Year 1 Cosmology Results from the Dark Energy Survey

Elisabeth Krause (Caltech/JPL)
on behalf of the Dark Energy Survey collaboration

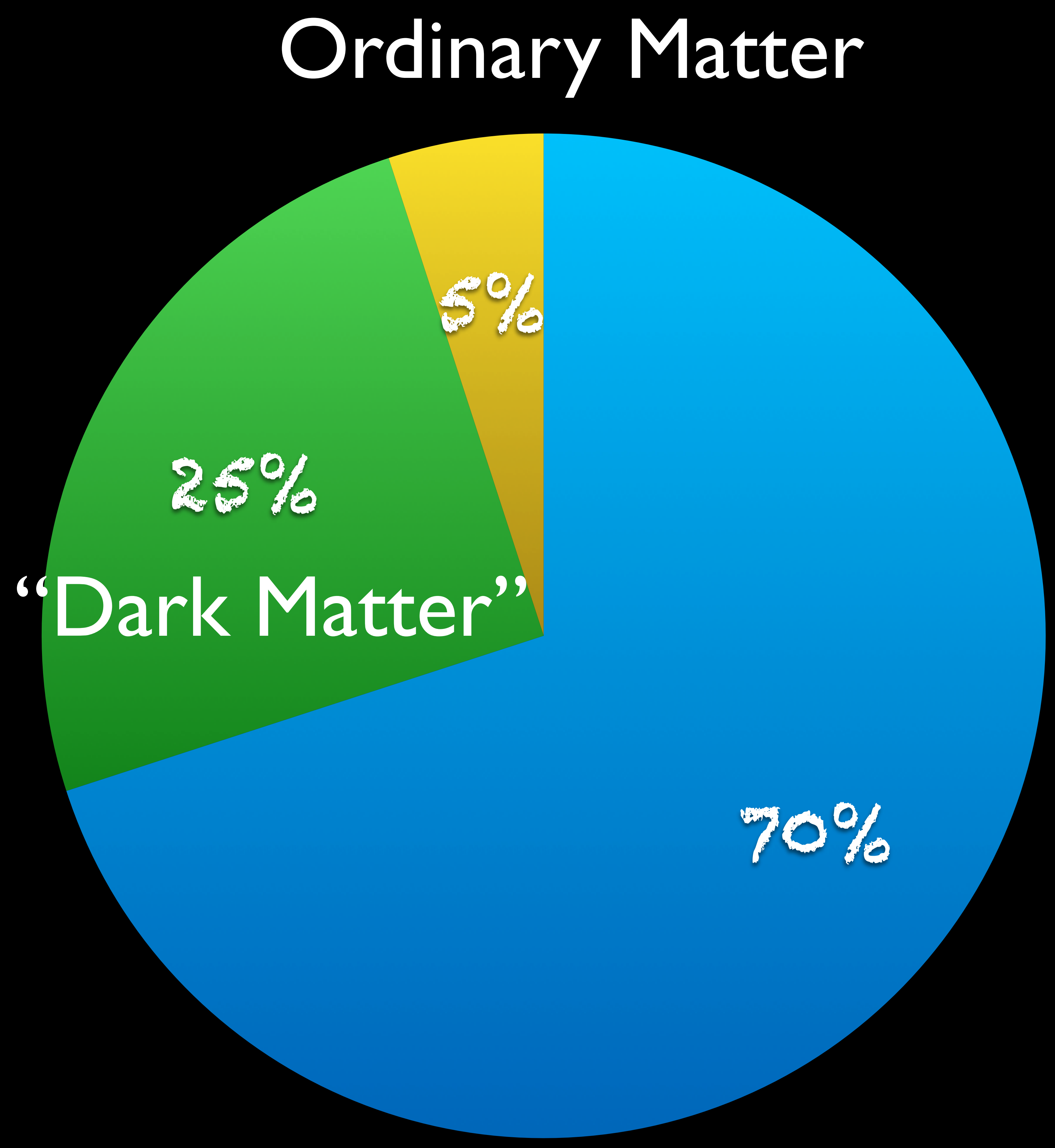
Dark Matter 2018, UCLA

Our Simple Universe

- On large scales, the Universe can be modeled with remarkably few parameters
  - age of the Universe
  - geometry of space
  - density of atoms
  - density of matter
  - amplitude of fluctuations
  - scale dependence of fluctuations

[of course, details often not quite as simple]

Our Puzzling Universe



"Dark Energy"

- accelerates the expansion
- dominates the total energy density
- smoothly distributed

acceleration first measured by SN 1998

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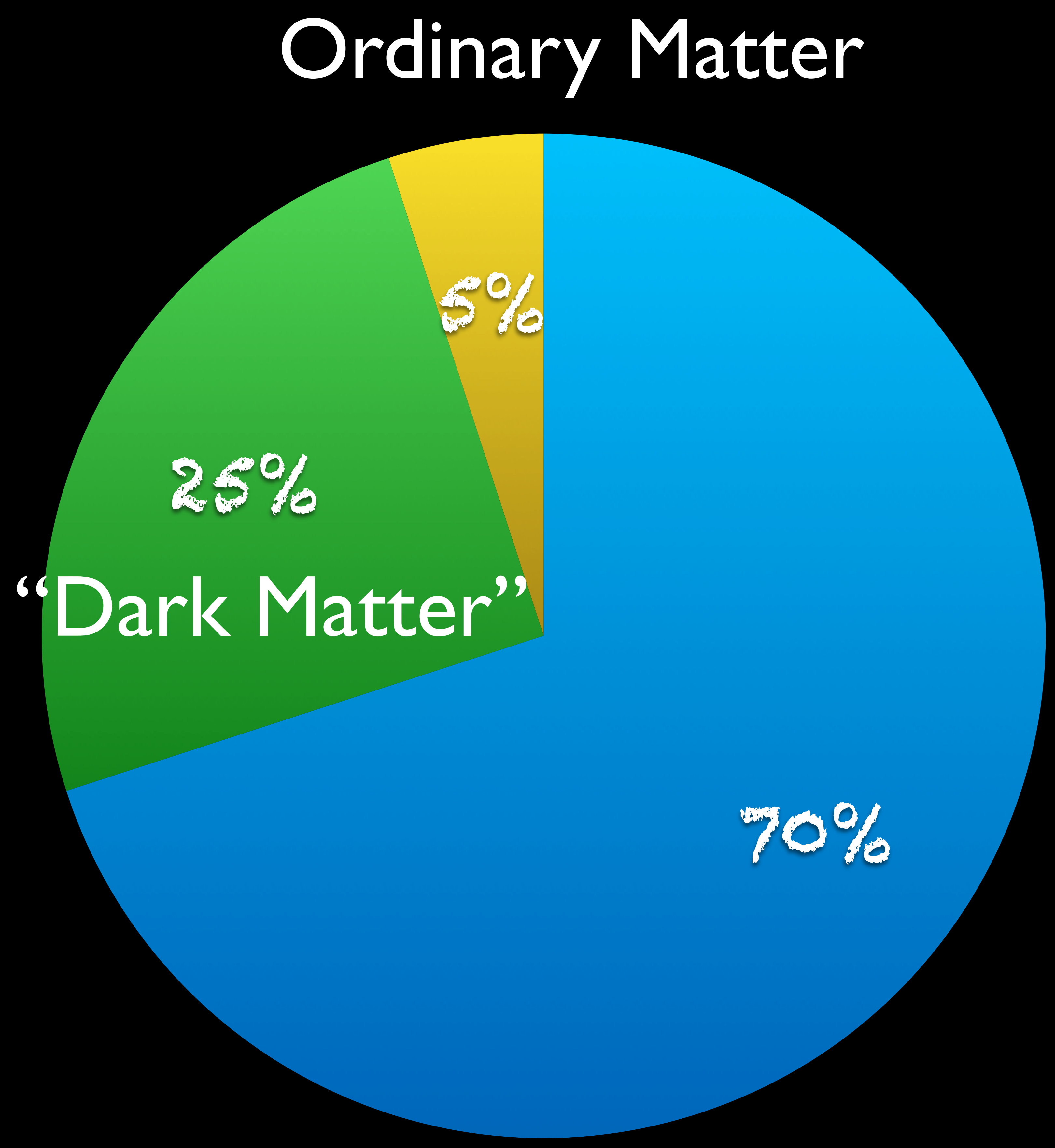

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next frontier: understand

- cosmological constant Λ: $w \equiv P/\varrho = -1$?
  - magnitude of Λ very surprising
- dynamic dark energy varying in time and space, w(a)?
- breakdown of GR?

Theoretical Alternatives to Dark Energy

Many new DE/modified gravity theories developed over last decades

Most can be categorized based on how they **break GR**:

The only local, **second-order** gravitational field equations that can be derived from a **four-dimensional action** that is constructed **solely from the metric tensor**, and admitting Bianchi identities, are GR + Λ.

Lovelock's theorem (1969)

[subject to viability conditions]

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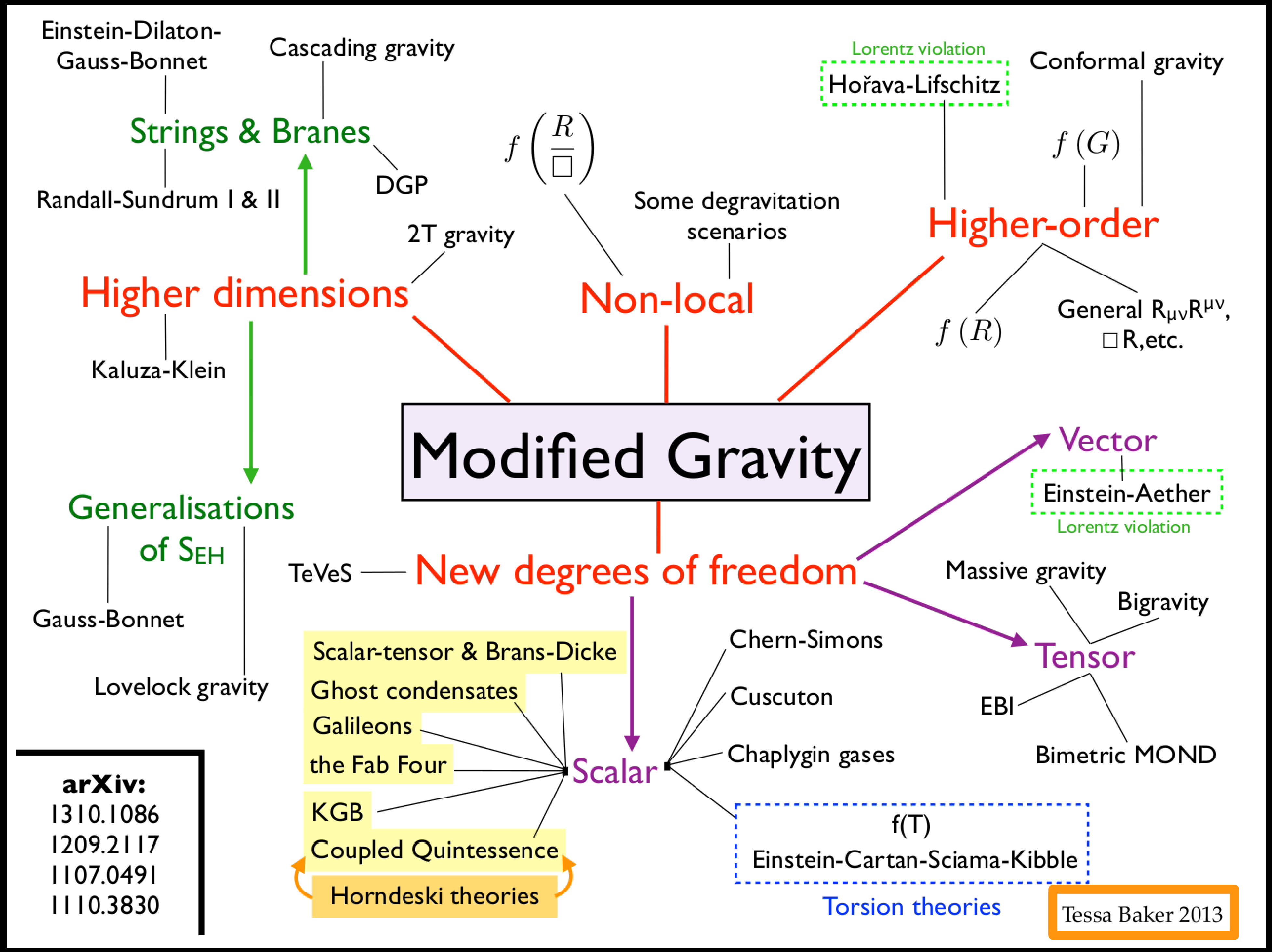

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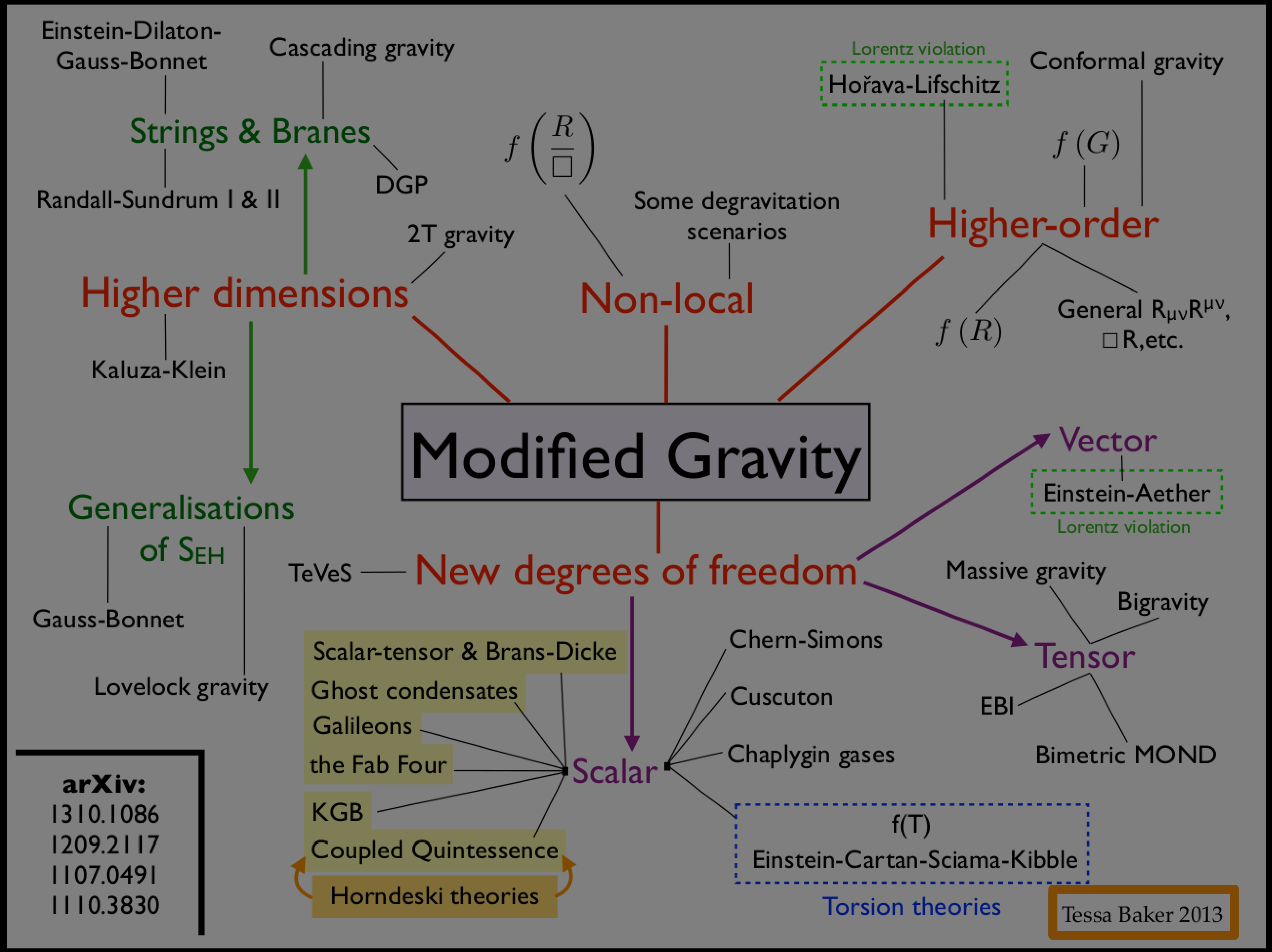


