

GAPS

Hunting for Dark Matter with Cosmic-Ray Antimatter

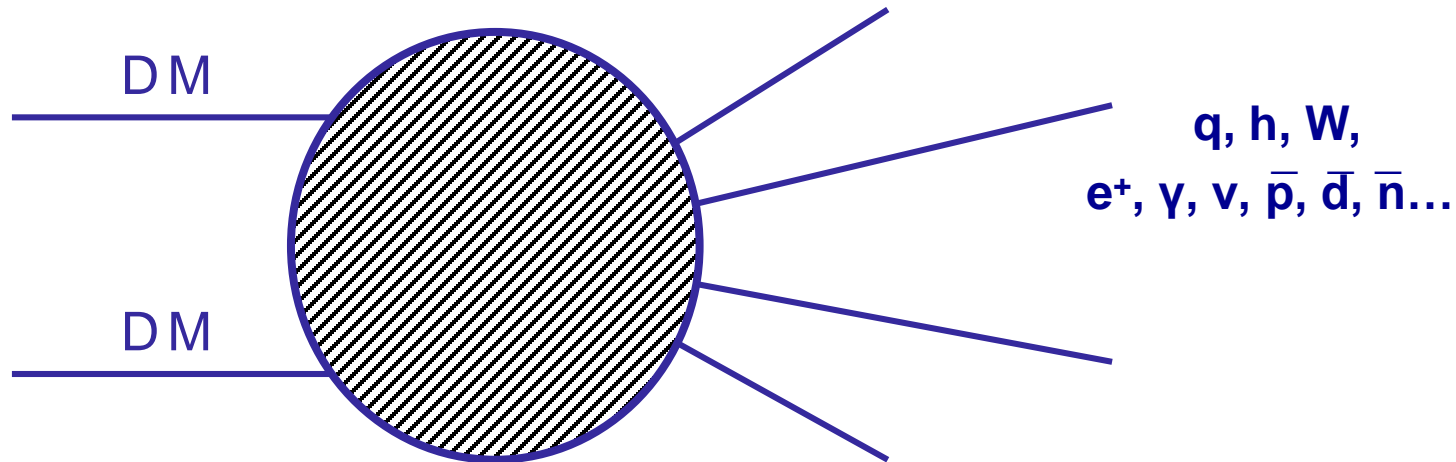


TSUGUO ARAMAKI,
SLAC
For GAPS
Collaboration

- **Indirect DM Search**
 - Recent results from Fermi, AMS-02
- **GAPS Mission**
 - Why antideuterons?
 - Detection method/particle identification technique
 - Detector design
 - Recent status
 - Sensitivity and complementarity to other DM experiments
- **Summary**

