

The Analytic Wavefunction

(with Mang Hei Gordon Lee, Scott Melville, Enrico Pajer)

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Introduction

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Can we use the amplitudes technology to study cosmological correlators?

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$$\langle \hat{\phi}_{\mathbf{k}_1} \hat{\phi}_{\mathbf{k}_2} \rangle = \frac{\int D\phi |\Psi[\phi]|^2 \phi_{\mathbf{k}_1} \phi_{\mathbf{k}_2}}{\int D\phi |\Psi[\phi]|^2} \quad (5)$$

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How do we compute $\Psi[\phi]$?

Path integrals in Cosmology

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$$\text{Log}(\Psi[\phi]) = \frac{1}{2} \int_{\mathbf{k}_1 \mathbf{k}_2} \psi^{(2)}(\mathbf{k}) \phi_{\mathbf{k}_1} \phi_{\mathbf{k}_2} + \frac{1}{6} \int_{\mathbf{k}_1 \mathbf{k}_2 \mathbf{k}_3} \psi^{(3)}(\mathbf{k}) \phi_{\mathbf{k}_1} \phi_{\mathbf{k}_2} \phi_{\mathbf{k}_3} + \dots \quad (7)$$

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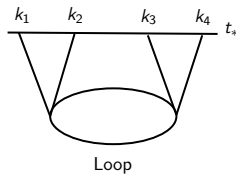
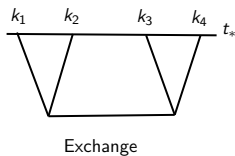
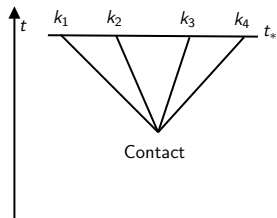
$\psi^{(n)}(\mathbf{k})$ are the wavefunction coefficients. They can be computed in perturbation theory using Feynman diagrams. Its analytic structure is deeply related to that of cosmological correlators.

Wavefunction coefficients

Feynman diagrams for $\psi^{(n)}(\mathbf{k})$ involve a boundary at $t = t_*$

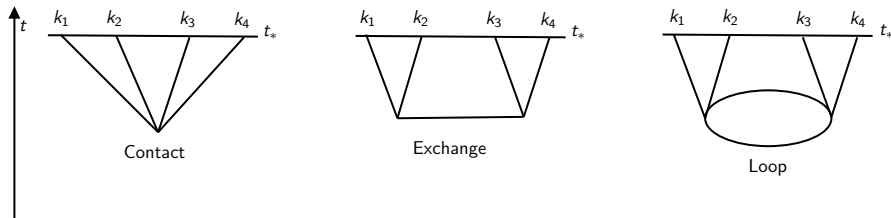
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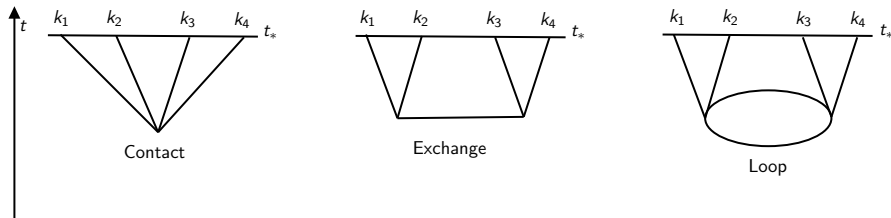


External lines are bulk-to-boundary propagators. They obey the free theory equations of motion and Dirichlet boundary conditions:

$$(\partial_t^2 + \Omega_{\mathbf{k}}^2)K_{\mathbf{k}}(t) = 0, \quad K_{\mathbf{k}}(t_*) = 1, \quad \Omega_{\mathbf{k}}^2 = \mathbf{k}^2 + m^2 \Rightarrow K_{\mathbf{k}}(t) = e^{i\Omega_{\mathbf{k}}(t-t_*)} \quad (8)$$

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Internal lines are bulk-to-bulk propagators.

$$(\partial_{t_1}^2 + \Omega_{\mathbf{k}}^2)G_{\mathbf{k}}(t_1, t_2) = -\delta(t_1 - t_2), \quad G_{\mathbf{k}}(t_*, t_2) = G_{\mathbf{k}}(t_1, t_*) = 0 \quad (9)$$

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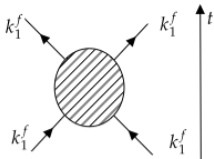
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α_n are the Wilson coefficients of the EFT and are fixed by UV physics. The EFT interactions are constrained by the IR symmetries and the boundary conditions.