Need simple tests to confront classes of models with data

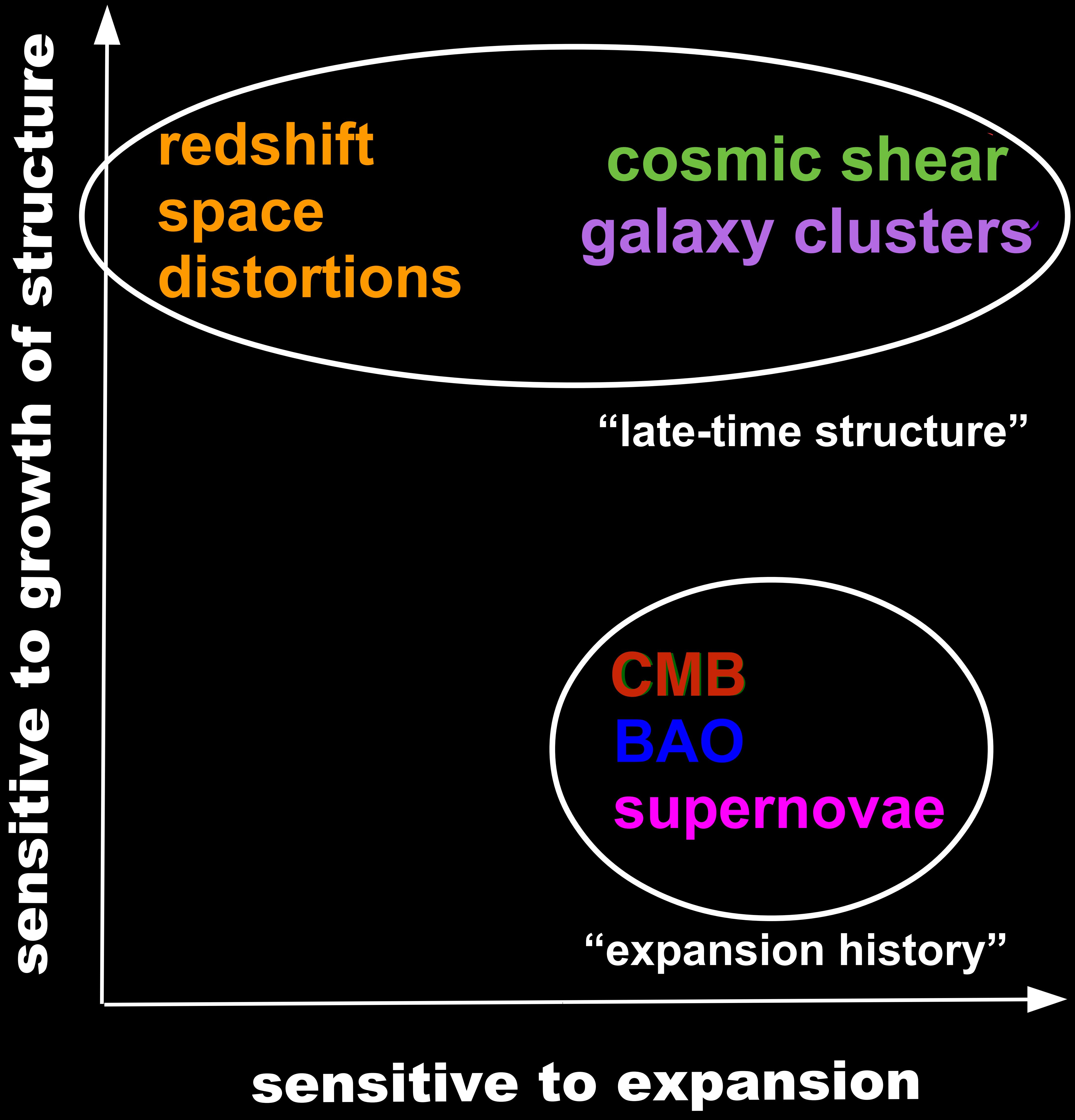
Are data from **early Universe** and **late Universe** fit by the same parameters?

Do measurements of **cosmic distances** and growth of **structure** agree?

Does the dark energy density change as space expands?

"Equation of state" parameter **w=pressure/density**

How to survey Dark Energy



Q: Do all these measurements agree with predictions in the same, fiducial ΛCDM model?

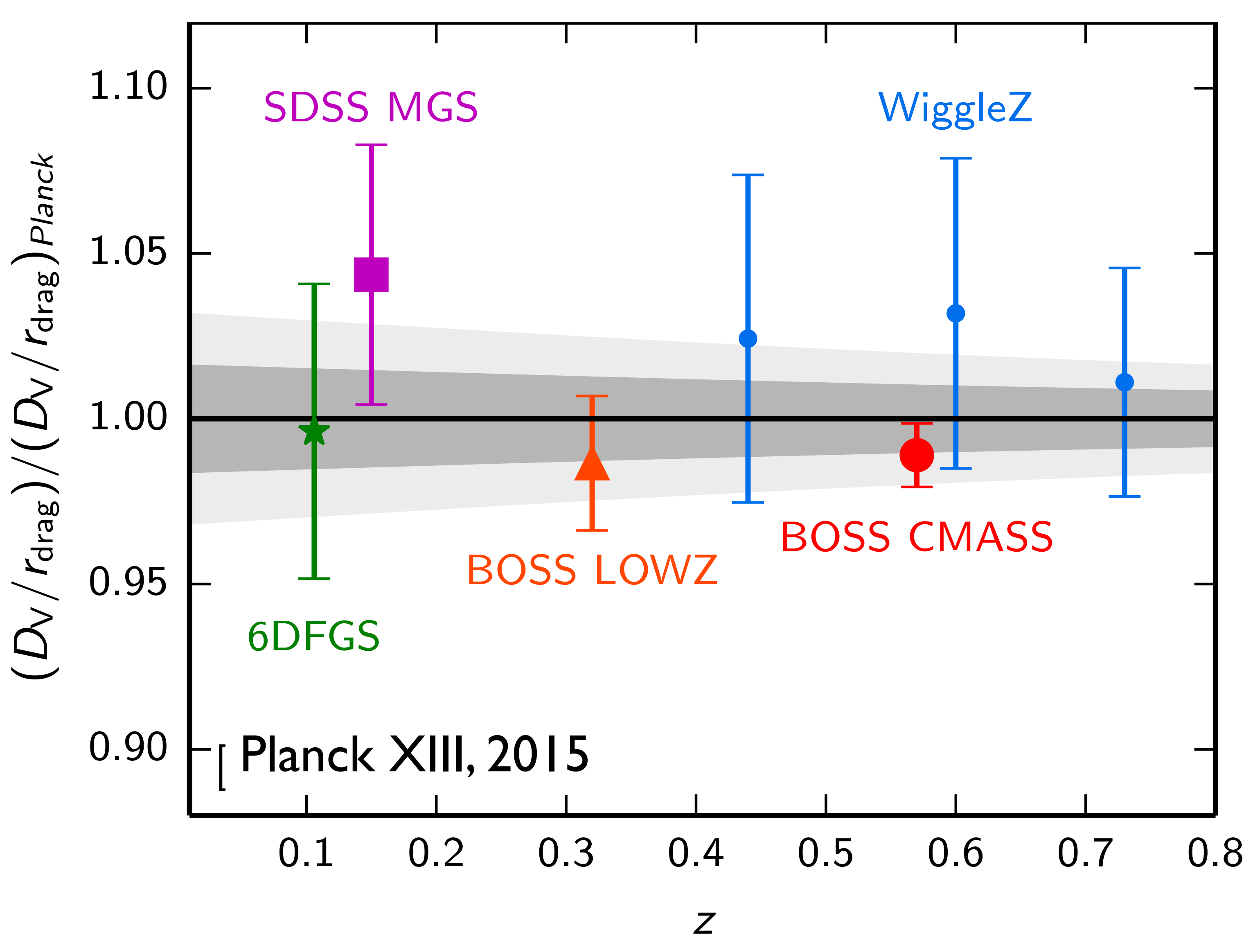
Testing Dark Energy I

Expansion history

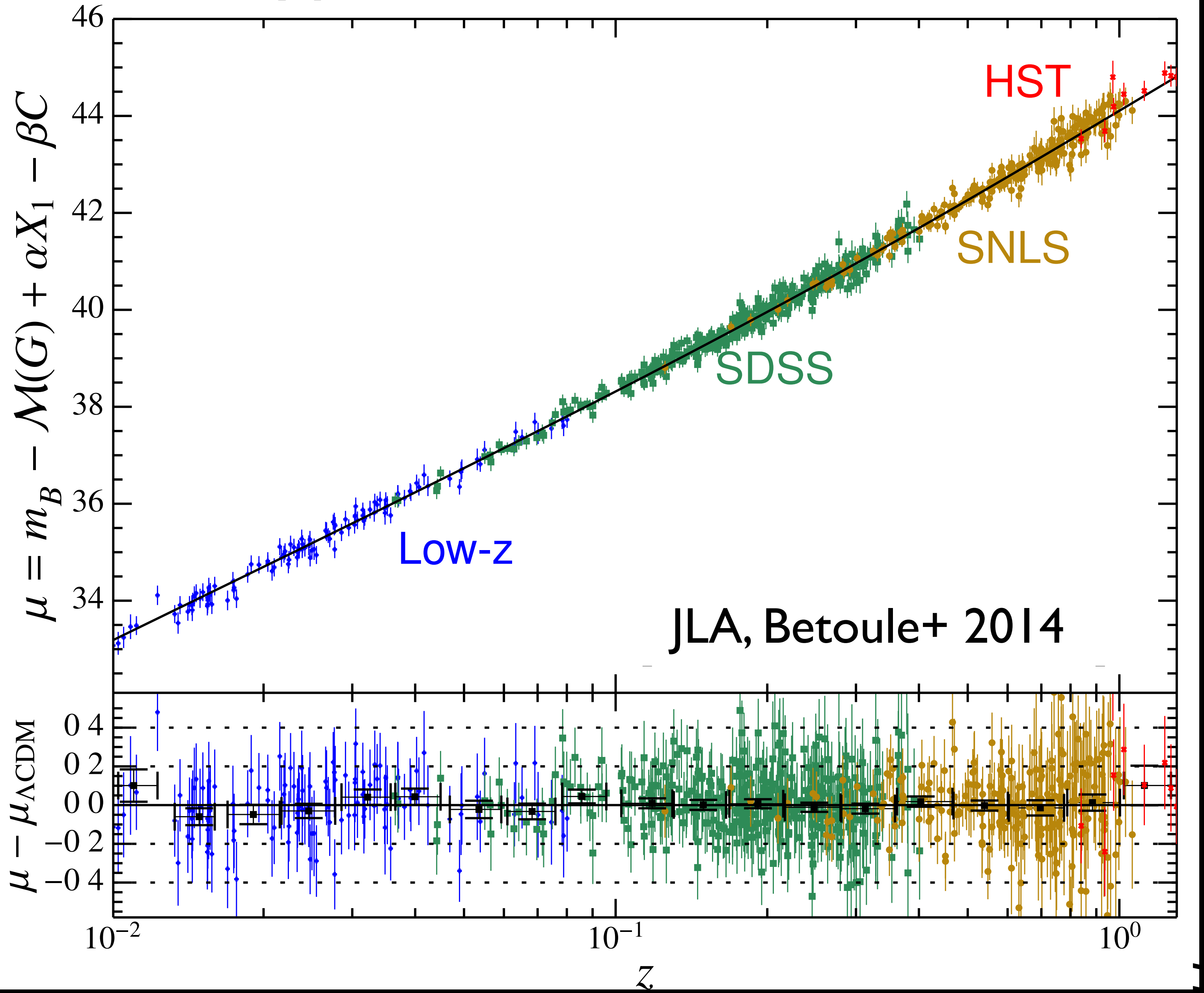
comparison of distance and redshift

- **standard rulers**: angle subtended by known scale
  - –CMB: sound horizon in early Universe
  - –BAO: same scale imprinted in late Universe
- **standard candle**: brightness of source with known luminosity
  - –supernovae
- **excellent agreement with ΛCDM**
- limited information on dark energy/modified gravity
  - –at most w0, wa

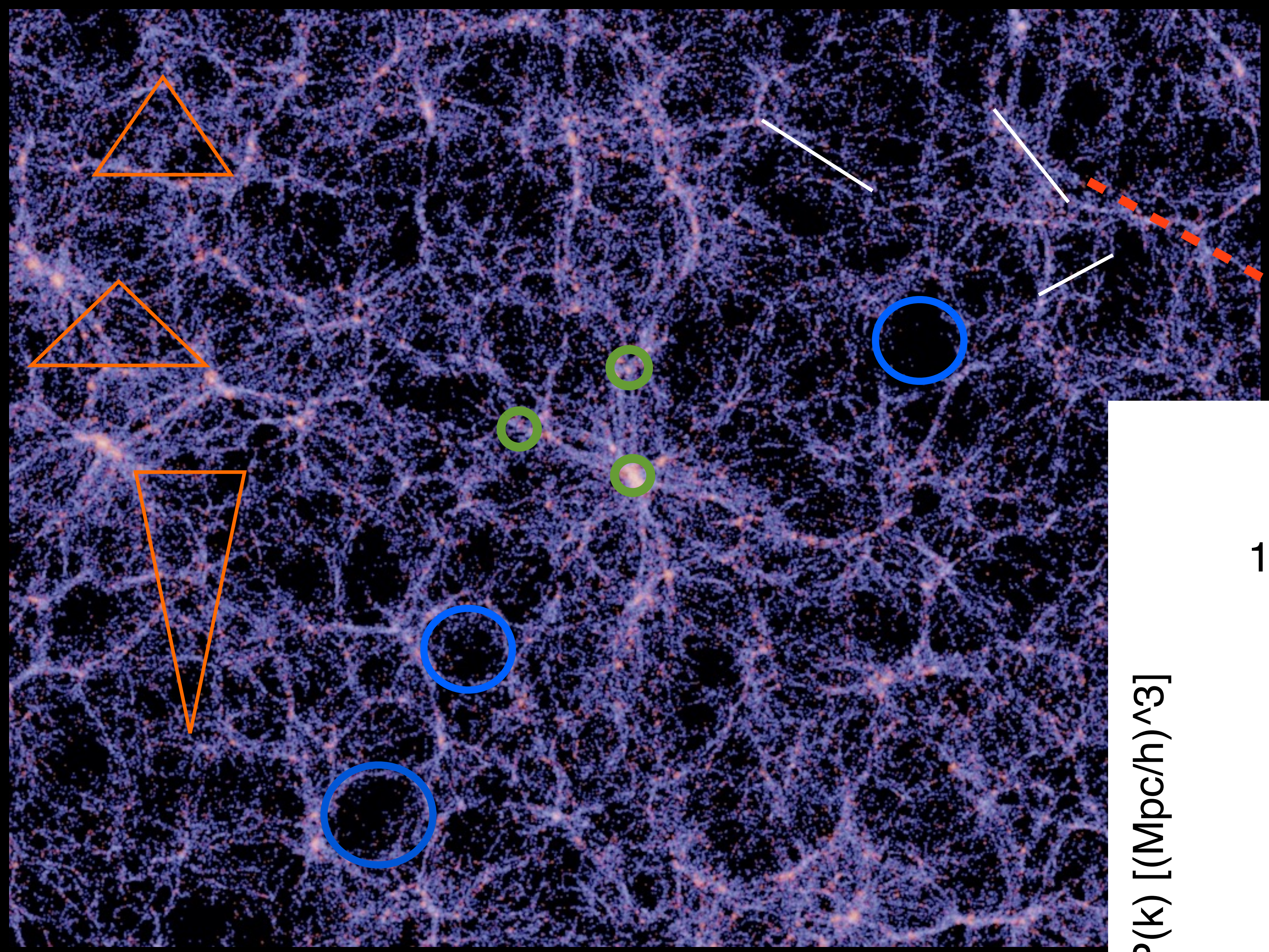
BAO state of the art



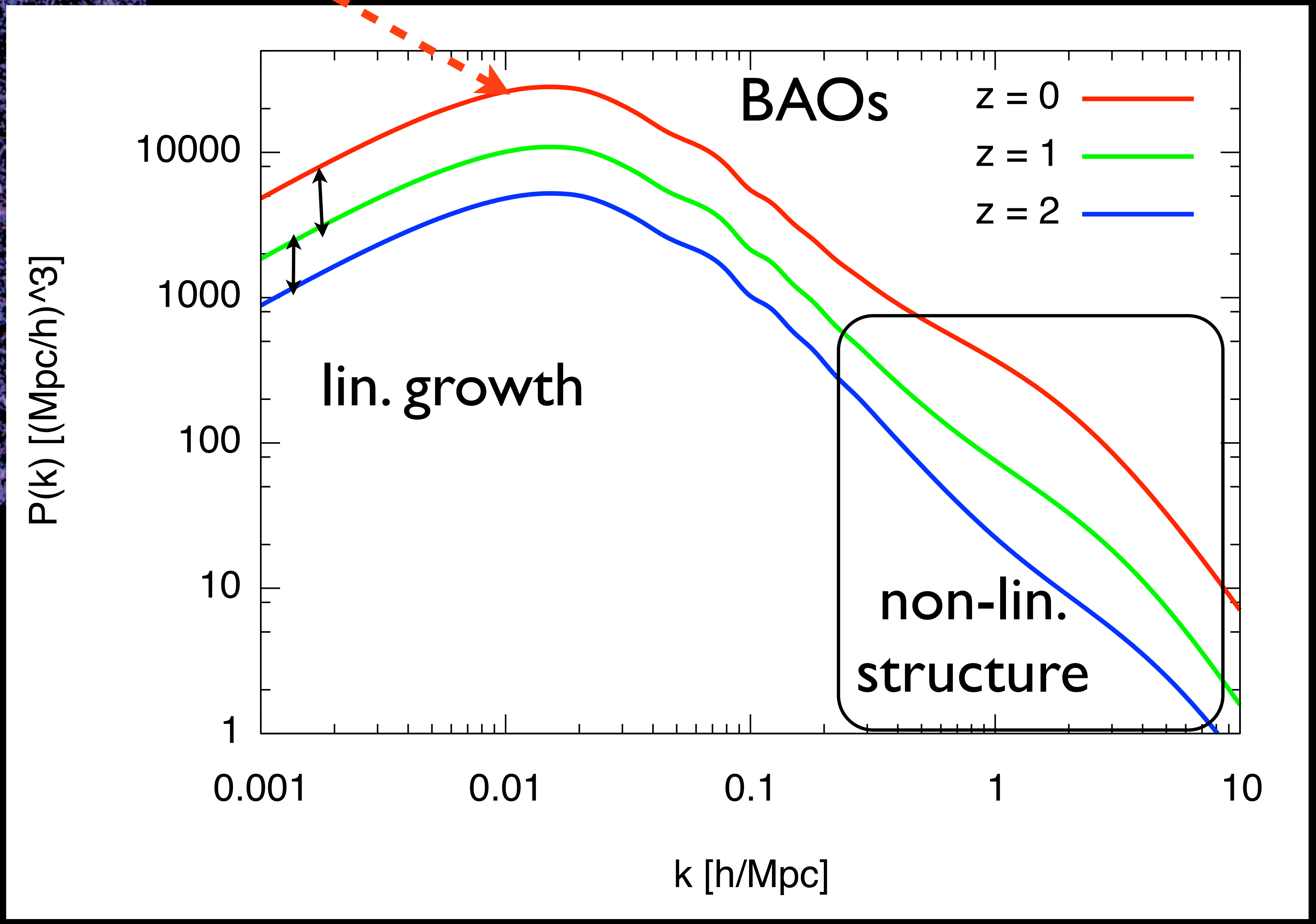
Type IA SN state of the art



Testing Dark Energy with Galaxies



need redshift, understand galaxy bias



- clusters (over densities),
- voids (under densities)
- two-point correlations
- three-point correlations,...

