

Running couplings in Higher derivatives theories

Diego Buccio

D.B., J. Donoghue, G. Menezes, R. Percacci, *Physical running of couplings in quadratic gravity*,
Phys.Rev.Lett. 133 (2024) 2, 021604, arXiv:2403.02397

D. B., J. F. Donoghue, G. Menezes and R. Percacci, *Renormalization and running in the 2D CP(1) model*,
arXiv: 2408.13142

D. B., L. Parente and O. Zanusso, *Physical Running in Conformal Gravity and Higher Derivative Scalars*,
arXiv: 2410.21475

Why higher derivative?

- Einstein general relativity as a QFT is perturbatively non renormalizable
- Higher derivative operators R^2 , $R_{\mu\nu}R^{\mu\nu}$ and $R_{\mu\nu\rho\lambda}R^{\mu\nu\rho\lambda}$ contain a fourth derivative kinetic term for the metric



The theory is now perturbatively renormalizable [Stelle, '77]

$$S = \int d^4x \sqrt{-g} \left[+2\Lambda - Z_N R + \frac{1}{2\lambda} C^2 + \frac{1}{\xi} R^2 - \frac{1}{\rho} E \right]$$

- Ostrogradsky instability, ghosts and breakdown of unitarity

Historical overview

RG equations of quadratic gravity:

- Julve & Tonin, '78: perturbatively Asymptotically Free (AF) only with scalar tachyon

No Nakanishi-Lautrup ghost



- Fradkin & Tseytlin, '82: AF without scalar tachyon

Numerical error



- Avramidi & Barvinski, '85: again AF only with scalar tachyon

Problem:

Do these RG equations reproduce
the momentum dependence of
scattering amplitudes?

Running couplings from amplitudes (Gell-Mann & Low)

Consider a scattering amplitude renormalized with a momentum subtraction scheme at the mass scale

$$A(g_i, m_j, p) \sim g_i + g_i^2 \log\left(\frac{p^2}{m_j^2}\right)$$

if $p \gg m_j$
Breakdown of perturbativity

Large logs are reabsorbed in $g_i(\bar{p}) = g_i + g_i^2 \log(\bar{p}^2/m_j^2)$ at a new energy scale $\bar{p} \sim p$

$$A(g_i, m_j, p, \bar{p}) \sim g_i(\bar{p}) + g_i^2(\bar{p}) \log\left(\frac{p^2}{\bar{p}^2}\right)$$

The “physical” running of g is defined by integrating an infinitesimal shift of \bar{p} :

$$\bar{p} \frac{d}{d\bar{p}} g(\bar{p}) = \beta^{\bar{p}}(g) \quad g(\bar{p}') = \frac{g(\bar{p})}{1 - \beta^{\bar{p}}(g) \log\left(\frac{\bar{p}'^2}{\bar{p}^2}\right)}$$

Extended perturbative regime

Running of coupling from regulators (Wilson)

Regularization and renormalization introduce an unphysical energy scale (Λ, k, μ, \dots)

A typical renormalized amplitude is

$$A(g_i, \mu, p) \sim g_i(\mu) + g_i^2(\mu) c \log\left(\frac{f(m_i, p)}{\mu^2}\right)$$

Physics independent of μ  Callan-Symanzik equations

$$\frac{d}{d\mu} A(g_i, \mu, p) = 0, \quad \beta^\mu(g) := \mu \frac{\partial}{\partial \mu} g(\mu)$$

In massless theories with only one regulator, only one dimensionless quantity p^2/μ^2

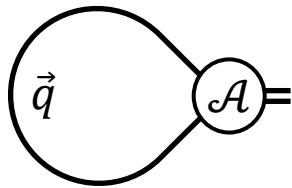


$$\beta^{\bar{p}}(g) = \beta_{m_i=0}^\mu(g)$$

Is this relation true in general?

