

Tunneling amplitudes in QFT

a Lorentzian worldline perspective

Karthik Rajeev

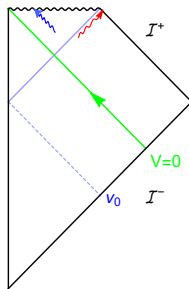
with

Anton Ilderton & William Lindved

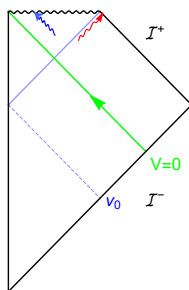


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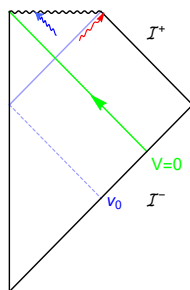


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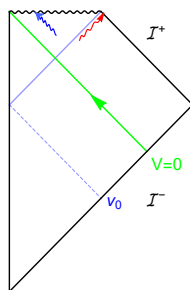
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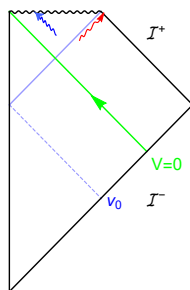
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- Analytic in the upper-half of the complex- v plane.
- Implies $\tilde{\mathcal{M}}(v \rightarrow \mathcal{E}')$ cannot vanish for all of $v > v_0$: **Tunneling!**

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- **From here on:** scalar QED for Schwinger effect and massless scalar for Hawking radiation.

Schwinger effect

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- Number of pairs $\sim e^{-\frac{(p_{\perp}^2 + m^2)\pi}{|eE|}}$ \rightarrow Non-perturbative.

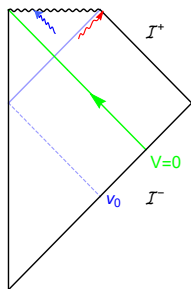
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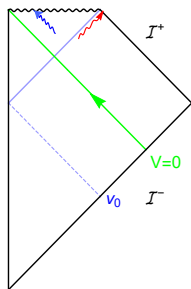
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- $|\text{Pair creation amplitude}|^2 \sim e^{-8\pi GM\mathcal{E}} \rightarrow \text{Thermal distribution.}$

Classical physics: worldlines in constant electric field

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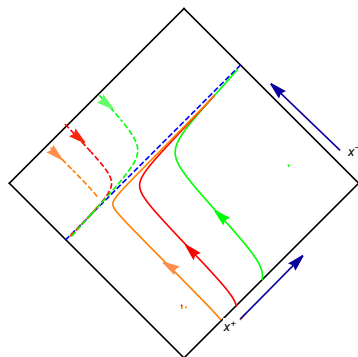
$$x_{\text{cl}}^-(\tau) = x_h^-(1 - e^{-2eE\tau}) \quad ; \quad x_h^- = \frac{p_+}{eE}$$

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Horizon? [...Ilderton & Lindved (2023)]



Schwinger effect

$$\varphi_p(x) = e^{-ip_+x^+ - ip_\perp x^\perp} e^{-\left(\frac{1}{2} - i\frac{m^2 + p_\perp^2}{2eE}\right) \log\left(1 - \frac{x^-}{x_h^-} + i0^+\right)} \quad (1)$$

- Positive-energy w.r.t x^- .

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- Tunneling exponential

$$\left| \frac{\text{Pre-exp factor in forbidden regn.}}{\text{Pre-exp factor in allowed regn.}} \right|^2 = \left| \frac{\mathcal{M}_{0 \rightarrow 2}}{\mathcal{M}_{1 \rightarrow 1}} \right|^2 = e^{-\frac{\pi(p_\perp^2 + m^2)}{|eE|}} \quad (2)$$

- Tunneling mode-function can be written as

$$\varphi_p^+(x) = \int_{-\infty}^{\infty} dT \int^{z(\tau_2)=x} \mathcal{D}[z] e^{iS_{\text{WL}}[z, T]},$$

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- QM analogue: fixed-energy wavefunction from PI, introduce a Lagrange multiplier T .

