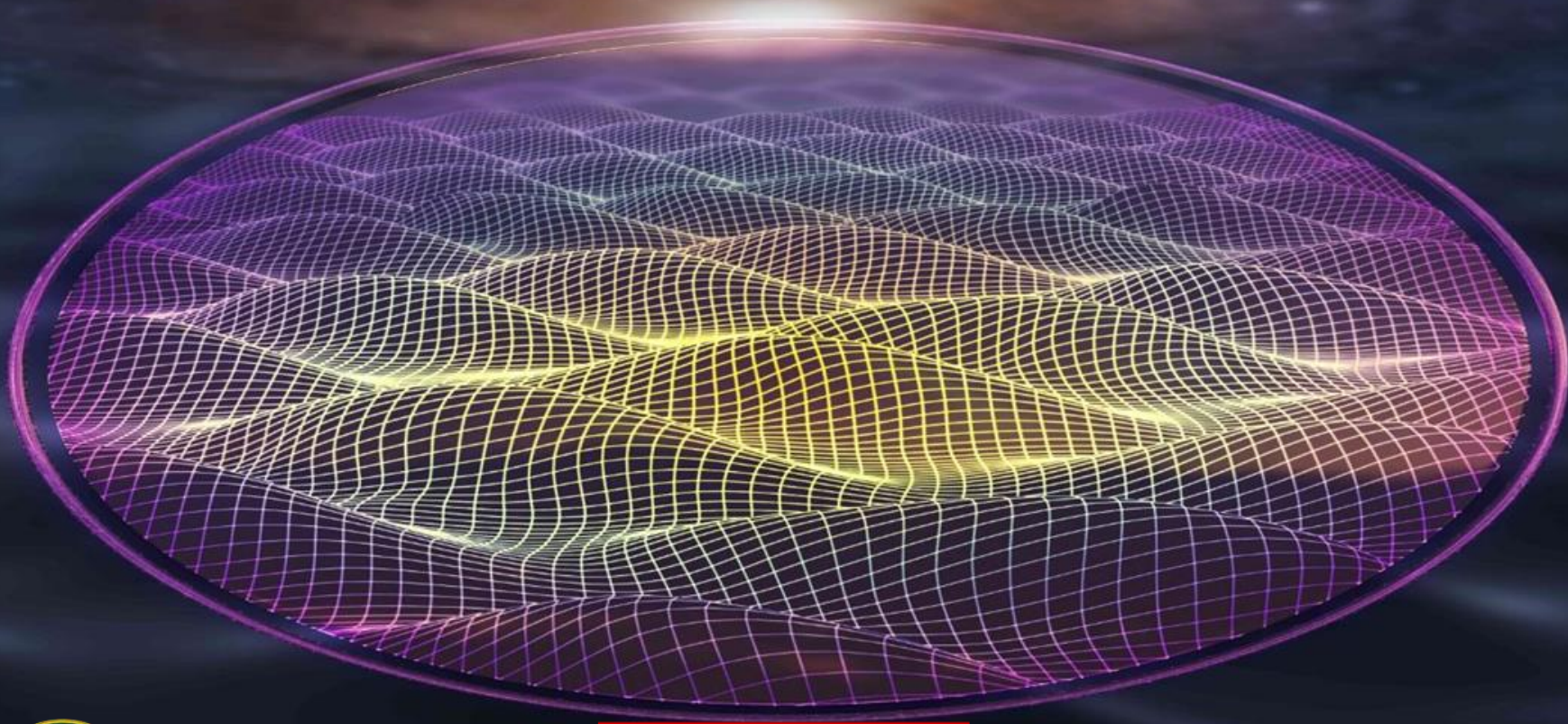


# Searching for axion-like dark matter with precision NMR: the Cosmic Axion Spin Precession Experiments

Alex Sushkov (CASPER collaboration)



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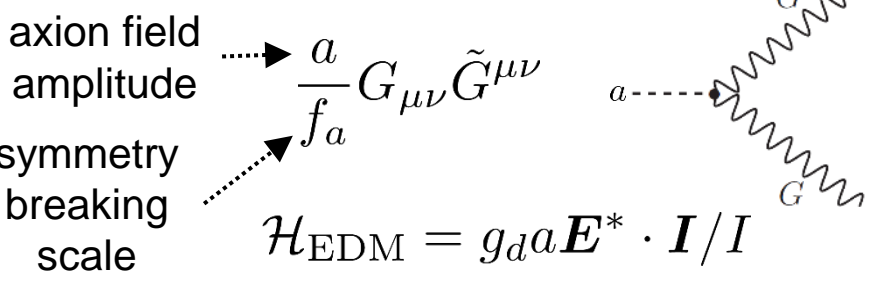
Alfred P. Sloan  
FOUNDATION



# Searching for interactions of axions and axion-like particles

1. Proposed to solve the **strong CP problem** of Quantum Chromodynamics [Phys. Rev. Lett. **38**, 1440 (1977)]
2. Well-motivated and thoroughly-studied **dark matter** candidate:  $a(t) \approx a_0 \cos \omega_a t$
3. 3 possible (non-gravitational) interactions with standard model particles:

## defining QCD axion interaction with gluons: (solves strong-CP problem)

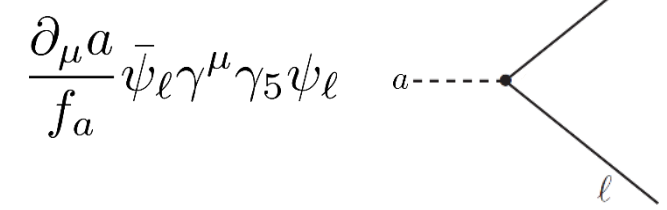


$$\mathcal{H}_{\text{EDM}} = g_d a \mathbf{E}^* \cdot \mathbf{I} / I$$

→ nuclear spin  $\mathbf{I}$  interacts with an oscillating electric dipole moment (EDM)  $d_n = g_d a$  in presence of effective electric field  $\mathbf{E}^*$ .