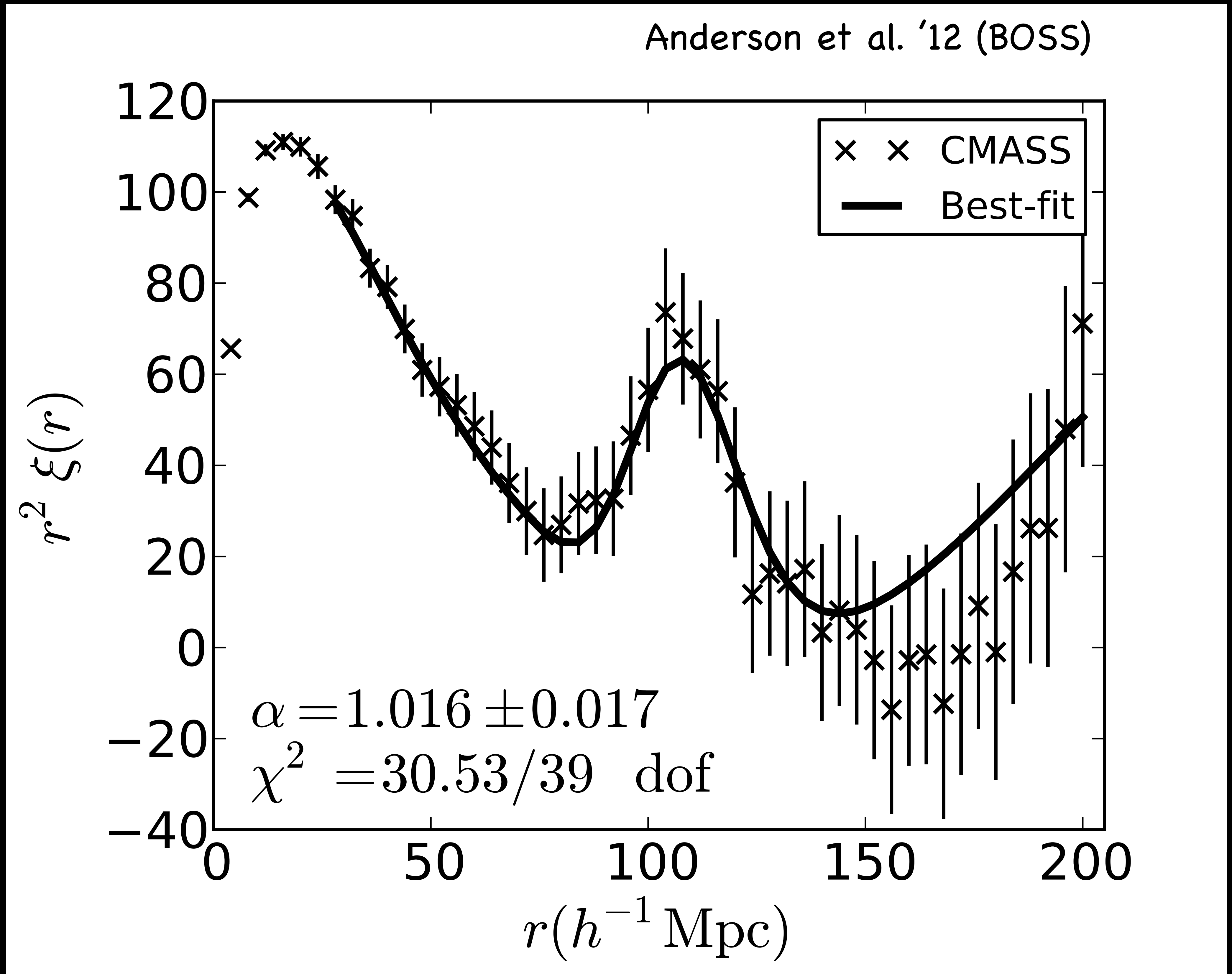
LSS Probes of Dark Energy

Galaxy Clustering

- measure BAOs + shape of correlation function

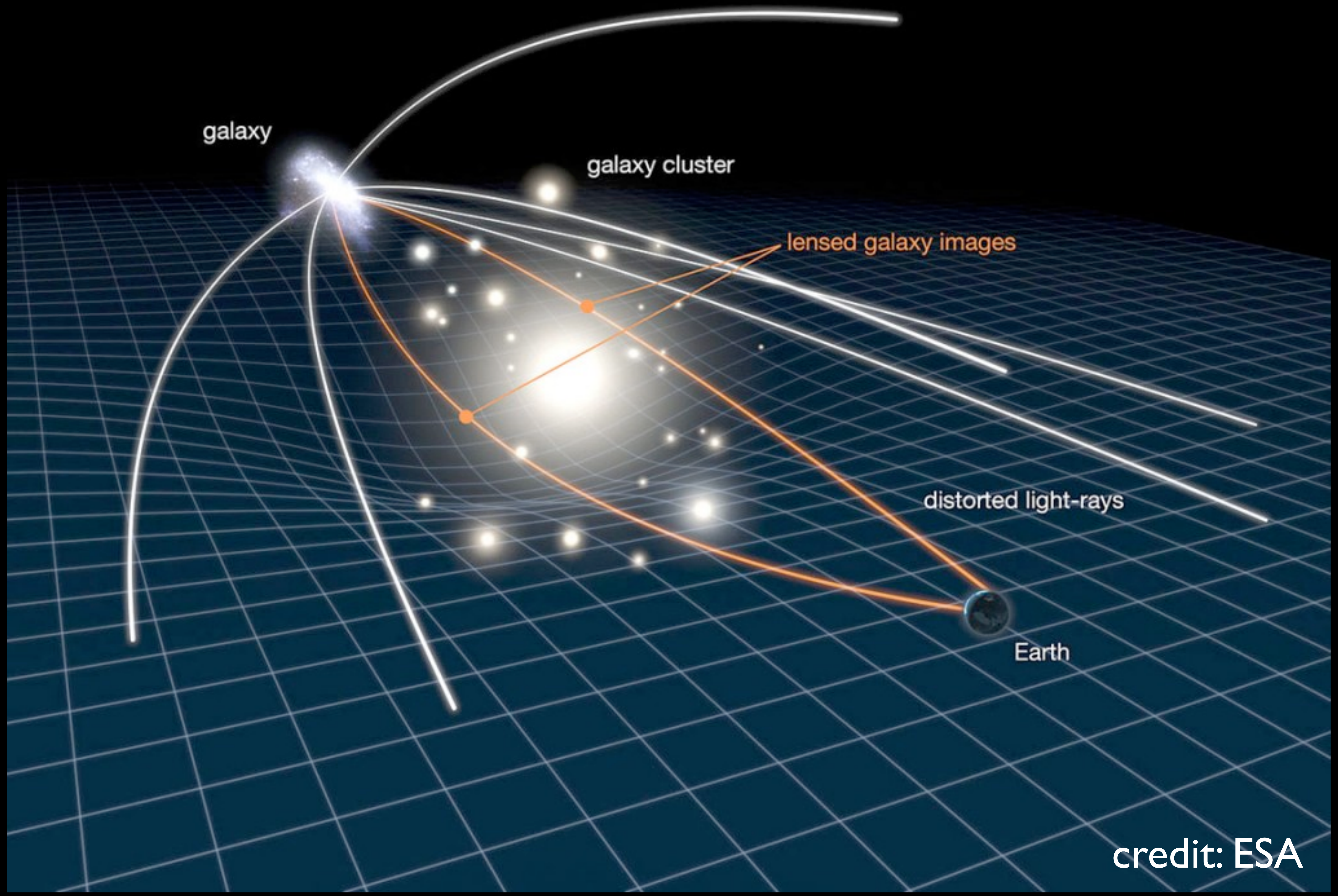
  → growth of structure, expansion history

- key systematic: **galaxy bias**



LSS Probes of Dark Energy

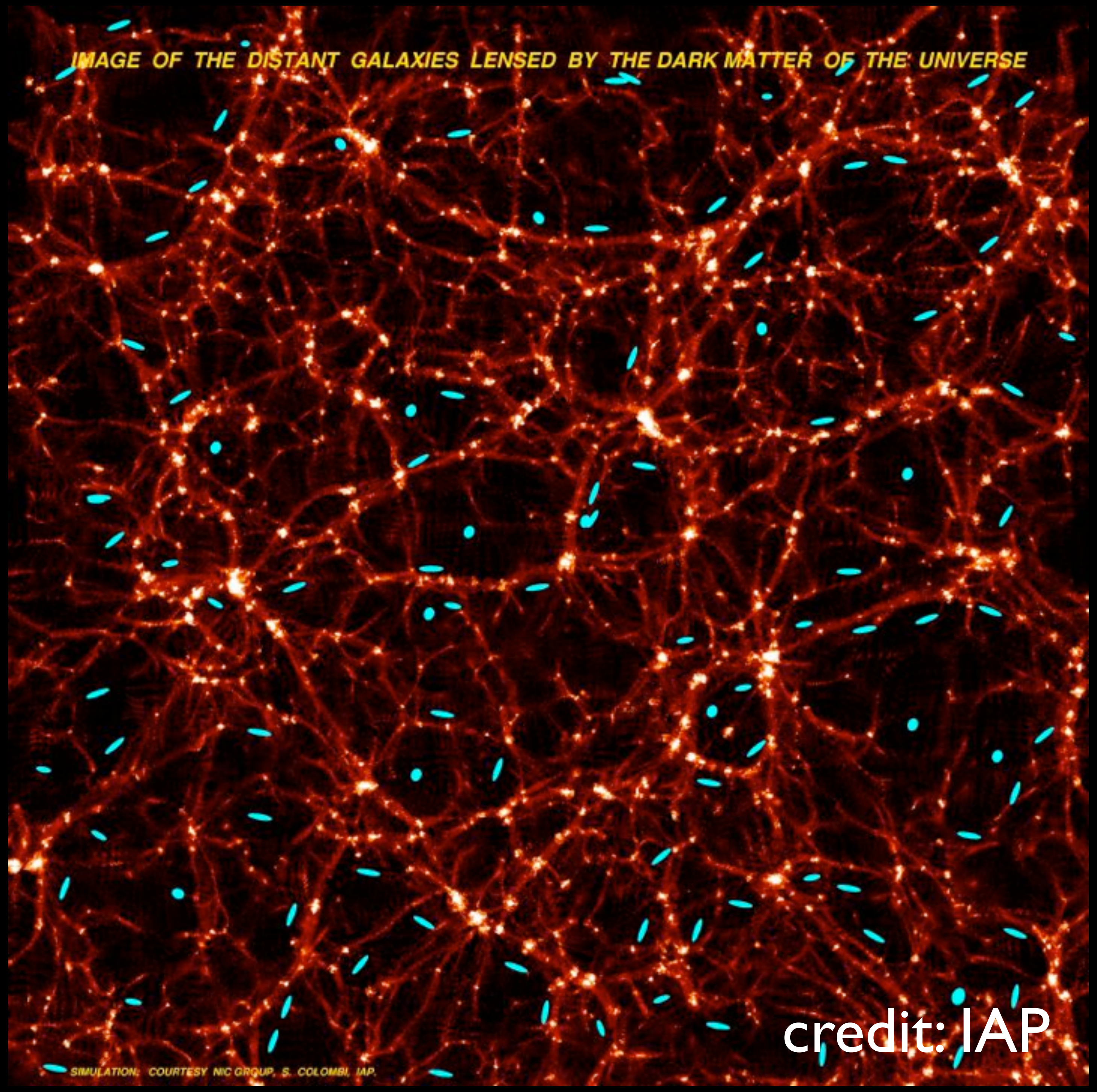
Weak Gravitational Lensing



LSS Probes of Dark Energy

Weak Gravitational Lensing

- light deflected by tidal field of LSS
  - coherent distortion of galaxy shapes "shear"
- shear related to projected matter distribution
- key systematics
  - shape measurements
  - assume random intrinsic orientation, average over many galaxies

