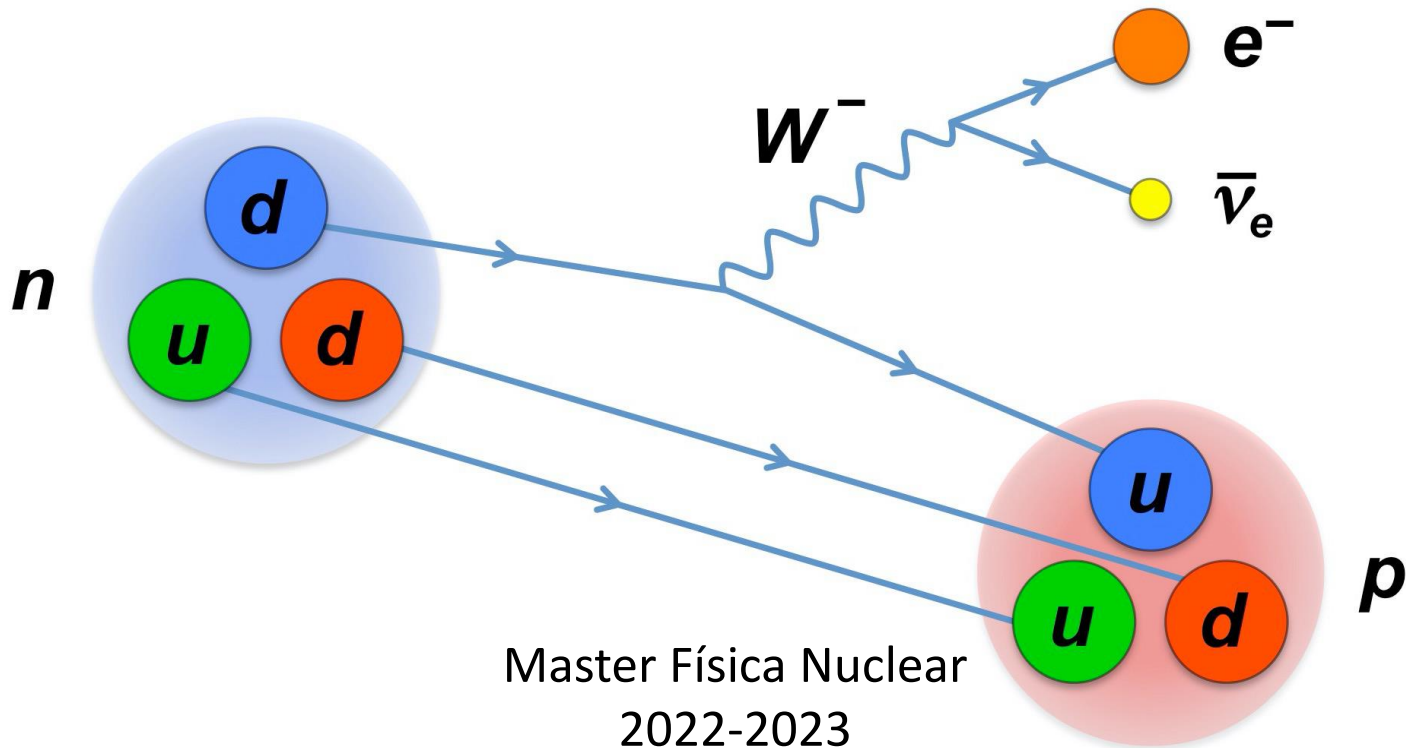


# Beta decay experiments



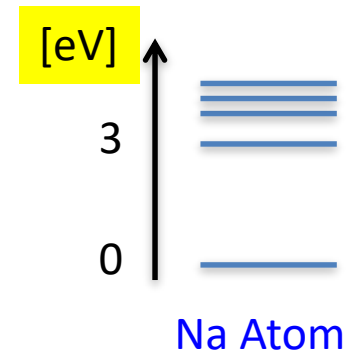
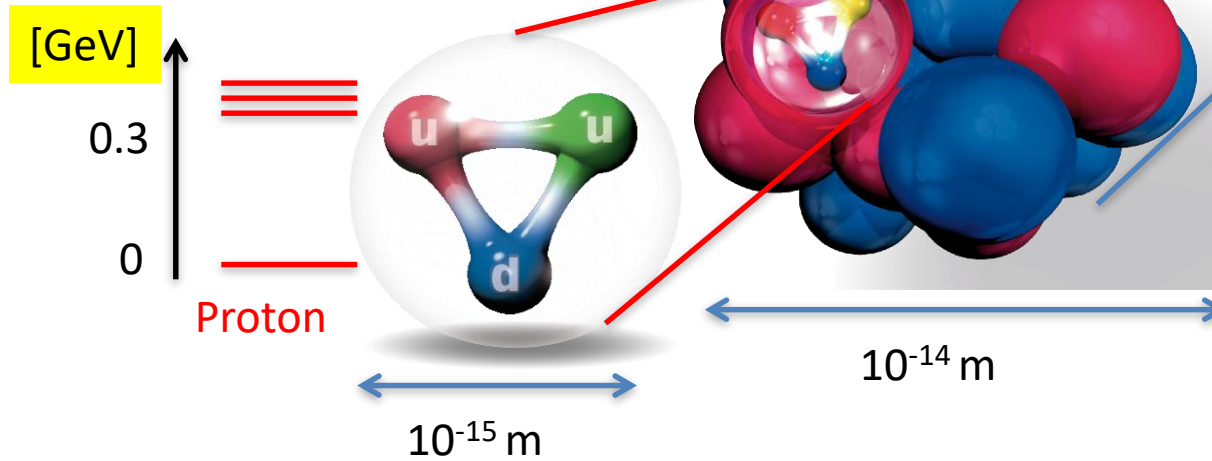
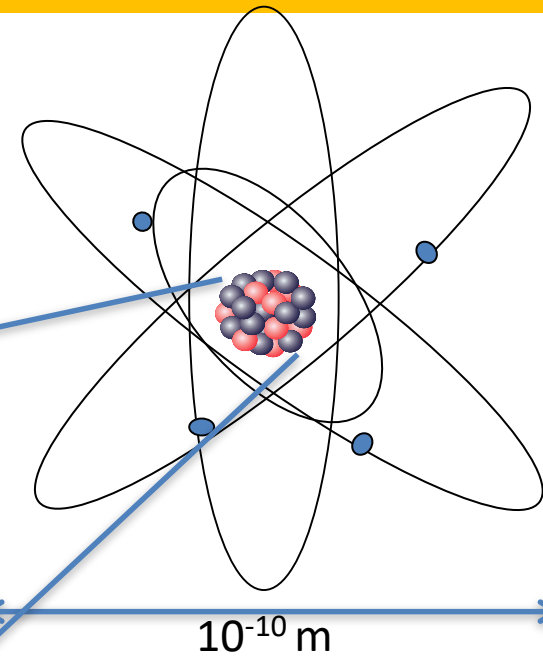
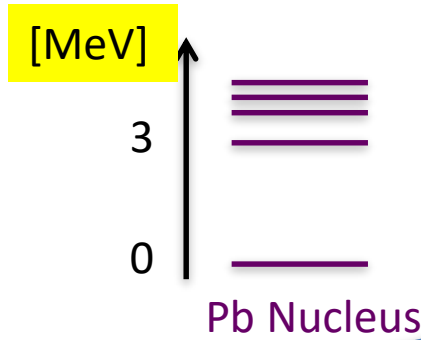
Master Física Nuclear  
2022-2023  
Bruno Olaizola  
CSIC-IEM

# Table of contents

- Basic concepts of beta decay theory
  - Mass parabola
  - Beta particle energy spectrum
  - Neutrino mass
  - Neutrinoless double beta decay
  - $\text{Log}(ft)$
- Beta decay experimental setup
  - GRIFFIN
    - Main array
    - Ancillary detectors
- Some experiments using beta decay

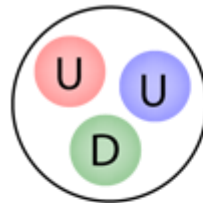
# Subatomic Structure

The nuclei: **protons** (red)  
**neutrons** (blue), each  
consists of 3 quarks  
connected by gluons.



# Particle masses

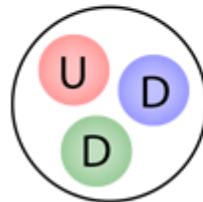
- Mass 938.3 MeV\*



Proton

U = "up" quark  $+\frac{2}{3} e$   
D = "down" quark  $-\frac{1}{3} e$

- Mass 939.6 MeV\*

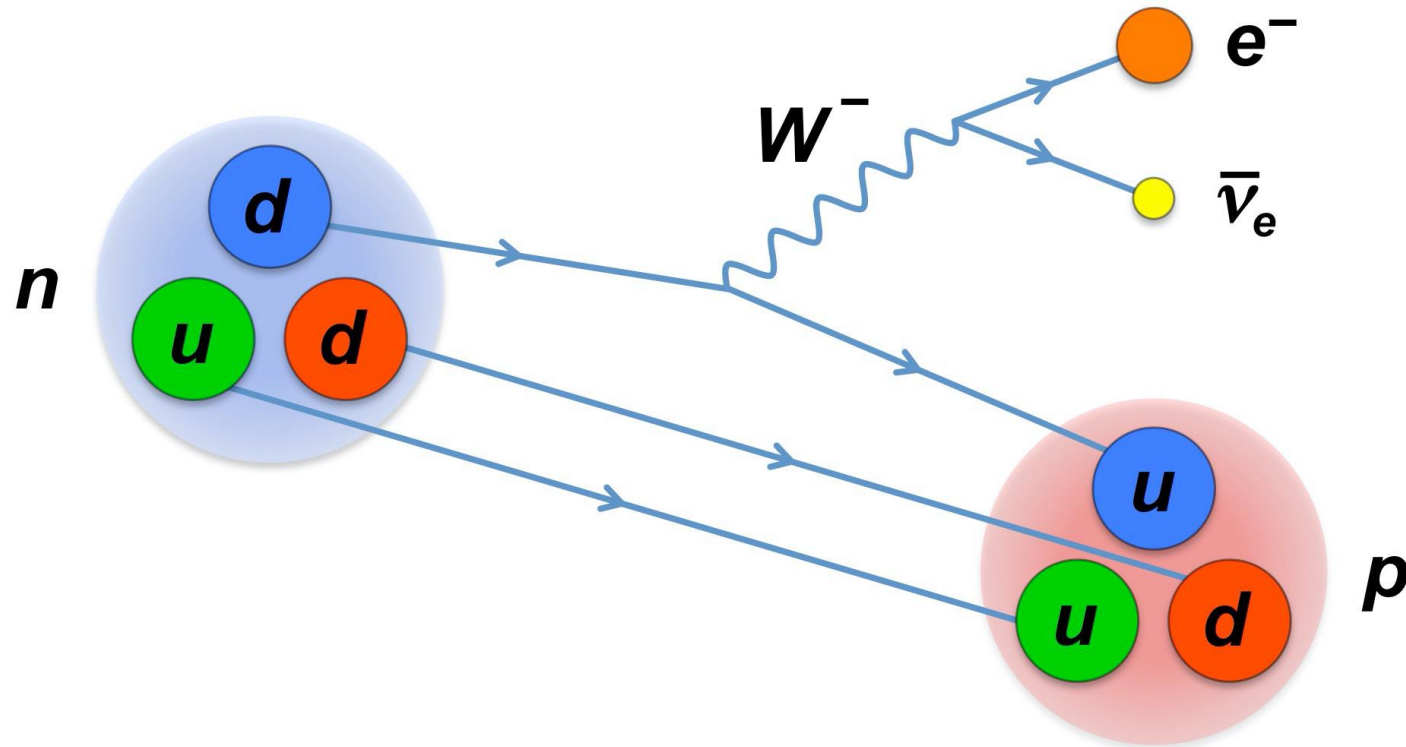


Neutron

- Mass 1.7-3.3 MeV\*
- Mass 4.1-5.8 MeV\*

\*Independent particle rest mass which, in the case of quarks, is meaningless

# Beta decay



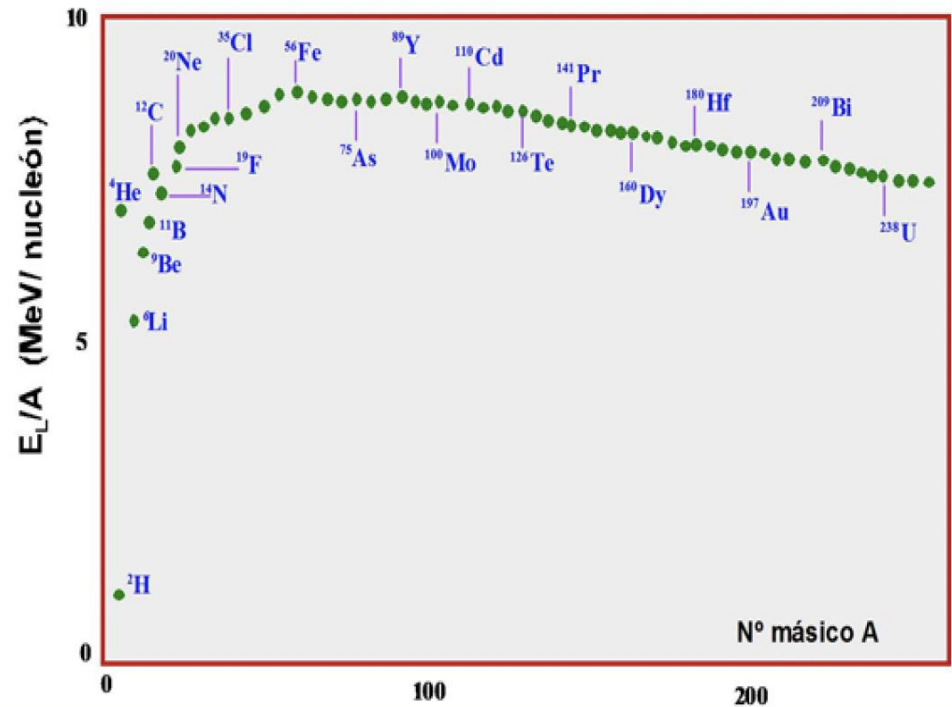
- Free neutrons decay into protons with  $\tau \sim 900$  s
- Due to charge conservation, an electron is emitted  $e^-$
- Due to lepton number conservation, and antineutrino is also emitted  $\bar{\nu}_e$
- Free protons are stable

# Nuclear mass

$$m_N < Z \cdot m_p + N \cdot m_n$$

$$m_N = Z \cdot m_p + N \cdot m_n - B/c^2$$

- B is the nuclear binding energy
- It is a residue of the strong interaction between quarks
- $B/A \sim 8$  MeV on average
- Maximum around  $^{62}\text{Ni}$
- Explains stellar nuclear reactions
  - Element abundance
- Nuclear reactors



# Semi-empirical mass formula

$$B(A, Z) = a_v A - a_s A^{2/3} - a_c Z(Z-1)A^{-1/3} - a_{sym} \frac{(A-2Z)^2}{A} + \delta a_p A^{-3/4}$$

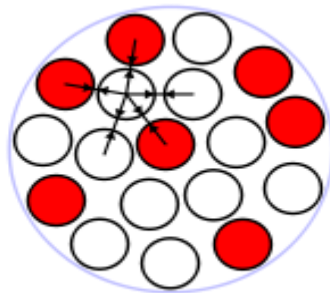
↗  
volume

↑  
surface

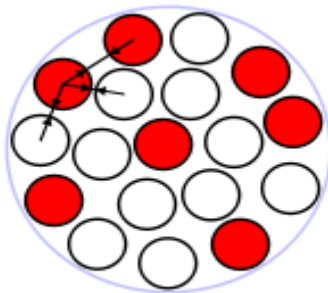
↑  
Coulomb

↑  
symmetry

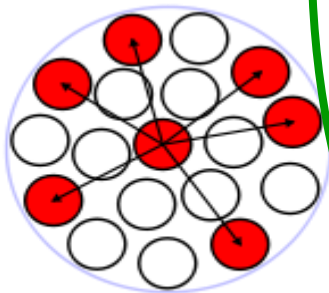
↖  
pairing



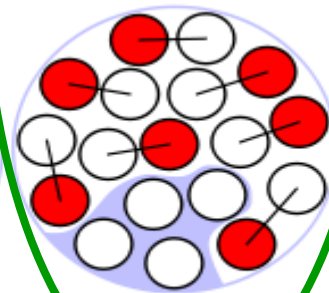
Volume



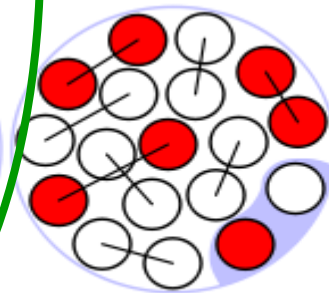
Surface



Coulomb



Asymmetry



Pairing

[https://en.wikipedia.org/wiki/Semi-empirical\\_mass\\_formula](https://en.wikipedia.org/wiki/Semi-empirical_mass_formula)

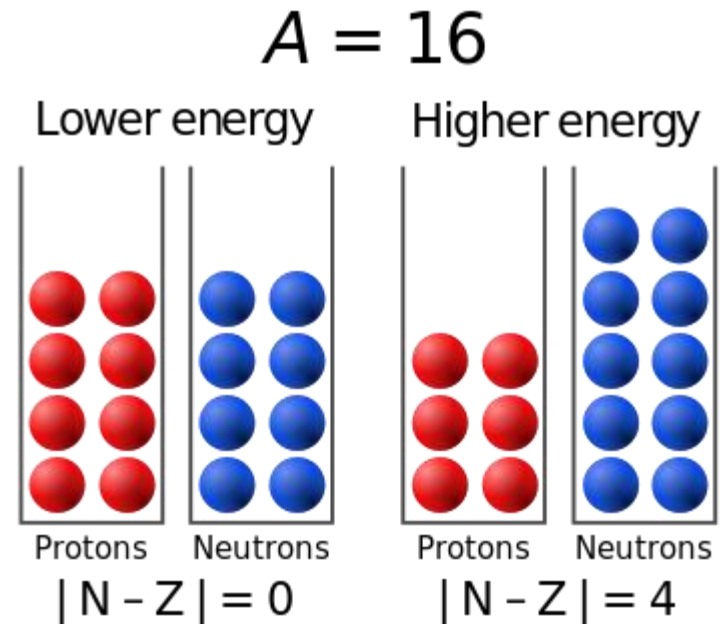
# Assymetry term

- Also known as *Pauli term*
- It originates from the Pauli exclusion principle and the Shell structure
- An excess of neutrons would need to occupy higher energy orbitals

- It can be expressed as:

$$a_A \frac{(N-Z)^2}{A}$$

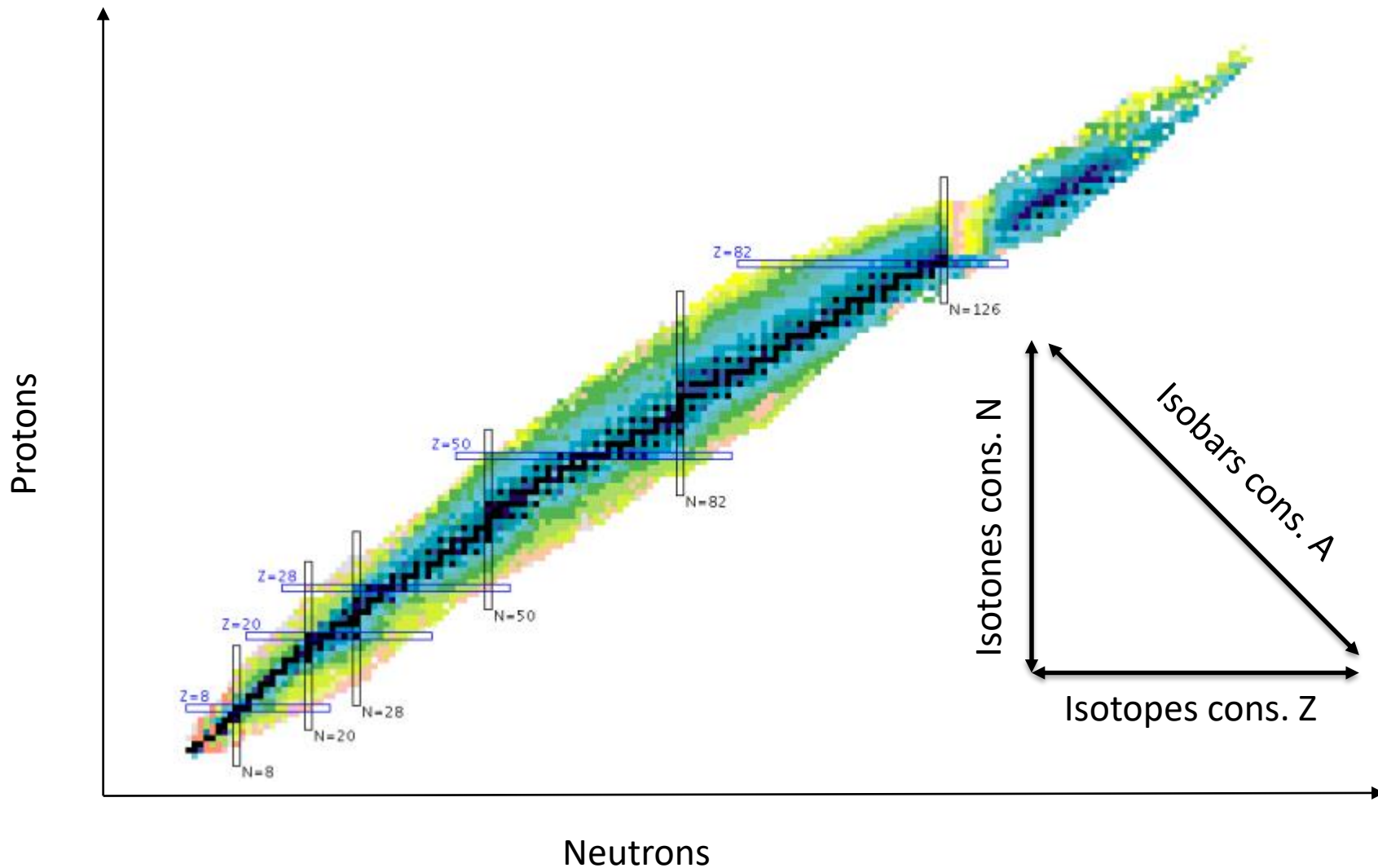
- Notice the square dependence with the difference between protons and neutrons
- This can also be derived by modeling the nucleus as a Fermi ball of nucleons