### CASPEr-electric

## interaction with fermions:

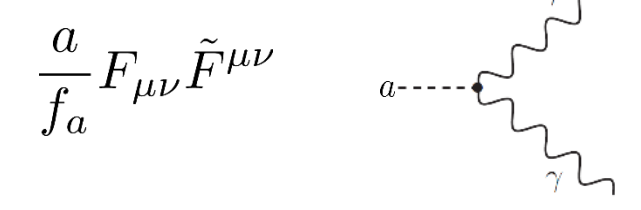


$$\mathcal{H}_{aNN} = g_{aNN} \nabla a \cdot \mathbf{I}$$

→ nuclear spin  $\mathbf{I}$  interacts with an effective magnetic field  $\nabla a$ .  
co-magnetometers  
force mediator → ARIADNE  
electron spin → QUAX

### CASPEr-gradient

## interaction with photons:



$$\mathcal{L}_{a\gamma\gamma} = g_{a\gamma\gamma} a \mathbf{E} \cdot \mathbf{B}$$

→ ALP ↔ electromagnetic field mixing  
→ precision electromagnetic sensors

ADMX, HAYSTAC, DMradio, SHAFT, ABRA, ALPS, CAST, IAXO, CAPP, ORGAN, BREAD, SLIC, LC circuit, MADMAX, KLASH, BRASS, many others

**SHAFT** → a kHz-MHz search using SQUIDs and ferromagnetic toroidal cores

CASPEr (Cosmic Axion Spin Precession Experiments) search for experimental signatures of these interactions using precision magnetic resonance

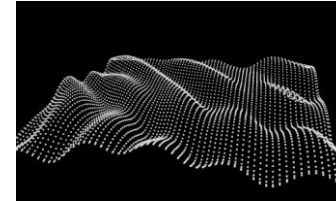
[D. Aybas et al., Phys. Rev. Lett. **126**, 160505 (2021)]  
[D. Aybas et al., Quant. Sci. Tech. **6**, 034007 (2021)]

[D. Budker et al., Phys. Rev. X **4**, 021030 (2014)]  
[A. Garcon et al., Sci. Adv. **5**, eaax4539 (2019)]

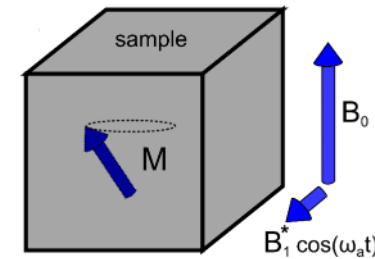
[A. Gramolin et al., Nature Physics **17**, 79 (2021)]



# Searching for interactions of axions and axion-like particles



axion-like field



spin ensemble acts as the transducer



electromagnetic sensor measures spin ensemble evolution

flexibility to optimize spin ensemble parameters to optimize transducer efficiency → maximize sensitivity

defining QCD axion interaction with gluons:  
(solves strong-CP problem)

axion field amplitude  $\rightarrow \frac{a}{f_a} G_{\mu\nu} \tilde{G}^{\mu\nu}$

symmetry breaking scale  $\rightarrow \mathcal{H}_{\text{EDM}} = g_d a \mathbf{E}^* \cdot \mathbf{I}/I$

→ nuclear spin  $\mathbf{I}$  interacts with an oscillating electric dipole moment (EDM)  $d_n = g_d a$  in presence of effective electric field  $\mathbf{E}^*$ .

**CASPEr-electric**

interaction with fermions:

$$\frac{\partial_\mu a}{f_a} \bar{\psi}_\ell \gamma^\mu \gamma_5 \psi_\ell$$

$$\mathcal{H}_{aNN} = g_{aNN} \nabla a \cdot \mathbf{I}$$

→ nuclear spin  $\mathbf{I}$  interacts with an effective magnetic field  $\nabla a$ .  
co-magnetometers  
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**CASPEr-gradient**

CASPEr (Cosmic Axion Spin Precession Experiments) search for experimental signatures of these interactions using precision magnetic resonance

[D. Aybas et al., *Phys. Rev. Lett.* **126**, 160505 (2021)]

[D. Aybas et al., *Quant. Sci. Tech.* **6**, 034007 (2021)]

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# Searching for axionic coupling to spin with magnetic resonance

