



CDEX experiment: status and prospect



Li, Hau-Bin on behalf of CDEX collab.
Academia Sinica 中央研究院
(Taiwan)



OUTLINE

- CDEX experiments: overview.
- CDEX-1, CDEX-10 results.
- Plan: CDEX-50 for Dark Matter, CDEX-300v for $0\nu\beta\beta$.
- summary



Dark Matter 2023

The 14th symposium will be held March 29 - April 1, 2023

31 March 2023
at UCLA, Los Angeles

CDEX: China Dark matter EXperiment



Established in 2009

- Tsinghua University (THU)
- Sichuan University (SCU)
- Beijing Normal University (BNU)
- Nankai University (NKU)
- Sun Yat-Sen University (SYSU)
- Peking University (PKU)
- China Institute of Atomic Energy (CIAE)
- Yalong River Company



- Academia Sinica, Taiwan
- Banaras Hindu University, India
- Dokuz Eylül University, Turkey



PCGe data/physics analysis, etc.

- Ge: good resolution, low threshold, ideal for low mass WIMP.

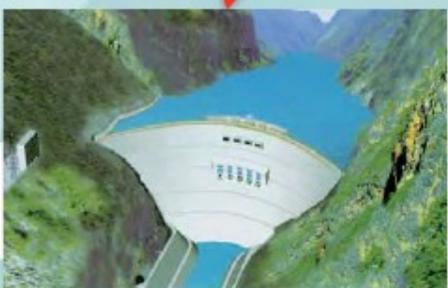
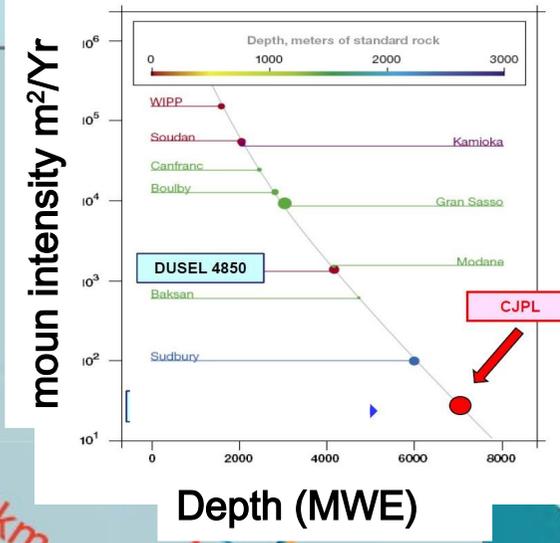
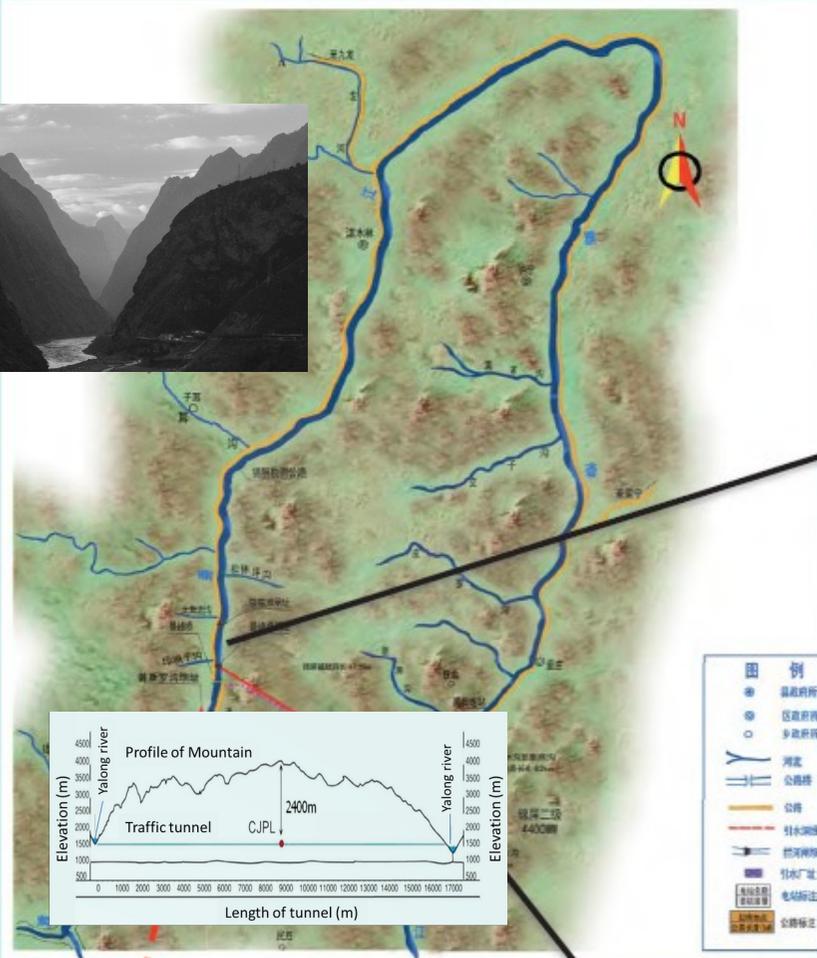


Jinping Hydroelectric Power Plants



4 hydraulic tunnels
 $\Phi 13\text{m} \times 16.6\text{km}$

Jinping-II Power Plant
 4800MW
 (8×600MW)

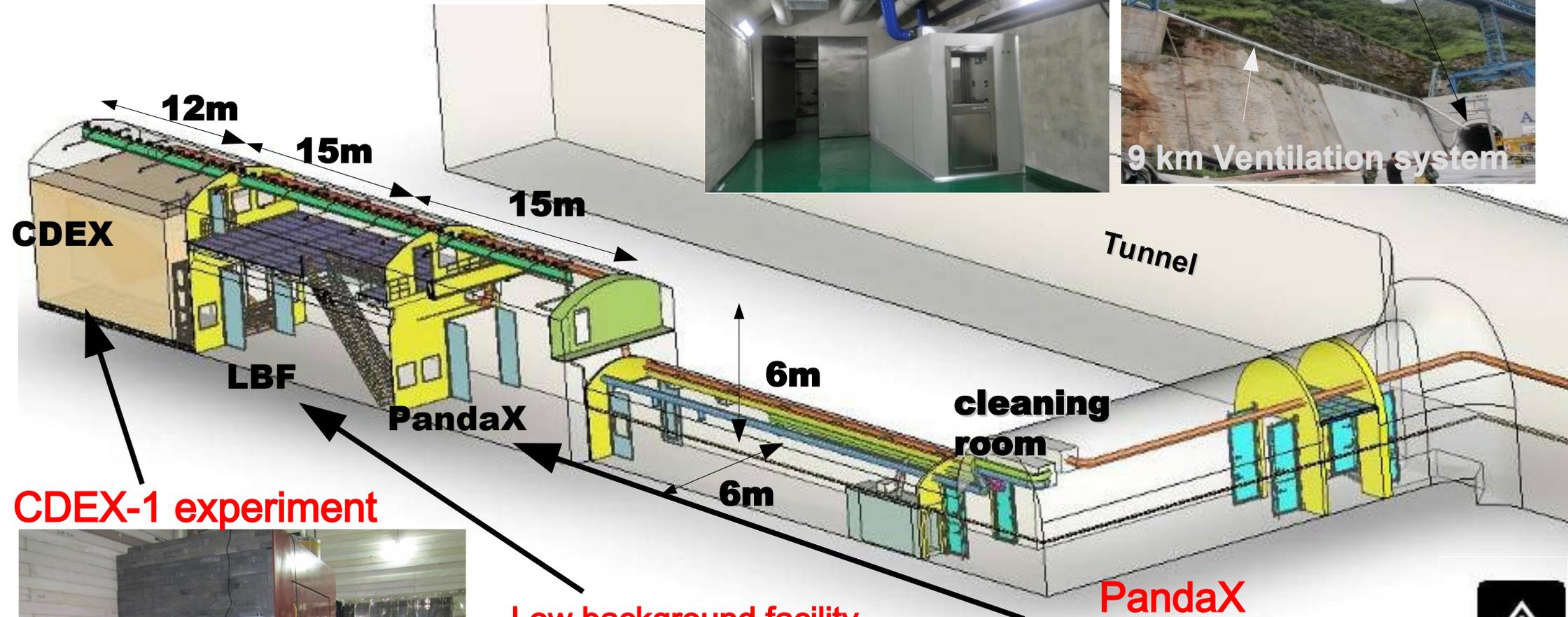
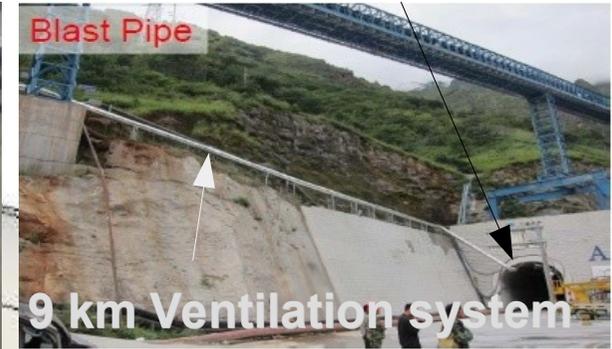


Jinping-I Power Plant
 3600MW
 (6×600MW)



CJPL-I

tunnel entrance





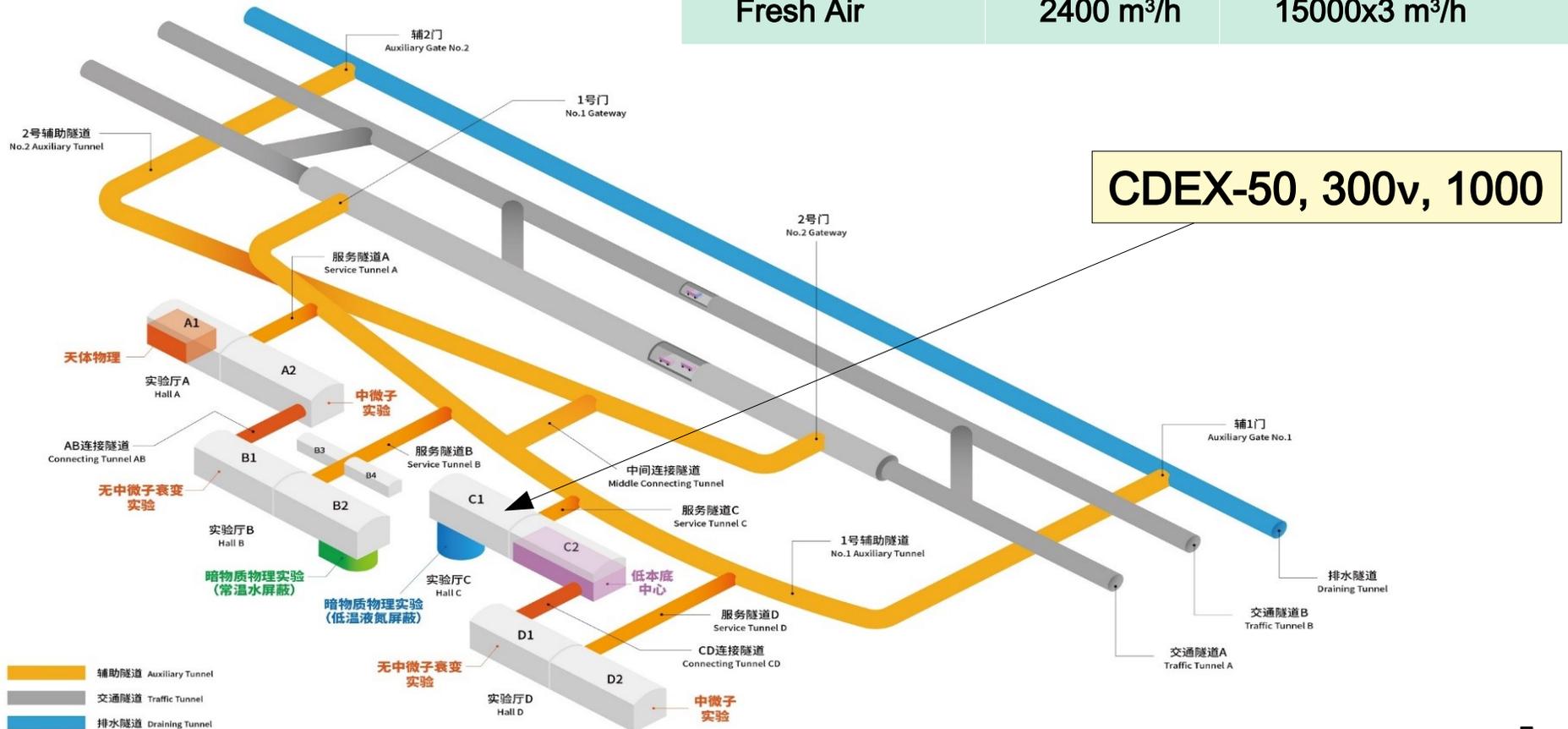
CJPL-II

- 4 main halls : 14m(H)×14m(W)×130m(L);
- Total Volume: 300K m³;
- Two expanded spaces:

