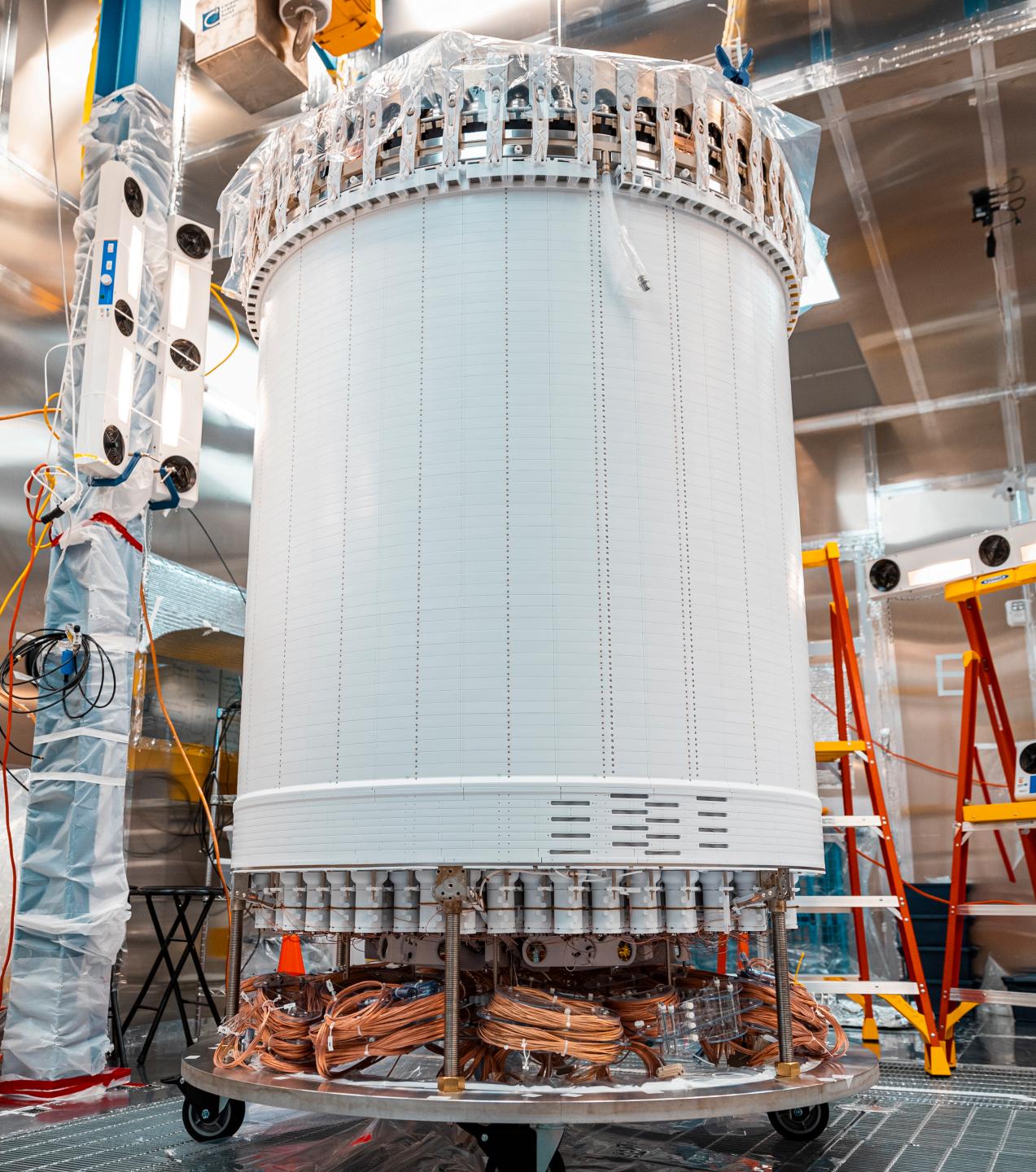


Status of the LUX-ZEPLIN Experiment



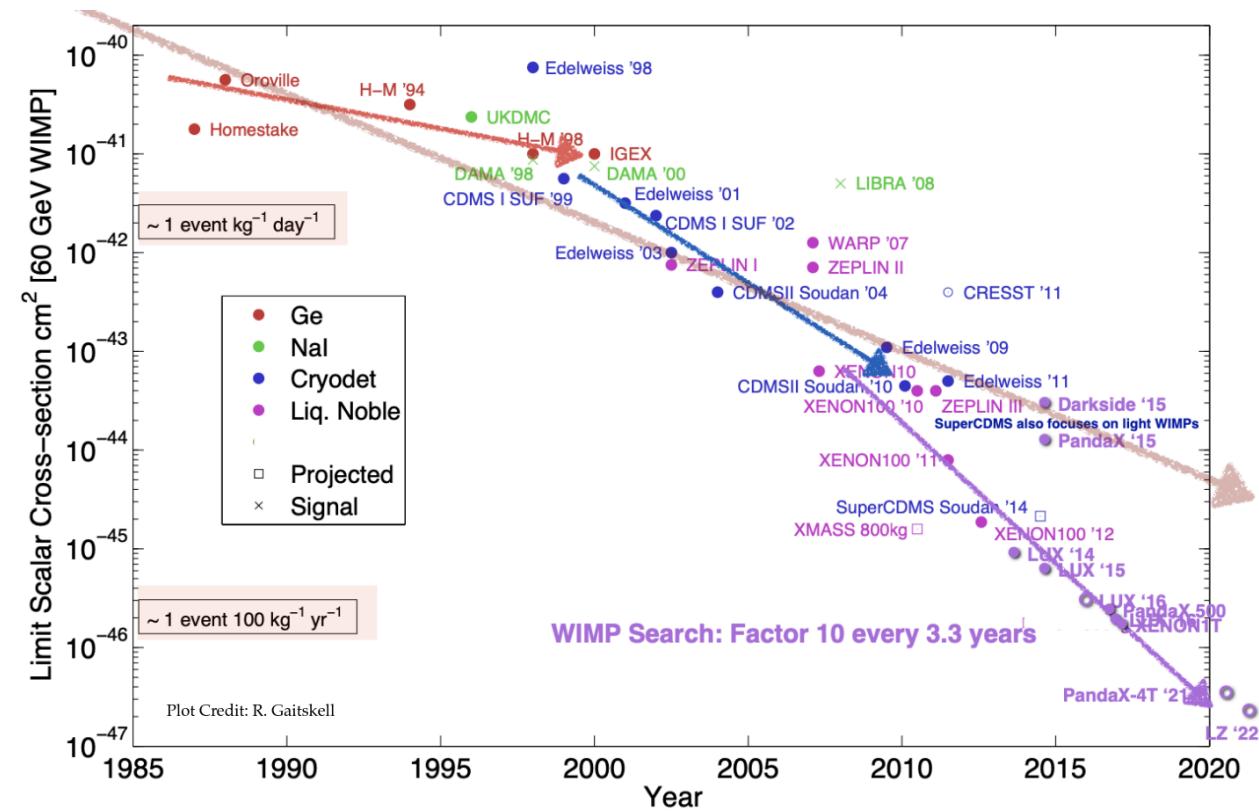
