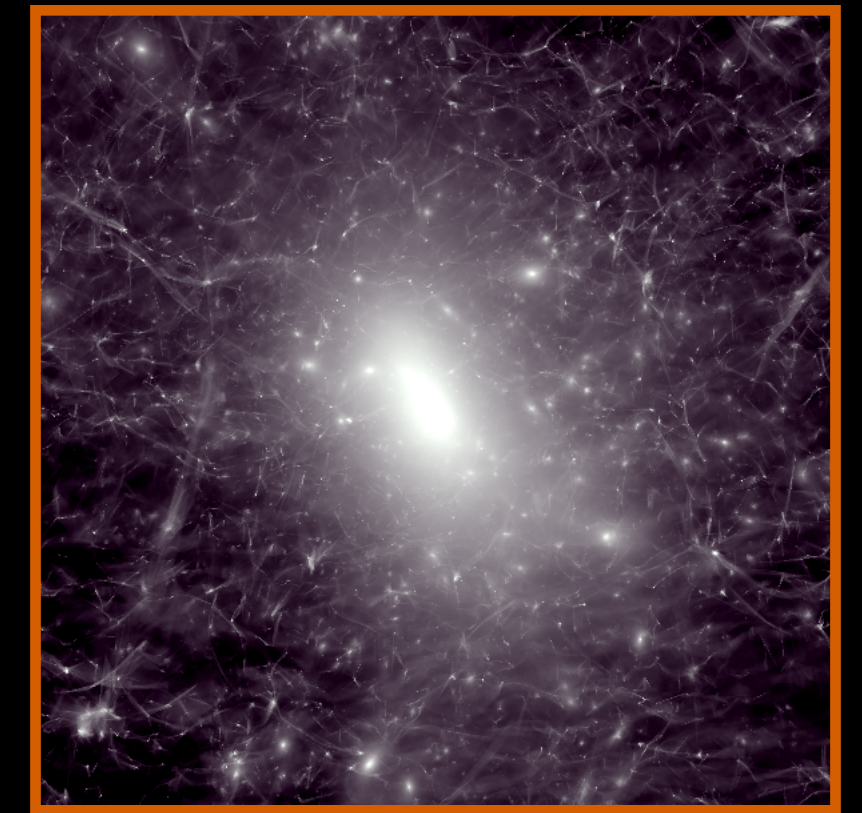
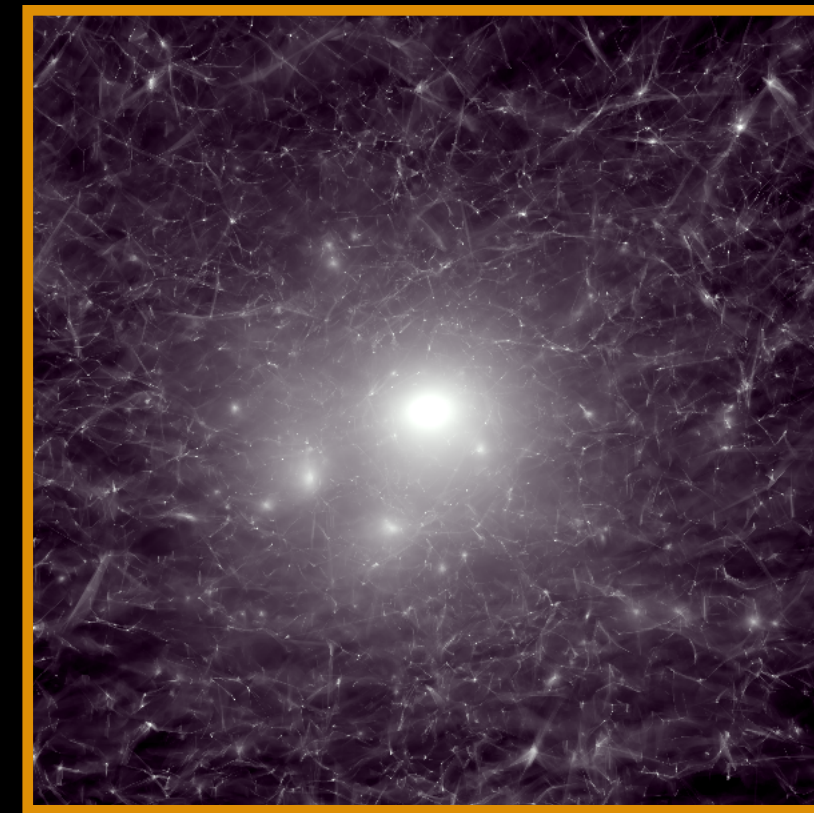
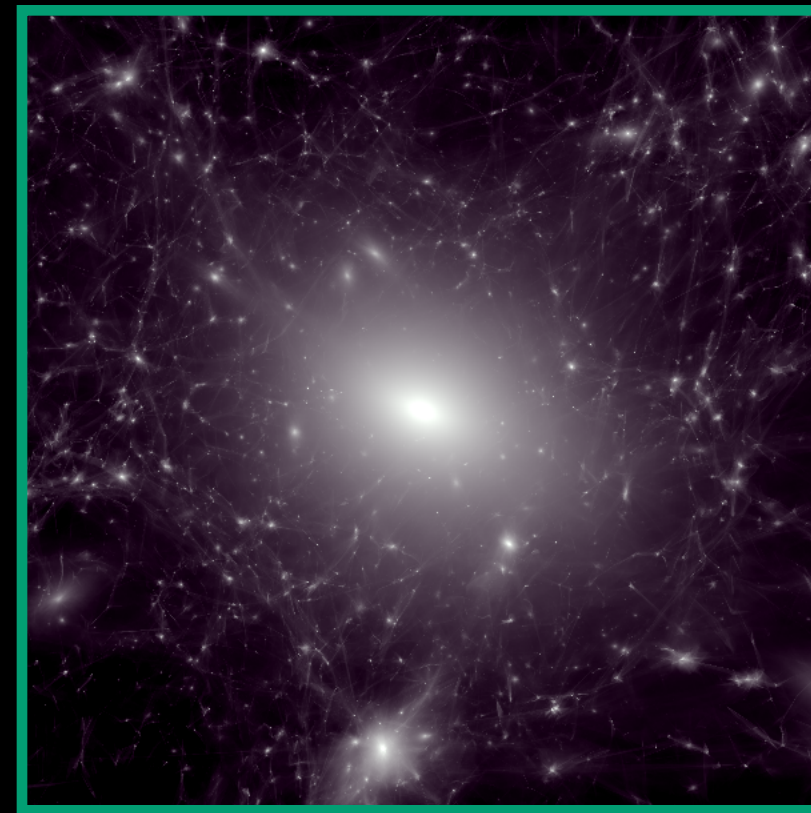
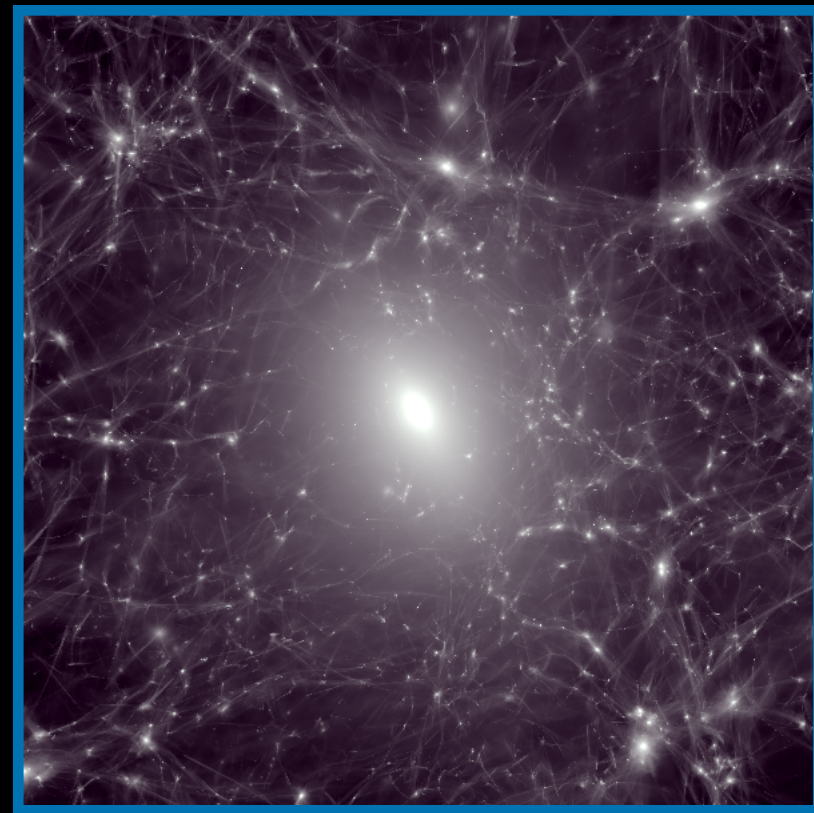
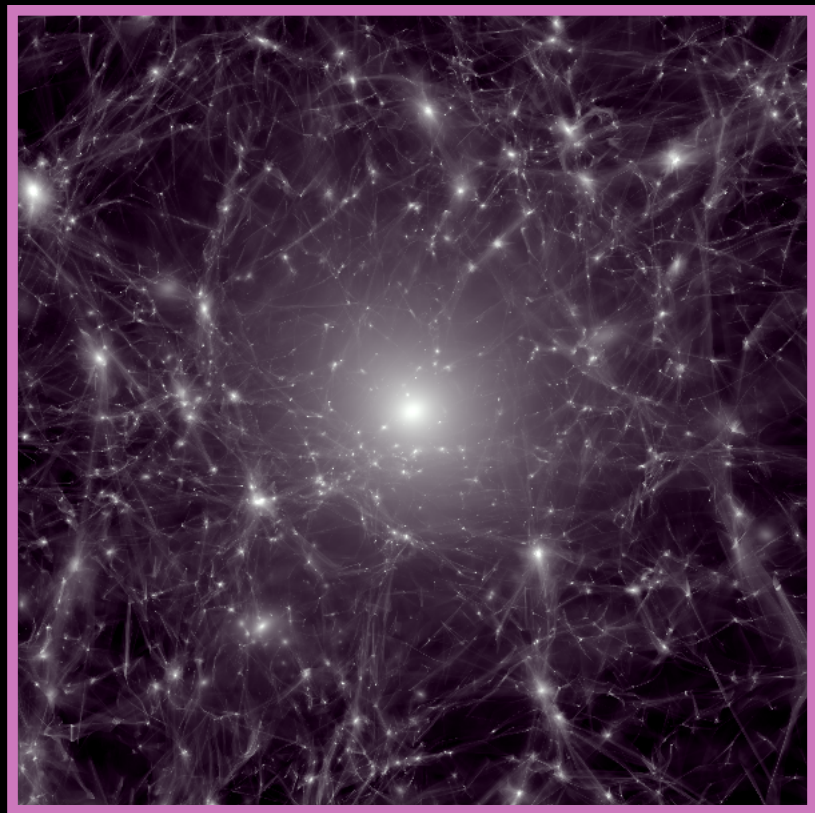


