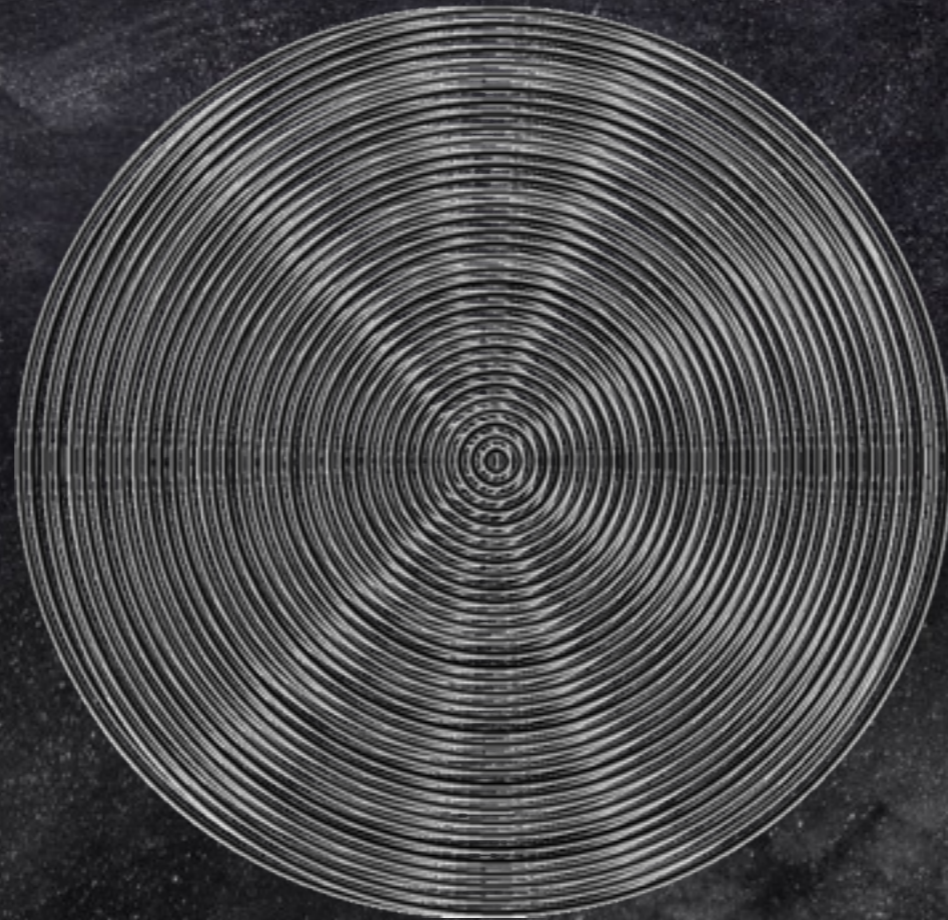


Di-boson Interference: Death and Resurrection

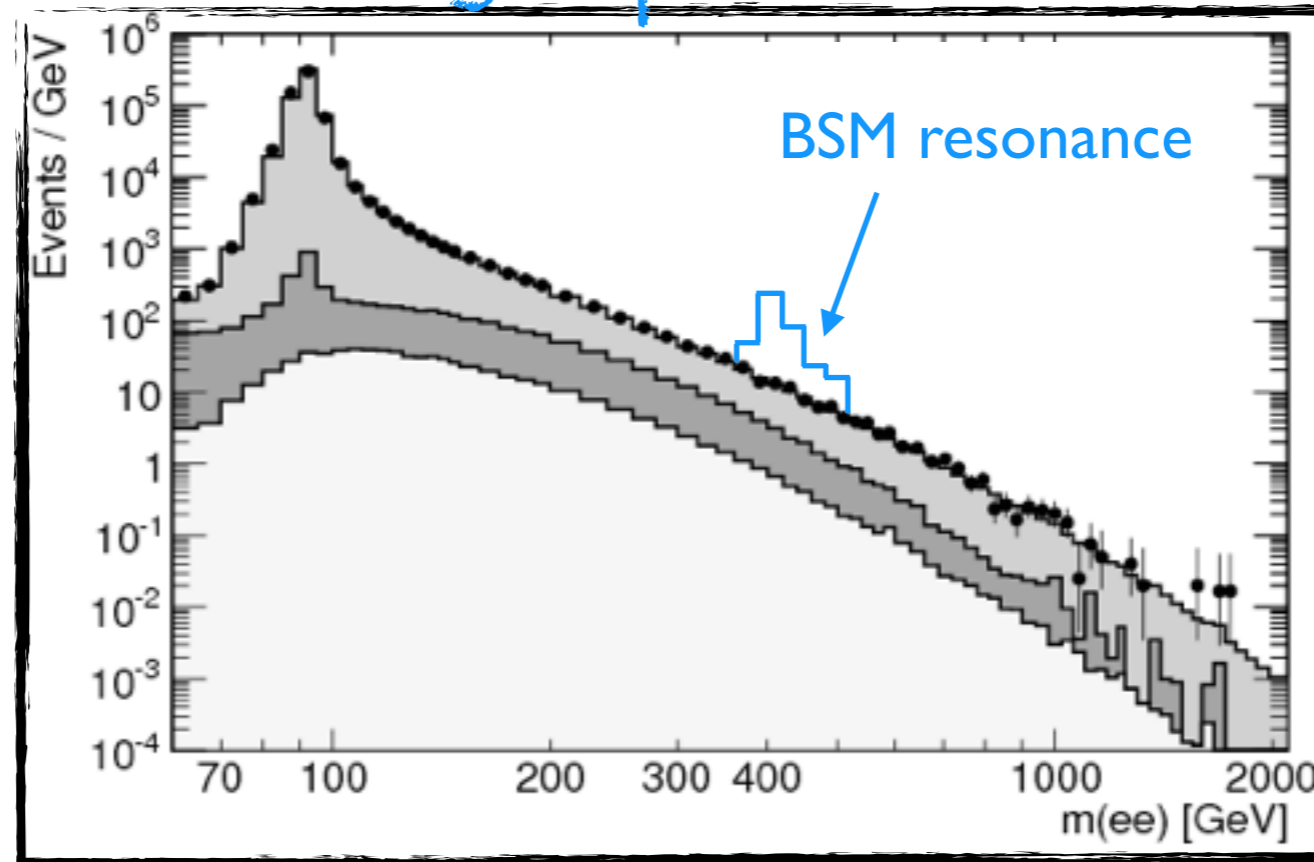


Francesco Riva
(CERN)

In collaboration with
Panico, Wulzer 1708.07823,
Azatov, Contino, Machado 1607.05236
Liu, Pomarol, Rattazzi 1603.03064

LHC Exploration (so far 2009-2015)

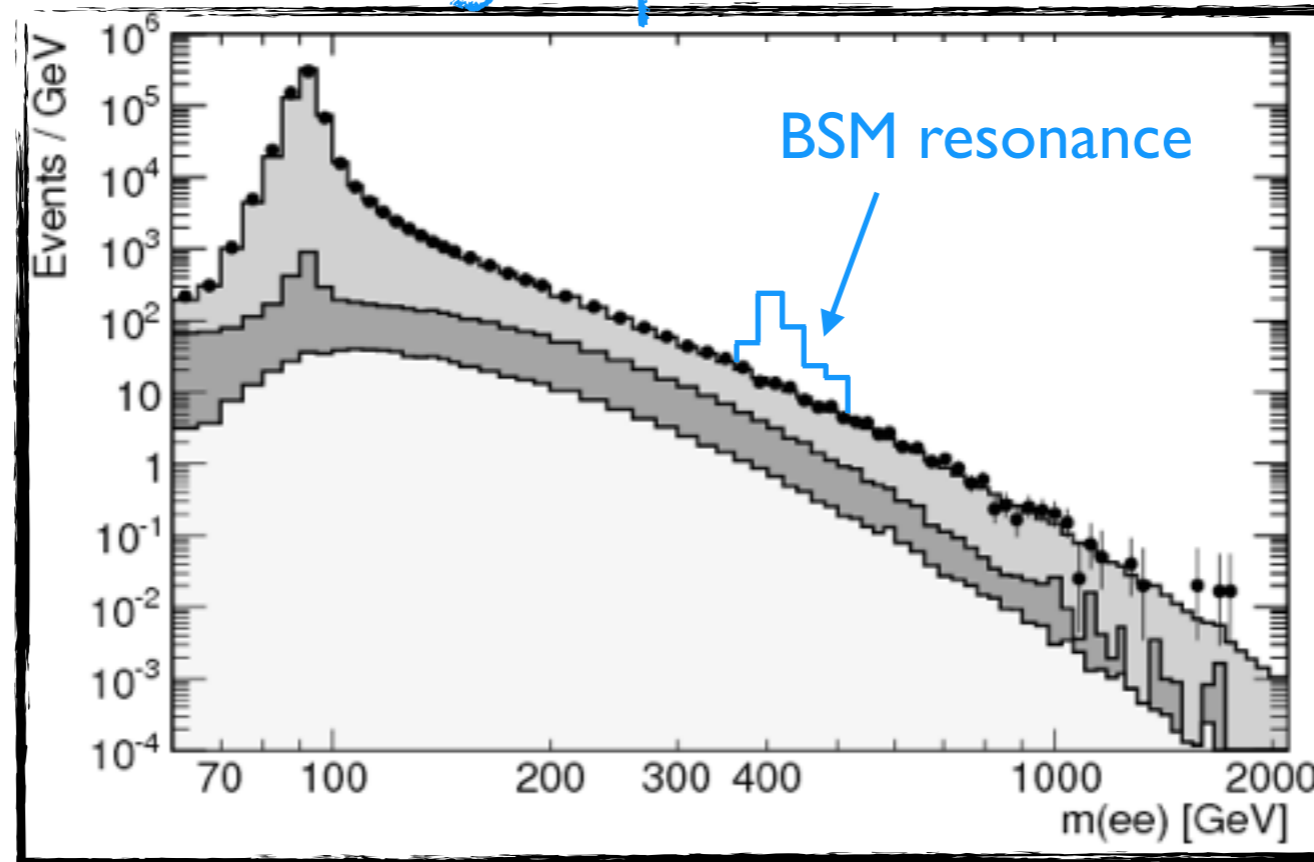
Focus: Search for new light particles



Energy frontier (13 TeV)

LHC Exploration (so far 2009-2015)

Focus: Search for new light particles

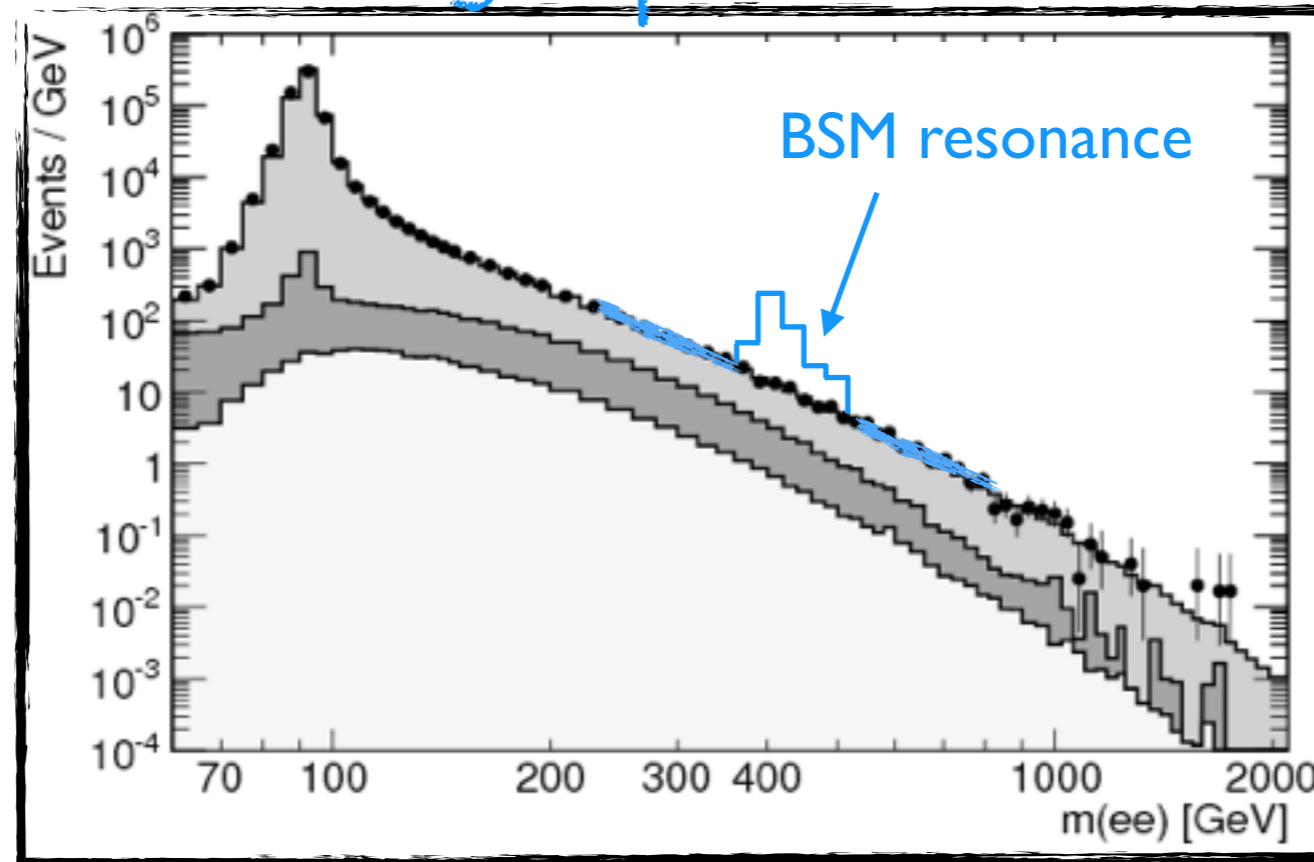


Energy frontier (13 TeV)

► Experimentally: First accessible signal/Easy to study

LHC Exploration (so far 2009-2015)

Focus: Search for new light particles



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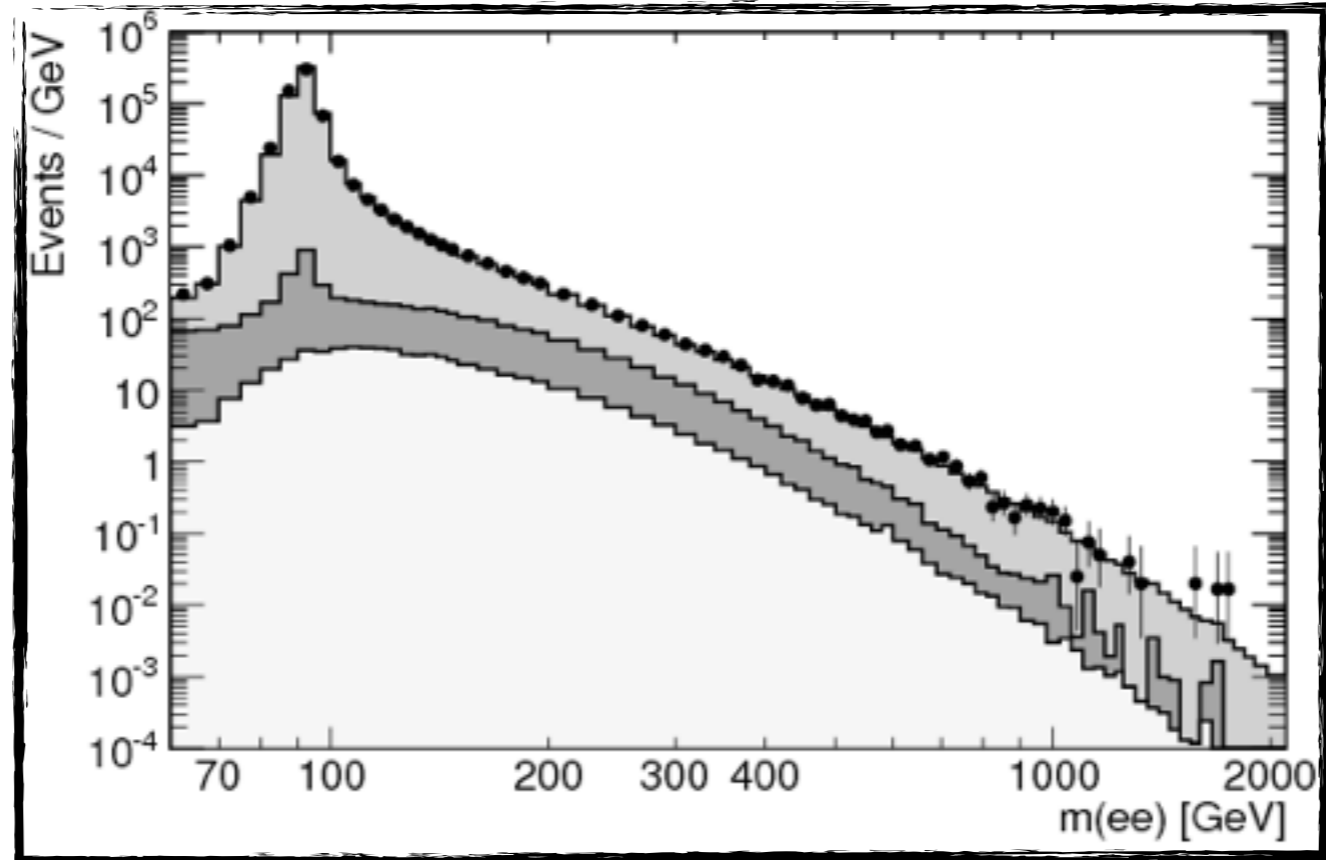
LHC Exploration (now \rightarrow 2030's)

Focus: Standard Model Precision Tests

(2035: 3000 fb^{-1})

intensity
frontier

(2016: 40 fb^{-1})



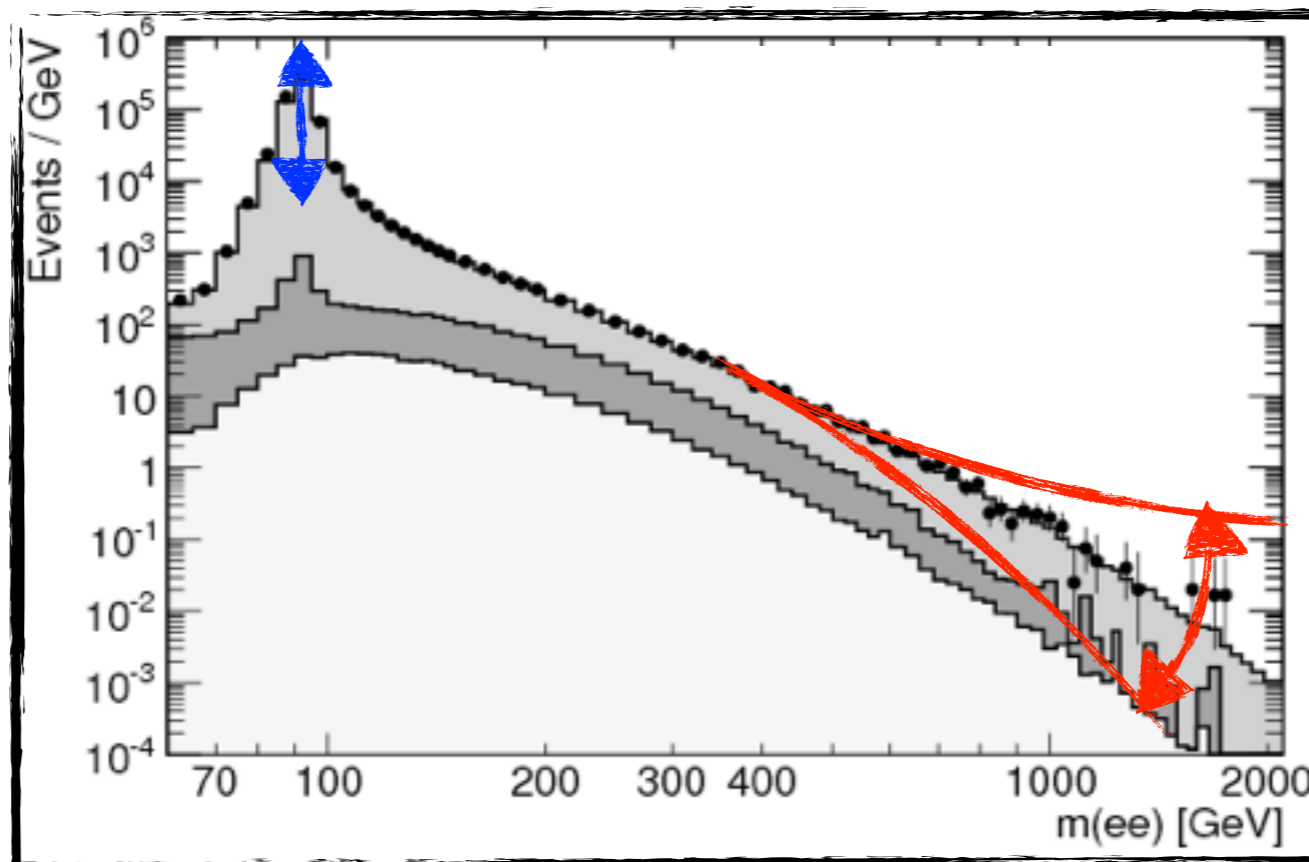
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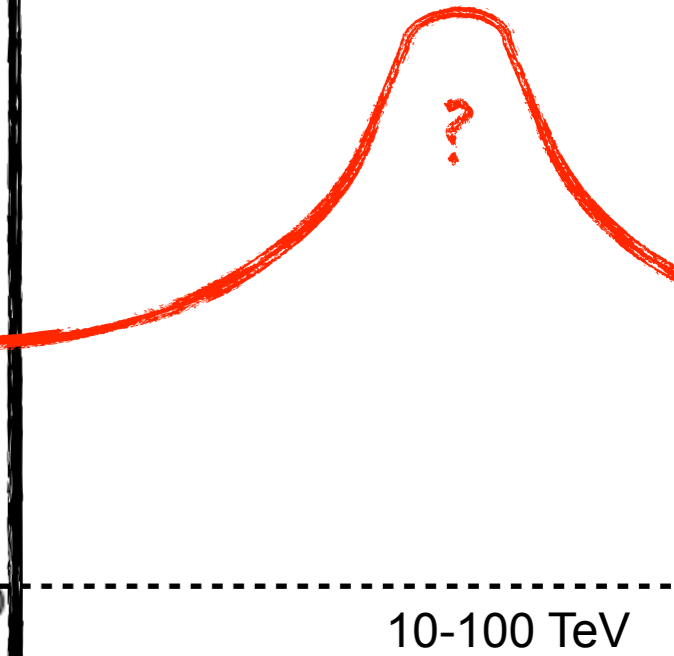
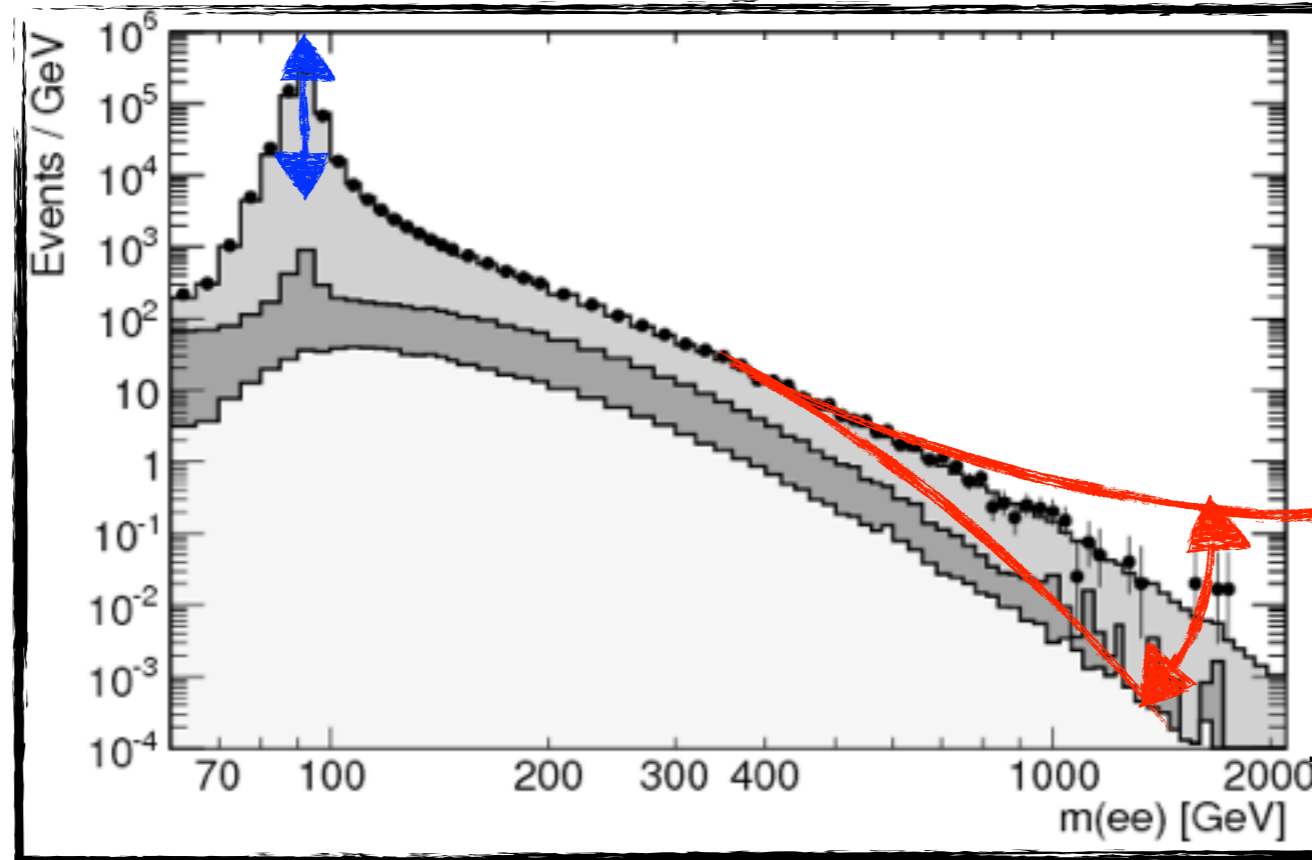
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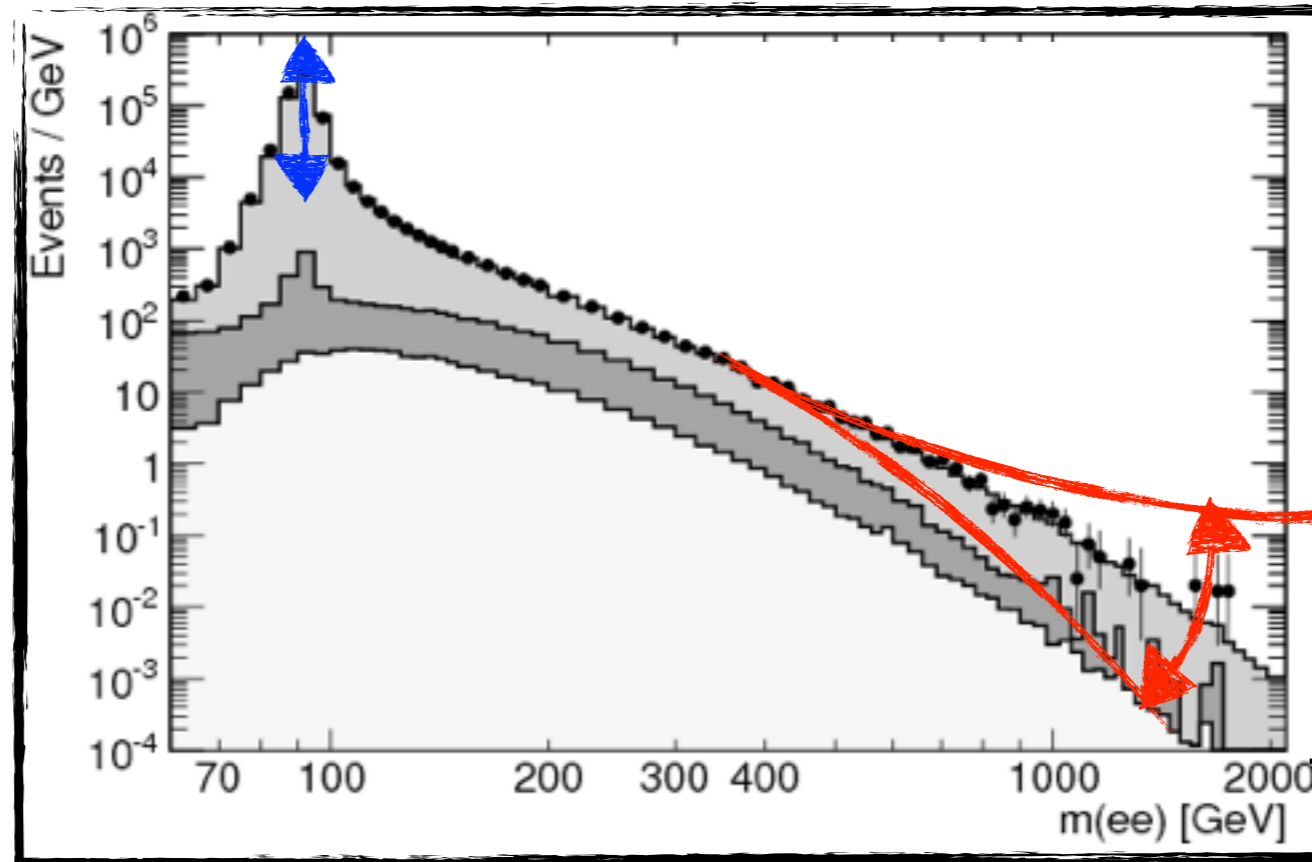
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10-100 TeV

