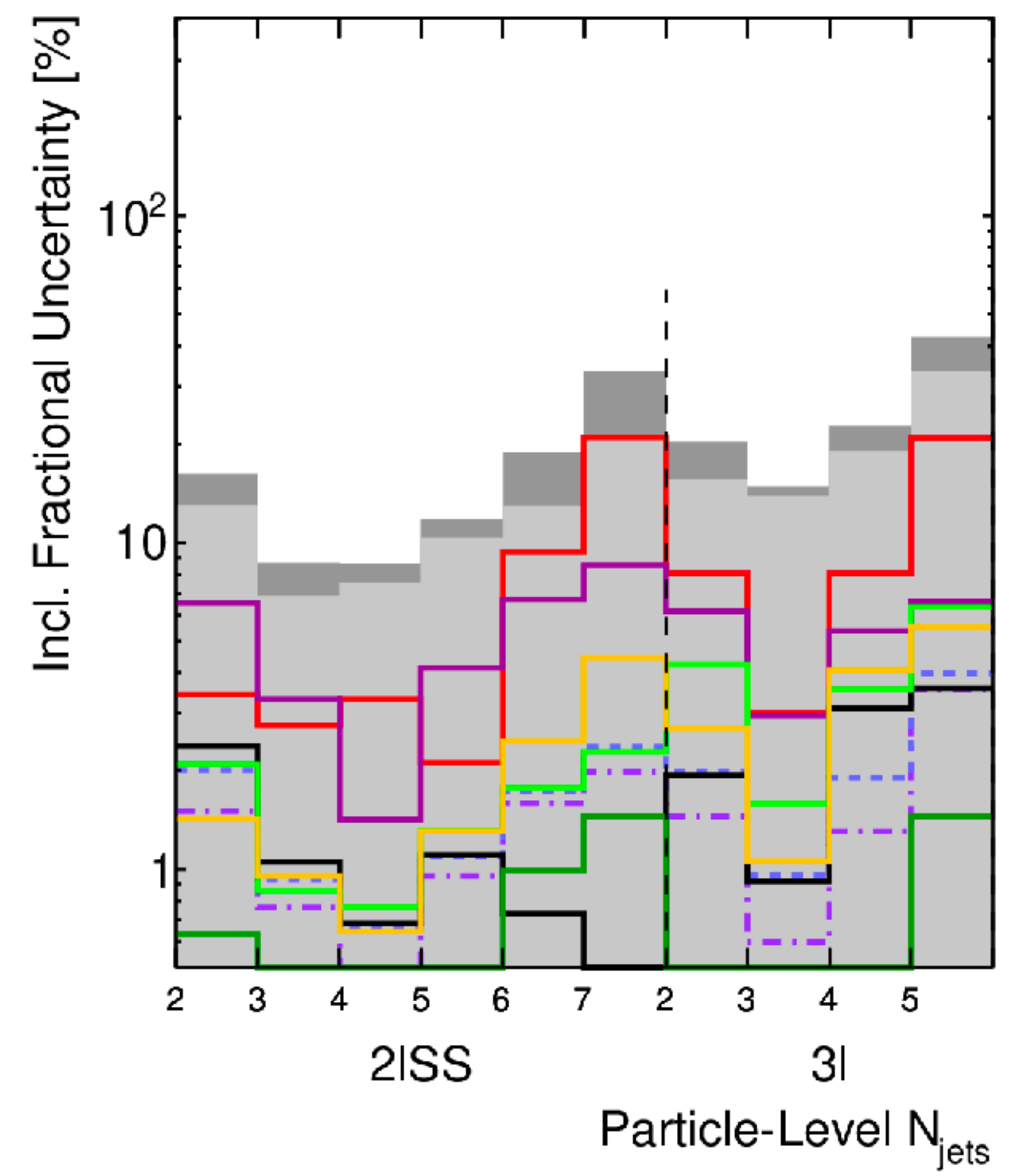
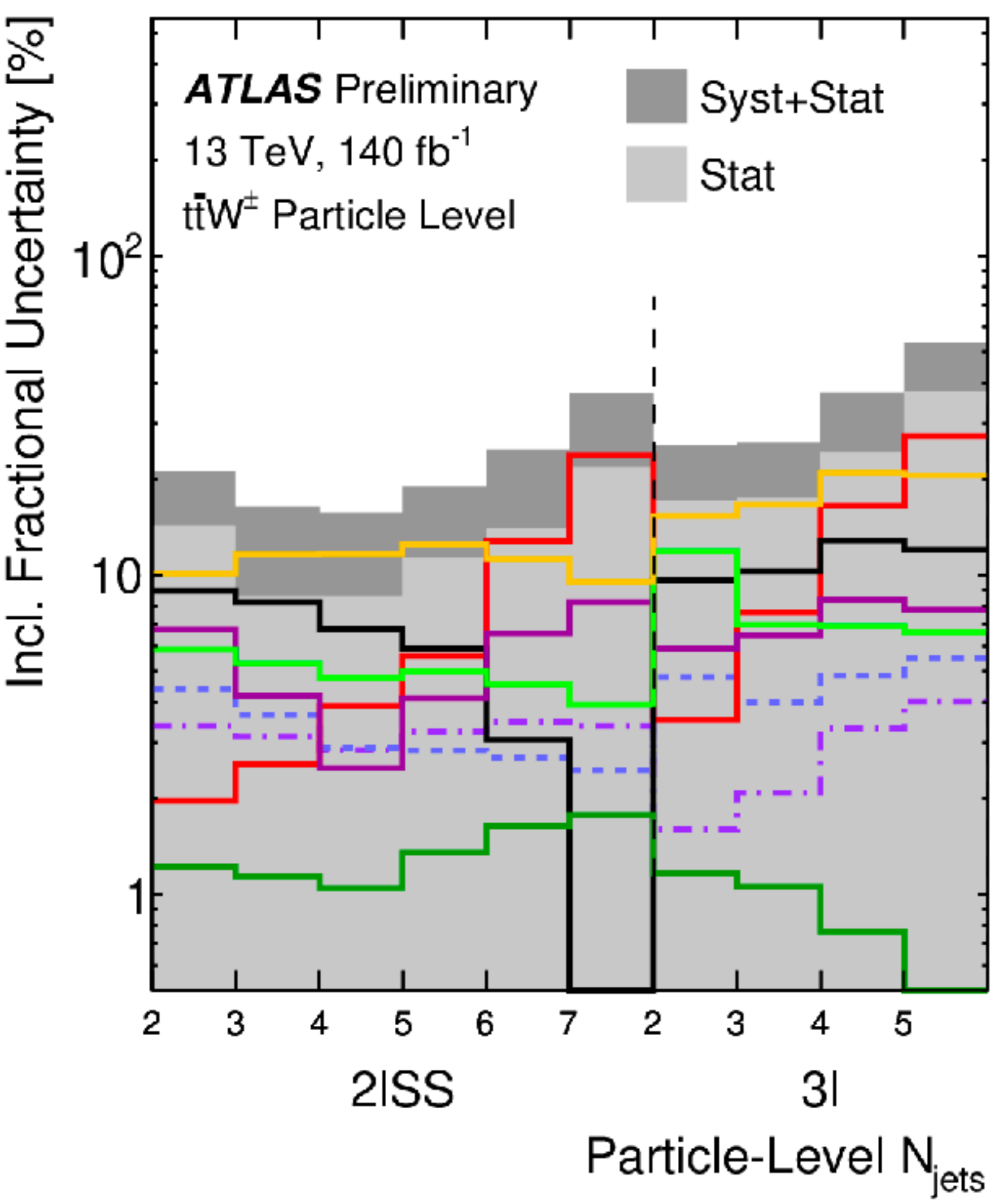
Differential results | njets

► Unfold 9 observables. Showing here njets:



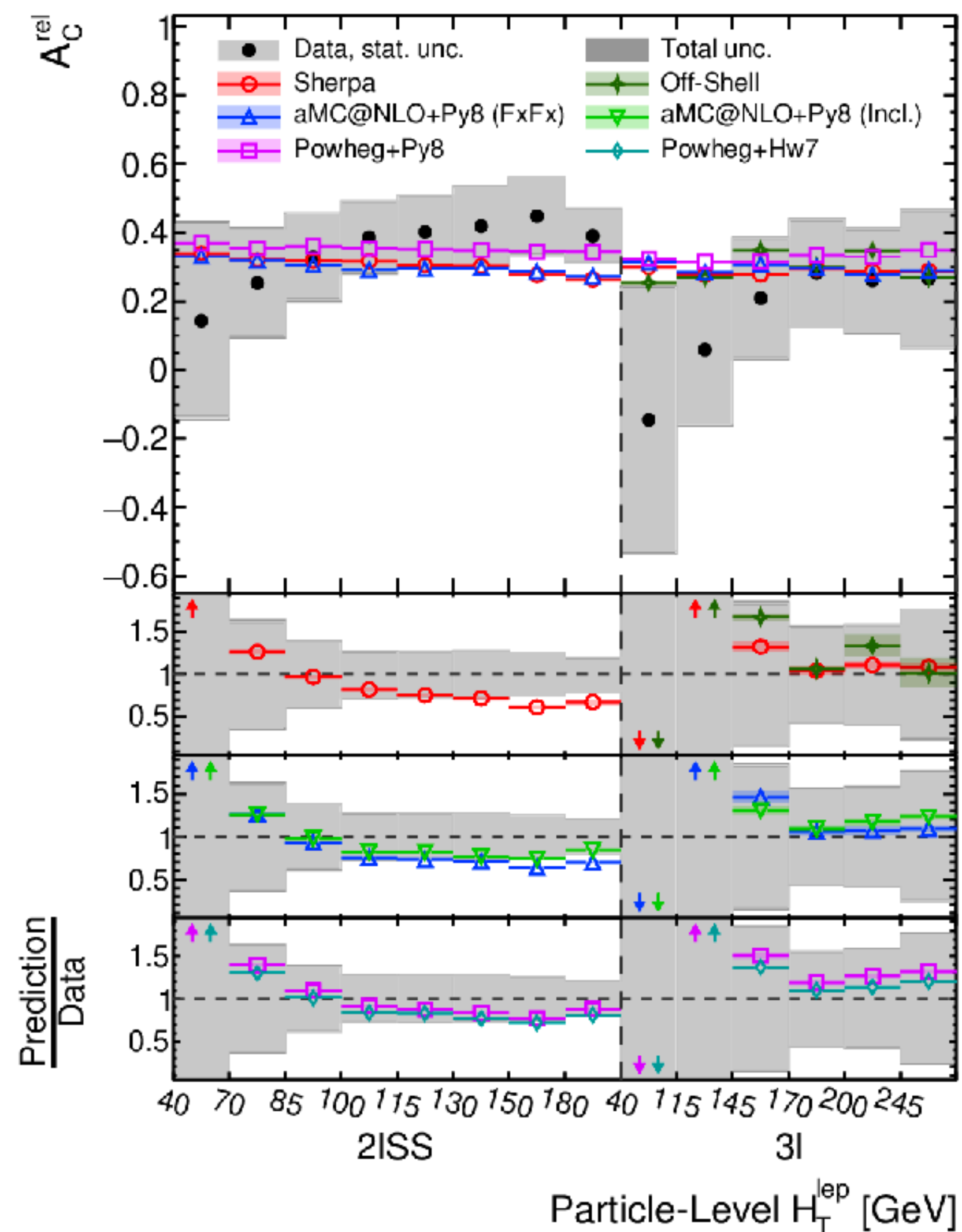
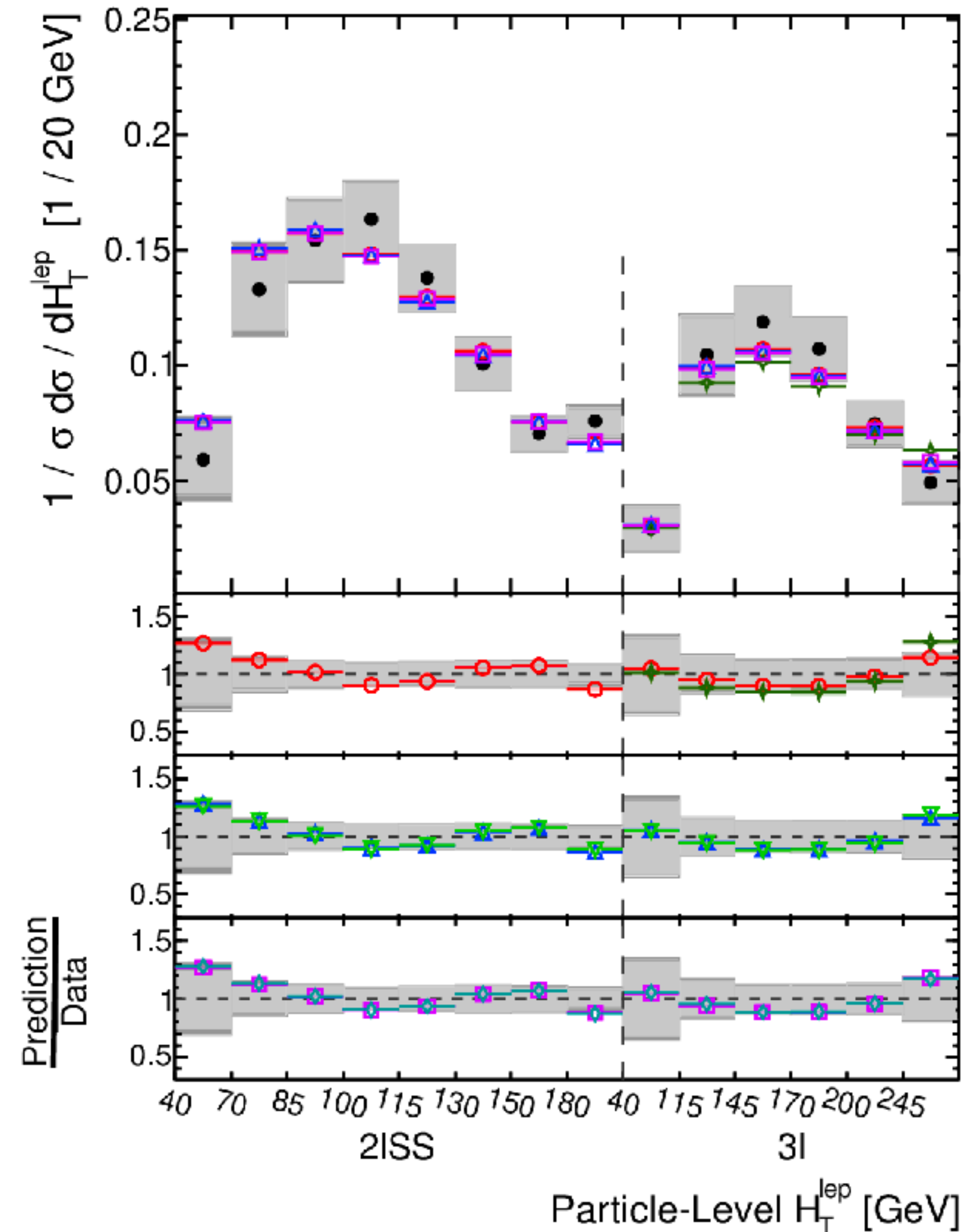
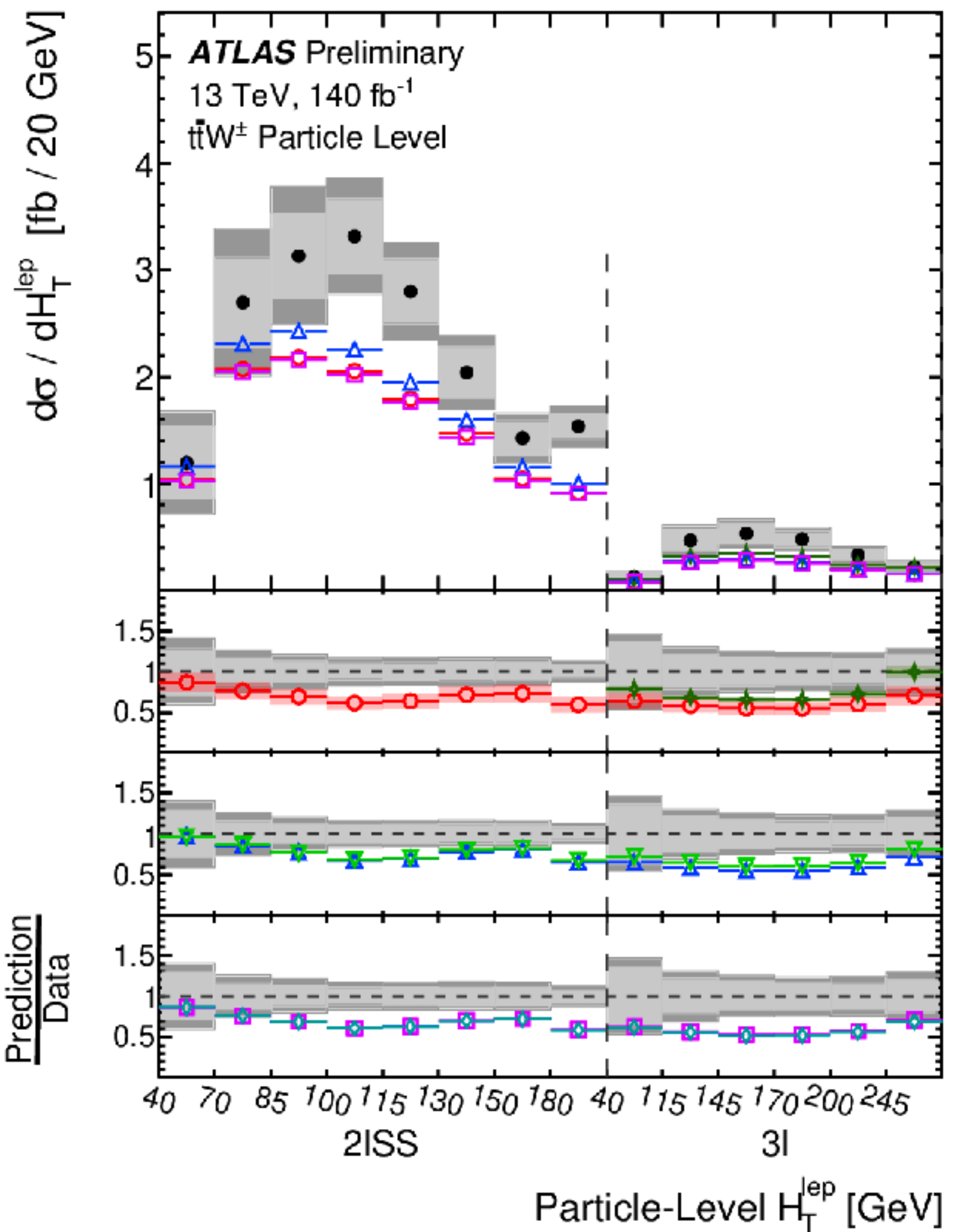


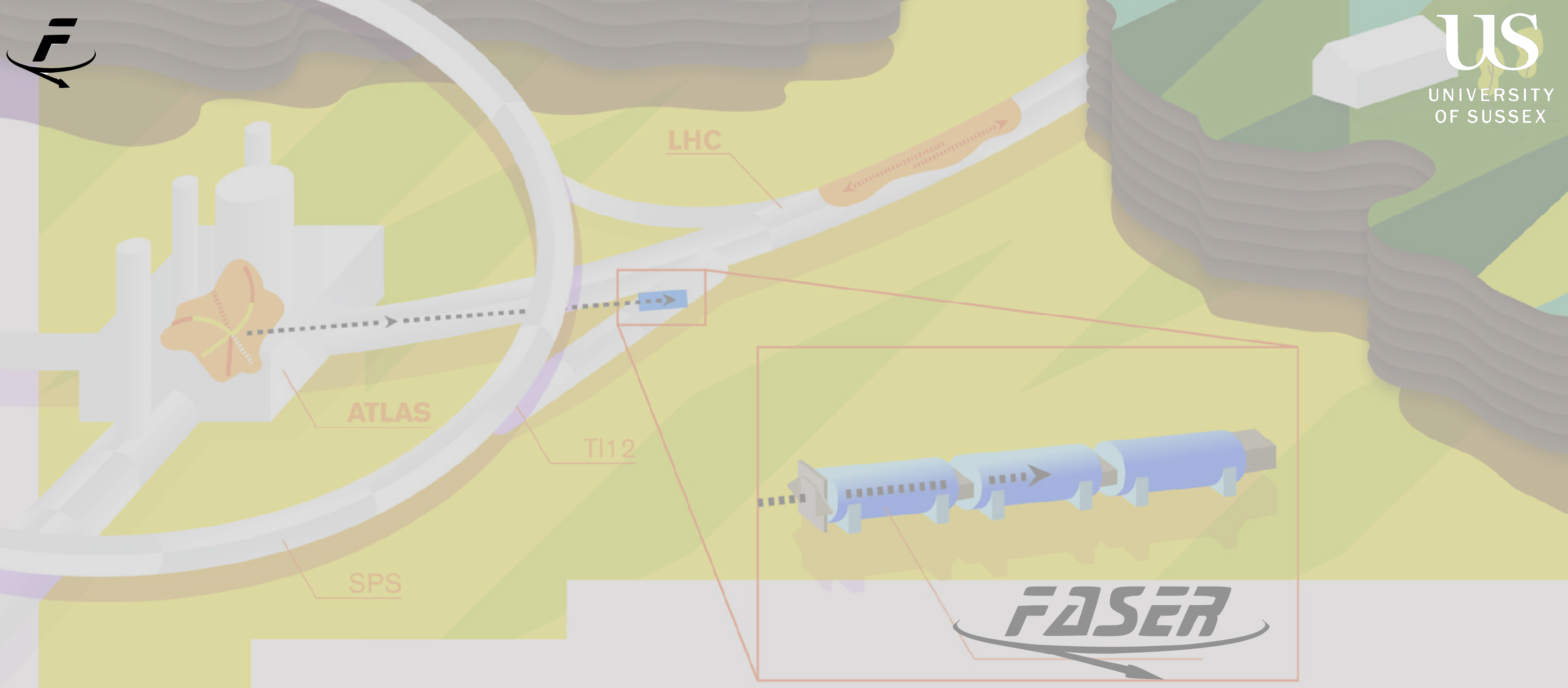
Differential results | Uncertainties



Differential results

► Off-shell theory predictions provided for 3L channel

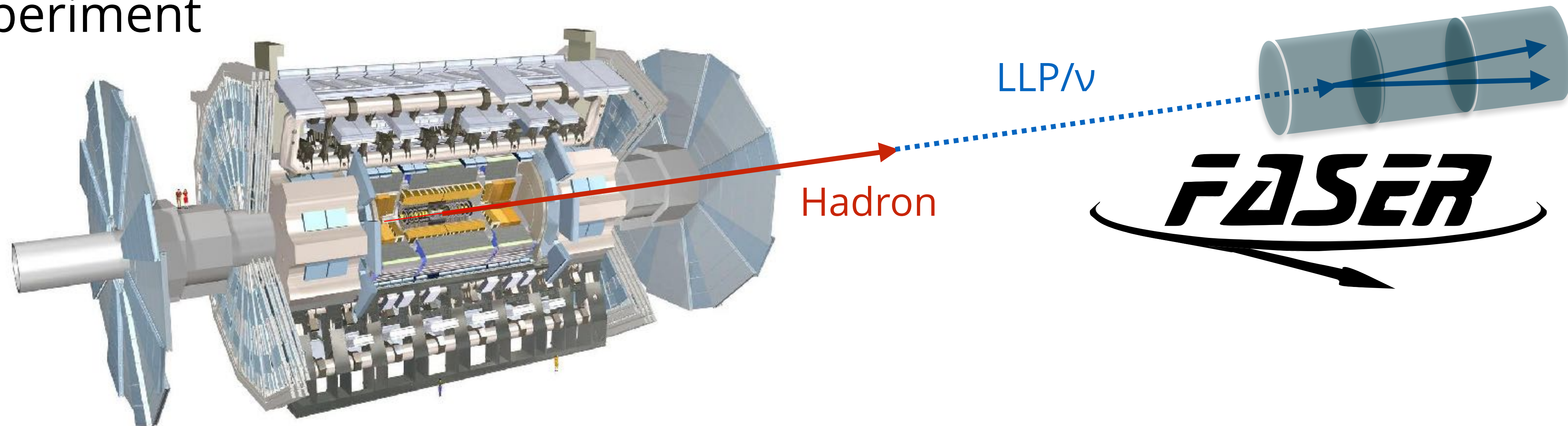




First FASER Results

Overview

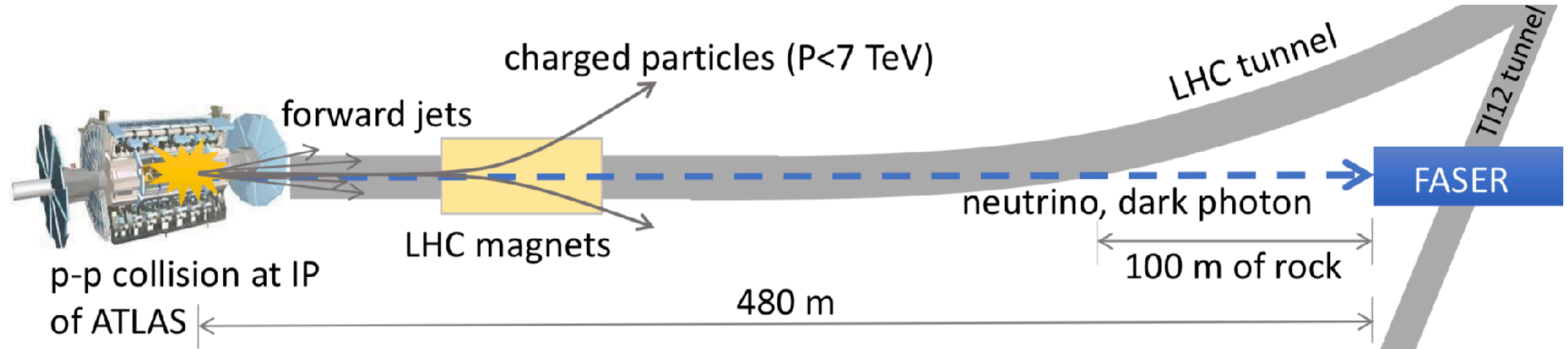
- ▶ A new experiment at CERN!



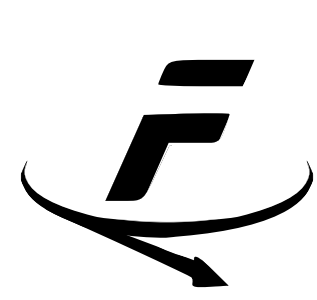
- ▶ Detector is 480m from ATLAS IP1
 - ▶ In line with beam collision axis. Transverse size of 10cm \rightarrow mrad regime ($\eta > 9.1$)
- ▶ Inelastic pp cross section is huge $\rightarrow 10^{16}$ collisions in Run 3 $\rightarrow 10^{17} \pi, 10^{13} B$
 - ▶ From only 10^{-8} of solid angle 1% of π_0 s are in acceptance.
- ▶ **Neutrinos produced copiously in decays of forward hadrons**
- ▶ **Very weakly interacting LLPs could be produced in significant numbers**

Location

- ▶ The T112 service tunnel happens to be in just the right place for FASER:

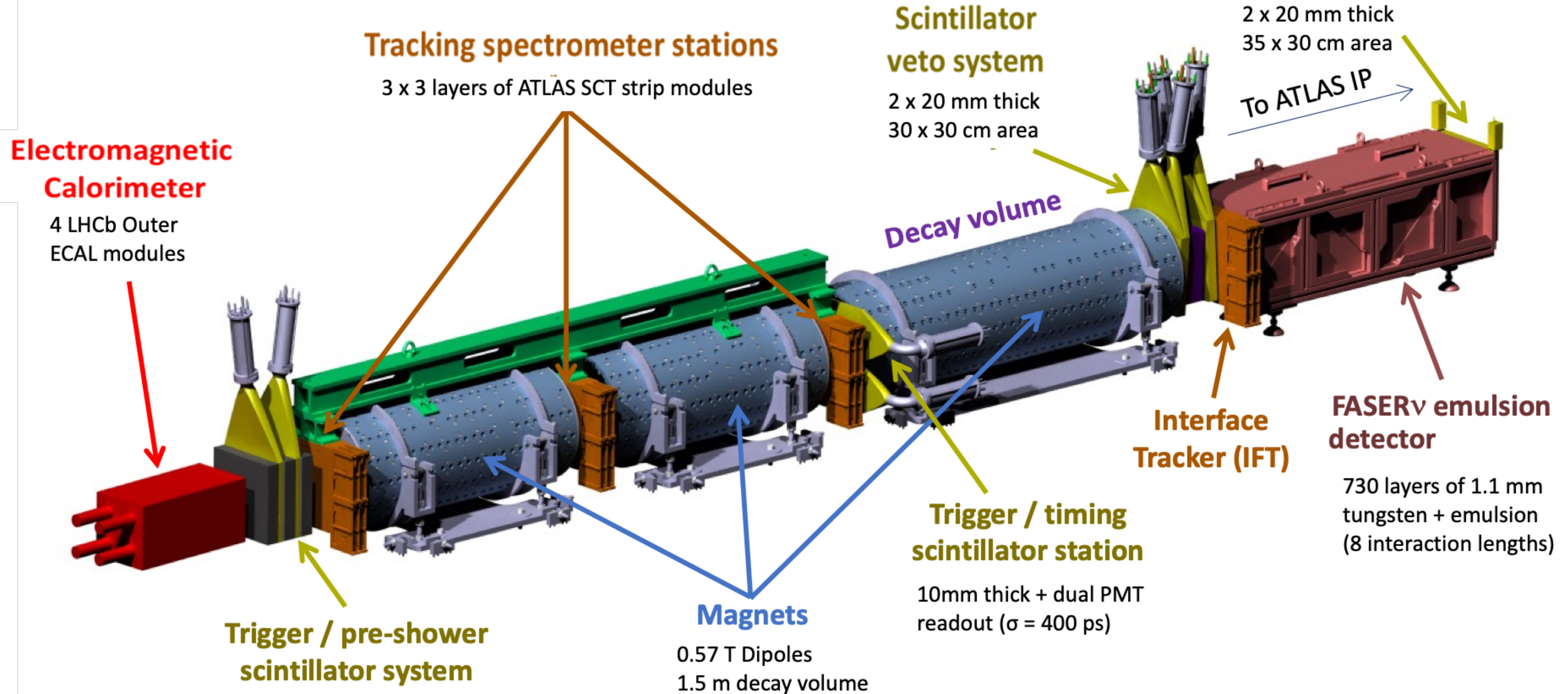


- ▶ Old SPS → LEP tunnel
 - ▶ On line-of-sight (with some digging)
 - ▶ Shielded by ~100m rock/concrete
 - ▶ Low beam backgrounds
 - ▶ Charged particles bent by LHC magnets



Detector design

▶ Small inexpensive design [2207.11427]



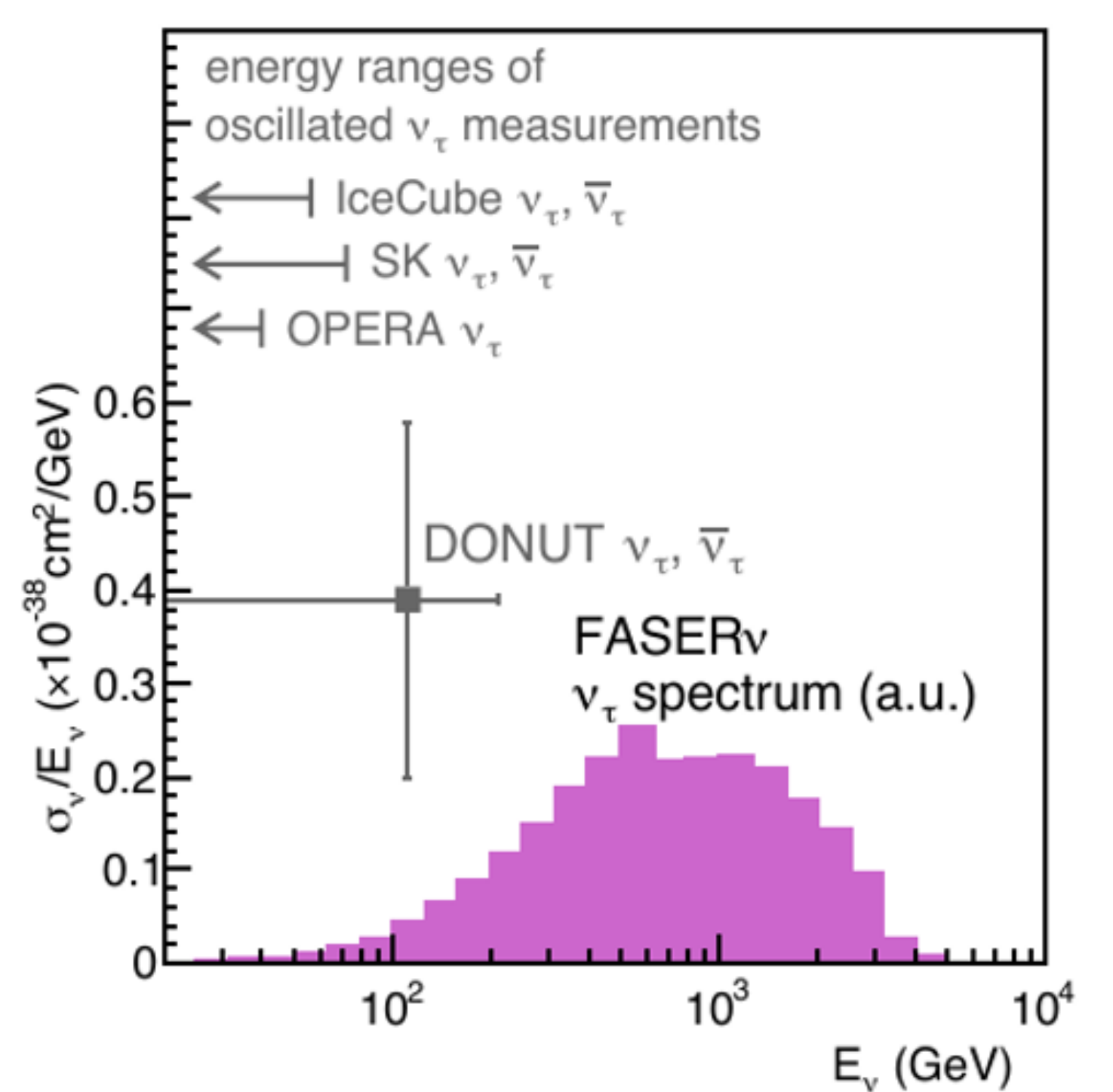
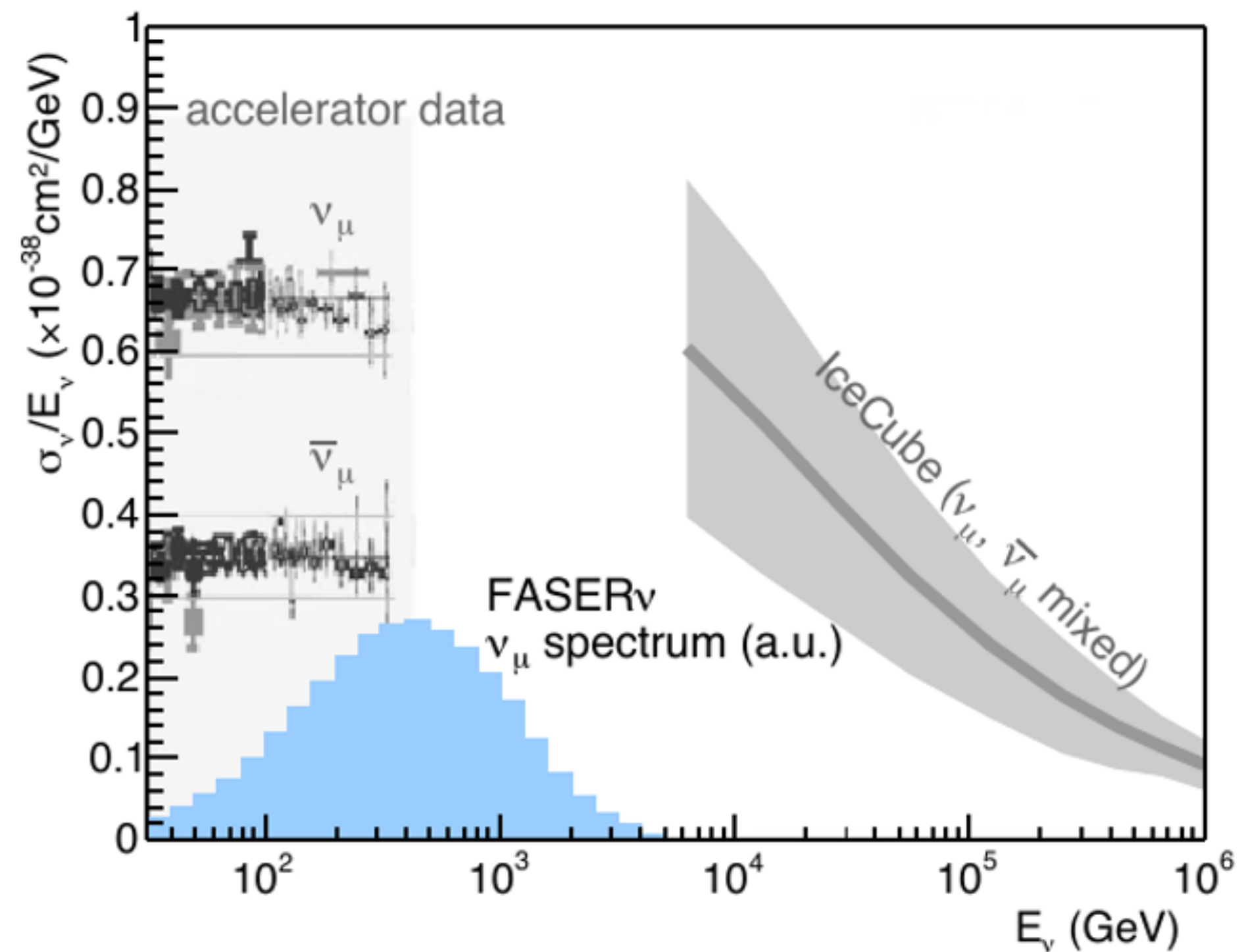
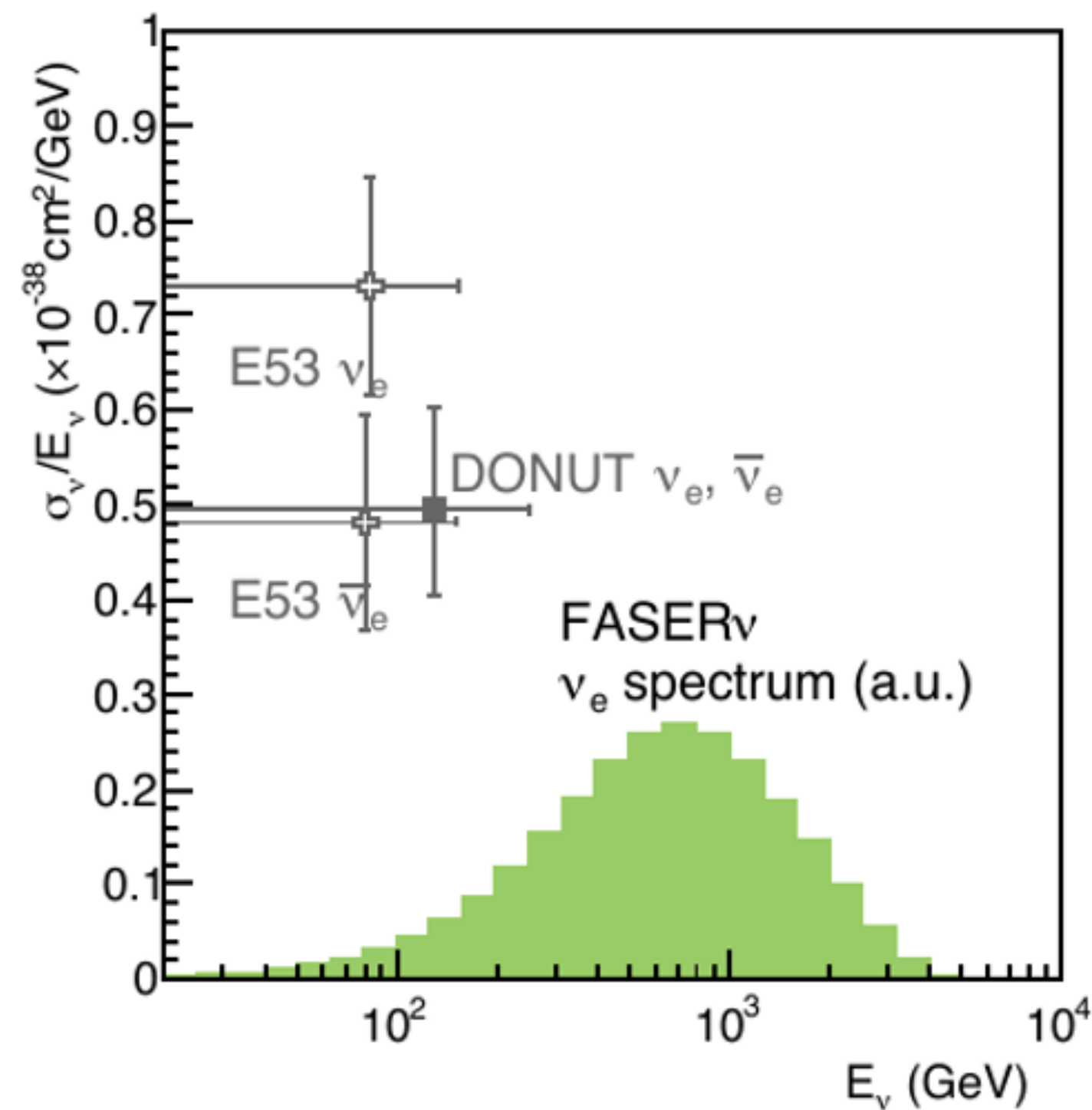


Installation



Neutrino analysis

- ▶ Neutrinos produced copiously in decays of forward hadrons
 - ▶ Highly energetic (TeV scale) → high interaction cross section
- ▶ Extends FASER physics program into SM measurements
 - ▶ Targets measurement of highest energy man-made neutrinos
 - ▶ Energy range complementary to existing neutrino experiments



Neutrino analysis

- ▶ Neutrinos produced copiously in decays of forward hadrons
 - ▶ Highly energetic (TeV scale) → high interaction cross section
- ▶ Extends FASER physics program into SM measurements
 - ▶ Targets measurement of highest energy man-made neutrinos
 - ▶ Energy range complementary to existing neutrino experiments

For 35 fb ⁻¹	ν_e	ν_μ	ν_τ
Main source	Kaons	Pions	Charm
# traversing FASERv	$\sim 10^{10}$	$\sim 10^{11}$	$\sim 10^8$
# interacting in FASERv	≈ 200	≈ 1200	≈ 4

[[PRD 104, 113008](#)]

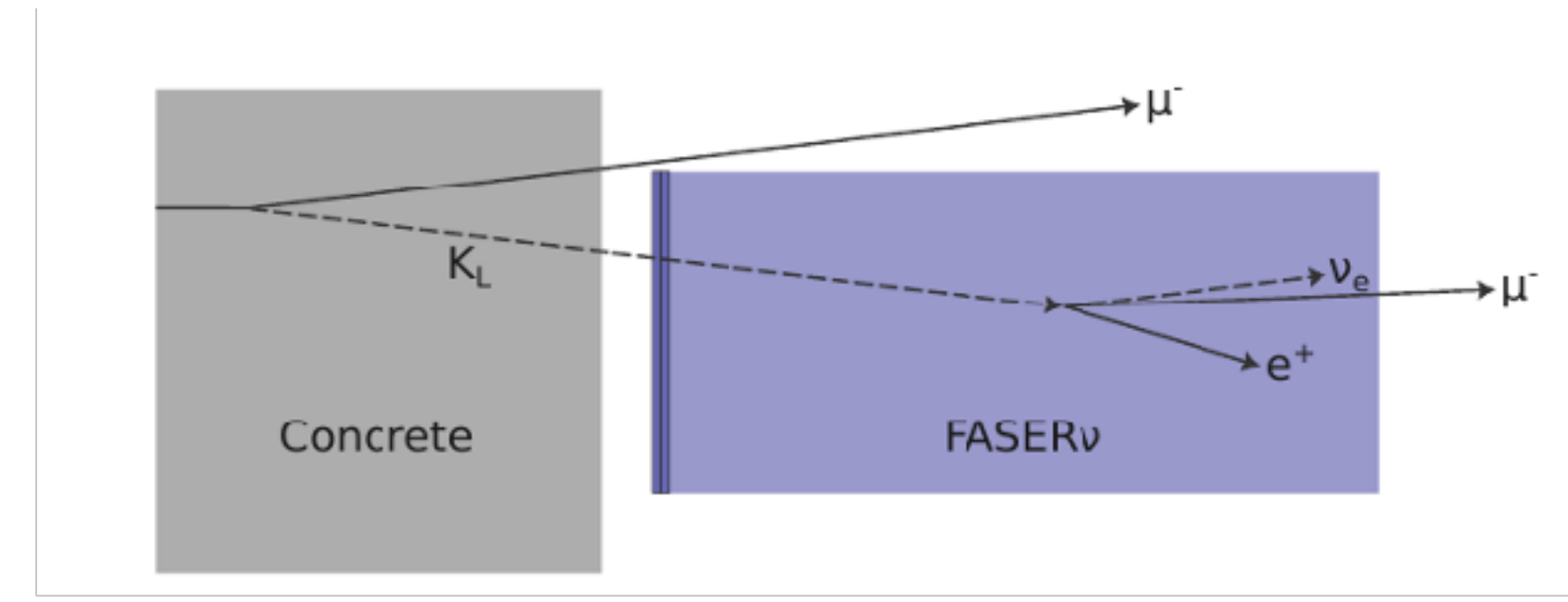
- ▶ Study at colliders originally proposed by Rújula and Rückl in 1984!

Neutrino analysis | Backgrounds

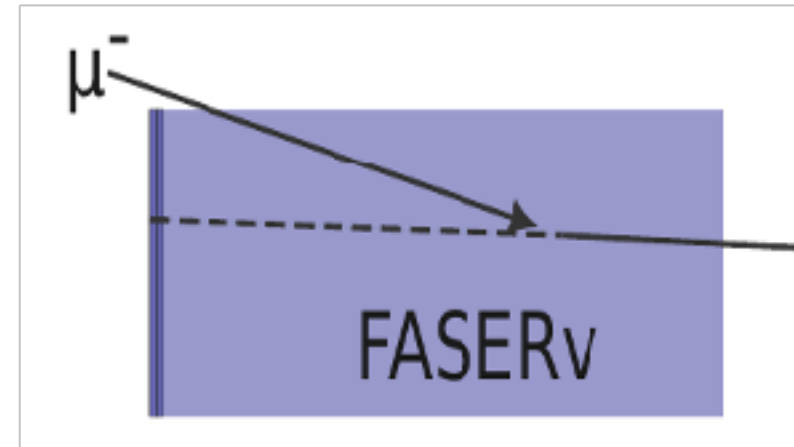
[2303.14185]



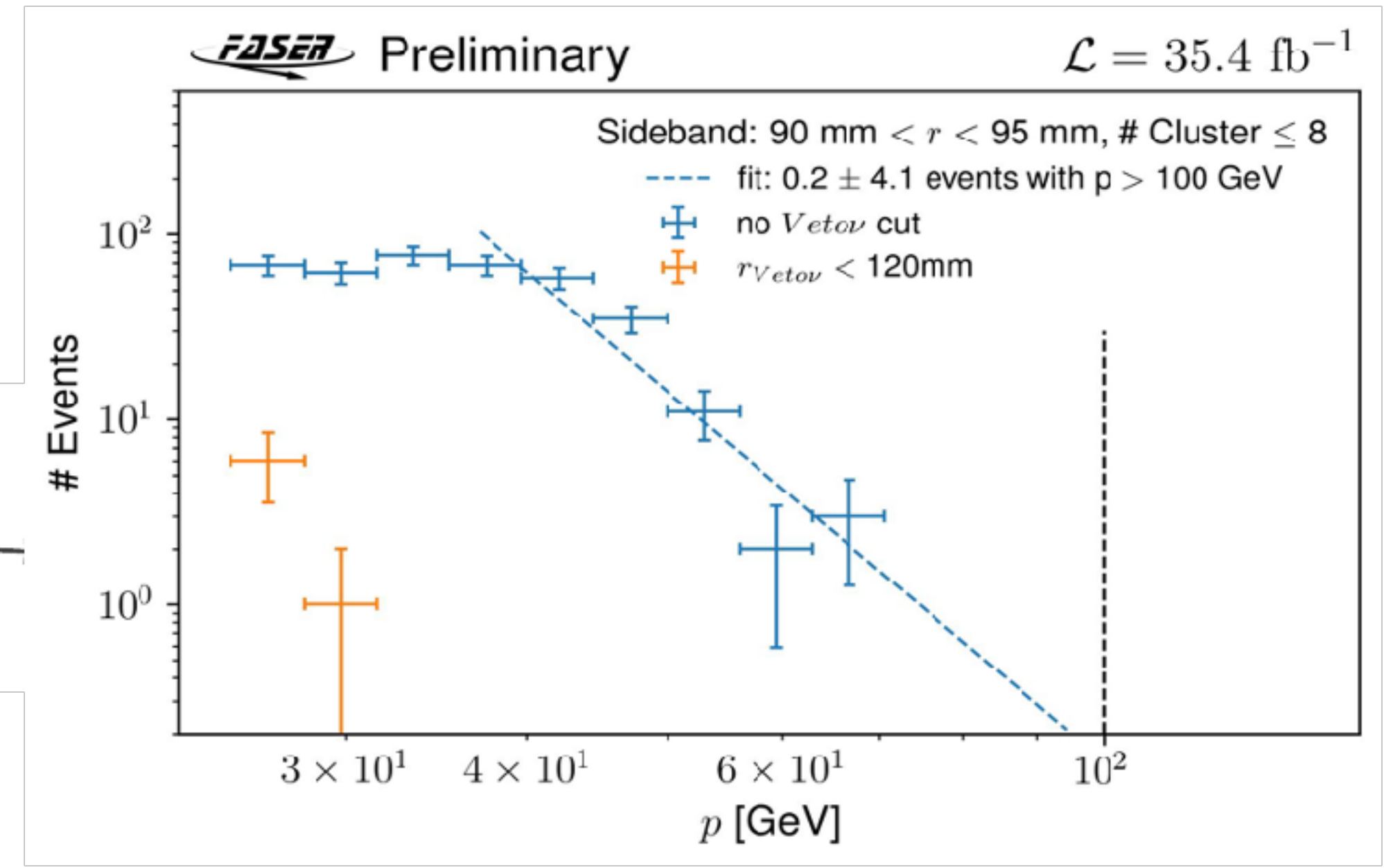
- ▶ Neutral hadrons estimated from 2-step simulation
 - ▶ Expect ~300 neutral hadrons with $E > 100$ GeV reaching FASERv
 - ▶ Most accompanied by μ but conservatively assume missed
 - ▶ Estimate fraction of these passing event selection
 - ▶ Most are absorbed in tungsten with no high-momentum track
 - ▶ Predict $N = 0.11 \pm 0.06$ events



- ▶ Scattered muons estimated from data SB
 - ▶ Take events w/o front veto radius requirement and single track segment in first tracker station with $90 < r < 95$ mm



- ▶ Fit to extrapolate to higher momentum
- ▶ Scale by # events with front veto cut
 - ▶ Use MC to extrapolate to signal region
- ▶ Predict $N = 0.08 \pm 1.83$ events
- ▶ Uncertainty from varying selection



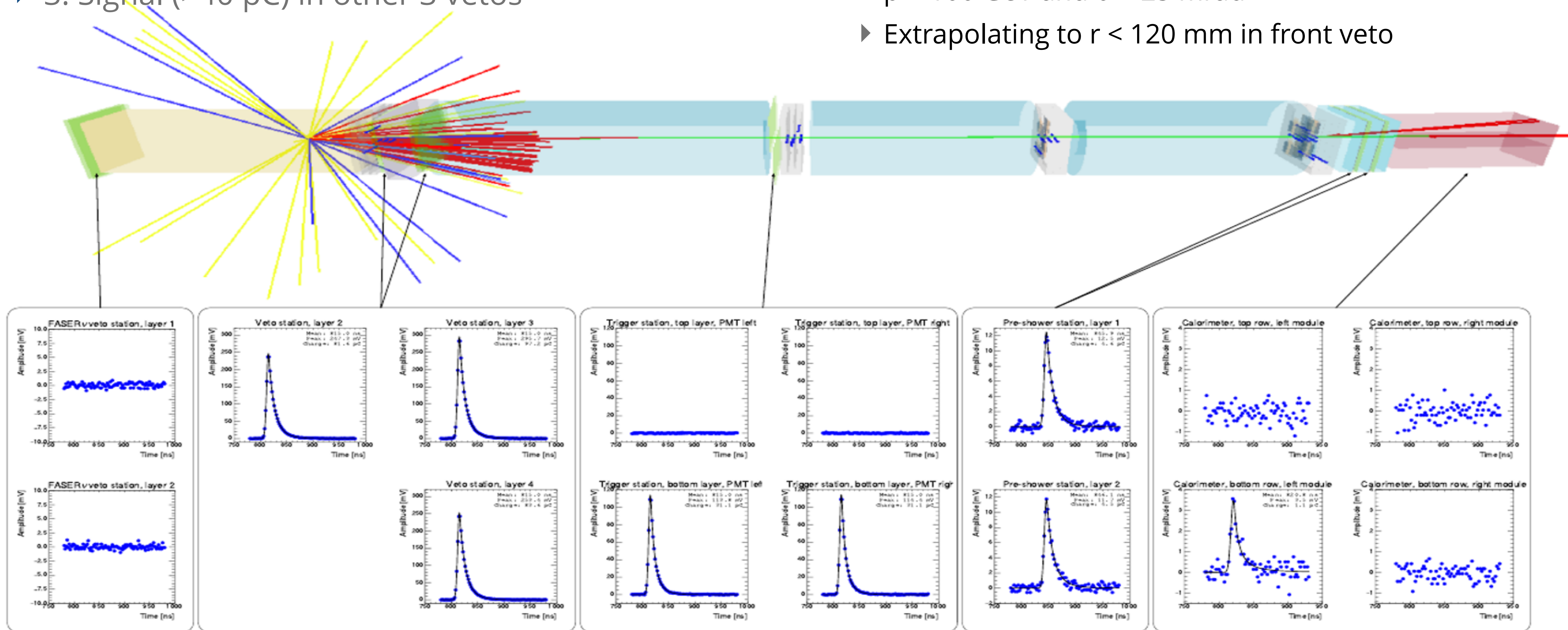
- ▶ Veto inefficiency estimated from final fit
 - ▶ Fit events with 0 (SR) and also 1 (1st or 2nd) or 2 front veto layers firing
 - ▶ Final negligible background due to very high veto efficiency



Neutrino analysis | Selection

[2303.14185]

- ▶ 1. Collision event with good data quality
- ▶ 2. No signal (<40 pc) in 2 front vetos
- ▶ 3. Signal (>40 pC) in other 3 vetos
- ▶ 4. Timing and preshower consistent with ≥ 1 MIP
- ▶ 5. Exactly 1 good fiducial ($r < 95$ mm) track
- ▶ $p > 100$ GeV and $\theta < 25$ mrad
- ▶ Extrapolating to $r < 120$ mm in front veto



Neutrino analysis | Results

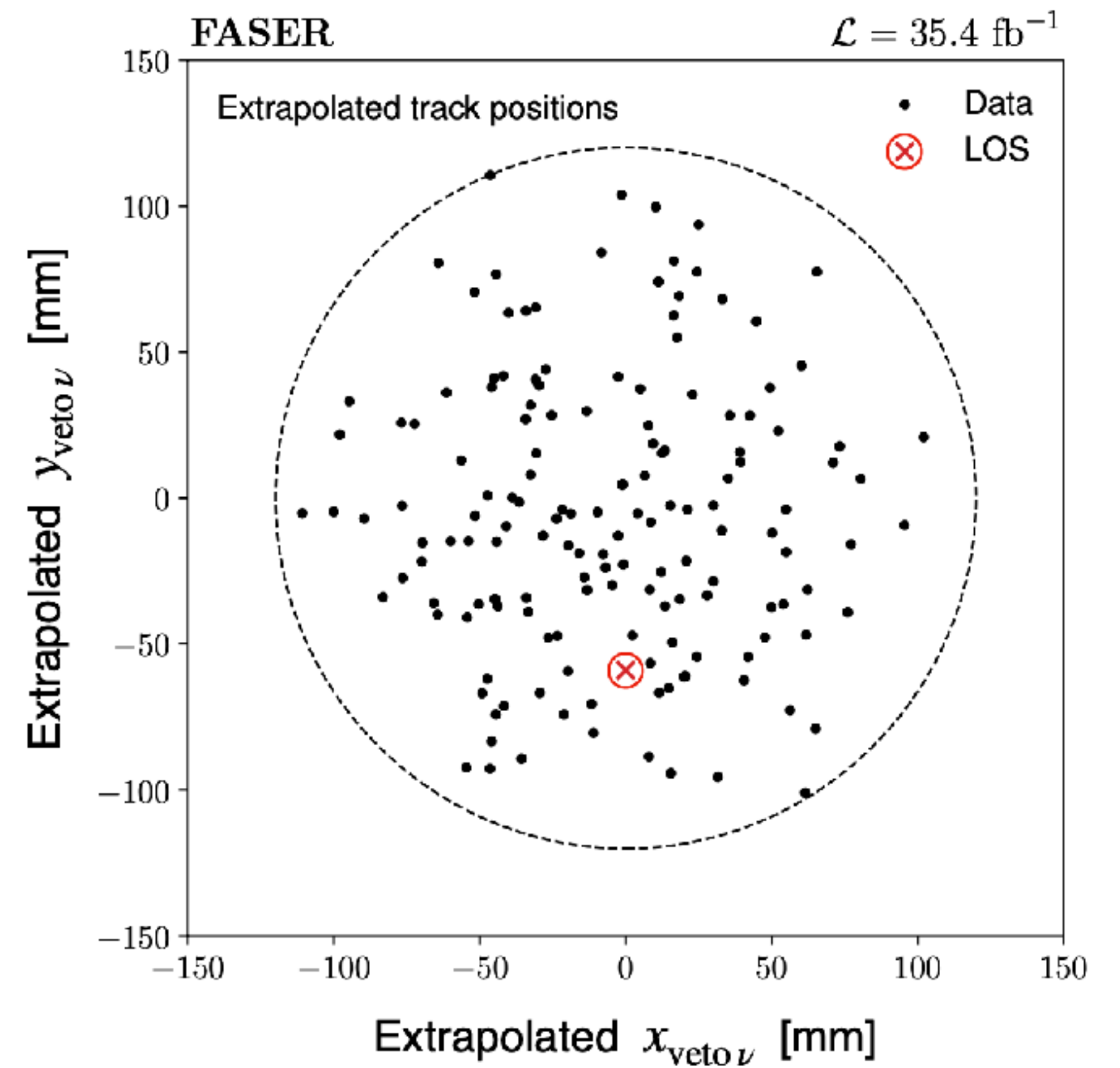
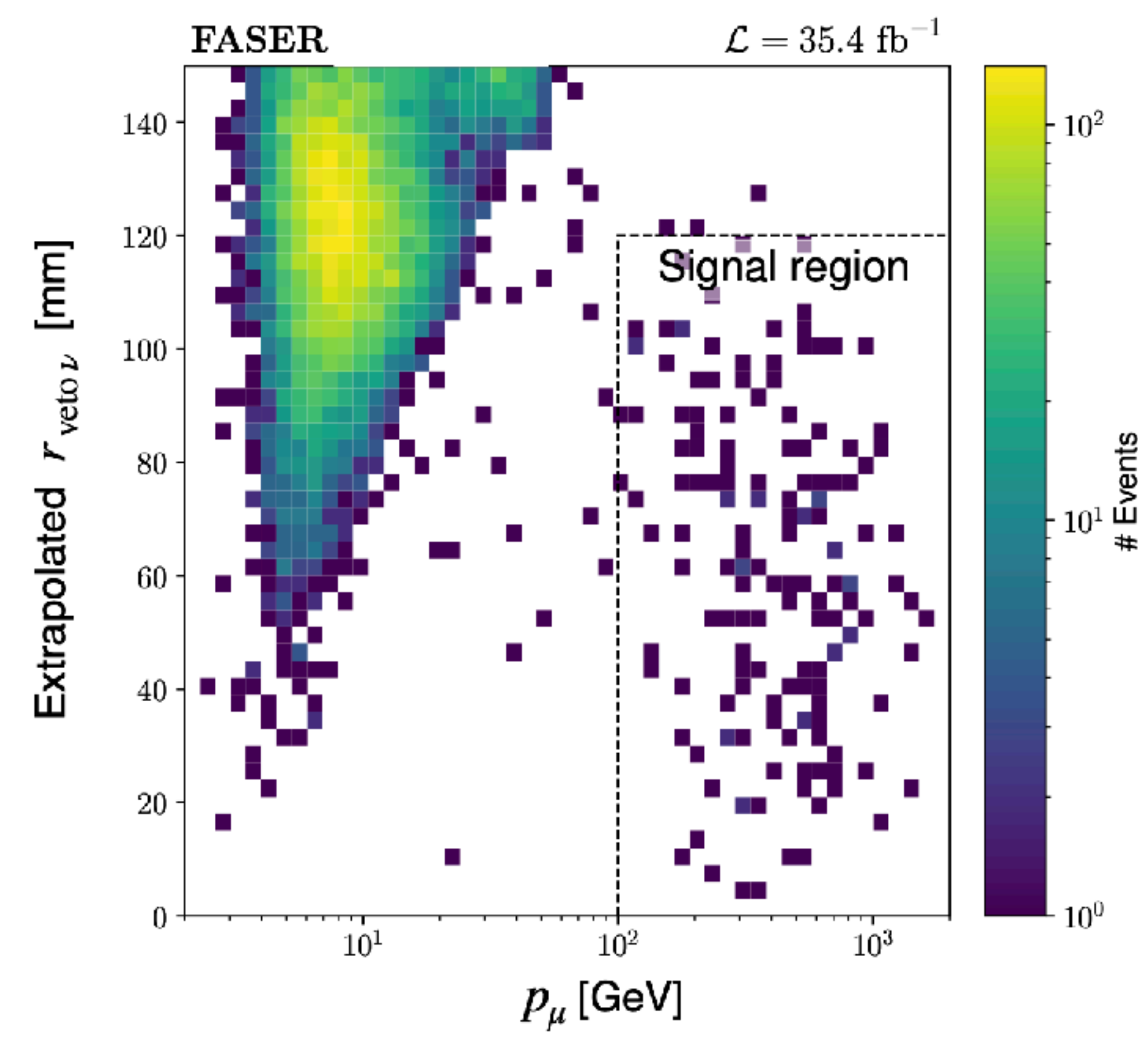
[2303.14185]