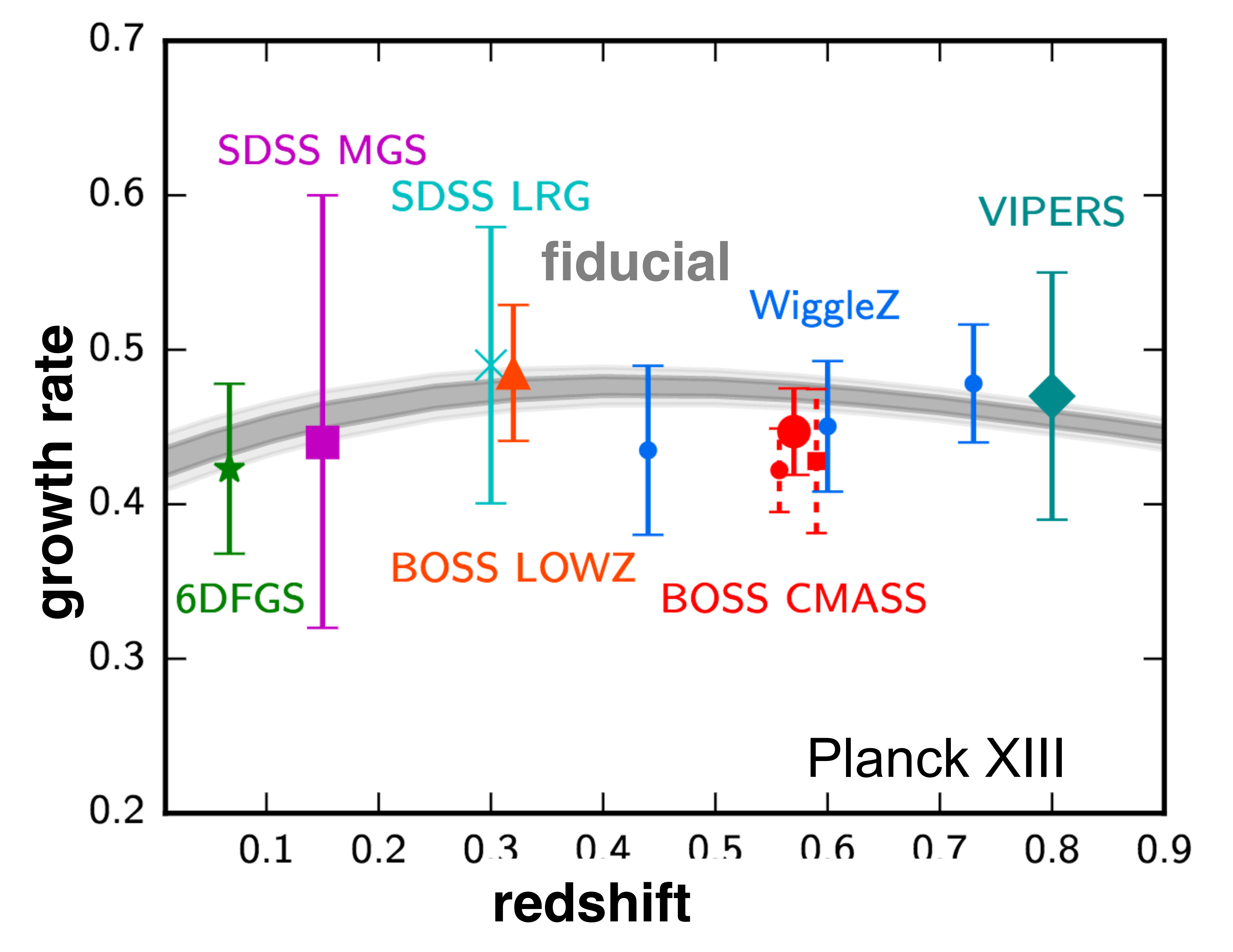


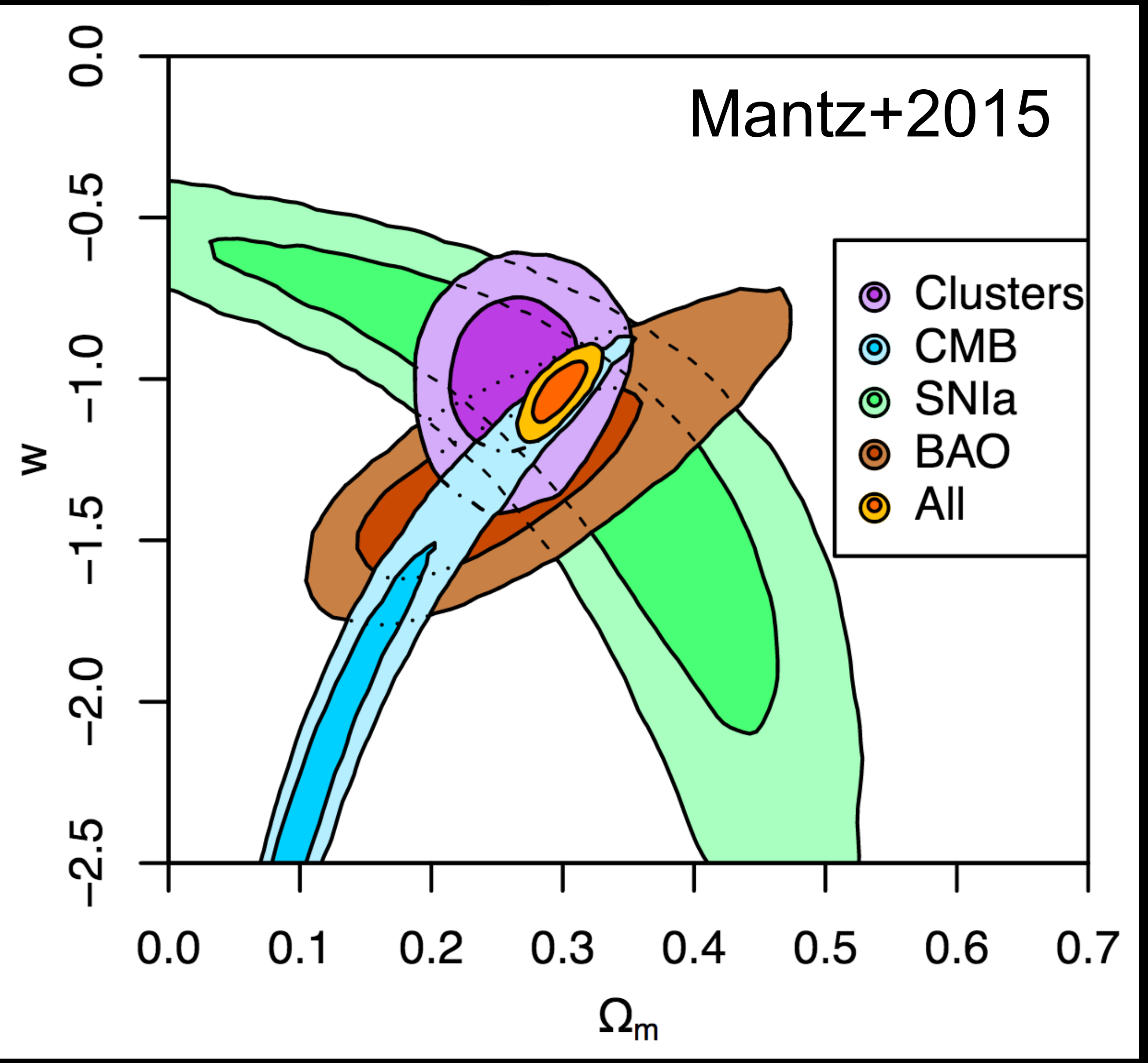
- measure **shear correlation function/power spectrum**
  - probes *total* matter power spectrum (w/ broad projection kernel)
- measure average (tangential) **shear around galaxies**
  - probes halo mass

Testing Dark Energy II

Late-Time Structure Growth

- redshift space distortions ✓
- galaxy clusters (✓)
  counts as functions of mass and redshift





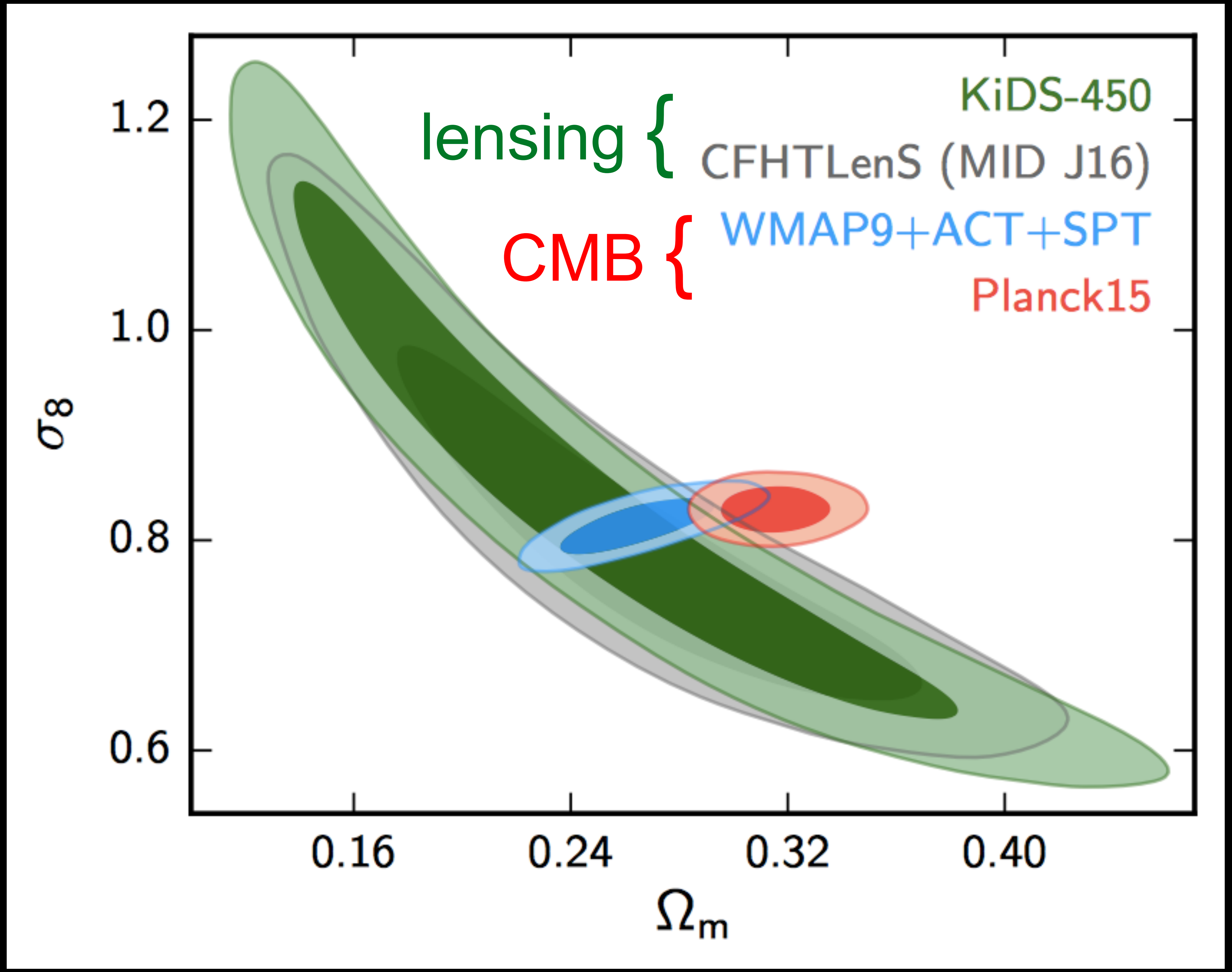
Testing Dark Energy II

Late-time Structure Growth

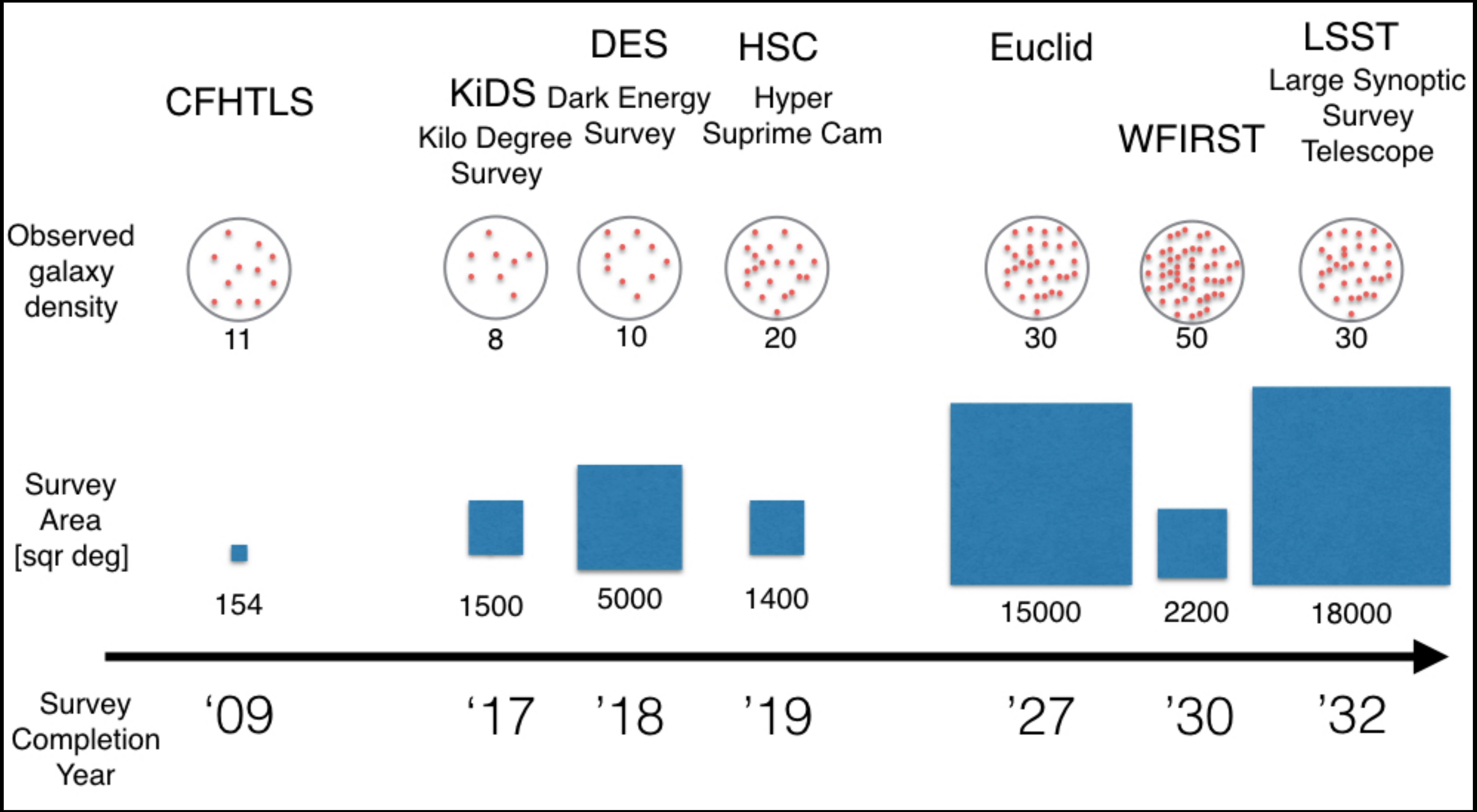
- redshift space distortions ✓
- galaxy clusters (✓)

  counts as functions of mass and redshift
- weak lensing

  recent studies have claimed 2-3σ tension with Planck
  - a fluke/non-issue?
  - a crack in LCDM?
  - a systematic error?