M

$$Amp = SM \left(1 + c \frac{E^2}{M^2} \right)$$

Effective
Field
Theory



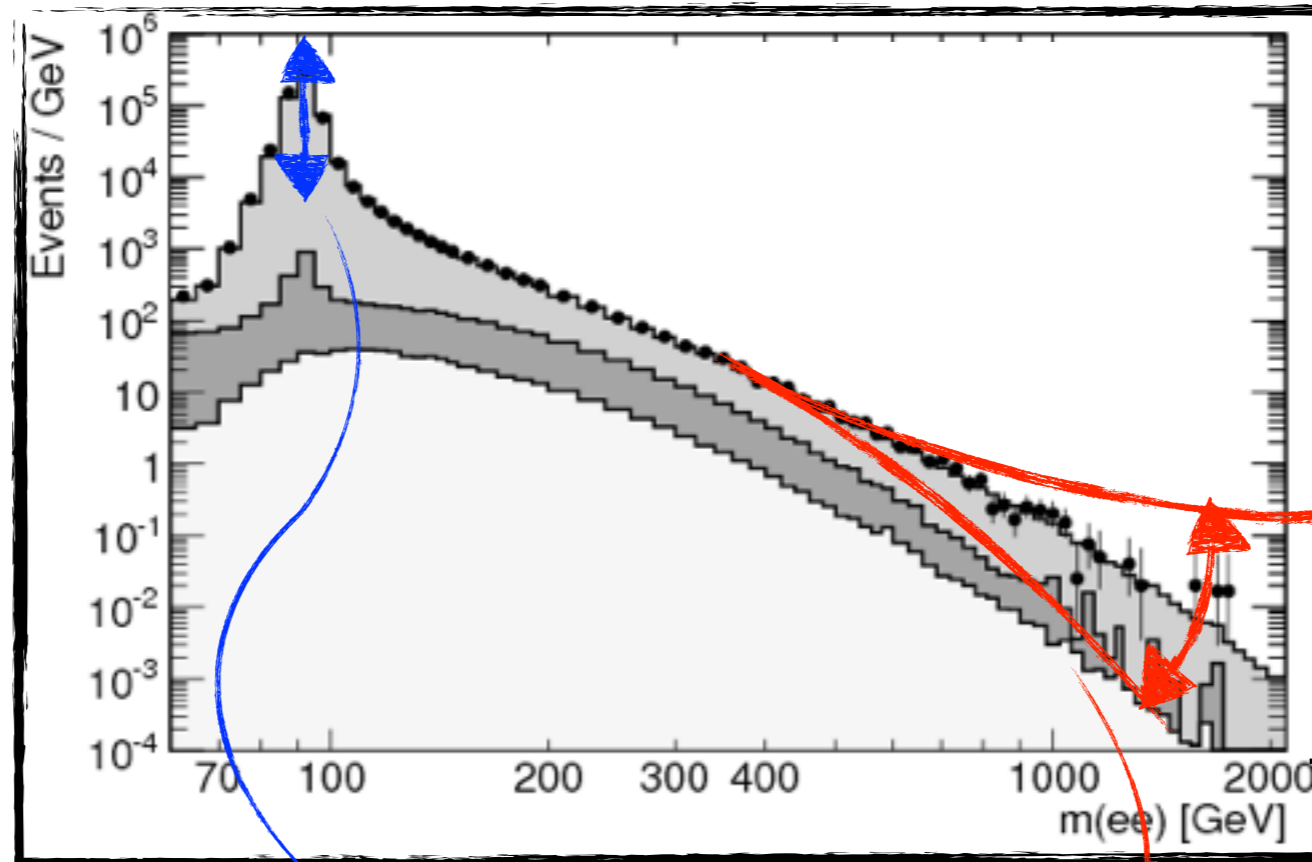
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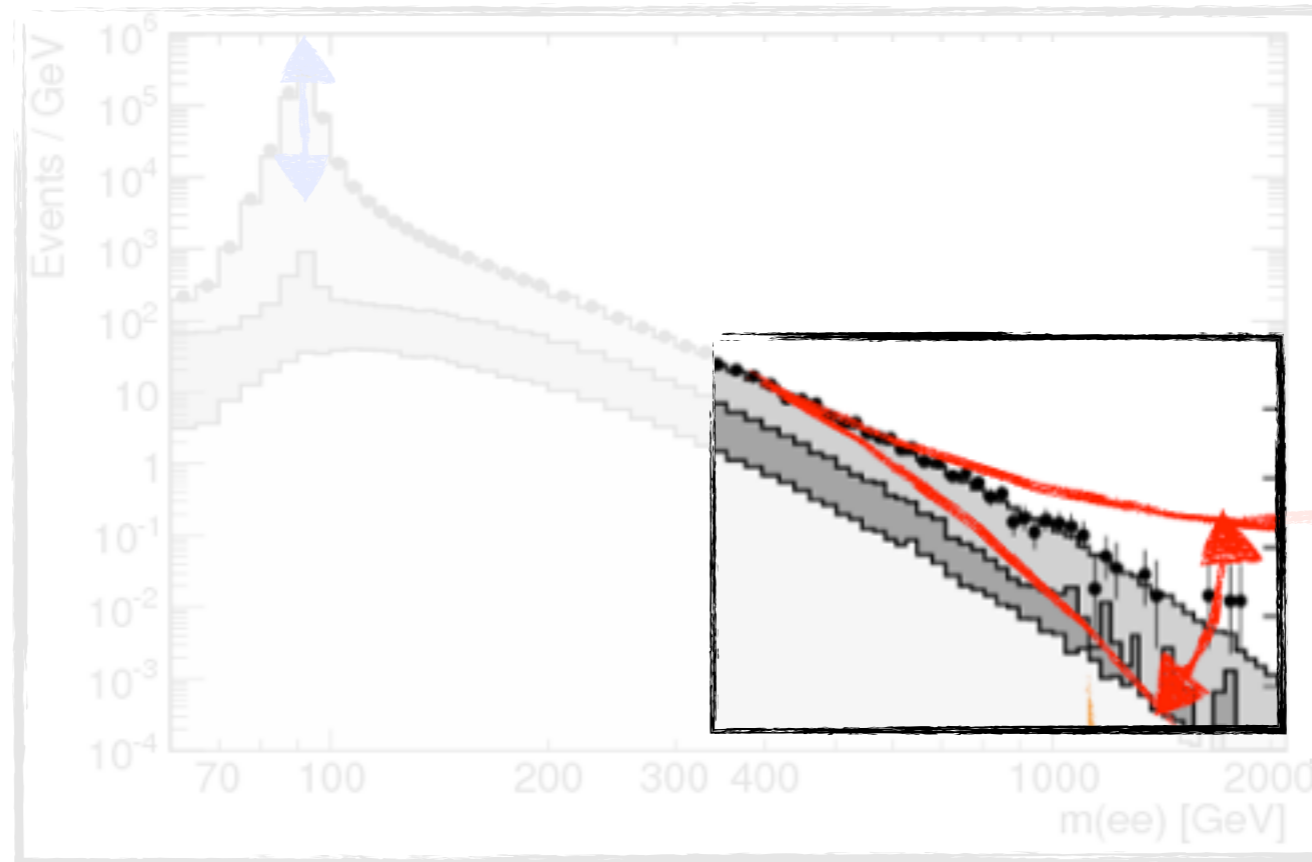
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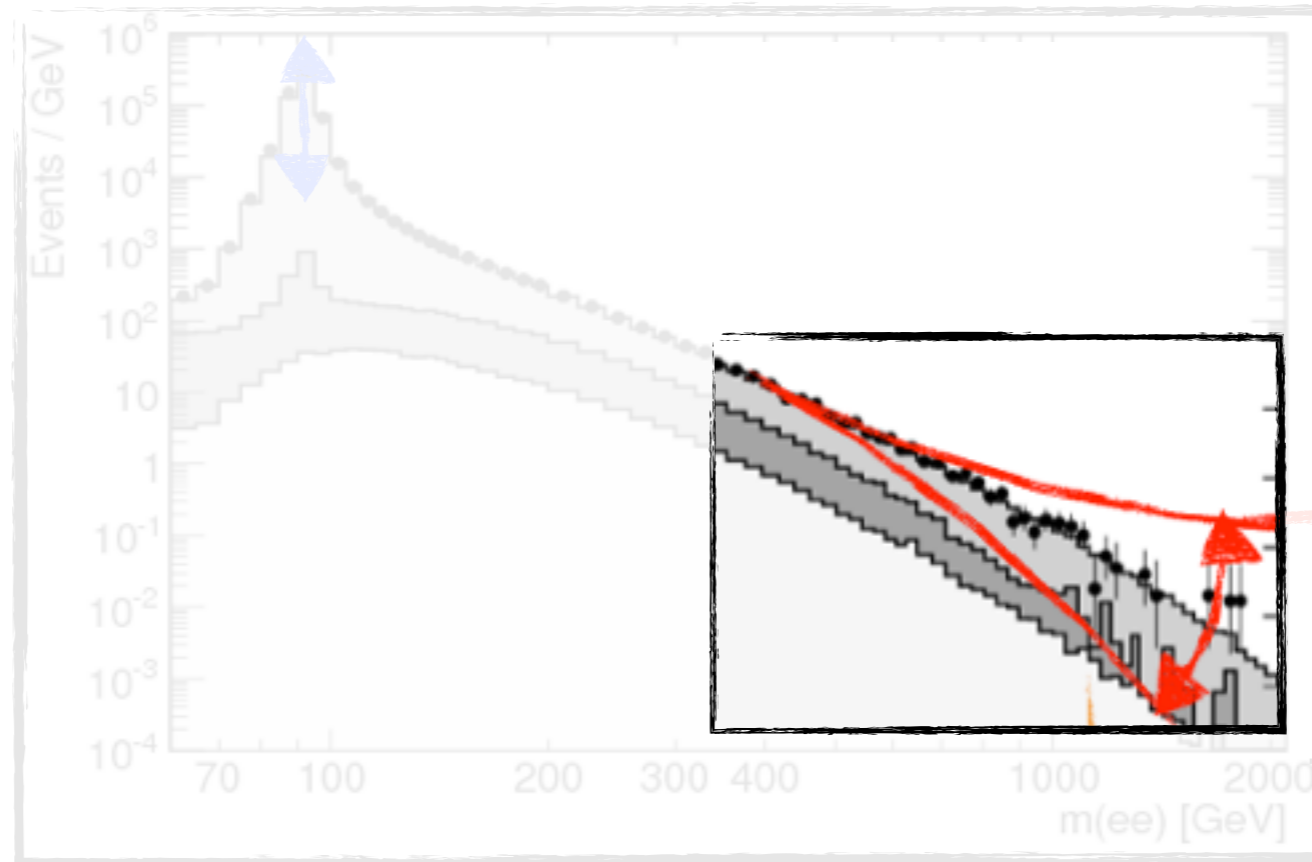
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$$Amp = SM \left(1 + c \frac{E^2}{M^2} \right)$$

Effective
Field
Theory

M

► Experimentally: Already very difficult...

and comes with a fundamental challenge: non-interference

Outline

- 1- Why SM/BSM Interference?
- 2- No Interference
- 3- Interference

Why Interference?

When SM and BSM contribute to the same amplitude:

$$\text{Amp} = SM + BSM = SM(1 + \delta_{BSM})$$

$\delta_{BSM} = c \frac{E^2}{M^2}$

Why Interference?

When SM and BSM contribute to the same amplitude:

$$Amp = SM + BSM = SM(1 + \delta_{BSM})$$

$\delta_{BSM} = c \frac{E^2}{M^2}$

► $\sigma \propto |Amp|^2 \simeq SM^2(1 + \delta_{BSM} + \delta_{BSM}^2)$

For **small** BSM effects $1 \gg \delta_{BSM}$,

interference dominates $\delta_{BSM} \gg \delta_{BSM}^2$

Non-Interference?

If SM and BSM contribute to different amplitudes:

$$\blacktriangleright \sigma \propto \sum |Amp|^2 \simeq SM^2 \left(1 + \cancel{c_i \frac{E^2}{\Lambda^2}} + c_i^2 \frac{E^4}{\Lambda^4} \right)$$

interference
vanishes

Non-Interference?

If SM and BSM contribute to different amplitudes:

interference vanishes

$$\sigma \propto \sum |Amp|^2 \simeq SM^2 \left(1 + \cancel{c_i \frac{E^2}{\Lambda^2}} + c_i^2 \frac{E^4}{\Lambda^4} \right)$$

The leading effects BSM are $O\left(\frac{1}{\Lambda^4}\right)$:

(the same order as dimension-8 that do interfere)

► Small effects, even smaller!

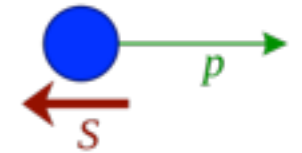
Interference is important... can we always count on it?

Non-Interference for BSM₆ amplitudes

Azatov, Contino, Machado, FR'16

Exploit:

For $E \gg m_W$ states have well defined helicity



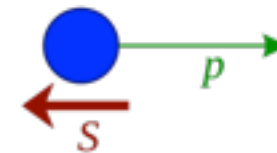
Amplitudes for $2 \rightarrow 2$ with different total h don't interfere

Non-Interference for BSM₆ amplitudes

Azatov, Contino, Machado, FR'16

Exploit:

For $E \gg m_W$ states have well defined helicity



Amplitudes for 2→2 with different total h don't interfere

Theorem:

A_4	$ h(A_4^{\text{SM}}) $	$ h(A_4^{\text{BSM}}) $
VVVV	0	4,2
VV $\phi\phi$	0	2
VV $\psi\psi$	0	2
V $\psi\psi\phi$	0	2
$\psi\psi\psi\psi$	2,0	2,0
$\psi\psi\phi\phi$	0	0
$\phi\phi\phi\phi$	0	0

Any BSM dim-6 operator

Massless limit + tree level + at least one transverse vector

- ▶ SM and BSM₆ contribute to different helicity amplitudes
- ▶ No interference

(

proof

Non-Interference for BSM_6 amplitudes

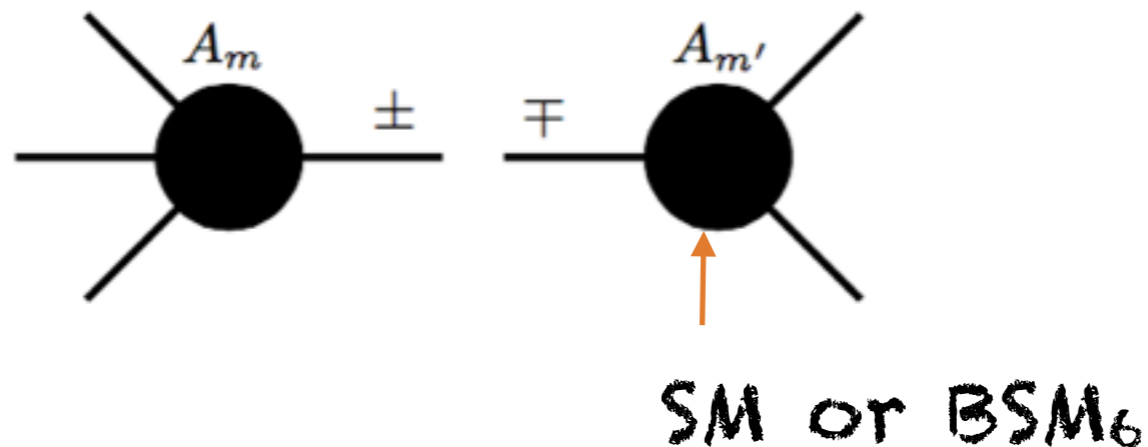
Azatov, Contino, Machado, FR'16

How?

i) Helicity sums for amplitudes on-shell:

$$h(A_n) = h(A_m) + h(A_{m'})$$

$n = m + m' - 2$ legs



Non-Interference for BSM_6 amplitudes

Azatov, Contino, Machado, FR'16

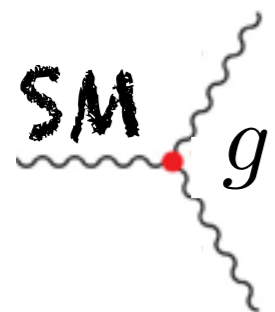
How?

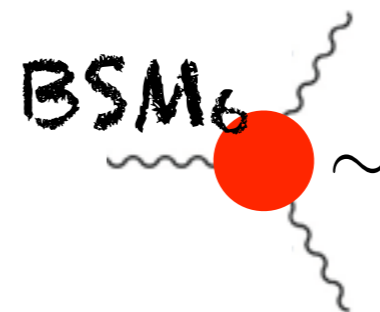
ii) For massless theories 3-point amplitude determined:

Poincaré
Dim. Analysis
3-point kinematics

Helicity of 3-point \leftrightarrow coupling dimension,

$$|h(A_3)| = 1 - [g]$$

SM  $g \Rightarrow |h^{SM}| = 1$

BSM_6  $\sim \frac{1}{\Lambda^2} \Rightarrow |h^{BSM}| = 3$

Non-Interference for BSM₆ amplitudes

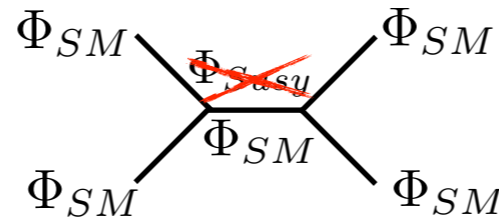
Azatov, Contino, Machado, FR'16

How?

iii) SUSY* Ward Identities: $|h(A_4^{SM})| < 2$ (except $\psi^+\psi^+\psi^+\psi^+$)
(maximal helicity violation)

*For $y_u=0$ or $y_d=y_L=0$: SM upliftable to SUSY+R-parity (1 Higgs doublet)

Grisaru, Pendleton, vanNieuwenhuizen'77



Non-Interference for BSM₆ amplitudes

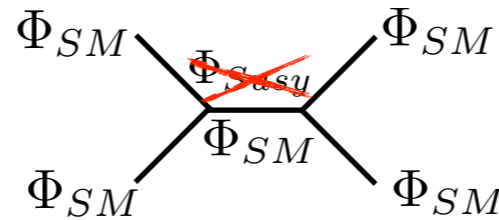
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Grisaru, Pendleton, vanNieuwenhuizen'77



e.g. $0 = \langle 0 | [Q, \Psi^+ V^+ V^+ V^+] | 0 \rangle = \sum_i \langle 0 | \Psi^+ \dots [Q, V^+] \dots V^+ | 0 \rangle \propto \langle 0 | V^+ V^+ V^+ V^+ | 0 \rangle$

$[Q, \psi] \sim V^i$ $[Q, V] \sim \psi$ $\psi^+ \psi^+ \dots = 0$

▶ $A(V^+ V^+ V^+ V^-) = A(V^+ V^+ \psi^+ \psi^-) = A(V^+ V^+ \phi \phi) = A(V^+ \psi^+ \psi^+ \phi) = 0$

Non-Interference for BSM₆ amplitudes

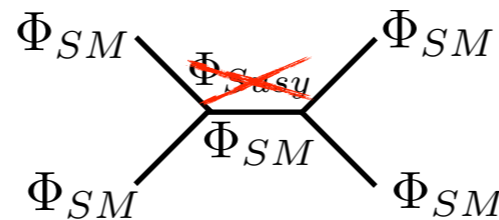
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$$\text{e.g. } 0 = \langle 0 | [Q, \Psi^+ V^+ V^+ V^+] | 0 \rangle = \sum_{[Q, \psi] \sim V^i} \langle 0 | \Psi^+ \dots [Q, V^+] \dots V^+ | 0 \rangle \propto \langle 0 | V^+ V^+ V^+ V^+ | 0 \rangle$$

$\psi^+\psi^+ \dots = 0$

▶ $A(V^+V^+V^+V^-) = A(V^+V^+\psi^+\psi^-) = A(V^+V^+\phi\phi) = A(V^+\psi^+\psi^+\phi) = 0$

BSM not supersymmetrizable in general

Non-Interference for BSM₆ amplitudes

Azatov, Contino, Machado, FR'16

How?

i) Helicity sums $h(A_n) = h(A_m) + h(A_{m'})$

ii) Helicity of 3-point \leftrightarrow coupling dimension $|h(A_3)| = 1 - [g]$

iii) SUSY* Ward Identities $|h(A_4^{SM})| < 2$

Non-Interference for BSM_6 amplitudes

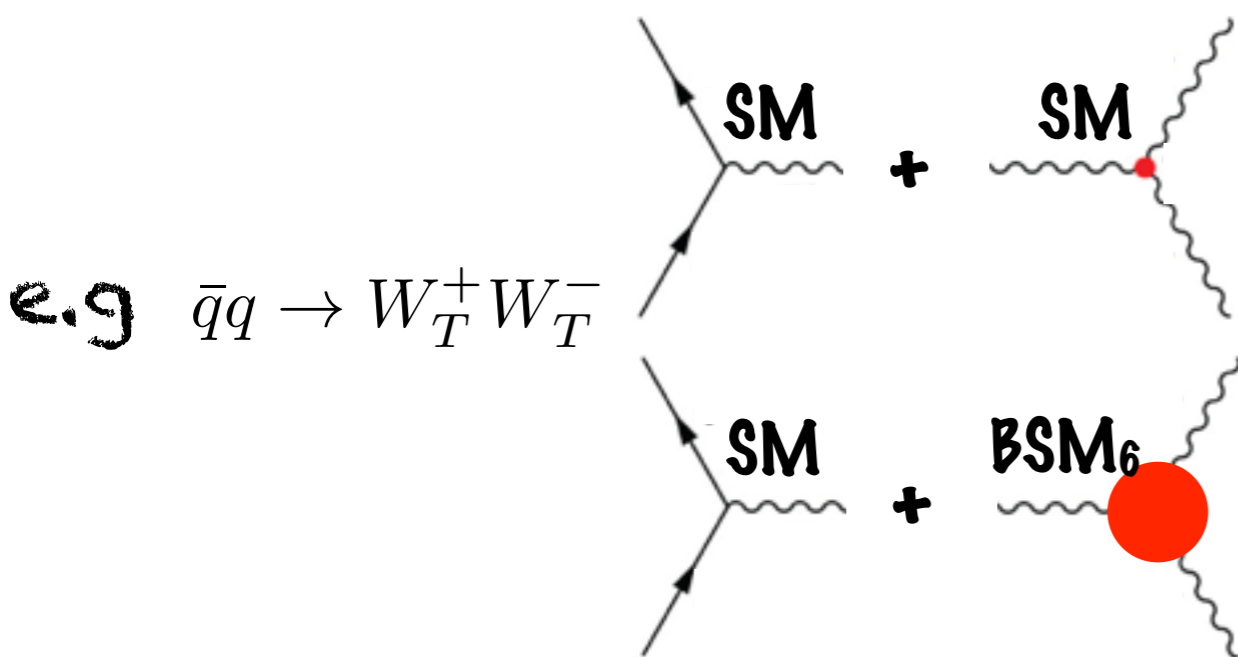
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$$h^{tot} = 1 - 1 = 0$$

$$h^{tot} = 1 \pm 3 = 2, 4$$

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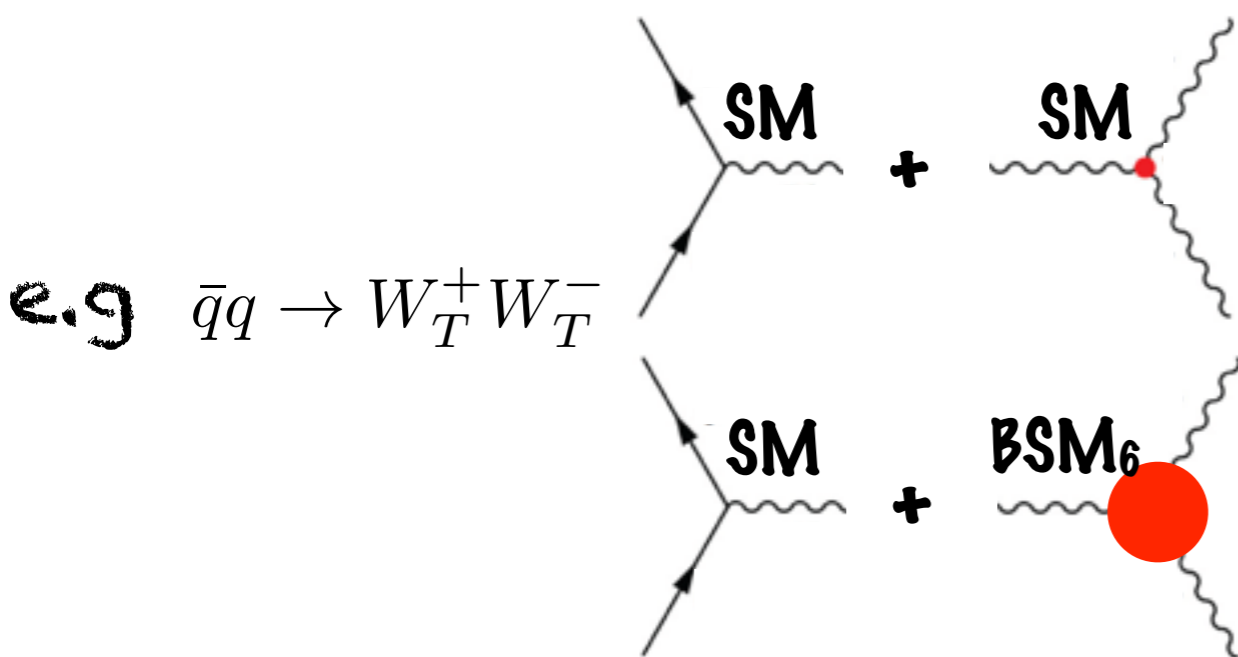
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(dim-6, 4-point)

Non-Interference for BSM_6 amplitudes

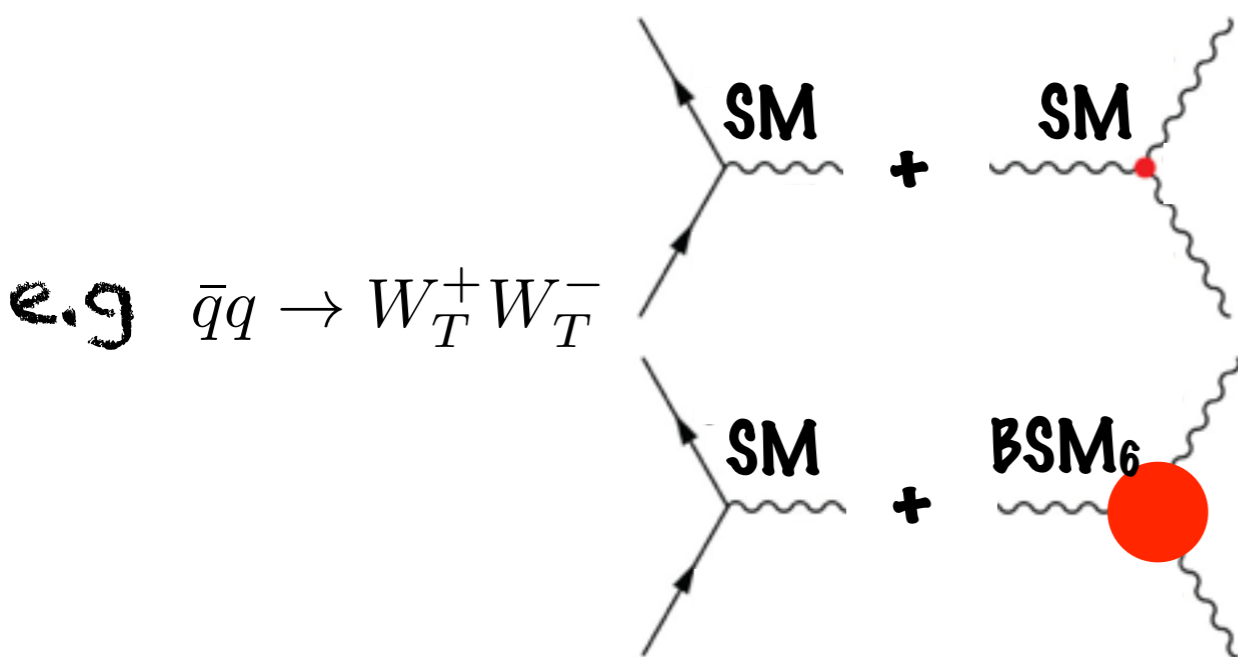
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No Interference
(dim-6, 4-point)

Massless limit + tree level + at least one transverse vector

► SM and BSM_6 no interference

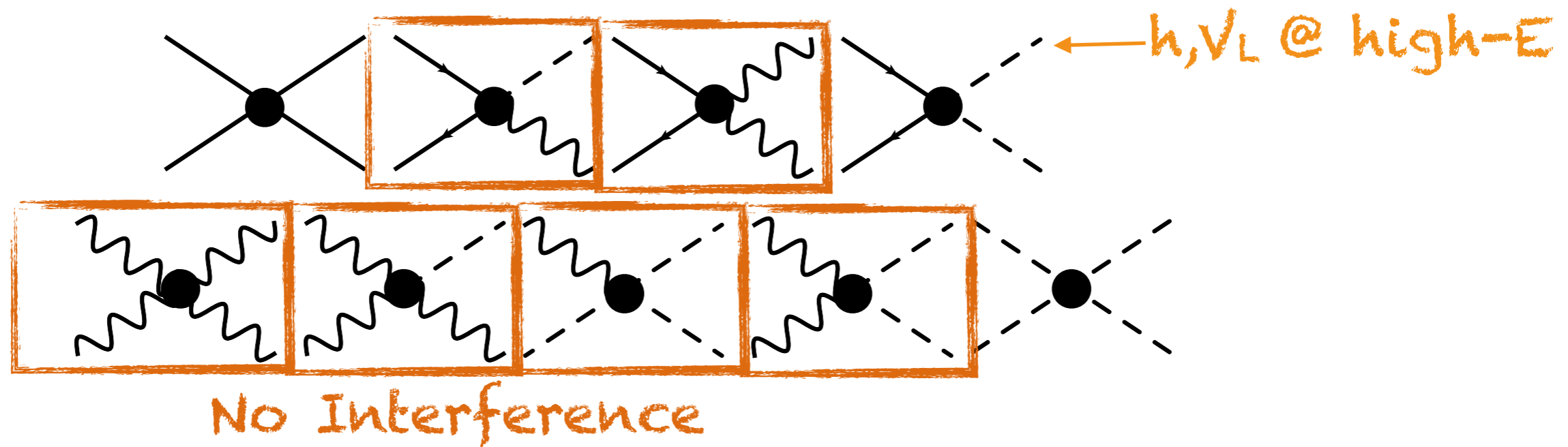
proof

)

Non-Interference for BSM₆ amplitudes

Azatov, Contino, Machado, FR'16

2→2 processes:



What to do to test transverse vectors?