INDIRECT DM SEARCH

Measure DM Annihilation/Decay Products



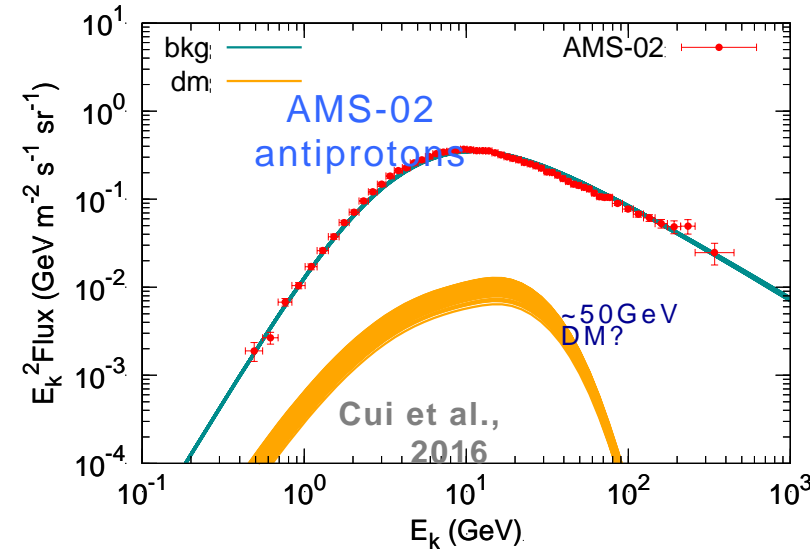
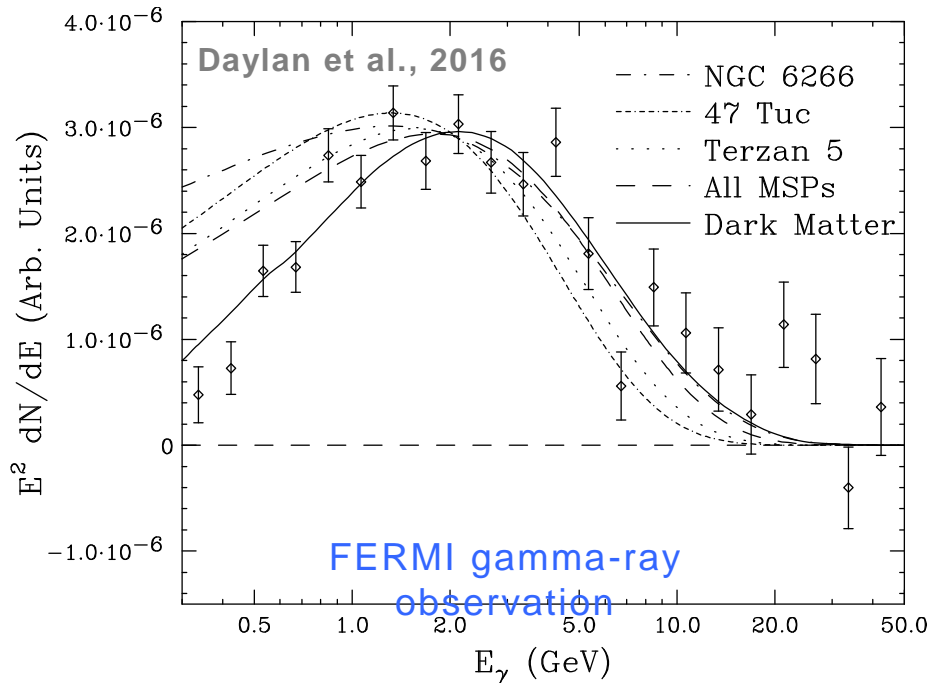
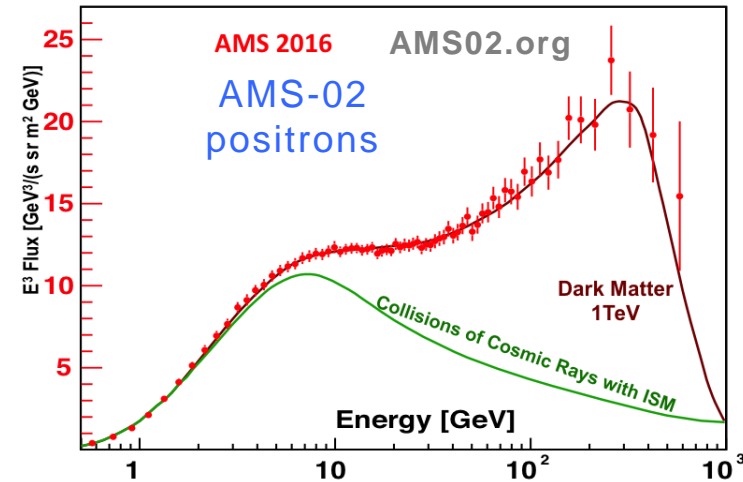
- Positron: AMS-02, Fermi-LAT, PAMELA, ...
- Gamma-ray: Fermi-LAT, CTA, HESS, VARITAS...
- Neutrino: IceCube, ANTARES...
- Antiproton: AMS-02, PAMELA, **GAPS**
- Antideuteron : AMS-02, **GAPS**
- Antihelium: AMS-02, **GAPS**

Complementary searches with different detection methods/background models are crucial to validate

INDIRECT DM SEARCH – FERMI AND AMS-02 RESULTS

Possible DM Signatures?

- Fermi Galactic Center Excess
 - ~50GeV DM/Astrophysical objects?
 - Similar excess in Fermi dSphs?
- AMS-02 positrons/antiprotons
 - DM/pulsars/propagation uncertainty?
- AMS-02 antihelium detection?