# Cosmological Simulations with Novel Dark Matter Physics



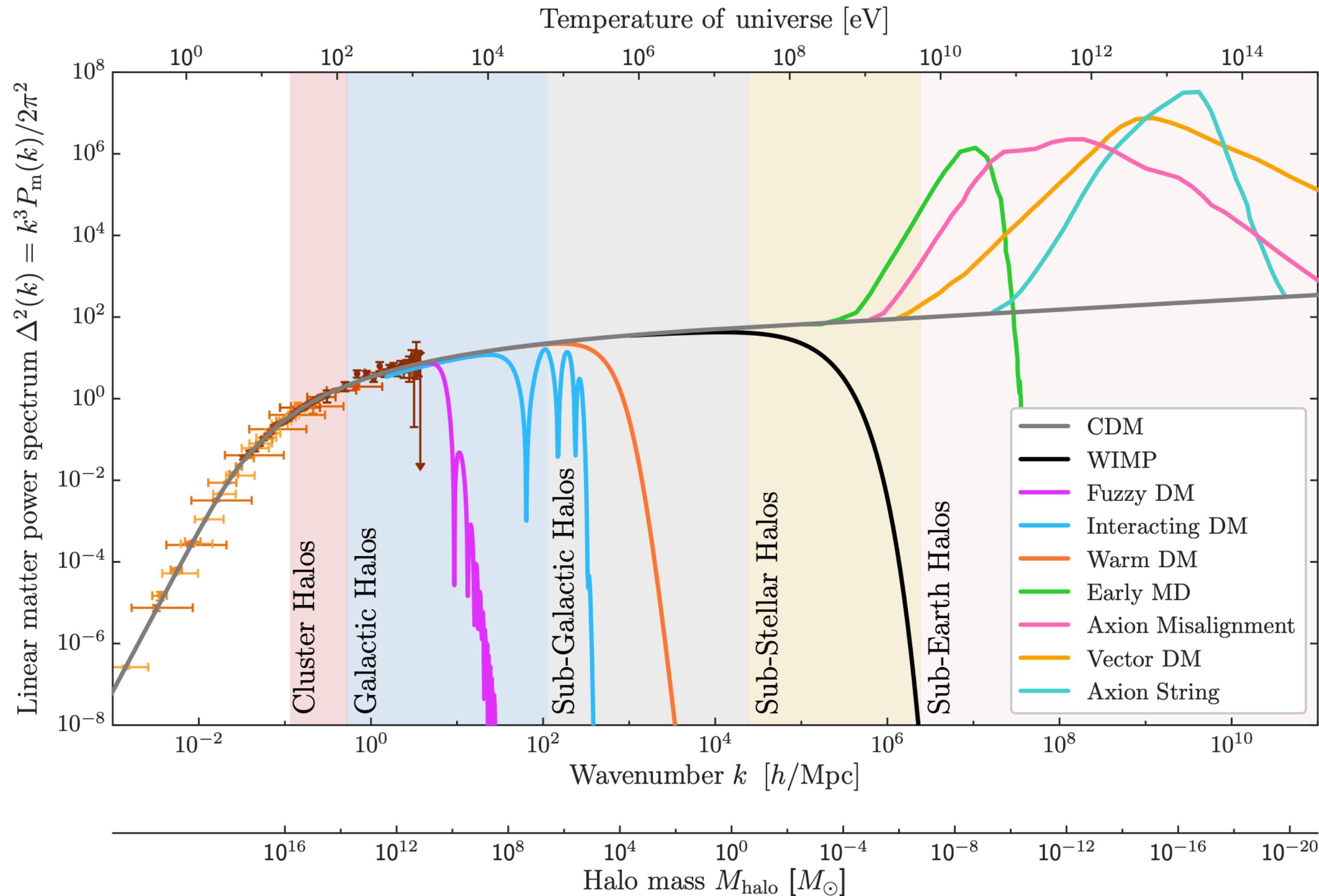
Ethan Nadler

UCLA Dark Matter 2023

3/30/2023



# Dark Matter Physics on Small Scales

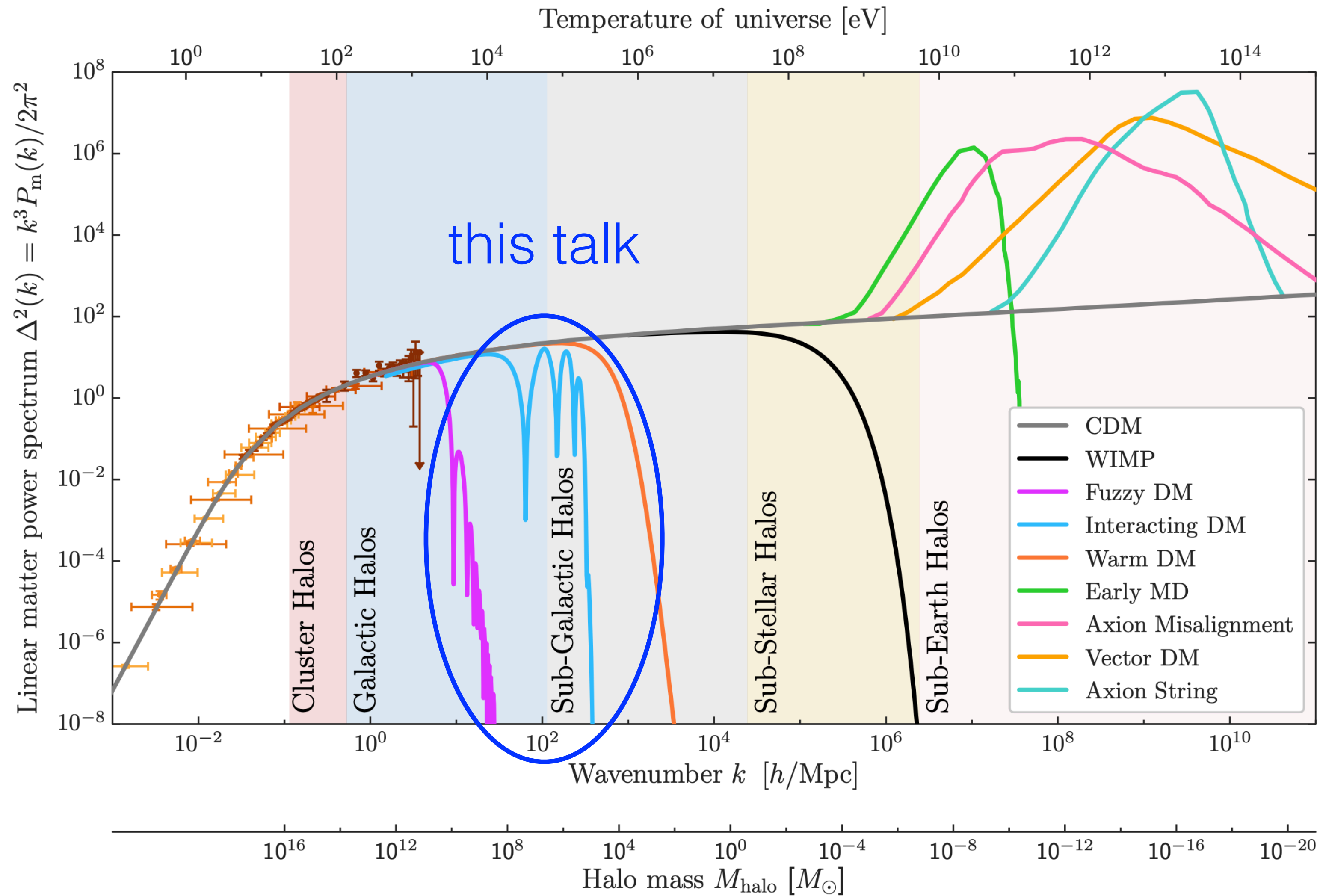


Dark matter physics affects structure formation throughout cosmic history:

- Clustering on scales smaller than  $\sim 1$  Mpc is mostly unconstrained
- New DM physics affects abundance & density profiles of low-mass halos
- **Simulations are needed** to robustly explore DM physics near and below the galaxy formation threshold



# Dark Matter Physics on Small Scales



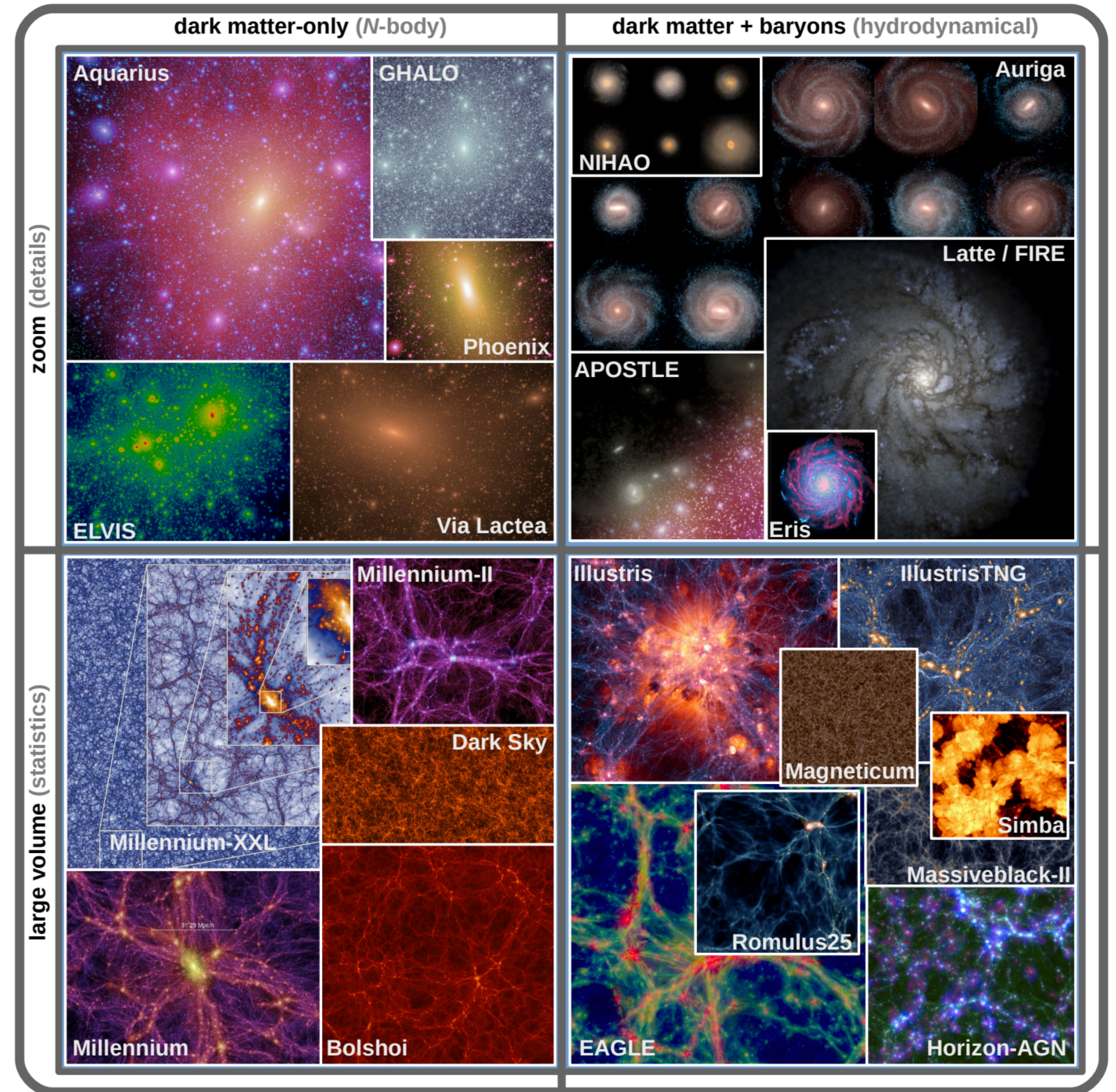
Dark matter physics affects structure formation throughout cosmic history:

- Clustering on scales smaller than  $\sim 1$  Mpc is mostly unconstrained
- New DM physics affects abundance & density profiles of low-mass halos
- **Simulations are needed** to robustly explore DM physics near and below the galaxy formation threshold



# The Landscape of Cosmological Simulations

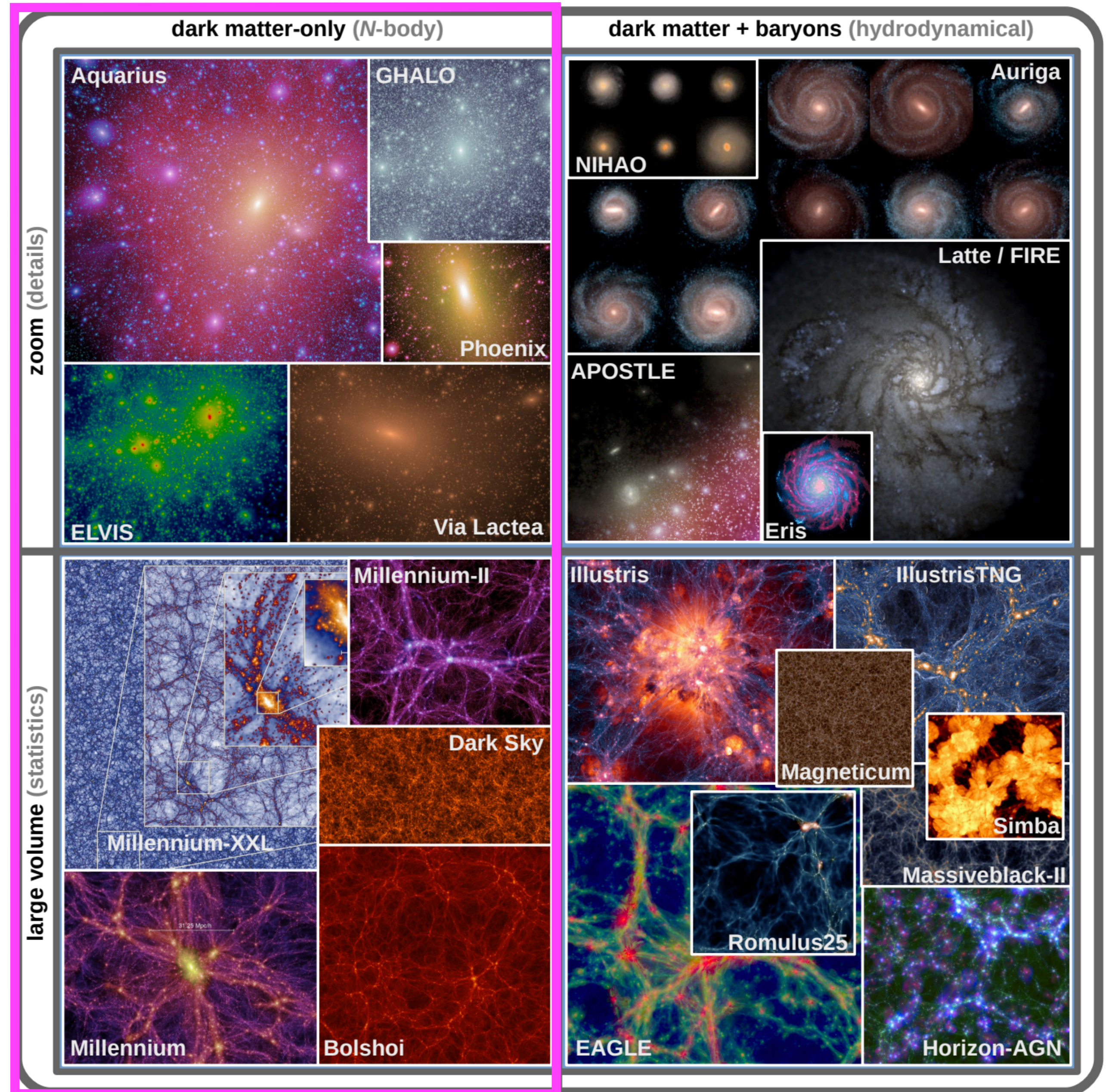
- Cosmological simulations are the **most accurate method** for modeling structure formation





# The Landscape of Cosmological Simulations

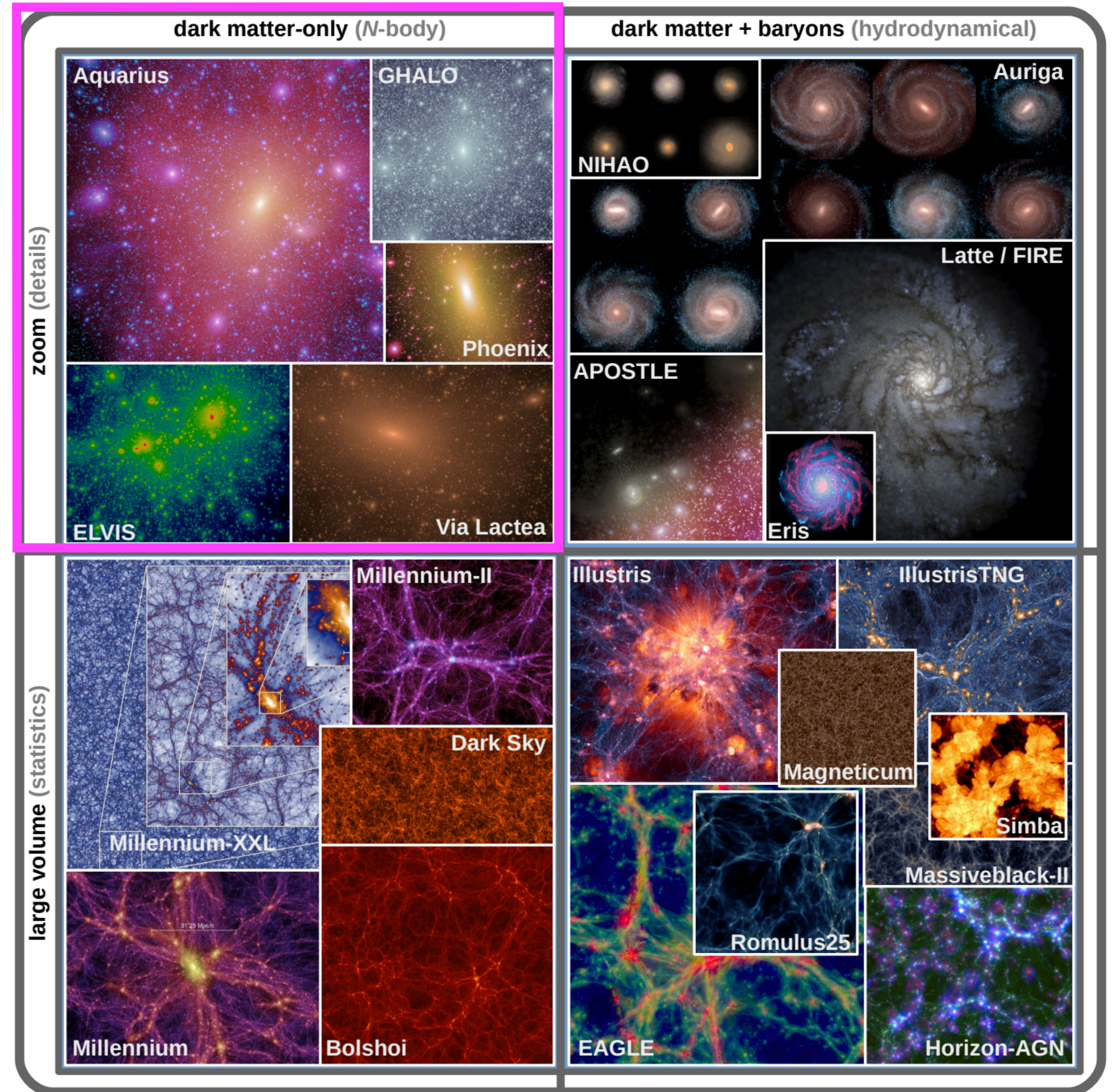
- Cosmological simulations are the **most accurate method** for modeling structure formation
- Dark matter-only simulations provide a **robust template** for more complex modeling