[https://en.wikipedia.org/wiki/Semi-empirical\\_mass\\_formula](https://en.wikipedia.org/wiki/Semi-empirical_mass_formula)



# Nuclear chart



<https://www.nndc.bnl.gov/>

# Nuclear stability

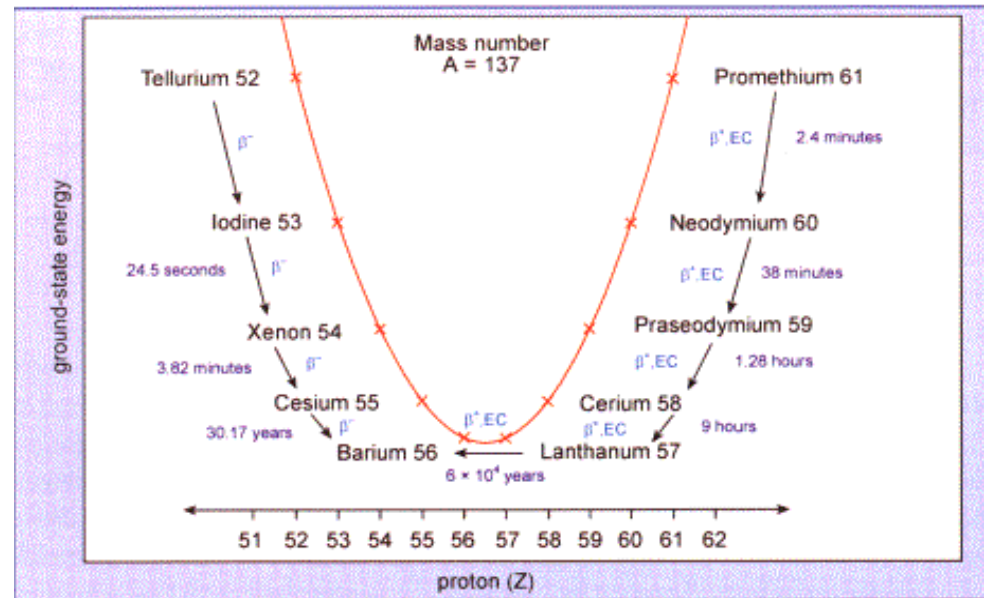
- Due to the asymmetry term, in isobar lines the masses will follow a parabola
- Nuclei will decay into the valley of stability
- This valley is not at  $N=Z$  mainly due to the Coulomb repulsion between protons
- Protons inside the nucleus can be unstable, they decay into a neutron by  $\beta^+$

Thus for each A-value one can calculate the nucleus with lowest mass (largest binding energy):

For a given A a parabolic behaviour of the nuclear masses show up.

**odd-A** only one stable nucleus. The rest  $\beta^\pm$  decay towards the only stable nucleus.

**even A** both even-even and odd-odd  $\Rightarrow 2$  parabolas implied by the mass equation.



# Beta Decay: definition

**Beta Decay:** universal term for all weak-interaction transitions between two neighboring isobars

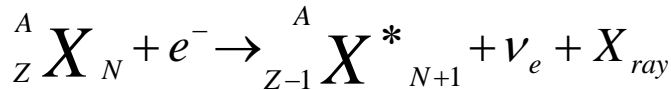
Takes place in 3 different forms

$\beta^-$ ,  $\beta^+$  & EC (capture of an atomic electron)

$\beta^+$ :  $p \rightarrow n + e^+ + \nu$



EC:  $p + e^- \rightarrow n + \nu$



|  |                                  |   |
|--|----------------------------------|---|
| ${}^{185}_{93}\text{Os}$<br>93.0 d<br>$\beta^+$  | ${}^{186}_{93}\text{Os}$<br>1.59 | ${}^{187}_{93}\text{Os}$<br>1.6                 |
| ${}^{184}_{75}\text{Re}$<br>38.00 d<br>$\beta^+$ | ${}^{185}_{75}\text{Re}$<br>37.4 | ${}^{186}_{75}\text{Re}$<br>3.72 d<br>$\beta^-$ |
| ${}^{183}_{74}\text{W}$<br>14.31                 | ${}^{184}_{74}\text{W}$<br>30.64 | ${}^{185}_{74}\text{W}$<br>7.10 d<br>$\beta^-$  |

$\beta^-$ :  $n \rightarrow p + e^- + \tilde{\nu}$

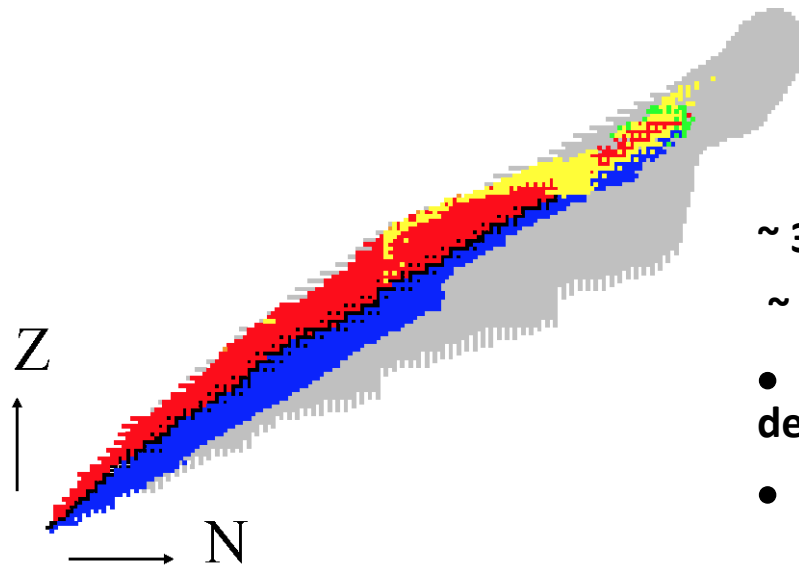


a nucleon inside the nucleus is transformed into another

**No preformation!**

# Atomic Mass Model

## Relationship with Nuclear Decay Models



- 167 e-e
- 254 Stable nuclei
- 4 o-o (2H, 6Li, 10B, <sup>14</sup>N)
- 83 e-o

34 Primordial ( $T_{1/2} > 10^9\text{y}$ )

~ 3000 produced in nuclear reactions

~ 7000 predicted to be bound

- Decay characteristics of most radioactive nuclei determined by  $\beta$ -decay i.e. weak interaction
- For heavier nuclei  $\rightarrow$  Electromagnetic interaction important  $\rightarrow$ 
  - $\alpha$ -decay
  - fission

- Moving away from stable nuclei by adding protons or neutrons  $\rightarrow$   
until the particle drip-lines ( $S_p = 0$  or  $S_n = 0$ ).

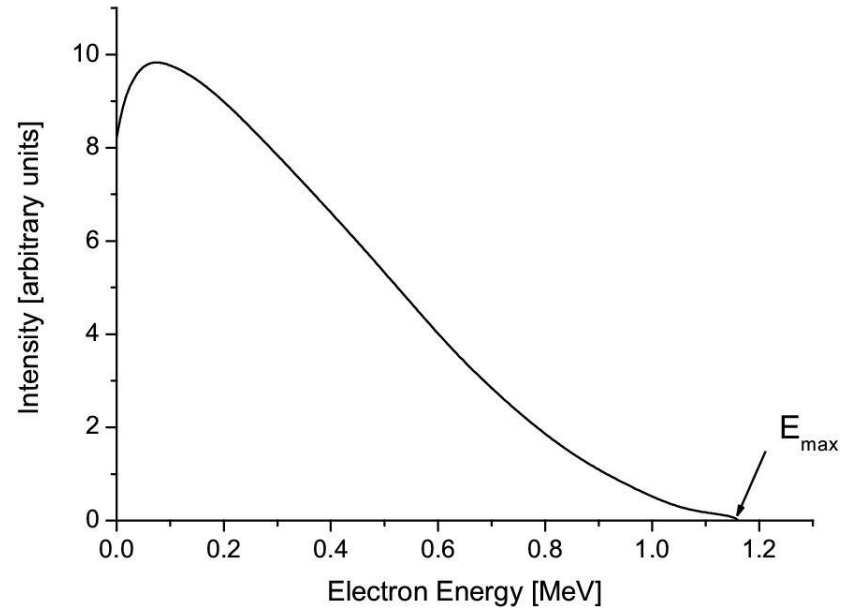
Nuclei beyond drip-line are unbound to nucleon emission, i.e. Strong interaction cannot bind one more nucleon to the nucleus

# Beta decay energy spectrum

- Energy spectrum of beta particle is continuous
- No discrete peaks
- This is why neutrinos were suggested
- 3-body decay: beta, (anti)neutrino and recoiling nucleus
- Energy mass difference is split among the 3 products
- Total energy available:

$$Q = (m_X - m_{X'} - m_\nu - m_e)c^2 = E_{X'} + E_\nu + E_e$$

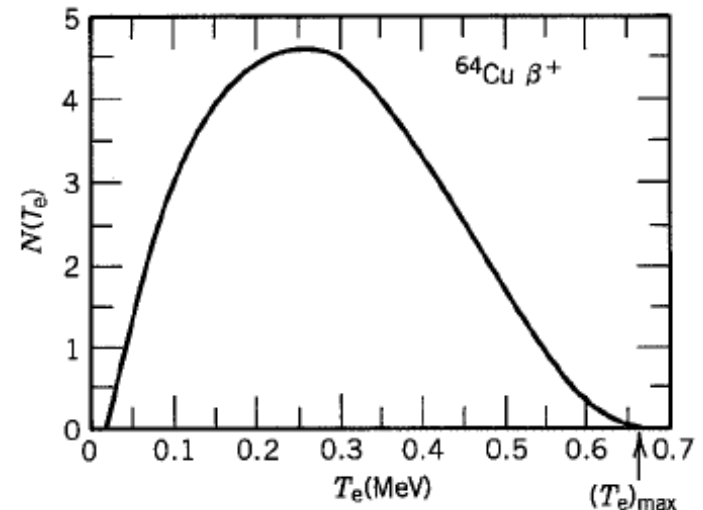
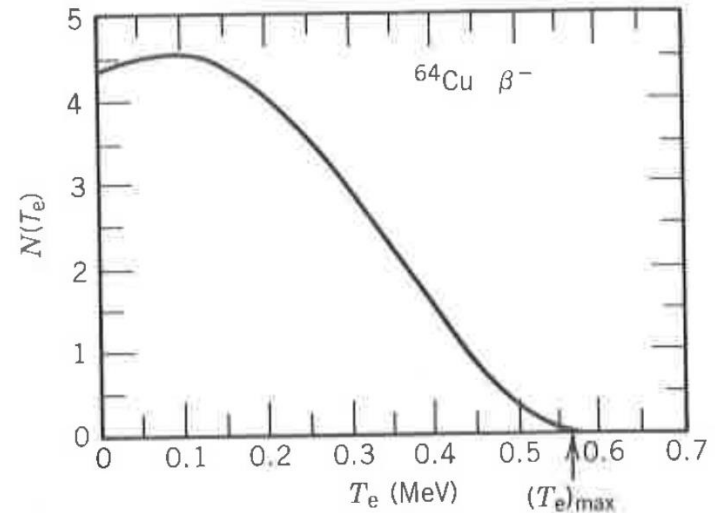
Typical Q is few MeV



# Energy spectra features

- $E_{\max} \sim Q_{\beta}$ , the  $\beta$  particle takes all the available energy; daughter nucleus and neutrino at rest
- There is a most likely value for the  $\beta$  particle energy  $0.3-0.4Q_{\beta}$
- $\beta^{-}$  can be emitted at  $\sim 0$  energy
- $\beta^{+}$  can not be emitted at  $\sim 0$  energy
  - The attraction or repulsion of the protons in the nucleus is not negligible

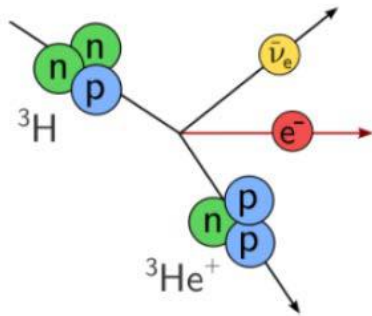
$$N(T_e) = \frac{C}{c^5} F(Z, Q) |V_{fi}|^2 [Q - T_e]^2 \sqrt{T_e^2 + 2T_e m_e c^2} (T_e + m_e c^2)$$



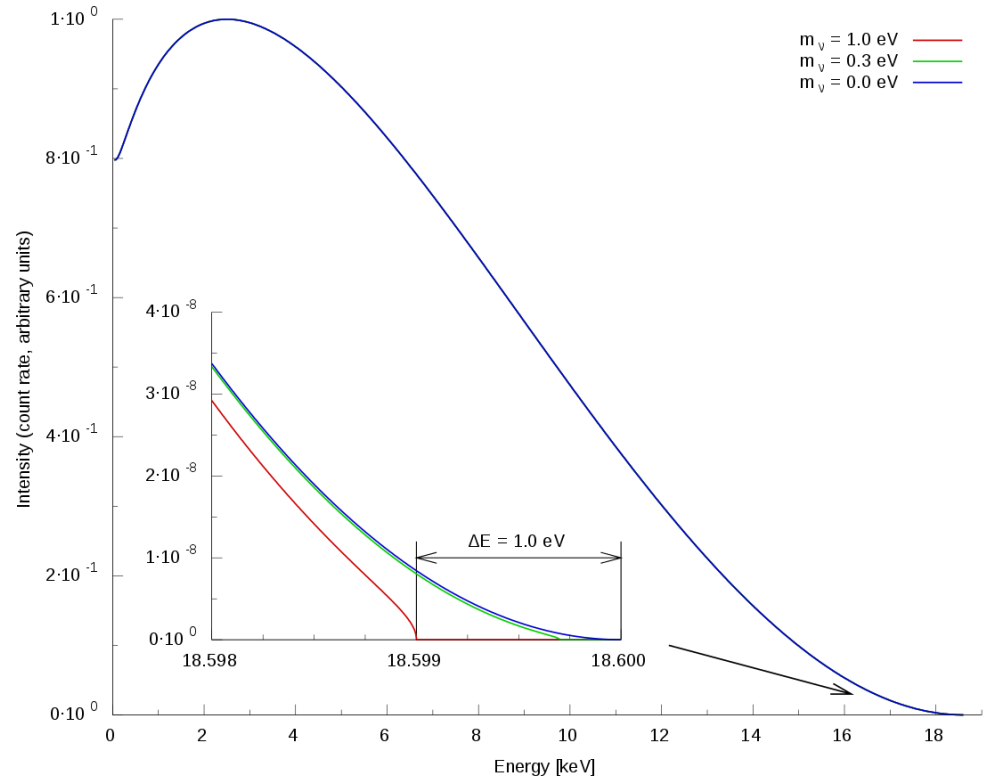
# Neutrino mass measurement

$$\frac{dN}{dE} = C \cdot F(E,Z) \cdot p(E+m_e) \cdot (E_0 - E) \cdot \sqrt{(E_0 - E)^2 - m_\nu^2}$$

$$m_{\nu_e}^2 = \sum_{i=1}^3 |U_{ei}|^2 m_i^2$$



- If we know with great precision the nuclear masses, we know  $Q_\beta$
- A precise measurement of the end point is sensitive to  $m_\nu$
- Model independent method



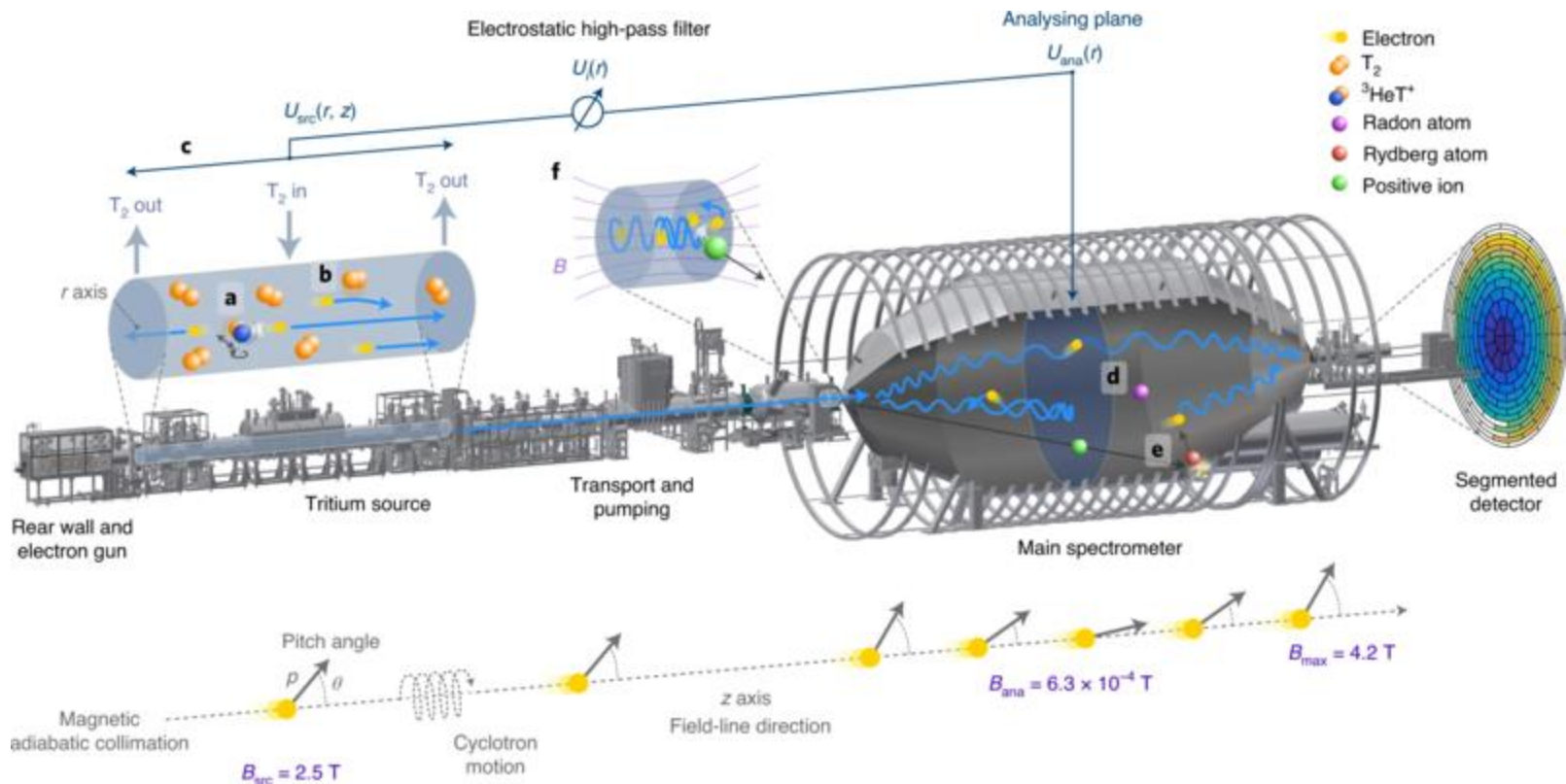
[https://indico.cern.ch/event/572149/contributions/2488095/attachments/1446803/2228876/ALPS2017\\_fraenkle.pdf](https://indico.cern.ch/event/572149/contributions/2488095/attachments/1446803/2228876/ALPS2017_fraenkle.pdf)