Difficult to verify DM signatures due to

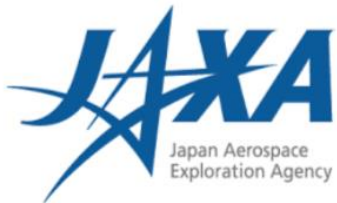
GAPS COLLABORATION

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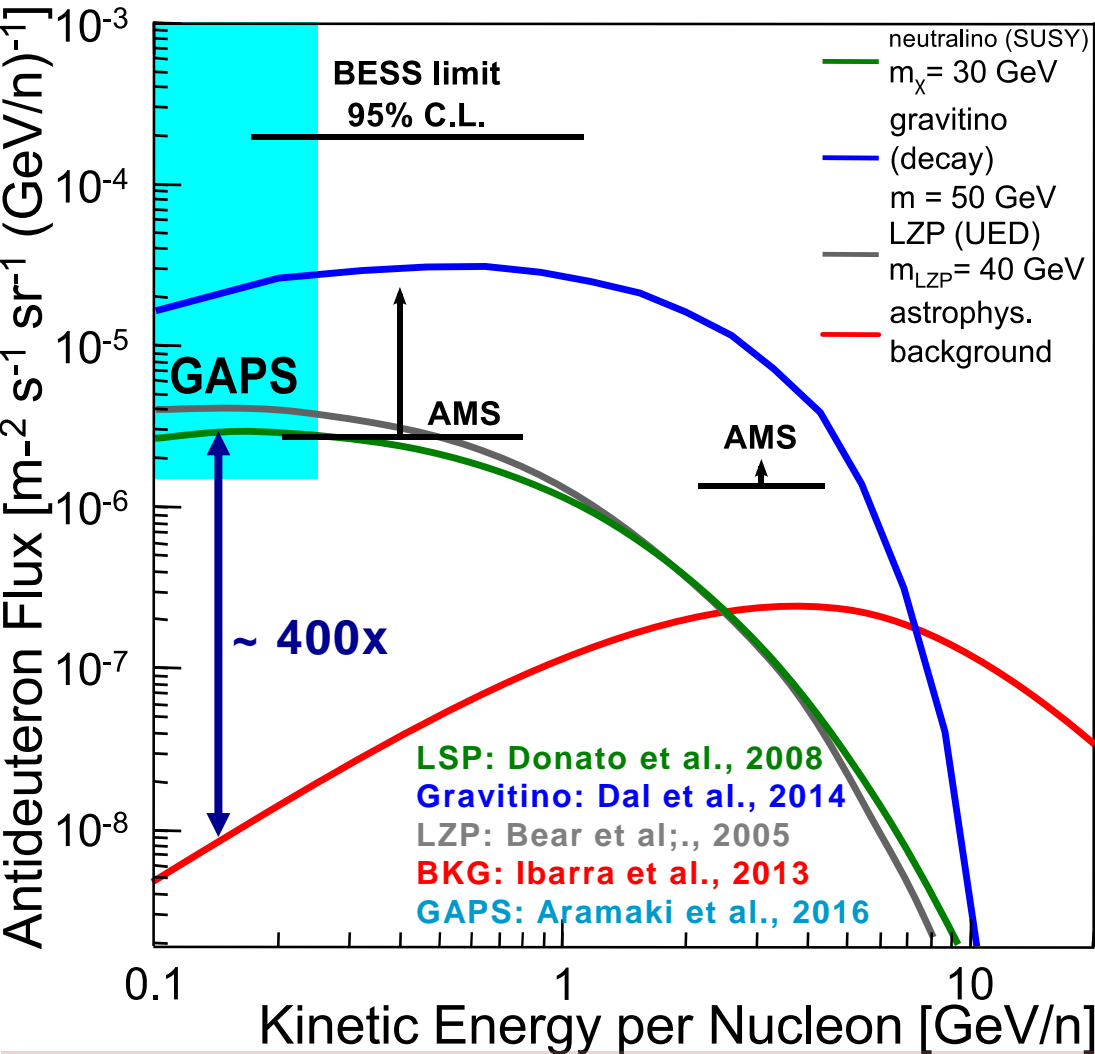
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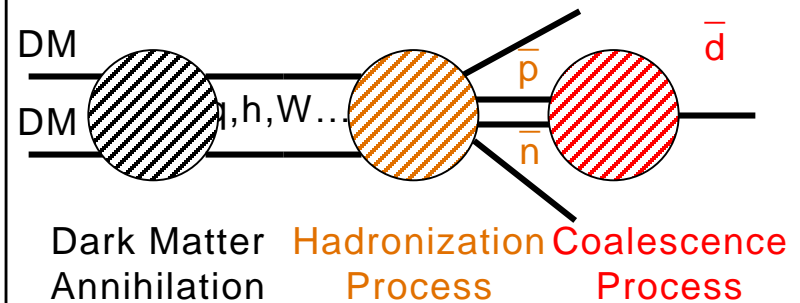
WHY ANTIDEUTERONS?