# The Landscape of Cosmological Simulations

- Cosmological simulations are the **most accurate method** for modeling structure formation
- Dark matter-only simulations provide a **robust template** for more complex modeling
- Zoom-in simulations focus on **small regions** of the universe at **high resolution**, allowing small-scale structure to be resolved

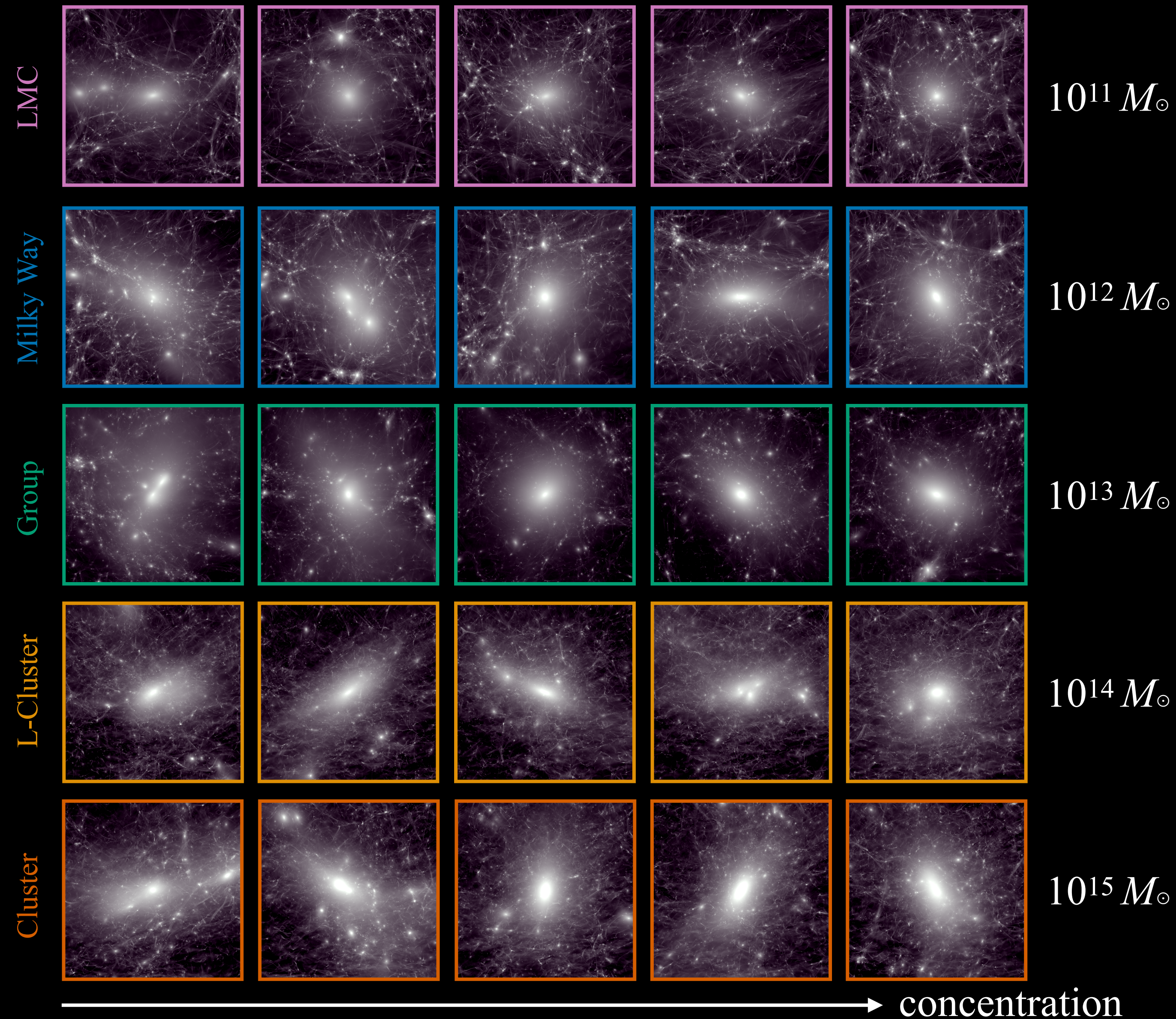




# Symphony Zoom-in Simulations

- **262** high-resolution cosmological zoom-in simulations spanning **4 decades of host halo mass**
- Includes the first large suites of **LMC** and **strong lens analog** host halos
- Run with a unified simulation and analysis code pipeline; all data is **publicly available!**

[web.stanford.edu/group/gfc/symphony](http://web.stanford.edu/group/gfc/symphony)





# Milky Way-est Zoom-in Simulations

- **25** high-resolution cosmological zoom-in simulations of **Milky Way-like systems**
- All realizations include **LMC analogs** on first infall and analogs of the ***Gaia-Enceladus*** merger

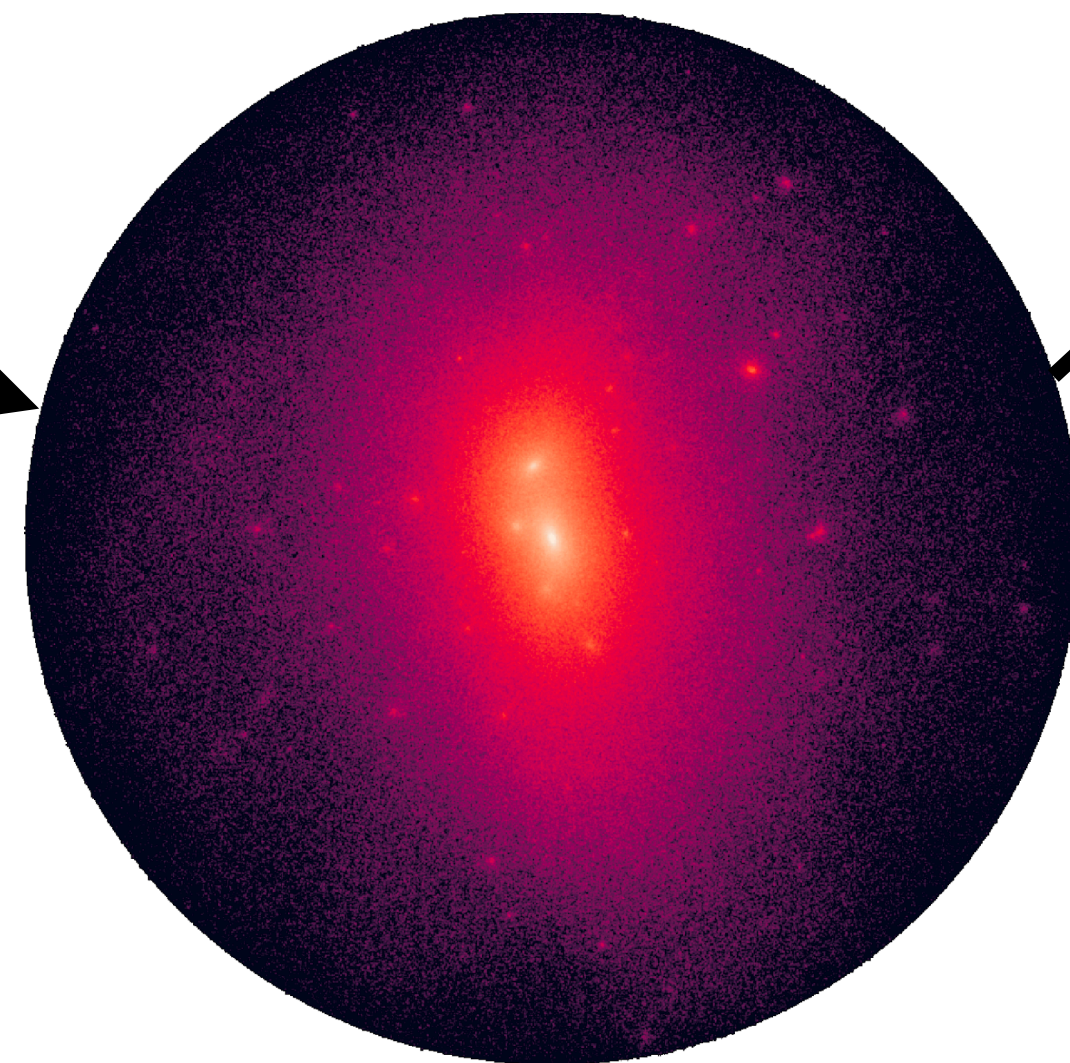
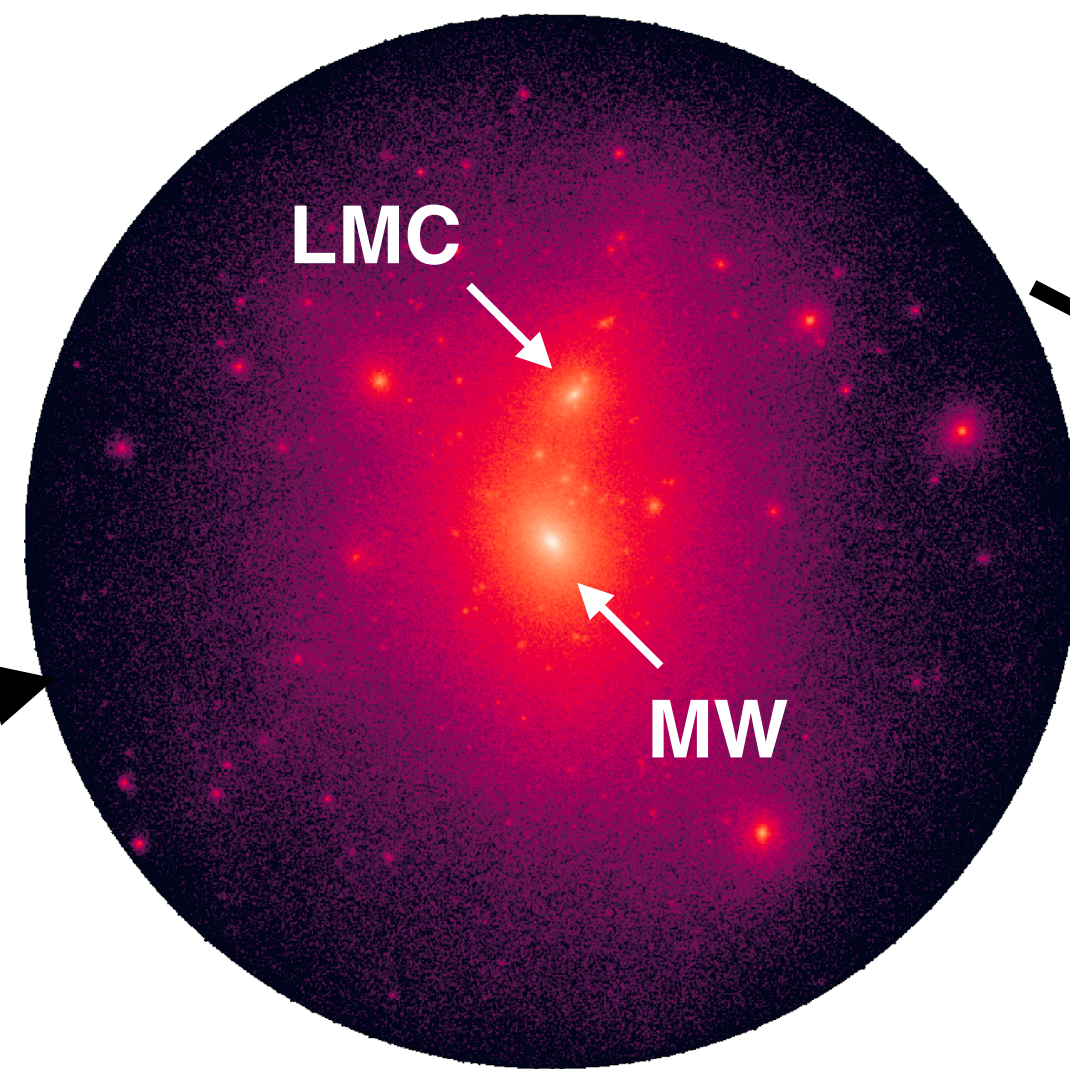
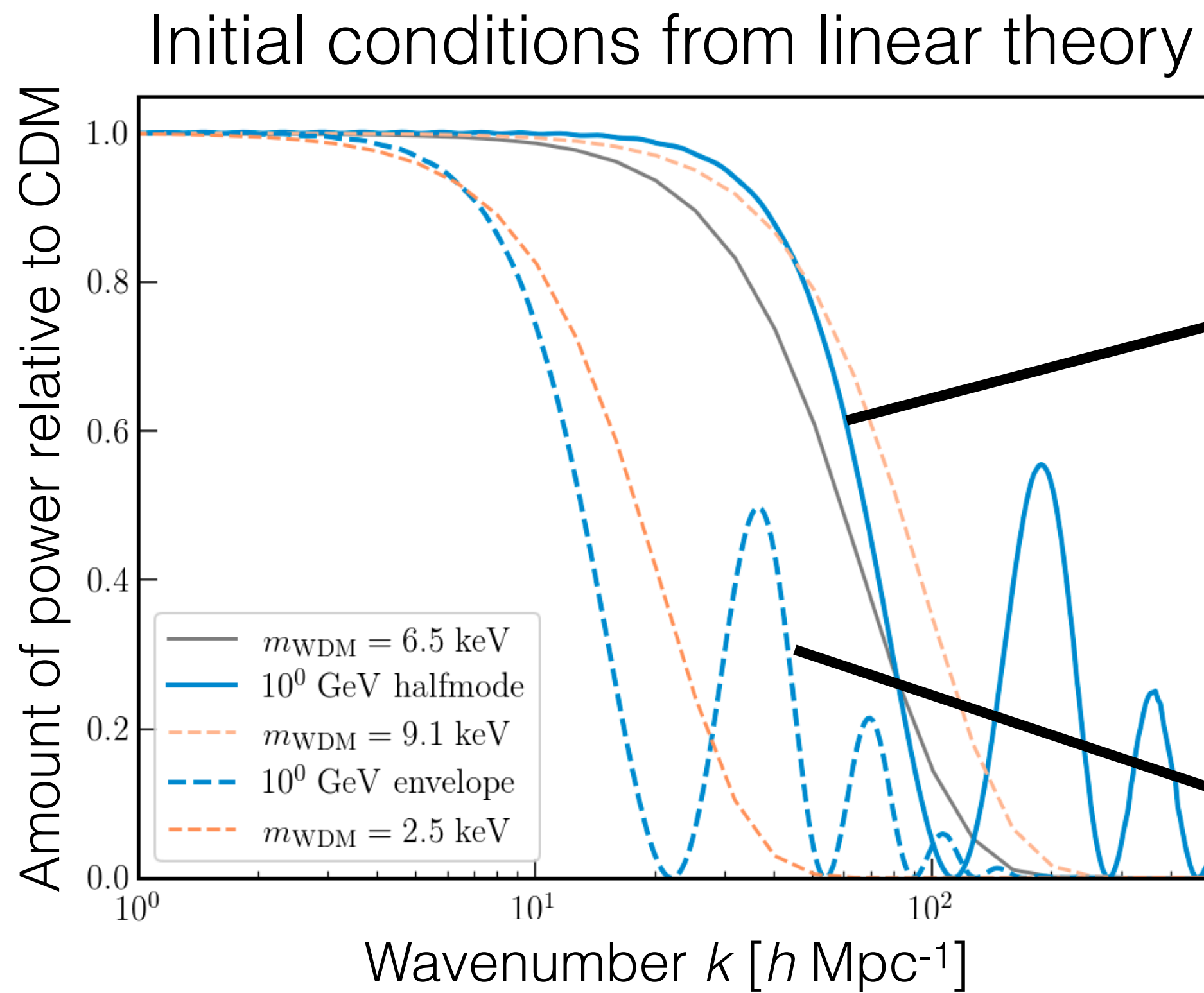