Photometric Dark Energy Surveys



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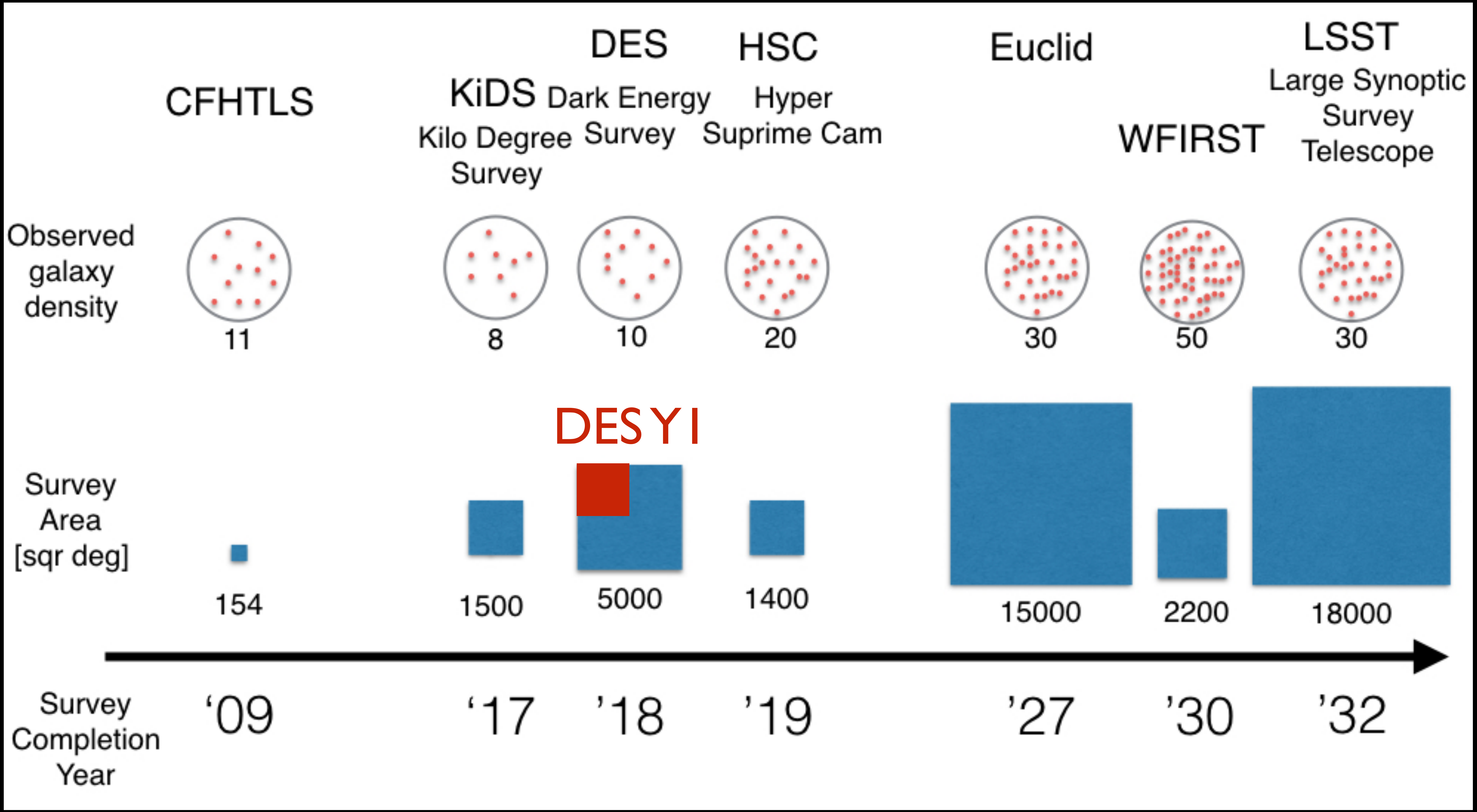
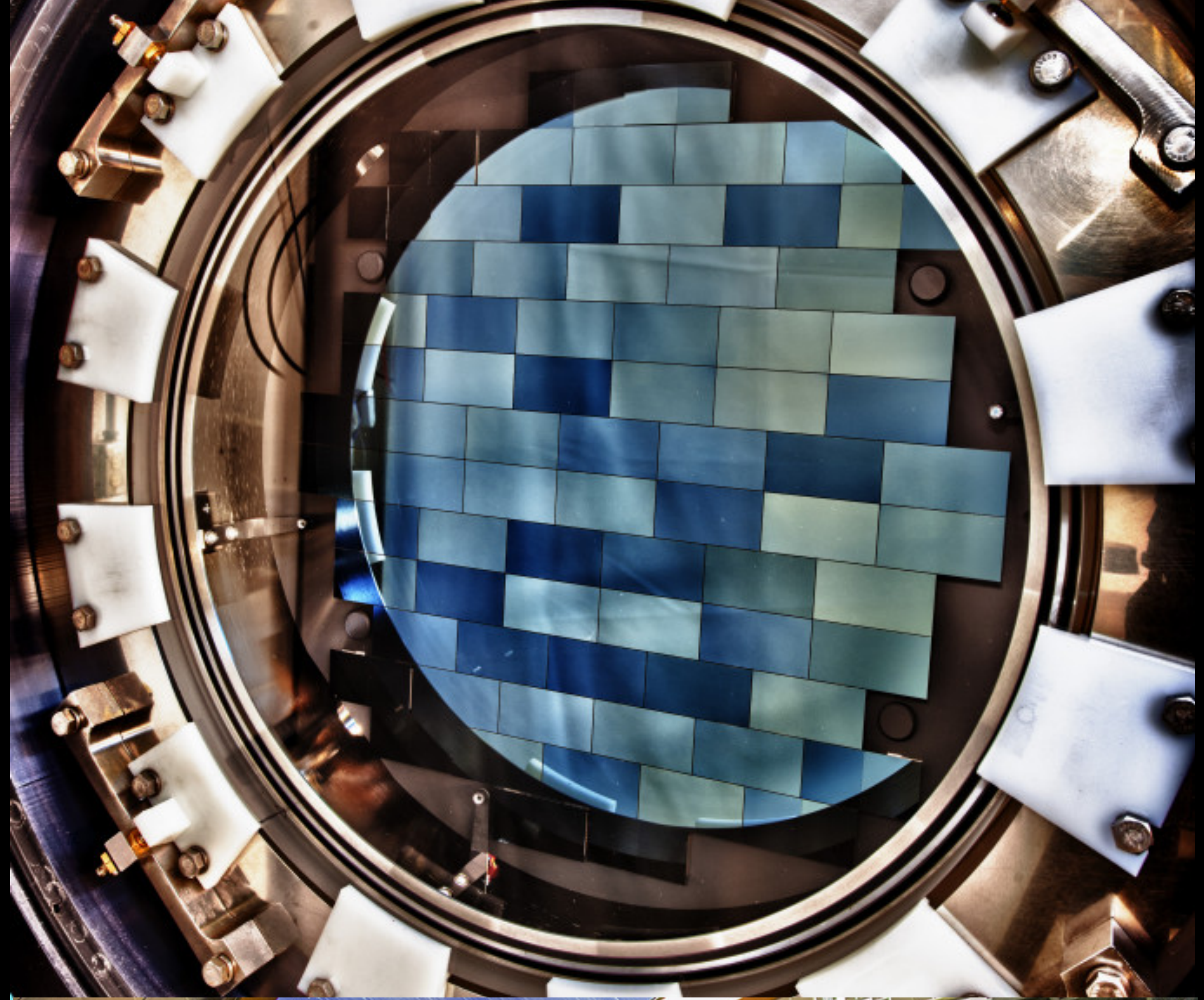
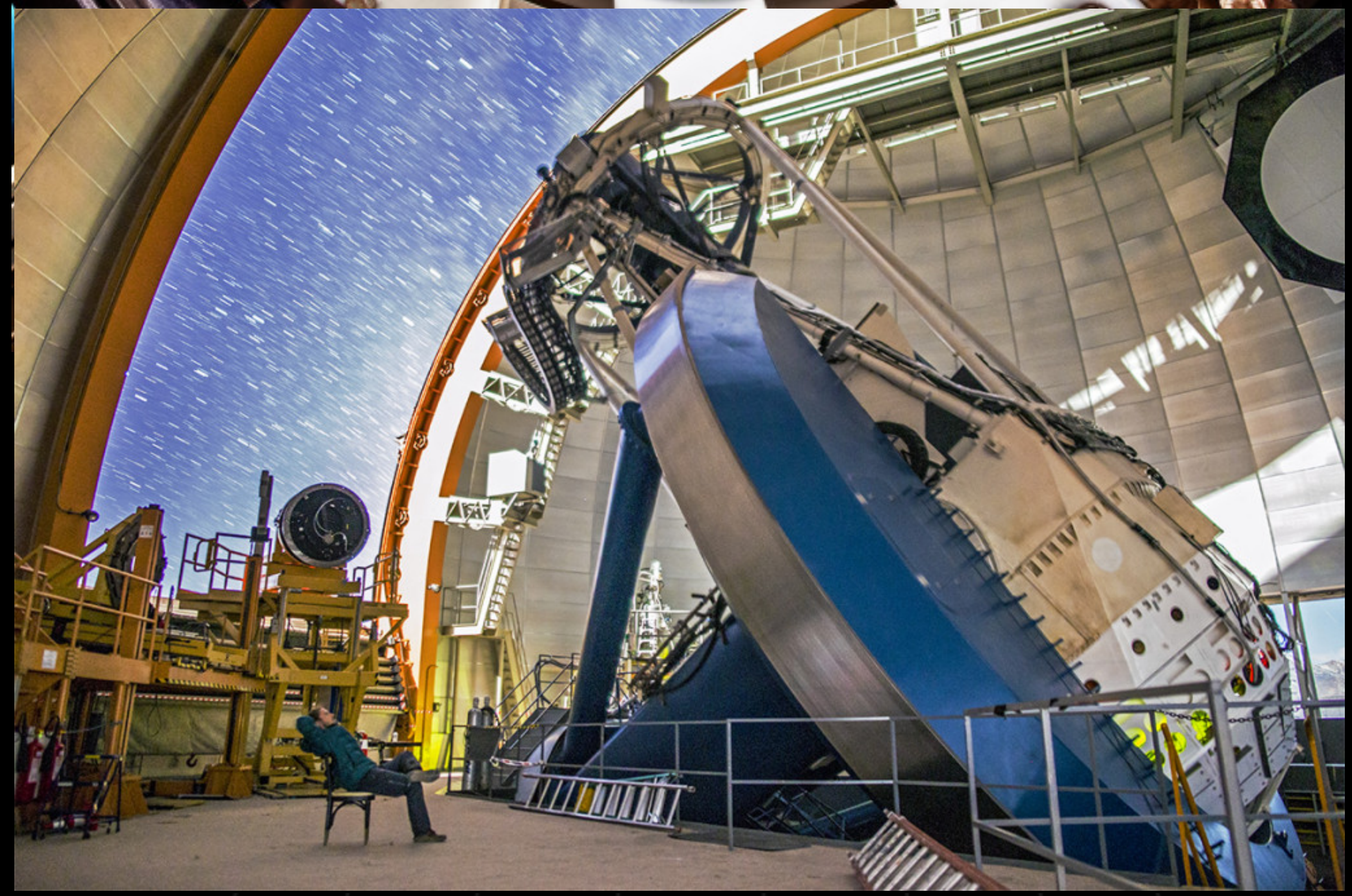
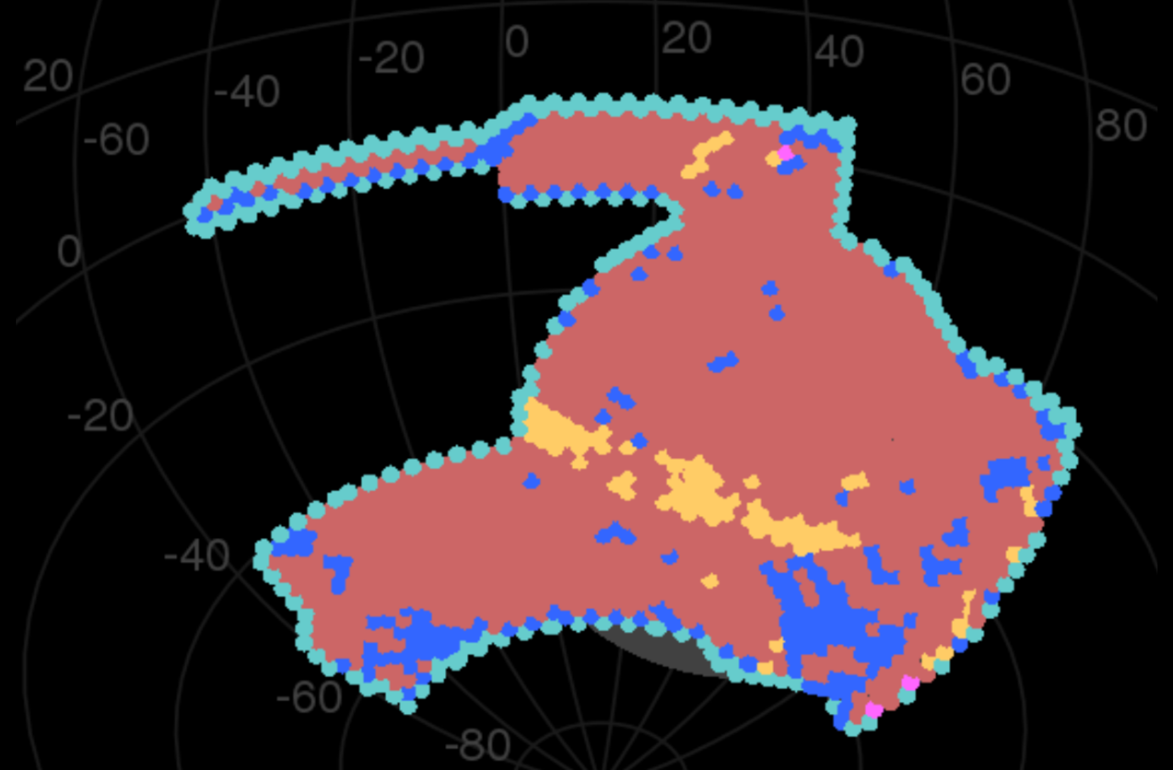


The Dark Energy Survey

- 5000 sq. deg. survey in grizY from Blanco @ CTIO
  - 10 exposures, 5 years, >400 scientists
- Primary goal: dark energy equation of state
- Probes: Galaxy Clustering, Supernovae,
  Cluster counts, Gravitational lensing
- Status:
  - SV (150 sq. deg, full depth):
    most science done, catalogs public
  - Y1 (1500 sq. deg, 40% depth):
    data processed, results on cosmology last week
  - Y3 (5000 sq. deg, 50% depth):
    data processed, vetting catalogs
  - Y4: data taking finished (70% depth)

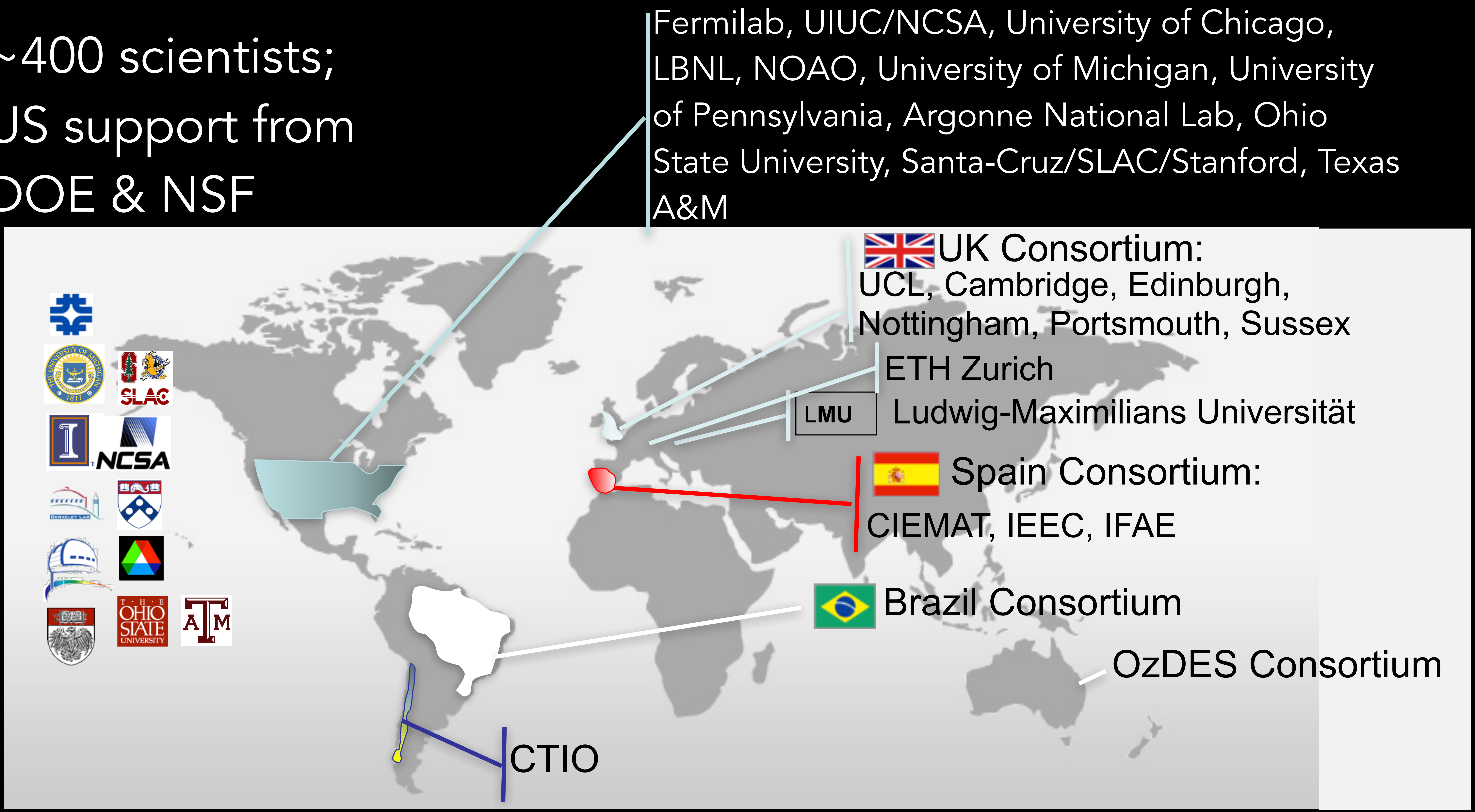




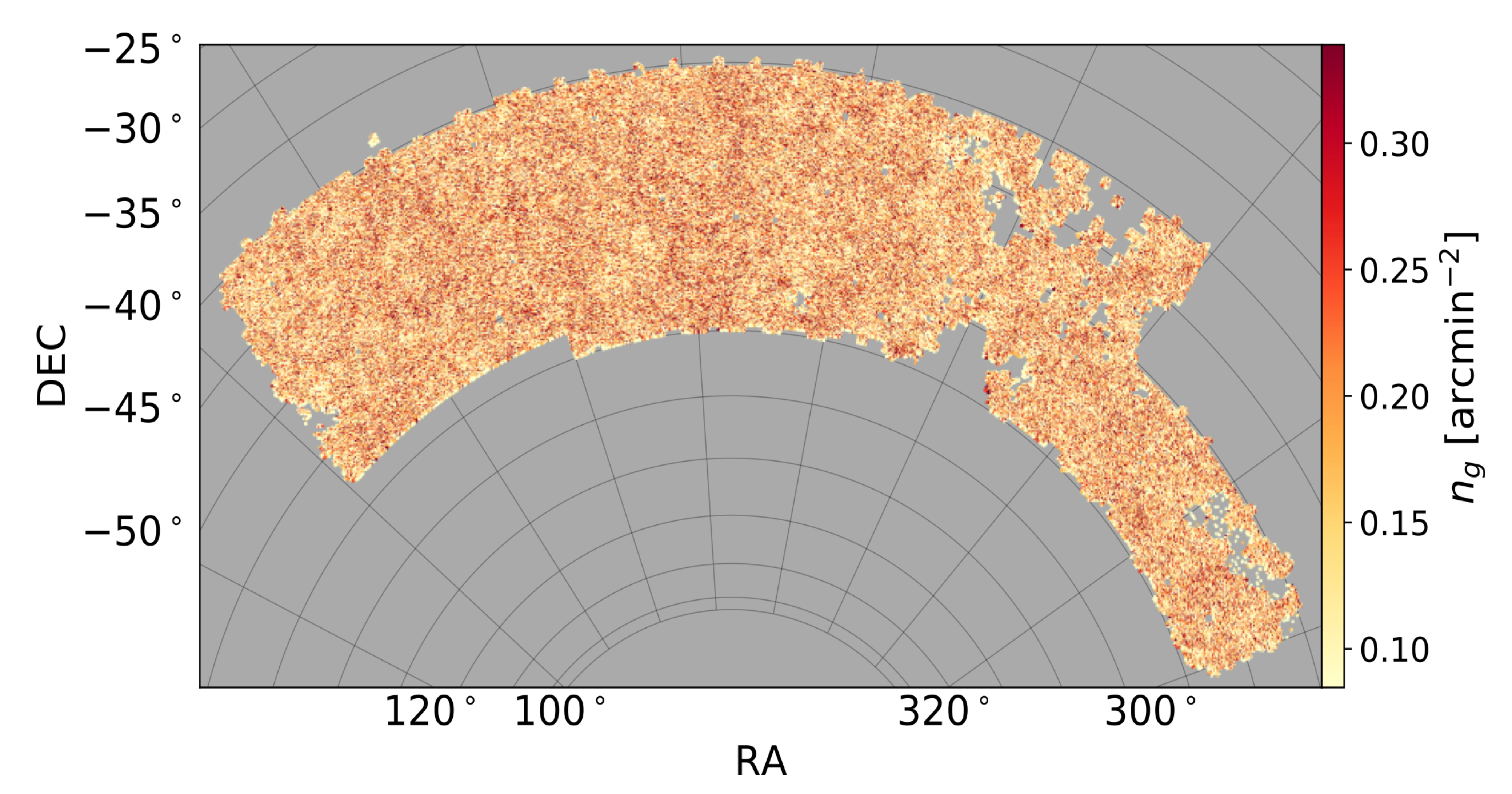


The Dark Energy Survey Collaboration

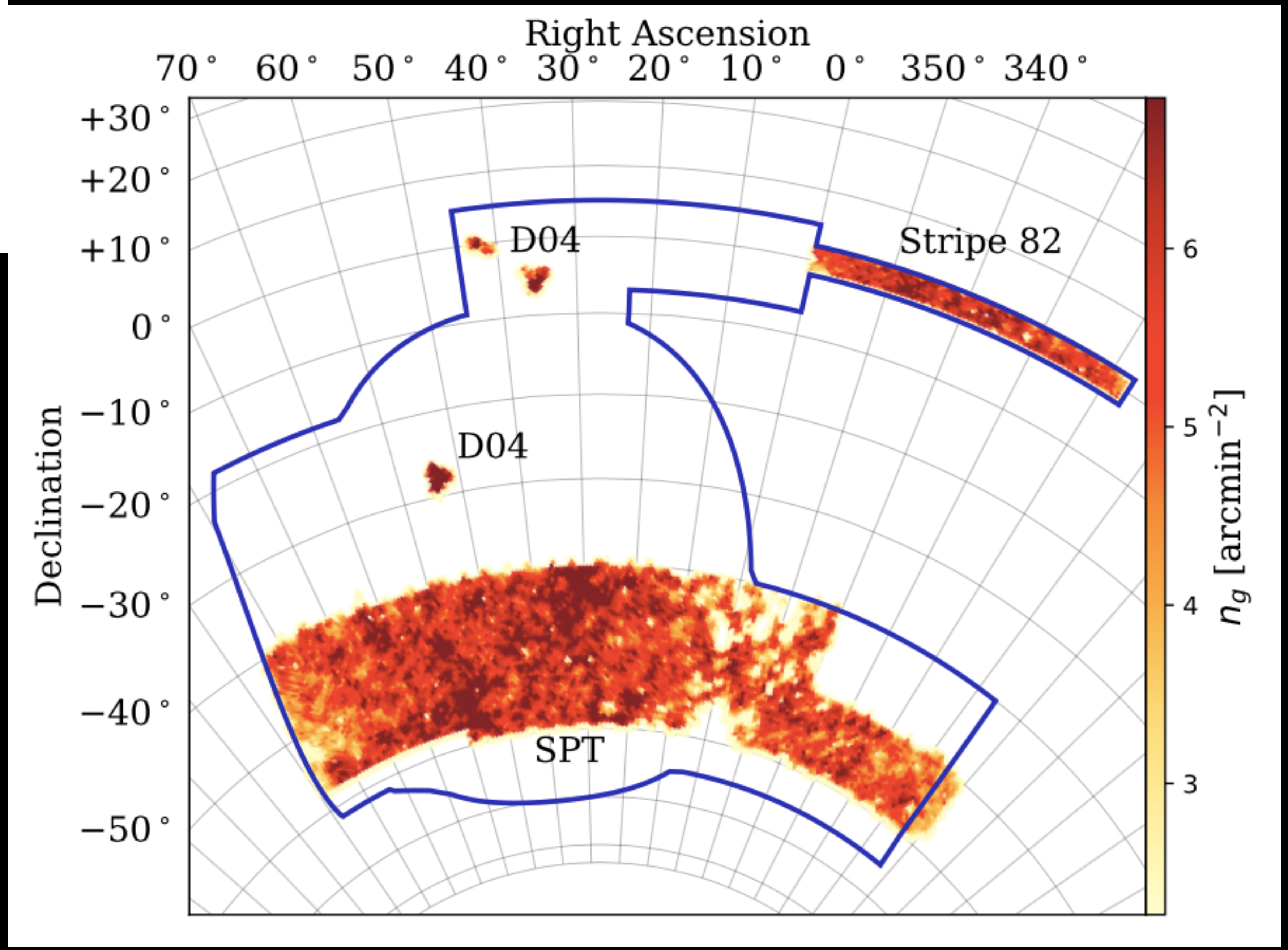
~400 scientists;
US support from
DOE & NSF