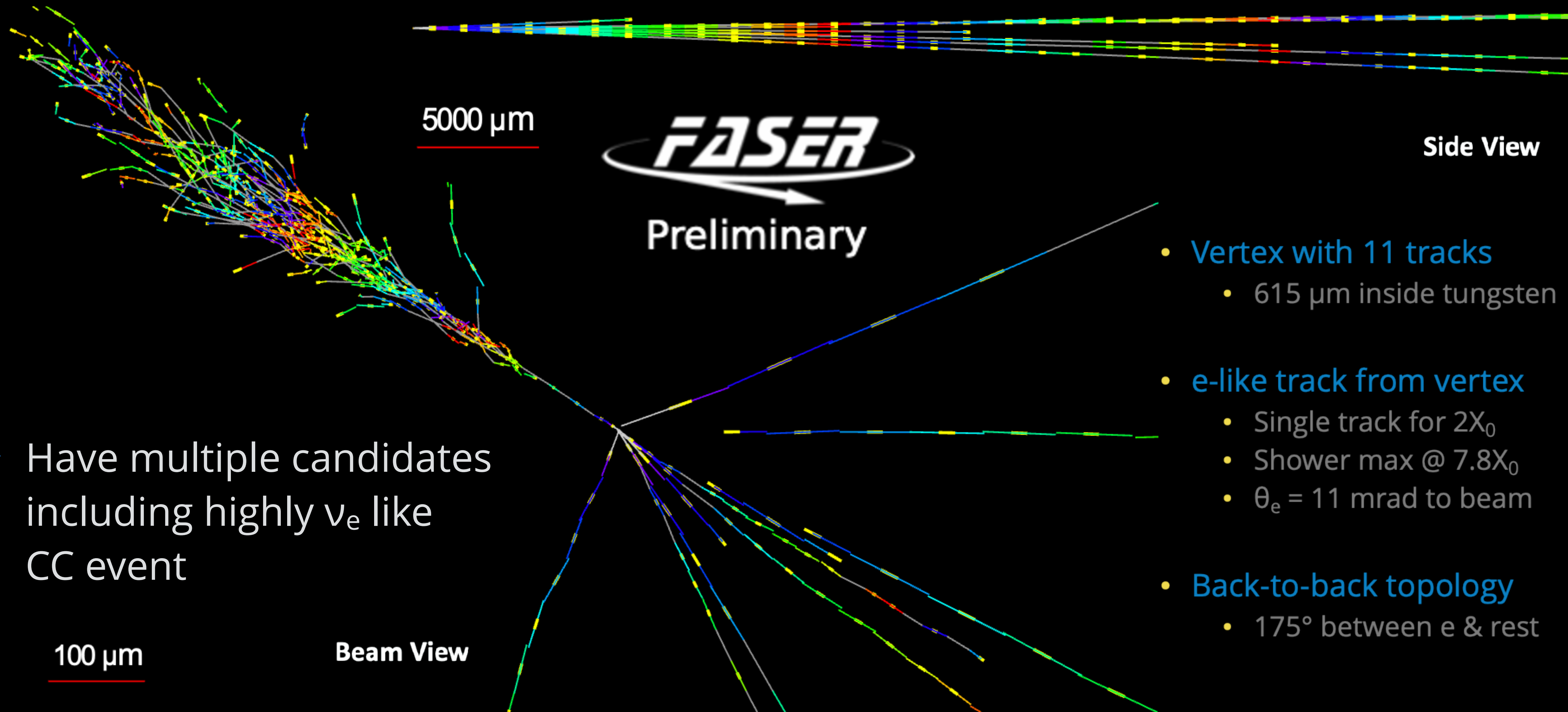
- ▶ Unblinded to find 153 events with no veto signal
- ▶ Just 10 events with one veto signal
- ▶ **First direct detection of collider neutrinos!**

Candidate	Events
n_0	153
n_{10}	4
n_{01}	6
n_2	64014695

- ▶ With signal significance of 16σ
- ▶ Expected 151 ± 41 events from GENIE simulation



- ▶ Analysis of FASERν emulsion detector underway

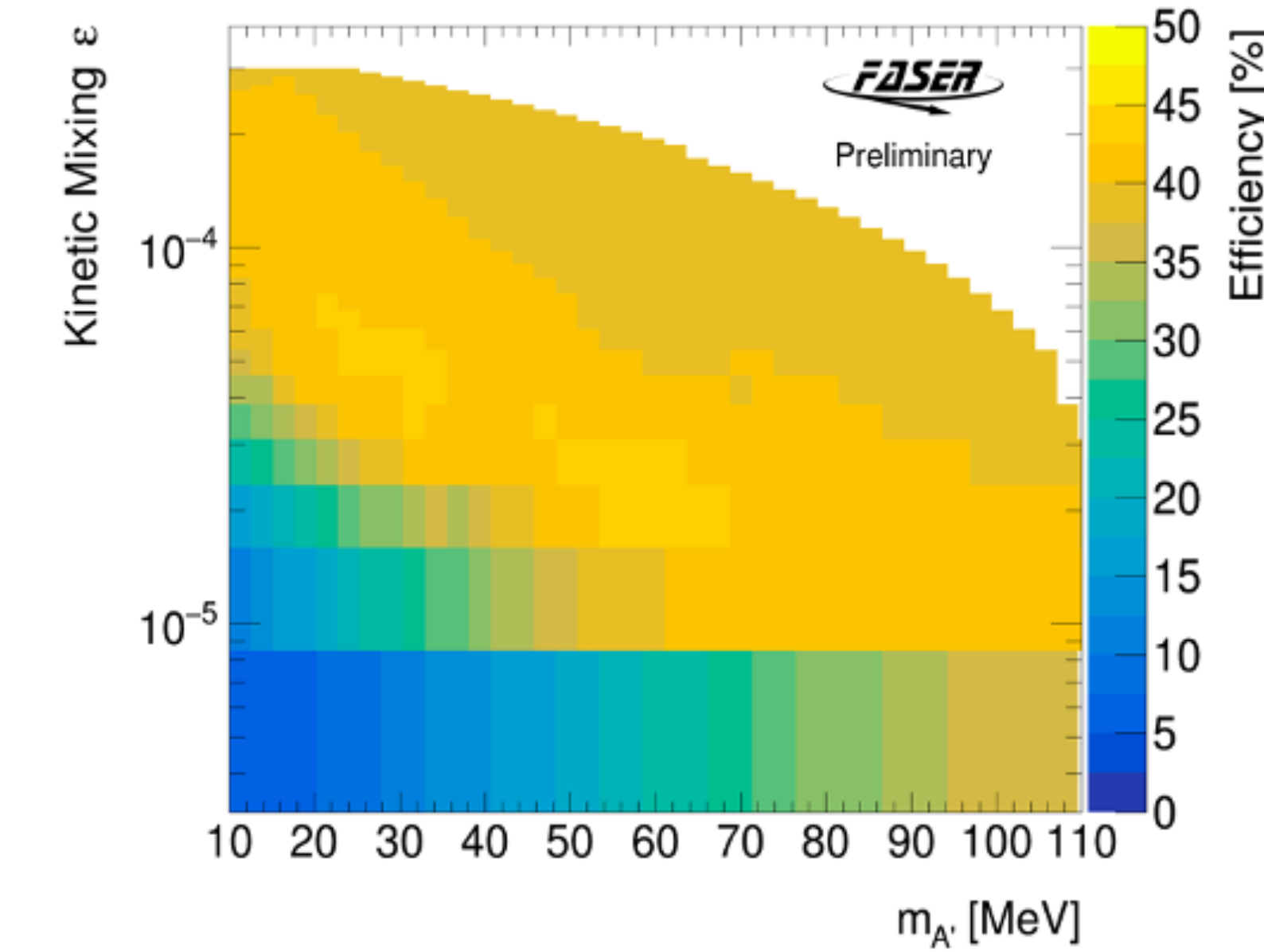


- ▶ Have multiple candidates including highly ν_e like CC event

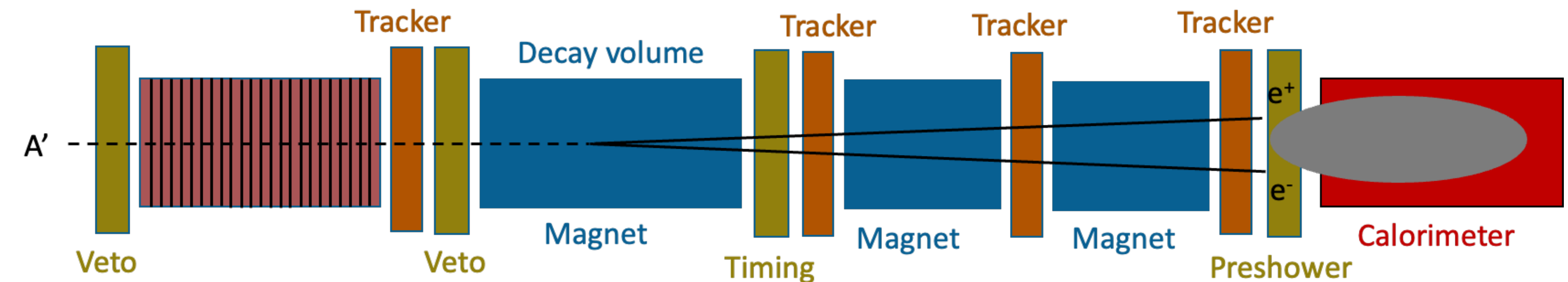
- Vertex with 11 tracks
 - 615 μm inside tungsten
- e-like track from vertex
 - Single track for $2X_0$
 - Shower max @ $7.8X_0$
 - $\theta_e = 11$ mrad to beam
- Back-to-back topology
 - 175° between e & rest

Dark Photon Search | Selection

- ▶ Simple and robust $A' \rightarrow e^+e^-$ selection
 - ▶ Blind events with no veto signal and $E(\text{calo}) > 100 \text{ GeV}$
- ▶ Selection
 - ▶ 1. Collision event with good data quality
 - ▶ 2. No signal ($< 40 \text{ pc}$) in any veto scintillator
 - ▶ 3. Timing and preshower consistent with $\geq 2 \text{ MIPs}$
 - ▶ 4. Exactly 2 good fiducial tracks
 - ▶ $p > 20 \text{ GeV}$ and $r < 95 \text{ mm}$ && Extrapolating to $r < 95 \text{ mm}$ at vetos
 - ▶ 5. Calo $E > 500 \text{ GeV}$

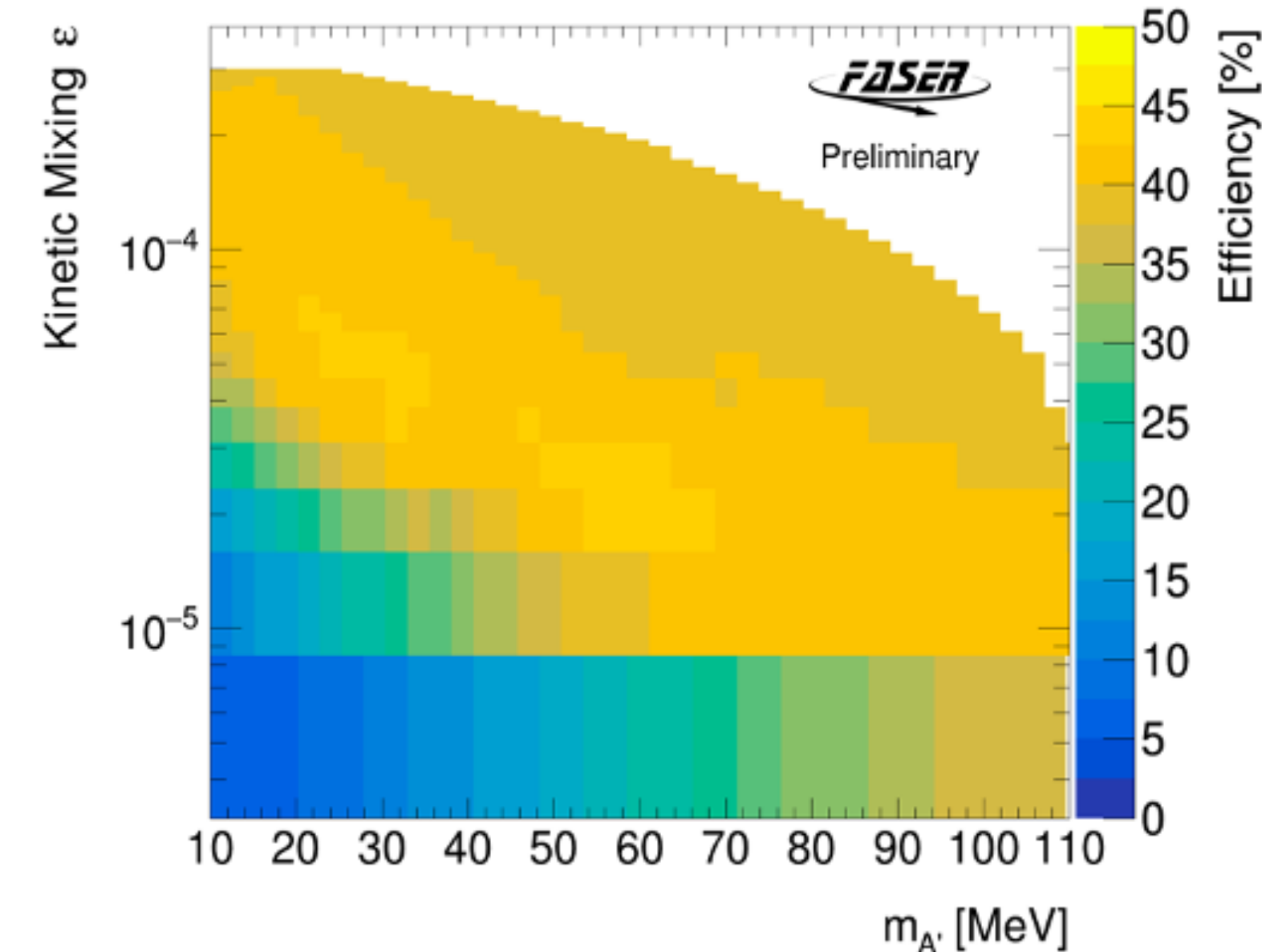


- ▶ Efficiency of $\sim 40\%$ across sensitive region

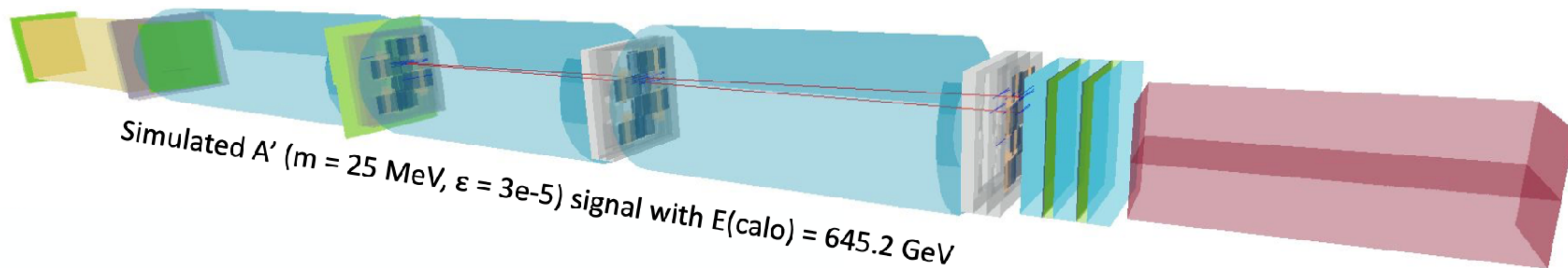


Dark Photon Search | Selection

- ▶ Simple and robust $A' \rightarrow e^+e^-$ selection
 - ▶ Blind events with no veto signal and $E(\text{calo}) > 100 \text{ GeV}$
- ▶ Selection
 - ▶ 1. Collision event with good data quality
 - ▶ 2. No signal ($< 40 \text{ pc}$) in any veto scintillator
 - ▶ 3. Timing and preshower consistent with $\geq 2 \text{ MIPs}$
 - ▶ 4. Exactly 2 good fiducial tracks
 - ▶ $p > 20 \text{ GeV}$ and $r < 95 \text{ mm}$ && Extrapolating to $r < 95 \text{ mm}$ at vetos
 - ▶ 5. Calo $E > 500 \text{ GeV}$

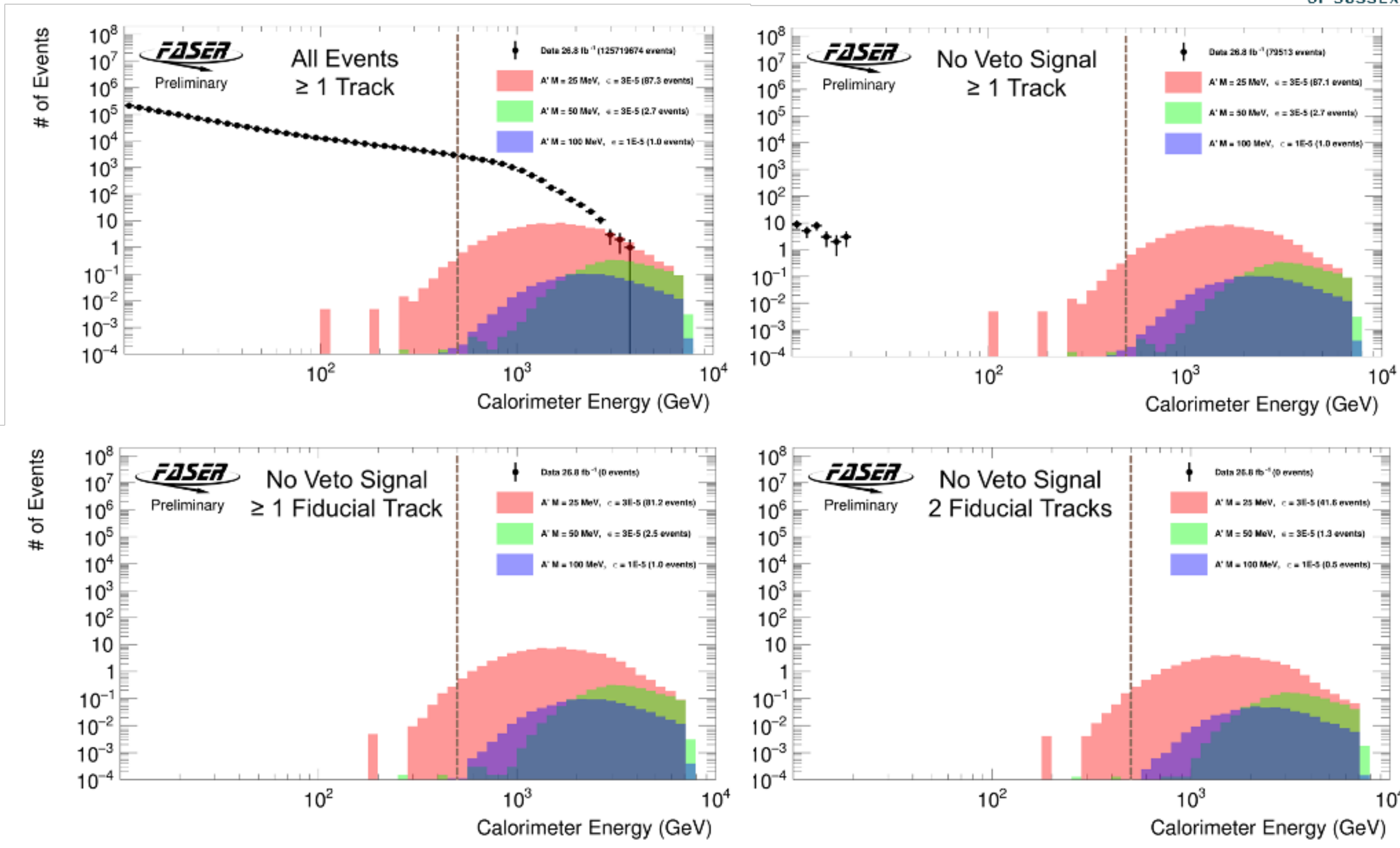


- ▶ Efficiency of $\sim 40\%$ across sensitive region



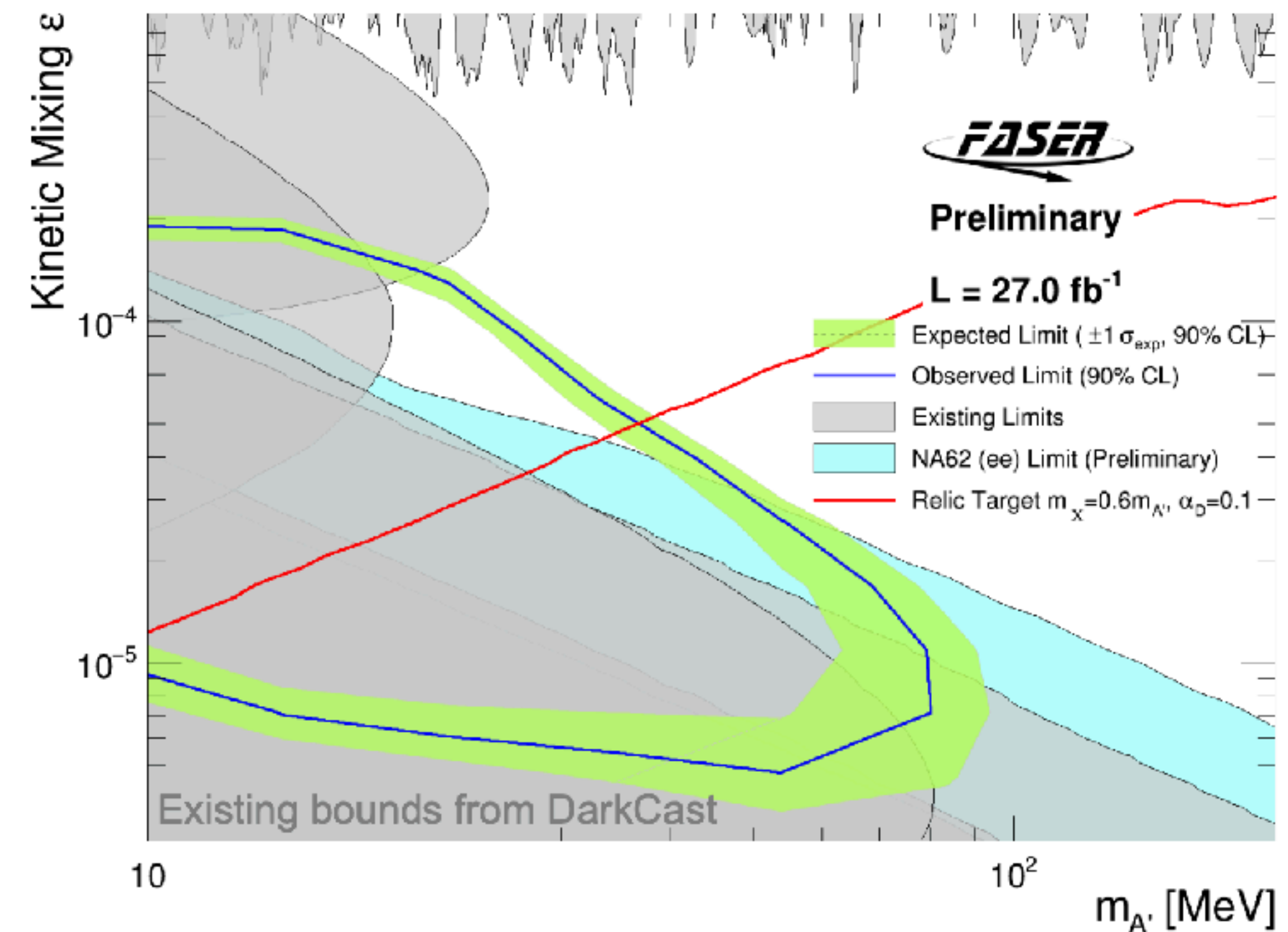
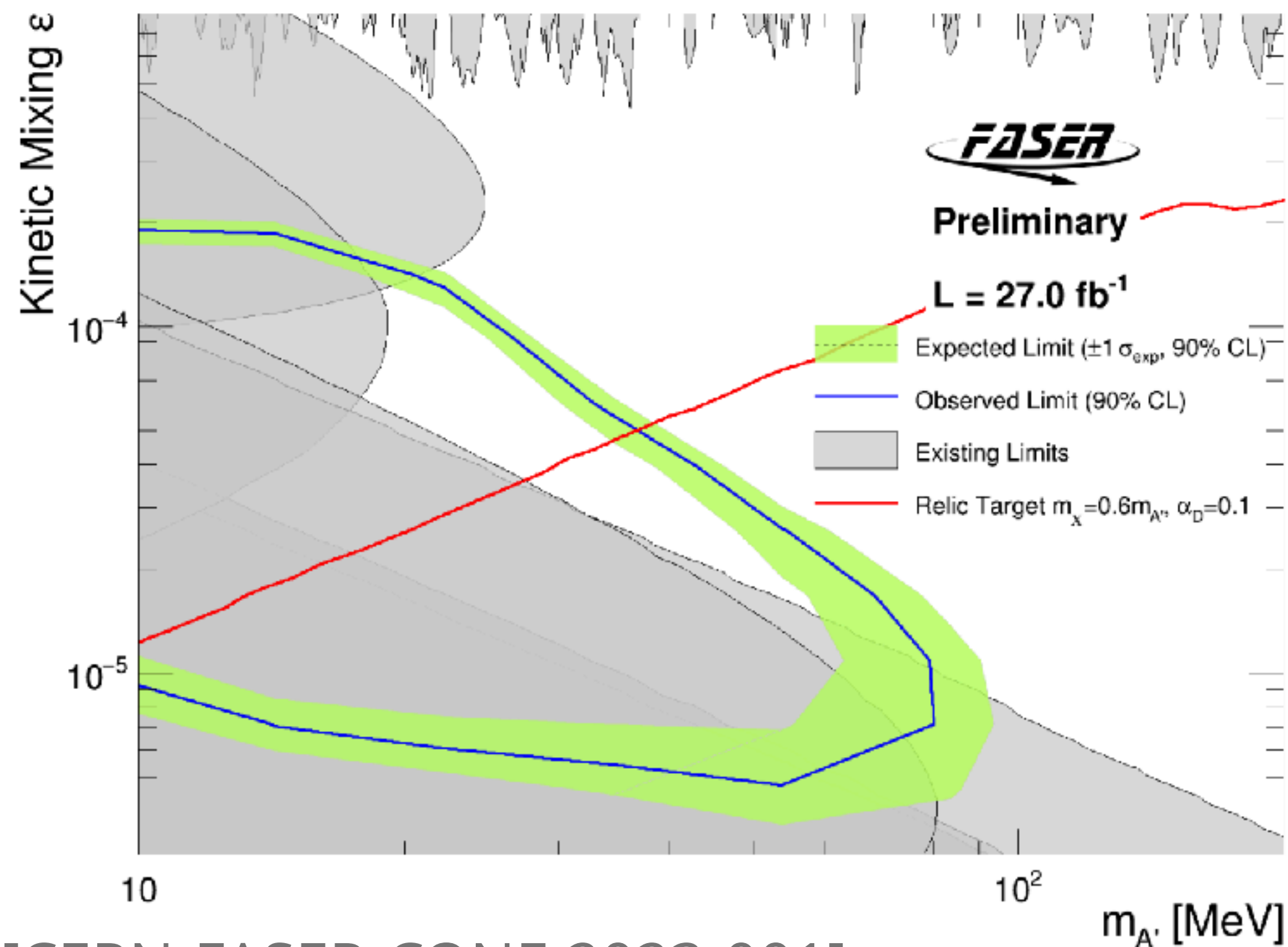
F Dark Photon Search | Results

- ▶ No events in unblinded signal region
- ▶ Not even any with ≥ 1 fiducial track

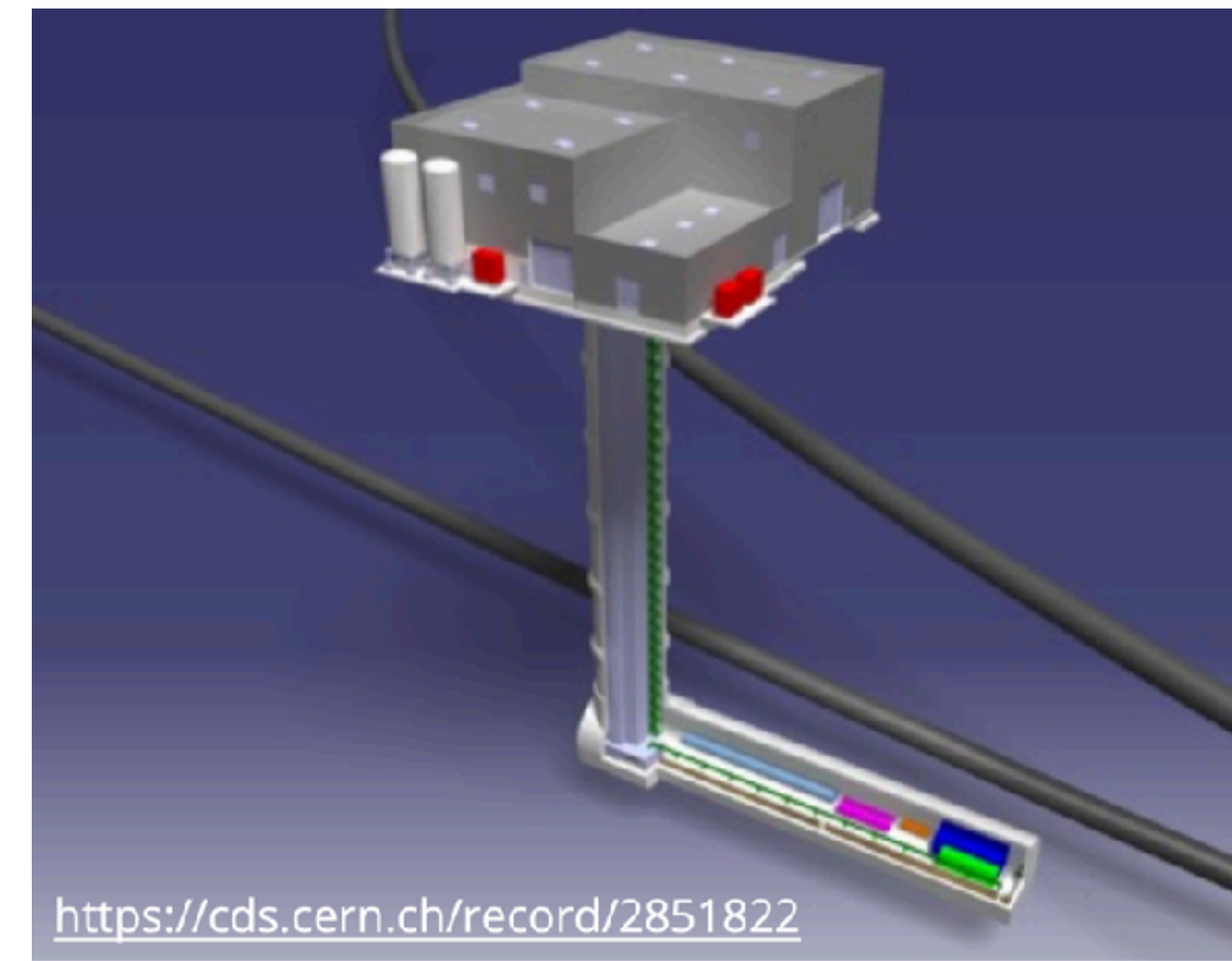


Dark Photon Search | Limits

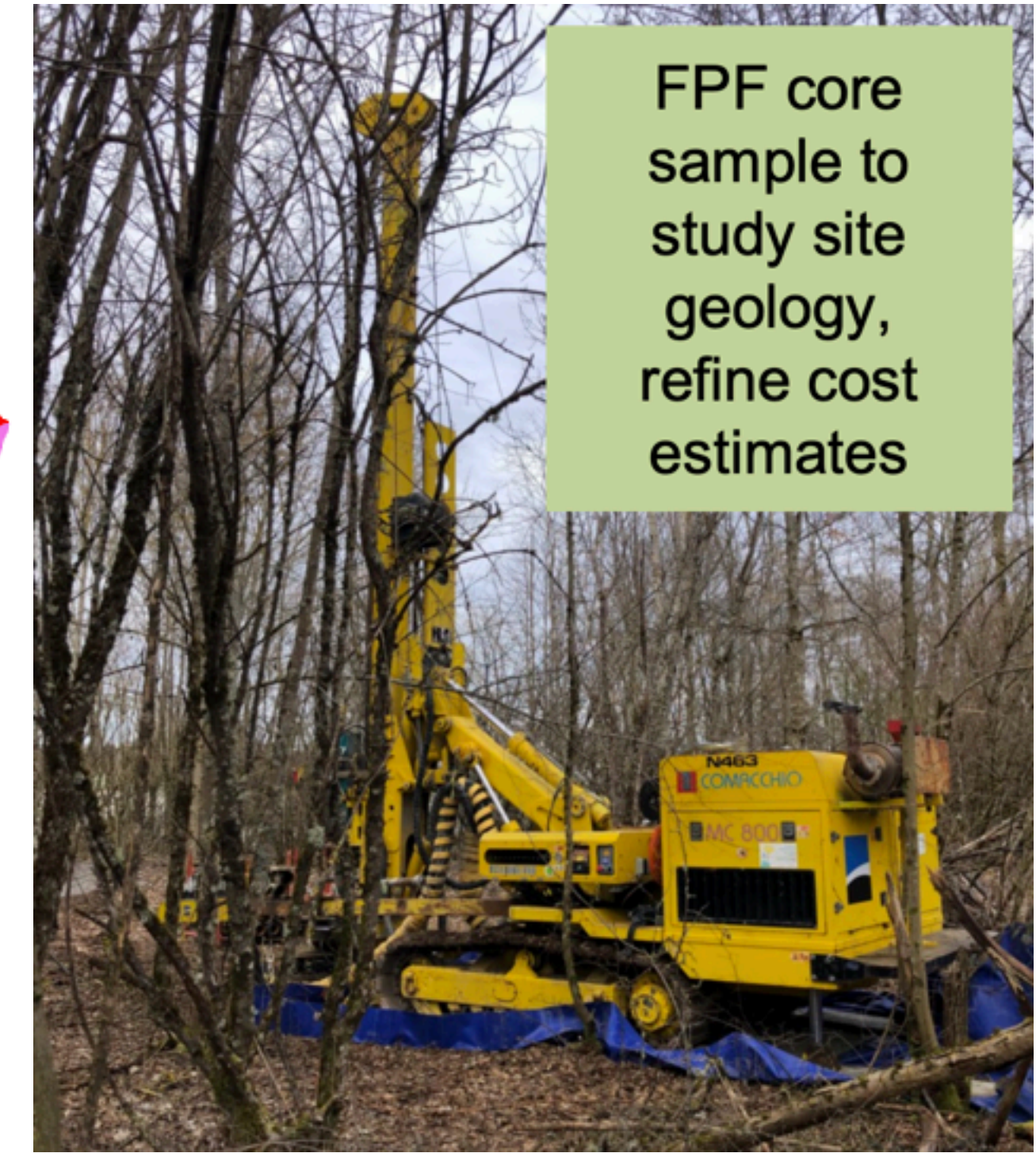
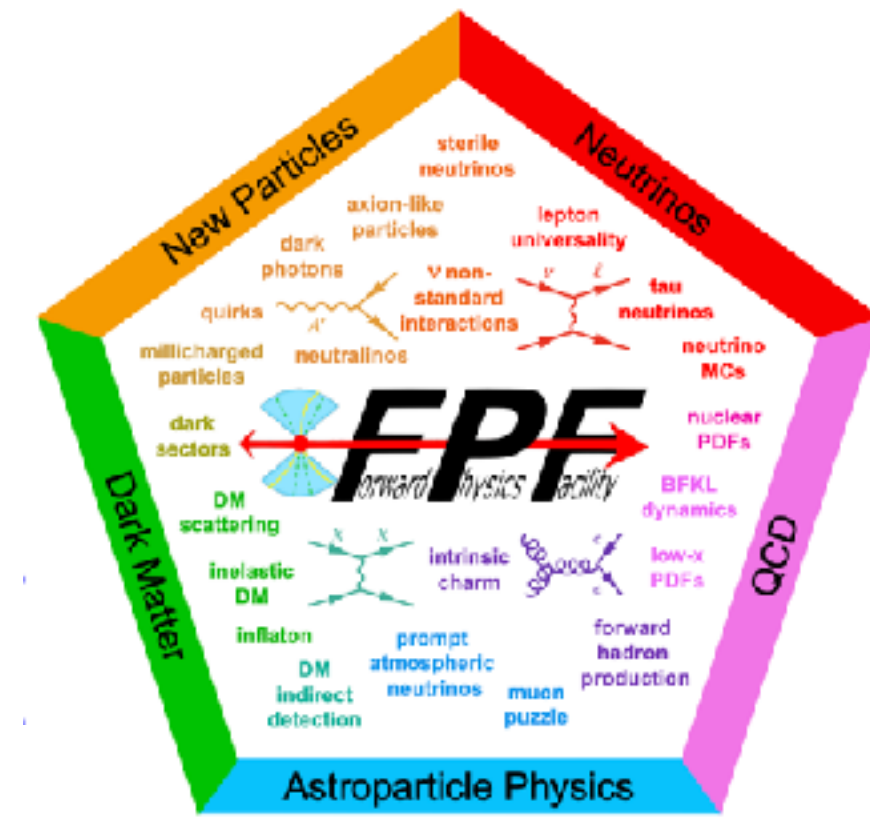
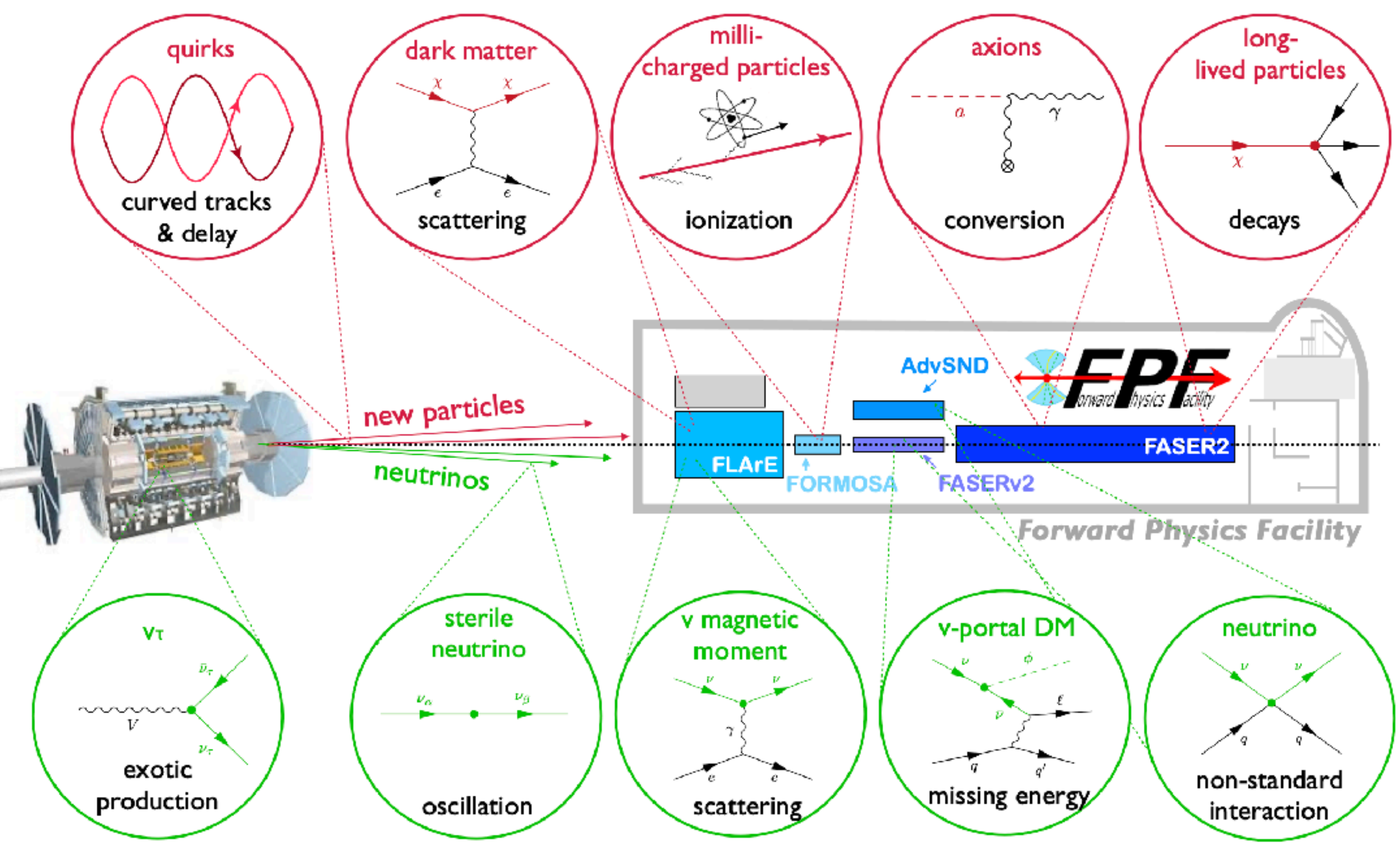
- ▶ After unblinding, no events seen in signal region, FASER sets limits on previously unexplored parameter space.
- ▶ First incursion (with NA62) into thermal relic region from low ϵ since 1990's.
- ▶ Background-free analysis bodes well for future sensitivity.
- ▶ Expect factor of ~ 10 more luminosity in Run 3 from 2022-25.



Forward Physics Facility

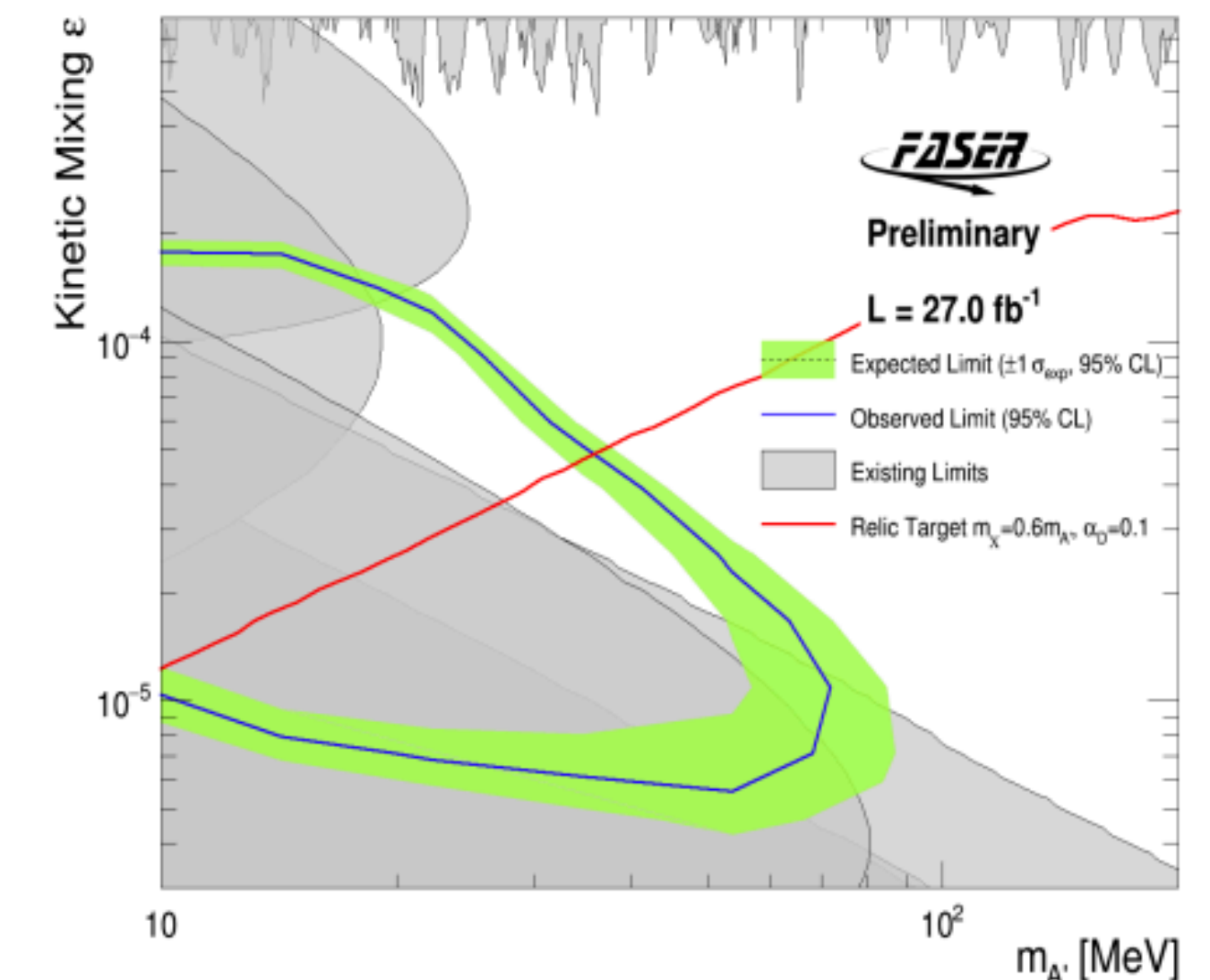
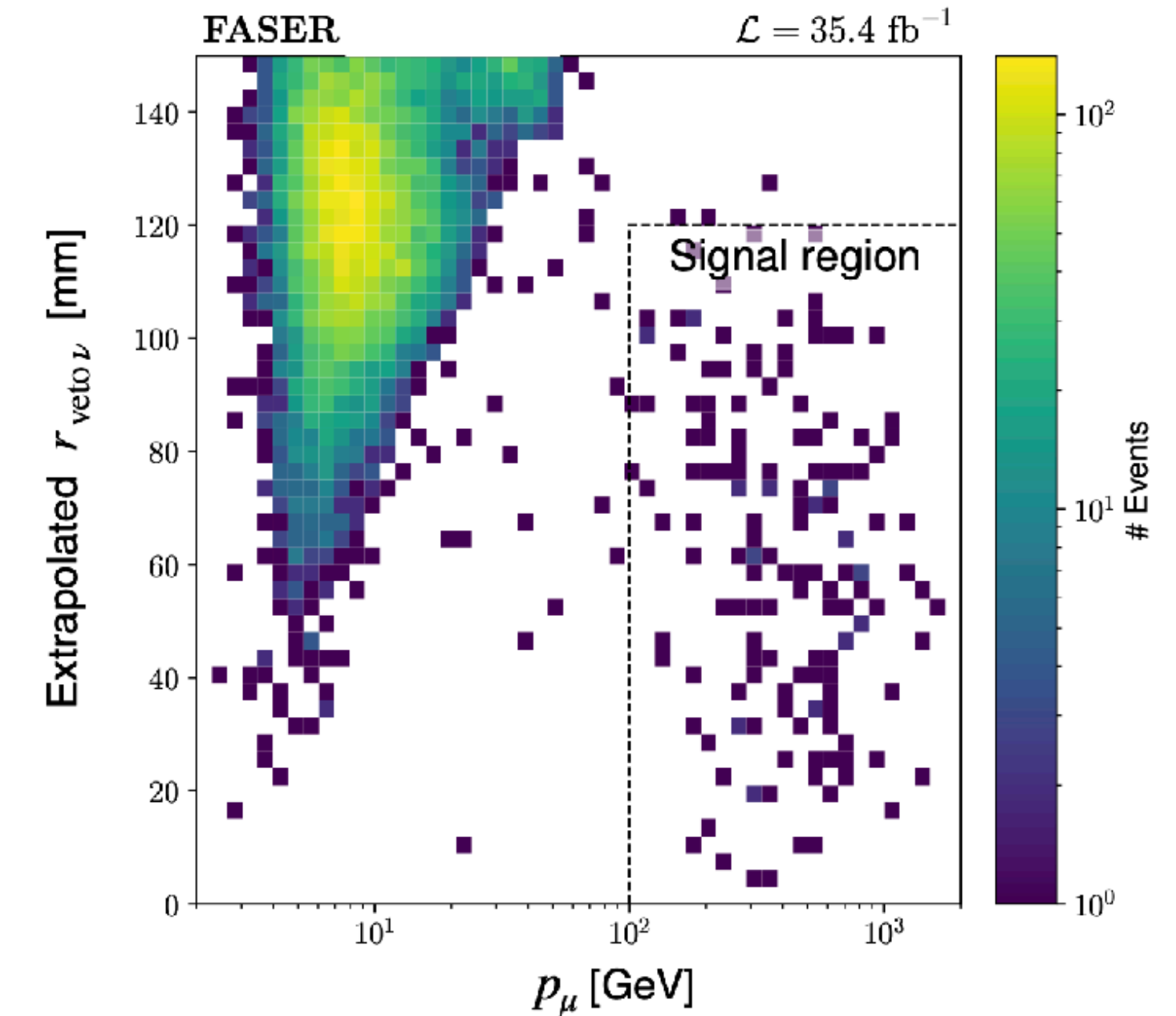


- ▶ FPF Papers:
 - ▶ FPF "Short" Paper: Phys. Rept. 968, 1 (2022)
 - ▶ FPF White Paper: J. Phys. G (2022)



Summary

- ▶ FASER successfully took data in 1st year of Run 3
 - ▶ Running with fully functional detector and very good efficiency
- ▶ First direct observation of collider neutrinos
 - ▶ Opens a new field: neutrino physics at the LHC
 - ▶ Submitted to PRL [[2303.14185](#)]
- ▶ Excluded A' in region of 10-100 MeV mass & very small ϵ
 - ▶ First limit in thermal relic region from low coupling for 30 yrs
 - ▶ [[CERN-FASER-CONF-2023-001](#)]
- ▶ More neutrino studies and BSM searches to come
 - ▶ Including first results from emulsion detector
 - ▶ Searches for ALPs, light gauge bosons, ...
- ▶ Strongly motivates FPF for the HL-LHC era



Acknowledgements

▶ FASER is supported by:

SIMONS
FOUNDATION

HEISING-SIMONS
FOUNDATION



Swiss National
Science Foundation



科研費
KAKENHI



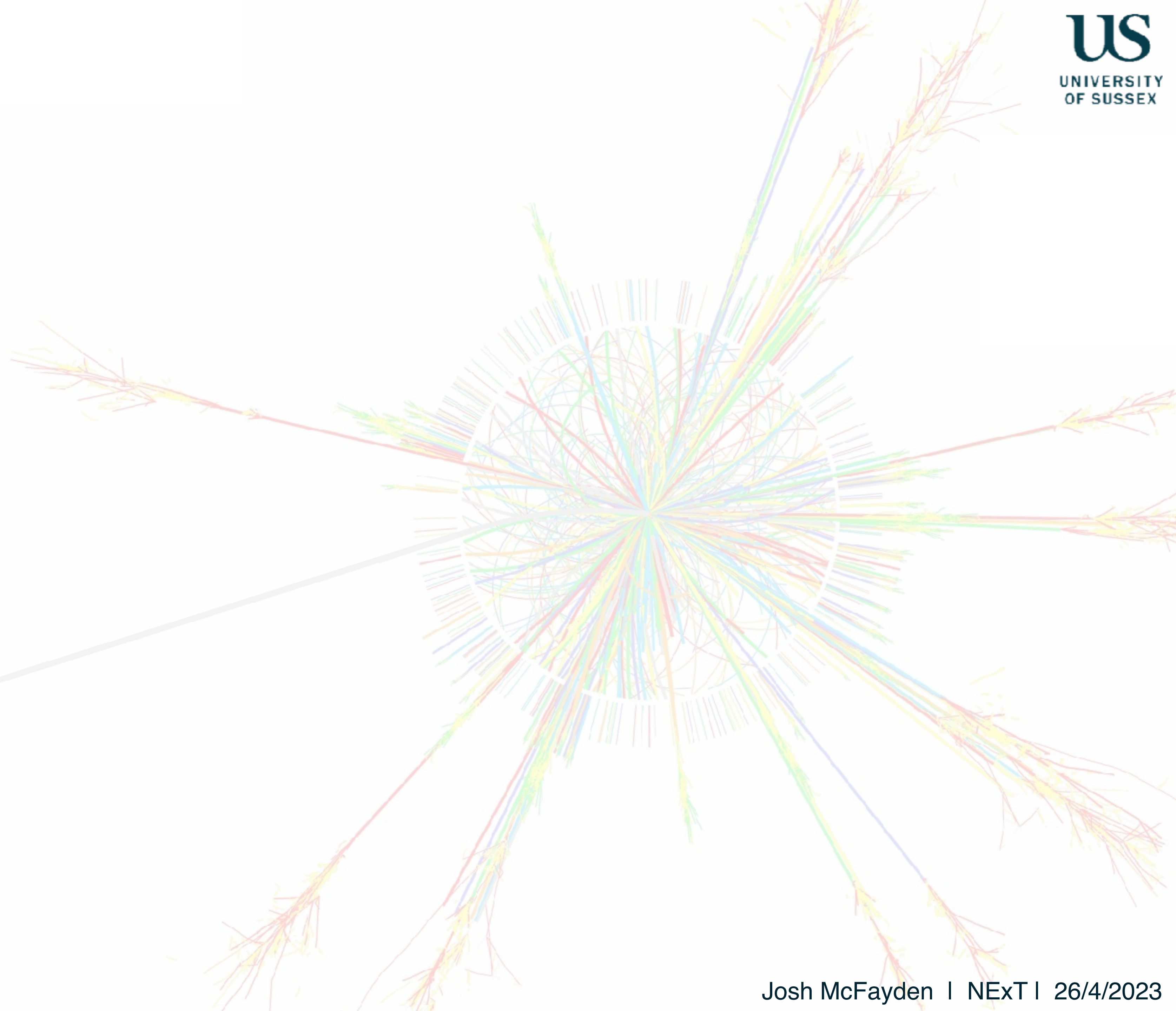
国家自然科学基金委员会
National Natural Science Foundation of China

- ▶ And would additionally like to thank
 - ▶ LHC for the excellent performance in 2022
 - ▶ ATLAS for providing luminosity information
 - ▶ ATLAS for use of ATHENA s/w framework
 - ▶ ATLAS SCT for spare tracker modules
 - ▶ LHCb for spare ECAL modules
 - ▶ CERN FLUKA team for background sim
 - ▶ CERN PBC and technical infrastructure groups for excellent support during design construction and installation

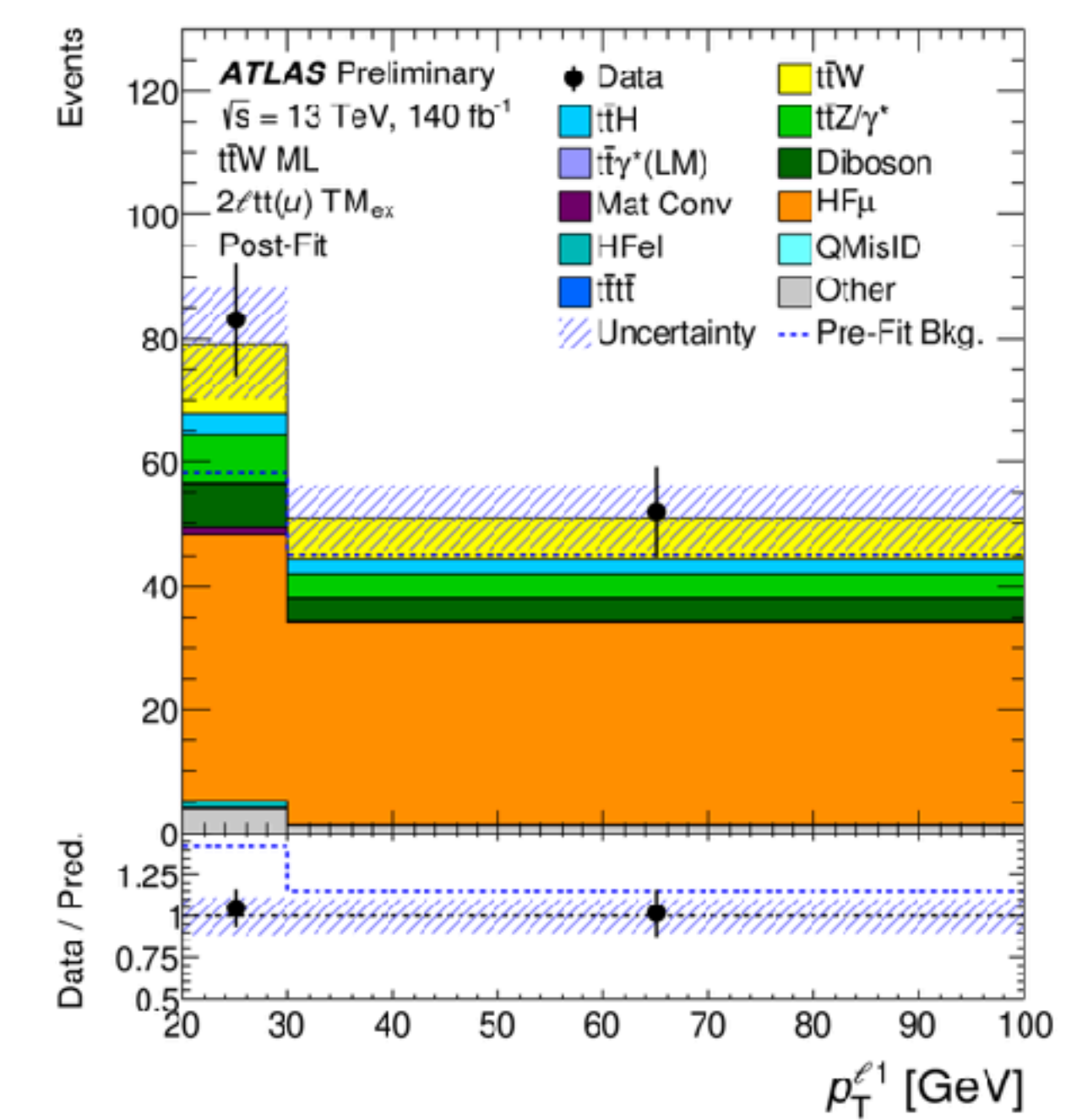
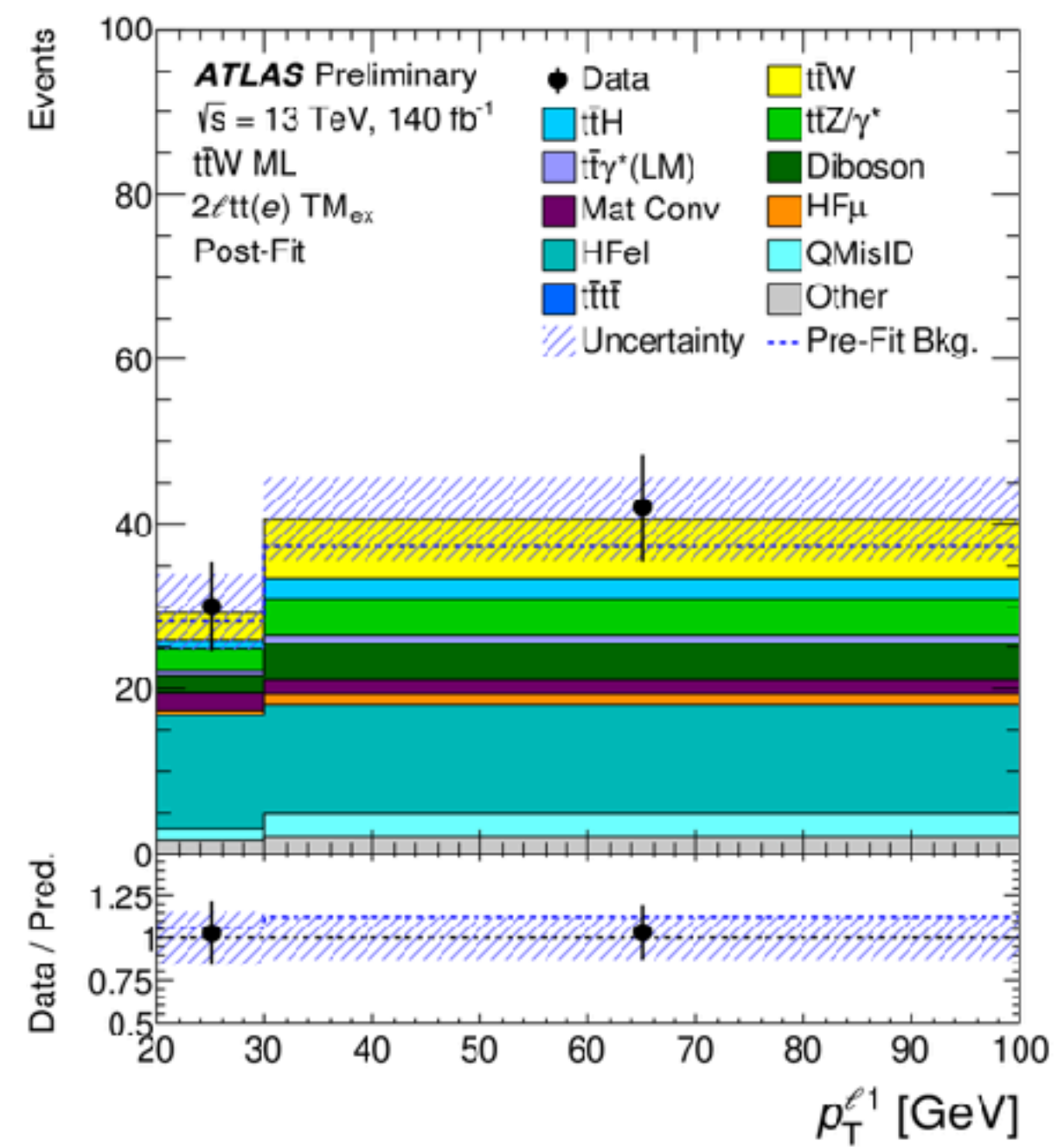
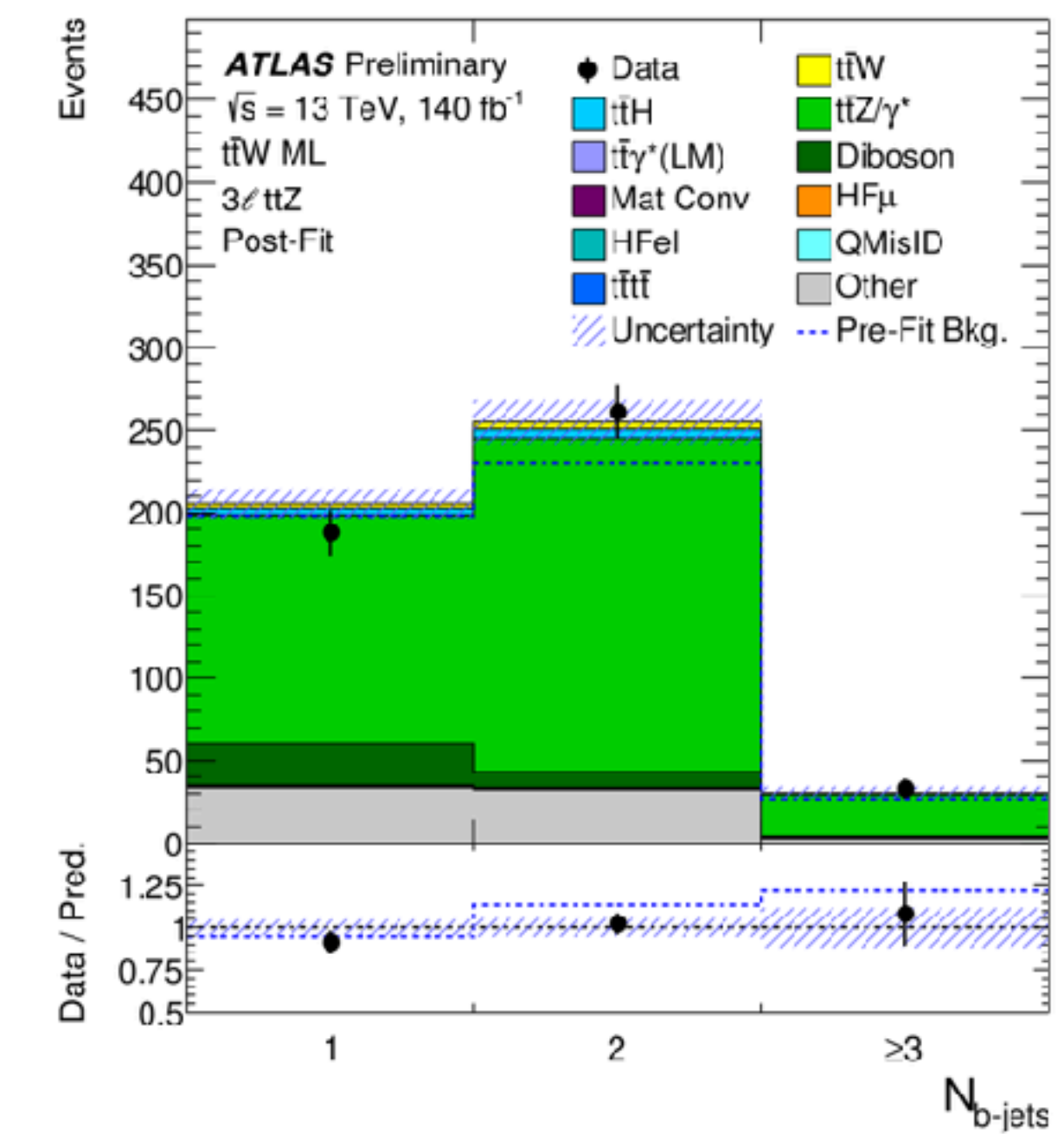
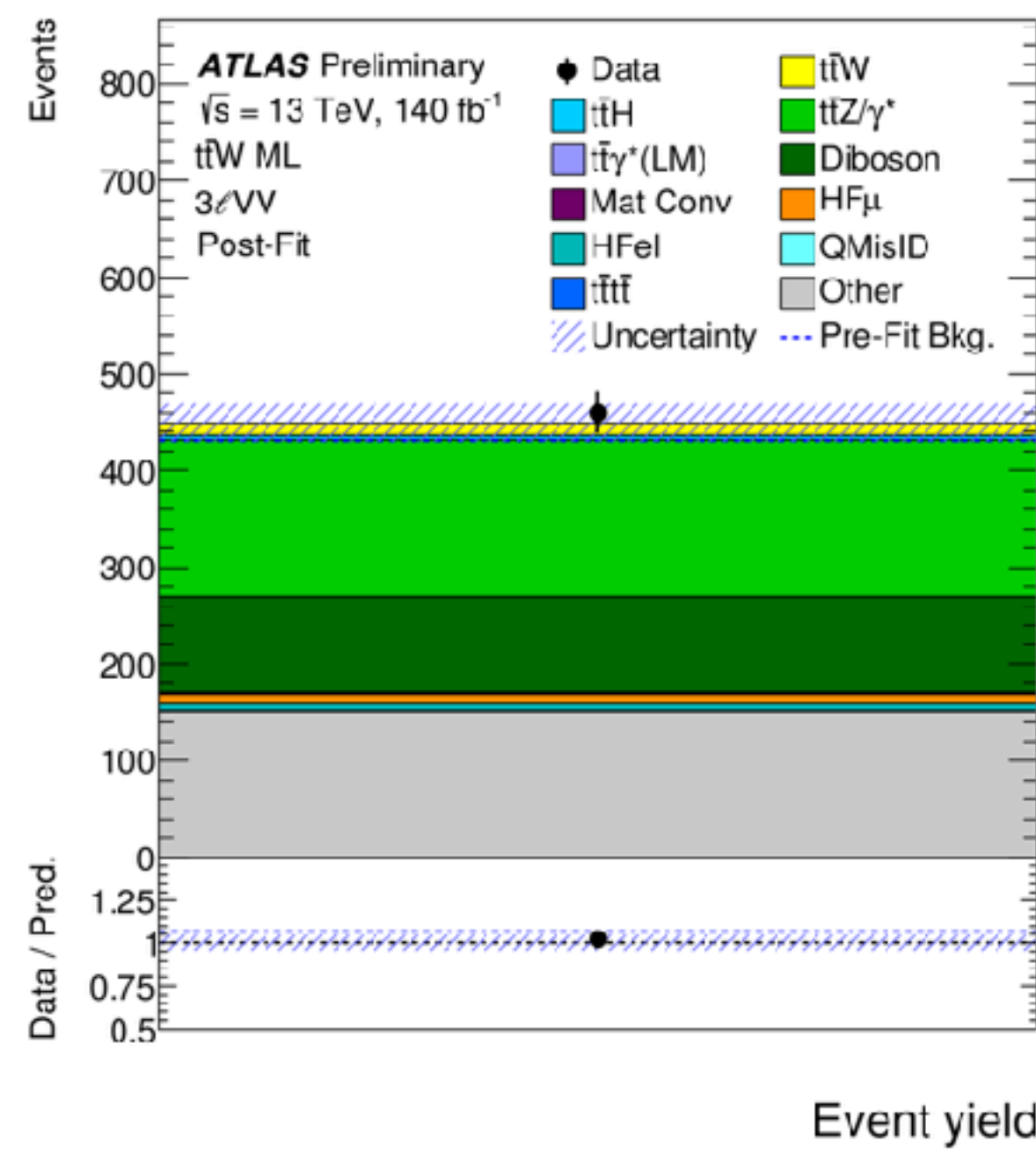