C1-- $\phi 18\text{m} \times 32\text{m(H)}$ →CDEX-1T

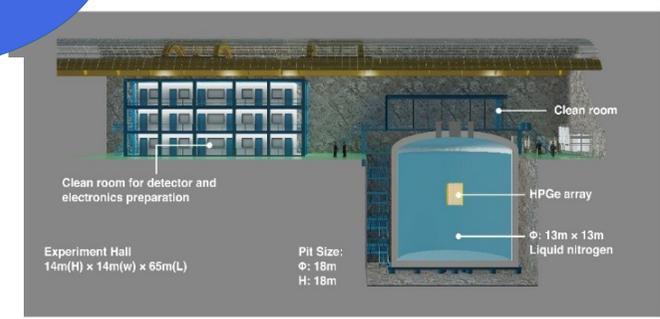
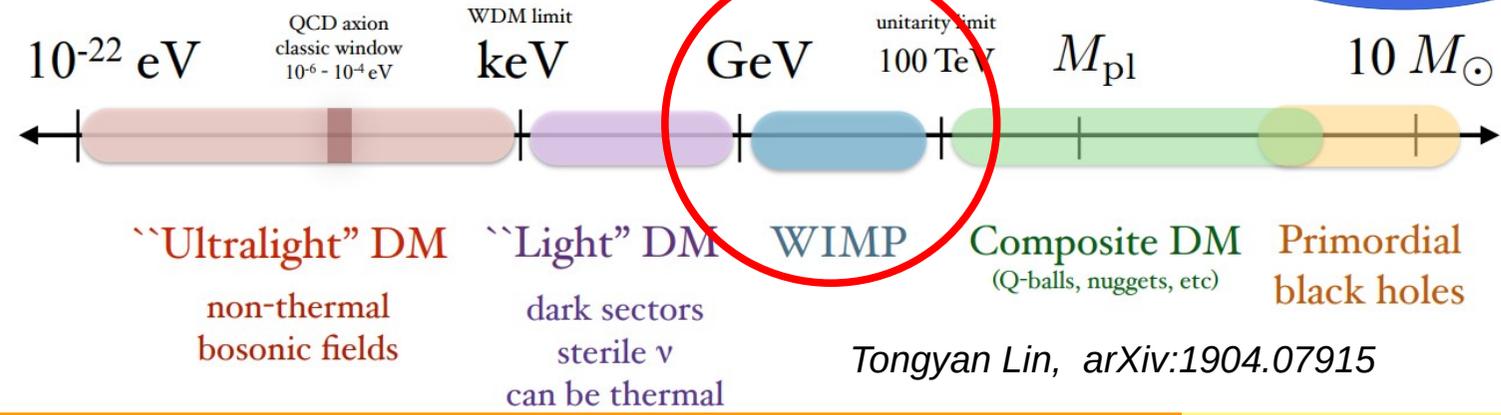
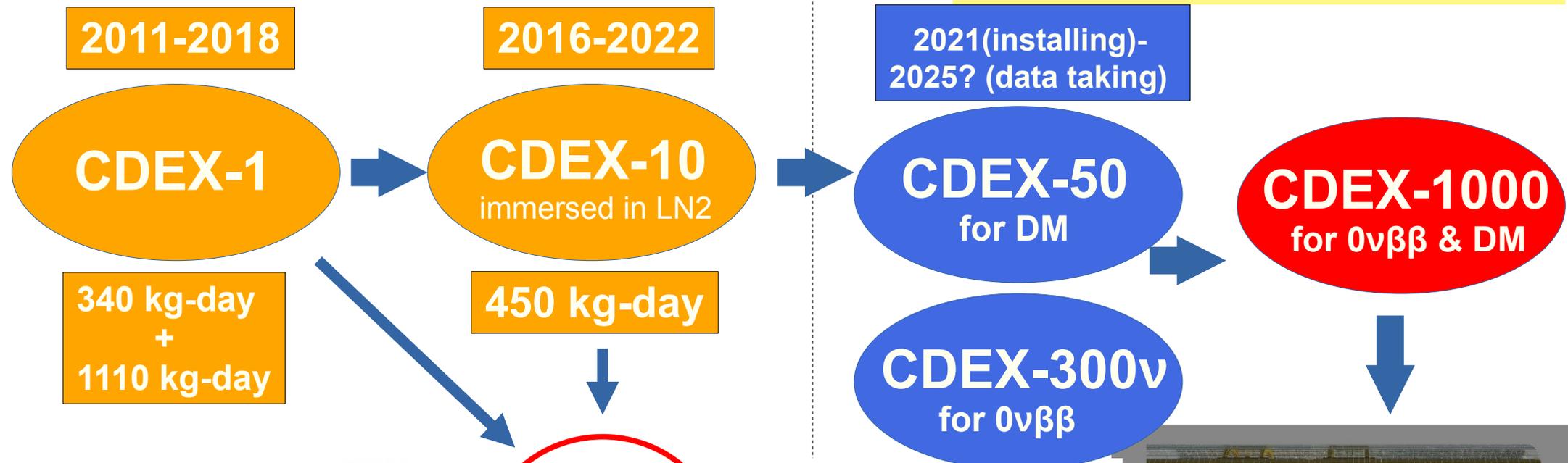
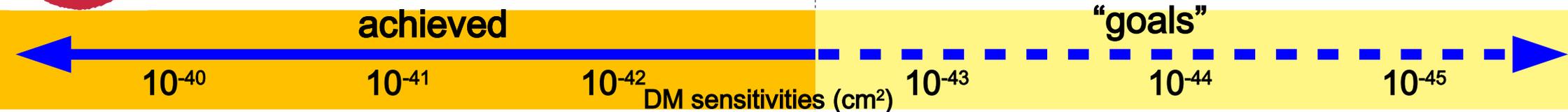
B2-- $27\text{m(L)} \times 14\text{m(W)} \times 30\text{m(H)}$

	CJPL-I	CJPL-II
Rock Work	4100 m ³	210000+151000m ³
Electric Power	70x2 kVA	10x2 MVA
Fresh Air	2400 m ³ /h	15000x3 m ³ /h





CDEX history and future



Tongyan Lin, arXiv:1904.07915



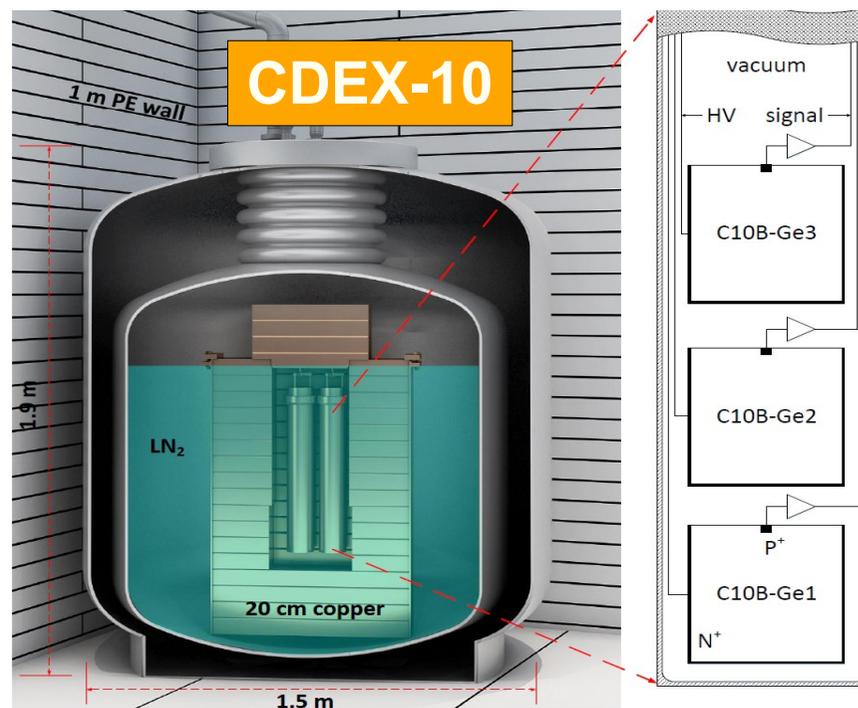
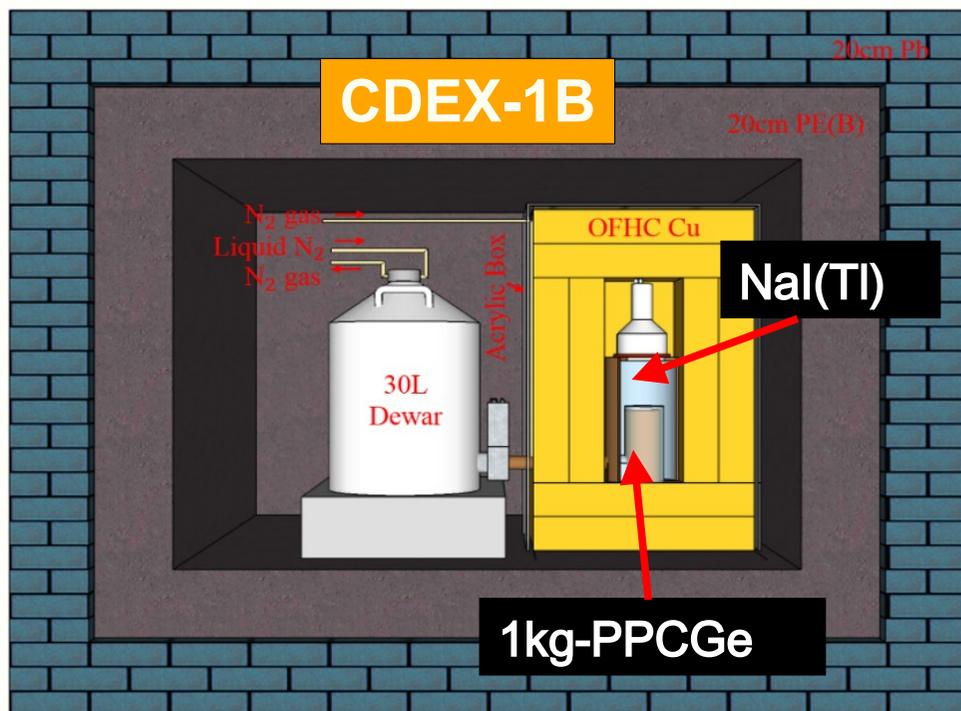
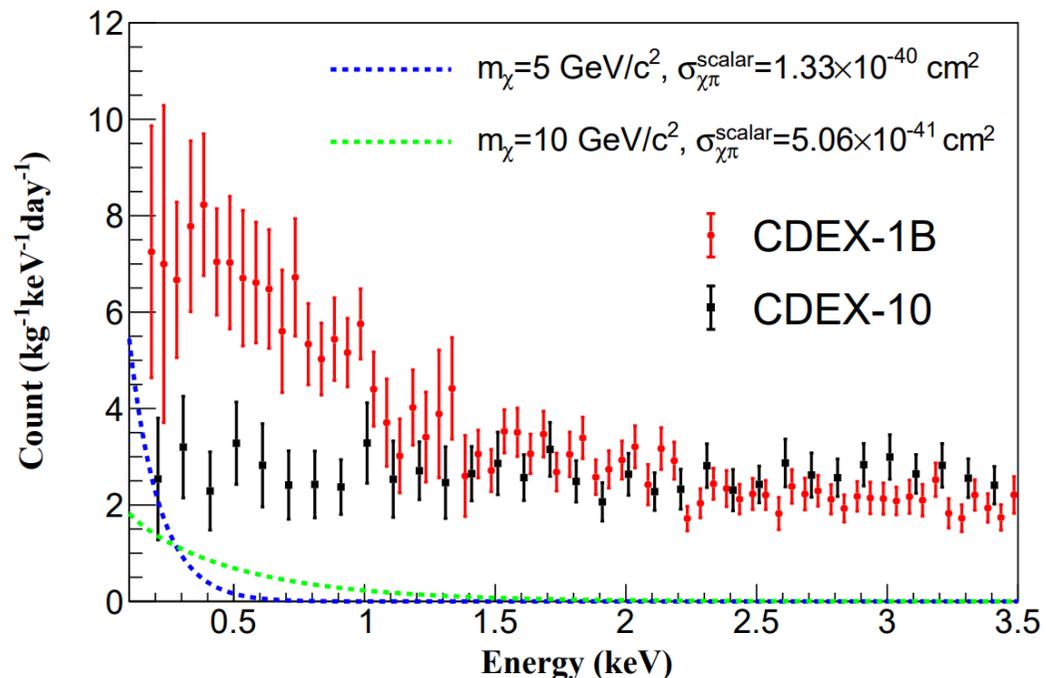
CDEX-1 & 10 configuration

1. CDEX-1B

- ✓ with or w/o NaI(Tl), 1kg point-contact-Ge
- ✓ heat-guide cooling.
- ✓ 700 kg-day, threshold~160-250 eV
- ✓ annual modulation results.

