

Regular black holes and their relationship to polymerised models



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Motivation

Investigate quantum black hole models interesting for quantum gravity

Many people contribute:

[Ashtekar, Alonso-Bardaji, Bojowald, Brizuela, Modesto, Cartin, Khanna, Boehmer, Vandersloot, Chiu, Campiglia,, Gambini, Pullin, Sabharwal, Brannlund, Kloster, De Benedictis, Olmedo, Dadhich, Joe, Singh, Haggard, Rovelli, Vidotto, Corichi, Saini, Cortez, Cuervo, Morales-Técol, Ruelas, Pawłowski, Bianchi,, Christodoulou, D'Ambrosio, Alesci, Bahrami, Pranzetti, Husain, Kelly, Santacruz, Wilson-Ewing, Lewandowski, Zhang, Ma, Song, Bodendorfer, Mele, Münch, Navascués, Mena Marugán, García-Quismondo, Perez, Speziale, Viollet, Han, K.G., Liu, Li, Vera, Weigl,...]

Recent reviews: [Gambini, Olmedo, Pullin '22], [Ashtekar, Olmedo, Singh '23]

Dynamical formulation of gravitational collapse: consider spherically symmetric models with dust

LTB models, Oppenheimer-Snyder collapse also special case of vacuum solution

Here we will consider effective models to formulate such models which involve (LQG inspired) quantum corrections

Aim: Develop formalism that allows to investigate a broad class of effective models [seminal work by Bojowald, Harada, Reyes, Tibrewala '08 '09]

I. Classical LTB models and vacuum solutions

I. Classical LTB models

LTB: Spherically symmetric solution with dust

We consider Ashtekar-Barbero variables for spherical symmetry (A_a^j, E_j^a)

After implementing the Gauss constraint: [Bojowald Kastrup '00], [Bojowald, Swiderski '03]

$$\begin{aligned} A_a^j \tau_j \, dX^a = & 2\beta K_x(x) \tau_1 \, dx + \left(\beta K_\phi(x) \tau_2 + \frac{\partial_x E^x(x)}{2E^\phi(x)} \tau_3 \right) d\theta \\ & + \left(\beta K_\phi(x) \tau_3 - \frac{\partial_x E^x(x)}{2E^\phi(x)} \tau_2 \right) \sin(\theta) d\phi + \cos(\theta) \tau_1 \, d\phi \\ E_j^a \tau_j \frac{\partial}{\partial X^a} = & E^x(x) \sin(\theta) \tau_1 \partial_x + (E^\phi(x) \tau_2) \sin(\theta) \partial_\theta + (E^\phi(x) \tau_3) \partial_\phi, \end{aligned} \quad (x, \theta, \phi)$$

Reduced phase space including dust:

$$\{K_x(x), E^x(y)\} = G\delta(x, y) \quad \{K_\phi(x), E^\phi(y)\} = G\delta(x, y) \quad \{T(x), P_T(y)\} = \delta(x, y)$$

I. Classical LTB models: LTB condition

General spherically symmetric metric

$$ds^2 = -N(x, t)^2 dt^2 + \frac{(E^\phi)^2}{|E^x|} (dx + N^x dt)^2 + |E^x| d\Omega^2$$

Consider the form of the LTB metric [Lemaître '33], [Tolman '34], [Bondi '47]

$$ds^2 = -dt^2 + \frac{((E^x)')^2}{4|E^x|(1 + \mathcal{E}(x))} dx^2 + |E^x| d\Omega^2$$

To match both metric we need [Bojowald, Harada, Reyes, Tibrewala '08 '09]

$$N = 1 \quad N^x = 0 \quad G_x(x) = \frac{E^{x'}}{2E^\phi}(x) - \sqrt{1 + \mathcal{E}(x)} = 0$$

marginally bound case $\mathcal{E}(x) = 0$ shells decouple classically

$$\partial_t E^x(x) = -2\sqrt{E^x}(x) K_\phi(x), \quad \partial_t K_\phi(x) = \frac{(K_\phi)^2}{2\sqrt{E^x}}(x)$$