# KATRIN (Karlsruhe Tritium Neutrino Experiment)





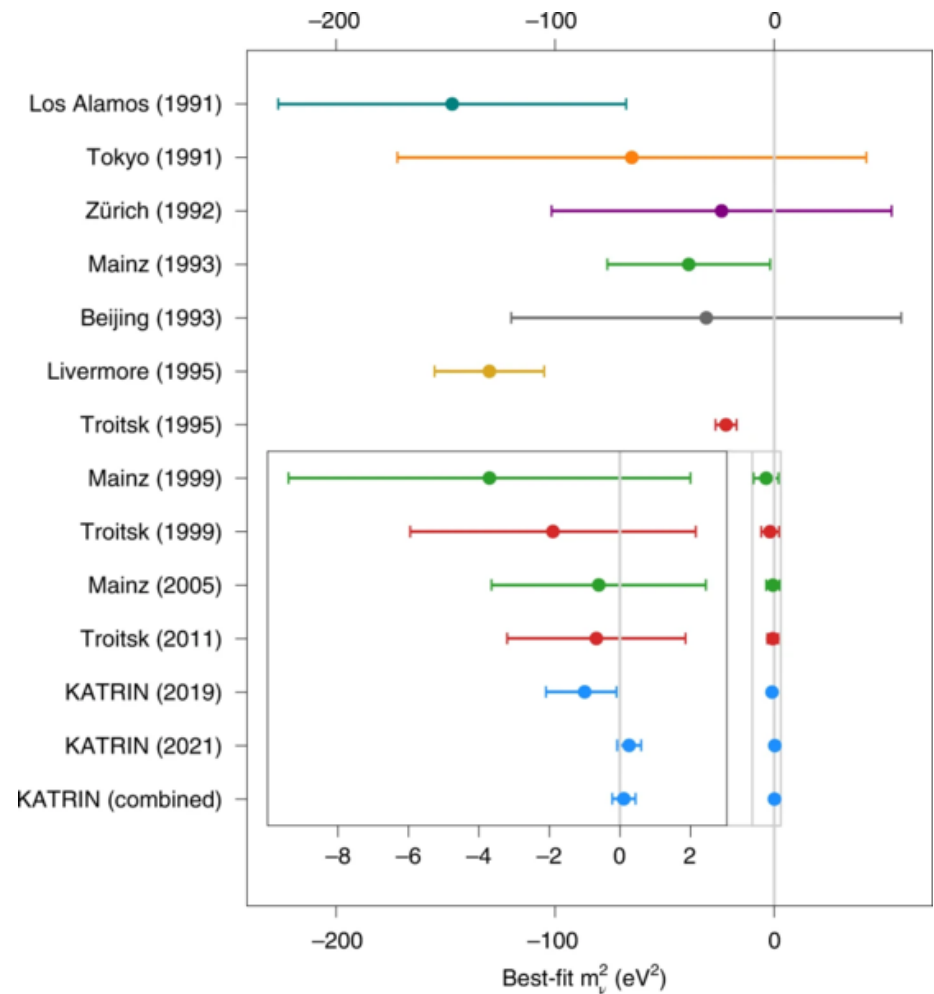
# KATRIN (Karlsruhe Tritium Neutrino Experiment)



- Molecular tritium source ( $T_{1/2}=12.3$  years and  $Q_{\beta}=18.59202(6)$  keV)
- Electric field to filter low energy  $\beta$  particles
- Magnetic field to conduct them to the detector

# KATRIN (Karlsruhe Tritium Neutrino Experiment)

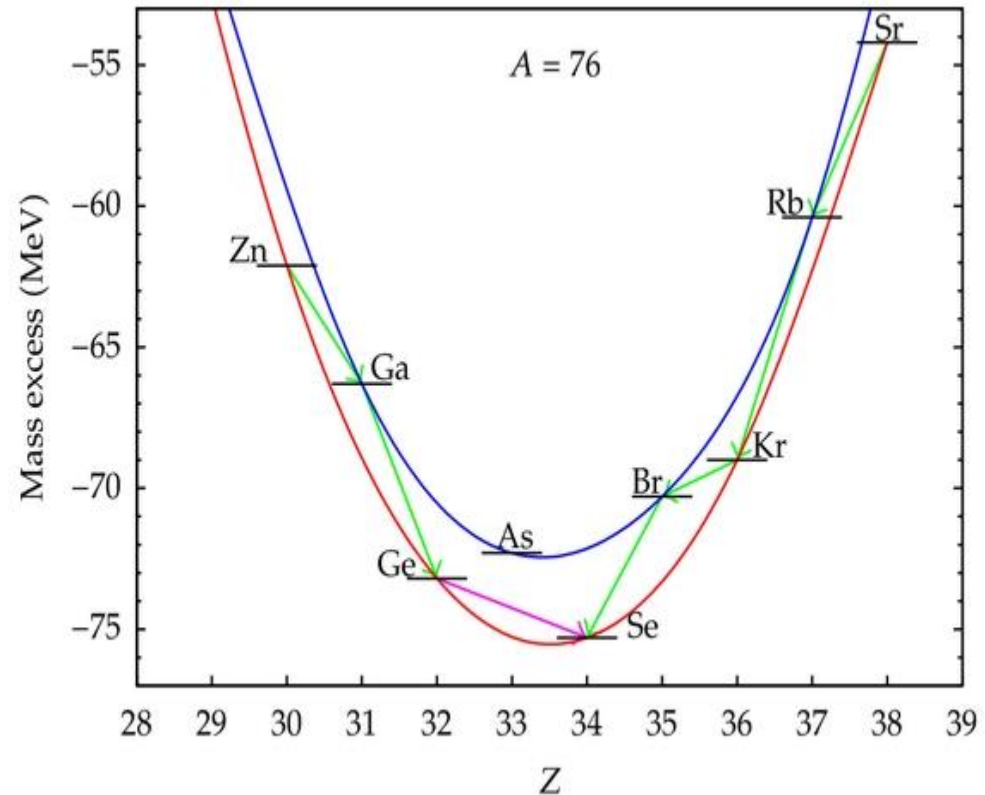
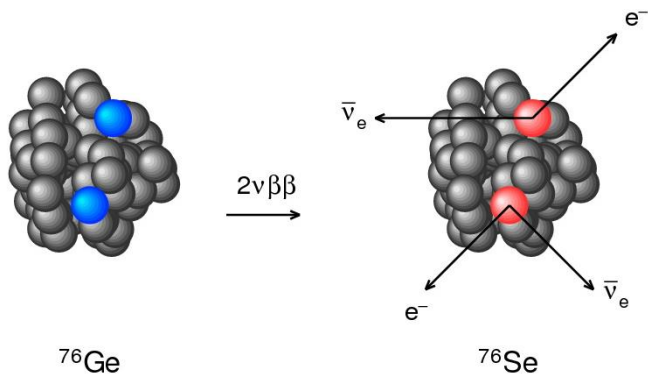
- Up to now ~65 days of measurement
- Out of ~1000 days planned
- Current upper limit of  $m_\nu < 0.8$  eV
- Plan to reach  $m_\nu < 0.2$  eV
- Model independent measurement, but relies on theoretical calculation
  - i.e. molecular dynamics



*Nat. Phys.* **18**, 160–166 (2022). <https://doi.org/10.1038/s41567-021-01463-1>

# Double beta decay

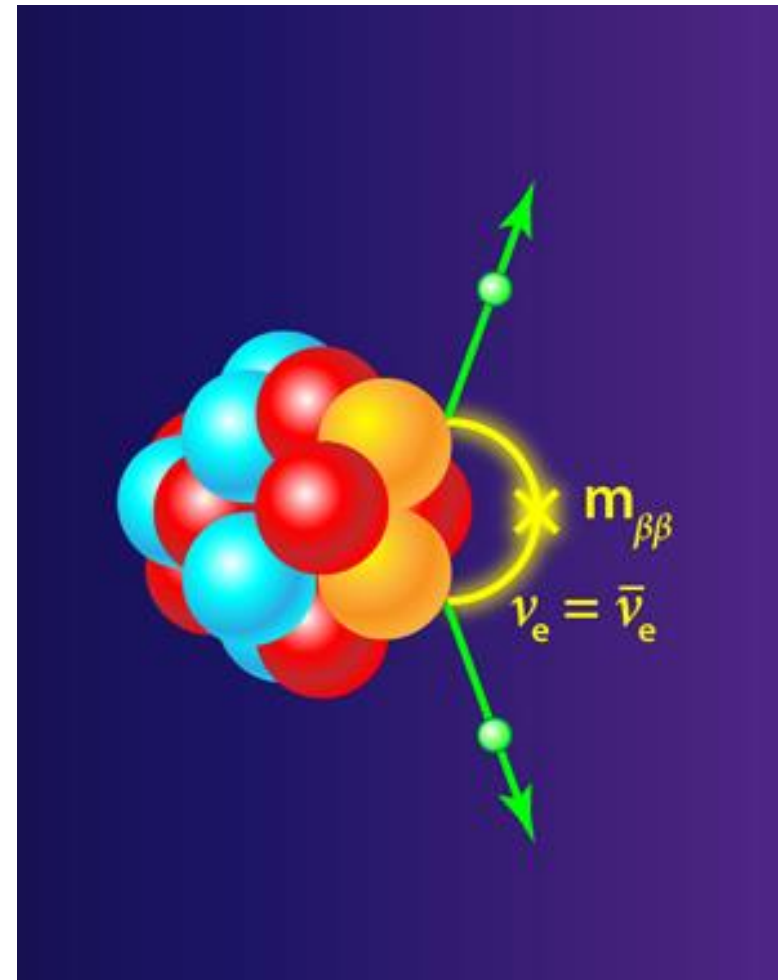
- There are a few cases for which simple  $\beta$  decay is not energetically allowed, but double  $\beta$  decay is.
- Has been observed in about dozen cases, i.e.,  ${}^{76}\text{Ge} \rightarrow {}^{76}\text{Se}$
- $T_{1/2} \sim 10^{18}\text{-}10^{21}$  years



*J. Mendez PhD thesis*

# Neutrinoless double beta decay

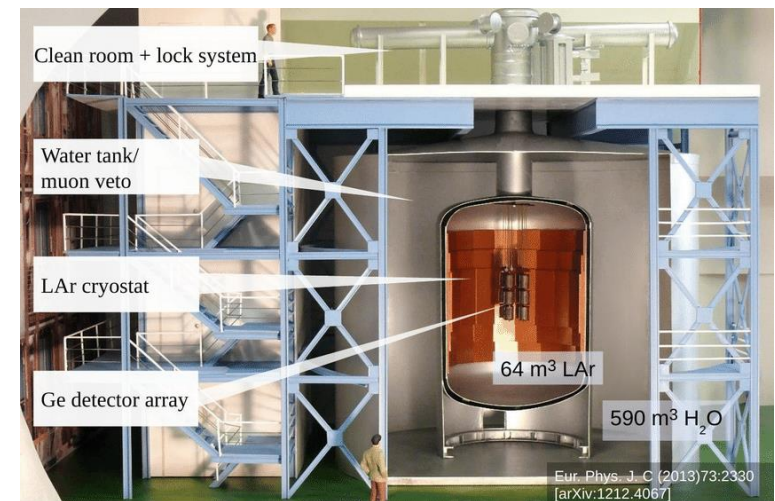
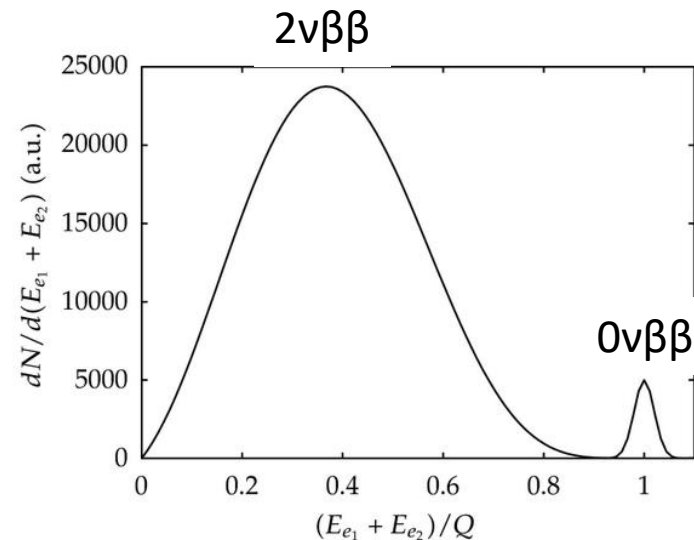
- If neutrinos have masses (they oscillate, so yes) they could be Majorana particles
- Neutrinos would be their own antiparticle  $\nu_e = \bar{\nu}_e$
- This decay would violate lepton number conservation (+2 from the emitted  $e^-$ )
- Would allow for absolute measurement of  $m_\nu$
- Imply physics beyond the standard model



Physics 11, 30 (2018)

# Neutrinoless double beta decay

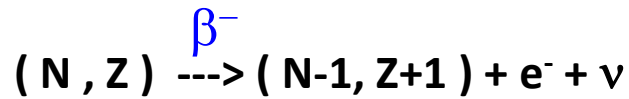
- In neutrinoless double beta decay, both electrons carry all the energy
  - Sum of both energies = total energy
- Very low probability
  - Large amount of radioactive material
  - Superb sensitivity and efficiency
  - Long collection time
- Ultra low background (i.e. underground)
- No event has been observed so far
  - $T_{1/2} > 10^{26}$  years (universe age  $\sim 10^{10}$  years)



<https://www.mpi-hd.mpg.de/gerda/>

# Beta decay process

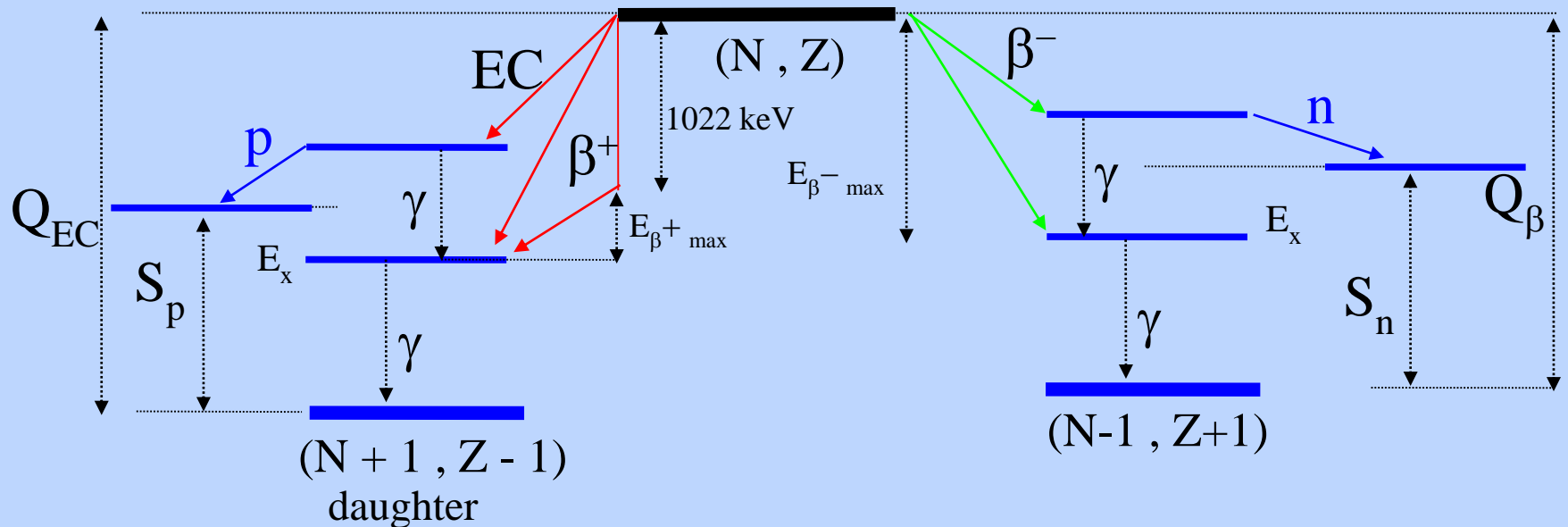
Process mediated by the weak interaction between two isobars



$$M(Z) - M(Z+1) = E_\beta + E_\nu + E_x$$



$$M(Z) - M(Z-1) = E_{\beta^+} + E_\nu + 1022 + E_x$$



# Fermi's golden rule

$$\lambda_{if} = \frac{2\pi}{\hbar} |M_{if}|^2 \rho_f$$

Transition rate or probability of decay per second

Matrix element for the (weak) interaction

Density of final states

$$M_{if} = \langle f|M|i\rangle = \int \psi_f^* \psi_e^* \psi_\nu^* M \psi_i dV$$

# Beta decay $ft$ value

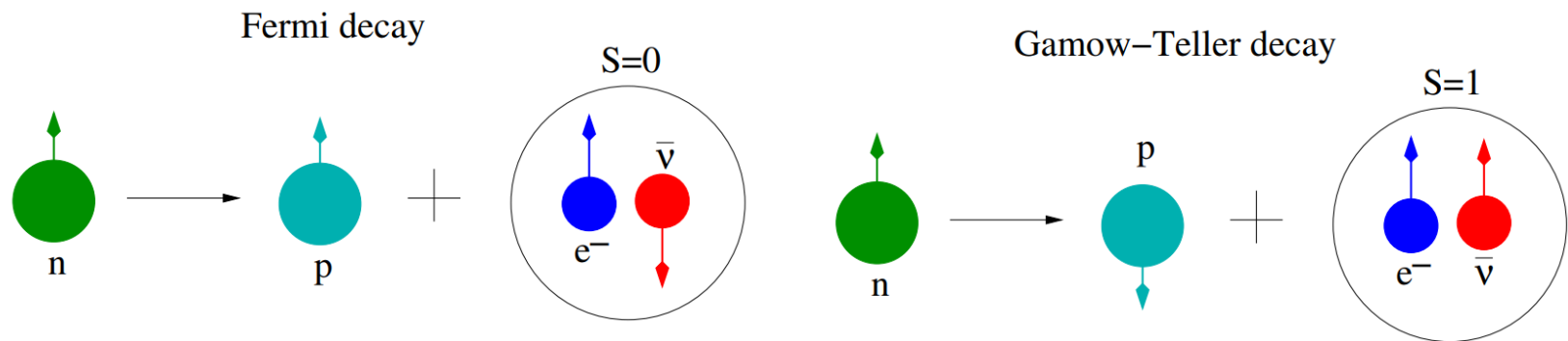
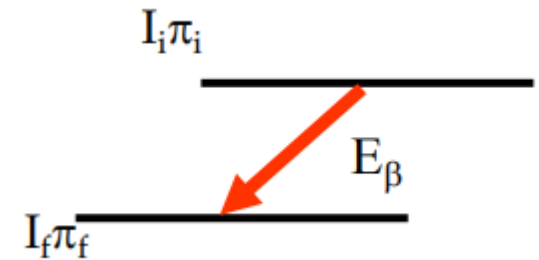
$$ft = f(Q) \frac{t_{1/2}}{BR} = \frac{2\pi^3 \hbar^7 \ln(2)}{g^2 m_e^5 c^4 |\overline{M'}_{if}|^2} = \frac{K}{g^2 |\overline{M'}_{if}|^2}$$

- $f$  is the phase-space integral, which is related to the statistical rate factor, and is a function of the  $Q$  value. It is tabulated. It favours populating low energy states.
- Partial half-life of a specific decay branch  $t_{1/2}^{\beta_i} = \frac{T_{1/2}^{exp}}{P_{\beta_i}}$
- If we measure  $Q$ ,  $BR$  and  $t_{1/2}$ , we have  $\log(ft)$ .