Deveshi Buch  
(Stanford)

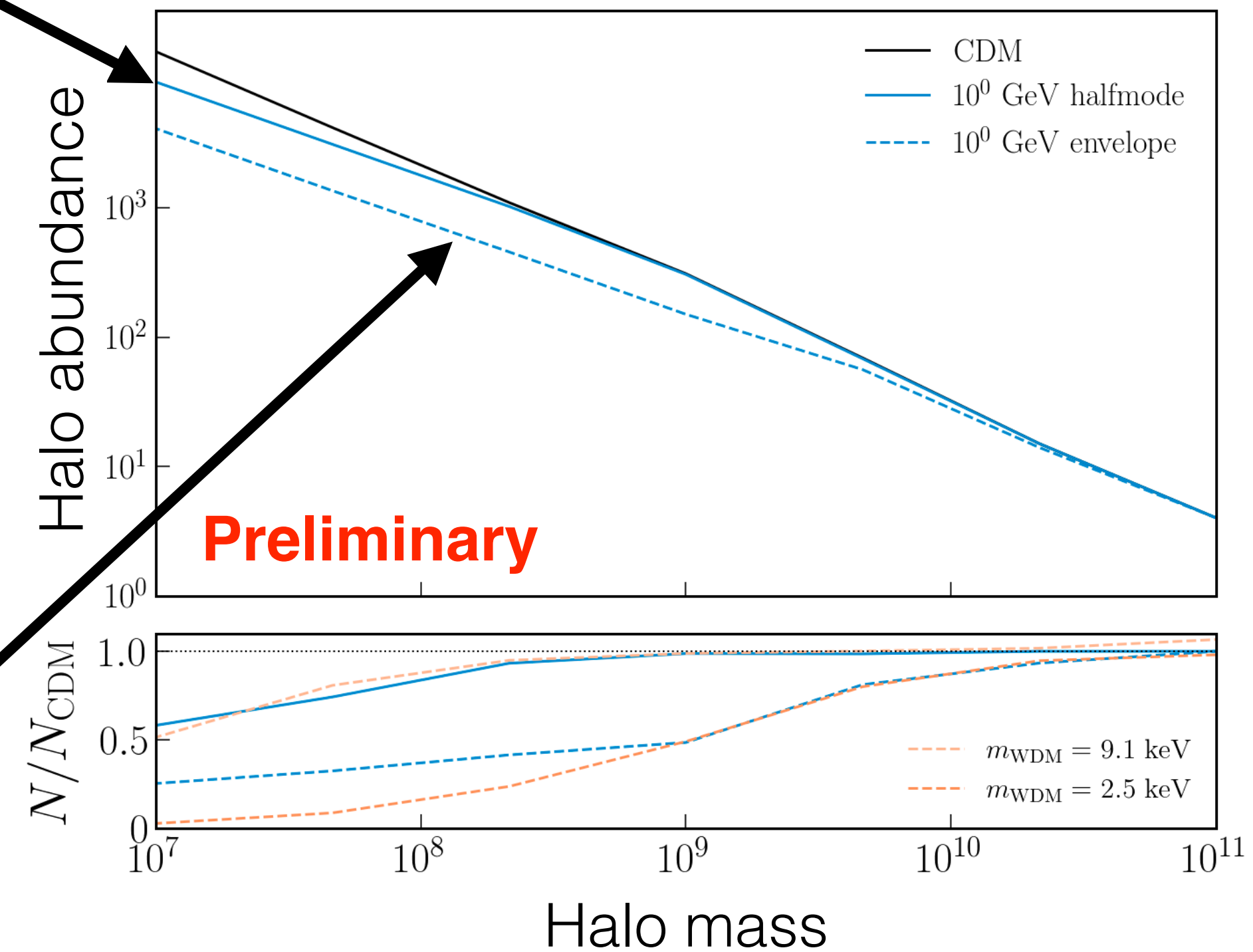




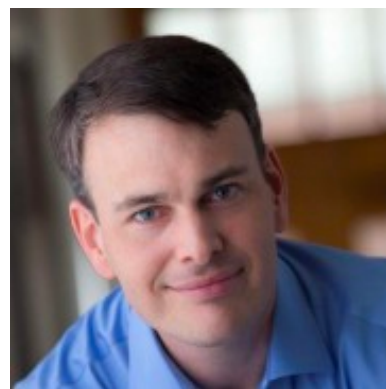
# Beyond-CDM Zoom-in Simulations



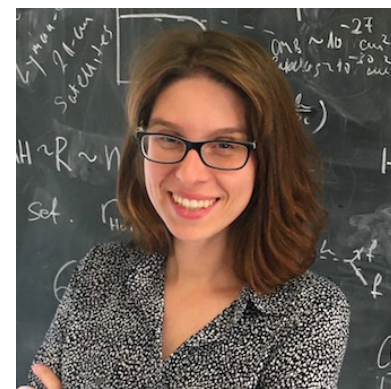
Halo and subhalo populations



Rui An  
(USC)



Andrew Benson  
(Carnegie)



Vera Gluscevic  
(USC)

**72 cosmological simulations** of Milky Way-like systems with initial conditions for warm, interacting, fuzzy dark matter



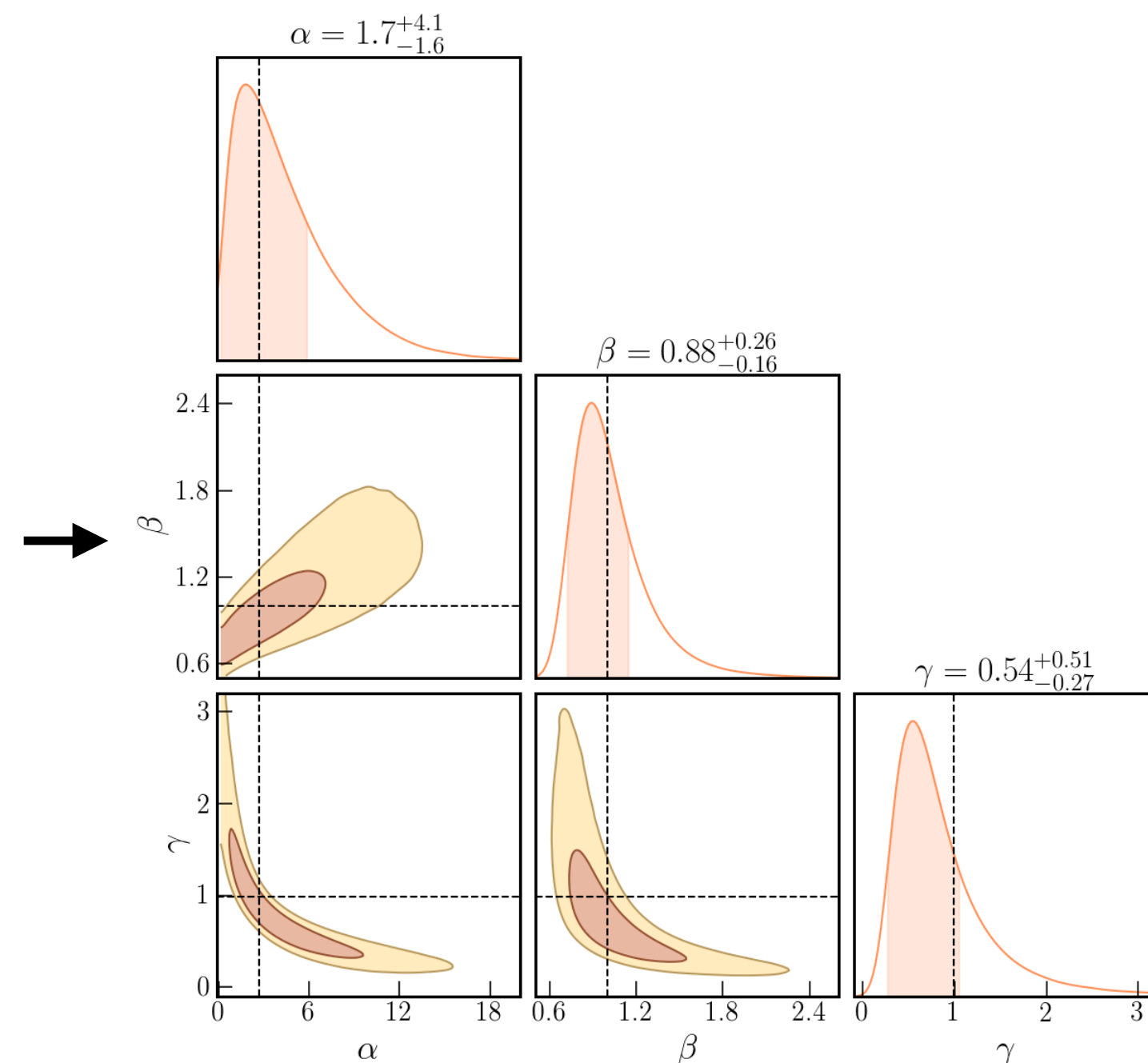
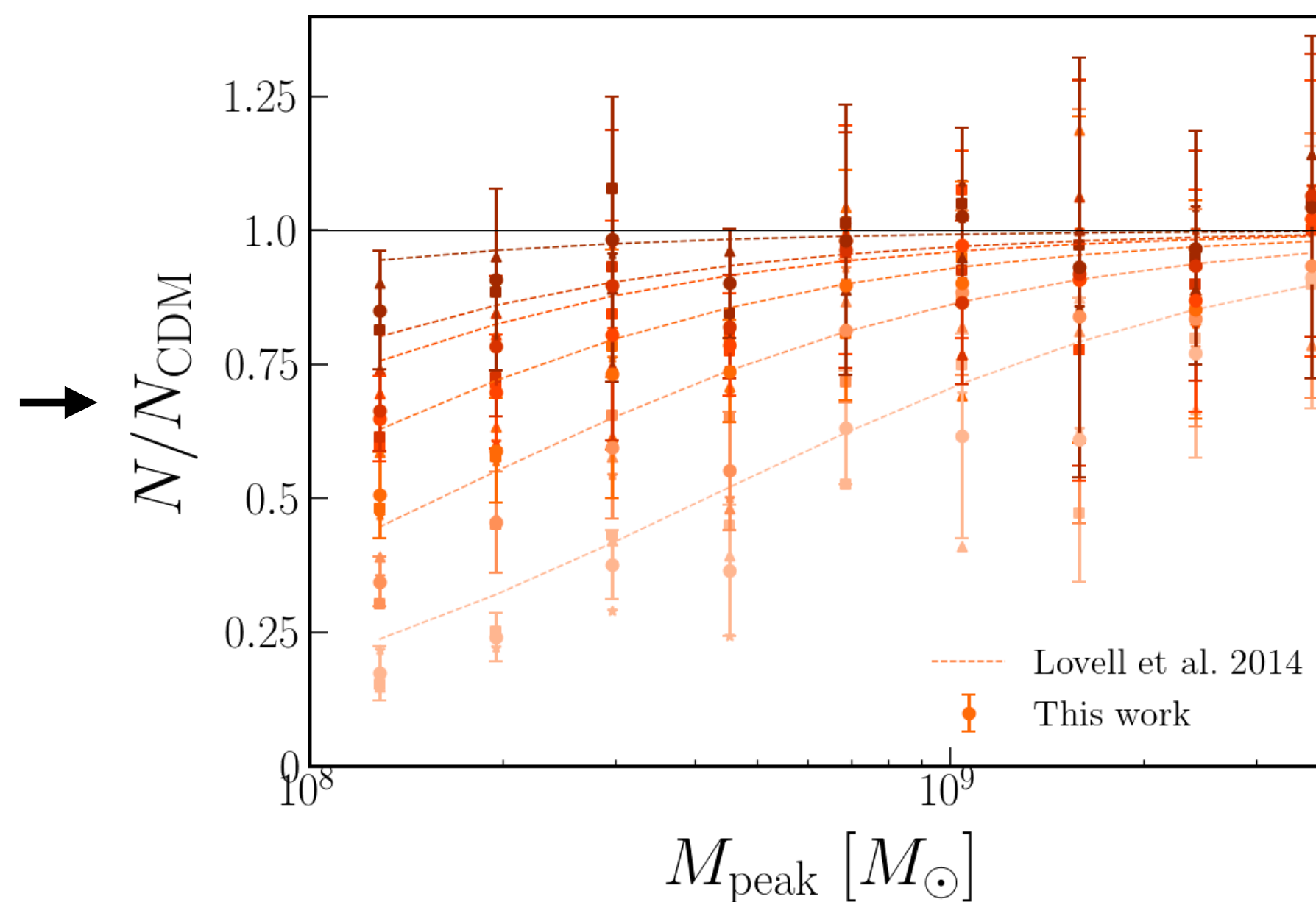
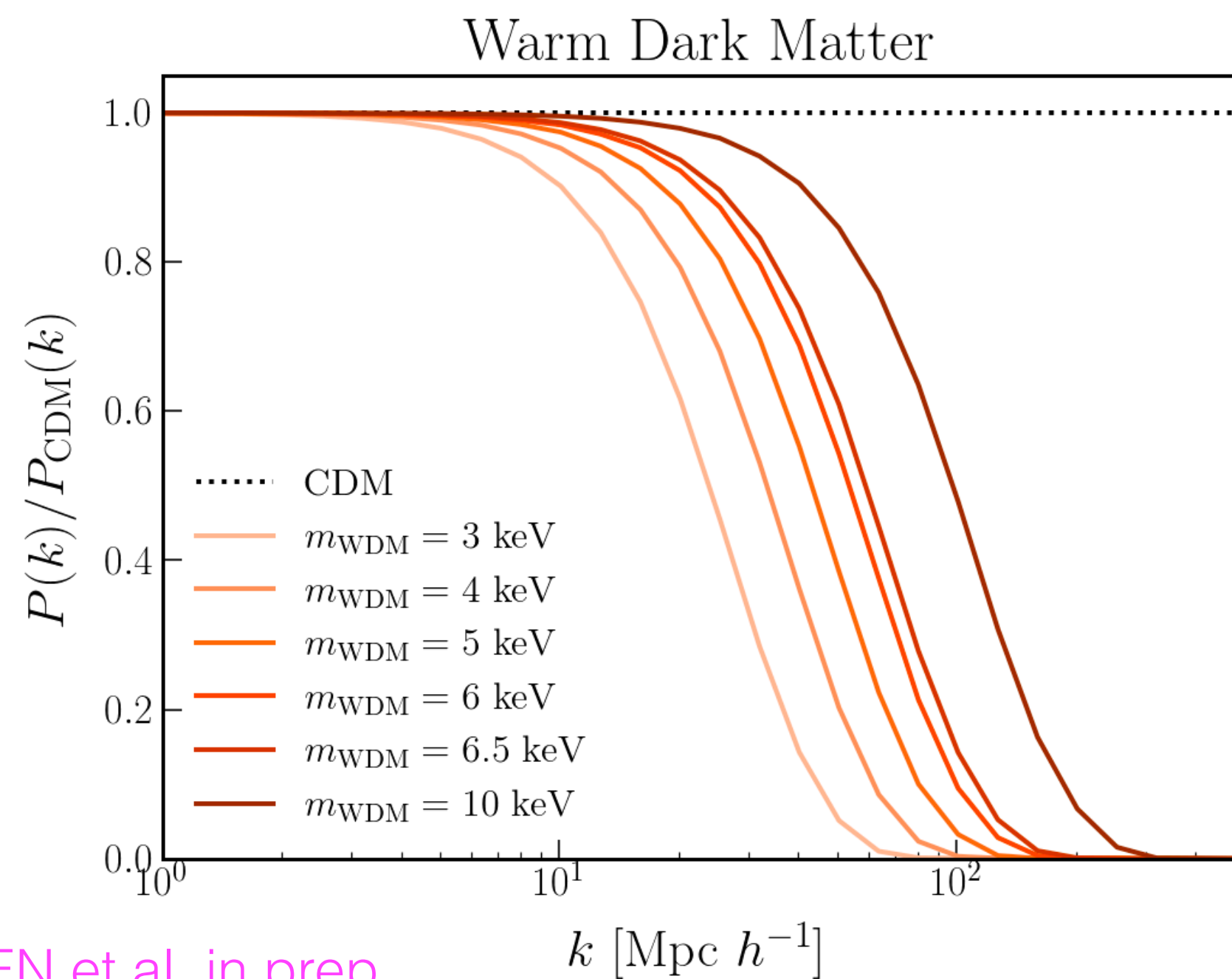
# Beyond-CDM Zoom-in Simulations