I. Classical Vacuum Solutions

Gravitational contribution to the scalar constraint, conserved quantity $M(x)$

$$C(x)|_{LTB} = \partial_x \tilde{H}(x), \quad -M(x) = \tilde{H}(x) := -\frac{1}{2G} \sqrt{E^x(x)} (K_\phi)^2(x)$$

We can rewrite EOMs in form of Friedmann equation $R(x) = \sqrt{E^x(x)}$

$$\frac{\dot{R}^2}{R^2}(x) = \frac{8\pi G}{3} \rho(x) \quad \text{with} \quad \rho(x) = \frac{3}{4\pi} \frac{M}{R^3}(x).$$

Classical vacuum solution: choose dust mass profile $M(x) = m = \text{const}$
using that M is conserved: $s(x)$ integration constant

$$K_\phi(x) = \frac{\sqrt{2Gm}}{(E^x)^{\frac{1}{4}}(x)} \longrightarrow \partial_t E^x(x) = -2\sqrt{2Gm} (E^x)^{\frac{1}{4}}(x) \longrightarrow E^x = \left[\frac{3}{2} \sqrt{2Gm} (s(x) - t) \right]^{\frac{4}{3}} \longrightarrow E^\phi$$

Transforming to Schwarzschild coordinates $r = \sqrt{E^x}$, $r_s := 2Gm$

$$ds^2 = - (1 - \mathcal{G}(r)^2) d\tau^2 + \frac{1}{(1 - \mathcal{G}(r)^2)} dr^2 + r^2 d\Omega^2 \quad \mathcal{G}(r)^2 := \frac{(\partial_x E^x)^2}{4E^x} = \frac{r_s}{r},$$

II. Effective models

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1. Can we construct polymerised vacuum solutions in a similar way?
2. Given a Schwarzschild-like metric can we reconstruct the associated polymerised model?

II. How to construct solutions here?

Class II: Compatible LTB condition exists + conserved energy density restricts possible polymerisation functions and form of LTB condition

Dynamics of effective model:

$$\text{EOM: } \partial_t E^x = -2\sqrt{E^x} f^{(2)}, \quad \partial_t K_\phi = \frac{1}{2\sqrt{E^x}} \left(f^{(1)} - g_{(\alpha)}^2 (2h_2 + 4E^x \partial_{E^x} h_2 - h_1) + h_1 \right)$$

$$\text{Cons. } \tilde{H}^{(\alpha)}(x) := \frac{1}{2G} \left[\frac{\sqrt{E^x}}{g_{(\alpha)}} \left(-F + h_2 \left(g_{(\alpha)}^2 - 1 \right) \right) \right] (x), \quad C^{(\alpha)}(x) \Big|_{LTB} = \partial_x \tilde{H}^{(\alpha)}(x),$$

energy dens.

Requirements for polarisation and inverse triad corrections in class II:

$$\partial_{K_\phi} F(K_\phi, E^x) = 2f^{(2)}(K_\phi, E^x) \quad \text{and} \quad \frac{h_1 - 2E^x \partial_{E^x} h_2}{h_2} = \frac{-2E^x \partial_{E^x} F + f^{(1)}}{F}$$

General effective LTB condition: $C_{LTB}^{(\alpha)}(x) := \left[\frac{E^{x'}}{2E^\phi} - \tilde{g}_{(\alpha)}(K_\phi, E^x, \Xi) \right] (x)$ $\Xi(x) := \sqrt{1 + \mathcal{E}(x)}$
 Here class II marginally bound: $g_{(\alpha)} = g_{(\alpha)}(E^x)$ classical case: $g_{(\alpha)} = 1$

II. Can we still construct a solution for E^x ?

We need the inverse of F :

$$\tilde{H}^{(\alpha)}(x) := \frac{1}{2G} \left[\frac{\sqrt{E^x}}{g_{(\alpha)}} \left(-F + h_2 \left(g_{(\alpha)}^2 - 1 \right) \right) \right] (x)$$

Given this we can solve
for $K_\phi(t, x)$

$$K_\phi(t, x) = F_{(i)}^{-1} \left[h_2 \left(g_{(\alpha)}^2 - 1 \right) - 2GM(x) \frac{g_{(\alpha)}}{\sqrt{E^x}}, E^x \right] =: F_{(i), M(x)}^{-1} [E^x]$$

Class II important: LTB
condition only depends
on E^x and likewise h_2

Why label (i) ?