# Beta decay momentum conservation

- $\vec{L}_\beta = \vec{l}_e + \vec{l}_\nu$ ,  $\vec{S}_\beta = \vec{s}_e + \vec{s}_\nu$  and  $\vec{J}_\beta = \vec{L}_\beta + \vec{S}_\beta$
- $\vec{J}_i = \vec{J}_f + \vec{L}_\beta$
- J is the total angular momentum
- L is the orbital angular momentum
- S is the spin
  - Fermi  $\vec{S}_e + \vec{S}_\nu = 0$  (antiparallel)
  - Gamow-Teller  $\vec{S}_e + \vec{S}_\nu = 1$  (parallel)



# Beta decay selection rules

## $L$ Classification

$L_\beta = 0 \rightarrow$  Allowed

$L_\beta = 1 \rightarrow$  First Forbidden

$L_\beta = 2 \rightarrow$  Second Forbidden

$L_\beta = 3 \rightarrow$  Third Forbidden

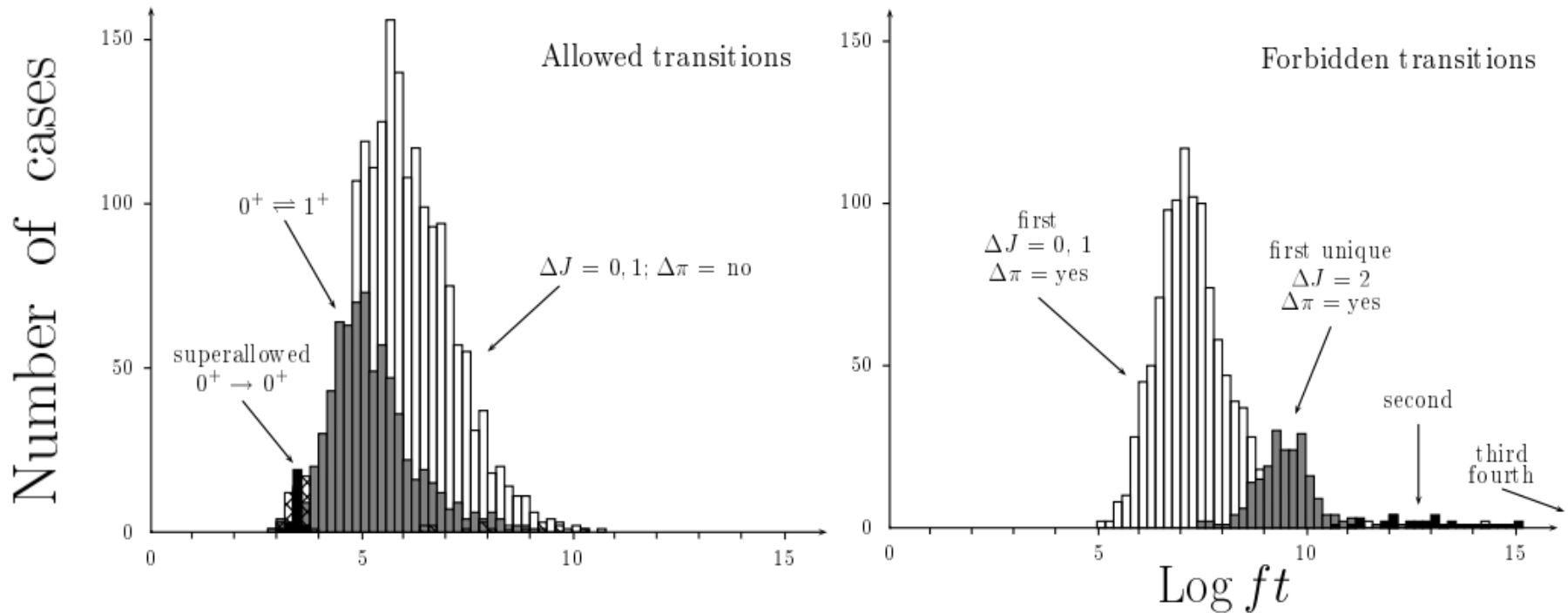
## $S$ Classification

$$\vec{S}_\beta = \vec{S}_{e^+} + \vec{S}_{\nu_e} = \begin{cases} 0 & \uparrow\downarrow \text{ or } \downarrow\uparrow & V (F) \\ 1 & \uparrow\uparrow \text{ or } \downarrow\downarrow & A (GT) \end{cases}$$

- $0^+ \rightarrow 0^+$  not allowed in GT, because  $S_\beta=1$  and therefore  $\Delta L \neq 0$

# Beta decay selection rules

B. Singh et al., Nucl. Data Sheets 84, 487 (1998)



- We can use  $\log(ft)$  to infer the spin-parity of daughter nucleus states
- Learn about nuclear structure

# Beta decay experiment

- Selection rules tell us which states we will “most likely” populate in the daughter nucleus
- $\text{Log}(ft)$  favours populating low energy states
- No need to re-accelerate beams
  - Higher intensity
  - No Doppler corrections
  - Easier experiment

# Beta decay experiment

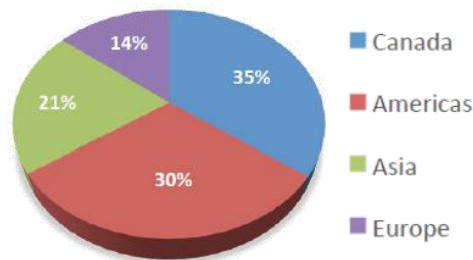
- Now we have all the ingredients to perform beta decay experiments and learn nuclear structure
- I will use GRIFFIN at TRIUMF as example, current state-of-the-art spectrometer
- More facilities are following its design



400 staff  
200 students &  
post-doctoral  
researchers

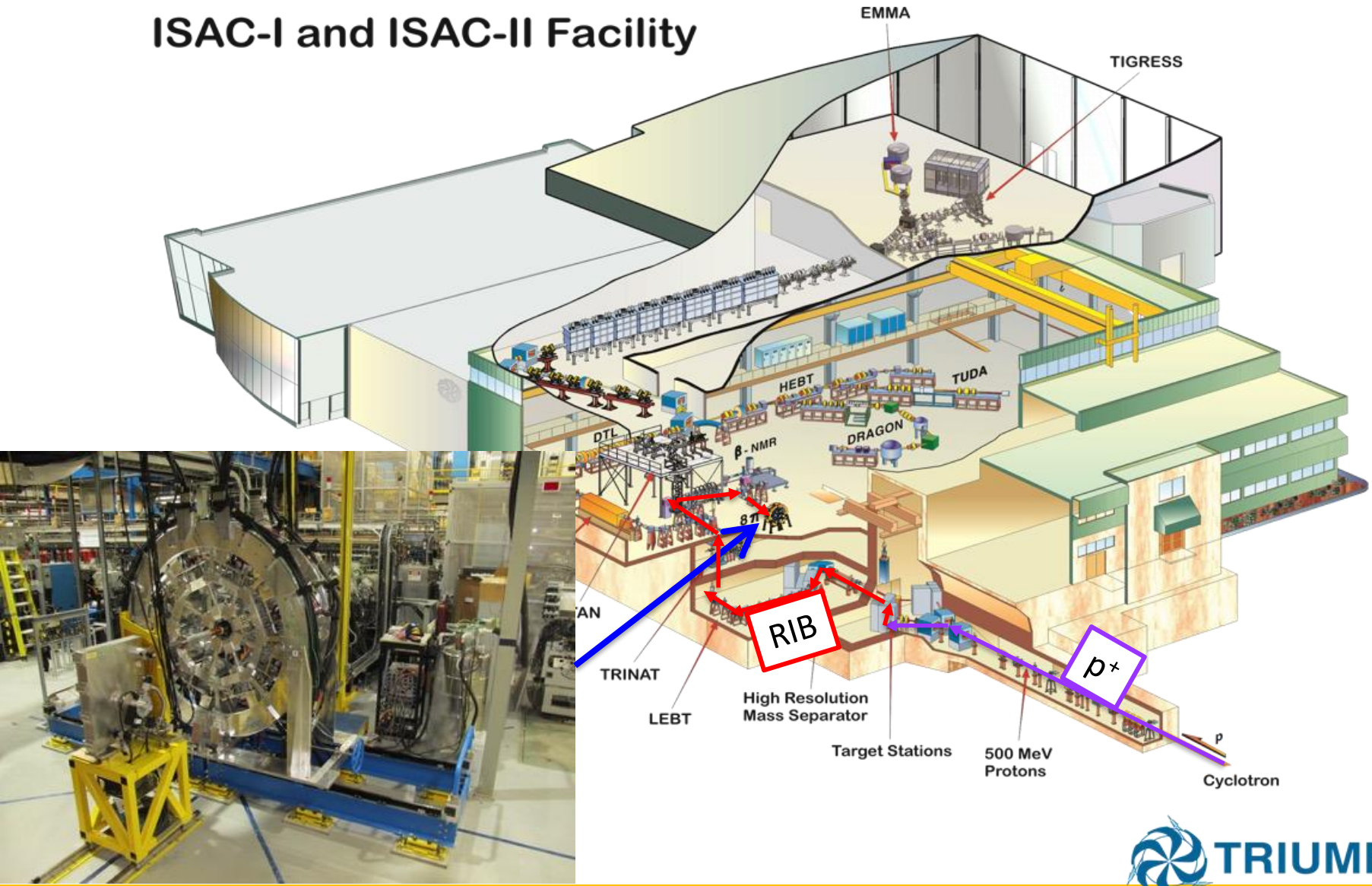


650 scientist &  
student researcher  
visits per year

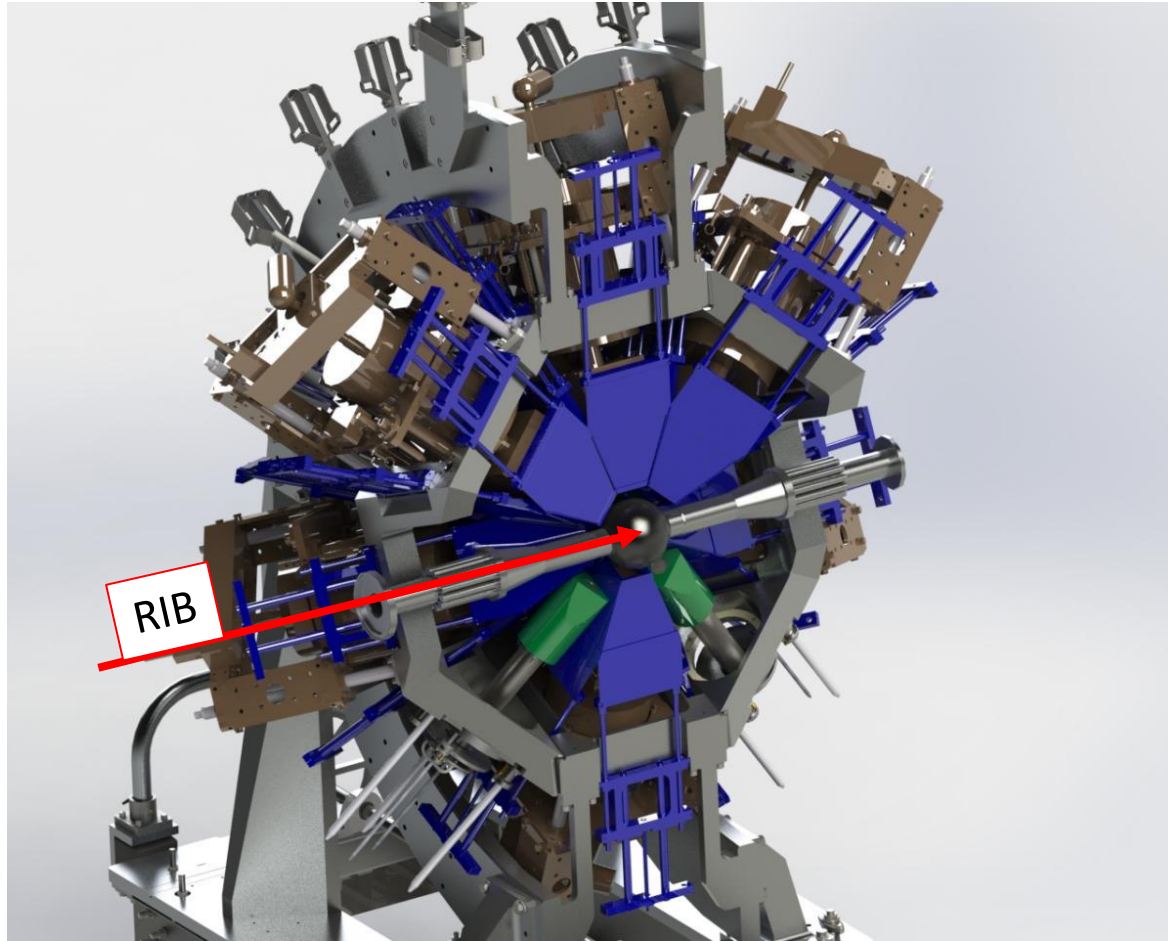


# ISAC@TRIUMF

## ISAC-I and ISAC-II Facility



# GRIFFIN

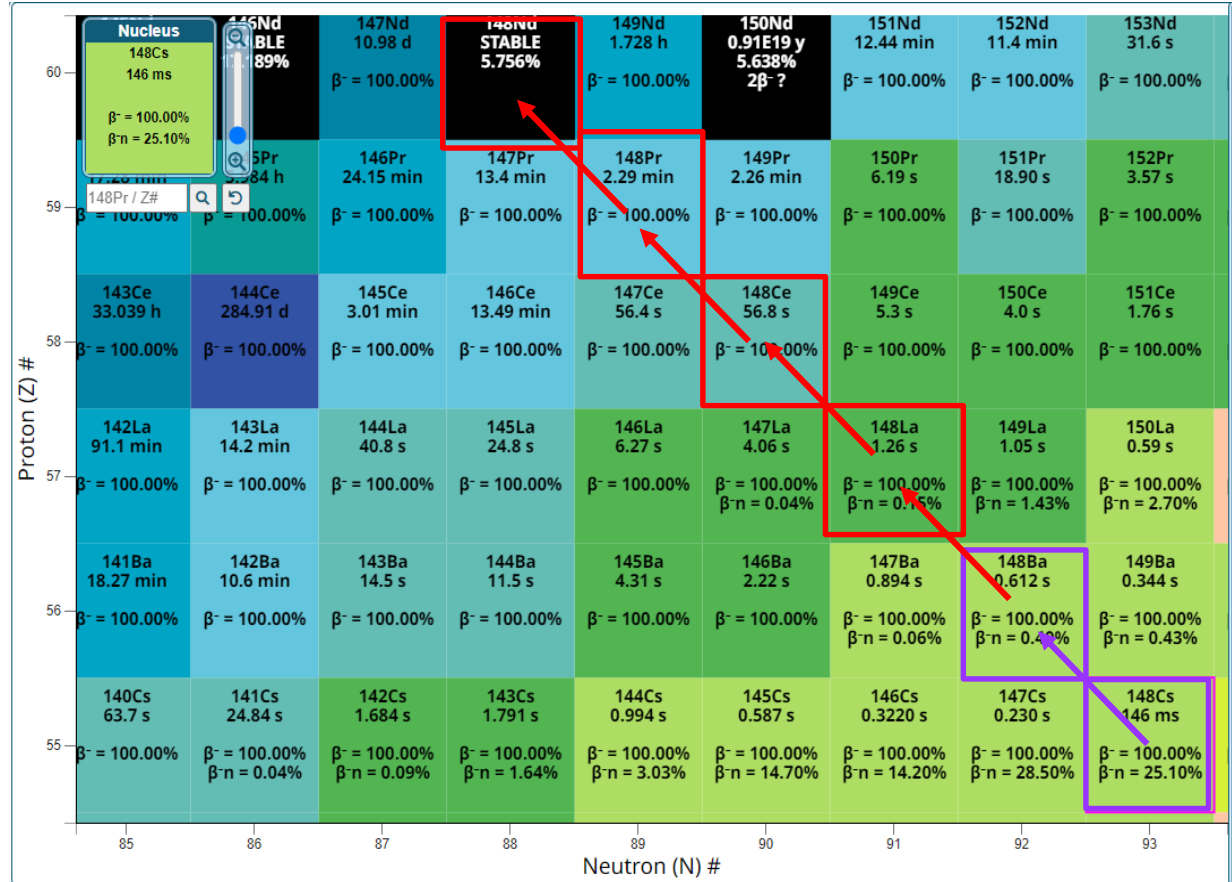


# Moving tape

Mylar tape

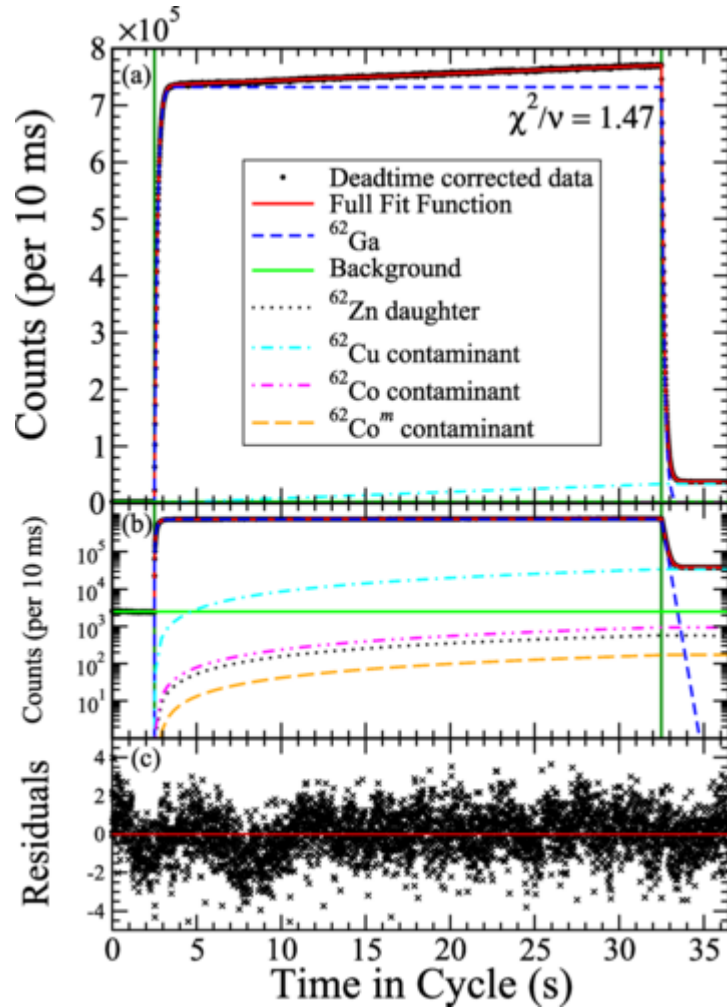


Lead wall

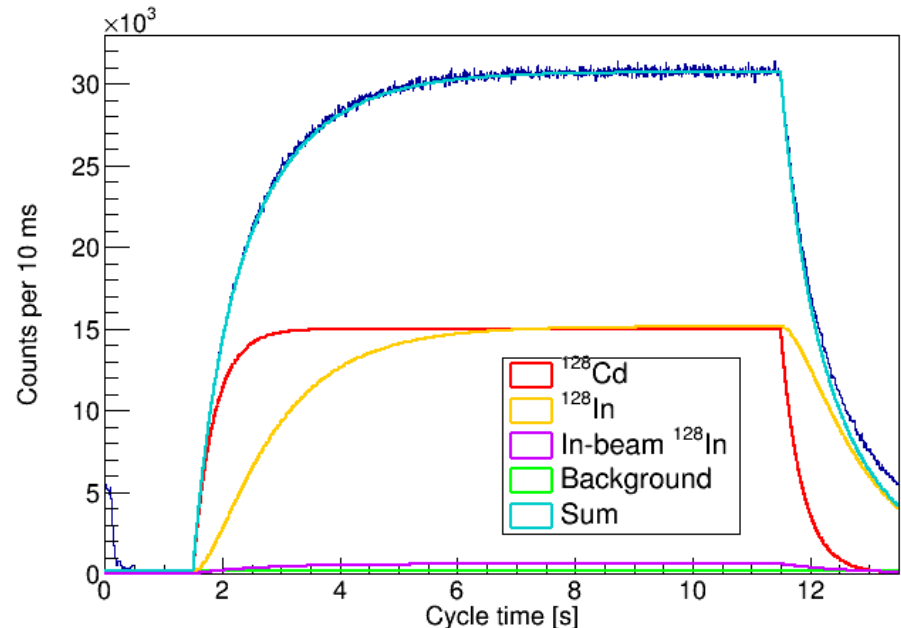




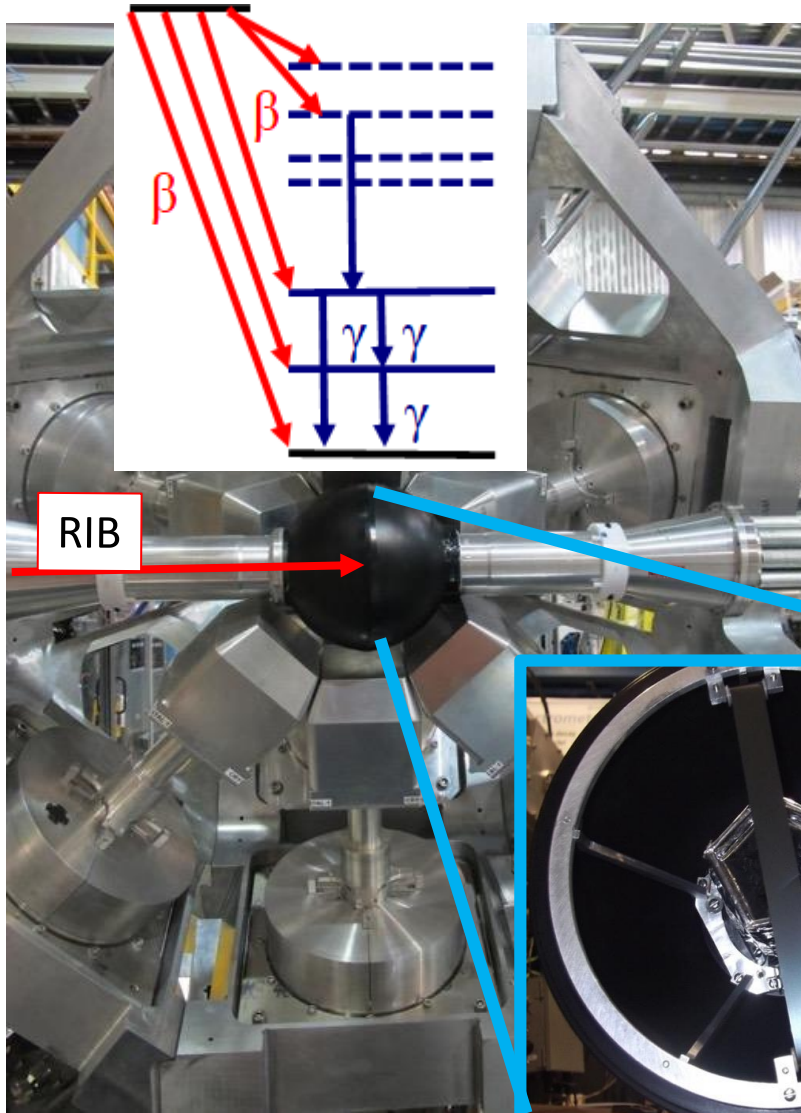
# Moving tape



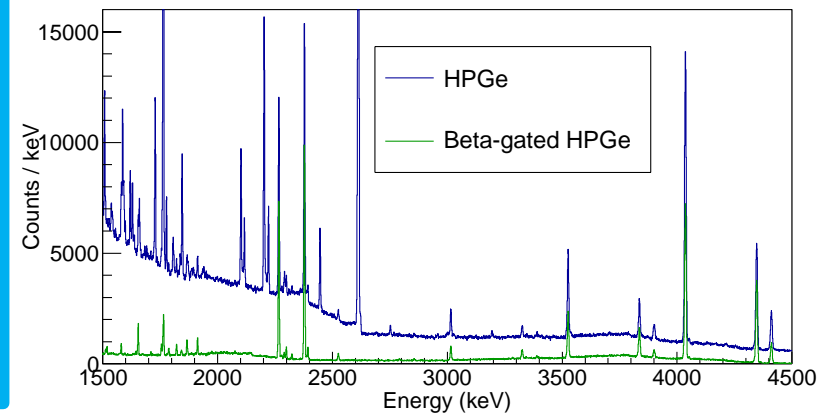
- You implant RIB for some time until your activity saturates
- Block RIB and measure decay
- Thanks to very different lifetimes, you observe different decay curves
- Move tape behind lead wall
- Start fresh implantation



