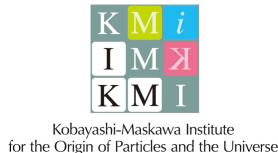


Dynamical Domain Wall and Localization

Shin'ichi Nojiri

Department of Physics &
Kobayashi-Maskawa Institute
for the Origin of Particles and the Universe (KMI),
Nagoya Univ.



Based on

Y. Toyozato, K. Bamba and S. Nojiri, “Scalar Domain Wall as the Universe”, Phys. Rev. D 87 (2013) 6, 063008 [arXiv:1202.5375 [hep-th]].

M. Higuchi and S. Nojiri, “Reconstruction of Domain Wall Universe and Localization of Gravity”, Gen. Rel. Grav. 46 (2014) 11, 1822 [arXiv:1402.1346 [hep-th]].

Y. Toyozato, M. Higuchi and S. Nojiri, “Dynamical Domain Wall and Localization”, arXiv:1510.01099.

Construction of the dynamical domain wall universe, where the four dimensional FRW universe is realized on the domain wall in the five dimensional space-time.

Localization of the fields (graviton, chiral spinor, vector, ...)

Introduction

Our universe could be a brane or domain wall embedded in a higher dimensional space-time.

K. Akama, “An Early Proposal of ‘Brane World’,” Lect. Notes Phys. 176 (1982) 267 [[hep-th/0001113](#)].

V. A. Rubakov and M. E. Shaposhnikov, “Do We Live Inside a Domain Wall?,” Phys. Lett. B 125 (1983) 136.

D-brane in string theories.

J. Dai, R. G. Leigh and J. Polchinski, “New Connections Between String Theories,” Mod. Phys. Lett. A 4 (1989) 2073.

J. Polchinski, “Tasi lectures on D-branes,” [hep-th/9611050](#).

Brane world scenario.

L. Randall and R. Sundrum, “A Large mass hierarchy from a small extra dimension,” Phys. Rev. Lett. 83, 3370 (1999) [hep-ph/9905221].

L. Randall, R. Sundrum, “An Alternative to compactification,” Phys. Rev. Lett. 83, 4690-4693 (1999). [hep-th/9906064].

Dynamics of brane?

Brane: a limit where the thickness of domain wall vanishes

⇒ Dynamics of domain wall

Domain wall model with two scalar fields

General FRW in the five dimensional space-time, the metric is given by

$$ds^2 = dw^2 + L^2 e^{u(w,t)} ds_{\text{FRW}}^2.$$
$$ds_{\text{FRW}}^2 = -dt^2 + a(t)^2 \left\{ \frac{dr^2}{1 - kr^2} + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2 \right\}.$$

Action with two scalar fields ϕ and χ :

$$S_{\phi\chi} = \int d^5x \sqrt{-g} \left\{ \frac{R}{2\kappa^2} - \frac{1}{2} A(\phi, \chi) \partial_\mu \phi \partial^\mu \phi - B(\phi, \chi) \partial_\mu \phi \partial^\mu \chi \right. \\ \left. - \frac{1}{2} C(\phi, \chi) \partial_\mu \chi \partial^\mu \chi - V(\phi, \chi) \right\}.$$

Energy-momentum tensor

$$T_{\mu\nu}^{\phi\chi} = g_{\mu\nu} \left\{ -\frac{1}{2} A(\phi, \chi) \partial_\rho \phi \partial^\rho \phi - B(\phi, \chi) \partial_\rho \phi \partial^\rho \chi - \frac{1}{2} C(\phi, \chi) \partial_\rho \chi \partial^\rho \chi - V(\phi, \chi) \right\}$$

$$+ A(\phi, \chi) \partial_\mu \phi \partial_\nu \phi + B(\phi, \chi) (\partial_\mu \phi \partial_\nu \chi + \partial_\nu \phi \partial_\mu \chi) + C(\phi, \chi) \partial_\mu \chi \partial_\nu \chi .$$