ttW back-ups



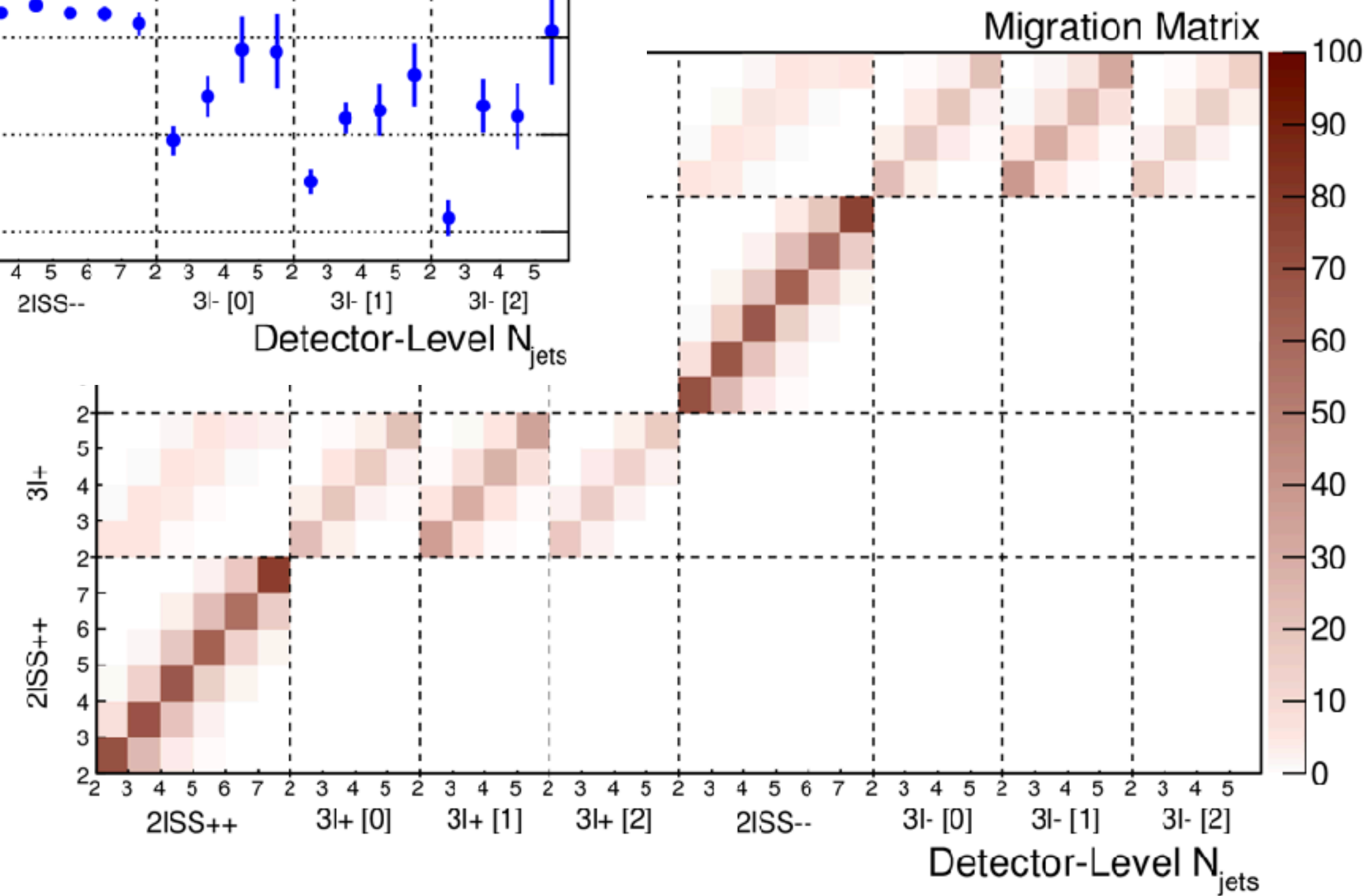
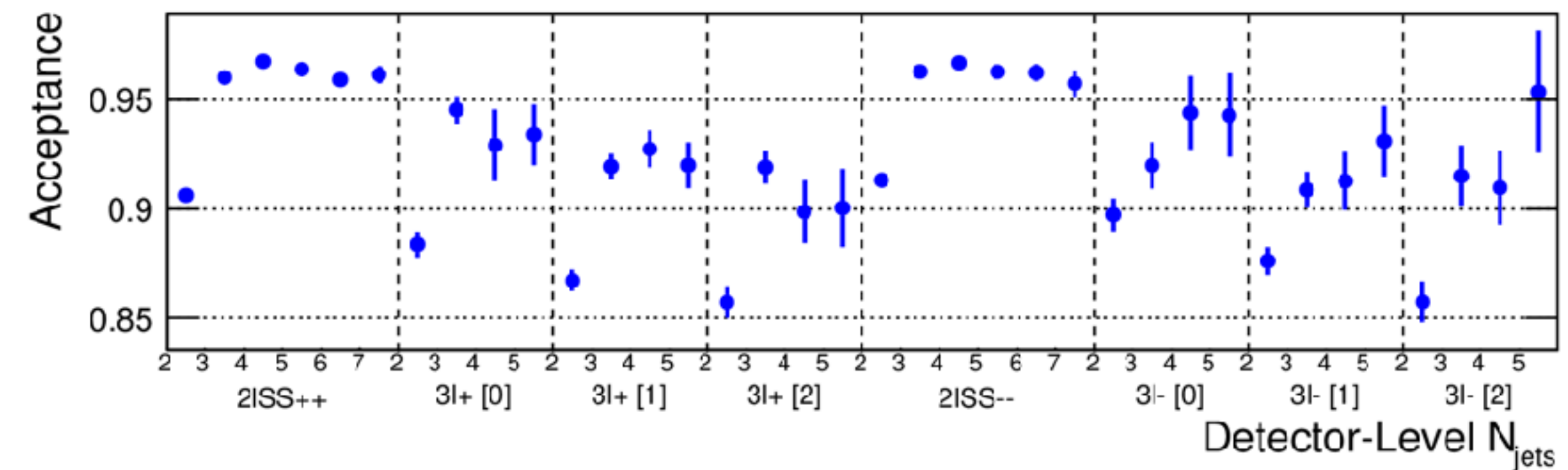
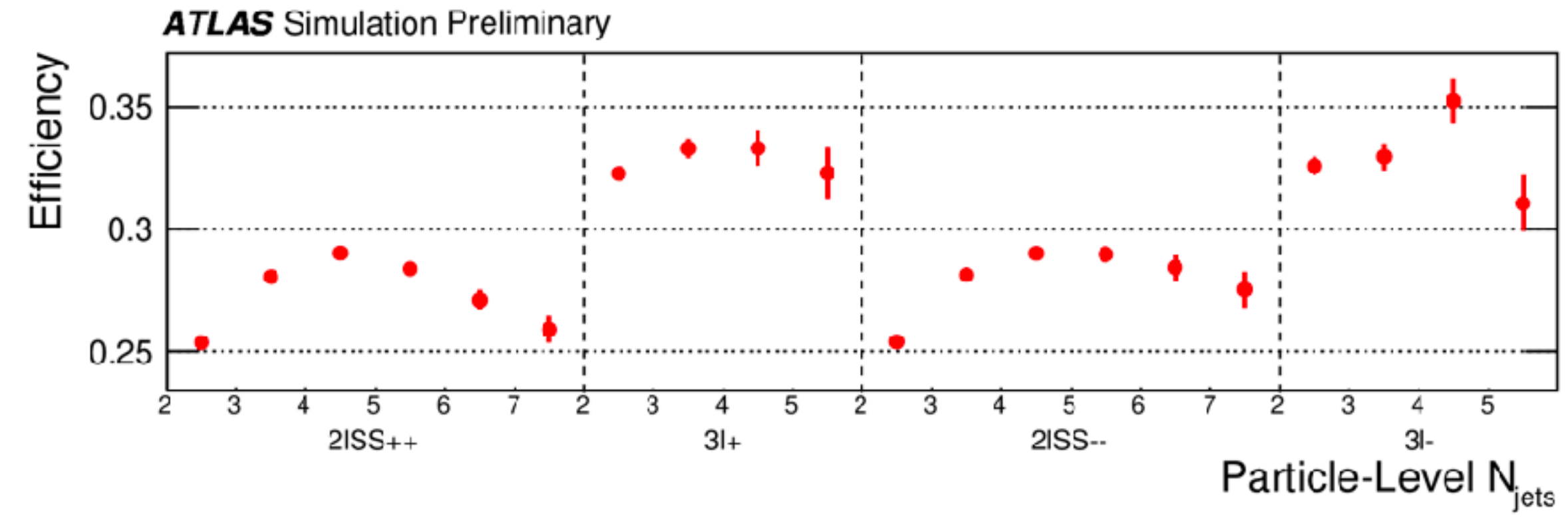
Backgrounds





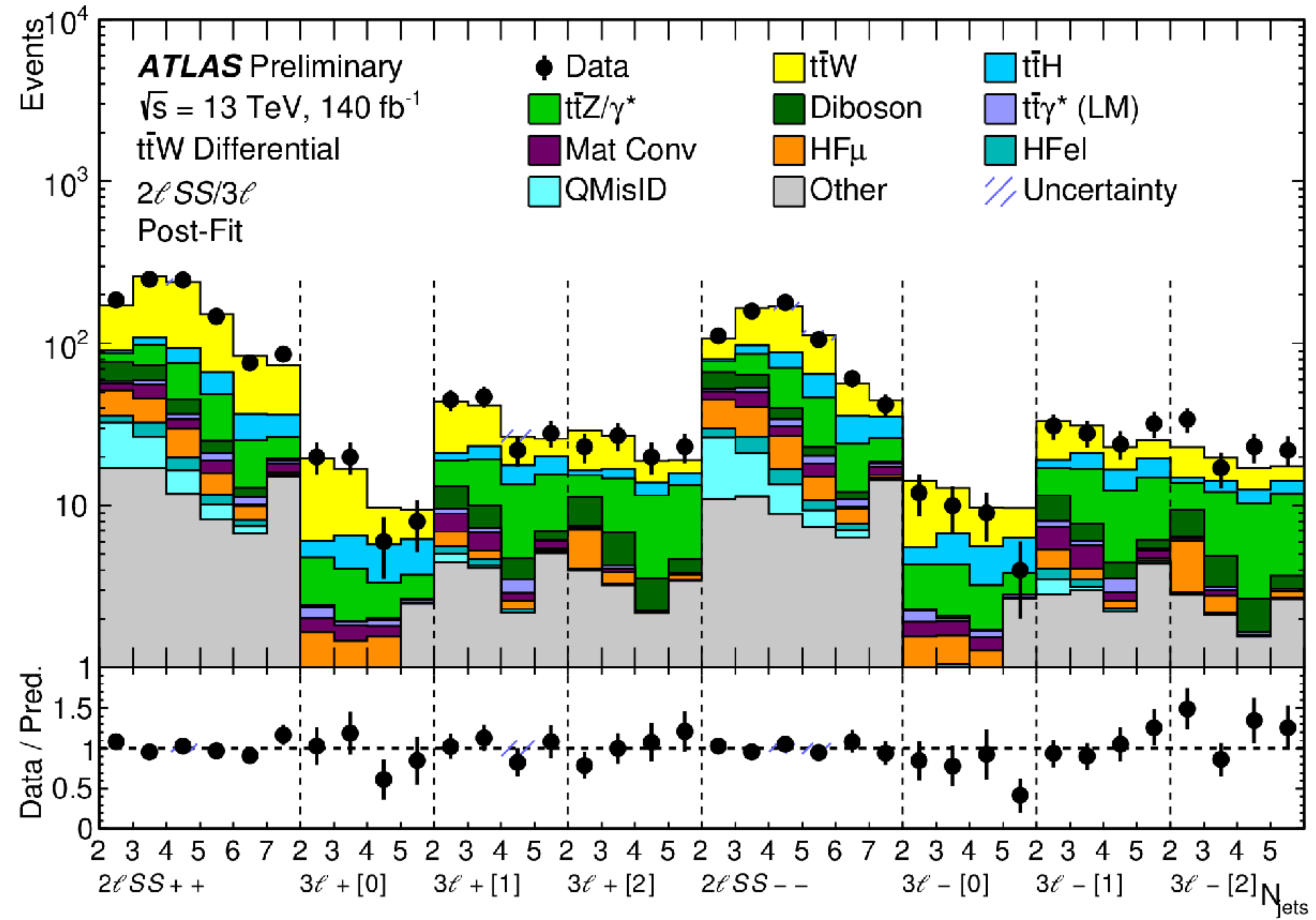
Differential results | Unfolding

► Profile likelihood unfolding technique used.



Differential results

- ▶ Background estimate identical
- ▶ SR split by observable



Contribution	Orders	% change	Cross section [fb]	Reference
LO QCD	$\alpha\alpha_s^2$	–	$363^{+24\%}_{-18\%}$	[30]
+NLO QCD	$+\alpha\alpha_s^3$	+50.0%	$544^{+11\%}_{-11\%}$	[30]
+ “leading” NLO EW	$+\alpha^2\alpha_s^2$	-4.2%	–	[30]
+ “subleading” NLO EW	$+\alpha^3\alpha_s$	+12.2%	–	[30]
+ “all” NLO EW	$+\alpha^3 + \alpha^4$	1.1%+1.3%	$577^{+11\%}_{-11\%}$	[30]
$t\bar{t}W + 0, 1, 2j$ FxFx	$\alpha\alpha_s^2 + (\text{at 1-loop})\alpha\alpha_s^3 + \alpha\alpha_s^4$	–	$691.1^{+9.5\%}_{-10.7\%}$	[28]
+ “subleading” NLO EW	$+\alpha^3\alpha_s^2$	+6.9%	$738.8^{+10.1\%}_{-11.0\%}$	[28]
+ “leading” NLO EW	$+\alpha^2\alpha_s^2$	-2.4%	$722.4^{+9.7\%}_{-10.8\%}$	[28]
NLO+NNLL QCD	–	–	$571.43^{+8.6\%}_{-5.7\%}$	[33]
NLO+NNLL QCD+EW	–	–	$606.13^{+8.9\%}_{-5.8\%}$	[33]
Offshell LO QCD ($\sigma(\text{fiducial}3\ell2b)$)	$\alpha^6\alpha_s^2$	–	$0.2218^{+25.3\%}_{-18.8\%}$	[38]
+NLO QCD	$+\alpha^6\alpha_s^3$	+6.6%	–	[38]
+ “leading” NLO EW	$+\alpha^7\alpha_s^2$	-5.5%	–	[38]
+ “subleading” NLO EW	$+\alpha^8\alpha_s$	+13.1%	–	[38]
+ “other” NLO EW	$+\alpha^8$	+1.0%	$0.2554^{+4.0\%}_{-6.5\%}$	[38]

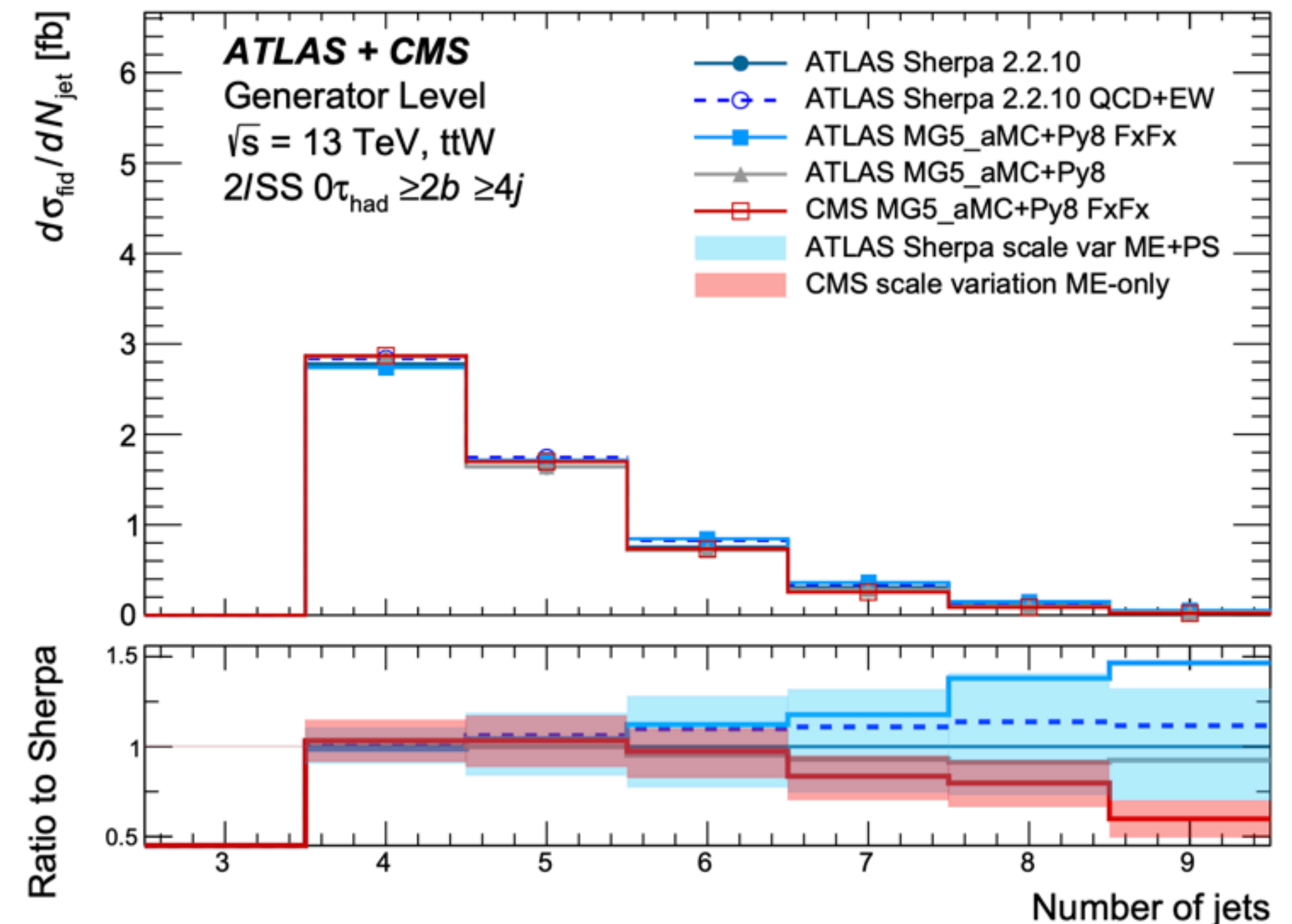
Table 1: Overview of theoretical predictions and their impact on the inclusive $t\bar{t}W$ cross section.

- [28] R. Frederix and I. Tsinikos, *On improving NLO merging for $t\bar{t}W$ production*, [JHEP 11 \(2021\) 029](#), arXiv: [2108.07826 \[hep-ph\]](#) (cit. on p. 5).
- [30] R. Frederix, D. Pagani and M. Zaro, *Large NLO corrections in $t\bar{t}W^\pm$ and $t\bar{t}t$ hadroproduction from supposedly subleading EW contributions*, [JHEP 02 \(2018\) 031](#), arXiv: [1711.02116 \[hep-ph\]](#)
- [33] A. Broggio et al., *Top-quark pair hadroproduction in association with a heavy boson at NLO+NNLL including EW corrections*, [JHEP 08 \(2019\) 039](#), arXiv: [1907.04343 \[hep-ph\]](#)
- [38] A. Denner and G. Pelliccioli, *Combined NLO EW and QCD corrections to off-shell $t\bar{t}W$ production at the LHC*, [Eur. Phys. J. C 81 \(2021\) 354](#), arXiv: [2102.03246 \[hep-ph\]](#) (cit. on p. 5).



Path to reference cross section

- ▶ From LHC Top WG discussion last summer came proposal to Frederix et al. [[2108.07826](#)]
- ▶ This is the largest of all ttW prediction cross sections at 722 fb
 - ▶ Does not agree well with ~equivalent Sherpa prediction of 616 fb
 - ▶ 15% difference is larger than scale uncertainties
- ▶ LHC Higgs WG ttH/tH subgroup note with generator comparisons
 - ▶ Comparison between ATLAS & CMS setups
 - ▶ Comparison between generators i.e. systematics





Path to reference cross section?

- ▶ Discussion in Joint LHC Top — LHC Higgs WGs meeting in December (agenda)
- ▶ Summary from experimental side
- ▶ Summary from theory side
 - ▶ Including direct input from MG5_aMC & Sherpa authors
- ▶ Stopped short of being able to define a WG-level **recommendation**
- ▶ Did get **agreement** between ATLAS and CMS to use Frederix et al. as reference cross-section
- ▶ But did not discuss if/how to normalise QCD & EW parts separately

Joint session of LHC Top and Higgs working groups: ttW modeling in light of ttH measurements

Friday 9 Dec 2022, 13:00 → 17:00 Europe/London

Virtual meeting

Speakers: Fabio Maltoni (Universite Catholique de Louvain (UCL) (BE) and Università di Bologna), Josh McFayden (University of Sussex), Malgorzata Maria Worek (Rheinisch-Westfälische Tech. Hoch. (DE)), Marbo Zaro (Università degli Studi e INFN Milano (IT)), Maria Aldaya Martin (DPSY), Sergio Sanchez Cruz (Universität Zuerich (CH)), Wolfgang Wagner (Bergische Universität Wuppertal (DE))

Registration: You are registered for this event. 93 participants. Check details

Participants: Aman Desai, Amartya Raj, Amitabh Lath, Andrea Helen Knue, Angela Giraldi, Arpan Ghosal, Barbara Alvarez Gonzalez, Brendon Bullard, Carlo Oleari, Carlos Vico Villalba, Carmen Diez Pardos

There are minutes attached to this event. Show them.

13:00 → 13:20	Issues in multilepton final states in ttW production	20m
Speakers: Didar Dobur (Ghent University (BE)), Elizaveta Shabalina (Georg-August-Universität Göttingen (DE))		
ttW_TOPLHCWG.pdf		
13:25 → 13:45	Reference cross-sections and methods used in analyses at ATLAS and CMS for ttW production	20m
Speakers: Clara Ramon Alvarez (Universidad de Oviedo (ES)), Tamara Vazquez Schroeder (CERN)		
120922_ttW_XSand...		
13:50 → 14:10	NLO QCD and EW corrections to off-shell ttW production	20m
Speakers: Giovanni Pelliccioli (Max-Planck-Institut für Physik), Giovanni Pelliccioli (Wurzburg University)		
gp_ttW_09_12_2022...		
14:15 → 14:35	Modelling uncertainties of ttW multilepton signatures	20m
Speaker: Laura Reina (Florida State University (US))		
ttW_modelling.pdf		
14:40 → 15:10	Coffee break	30m
15:10 → 15:30	Improving NLO merging for ttW production	20m
Speaker: Rikkert Frederix (Lund University)		
frederix.pdf		
15:35 → 15:55	NLO multi-jet merging for ttW production including electroweak corrections in Sherpa	20m
Speakers: Enrico Bothmann (University of Göttingen), Enrico Bothmann		
bothmann.pdf		
16:00 → 17:00	Discussion session	1h
M_Worek.pdf		

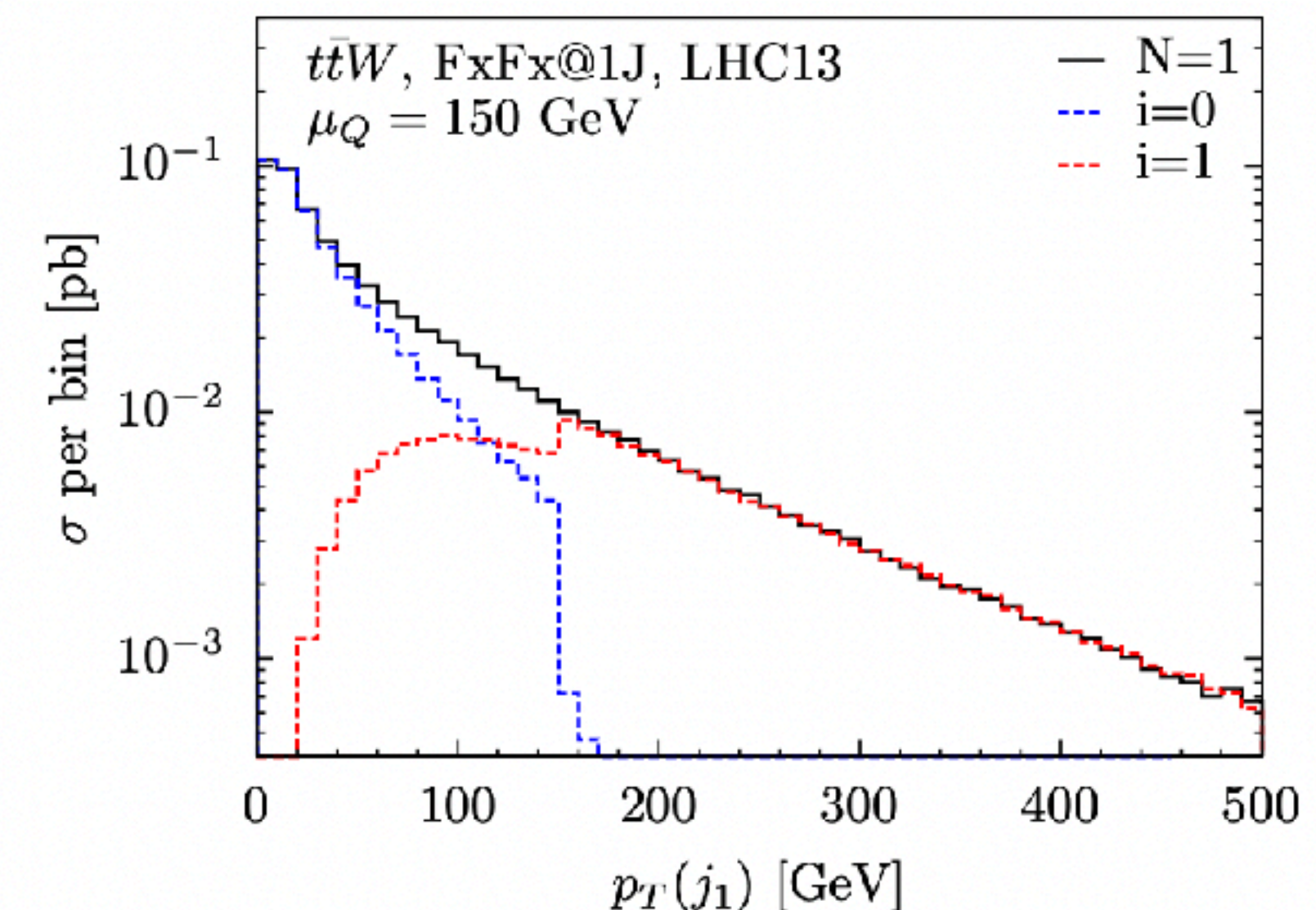
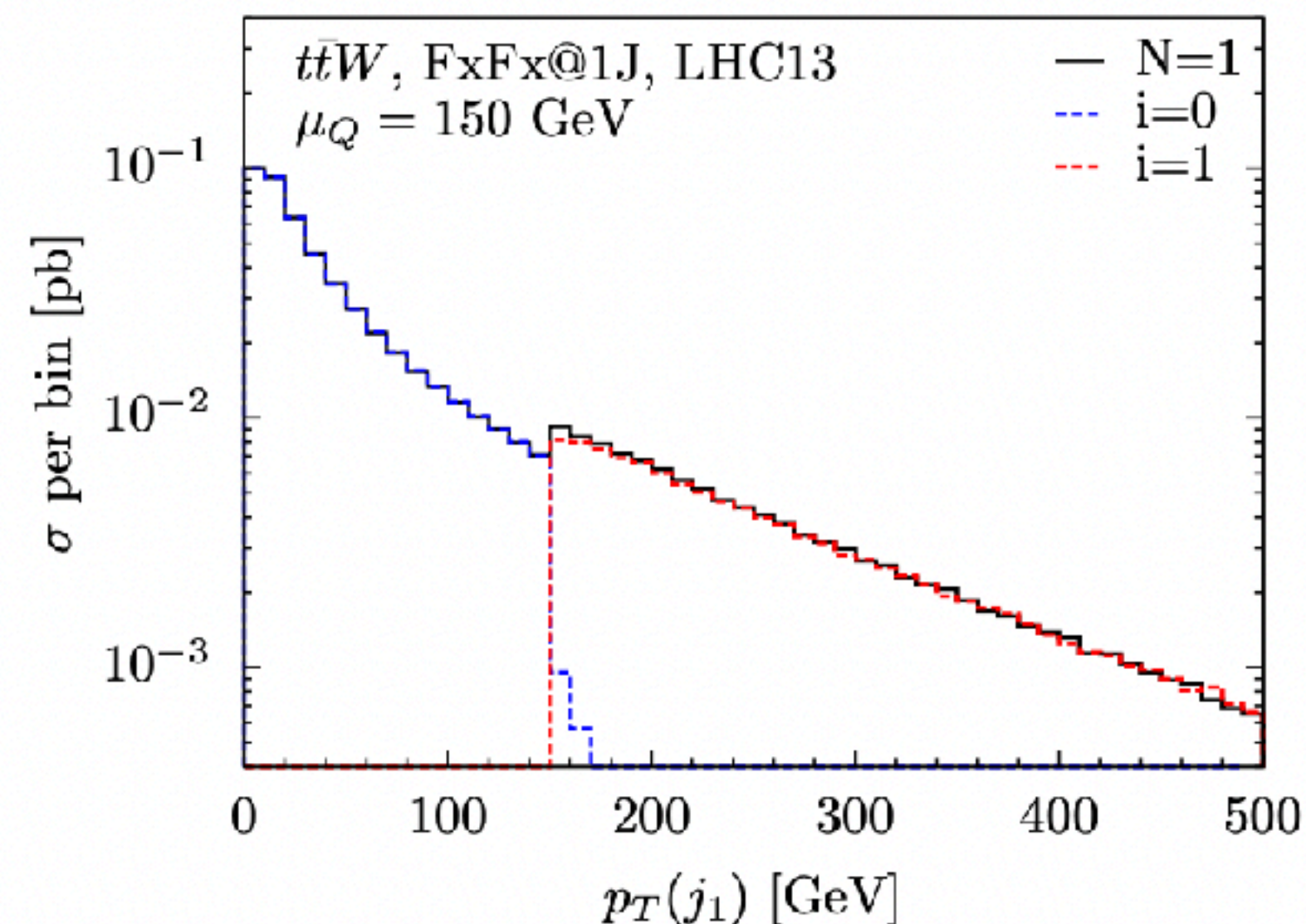
- ▶ NLO QCD corrections to qg-initiated ttWj contributions shown to be large [[1405.0301](#)]
- ▶ ttWjj contribution also potentially large [[2009.00032](#)]
- ▶ **NLO merging very important (especially in the absence of NNLO predictions)**
- ▶ NLO-merged calculation available with MG5_aMC+Py8 FxFx and Sherpa

▶ **ttW+0j@NLO**

→ **ttW+0,1j@NLO**

“k-factor” = 1.11

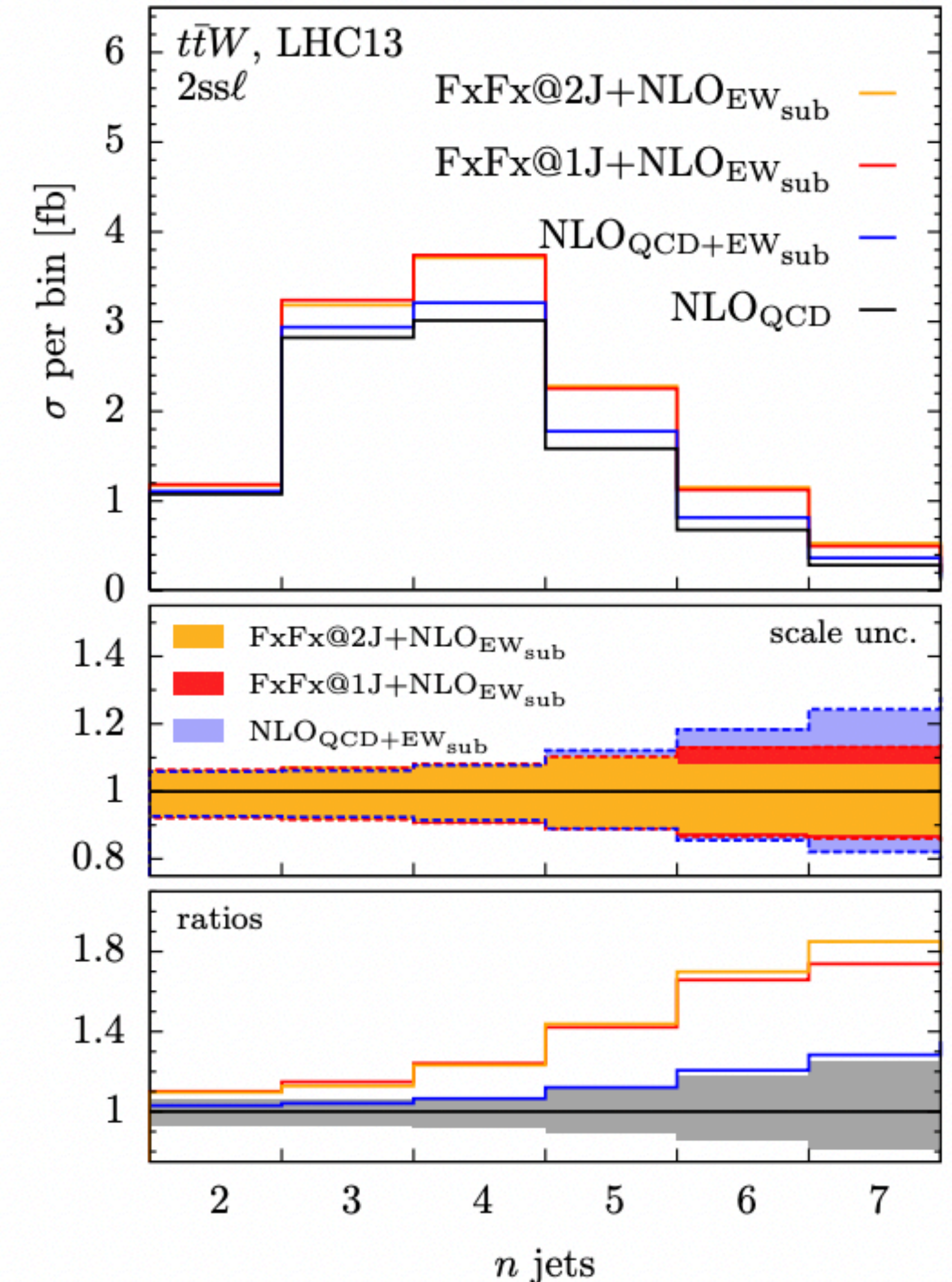
- ▶ But easy to get wrong...
- ▶ Care needed with “weak jet” merging [[2108.07826](#)]





- ▶ NLO + EW corrections from Powheg, MG5_aMC & Sherpa [[2101.11808](#)]
- ▶ Multi-leg merged 0-2j@NLO + EW corrections from Frederix et al. [[2108.07826](#)]

Order (default scale)	$\sigma \pm \text{scale} \pm \text{PDF}$ [fb]
FxFx@2J	$691.1(8)^{+65.7(+9.5\%)}_{-74.1(-10.7\%)} +7.3(+1.1\%)_{-7.3(-1.1\%)}$
FxFx@2J+NLO _{EW} ^{sub}	$738.8(8)^{+75.0(+10.1\%)}_{-81.3(-11.0\%)} +7.5(+1.0\%)_{-7.5(-1.0\%)}$
FxFx@2J+NLO _{EW} ^{lead} +NLO _{EW} ^{sub}	$722.4(8)^{+70.2(+9.7\%)}_{-77.7(-10.8\%)} +7.2(+1.0\%)_{-7.2(-1.0\%)}$



- ▶ New ttW reference cross-section now agreed between ATLAS and CMS (from Joint LHC Higgs—LHC Top WGs meeting)
 - ▶ Frederix et al. prediction of 722 fb (2jNLO+EW-lead+EW-sub)
- ▶ No agreement in how to treat EW and QCD components individually
 - ▶ Proposal today on how to deal with this for Sherpa and MG5_aMC+Py8
 - ▶ If we agree this should be adopted across Top WG and ATLAS
 - ▶ Can propagate k-factors to AMI/TopDataPrep
- ▶ The interplay between samples that have their own EWK corrections (Sherpa and Powheg) and those that don't is rather complicated
 - ▶ For ATLAS-wide recommendations we should perhaps try to be really specific about this on TopProcesses page?
- ▶ Thoughts?!



- ▶ Agree to adopt the total cross-section predicted by MadGraph5_aMC@NLO_FxFx_improved as a reference cross-section:
 $\sigma = 722 +70 / -78$ (scale) ± 7 (PDF) fb. Reference: JHEP 11 (2021) 29.
- ▶ Rikkert will work on quantifying the effect of the electroweak-jet treatment (improvement to FxFx)
- ▶ Enrico will work towards including the quark-gluon diagrams in Sherpa.
- ▶ Timescale for both theory items is the beginning of March 2023. We plan to have a next joint meeting to discuss these results and comparisons between improved FxFx and Sherpa.
- ▶ Experimentalists from ATLAS and CMS will provide a Rivet routine defining a common fiducial phase. Time line: before Christmas 2022.
- ▶ Plan to use joint-session format to continue work on ttW modeling and modeling uncertainties.
- ▶ Plan to establish smaller task force groups to deal with various aspects of the ttW modeling and modeling uncertainties (to be discussed during the meeting in March 2023 or offline).
- ▶ We think that regular communication and cooperation which facilitates a close working relationship between theoretical physicists and experimental physicists working for the LHC Top and Higgs working groups is essential if we want to profit from the current state-of-the-art theoretical predictions and theoretical advances in general for ttW.
- ▶ Setups for the particular ttW measurements should be communicated to the theory groups at the early stage of measurements. The latter included: SM input parameters, scale settings, PDFs, cuts in the fiducial phase space regions, choice of the observables, binning, details of the MC simulations used for unfolding and comparisons to the LHC data etc.



LHC Higgs WG ttH/tH | ttbb+ttW note

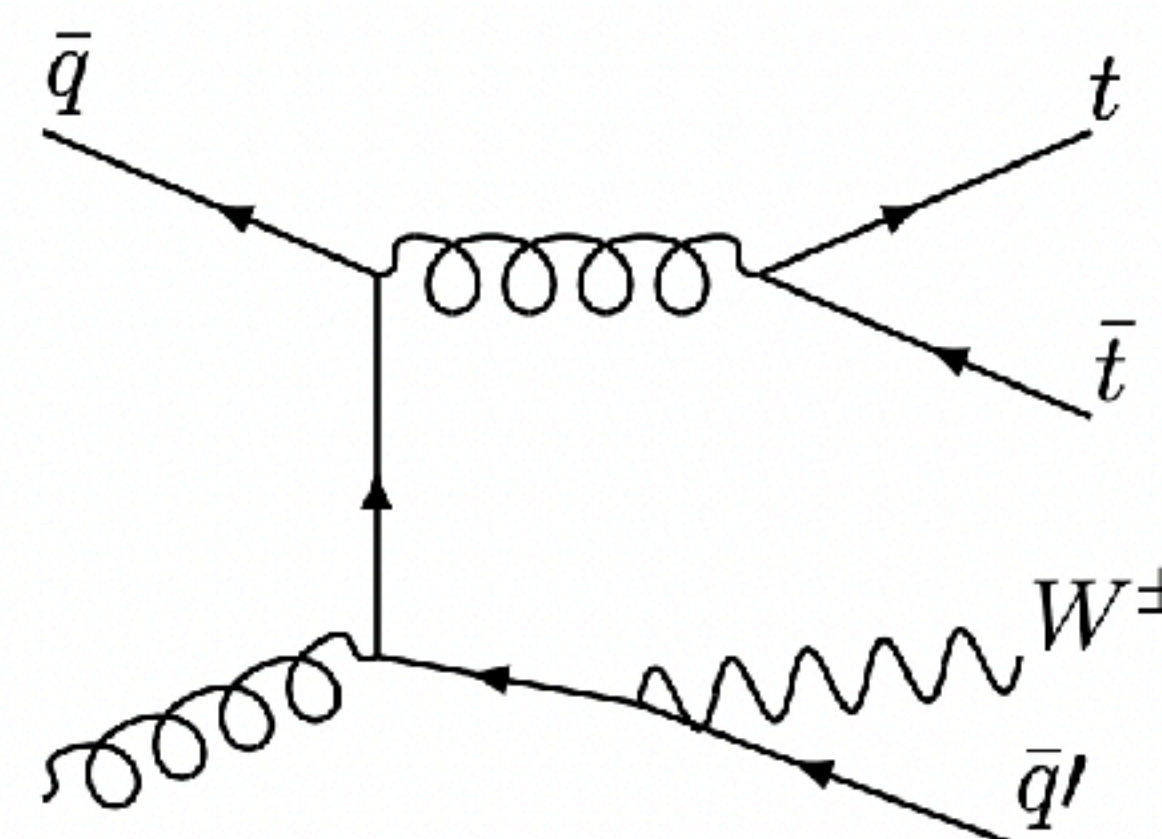
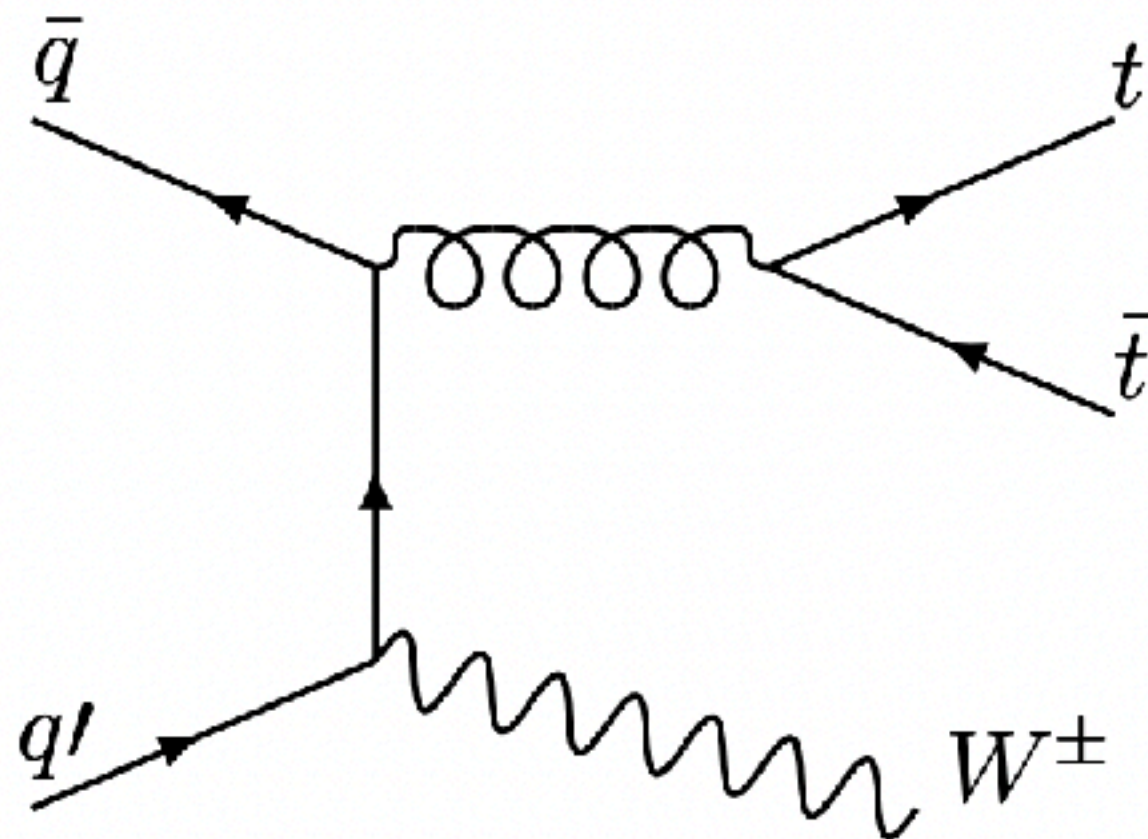


Label	ATLAS Sherpa 2.2.10	ATLAS Sherpa 2.2.10 QCD+EW	ATLAS MG5_aMC+Py8 FxFx	ATLAS MG5_aMC+Py8	CMS MG5_aMC+Py8 FxFx
Process	$t\bar{t}W$ inclusive	$t\bar{t}W$ inclusive	$t\bar{t}W$ inclusive	$t\bar{t}W$ inclusive	$t\bar{t}l\nu$ ($t\bar{t}W$ inclusive)
Generator	SHERPA 2.2.10 [27]	SHERPA 2.2.10 [27]	MG5_AMC@NLO 2.9.3 [69]	MG5_AMC@NLO 2.3.3 [70]	MG5_AMC@NLO 2.4.2
order of QCD ME	0,1 j @NLO ^a	0,1 j @NLO ^a	0,1 j @NLO	NLO	0,1 j @NLO
ME or core scale	$\mu_R = \mu_F = H_T/2$	$\mu_R = \mu_F = H_T/2$	dynamic scale choice [24, 67, 68]	$\mu_R = \mu_F = H_T/2$	dynamic scale choice [24, 67, 68]
order of EW corr.	-	$\alpha^3, \alpha^2\alpha_s^2, \alpha^3\alpha_s$	-	-	-
Parton Shower	SHERPA 2.2.10	SHERPA 2.2.10	PYTHIA 8.245 [8]	PYTHIA 8.210 [8]	PYTHIA 8.226
Merging Scheme	MEPs@NLO [62]	MEPs@NLO [62]	FxFx [24]	-	FxFx
Merging Scale	30 GeV	30 GeV	30 GeV	-	42 GeV
PDF	NNPDF3.0 NNLO [71]	NNPDF3.0 NNLO	NNPDF3.0 NLO	NNPDF3.0 NLO	NNPDF3.1 NLO [72]
Tune	SHERPA default	SHERPA default	A14 [33]	A14	CP5 [34]
Cross section ^b	597 fb	615 fb	613 fb	548 fb	220 fb (666 fb ^c)

^aIn addition to the implicit $2j$ @LO contribution from the real emission part of the $1j$ @NLO calculation, Sherpa adds the $2j$ @LO as an explicit separate process within the merging such that the ME is supplemented with higher-order improvements such as the CKKW scale choice and Sudakov factors.”

^b $\sigma_{tot}=600.8$ fb from YR4 is used for all samples in the generator comparisons in section 3.3.2 except for SHERPA QCD+EW

^ccalculated from $t\bar{t}l\nu$ as $0.2198 \times (1/ (3 \times 0.11))$



- ▶ qq-initiated process

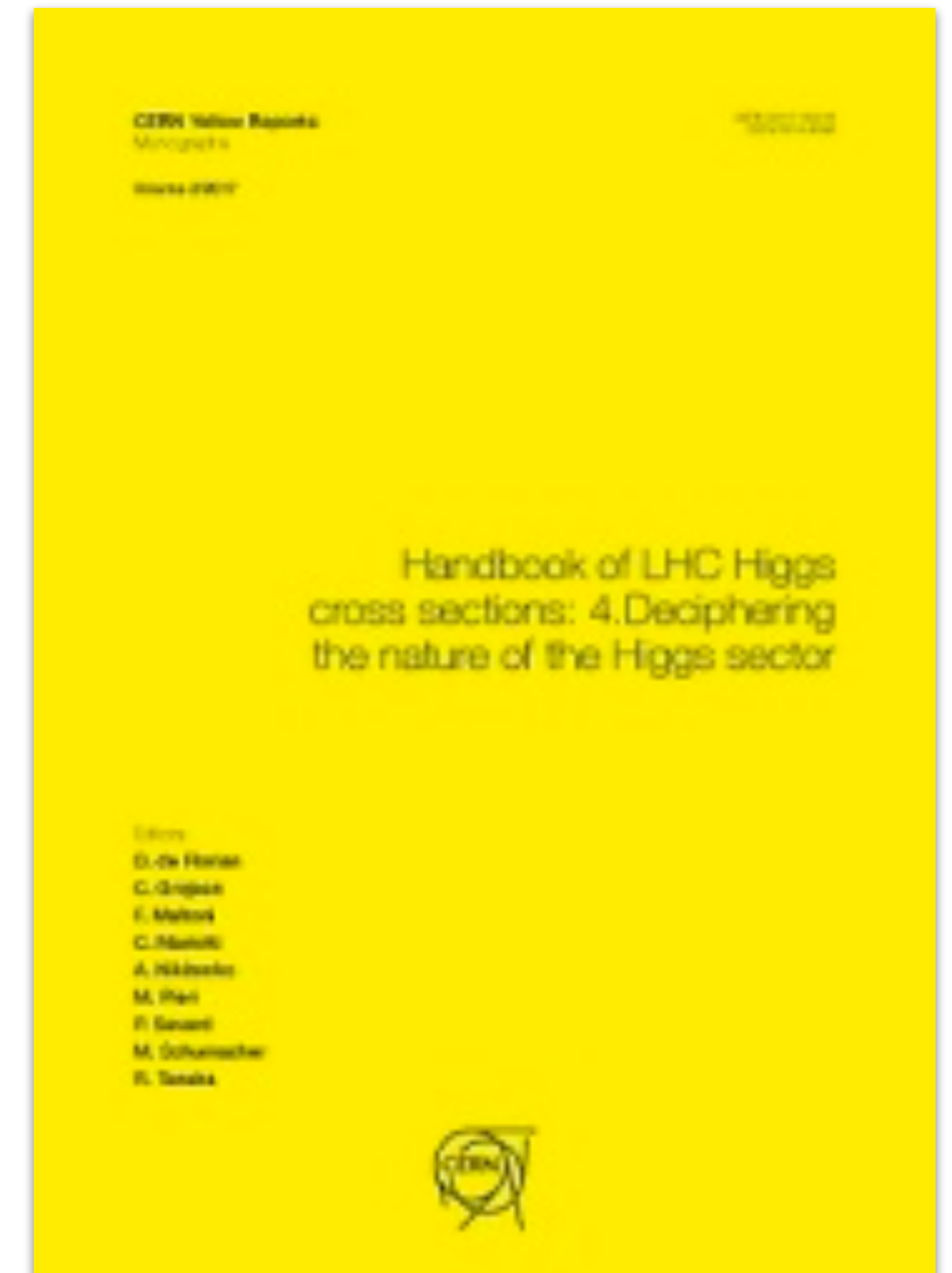
- ▶ Spin correlations are very important as W emission polarises incoming quark [[1406.3262](#)].

- ▶ First NLO calculation shows large ~30-50% k-factor [[1204.5678](#)]

- ▶ Large fraction coming from opening qg-initiated $ttWj$ production mode

First recommendation...

