



How could we probe the angular dependence of dark matter self-interactions?

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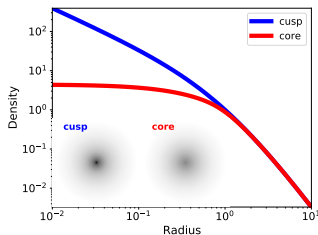
Collaborators:

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SIDM and Small-Scale Problems

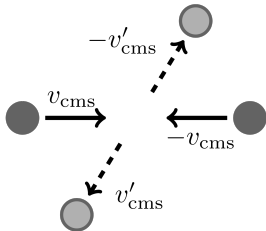
- Λ CDM can explain cosmological large-scale structure remarkably well
- There are several issues on small scales (small-scale crisis)
- Self-interacting dark matter (SIDM) is promising, can solve or at least mitigate small-scale problems.
 - e.g. SIDM can form density cores



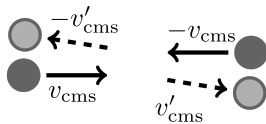


The Collision Term

We distinguish two regimes:



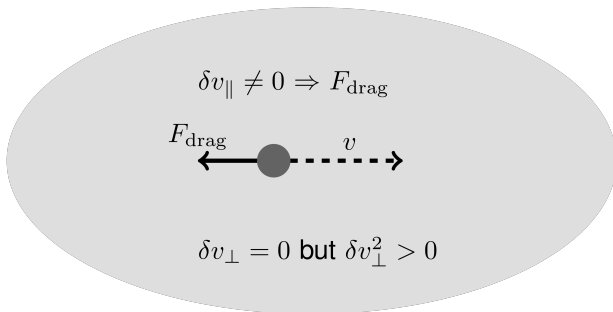
large-angle scattering
– rare –



small-angle scattering
– frequent –



Effective Description: Drag Force

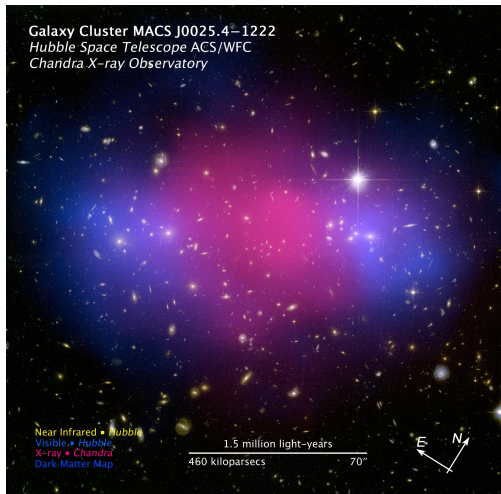


Description of drag force from Kahlhoefer et al. 2014



Galaxy Cluster Merger

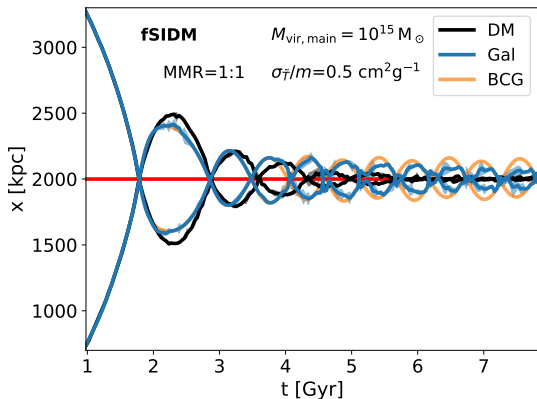
Credits: NASA, ESA, CXC, M. Bradac (University of California, Santa Barbara), and S. Allen (Stanford University)