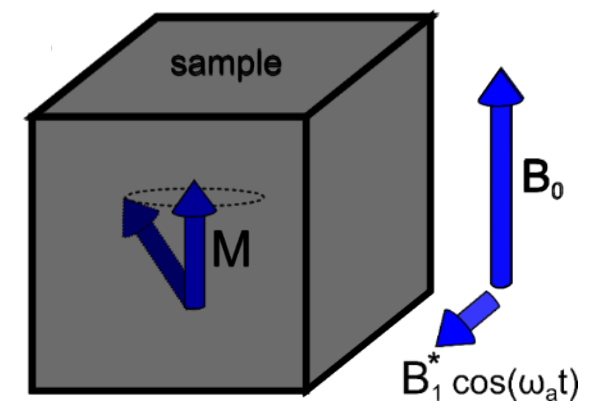
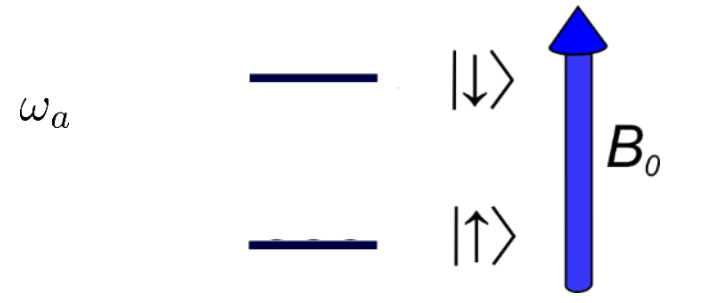
effective interaction:  $\mathcal{H}_{\text{CASPEr}} = -(\hbar\gamma_I \mathbf{B}_1^* \cos \omega_a t) \cdot \mathbf{I}$

$$\mathcal{H} = -\hbar\gamma_I \mathbf{B}_0 \cdot \mathbf{I} - (\hbar\gamma_I \mathbf{B}_1^* \cos \omega_a t) \cdot \mathbf{I}$$

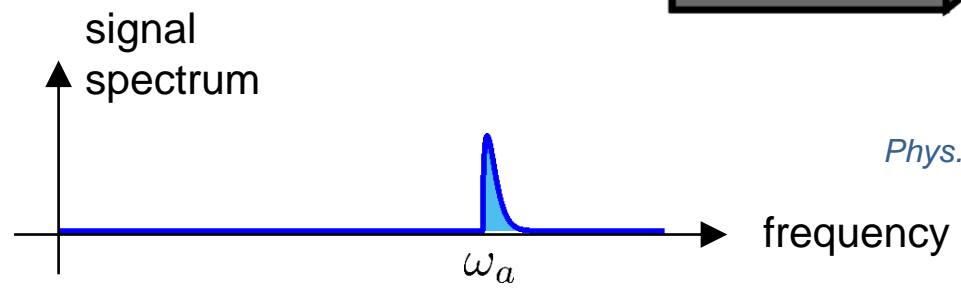


- 1) placing a spin-1/2 into an external magnetic field splits the spin states by  $\gamma_I B_0$
- 2) spin polarization (thermal or optical) in a  $\text{cm}^3$  sample
- 3) resonance:  $\omega_a = \gamma_I B_0$ 
  - ➔ axion-spin interaction can now flip spins!
  - ➔ sample magnetization tilts and precesses
- 4) a magnetometer next to the sample detects the magnetic field created by this precessing magnetization
- 5) search for unknown frequency  $\omega_a$  by sweeping bias magnetic field  $B_0$ , look for resonance

- constant bias magnetic field  $\mathbf{B}_0$
- spin-axion interaction plays the role of the RF field  $\mathbf{B}_1$



an NMR experiment with no RF magnetic field, instead axion-like dark matter flips spins



[D. Budker et al., *Phys. Rev. X* 4, 021030 (2014)]



# Searching for axionic coupling to spin with magnetic resonance

effective interaction:  $\mathcal{H}_{\text{CASPEr}} = -(\hbar\gamma_I \mathbf{B}_1^* \cos \omega_a t) \cdot \mathbf{I}$

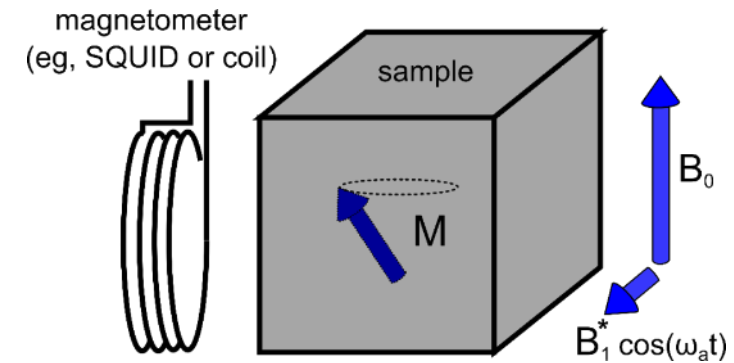
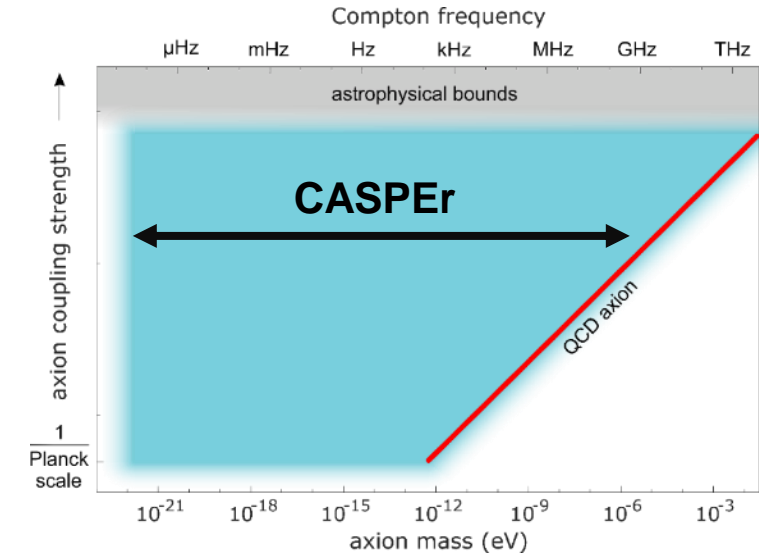
$$\mathcal{H} = -\hbar\gamma_I \mathbf{B}_0 \cdot \mathbf{I} - (\hbar\gamma_I \mathbf{B}_1^* \cos \omega_a t) \cdot \mathbf{I}$$