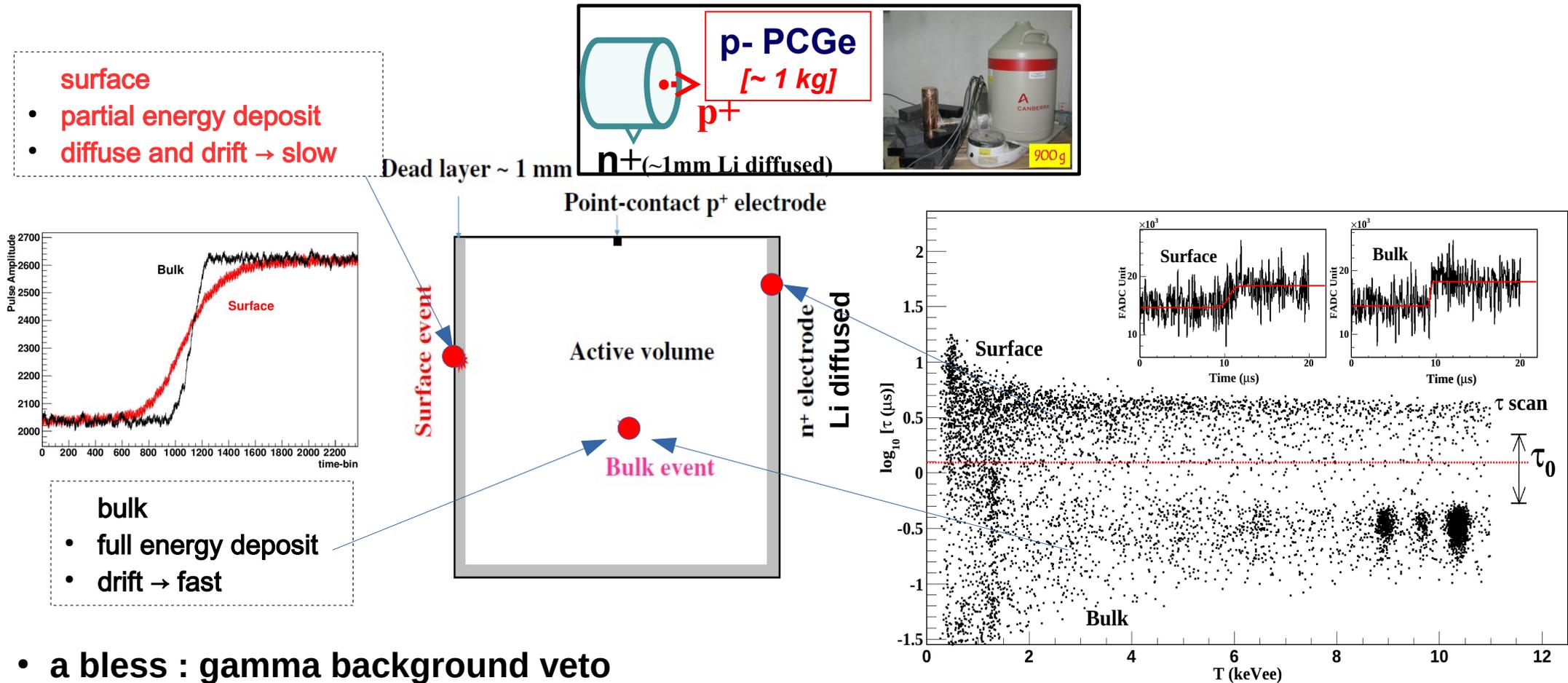
2. CDEX-10

- ✓ immersed in LN2, array format.
- ✓ threshold ~ 160 eV.
- ✓ results based on 200 kg-day on 1 kg Ge.





largest sources of uncertainties: bulk/surface separation

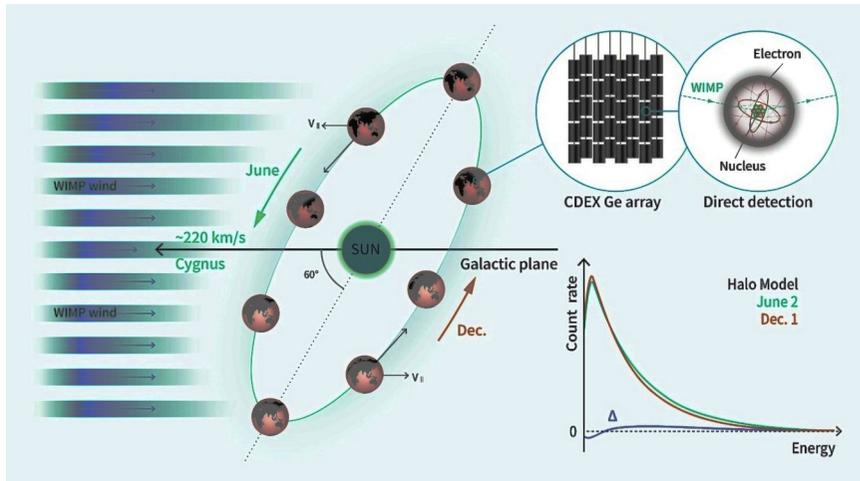


- a bless : gamma background veto and
- a curse : contaminate low energy spectrum

Astropart. Phys. 56, p1-8 (2014) : based on understanding of energy spectrum
NIMA 886, p13-23 (2018) : comparing rise-time of different sources
+ ongoing; Machine learning and improve τ measurement for bulk/surface

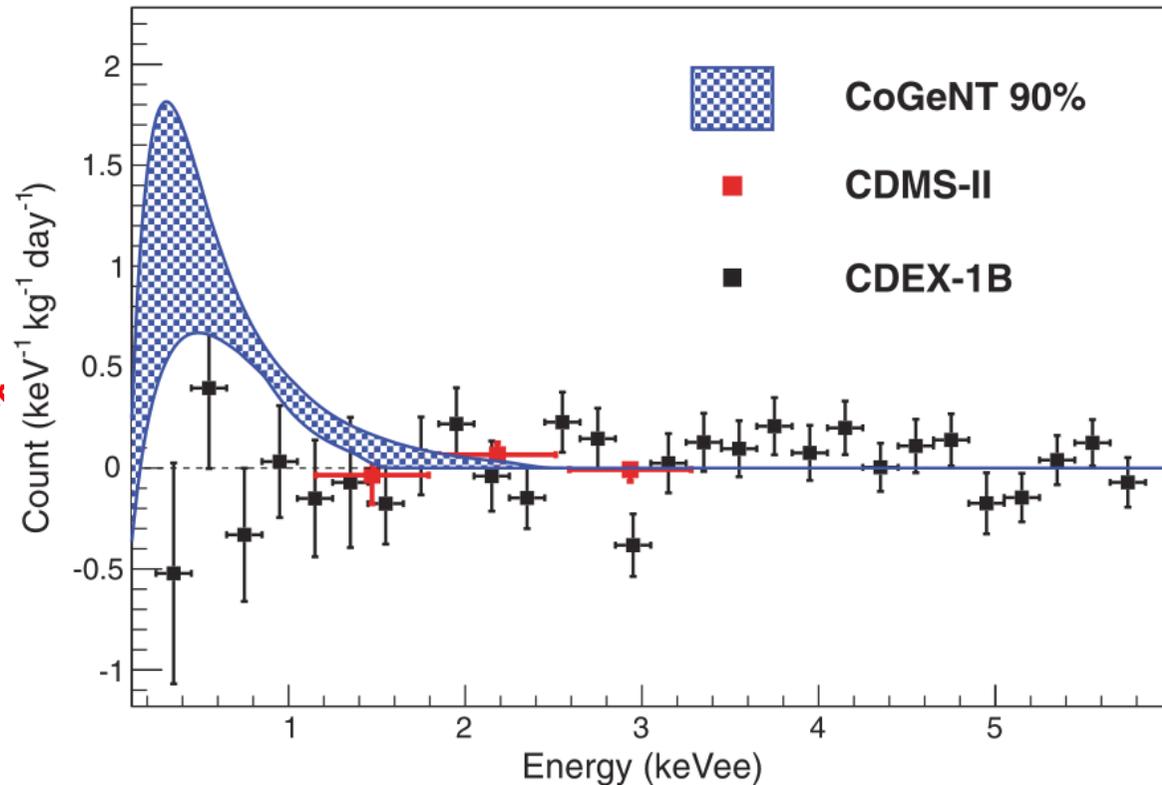
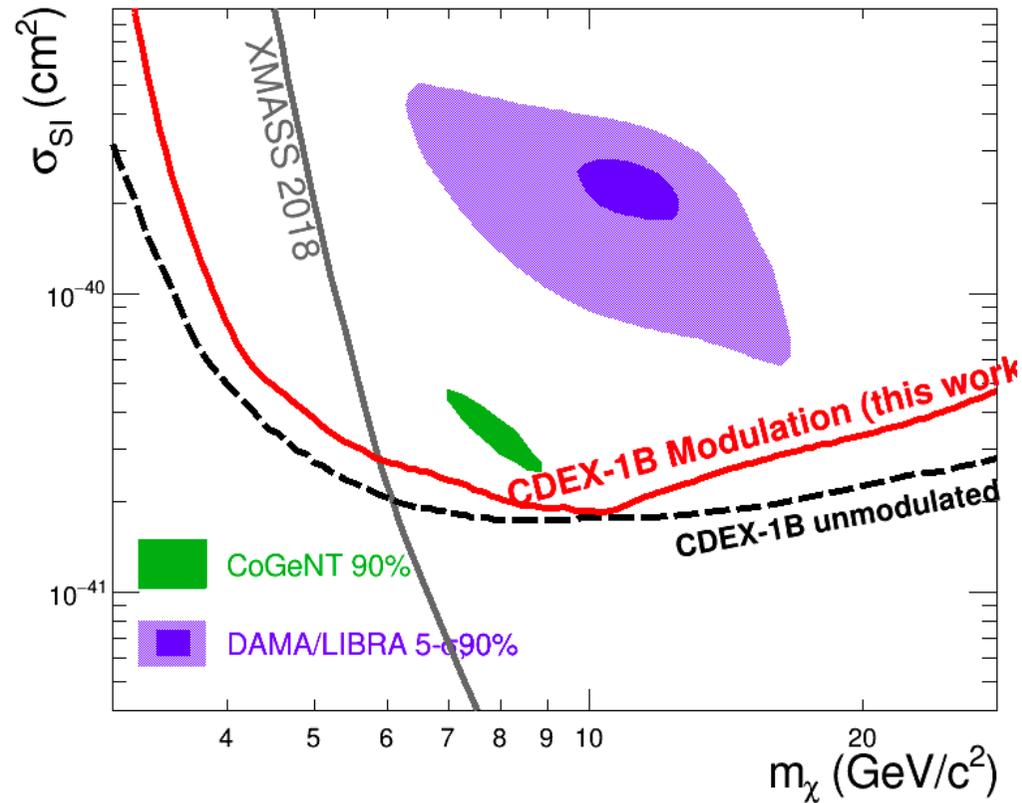