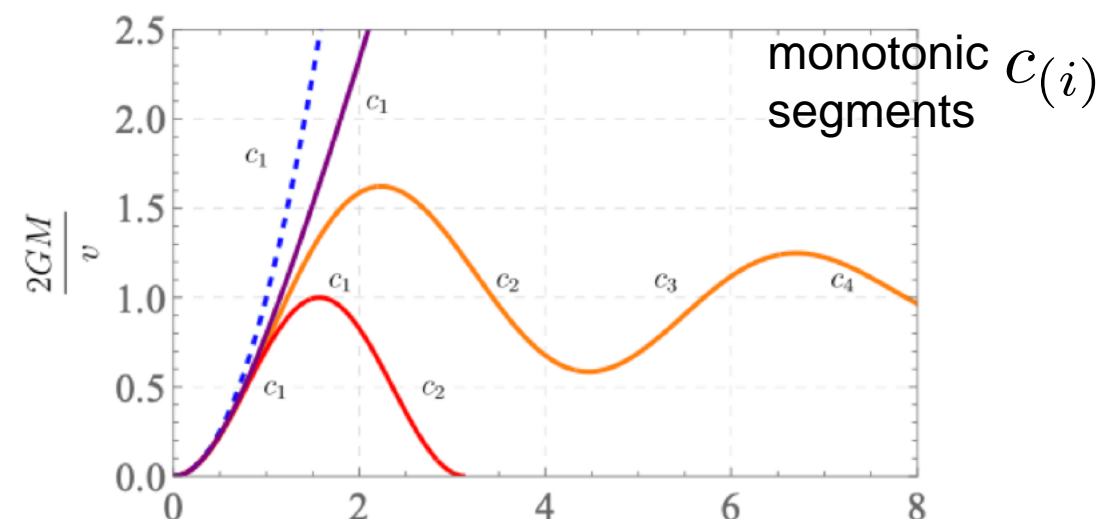
Specialise to models where classical LTB condition is compatible + no inverse triad corrections (mimetic model exists)

Phase space trajectories: level sets

$$c(K_\phi, E^x) := \tilde{H}^{(\alpha)}(x) = -M(x)$$

$$c(b, v) \quad b := \frac{K_\phi}{\sqrt{E^x}}, v := (E^x)^{\frac{3}{2}}$$

$$\tilde{H}^\alpha(x) = -M(x) = -\frac{1}{2G} v \tilde{F}(b), \quad F = E^x \tilde{F}(b)$$



II. Now we can construct the solution

Given the inverse F^{-1} we obtain for the EOM and its solution:

$$\frac{\partial_t E^x}{2\sqrt{E^x}}(t, x) = -f^{(2)} \left[F_{(i), M(x)}^{-1} [E^x], E^x \right]$$

$$E_{(i)}^x(t, x) = \mathcal{F}_{(i)}^{-1}(s(x) - t)$$

with

$$\mathcal{F}_{(i)} = \int \frac{dE^x}{2\sqrt{E^x} f^{(2)} \left[F_{(i), M(x)}^{-1} [E^x], E^x \right]}$$

and $E^x = E_{(1)}^x(t, x) \circ E_{(2)}^x(t, x) \circ \dots$
continuous by suitable choice of $s(x)$

Form of the solution in Schwarzschild-like coordinates

$$ds^2 = - \left(1 - \mathcal{G}_{(i)}(r)^2 \right) d\tau^2 + \frac{1}{g_{(\alpha)}(r)^2 \left(1 - \mathcal{G}_{(i)}(r)^2 \right)} dr^2 + r^2 d\Omega^2$$

Trafo $r = \sqrt{E^x}$ only
defined for each
piecewise segment

with

$$\mathcal{G}_{(i)}(r)^2 := \frac{\left(\partial_x E_{(i)}^x \right)^2}{4g_{(\alpha)}^2 E^x(i)} = \frac{\left(f^{(2)} \left[F_{(i), m}^{-1} [E^x], E^x \right] \right)^2}{g_{(\alpha)}^2}(r)$$

In general can be different $\mathcal{G}_{(i)}$ for each i
example later

II. Reconstruction algorithm

So far we chose some polymerisation determined F^{-1} and from solution $\mathcal{G}_{(i)}$

Can we also obtain the polymerised model given some choice of \mathcal{G} ?