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Proper time integral \Rightarrow on-shell condition

- The easiest way to see this is to consider the free case

$$\begin{aligned}\varphi_p^+(x) &= \int_{-\infty}^{\infty} dT e^{-ip \cdot x + i(p^2 - m^2)T}, \\ &\propto \delta(p^2 - m^2) e^{-ip \cdot x}\end{aligned}$$

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- 1 Exponential is nothing but the classical worldline action S_{WL}
- 2 The on-shell condition is equivalent to $\partial_T S_{WL} = 0$.

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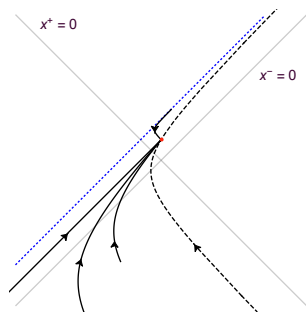
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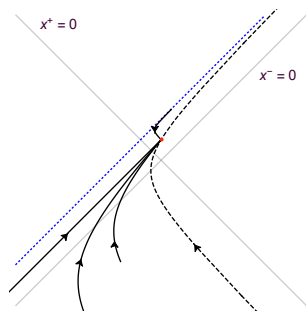
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- $\int dT (\dots) \leftrightarrow$ sum over WLs parametrised by T .

Classically allowed region: $x^- < x_h^-$



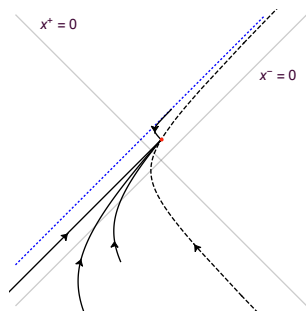
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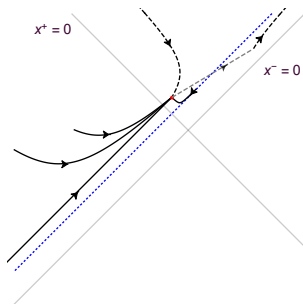
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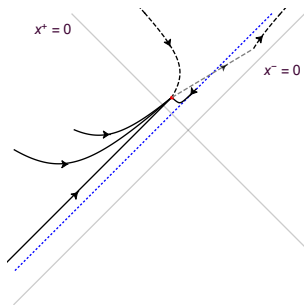
$$\varphi_p(x) = \mathcal{N} e^{-ip_{\perp}x^{\perp} - ip_{+}x^{+}} e^{-\left(\frac{1}{2} - i\frac{p_{\perp}^2 + m^2}{2eE}\right) \log\left(1 - \frac{x^-}{x_h^-}\right)} \quad (3)$$

Classically forbidden region: $x^- > x_h^-$



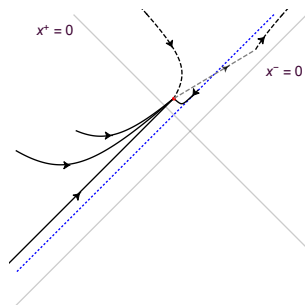
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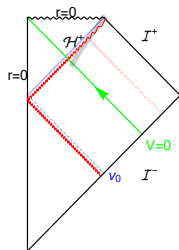


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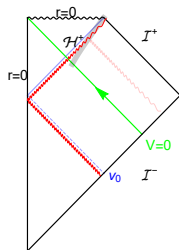
$$\varphi_p(x) = e^{-\pi \frac{p_{\perp}^2 + m^2}{2eE}} e^{-i\frac{\pi}{2}} \mathcal{N} e^{-ip_{\perp}x^{\perp} - ip_+x^+} e^{-\left(\frac{1}{2} - i\frac{p_{\perp}^2 + m^2}{2eE}\right) \log\left(\frac{x^-}{x_h^-} - 1\right)} \quad (4)$$

Hawking radiation: standard approach

Out-modes



- Near \mathcal{I}^+ , outmodes take the form $\varphi^{(\text{out})} \sim \frac{1}{r} e^{-i\mathcal{E}u}$



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- Near $v \sim v_0$ in \mathcal{I}^- , we get

$$\varphi^{(\text{out})} \sim \frac{1}{r} e^{-i\mathcal{E}v + 4GMi \log[(v_0 - v)/|v_0]} \Theta[v_0 - v] \quad (5)$$

$$= \sum_{\mathcal{E}'} \alpha_{\mathcal{E}\mathcal{E}'} e^{-i\mathcal{E}'v} + \sum_{\mathcal{E}'} \beta_{\mathcal{E}\mathcal{E}'} e^{i\mathcal{E}'v} \quad (6)$$

The Feynman modes

- Out(in) modes encode information about **Bogoliubov coefficients**.