Background-Free DM Search at Low-Energy



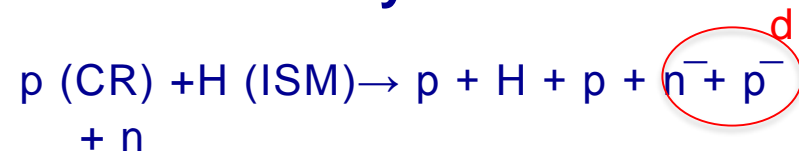
Primary flux

DM annihilation/decay



Secondary flux

Cosmic ray interaction

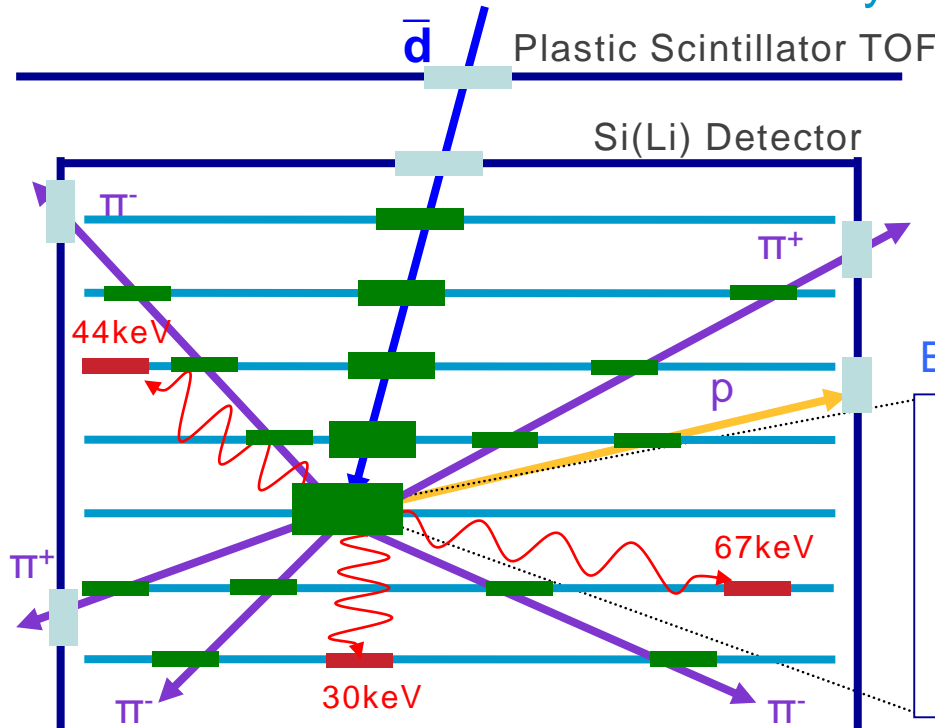


GAPS antideuteron measurement plays an important

GAPS DETECTION CONCEPT

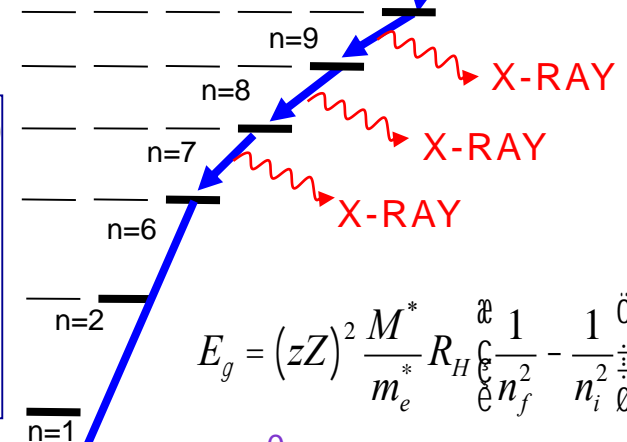
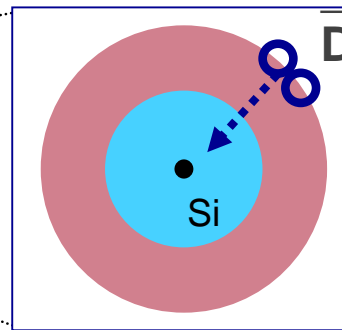
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Measure Atomic X-rays and Annihilation Products



ATOMIC TRANSITIONS

EXOTIC ATOM



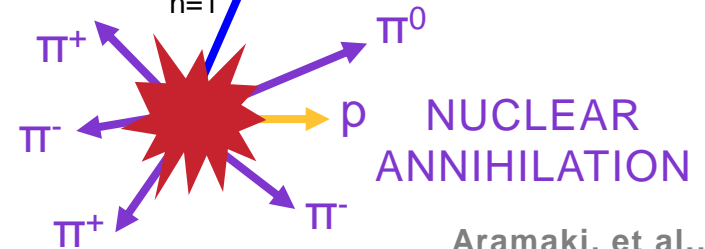
$$E_g = (zZ)^2 \frac{M^*}{m_e} R_H \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

A time of flight (TOF) system tags candidate events and records velocity

The antiparticle slows down & stops, forming an excited exotic atom

Deexcitation X-rays provide signature

Annihilation products provide added background suppression



Aramaki, et al., 2013

Concept proven with accelerator beam test
 Measured/verified X-ray yields with different targets
 Developed cascade model to predict X-ray yields

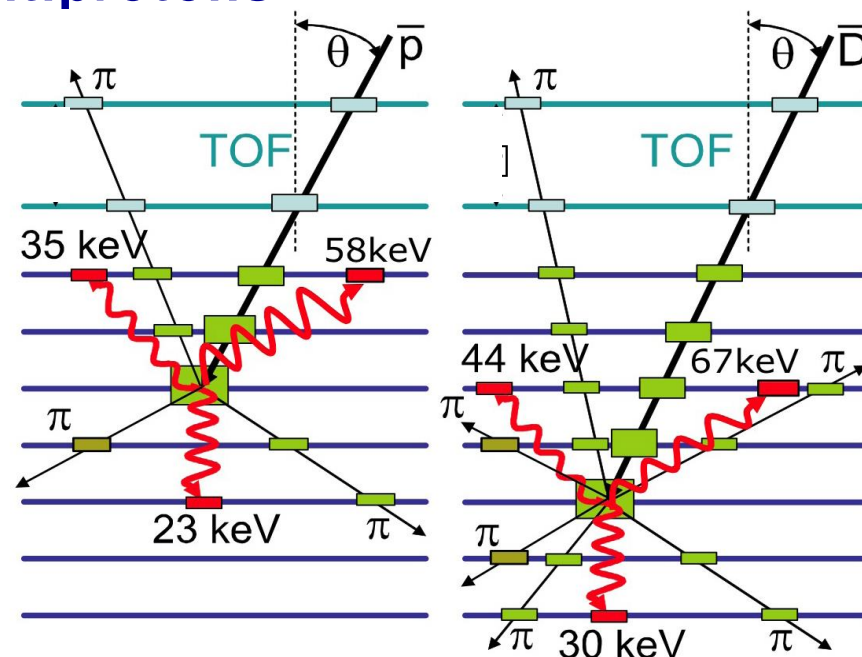
GAPS ANTIDEUTERON IDENTIFICATION TECHNIQUE

CR p, e^\pm rejection: antiproton and antideuteron selection

- Select **slow particles** with TOF
- Simultaneous detection of **annihilation products**
 - relativistic pions/slow protons