- Recalibrate WDM halo mass function suppression: full treatment of **statistical uncertainties, halo-to-halo scatter, fit degeneracies**; integrated with **CLASS**

- Halo mass function suppression slightly enhanced relative to previous fits

- Extremely small contamination from artificial halos

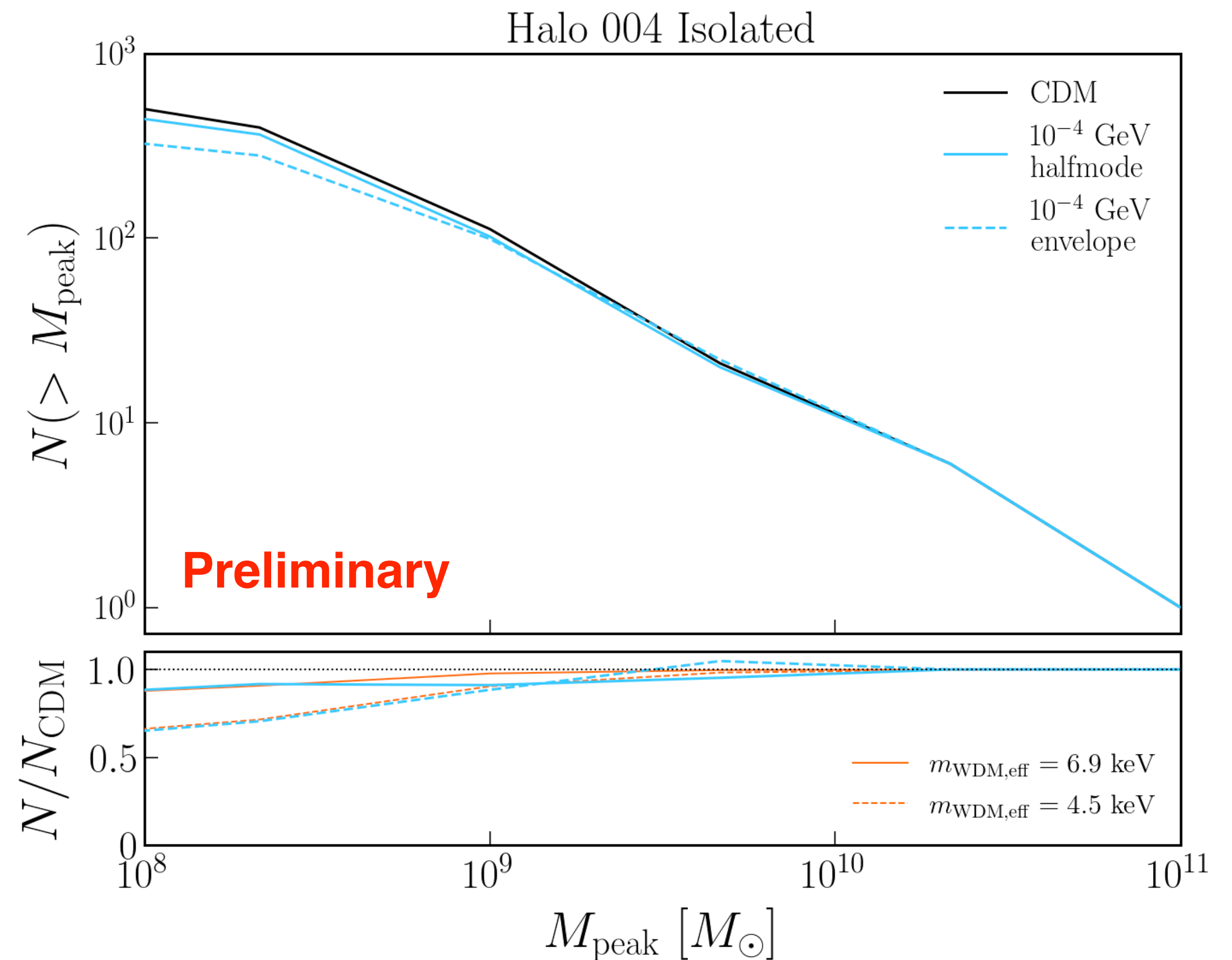
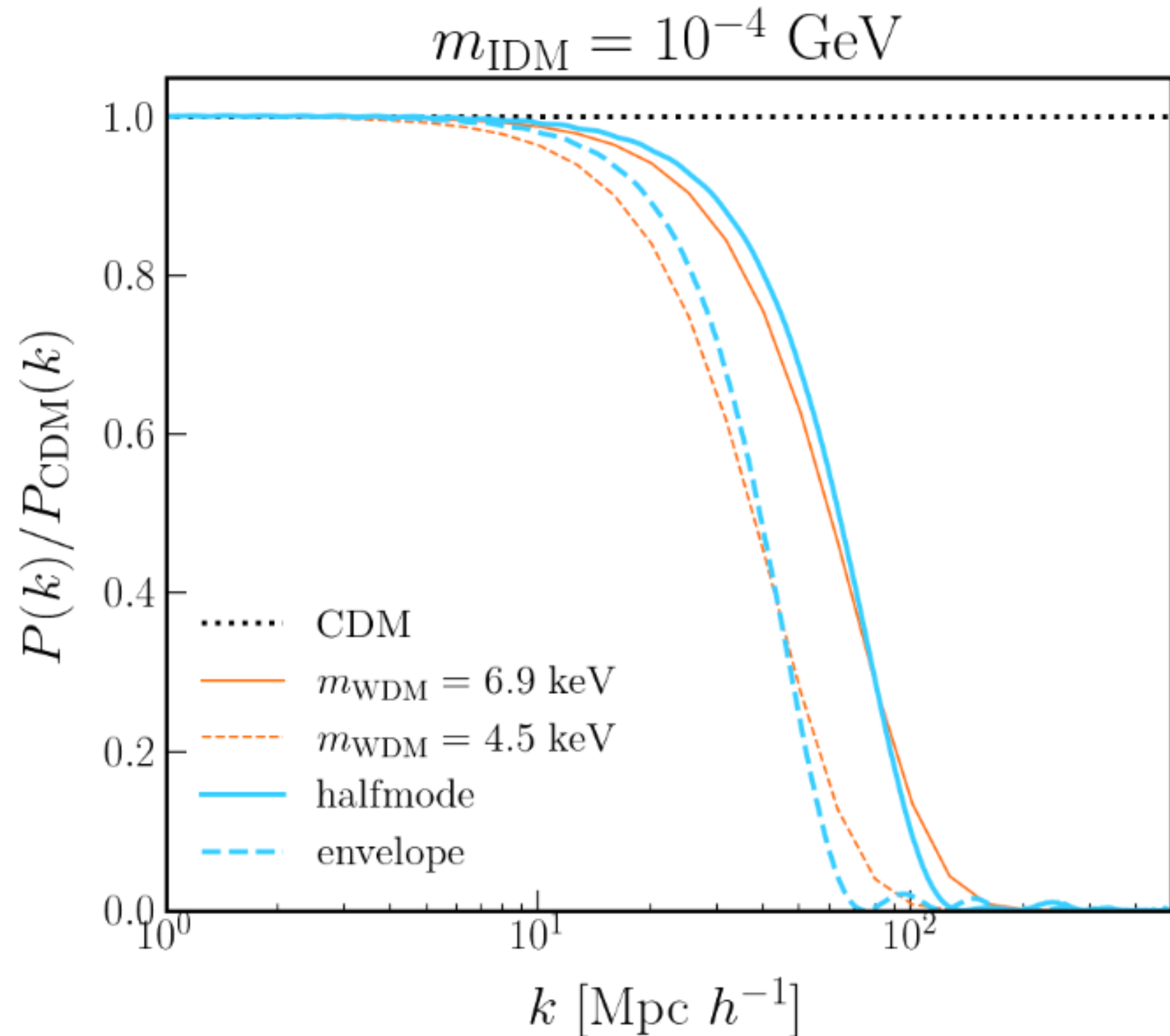
$$\frac{(dn/dM)_{\text{WDM}}}{(dn/dM)_{\text{CDM}}} = \left[ 1 + \left( \frac{\alpha M_{\text{hm}}(m_{\text{WDM}})}{M} \right)^{\beta} \right]^{-\gamma}$$