CDEX-1B: annual modulation



$$\chi_{ik}^2 = \sum_{j \in \text{Time}} \frac{(n_{ijk} - P_{ijk} - B_{ik} - A_{ik} \cos(\frac{2\pi(t_j - \phi)}{T_{yr}}))^2}{\Delta_{ijk}^2}$$

- not exclude I-recoil of DAMA/LIBRA

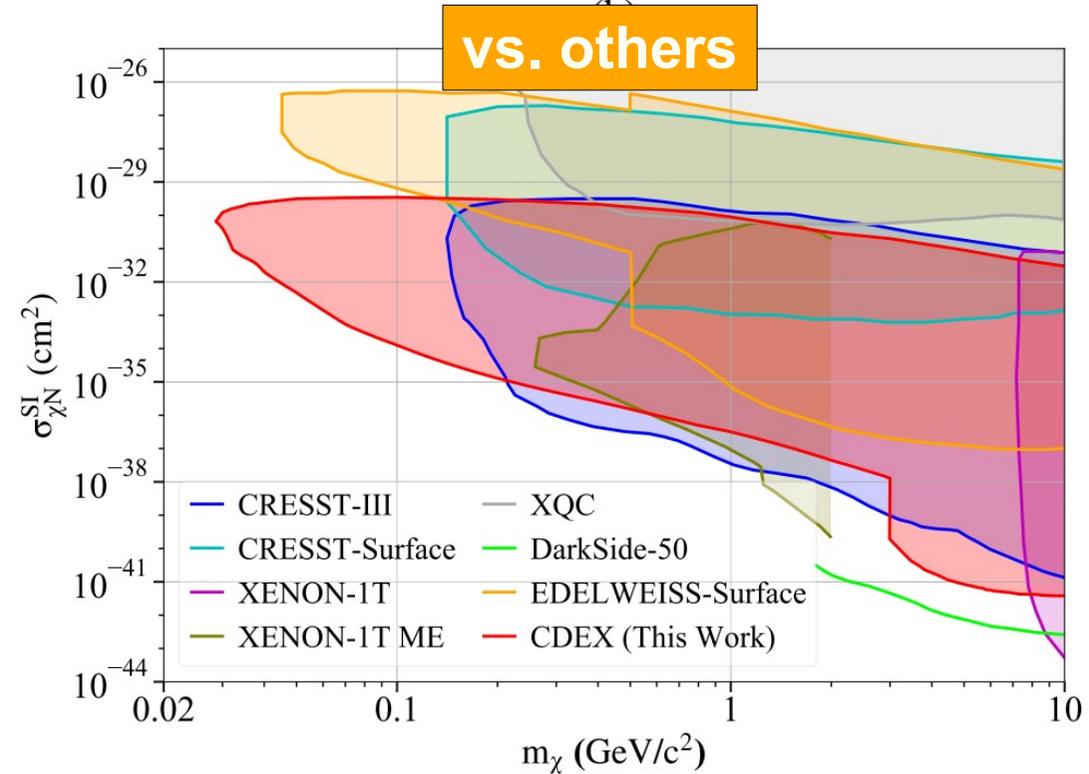
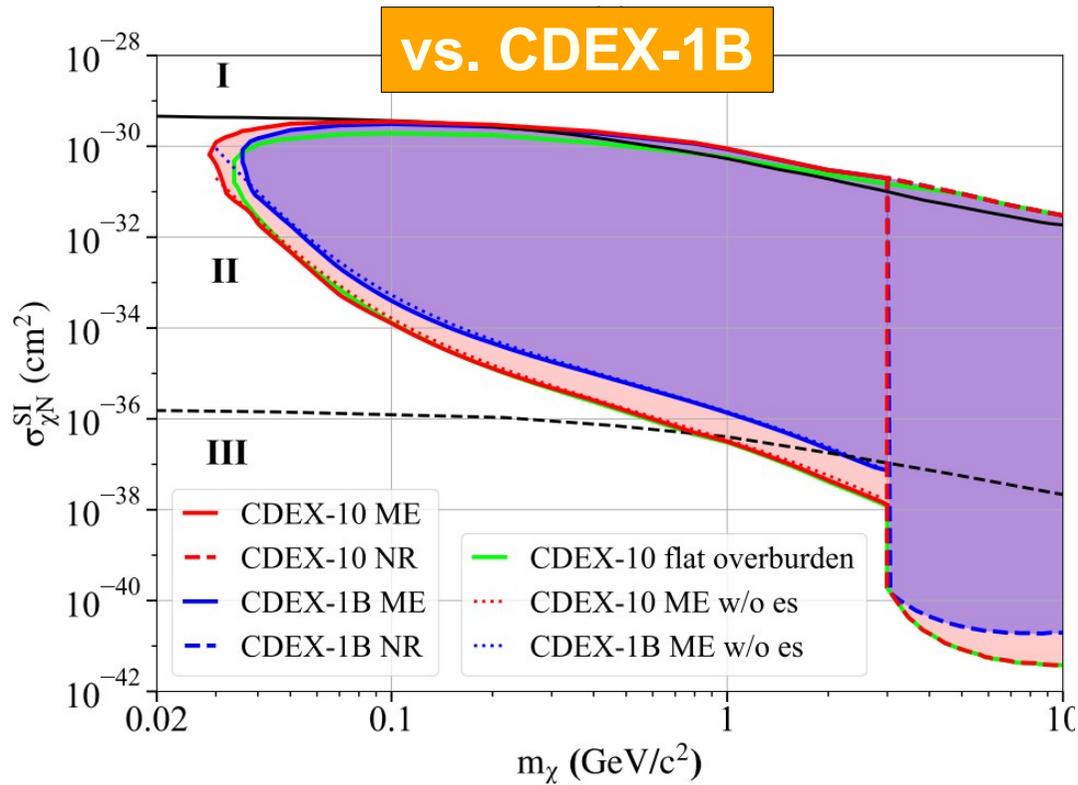
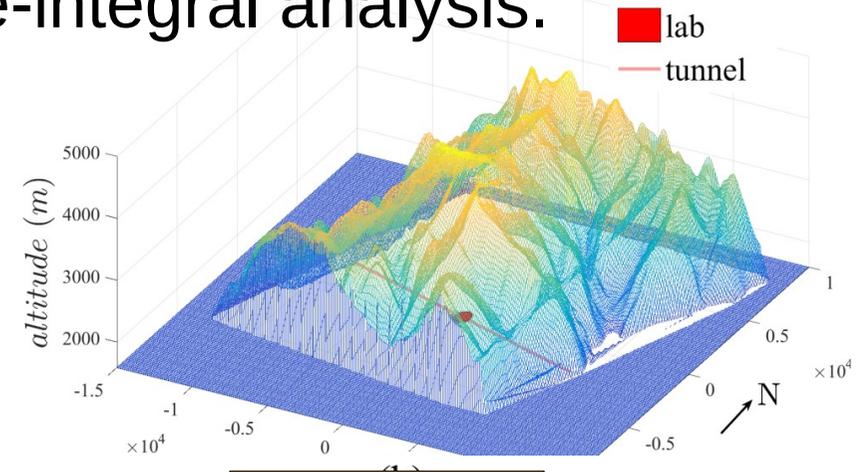
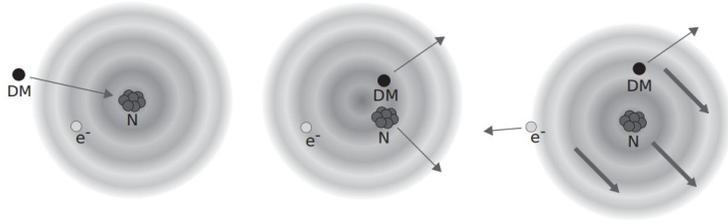


Phys. Rev. Lett. 123, 221301 (2019)



CDEX-10

- + Migdal-effect and earth scattering, time-integral analysis.

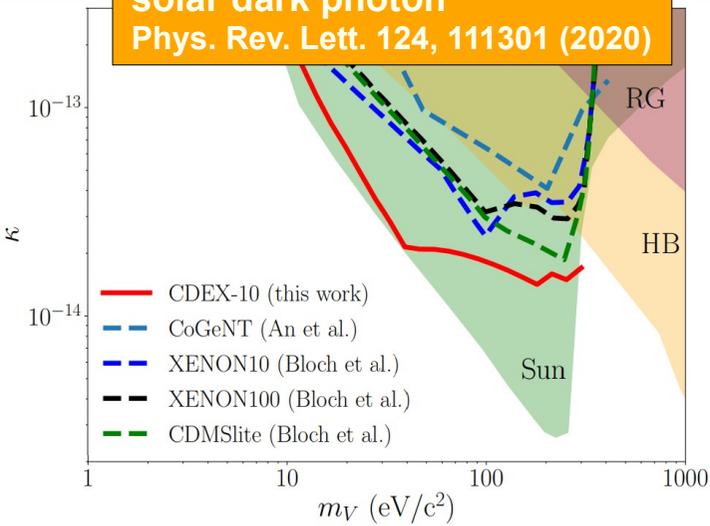


Phys. Rev. D 105, 052005 (2022)

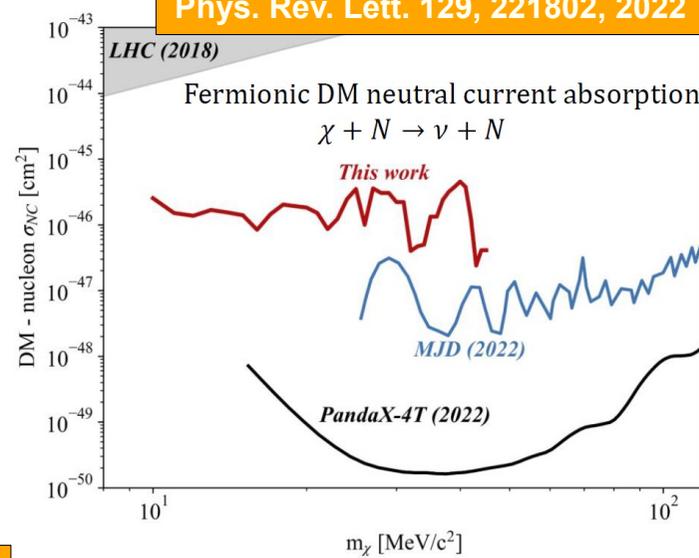


DM related analysis on CDEX-1B and CDEX-10

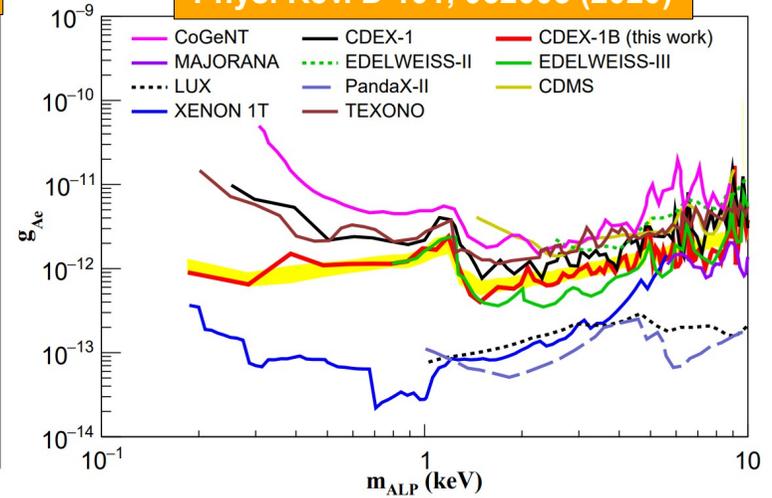
solar dark photon
Phys. Rev. Lett. 124, 111301 (2020)