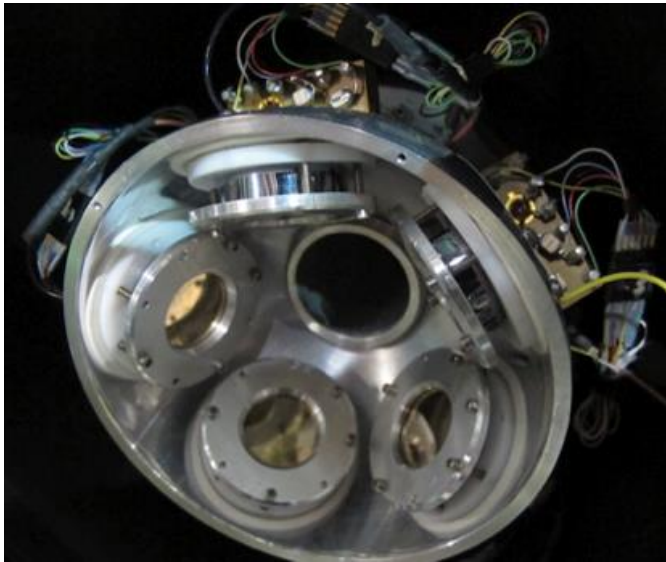
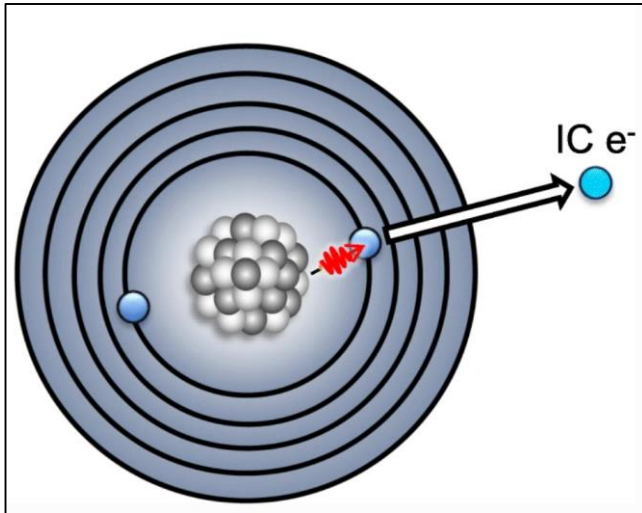
# Beta tagging



- Decay of interest begins with a  $\beta$  emission
- Surround point of implantation with plastic scintillators
- Require detection of a  $\beta$  particle
- Will not detect  $\beta$  particle from ambient radiation
- Suppresses room background by 5 orders of magnitude

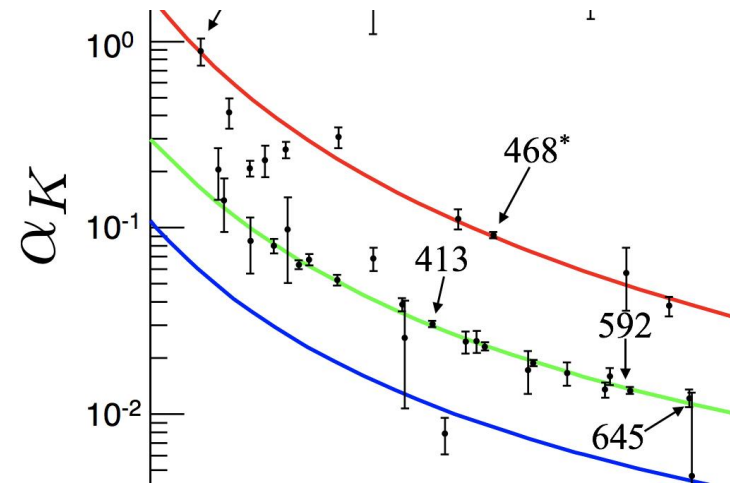


# Conversion electrons

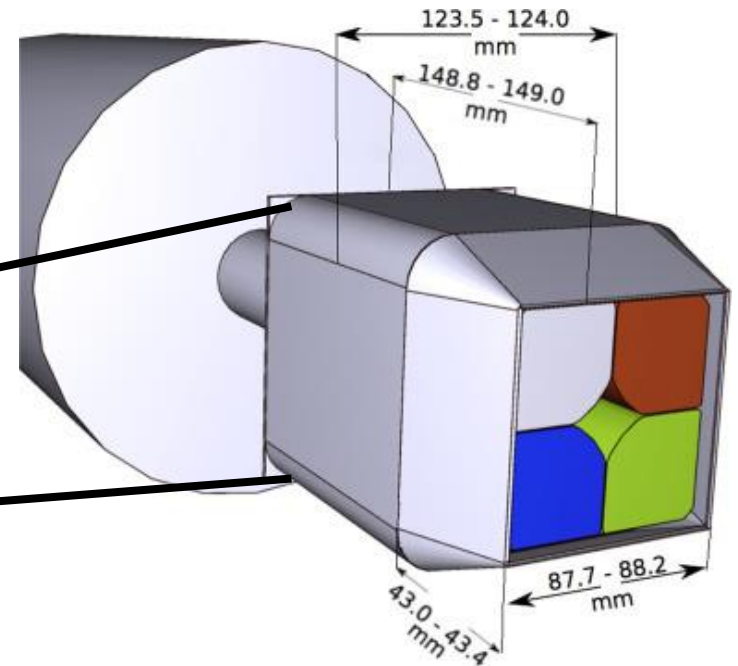
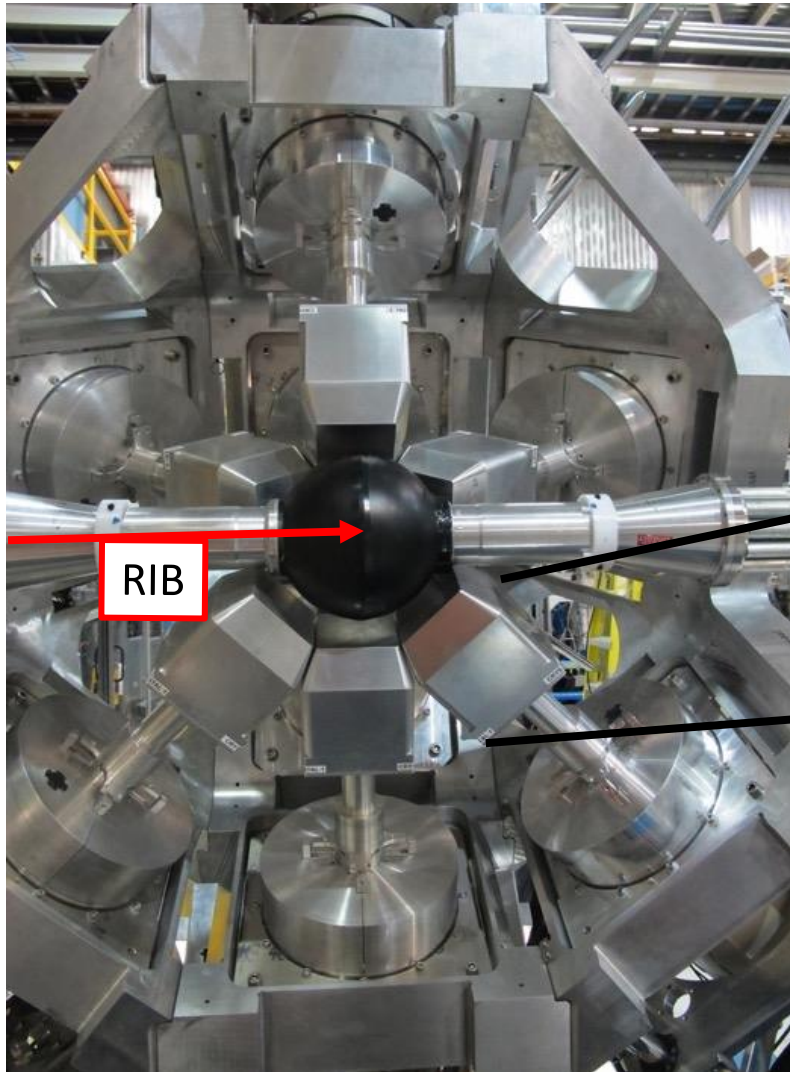


J. Park *et al.*, PRC 96, 014315 (2016).

- Whenever a low energy gamma is emitted, an inner shell atomic orbital can absorb it
- Probability depends on multipolarity of gamma and energy
- Electron emitted with discrete energy
- Si detectors, cooled to reduce electronic noise
- PACES in GRIFFIN, SPEDE in IDS



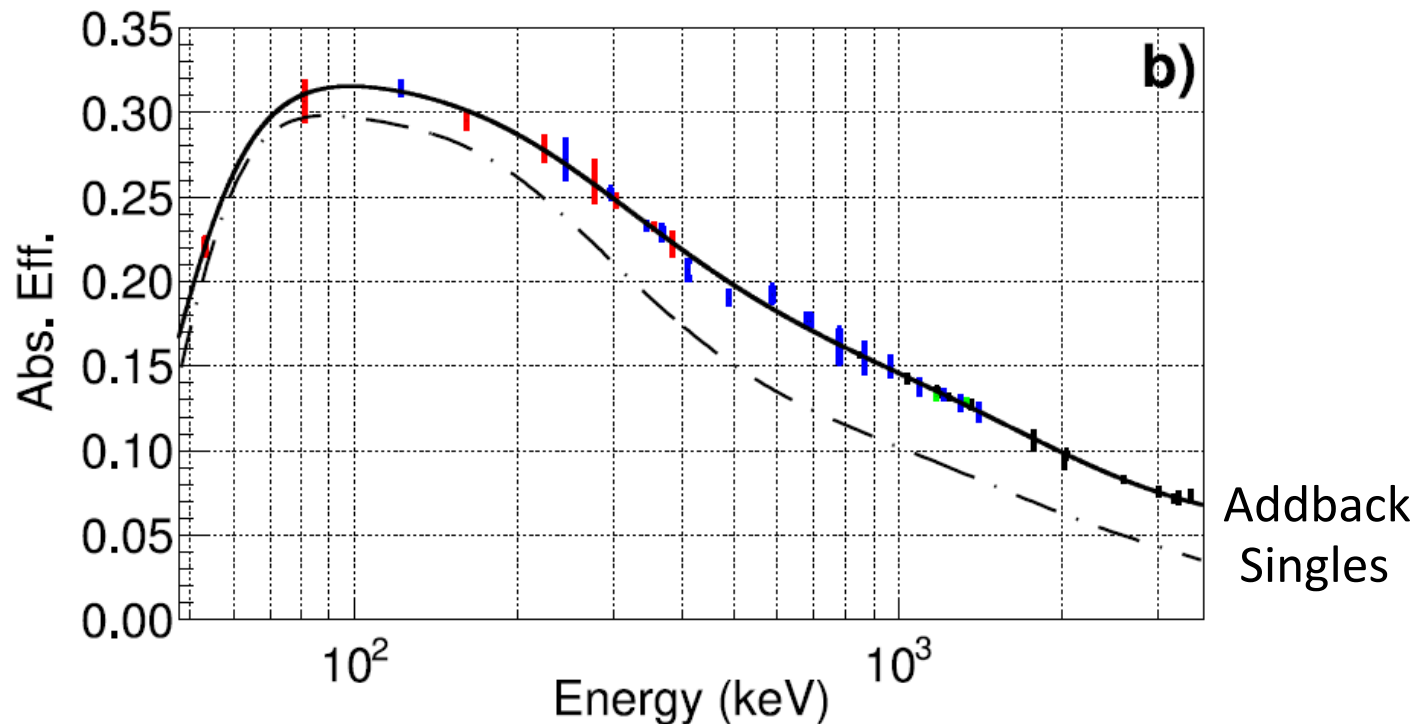
# HPGe clovers



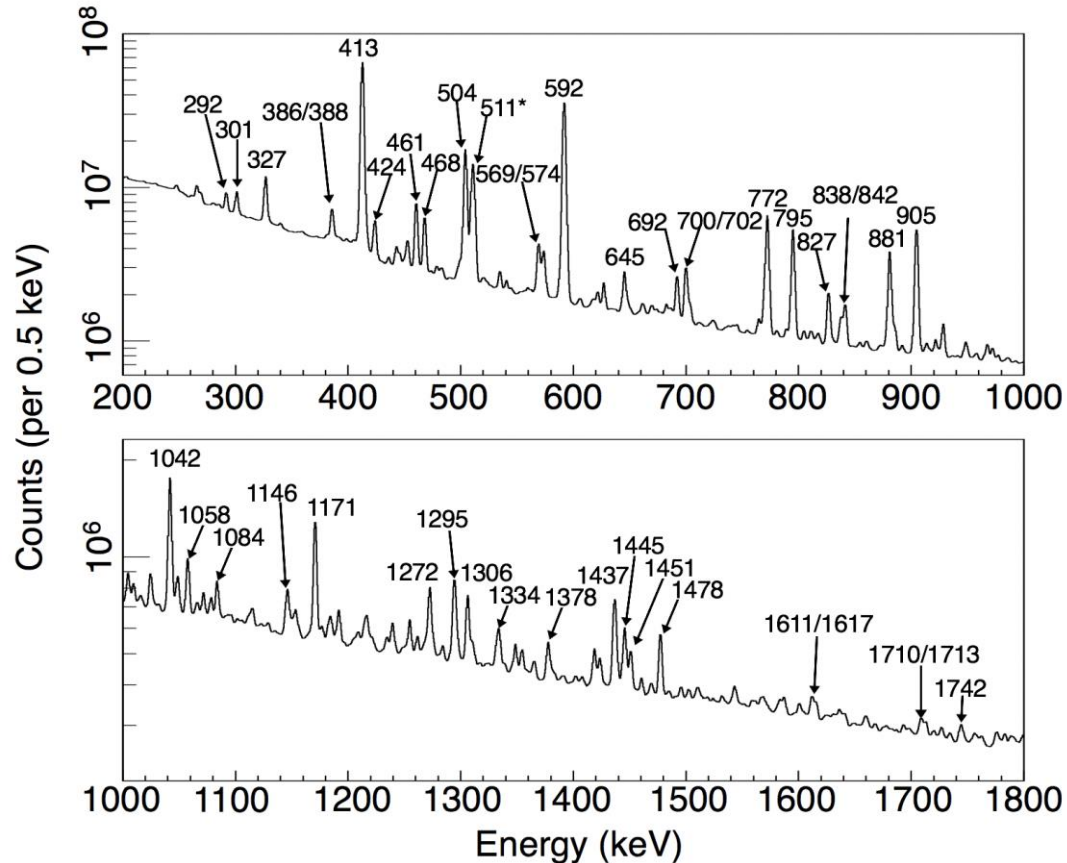
A.B. Garnsworthy, C.E. Svensson, M. Bowry et al. Nuclear Inst. and Methods in Physics Research, A 918 (2019) 9–29

# HPGe clovers

- 16 clovers, 4 crystal each, 64 total
- ~30% efficiency at ~120 keV
- Energy resolution@ 1.3MeV = 1.89(6) keV



# Statistics

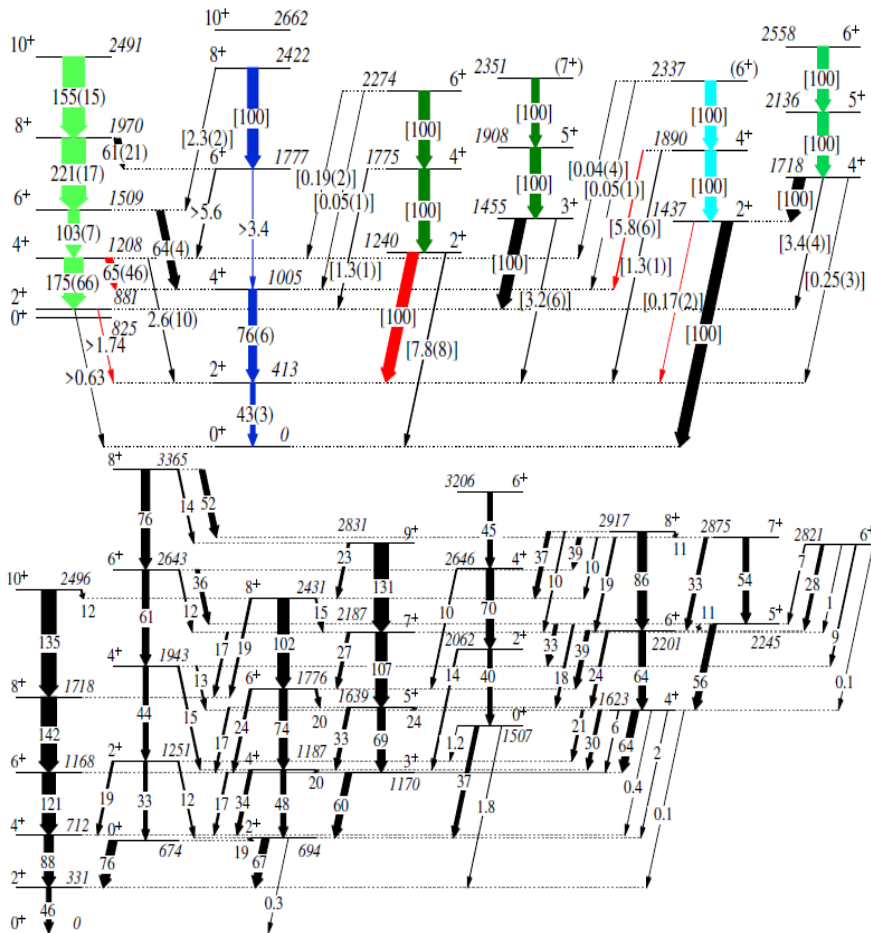


- High efficiency allows for high statistics
- Detects more and weaker  $\gamma$  rays
- Wealth of information, but more complex analysis