- ▶ NLO+PS predictions become available [[1208.2665,1507.05640](#)]
- ▶ Augmented with “leading” NLO EW corrections ($\alpha^3 + \alpha_s^2\alpha^2$) [[1504.03446](#)]
- ▶ Led to best “**agreed upon**” calculation:
Yellow Report 4 [[1610.07922](#)]
- ▶ Has been the recommended value for a number of years
- ▶ **But misses some crucial inputs now available...**
 - ▶ NLO QCD merging
 - ▶ “Subleading” NLO EW corrections
 - ▶ NNLL resummation
 - ▶ Off-shell effects



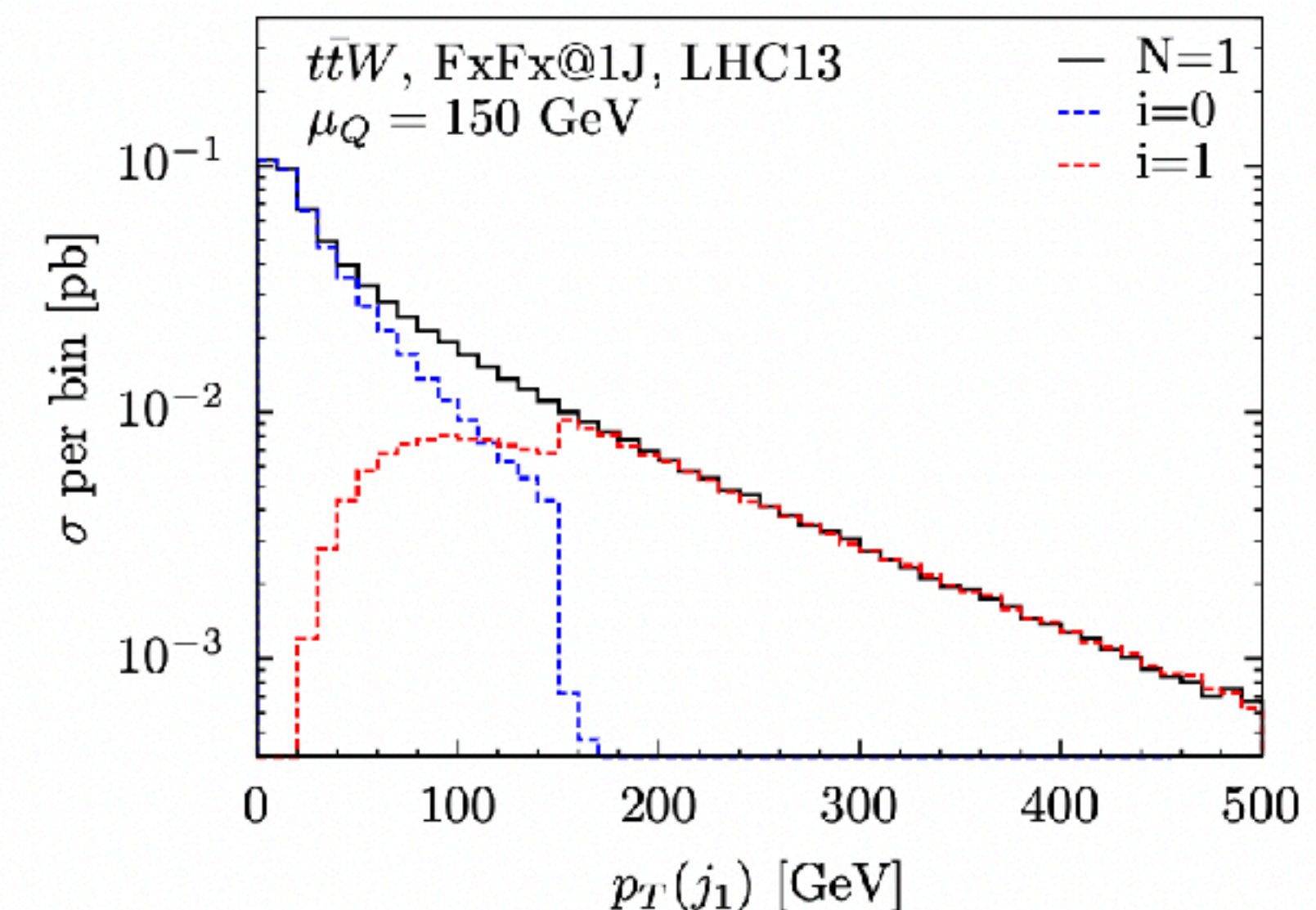
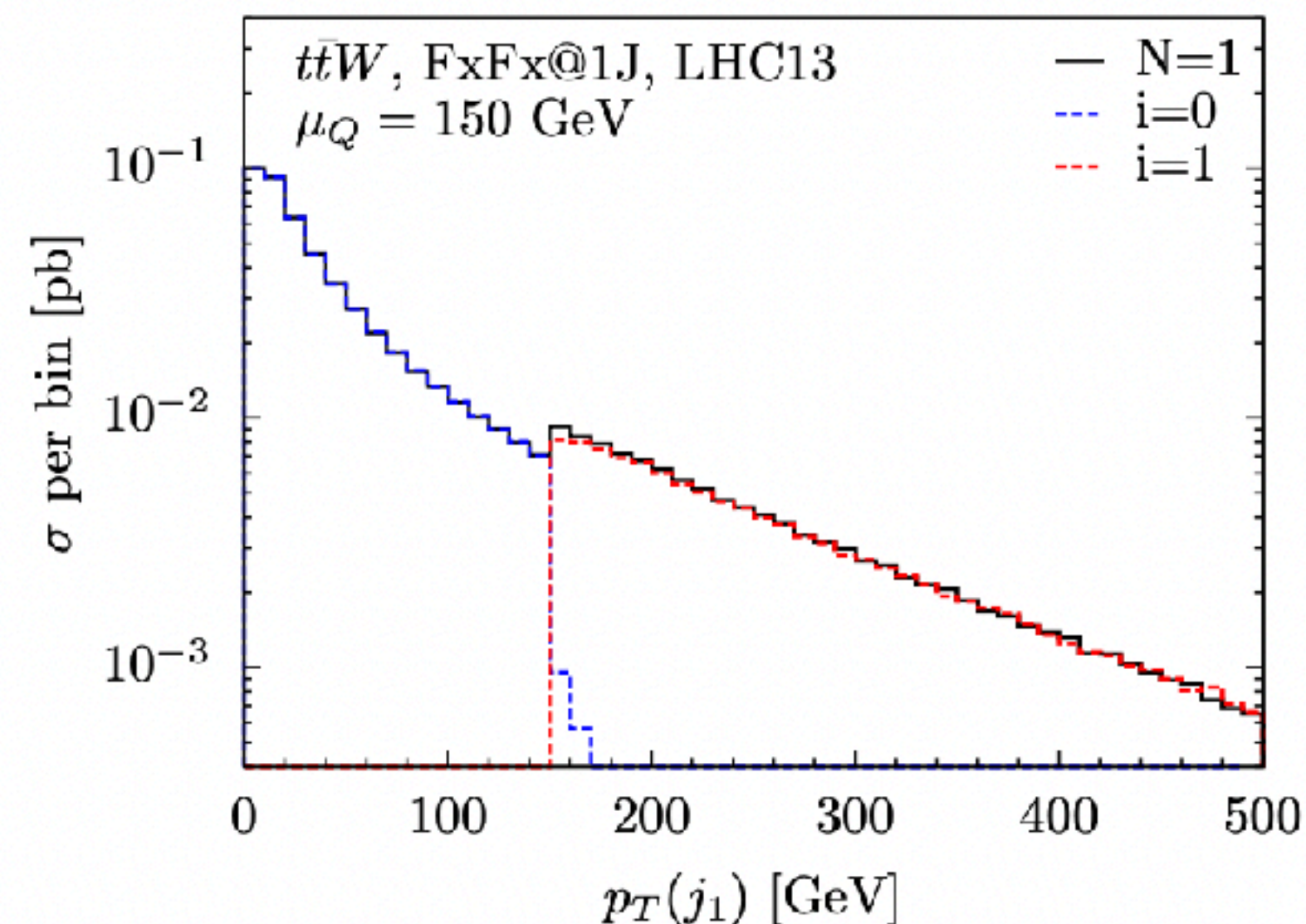
- ▶ NLO QCD corrections to qg-initiated $ttWj$ contributions shown to be large [[1405.0301](#)]
- ▶ $ttWjj$ contribution also potentially large [[2009.00032](#)]
- ▶ **NLO merging very important (especially in the absence of NNLO predictions)**
- ▶ NLO-merged calculation available with MG5_aMC+Py8 FxFx and Sherpa

▶ $ttW+0j@NLO$

→ $ttW+0,1j@NLO$

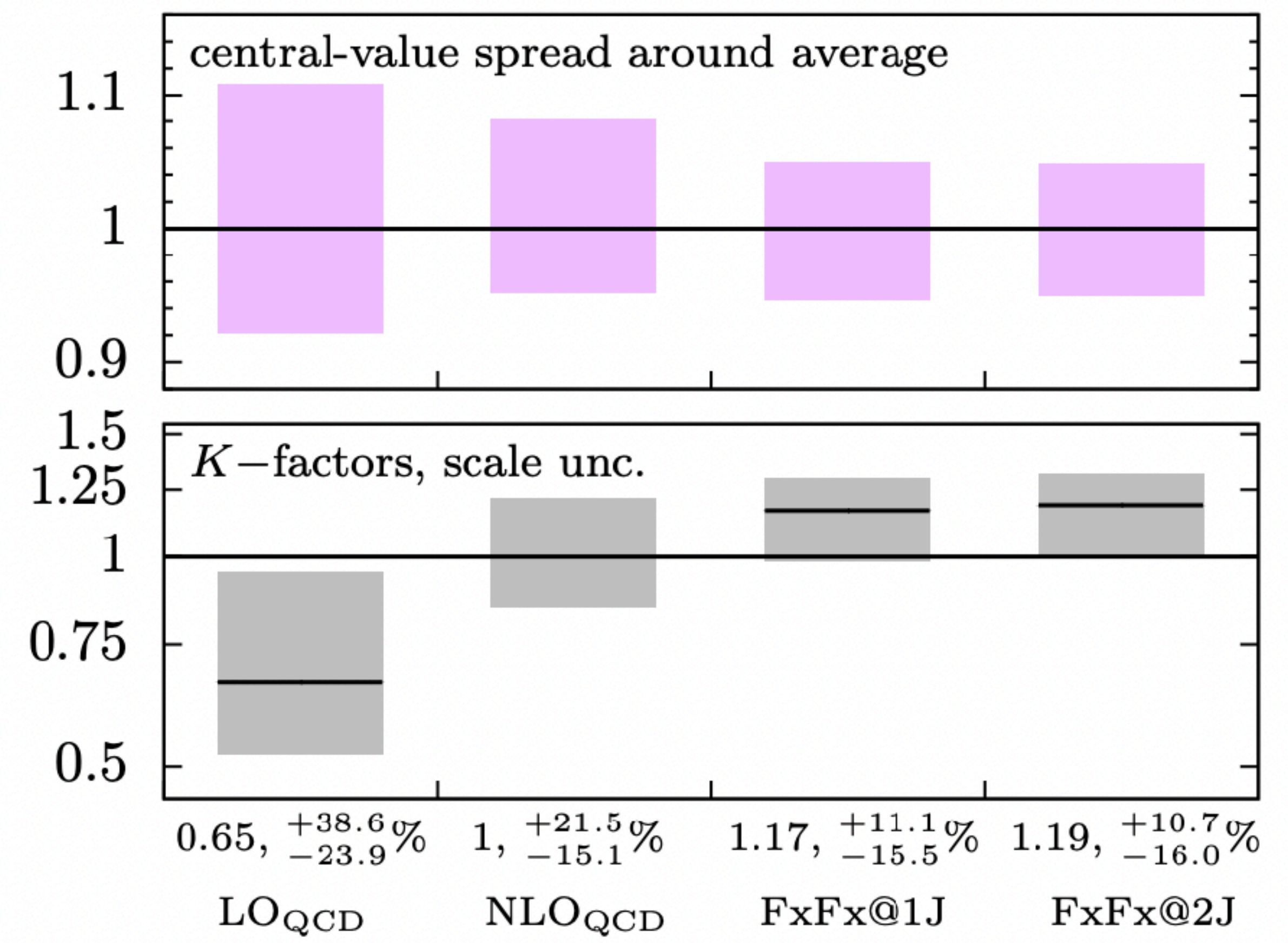
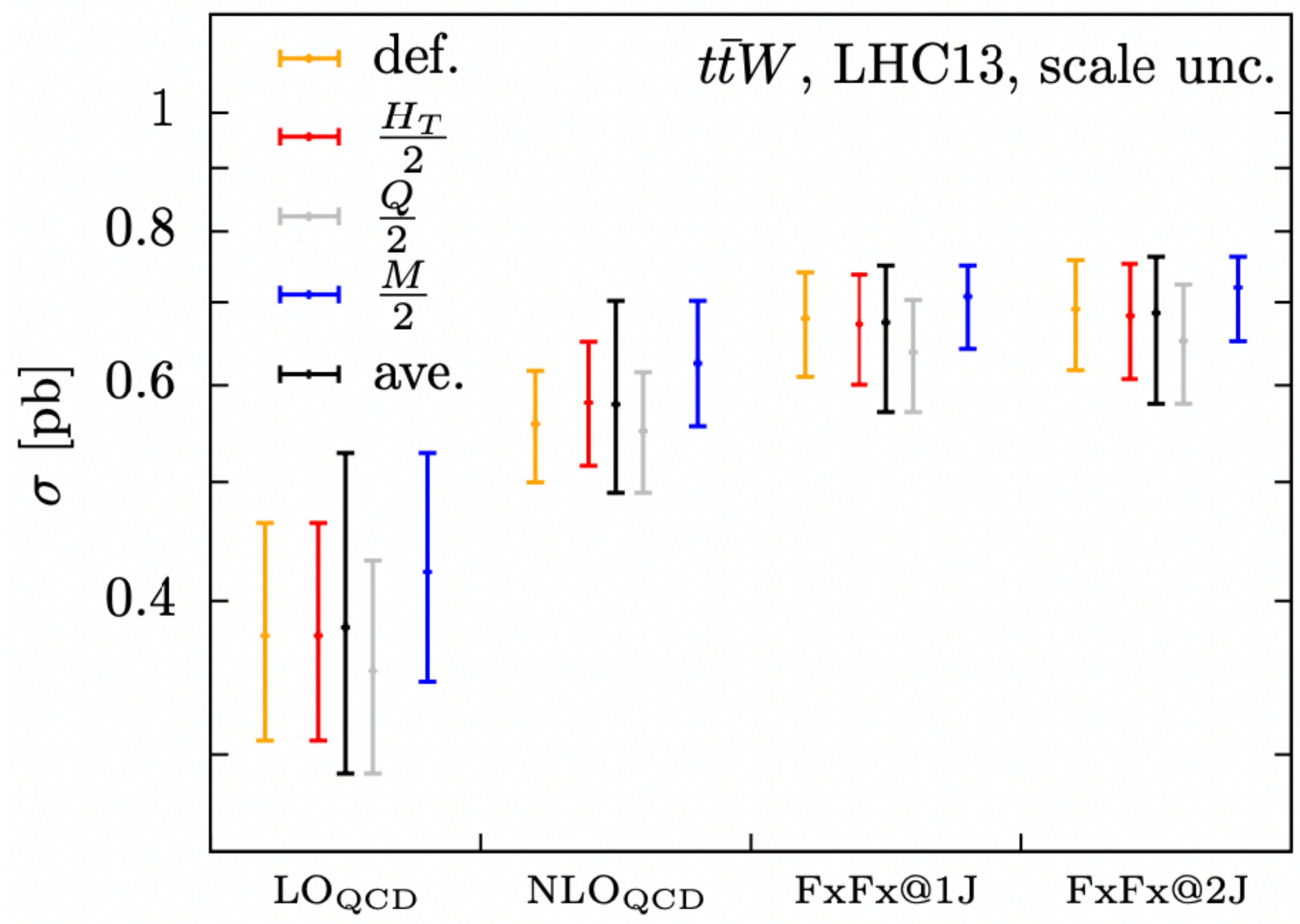
“k-factor” = 1.11

- ▶ But easy to get wrong...
- ▶ Care needed with “weak jet” merging [[2108.07826](#)]



NLO merging

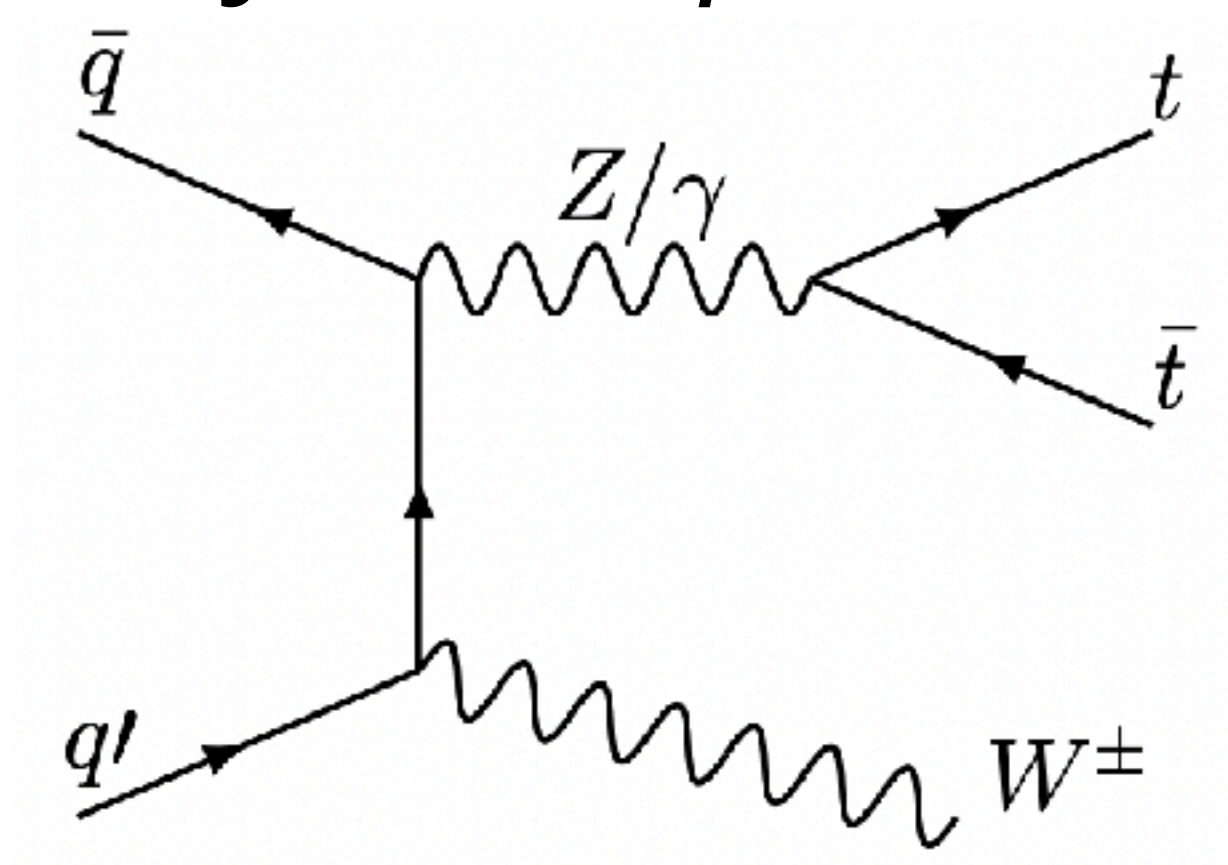
► Multi-leg merged $t\bar{t}W+0-2j@NLO$ from Frederix et al. [[2108.07826](#)]





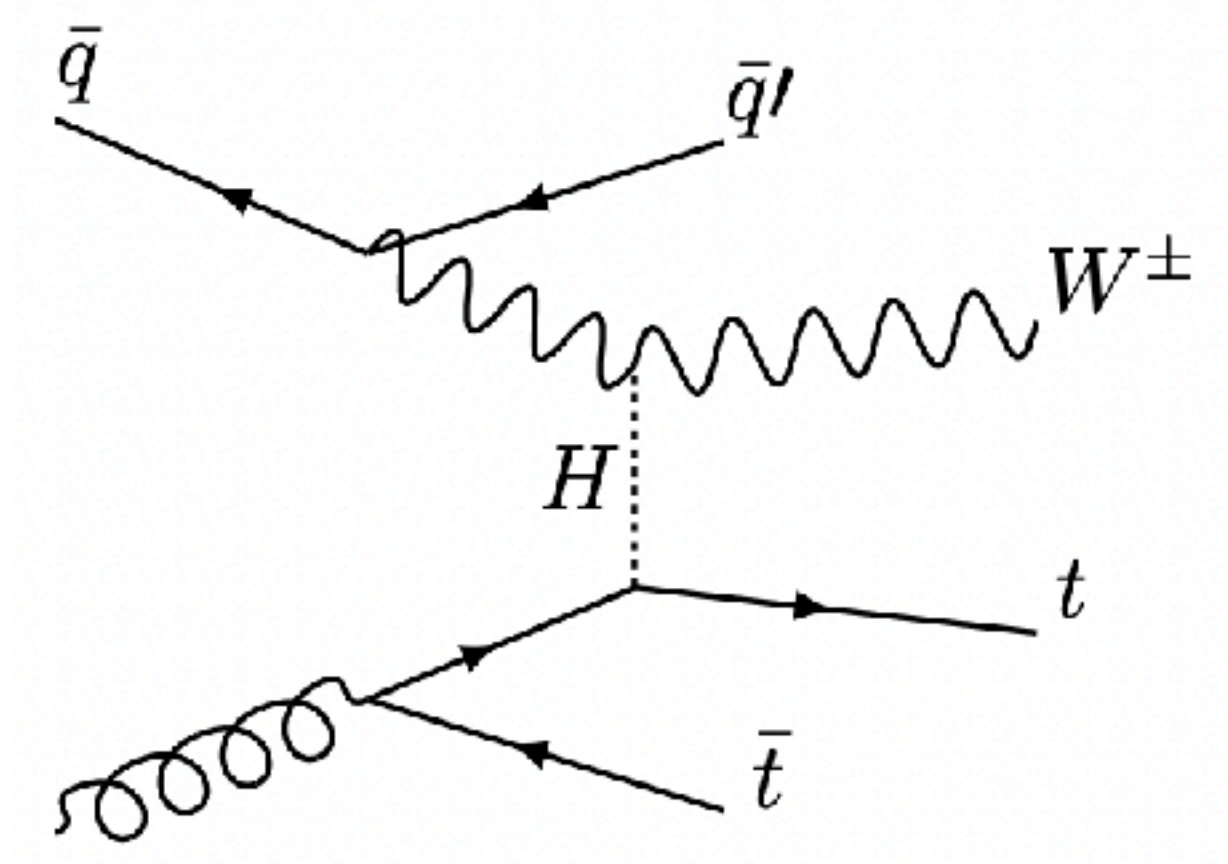
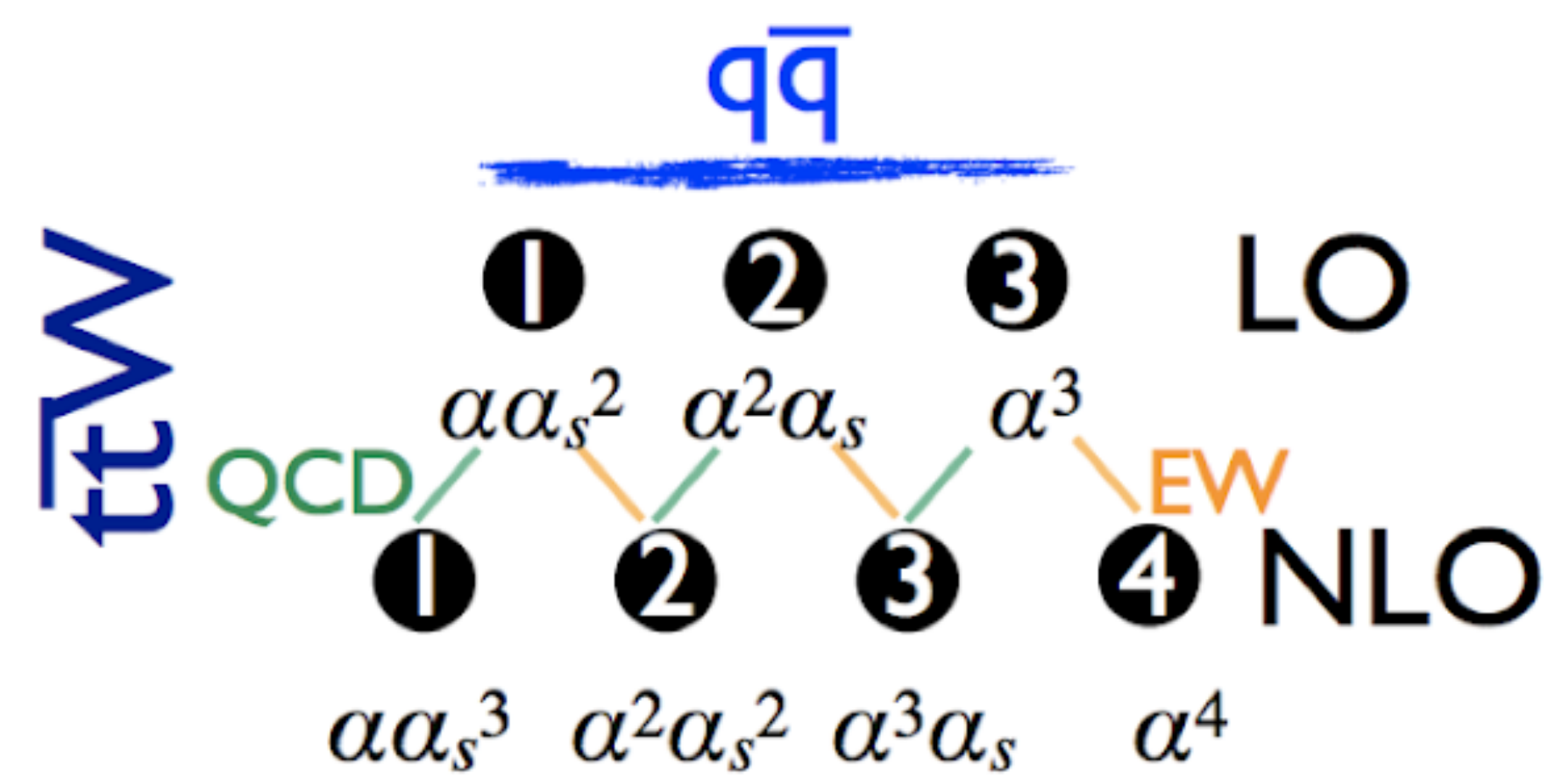
EWK corrections

- ▶ YR4: "Subleading NLO terms of $O(\alpha_s \alpha^3)$... are not included as they are expected to be strongly suppressed."
- ▶ [1711.02116] shows that these "subleading" NLO EWK corrections are larger than "leading" corrections!
- ▶ Resulting inclusive correction is 1.09



13 TeV

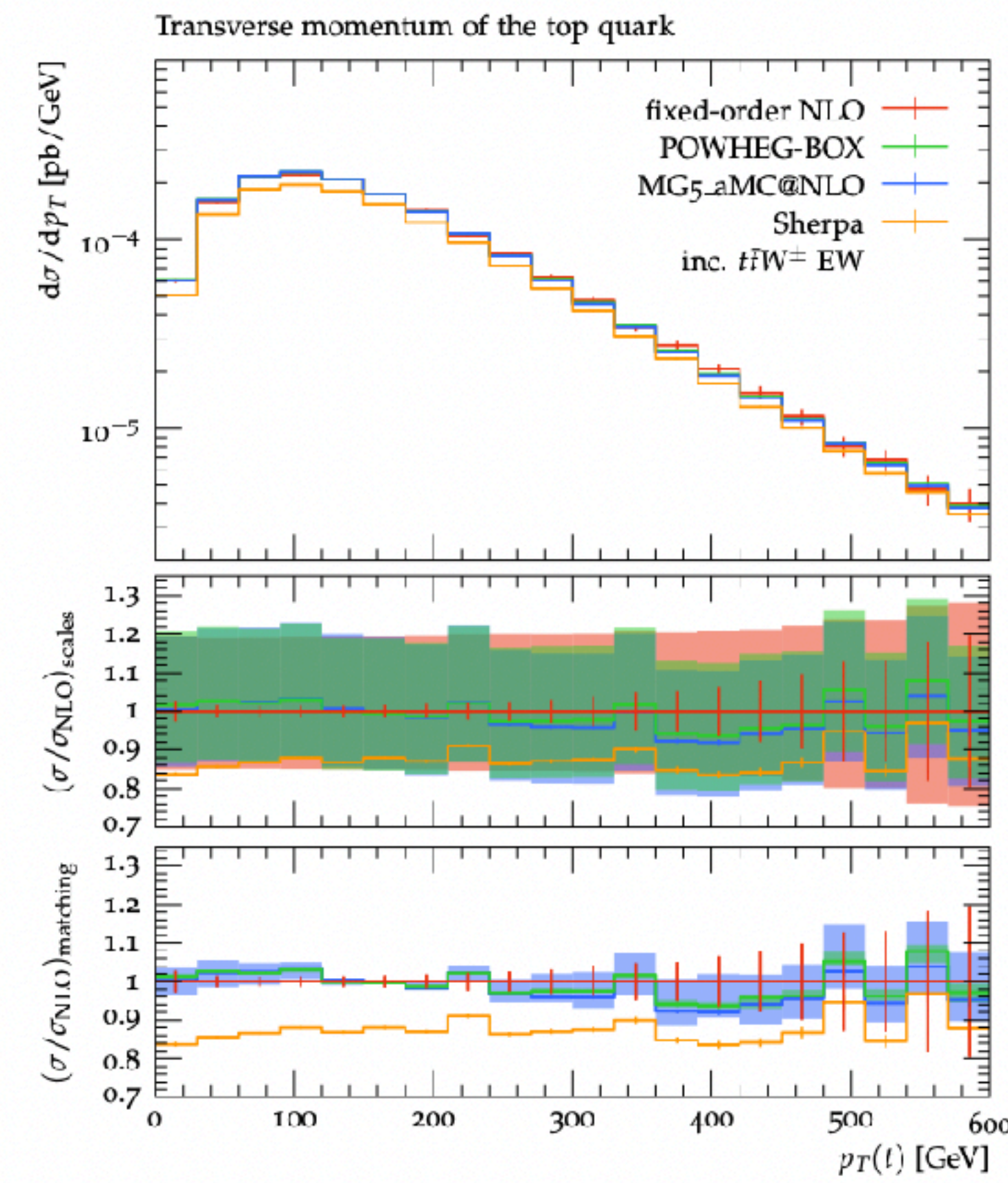
$\delta[\%]$	$\mu = H_T/2$
LO ₂	-
LO ₃	0.9
NLO ₁	50.0 (25.7)
NLO ₂	-4.2 (-4.6)
NLO ₃	12.2 (9.1)
NLO ₄	0.04 (-0.02)





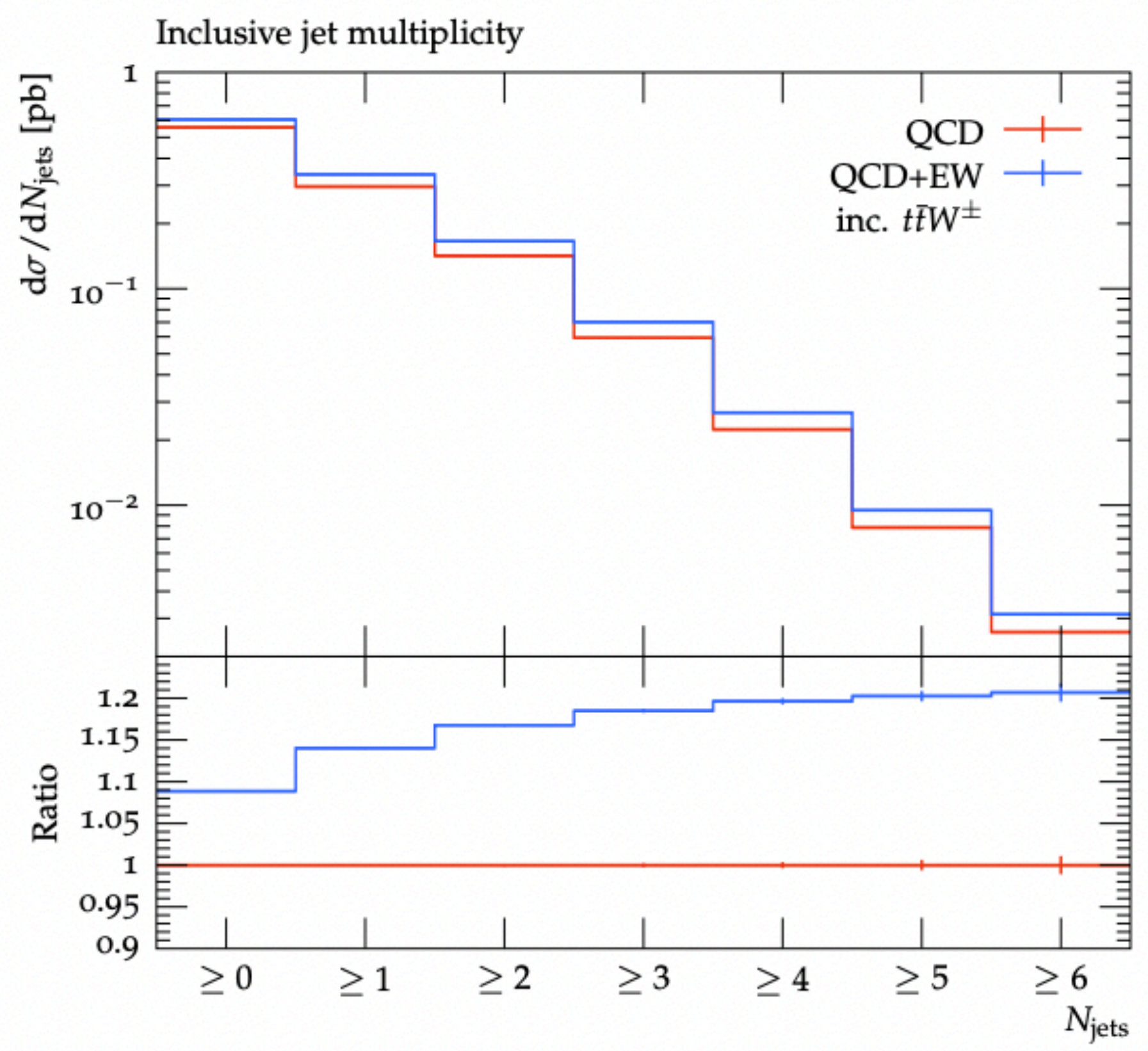
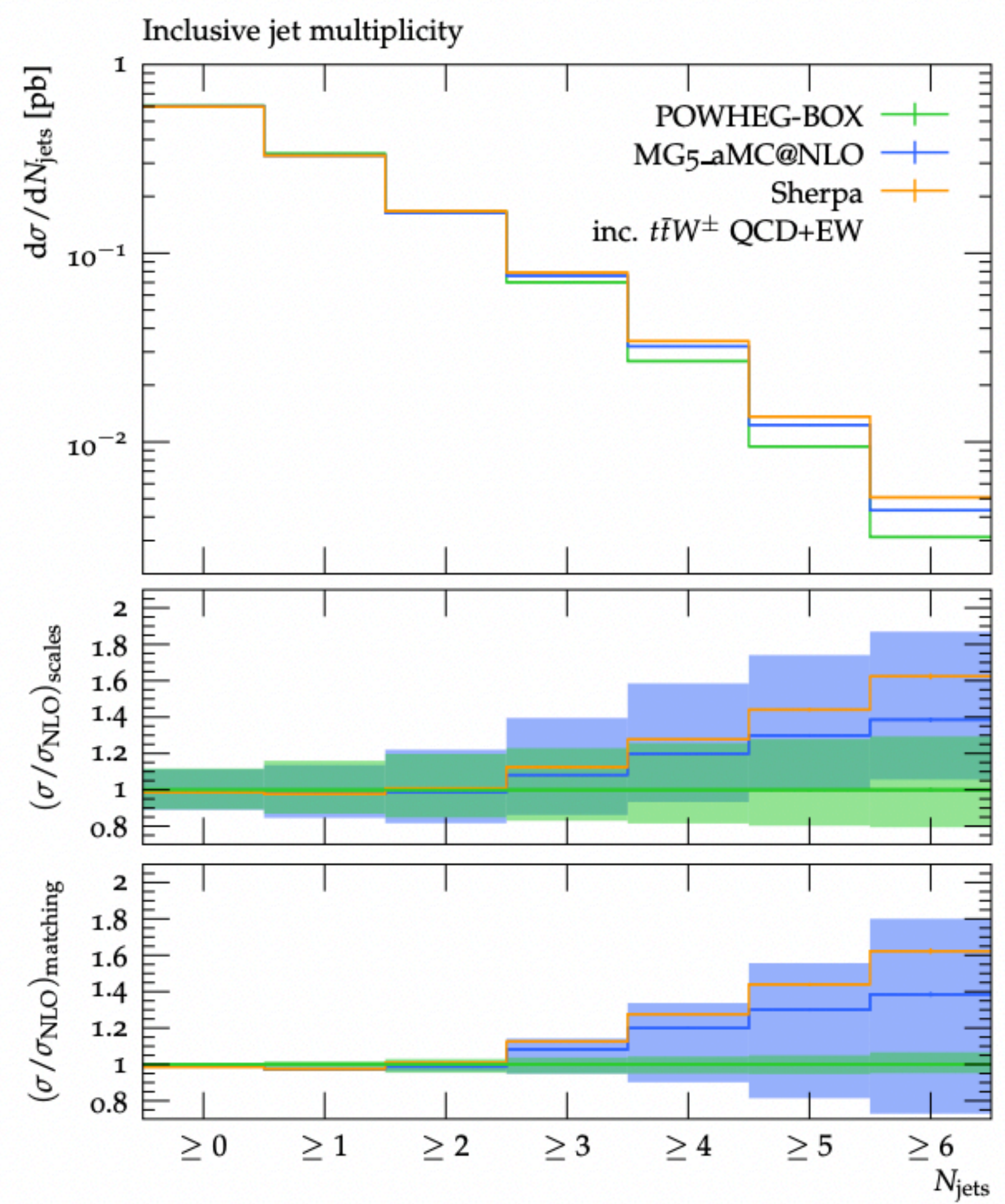
EWK corrections

- ▶ YR4: “Subleading NLO terms of $O(\alpha_s \alpha^3)$... are not included as they are expected to be strongly suppressed.”
- ▶ [1711.02116] shows that these “subleading” NLO EWK corrections are larger than “leading” corrections!
- ▶ Resulting inclusive correction is 1.09
- ▶ In practise not trivial to produce
 - ▶ Sherpa event weights vs dedicated LO sample differ ($\sim 2x$)
 - ▶ Sherpa and MG LO sample cross sections differ by $\sim 25\%$.
 - ▶ Powheg similar to MG [2101.11808]
- ▶ Includes very phenomenologically interesting tW -scattering vertex!



Putting it together

► NLO + EW corrections from Powheg, MG5_aMC & Sherpa [2101.11808]

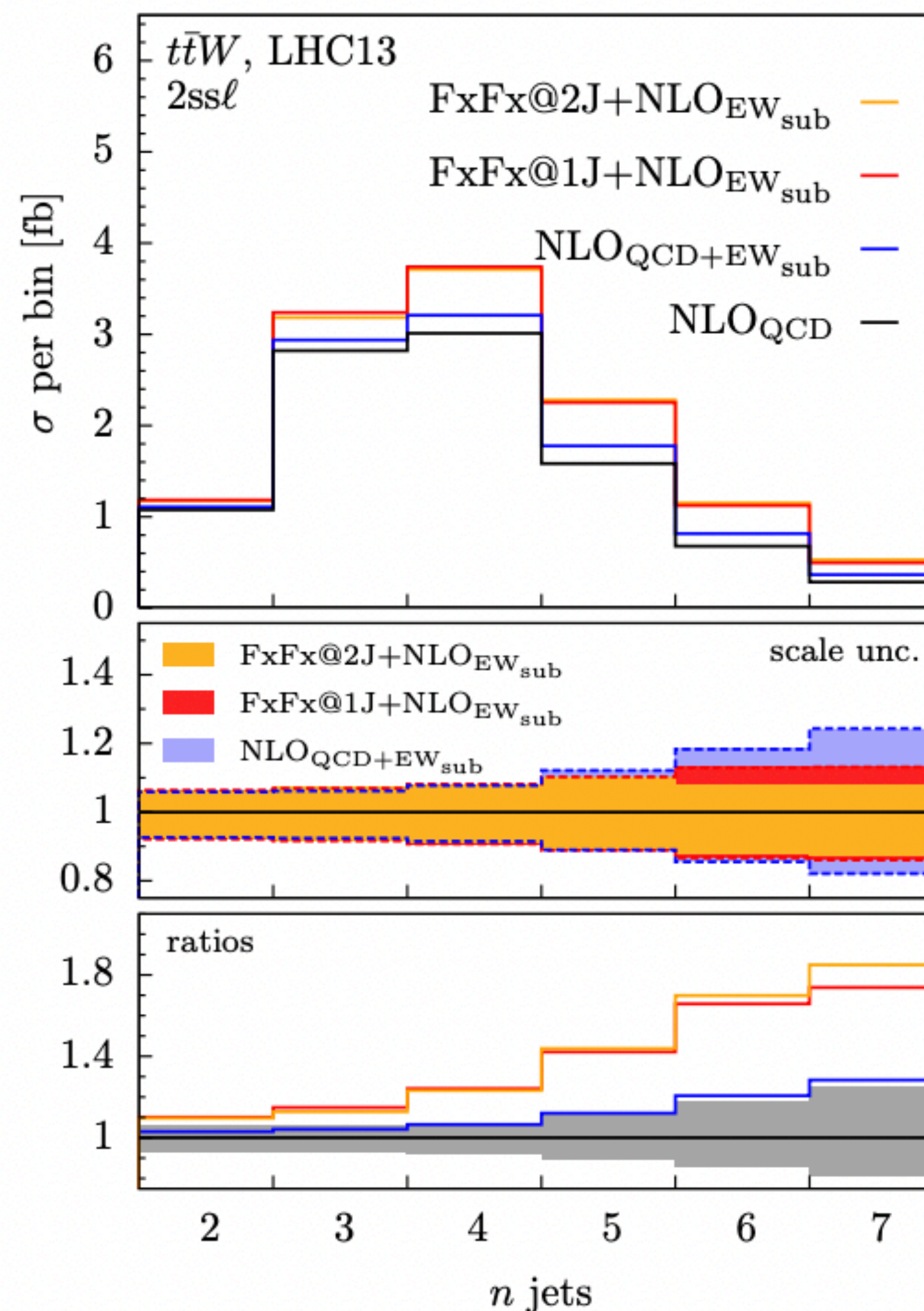


Putting it together

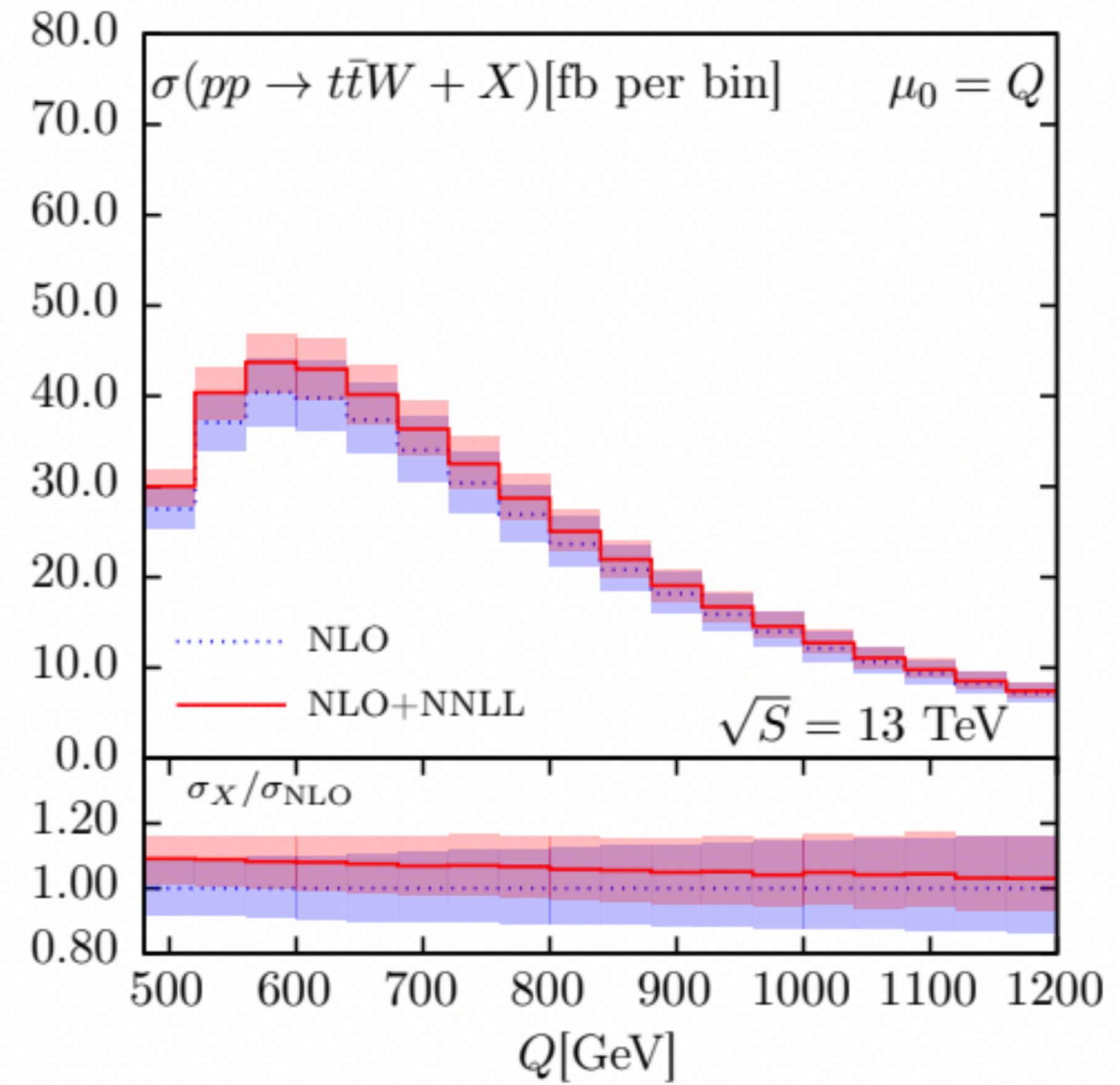
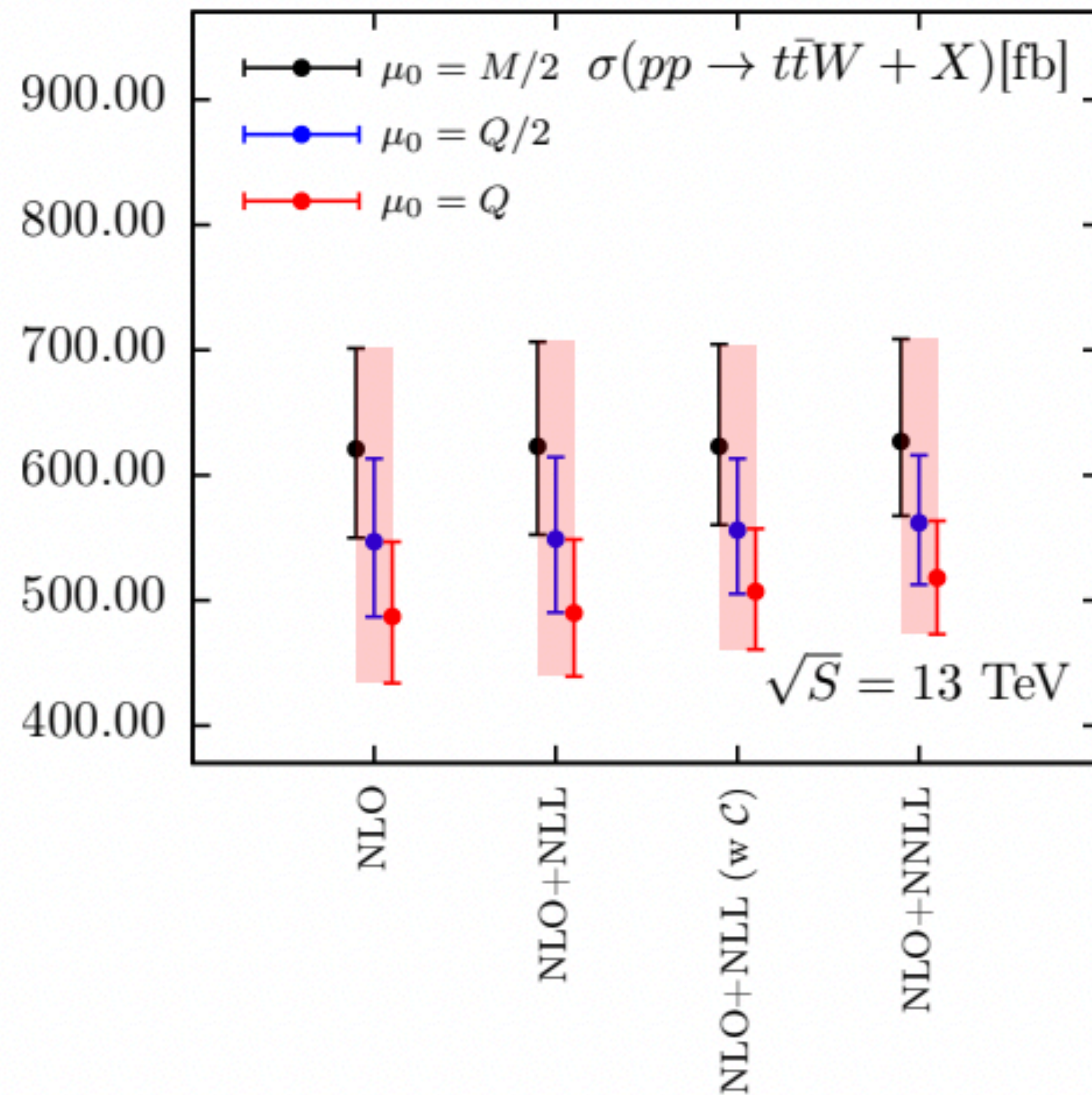
► NLO + EW corrections from Powheg, MG5_aMC & Sherpa [[2101.11808](#)]

► Multi-leg merged 0-2j@NLO + EW corrections from Frederix et al. [[2108.07826](#)]

Order (default scale)	$\sigma \pm \text{scale} \pm \text{PDF}$ [fb]
FxFx@2J	$691.1(8)^{+65.7(+9.5\%)}_{-74.1(-10.7\%)} +7.3(+1.1\%)_{-7.3(-1.1\%)}$
FxFx@2J+NLO _{EW} ^{sub}	$738.8(8)^{+75.0(+10.1\%)}_{-81.3(-11.0\%)} +7.5(+1.0\%)_{-7.5(-1.0\%)}$
FxFx@2J+NLO _{EW} ^{lead} +NLO _{EW} ^{sub}	$722.4(8)^{+70.2(+9.7\%)}_{-77.7(-10.8\%)} +7.2(+1.0\%)_{-7.2(-1.0\%)}$

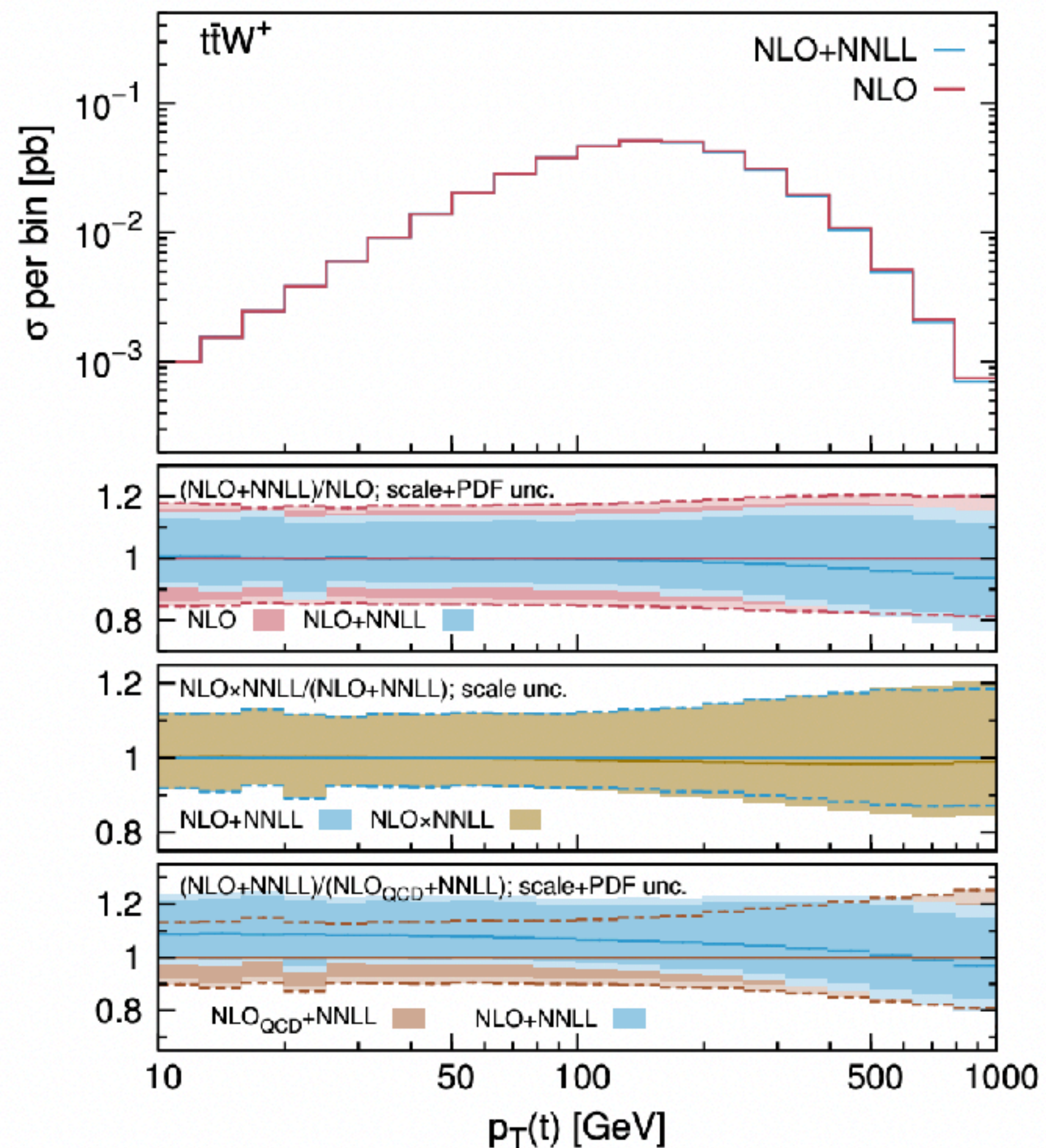
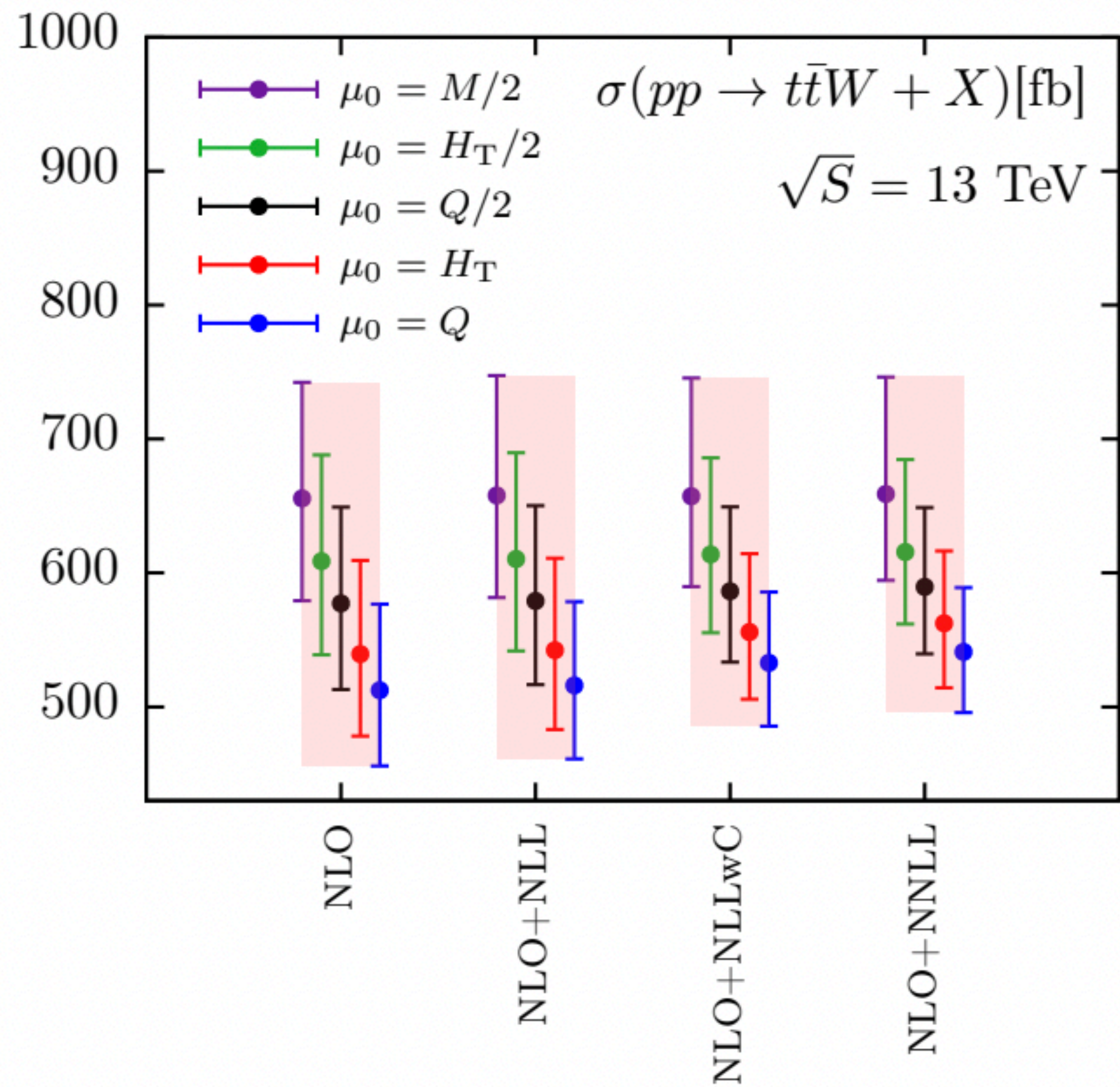


► First calculations at NLO+NNLL become available [[1812.08622](#)]

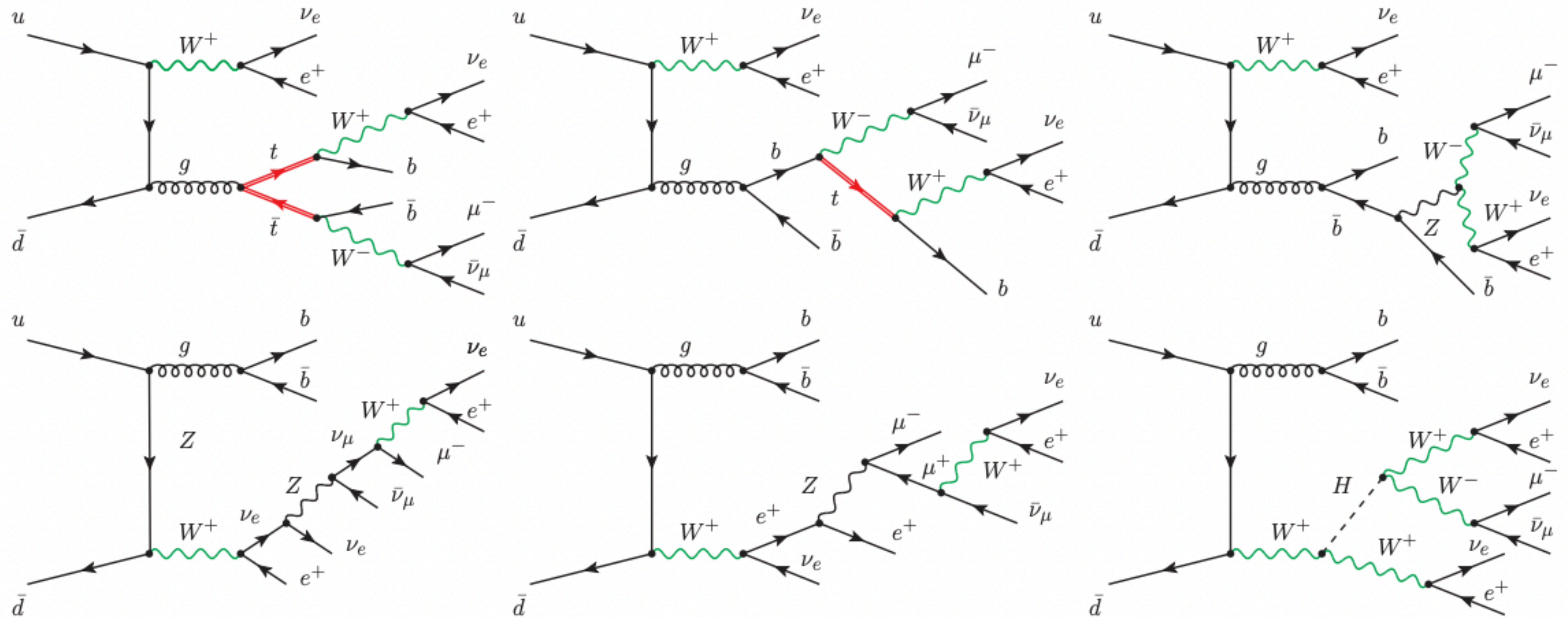


► First calculations at NLO+NNLL become available [[1812.08622](#)]

► NLO EW corrections are added on top [[1907.04343](#), [2001.03031](#)]



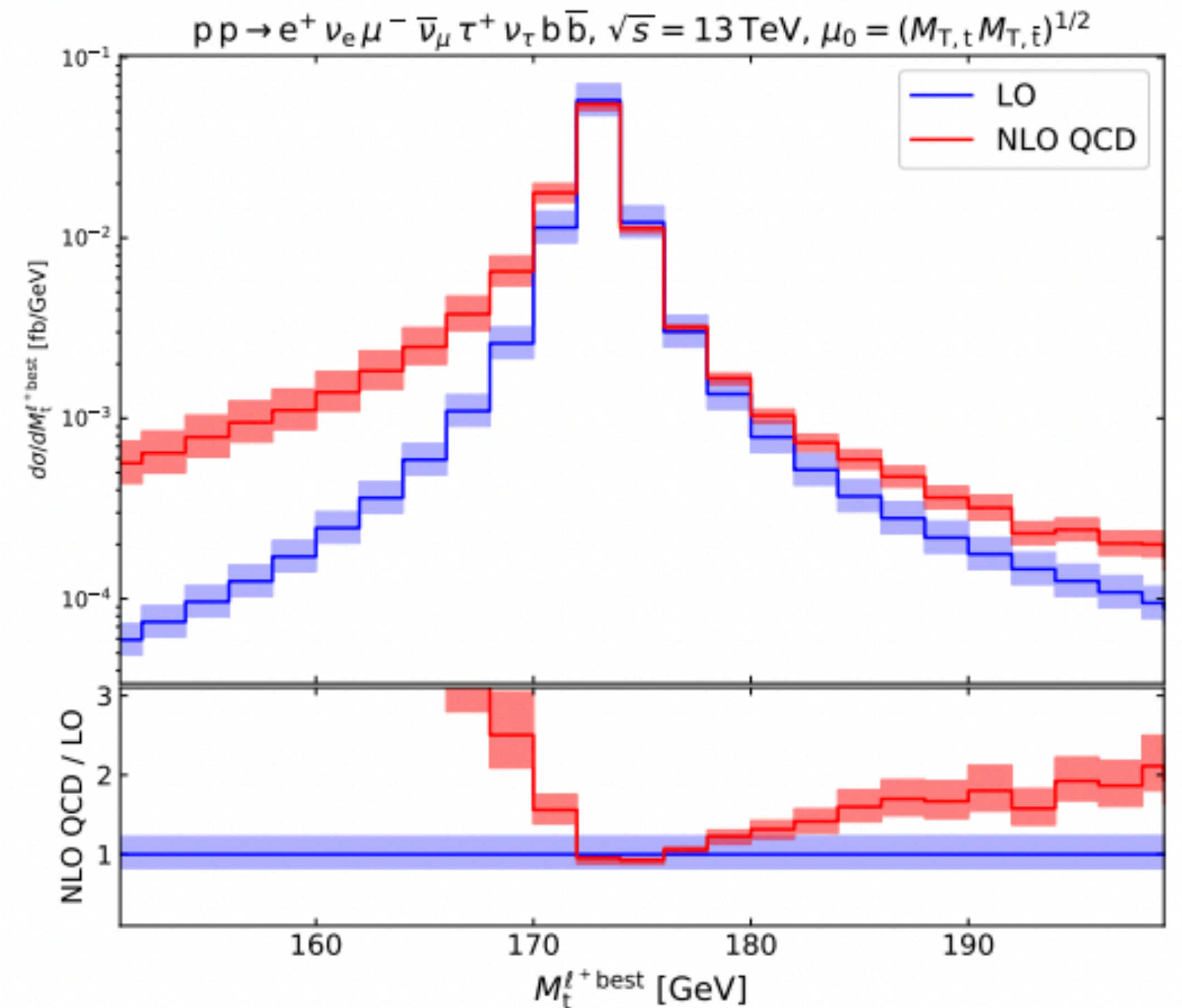
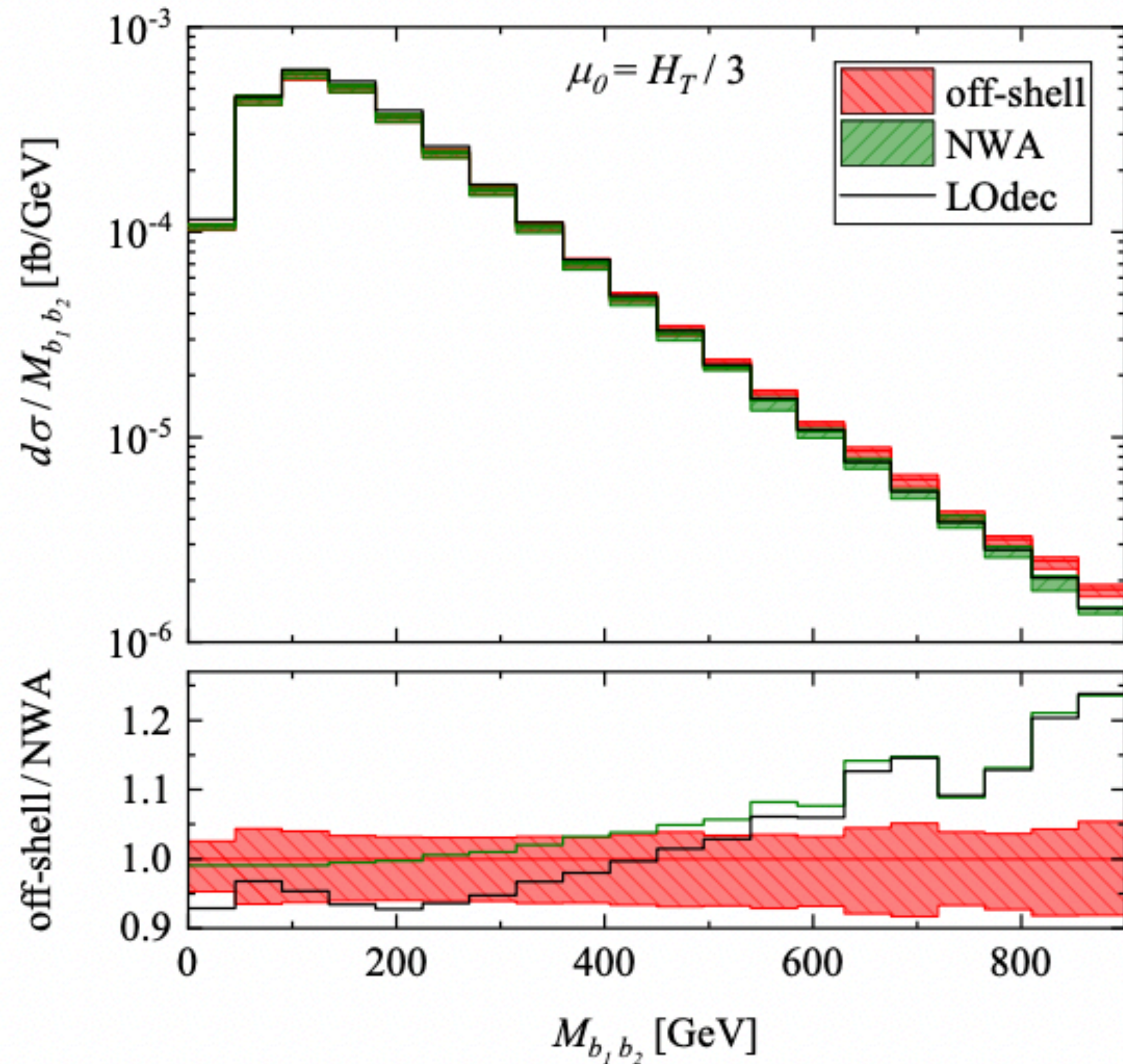
- ▶ NLO off-shell calculations also become available
[2005.09427, 2007.12089, 2012.01363]





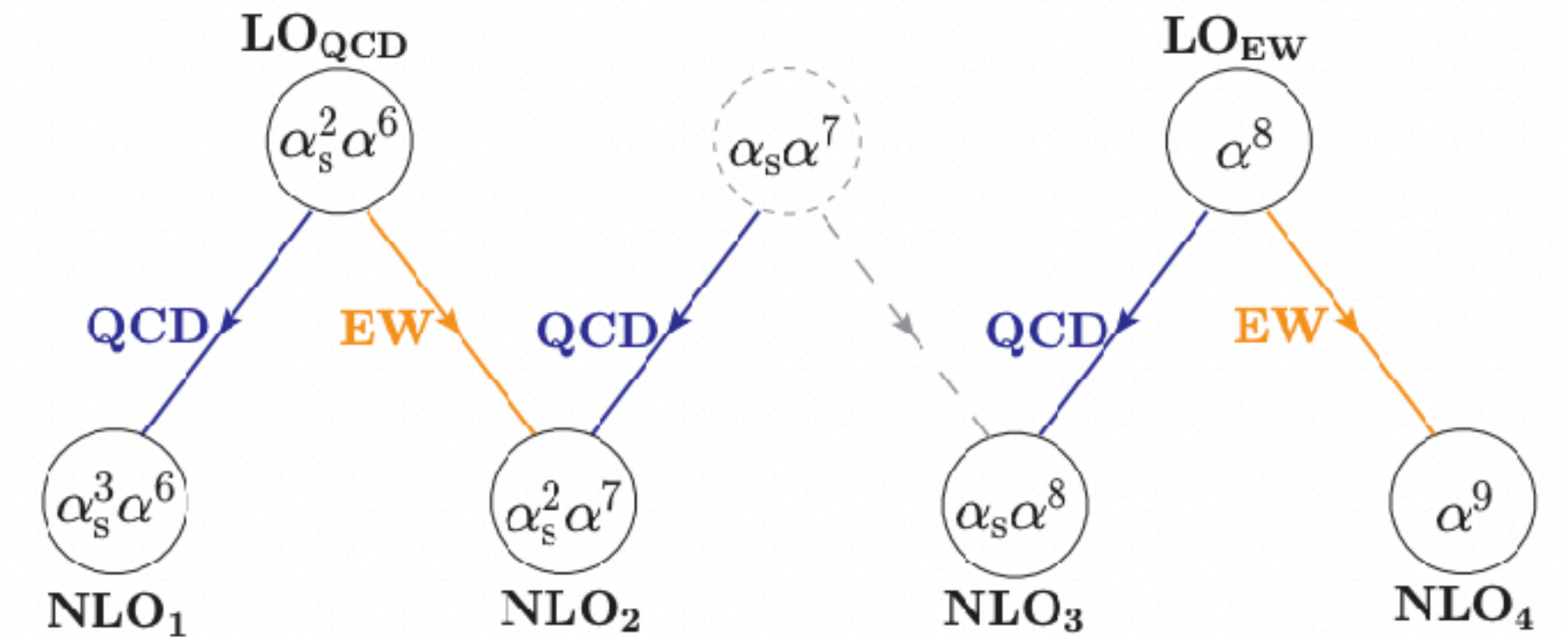
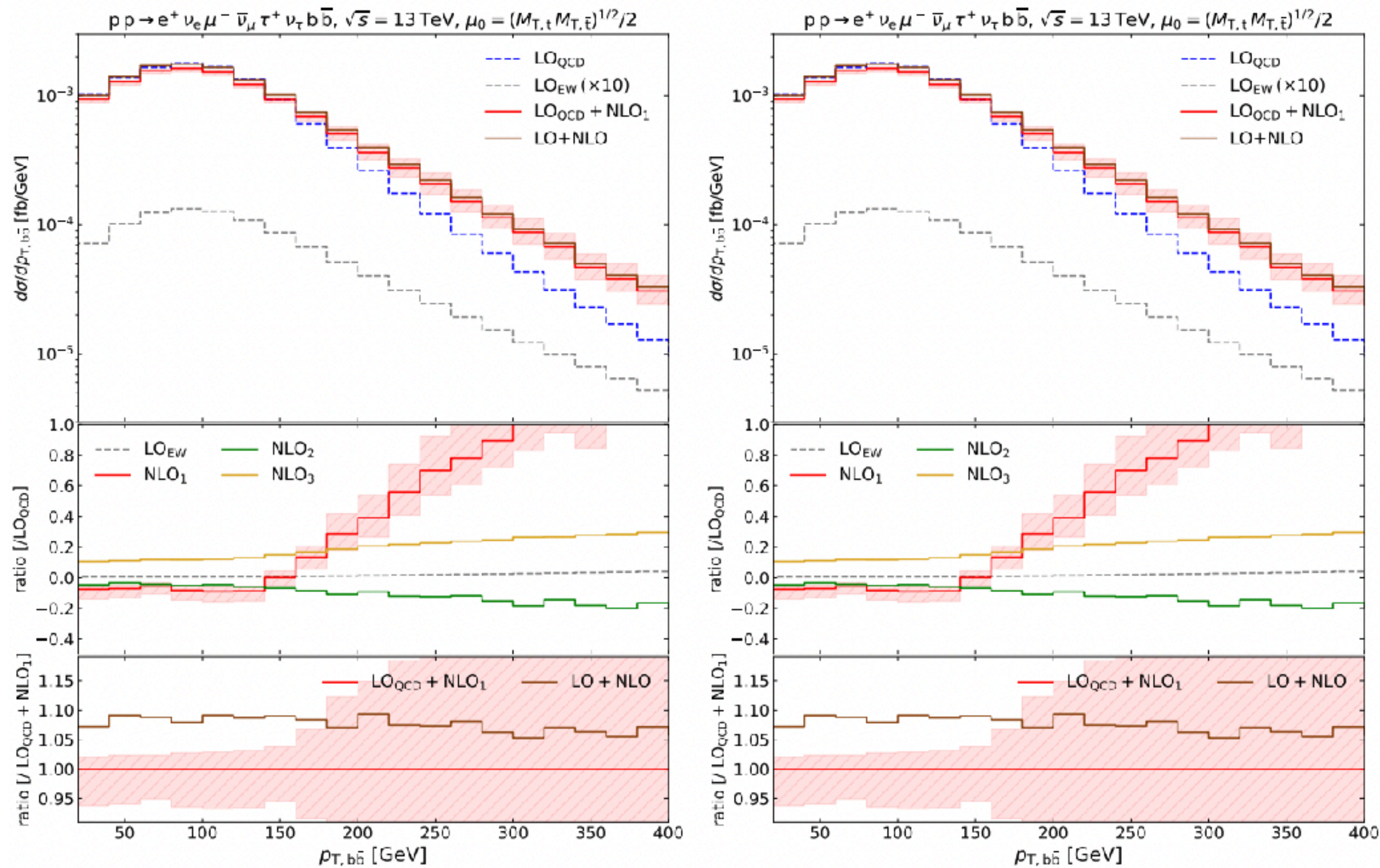
Off-shell