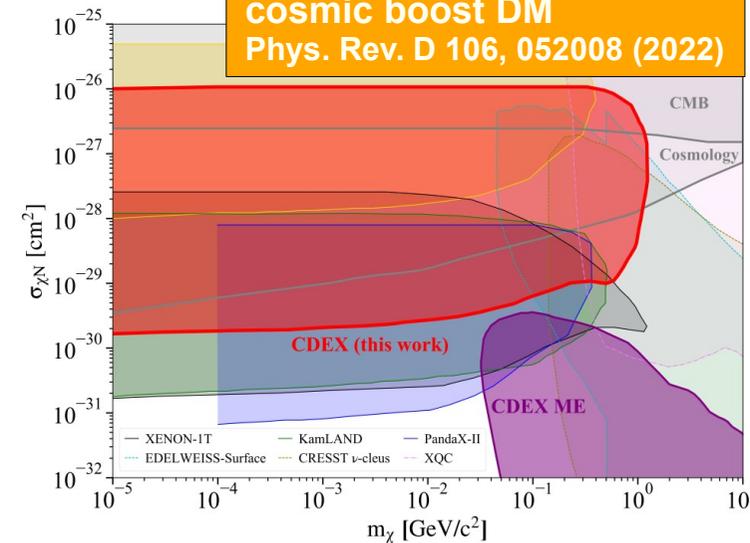
Fermionic DM
Phys. Rev. Lett. 129, 221802, 2022



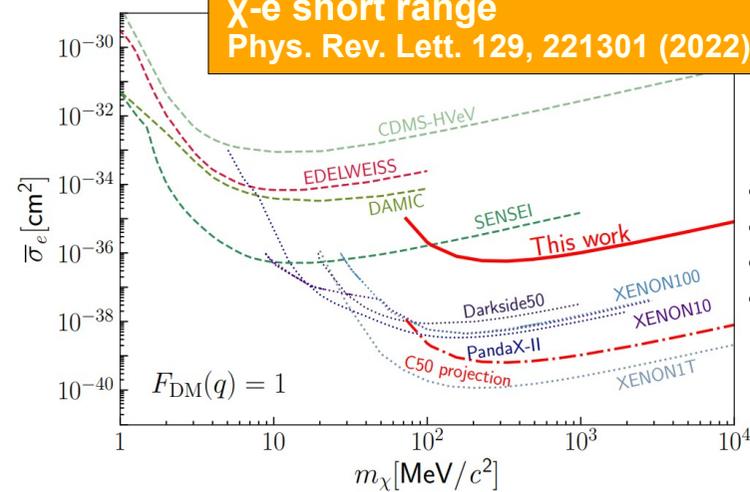
galactical Axion-like particle
Phys. Rev. D 101, 052003 (2020)



cosmic boost DM
Phys. Rev. D 106, 052008 (2022)



chi-e short range
Phys. Rev. Lett. 129, 221301 (2022)



other:

- CDEX-10: Phys. Rev. Lett. 120, 241301 (2018).
- sterile neutrino: arXiv:2210.01604.
- Dark Matter from black hole: arXiv:2211.07477.
- $0\nu\beta\beta$: Phys. Rev. D 106, 032012 (2022)

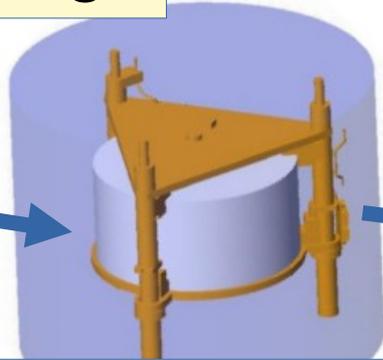


CDEX-50 at CJPL-II

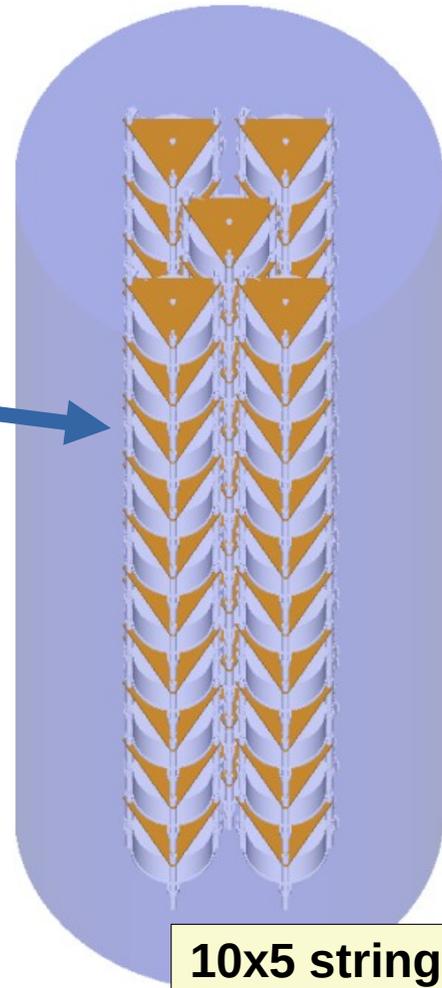
under technical design



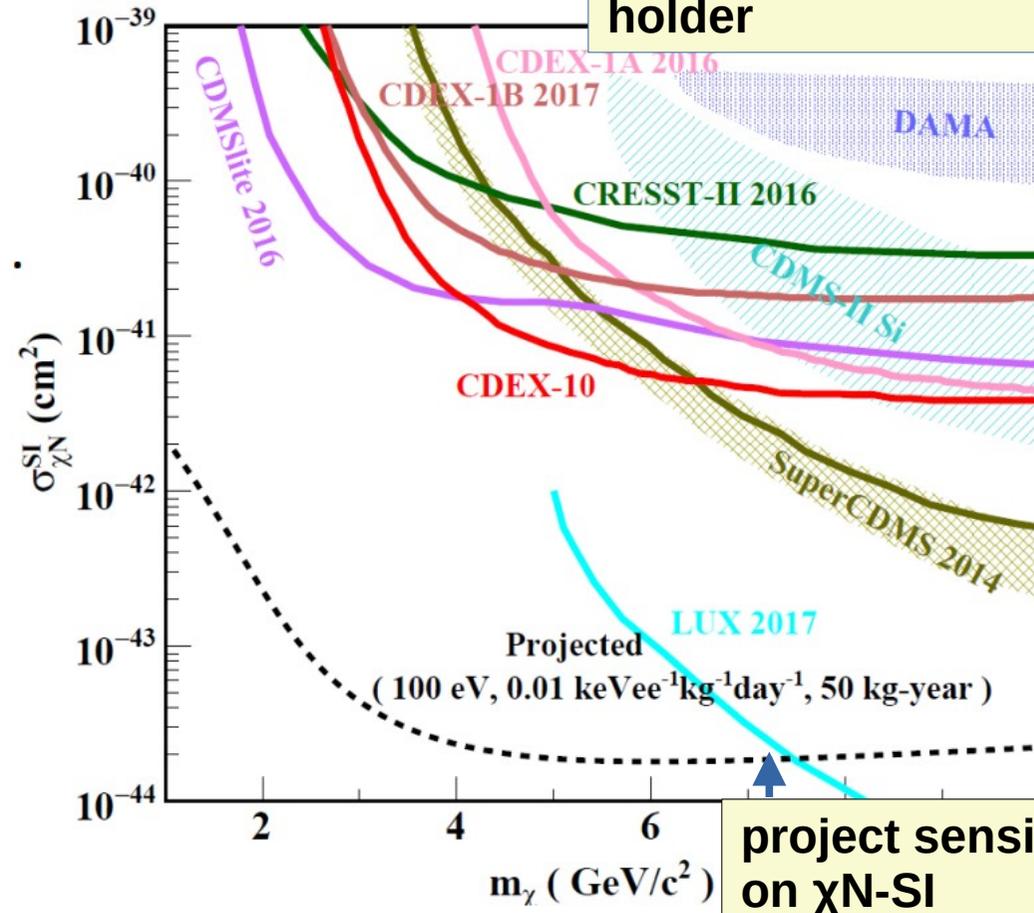
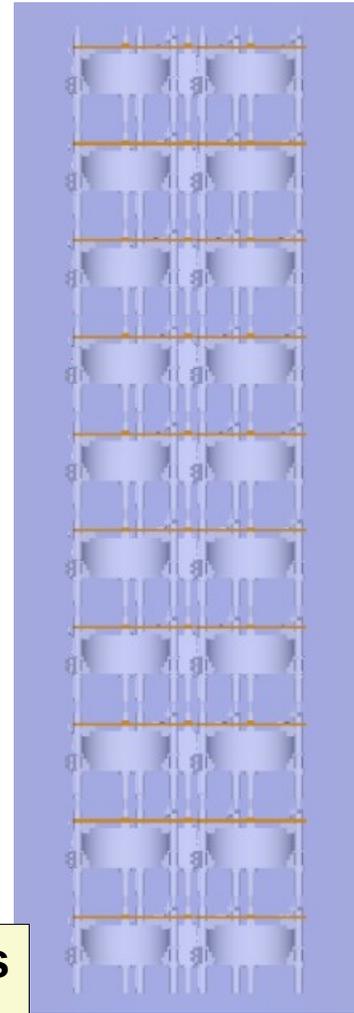
1kg PCGe or BEGe



79 g Cu + 10 g PTFE holder



10x5 strings in LN2



expect:

- 0.01 count $\text{keV}^{-1} \cdot \text{kg}^{-1} \cdot \text{day}^{-1}$ background at 1 keV
- threshold ~ 100 eV
- $\sim 10^{-44} \text{cm}^2$ for spin-independent interaction.



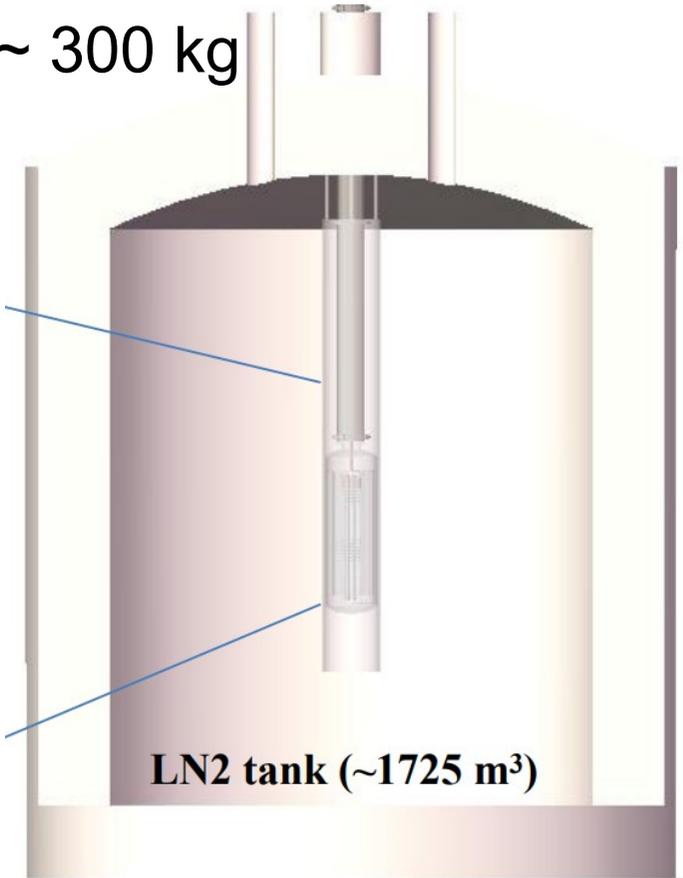
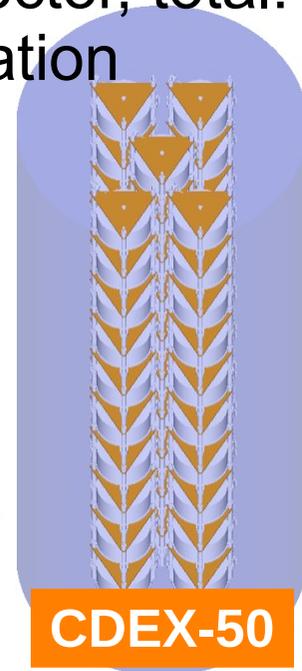
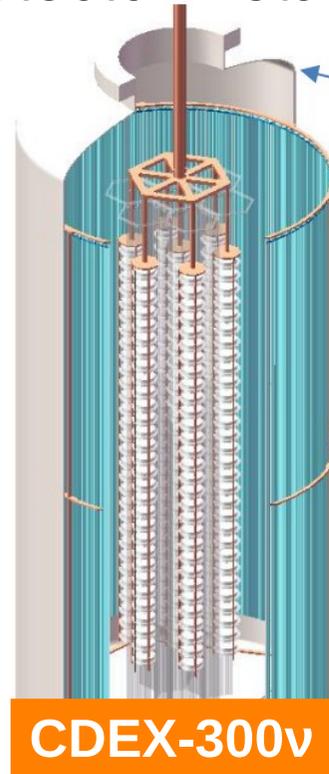
CDEX-50 & CDEX-300v at CJPL-II