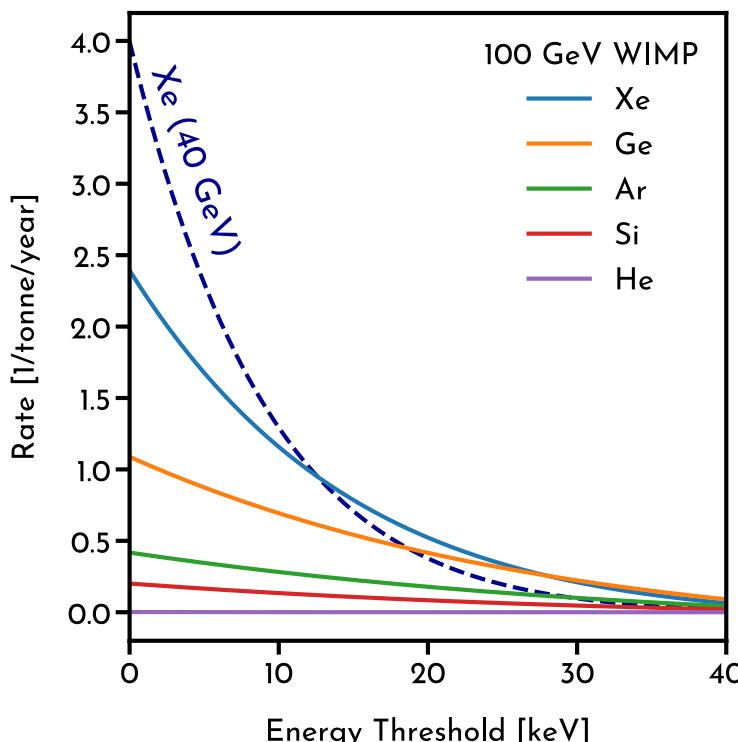
Albert Baker
on behalf of the LZ collaboration