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an NMR experiment with no RF magnetic field, instead axion-like dark matter flips spins



ultimate goal:  
QCD axion



$$B_1^* \approx 10^{-20} \text{ T}$$

$$\gamma_I B_1^* \approx 10^{-13} \text{ rad/s}$$

$$\gamma_I B_1^* T_2 \approx 10^{-15} \text{ rad}$$

in SHAFT  
we achieved  
sensitivity  
 $\approx 10^{-18} \text{ T}$

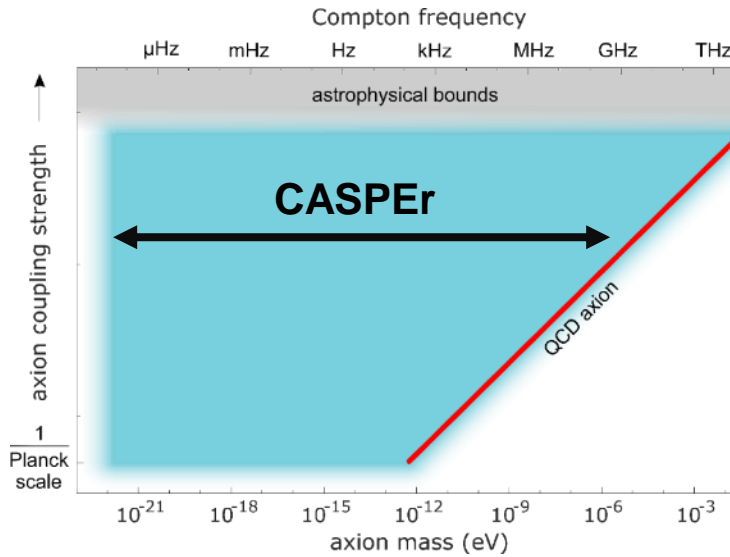


# CASPER program

## Boston University:

CASPER-electric using spins in solids

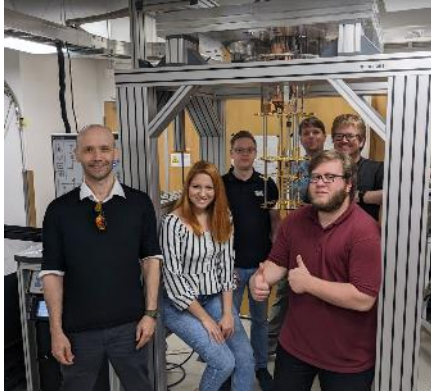
→ sensitive to both  $\mathcal{H}_{EDM} = g_d a \mathbf{E}^* \cdot \mathbf{I} / I$   
 $\mathcal{H}_{aNN} = g_{aNN} \nabla a \cdot \mathbf{I}$



## Mainz:

CASPER-gradient using hyperpolarized liquids

→ sensitive to  $\mathcal{H}_{aNN} = g_{aNN} \nabla a \cdot \mathbf{I}$



Janos Adam  
Stephen Kuenstner  
Glenn Randall  
Andrew Winter  
Tanja Maric

Hendrik Bekker  
Arne Wickenbrock  
Yuzhe Zhang  
John Blanchard  
Deniz Aybas



HEISING - SIMONS  
FOUNDATION

SIMONS  
FOUNDATION

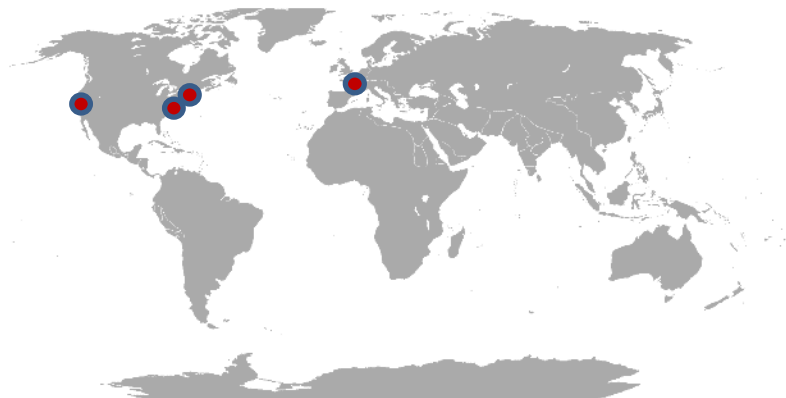
GORDON AND BETTY  
MOORE  
FOUNDATION



Alfred P. Sloan  
FOUNDATION



Deutsche  
Forschungsgemeinschaft  
DFG



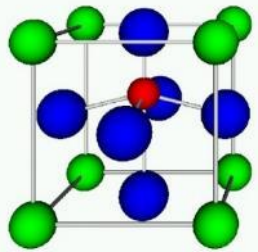
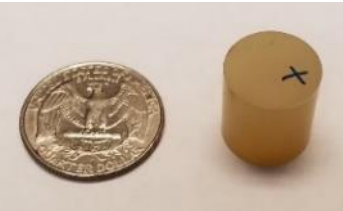
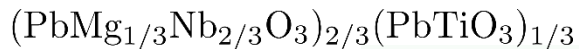
Dmitry Budker, Peter Graham, Derek Kimball, Surjeet Rajendran, Alex Sushkov