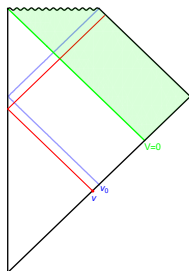
The Feynman modes

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- The Feynman-modes encode **amplitudes**

$$\bar{\varphi}^{(F)}(x; p) \approx \begin{cases} \int_{p'} \mathcal{M}_{1 \rightarrow 1}(p' \rightarrow p) e^{ip' \cdot x} & ; \quad x \stackrel{\rightarrow}{\in} \text{Asym. past} \\ e^{ip \cdot x} + \int_{p'} \mathcal{M}_{0 \rightarrow 2}(p, p') e^{-ip' \cdot x} & ; \quad x \stackrel{\rightarrow}{\in} \text{Asym. future} \end{cases} \quad (7)$$

2. Hawking radiation and tunneling

[Hartle & Hawking (1976), Damour & Ruffini (1976), Srinivasan & Padmanabhan (1999), Parikh & Wilczek (2000),....]



$$\mathcal{E}' - \mathcal{E} \equiv \Delta\mathcal{E} = - \left. \frac{2G \tilde{p} \cdot p}{r} \right|_{v \rightarrow 0^-} = \frac{4GM\mathcal{E}}{v} \quad (8)$$

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- Therefore

$$S_{\text{WL}}(\mathcal{E}', \nu) = \mathcal{E}'\nu - 4GM\mathcal{E}' \log\left(\frac{\nu_0 - \nu - i0^+}{r_0}\right) \quad (10)$$

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- Imaginary part of S_{WL} for $v > v_0$ leads to $\mathcal{P}_{\mathcal{E}'} = e^{-8\pi GM\mathcal{E}'}$

Path integral formalism for Hawking radiation

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- **This work:** Systematic treatment of the dynamical background in *Lorentzian* WL approach.

- Feynman mode for Hawking radiation

$$\varphi_{\mathcal{E}'}(x) \stackrel{?}{=} \int_{-\infty}^{\infty} dT \int^{z(\lambda_2)=x} \mathcal{D}[z] e^{iS_{\text{WL}}[z, T]},$$

$$S_{\text{WL}} = - \int_{\lambda_1}^{\lambda_2} \frac{1}{4} g_{\mu\nu}(z) \dot{z}^\mu \dot{z}^\nu d\lambda + \underbrace{S_{\text{bdy.}}}_{\text{Boundary term}},$$

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Boundary conditions

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$$z(\lambda_2) = (r \rightarrow \infty, \nu, \theta, \phi)$$

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$$z(\lambda_2) = (r \rightarrow \infty, v, \theta, \phi)$$

- Outgoing and zero-angular-momentum conditions

$$\underbrace{\dot{V}(\lambda_2) - \frac{\dot{R}(\lambda_2)}{1 - \frac{2GM}{R(\lambda_2)}}}_{U(\lambda_2)=0 \text{ (outgoing)}} = 0 \quad ; \quad \underbrace{\frac{\dot{R}(\lambda_2)}{2}}_{\text{fixes energy}} = \mathcal{E}' > 0$$

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- Boundary term turns out to be

$$S_{\text{bdy.}} = -\mathcal{E}' \left[V - 2R - 4GM \log \left(\frac{R}{2GM} - 1 \right) \right] \equiv -\mathcal{E}' U$$

- Semiclassical approximation gives

$$\varphi_{\mathcal{E}'} \approx \frac{1}{r} e^{iS_{\text{WL}}[z_{\text{cl}}]}$$

Geometric optics approx = semiclassical approx.

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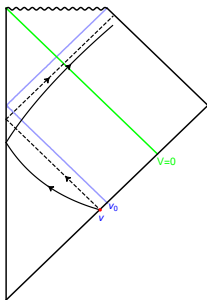
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- Hence $S_{\text{WL}} = S_{\text{bdy.}} = -\mathcal{E}' U$
- Consistent with Hawking's ray-tracing argument.