A. MacLean PhD thesis, UoGuelph 2021

# Build complex level scheme

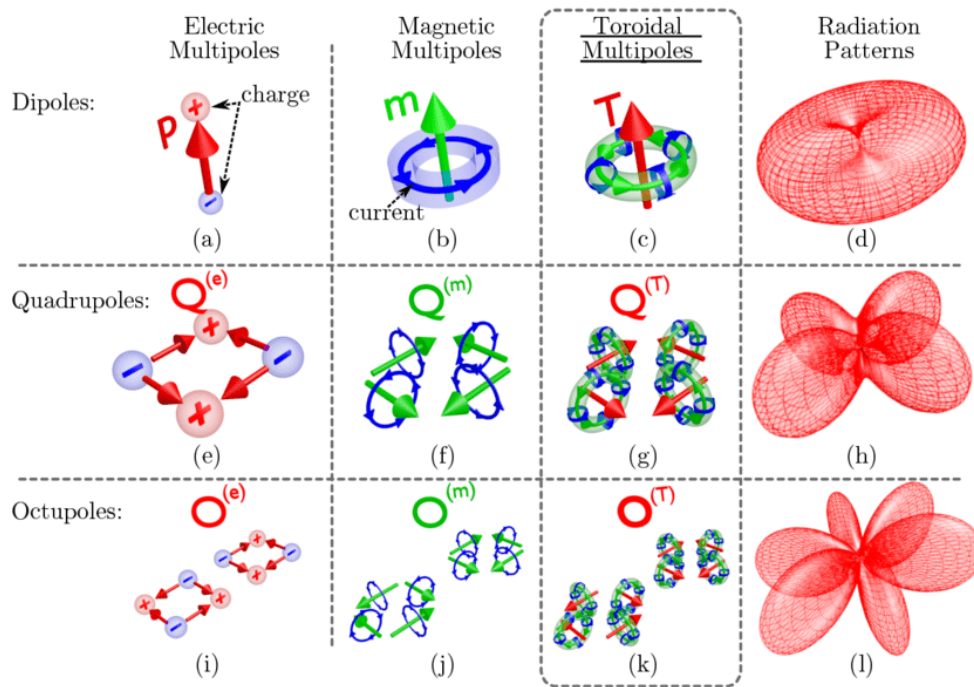
5



A. MacLean PhD thesis, UoGuelph 2021

- Use  $\gamma$ - $\gamma$  coincidences to build level scheme
- Coincidences require large statistics and efficiency
- This is a small fraction of a level scheme, can get a lot more complex
- Reveals inner structure of the nucleus

# $\gamma$ emission

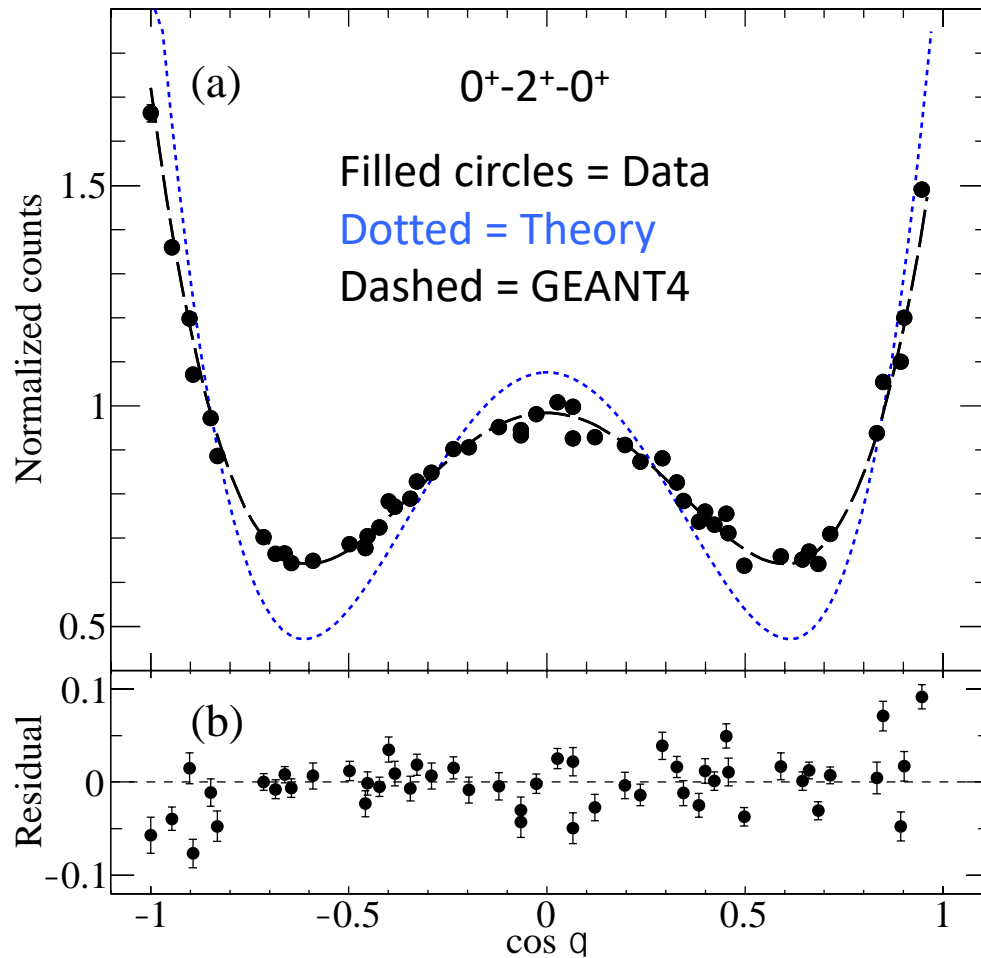


- $\gamma$ -ray emission is not isotropic
- Different for a dipole than a quadrupole
- Depends on  $J_i$ ,  $J_f$  and E/ML involved
- Powerful tool to firmly assign  $J$  to levels and E/ML to  $\gamma$ -rays

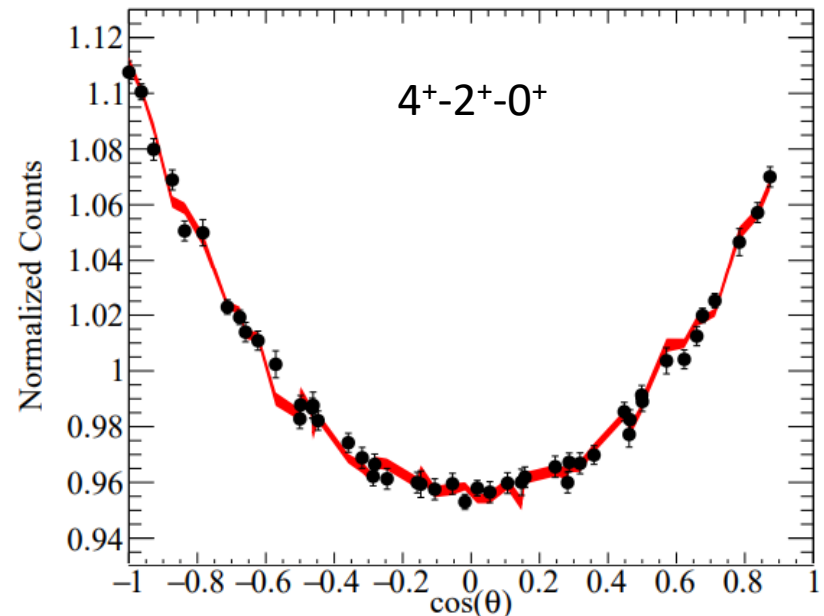
N. Papasimakis Nature Materials volume 15, pages263–271 (2016)



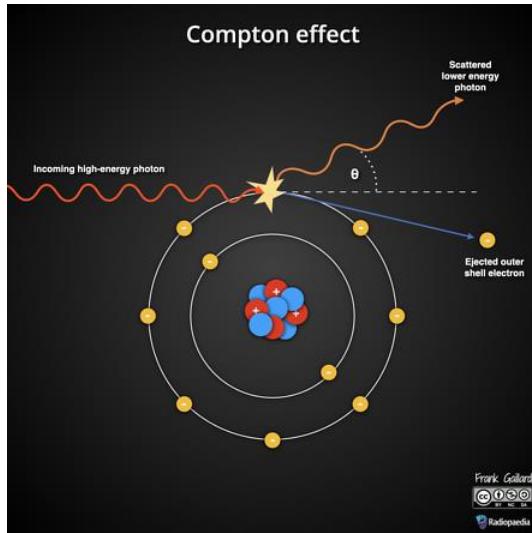
# $\gamma$ - $\gamma$ angular correlations



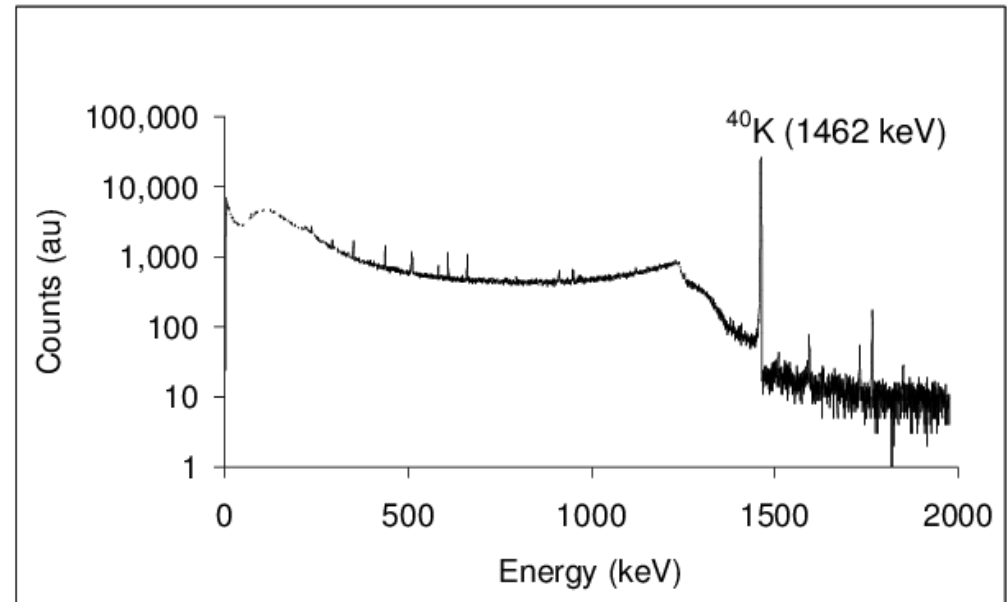
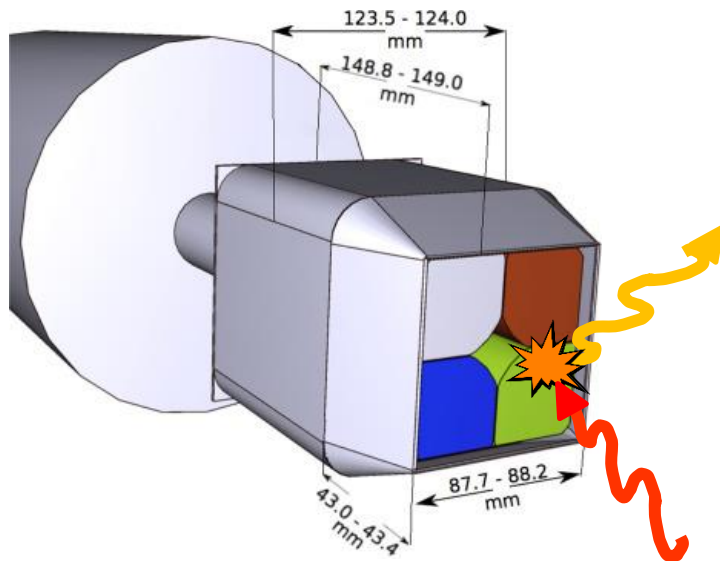
- With 64 crystals, there are 4096 pair combinations
- Very high granularity
- 52 unique angles
- Great angular resolution
- Precise angular correlations



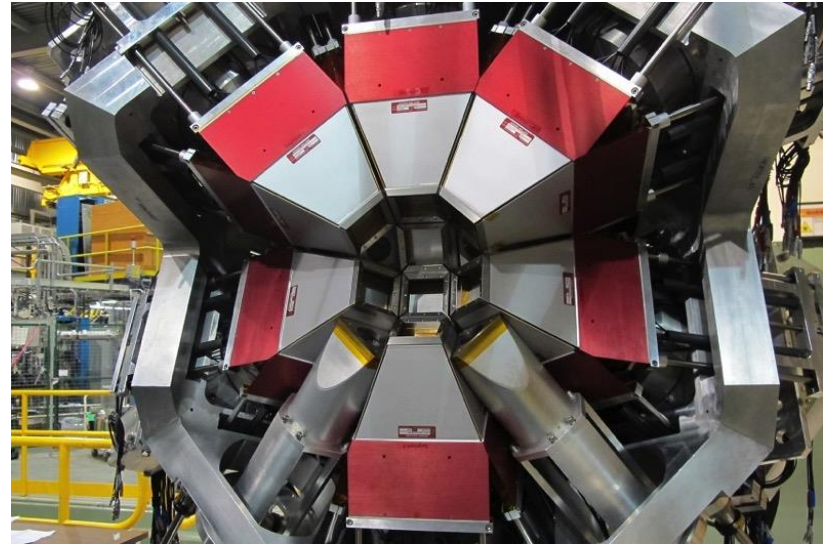
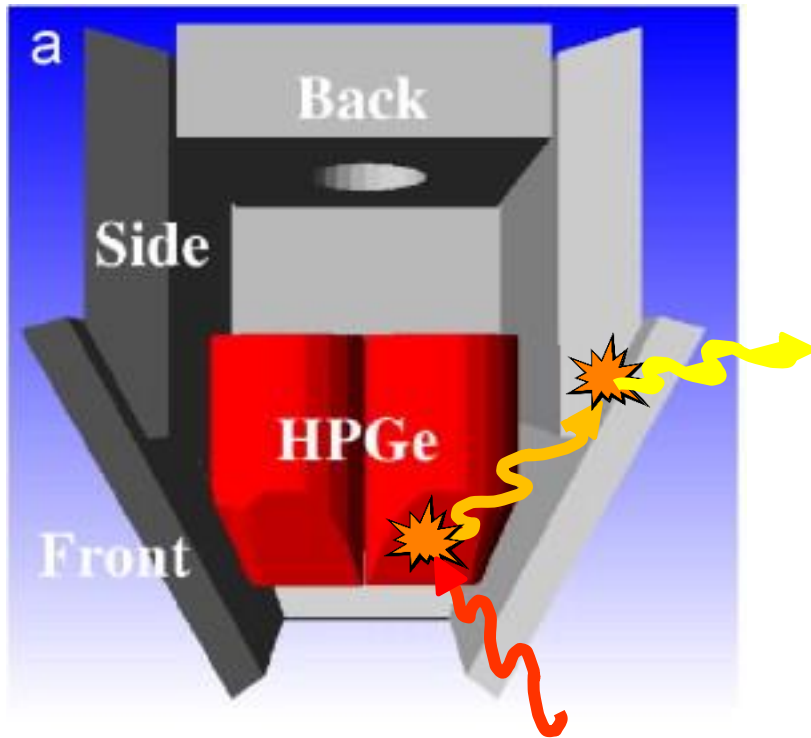
# Compton scattering



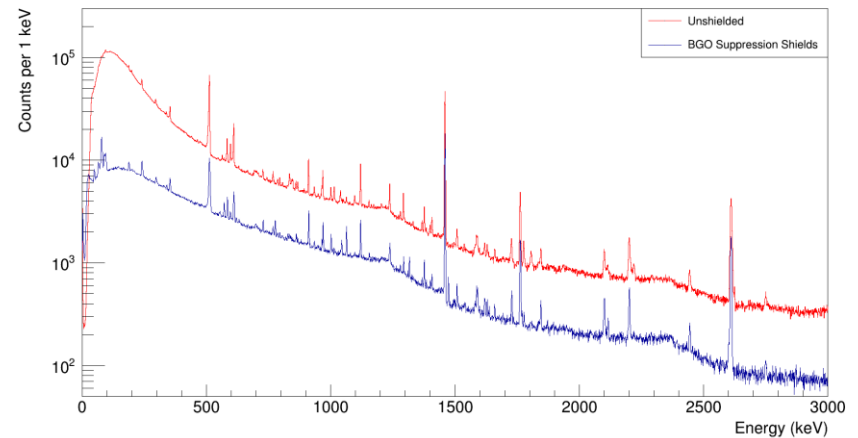
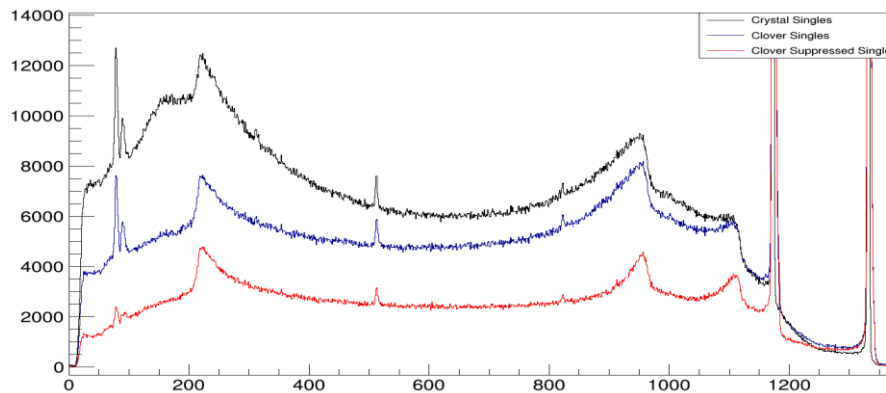
- A photon interacts with an atom depositing a fraction of its energy (not all)
- If after the Compton event, it escapes the detector, we will not detect the full energy
- Compton background difficult measurement of weaker gammas



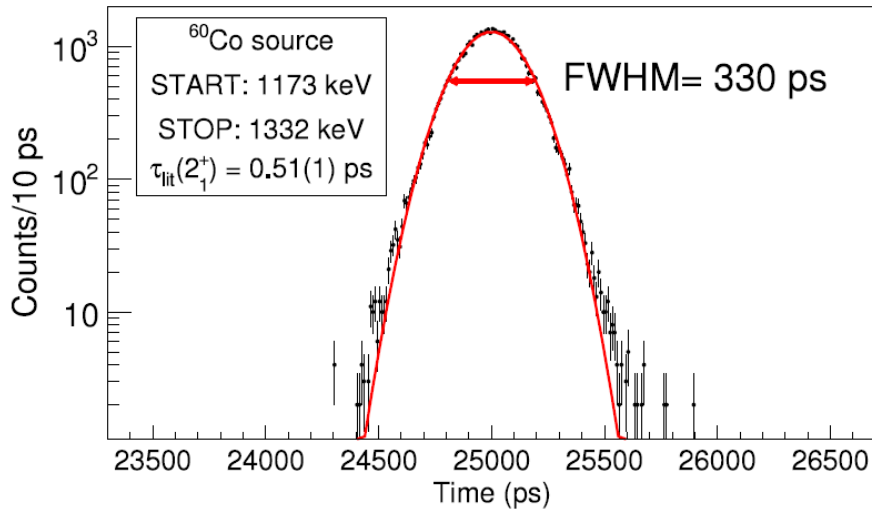
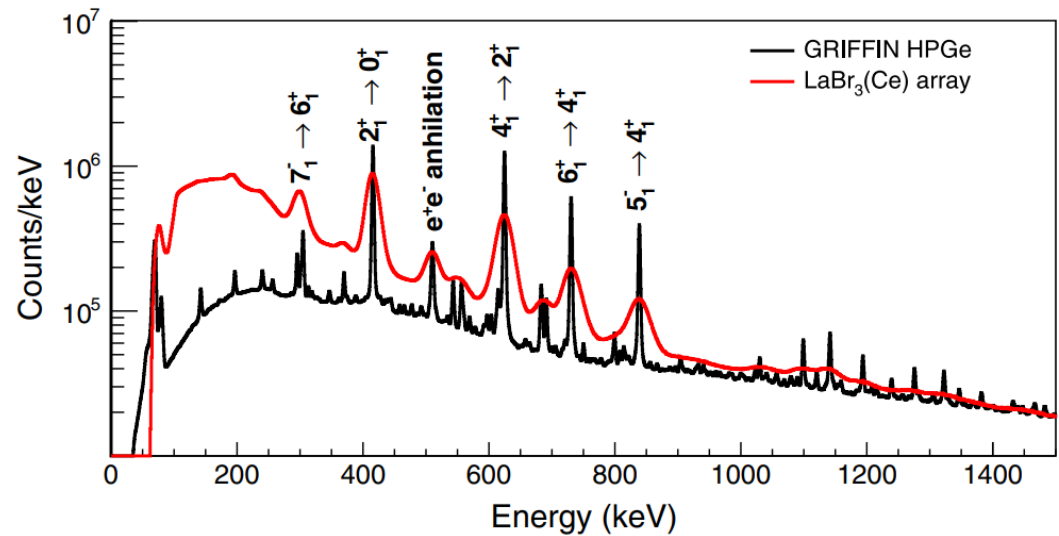
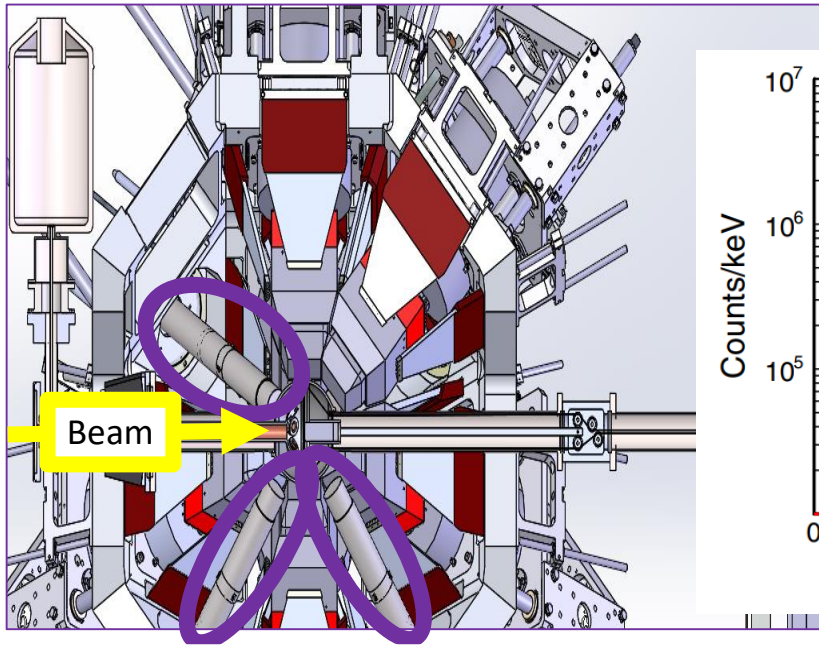
# BGO Compton suppressors