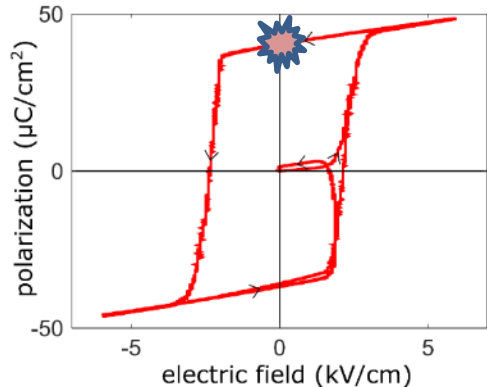
# Millimeter-scale CASPER-e axion-like dark matter search

sample: 4mm

<sup>207</sup>Pb nuclear spins in ferroelectrically-polarized PMN-PT



$3 \times 10^{20}$  spins



$E^* = 340 \text{ kV/cm}$

[Phys. Rev. A 72, 34501 (2005)]

(similar to a polar molecule)

ACME [Science 343, 269 (2013)]

[Nature 562, 355 (2018)]

sensor:

low-noise radiofrequency amplifier

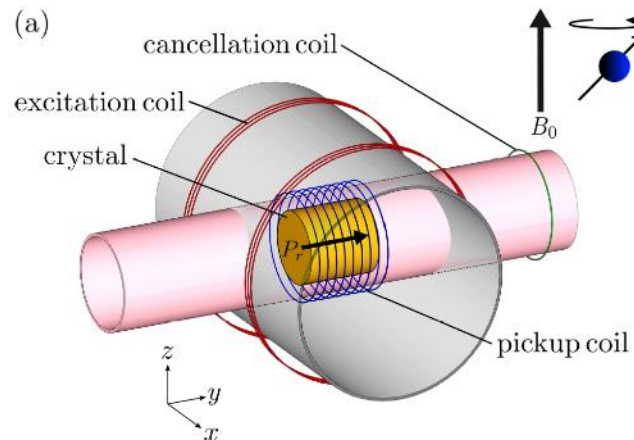


SQUID

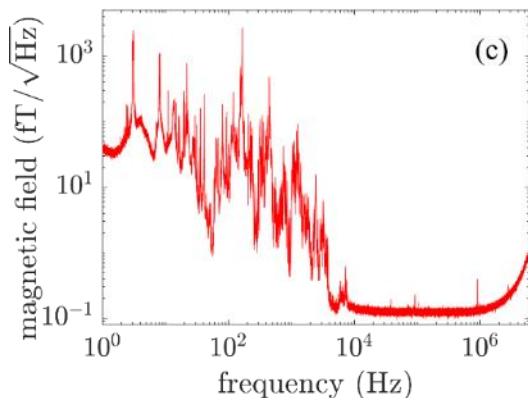


NMR calibration

crossed excitation and pickup coils



liquid helium (4 K) bath cryostat with 9T magnet



millimeter-scale CASPER-e search based on nuclear magnetic resonance

[D. Aybas et al., Phys. Rev. Lett. 126, 160505 (2021)]



# Millimeter-scale CASPER-e axion-like dark matter search

CASPER-e limits on nucleon EDM and gradient interactions of axion-like dark matter

$$\begin{aligned} \Rightarrow \mathcal{H}_{\text{EDM}} &= g_d a \mathbf{E}^* \cdot \mathbf{I} / I \Rightarrow \\ \mathcal{H}_{aNN} &= g_{aNN} \nabla a \cdot \mathbf{I} \end{aligned}$$

→ limits on oscillation amplitudes of neutron EDM and  $\theta_{\text{QCD}}$ :

$$\begin{aligned} |d_n| &< 1.0 \times 10^{-21} \text{ e} \cdot \text{cm} \\ |\theta| &< 4.3 \times 10^{-6} \end{aligned}$$