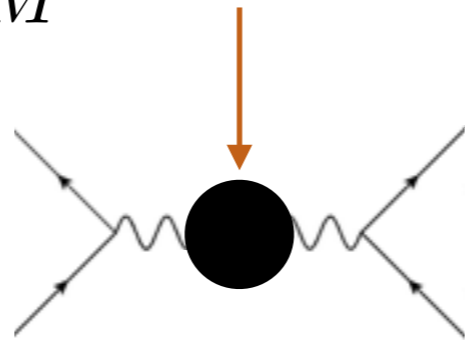
- 1) BSM with large effects can be tested consistently
see Pomarol's talk
- 2) Vectors in interfering amplitudes
- 3) NLO effects that do interfere
- 4) Interference Resurrection

2. Vectors in interfering amplitudes

- ▶ Effects in transverse 2-point function affect interfering 4-point amplitude

Remedios Scenario $\rightarrow \frac{1}{M^2} (D_\rho W_\mu^{a,\nu})^2$
Liu, Pomarol, Rattazzi, FR'16

See Wulzer talk

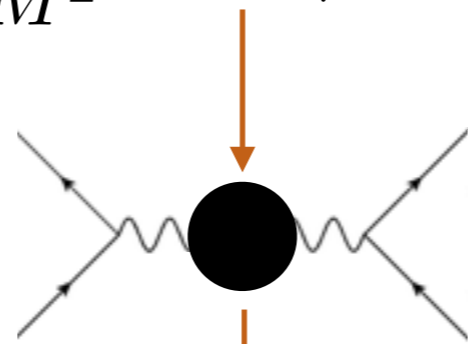


2. Vectors in interfering amplitudes

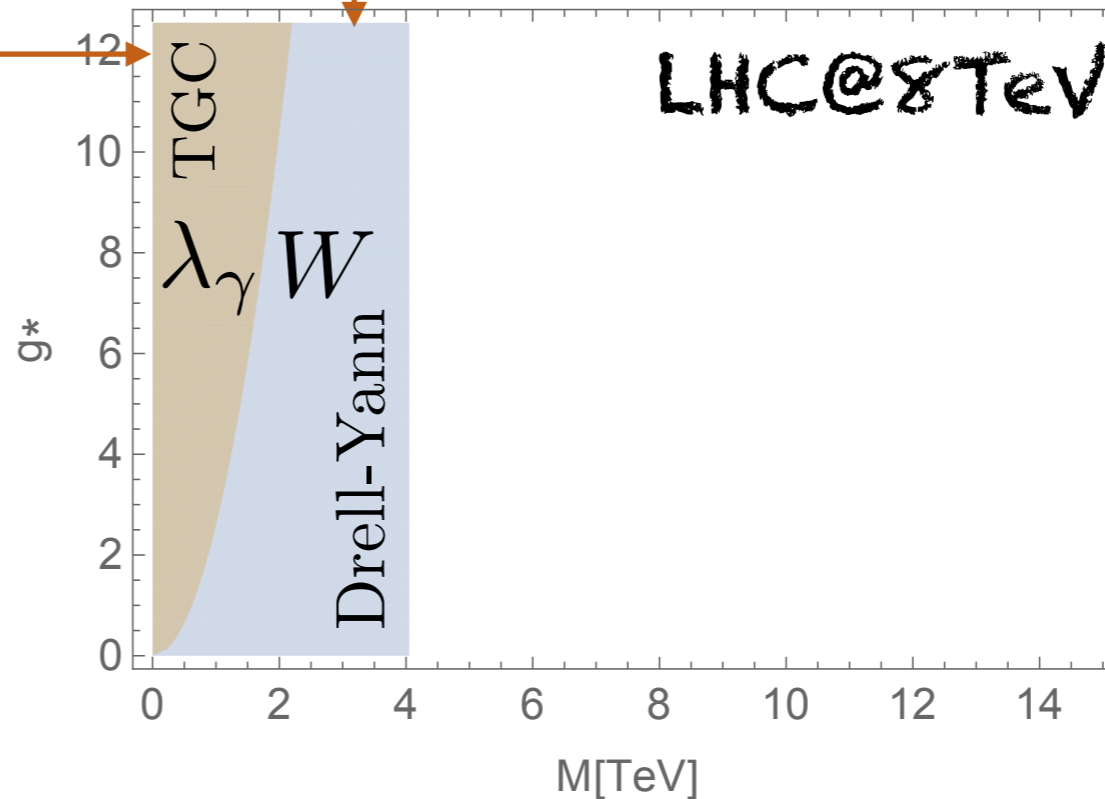
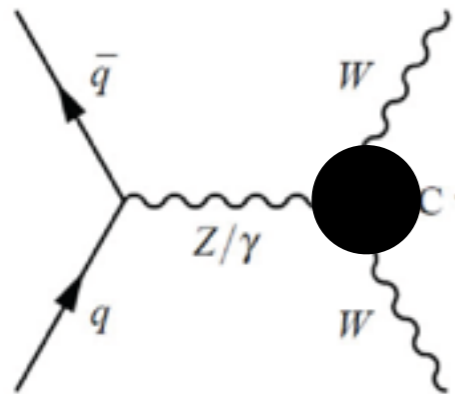
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Remedios Scenario $\rightarrow \frac{1}{M^2} (D_\rho W_\mu^{a,\nu})^2$
 Liu,Pomarol,Rattazzi,FR'16



$\frac{g_*}{M^2} \epsilon_{abc} W_\mu^{a\nu} W_{\nu\rho}^b W^{c\rho\mu}$



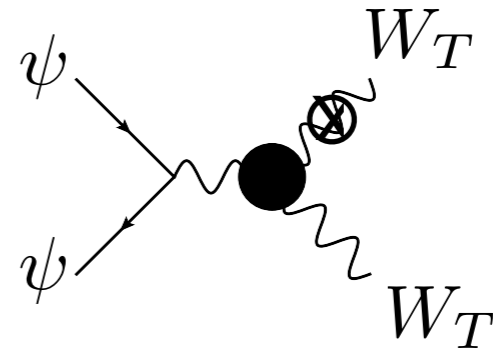
3. NLO

Non-interference only for massless/tree-level/2→2 processes!

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Non-interference only for massless/tree-level/2- \rightarrow 2 processes!

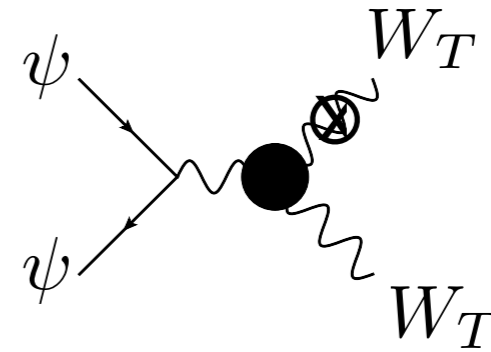
▶ EW finite mass effects $\sim \frac{m_W^2}{E^2}$



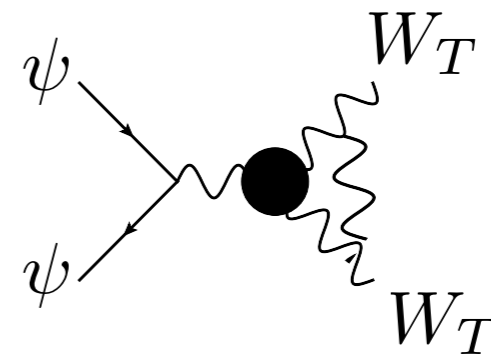
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▶ Loop effects $\sim \frac{\alpha_{s,em}}{4\pi}$ (no soft-limit enhancement, at high-energy)

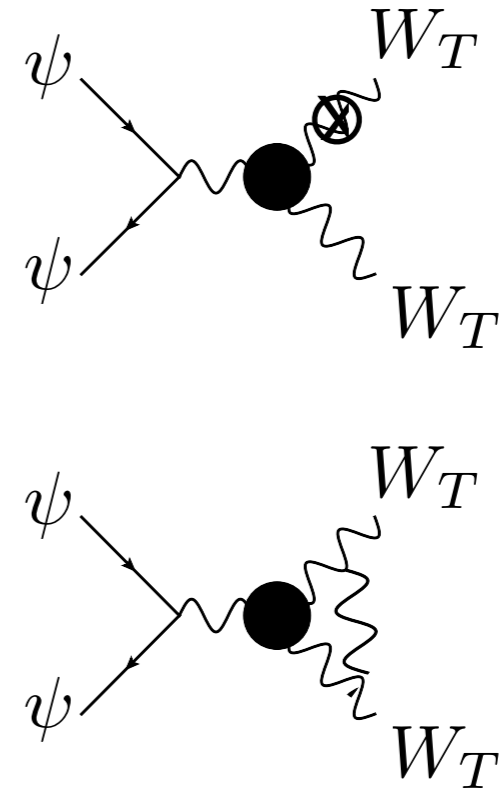


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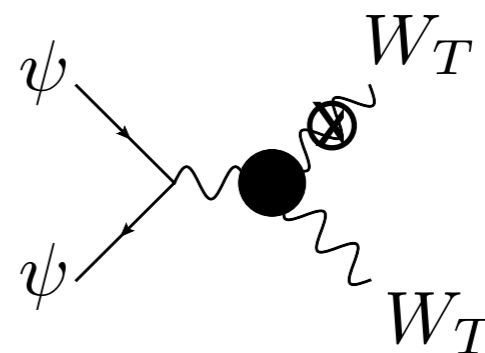


} Only signal suppressed

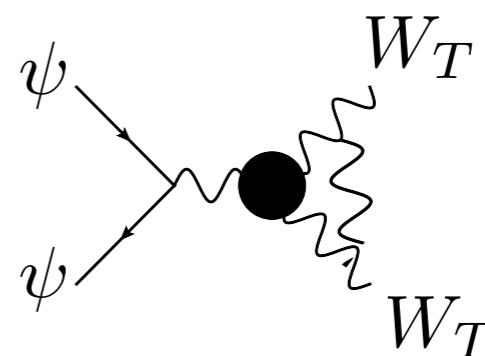
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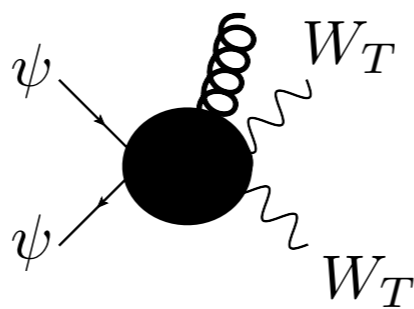
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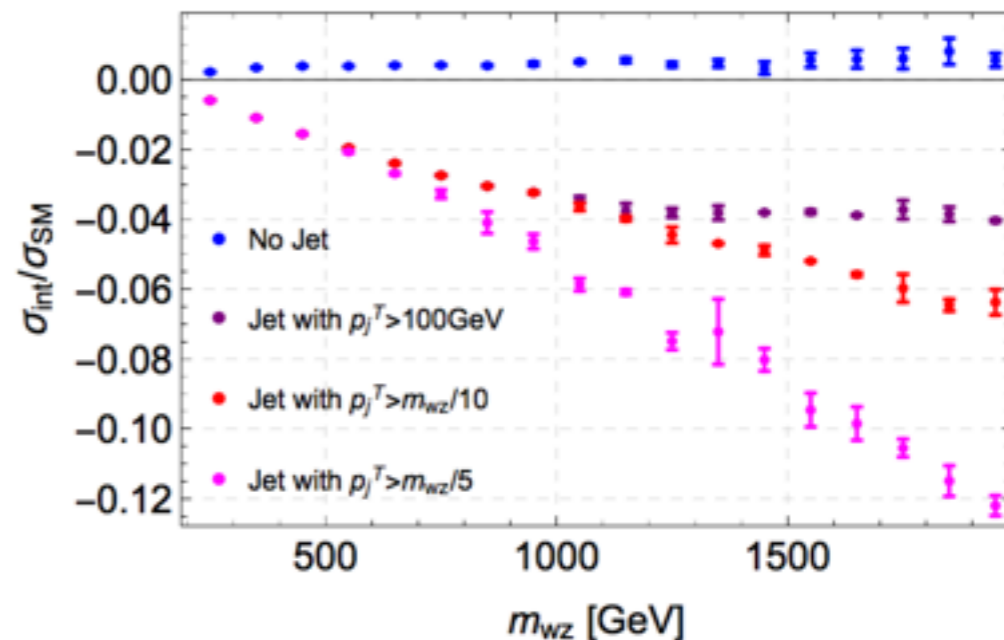
Only signal suppressed

▶ Hard radiation 2→3

$$\frac{S}{\sqrt{B}} \sim \sqrt{\frac{4\pi}{\alpha_s}} \times \frac{\alpha_{s,em}}{4\pi}$$



for gluons see Dixon,Shadmi'93;



3. NLO

Non-interference only for massless/tree-level/2→2 processes!

▶ EW finite mass effects $\sim \frac{m_W^2}{E^2}$

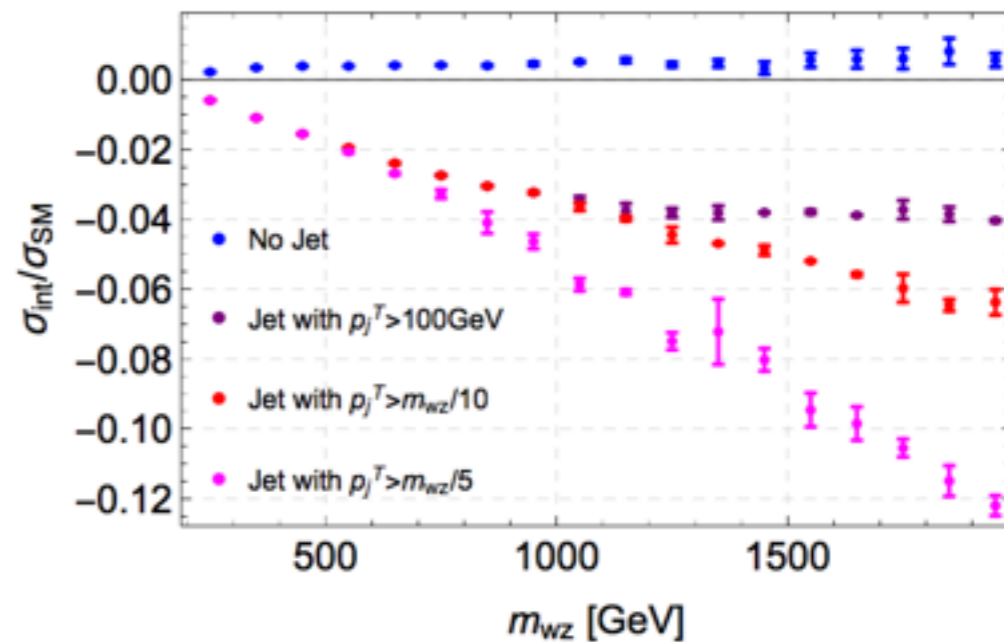
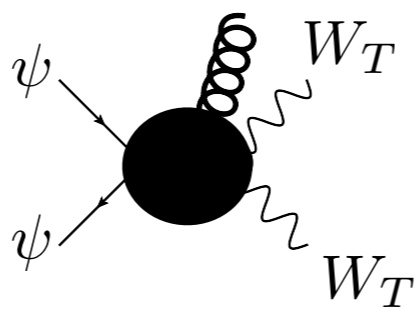
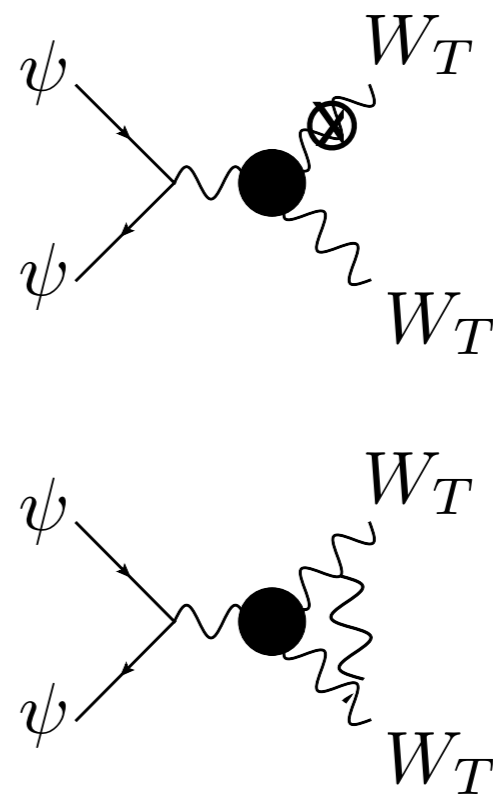
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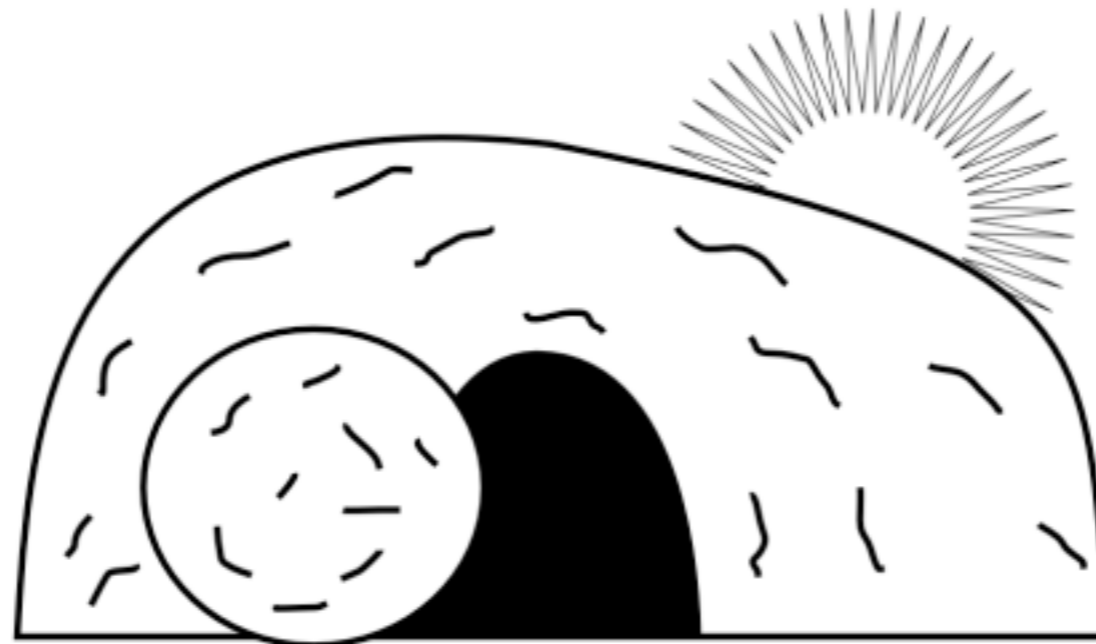
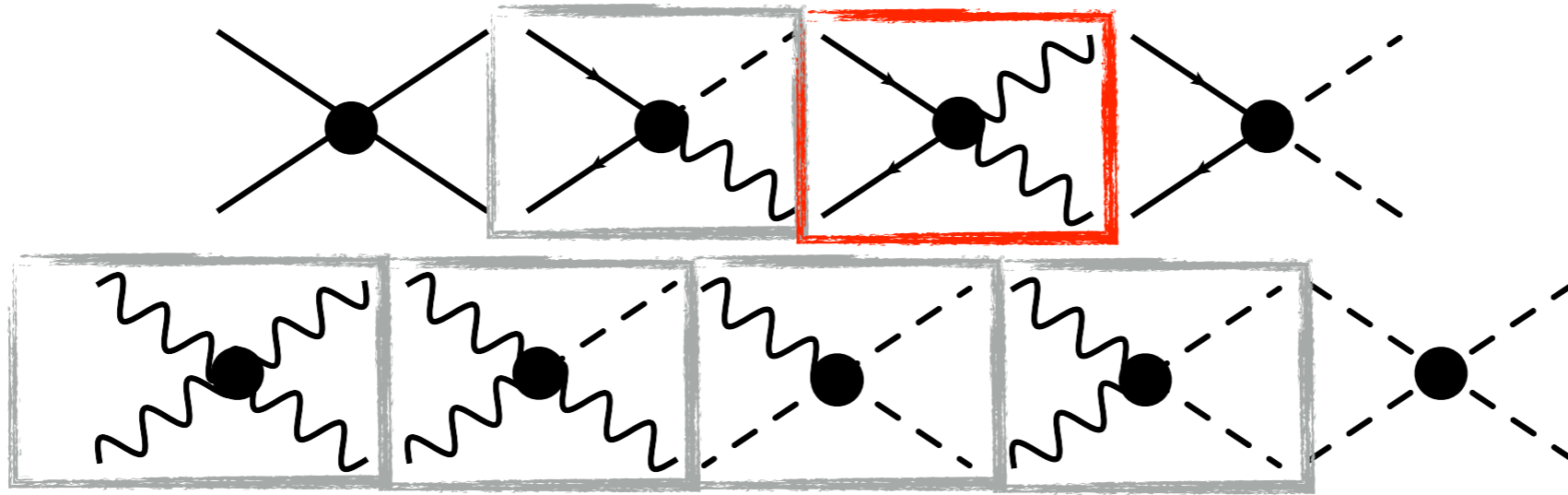
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for gluons see Dixon,Shadmi'93;

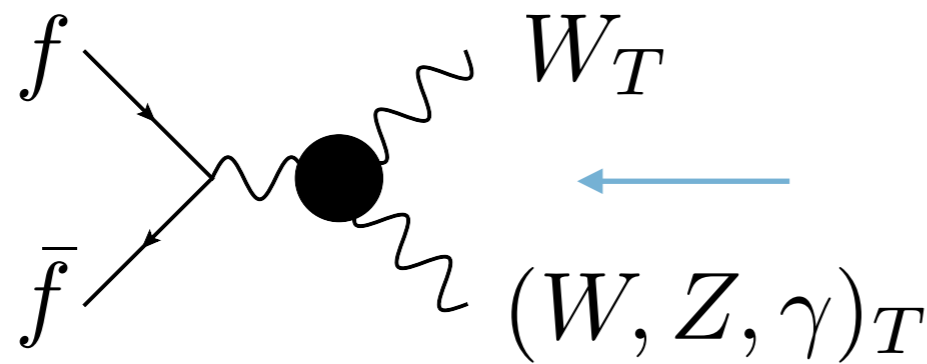
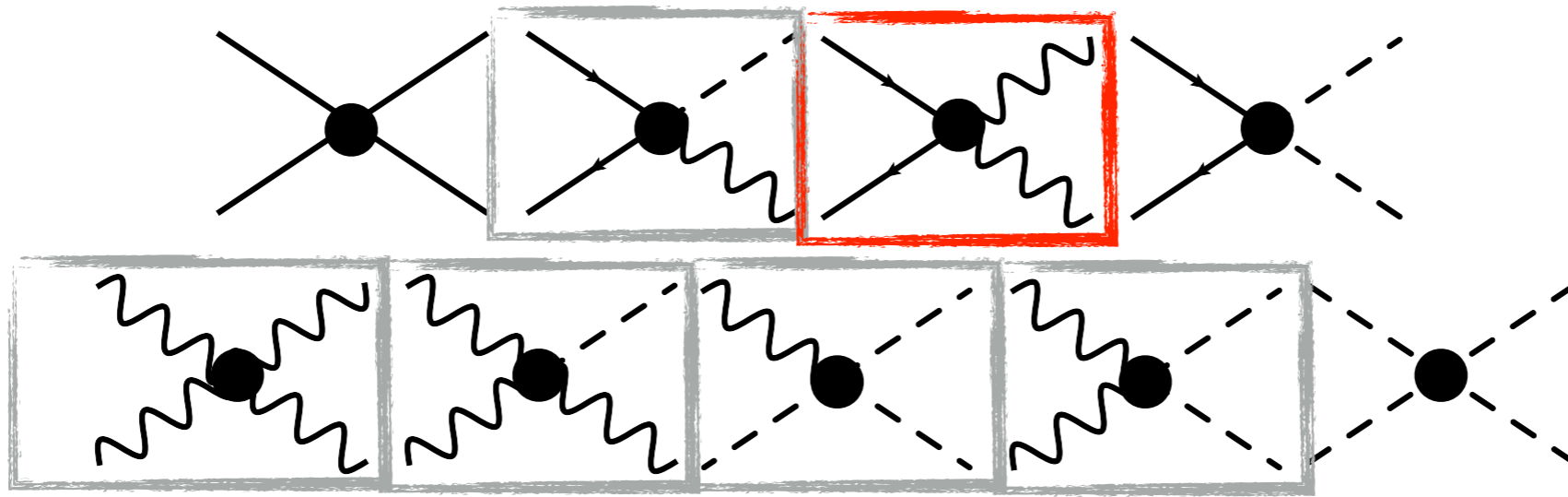
...but small statistics for EW processes!