Field equations read

$$0 = \frac{1}{2} A_\phi \partial_\mu \phi \partial^\mu \phi + A \nabla^\mu \partial_\mu \phi + A_\chi \partial_\mu \phi \partial^\mu \chi + \left(B_\chi - \frac{1}{2} C_\phi \right) \partial_\mu \chi \partial^\mu \chi$$

$$+ B \nabla^\mu \partial_\mu \chi - V_\phi ,$$

$$0 = \left(-\frac{1}{2} A_\chi + B_\phi \right) \partial_\mu \phi \partial^\mu \phi + B \nabla^\mu \partial_\mu \phi + \frac{1}{2} C_\chi \partial_\mu \chi \partial^\mu \chi + C \nabla^\mu \partial_\mu \chi$$

$$+ C_\phi \partial_\mu \phi \partial^\mu \chi - V_\chi .$$

$A_\phi = \partial A(\phi, \chi) / \partial \phi$, **etc.**

We now choose $\phi = t$ **and** $\chi = w$.

We can construct a model to realize the arbitrary metric in the form,
 $ds^2 = dw^2 + L^2 e^{u(w,t)} ds_{\text{FRW}}^2$,
 $ds_{\text{FRW}}^2 = -dt^2 + a(t)^2 \left\{ \frac{dr^2}{1-kr^2} + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2 \right\}$,
by choosing ($H \equiv (1/a)(da/dt)$)

$$A = \frac{1}{\kappa^2} \left(\frac{2k}{a^2} - \ddot{u} - 2\dot{H} + \frac{(\dot{u})^2}{2} + \dot{u}H \right) ,$$

$$B = -\frac{3u'}{2\kappa^2 L^2 e^u} (\dot{u} + 2H) ,$$

$$C = \frac{1}{\kappa^2} \left(-\frac{3}{2}u'' + \frac{2k}{L^2 e^u a^2} - \frac{1}{2e^u} (\ddot{u} + 2\dot{H} + (\dot{u})^2 + 5\dot{u}H + 6H^2) \right) ,$$

$$V = \frac{1}{\kappa^2} \left(-\frac{3}{4} \left(u'' + 2(u')^2 \right) + \frac{3k}{L^2 e^u a^2} + \frac{1}{4L^2 e^u} (3\ddot{u} + 6\dot{H} + 3(\dot{u})^2 + 15\dot{u} + 18H^2) \right) .$$

Einstein equation is satisfied.

Field equations are nothing but the Bianchi identities.

If any eigenvalue of the matrix $\begin{pmatrix} A & B \\ B & C \end{pmatrix}$ is negative, there appears a ghost.

Example without ghost

$$a(t) \propto t^{h_0}, \quad e^{u(w,t)} = W(w) T(t),$$

$$T(t) = T_1 t^{1-3h_0} + T_2 t^{-2h_0}, \quad W(w) = e^{-\sqrt{1+\frac{w^2}{w_0^2}}},$$

$$\Rightarrow A = \frac{3}{2\kappa^2} \left(\frac{\dot{T}(t)}{T(t)} + \frac{2h_0}{t} \right)^2 > 0, \quad B = 0, \quad C = \frac{3}{2\kappa^2} \frac{\frac{1}{w_0^2}}{\left(1 + \frac{w^2}{w_0^2}\right)^{\frac{3}{2}}} > 0.$$

... No ghost.

General FRW universe can be realized by the Brans-Dicke type model.

Localization of graviton

Perturbation $g_{\mu\nu} \rightarrow g_{\mu\nu} + h_{\mu\nu}$.

Assume $h_{\mu\nu} = 0$ if μ or $\nu \neq 1, 2, 3$.

Einstein eq. \Rightarrow **equation for graviton in five dimensions:**

$$0 = \left[\partial_w^2 - u'' - u'^2 + L^{-2} e^{-u} \left(\ddot{u} + \frac{\dot{a}\dot{u}}{a} + \dot{u}\partial_t + 2\frac{\ddot{a}}{a} + \frac{\dot{a}}{a}\partial_t - \partial_t^2 + \frac{\Delta}{a^2} \right) \right] h_{ij} .$$

By assuming $h_{ij}(w, x) = e^{u(w,t)} \hat{h}_{ij}(x)$,

$$0 = \left(2\frac{\dot{a}\dot{u}}{a} - \dot{u}\partial_t + 2\frac{\ddot{a}}{a} + \frac{\dot{a}}{a}\partial_t - \partial_t^2 + \frac{\Delta}{a^2} \right) \hat{h}_{ij} .$$

If u goes to minus infinity sufficiently rapidly for large $|w|$, $h_{ij}(w, x)$ is normalized in the direction of w and therefore there occurs the localization of graviton.

