Position-Sensitive Detectors Applications in Astro-particle Physics

Robert P. Johnson University of California at Santa Cruz Santa Cruz Institute for Particle Physics

Introduction

Applications of Position-Sensitive-Detectors in Astro-particlephysics is a very broad field, which I could not possibly survey completely!

Most astrophysics detectors are position-sensitive to some degree, from precision cosmic-ray tracking detectors, to position-sensitive calorimeters, to pixelated detectors on photon imaging planes.

I will concentrate on areas and examples most familiar to me:

- Orbiting gamma-ray telescopes
- Orbiting or suborbital cosmic-ray telescopes

And I will touch on some technology transfer to medical physics that has arisen from work in these areas.

Gamma-Ray Pair-Conversion Telescopes







AGILE

Pair-Conversion Telescope Concept



• Tracker:

• Silicon sensors detect the e^+e^- to reveal the γ direction.

• Calorimeter:

- Dense CsI scintillating crystals measure the energy.
- Veto counters:
 - Plastic scintillators detect cosmic-ray background.

Unavoidable random scattering in the material limits the angular resolution of the telescope.

Fermi-LAT Si Tracker Module (1 of 16)

73.8 m² of Si sensors! 9216 wafers, 884,763 channels, 14,976 ASICs

Module Structure Components: Composite Sidewalls: Italy (Plyform) Other parts: SLAC



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PSD12 R.P. Johnson W.B. Atwood et al, Astropart. Phys. 28 (2007) 422.

Fermi-LAT Performance after 10 Years on Orbit

- Fermi began operations on orbit 13 years ago.
- The LAT remains fully operational today, with negligible degradation in detector performance. *Excellent demo of the power of solid-state tech.*
 - See arXiv:2106.12203v2, July 8, 2021
- Of the ≈15,000 ASICs in the tracking system, only one has failed (before launch). Full operation of all channels in that module was maintained because of the left-right readout redundancy.
- The MIP peak (from time-over-threshold) trended upward by ≈11% over 10 years, with most of the change occurring in the first few months.
- Bad channels:
 - 0.31% at launch (concentrated in the first tracker module assembled)
 - 0.46% after 10 years
 - After 10 years, 58% of the 598 masked noisy channels were on 2 of the 2304 silicon-strip ladders
- Radiation damage: visible, but negligible effect on performance. Overall hit efficiency dropped from 99.73% to 99.70% in 10 years.
 - Very high S/N of Si sensors ensures ≈100% hit efficiency with ≈0 noise!



Computer Memory Upsets and the SAA

The radiation environment is relatively benign in low-Earth orbit, but some care has to be taken for single-event upsets (probably from heavy ions).



Compton Gamma-ray Telescopes









Compton Spectrometer & Imager (COSI)



3D imaging of Compton scatters within cold germanium detectors:

- Double-sided strips at 90degree stereo
- Time difference of electron and hole collection. 9/15/2021

Two successive Compton scatte give information needed to reconstruct a circle on the sky. **Redundant information from** scattering energies allows the interaction order to be inferred.

Selected in 2020 for a SMEX Phase-A study.

400

200

600

PI: John Tomsick, UCB Space Sciences Lab

150

100

50 F

Compact Compton Telescope Concept



- 1.5 m lever arm
- PMT readout with fast timing, to measure TOF.
- Positions within modules derived from relative PMT signals.



- Compact: photon scatterings and absorption can be confined to a single Ge detector.
- No TOF. Redundant kinematic information is used to determine interaction ordering.

COSI-APRA Sub-Orbital Payload (arXiv:1701.05558)



Two Galactic point sources, an AGN, and one Gamma-Ray-Burst (GRB) were observed.





Medium Energy Gamma-ray Astronomy (MEGA) 0.4 to 50 MeV arXiv:astro-ph/0110129





- Above ≈ 1.5 MeV the scattered electron can be tracked.
- Above ≈ 8 MeV pair conversion starts to dominate.
 - With no heavy metal converter in the tracker, the lowenergy conversions are more useful than in Fermi-LAT
- A large observatory using this technology, eAstrogam, was proposed (arXiv:1711.01265), for 0.15 to 3000 MeV.



All-Sky Medium Energy Gamma-ray Observatory

- The CsI calorimeter and plastic-scintillator anti-coincidence shield are similar to MEGA & eAstrogam.
 - Horizontal CsI "logs" (*a la* Fermi-LAT) instead of vertical pixels
- The tracking detector is called "AstroPix", based on the ATLASPix monolithic silicon detector technology (HV CMOS)
 - Square pixels 1 to 2 mm in size
 - 0.5 mm thick
- Low pixel capacitance, compared to long double-sided Si strips, could allow much lower thresholds for detection of Compton scattering, and improved energy resolution.
 - 25 keV trigger threshold
 - 5% resolution at 100 keV



Compton angular resolution is about 5° at 1 MeV. Pair-conversion resolution is about 5° at 40 MeV.