DMUK
7th January 2025



Xenon detectors for dark matter searches

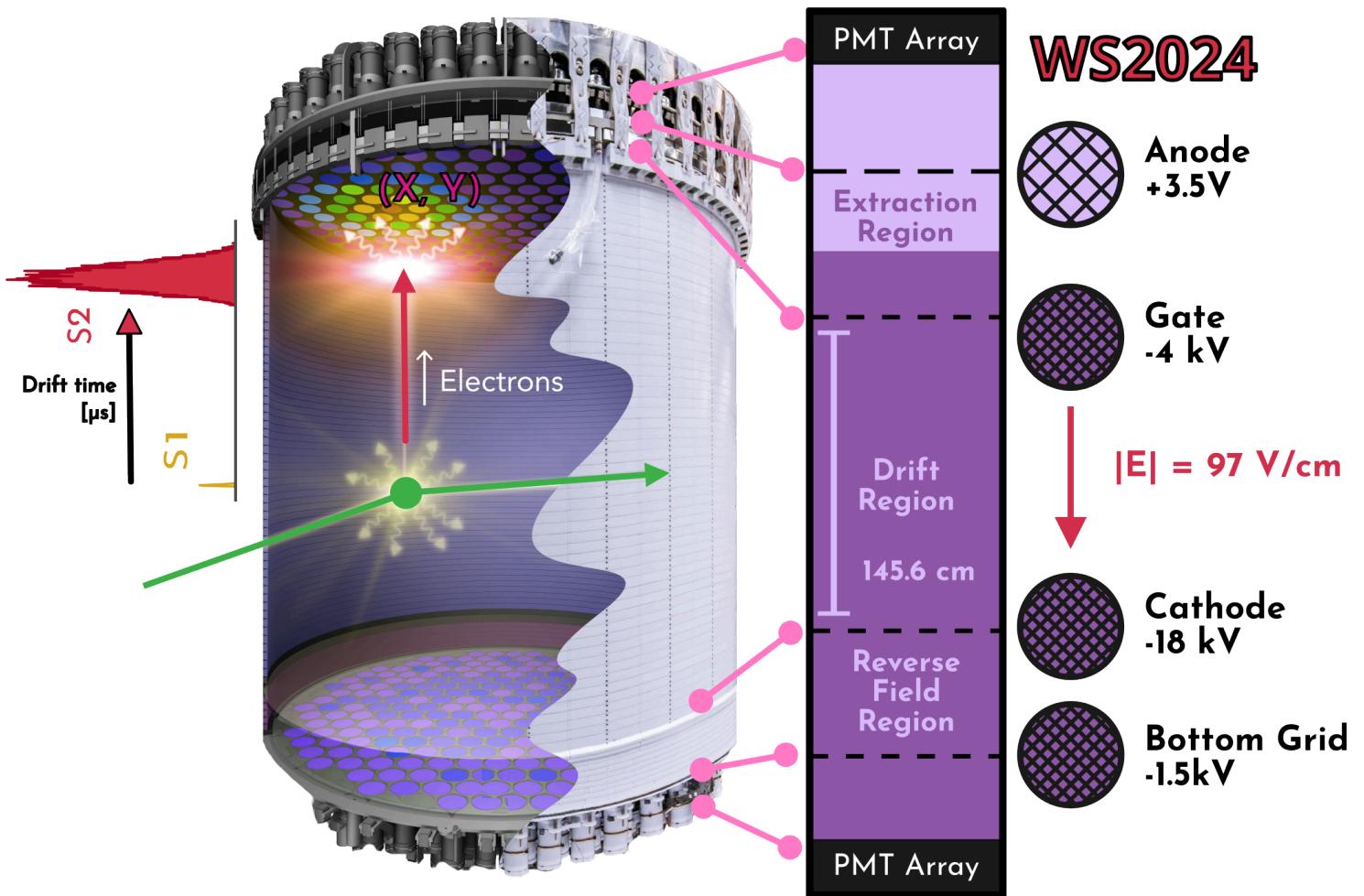
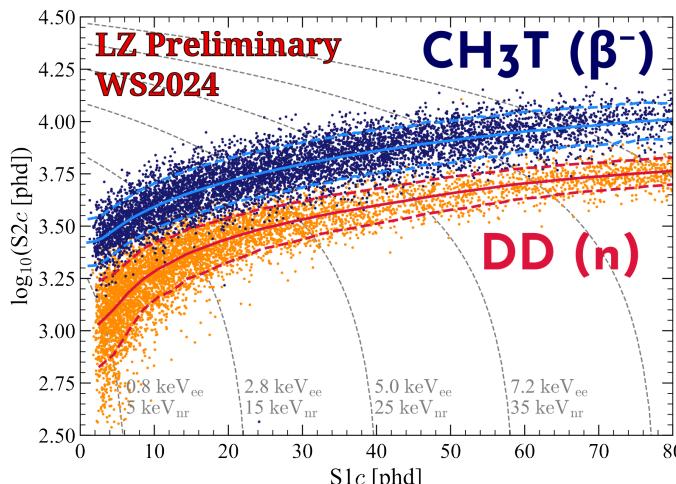
- Large nucleus
 - Significant coherent nuclear scaling ($\sigma \sim A^2$)
- Significant self shielding from high liquid density ($\sim 3 \text{ g/cm}^3$)
- Noble gasses are easy to purify
 - Dedicated processes for Rn and Kr removal