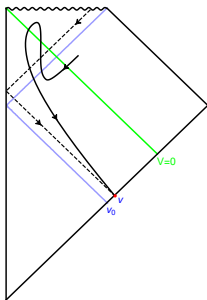
Case 1: Classically allowed ($v < v_0$)



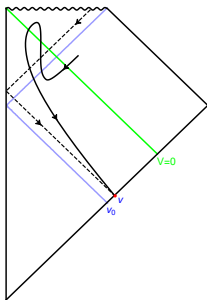
- The Feynman mode takes the form

$$\varphi_{\mathcal{E}'}(x) \sim \frac{1}{r} e^{-i\mathcal{E}'v + i4GM\mathcal{E}' \log\left(\frac{v_0 - v}{2GM}\right)}$$

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Amplitudes from Feynman-modes

- Recall

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- We have

$$\bar{\varphi}_{\mathcal{E}'}(x) = \begin{cases} \frac{e^{i\mathcal{E}'v - i4GM\mathcal{E}' \log\left(\frac{v_0 - v - i0^+}{\mu}\right)}}{\mathcal{E}'r} & ; \quad x \in \mathcal{I}^- \\ \frac{e^{i\mathcal{E}'u}}{\mathcal{E}'r} & ; \quad x \in \mathcal{I}^+ \\ e^{-4\pi GM\mathcal{E}'} \phi^{\mathcal{H}^+}(x; \mathcal{E}') & ; \quad x \in \mathcal{H}^+ \end{cases} \quad (12)$$

Schwinger Effect

Hawking radiation

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- **Exact** PI evaluation of tunneling wavefunction.

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- Picard-Lefschetz analysis? (However, T -integral is still just a normalization.)

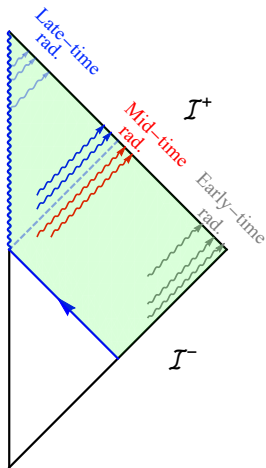


THANK YOU!²

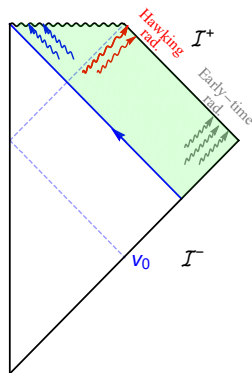
²Bonus slides below!

Hawking radiation from the double copy

[A. Ilderton, W. Lindved, [KB](#), [PRL. 136 (2026)]]



(a) $\sqrt{\text{Vaidya}}$



(b) \square Vaidya

Classical Double Copy of Geodesics

- For $\sqrt{\text{Vaidya}}$:

$$\Delta\mathcal{E} = \frac{2g^2 qQ}{v}. \quad (13)$$

³In the spirit of [Goldberger and Ridgway, (2017)], [Gonzo and Shi, (2021)], ... 

Classical Double Copy of Geodesics

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- Double-copy rule!³

$$c^a \tilde{c}^a \rightarrow -p \cdot \tilde{p} \quad ; \quad g^2 \rightarrow 2G \quad (14)$$

where $\tilde{p} = Mdt$ is the BH and p is the *initial* probe momenta.

³In the spirit of [Goldberger and Ridgway, (2017)], [Gonzo and Shi, (2021)], ...

Worldline Representation of Modes

- The wavefunctions, for $x \rightarrow \mathcal{I}^-$, in $\sqrt{\text{Vaidya}}$ obtained from WL formalism:

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- Under the double copy rules, we get the correct out-modes in Vaidya

$$\varphi_{\mathcal{E}'}(x) \rightarrow \frac{1}{r} e^{-i\mathcal{E}'v + 4iGM\mathcal{E}' \log\left(\frac{v_0 - v}{r_0}\right)}. \quad (15)$$

Pair-creation amplitude

- For $\sqrt{\text{Vaidya}}$:

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- Under the double-copy rule

$$|\alpha^{-1}\beta| \rightarrow e^{-4\pi GME'}.$$

This is exactly Hawking's result!



THANK YOU!