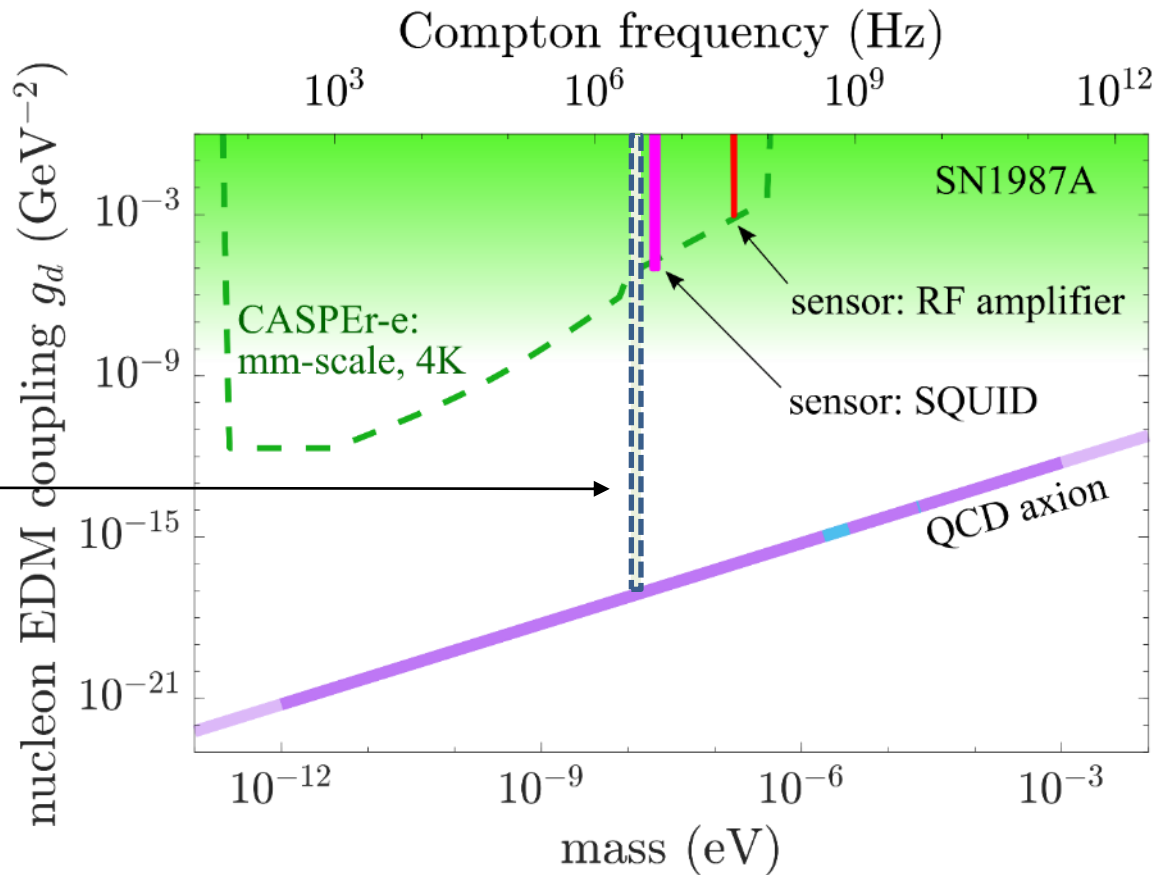
one possible path: scale up experimental volume

our next goal

→ our current approach: 1) optimize material properties  
2) maximize sensitivity with resonant circuit

unique features of the magnetic resonance approach →

- **broad-band**: tune frequency by changing magnetic field (i.e. current) → can cover 4-5 decades in axion mass
- **flexible**: sensitivity can be improved by choice of material (→  $E^*$ ), as well as by scaling up the experimental volume
- **searching for QCD interaction**: this is the defining interaction of the QCD axion



**our next goal:**  
**reach QCD axion sensitivity with a cm-scale search:**  
**1) resonant detection scheme in our dilution fridge**  
**2) nuclear spin hyperpolarization**

[D. Aybas et al., *Phys. Rev. Lett.* **126**, 160505 (2021)]  
 [D. Aybas et al., *Quant. Sci. Tech.* **6**, 034007 (2021)]



# Towards the centimeter-scale CASPER-e QCD axion dark matter search

CASPER-e limits on nucleon EDM and gradient interactions of axion-like dark matter

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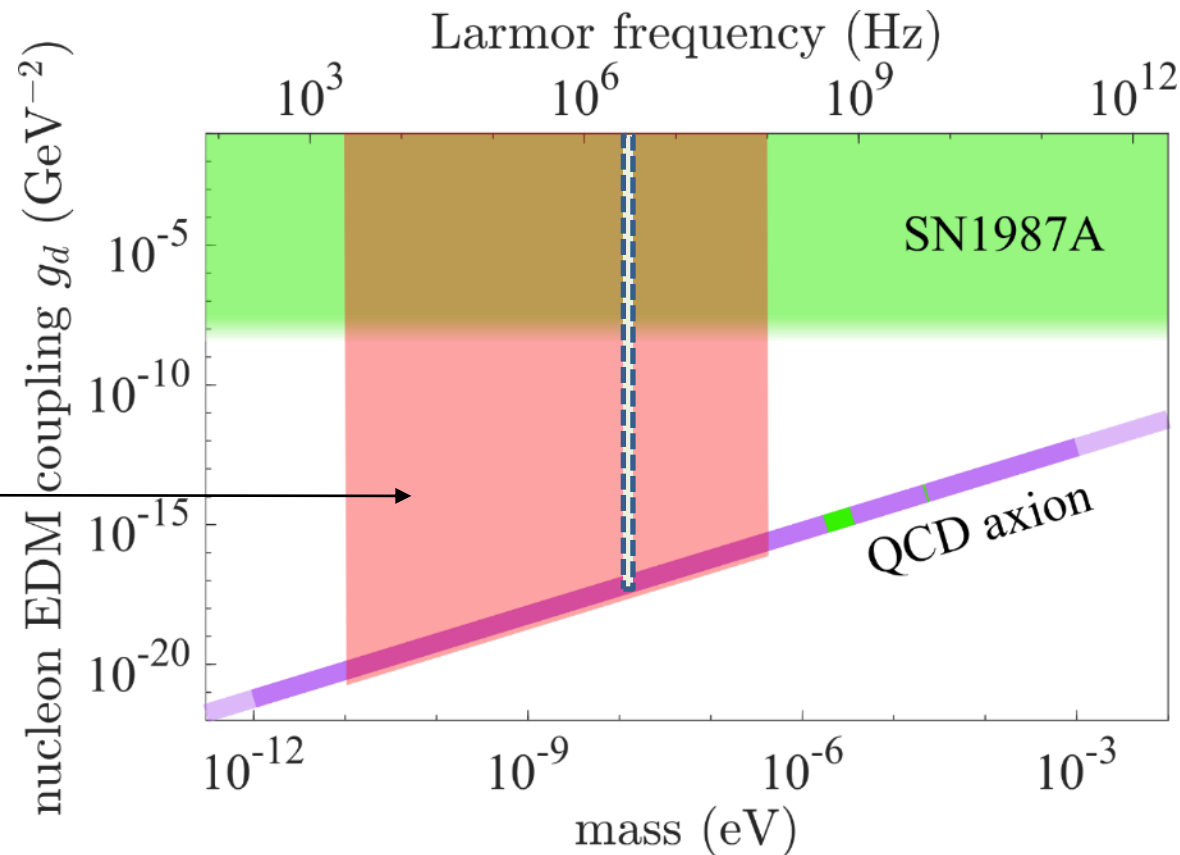
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**our long-term goal**