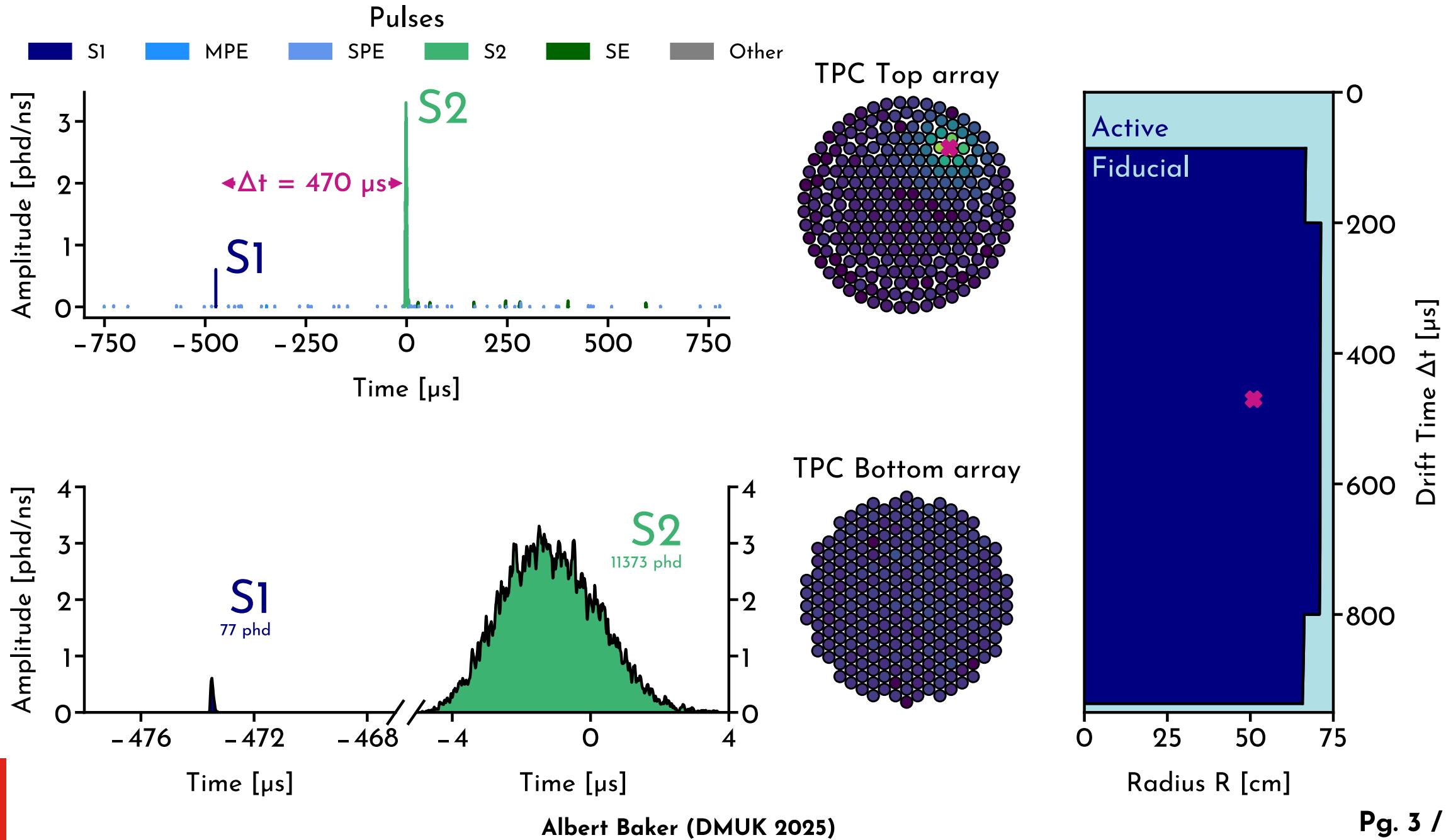


- Proven track record for ~ 2 decades
- Scalable technology
 - kg to multi-tonne scale

Two-phase Time Projection Chambers (TPCs)

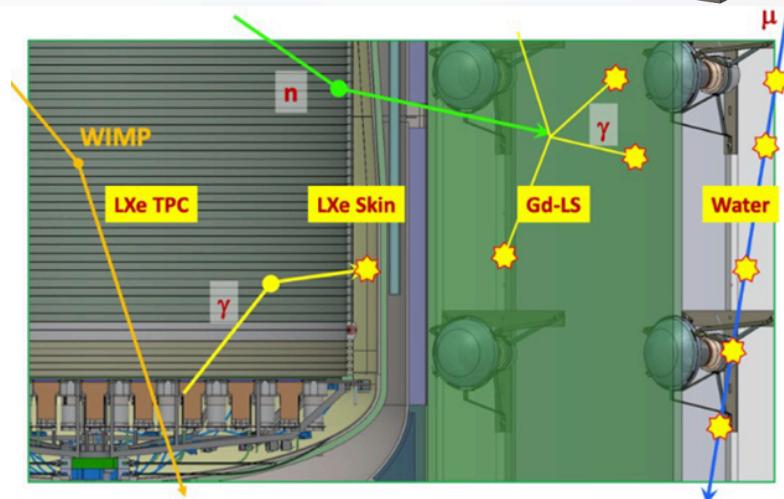
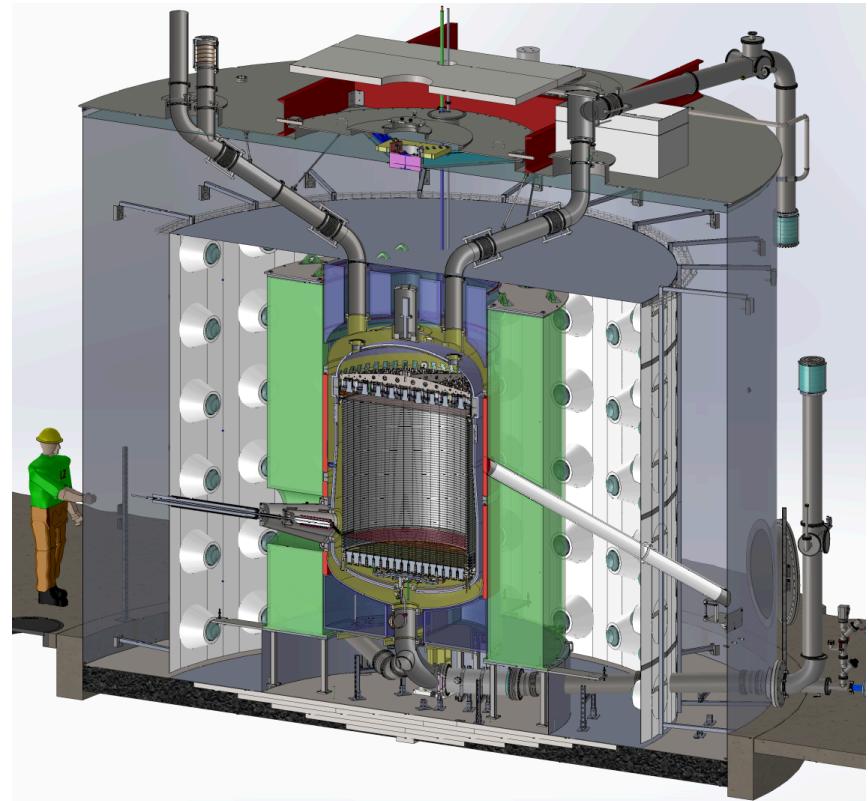
- Liquid target with thin vapour layer
- Time resolved scintillation and ionisation signals
 - Light = prompt scintillation (**S1**)
 - Charge = delayed electroluminescence of ionisation electrons (**S2**)
- 3D vertex reconstruction:
 - (X, Y) from S2 hit pattern
 - Z from the electrons drift time
- Discriminate electron (ER) and nuclear (NR) recoils using S2/S1