Antideuteron identification from antiprotons

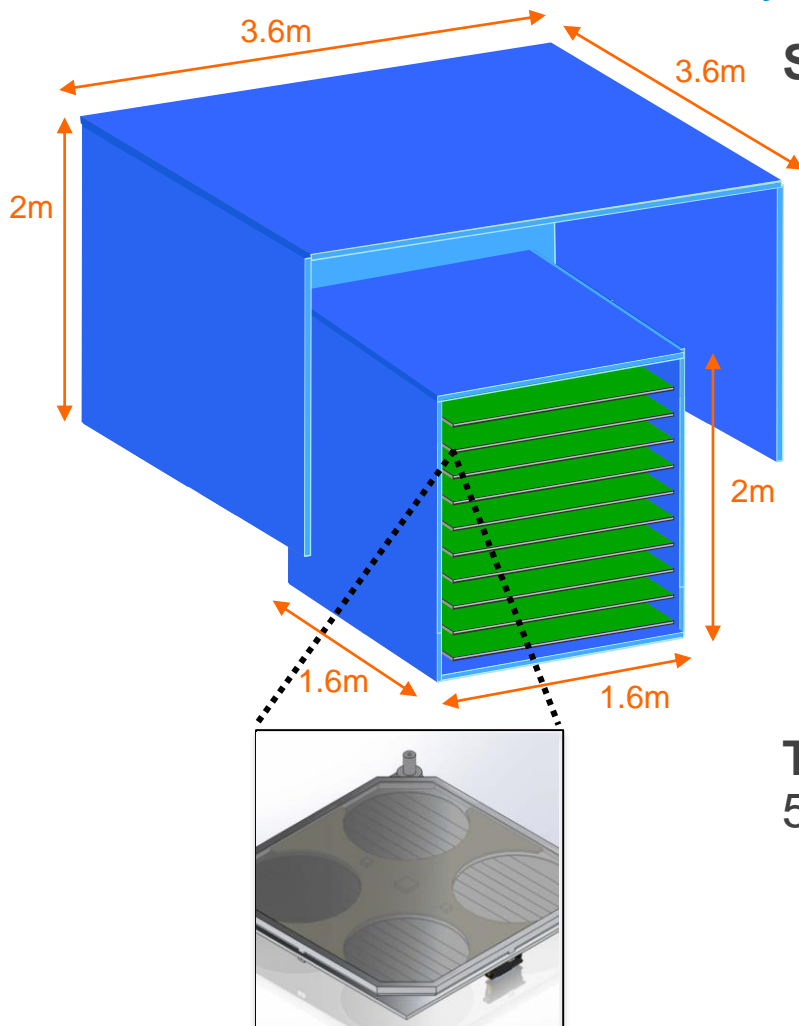
- **atomic X-rays** from exotic atom
 - different energy
- **pion/proton multiplicity**
 - more for antideuterons
- **stopping range** (depth sensing)
 - antideuterons go deeper
- **dE/dX** energy deposit in layers
 - more for antideuterons



Background/Mimic Events ~ 0.01

GAPS INSTRUMENT – DESIGN CONCEPT

Well-studied, widely-used Si(Li) & plastic scintillators



Si(Li) detector: 4 inch, 2.5mm thick wafer

- 10 layers, ~140 Si(Li) detectors/layer
- Segmented into 8 strips
-> 3D particle tracking
- Energy/timing resolution: ~4 keV, ~100 ns
- Operation temperature: -40C
cooled with oscillating heat pipe (OHP)
- Dual channel electronics
20-80 keV: X-rays
0.1-50MeV: charged particles

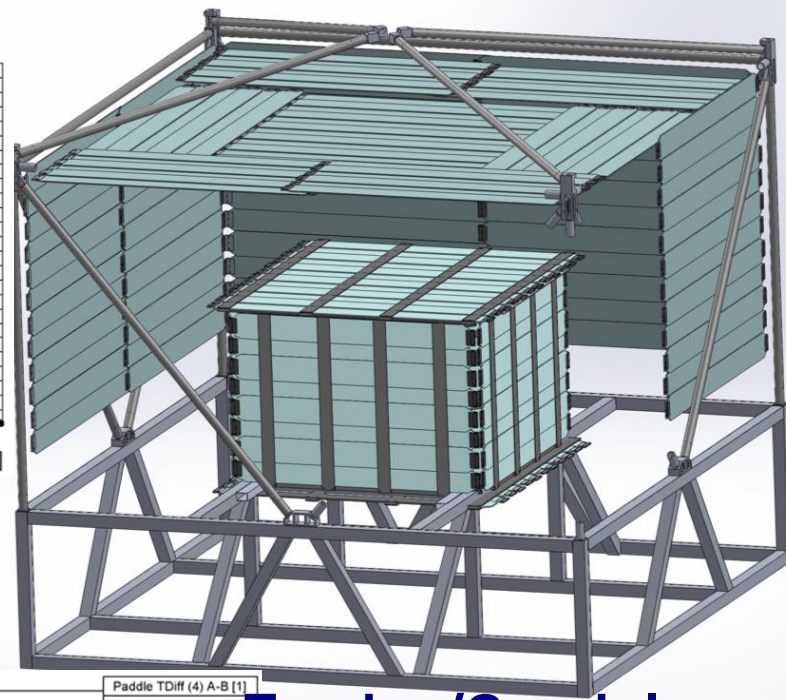
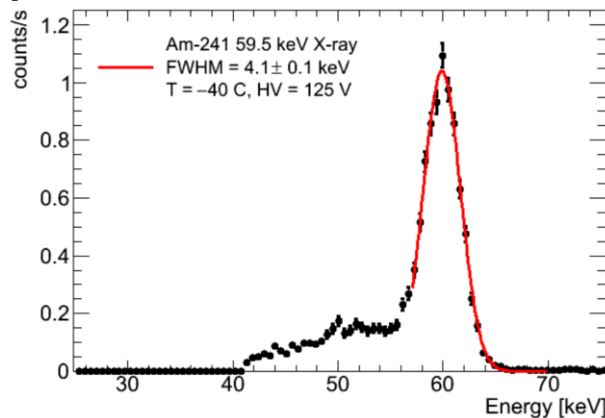
TOF plastic scintillators: 18cm x 1.6m x 5mm