In particular interesting for regular black holes like Bardeen and Hayward

$$\text{Bardeen} \quad \mathcal{G}(r)^2 = \frac{r_s r^2}{(r^2 + l^2)^{\frac{3}{2}}} \quad \text{Hayward} \quad \mathcal{G}(r)^2 = \frac{r_s r^2}{r^3 + l^2 r_s}$$

Reconstruction: special case of class II with classical compatible LTB condition+no inverse triad corrections, then factorisation in M

$$-M(x) = \tilde{H}^{(\alpha)}(x) = -\frac{1}{2G} [v \tilde{F}(b)](x), \quad v := (E^x)^{\frac{3}{2}}$$

Using the result for $\mathcal{G}_{(i)}$ from before we obtain it as a function of $\frac{2Gm}{v} = \frac{r_s}{r^3}$

$$\mathcal{G}(r) = \frac{r}{2} \tilde{F}' \left[\tilde{F}^{-1} \left[\frac{r_s}{r^3} \right] \right] = -\frac{3r_s}{2r^3 \frac{d}{dr} \left(\tilde{F}^{-1} \left[\frac{r_s}{r^3} \right] \right)} \longrightarrow \tilde{F}^{-1} \left[\frac{r_s}{r^3} \right] = -\int dr \frac{3r_s}{2r^3 \mathcal{G}(r)} = \int d \left(\frac{r_s}{r^3} \right) \frac{r}{2\mathcal{G}(r)}$$

From this we can derive the polymerisations of Bardeen and Hayward metrics

III. Examples

(i). Standard LQC polymerisation

Polymerised vacuum solutions with a symmetric bounce: (no inverse triad corr.)

$$\tilde{F}(b) = \frac{\sin^2(\alpha_\Delta b)}{\alpha_\Delta^2}, \quad \alpha_\Delta := \gamma\sqrt{\Delta}, \quad \Delta := 4\pi l_P \quad \text{LTB condition} \quad g(\alpha) = g_\Delta = 1$$

Then we obtain in Schwarzschild-like coordinates

Since $\mathcal{G}(r)$ is real, minimal radius $r_{\min} = (r_s \alpha_\Delta^2)^{\frac{1}{3}}$

This coordinate is only defined for $b \in (0, \frac{\pi}{2})$ or $b \in (\frac{\pi}{2}, \pi)$ two monotone segments have same

Taking $\mathcal{G}(r)$ as starting point one can also obtain by reconstruction $\tilde{F}(b) = \frac{\sin^2(\alpha_\Delta b)}{\alpha_\Delta^2}$

Next: underlying effective spherically symmetric model

$$\tilde{f}^{(1)}(b) = \frac{3 \sin^2(\alpha_\Delta b) - \alpha_\Delta b \sin(2\alpha_\Delta b)}{\alpha_\Delta^2}, \quad \tilde{f}^{(2)}(b) = \frac{\sin(2\alpha_\Delta b)}{2\alpha_\Delta}$$

mimetic Lagrangian

Effective Hamiltonian:

$$C^\Delta = -\frac{E^\phi \sqrt{E^x}}{2G} \left[\frac{3 \sin^2(\alpha_\Delta b)}{\alpha_\Delta^2} + \left(\frac{2\sqrt{E^x} K_x}{E^\phi} - b \right) \frac{\sin(\alpha_\Delta b)}{\alpha_\Delta} + \frac{1 - \left(\frac{E^{x'}}{2E^\phi} \right)^2}{E^x} - \frac{2}{E^\phi} \left(\frac{E^{x'}}{2E^\phi} \right)' \right].$$

*J. G. Kelly, R. Santacruz, and E. Wilson-Ewing '20,
Lewandowski, Y. Ma, J. Yang, and C. Zhang '22]*