CDEX-50 (~2025-)

- 10 strings, with 10×1 kg detectors per string, total:~50kg

CDEX-300v (~2028-)

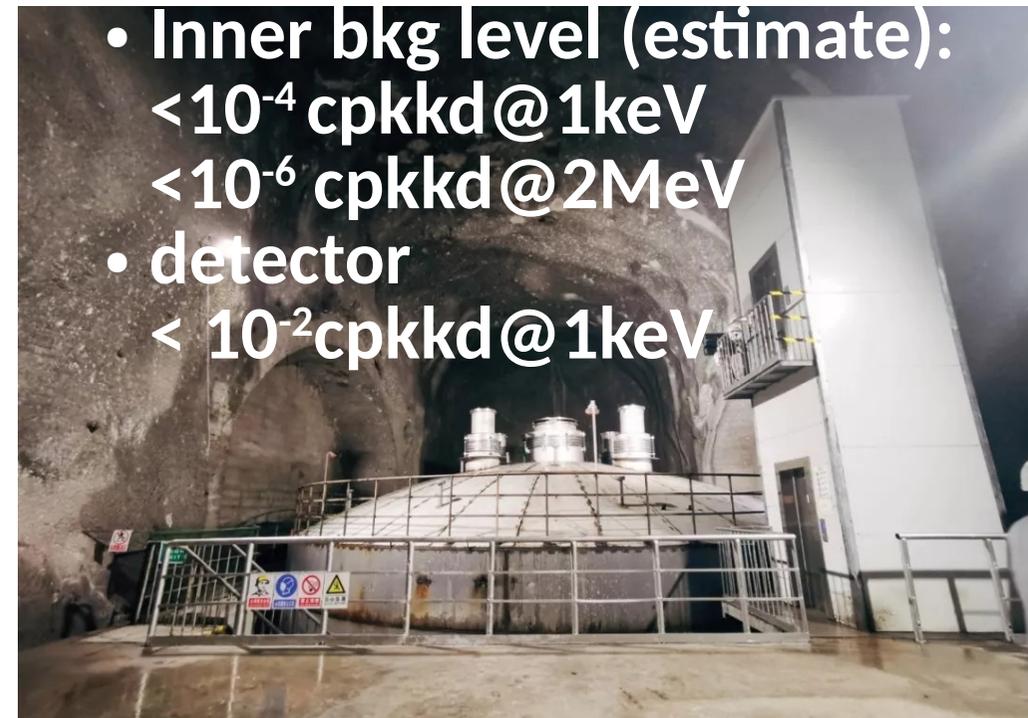
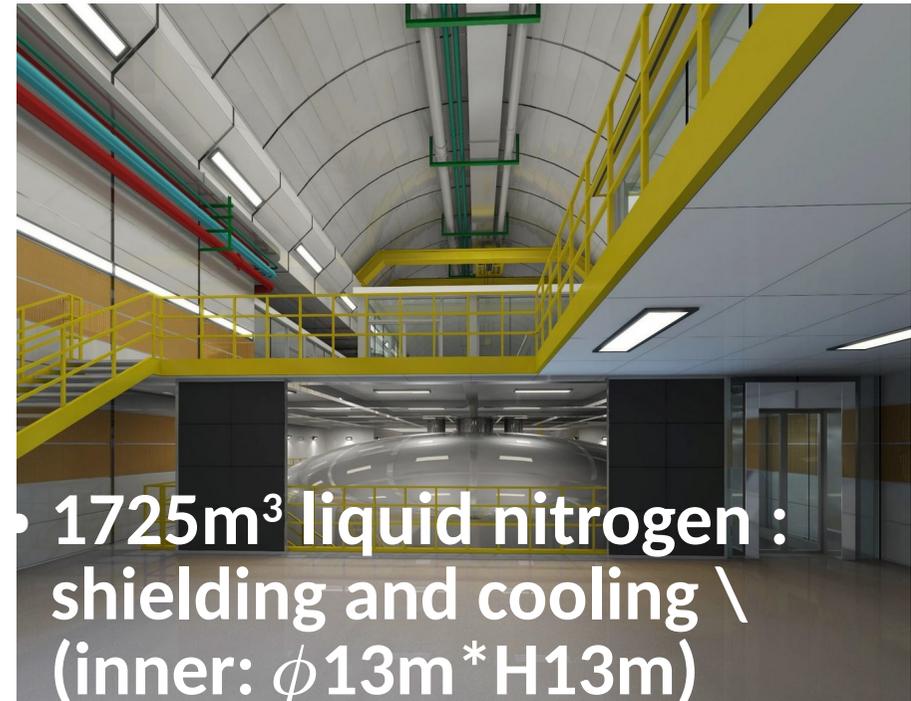
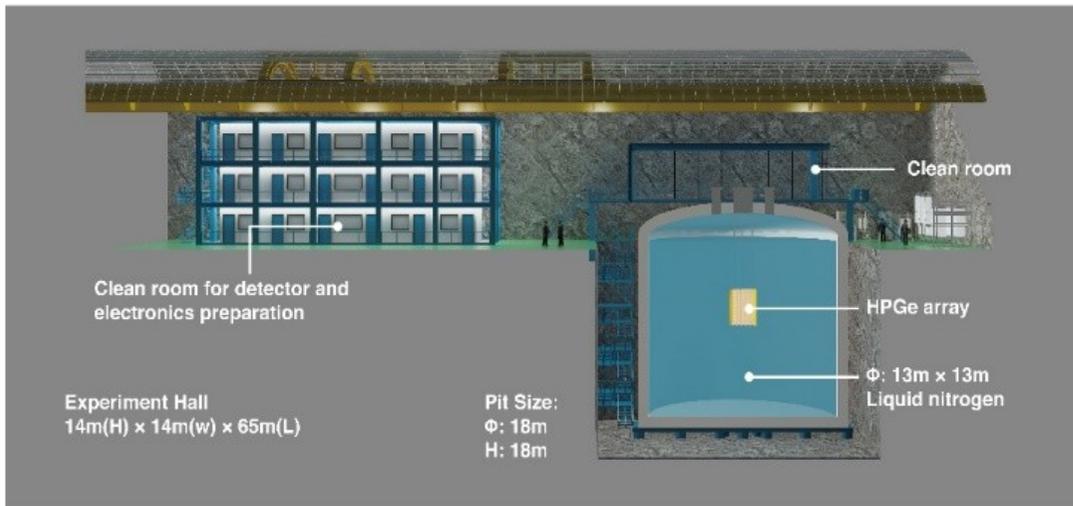
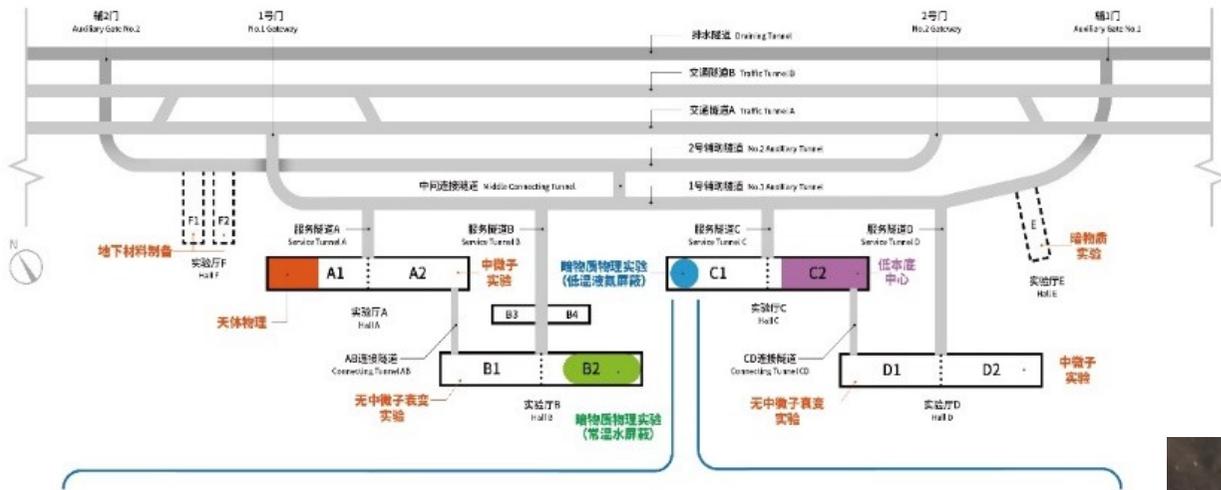
- 3 module × 7 string × 15×1 kg detector, total: ~ 300 kg
- Top clean room for detector installation
- immersed in LN2



LAr cryostat (~keV background?)
with PMT/SiPM readout



CJPL-II: the lab.





Ge detector : Dark matter vs. $0\nu\beta\beta$

- detector suitable for dark matter is also suitable for $0\nu\beta\beta$ (?)

	Dark Matter experiment	$0\nu\beta\beta$ experiment
low background good resolution	at $<\sim$ keV	at ~ 2 MeV
low threshold	noise edge reduction needed	not needed
^{76}Ge : least cosmogenic isotope	could reduce cosmogenic induced background.	reduce cosmogenic induced background, and $\beta\beta$ sources.
underground Ge growth	crucial for ~ 1 keV	crucial for ~ 2 MeV

- CDEX-50 will use natural Ge (mostly) $\sim 10^{-2}$ cpkkd (control transportation & fabrication).
- explore: underground Ge growth and manufacturing $\sim 10^{-4}$ - 10^{-5} cpkkd.



summary

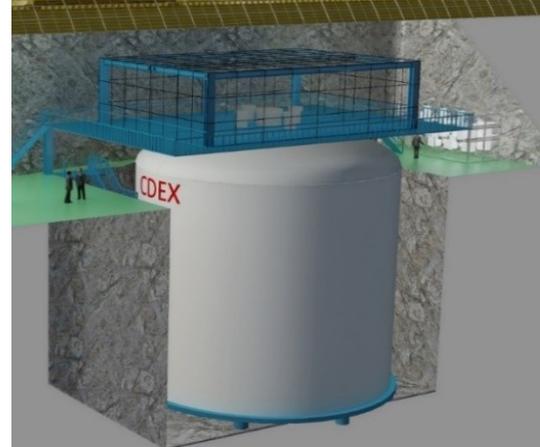
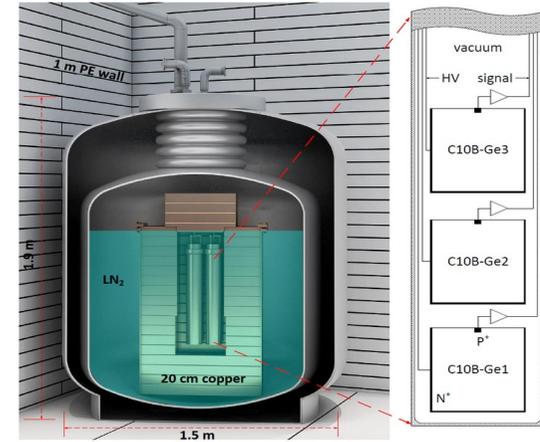
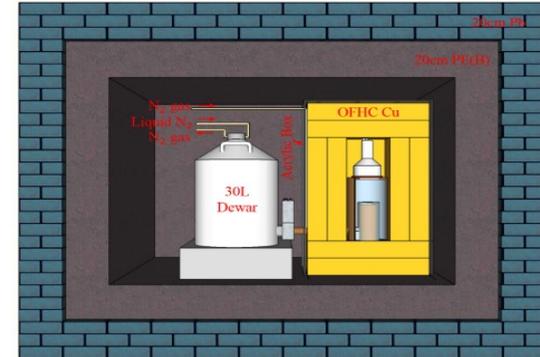
- low threshold Ge detector dark matter searches.
- CDEX-1 + CDEX-10 results.
- CDEX-50 with 50 kg Ge immersed in LN2: installing.
- CDEX-300v for $0\nu\beta\beta$ (+DM ?): soon.
- CDEX-1000 for dark matter and $0\nu\beta\beta$: future plan.