- Identify incoming charged particles
- 1m separation between inner/outer TOF
- Timing resolution: ~0.5ns, SiPM on each

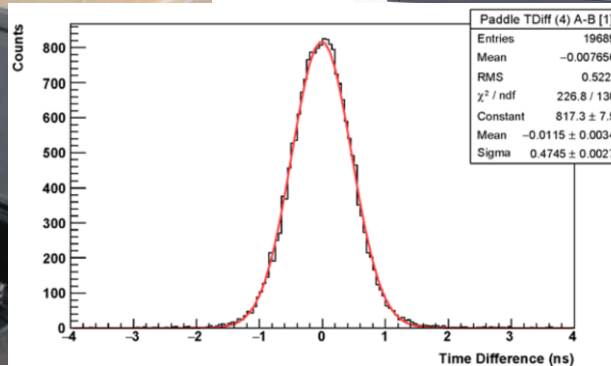
Key instruments were tested/validated in the ⁹

GAPS INSTRUMENT – RECENT STATUS

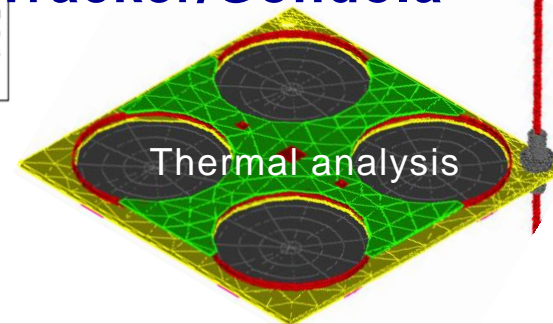
See Field Rogers' poster for more details
Si(Li) detector