4. Interference Resurrection



4. Interference Resurrection

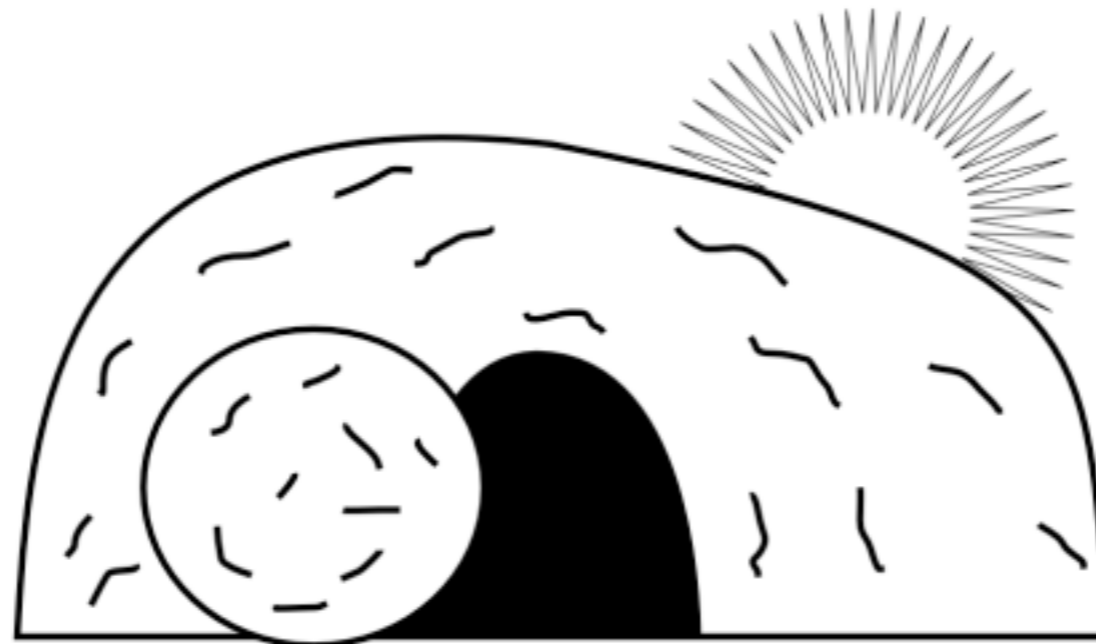


$$\epsilon_{abc} W_{\mu}^{a\nu} W_{\nu\rho}^b W^{c\rho\mu}$$

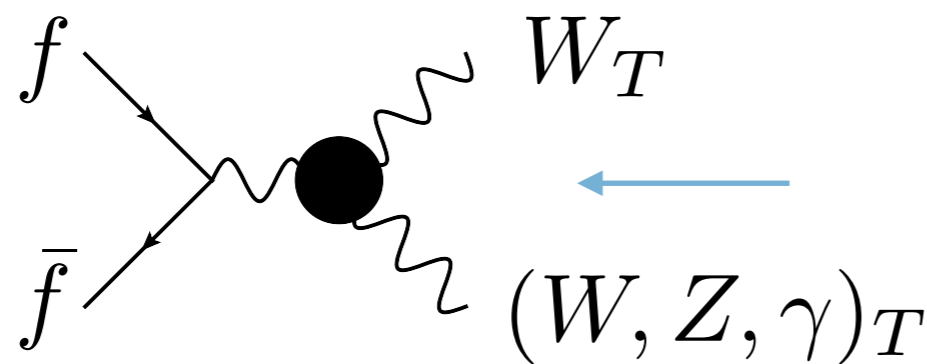
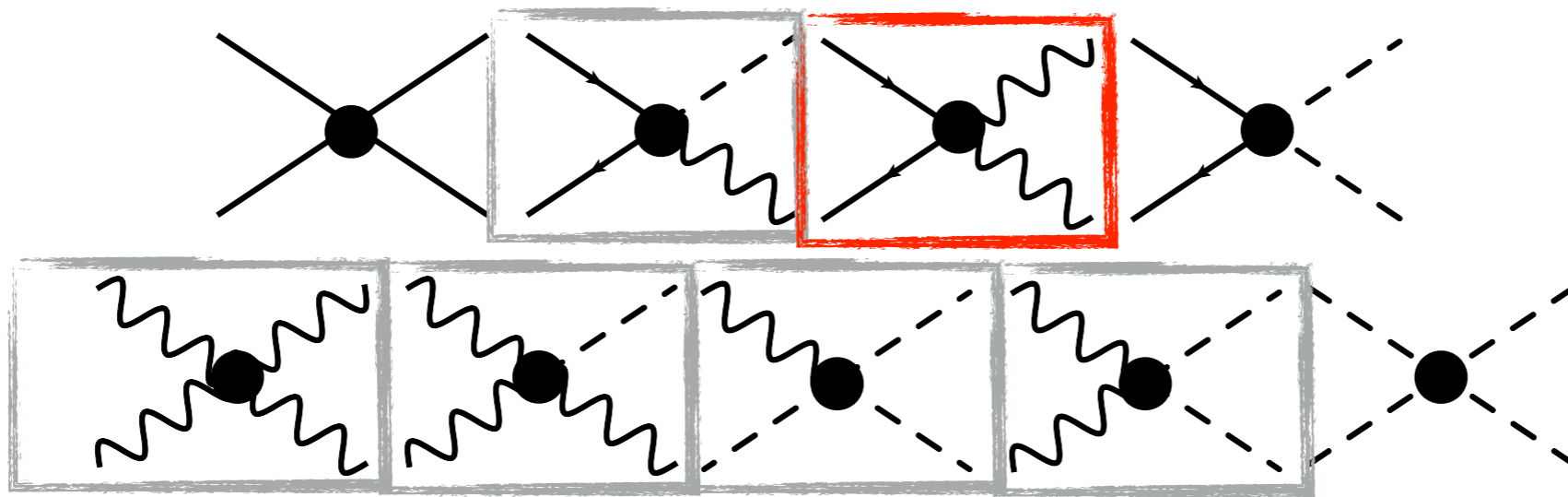
$$\epsilon_{abc} W_{\mu}^{a\nu} W_{\nu\rho}^b \widetilde{W}^{c\rho\mu}$$

CP-even

CP-odd



4. Interference Resurrection

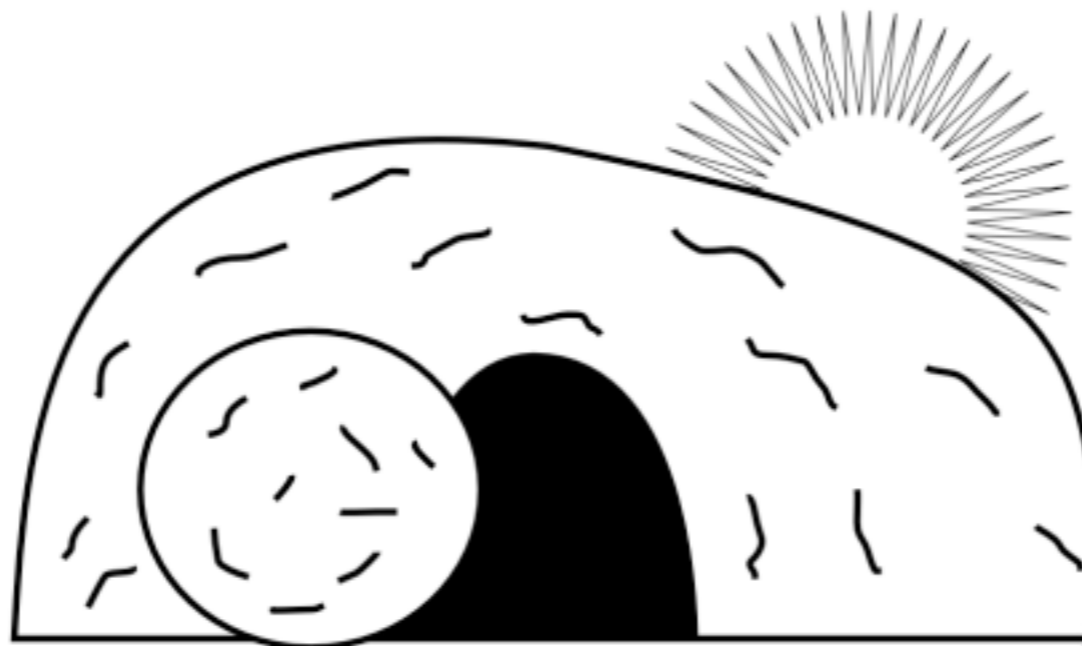


$$\epsilon_{abc} W_{\mu}^{a\nu} W_{\nu\rho}^b W^{c\rho\mu}$$

$$\epsilon_{abc} W_{\mu}^{a\nu} W_{\nu\rho}^b \widetilde{W}^{c\rho\mu}$$

CP-even

CP-odd



...resurrecting ideas from 1986

PROBING THE WEAK BOSON SECTOR IN $e^+e^- \rightarrow W^+W^-$

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Received 23 June 1986

4. Angular correlations for final state fermions

In this section we present the most general angular distributions of the decay products in the process

$$\begin{aligned} e^-(k, \sigma) + e^+(\bar{k}, \bar{\sigma}) &\rightarrow W^-(q, \lambda) + W^+(\bar{q}, \bar{\lambda}), \\ W^-(q, \lambda) &\rightarrow f_1(p_1, \sigma_1) + \bar{f}_2(p_2, \sigma_2), \\ W^+(\bar{q}, \bar{\lambda}) &\rightarrow \bar{f}_3(p_3, \sigma_3) + f_4(p_4, \sigma_4), \end{aligned} \quad (4.1)$$

王百凶

MEASURING THE WWZ COUPLING AT THE TEVATRON

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ABSTRACT

Hadroproduction of weak boson pairs, $W^\pm Z$ and W^+W^- , is studied quantitatively at Tevatron energies. Although the cross sections are expected to be

VOLUME 55, NUMBER 8

PHYSICAL REVIEW

New Standard-Model Test for Future Colliders

M. J. Duncan and G. L. Kane

Randall Physics Laboratory, University of Michigan, Ann Arbor, Michigan 48109

and

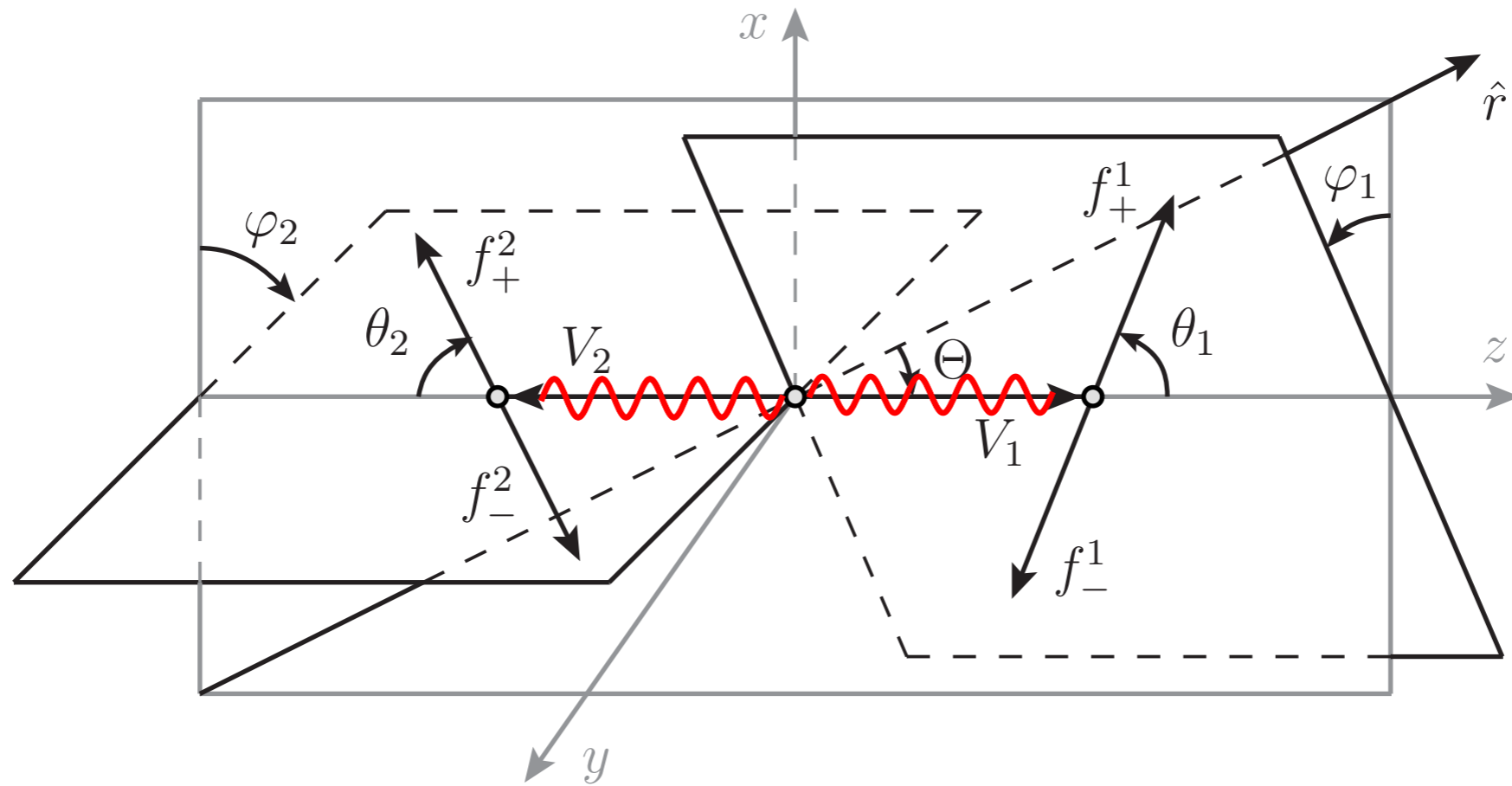
Wayne W. Repko

Department of Physics and Astronomy, Michigan State University, East Lansing, Michigan 48824

(Received 28 May 1985)

We point out that for W -pair production from e^+e^- or $q\bar{q}$ beams, the correlation between the decay planes of the W 's is numerically negligible (at tree level) in the standard model. This occurs for the production of WW via intermediate fermions or gauge bosons and independently for WW produced via Higgs-boson exchange, and is not restricted to two-body diagrams. In general, the correlation need not vanish, and so a nonzero correlation would be a clear signal of physics beyond the standard model. Definition of the W -decay plane should be fairly straightforward from the decay $W \rightarrow ff$ and, hence, the correlation measurement is experimentally feasible.

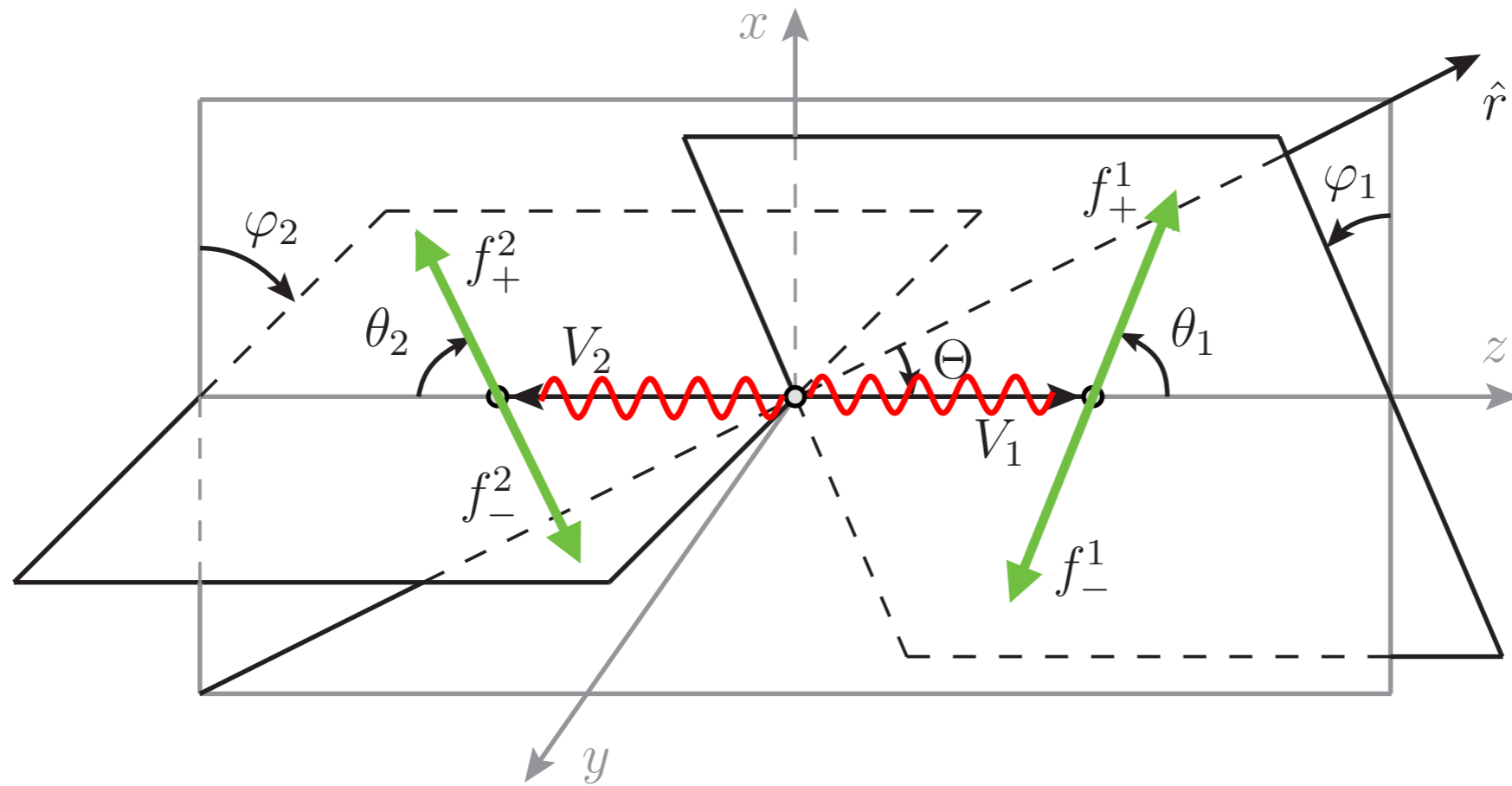
Differential measurements WW, WZ



$V_{1,2}$: Helicity $\pm\mp/\pm\pm$ in SM/BSM

- ▶ Quantum mechanically **different**, **no** interference

Differential measurements WW, WZ



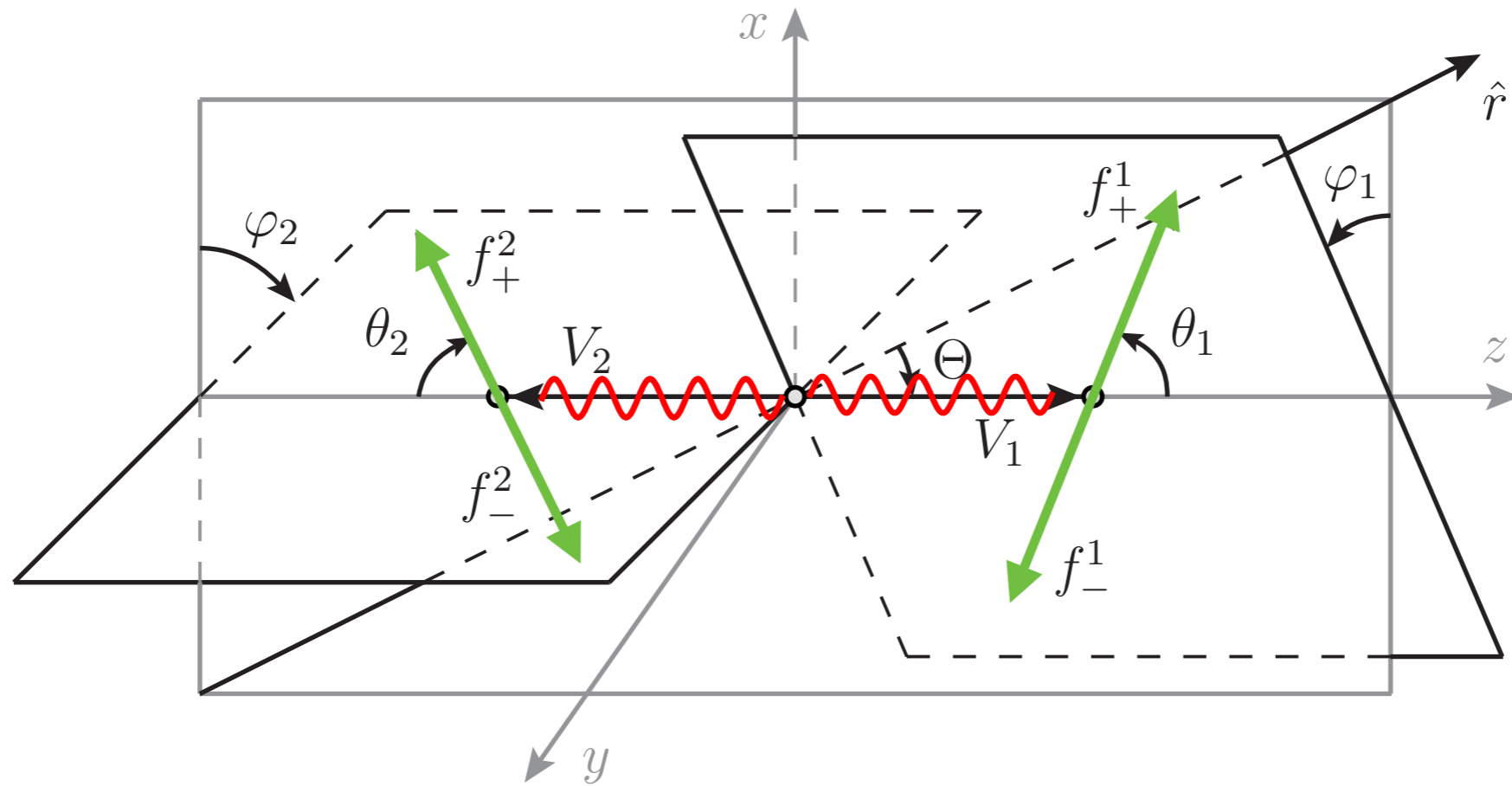
$V_{1,2}$: Helicity $\pm\mp/\pm\pm$ in SM/BSM

- ▶ Quantum mechanically **different**, **no** interference

$f_{(1,3)} f_{(2,4)}$: Helicity $+1/2 -1/2$ in SM **and** in BSM

- ▶ QM **same**, interference possible

Differential measurements WW, WZ

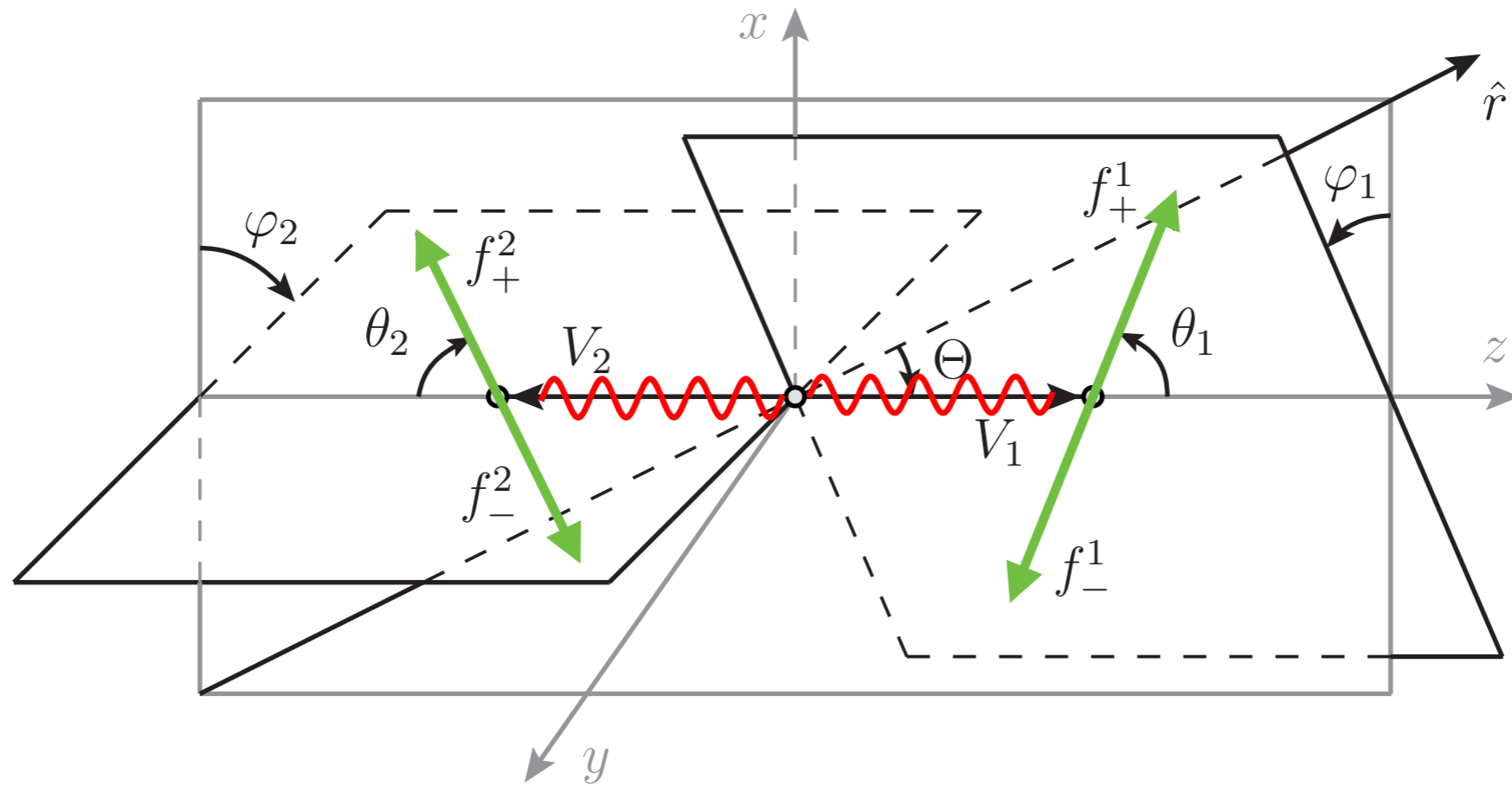


$$Int^{CP} \propto \mathcal{A}_{\mathbf{h}}^{SM (+1, -1)} \mathcal{A}_{\mathbf{h}'}^{BSM+ (+1, +1)} \cos [\Delta \mathbf{h} \cdot \boldsymbol{\varphi}]$$

$(h_1 - h'_1, h_2 - h'_2)$
 (φ_1, φ_2)

$$Int^{CP} \propto \mathcal{A}_{\mathbf{h}}^{SM} \mathcal{A}_{\mathbf{h}'}^{BSM-} \sin [\Delta \mathbf{h} \cdot \boldsymbol{\varphi}]$$

Differential measurements WW, WZ



$$Int^{CP} \propto \mathcal{A}_{\mathbf{h}}^{SM (+1, -1)} \mathcal{A}_{\mathbf{h}'}^{BSM+ (+1, +1)} \cos [\Delta \mathbf{h} \cdot \boldsymbol{\varphi}]$$

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 (φ_1, φ_2)