Thanks

backup

CDEX stages

- Light WIMP mass searches on Ge
- CDEX-1: Development of pPC-HPGe detector, its background understanding, annual modulation results.
- CDEX-10: Ge immersed on LN2, various DM results.
- CDEX-50: Ge immersed on LN2 array for DM, plan.
- CDEX-300v: ^{76}Ge enriched array for $0\nu\beta\beta$, plan.
- CDEX-1000: Multi-purpose experiment for dark matter and double beta decay.



CDEX-1

CDEX-1A 1kg PCGe



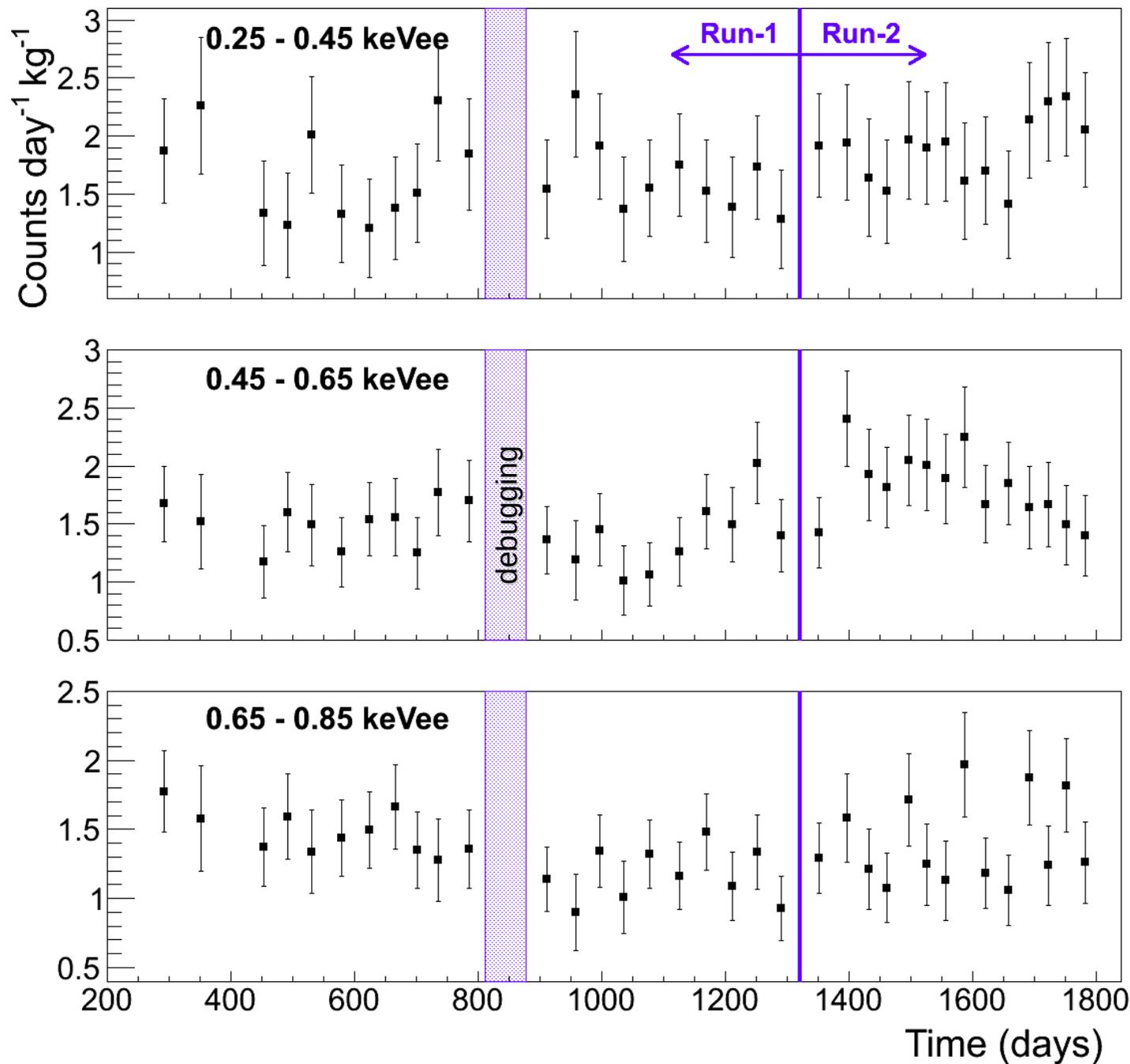
CDEX-1B 1kg PCGe



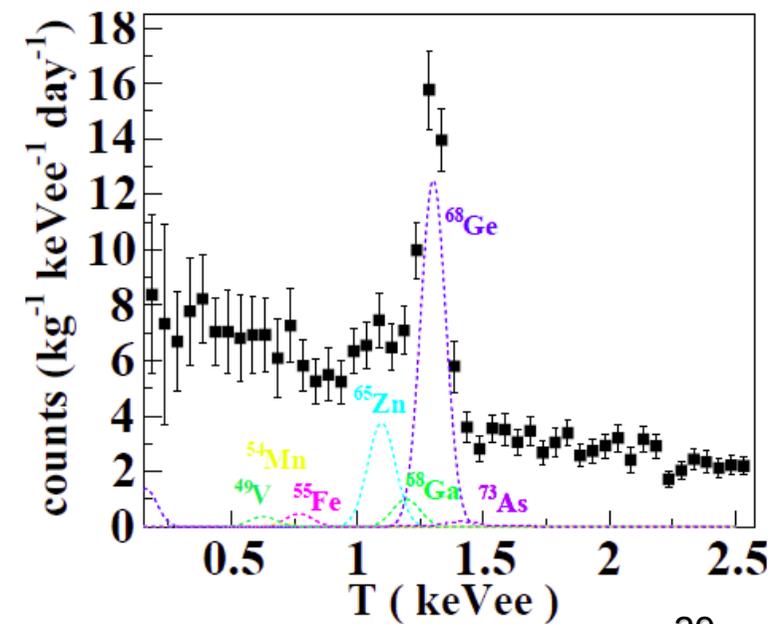
20cm OFHC Copper
+20cm Lead



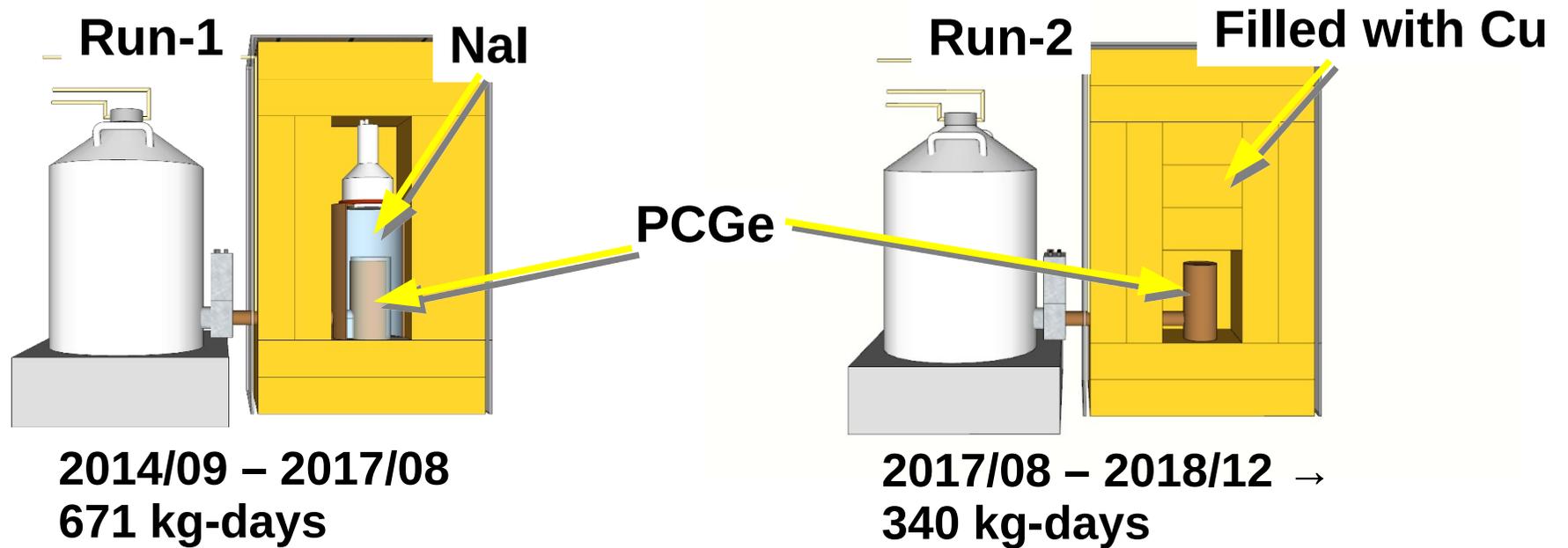
CDEX-1B data < 0.85 keV



- run-1 to run-2: change of shielding.
- 0.25 – 0.85 keV: most important region for low mass WIMP
- χ^2 test consistent with null-hypothesis.

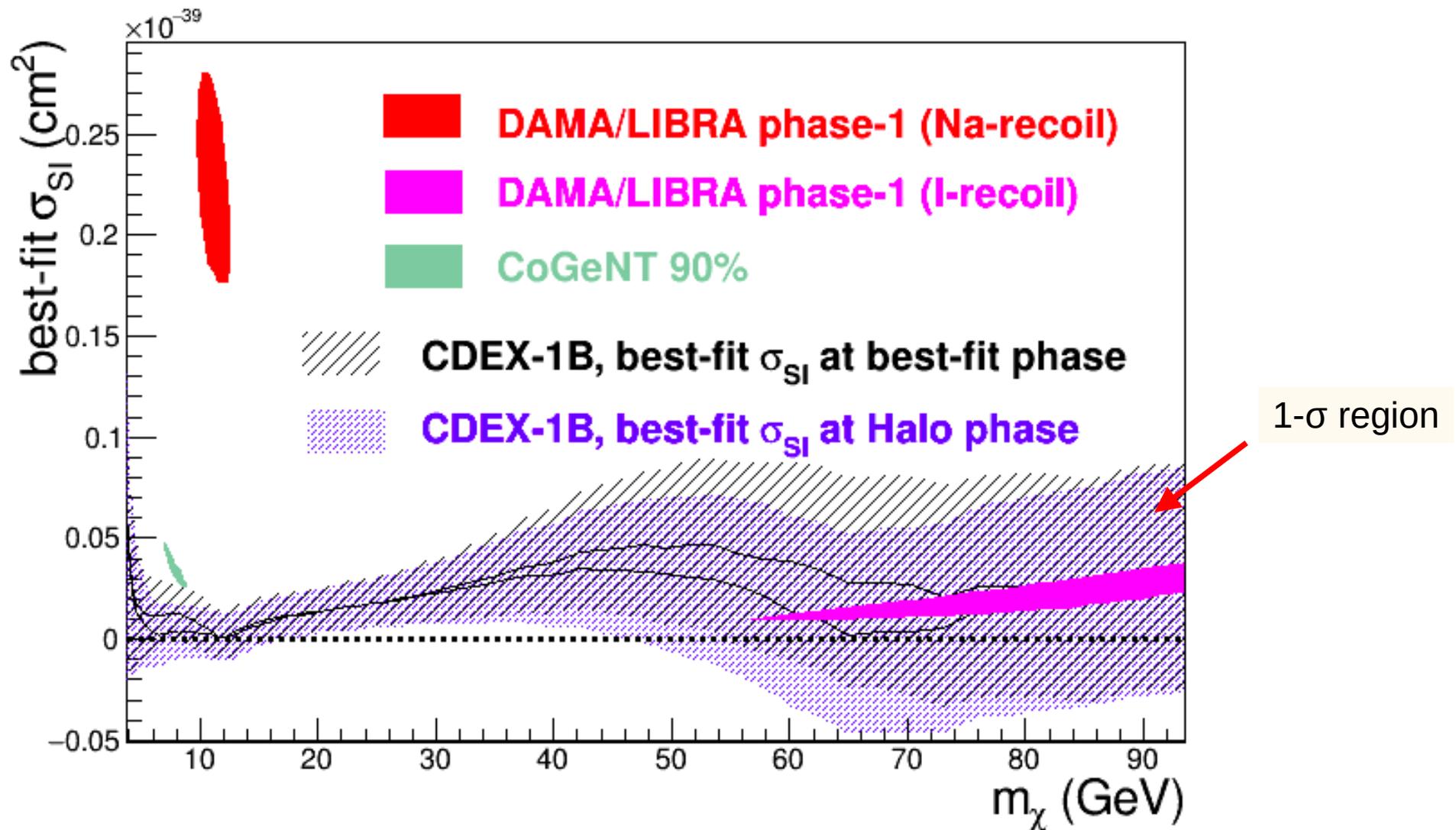


CDEX-1B experiment



- 1 kg-scale-mass HPGe detector, cooled by cold finger.
- A NaI(Tl) detector is used as active shielding to veto the gamma-ray induced background events.
- The detector has been under stable data taking conditions since March 27th, 2014.
- Threshold ~ 160 eVee. For modulation analysis, threshold ~ 250 eVee.
- Largest analysis uncertainties: bulk/surface separation at low energy.