# Beyond-CDM Zoom-in Simulations

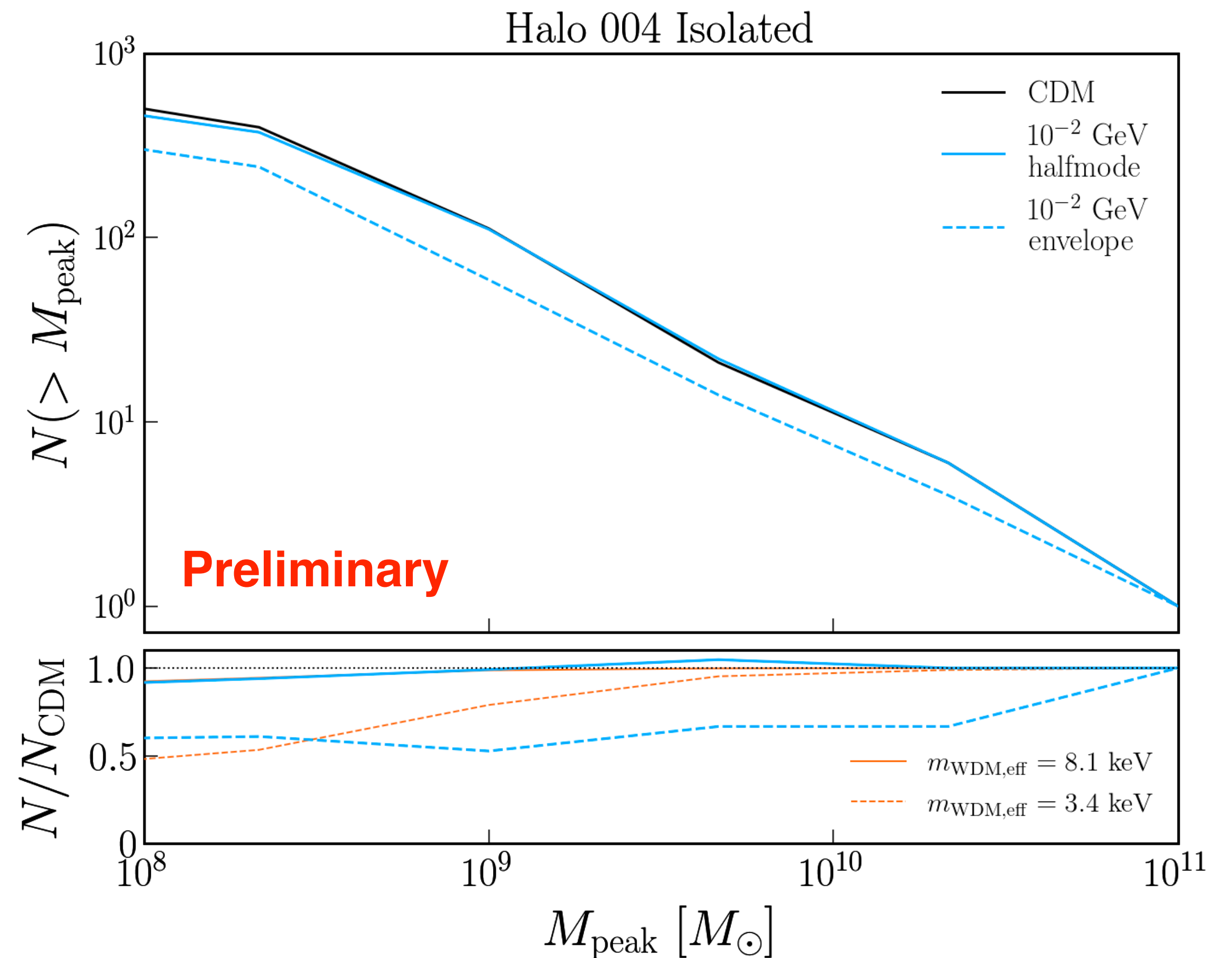
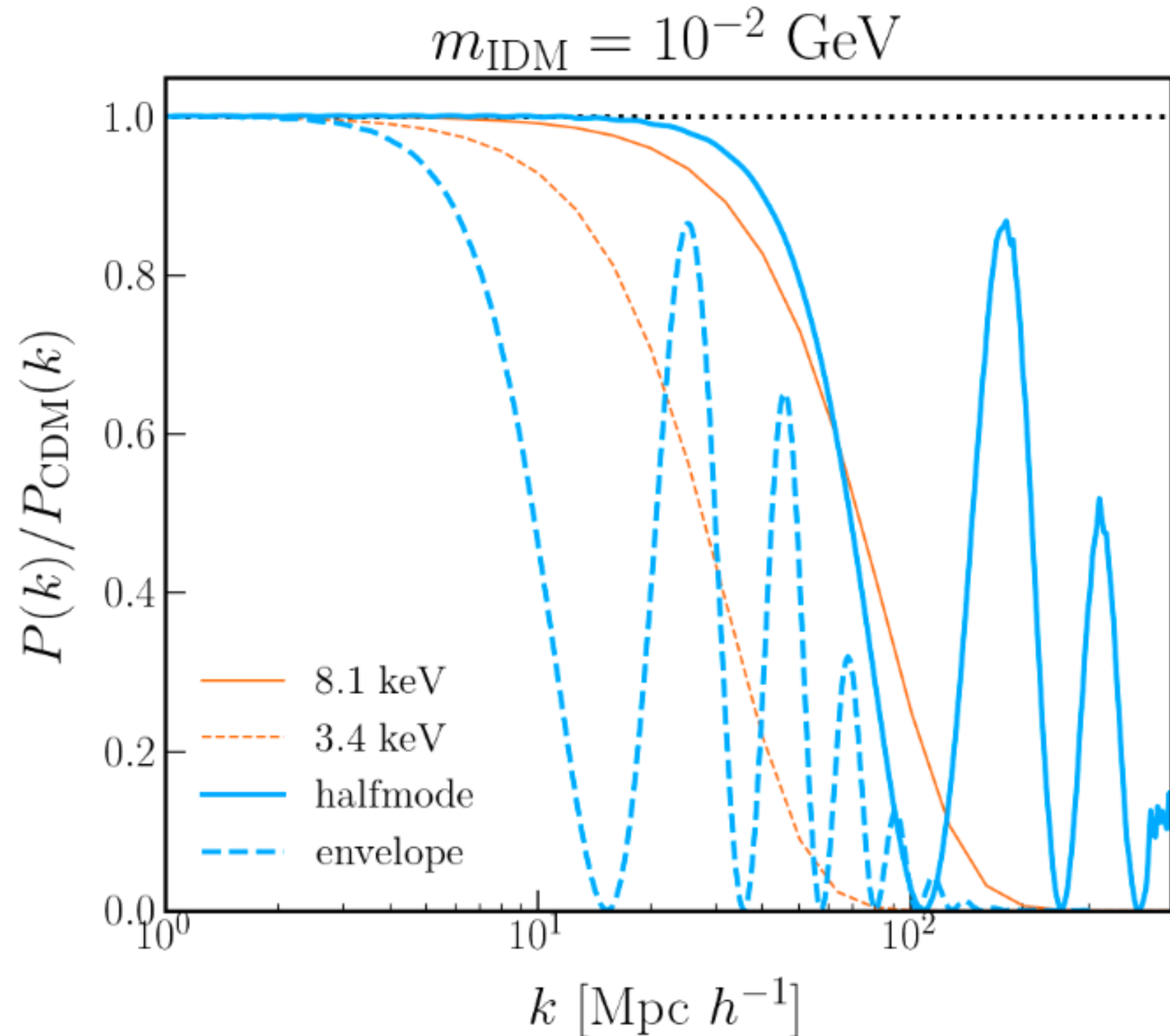
- Interacting dark matter models with **small dark acoustic oscillations** map to effective WDM models:





# Beyond-CDM Zoom-in Simulations

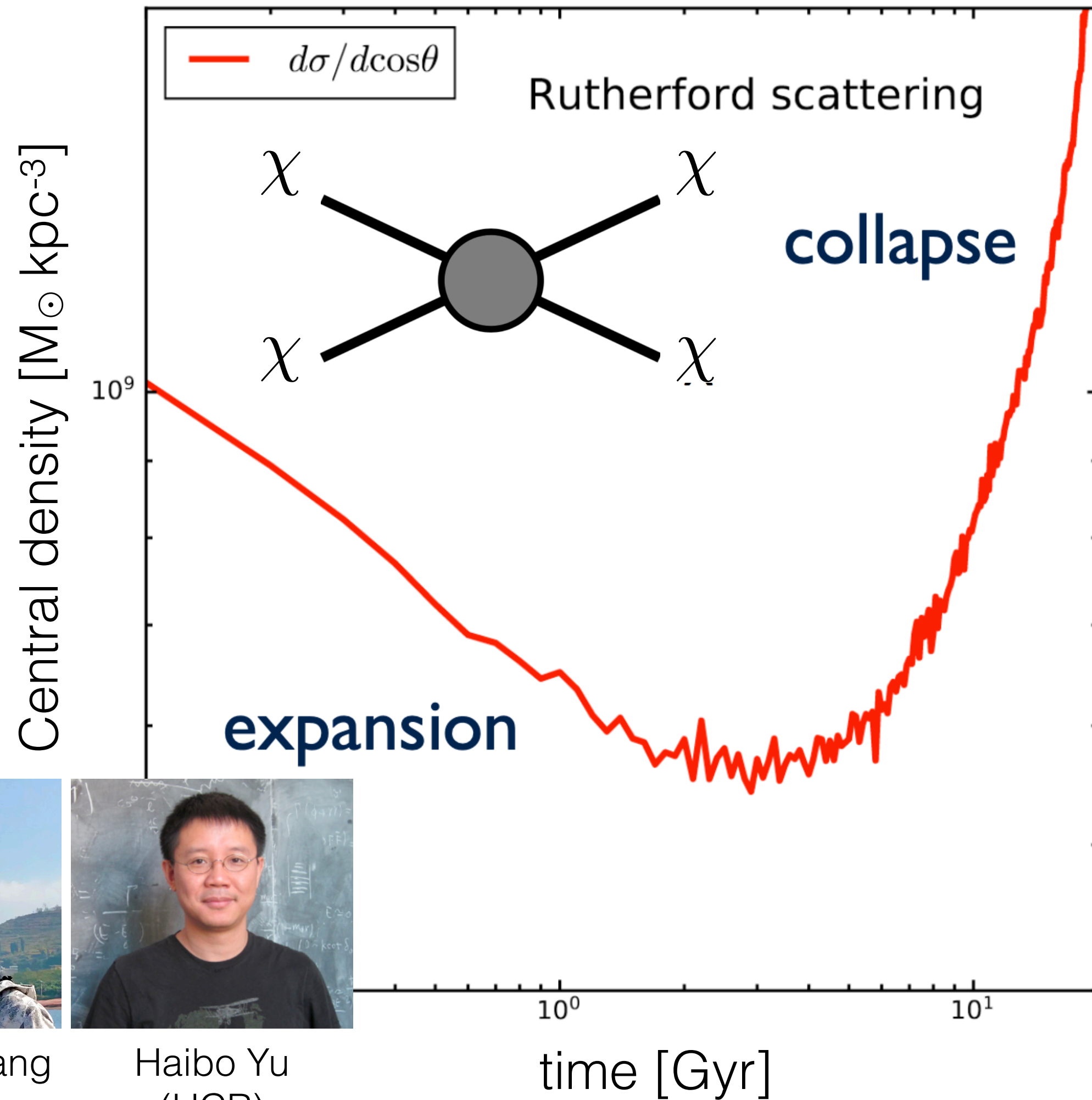
- Interacting dark matter models with **large dark acoustic oscillations** are “colder” than WDM models with the same initial cutoff:



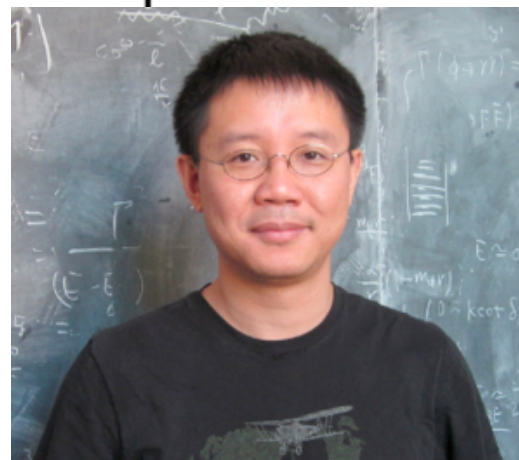


# Signatures of Strong Dark Matter Self-interactions

Yang & Yu 2022 (2205.03392)

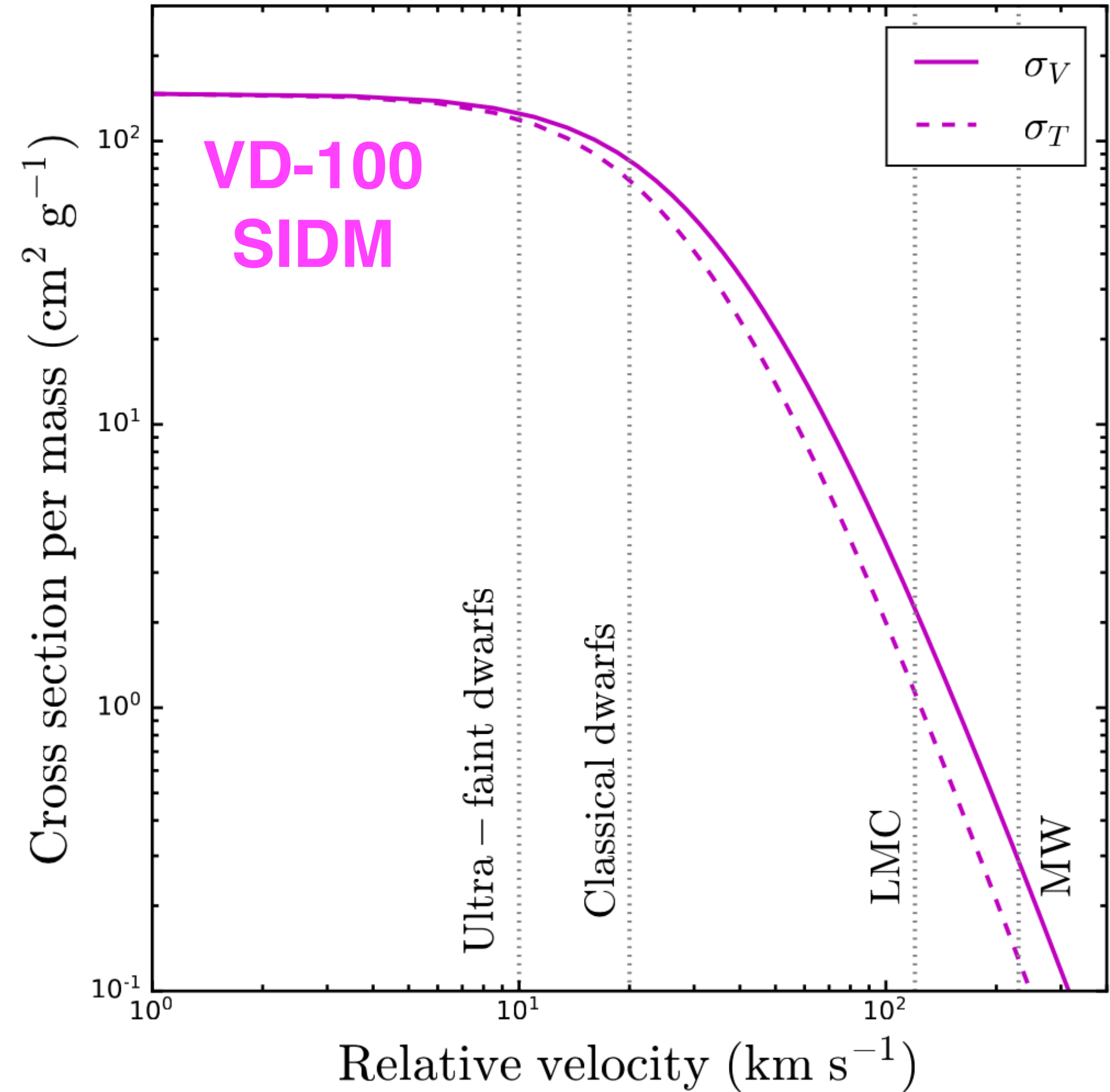


Daneng Yang  
(UCR)



Haibo Yu  
(UCR)

Yang, EN, Yu 2023 (2211.13768)



Strong, velocity-dependent self-interactions  $\rightarrow$  **core-collapse** in small halos & core-formation in large halos