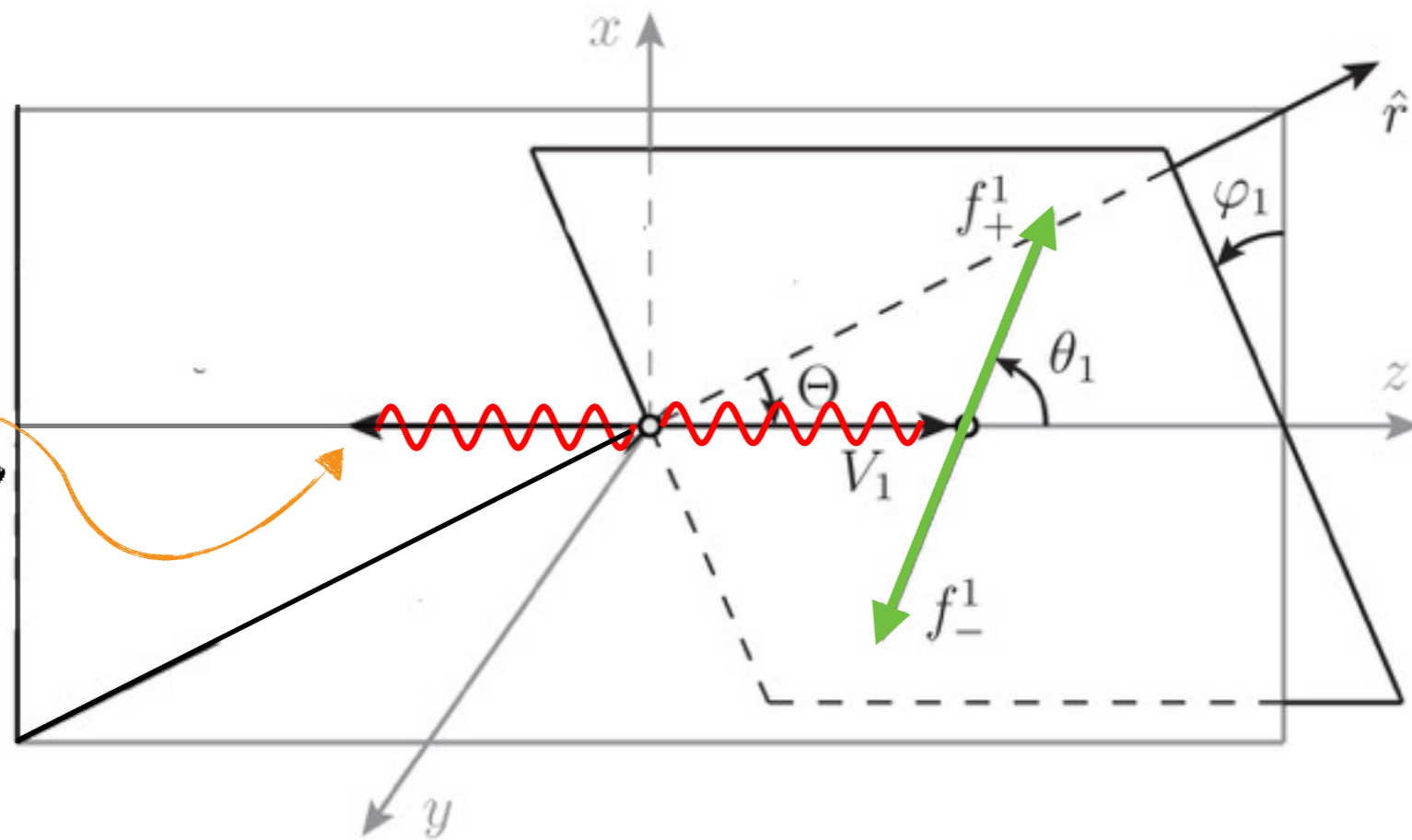
$$Int^{CP} \propto \mathcal{A}_{\mathbf{h}}^{SM} \mathcal{A}_{\mathbf{h}'}^{BSM-} \sin [\Delta \mathbf{h} \cdot \boldsymbol{\varphi}]$$

► Cancels when integrated over $\varphi \in [-\pi, \pi]$

Differential measurements $W\gamma$

$W\gamma$

No (leptonic)
Branching Ratio



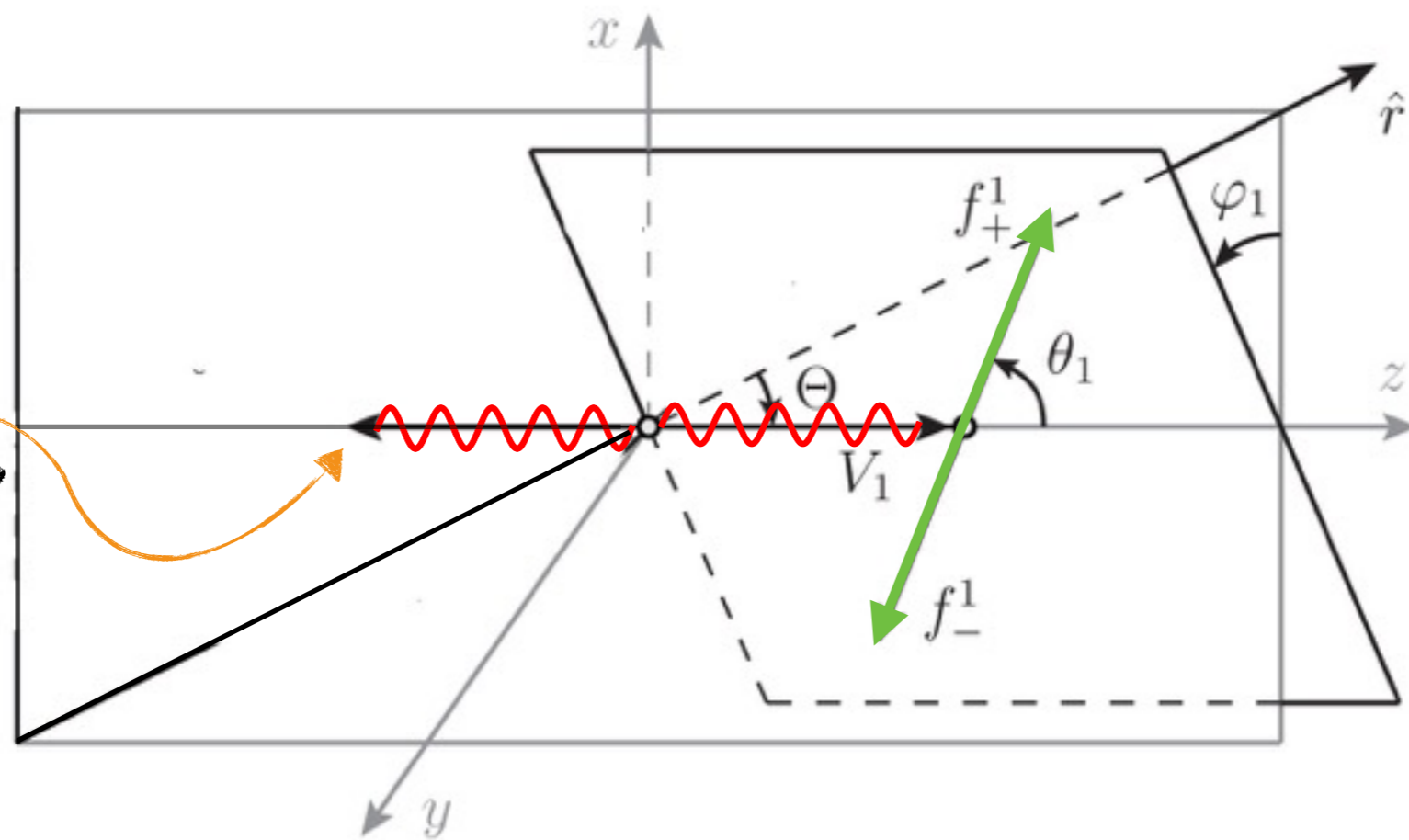
$$Int^{CP} = 2g^2 \sin^2 \theta \mathcal{A}_{++}^{BSM+} [\mathcal{A}_{-++}^{SM} + \mathcal{A}_{+-}^{SM}] \cos 2\varphi,$$

$$Int^{CP} = 2ig^2 \sin^2 \theta \mathcal{A}_{++}^{BSM-} [\mathcal{A}_{-+-}^{SM} - \mathcal{A}_{+-}^{SM}] \sin 2\varphi$$

Differential measurements $W\gamma$

$W\gamma$

No (leptonic)
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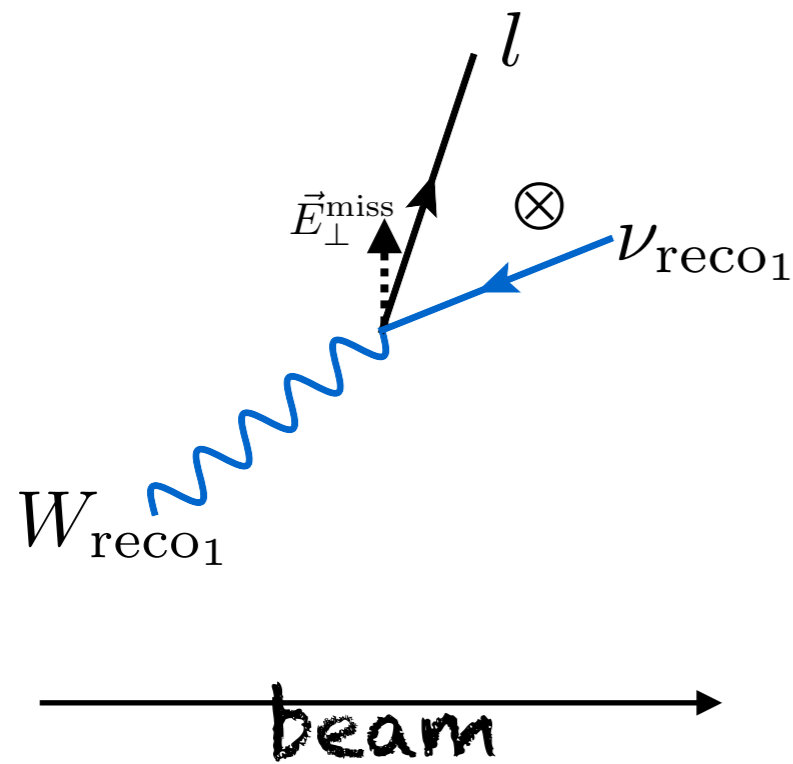
$$Int^{CP} = 2ig^2 \sin^2 \theta \mathcal{A}_{++}^{BSM-} [\mathcal{A}_{-+}^{SM} - \mathcal{A}_{+-}^{SM}] \sin 2\varphi$$

Differential azimuthal distributions = SM-BSM interference

Azimuthal Angle... in reality

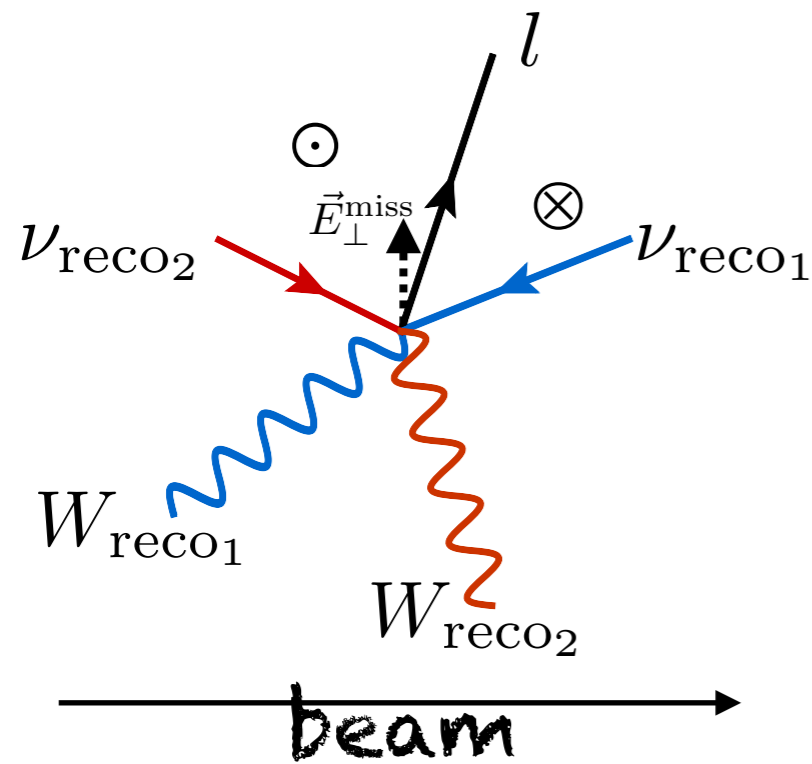
Neutrino: from missing energy + reconstruct W mass

1)



