# Introduction

Process mediated by the weak interaction between two isobars

$$\beta^{-}$$
 (N,Z) ---> (N-1,Z+1) + e<sup>-</sup> + v   
 M(Z) - M(Z+1) = E<sub>β</sub> + E<sub>ν</sub> + Ex

M(Z) - M(Z-1)=  $E_{\beta^+} + E_{\nu} + 1022 + Ex$ 







## Beta Delayed Particle Emission







# **Planning and References**

Beta delayed Particle emission

- Mechanisms of Breakup
- Analysis technique

#### References:

"Particle Emission from Nuclei" Ed. D.N. Poenaru & M.S. Ivaçcu CRD 1989 Vol I, II, III
B. Blank and M.J.G. Borge, Prog Part and Nuc. Phys 60 (2008) 403
M. Pfützner, L.V. Grigorenco, M. Karny & K. Riisager, Rev. Mod. Phys, ArXiV:1111.0482
Euroschool on Exotic Beams, Lectures Notes:
"Decay Studies of N~Z Nuclei", E. Roeckl, École Joliot-Curie de Physique Nucleaire, 2002



# Reminder

#### Spectra $\beta^{\pm}$



Expand in a large E-scale  $E_{\beta}$ -= 2,6 keV (<sup>187</sup>Re,  $\beta^{-}$ )  $E_{\beta}$ -= 22800 keV (<sup>22</sup>N,  $\beta^{-}$ ) Q=M(Z,N) - M'(Z+1,N-1)  $m_ec^2 = T_{M'}$ + Te +Tv N(p)  $\propto p^2$ (Q-Te-Tv)

#### Half-life

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**Emission of delayed particles** 

$$P_p = 6 \ 10^{-6} \ (^{151}Lu)$$
 to 100 % (<sup>31</sup>Ar)

$$P_n = 5.5 \ 10^{-4} \ (^{79} \, \text{Ge}) \text{ to } 99 \ \% \ (^{11} \, \text{Li})$$

<sup>35</sup> Na,  $T_{1/2} = 1,5$  ms <sup>50</sup>V,  $T_{1/2} = 10^{17}$  y; (<sup>115</sup>In, 10<sup>14</sup> y ; <sup>113</sup>Cd, 10<sup>15</sup> y) β p, β2p, β3p, ...βn, β2n ...



# Classification of β-decay transitions



 $L_{\beta}$ >0 and/or  $\pi_{i}\pi_{f}$ =-1



### **Practical example**

$$t = T_{1/2}^{\beta_i} = \frac{T_{1/2}^{\exp}}{P_{\beta_i}}$$

$$P_{\beta_i} = \eta [I^{tot} (out) - I^{tot} (in)]$$

$$I^{2} = \frac{\alpha_1(M_1) + \delta^2 \alpha_1(E_2)}{P_{\beta_i}}$$

$$P_{\beta_i} = \eta [I^{tot} (out) - I^{tot} (in)]$$

$$I^{2} = \frac{\alpha_2(M_1) + \delta^2 \alpha_1(E_2)}{P_{\beta_i}}$$

$$I^{(N)} = \frac{\alpha_1(M_1) + \delta^2 \alpha_1(E_2)}{P_{$$

Maria J.Gª Borge, Beta decay studies



### **Beta Transitions**





#### Decay properties of exotic nuclei





### Beta-proton emitters

✓ More than 160 precursors identified

 $\checkmark$  For every element up to Z = 73 at least one proton precursor

 $\checkmark$  The βp spectrum depends on the Z and A of the precursor and differs in the different mass region due to differences in level density in the Q-Sp window

✓ Properties of  $\beta p$  well understood → large variety of spectroscopic information



# $^{32}Ar(Z=18, N=14) \rightarrow ^{32}CI(Z=17, N=15)$



- By  $\beta p \rightarrow$  study up to 8153 keV  $\rightarrow$  73% of energy window
- 22.6% feeding IAS→ logft = 3.19

$$\Gamma_{p} / \Gamma_{\gamma}(IAS) = 11.2(11)$$
  
$$\Gamma = \Gamma_{p} + \Gamma_{\gamma} = 20(5) \text{ eV}$$

The width of the IAS is very narrow as the proton emisión is isospín forbidden, facilitating the emisión of M1 transition of the 5.046 MeV IAS  $B(F) = T(T+1)-T_{zi}T_{zf} = 2x3-(-2)(-1)$ = 6-2 = 4Predicted mixing with 0<sup>+</sup> T= 1 states  $\rightarrow$  No strong feeding observed to states nearby. Closest E = 5425 keV ( $\Delta E$  = 379 keV)