- ▶ NLO off-shell calculations also become available
[2005.09427, 2007.12089, 2012.01363]



► NLO off-shell calculations also become available
 [[2005.09427](#),[2007.12089](#),[2012.01363](#)]

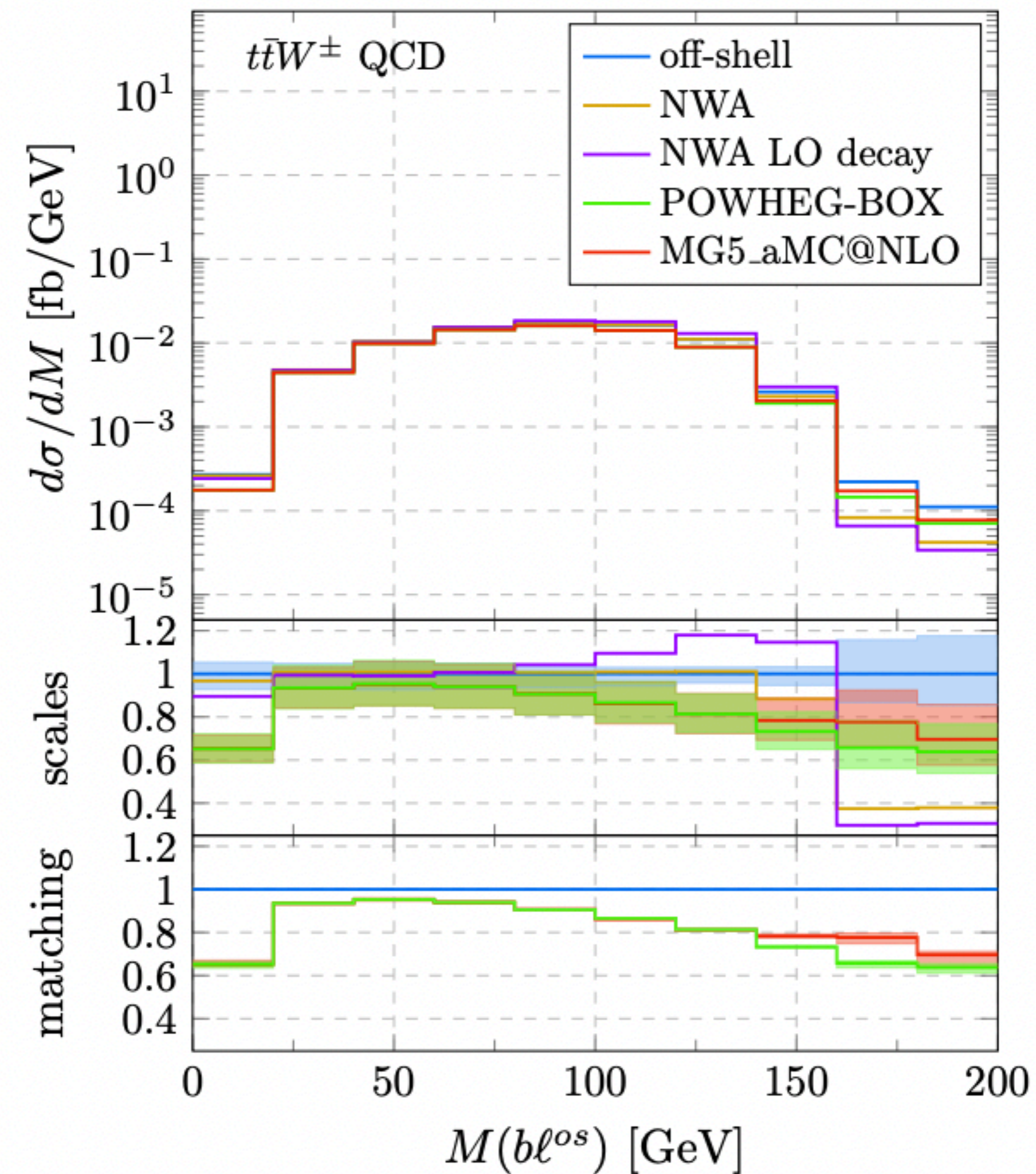
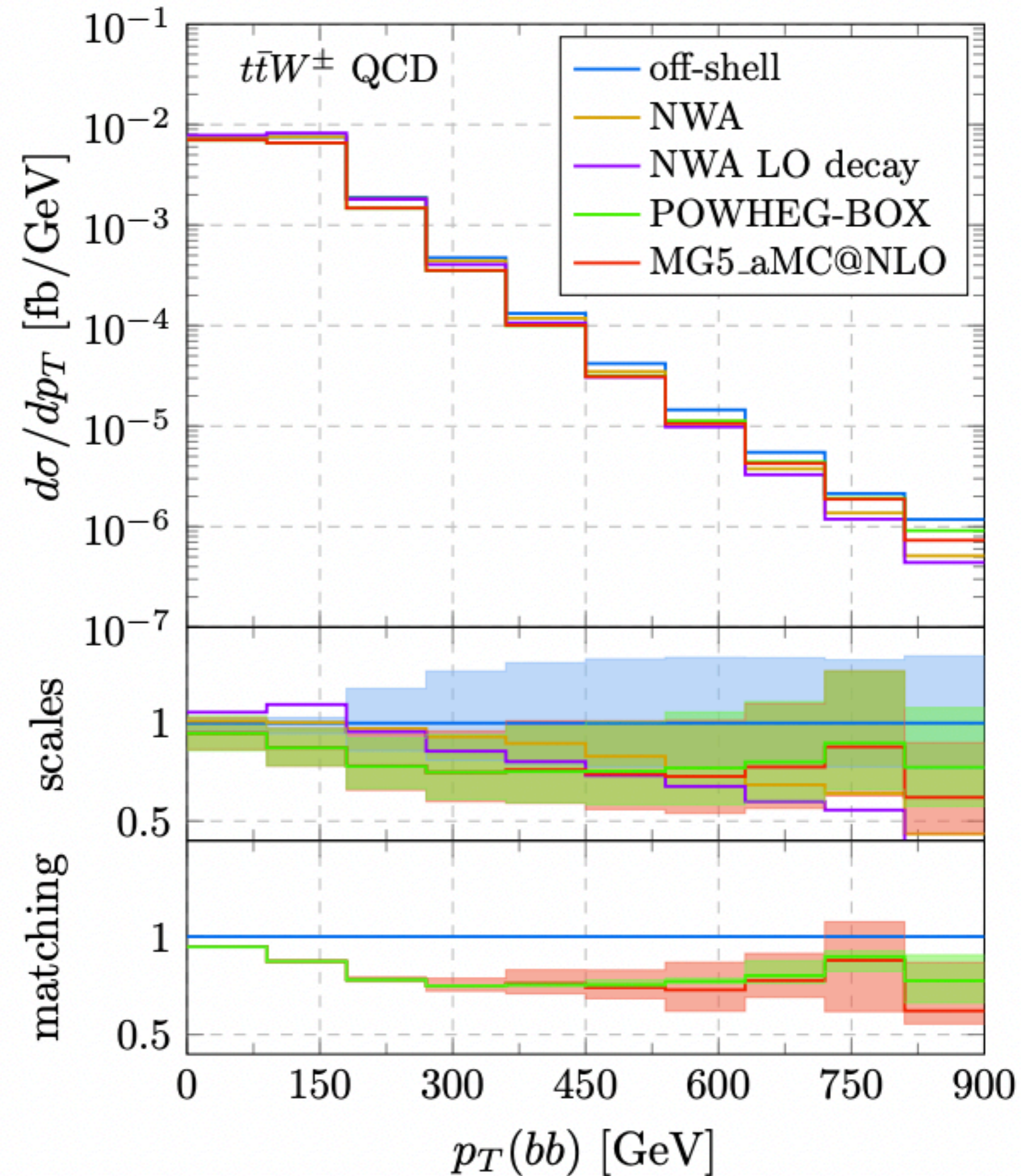
► NLO EW corrections are added to the off-shell predictions [[2102.03246](#)]





Off-shell vs NWA vs LO decay

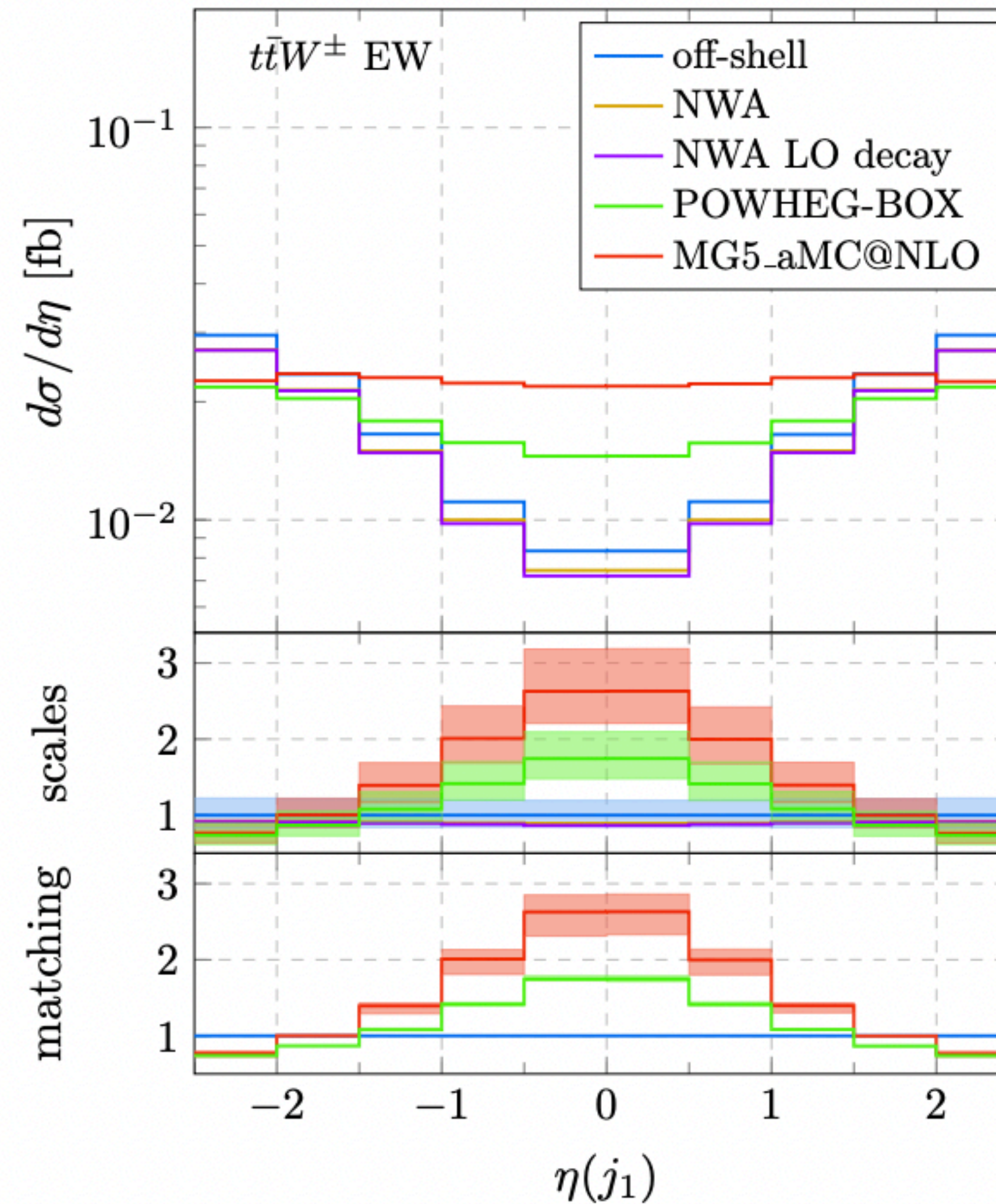
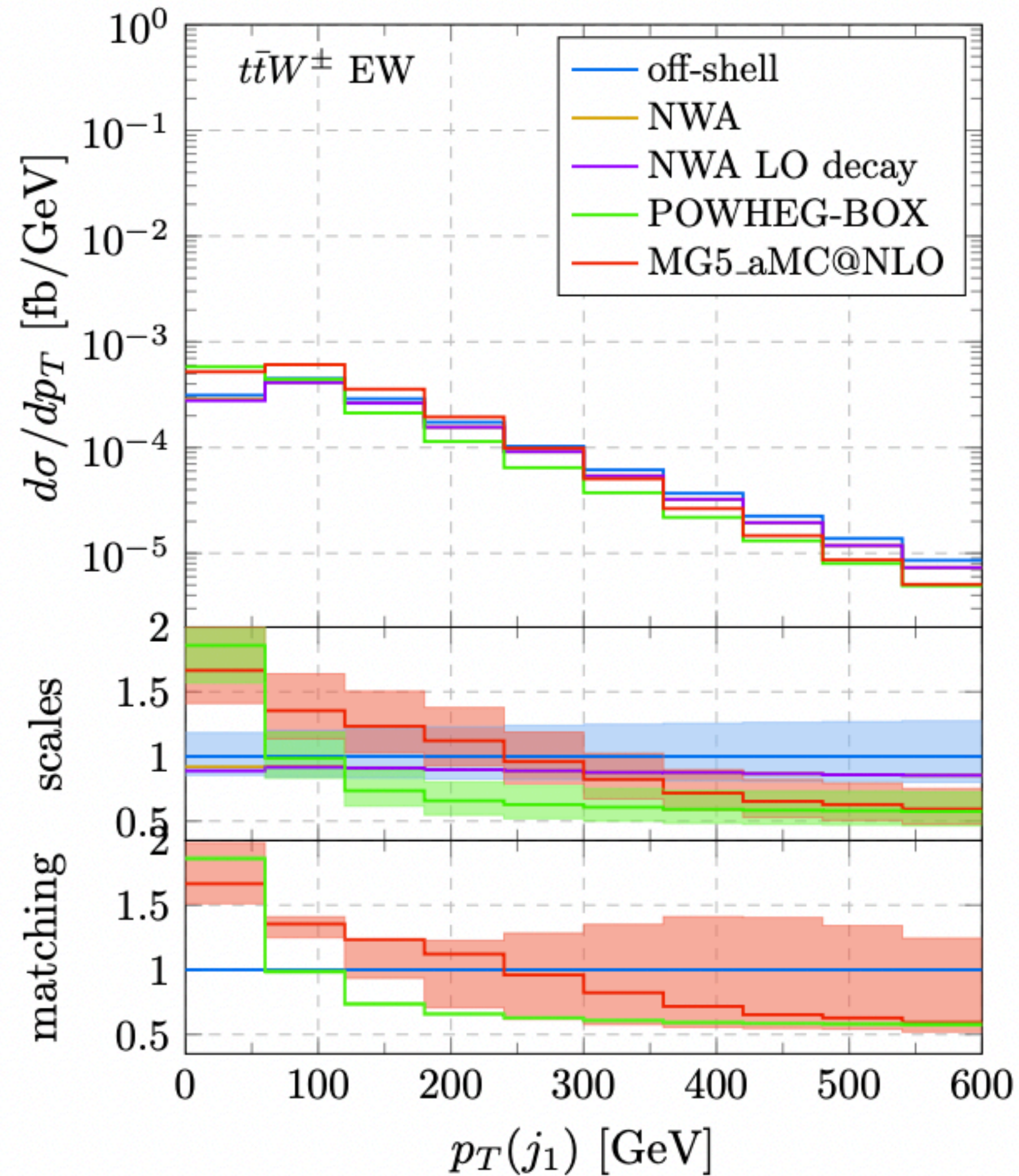
► Systematic look at modelling uncertainties [[2109.15181](#)]





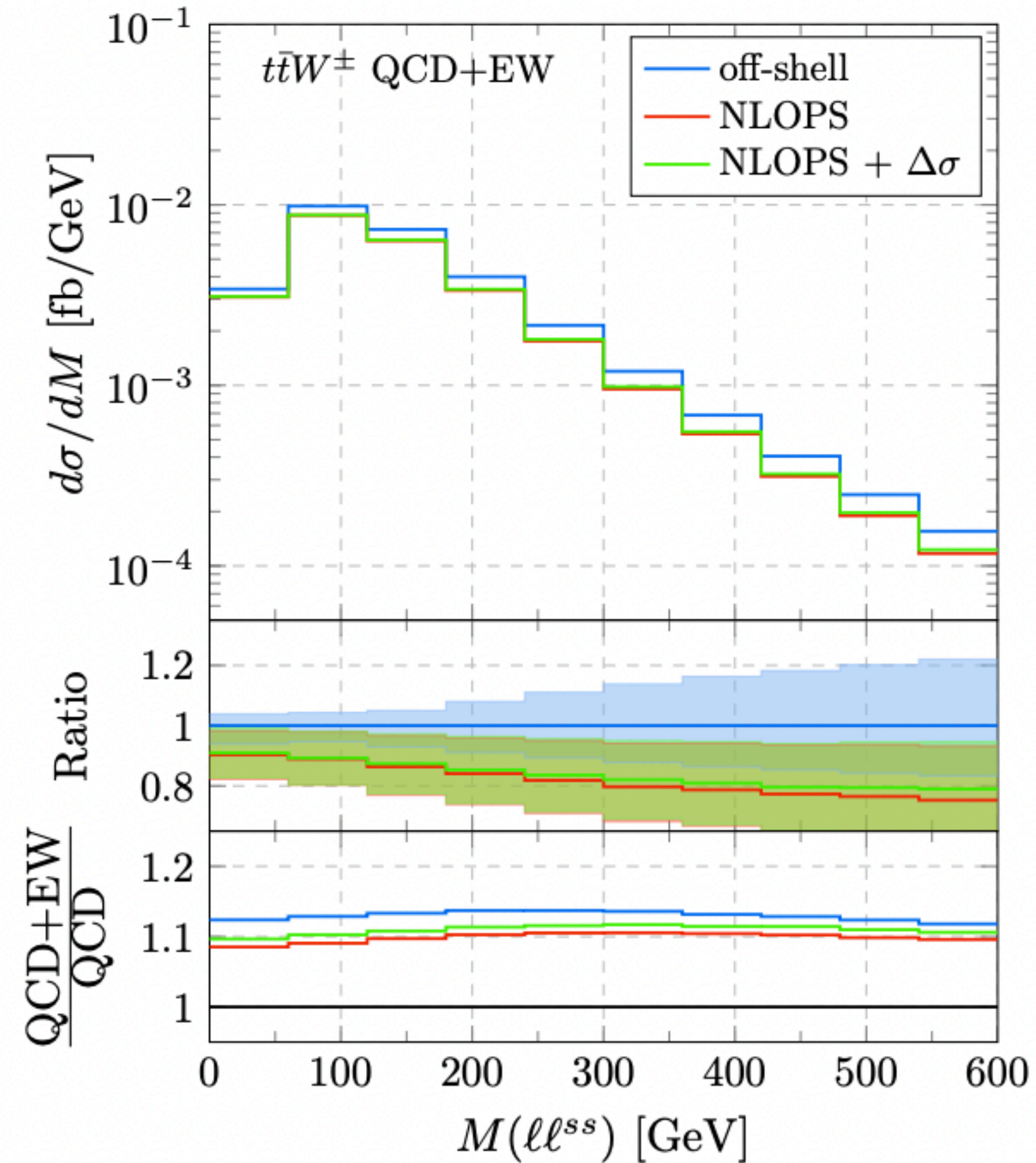
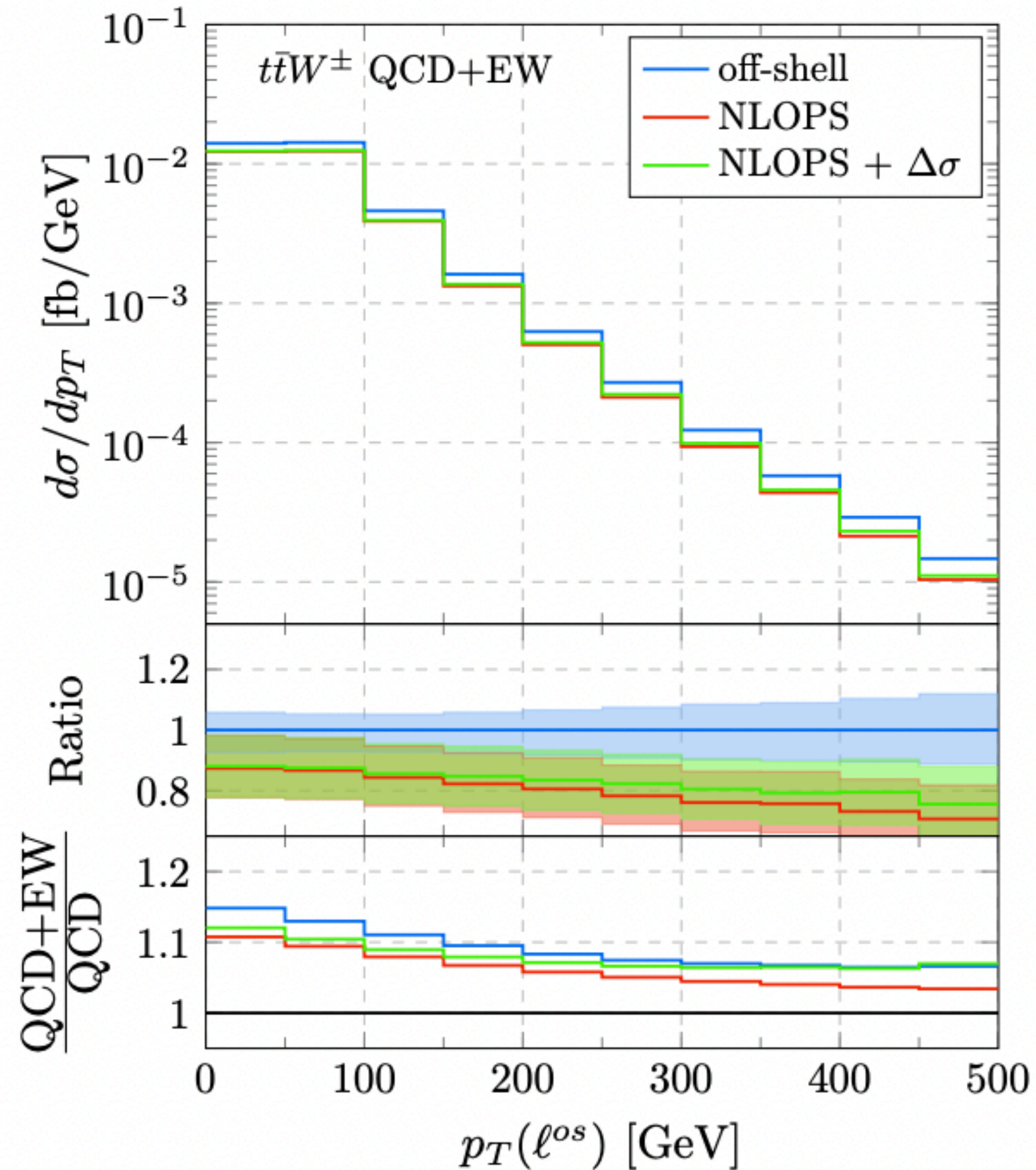
Off-shell vs NWA vs LO decay

► Systematic look at modelling uncertainties [2109.15181]



Off-shell vs NWA vs LO decay

- ▶ Systematic look at modelling uncertainties [[2109.15181](#)]
- ▶ Including method to add off-shell effects to NLOPS sample:

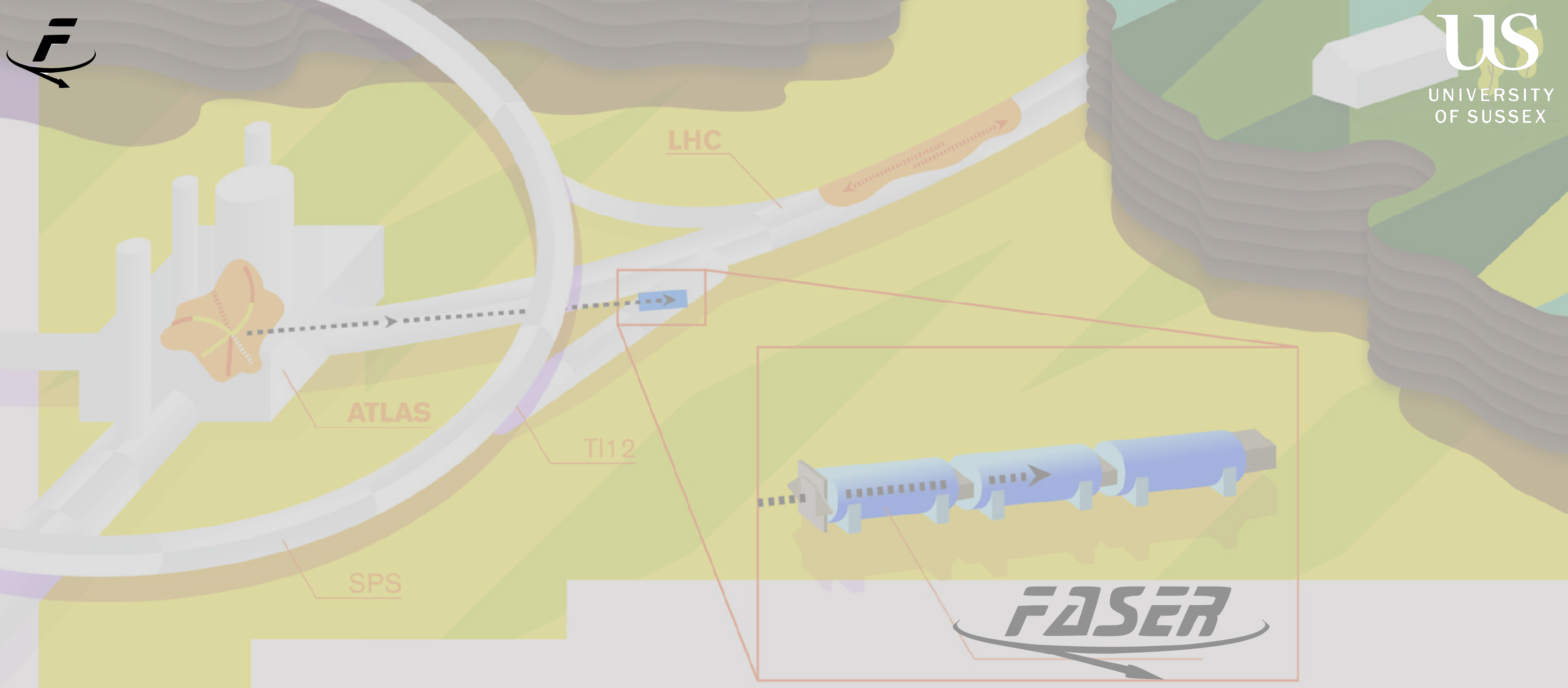


- ▶ So far doesn't seem that any of these improvements in calculations are able to resolve tensions between predictions and measurements
- ▶ General consensus that NLO+NNLL predictions are not as important at NLO-merged predictions
 - ▶ *“a large component of NLO_1 corrections, and therefore the associated scale uncertainties, originates from hard radiation in the $gq \rightarrow ttWq'$ channel. Therefore, the threshold resummation in the $qq \rightarrow ttW$ channels is not expected to drastically reduce the total scale uncertainty.” - [1907.04343]*
 - ▶ Uncertainty is still smaller, but misses crucial $ttWj@NLO$ and $ttWjj@LO$ contributions
- ▶ NLO off-shell calculations show important contributions in certain regions of phase space, but affect on inclusive cross section is small ($\sim 5\%$)

Off-shell effects inclusively

► <https://arxiv.org/pdf/2005.09427.pdf>

MODELLING APPROACH	σ^{LO} [ab]	σ^{NLO} [ab]
full off-shell ($\mu_0 = m_t + m_W/2$)	$106.9^{+27.7(26\%)}_{-20.5(19\%)}$	$123.2^{+6.3(5\%)}_{-8.7(7\%)}$
full off-shell ($\mu_0 = H_T/3$)	$115.1^{+30.5(26\%)}_{-22.5(20\%)}$	$124.4^{+4.3(3\%)}_{-7.7(6\%)}$
NWA ($\mu_0 = m_t + m_W/2$)	$106.4^{+27.5(26\%)}_{-20.3(19\%)}$	$123.0^{+6.3(5\%)}_{-8.7(7\%)}$
NWA ($\mu_0 = H_T/3$)	$115.1^{+30.4(26\%)}_{-22.4(19\%)}$	$124.2^{+4.1(3\%)}_{-7.7(6\%)}$
NWA _{LOdecay} ($\mu_0 = m_t + m_W/2$)		$127.0^{+14.2(11\%)}_{-13.3(10\%)}$
NWA _{LOdecay} ($\mu_0 = H_T/3$)		$130.7^{+13.6(10\%)}_{-13.2(10\%)}$



Back-ups

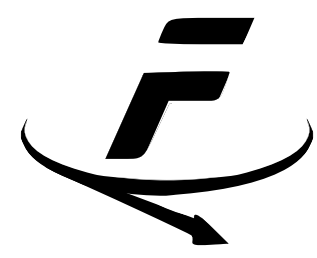
Collaboration

- ▶ 87 members
- ▶ 24 institutions
- ▶ 10 countries



International laboratory covered by a cooperation agreement with CERN





Installation



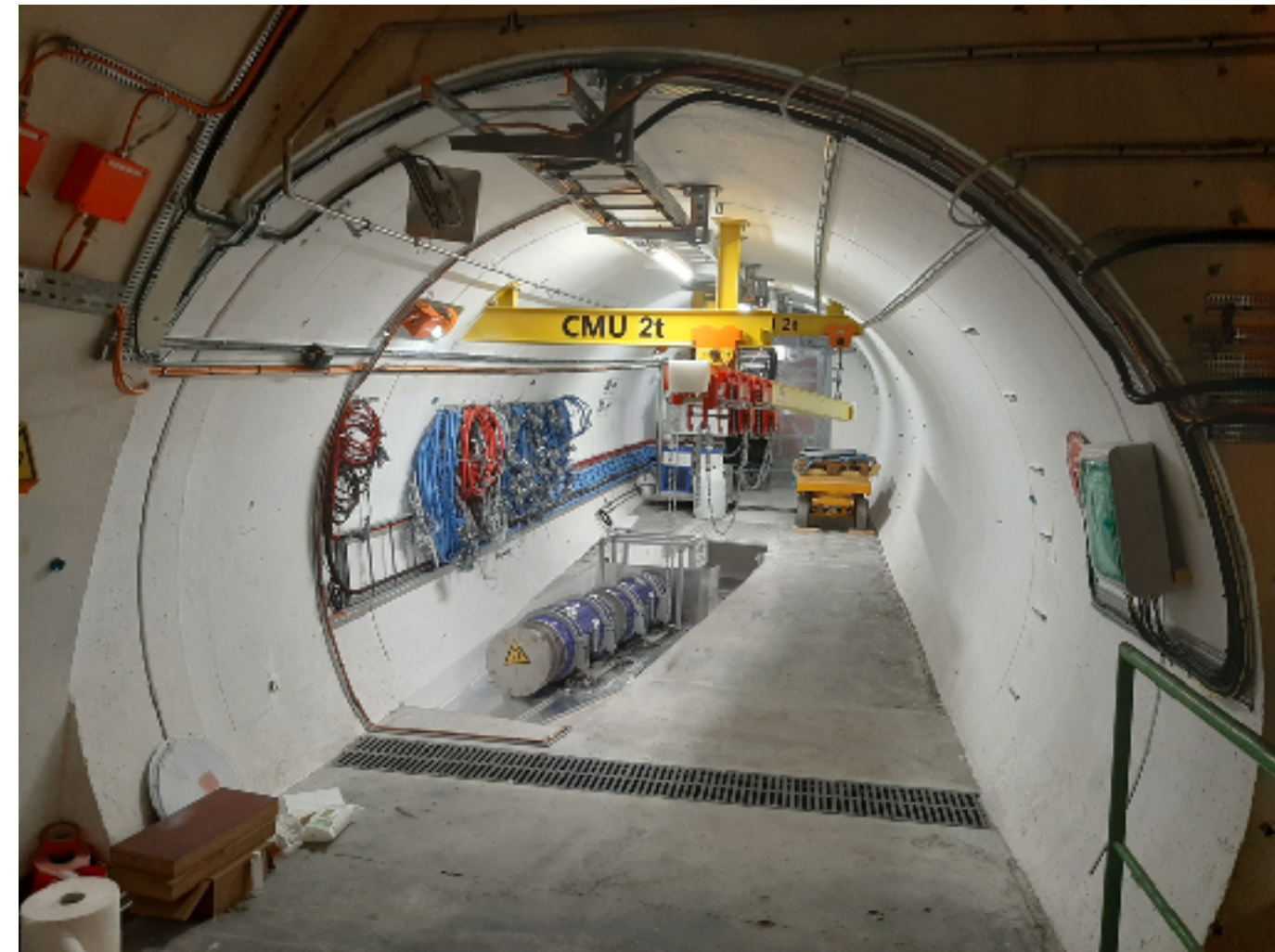
8/18



8/19



4/20



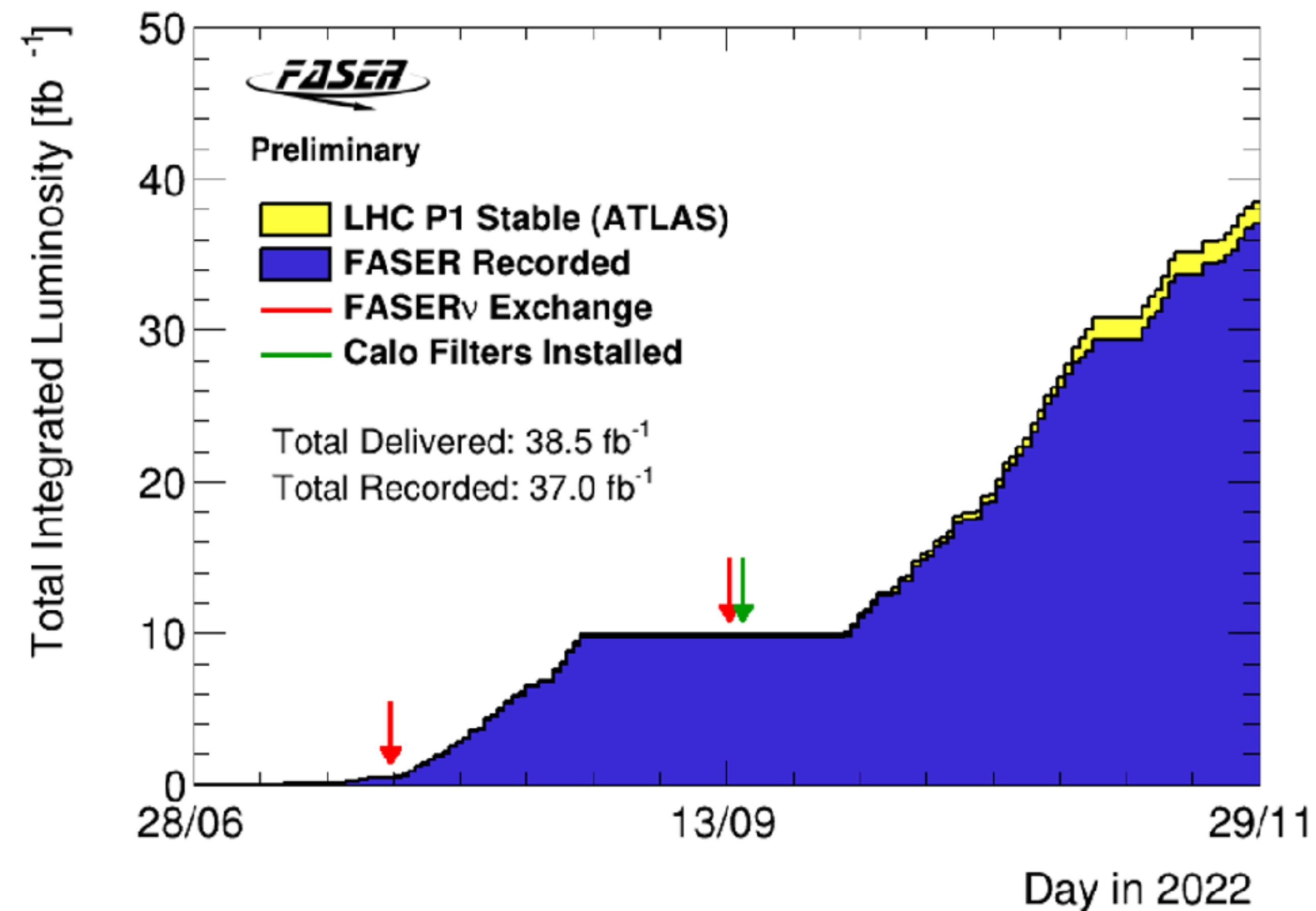
11/20



3/21

Operations

- ▶ Successfully operated throughout 2022
 - ▶ Continuous data taking
 - ▶ Largely automated
 - ▶ Up to 1.3 kHz
- ▶ Recorded 96.1% of delivered lumi.
 - ▶ DAQ dead-time of 1.3%
 - ▶ couple of DAQ crashes
- ▶ Emulsion detector exchanged twice
 - ▶ Needed to manage occupancy
 - ▶ First box only partially filled
- ▶ Calorimeter gain optimised for:
 - ▶ Low E (<300 GeV) before 2nd exchange
 - ▶ High E (up to 3 TeV) after this exchange



Analyses presented use 27.0 fb⁻¹ or 35.4 fb⁻¹



Dark Photon (A') Search

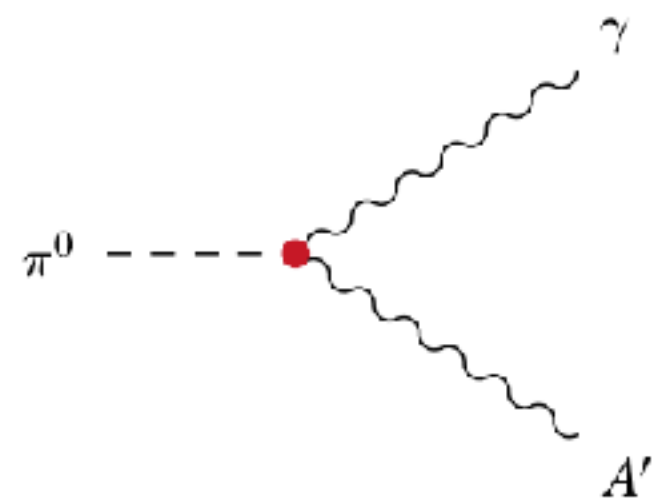


▶ Dark photon is a common feature of hidden sector models:

▶ Weakly coupling to SM via kinetic mixing (ϵ) with SM photon

$$\mathcal{L} \supset \frac{1}{2} m_{A'}^2 A'^2 - \epsilon e \sum_f q_f \bar{f} A' f$$

▶ MeV A' 's produced mainly in meson decays at LHC



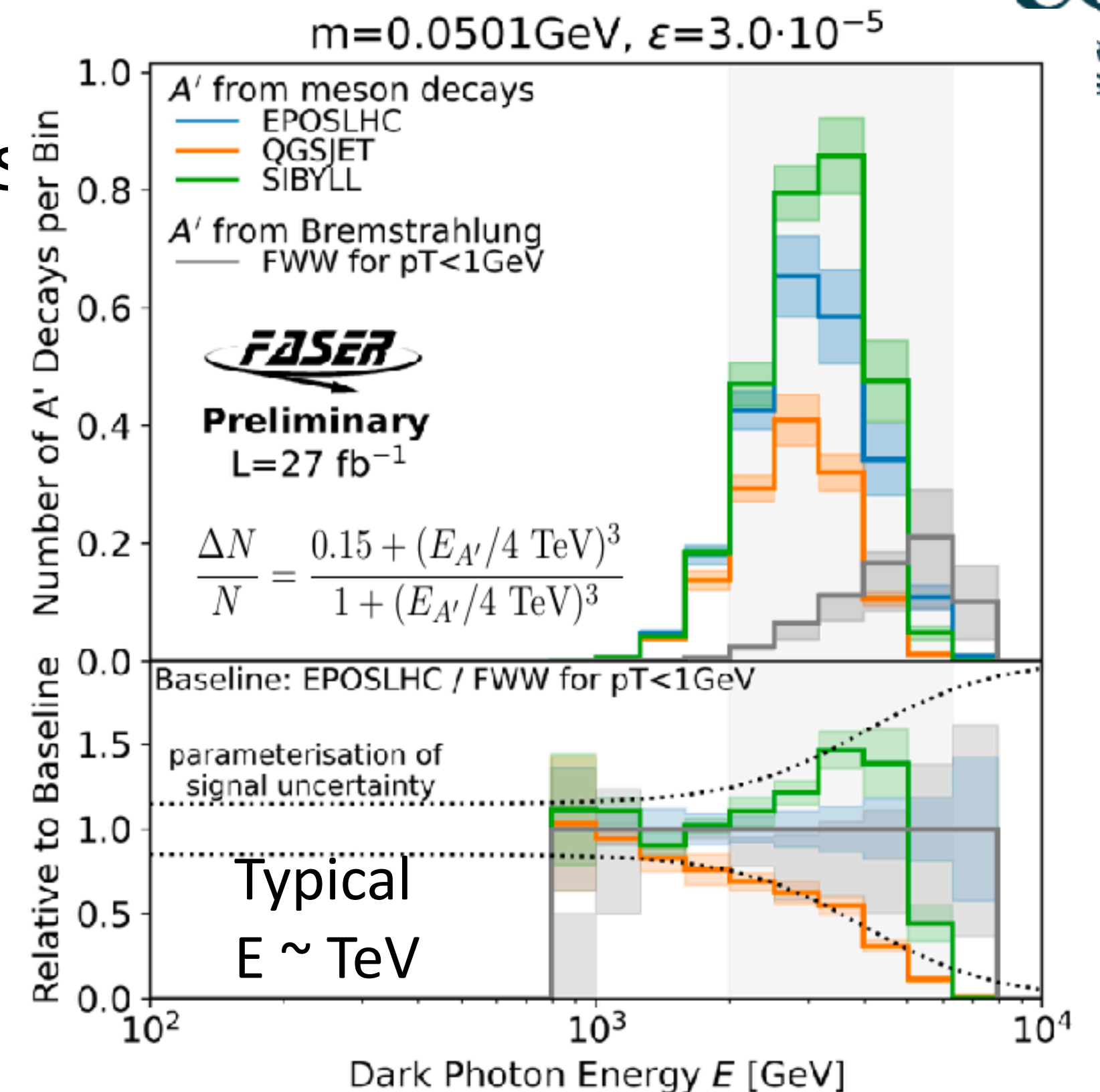
$$B(\pi^0 \rightarrow A' \gamma) = 2\epsilon^2 \left(1 - \frac{m_{A'}^2}{m_{\pi^0}^2}\right)^3 B(\pi^0 \rightarrow \gamma \gamma)$$

▶ FASER targets small ϵ , where A' has long decay length

$$L = c\beta\tau\gamma \approx (80 \text{ m}) \left[\frac{10^{-5}}{\epsilon}\right]^2 \left[\frac{E_{A'}}{\text{TeV}}\right] \left[\frac{100 \text{ MeV}}{m_{A'}}\right]^2$$

▶ Below $2m_{\mu}$, A' has 100% decay to e^+e^- pair

* arXiv:2105.07077



▶ $A' \rightarrow e^+e^-$ simulated with FORESEE*

- ▶ π^0 and η via EPOS-LHC generator
- ▶ Subdominant dark brem. via FWW

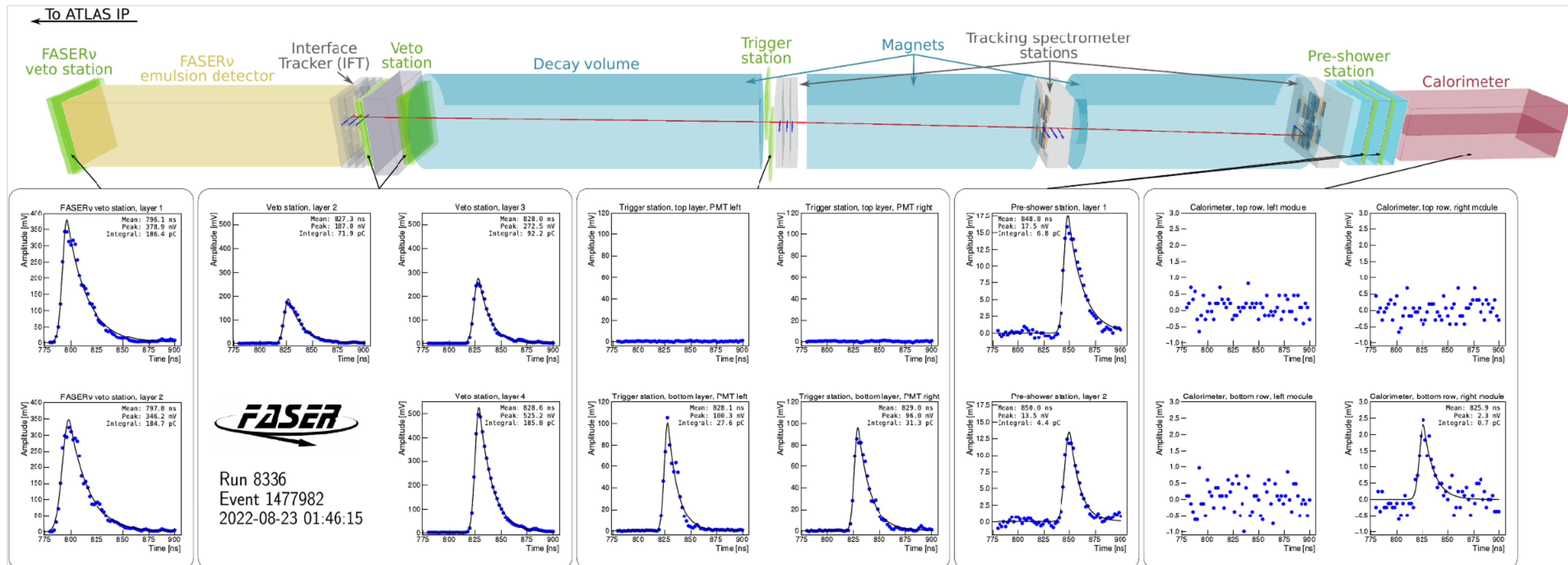
▶ Generator uncertainty dominates

▶ Difference to QGSJET/SIBYLL

▶ Parameterised based on A' energy

Operations

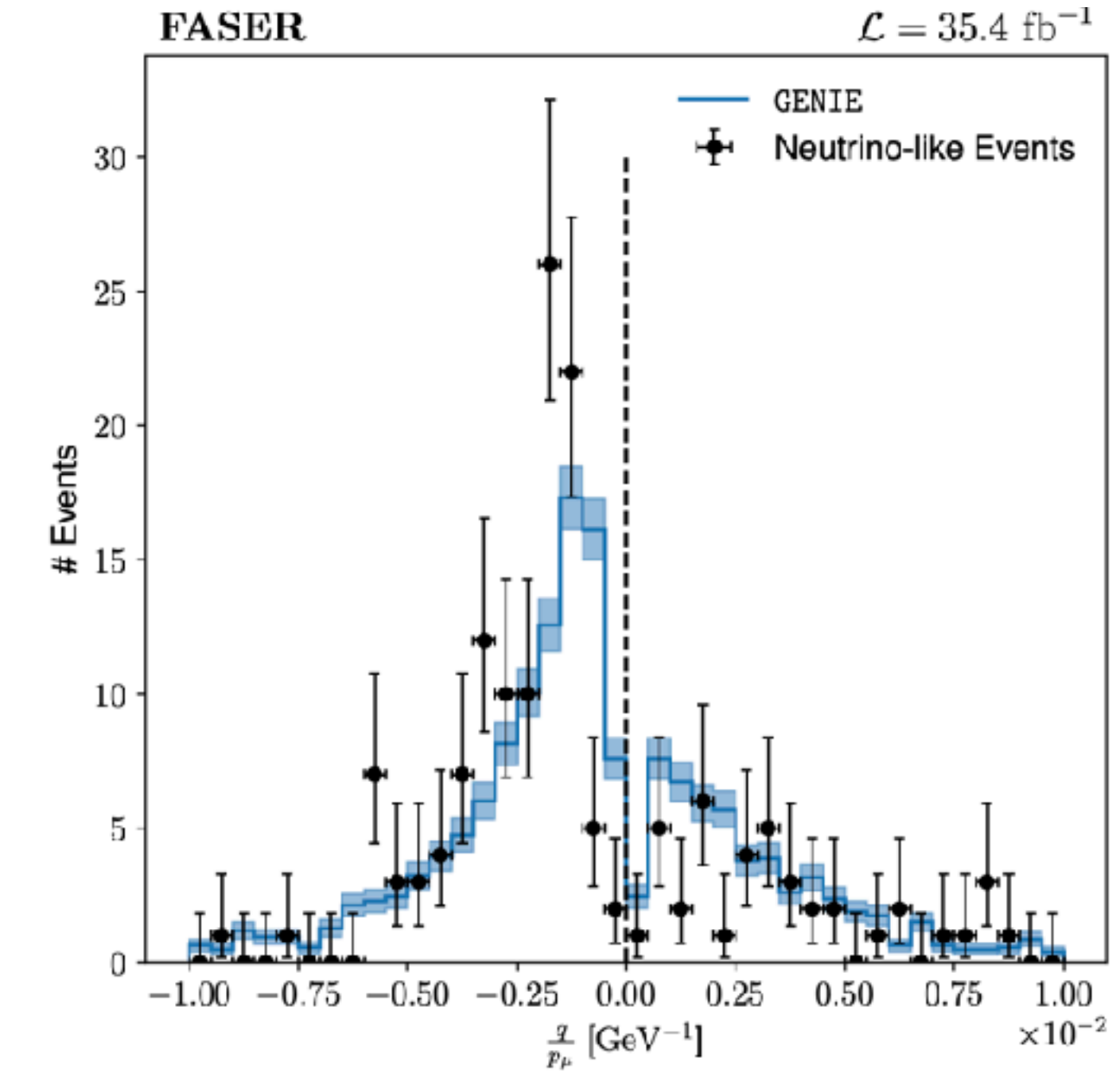
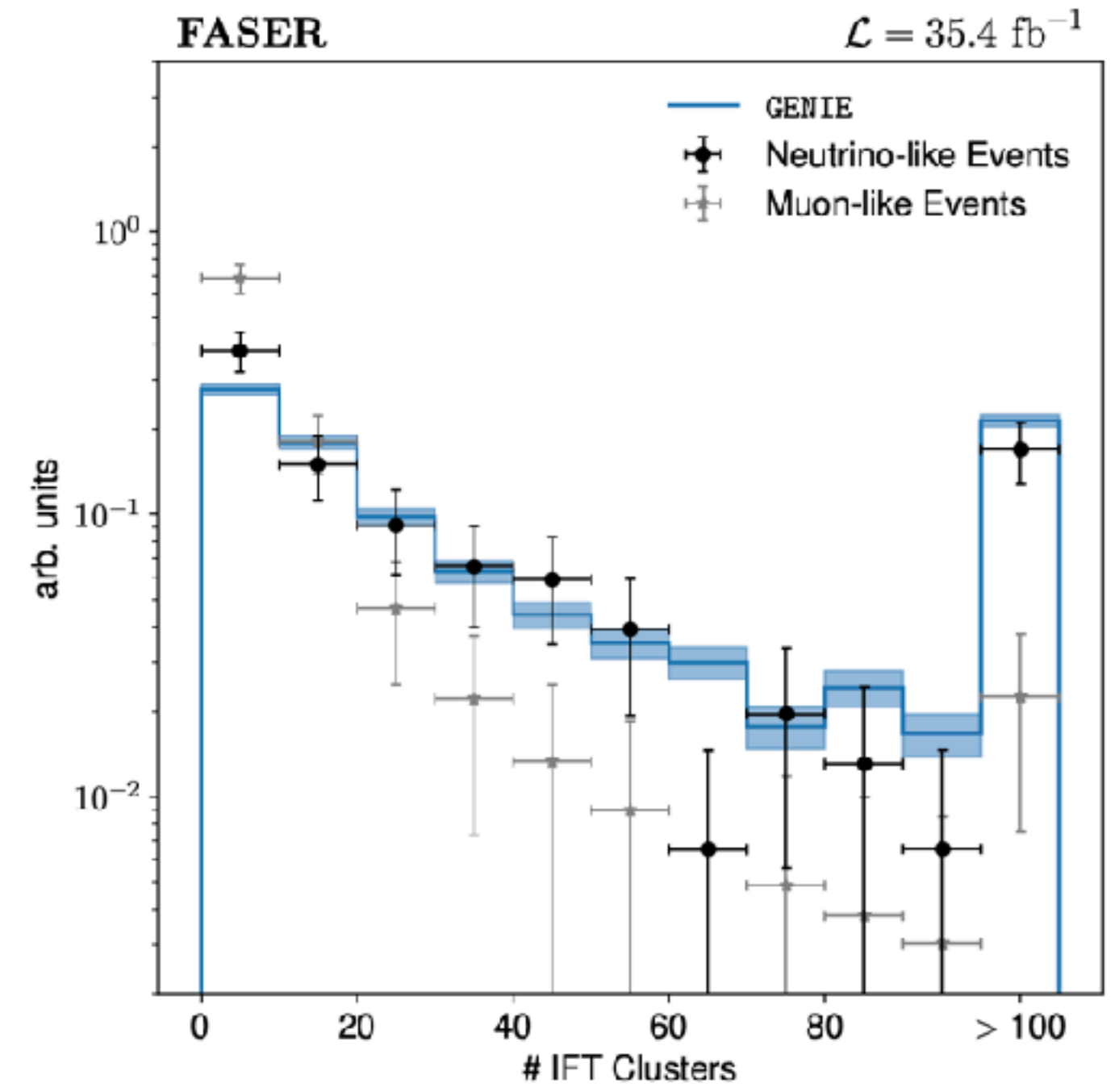
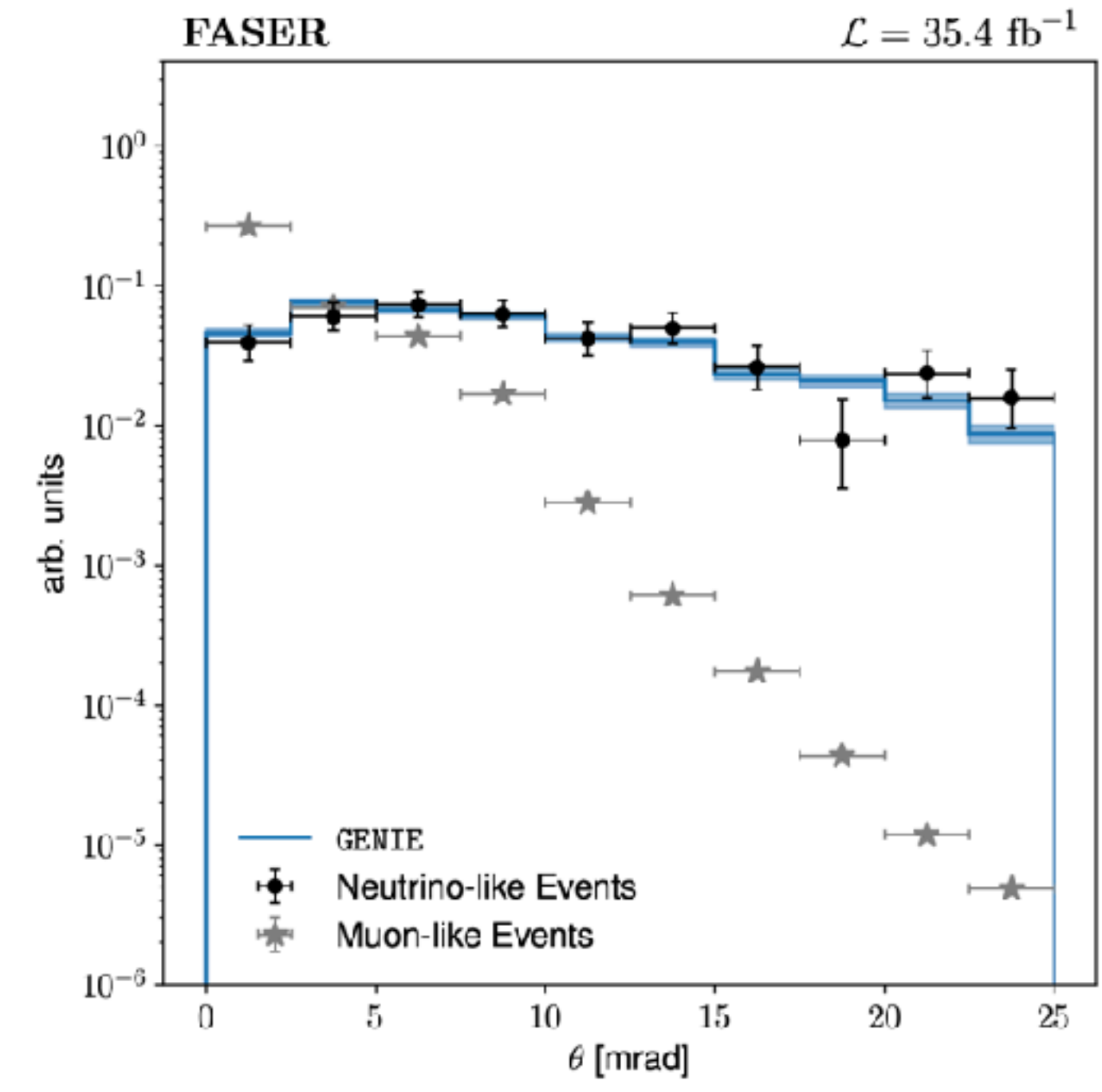
- ▶ All detector components performing excellently
 - ▶ More than 350M single-muon events recorded
 - ▶ Example: muon leaving track passing through full detector - consistent with MIP