Graviton localized on the domain wall

$$0 = \left(2\frac{\dot{a}\dot{u}}{a} - \dot{u}\partial_t + 2\frac{\ddot{a}}{a} + \frac{\dot{a}}{a}\partial_t - \partial_t^2 + \frac{\Delta}{a^2} \right) \hat{h}_{ij} .$$

In the four dimensional FRW space-time,

$$0 = \left(2\frac{\ddot{a}}{a} + \frac{\dot{a}}{a}\partial_t - \partial_t^2 + \frac{\Delta}{a^2} \right) h_{ij} .$$

If $\dot{u} \left(2\frac{\dot{a}}{a} - \partial_t \right) \hat{h}_{ij} = 0$, two expressions coincide with each other.
If $\dot{u} \neq 0$, there could appear some corrections proportional to \dot{u} .

Localization of chiral spinor

There was an attempt to localize the fermion on the domain wall has been failed.

D. P. George, M. Trodden and R. R. Volkas, “Extra-dimensional cosmology with domain-wall branes,” JHEP 0902 (2009) 035 [arXiv:0810.3746 [hep-ph]].

Dirac equation in five dimensions

$$\Gamma^M \nabla_M \Psi + \tilde{f} \chi(w) \Psi = 0.$$

$\tilde{f} \chi$: a function of the scalar field $\chi = w$, general yukawa interaction.

∇_M : covariant derivative:

$$\nabla_M \stackrel{\text{def}}{=} \partial_M + \frac{1}{4} \omega_{ABM} \Gamma^{AB}, \quad \Gamma^{AB} \stackrel{\text{def}}{=} \frac{1}{2} [\Gamma^A, \Gamma^B].$$

$\Psi = \eta(t, w) \psi(x)$ (ψ : **four dimensional Dirac spinor**)

$$\begin{aligned} & \Gamma^M \nabla_M \Psi + \tilde{f} \chi(w) \Psi \\ &= L^{-1} e^{-u/2} \left\{ \Gamma^{\hat{0}} \partial_0 \psi + \frac{1}{a(t)} \Gamma^{\hat{i}} \partial_i \psi \right\} \eta + \{ \partial_5 \eta + u'(t, w) \eta \} \Gamma^{\hat{5}} \psi \\ &+ \tilde{f} \chi(w) \eta \psi + L^{-1} e^{-u/2} \left\{ \partial_0 \eta + \frac{3}{2} \left(\frac{1}{2} \dot{u} + \frac{\dot{a}}{a} \right) \eta \right\} \Gamma^{\hat{0}} \psi. \end{aligned}$$

Let $\tilde{C}_1(w)$ and $D_1(t)$ be arbitrary functions.

$$\begin{aligned} \eta(t, w) &\stackrel{\text{def}}{=} \zeta(t) \lambda(w) g(t, w) \\ g(t, w) &\stackrel{\text{def}}{=} \exp \left[\tilde{C}_1(w) + D_1(t) - u(t, w) \right], \end{aligned}$$

By assuming $\Gamma^5\psi = \pm\psi$,

$$\begin{aligned}
& \Rightarrow \Gamma^M \nabla_M \Psi + \tilde{f} \chi(w) \Psi \\
&= L^{-1} e^{-u/2} \left\{ \Gamma^{\hat{0}} \partial_0 \psi + \frac{1}{a(t)} \Gamma^{\hat{i}} \partial_i \psi \right\} \eta \\
&\quad \pm \zeta g \left\{ \partial_5 \lambda(w) + \tilde{C}'_1(w) \lambda(w) \pm \tilde{f} \chi(w) \lambda(w) \right\} \psi \\
&\quad + L^{-1} e^{-u/2} \lambda g \left\{ \partial_0 \zeta(t) + \left(\frac{3\dot{a}}{2a} - \frac{1}{4}\dot{u}(t, w) + \dot{D}_1(t) \right) \zeta(t) \right\} \Gamma^{\hat{0}} \psi.
\end{aligned}$$