DES Year 1 Galaxy Samples



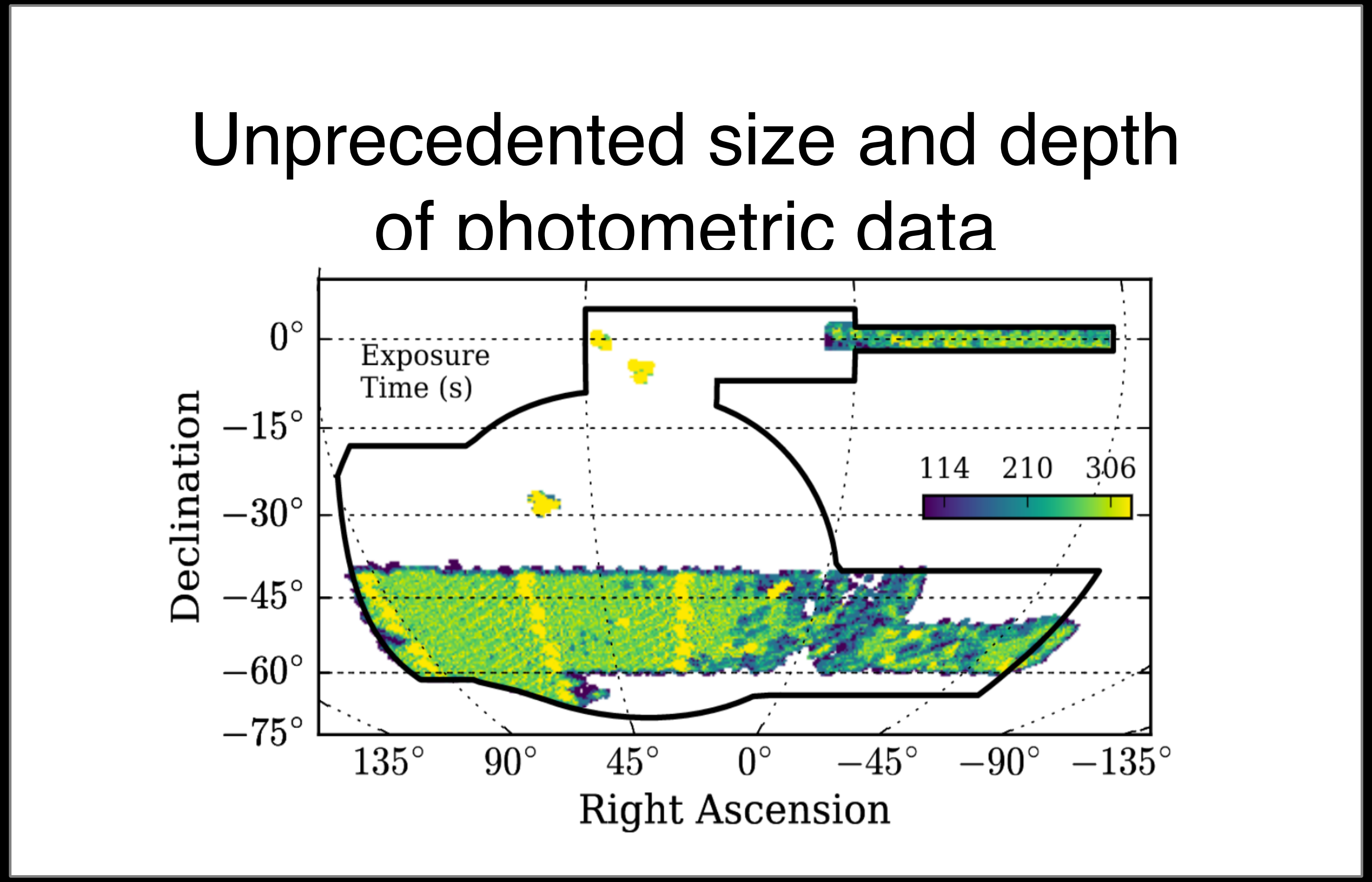
- 26 million source galaxies
- 4 redshift bins
- Sources for cosmic shear & galaxy-galaxy lensing



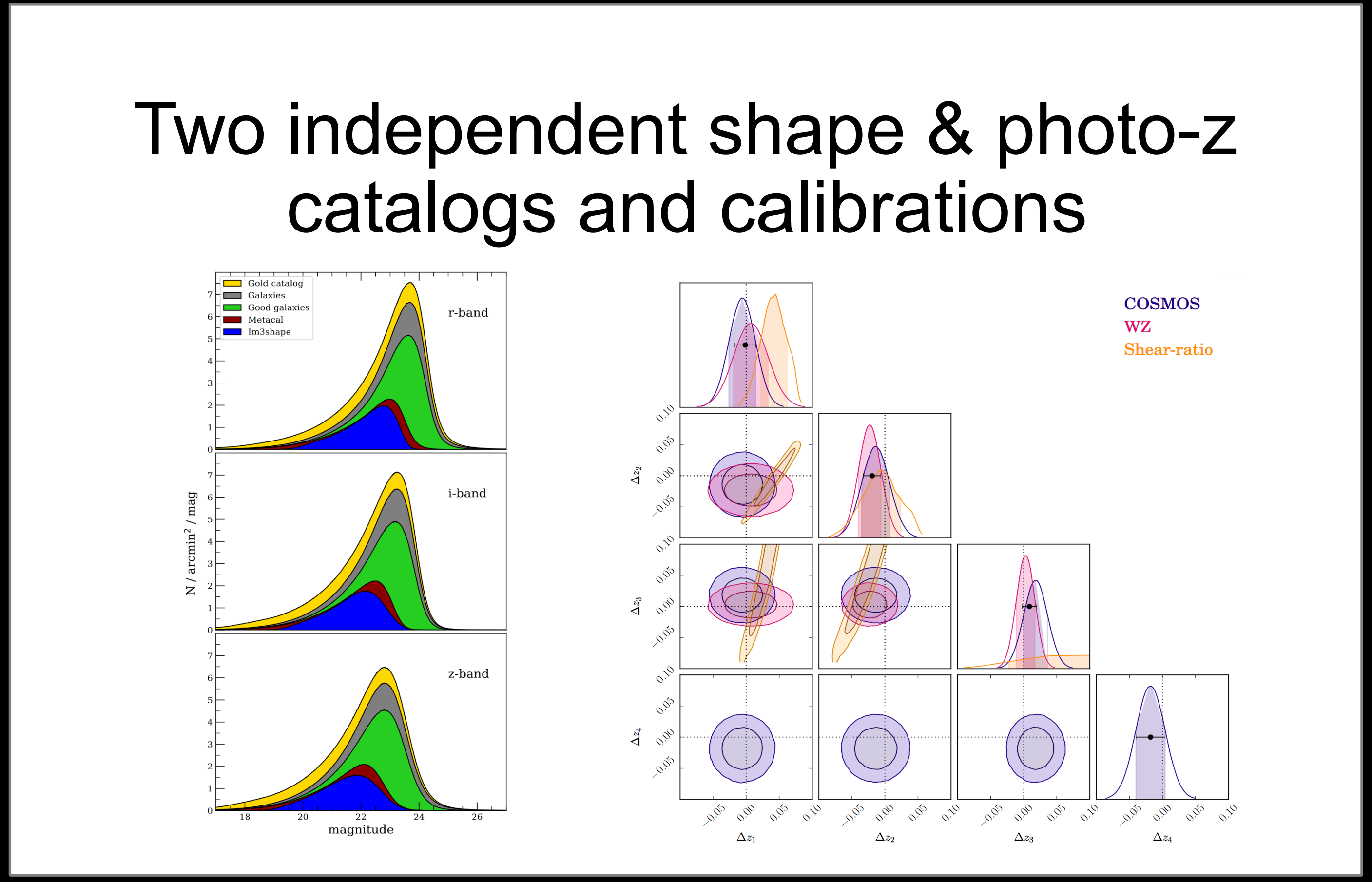
- 660,000 redMaGiC galaxies with excellent photo-z's
- Measure angular clustering in 5 redshift bins
- Use as lenses for galaxy-galaxy lensing

First Year of Data: ~1800 sq. deg. Analyzed 1321 s.d. after cuts

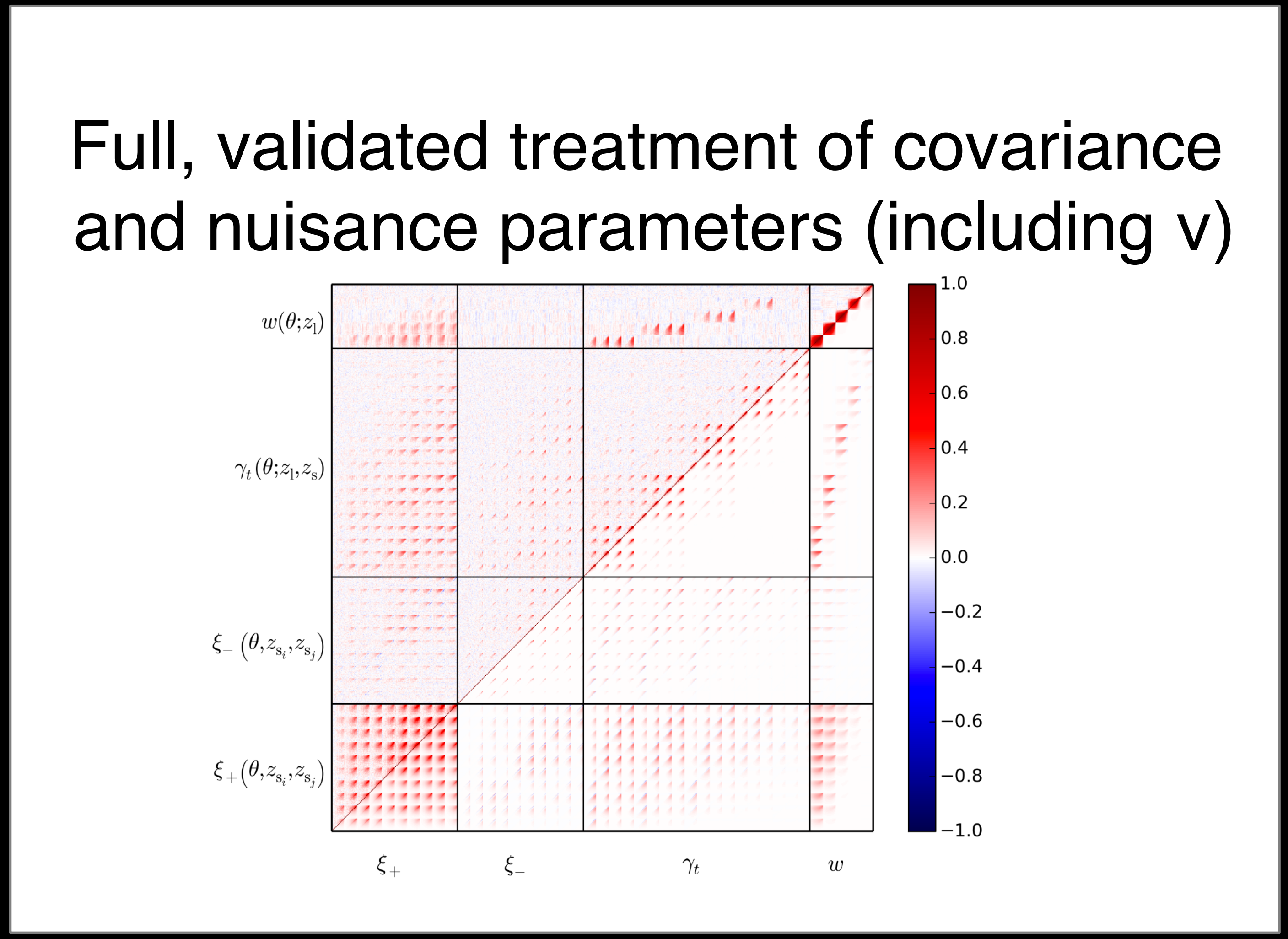
With great statistical power comes great systematic responsibility

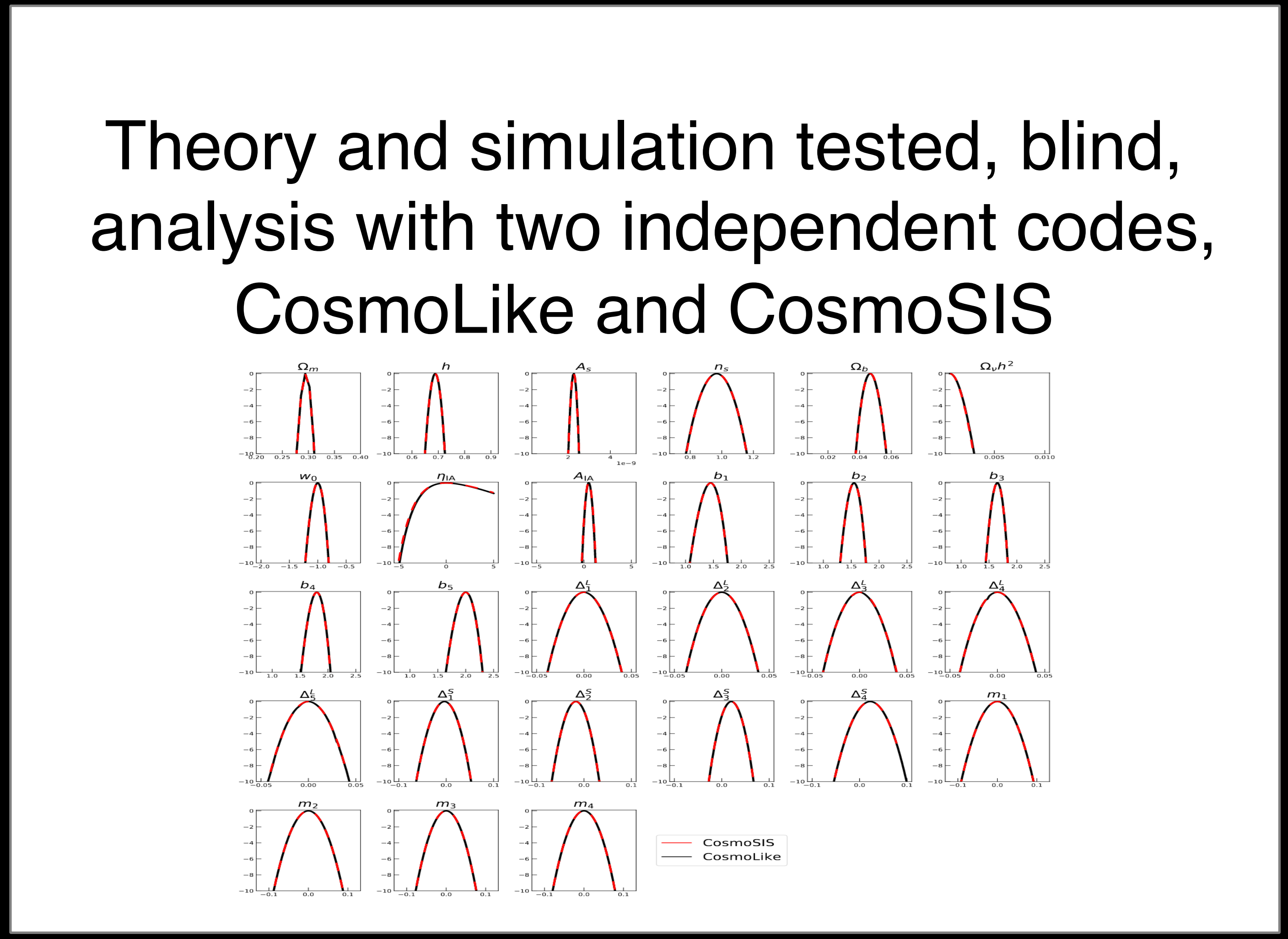


Drlica-Wagner, Rykoff, Sevilla+ 2017



Zuntz, Sheldon+; Samuroff+; Hoyle, Gruen+ 2017;
Davis+, Gatti, Vielzeuf+, Cawthon+ 2017