3 Hour Room Background

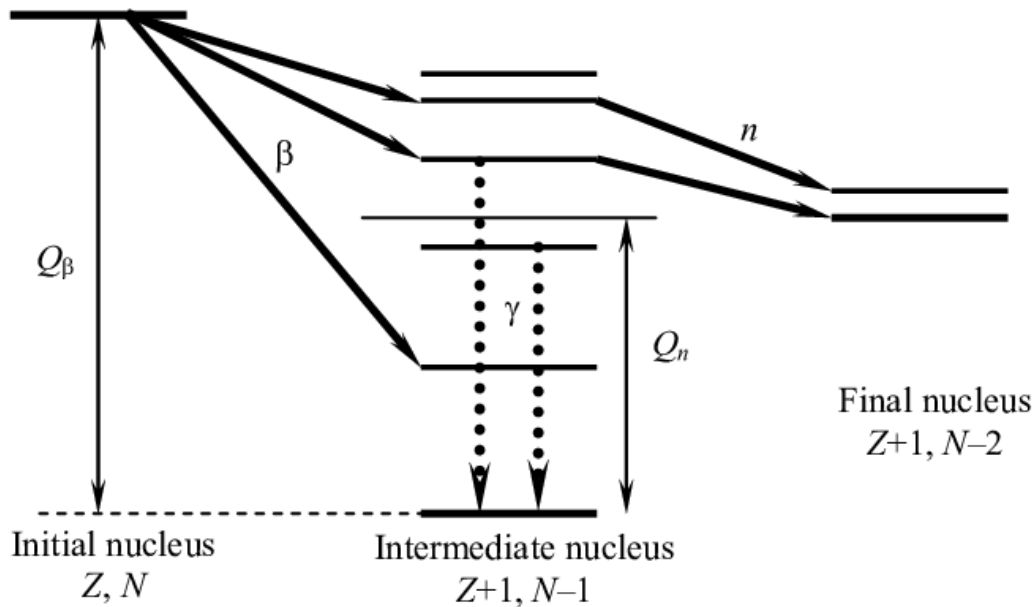


# Timing with scintillators



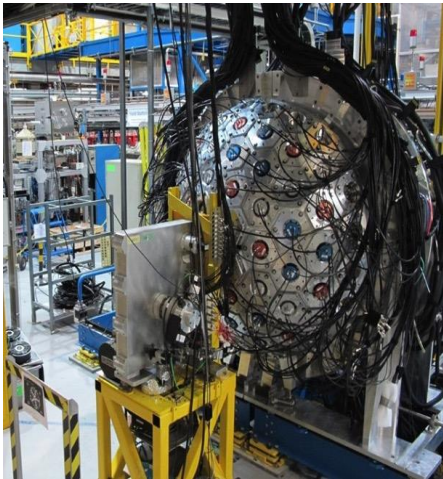
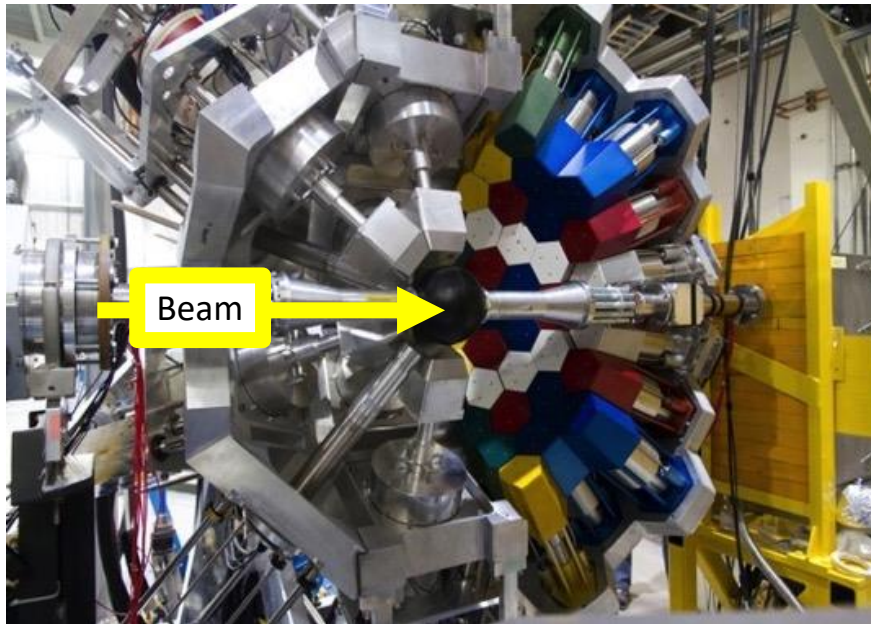
- We can add scintillators, such as  $\text{LaBr}_3(\text{Ce})$
- Worse energy resolution, superb timing properties
- Lifetime measurements down to ps

# Beta delayed neutron emission



- After a beta decay, the daughter nucleus can be left in a state above the neutron separation energy
- The intermediate nucleus can emit a neutron instead of a gamma.

# Neutron measurement

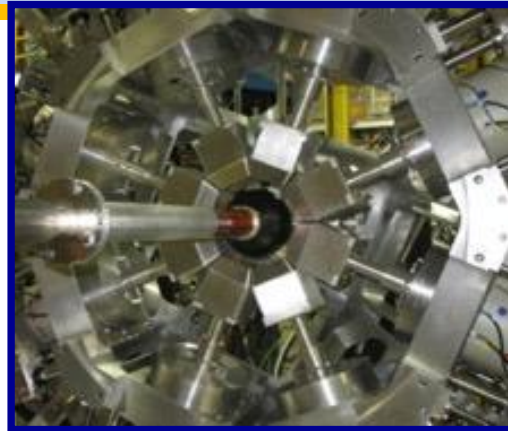
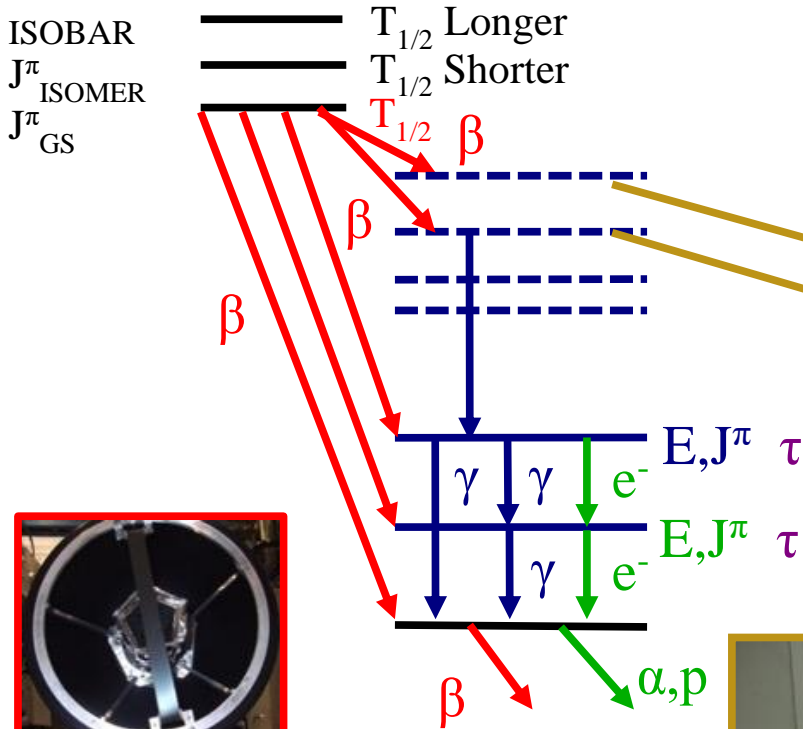


- We can use DESCANT to detect neutrons
- Liquid scintillators
- VANDLE (at ISOLDE) are plastic scintillators strips
- By neutron coincidences, we can distinguish if we are populating the N-1 nucleus or the N-2
- Without neutron detection, is not always trivial

# Beta decay experiment



Fast, in-vacuum tape system  
Enhances decay of interest



HPGe: 16 Clovers  
Detect gamma rays and  
determines branching  
ratios, multipolarities  
and mixing ratios



LaBr<sub>3</sub>:  
8xLaBr<sub>3</sub>  
Fast-timing to  
measure level  
lifetimes



SCEPTAR: 10+10 plastic  
scintillators  
Detects beta decays and  
determines branching ratios



PACES: 5 Cooled Si(Li)s  
Detects Internal Conversion  
Electrons and alphas/protons



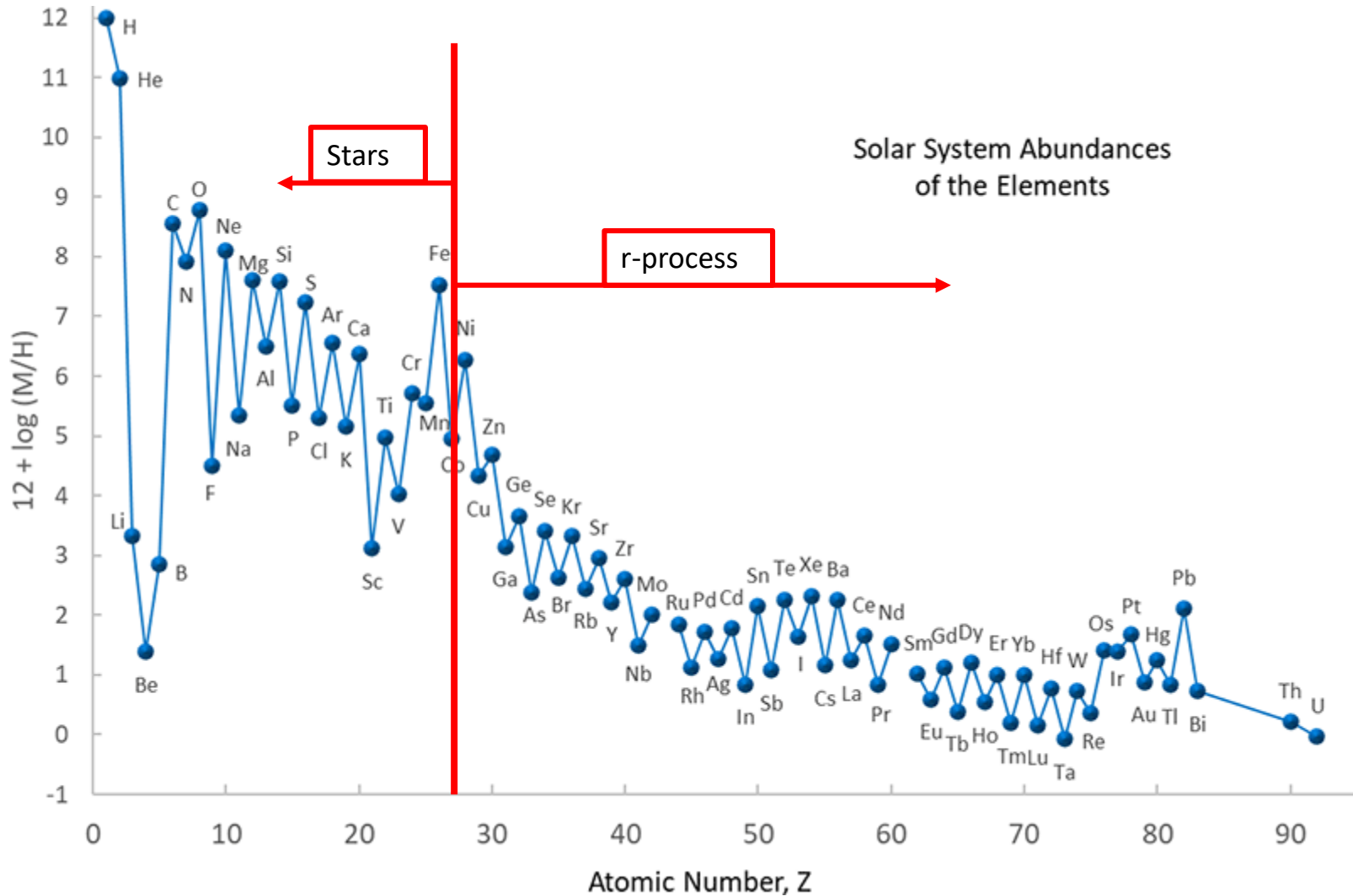
# Beta decay experiment

## Nuclear structure

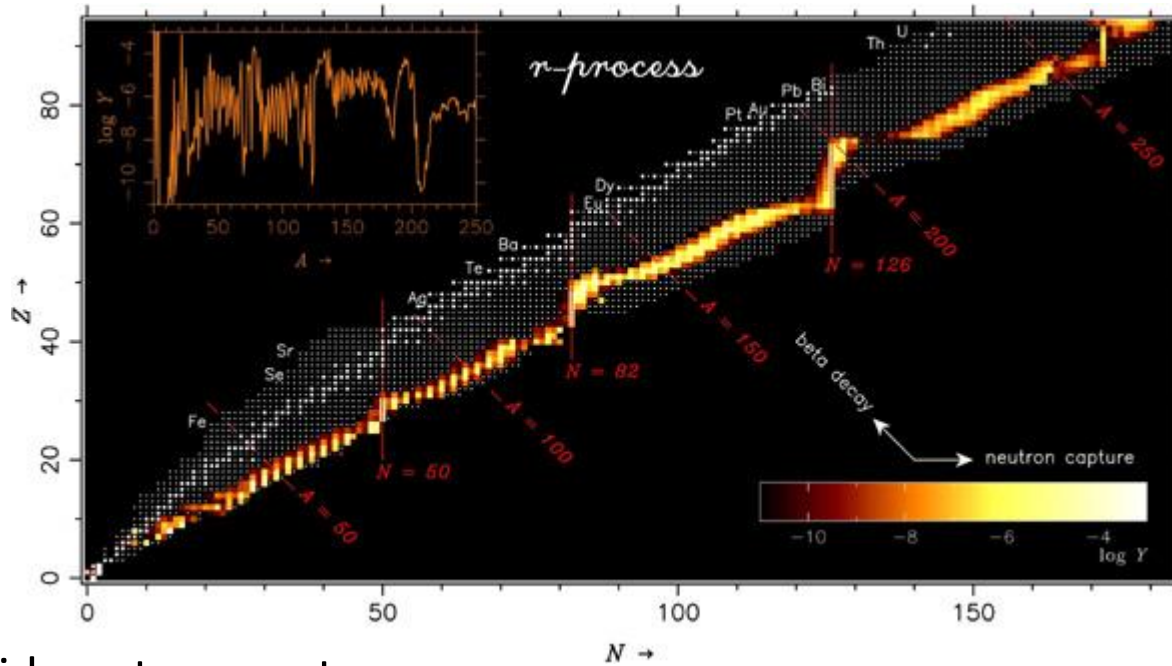
- Measure beta decay half-life
- Construct level scheme with  $\gamma\gamma$  coincidences
- Assign  $J^\pi$  to excited states using:
  - $\text{Log}(ft)$  from beta decay
  - $\gamma\gamma$  angular correlations
  - $\alpha$  from conversion electrons
  - $B(XL)$  from lifetime measurements
- Firmly assigning  $J^\pi$  is key to understand the structure of a nucleus
- **NOTE:** none of this techniques is exclusive from beta decay experiments, but, in general, it would be the simplest and cleanest experiment to use them



# Beta decay experiment



# r-process

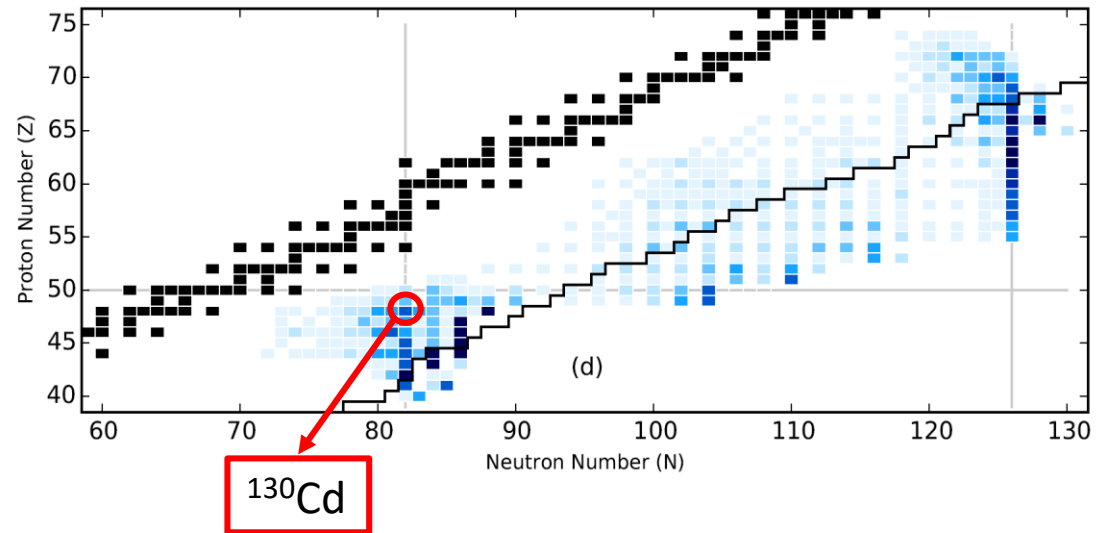
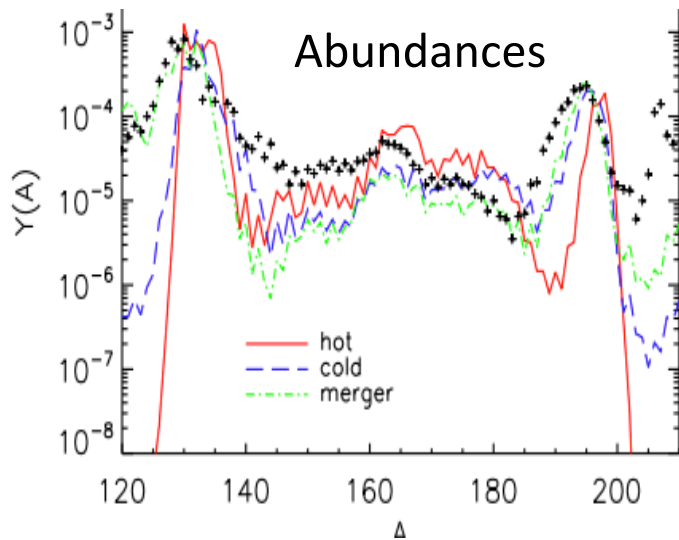


- Rapid neutron-capture process
- Supernovae explosions
- Competition between beta decay and neutron absorption
- Most relevant observables:
  - Beta decay half life (g.s. and isomers)
  - Neutron emission
  - Neutron capture cross section

<http://www.ph.sophia.ac.jp/~shinya/research/research.html>

# Half-Lives of Neutron-Rich $^{128-130}\text{Cd}$

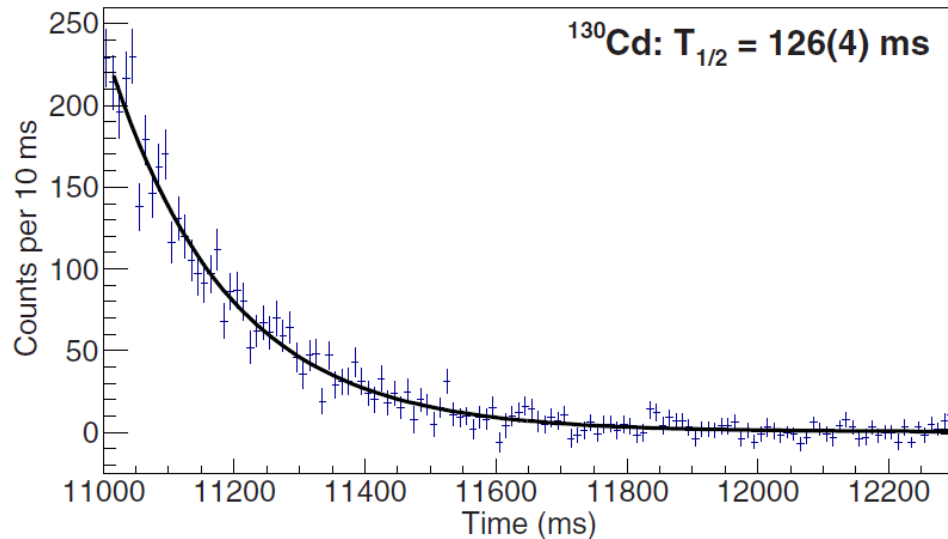
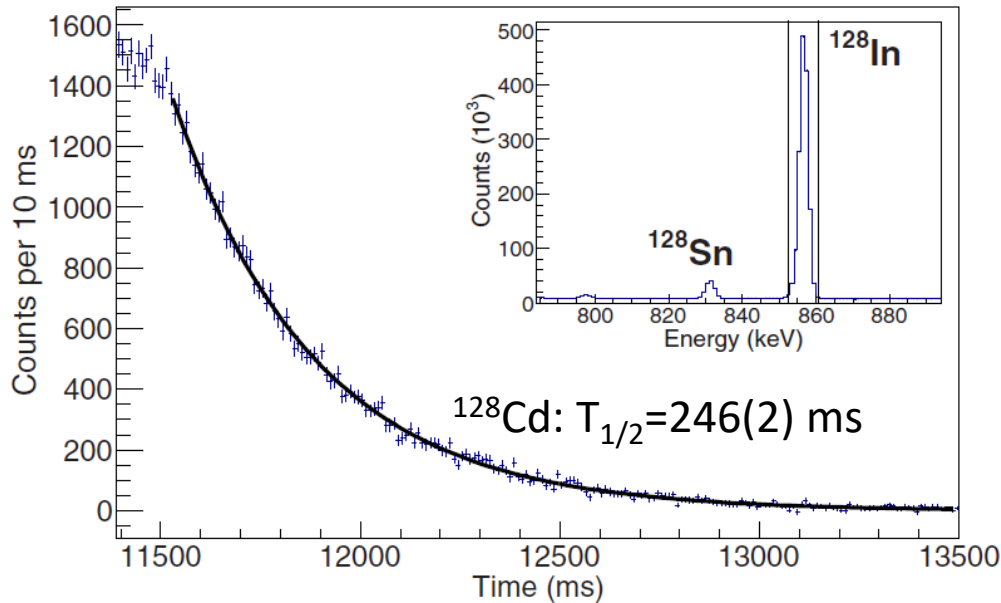
Nuclei near  $N = 82$  are responsible for the  $A \sim 130$  r-process abundance peak. These ‘waiting point’ nuclei are important in calculations of all astrophysical environments.