Neutrino analysis | Results

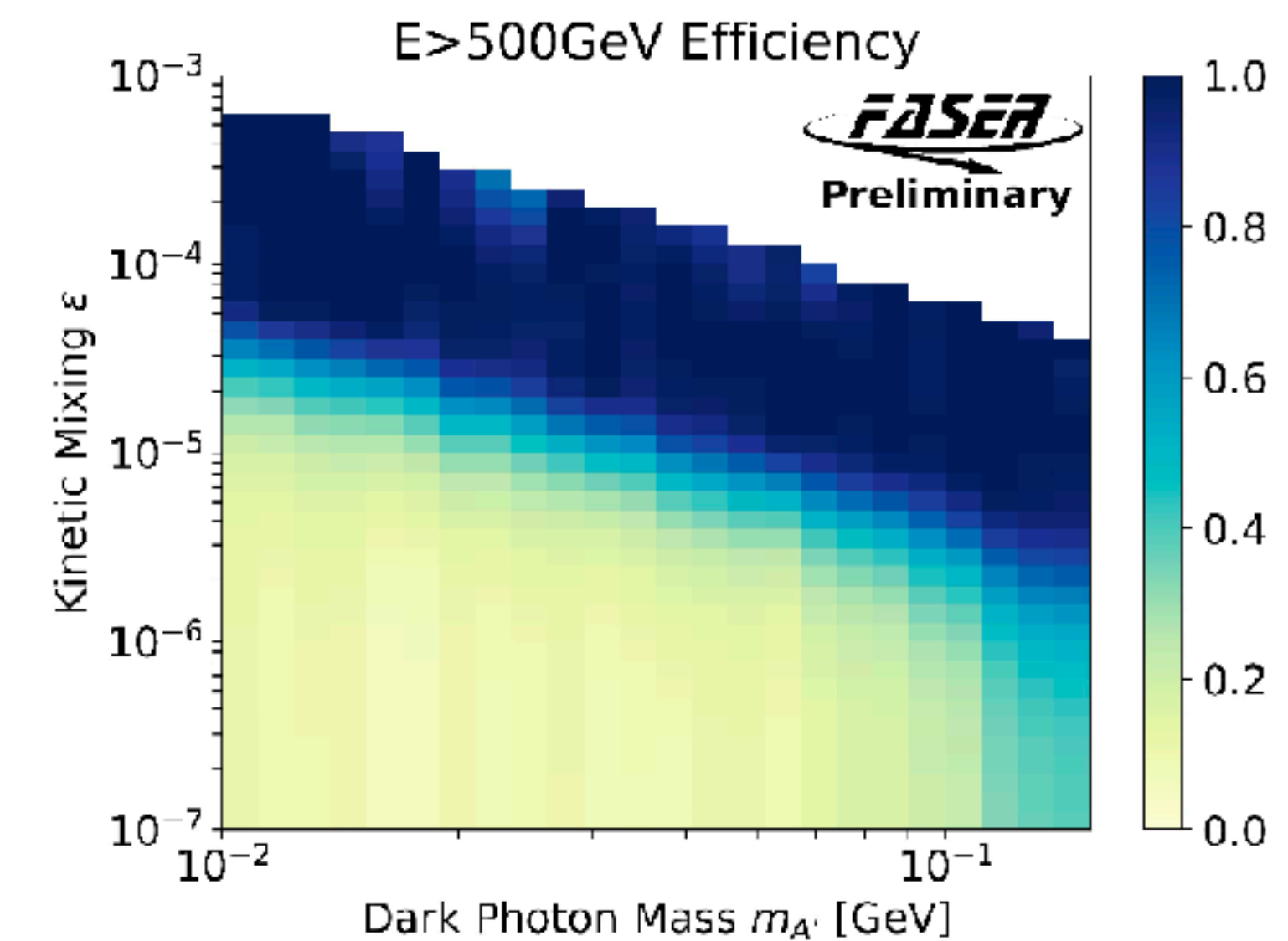
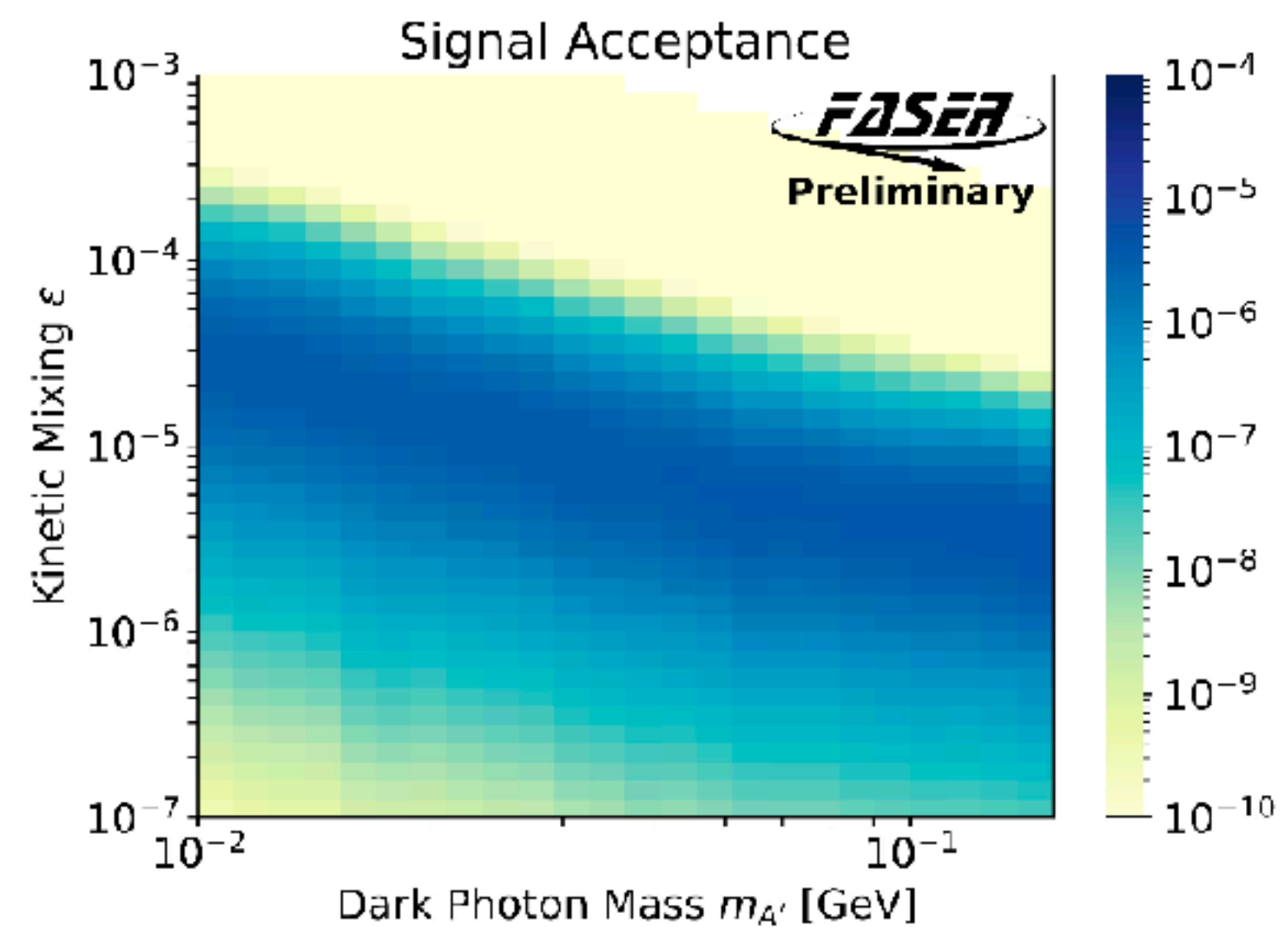
- ▶ Candidate neutrino events match expectation from signal
- ▶ High occupancy in front tracker station
- ▶ Most events have high μ momentum
- ▶ More $\nu\mu$ than anti- $\nu\mu$

▶ NB: no acceptance corrections nor any systematic uncertainties in these plots

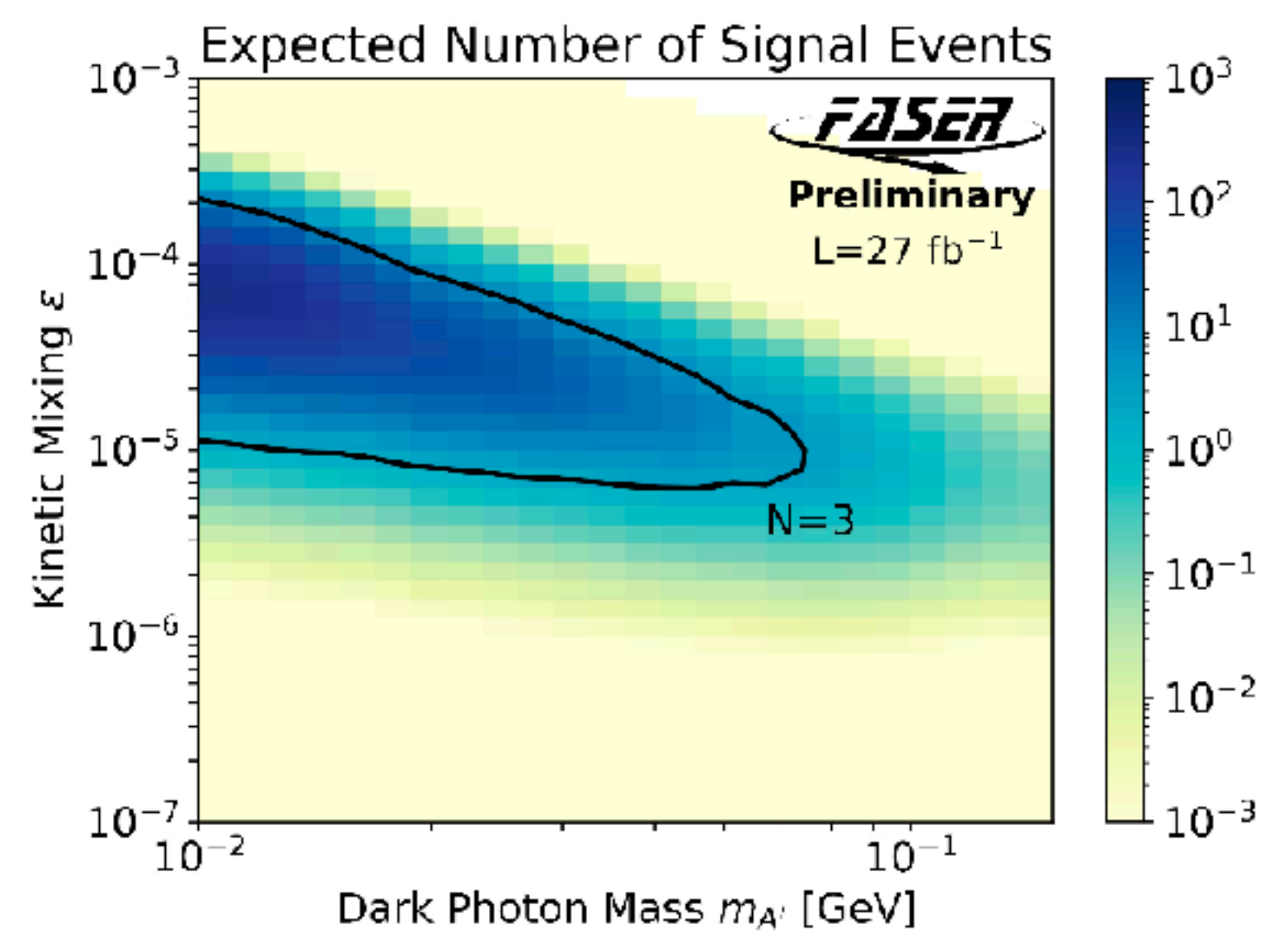


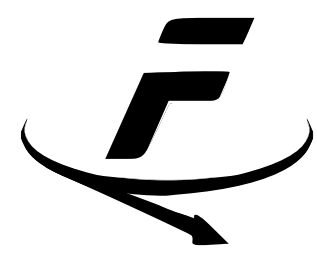
Dark Photon | Signal

- ▶ Acceptance 10^{-6}
- ▶ Decay volume 10^{-8} solid angle
- ▶ $P(\text{decay in FASER}) = 10^{-3}$



$$L = c\beta\tau\gamma \approx (80 \text{ m}) \left[\frac{10^{-5}}{\epsilon} \right]^2 \left[\frac{E_{A'}}{\text{TeV}} \right] \left[\frac{100 \text{ MeV}}{m_{A'}} \right]^2$$

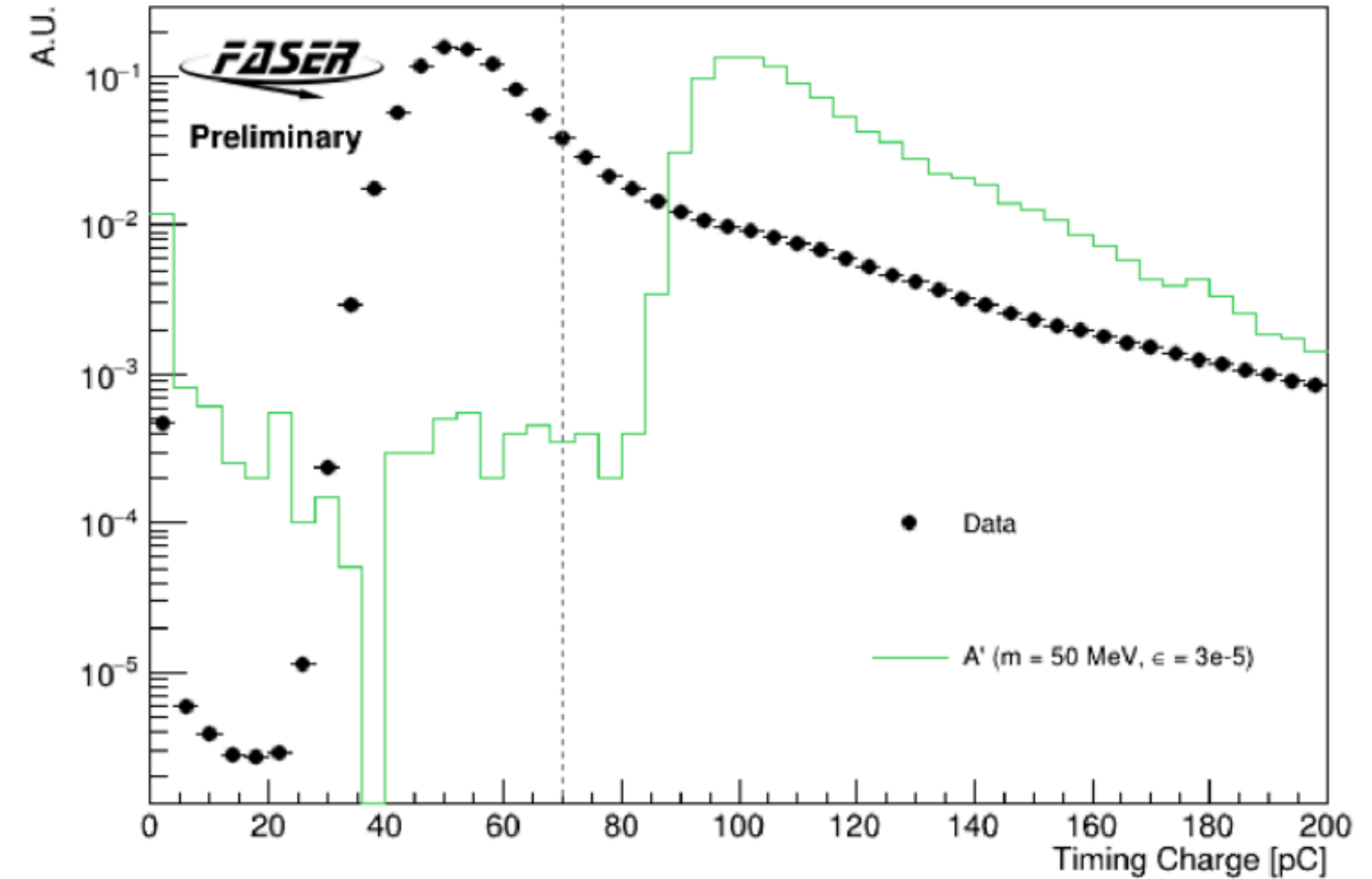




Dark Photon | Selection

Description	Value
Pre-selecton	
Time consistent with a colliding bunch identifier Timing scintillator trigger	
Scintillator	
Timing station: Top or Bottom Scintillator charge OR Top and Bottom charge	> 70 pC > 30 pC
Each Preshower scintillator charge Each Veto scintillator charge	>2.5 pC <40 pC
Tracking	
Exactly 2 Good Tracks Momentum χ^2/NDF	> 20 GeV < 25
Number of tracker layers on track Number of tracker hits on track	≥ 7 ≥ 12
Fiducial selection	
Track extrapolated to all scintillators and tracking stations	< 95mm
Calorimeter	
Calorimeter energy (sum of four channels)	> 500 GeV

TABLE I. Summary of selection requirements.



Selection Criteria	Efficiency
Good collision event	99.7%
No Veto Signal	98.4%
Timing/Preshower Signal	97.3%
≥ 1 good track	89.2%
= 2 good tracks	44.5% *
Track radius < 95 mm	42.3% *
Calo energy > 500 GeV	41.6% *

$$\epsilon = 3 \times 10^{-5} \quad m_{A'} = 25.1 \text{ MeV}$$

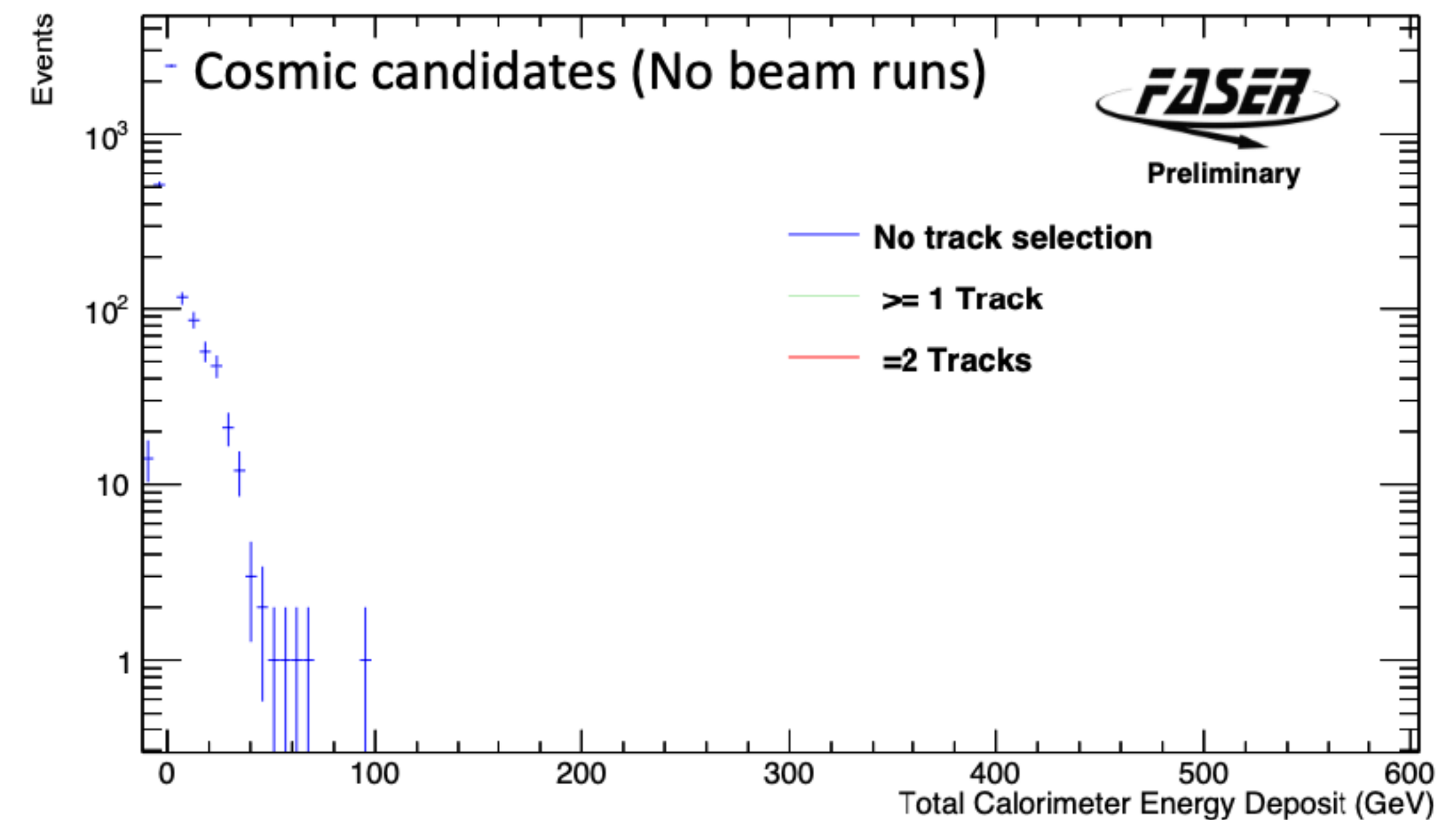
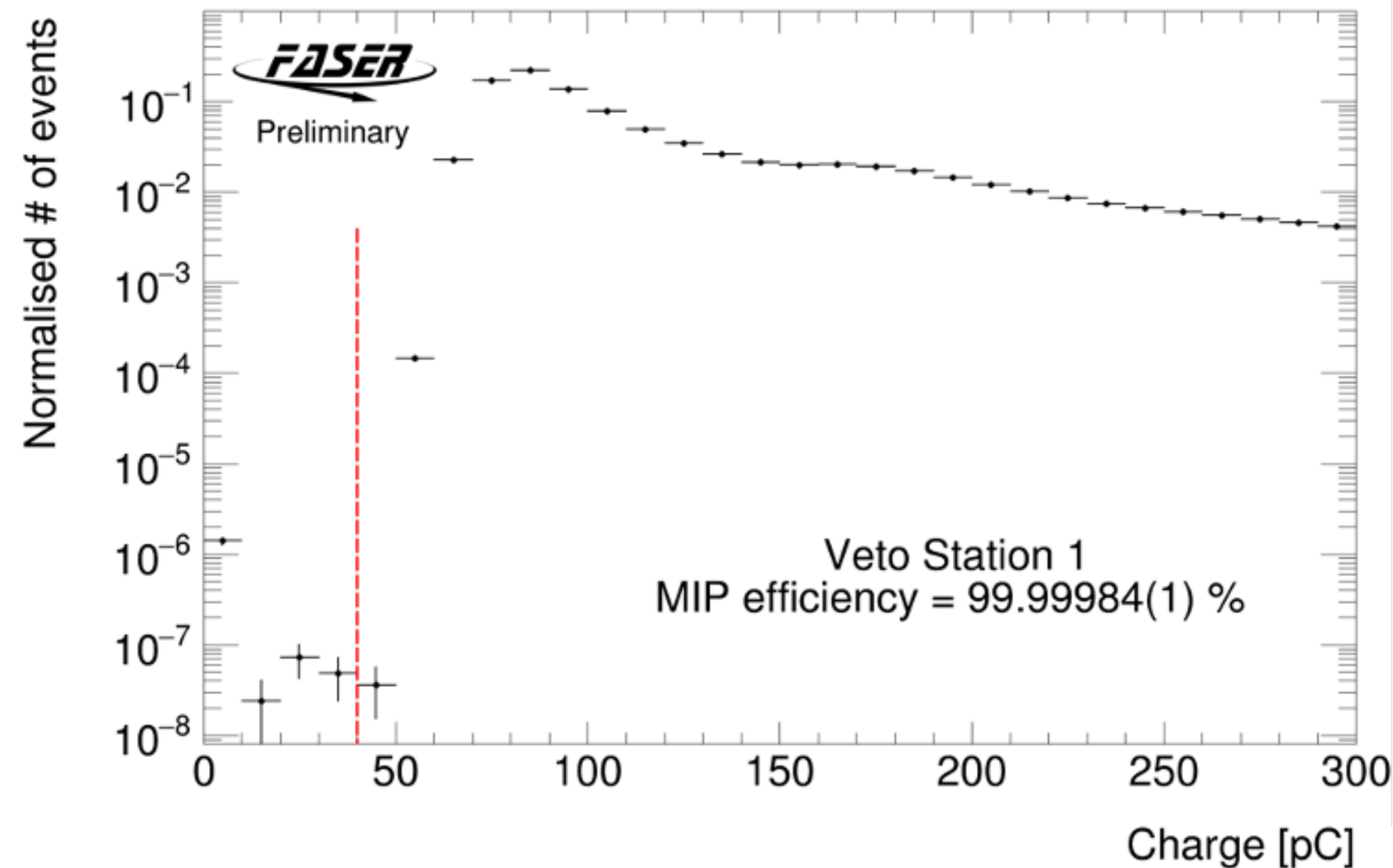
Dark Photon | Backgrounds

▶ Veto inefficiency

- ▶ Measured layer-by-layer via muons with tracks pointing back to vetos
- ▶ Layer efficiency $> 99.9998\%$
- ▶ 5 layers reduce exp. 10^8 muons to negligible level (even before cuts) ($< 10^{-20}$ inefficiency)

▶ Non-collision backgrounds

- ▶ Cosmics measured in runs with no beam
- ▶ Near-by beam debris measured in non-colliding bunches
- ▶ No events observed with ≥ 1 track or $E(\text{calo}) > 500$ GeV individually



Dark Photon | Backgrounds

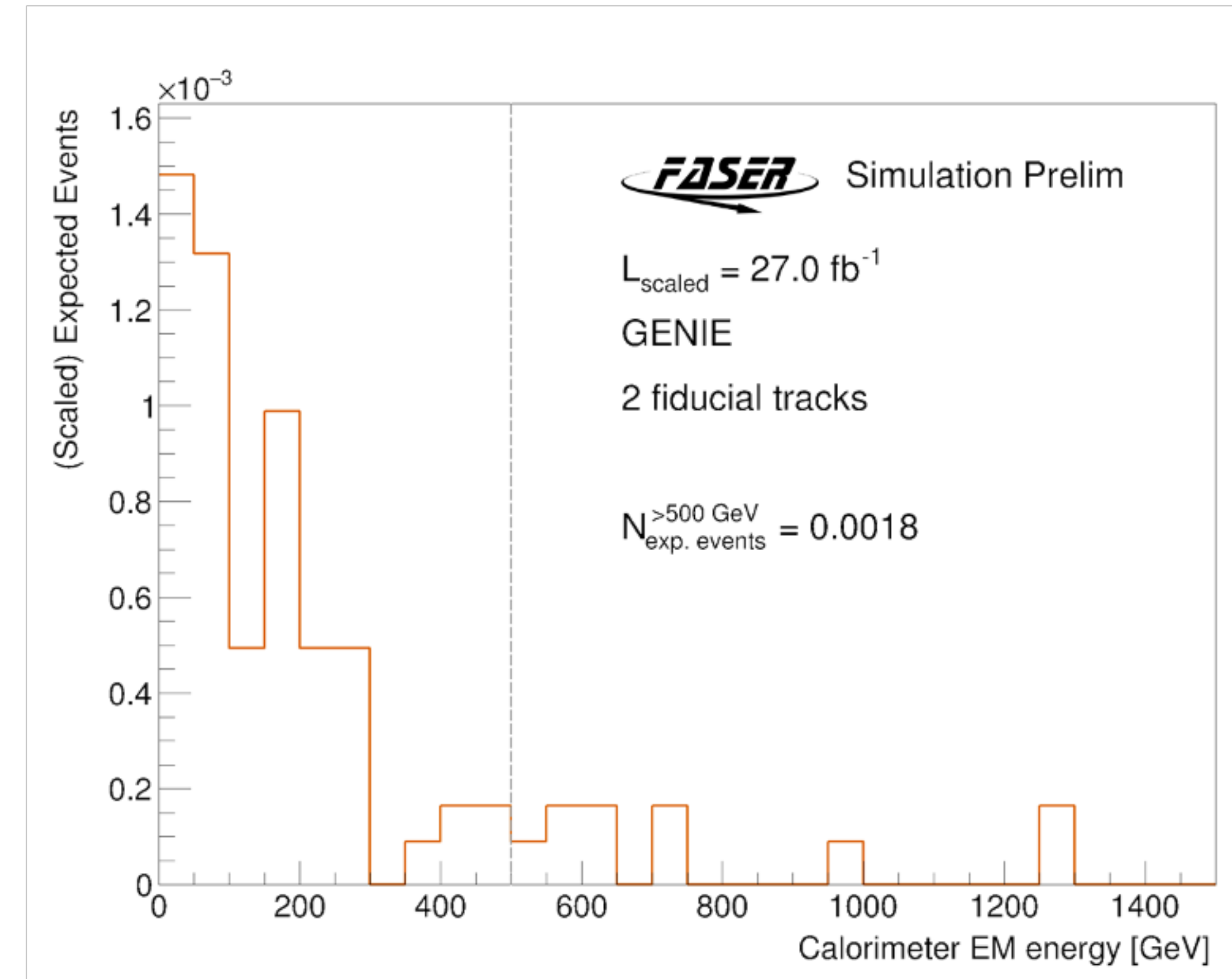
- ▶ Main background is from Neutrino interactions
 - ▶ Primarily coming from vicinity of timing detector
 - ▶ Estimated from GENIE simulation (300 ab⁻¹)
 - ▶ Uncertainties from neutrino flux & mismodelling
 - ▶ Predicted events with E(calor) > 500 GeV

$$N = (1.8 \pm 2.4) \times 10^{-3}$$

- ▶ Neutral hadrons (e.g. Ks) from upstream muons interacting in rock in front of FASER

- ▶ Heavily suppressed since:
 - ▶ muon nearly always continues after interaction
 - ▶ has to pass through 8 interaction lengths (FASERv)
 - ▶ decay products have to leave E(calor) > 500 GeV
- ▶ Estimated from lower energy events with 2/3 tracks and different veto conditions

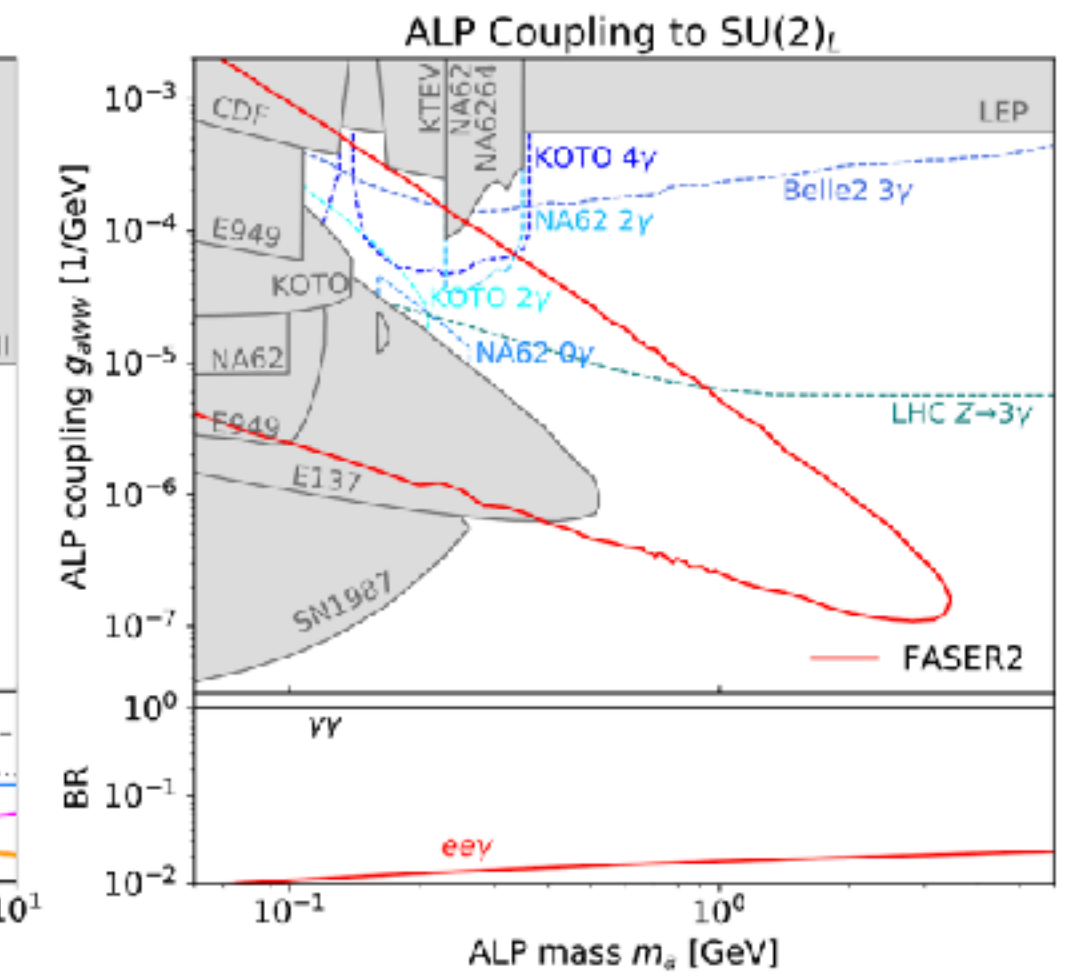
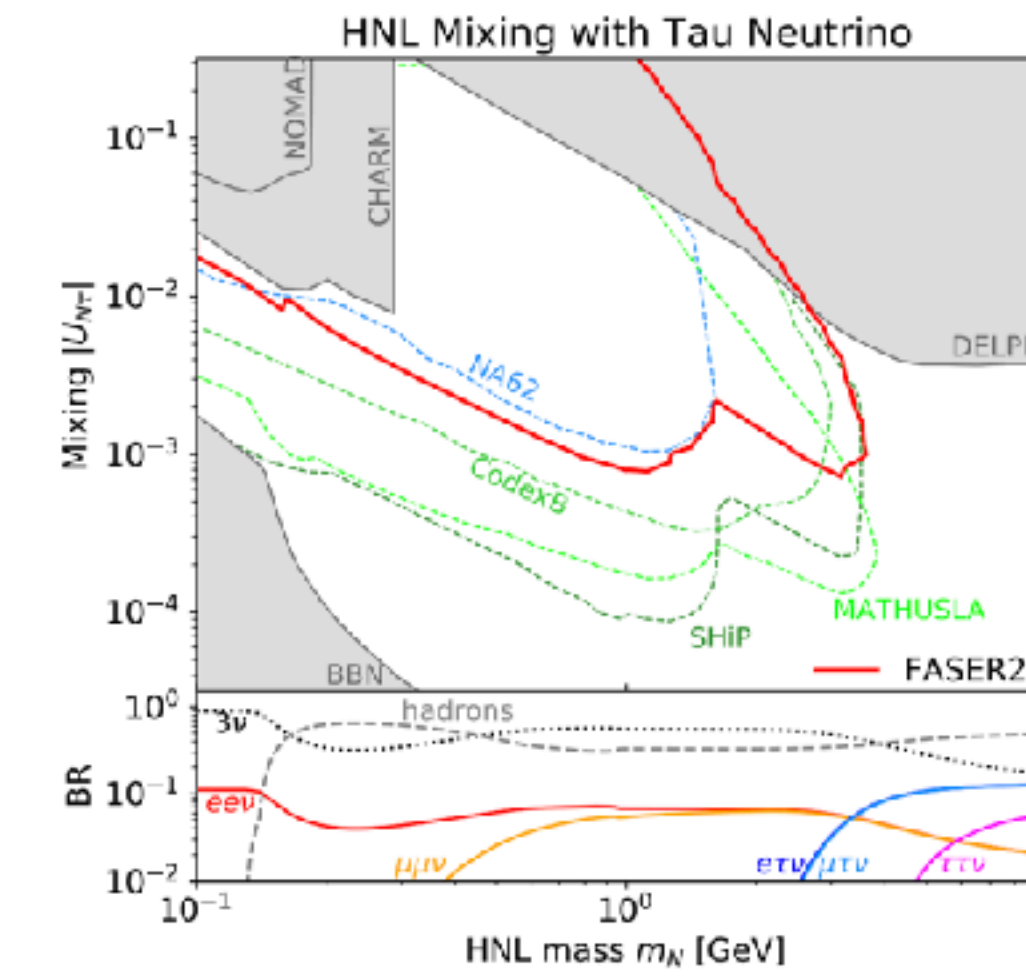
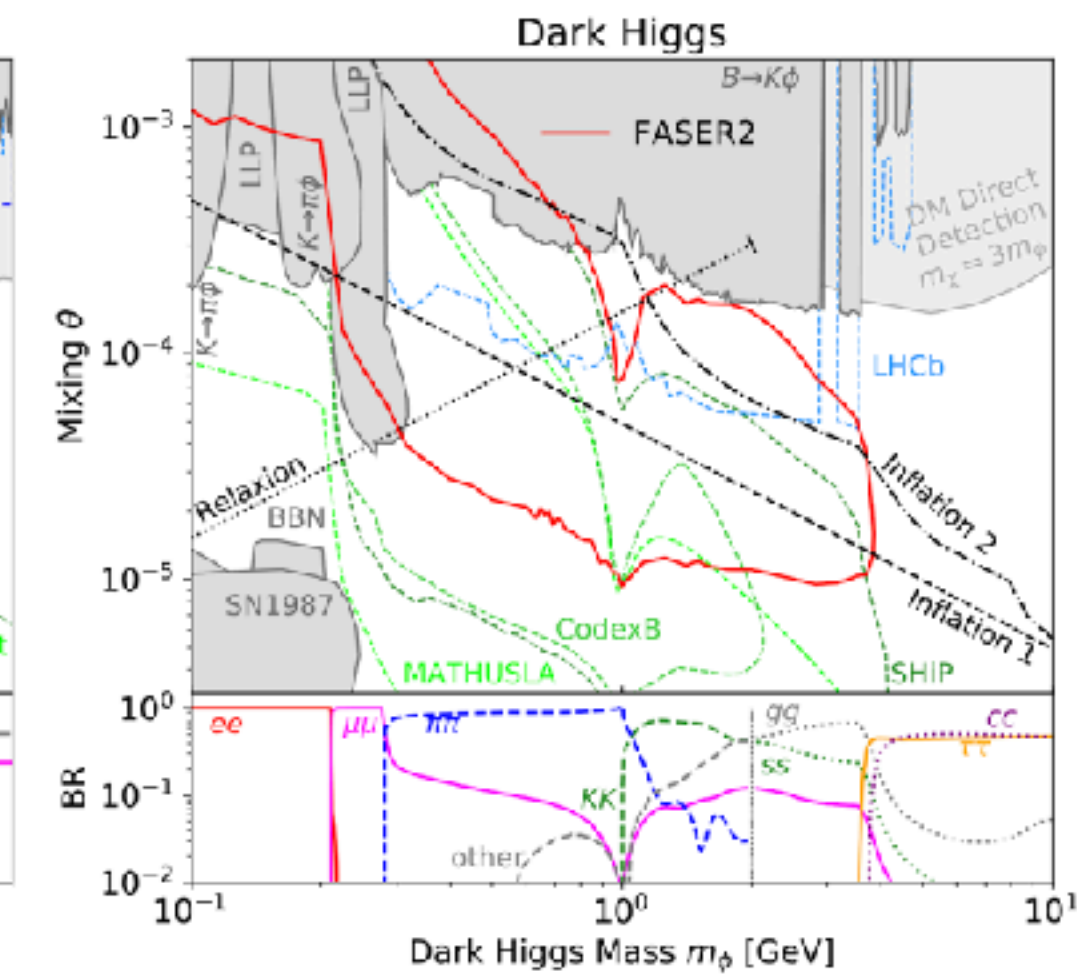
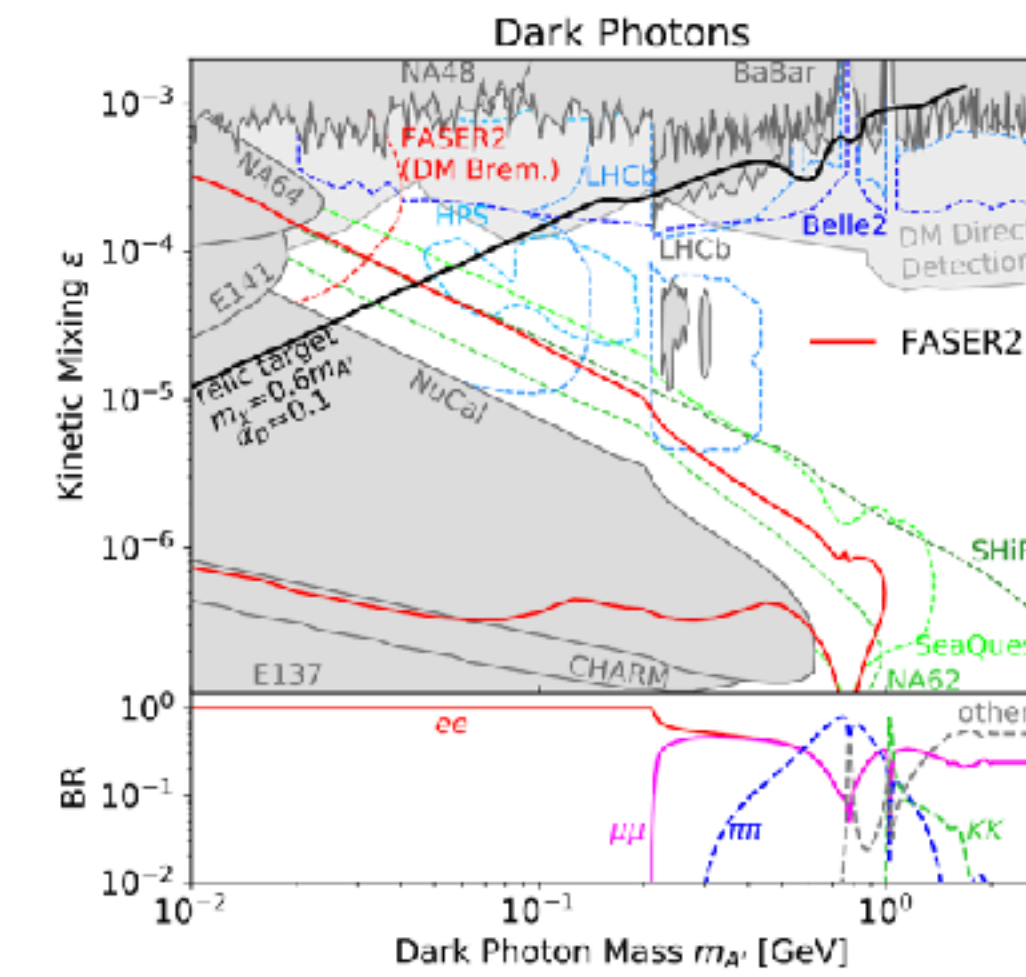
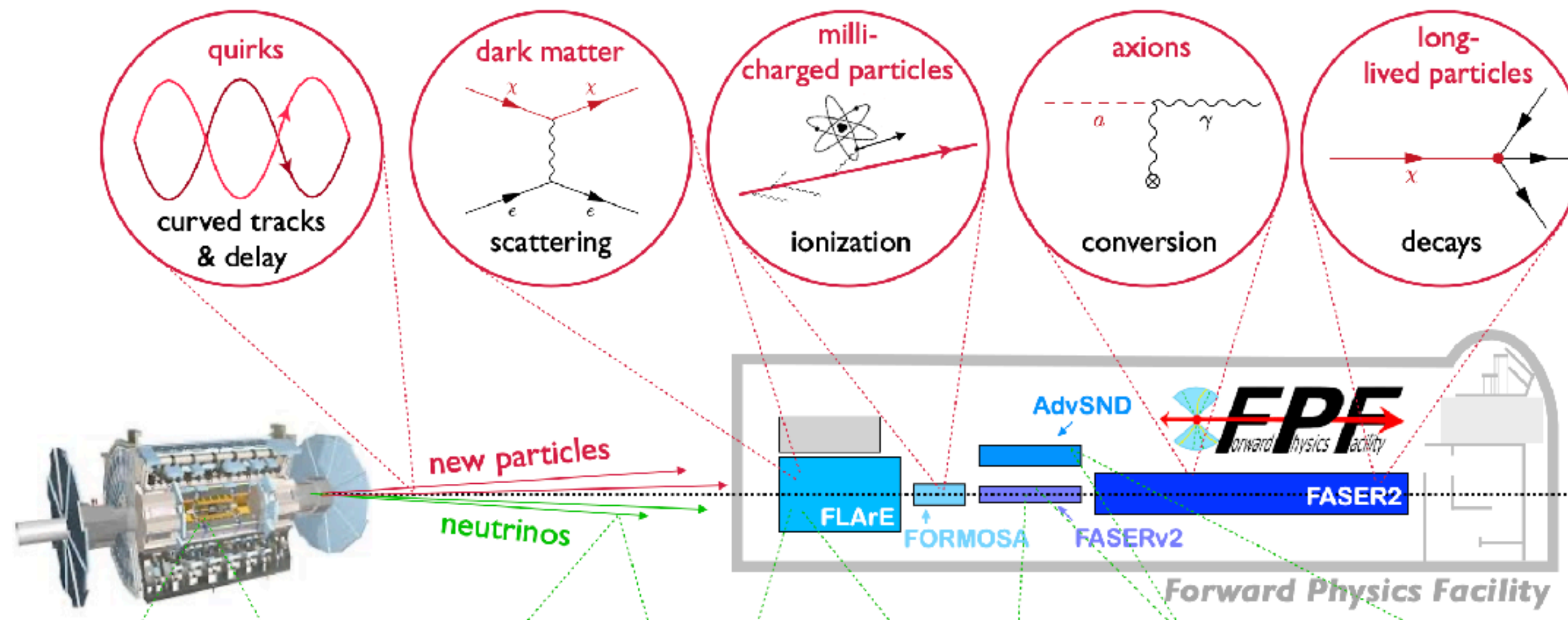
$$N = (2.2 \pm 3.1) \times 10^{-4}$$



Forward Physics Facility | BSM

► FPF Papers:

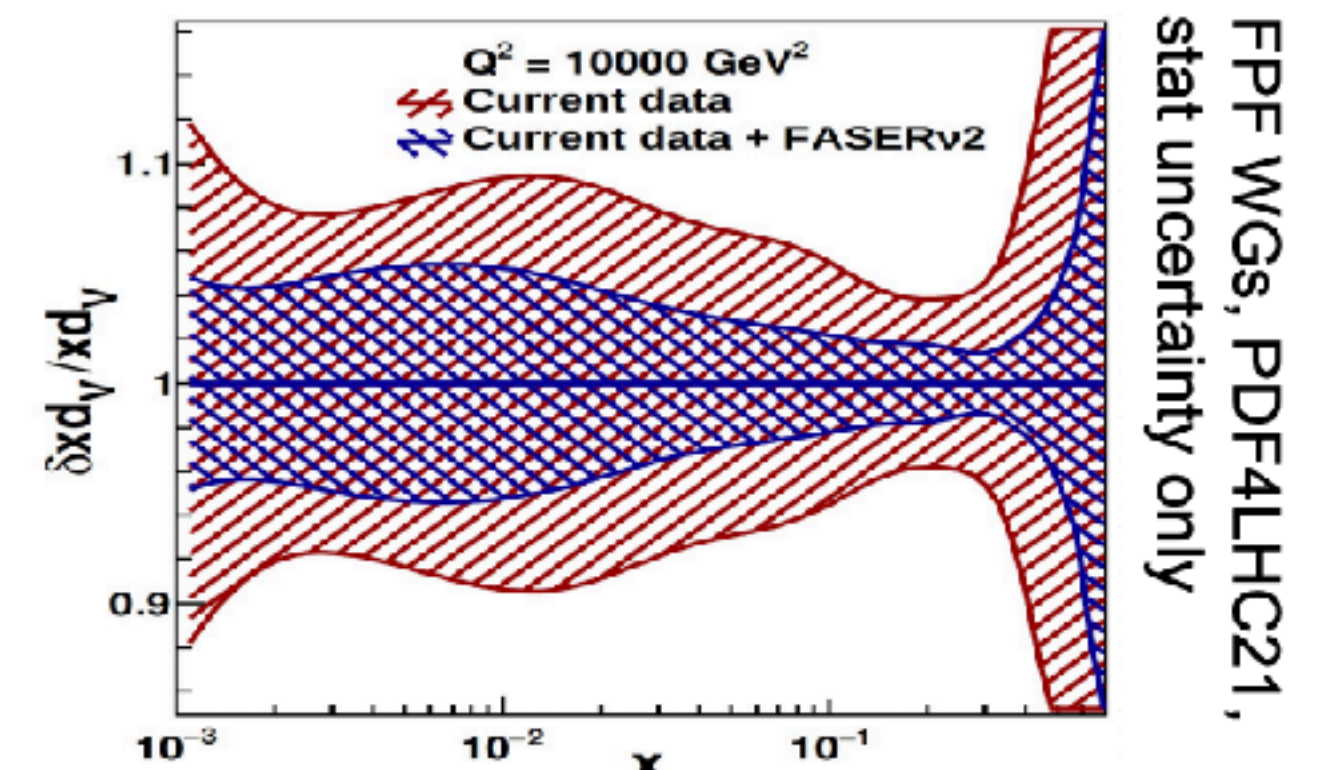
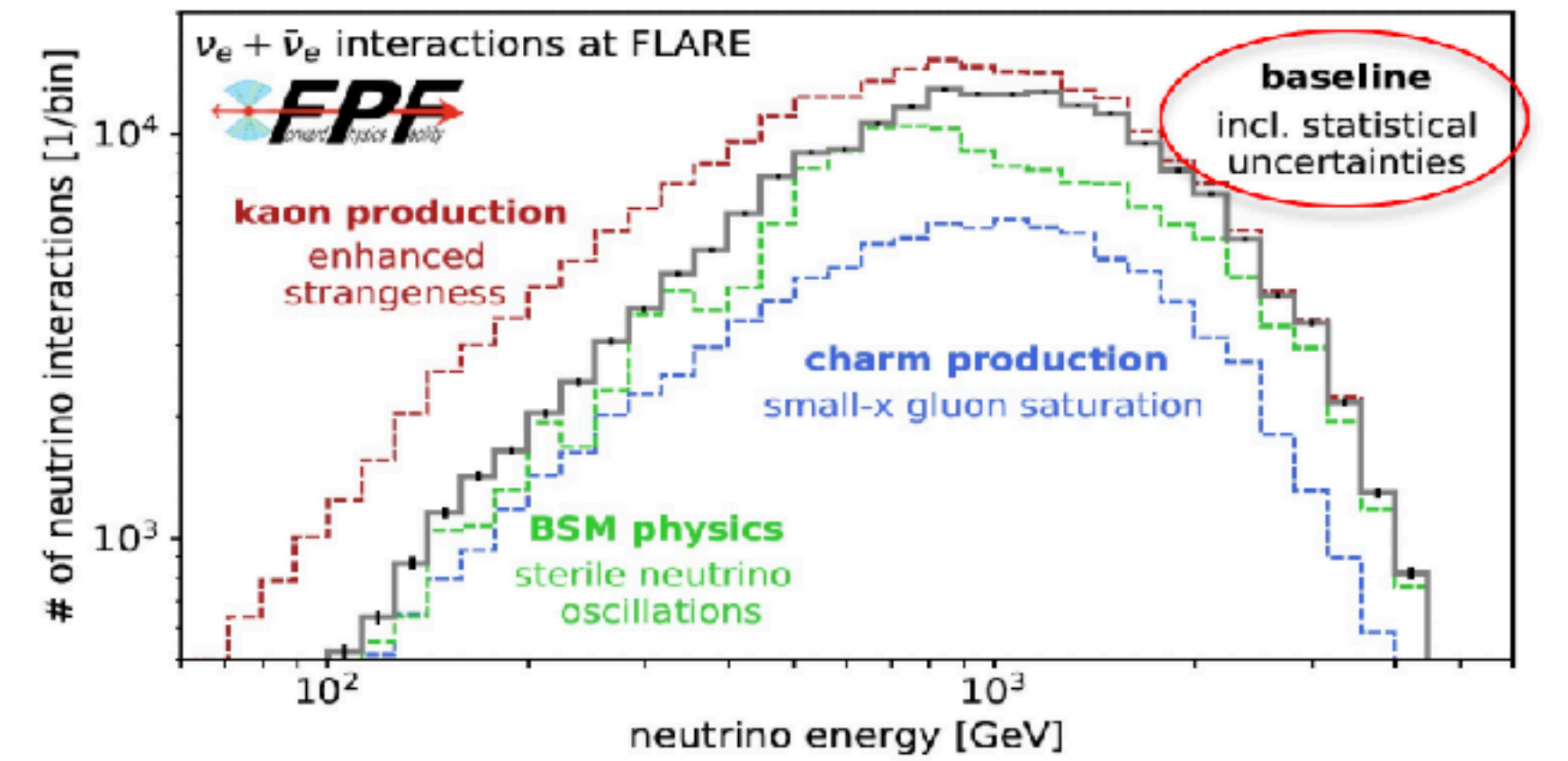
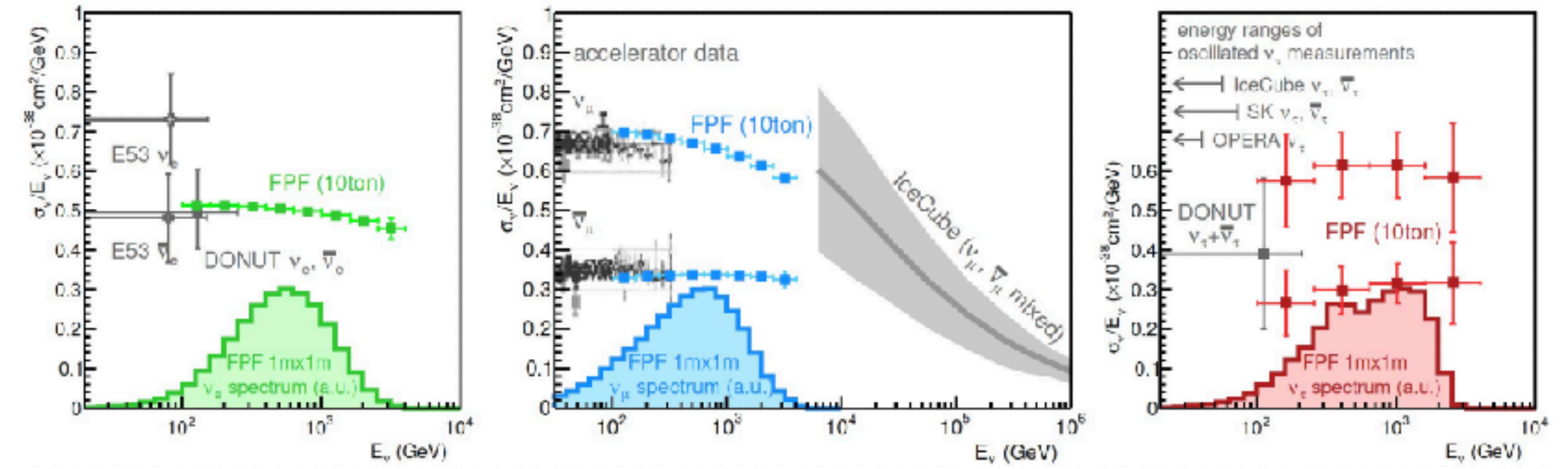
- FPF "Short" Paper: Phys. Rept. 968, 1 (2022)
- FPF White Paper: J. Phys. G (2022)



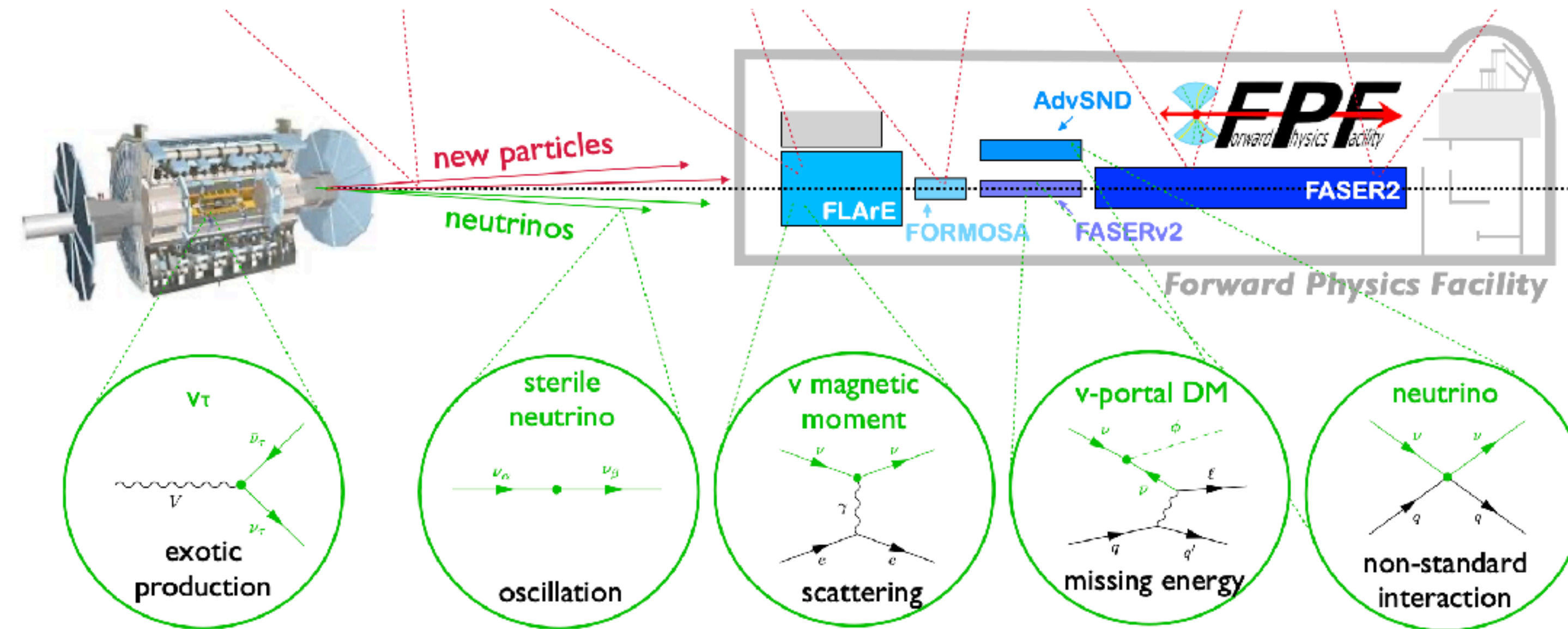
Forward Physics Facility | Neutrinos

FPF Papers:

- FPF "Short" Paper: Phys. Rept. 968, 1 (2022)
- FPF White Paper: J. Phys. G (2022)

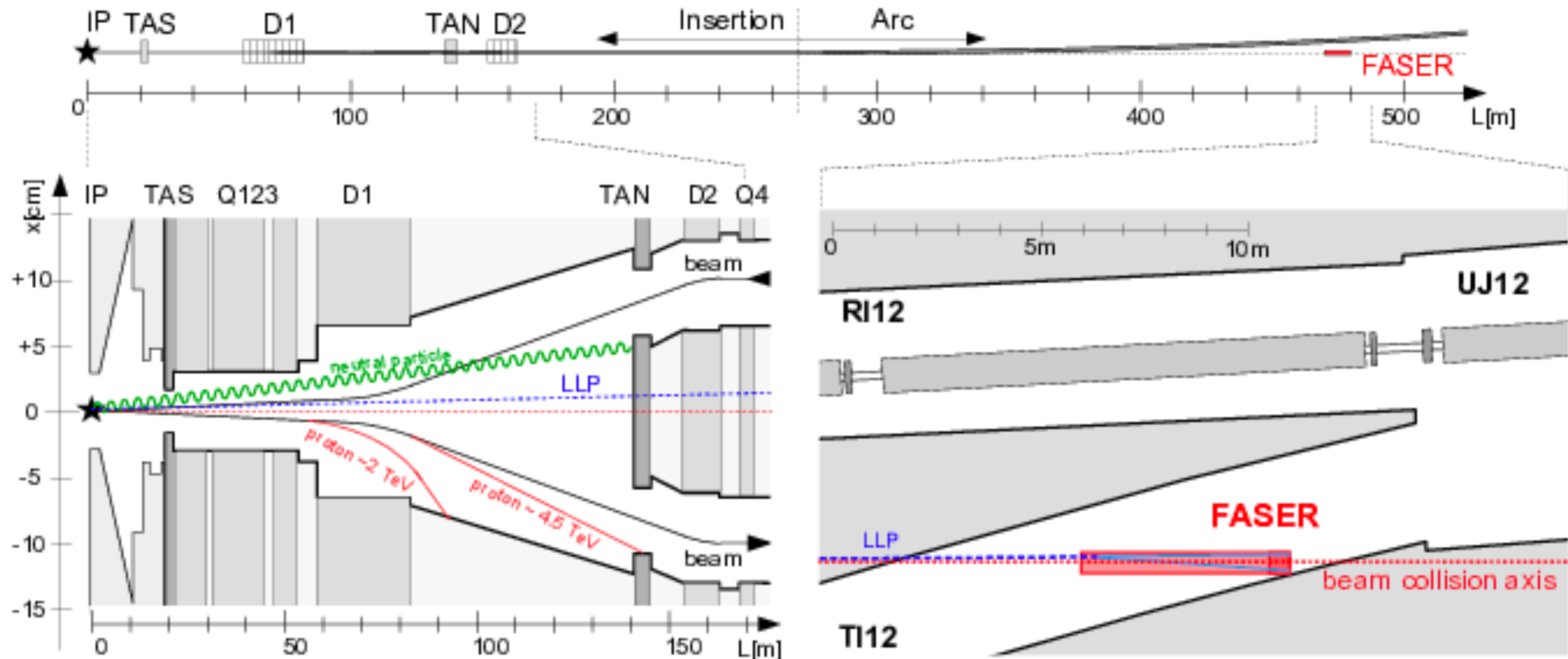


FPF WGs, PDF4LHC21, stat uncertainty only



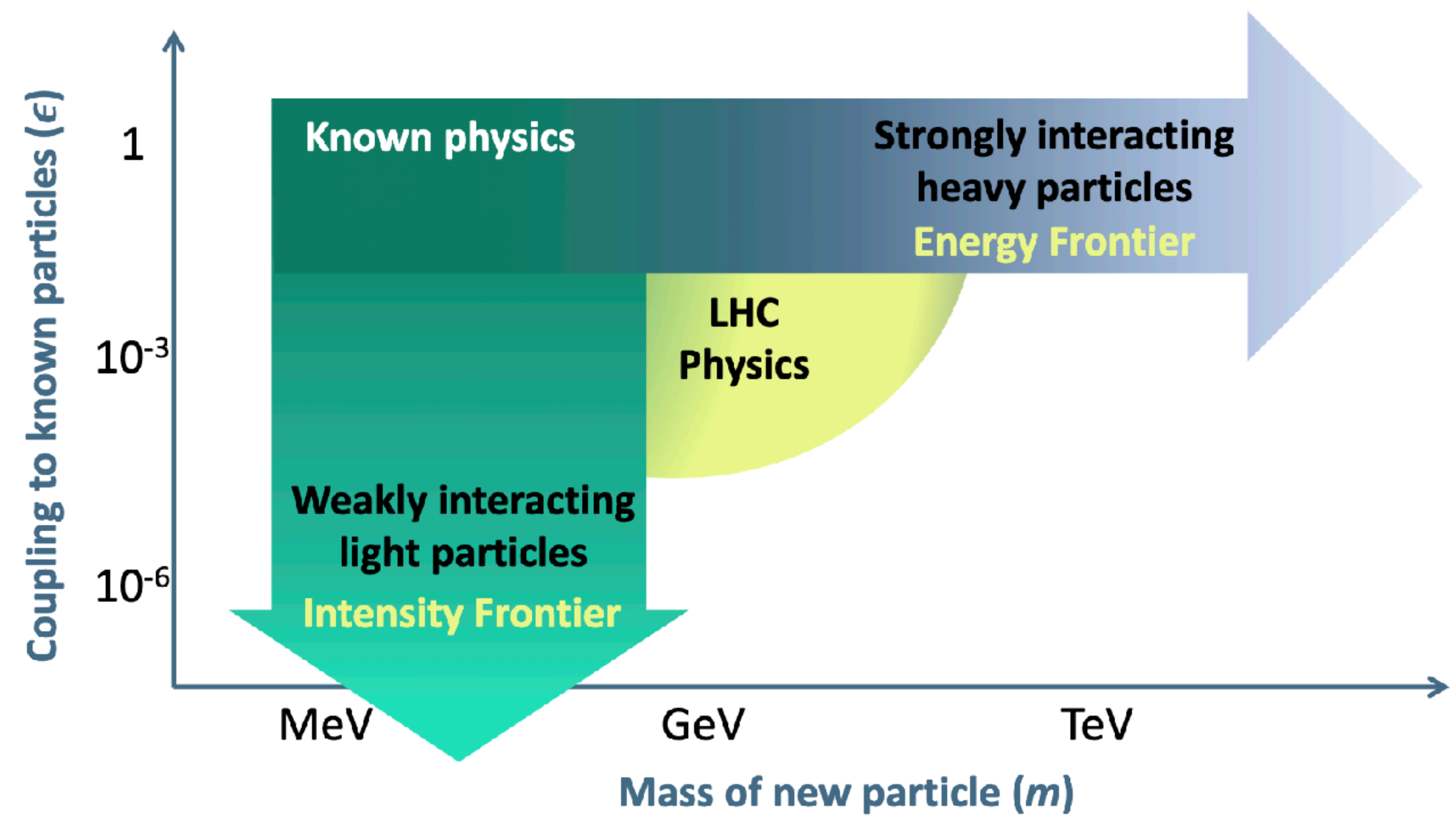
FASER Location

► A closer look at the LHC infrastructure on the line-of-sight:



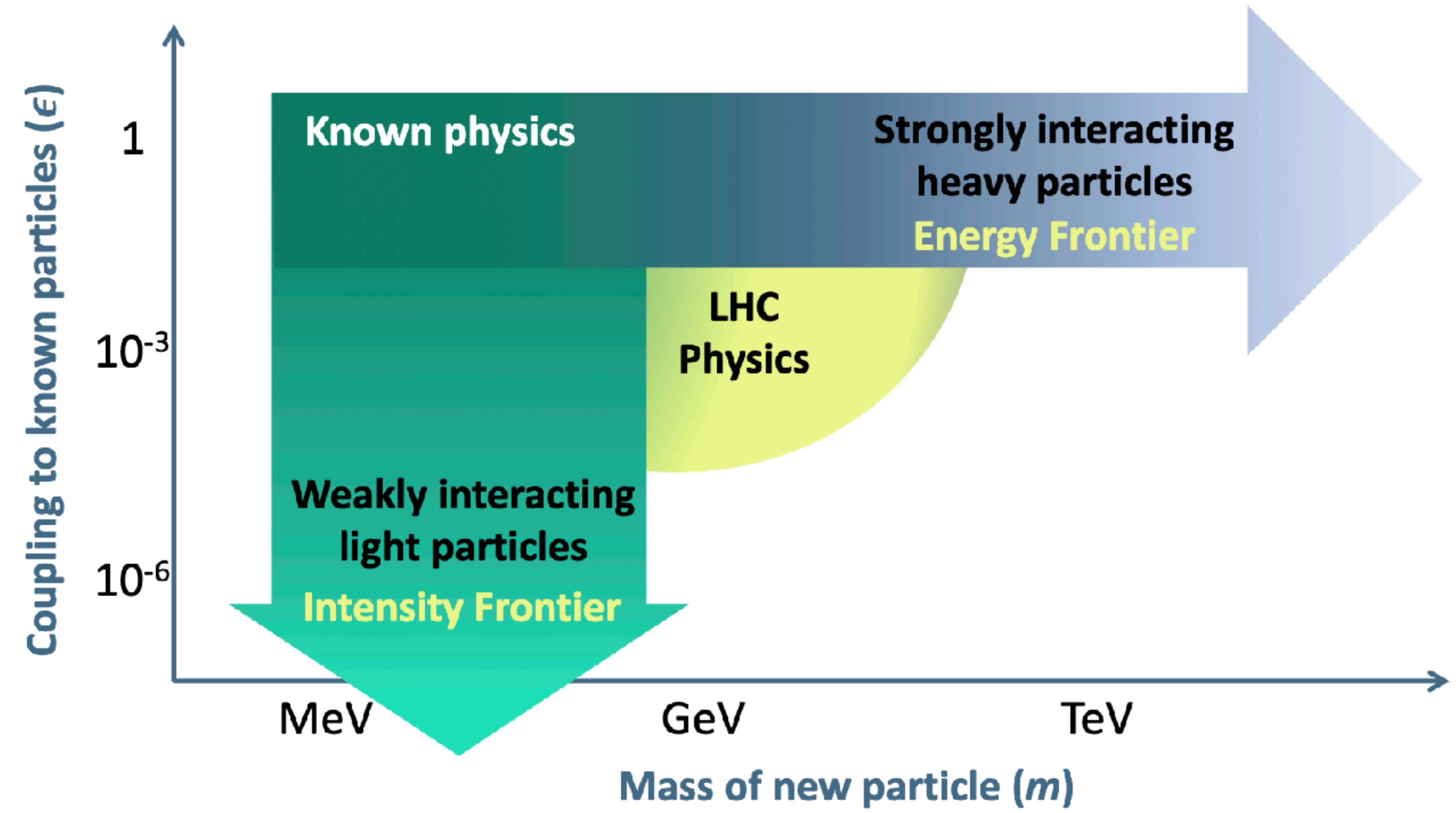
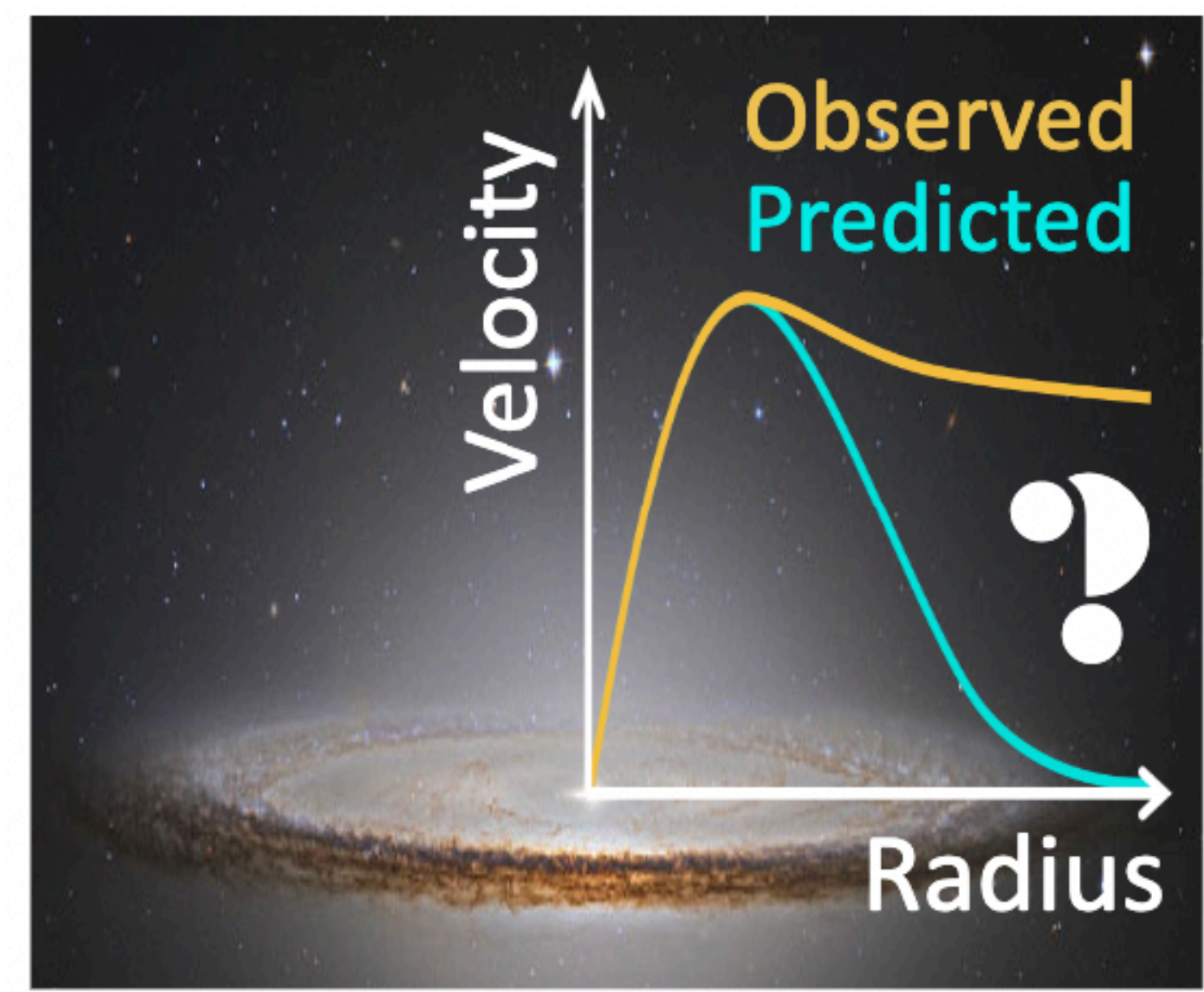
Physics Motivation

- ▶ The **LHC experiments** are producing incredible results, searching in previously unexplored phase spaces and performing increasingly precise measurements.
- ▶ But the lack of any observation of BSM physics motivates **looking elsewhere** too.