See also "The CMOS pixel sensors particle tracker for the CSES-02 space experiment", S. Beole, in this session.

Technology Transfer to Proton (& He) CT

- Make 3D images of proton stopping power for planning proton cancer therapy.
- Measures directly proton energy loss, instead of inferring it from X-ray images.
- But, more difficult and expensive than X-ray CT.
- Track protons entering and exiting the phantom.
 - Each tracker has 4 layers of Fermi-LAT single-sided SSD.
 - Front-end readout by a new, custom VLSI ASIC.
- Measure the residual proton energy by stopping them in plastic scintillators.
- *R. Johnson et al, IEEE Trans Nucl Sci 63, 2016, p 5.* 9/15/2021



PSD12 R.P. Johnson

UCSC/Loma-Linda pCT Scanner



See also the posters "Bergen proton-CT project", V. Eikeland, 9/16/21 at 11:00 am, and "Design and integration of CMOS tracker layers in digital tracking calorimeter for pCT application", S. Mehednale, 9/17/21 at 8:18 am.



- About 360 million proton measurements make one 3D image.
 - 6 minutes of acquisition, at 1 million protons per second.
- This device has been used by several groups in the USA and Germany to study extensively this imaging modality, including imaging with He ions in Heidelberg, Germany.

Proton-VDA Scanner



- Start-up company funded by US government SBIR funds.
- Only existing effort that I know of that is currently targeting clinical applications.
 - Scintillating-fiber tracking.
 - Thin monolithic calorimeter read out by 16 PMTs.
 - Adjust beam energy according to the phantom thickness.



https://protonvda.com/



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Technology Transfer to Prompt-Gamma Imaging

- *Real-time* monitoring of the Bragg-peak location during proton cancer therapy.
- MeV photons are emitted instantaneously from the Bragg peak and exit the patient with little interaction
- Many groups have worked on imaging the photons using Compton cameras.

A few different examples:



- Double-sided Si with LaBr₃ crystals: *S. Aldawood et al, Rad. Phys. & Chem. 140, 2017, p. 190.*
- CZT strip detectors plus LSO crystal: F Hueso-González et al, JINST 9, 2014, P05002.
- Multiple layers of commercial CZT detectors: J. Polf et al, Phys. Med. Biol. 60, 2015, p. 7085.
- Scintillating fibers of heavy material (e.g. LYSO:Ce) read by SiPM: A. Wronska et al, Acta Phys. Pol. B 51, 2020, p. 17.

I have not seen any use of Ge strip detectors such as in COSI.

No efforts have yet reached clinical, or pre-clinical, use. A typical problem is inadequate efficiency. Expense and complexity are other issues.

Cosmic-Ray Telescopes









AMS Tracking Detector

- 9 planes of *double-sided* SSD supported by carbon-fiber structures.
- Alignment precision of about 6 microns.
- Can measure the cosmic-ray rigidity up to 2 TeV.



10⁴

400 GeV/c Monte Carlo Simulation

CALorimetric Electron Telescope (CALET)

- Similar to the Fermi-LAT but
 - Thicker by about ×3, for very high energy electrons and photons
 - Scintillating fibers instead of silicon-strips
 - Much smaller area (about 1/20)
 - Lots of ISS material in parts of the field of view
- Cannot compete very much with the Fermi-LAT on photon science, but electron cosmicray measurements extend to higher energy

Astropart. Phys. (2018) 10029



AESOP-Lite High Altitude Balloon Experiment



- 8 Si tracking planes:
 - Fermi-LAT SSDs.
 - ASICs from pCT project.
- Permanent magnet (≈0.3 T)
- Cherenkov detector
- Scintillators, with TOF

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- Measure the electron and positron CR fluxes at low energy (down to 20 MeV).
- Flew from northern Sweden to Ellesmere Island, Canada in 2018.
- Second flight, with detector upgrade and a larger (higher) balloon in Antarctica, 2022.



Conclusions

- Position sensitive detectors are widely used in space-based particle detection and tracking.
 - Silicon strips are especially by now an old, standard, wellproven technology.
- One of the biggest contemporary challenges is a sensitive Compton telescope to explore the MeV universe.
 - Many proposals with Ge and Si, plus now pixel detectors.
- The same technologies have found their way into medical physics applications (see Reinhard Schulte's talk).
 - Proton Computed Tomography and Radiology are poised to begin an entry into clinical use.
 - Prompt gamma imaging could have an even greater medical impact, but there the challenges posed by Compton telescopes are again a big factor.