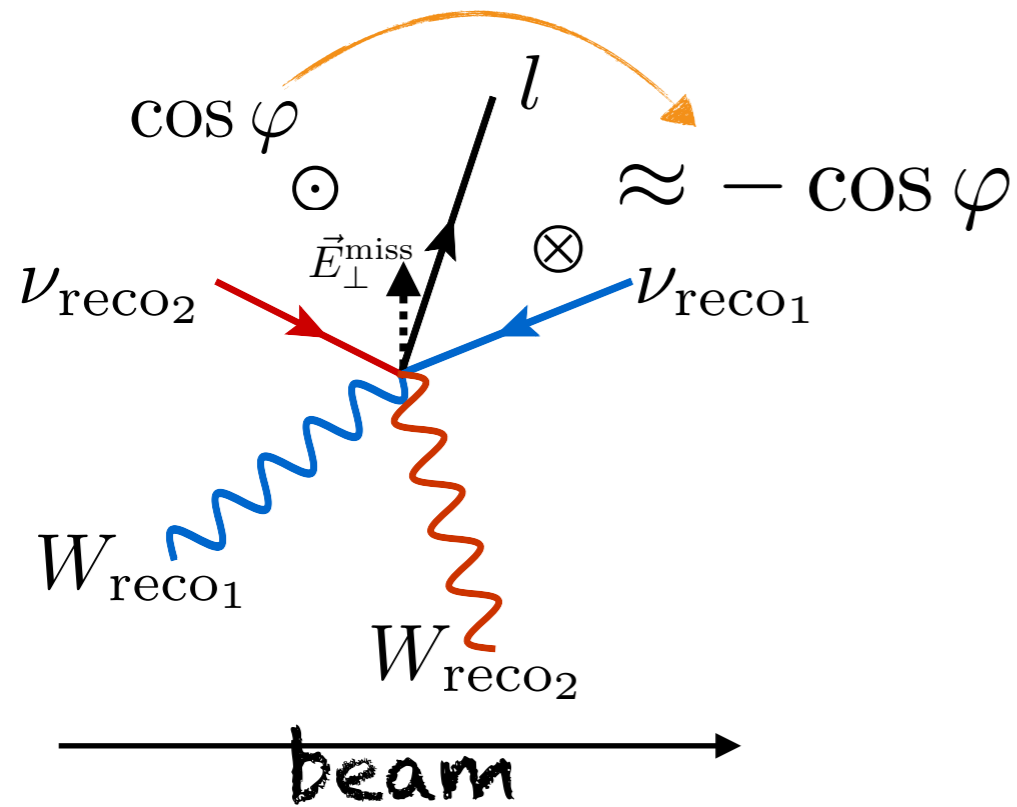
Azimuthal Angle... in reality

Neutrino: from missing energy + reconstruct W mass



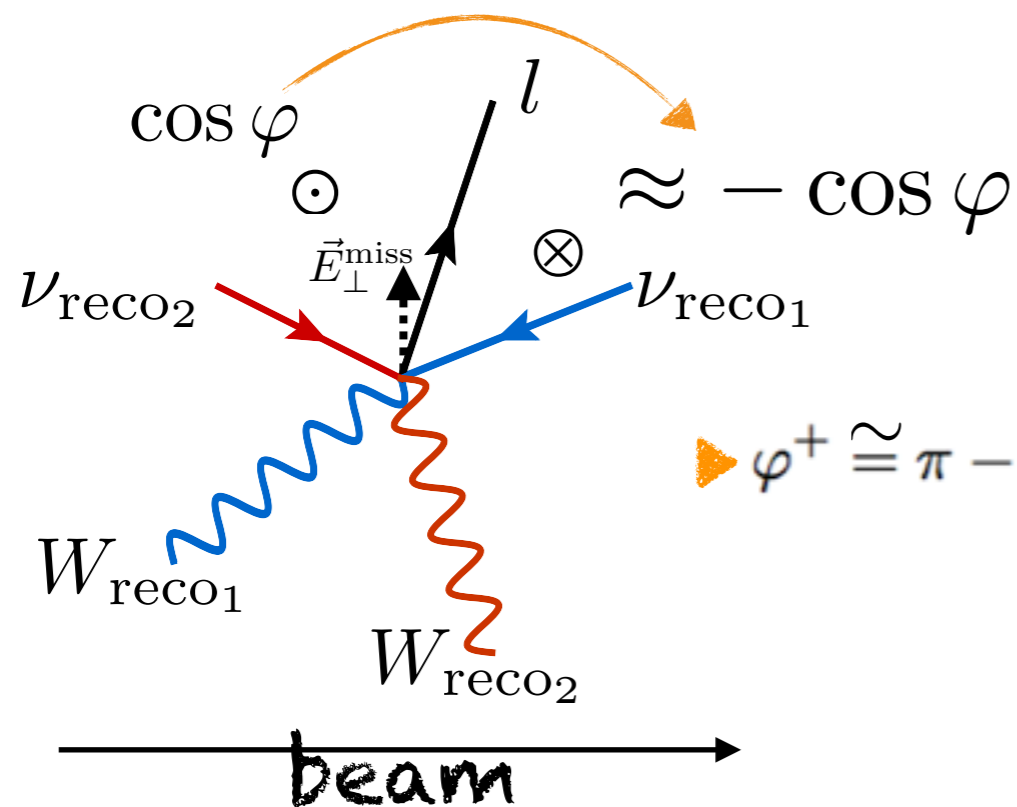
Azimuthal Angle... in reality

Neutrino: from missing energy + reconstruct W mass



Azimuthal Angle... in reality

Neutrino: from missing energy + reconstruct W mass

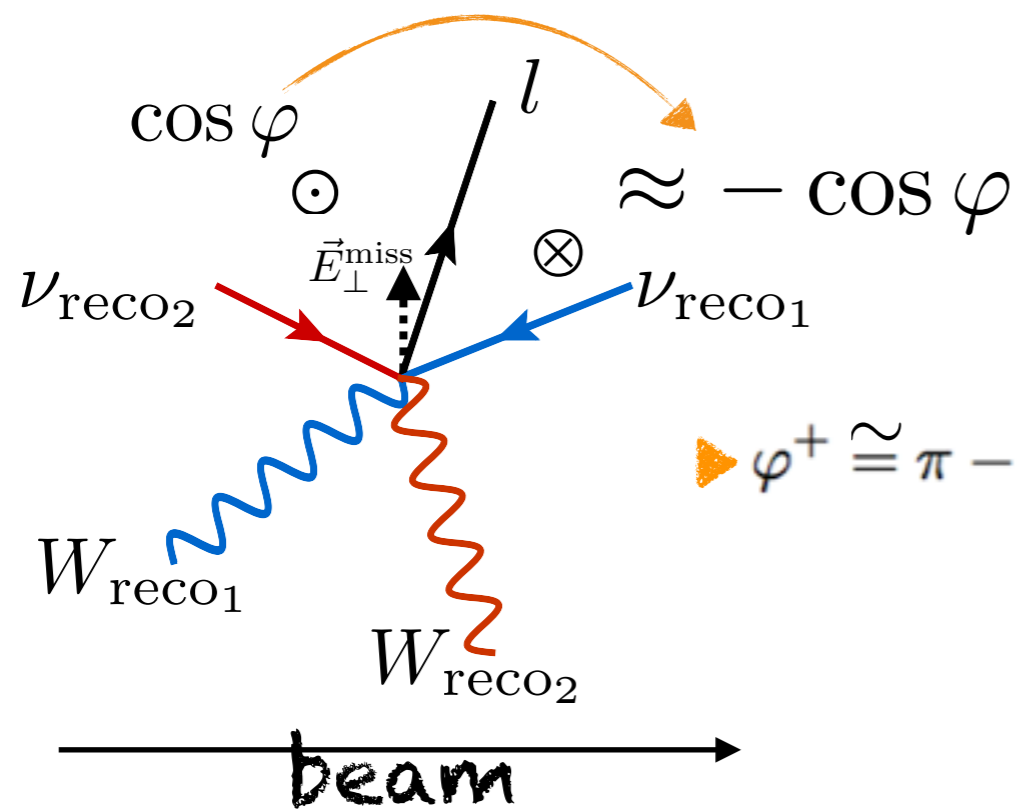


φ_{reco}

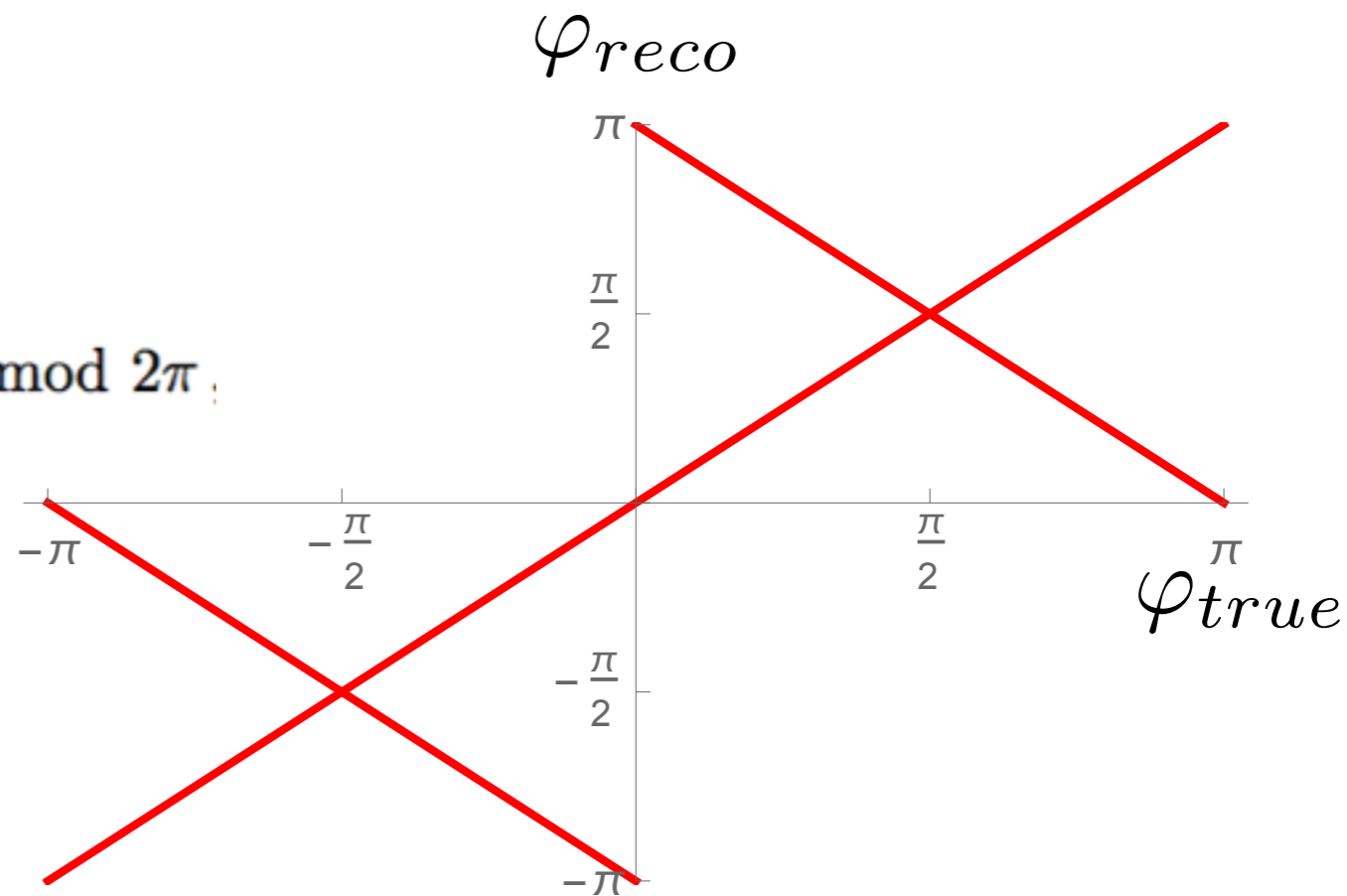
φ_{true}

Azimuthal Angle... in reality

Neutrino: from missing energy + reconstruct W mass

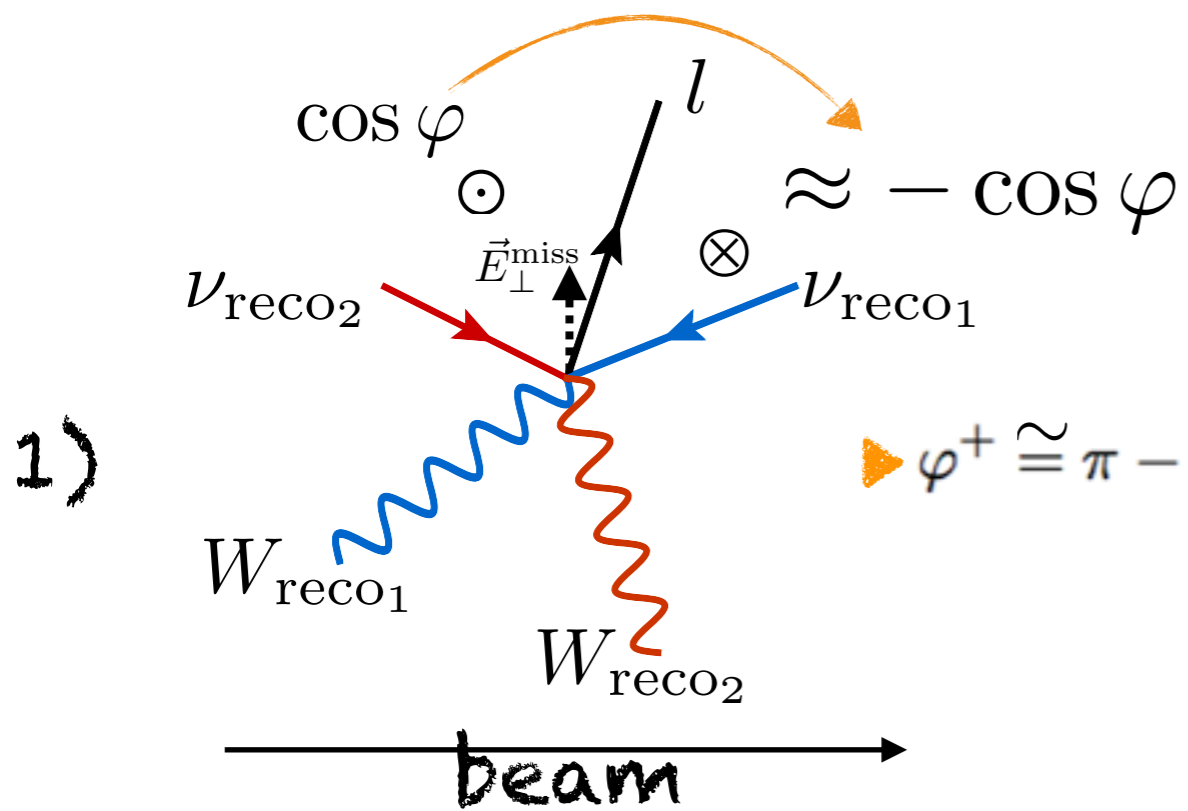


▶ $\varphi^+ \approx \pi - \varphi^- \pmod{2\pi}$

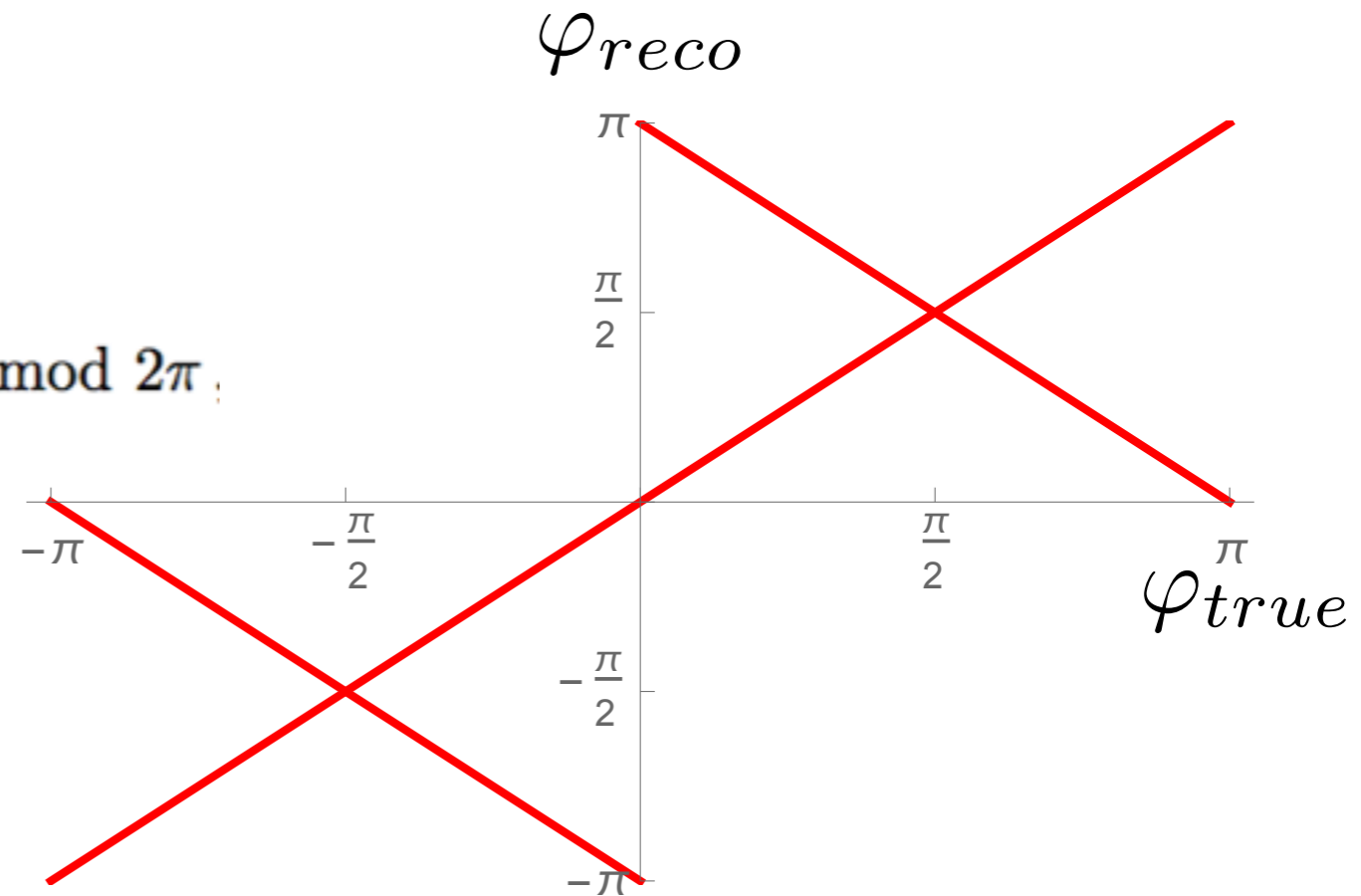


Azimuthal Angle... in reality

Neutrino: from missing energy + reconstruct W mass



► $\varphi^+ \approx \pi - \varphi^- \pmod{2\pi}$



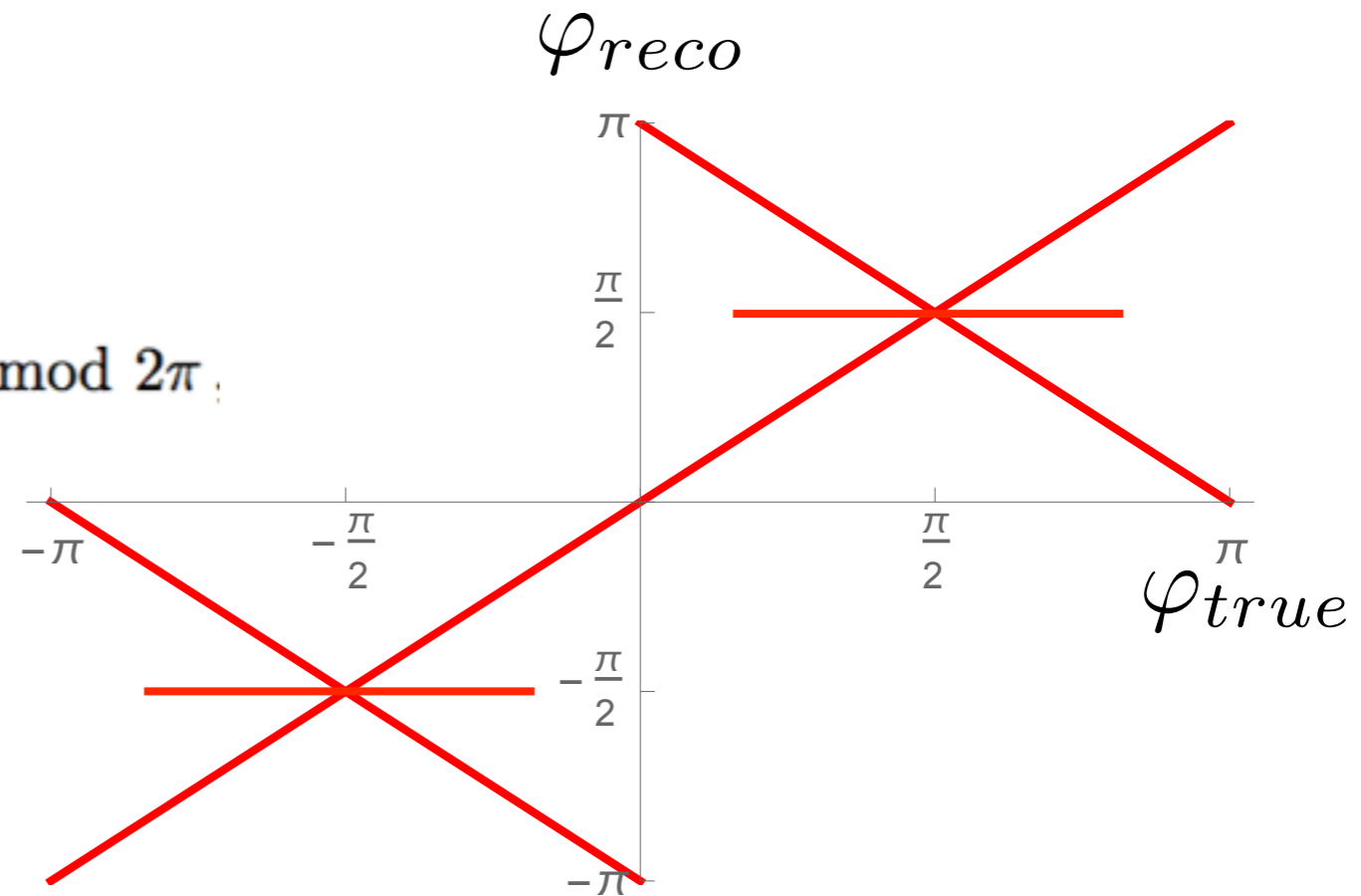
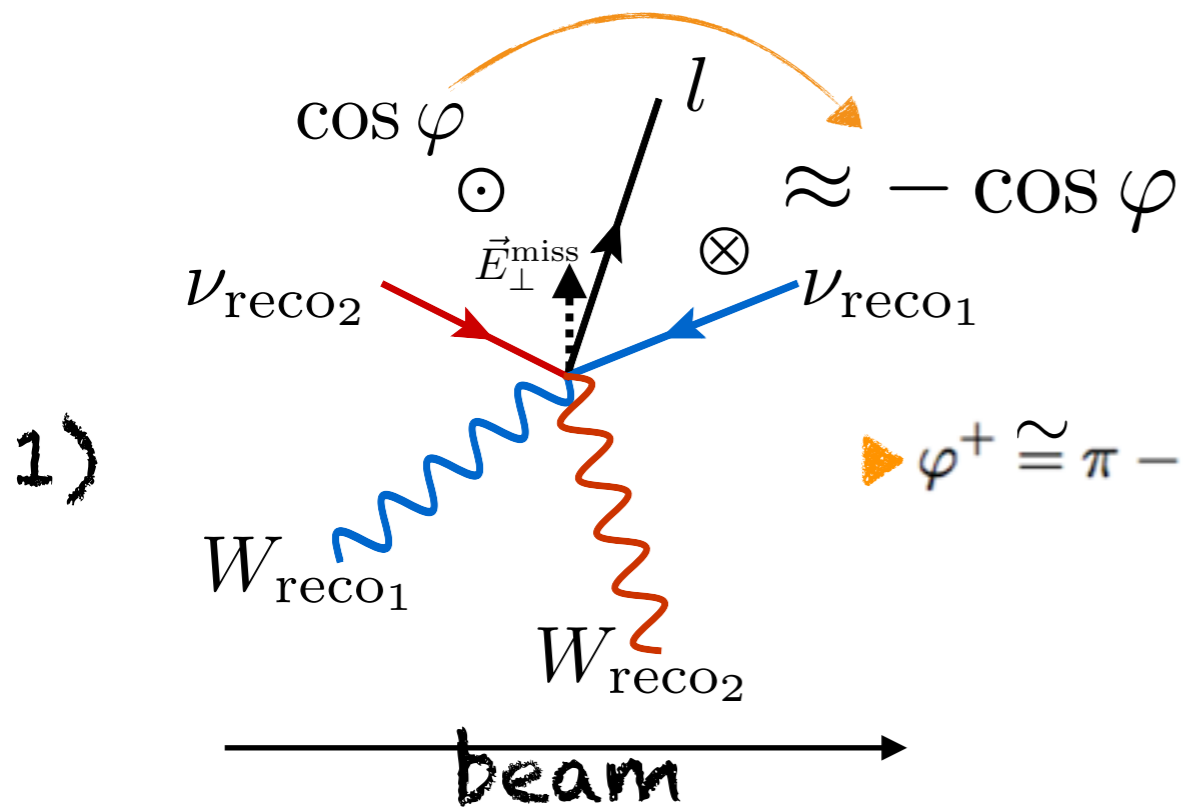
2) Some events: $m_{\perp}^2 > m_W^2$
(off-shell, exp.error)

reconstructed as $m_{\text{inv}}^2 = m_W^2$

► $\varphi = \pi/2$ or $\varphi = -\pi/2$.

Azimuthal Angle... in reality

Neutrino: from missing energy + reconstruct W mass



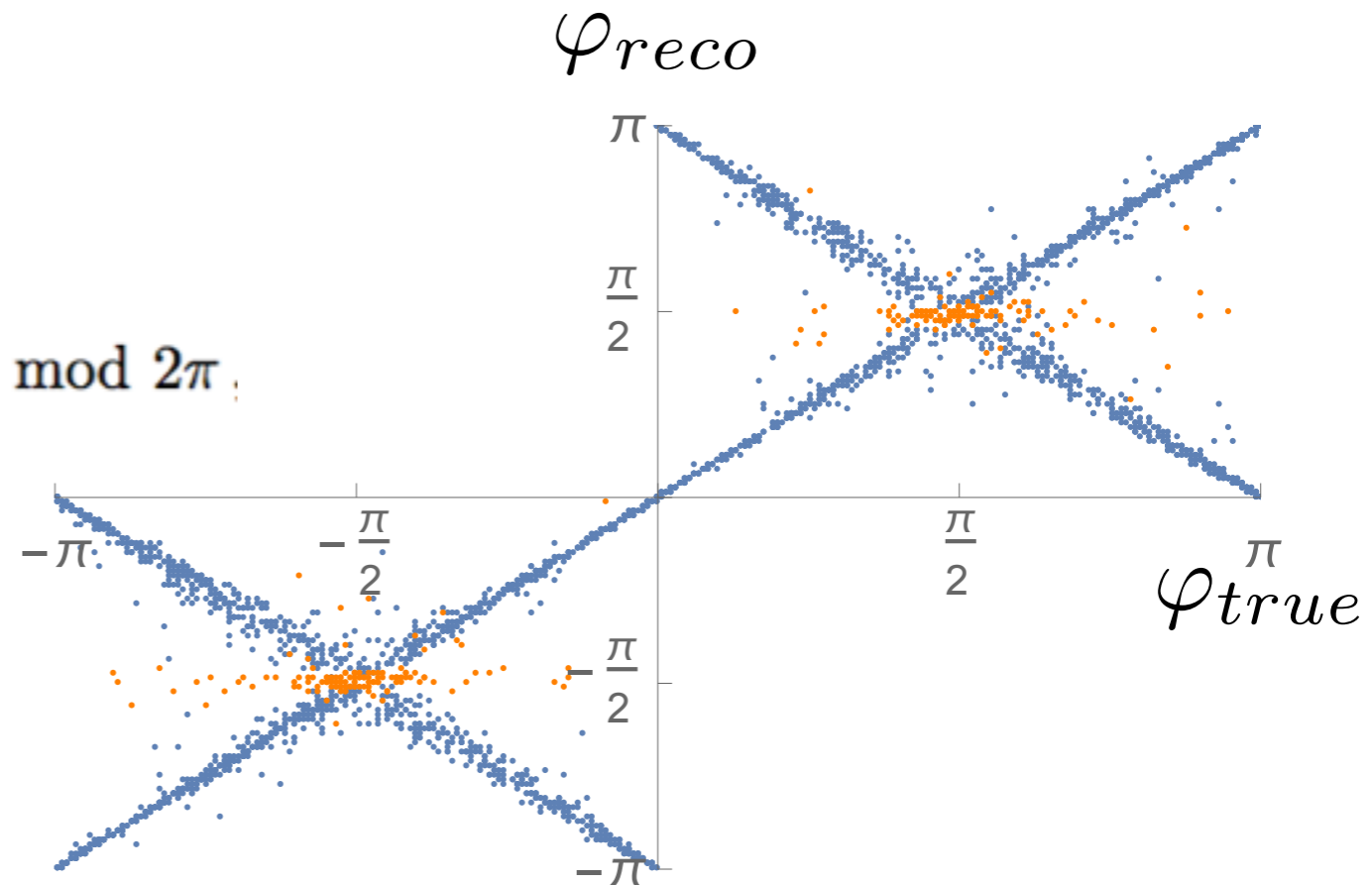
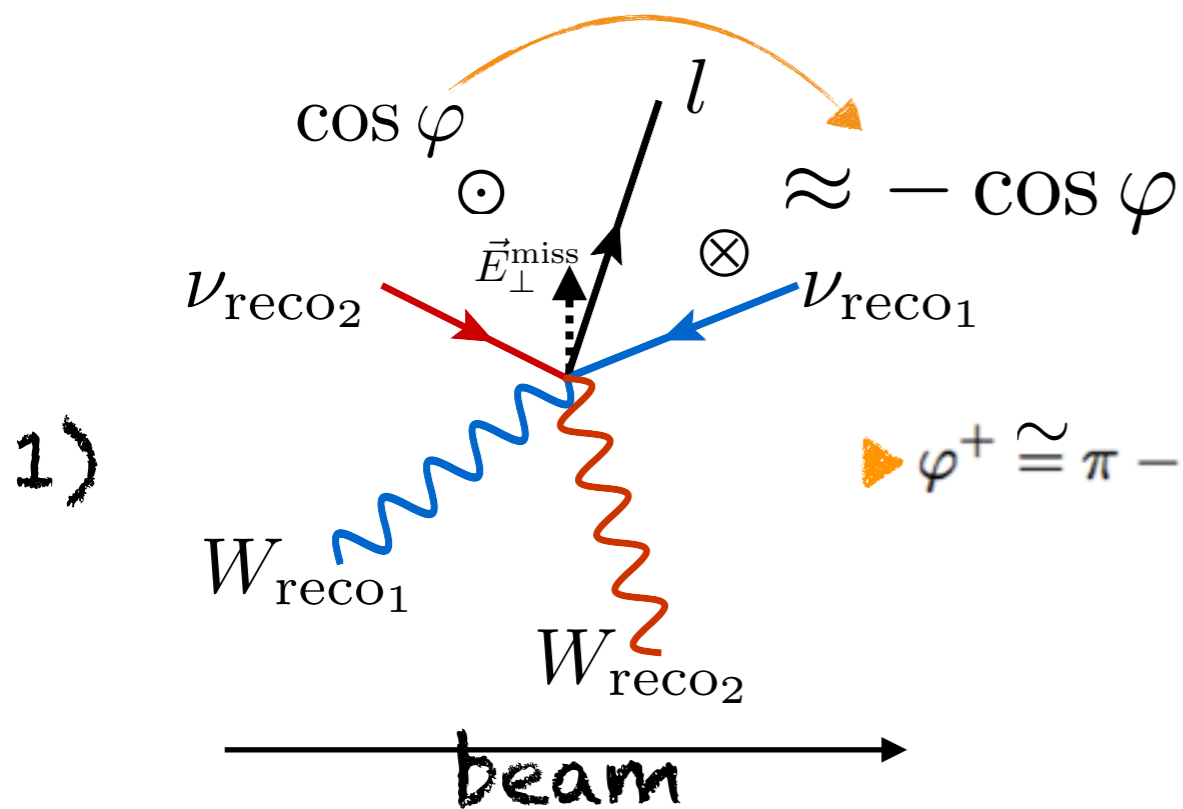
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Azimuthal Angle... in reality

Neutrino: from missing energy + reconstruct W mass



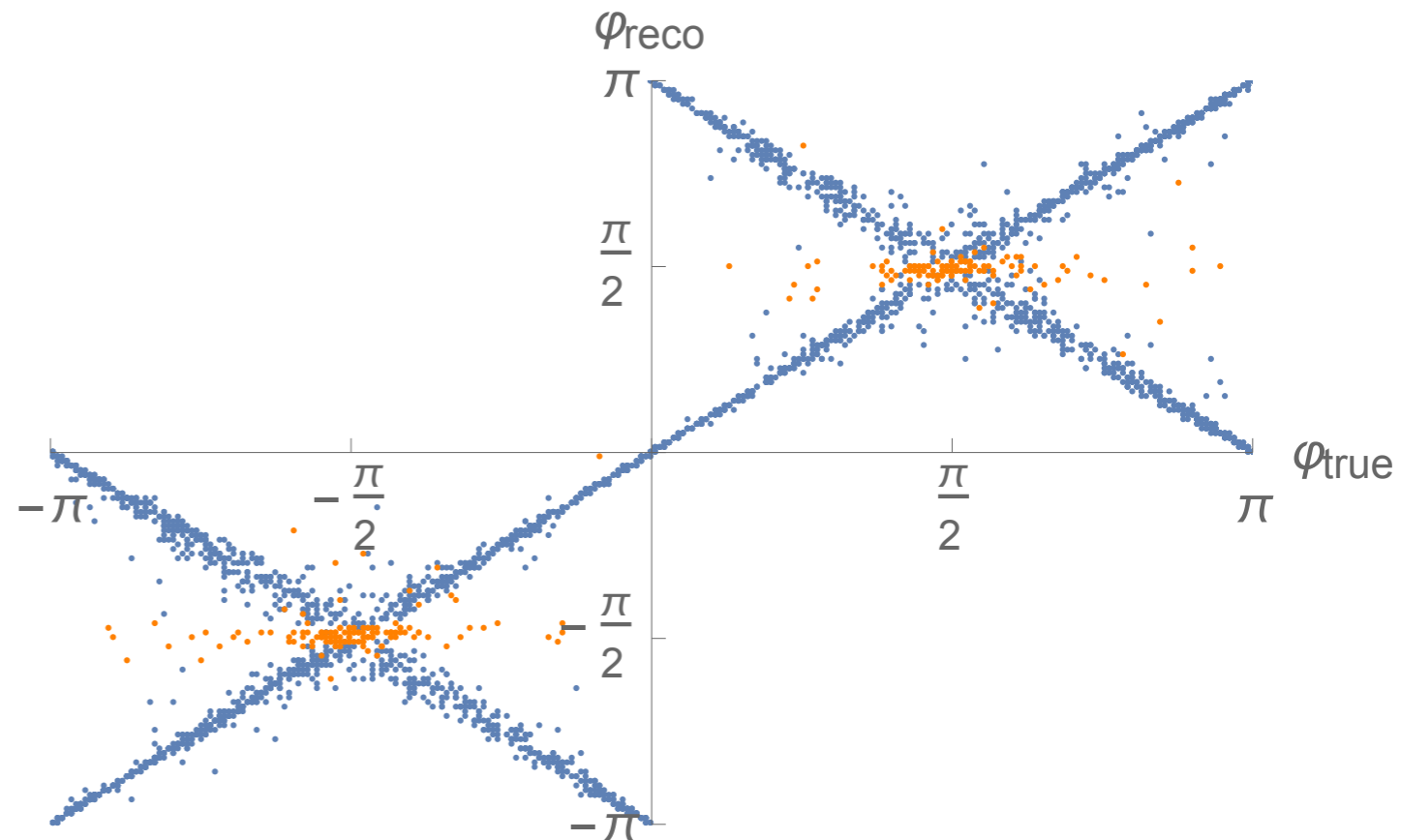
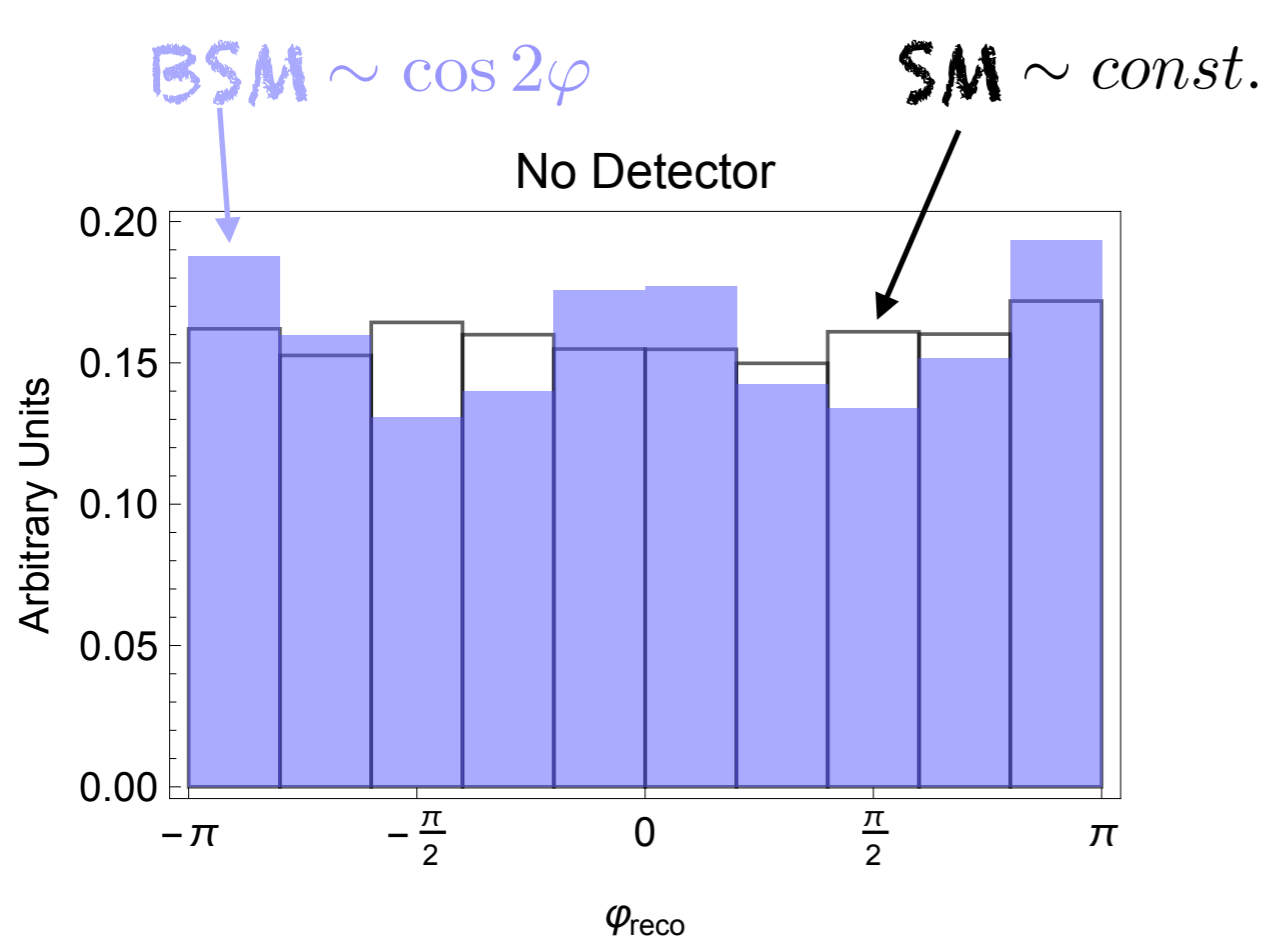
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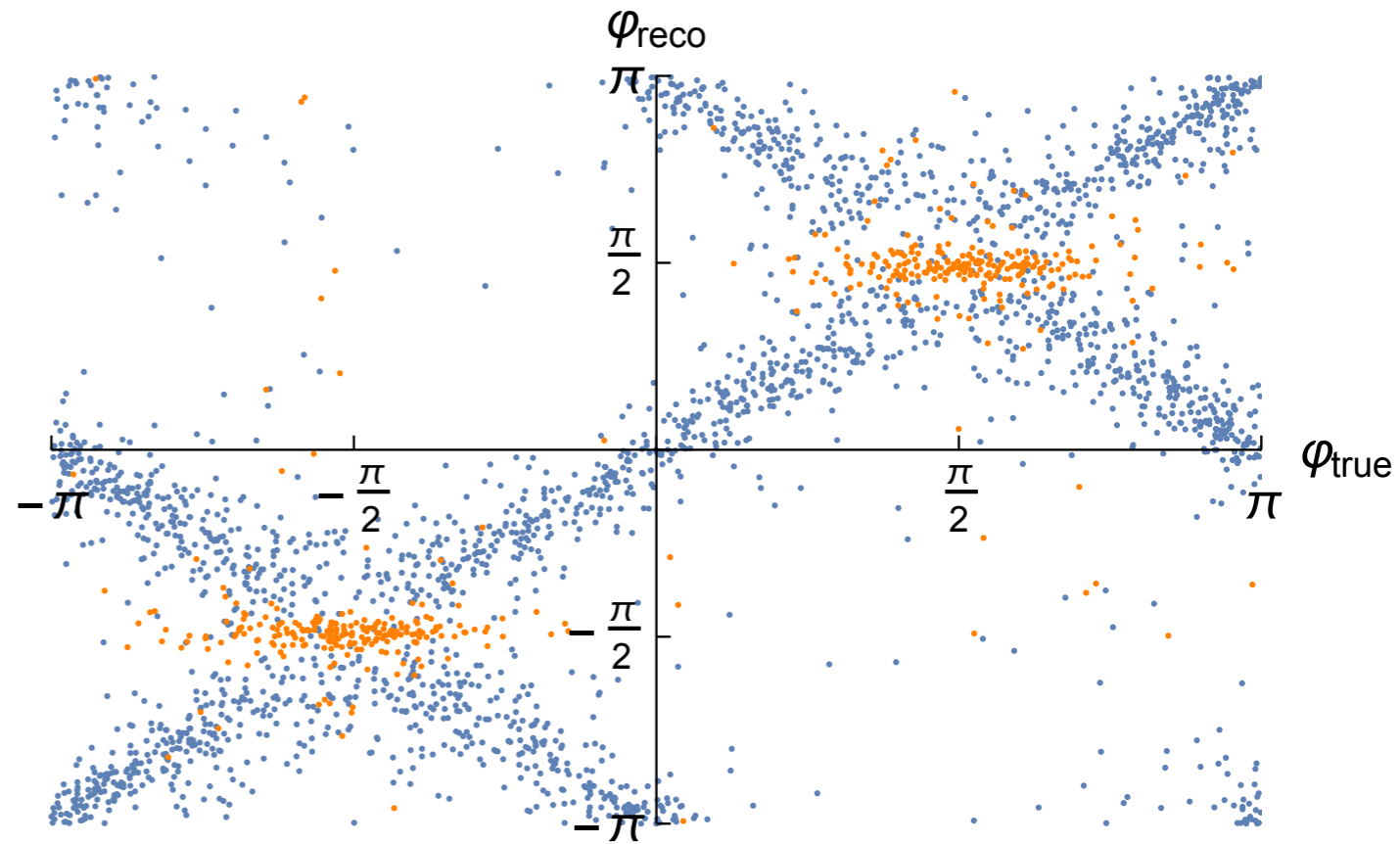
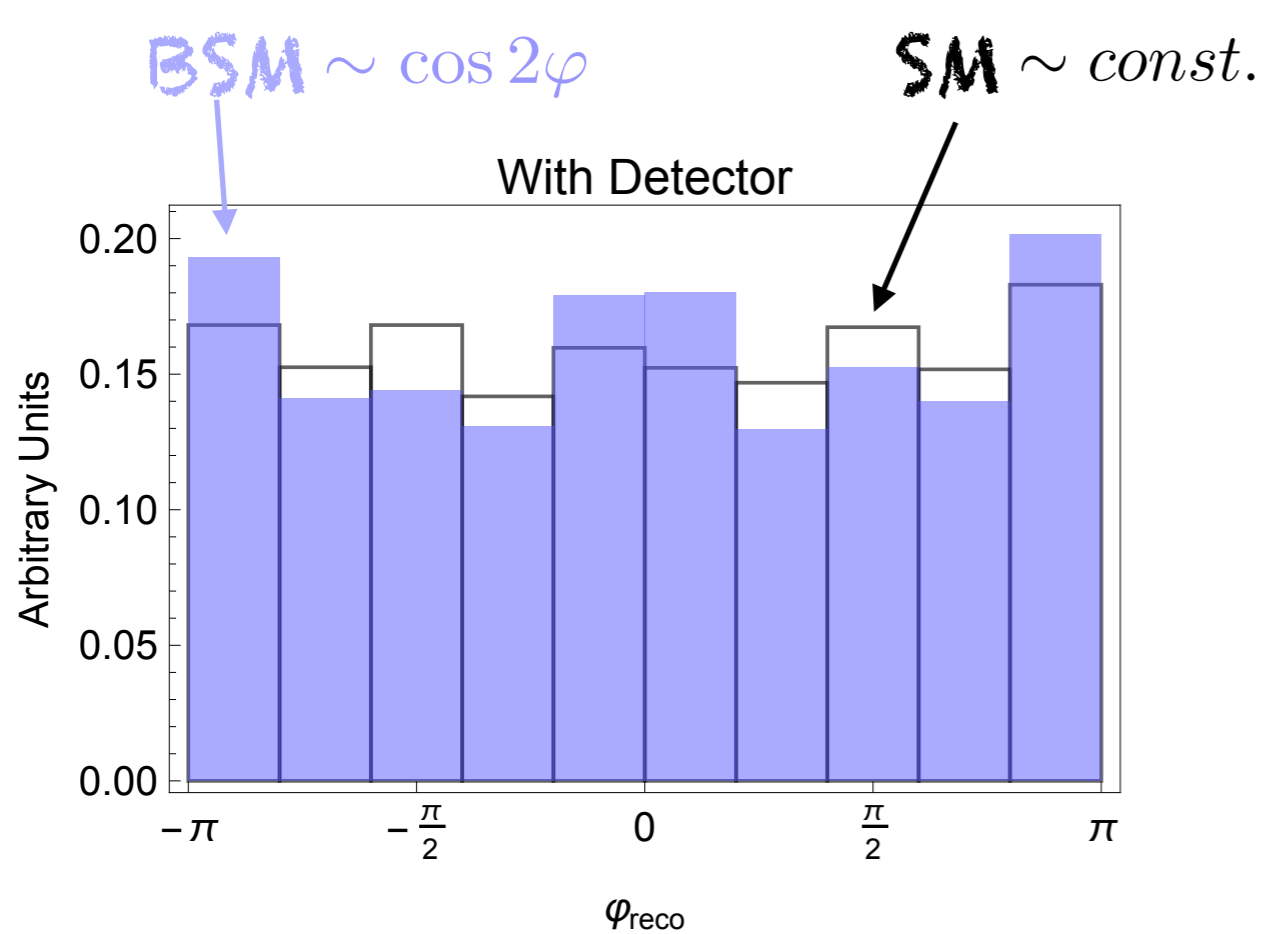
Azimuthal Angle... in reality