$e^{u(t,w)} = T(t) W(w)$ **and using the Dirac equation,**

$$\begin{aligned}
\lambda(w) &= C_3 \exp \left[-\tilde{C}_1(w) \mp \tilde{f} \int \mathbf{d}w \chi(w) \right] \\
\zeta(t) &= C_4 a^{-3/2} T^{1/4} \exp [-D_1(t)].
\end{aligned}$$

\Rightarrow

$$\begin{aligned}\eta(t, w) &= C_3 C_4 a^{-3/2} T^{1/4} \exp \left[-u(t, w) \mp \tilde{f} \int \mathbf{d}w \chi(w) \right] \\ &= C_5 a^{-3/2} T^{-3/4} W^{-1} \exp \left[\mp \tilde{f} \int \mathbf{d}w \chi(w) \right],\end{aligned}$$

$$\chi(w) = w \Rightarrow$$

$$\eta(t, w) = C_5 a^{-3/2} T^{-3/4} W^{-1} \exp \left[\mp \frac{\tilde{f}}{2} w^2 \right].$$

Condition of localization

$$I = \int_{-\infty}^{+\infty} \mathbf{d}w e^{3u/2} |\eta|^2 = C_5^2 a^{-3} \int_{-\infty}^{+\infty} \mathbf{d}w W^{-1/2} \exp \left[\mp \tilde{f} w^2 \right] < \infty.$$

Fermion can be localized on the domain wall and the localized fermion can be chiral or anti-chiral.

In

D. P. George, M. Trodden and R. R. Volkas, “Extra-dimensional cosmology with domain-wall branes,” JHEP 0902 (2009) 035
[arXiv:0810.3746 [hep-ph]]

It was assumed that the warp factor does not depend on the time but we have used the time-dependent warp factor.

Localization of vector field

Action of vector field

$$S_V = \int d^5x \sqrt{-g} \left\{ -\frac{1}{4} F_{MN} F^{MN} - \frac{1}{2} m(\chi)^2 A_M A^M \right\},$$
$$F_{MN} = \partial_M A_N - \partial_N A_M.$$

Background

$$ds^2 = dw^2 + L^2 W(w) T(t) ds_{\text{FRW}}^2,$$

$$ds_{\text{FRW}}^2 = -dt^2 + a(t)^2 \sum_{i=1}^3 (dx^i)^2.$$

\Rightarrow

$$\begin{aligned} S_V = \int d^5x & \left\{ \frac{1}{2}L^2W(w)T(t)a(t)^3F_{50}^2 - \frac{1}{2}L^2W(w)T(t)a(t)F_{5i}^2 \right. \\ & + \frac{1}{2}a(t)F_{0i}^2 - \frac{1}{4}a(t)^{-1}F_{ij}^2 \\ & - \frac{1}{2}m(\chi)^2 \left(L^4W(w)^2T(t)^2a(t)^3A_5^2 \right. \\ & \left. \left. - L^2W(w)T(t)a(t)^3A_0^2 + L^2W(w)T(t)a(t)A_i^2 \right) \right\}. \end{aligned}$$

Variation of A_5 , A_0 , and $A_i \Rightarrow$

$$0 = L^2 W(w) \partial_0 (T(t) a(t)^3 (\partial_5 A_0 - \partial_0 A_5))$$

$$- L^2 W(w) T(t) a(t) (\partial_5 \partial_i A_i - \partial_i^2 A_5)$$

$$- m(\chi)^2 L^4 W(w)^2 T(t)^2 a(t)^3 A_5 ,$$

$$0 = - L^2 T(t) a(t)^3 \partial_5 (W(w) (\partial_5 A_0 - \partial_0 A_5))$$

$$+ a(t) (\partial_0 \partial_i A_i - \partial_i^2 A_0) + m(\chi)^2 L^2 W(w) T(t) a(t)^3 A_0 ,$$

$$0 = L^2 T(t) a(t) \partial_w (W(w) (\partial_5 A_i - \partial_i A_5)) - \partial_0 (a (\partial_0 A_i - \partial_i A_0))$$

$$- a(t)^{-1} (\partial_i \partial_j A_j - \partial_j^2 A_i) - m(\chi)^2 L^2 W(w) T(t) a(t) A_i .$$