TOF Plastic scintillator



Tracker/Gondola

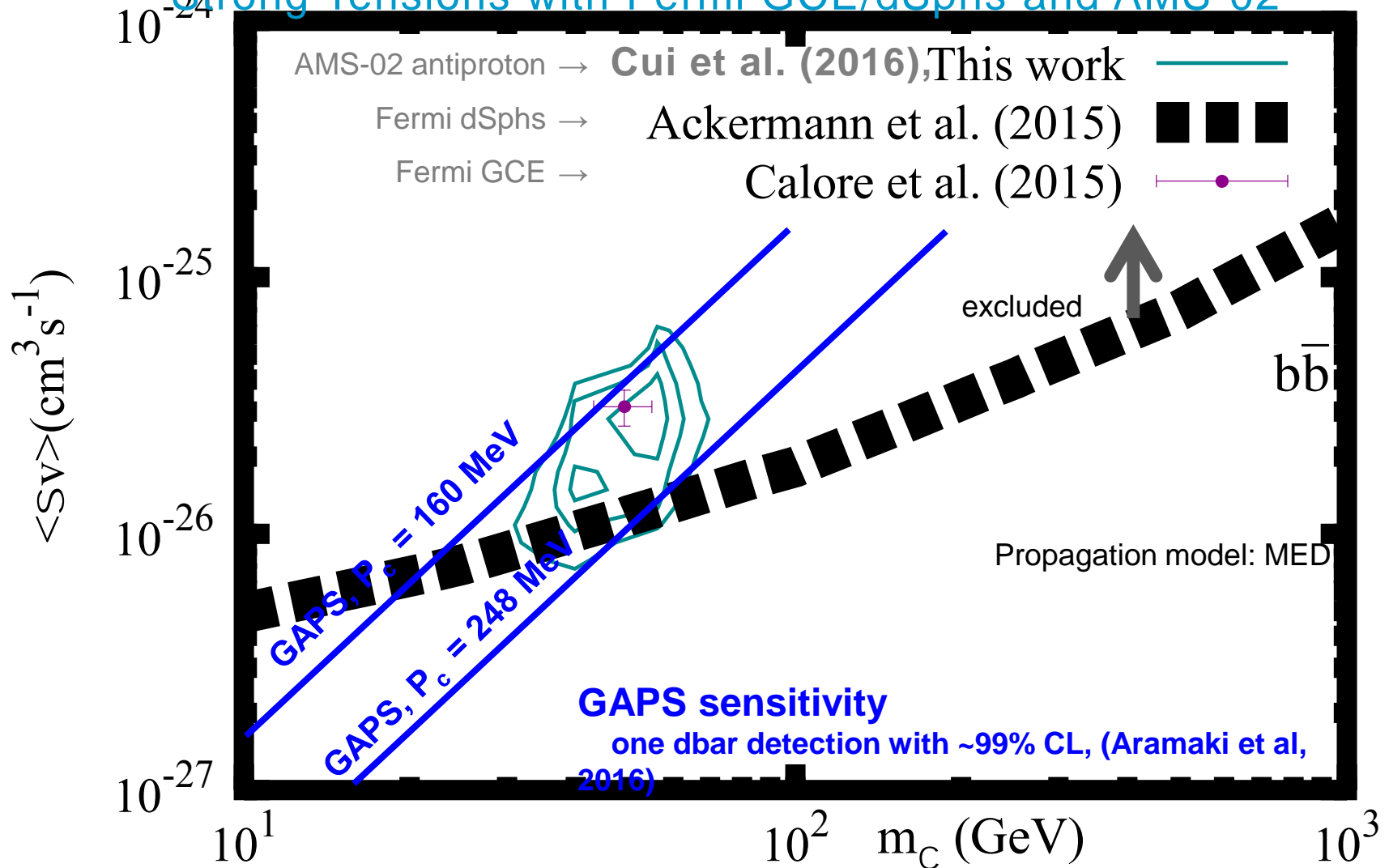


We are on track for upcoming PDR later this year

GAPS SENSITIVITY IN DM PARAMETER SPACE

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Strong Tensions with Fermi GCE/dSphs and AMS-02