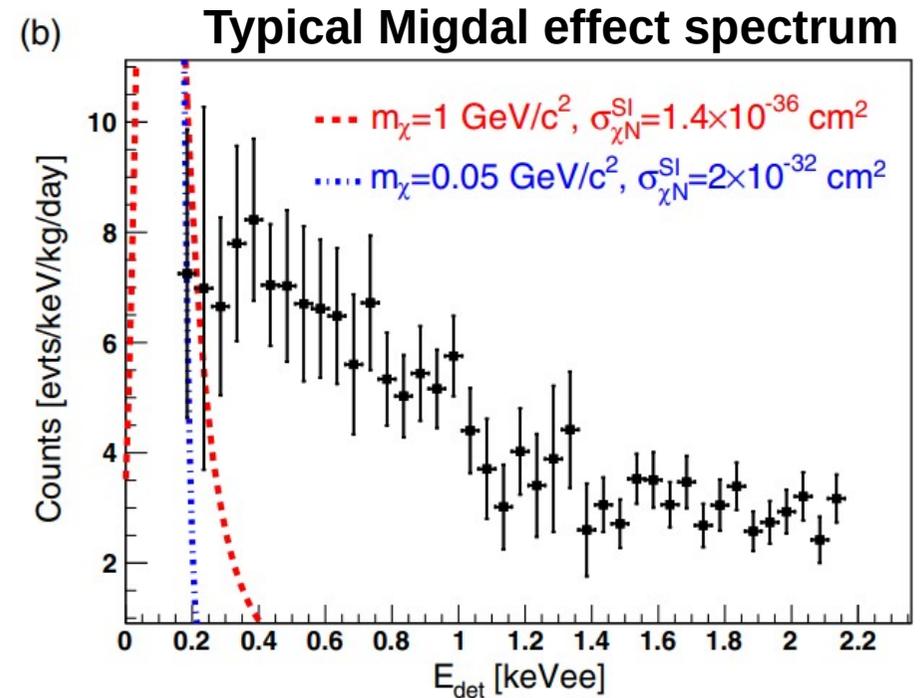
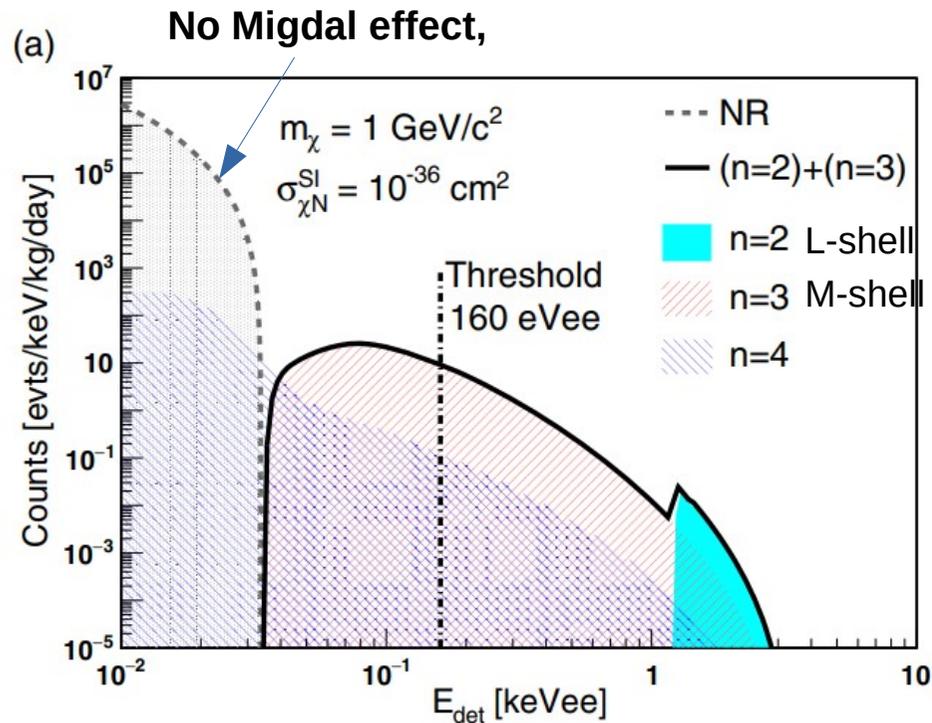
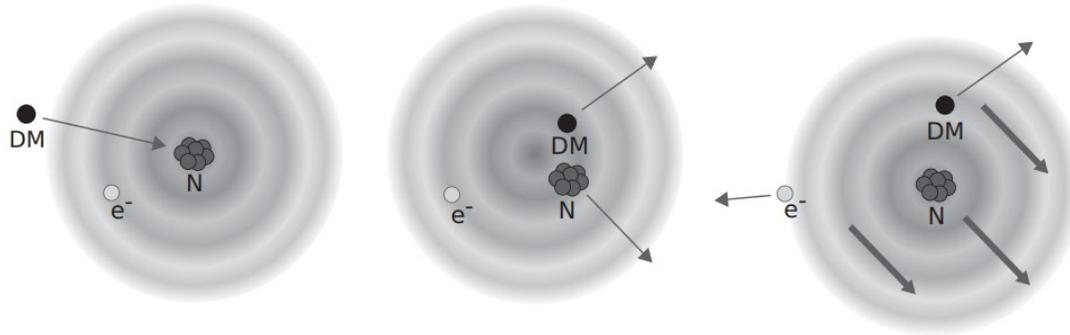
σ_{SI} at fixed/best-fit phase



- The results consistent with null-hypothesis at any phases (within 2- σ) up to 100 GeV.

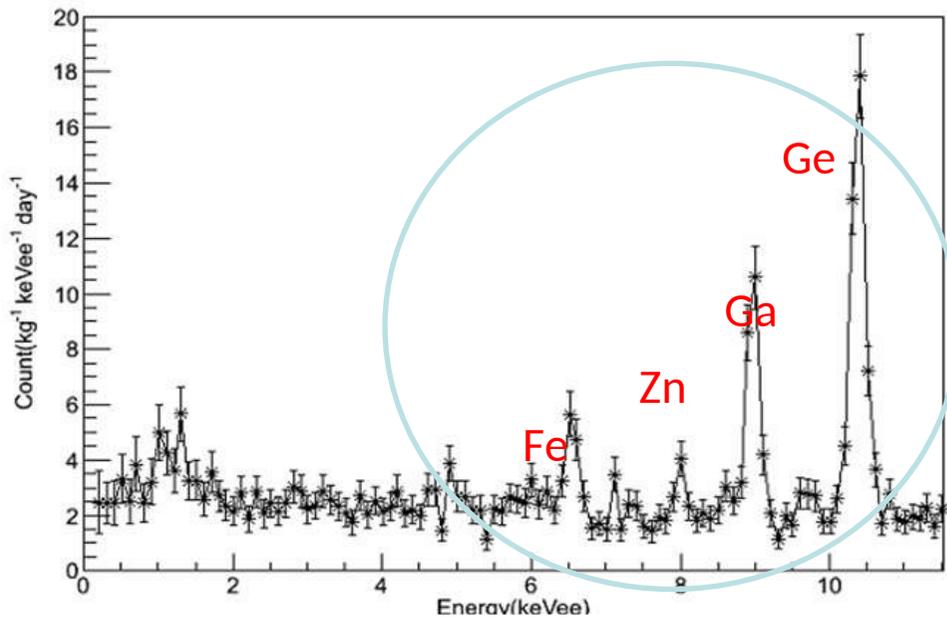
Migdal effect

nuclear recoil \rightarrow electrons cloud move, except one $e^- \rightarrow$ ionization

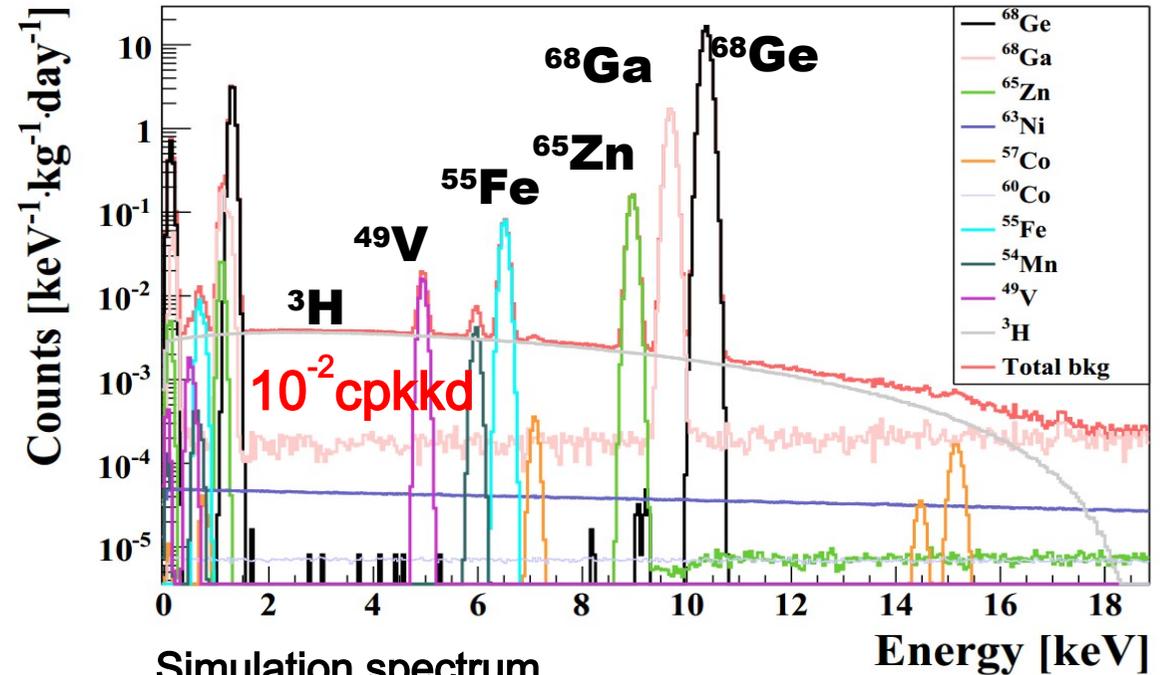


Cosmogenic Background of Ge detector

- Long-time ground preparation of detector induces high cosmogenic background level ($\sim 2 \text{ cpkkd}$ @ 2-4 keV);
- Based on simulation, **2 months ground fabrication and transportation** could decrease the ^3H continuous background level to $\sim 10^{-2} \text{ cpkkd}$ @ 2-4 keV.



CDEX-10 background spectrum



Simulation spectrum

Cosmogenic Background of U-Ge detector

- Underground germanium crystal growth and detector fabrication could dramatically decrease the cosmogenic backgrounds from non-Ge isotopes, such as ^3He , ^{65}Zn ;
- ^{76}Ge Enriched germanium material could help to decrease $^{68}\text{Ge}(^{68}\text{Ga})$ cosmogenic backgrounds too. (2nu background $\sim 10^{-5}\text{cpkkd}$ at 1keV)