Assuming

$$A_5 = 0, \quad A_\mu = X(w)C_\mu(x^\nu), \quad \mu, \nu = 0, 1, 2, 3,$$

and choosing $m(\chi = w)^2 = \frac{(W(w)X'(w))'}{W(w)X(w)},$

$$\Rightarrow 0 = \partial_5 X(w) \left\{ \partial_0 (T(t)a(t)^3 C_0) - T(t)a(t)\partial_i C_i \right\},$$

$$0 = \partial_0 \partial_i C_i - \partial_i^2 C_0,$$

$$0 = \partial_0 (a(t)(\partial_0 C_i - \partial_i C_0)) + a(t)^{-1} (\partial_i \partial_j C_j - \partial_j^2 C_i).$$

Last two eqs. : nothing but the field equations of the vector field in four dimensions.

First eq. : a gauge condition, which is a generalization of the Landau gauge, $\partial^\mu A_\mu = 0$.

If $X(w) \rightarrow 0$ sufficiently when $|w| \rightarrow \infty$, A_μ : normalizable.
 \Rightarrow **Localization of vector field on the domain wall.**

In case

$$(W(w)X'(w))' = 0 \Rightarrow m(\chi) = 0.$$

\Rightarrow **non-abelian gauge theory?**
(In general, non-normalizable)

Summary

- We have constructed the dynamical domain wall, where arbitrary four dimensional FRW universe is realized on the domain wall in the five dimensional space-time.
- Graviton, Chiral Spinor, and Vector can localize on the brane.

Graviton in FRW Universe in Four Dimensions

Perturbation $g_{\mu\nu} \rightarrow g_{\mu\nu} + h_{\mu\nu} \Rightarrow$

$$\begin{aligned}\delta R_{\mu\nu} = & \frac{1}{2} [\nabla_\mu \nabla^\rho h_{\nu\rho} + \nabla_\nu \nabla^\rho h_{\mu\rho} - \nabla^2 h_{\mu\nu} - \nabla_\mu \nabla_\nu (g^{\rho\lambda} h_{\rho\lambda}) \\ & - 2R^\lambda_\nu h_{\lambda\rho} + R^\rho_\mu h_{\mu\nu} + R^\rho_\nu h_{\rho\mu}] , \\ \delta R = & - h_{\mu\nu} R^{\mu\nu} + \nabla^\mu \nabla^\nu h_{\mu\nu} - \nabla^2 (g^{\mu\nu} h_{\mu\nu}) .\end{aligned}$$

Imposing the gauge condition $\nabla^\mu h_{\mu\nu} = g^{\mu\nu} h_{\mu\nu} = 0$

Einstein equation $R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = \kappa^2 T_{\mu\nu} \Rightarrow$

$$\begin{aligned}\frac{1}{2} [& -\nabla^2 h_{\mu\nu} - 2R^\lambda_\nu h_{\lambda\rho} + R^\rho_\mu h_{\mu\nu} + R^\rho_\nu h_{\rho\mu} \\ & - h_{\mu\nu} R + g_{\mu\nu} R^{\rho\lambda} h_{\rho\lambda}] = \kappa^2 \delta T_{\mu\nu} .\end{aligned}$$

By using the formulation in

S. Nojiri and S. D. Odintsov, “Unifying phantom inflation with late-time acceleration: Scalar phantom-non-phantom transition model and generalized holographic dark energy”, Gen. Rel. Grav. 38 (2006) 1285 [hep-th/0506212].