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#### Electron-Neutrino Correlations (pp emitters)

✓ βν correlation depends of the type of the transition

✓ Important probe of the nature of weak interaction

The V-A character of  $\beta$ -decay was determined by measuring the recoil energy spectrum of <sup>6</sup>He [Johnson et al, PR132(63)1149]







- Isospin mixing in Fermi decays
- Contaction sing
  32Cl
  31S
- Lever interferences
- Spin assignment
- Excitation energies

Schardt & Riisager, Z. Phys. A 345 (1993) 265

Adelberger & Garcia, Hyp. Int 129 (2000) 237



#### βv correlation studies: Search for New Physics



• If  $\beta p$  emitter  $\Rightarrow$  measurement of e-v correlation  $\Rightarrow$ F/GT nature of transition from the broadening of proton peak.

• Limit for scalar component in beta decays  $M_S \ge 4.1 M_W$ 

Adelberguer et al., PRL (1999)





# $^{32,33}$ Ar – $\beta$ decay



E-resolution = 8 keV



Angular correlations between e+ and vin the Fermi and GT transitions  $\rightarrow$  the Doppler effect larger recoil broadening of the proton lines for Fermi than for GT decay (Emission before the recoil daughter comes to rest)

 $E^* = Ep(CM) + Sp(33CI) = Ep\frac{\Delta M(32S) + \Delta M(p)}{\Delta M(32S)} + Sp(33CI)$ 



# The decay of <sup>31</sup>Ar

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### $\beta$ -delayed 2p emission from <sup>31</sup>Ar



Diagonal from decays via single intermediate state from many initial states fed in beta-decay

$$E_1 = \frac{M_{D1}}{M_{D1} + m_p} Q_1$$



2p

IAS

β

β

<sup>31</sup>Ar

<sup>29</sup>S+2p

Зр

2p

β

IAS

2.42 1.95 1.39

p

Stable <sup>28</sup>Si+3p

# 2p emission from <sup>31</sup>Ar IAS



# Individual proton projections from $\beta$ 2p in <sup>31</sup>Ar

In sequential two-proton emission the energy of the first proton is

$$E_1 = \frac{M_{D1}}{M_{D1} + m_p} Q_1$$

 $M_{D1} = M(^{30}S)$ 

$$Q_1 = E\left({}^{31}Cl\right) - E\left({}^{30}S\right) - S_{pl}$$
$$E'_2 = \frac{M_{D2}}{M_{D2} + m_p}Q_2$$

$$Q_2 = E\left({}^{30}S\right) - E\left({}^{29}P\right) - S_{p2}$$

$$E_{2} = E_{2}' + \left(\frac{m_{p}}{M_{D1}}\right)^{2} E_{1} - \frac{m_{p}}{M_{D1}} \sqrt{E_{1}E_{2}'} \cos \theta_{2p}$$







#### Decay of the IAS of <sup>31</sup>Ar (Z =18,N=13)

#### E<sub>IAS</sub> = 12322(2)(50) keV from Q2p

 $Q_{FC} = E_{IAS} + \Delta Ec - \Delta np$  $\checkmark \Delta Ec = 1448.8 [ZA^{-1/3}] - 1026.3 \text{ keV}$ Antony & Pape [ADNDT 34 (86) 279] ✓ ∆Ec =7045 keV Leaving the coef. free and using exp. **Coulomb energy shifts** between <sup>32,33,34</sup>Cl and <sup>32,33,34</sup>Ar  $\Delta Ec = 6950(90) \text{ keV} \implies$  $Q_{rc} = 18,49(11) \text{ MeV}$  $f(E_{\beta | AS})t_{| AS} = 6145(4) s / [B(F) + B(GT)]$ b.r.(IAS) =  $T_{1/2} / t_{IAS}$  $B(F) = [T(T+1)-T_{i}T_{f}]\delta_{if} = 5$ Expected b.r. (IAS) = 4.35(31)%







Fynbo et.al. NP A677(2000)38



#### Mapping of Neutron Deficient nuclei 22 < Z < 28



#### From peaks to continua (Hardy, Cargese, 1976)

βp explored high excitation energy in the daughter => individual transition are not longer resolved

$$I(Ep) = \sum_{i,f} f(Z,Q-E^{i})S_{\beta}(E^{i}) \frac{\Gamma_{p}^{ij}}{\Gamma_{p}^{i}+\Gamma_{\gamma}^{i}}$$

To fit the proton spectrum average of the above quantities are considered.