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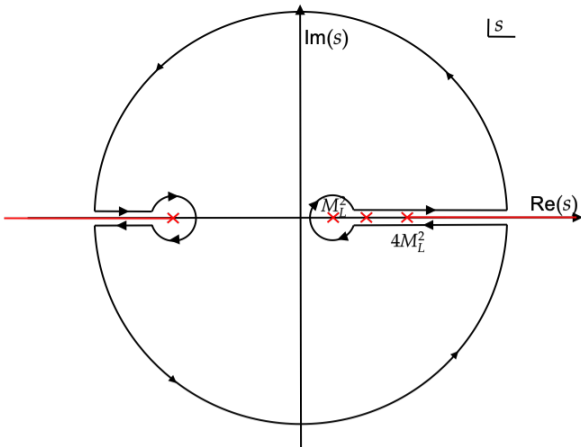
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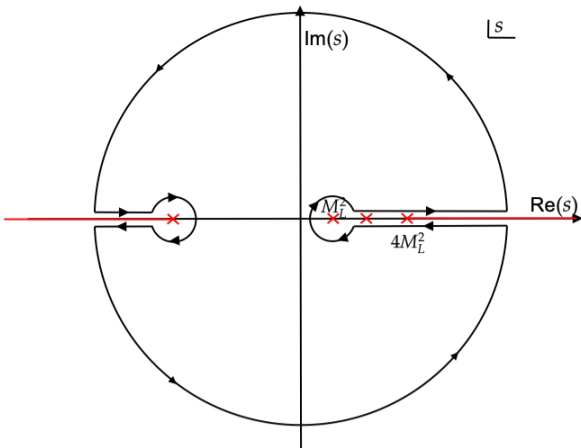
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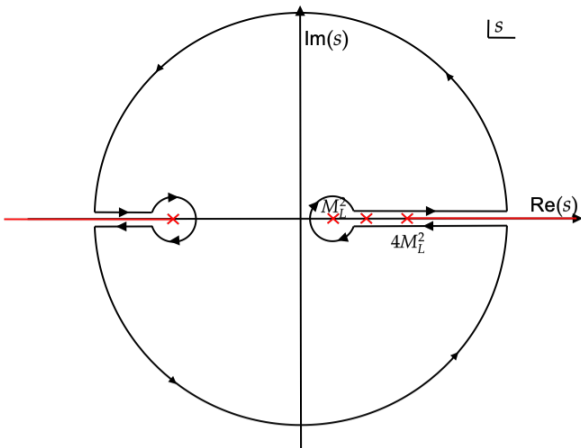


$\bar{A}(s,t)$ satisfies a dispersion relation:

$$\bar{A}(s,t) = \text{Res}_{\infty} \left(\frac{\bar{A}(s',t)}{s' - s} \right) + \int \frac{ds'}{2\pi i} \frac{\text{disc}(\bar{A}(s',t))}{s' - s}$$

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The residue at $s' = \infty$ is related to UV subtraction. We have included the isolated poles into the discontinuity along the real axis.

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$$\text{RHS} = \bar{A}_{\text{UV}}(s, t) = \text{Res}_{\infty} \left(\frac{\bar{A}_{\text{UV}}(s', t)}{s' - s} \right) + \int \frac{ds'}{2\pi i} \frac{\text{disc}(\bar{A}_{\text{UV}}(s', t))}{s' - s} = \bar{A}(s, t) \quad (11)$$

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Therefore, we can derive sum rules for the Wilson coefficients of the EFT evaluating the RHS integrals at $s = 0$ and taking derivatives on the LHS at $s = 0$:

$$\bar{A}(0, 0) = \alpha_0 = \text{Res}_{\infty} \left(\frac{\bar{A}_{\text{UV}}(s', 0)}{s'} \right) + \int \frac{ds'}{2\pi i} \frac{\text{disc}(\bar{A}_{\text{UV}}(s', 0))}{s'} \quad (13)$$

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$$K_k(t) = e^{i\Omega_k(t-t_*)} \Rightarrow K_\omega(t) = e^{i\omega(t-t_*)}, \quad \psi^{(n)}(\mathbf{k}) \Rightarrow \tilde{\psi}^{(n)}(\omega, \mathbf{k}) \quad (14)$$

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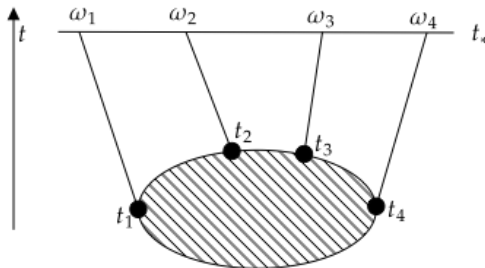
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Analytic structure I