Neutrino: from missing energy + reconstruct W mass



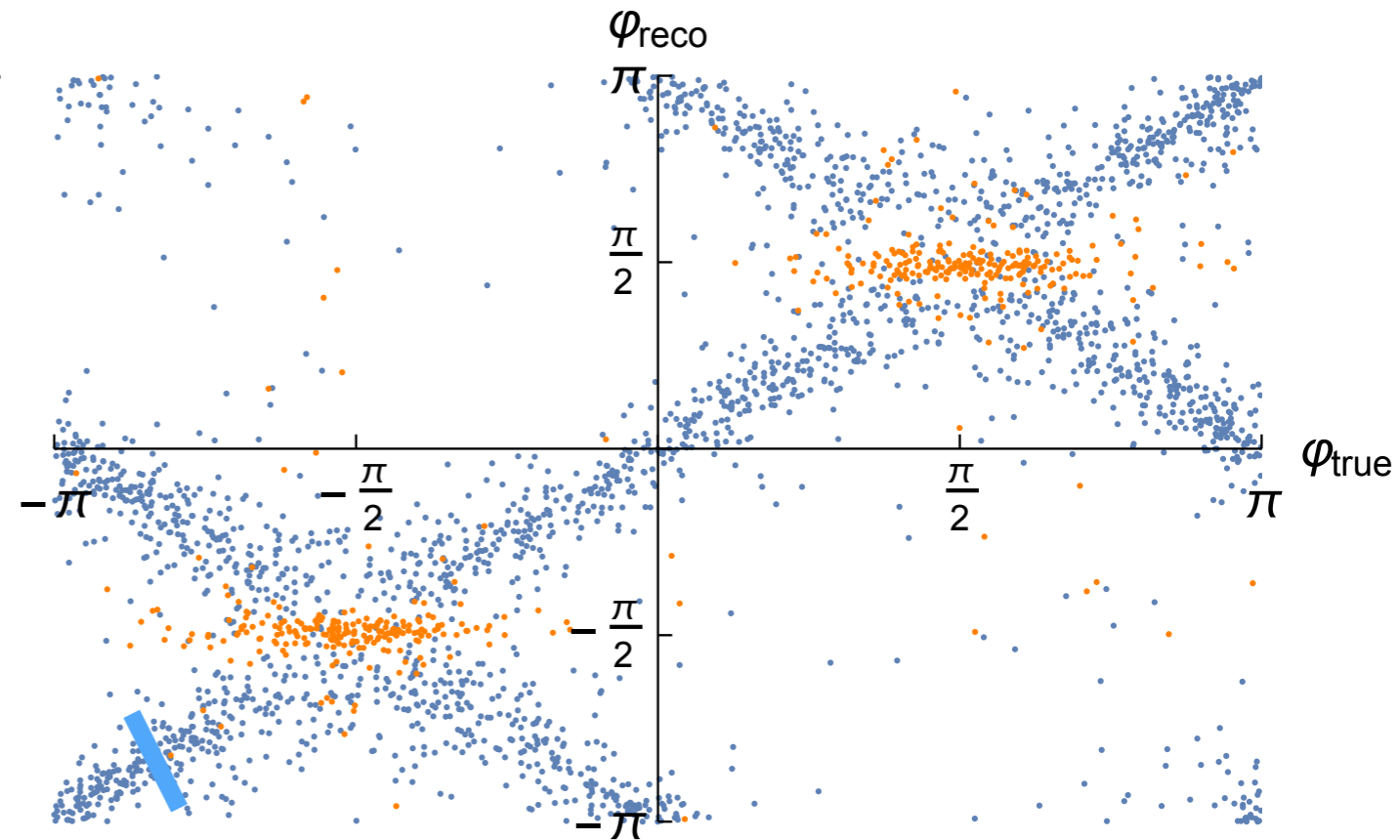
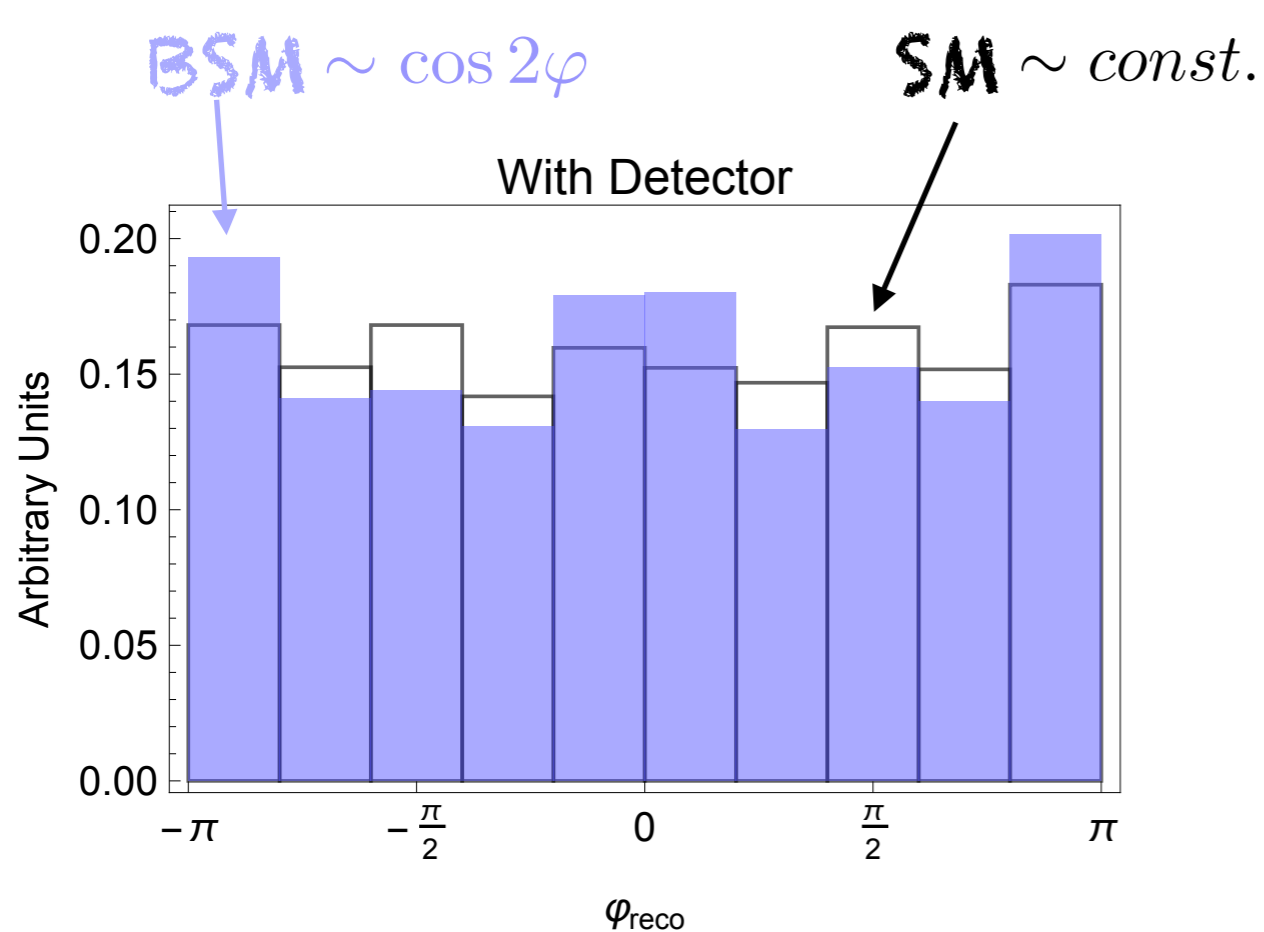
Azimuthal Angle... **more** in reality

Neutrino: from missing energy + reconstruct W mass
With (DELPHES) detector simulation



Azimuthal Angle... **more** in reality

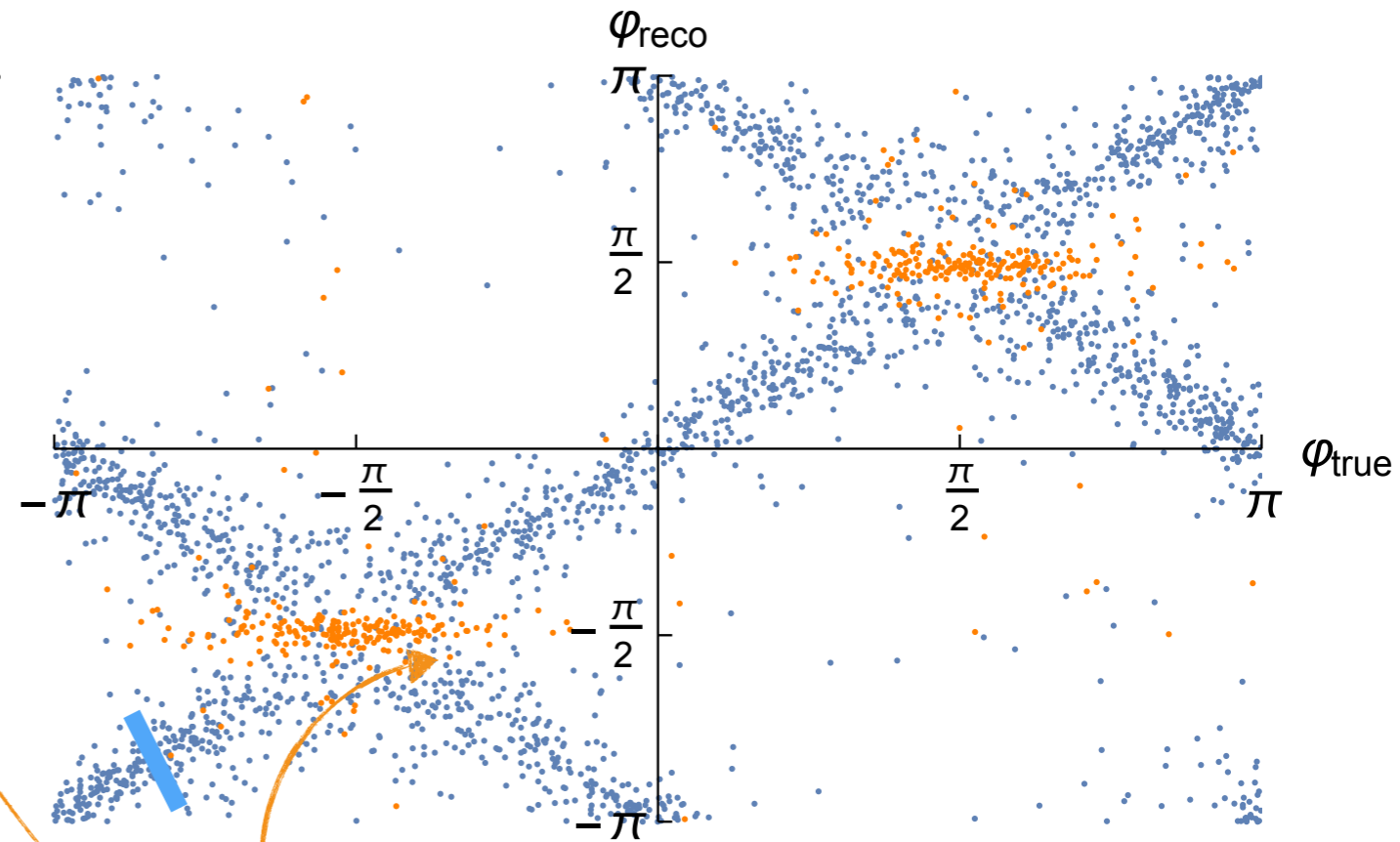
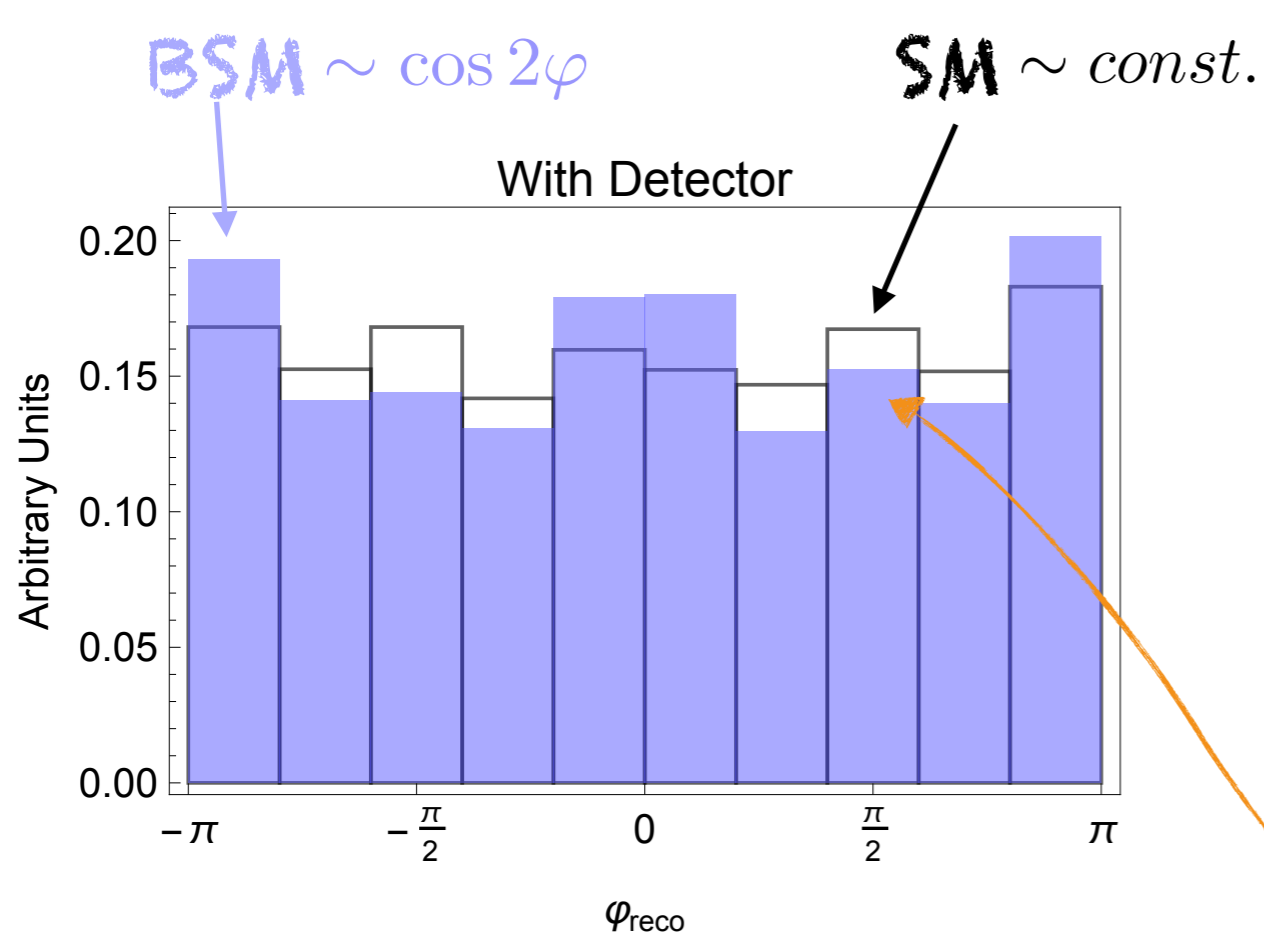
Neutrino: from missing energy + reconstruct W mass
With (DELPHES) detector simulation



Spread under control

Azimuthal Angle... **more** in reality

Neutrino: from missing energy + reconstruct W mass
With (DELPHES) detector simulation

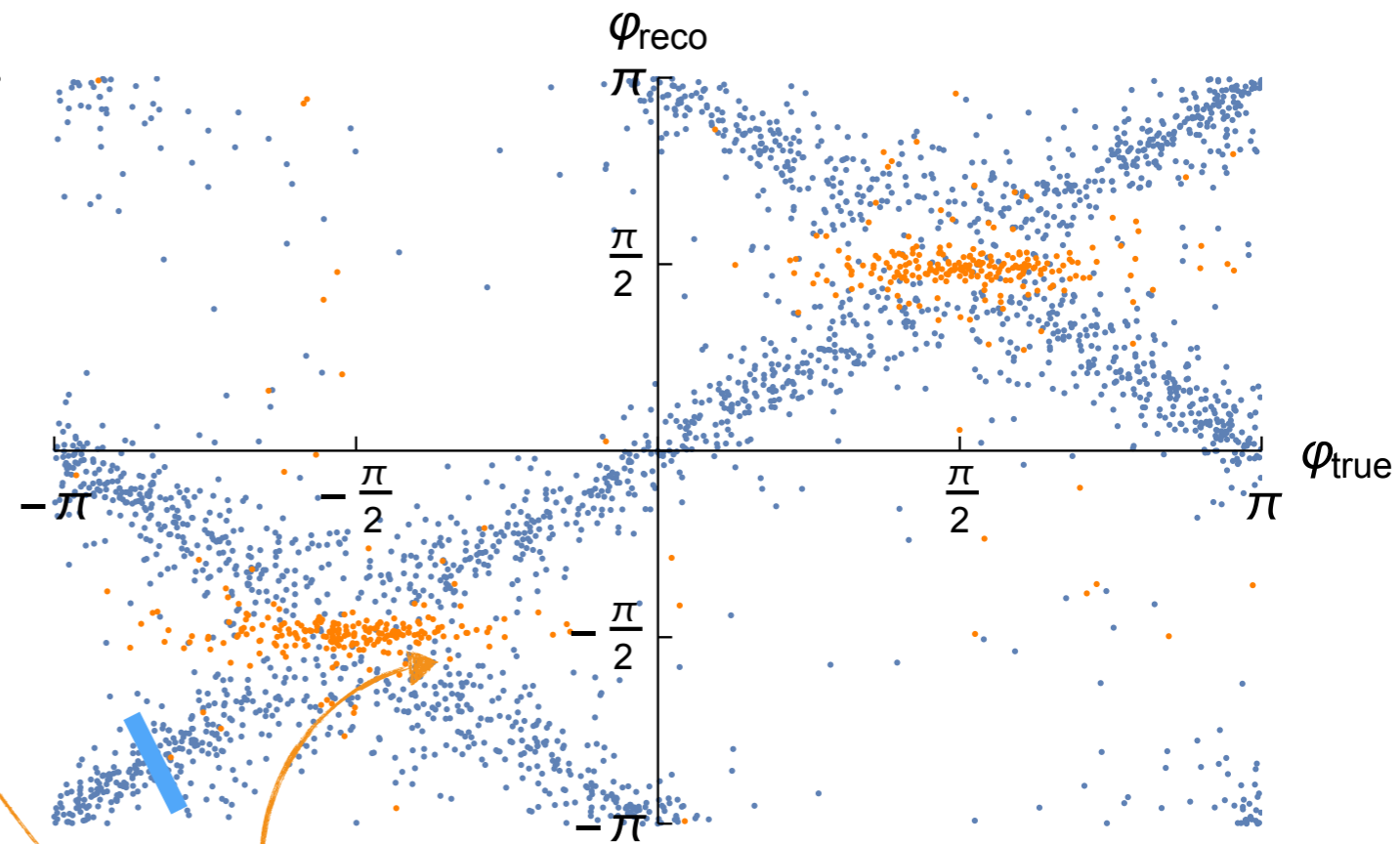
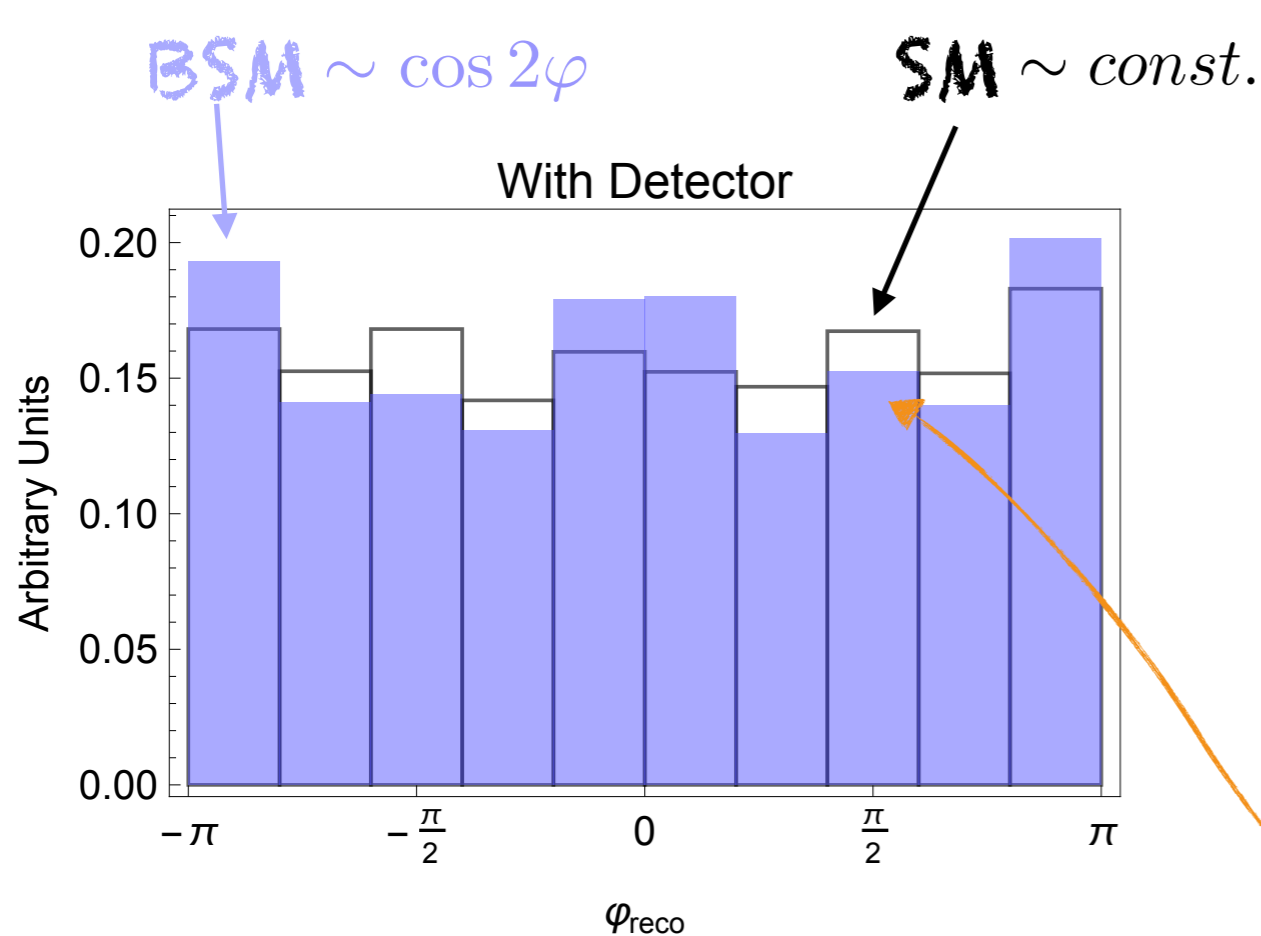


Spread under control

More events with $m_{\perp}^2 > m_W^2$

Azimuthal Angle... **more** in reality

Neutrino: from missing energy + reconstruct W mass
with (DELPHES) detector simulation