(i). Standard LQC polymerisation

Modified Friedmann equation gives the following polymerised vacuum solution:

$$R(t, x) = (2Gm)^{\frac{1}{3}} \left(\frac{9}{4}(s(x) - t)^2 + \alpha_{\Delta}^2 \right)^{\frac{1}{3}}$$

bouncing solution with minimal radius

$$r_{\min} = (2Gm\alpha_{\Delta}^2)^{\frac{1}{3}}$$

We can further extend the solution to the marginally bound case LTB solution:

$$R(t, x) = (2GM(x))^{\frac{1}{3}} \left(\alpha_{\Delta}^2 + \frac{9}{4}z^2 \right)^{\frac{1}{3}}, \quad z = s(x) - t$$

minimal radius

$$r_{\min} = (2GM(x)\alpha_{\Delta}^2)^{\frac{1}{3}}$$

Curvature scalars for generic dust profile $M(x)$

$$\mathcal{R} = \frac{\mathcal{A}}{(9z^2 + 4\alpha_{\Delta}^2)^2 \mathcal{S}}, \quad \mathcal{K} = \frac{\mathcal{B}}{(9z^2 + 4\alpha_{\Delta}^2)^4 \mathcal{S}^2},$$

$$\mathcal{S} = M'(x) (9z^2 + 4\alpha_{\Delta}^2) + 18M(x)s'(x)z$$

at the bounce $z=0$, no dependence on M

$$M'(x) \neq 0 : \quad \mathcal{R}|_{z=0} = \frac{9}{\alpha_{\Delta}^2}, \quad \mathcal{K}|_{z=0} = \frac{27}{\alpha_{\Delta}^4}$$

Central singularity is resolved, shell crossing if $9M(x)^2 s'(x)^2 - 4\alpha_{\Delta}^2 M'(x)^2 \geq 0$ (real roots of $\mathcal{S}(x)$)

(ii). Non-standard LQC polymerisation

[Dapor-Liegener '17], [Ying, Ding, Ma '09]

Polymerised vacuum solutions with an asymmetric bounce: (no inverse triad corr.)

$$\tilde{F}(b) = \frac{\sin^2(\alpha_\Delta b \gamma) (1 - (\gamma^2 + 1) \sin^2(\alpha_\Delta b \gamma))}{(\alpha_\Delta \gamma)^2} \quad \text{LTB condition} \quad g_{(\alpha)} = g_\Delta = 1$$

Then we obtain in Schwarzschild-like coordinates

$$\dot{R}(t, x) = \mathcal{G}_\pm(r)|_{r=R(t, x)} = \frac{R(t, x) \sqrt{\frac{1}{2} - \gamma^2 x_0} \sqrt{x_0 \pm \sqrt{1 - 2\gamma^2 x_0} + 1}}{\alpha_\Delta (\gamma^2 + 1)}$$

Here example of two different $\mathcal{G}(r)$:

Taking $\mathcal{G}_\pm(r)$ as starting point one reconstruct $\tilde{F}(b)$

Next: underlying effective spherically symmetric model (+mimetic Lagrangian)

$$\begin{aligned} \tilde{f}^{(1)}(b) &= \frac{\sin(\alpha_\Delta b \gamma) (\sin(\alpha_\Delta b \gamma) ((\gamma^2 + 1) (2\alpha_\Delta b \gamma \sin(2\alpha_\Delta b \gamma) - 3 \sin^2(\alpha_\Delta b \gamma)) + 3) - 2\alpha_\Delta b \gamma \cos(\alpha_\Delta b \gamma))}{\alpha_\Delta^2 \gamma^2} \\ \tilde{f}^{(2)}(b) &= \frac{\sin(2\alpha_\Delta b \gamma) ((\gamma^2 + 1) \cos(2\alpha_\Delta b \gamma) - \gamma^2)}{2\alpha_\Delta \gamma} \end{aligned} \quad \longrightarrow \quad \text{effective Hamiltonian}$$

(ii). Non-standard LQC polymerisation

[Dapor-Liegener '17], [Ying, Ding, Ma '09]