Giovinazzo et al, NPA674 (2000) 394

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βp + X-Ray ratio strongly constraint the level density distribution

Good estimate of proton and gamma widths for exotic nuclei of interest for nucleosynthesis The Porter Thomas distribution accounts of the fluctuations observed in the spectrum





<sup>21</sup> IEM

### Exotic Radioactivities



#### 2p-correlation measured for first time in <sup>45</sup>Fe



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# **Beta-delayed Neutron Emission**







# **Beta-delayed Neutron emitters**



About 220 cases measured, Mainly  $T_{1/2}$  and  $P_n$ -values Spectroscopy hampered by Detection system.

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Compilation for fission products 26 < Z < 58, Pfeiffer, Kratz, Möller, Prog. Nucl. Energy 41(2002)39

$$1/T_{1/2} = \sum_{E_i \ge 0}^{E_i \le Q_{\beta}} S_{\beta}(E_i) \times f(Z, Q_{\beta} - E_i)$$

$$P_n = \frac{\sum_{B_n}^{Q_{\beta}} S_{\beta}(E_i) f(Z, Q_{\beta} - E_i)}{\sum_{0}^{Q_{\beta}} S_{\beta}(E_i) f(Z, Q_{\beta} - E_i)}$$
Kratz-Hermann formula

Beta-Delayed Neutron-Emission Where C is the parameter of pairing, depending of even or odd character of daughter nucleus



	a	b	regr	а	b	χ2
$29 \le Z \le 57$	85.16	3.99	0.83	80.58	4.72	78.23
				$\pm 20.72$	$\pm 0.34$	

# Measurement of Neutrons & βn

- Long Counter: reduced energy to termal values by scattering in parafine.
- Time-of-Flight, giving signals in plastic scintillator. Energy of neutron deduced.
- βn can be deduce by obsevation of γ-ray transition in the beta-delayed neutron daughter.

Talk on Neutron deteccion by JL Tain



#### Beta decay of an exotic n-rich nuclei





#### Beta delayed particle emitters









#### Decay Scheme $\rightarrow$ Structure Information (N= 20)



S. Nummela et al PRC64 054313 (2001)

#### Intruder states & Effective interaction in sd-pf shell





#### $\beta$ n from <sup>133g</sup>In(9/2<sup>+</sup>) and <sup>133m</sup>In(1/2<sup>-</sup>) $\rightarrow$ <sup>133</sup>Sn study

<sup>133</sup>In is a key nucleus for

- Astrophysics due to its placement on the r-process bottleneck regions in most scenarios.
- its proximity to the doubly magic <sup>132</sup>Sn (50 protons and 82 neutrons) offers a uniquely simple β-decay system to validate nuclear theories.

133mIn<sub>84</sub>



N = 82

h11/2

99/2 **9** 

g7/2

339In<sub>84</sub>

High resolution laser tuning allowed Separate the contribution from gs and isomer <sup>133</sup>In

Neutron ToF Spectrum from VANDLE



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# **Neutron Detection systems**





#### Halo nuclei



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✓ Energy threshold effect
✓ Highlight by nuclear reactions
✓ Effects in beta decay



#### Beta-delayed deuterons



#### Beta decay of an exotic nuclei







#### <sup>11</sup>Li, gamma rays







# <sup>11</sup>Li βd spectrum finally measured @ TRIUMF!!



DSSSD 16x16 mm<sup>2</sup>, 70μm thick 48x48 strips, 300 μm, 2304 pixels J. Büscher et al., NIM B 266 (2008) 19

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# Stringent test of Nuclear Models

✓ Test Isobaric Multiplet Mass eq.  $M(A,T,T_z) = a + bT_z + cT_z^2 + \delta(dT_z^3 + eT_z^4)$ ✓ If 2-body forces responsible of charge dependence in nuclei IMME to  $T_z^2$ ✓  $B_F = T(T+1) - T_{Zi}T_{Zf}$ ✓ If strength to IAS  $\neq B_F \iff$  Mixing