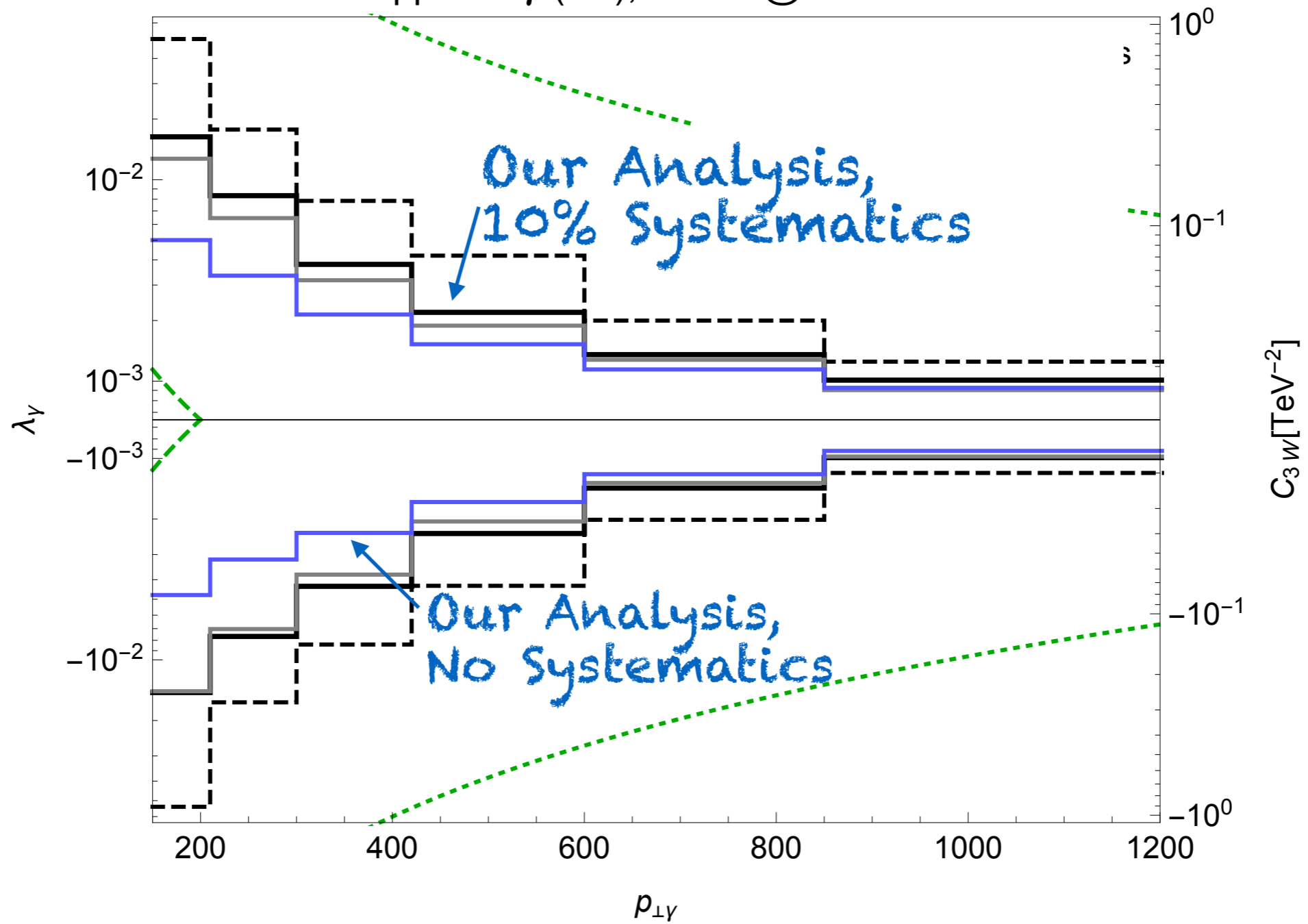
Spread under control

More events with $m_{\perp}^2 > m_W^2$

► Resurrection is real

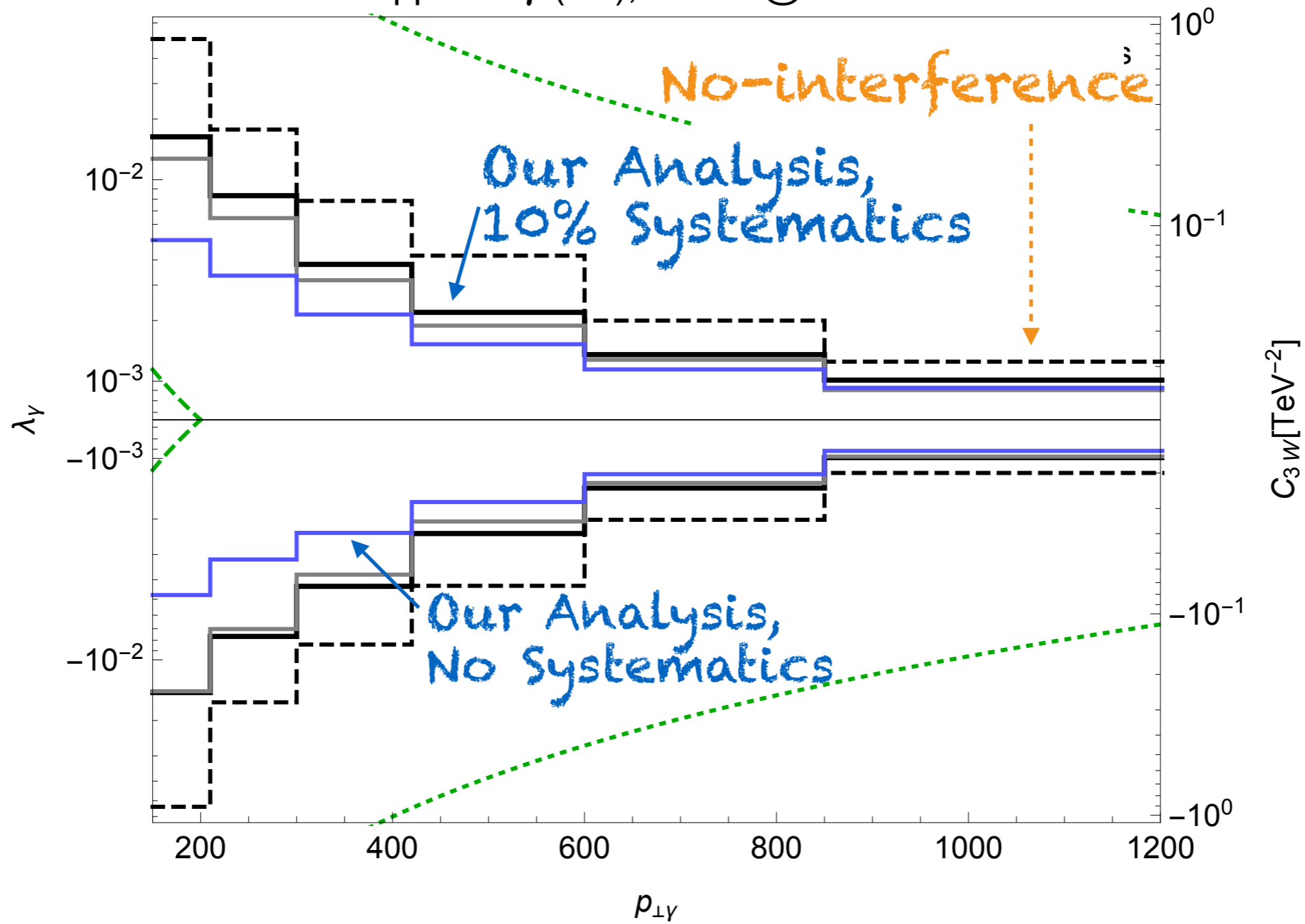
Results

$pp \rightarrow W\gamma$ (LO), 3ab^{-1} @14 TeV



Results

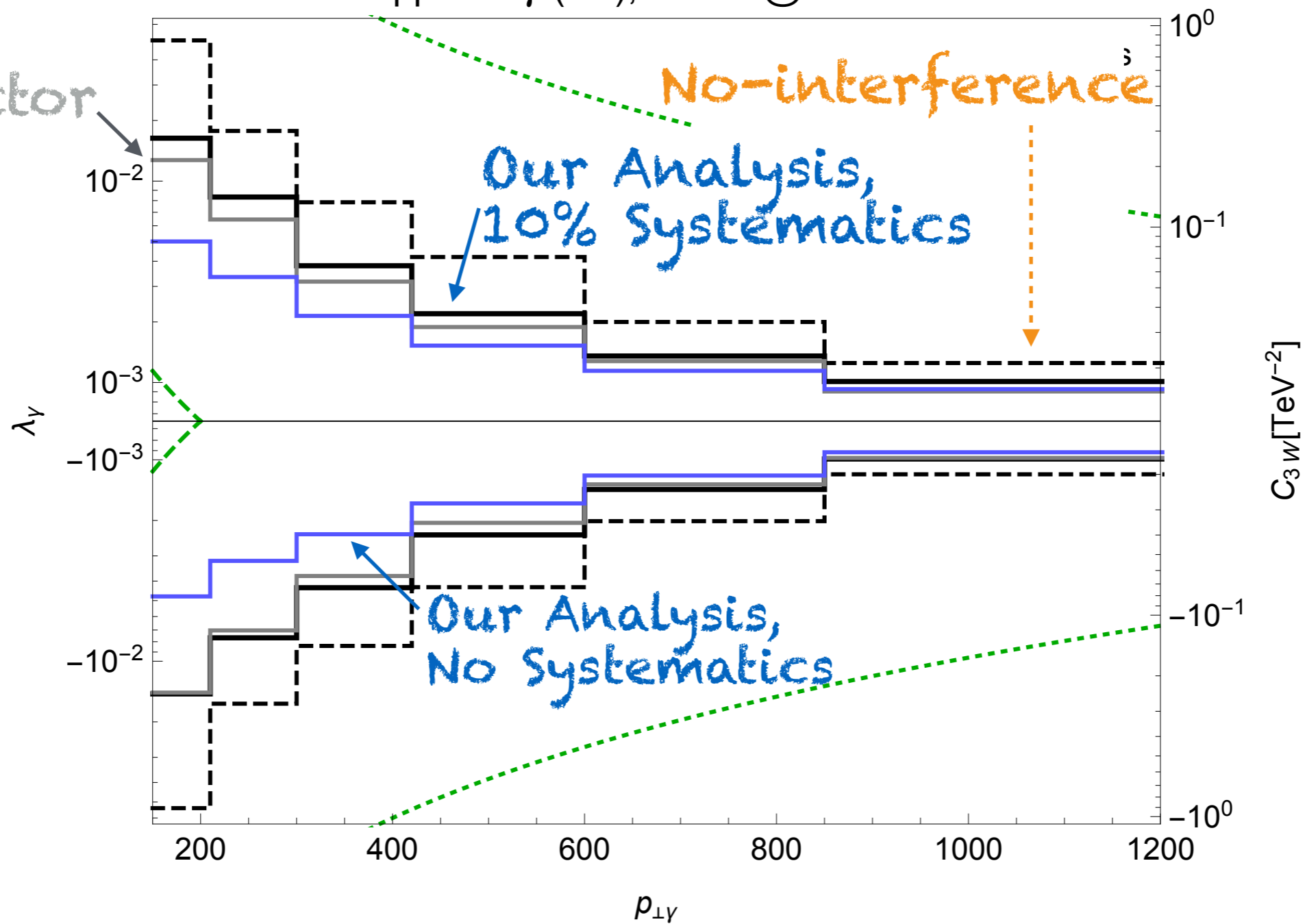
$pp \rightarrow W\gamma$ (LO), 3ab^{-1} @14 TeV



Results

$pp \rightarrow W\gamma$ (LO), 3ab^{-1} @14 TeV

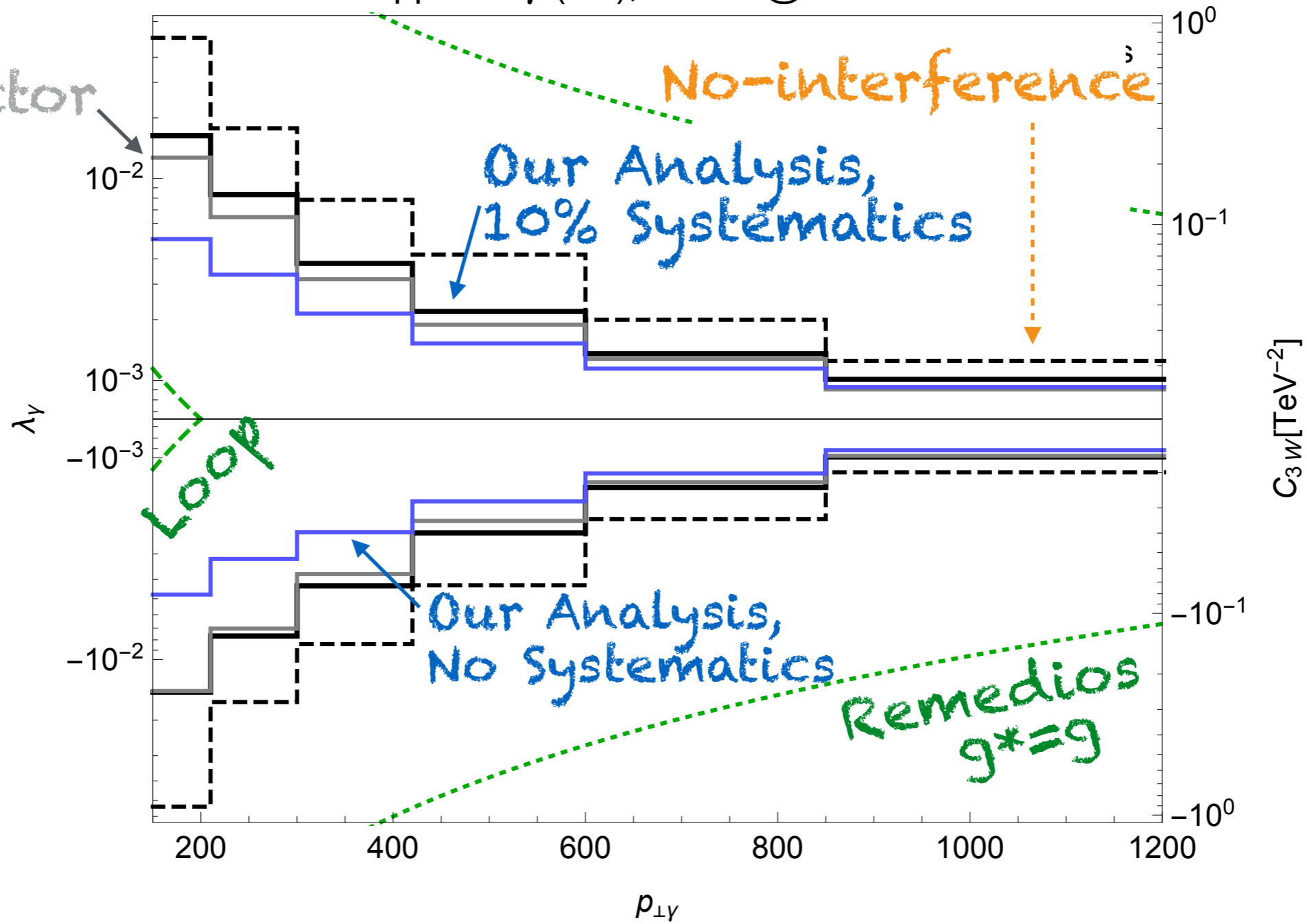
No detector effects



Results

pp → Wγ (LO), 3ab⁻¹@14 TeV

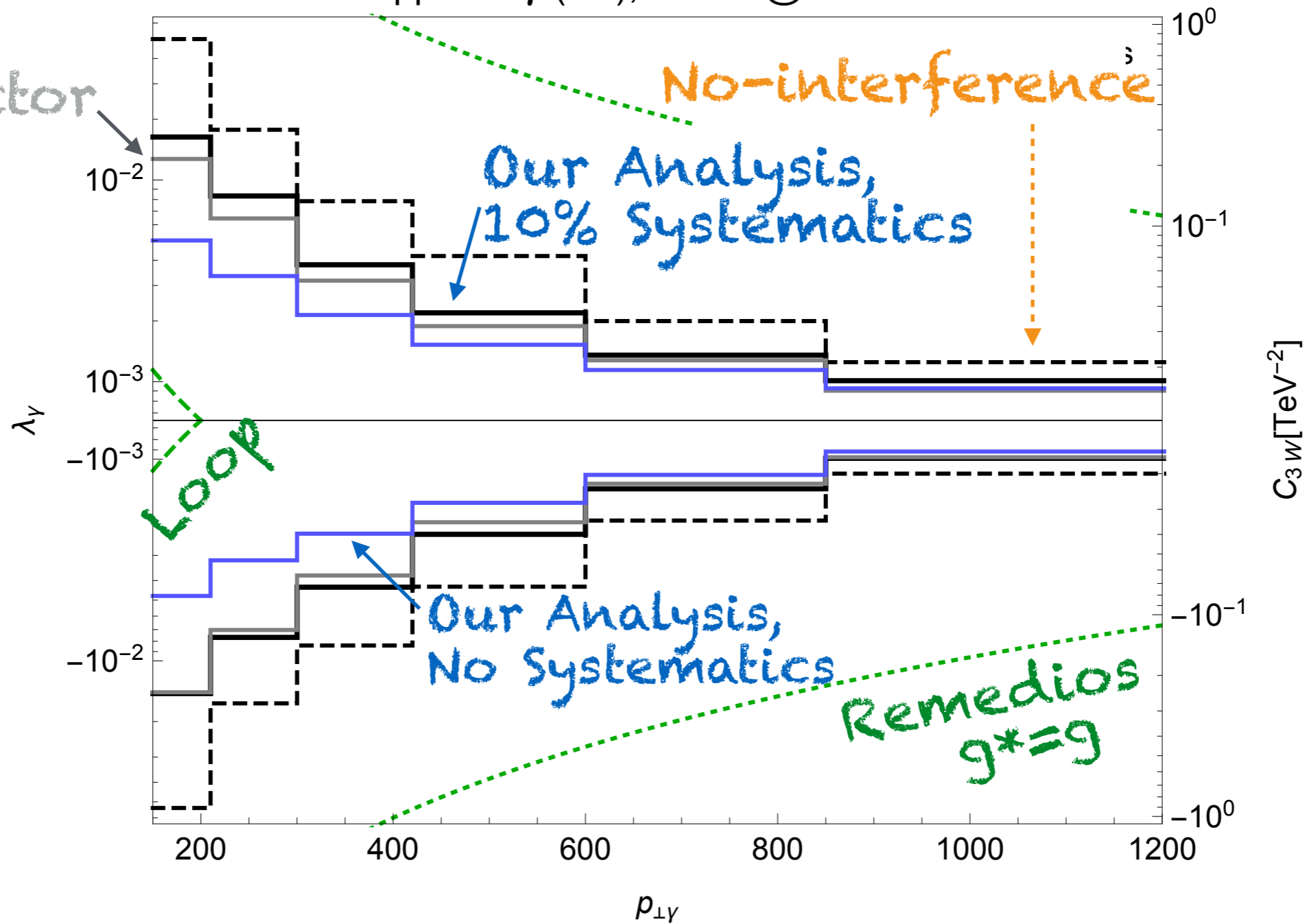
No detector effects



Results

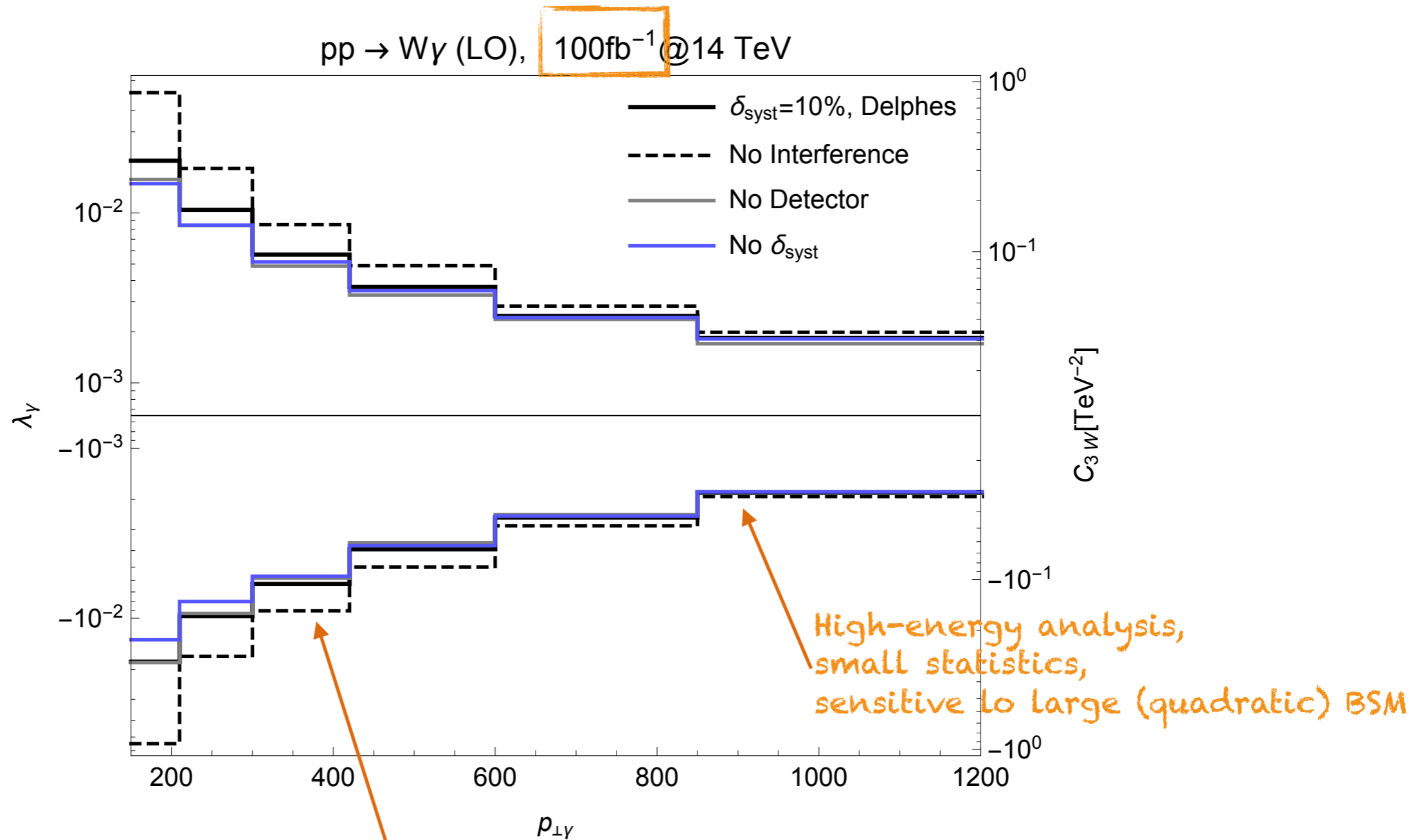
pp \rightarrow W γ (LO), 3ab⁻¹@14 TeV

No detector effects



▶ Important improvement, though not yet there for weakly coupled/loop-generated new physics

Results

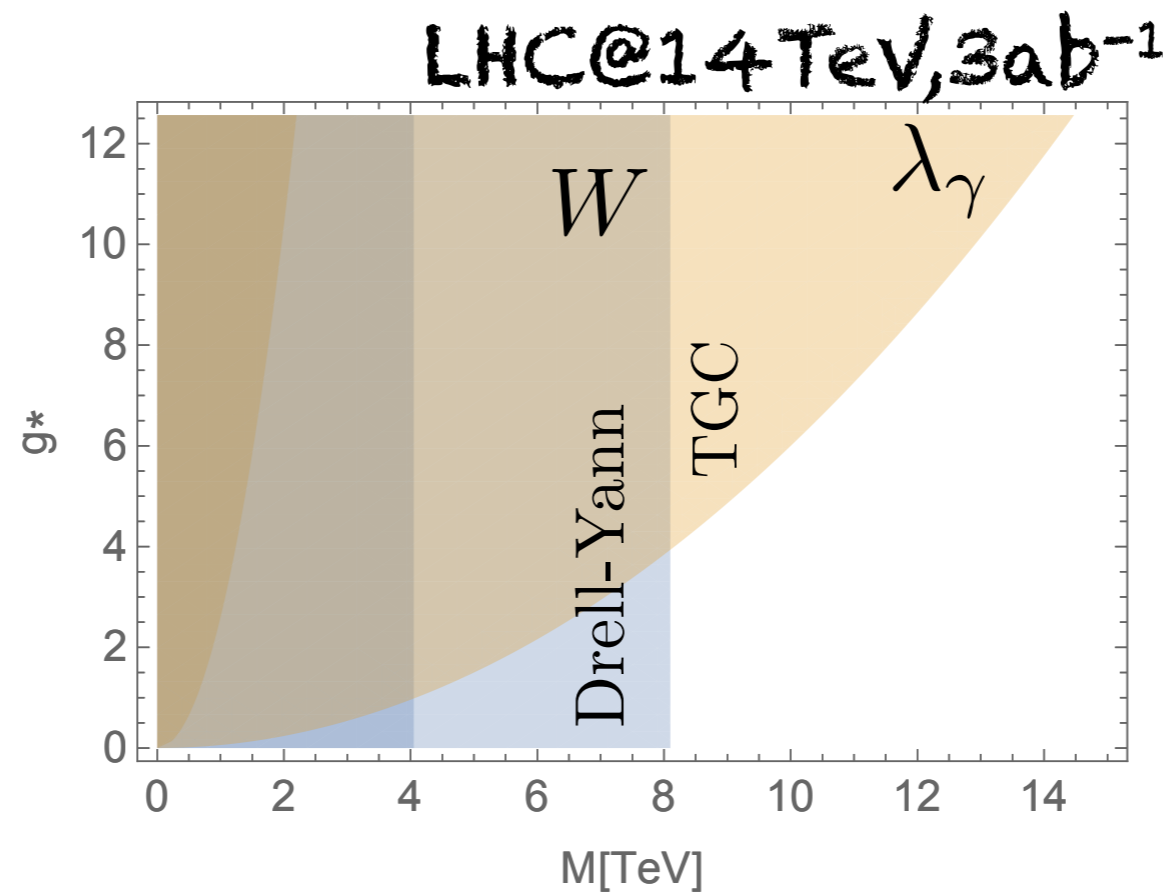


At small energy, interference has impact already now.
(improving low-energy measurement, important for validity)

Explicit Model (Remedios)

see Pomarol's talk

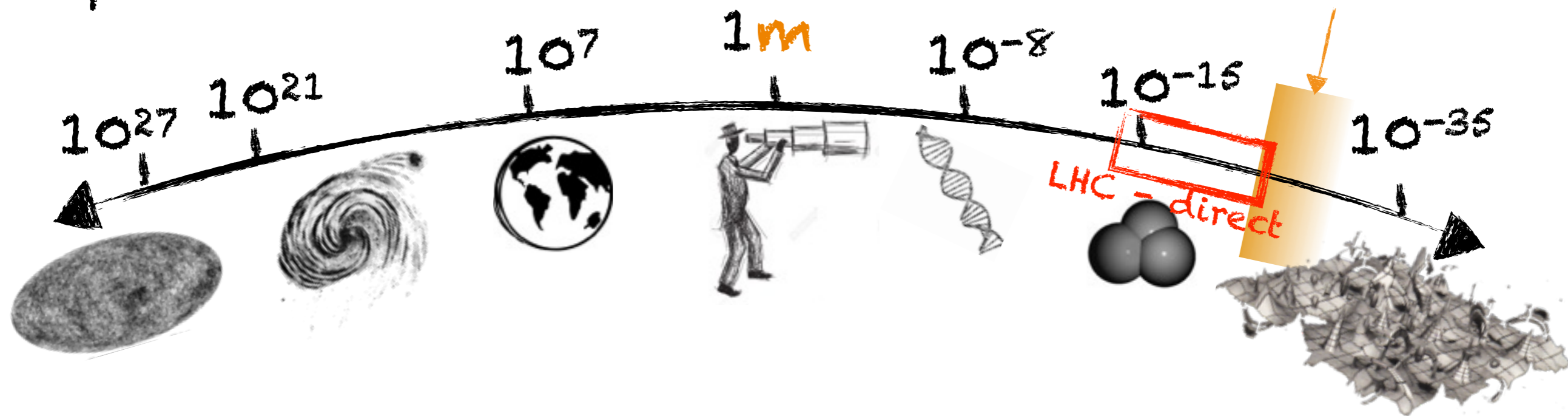
Liu, Pomarol, Rattazzi, FR'16



Interference Resurrection makes the difference.

Message

SM precision tests will define the new distance frontier



► LHC good in High-E $2\gamma 2$ processes

Challenge:

non-interference limits precision in learning about transverse vectors

► Azimuthal distributions crucial

► Realistic in other processes? WZ? VBF?