Veto detectors

- Outer detector (OD):
 - Gd-loaded liquid scintillator
 - Detects neutrons via γ -rays from neutron capture
- Skin xenon veto:
 - Instrumented xenon outside the TPC
 - Detects primary γ -rays from components/target
- Vital to measure and constrain neutron backgrounds
 - Achieved high neutron veto efficiency in WS2024
 - $89 \pm 3\%$ derived from AmLi calibrations
 - **$92 \pm 4\%$** from background neutron simulations





TPC Underground 2019

Installation complete 2020

Science data from 2021

World leading WIMP limits 2022

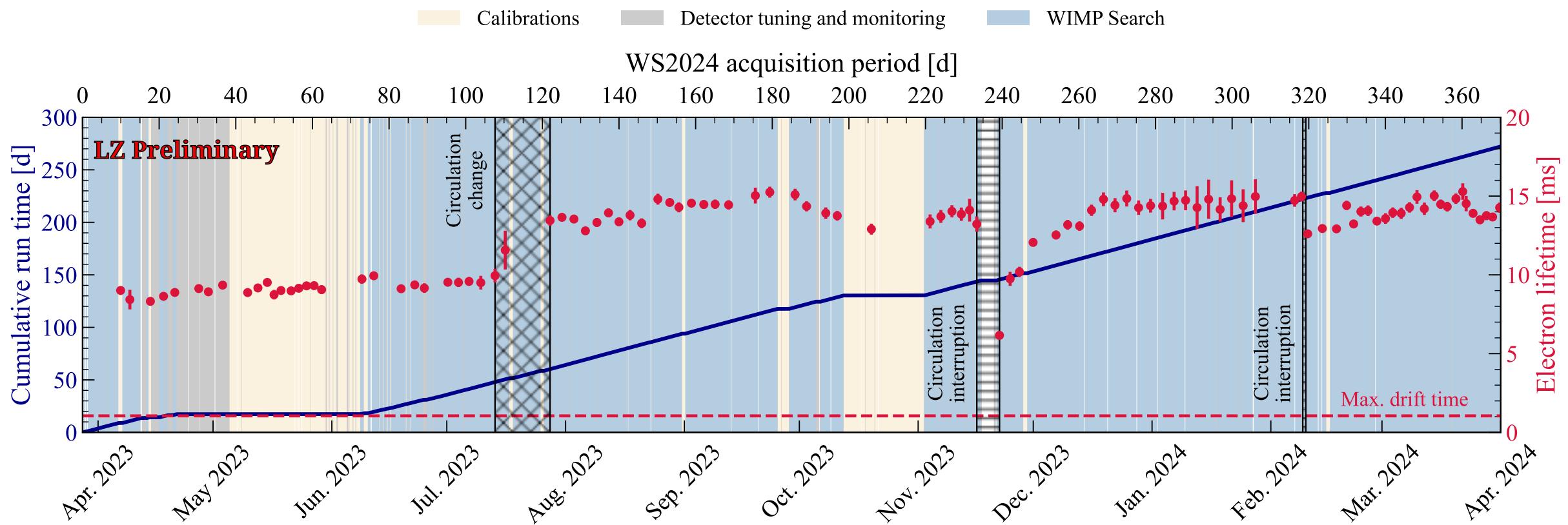


2024 Science Run (WS2024)

WIMP Search 2024 paper arXiv:2410.17036

- Acquired data for ~ 370 days with 95.2% detector up-time
 - **220 day exposure after data quality cuts**