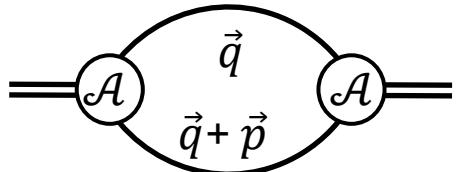
1-loop beta functions are universal in 2-derivative theories if $\bar{p}, \mu \gg m_i$

Higher derivatives theories



$$\begin{aligned} & \mu^{2\epsilon} \int d^{4-2\epsilon} q \frac{\mathcal{A}}{m^2 q^2 + q^4} \\ & \sim \frac{\mathcal{A}}{\epsilon} + \mathcal{A} \log\left(\frac{\mu^2}{m^2}\right) \end{aligned}$$

No running
with p



$$\begin{aligned} & \mu^{2\epsilon} \int d^{4-2\epsilon} q \frac{\mathcal{A}^2}{(m^2 q^2 + q^4)(m^2(q+p)^2 + (q+p)^4)} \\ & \sim \frac{\mathcal{A}^2}{p^4} \log\left(\frac{m^2}{p^2}\right) + O\left(\frac{m}{p}\right) \end{aligned}$$

p-running
without $\frac{1}{\epsilon}$ poles
if $\mathcal{A} \sim p^4$

General diagram

In higher derivatives theories there are off-shell IR divergences:

$$\sim \int d^4q \frac{N(p, q_i)}{q^4(q + p_1)^4 \times \dots \times (q + p_1 + \dots + p_{n-1})^4}$$

$\frac{1}{p_1^4 \times \dots \times (p_1 + \dots + p_{n-1})^4} \int d^4q \frac{N(p, q_i)}{q^4} + \dots$

$\int d^4q \frac{N(p, q_i)}{q^{4n}} + \dots$

$q \ll p_i$ $q \gg p_i$

If IR regulator \neq UV regulator (ex. $m_i \neq 0$, IR cutoff λ)

$$\beta^{\bar{p}}(g) \neq \beta^\mu(g) \quad \text{even when} \quad \bar{p} \gg m_i$$

[D.B., John Donoghue, Roberto Percacci, '23]
 [John Donoghue, Gabriel Menezes, '23]

HD scalar in curved spacetime

$$S_{HDS} = \frac{1}{2} \int d^4x \sqrt{\bar{g}} [\phi \bar{\square}^2 \phi + \phi \bar{\nabla}_\mu ((\xi_1 \bar{R}^{\mu\nu} + \xi_2 \bar{g}^{\mu\nu} \bar{R}) \bar{\nabla}_\nu \phi) + U \phi]$$

$$U = \lambda_1 \bar{C}^2 + \lambda_2 \bar{R}_{\mu\nu} \bar{R}^{\mu\nu} + \lambda_3 \bar{R}^2 + \lambda_4 \bar{\square} \bar{R}$$

$$\Gamma_{1-loop}[\bar{g}, \varphi] = \frac{1}{2} Tr \log \left(\frac{\delta^2 S}{\delta^2 \phi} \right)$$

$$\Gamma_{1-loop}[\bar{g}, 0] = \int d^4x \sqrt{\bar{g}} [\bar{C}_{\mu\nu\rho\sigma} f_\lambda(\bar{\square}; \mu^2, m^2) \bar{C}^{\mu\nu\rho\sigma} + \bar{R}_{\mu\nu} f_\xi(\bar{\square}; \mu^2, m^2) \bar{R}^{\mu\nu} + O(R^3)]$$

when $\bar{\square} \gg m^2$,

$$f_i(\bar{\square}, \mu, m) \sim b_i \log \left(\frac{\mu^2}{m^2} \right) + c_i \log \left(\frac{\bar{\square}}{m^2} \right)$$

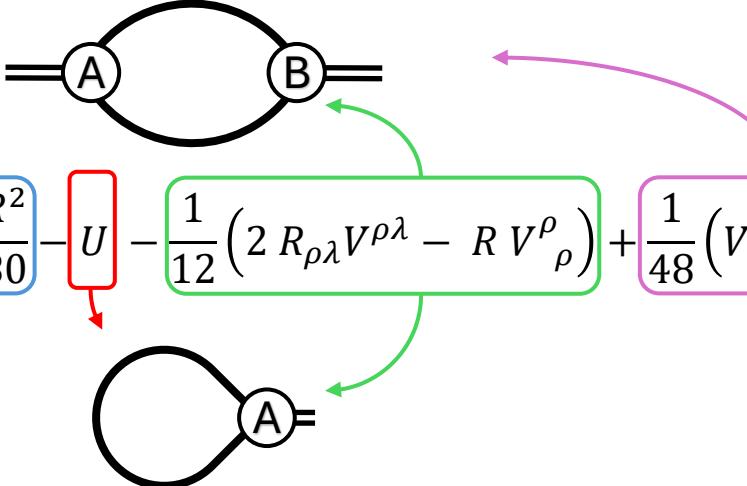
$$\beta_i^\mu \propto -b_i, \quad \beta_i^{\bar{p}} \propto c_i \quad b_i = -c_i?$$