GAPS can elucidate the tensions from a different¹

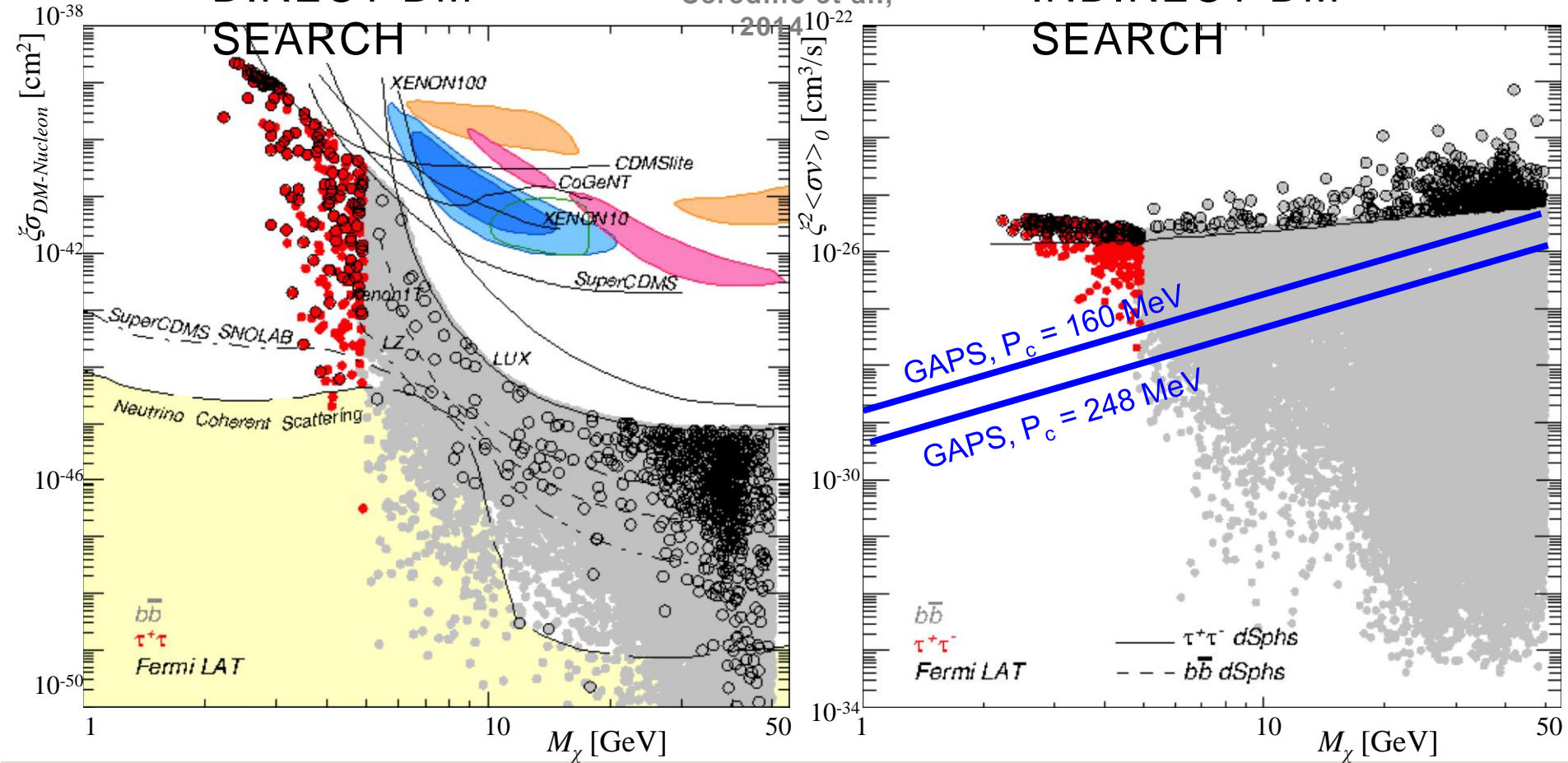
GAPS COMPLEMENTARY DM SEARCH

RH Sneutrinos in NMSSM

Ceredino et al., 2014

DIRECT DM SEARCH

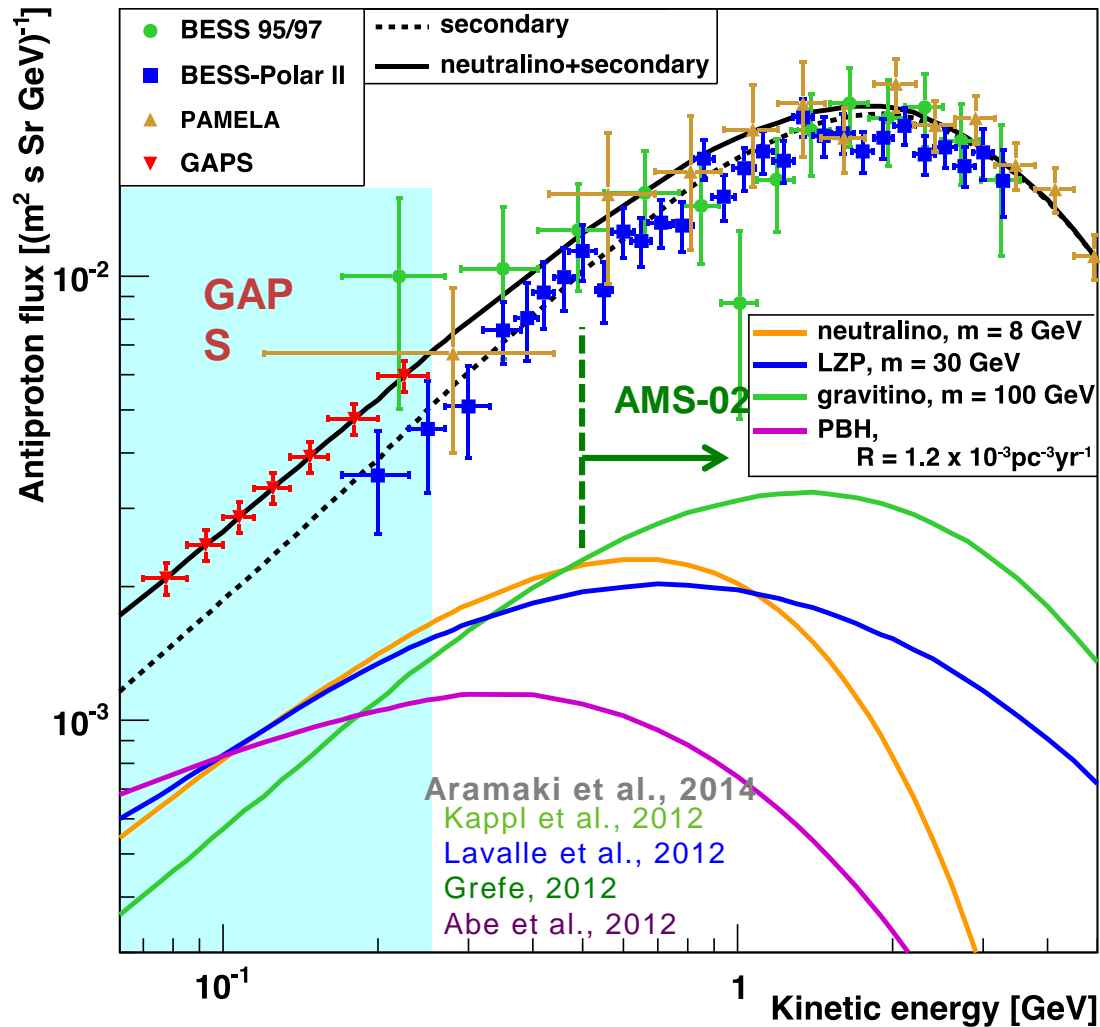
INDIRECT DM SEARCH



GAPS, SuperCDMS and LZ/XENON can complementarily investigate DM parameter space

GAPS ANTIPROTON MEASUREMENT

GAPS Can Exclusively Measure Low-Energy Antiprotons



Low-mass DM search

- **SUSY LSP**
neutralino, RH sneutrino
 - **LZP**
extra dimensions
 - **Gravitino**
small R-parity violation
 - **PBH Evaporation**
- Solar modulation**

- **BESS**
29 at $E \sim 0.2$ GeV
- **PAMELA**
7 at $E \sim 0.25$ GeV
- **AMS-02**
 $E > 0.5$ GeV
- **GAPS**
 ~ 1500 at $E < 0.25$ GeV

GeV

GAPS can uniquely explore low-mass DM, PBH and

SCHEDULE AND SUMMARY

SLAC

First Science Flight funded, Scheduled from Antarctic in Late-2020



- GAPS antideuteron measurement is considered as background-free DM search and can uniquely explore DM parameter space.
- GAPS antideuteron measurement can elucidate the tensions between Fermi GCE/dSphs and AMS-02 antiproton measurement.
- GAPS antiproton measurement can deeply investigate low-mass DM, PBH and solar modulation.