Krause, Eifler+2017; MacCrann, DeRose+ 2018

DES Y1 Shear Catalogs

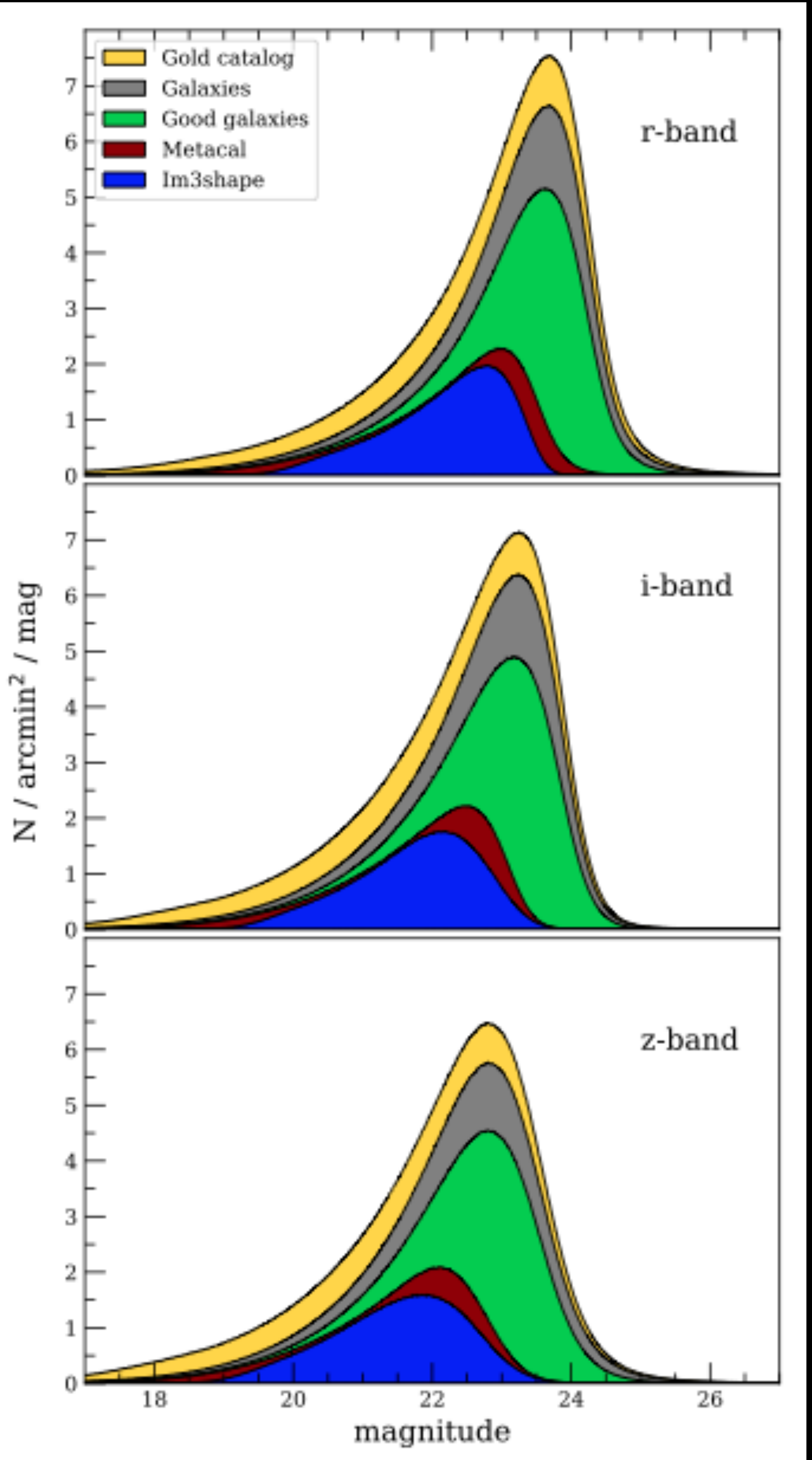
(Zuntz+17)

Metacalibration *(Huff+17, Sheldon+17)*:

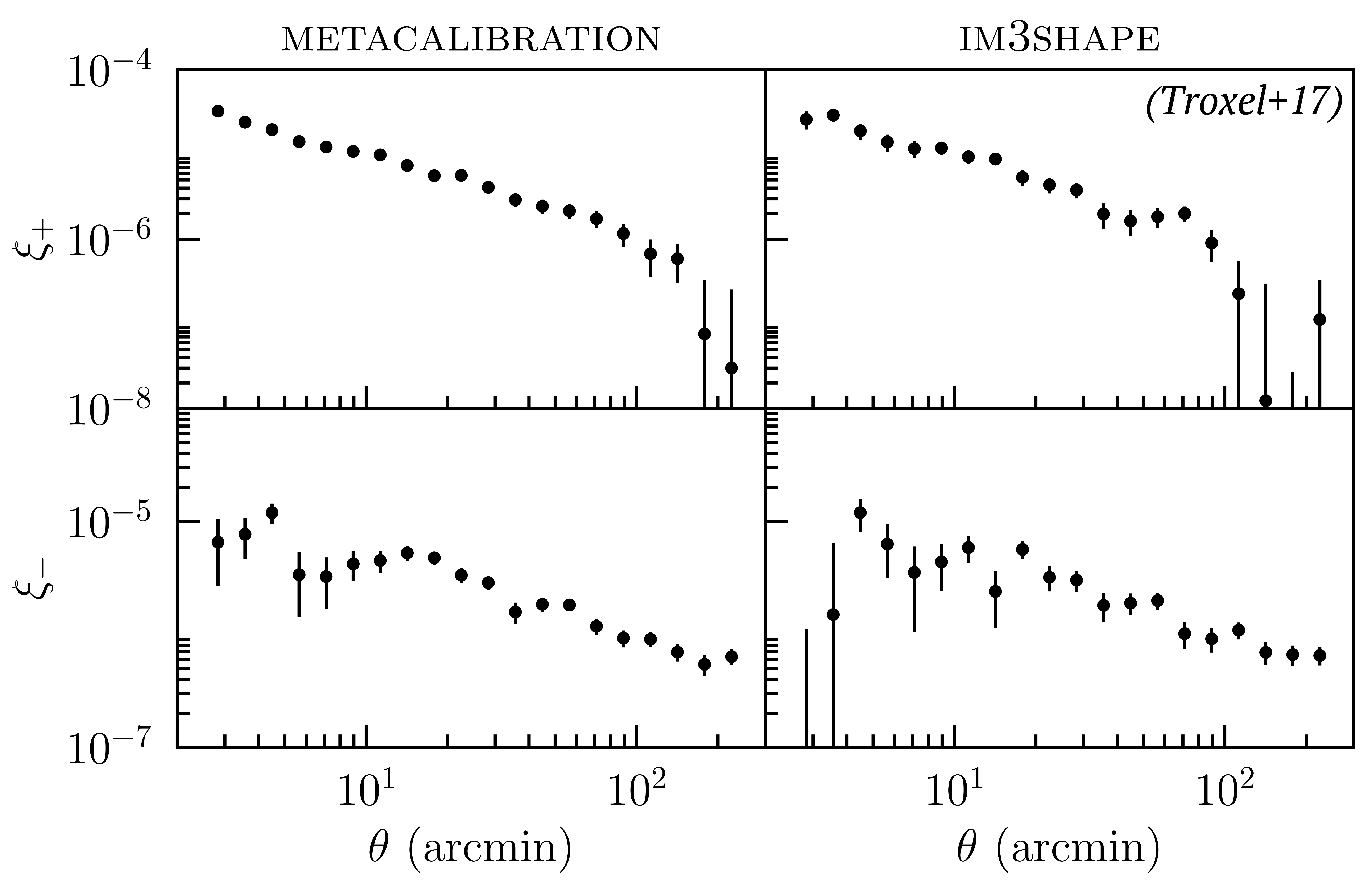
- New method measuring estimator shear response internally by deconvolving, shearing, reconvolving.
- It uses g, r, i bands.
- 35 M galaxies (26 M for cosmology).

im3shape:

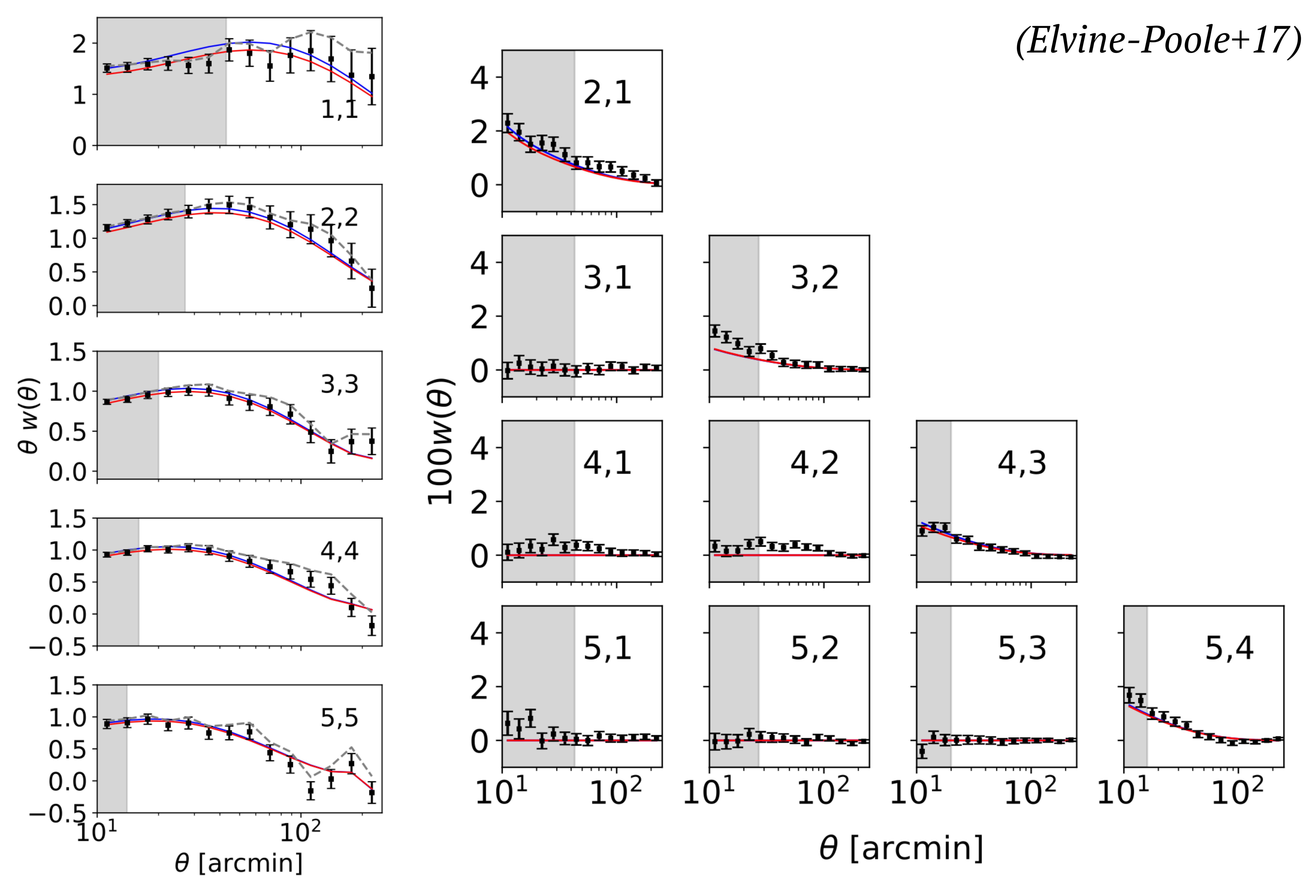
- Best-fit bulge & disc models, calibrated with simulations.
- Only r-band.
- 22 M galaxies (18 M for cosmology).



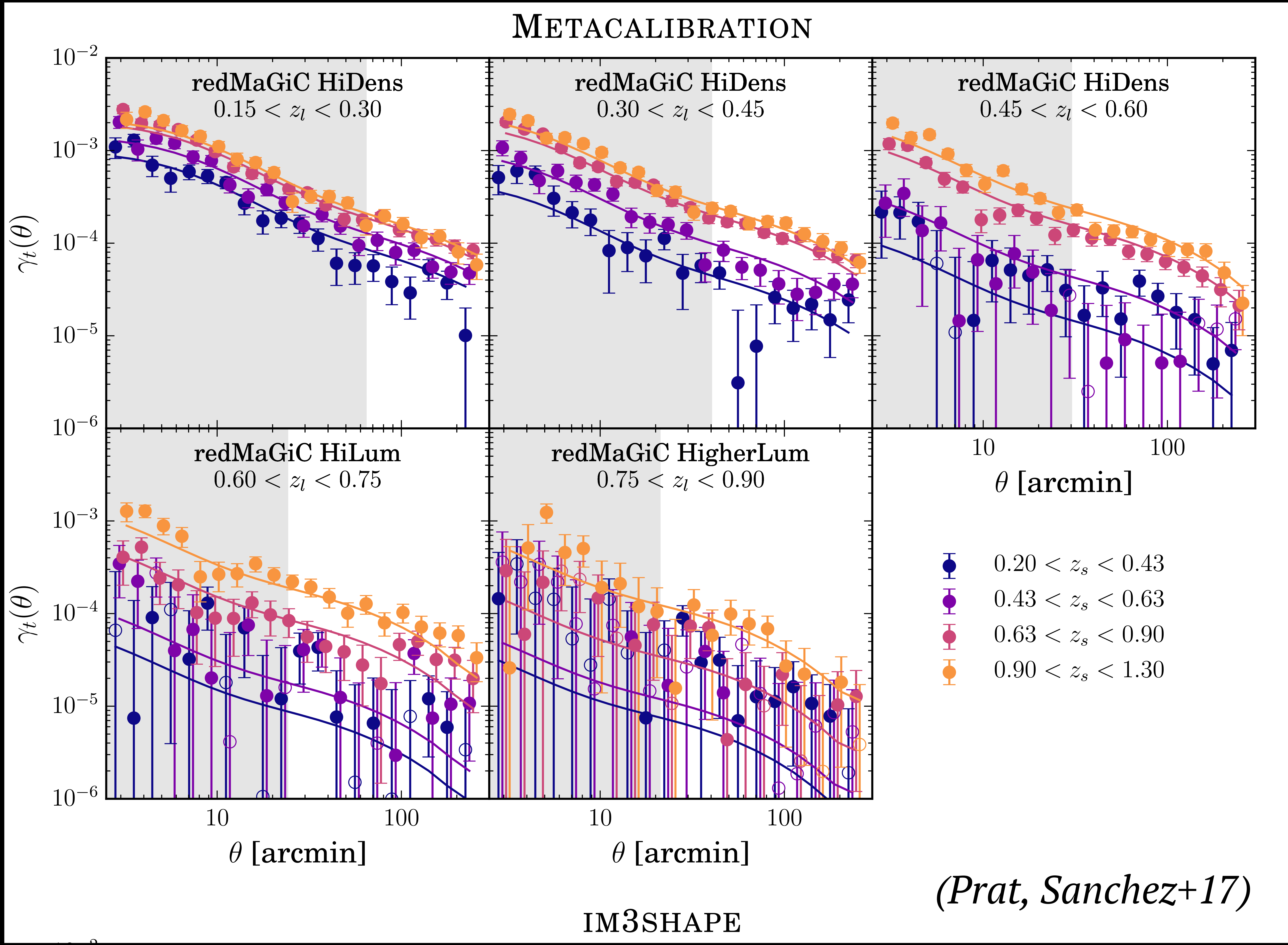
DESY1 Measurements: Cosmic Shear



DESY1 Measurements: Galaxy Clustering



DESY1 Measurements: Galaxy-Galaxy Lensing



Multi-Probe Methodology

(Krause, Eifler+17)

from data vector D to parameters p

$$L(\mathbf{D}|\mathbf{p}) \propto \exp\left(-\frac{1}{2}\left[(\mathbf{D}-\mathbf{M}(\mathbf{p}))^T \mathbf{C}^{-1}(\mathbf{D}-\mathbf{M}(\mathbf{p}))\right]\right)$$

- model data vector, incl. relevant systematics
  - implementation details should not contribute to error budget
  - are the systematics parameterizations sufficient for DES-Y1?
- covariance for ~450 data points
- sampler - don't get the last step wrong...

methods paper: validate model + implementation, covariance, sampling

Multi-Probe Blinding

(DES Collaboration 17)