Physics Motivation

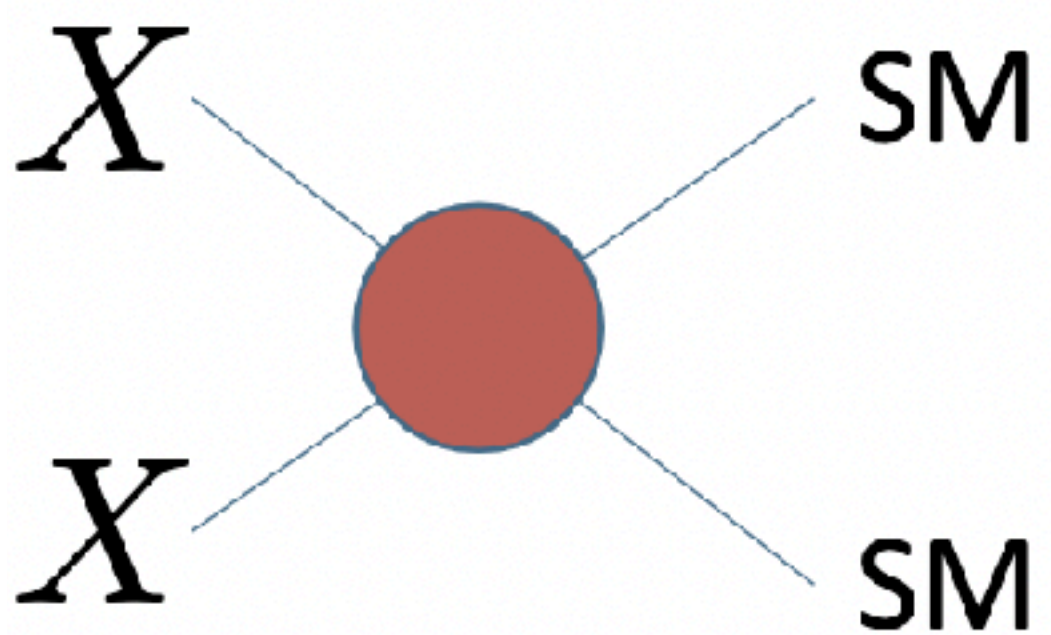
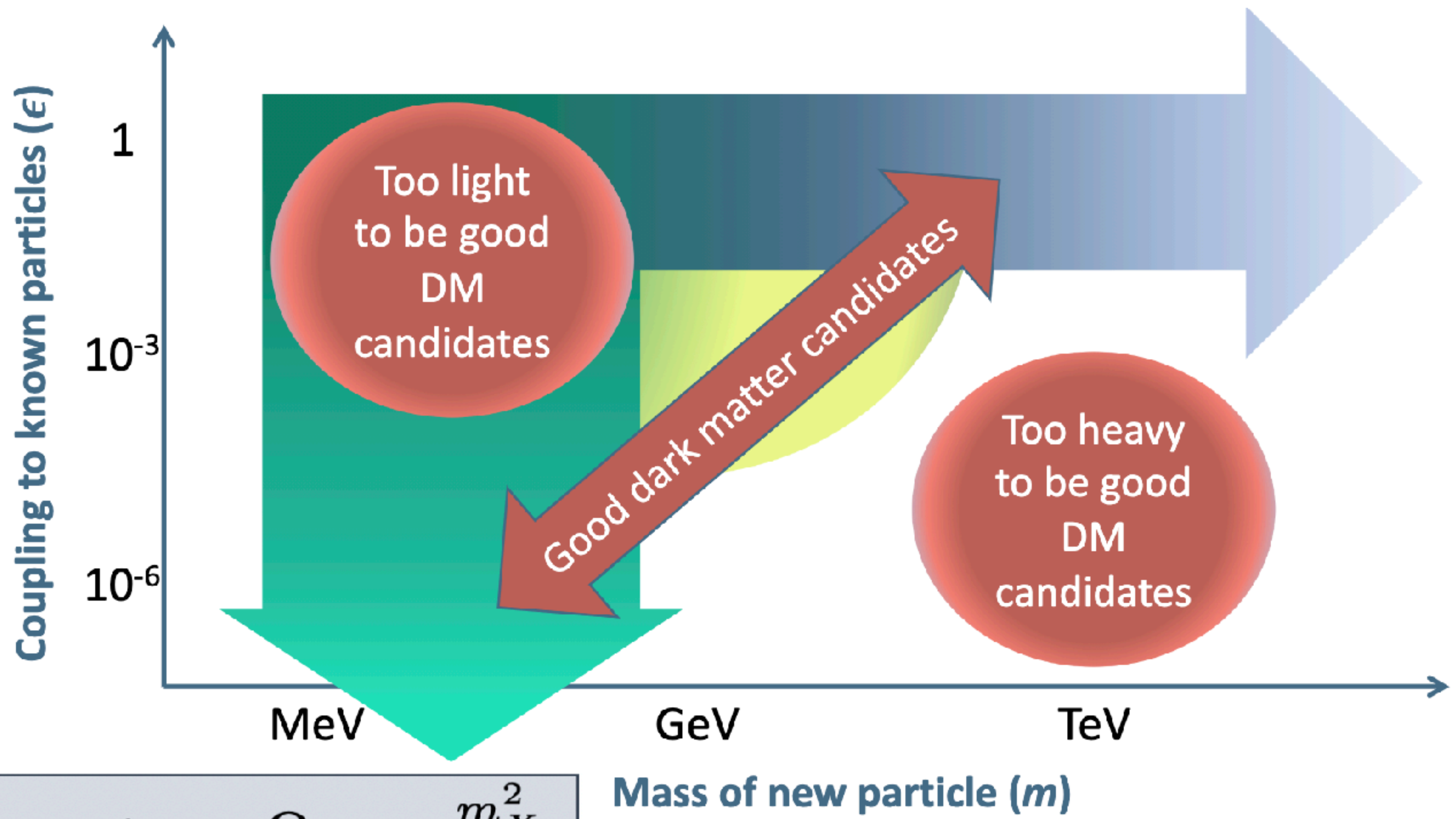
- ▶ The indirect observations of **dark matter** offers one of the most tangible indications of BSM physics and strongly motivates closer attention.



Physics Motivation

► The indirect observations of **dark matter** offers one of the most tangible indications of BSM physics and strongly motivates closer attention.

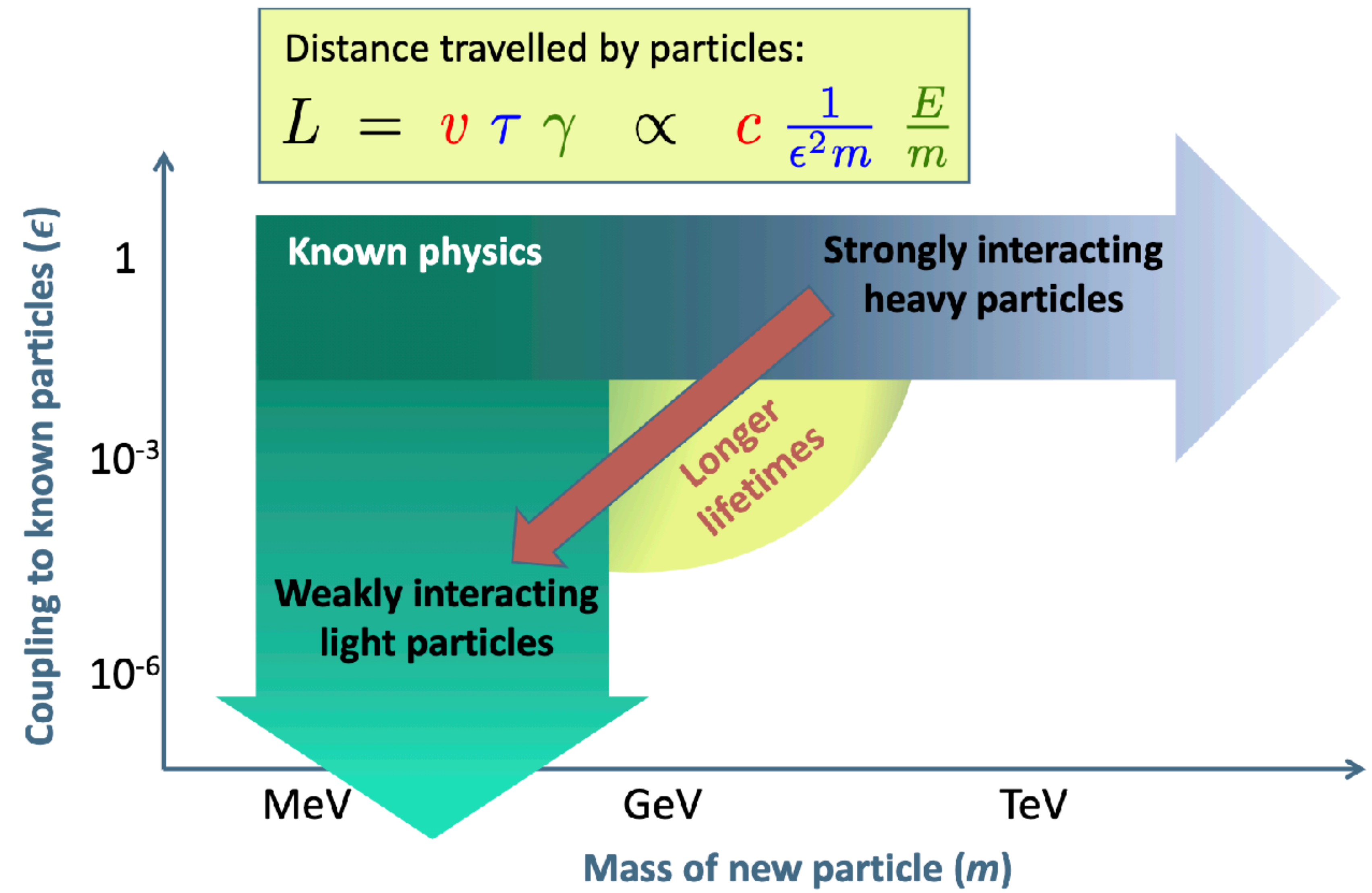
► Main region of interest is for new particles that satisfy DM relic density requirements.



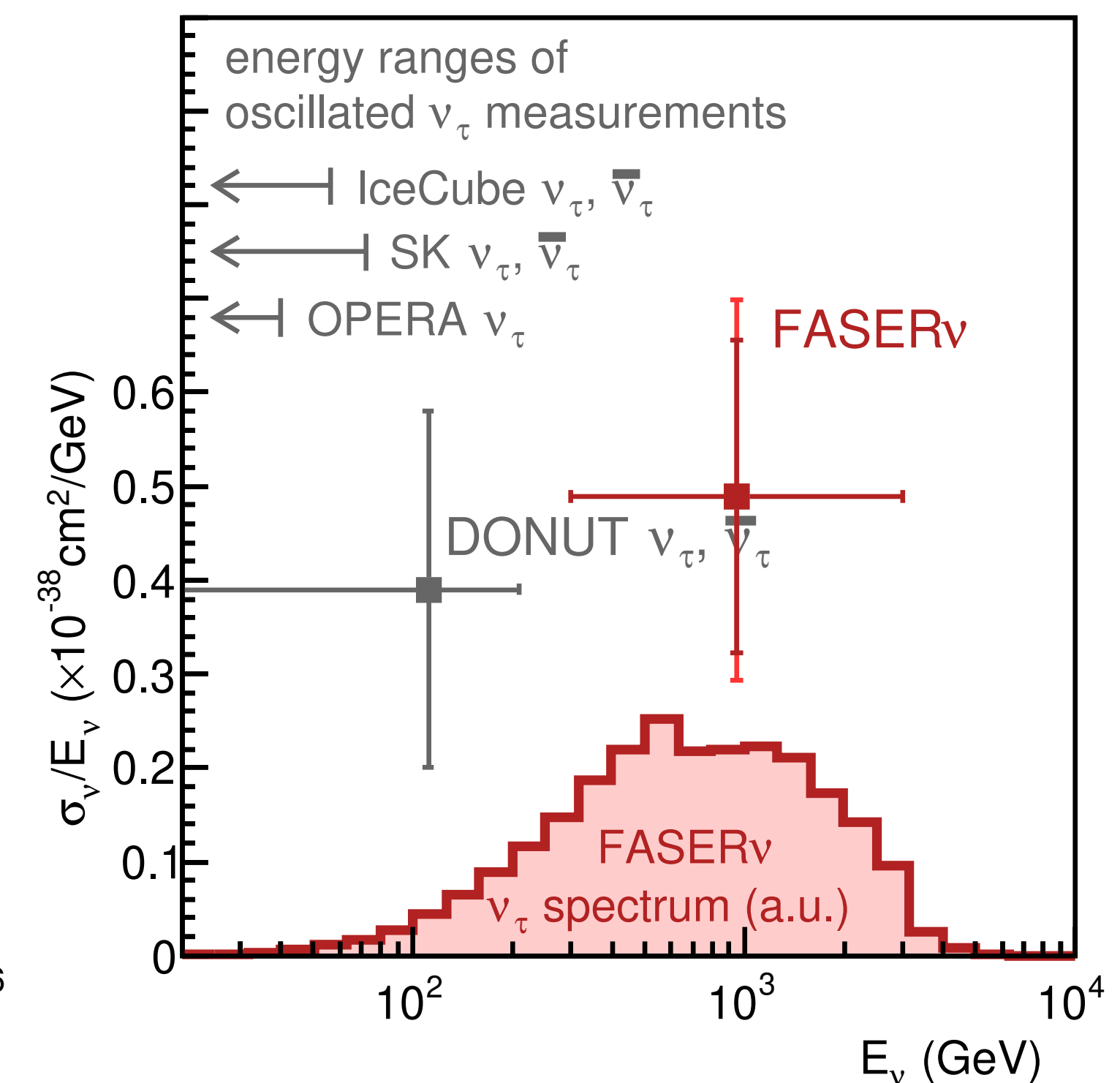
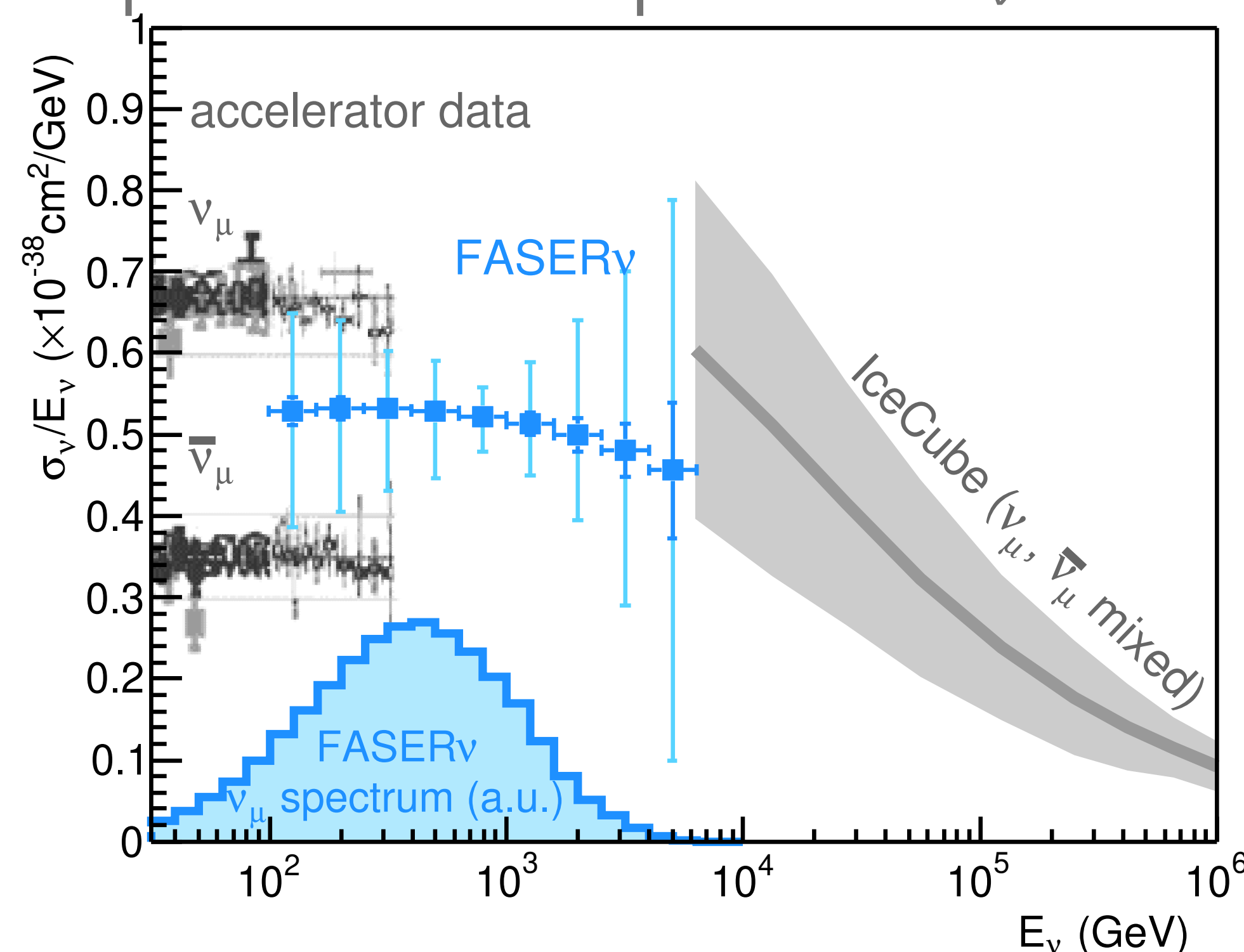
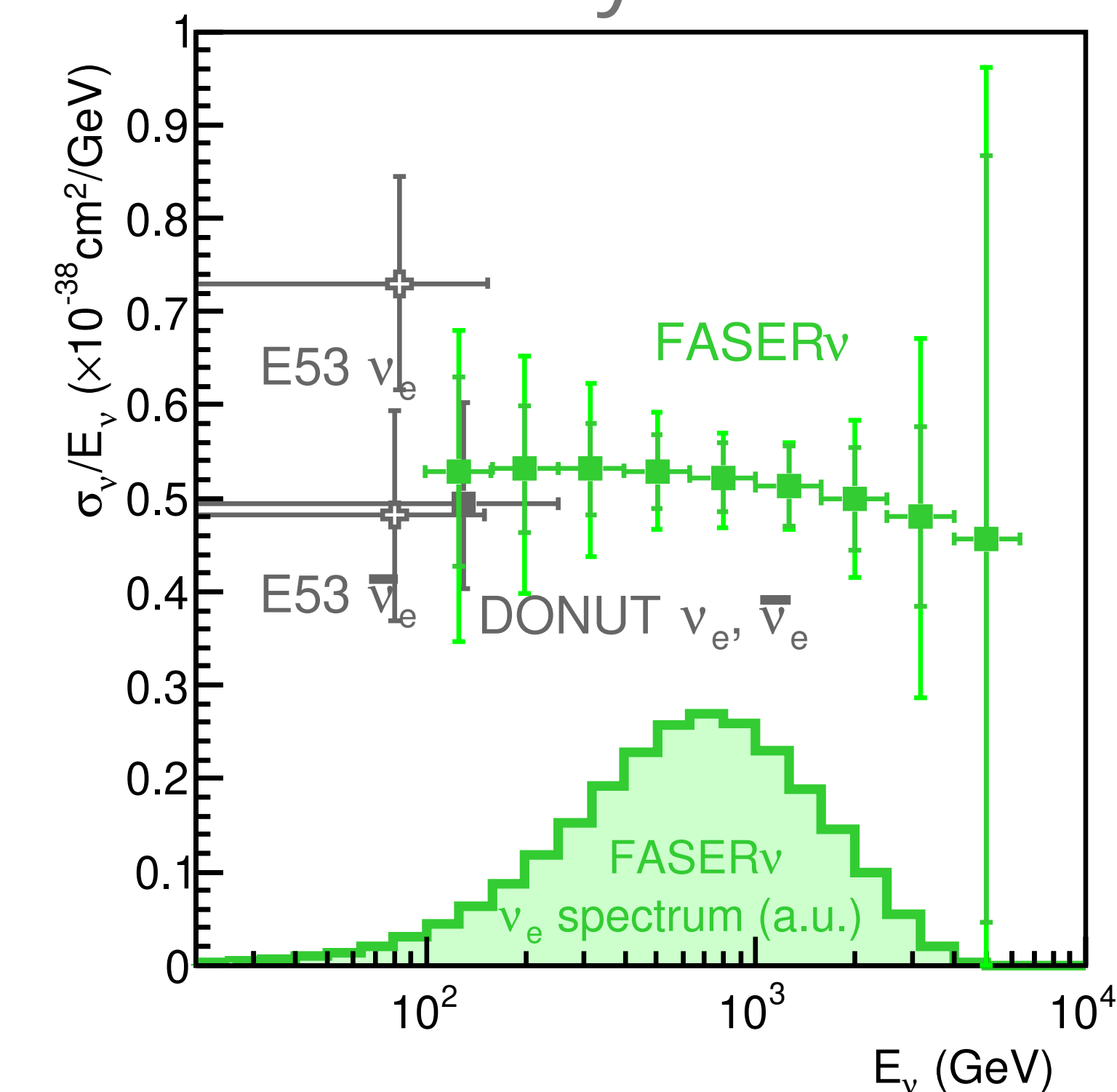
Surviving DM density: $\Omega_X \propto \frac{m_X^2}{\epsilon_X^4}$

Physics Motivation

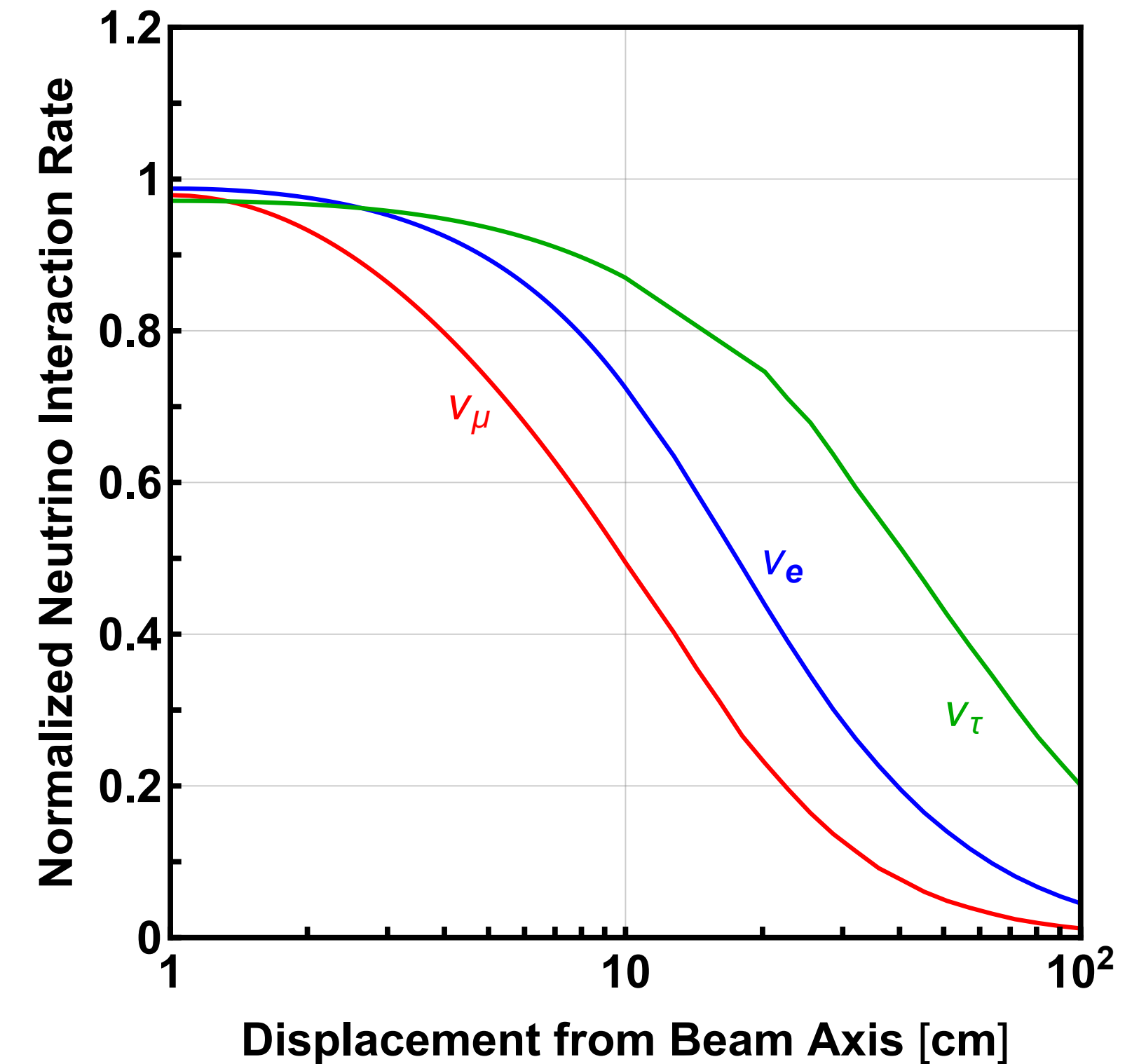
- ▶ One of the defining characteristics of weakly interacting light particles is their **long lifetime**.
- ▶ Distinct signatures
- ▶ Opportunity for exploration!



- ▶ The energy spectrum expected at FASERν is rather complementary to existing neutrino experiments
- ▶ Expected cross section sensitivity significantly extends current measurements during Run 3 (150 fb⁻¹):
- ▶ Uncertainty from neutrino production important. E_ν reco resolution ~30% (sim).



- ▶ The energy spectrum expected at FASERν is rather complementary to existing neutrino experiments
- ▶ Expected cross section sensitivity significantly extends current measurements during Run 3 (150 fb^{-1})
- ▶ Being located on line-of-sight FASERν is able to observe a maximum rate of all neutrino flavours:



FASERv | Rich neutrino physics program

► BSM physics

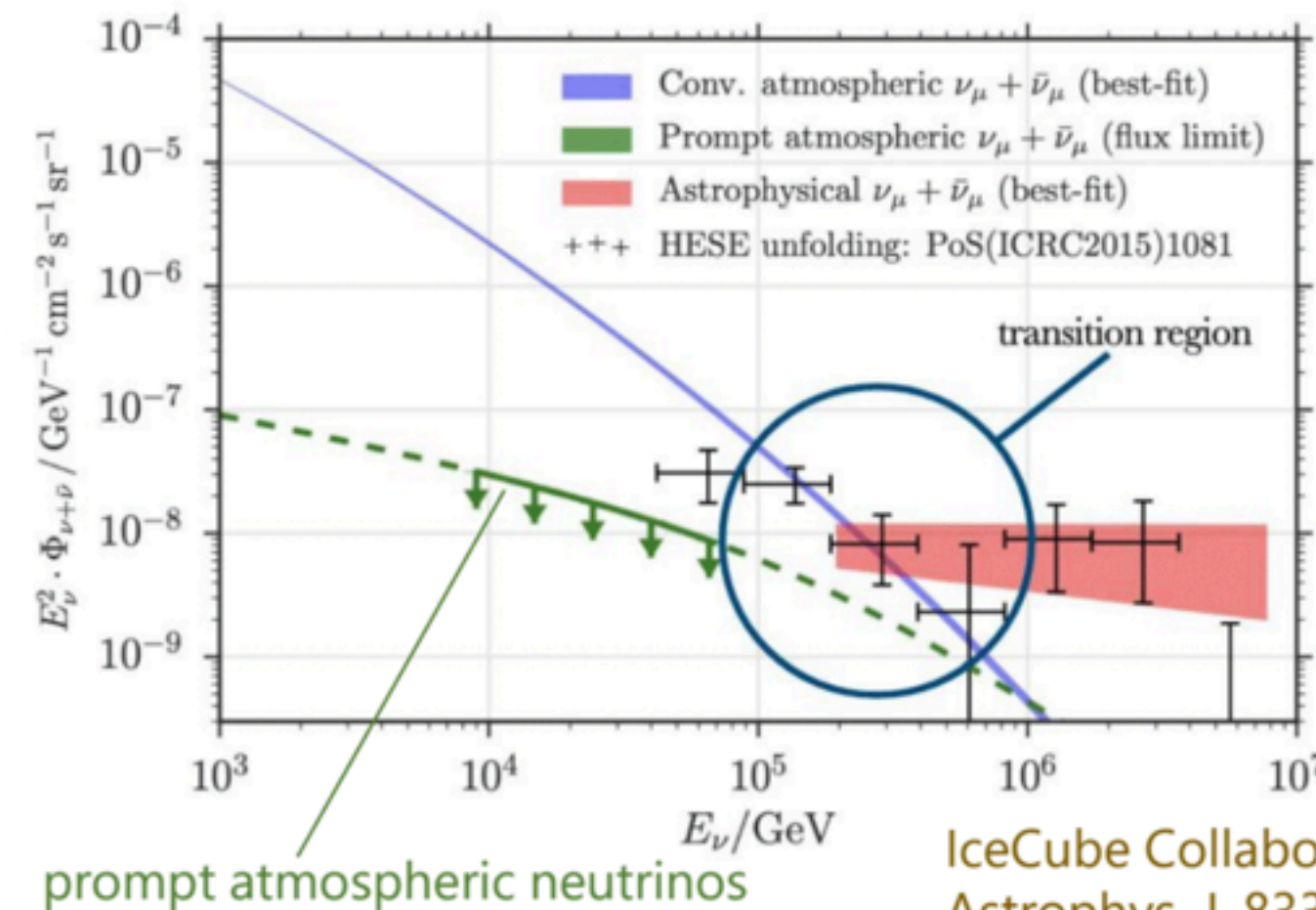
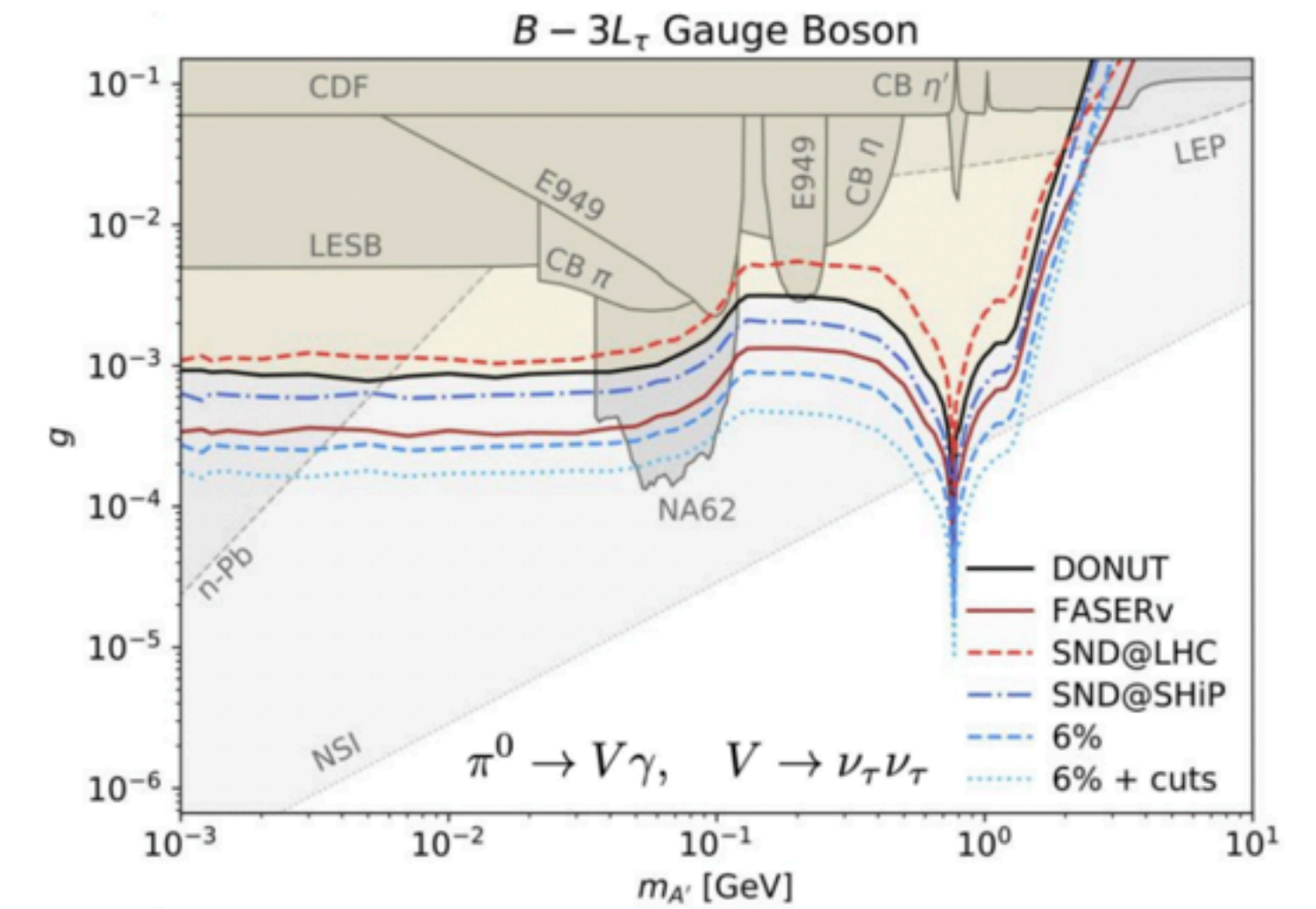
- New light weakly coupled gauge boson ($\rightarrow \nu_\tau$) could enhance ν_τ flux.
- Sterile neutrinos with mass ~ 40 eV can cause oscillations at FASER

► QCD

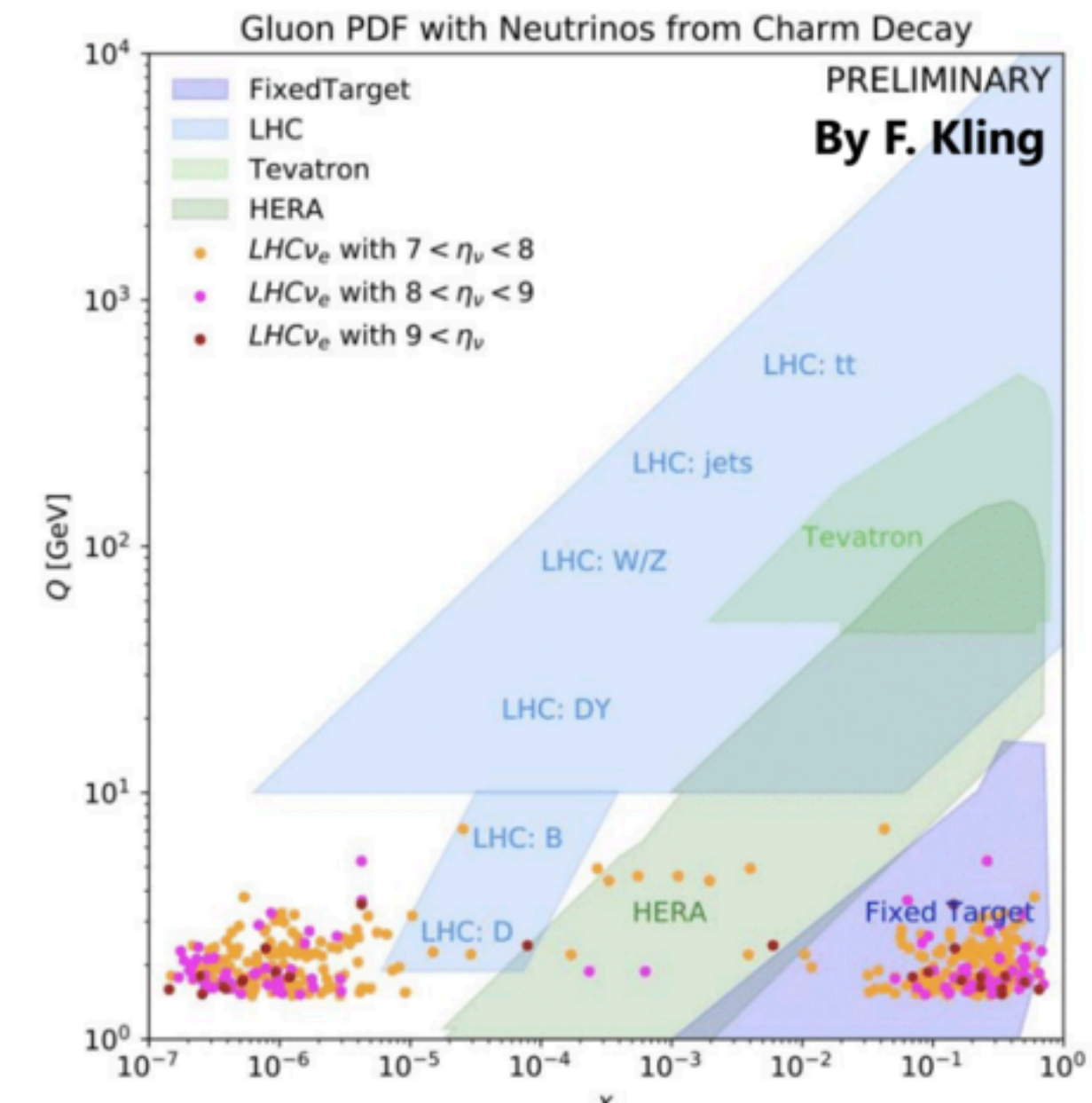
- FASER's neutrino flux measurements will provide novel complimentary constraints that can be used to validate/improve MC generator very forward particle production.
- Neutrinos from charm decay could allow to test transition to small-x factorisation, constrain low-x gluon PDF and probe intrinsic charm

► Cosmic rays and neutrinos

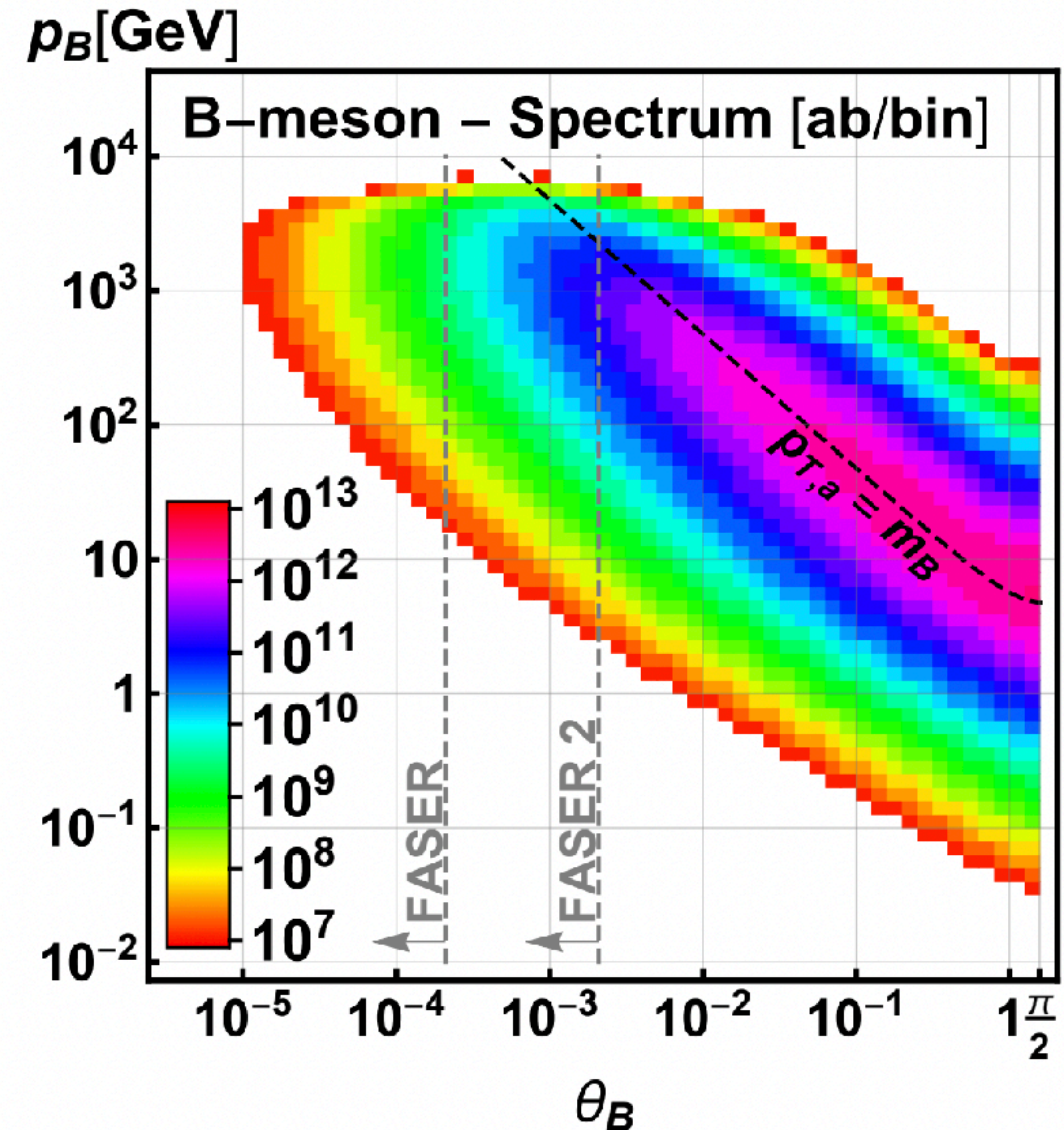
- IceCube needs measurements of high energy and large rapidity charm for precise measurements of cosmic neutrino flux.
- Direct measurement of prompt neutrino production at FASER would provide important data for current & future neutrino telescopes



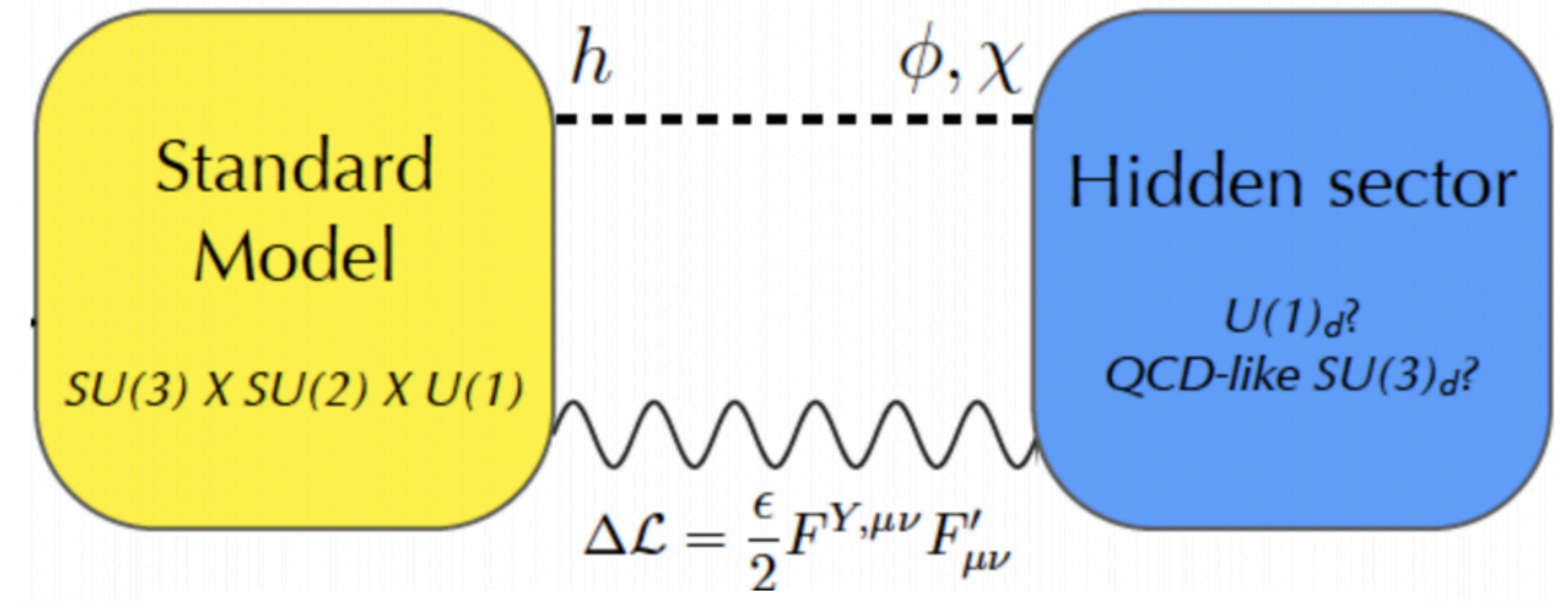
IceCube Collaboration,
Astrophys. J. 833 (2016)



- ▶ Increasing detector radius to 1m would allow sensitivity to new physics produced in heavy meson (B, D) decays increasing the physics case beyond just the increased luminosity.

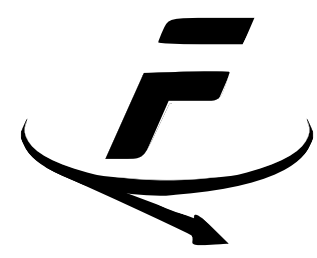


- ▶ Hidden sector physics:
 - ▶ New mediating particles, couplings to SM via mixing with SM “portal” operator
 - ▶ Related to nature of DM (mediator or candidate), baryogenesis, neutrino oscillations...
 - ▶ Can possibly resolve low-energy experiment anomalies (muon g-2, proton size, Be8)
 - ▶ Typically long-lived particles (LLPs) that travel macroscopic distances before decaying to SM particles



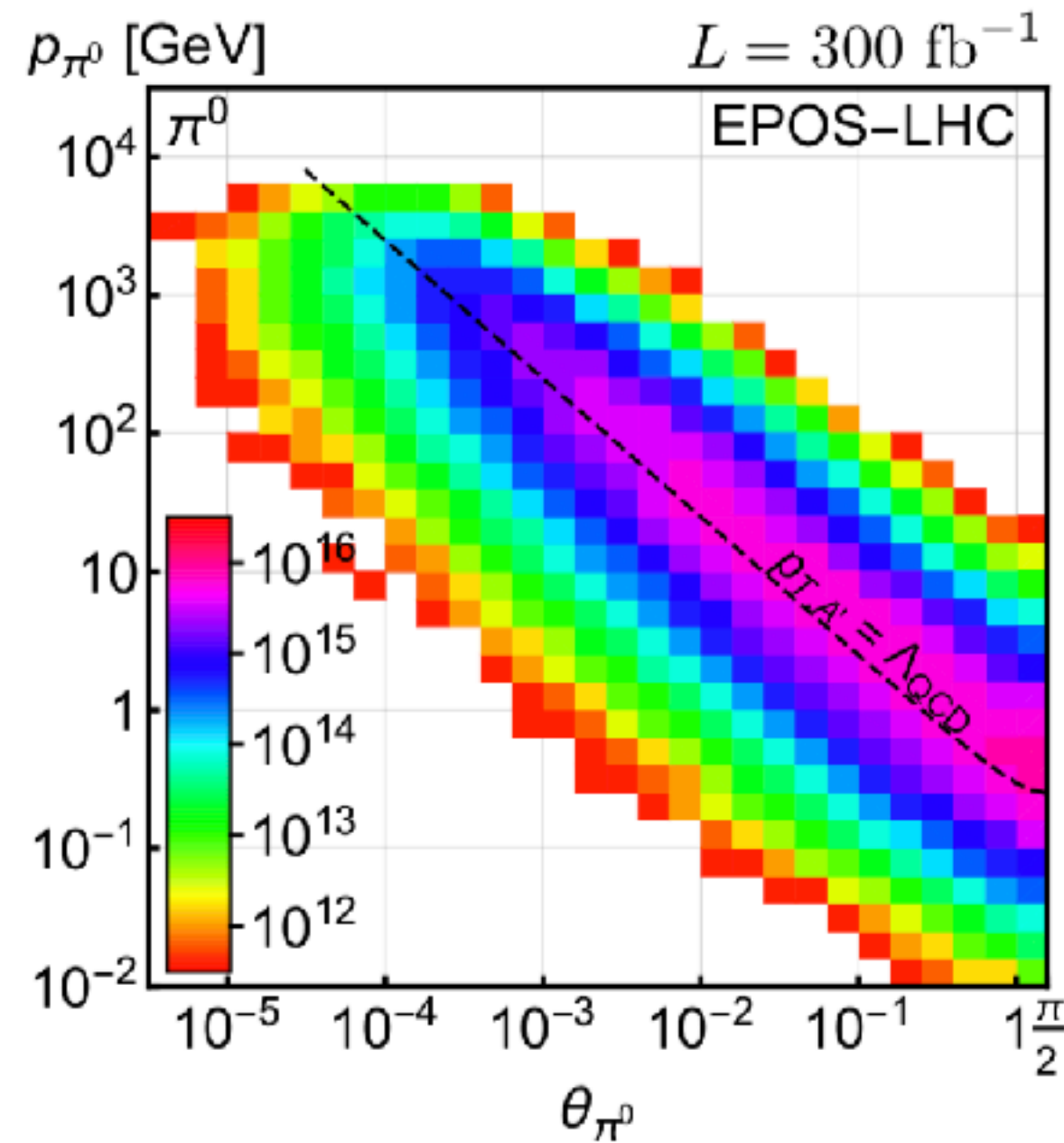
$$\mathcal{L}_{\text{portal}} = \sum O_{\text{SM}} \times O_{\text{DS}}$$

Portal	Coupling
Dark Photon, A_μ	$-\frac{\epsilon}{2 \cos \theta_W} F'_{\mu\nu} B^{\mu\nu}$
Dark Higgs, S	$(\mu S + \lambda S^2) H^\dagger H$
Axion, a	$\frac{a}{f_a} F_{\mu\nu} \tilde{F}^{\mu\nu}, \frac{a}{f_a} G_{i,\mu\nu} \tilde{G}_i^{\mu\nu}, \frac{\partial_\mu a}{f_a} \bar{\psi} \gamma^\mu \gamma^5 \psi$
Sterile Neutrino, N	$y_N L H N$



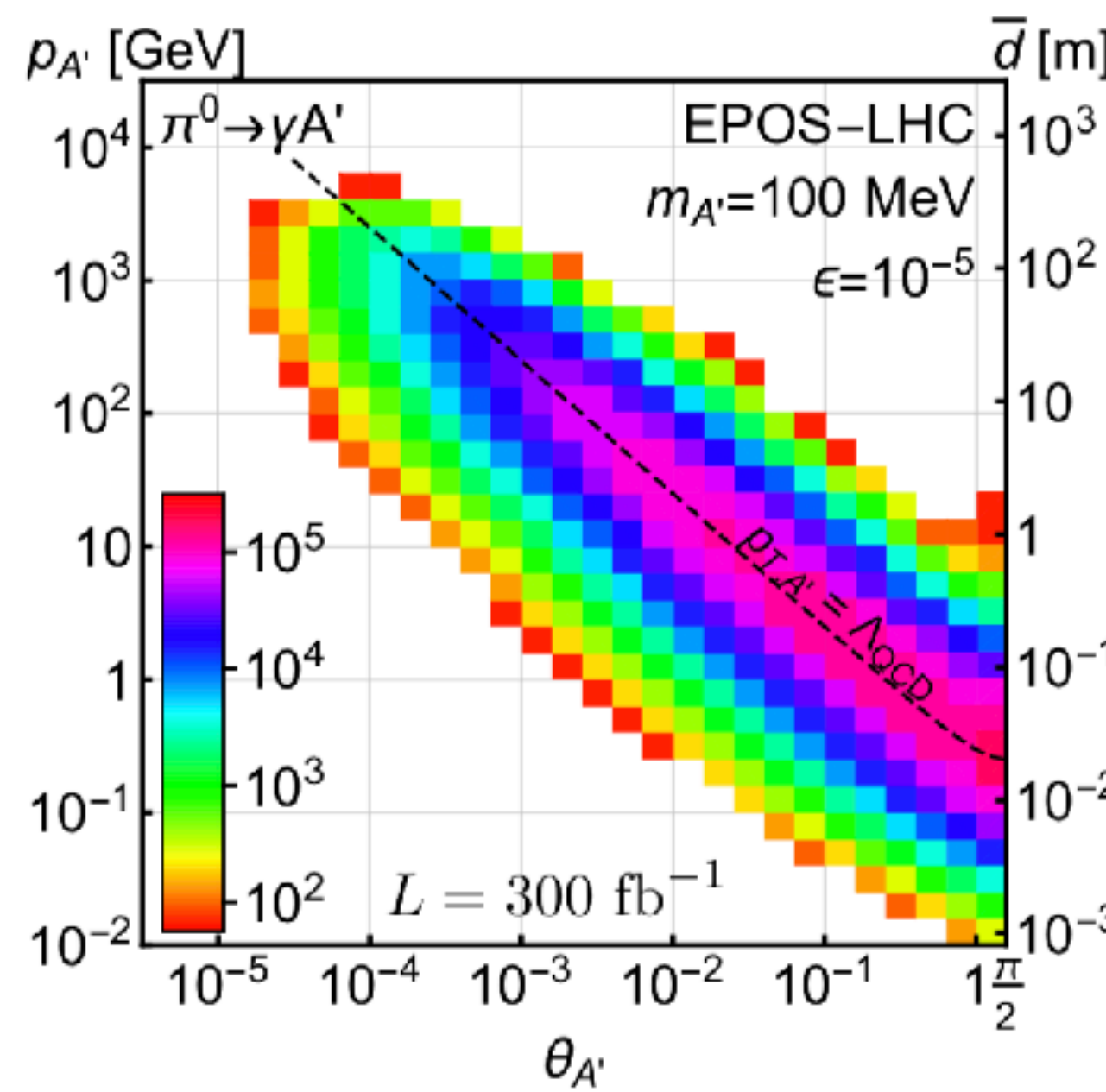
Overview | Dark photons

Pions at IP



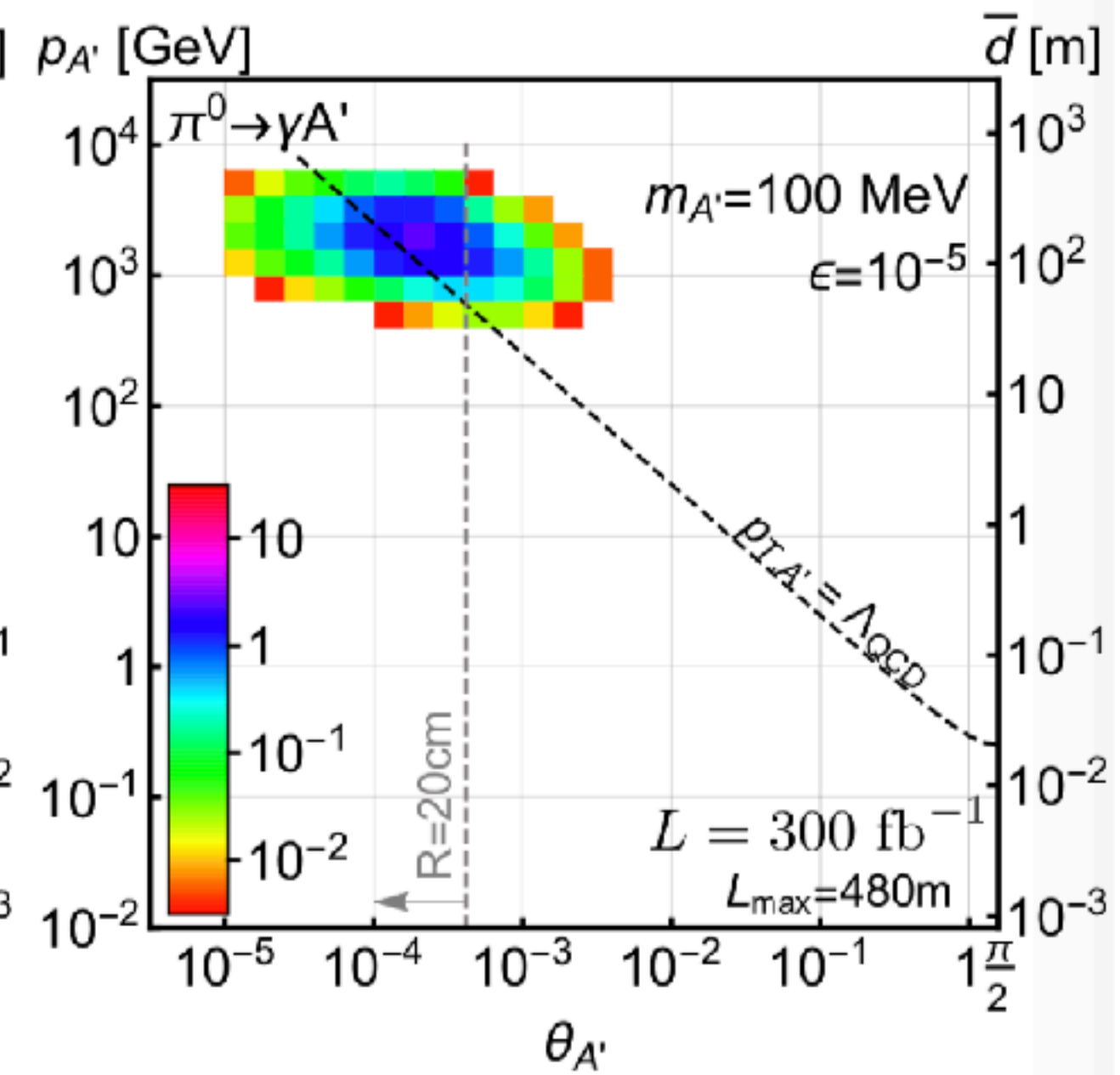
- dedicated hadronic interaction models, grounded on LHC data
- production peaks at $p_T \sim \Lambda_{QCD}$
- enormous event rates $N \sim 10^{15}$ per bin