consider scalar field theory

$$\begin{aligned} S_\phi &= \int d^4x \sqrt{-g} \mathcal{L}_\phi, \quad \mathcal{L}_\phi = -\frac{1}{2}\omega(\phi)\partial_\mu\phi\partial^\mu\phi - V(\phi), \\ \Rightarrow \quad T_{\mu\nu} &= -\omega(\phi)\partial_\mu\phi\partial_\nu\phi + g_{\mu\nu}\mathcal{L}_\phi, \\ \Rightarrow \quad \delta T_{\mu\nu} &= h_{\mu\nu}\mathcal{L}_\phi + \frac{1}{2}g_{\mu\nu}\omega(\phi)\partial^\rho\phi\partial^\lambda\phi h_{\rho\lambda}. \end{aligned}$$

Because we are now interested in the graviton, we may assume $h_{\mu\nu} = 0$ except the components with $\mu, \nu = 1, 2, 3$.

In the FRW universe ($k = 0$), assume $\phi = t$, FRW equations

$$\begin{aligned} \frac{3}{\kappa^2} \frac{\dot{a}^2}{a^2} &= \frac{\omega}{2} + V, \quad \frac{1}{\kappa^2} \left(2 \frac{\ddot{a}}{a} + \frac{\dot{a}^2}{a^2} \right) = \frac{\omega}{2} - V, \\ \Rightarrow \quad \omega &= \frac{1}{\kappa^2} \left(2 \frac{\ddot{a}}{a} + 4 \frac{\dot{a}^2}{a^2} \right), \quad V = -\frac{1}{\kappa^2} \left(\frac{\ddot{a}}{a} + \frac{\dot{a}^2}{a^2} \right). \\ \Rightarrow \quad 0 &= \left(2 \frac{\ddot{a}}{a} + \frac{\dot{a}}{a} \partial_t - \partial_t^2 + \frac{\Delta}{a^2} \right) h_{ij}. \end{aligned}$$

Explicit expressions of connections and curvatures in five dimensions

Metric

$$g_{AB} = \begin{pmatrix} -L^2 e^{u(w,t)} & & & & \\ & L^2 e^{u(w,t)} \frac{a(t)^2}{1-kr^2} & & & \\ & & L^2 e^{u(w,t)} a(t)^2 r^2 & & \\ & & & L^2 e^{u(w,t)} a(t)^2 r^2 \sin^2 \theta & \\ & & & & 1 \end{pmatrix}.$$

Connections

$$\Gamma_{tt}^t = \frac{1}{2}\dot{u}, \quad \Gamma_{tt}^w = \frac{1}{2}L^2 e^u u', \quad \Gamma_{rt}^r = \Gamma_{\theta t}^\theta = \Gamma_{\phi t}^\phi = \frac{\dot{a}}{a} + \frac{1}{2}\dot{u}, \quad \Gamma_{tw}^t = \Gamma_{rw}^r = \Gamma_{\theta w}^\theta = \Gamma_{\phi w}^\phi = \frac{1}{2}u',$$

$$\Gamma_{ij}^t = L^{-2} e^{-u} \left(\frac{\dot{a}}{a} + \frac{1}{2}\dot{u} \right) g_{ij}, \quad \Gamma_{rr}^r = \frac{kr}{1-kr^2}, \quad \Gamma_{ij}^w = -\frac{1}{2}u' g_{ij}, \quad \Gamma_{\theta r}^\theta = \Gamma_{\phi r}^\phi = \frac{1}{r},$$

$$\Gamma_{\theta\theta}^r = -r(1-kr^2), \quad \Gamma_{\phi\theta}^\phi = \cot \theta, \quad \Gamma_{\phi\phi}^r = -r(1-kr^2) \sin^2 \theta, \quad \Gamma_{\phi\phi}^\theta = -\cos \theta \sin \theta.$$