In this case the marginally bound LTB solution reads:

$$R(t, x) = \sqrt[3]{\frac{2GM(x) (4\alpha_\Delta^2 \gamma^2 + 9\eta^2)^2}{18\eta^2 - 8\alpha_\Delta^2 \gamma^4}}, \quad s(x) - t = \eta - \frac{2}{3}\alpha_\Delta (\gamma^2 + 1) \tanh^{-1} \left(\frac{2\alpha_\Delta \gamma^2}{3\eta} \right)$$

Bouncing solutions with bounce at η_0 and minimal radius $r_{\min} = 2^{2/3} \sqrt[3]{a^2 \gamma^2 (\gamma^2 + 1) M(x)}$

Modified FRW eqn defined for each segments C_\pm

$$\eta \geq \eta_0 \quad C_- \quad \text{and} \quad \frac{2}{3}\alpha_\Delta \gamma^2 < \eta \leq \eta_0 \quad C_+$$

Curvature scalars

$$\mathcal{R} = \frac{\mathcal{A}}{(9\eta^2 + 4\alpha_\Delta^2 \gamma^2)^4 \mathcal{S}}, \quad \mathcal{K} = \frac{\mathcal{B}}{(9\eta^2 + 4\alpha_\Delta^2 \gamma^2)^8 \mathcal{S}^2}$$

for vacuum case

$$\mathcal{K} \sim \frac{81\gamma^2}{16\alpha_\Delta^2 (\gamma^2 + 1)^2 (2\gamma^2 + 1) (\eta - \eta_0)^2}$$

$$\mathcal{S} = M'(x) (9\eta^2 + 4\alpha_\Delta^2 \gamma^2)^2 + 18M(x)s'(x)\eta (4\alpha_\Delta^2 \gamma^2 (2\gamma^2 + 1) - 9\eta^2)$$

Central singularity is resolved, shell crossing singularity still present

(iii). Bardeen and Hayward

Polymerised vacuum solutions with a symmetric bounce: (no inverse triad corr.)

	Bardeen	Hayward
metric	$\mathcal{G}(r)^2 = \frac{r_s r^2}{\left(r^2 + \alpha^{\frac{4}{3}} r_s^{\frac{2}{3}}\right)^{\frac{3}{2}}}$	$\mathcal{G}(r)^2 = \frac{r_s r^2}{(r^3 + \alpha^2 r_s)}$
Polymerization function	$\tilde{F}^{-1} = \frac{\alpha r_s}{2r^3} {}_2F_1\left(-\frac{3}{2}, -\frac{3}{4}, -\frac{1}{2}, -\left(\alpha^2 \frac{r_s}{r^3}\right)^{-\frac{2}{3}}\right) + \frac{\sqrt{\pi}\Gamma\left(\frac{3}{4}\right)}{\alpha\Gamma\left(-\frac{3}{4}\right)}$	$\tilde{F}^{-1} = \frac{2\eta\alpha + \sinh(2\alpha\eta)}{4\alpha},$ $\alpha\eta = \sinh^{-1} \sqrt{\frac{\alpha^2 r_s}{r^3}}$
Marginally bound solution	$R(t, x) = (2GM(x))^{\frac{1}{3}} \sqrt{\eta^{\frac{4}{3}} - \alpha^{\frac{4}{3}}}, \quad \eta \geq \alpha$ $s(x) - t = \frac{2}{3}\eta + \alpha \tan^{-1}\left(\eta^{\frac{1}{3}} \alpha^{-\frac{1}{3}}\right) - \alpha \operatorname{Re} \tanh^{-1}\left(\eta^{\frac{1}{3}} \alpha^{-\frac{1}{3}}\right)$	$R(t, x) = \left(\frac{2GM(x)\alpha^2}{\sinh^2(\alpha\eta)}\right)^{\frac{1}{3}}, \quad \eta \geq 0$ $s(x) - t = \frac{2}{3}\alpha(\coth(\alpha\eta) - \alpha\eta)$
Curvature scalars (see App. B)	$\mathcal{R} = \frac{\mathcal{A}}{\eta^{14/3}\mathcal{S}}, \quad \mathcal{K} = \frac{\mathcal{B}}{\eta^{28/3}\mathcal{S}^2}$ $\mathcal{S} = M'(x)\eta + 3M(x)s'(x)$	$\mathcal{R} = \frac{\mathcal{A}}{4\alpha^3\mathcal{S}}, \quad \mathcal{K} = \frac{\mathcal{B}}{16\alpha^6\mathcal{S}^2}$ $\mathcal{S} = M'(x) + 3M(x)s'(x) \frac{\tanh(\alpha\eta)}{r}$

shell crossing singularities: situation like in GR: can be avoided by choosing suitable dust profile $M(x)$ and $s(x)$, e.g. $M'(x) \geq 0$ and $s'(x) \geq 0$