Heat kernel computation

$$\frac{\delta^2 S}{\delta^2 \phi} = H = \square^2 + V^{\rho\lambda} \nabla_\rho \nabla_\lambda + N^\mu \nabla_\mu + U$$

- Heat kernel (HK) technique permits to do manifestly covariant computations in perturbation theory.
Divergencies are

$$-\frac{1}{\epsilon} \frac{1}{2(4\pi)^2} \int d^4x \text{tr} \left[\frac{R_{\mu\nu\rho\sigma}^2}{90} - \frac{R_{\mu\nu}^2}{90} + \frac{5R^2}{180} - U - \frac{1}{12} (2 R_{\rho\lambda} V^{\rho\lambda} - R V^\rho_\rho) + \frac{1}{48} (V^\rho_\rho V^\lambda_\lambda + 2 V_{\rho\lambda} V^{\rho\lambda}) \right]$$



HK introduces an unphysical IR regulator [Avramidi & Barvinski, '85]



Local HK expansion insensible to IR $\log(m^2/p^2)$

With Feynman diagrams

Expand around flat background ($\Lambda = 0$) [Julve & Tonin, '78]

$$\bar{g}_{\mu\nu} = \eta_{\mu\nu} + f_{\mu\nu}$$

Fourier space well defined \rightarrow I can use Feynman diagrams

$O(f^2)$ is enough to reconstruct R^2 terms in Γ_{1-loop}

Sensitive to
IR $\log(p^2)!$

$$H = \boxed{\Delta} + \mathcal{D}^{\rho\lambda\mu\nu}\partial_\rho\partial_\lambda\partial_\mu\partial_\nu + \mathcal{C}^{\rho\lambda\mu}\partial_\rho\partial_\lambda\partial_\mu + \mathcal{V}^{\rho\lambda}\partial_\rho\partial_\lambda + \mathcal{N}^\rho\partial_\rho + \mathcal{U}$$

Flat
propagator

$$Tr \log(H) = Tr \log(\Delta + A)$$

$$\approx Tr \left[\log \Delta + A \cancel{\frac{1}{\Delta}} - \boxed{\frac{1}{2}A \frac{1}{\Delta} A \frac{1}{\Delta}} + \dots \right]$$

$O(f^2)$ can be neglected in A

New IR divergent bubble diagrams

$$\text{If } \Delta = \square^2$$

$$-\frac{1}{2} \text{Tr} \left[A \frac{1}{\square^2} A \frac{1}{\square^2} \right] = \int d^{4-2\epsilon} q \frac{\text{Num}}{q^4 (q+p)^4}$$

vertices	numerator	$\log(p^2)$ term
UU	$-\frac{1}{2} \mathcal{U}_{AB} \mathcal{U}^{BA}$	$-\frac{\mathcal{U}^{AB} \mathcal{U}_{BA}}{16\pi^2 p^4}$
UN	$-\frac{i}{2} \mathcal{U}_{AB} \mathcal{N}^{\mu BA} q_\mu$	$i \frac{\mathcal{U}^{AB} \mathcal{N}_{BA}^\mu p_\mu}{32\pi^2 p^4}$
UV	$\frac{1}{2} \mathcal{U}_{AB} \mathcal{V}^{\mu\nu BA} q_\mu q_\nu$	$\frac{\mathcal{U}^{AB} \mathcal{V}_{BA}^{\mu\nu} p_\mu p_\nu}{32\pi^2 p^4}$
UC	$\frac{i}{2} \mathcal{U}_{AB} \mathcal{C}^{\mu\nu\rho BA} q_\mu q_\nu q_\rho$	$-i \frac{\mathcal{U}^{AB} \mathcal{C}_{BA}^{\mu\nu\rho} p_\mu p_\nu p_\rho}{32\pi^2 p^4}$
UD	$-\frac{1}{2} \mathcal{U}_{AB} \mathcal{D}^{\mu\nu\rho\sigma BA} q_\mu q_\nu q_\rho q_\sigma$	$-\frac{1}{32\pi^2} \left(\frac{\mathcal{U}^{AB} \mathcal{D}_{BA}^{\mu\nu\rho\sigma} p_\mu p_\nu p_\rho p_\sigma}{p^4} - \frac{\mathcal{U}^{AB} \mathcal{D}_{\mu\nu BA}^{\mu\nu}}{8} \right)$

Apparently nonlocal contributions!