→ our current approach: 1) optimize material properties  
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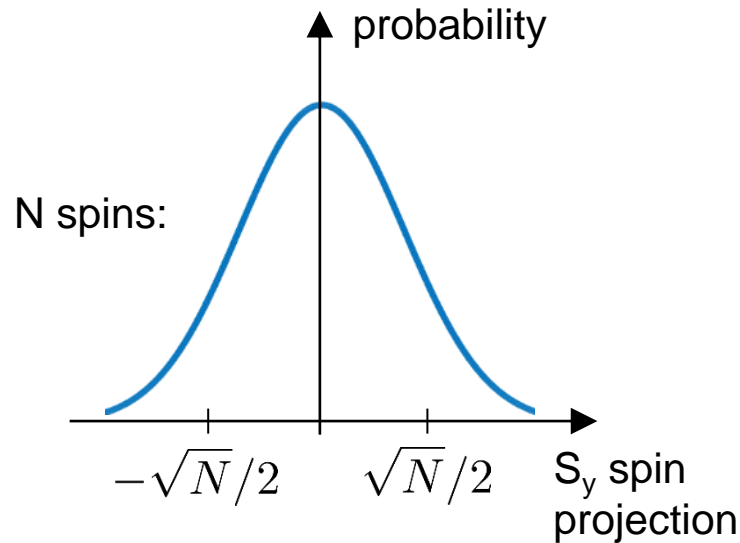
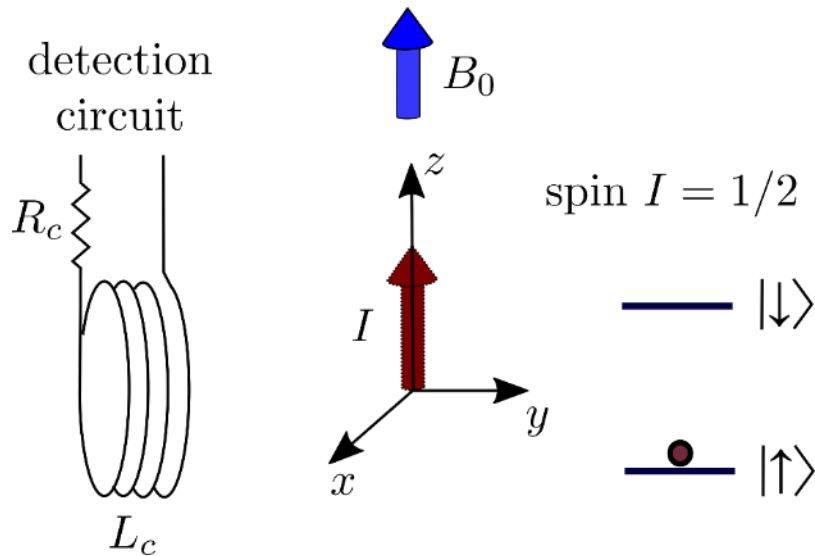
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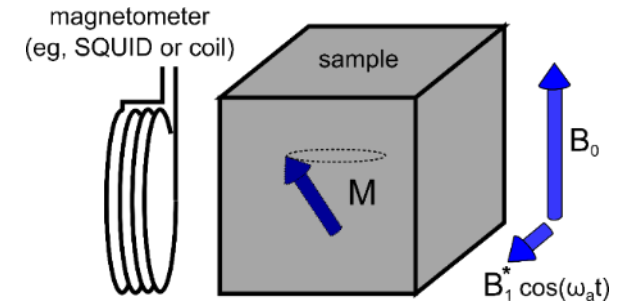
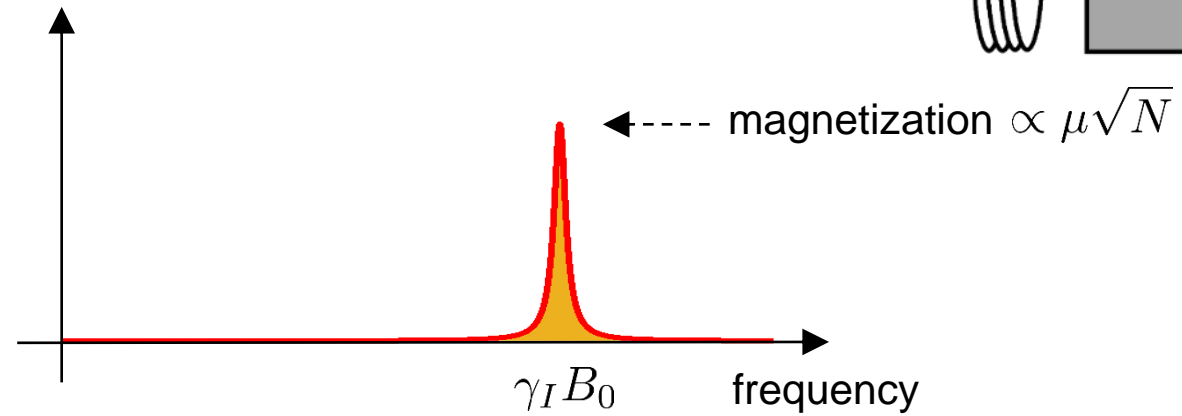
**our long-term goal:**  
**search for pre-inflationary QCD axion dark matter over multiple mass decades**