Goal: minimize confirmation bias

Implementation: two-staged blinding process

- shear catalogs scaled by unknown factor, until catalogs fixed
- cosmo params shifted by unknown vector, until full analysis fixed
- (do not overplot measurement + theory)
- (clearly state any post-unblinding changes in paper)

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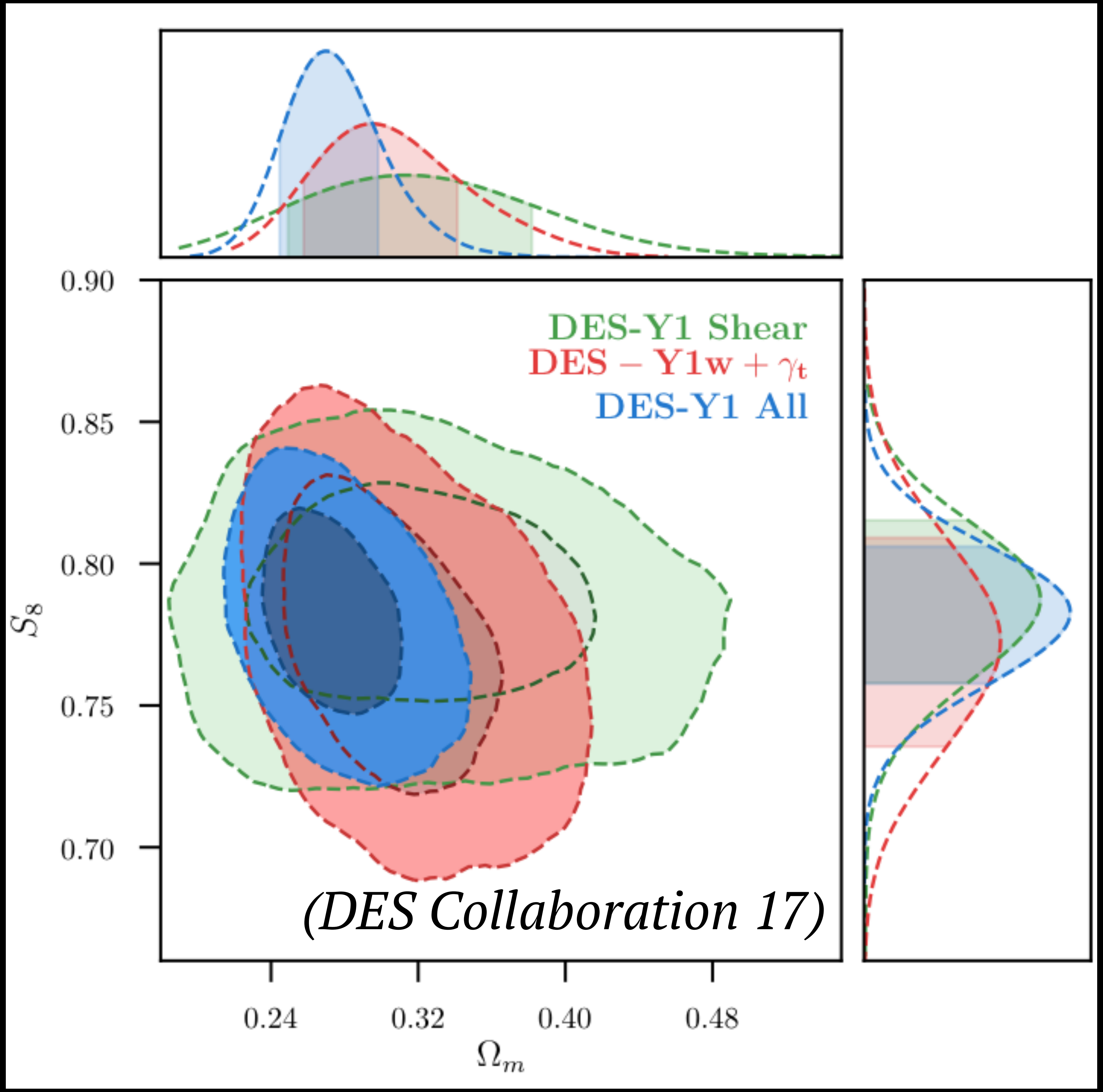
Lessons

- clearly define scope of blinding
    - e.g., parameter measurements vs. model testing
- make sure blinding scheme allows null tests
    - for parameter measurements, this may include consistency between probe
- *someone not knowing what they're doing, shouldn't be able to unblind intentionally; someone knowing what they're doing, shouldn't be able to unblind unintentionally*

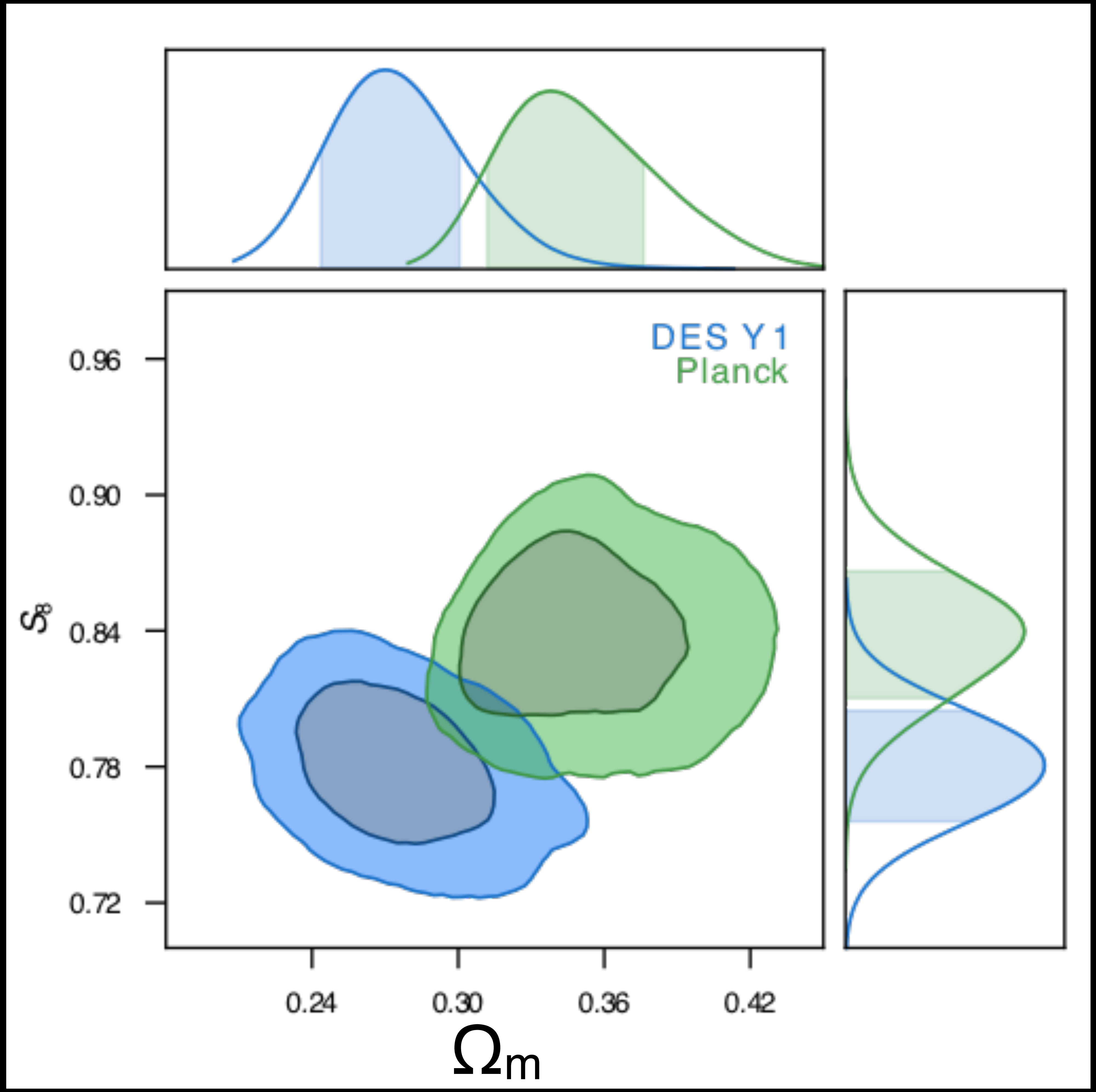
Multi-Probe Constraints: LCDM



- DES-Y1 most stringent constraints from weak lensing
- marginalized 4 cosmology parameters, 10 clustering nuisance parameters, and 10 lensing nuisance parameters
- consistent (R = 583) cosmology constraints from weak lensing and clustering in configuration space

Key Result: Consistency of late Universe with Planck in ΛCDM

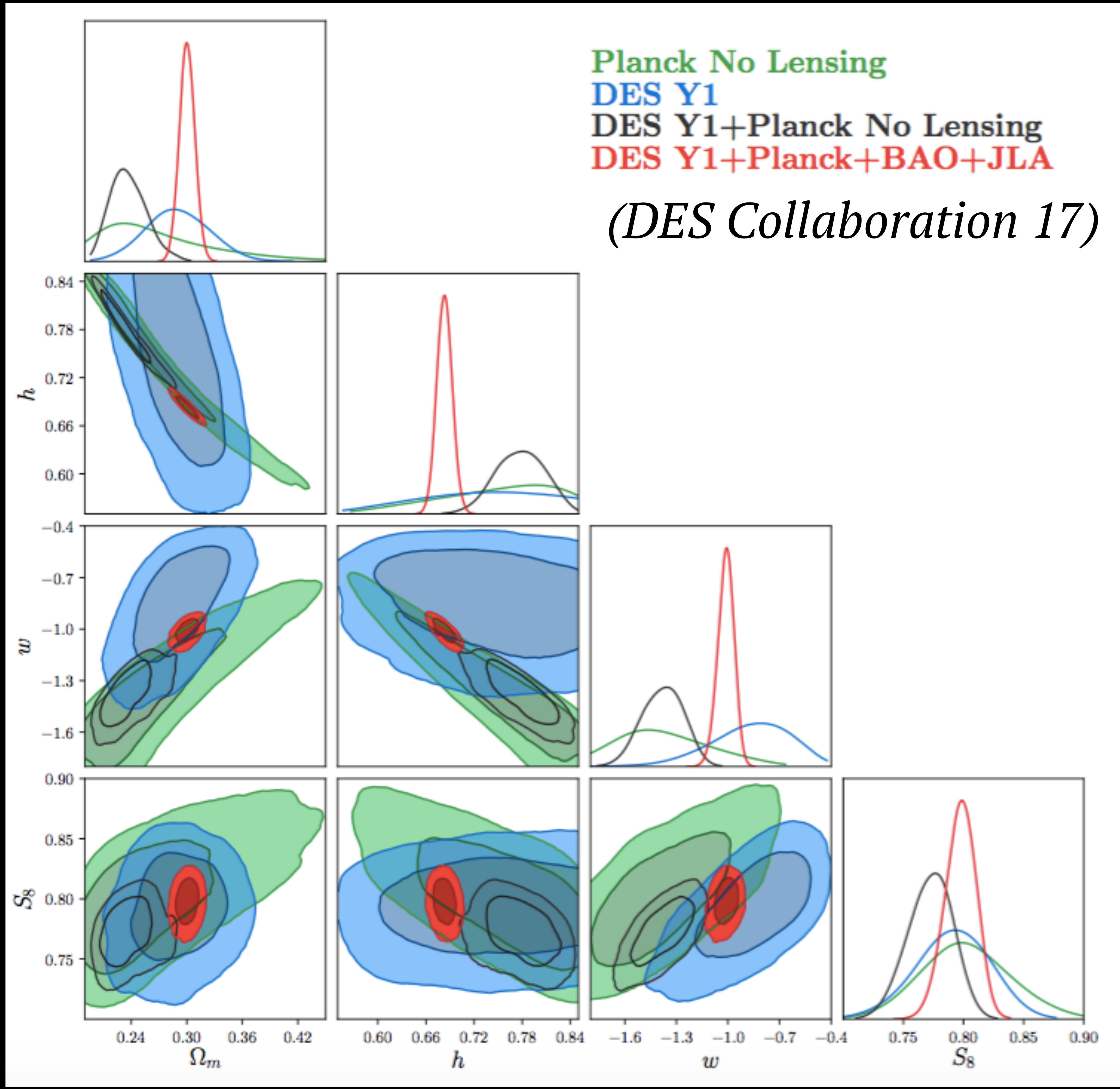
- DES and Planck (TT+lowP) constrain matter density and S_8 with comparable strength
- Central values differ by >1σ, in the same direction as earlier lensing results
- Bayes Factor R = 6.6 evidence for consistency in ΛCDM
- Still consistent (R=9.0) for joint low-z results + Planck, which is why we combine...



Key Result: Combined Constraints in wCDM

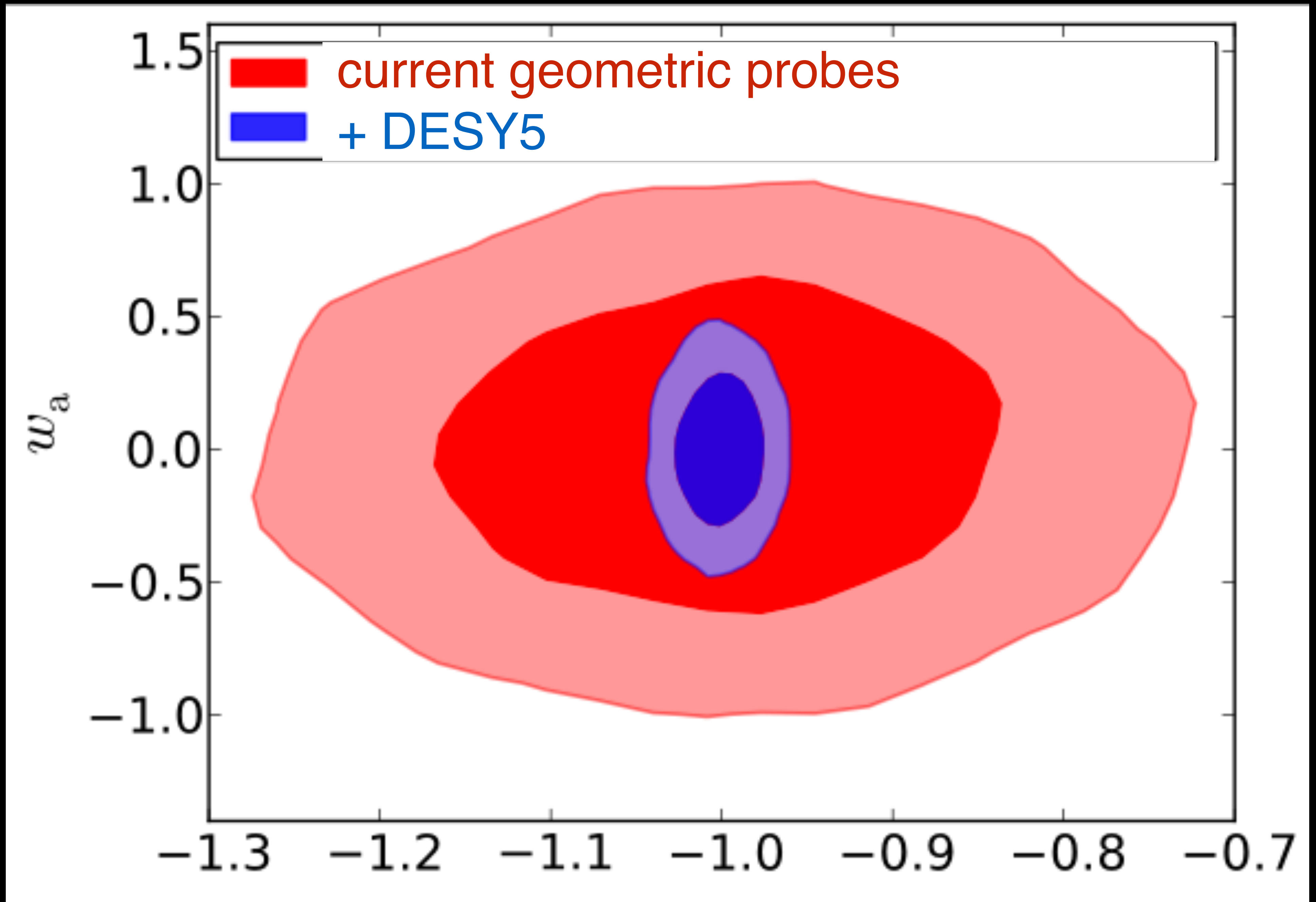
- consistent constraints from geometric probes (R=244)
- most precise parameter constraints from DES+Planck +BAO (BOSS) +SN (JLA)
- no evidence for w ≠-1

$$w = -1.00^{+0.04}_{-0.05}$$



Steps forward: more precise tests of broader range of models

- DESY1++ is a precise test of ΛCDM. Any potential discrepancies are smaller than its uncertainty.
- It does not explain ΛCDM.
- It is not very sensitive to models with time-varying Dark Energy equation of state (among others)
- Future joint analyses will be!



Forecast: Eifler++

Conclusions

- DES Y1 Cosmology results from galaxy clustering, galaxy-galaxy lensing, and cosmic shear (3x2) are out: 19 papers, more to follow.
- DES Y1 results consistent with Planck CMB *in context of ΛCDM*.
- DES Y1 results in combination with Planck, BAO, JLA SN provide stringent constraints on ΛCDM parameters.
- Precision will increase with larger data sets, and by bringing in more probes (clusters, SN, cross-correlations…), enabling tests of more complex models ($w_0 w_a$CDM, modified gravity,…)
- 2020s will be an exciting decade for survey cosmology!