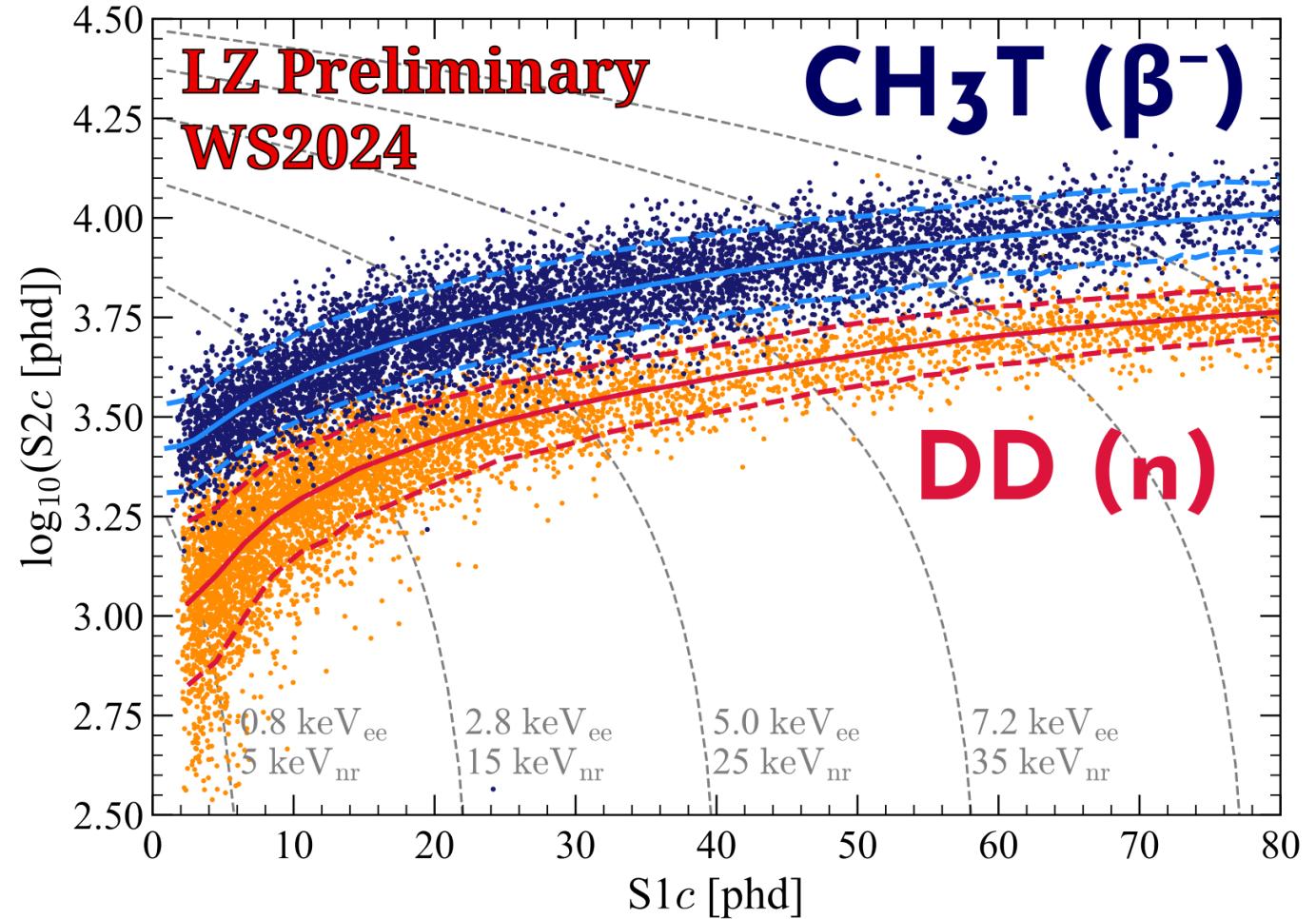
- Performed intermittent calibrations
- High target purity throughout
 - Minimal suppression of charge (S2) signals