Localization

$$\mathcal{U} \sim \partial\partial\partial\partial f, \mathcal{N} \sim \partial\partial\partial f, \mathcal{V} \sim \partial\partial f, \mathcal{C} \sim \partial f, \mathcal{D} \sim f$$

$$\frac{\mathcal{U}\mathcal{A}p^{4-\dim[\mathcal{A}]}}{p^4} \sim \frac{p^4 f \ p^4 f}{p^4}$$



“nonlocal logs” become local $+O(f^3)$ nonlocalities

$$\frac{\mathcal{U}\mathcal{A}p^{4-\dim[\mathcal{A}]}}{p^4} \sim p^2 f \ p^2 f$$

All f^2 terms can be rearranged in C^2 and R^2 up to topological terms

Comparing beta functions

μ -running

$$\begin{aligned}\beta_\lambda &= -\frac{\lambda^2}{(4\pi)^2} \left[\frac{1}{30} + \frac{\xi_1}{24} (\xi_1 - 4) - 2\lambda_1 - \lambda_2 \right] \\ \beta_\xi &= -\frac{\xi^2}{(4\pi)^2} \left[\frac{1}{18} + \frac{\xi_1}{18} + \frac{5\xi_1^2}{72} + \frac{\xi_2}{3} + \frac{\xi_1\xi_2}{3} + \xi_2^2 - \frac{2\lambda_2}{3} - 2\lambda_3 \right]\end{aligned}$$

Physical running

$$\begin{aligned}\beta_\lambda &= -\frac{\lambda^2}{(4\pi)^2} \left[\frac{1}{30} + \frac{\xi_1}{24} (\xi_1 - 4) \right] \\ \beta_\xi &= -\frac{\xi^2}{(4\pi)^2} \left[\frac{1}{18} + \frac{\xi_1}{18} + \frac{5\xi_1^2}{72} + \frac{\xi_2}{3} + \frac{\xi_1\xi_2}{3} + \xi_2^2 - 2\lambda_4^2 \right]\end{aligned}$$

Discrepancy given by λ_i

Universality with $\lambda_i = 0$ $\rightarrow U = 0 \rightarrow$ Shift symmetry

Conformal theory: $\lambda_i = 0, \xi_1 = 2, \xi_2 = \frac{3}{2}$

Conformal anomaly $\langle T \rangle = -\frac{1}{15(4\pi)^2} C^2$ is universal!

Quadratic gravity

$$S_g = \int d^4x \sqrt{-g} \left[+2\Lambda - Z_N R + \frac{1}{2\lambda} C^2 + \frac{1}{\xi} R^2 - \frac{1}{\rho} E \right]$$

$$Z_N = \frac{m_p^2}{16\pi}$$

The Euler density $E = R_{\mu\nu\rho\sigma}^2 - 4R_{\mu\nu}^2 + R^2$ is topological in d=4

particle	spin	Mass ²
graviton	2	0
ghost	2	$\lambda m_p^2/2$
scalar	0	$-\xi m_p^2/12$

Background field method

$$g_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu}$$

$$S = S_g + S_{GF} + S_{FP}$$

$$S = \bar{S}_g + \int d^4x \sqrt{-\bar{g}} \frac{1}{2} h \frac{\delta^2 S_g}{\delta h^2} \Big|_{g_{\mu\nu}=\bar{g}_{\mu\nu}} h + \frac{1}{2a} \int d^4x \sqrt{-\bar{g}} F^\mu(h) Y_{\mu\nu} F^\nu(h) + \bar{c}^\mu \Delta_{GH\mu\nu} c^\nu + O(h^3)$$

$$h \frac{\delta^2 (S_g + S_{GF})}{\delta h^2} \Big|_{g_{\mu\nu}=\bar{g}_{\mu\nu}} h = h H h$$

One-loop effective action:

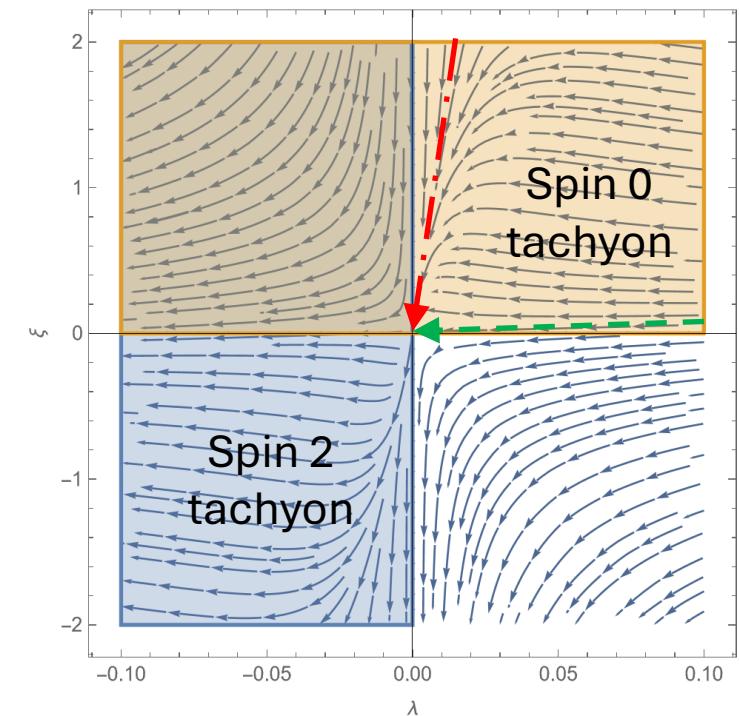
$$\Gamma_{1-loop} = \bar{S}_g + \frac{1}{2} Tr(\log H) - \frac{1}{2} Tr(\log Y) - Tr(\log \Delta_{GH})$$

Beta functions of quadratic gravity

μ -running

$$\beta_\lambda = -\frac{1}{(4\pi)^2} \frac{133}{10} \lambda^2$$

$$\beta_\omega = -\frac{1}{(4\pi)^2} \frac{25 + 1098\omega + 200\omega^2}{60} \lambda$$



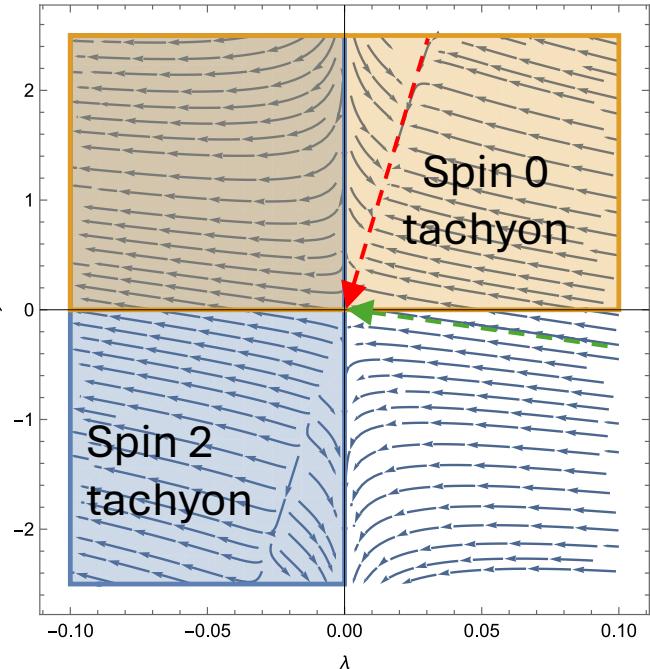
Spin 0 tachyon

$$\omega = -\frac{3\lambda}{\xi}$$

Physical running

$$\beta_\lambda = -\frac{\lambda^2(539\omega + 20)}{240\pi^2} \omega$$

$$\beta_\omega = \lambda \frac{1400\omega^2 - 1138\omega - 45}{480\pi^2}$$



Conformal Gravity

$$S = \int d^4x \sqrt{-g} \left[\frac{1}{2\lambda} C^2 - \frac{1}{\rho} E \right]$$

Topological terms do not contribute to scattering amplitudes



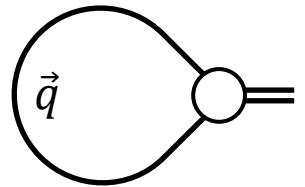
$\beta^{\bar{p}}(\rho)$ not defined

$$\beta_\lambda^{HK} = -\frac{1}{(4\pi)^2} \frac{199}{15} \lambda^2 \quad \beta_\lambda^{\bar{p}} = -\frac{1}{(4\pi)^2} \frac{93}{5} \lambda^2$$

No qualitative differences

Conformal anomaly not well defined with dynamical metric

$$\ln d = 2$$



$$\sim \int d^2 q \frac{1}{q^2} \log \text{divergent}$$

Potentially the same problem!