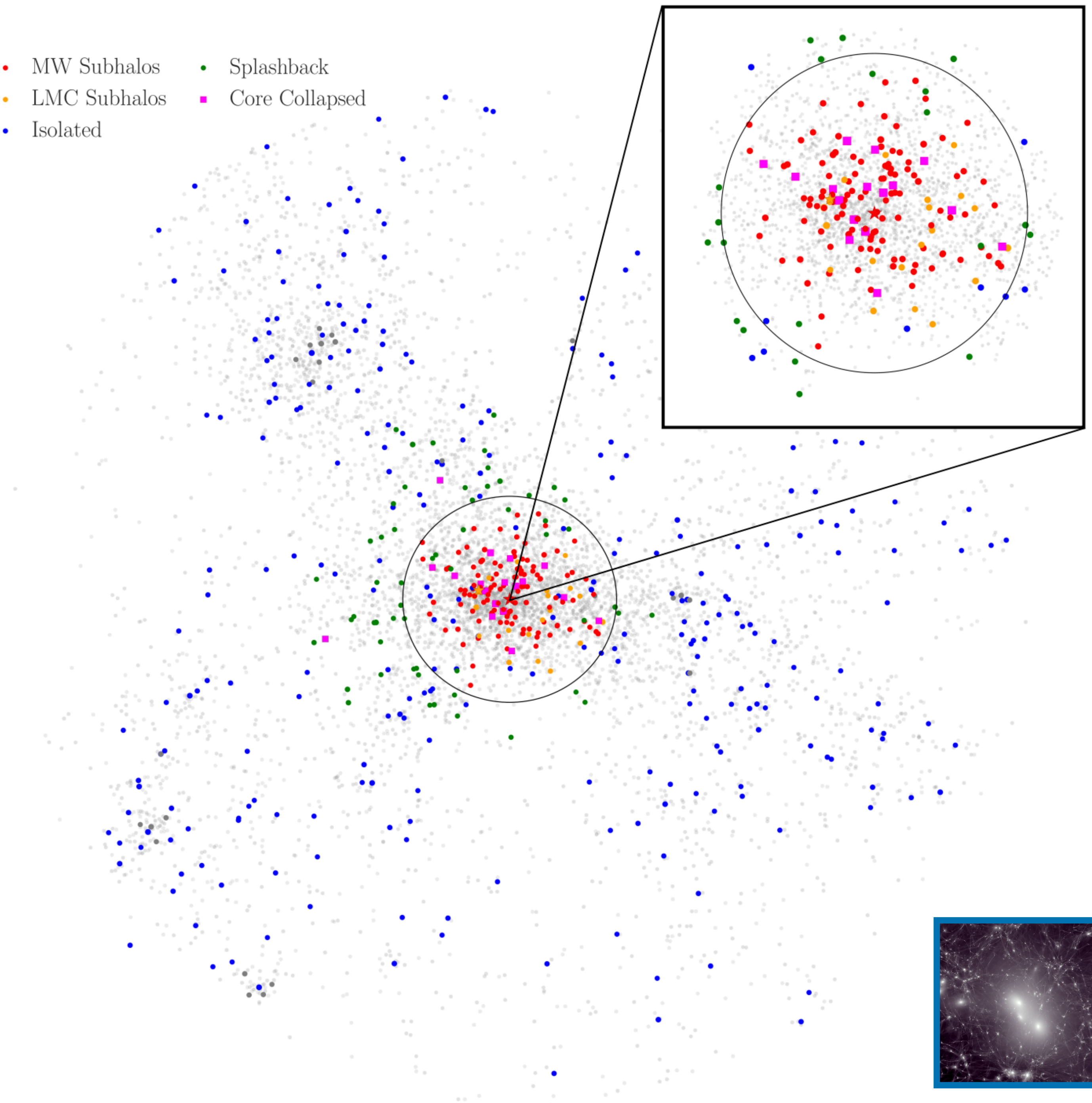


# VD-100 SIDM

## Milky Way Simulation

- Extremely high-resolution MW zoom-in with **strong, velocity-dependent self-interactions**
- Self-consistent analysis of halos in **all environments** throughout high-resolution volume
- **Deep core-collapse** in  $\sim 10\%$  of isolated halos,  $\sim 20\%$  of subhalos down to  $10^8 M_{\odot}$

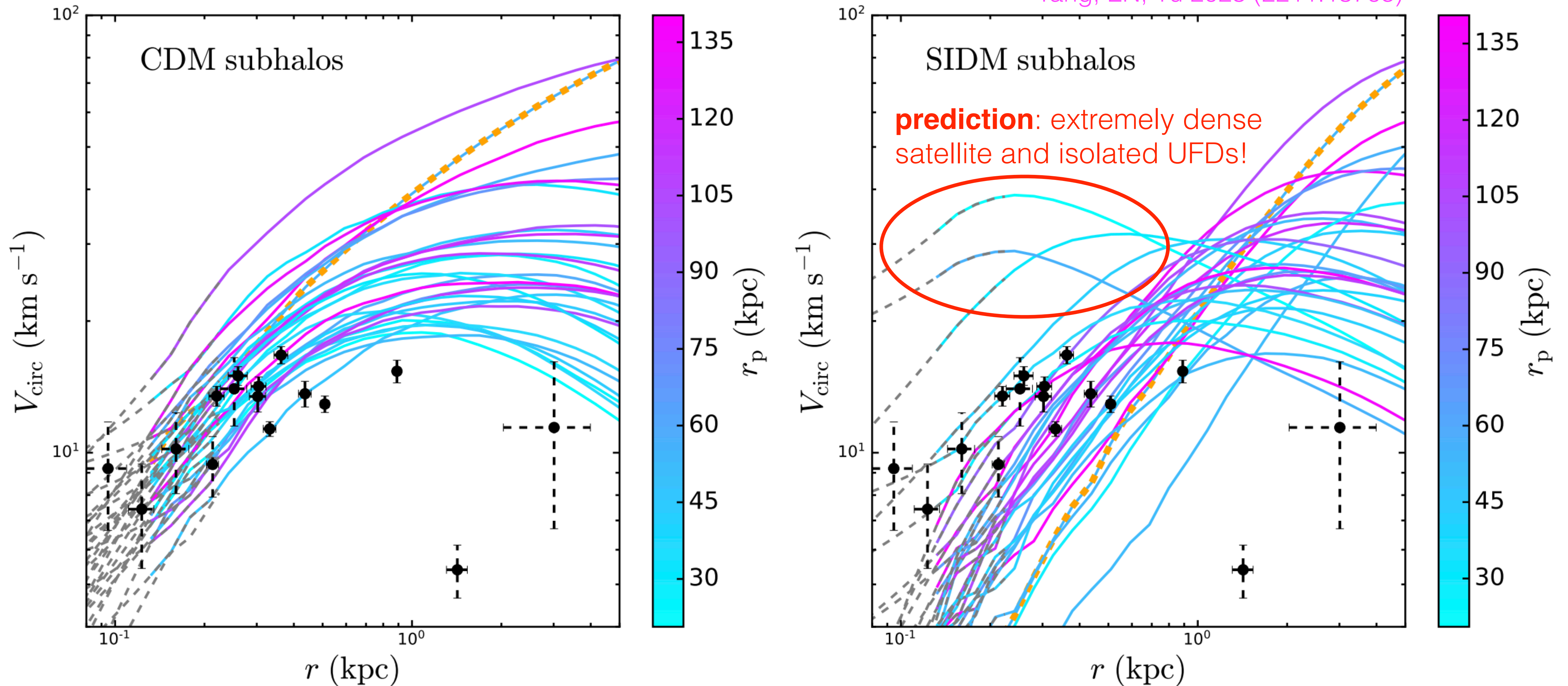
- MW Subhalos
- LMC Subhalos
- Isolated
- Splashback
- Core Collapsed





# VD-100 SIDM Milky Way Simulation

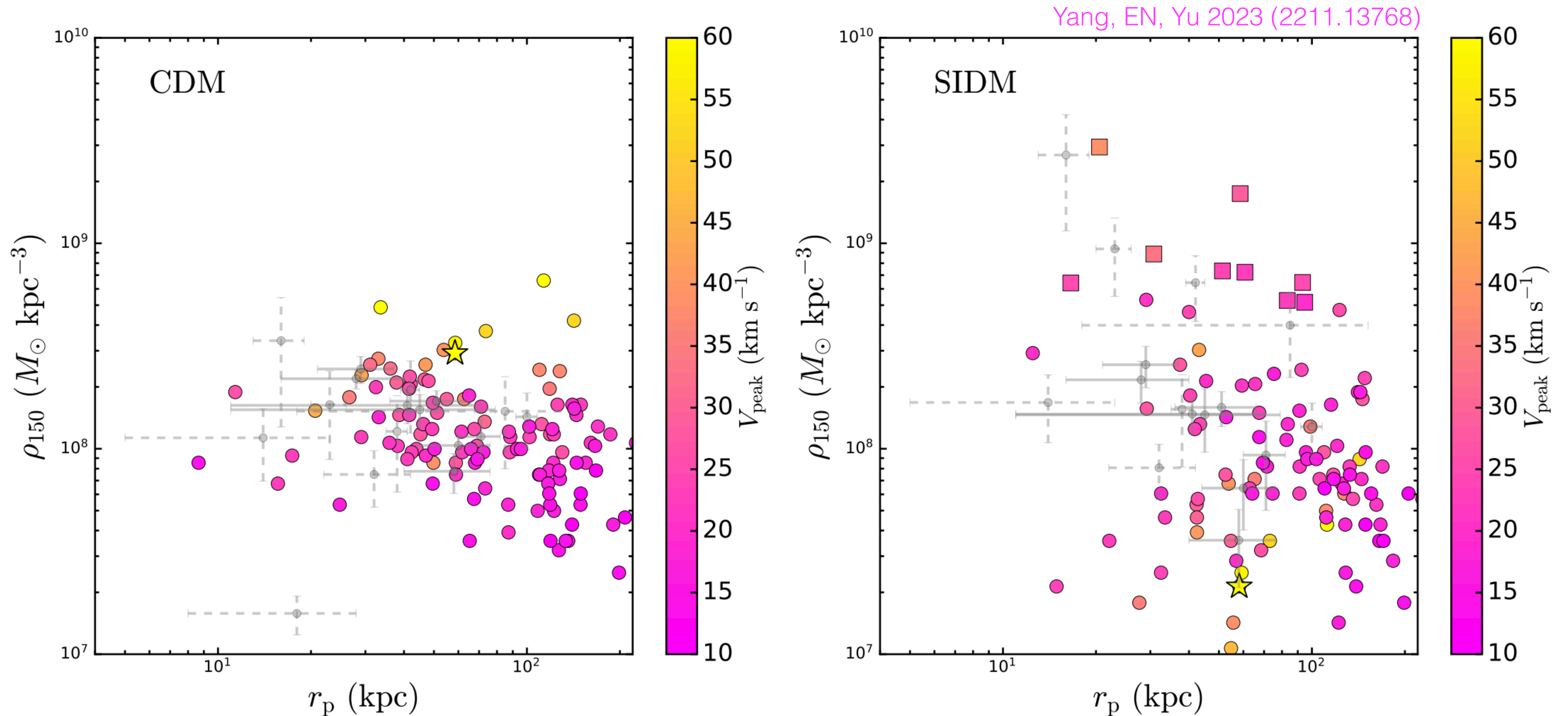
Yang, EN, Yu 2023 (2211.13768)



VD-100 subhalos are **more diverse** than in CDM, alleviating too big to fail problem for brightest systems



# VD-100 SIDM Milky Way Simulation

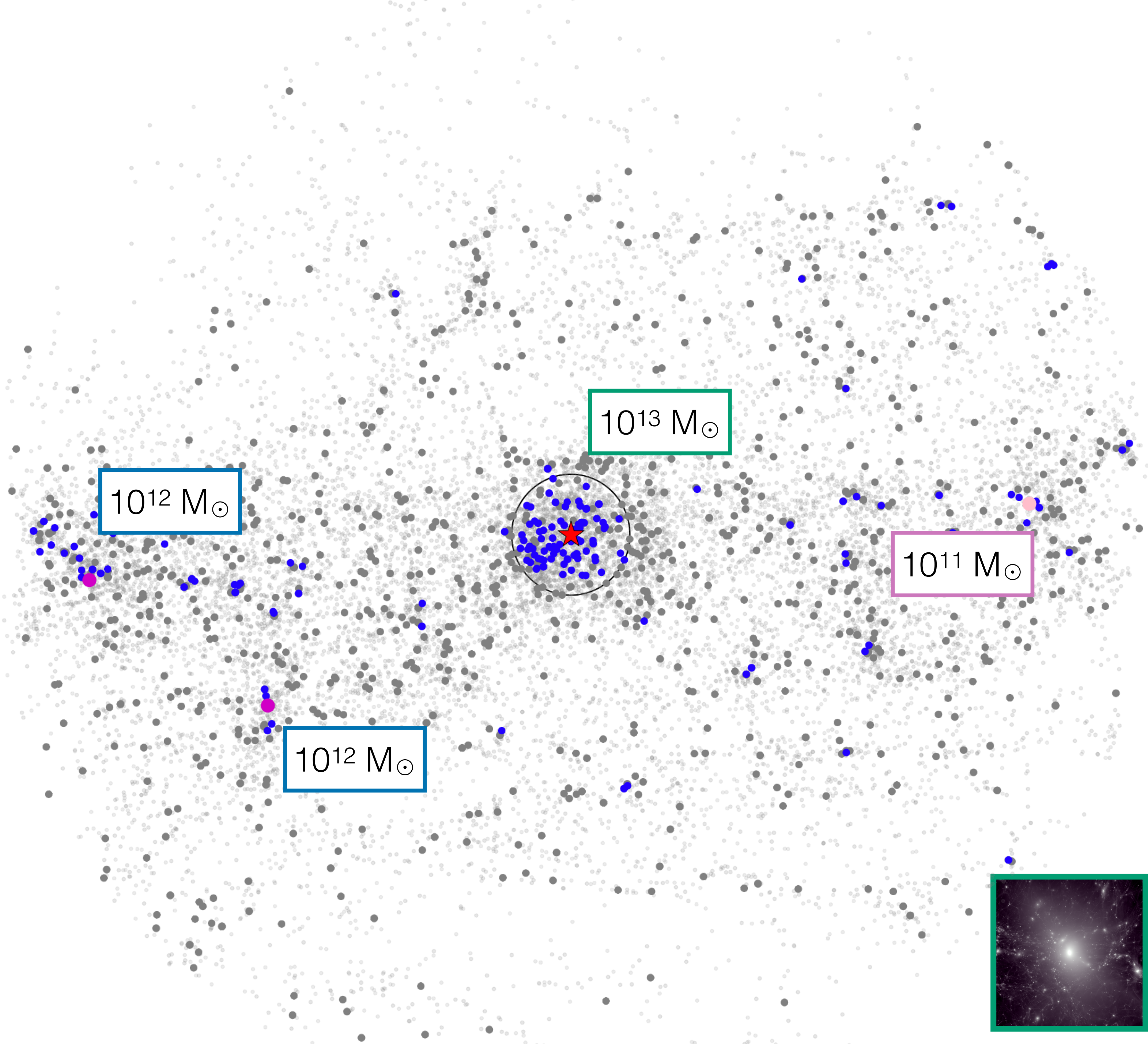
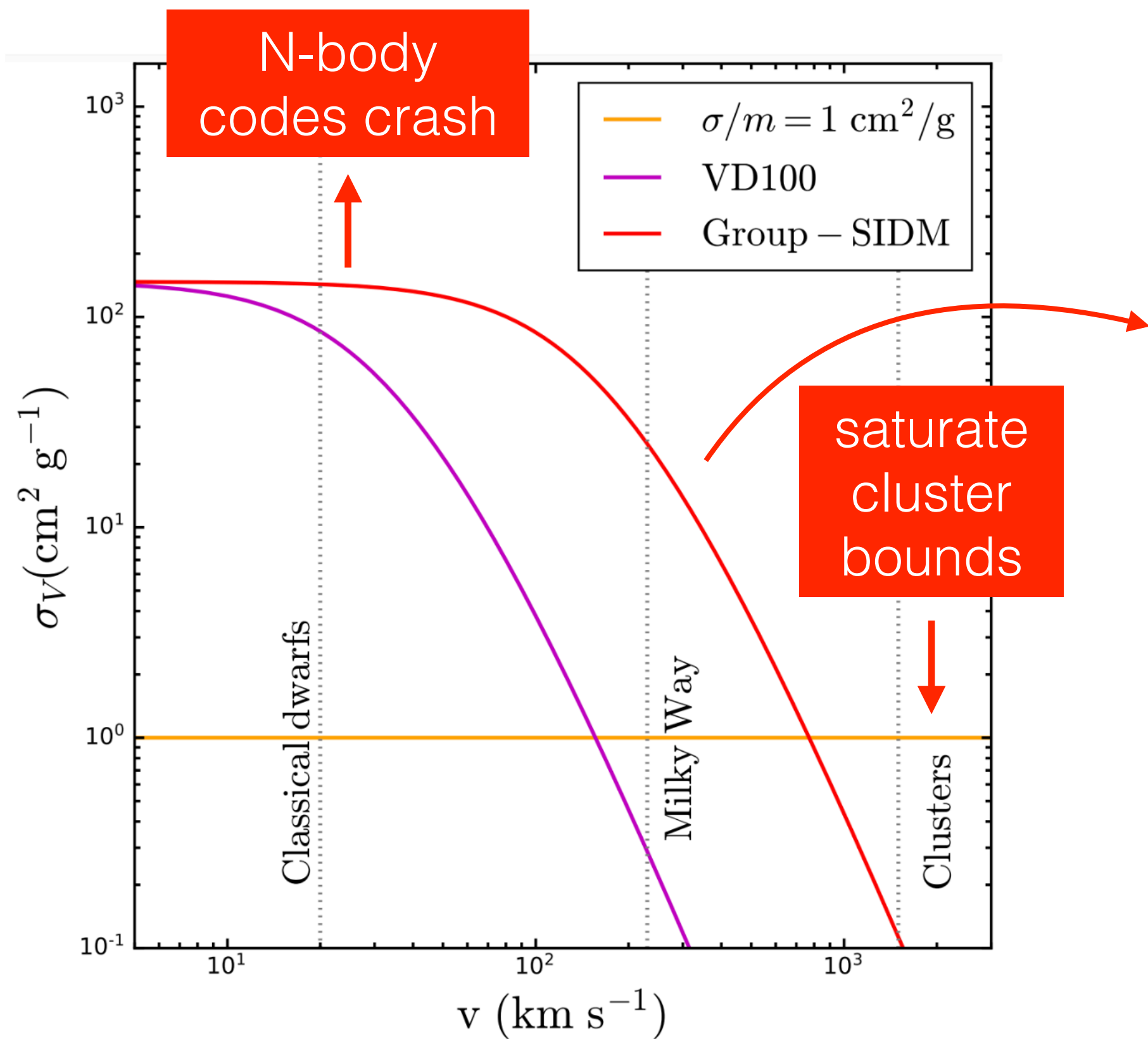


