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Structure Formation in the Universe

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In its most basic form, the highly successful Λ CDM cosmology can be encapsulated in six parameters. Once these parameters are specified, so too is a wide variety of phenomena, from fluctuations in the microwave background to the growth of structure to the evolution of the expansion rate of the Universe. I will review the predictions related to cosmological structure formation, focusing on areas where potential tensions have emerged: these include nonlinear scales at the earliest times and smallest mass scales and comparisons between measurements early and late in the Universe's history.

Author: BOYLAN-KOLCHIN, Mike **Presenter:** BOYLAN-KOLCHIN, Mike

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UCLA)

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