Ricci curvatures

$$R_{tt} = \left[-\frac{1}{2}u'' - u'^2 + \frac{3}{2}L^{-2}\mathrm{e}^{-u} \left(\ddot{u} + \frac{\dot{a}\dot{u}}{a} + 2\frac{\ddot{a}}{a} \right) \right] g_{tt},$$
$$R_{ij} = \left[-\frac{1}{2}u'' - u'^2 + \frac{1}{2}L^{-2}\mathrm{e}^{-u} \left(\ddot{u} + 5\frac{\dot{a}\dot{u}}{a} + 2\frac{\ddot{a}}{a} + 4\frac{\dot{a}^2}{a^2} + \dot{u}^2 + 4\frac{k}{a^2} \right) \right] g_{ij},$$

$$R_{ww} = -2u'' - u'^2$$

$$R_{tw} = -\frac{3}{2}\dot{u}'.$$

Scalar curvature

$$R = -4u'' - 5u'^2 + 3L^{-2}\mathrm{e}^{-u} \left(\ddot{u} + \frac{1}{2}\dot{u}^2 + 3\frac{\dot{a}\dot{u}}{a} + 2\frac{\ddot{a}}{a} + 2\frac{\dot{a}^2}{a^2} + 2\frac{k}{a^2} \right).$$

Explicit form of spin connection

Vierbein field $e^{\hat{B}}$, $g_{MN} = \overset{\text{def}}{=} e^{\hat{A}}_M \eta_{\hat{A}\hat{B}} e^{\hat{B}}_N$

$$\Rightarrow e^{\hat{A}}_M = \begin{bmatrix} L e^{u/2} & 0 & 0 & 0 & 0 \\ 0 & L e^{u/2} a & 0 & 0 & 0 \\ 0 & 0 & L e^{u/2} a & 0 & 0 \\ 0 & 0 & 0 & L e^{u/2} a & 0 \\ 0 & 0 & 0 & 0 & 1 \end{bmatrix} .$$

$$e_{\hat{A}M} \stackrel{\text{def}}{=} \eta_{\hat{A}\hat{B}} e^{\hat{B}}_M = \begin{bmatrix} -L e^{u/2} & 0 & 0 & 0 & 0 \\ 0 & L e^{u/2} a & 0 & 0 & 0 \\ 0 & 0 & L e^{u/2} a & 0 & 0 \\ 0 & 0 & 0 & L e^{u/2} a & 0 \\ 0 & 0 & 0 & 0 & 1 \end{bmatrix} ,$$

$$e_{\hat{A}}{}^M \stackrel{\text{def}}{=} e_{\hat{A}N} g^{MN}$$

$$= \begin{bmatrix} L^{-1} e^{-u/2} & 0 & 0 & 0 & 0 \\ 0 & L^{-1} e^{-u/2} a^{-1} & 0 & 0 & 0 \\ 0 & 0 & L^{-1} e^{-u/2} a^{-1} & 0 & 0 \\ 0 & 0 & 0 & L^{-1} e^{-u/2} a^{-1} & 0 \\ 0 & 0 & 0 & 0 & 1 \end{bmatrix}.$$

Ricci rotation coefficients and spin connections

$$\Omega_{ABM} \stackrel{\text{def}}{=} \left(\partial_A e^{\hat{C}}_B - \partial_B e^{\hat{C}}_A \right) e_{\hat{C}M} = -\Omega_{BAM},$$

$$\omega_{ABM} \stackrel{\text{def}}{=} -\frac{1}{2} (\Omega_{ABM} - \Omega_{BMA} - \Omega_{MAB}) = -\omega_{BAM}.$$

Non-vanishing components

$$\begin{cases} \Omega_{050} = -\Omega_{500} = \frac{1}{2}u'L^2e^u \\ \Omega_{0ij} = -\Omega_{i0j} = \left(\frac{1}{2}\dot{u}a^2 + \dot{a}a\right)L^2e^u\delta_{ij}, \\ \Omega_{5ij} = -\Omega_{i5j} = \frac{1}{2}u'L^2e^ua^2\delta_{ij}, \end{cases}$$

$$\begin{cases} \omega_{050} = -\omega_{500} = -\frac{1}{2}u'L^2e^u, \\ \omega_{0ij} = -\omega_{i0j} = -\left(\frac{1}{2}\dot{u}a^2 + \dot{a}a\right)L^2e^u\delta_{ij}, \\ \omega_{5ij} = -\omega_{i5j} = -\frac{1}{2}u'L^2e^ua^2\delta_{ij}. \end{cases}$$