VD-100 diversifies central density-pericenter relation; velocity-independent interactions erase anti-correlation



# Group-SIDM Strong Lens Analog Simulation

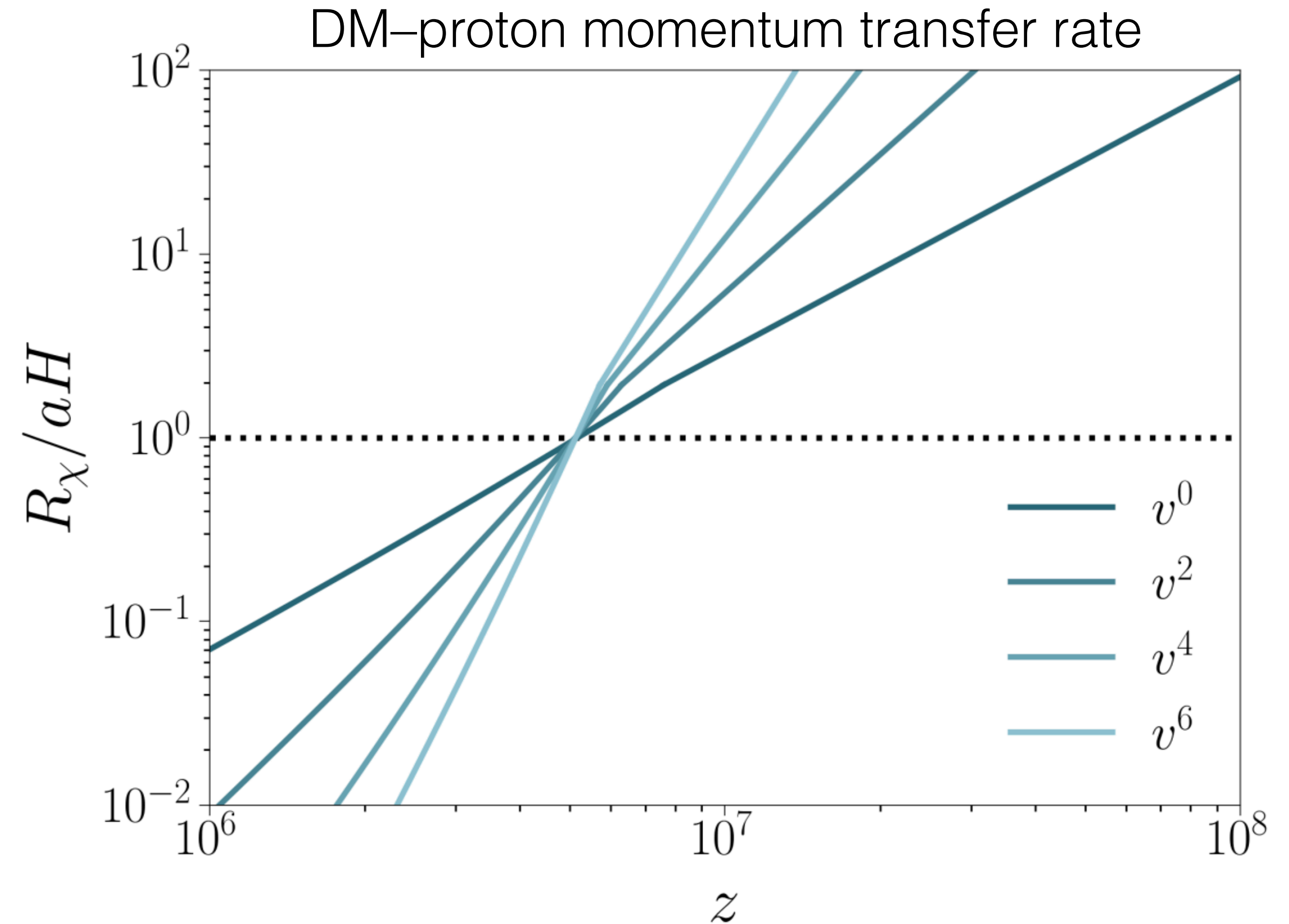
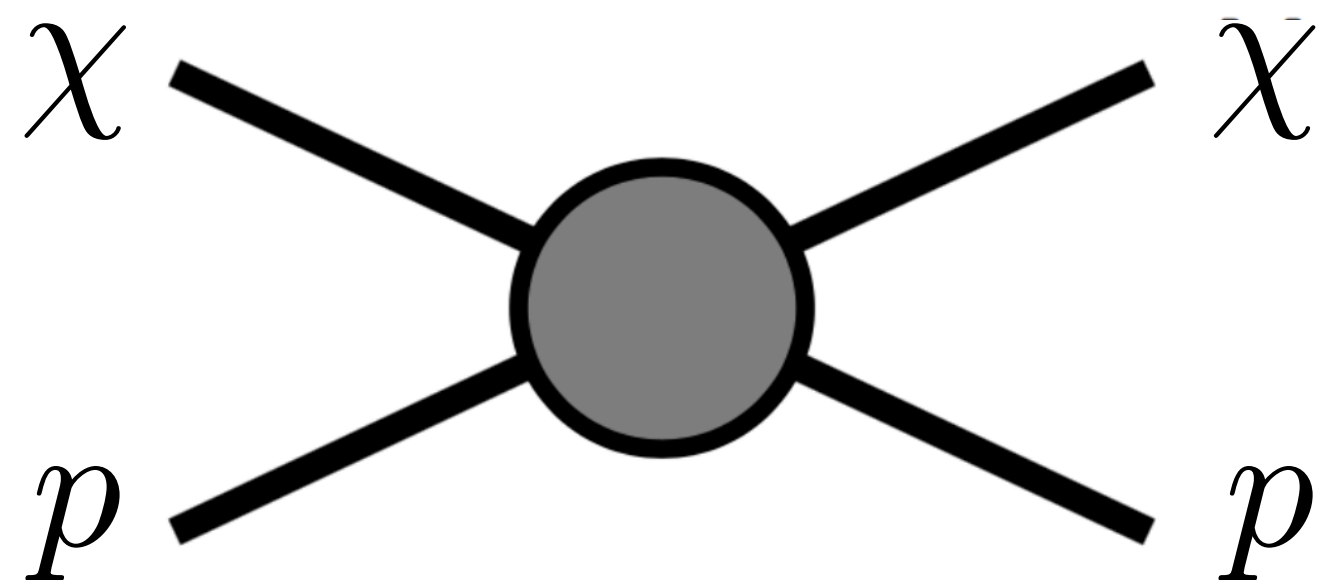
- **First** group-scale simulation with core-collapse (see Haibo Yu's talk)





# Simulating Dark Matter–Baryon Interactions

- Late-time dark matter–baryon scattering is constrained to be **rare**
- Idea: simulate DM–baryon scattering with an N-body algorithm, analogous to SIDM; **first implementation of this physics!**
- Unlike SIDM, these interactions couple DM to a hot, collisional species





# Simulating Dark Matter–Baryon Interactions



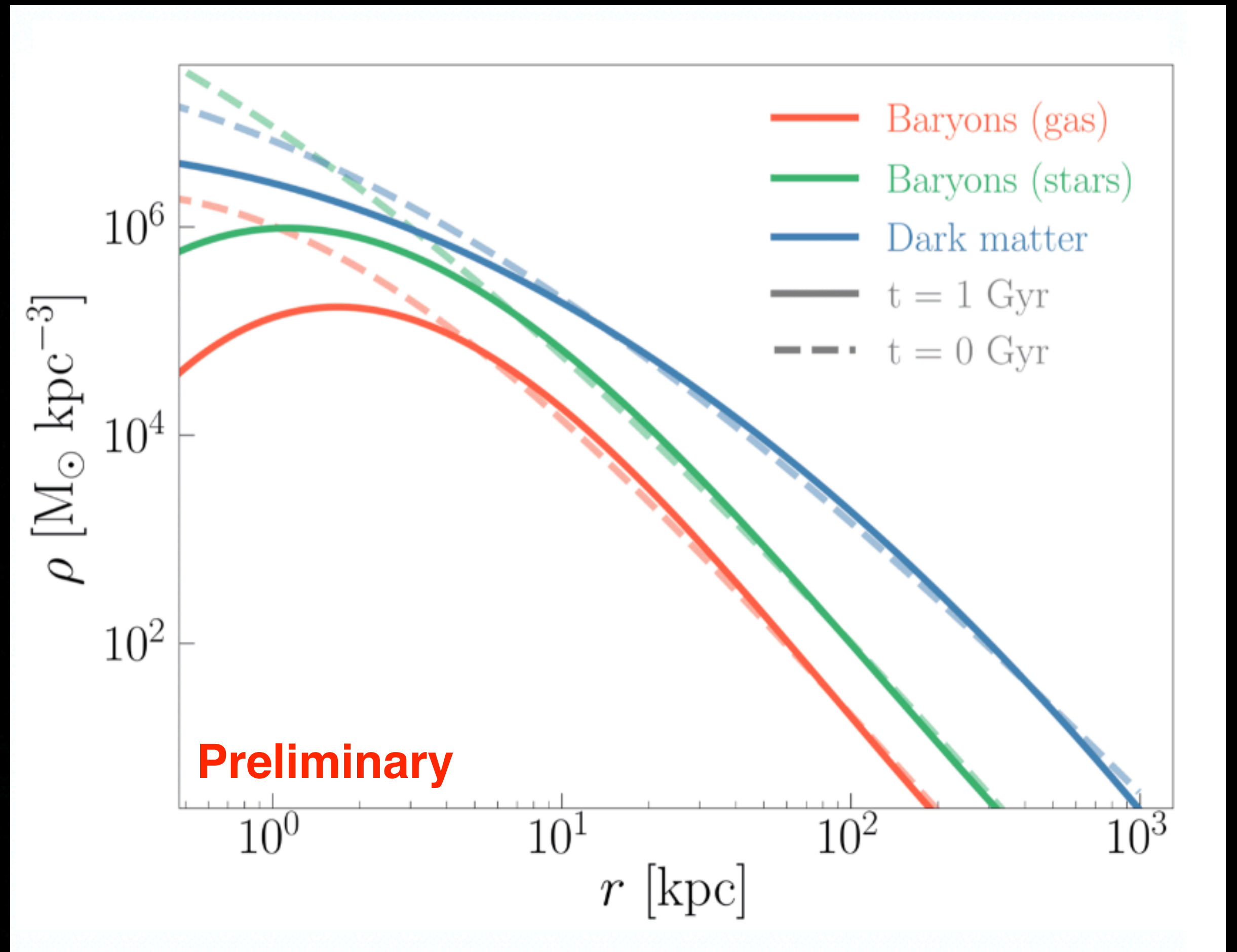
Karime Maamari  
(USC)

DM  
Star  
Gas

Maamari, EN, Gluscevic in prep.

$\sigma_{DM}$ :  $5e-3 \text{ cm}^2\text{g}^{-1}$   
Scale: 200 kpc  
npart: 100000, 100000  
Mass:  $1.3e8, 8.6e5 \text{ Msol}$

$t = 0.0000 \text{ Gyr/h}$



Dark matter–baryon interactions thermalize DM in inner regions, phenomenologically similar to SIDM



- **Symphony**: **262** high-resolution cosmological zoom-in simulations, spanning four decades of host halo mass, including the first suites of **LMC** and **strong lens analog** hosts
- **Milky Way-est**: **25** high-resolution cosmological zoom-in simulations of **Milky Way-like systems**, including realistic **LMC** and **Gaia-Enceladus analogs**
- **Beyond-CDM**: **72** high-resolution cosmological zoom-in simulations of Milky Way systems with initial conditions appropriate for **warm, interacting, fuzzy** DM
- **VD-100 SIDM**: extremely high-resolution Milky Way-like system with **strong, velocity-dependent self-interactions** yields **diverse halo populations**
- Simulating dark matter–baryon interactions: **first implementation** of this physics, with hints of SIDM-like signatures