Equal-Mass Merger

Offsets at later merger stages are much larger

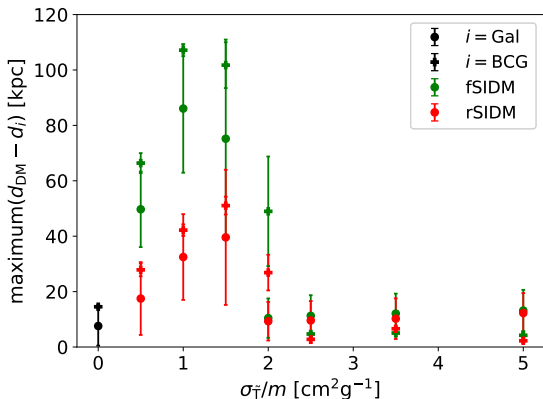


Fischer et al. 2021b



Maximum Offset

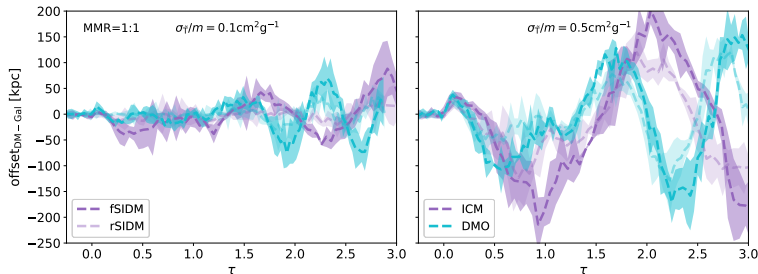
The maximum achievable offset is much larger for fSIDM than for rSIDM



Fischer et al. 2021a



Equal-Mass Merger: Offsets comparison

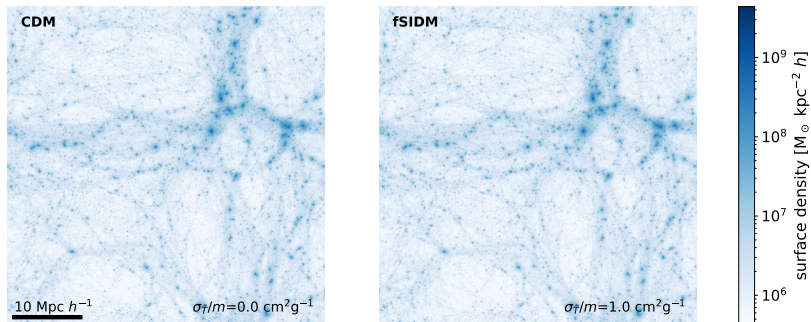


Fischer et al. 2023

Offsets at later times are much larger when including ICM



Cosmological Study

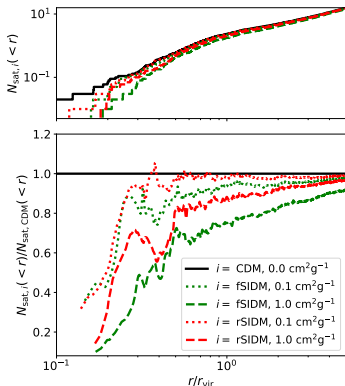


Fischer et al. 2022

No differences on large scales



Cosmological Study: Satellite Abundance



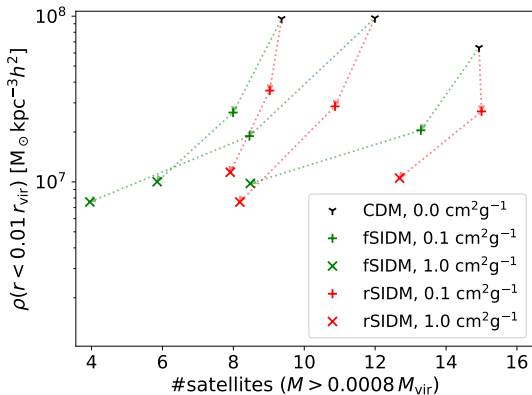
Fischer et al. 2022

Interestingly large suppression of satellites for fSIDM



Central Density vs. Number of Satellites

Qualitative difference between rare and frequent scattering



Fischer et al. 2022



Take Home Messages

N-body simulations of fSIDM are ...

1. possible

- We developed a new numerical scheme,
- based on an effective description (drag force).

2. important

- fSIDM and rSIDM have different phenomenology (offsets, satellite abundance),
- significant difference also at small cross-sections ($\lesssim 0.1 \text{ cm}^2/\text{g}$).



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