IV. Summary & Conclusions

Formalism allows to investigate a broad class of effective models with different kind of polymerisation

The formalism can be used in different ways:

- 1.) start with a given polymerisation in the LTB sector determine the underlying gauge-unfixed spherically symmetric model
- 2.) For regular black holes: start with a modified Schwarzschild-metric and derive the corresponding effective model or vice versa.

Here we considered examples for bounded and unbounded polymerisations

Saw that gauge fixing and/or coordinate choice is more subtle in effective models in general underlying covariant mimetic model helpful.

Next steps:

Investigate non-marginally bound case in a similar manner

Investigate more in detail shock solutions and whether we can construct polymerised model without the presence of shell-crossing singularities

Thank you for your time!

Underlying covariant Lagrangian

Extended mimetic gravity

$$S[g_{\mu\nu}, \phi, \lambda] = \frac{1}{8\pi G} \int_{M_4} d^4x \sqrt{-g} \left[\frac{1}{2} R^{(4)} + L_\phi(\phi, \chi_1, \dots, \chi_p) + \frac{1}{2} \lambda (\phi_\mu \phi^\mu + 1) \right]$$

mimetic field ϕ , λ is a Lagrange multiplier for the mimetic condition, 2+1 dof

$$\chi_n \equiv \sum \phi_{\mu_1}^{\mu_2} \phi_{\mu_2}^{\mu_3} \dots \phi_{\mu_{n-1}}^{\mu_n} \phi_{\mu_n}^{\mu_1}, \quad \phi_\mu = \nabla_\mu \phi, \quad \phi_{\mu\nu} = \nabla_\mu \nabla_\nu \phi.$$

Spherically symmetric model: Sufficient to have $L_\phi(\chi_1, \chi_2)$, $\psi = \ln(E^x)$ 2D action
[Achour, Lamy, Liu, Noui '18], [Han, Liu '22]

$$S_2 = \frac{1}{4G} \int_{\mathcal{M}_2} d^2x \det(e) e^{2\psi} \left\{ \mathcal{R} + L_\phi(X, Y) + \frac{\lambda}{2} (\phi_{,j} \phi^{,j} + 1) \right\}$$

(Smooth) mimetic field defines foliation into spacelike hyper surfaces $\phi = \text{const}$
Generalised Einstein's equation

$$G_{\mu\nu}^\Delta := G_{\mu\nu} - T_{\mu\nu}^\phi = -\lambda \partial_\mu \phi \partial_\nu \phi, \quad \partial_\mu \phi \partial^\mu \phi = -1$$

Underlying covariant Lagrangian

Now for models with have no inverse triad corrections + compatible with $\bar{\mu}$ -scheme we can relate the choice of the mimetic potential to specific choices of polymerisation function

[Achour, Lamy, Liu, Noui '18], [Han, Liu '22]

$$S_2 = \frac{1}{4G} \int_{\mathcal{M}_2} d^2x \det(e) e^{2\psi} \left\{ \mathcal{R} + L_\phi(X, Y) + \frac{\lambda}{2} (\phi_{,j} \phi^{,j} + 1) \right\}$$

Higher derivative couplings can be expressed in terms of X,Y and relate to extrinsic curvature

$$X = -\square_h \phi + Y = \frac{\partial_t E^\phi}{E^\phi}, \quad Y = -h^{ij} \partial_i \psi \partial_j \phi = \frac{\partial_t E^x}{2E^x} = \frac{\sin(2\alpha b)}{2\alpha}$$

Underlying covariant model allows to gauge-unfix temporal gauge with respect to mimetic field and consider coordinate taros in (t,x)

Interpretation: effective model has different clock than classical model