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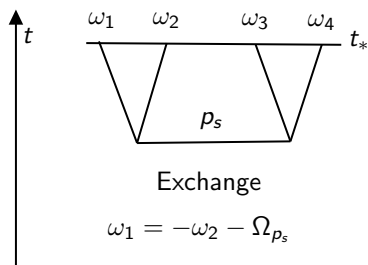
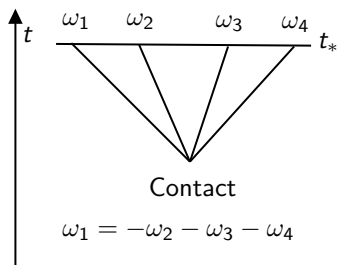
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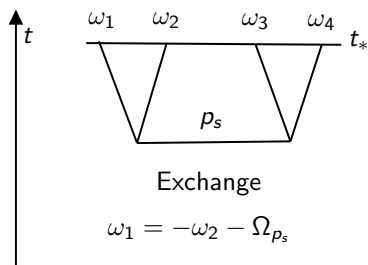
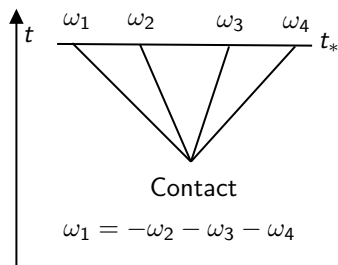
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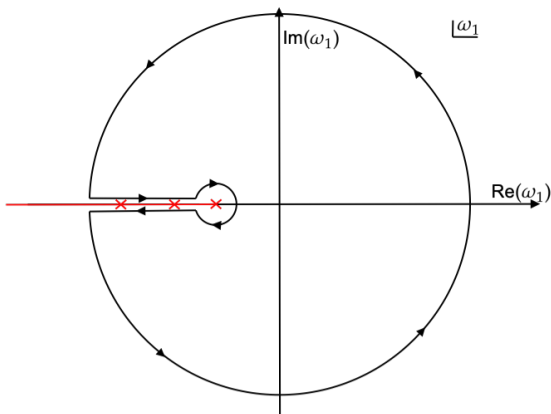
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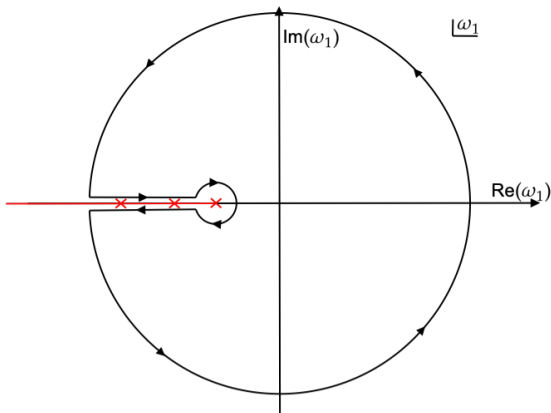
All singularities are located on the ω_1 negative real axis.

Analytic structure II



- $\omega_T \tilde{\psi}^{(n)}(\omega, \mathbf{k})$ is analytic in the lower half of the ω_1 complex plane.
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$\omega_T \tilde{\psi}^{(n)}(\omega, \mathbf{k})$ also satisfies a dispersion relation:

$$\omega_T \tilde{\psi}^{(n)}(\omega, \mathbf{k}) = \text{Res}_{\infty} \left(\frac{\omega_T \tilde{\psi}^{(n)}(\omega', \mathbf{k})}{\omega' - \omega_1} \right) + \int \frac{d\omega'}{2\pi i} \frac{\text{disc}(\omega_T \tilde{\psi}^{(n)}(\omega', \mathbf{k}))}{\omega' - \omega_1}$$

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$$\mathcal{L}_{\text{EFT}}^{\text{bulk}} \supset \frac{\alpha_0}{4!} \Phi^4 + \frac{\alpha_2}{4} \Phi^2 \square \Phi^2 + \frac{\alpha_4}{4} \Phi^2 \square^2 \Phi^2$$

$$\begin{aligned} \mathcal{L}_{\text{EFT}}^{\text{bdy}} \supset & \frac{\beta_{00}}{4!} \Phi^4(t_*) + \frac{\beta_{20}}{4} \Phi^2(t_*) \square \Phi^2(t_*) - \frac{\beta_{11}}{4} \Phi^2(t_*) \partial_t \Phi^2(t_*) \\ & - \frac{\beta_{22}}{4} \Phi^2(t_*) \partial_i^2 \Phi^2(t_*) - \frac{\beta_{31}}{4} \Phi^2(t_*) \partial_t \square \Phi^2(t_*) - \frac{\beta'_{31}}{4} \square \Phi^2(t_*) \partial_t \Phi^2(t_*) \end{aligned}$$

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$$\text{IR} \Rightarrow \omega_T \tilde{\psi}_{\text{IR}}^{(n)}(\omega, \mathbf{k}) = \int \frac{d\omega'}{2\pi i} \frac{\text{disc}(\omega_T \tilde{\psi}^{(n)}(\omega', \mathbf{k}))}{\omega' - \omega_1} \Leftarrow \text{UV} \quad (16)$$

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Like for amplitudes, we have to take derivatives with respect to $\omega_{a \neq 1}$ and the internal momenta p_s, p_t, p_u to obtain a complete list of the sum rules.

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- 1 The Wavefunction of the universe
- 2 To you, 60 years ago
- 3 Off-shell wavefunctions
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The future ahead:

- How is UV bulk unitarity encoded in the boundary Wilson coefficients?
- Can we build positivity bounds with these sum rules?
- How do we import more amplitudes technology to $\tilde{\psi}^{(n)}(\omega, \mathbf{k})$?