✓ Impressive reproduction of the  $B_{GT}$  distribution by Shell Model calculation

 Quenching factor close to one, sensitive to the placement of the GTGR



#### Kinematic identification of $\beta$ t emission in <sup>11</sup>Li



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# Viewing B-delayed triton emission



New branch  $\beta$ n from this state confirming the character of super-allowed transition

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# Beta-delayed Fission in the Pb-region



# Beta-alpha emission

- Ba-Identified first in Natural radiactivity
- Ba favoured in light nuclei with Tz = -1: <sup>8</sup>B, <sup>12</sup>N, <sup>20</sup>Na...
- In cases were both \u03b8a and \u03b8p are allowed =>\u03b8p dominates due to barrier penetrabilities
- Branches of  $\beta a > 1\%$  are only observed in nuclei A < 20 and <sup>118</sup>I
- Some of these states are of astrophysical relevance
  - For instance the  ${}^{16}N(\beta a)$  helped to elucidate the  ${}^{12}C(a,\gamma){}^{16}O$
- βpα or βαp is a decay mode open for <sup>9</sup>C, <sup>13</sup>O, <sup>17</sup>Ne, <sup>21</sup>Mg and <sup>23</sup>Si
   Only identified <sup>9</sup>C and <sup>17</sup>Ne

<sup>9</sup>C special as the daughter is unbound to p-emission. Expectacular asymmetry in some of the mirror transitions with <sup>9</sup>Li

<sup>17</sup>Ne astrophysical relevance to learn about the E2 capture rate in stellar  ${}^{12}C(\alpha,\gamma){}^{16}O$  reaction



# Beta-decay of <sup>17</sup>Ne



2

4

6

Energia (MeV)

8

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1/2-

3/2-

10



- Separation of  $\beta p$ ,  $\beta \alpha$  channels with  $\Delta E$ -E and ToF techniques, allowing independent analysis of the two spectra.
- Bp and Bα branches confimed previous values with higher precision.
- Treatment of the full proton spectrum using R-Matrix.
- Feeding to subthreshold <sup>16</sup>0 states, study of their partial α-widths on progress.

#### Beta Delayed Particle Emission











$$W( heta) = 1 + a rac{p_eta}{E_eta} cos( heta_{eta 
u}), \qquad with \quad a = rac{g_V^2 B_F - rac{1}{3} g_A^2 B_{GT}}{g_V^2 B_F + g_A^2 B_{GT}}.$$

If the decay is followed by particle emission, the recoil of the daughter shifts the energy of the delayed particle by about 10 keV that it is easy to measure.

 $\checkmark$  First used to deduced the nature of the decay of  $^{8}\text{Li}\ \beta2\alpha$ 

$$W = 1 + \frac{1}{2}(3a - A)\frac{p_{\beta}}{E_{\beta}}\cos\theta_{\beta\nu} + \frac{3}{2}(A - a)\frac{p_{\beta}}{E_{\beta}}\cos\theta_{p\beta}\cos\theta_{p\nu}$$
$$A = \frac{g_V^2 B_F - (\frac{1}{3} + \frac{2}{30}\tau\Theta)g_A^2 B_{GT}}{g_V^2 B_F + g_A^2 B_{GT}} \quad \Theta\tau \neq 0 \quad \begin{cases} \mathsf{GT} \\ \mathsf{Ip} \neq 0 \end{cases}$$

 $\beta - \nu$ correlations allow to extract Spins: <sup>31</sup>Ar Fermi / GT character Intrinsic widths of levels Final state (gs/excited) of delayed particle The energy shift in the delayed particle averaged over the neutrino angles

$$< t >_{\nu} = -k\cos\theta_{p\beta} \left( 1 + \frac{1}{3}A\frac{p_{\nu}c}{E_{\beta}} \right) p_{\beta}c$$

$$(1) \int_{2000}^{2000} \int_{1}^{31}Ar \int_{1}^{1} \int_{1}^{4} \int_{1}^{4} \int_{1}^{4} \int_{1}^{3170} \int_{1}^{33}Ar \int_{1}^{4} \int_{1}^{4}$$

Thaysen et al, Phys. Lett B 467 (1999) 194



# <sup>11</sup>Li (βd) @ ISOLDE





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