A' at IP



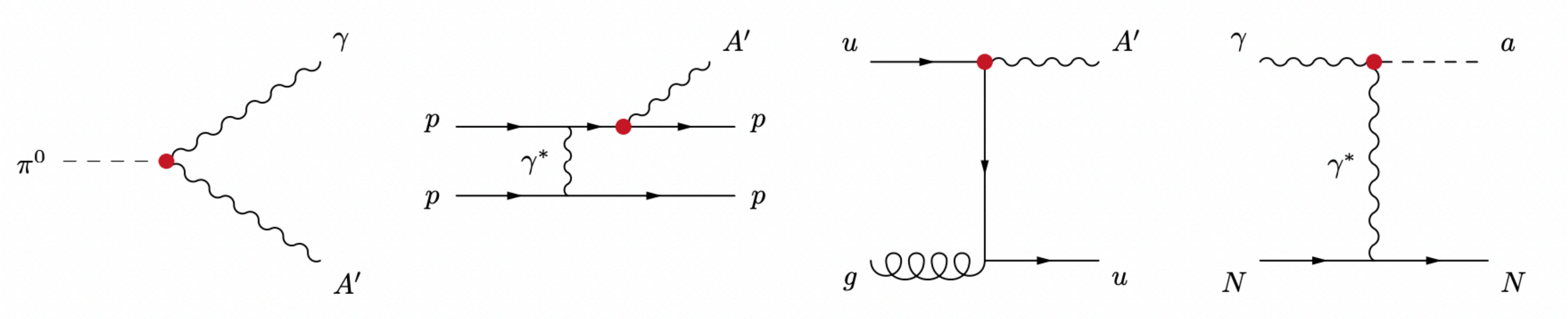
- production peaks at $p_T \sim \Lambda_{QCD}$
- rates highly suppressed by $\epsilon^2 \sim 10^{-10}$
- still rates $N \sim 10^5$ per bin: LHC could be dark a photon factory

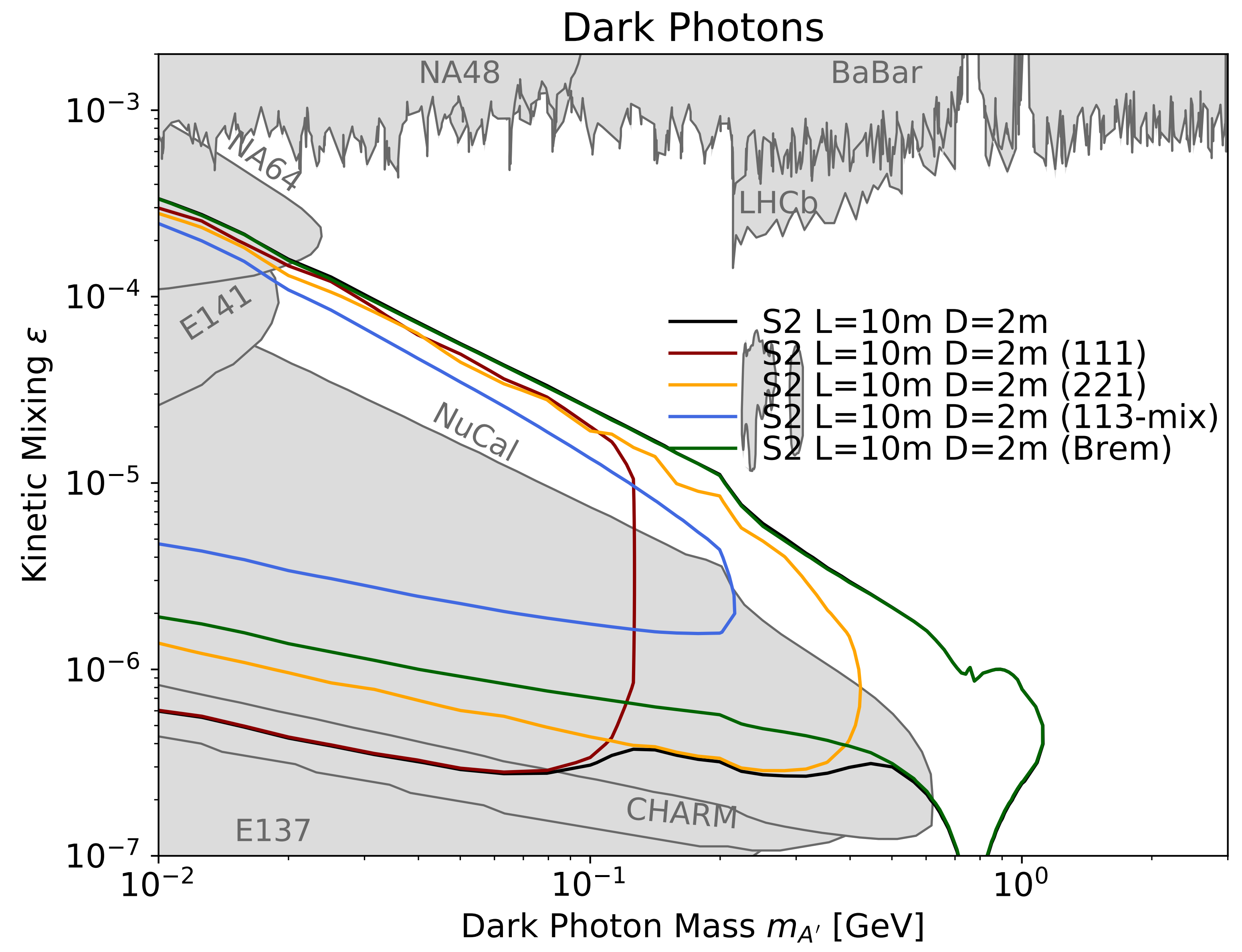
A' decay at FASER

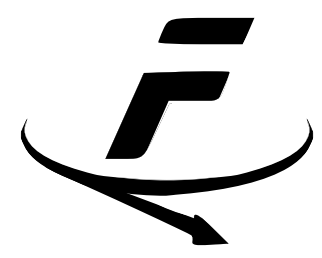


- only highly boosted $\sim \text{TeV } A'$ arrive at FASER
- rates suppressed by decay requirements
- still rates $N \sim 100$ signal events within 20cm of beam collision axis

Overview | LLP production modes



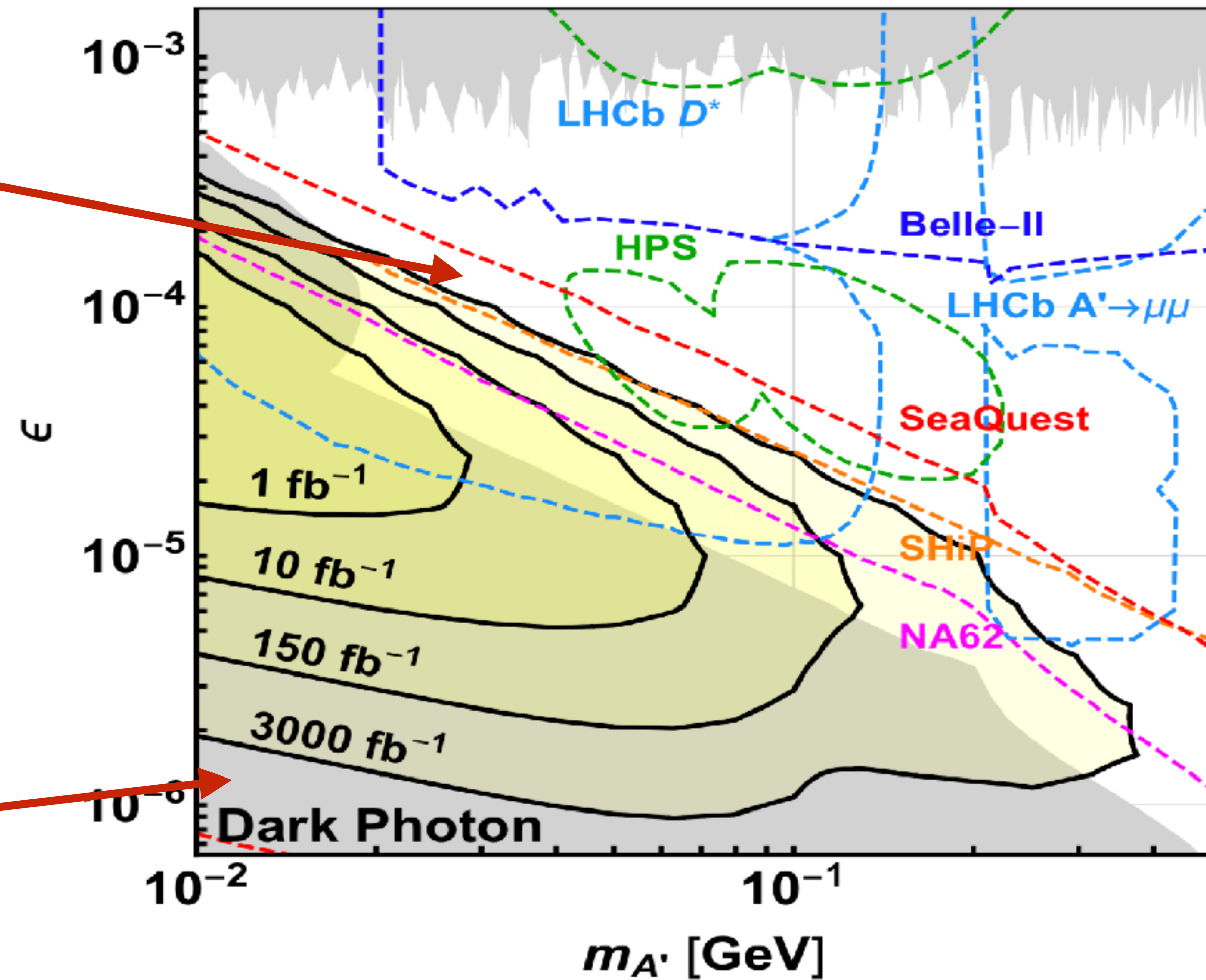




Overview | Dark photon reach

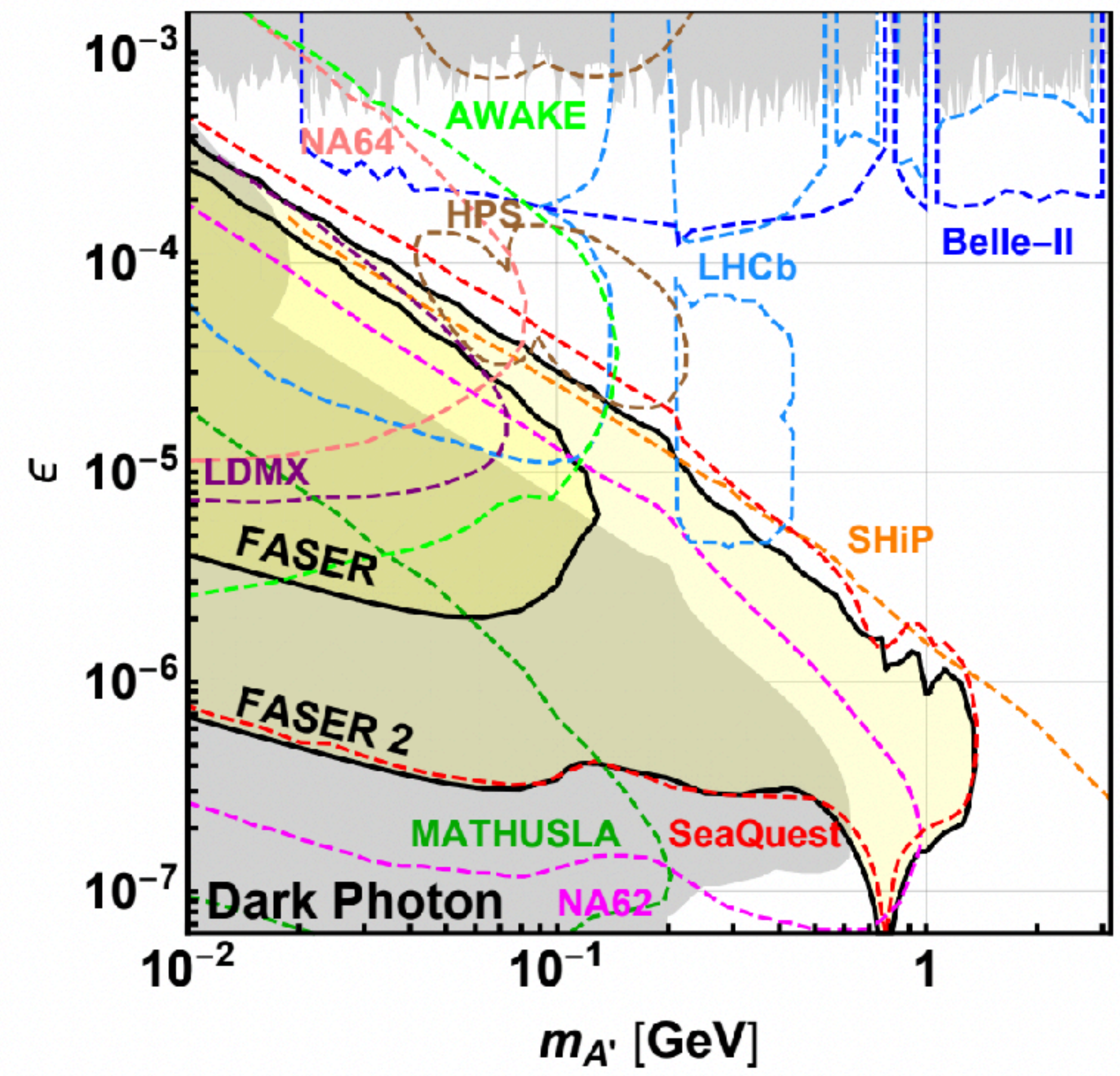
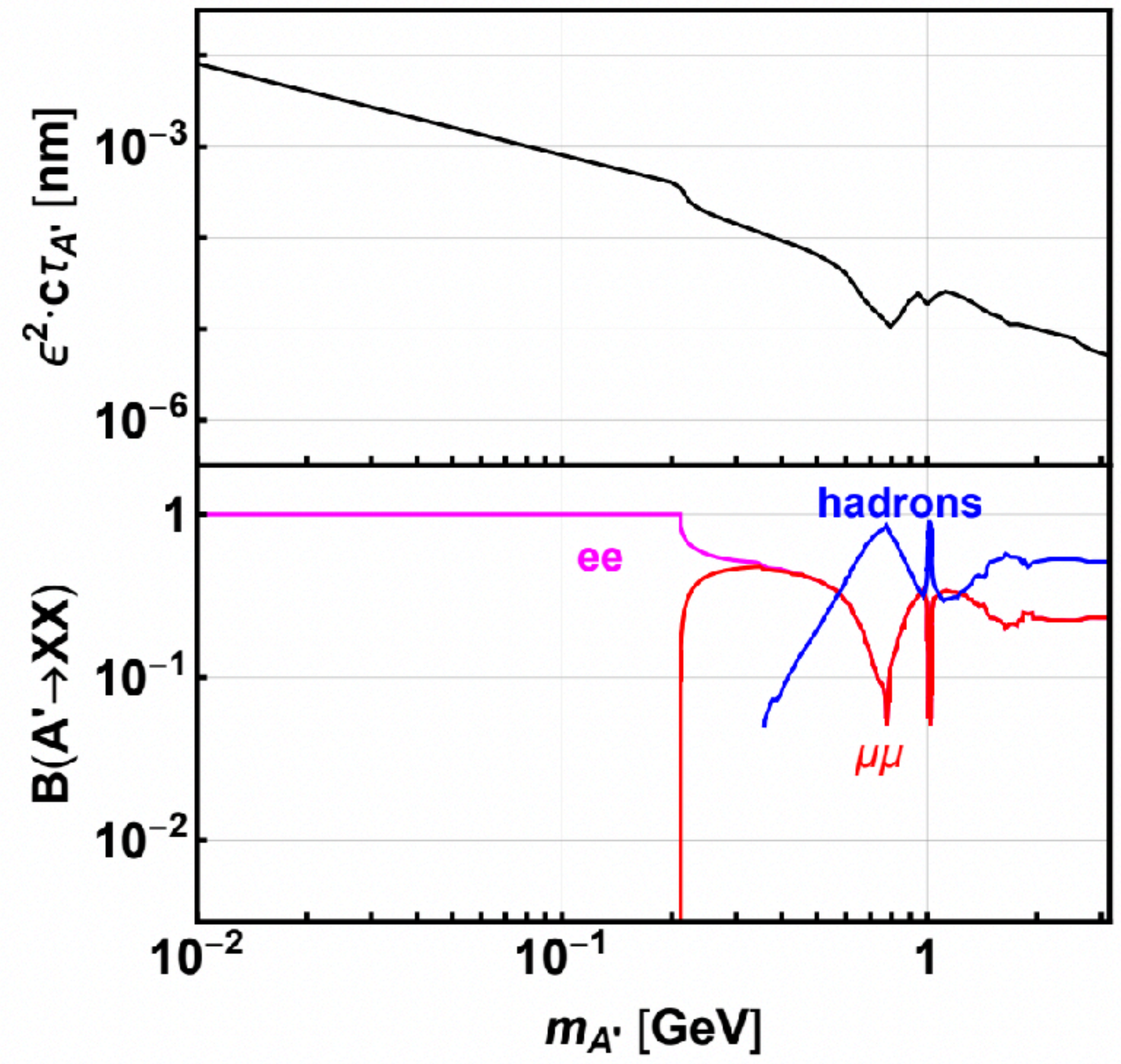
For lower lifetime the number of signal events becomes exponentially suppressed once the A' decay length drops below the distance to the detector

Combining dependence in both production rate and decay width, total number of signal events in the detector scales as ϵ^4

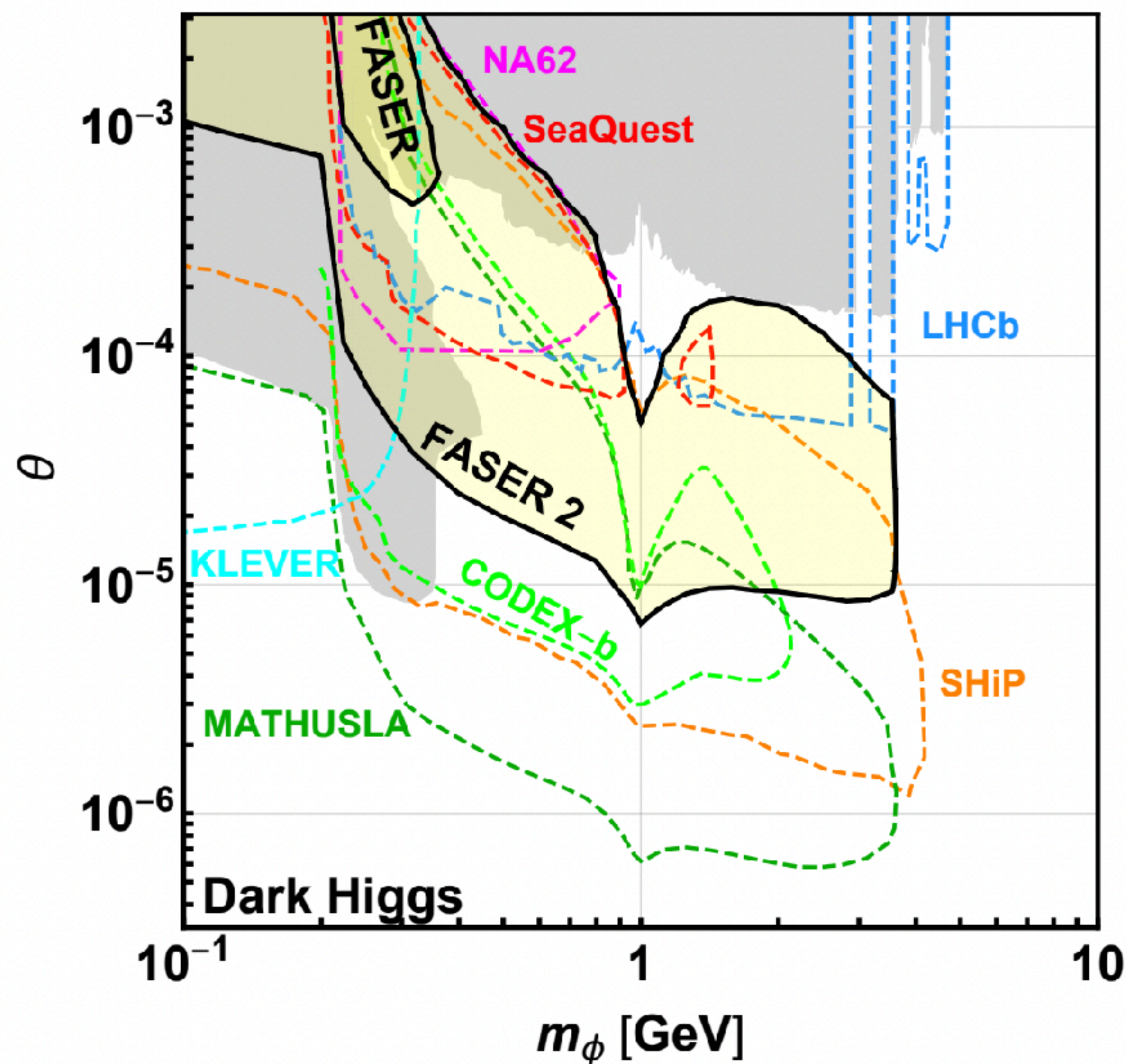
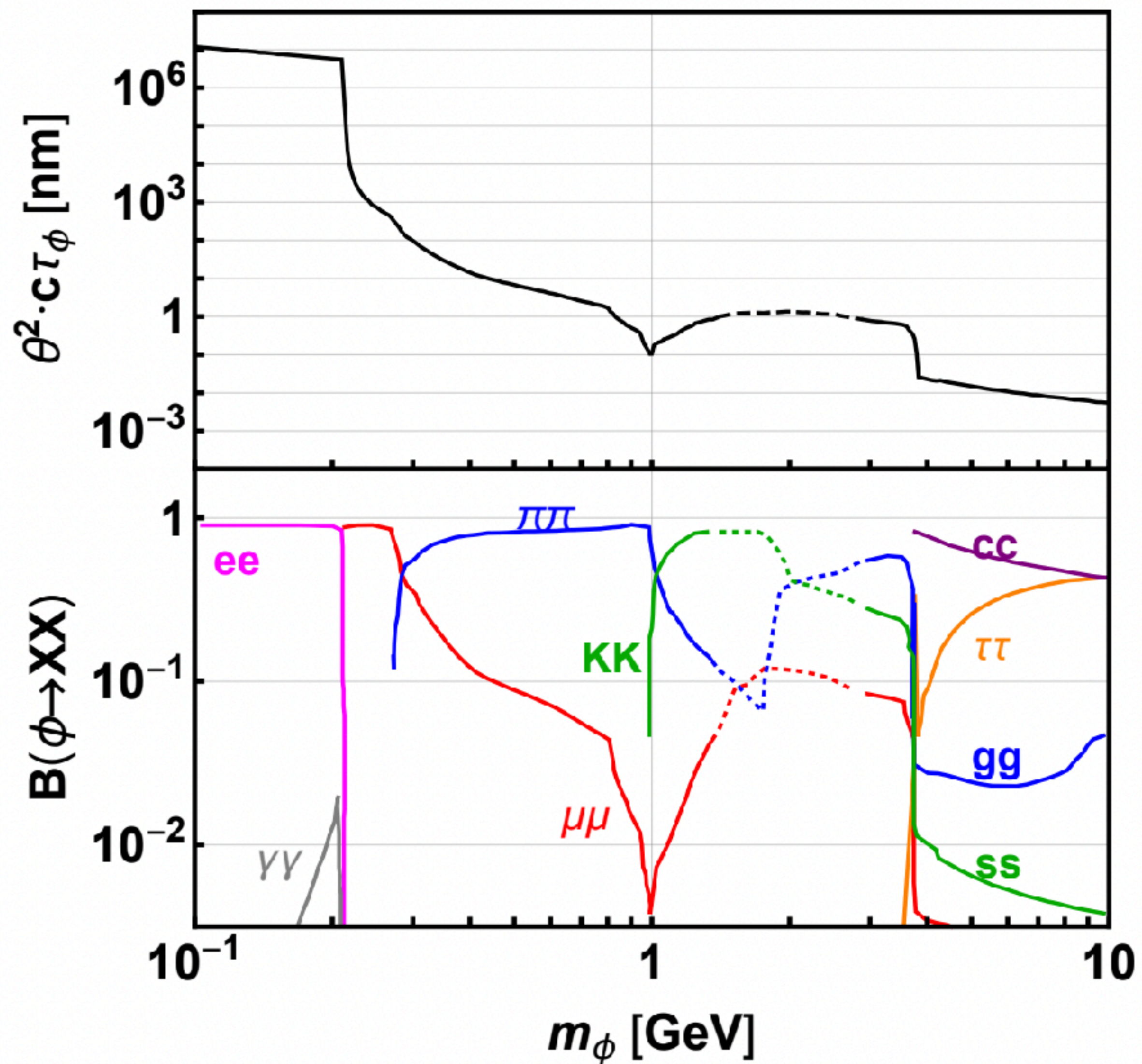




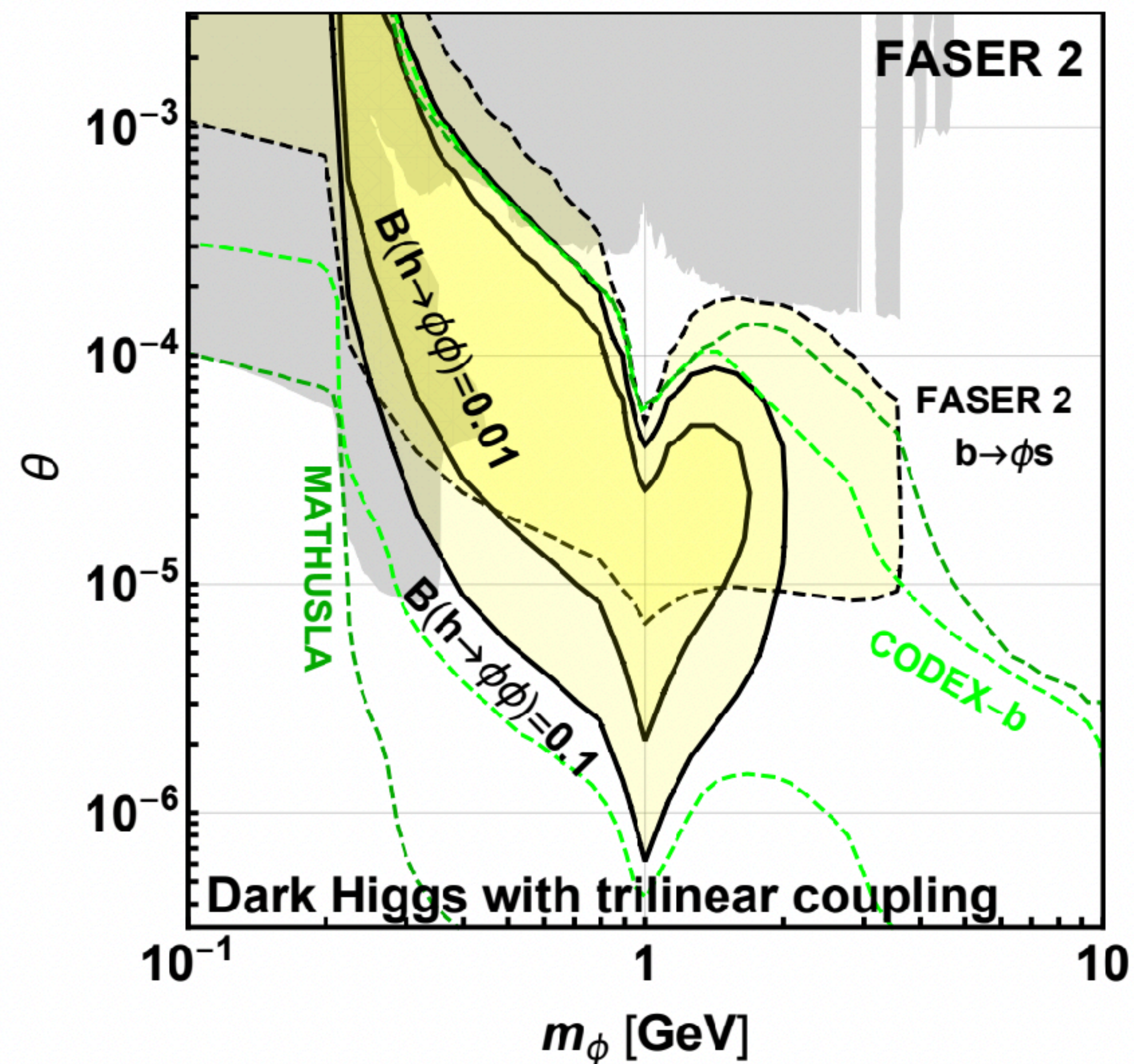
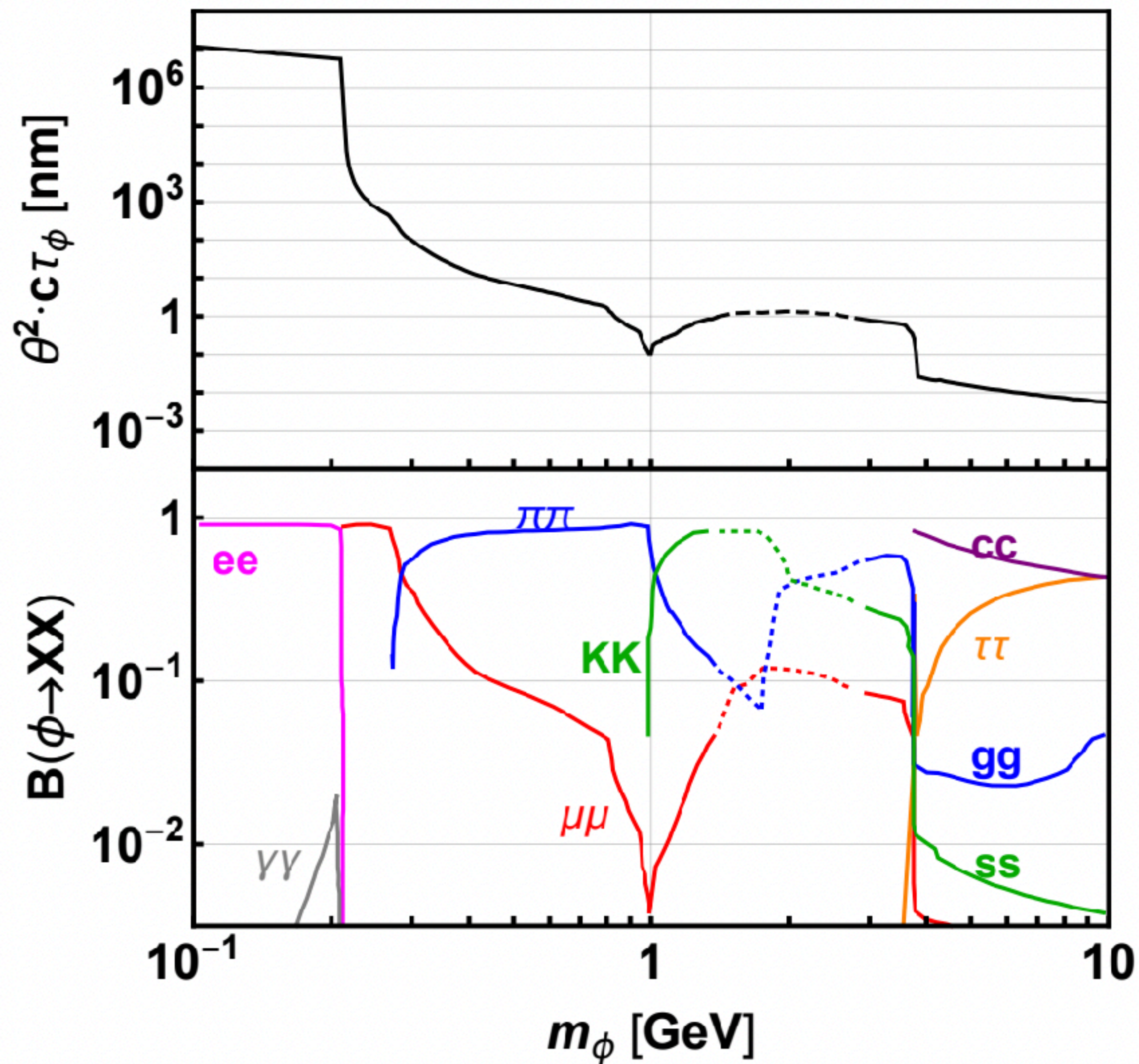
Target scenarios | Dark Photon

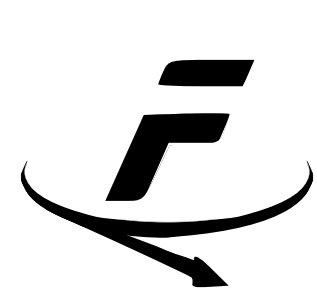


Target scenarios | Dark Higgs

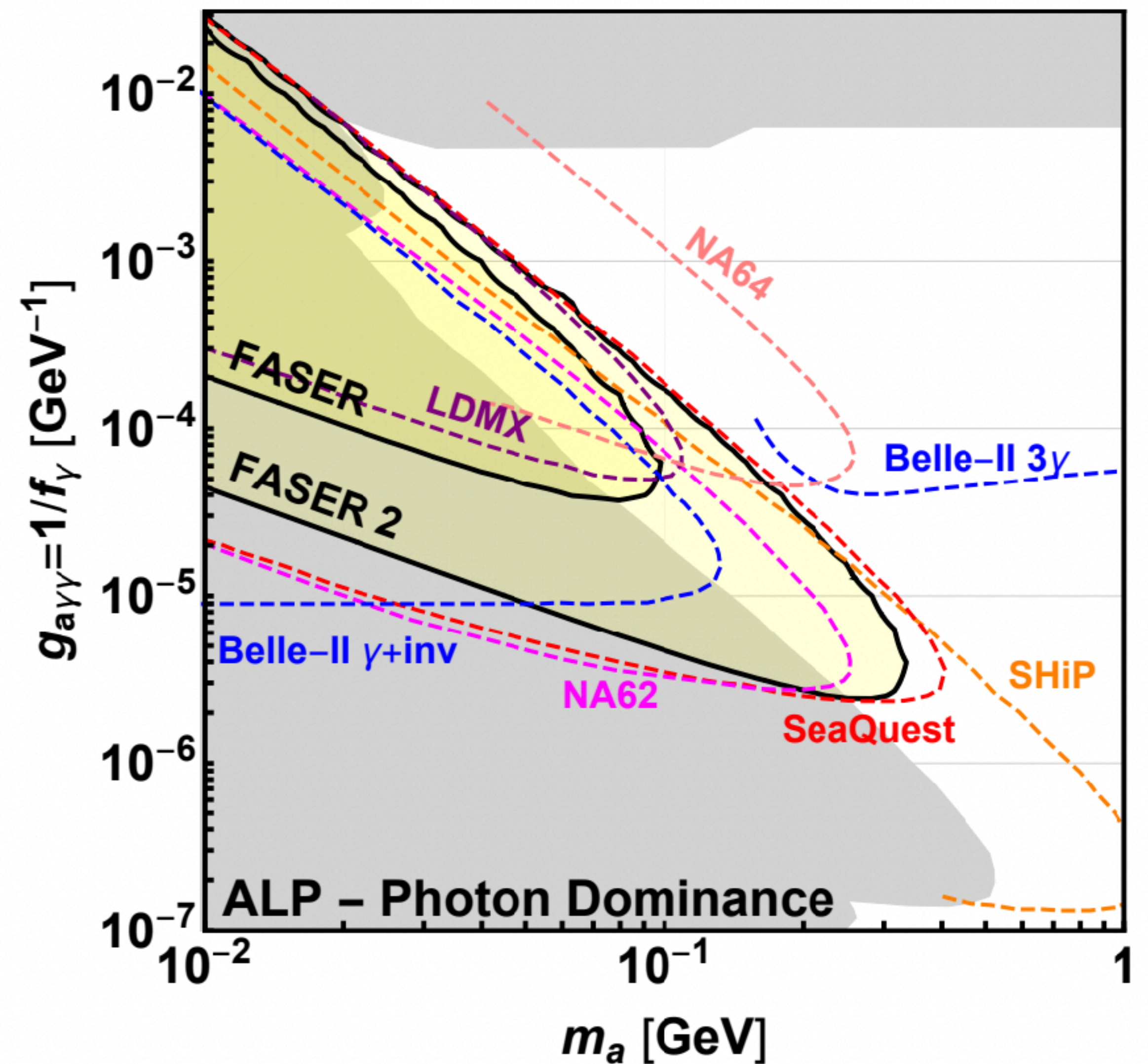
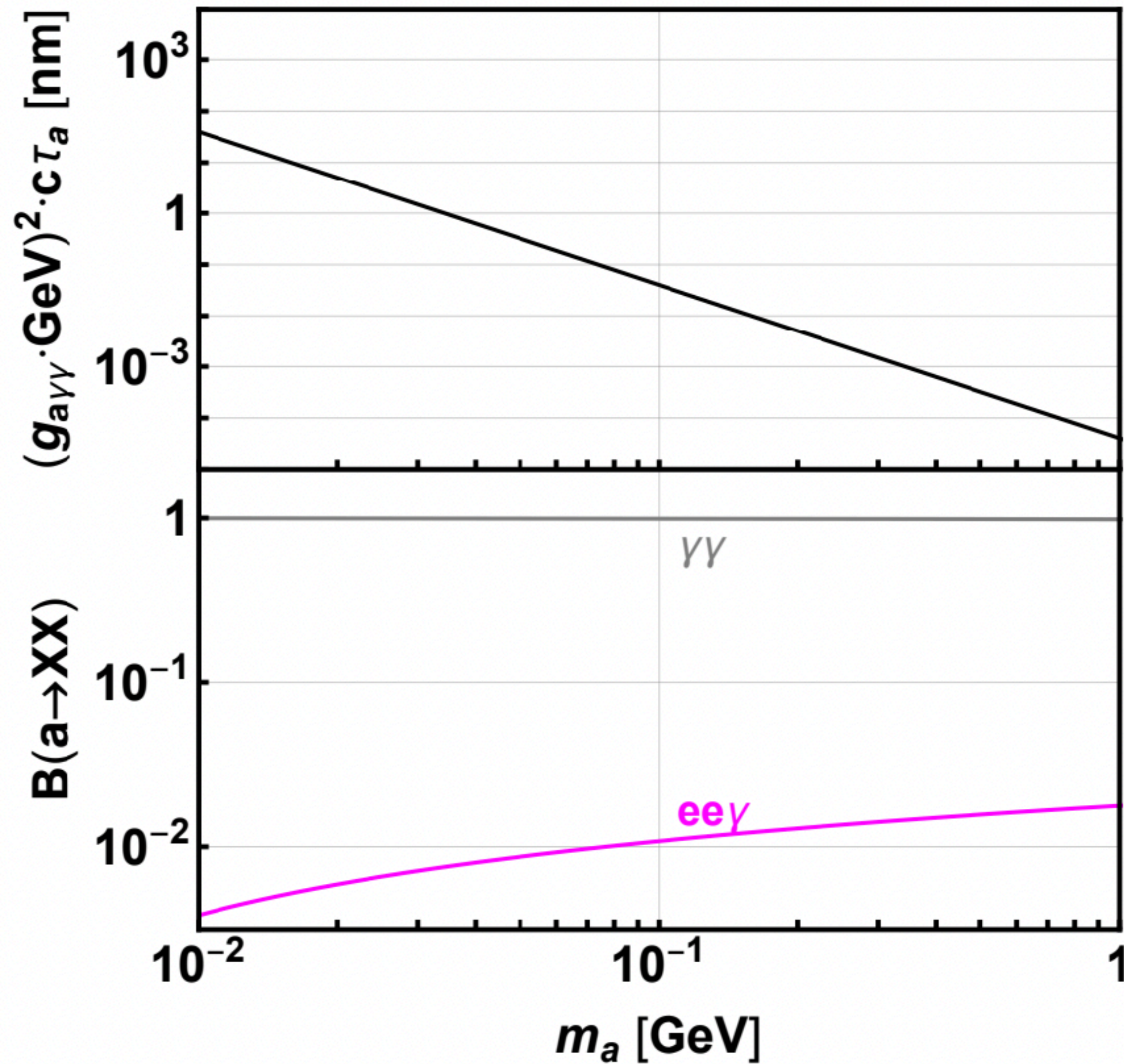


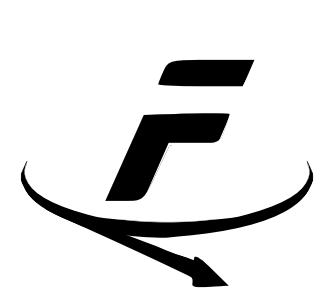
Target scenarios | Dark Higgs



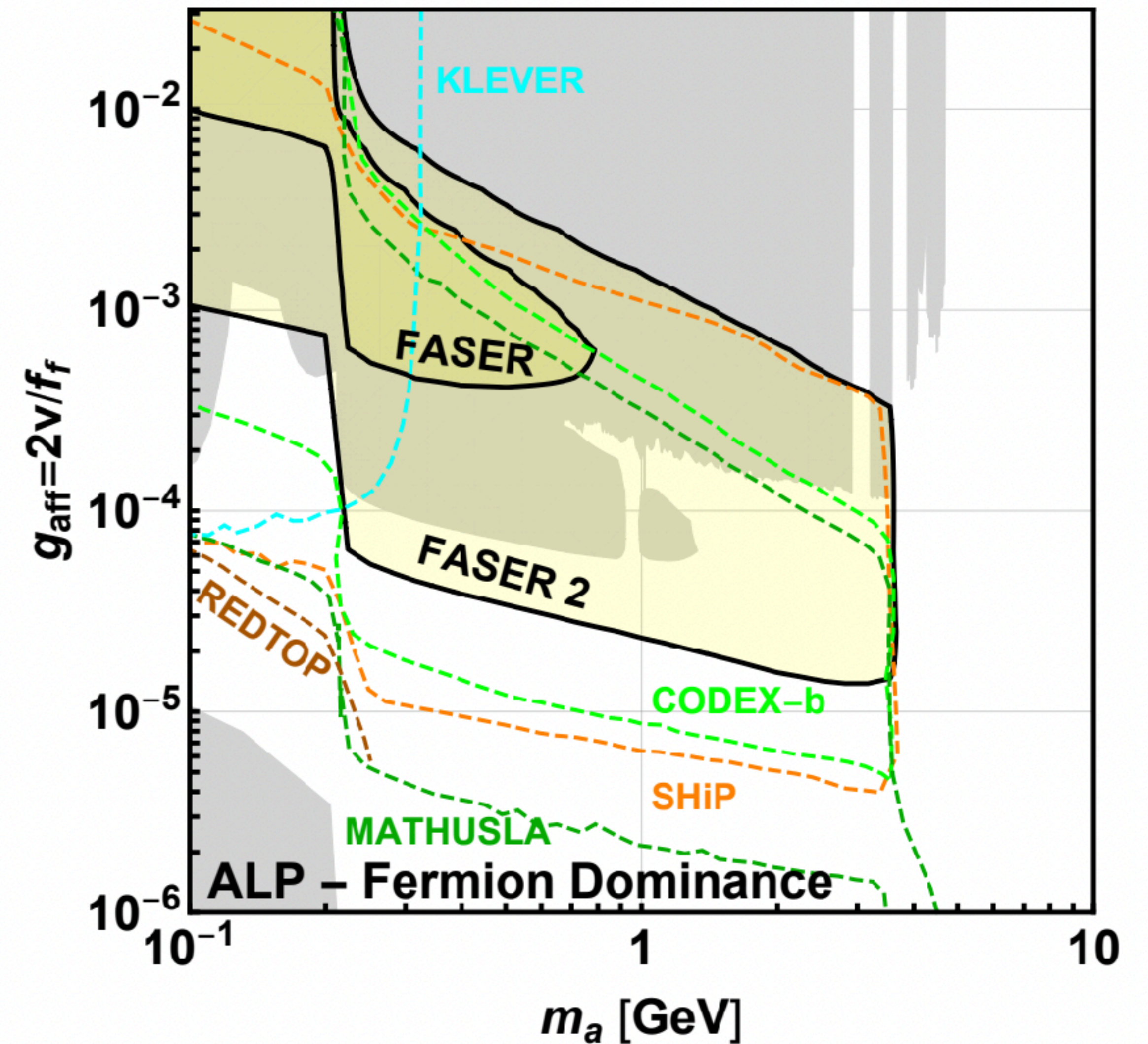
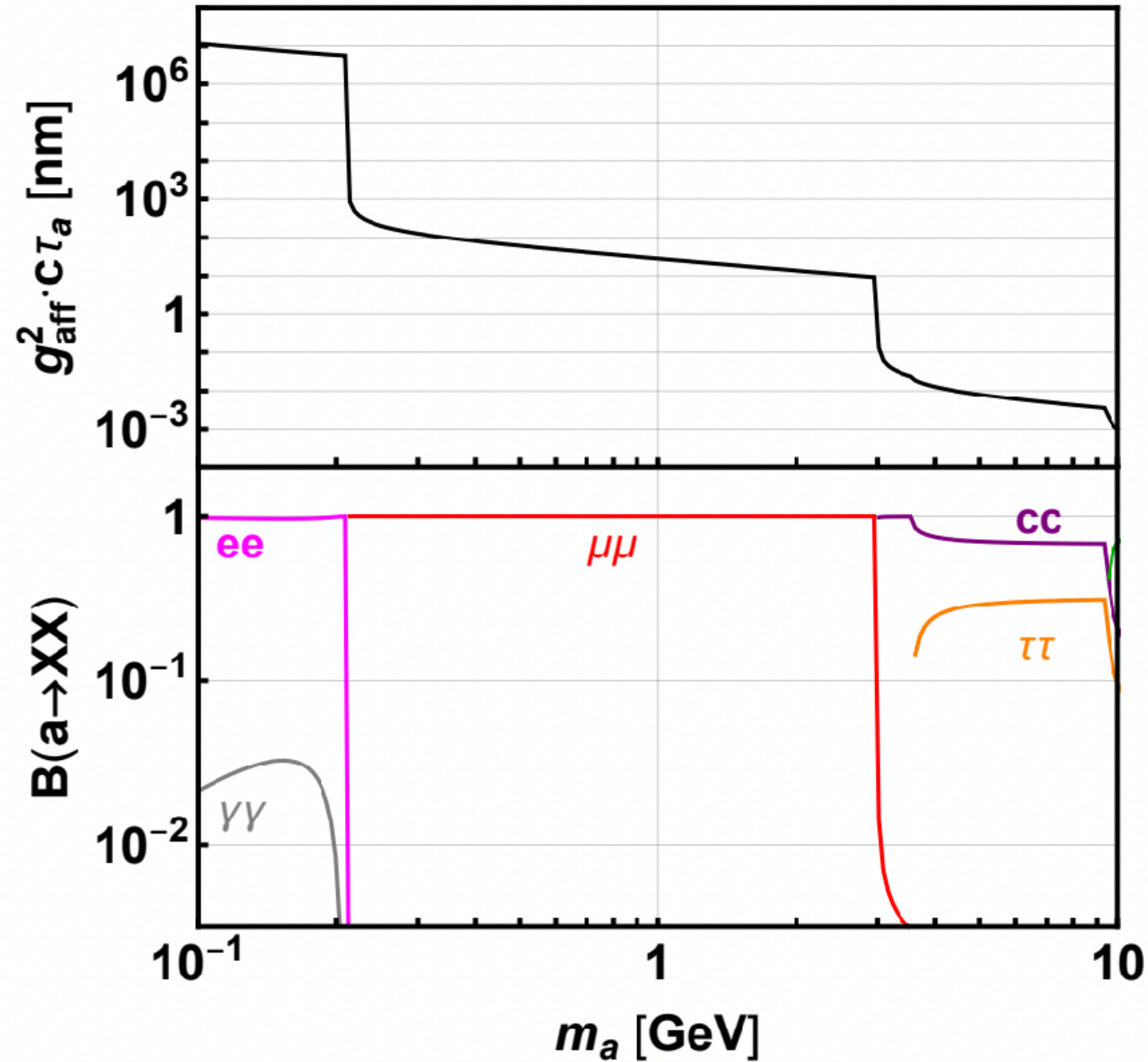


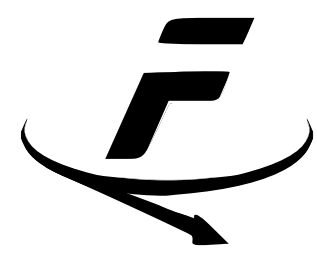
Target scenarios | ALP



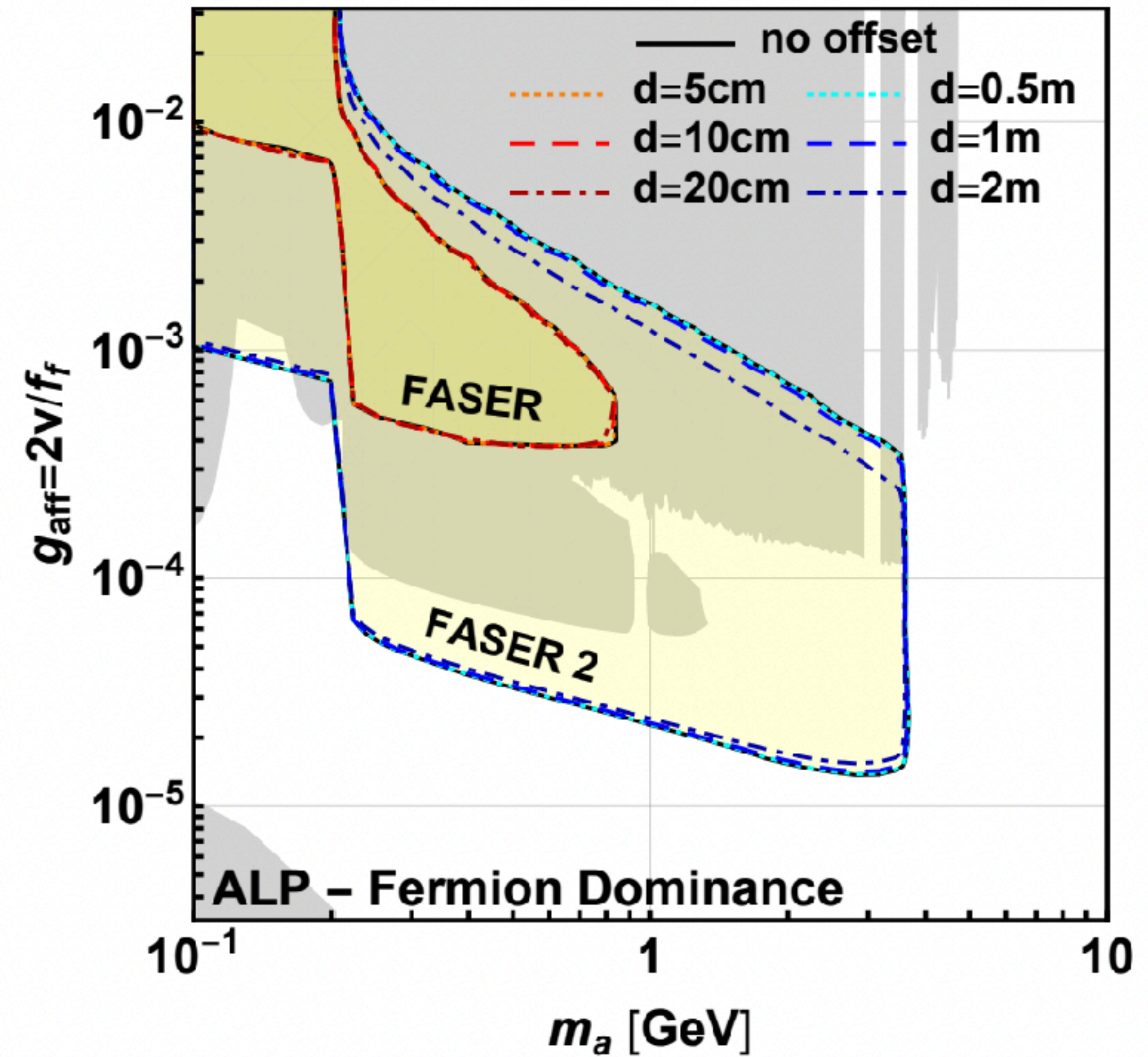
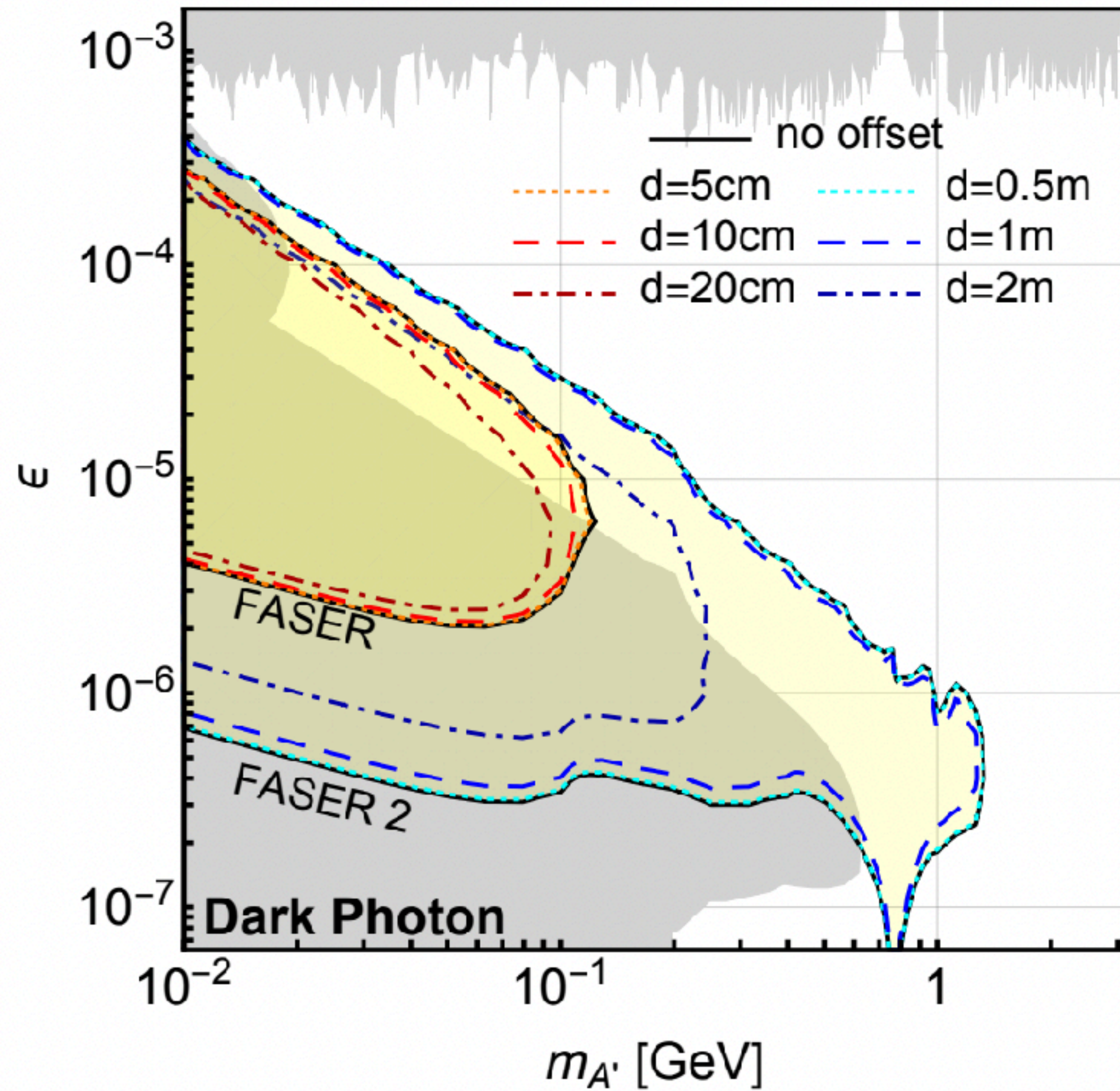


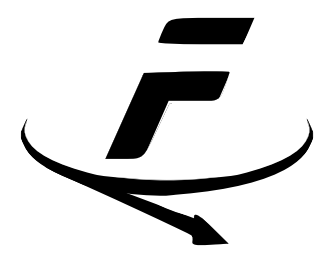
Target scenarios | ALP



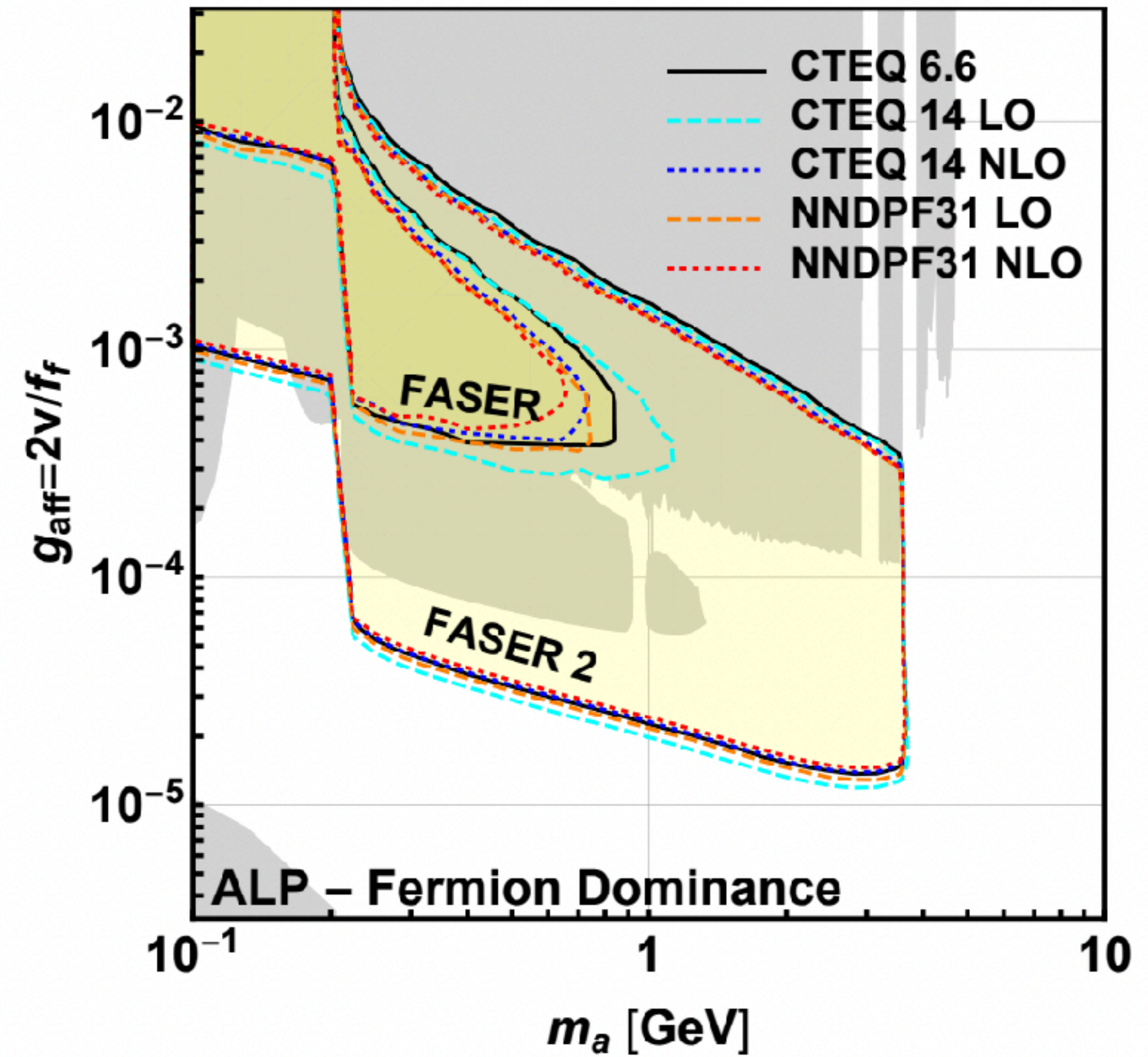
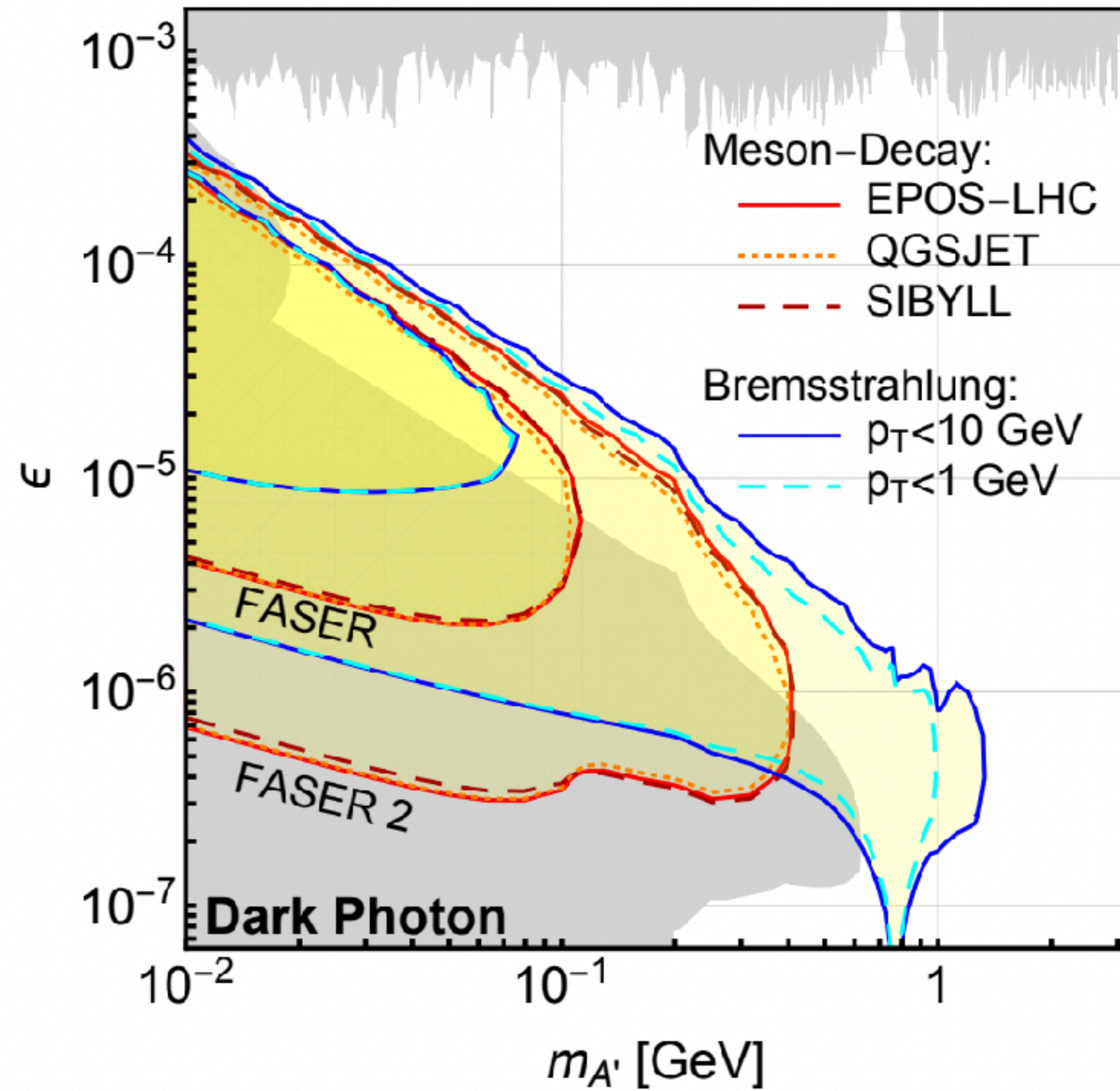


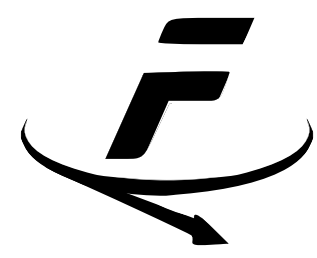
Beam offset





Modelling uncertainties





Energy threshold