$CP(1)$ (or $O(3)$) NLSM

$$S = \frac{1}{2g} \int d^2 x \frac{\partial_\mu \phi^a \partial^\mu \phi^a}{\left[1 + \frac{1}{4} \phi^a \phi^a \right]^2} \quad a = 1, 2$$

Usually β_g computed from tadpoles [Shifman '12,...]
unphysical?

$2 \rightarrow 2$ amplitude

The theory is integrable: all amplitudes writable in terms of $2 \rightarrow 2$ processes

$$\phi_1 + \phi_1 \rightarrow \phi_2 + \phi_2$$

is IR safe at one-loop

$$\mathcal{M} = g^2(\mu)s - \frac{g^4 s}{8\pi} \left[\log\left(-\frac{t}{\mu^2}\right) + \log\left(-\frac{u}{\mu^2}\right) \right] - \frac{g^4}{8\pi} (t-u) \log\left(\frac{t}{u}\right)$$

Logs of the artificial IR cutoff m from tadpole diagrams are cancelled by IR divergent bubble diagrams and replaced with $\log(p)$

$$\beta^{\bar{p}}(g) = \beta^\mu(g) = -\frac{g^3}{4\pi}$$

Conclusions

- Renormalization and running are not equivalent concepts
- In higher derivatives theories UV-IR mixing in beta functions
- The heat kernel misses IR running
- Taking in account also IR contributions, there exists a unique AF trajectory without tachyons in quadratic gravity
- In $d = 2$ $CP(1)$ NLSM universality recovered thanks to unitarity (Kinoshita–Lee–Nauenberg theorem)

Future perspectives

- **Beta functions of nonlocal partner of R and Λ**

Beta functions of $m_p^2 R \frac{1}{\square} R$ and $m_p^4 R \frac{1}{\square^2} R$

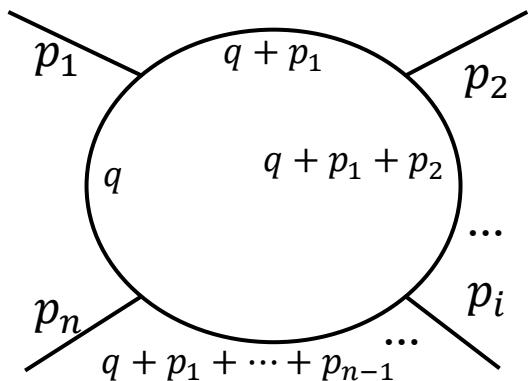
We perturbed around flat spacetime, extension to $\Lambda \neq 0$?

- **The same result can be reproduced with non-local HK expansion?**
- **The effective action is off-shell and gauge-fixing dependent: compare with running couplings extracted from scattering amplitudes in quadratic gravity**

Thank you!

2 derivatives theories

In general $\log(p^2/\mu^2)$ can arrive only from the UV region of loop integrals



$$\sim \int d^4q \frac{N(p, q_i)}{q^2(q + p_1)^2 \times \dots \times (q + p_1 + \dots + p_{n-1})^2}$$

\downarrow q \ll p_i \downarrow q \gg p_i

$$\frac{1}{p_1^2 \times \dots \times (p_1 + \dots + p_{n-1})^2} \int d^4q \frac{N(p, q_i)}{q^2} + \dots$$
$$\int d^4q \boxed{\frac{N(p, q_i)}{q^{2n}}} + \dots$$

The UV regions are equal in massive and massless theories, up to $O(m_i/p)$

$$\beta^{\bar{p}}(g) = \beta_{m_i=0}^\mu(g) = \beta_{m_i \neq 0}^\mu(g) \quad \text{when } \bar{p} \gg m_i$$