# Fundamental quantum noise: spin projection noise (standard quantum limit)



[F. Bloch, *Phys. Rev.* **7-8**, 460 (1946)]

→ detected spectrum with a noiseless detection circuit



standard quantum limit (SQL):  $\frac{\delta M_{\perp}}{M_0} \approx \frac{1}{\sqrt{N}}$  ← spin projection measurement uncertainty

spin projection noise has been detected in NMR experiments

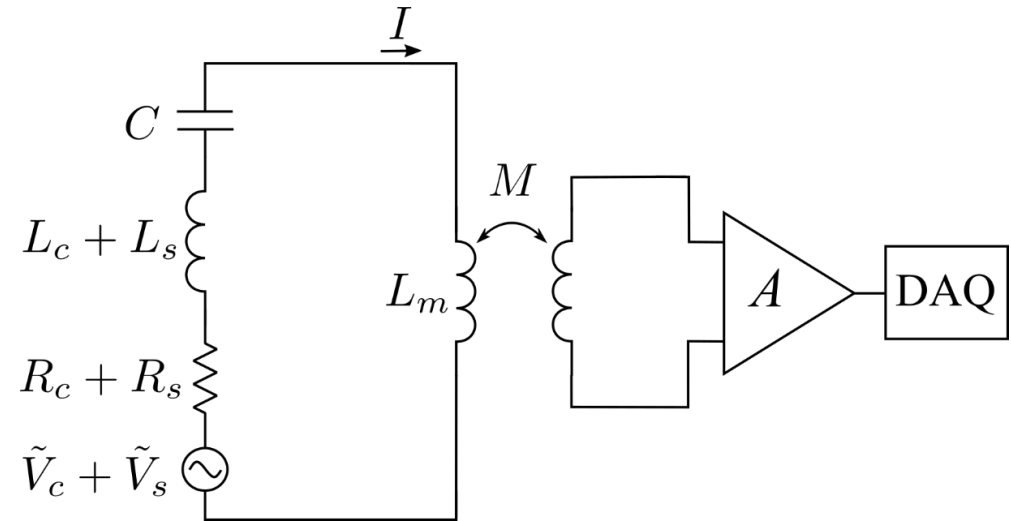
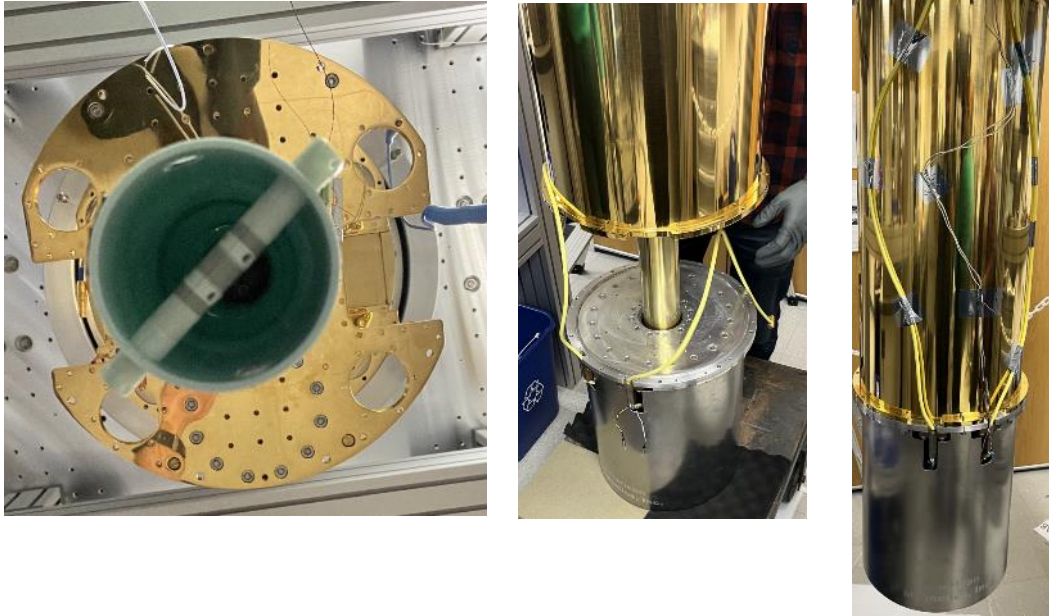
[T. Sleator et al., *Phys. Rev. Lett* **55**, 1742 (1985)]

[M.A. McCoy, R.R. Ernst., *Chem. Phys. Lett.* **159.5**, 587 (1989)]

[M. Gueron, J. L. Leroy., *J. Mag. Res.* **85.1**, 209 (1989)]

we are building CASPEr with sensitivity limited by spin projection noise

# Boston CASPEr-electric search for the EDM coupling



preliminary noise data:

➔ solid-state NMR apparatus in a dilution refrigerator

➔ spin projection noise-limited NMR search for axion-like dark matter

condition for an experiment limited by spin projection noise:

$$\theta_c = 50 \text{ mK}, \quad Q_c > 3 \times 10^4$$