*G. Lorusso et al. PRL 114 192501 (2015)*

*M. Mumpower et al., Prog. Part. Nucl. Phys. 86, 86 (2016)*

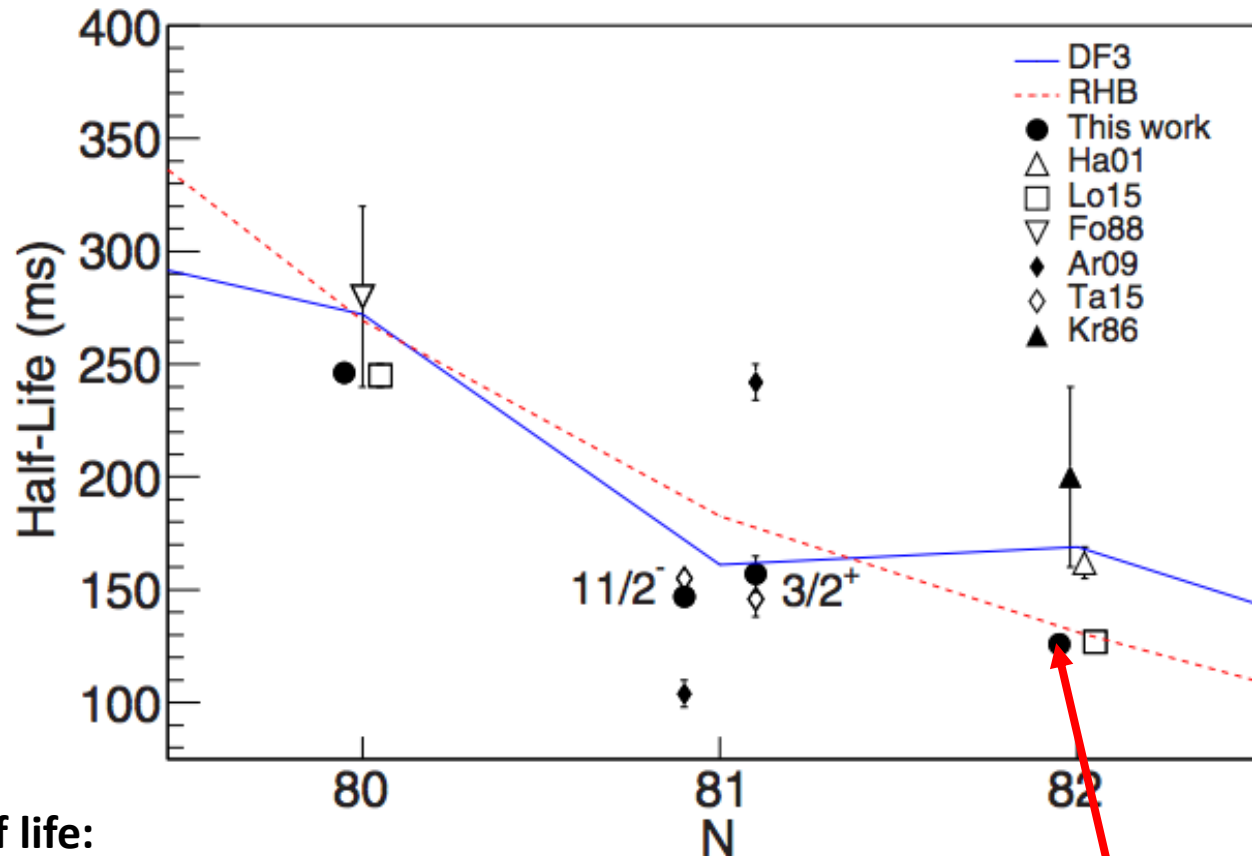
# Half-Lives of Neutron-Rich $^{128-130}\text{Cd}$



- Lifetimes were measured in  $\beta$ - $\gamma$  coincidences
- HPGe gate cleans the spectra
- Lifetime is extracted from SCEPTAR ( $\beta$ ) activity

R. Dunlop et al., PRC 93, 062801(R) (2016).

# Beta decay experiment



## <sup>130</sup>Cd half life:

195 (35) ms K.-L. Kratz *et al.*, Z. Phys.A325, 489 (1986).

162 (7) ms M. Hannawald *et al.*, NPA 688, 578 (2001).

127 (2) ms G. Lorusso *et al.* PRL 114 192501 (2015).

126 (4) ms R. Dunlop *et al.*, PRC 93, 062801(R) (2016).

This work

# SM half-lives prediction below N=82

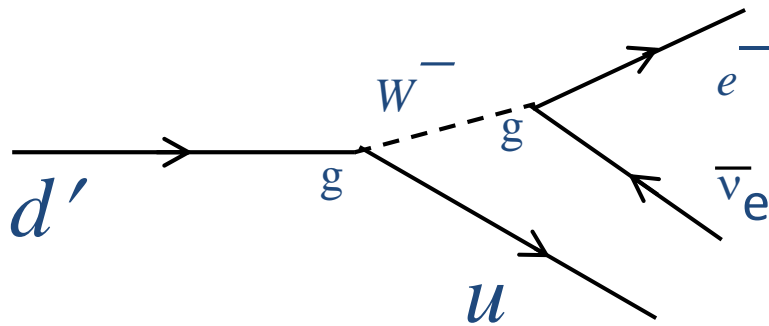
| Nucleus           | Measured<br>Half-life<br>(ms) [2] | SM<br>( $q = 0.66$ )<br>(ms) [5] | Scaled SM<br>( $q = 0.75$ )<br>(ms) |
|-------------------|-----------------------------------|----------------------------------|-------------------------------------|
| $^{131}\text{In}$ | 261(3)                            | 247.53                           | 194.1                               |
| $^{130}\text{Cd}$ | 127(2)                            | 164.29                           | 127                                 |
| $^{129}\text{Ag}$ | 52(4)                             | 69.81                            | 54.03                               |
| $^{128}\text{Pd}$ | 35(3)                             | 47.25                            | 36.57                               |
| $^{127}\text{Rh}$ | $20^{+20}_{-7}$                   | 27.98                            | 21.67                               |

- Previous SM prediction were  $\sim 25\%$  off
- New quenching precisely reproduces lifetimes
- Affects the N=82 waiting point and the  $A \sim 130$  peak abundance
- **Creates significant discrepancy for  $^{131}\text{In}$**
- R. Dunlop *et al.* Phys. Rev. C **93**, 062801(R) (2016)

# The Standard Model of particle physics

The CKM matrix plays a central role in the Standard Model and underpins all quark flavour-changing interactions:

→ weak interaction eigenstates  $\neq$  quark mass eigenstates



$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

$$|d'\rangle = V_{ud}|d\rangle + V_{us}|s\rangle + V_{ub}|b\rangle$$

In the Standard Model the CKM matrix describes a unitary transformation:

$$V_{ud}^2 + V_{us}^2 + V_{ub}^2 = 1$$

The first row of the CKM matrix provides the most demanding experimental test of the unitarity condition.

# Superaligned Fermi $\beta$ Decay

phase space (Q-value)  $\rightarrow$   $ft = \frac{K}{|M'_{if}|^2 g^2}$   $\leftarrow$  constants

half-life, branching ratio  $\rightarrow$   $|M'_{if}|^2 g^2$   $\leftarrow$  Weak coupling strength

$|M'_{if}|^2$   $\leftarrow$  matrix element

For the special case of  $0^+ \rightarrow 0^+$  (pure Fermi)  $\beta$  decays between isobaric analogue states (superaligned) the matrix element is that of an isospin ladder operator:

$$|M_{fi}|^2 = (T - T_z)(T + T_z + 1) = 2 \quad (\text{for } T=1)$$

Strategy: Measure superallowed  $ft$ -values, deduce  $G_V$  and  $V_{ud}$ :

Vector coupling constant  $\rightarrow$   $ft = \frac{K}{2 G_V^2}$

$|V_{ud}| = G_V / G_F$   $\leftarrow$  Fermi coupling constant



# Superaligned Fermi $\beta$ Decay

$$Ft = ft(1 + \delta'_R)(1 + \delta_{NS} - \delta_C) = \frac{K}{2G_V^2(1 + \Delta_R^V)} = \text{constant}$$

“Corrected” ft value      Experiment      Calculated corrections ( $\sim 1\%$ ) (nucleus dependent)      Inner radiative correction ( $\sim 2.4\%$ ) (nucleus independent)      CVC Hypothesis

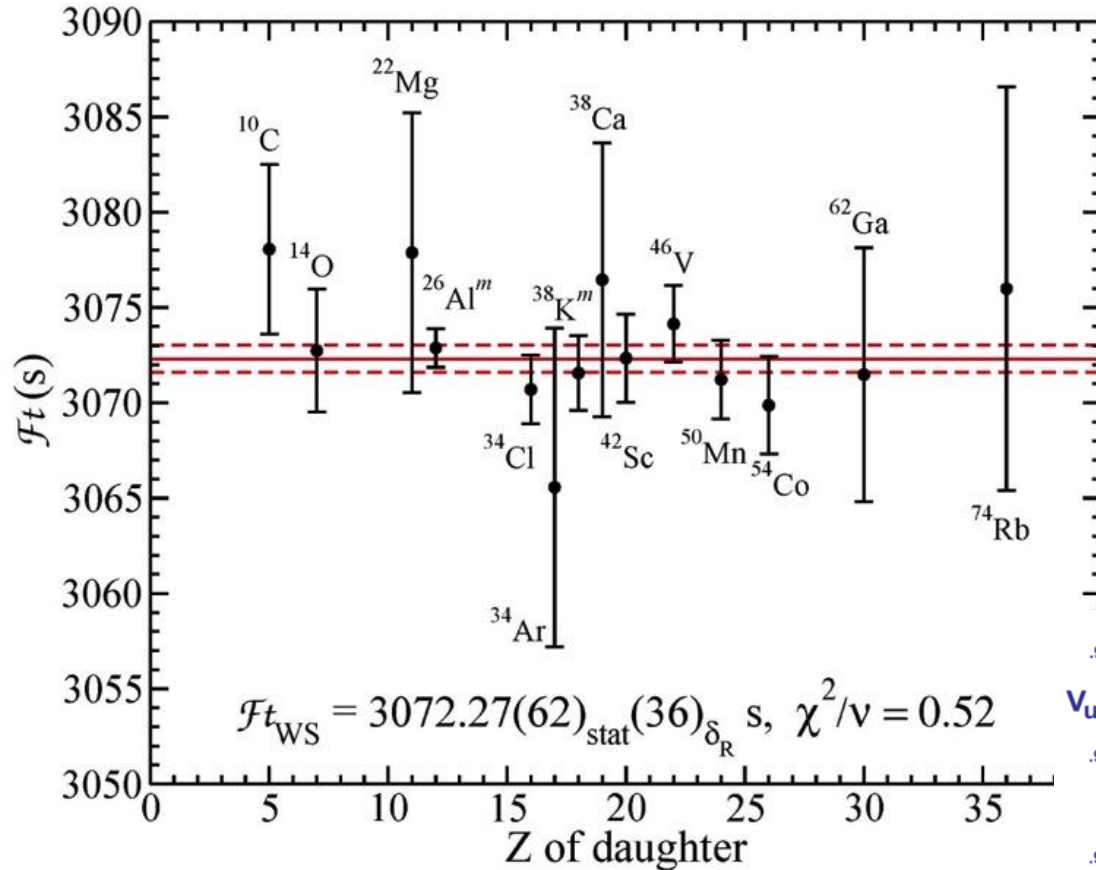
$\Delta_R^V$  = nucleus independent inner radiative correction: 2.361(38)%

$\delta_R$  = nucleus dependent radiative correction to order  $Z^2\alpha^3$ :  $\sim 1.4\%$   
 - depends on electron's energy and  $Z$  of nucleus

$\delta_{NS}$  = nuclear structure dependent radiative correction:  $-0.35\% - 0.05\%$

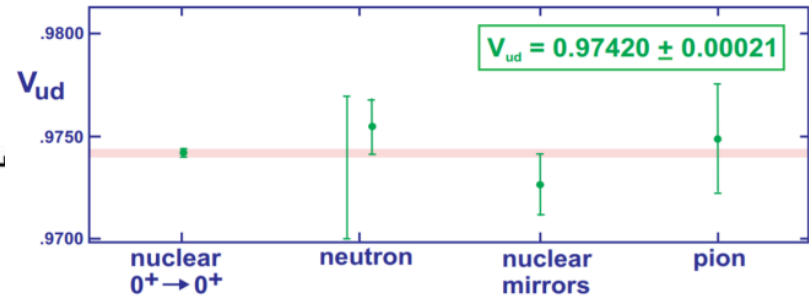
$\delta_C$  = nucleus dependent isospin-symmetry-breaking correction:  $0.2\% - 1.6\%$   
 - strong nuclear structure dependence

# Superaligned Fermi $\beta$ Decay

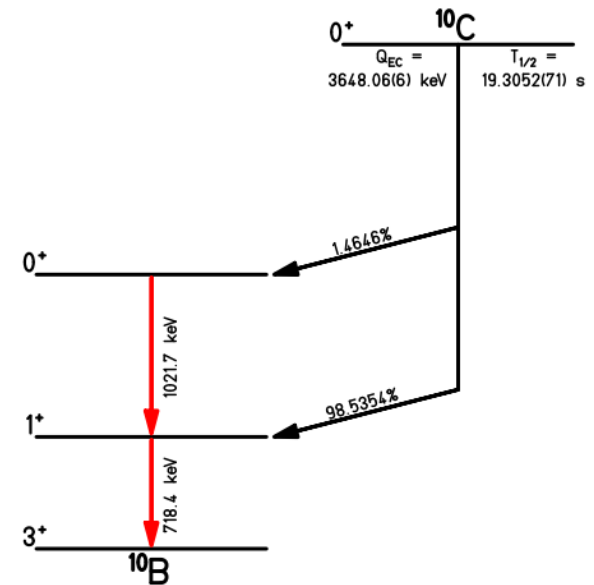
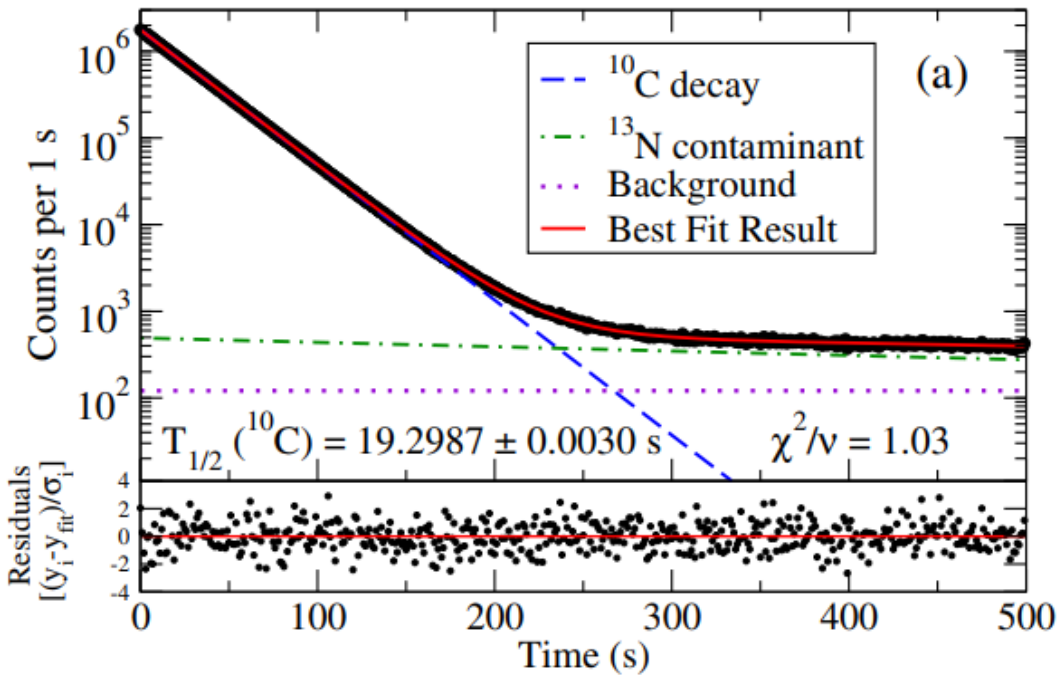


$|V_{\text{ud}}| = 0.97417 \pm 0.00021$

*Hardy and Towner,  
 Phys. Rev. C **91**, 025501 (2015)*



# Beta decay experiment



- The job of nuclear physics is to precisely measure  $Q_\beta$ ,  $T_{1/2}$  and branching ratio
- Extract  $Ft$  better than 1 part in  $10^4$
- If  $|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 \neq 1$  would imply physics beyond the standard model

# Beta decay experiments

- One of the cleanest and most precise probes we have in nuclear physics
- Versatile experiments able to measure multiple observables simultaneously
- Relevant for astrophysics
- Powerful tool to search for physics beyond the Standard Model

# Useful References

## Books

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- ✓ “Radiation detection and measurements”, G.F. Knoll, 1989
- ✓ “Alpha-, Beta- and Gamma-ray Spectroscopy”, Ed. K. Siegbahn, 1965
- ✓ “Introductory Nuclear Physics”, K. S. Krane, 1988
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- ✓ “Subatomic Physics”, Ernest M. Henley and Alejandro García, 3rd Edition, 2007, World Scientific Publishing,

## Journal Articles

- ✓ [Euroscool on Exotic Beams, Lectures Notes](#): “Decay Studies of N~Z Nuclei”, E. Roeckl, Vol I, “Beta Decay of exotic Nuclei”, B. Rubio & W. Gelletly, and B. Blank”1-2p radioactivity” in Vol III
- ✓ B. Blank and M.J.G. Borge, Prog Part and Nuc. Phys 60 (2008) 403
- ✓ M. Pfützner, L.V. Grigorenco, M. Karny & K. Riisager, Rev. Mod. Phys 84 (2012)567
- ✓ G. Benzoni, Eur. Phys. J plus 131 (2016) 99
- ✓ A. Algora et al., Eur. Phys A 57 (2021) 85
- ✓ École Joliot-Curie de Physique Nucleaire, 2002

Bibliography courtesy of MJ Borge

Any questions?

# Beta decay experiment