Calibrations in WS2024

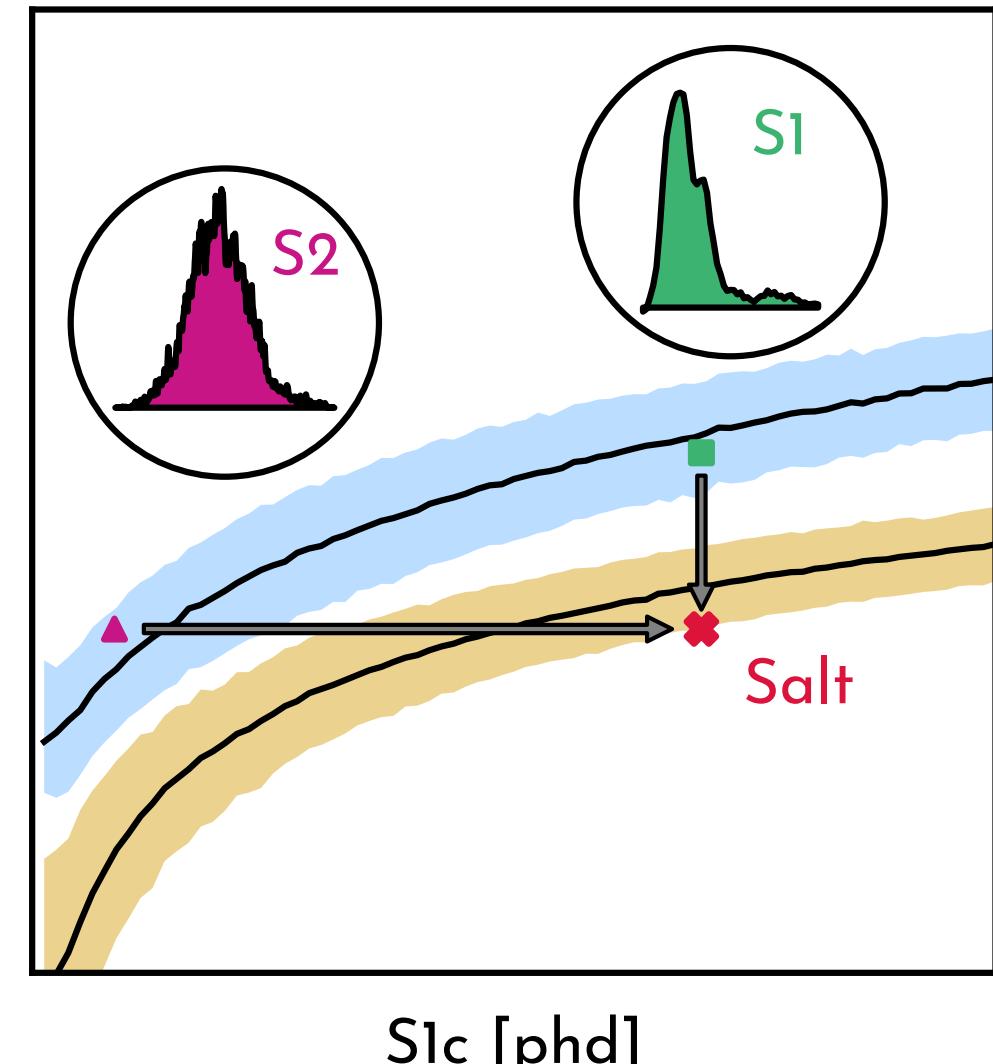
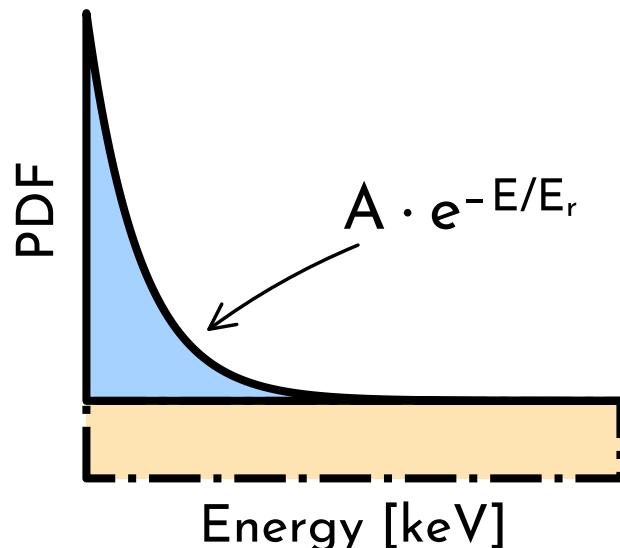
LZ calibration systems 2024 JINST 19 P08027

- Electronic recoils (background)
 - Tritium radio-labelled methane & ^{14}C
 - Mono-energetic $^{83\text{m}}\text{Kr}$
- Nuclear recoils (signal)
 - DD neutron generator (2.45 MeV neutrons)
 - An AmLi source, which emits low energy (<1.5 MeV) neutrons, can be positioned at nine different depths
- WS2024 NEST model:
 - Light gain (\mathbf{g}_1): $0.112 \pm 0.002 \text{ phd/photon}$
 - Charge gain (\mathbf{g}_2): $34.0 \pm 0.9 \text{ phd/electron}$
 - Single electron amplification: 44.5 phd
 - **99.9% discrimination** of β^- below 40 GeV WIMP median



Bias Mitigation (Salting)

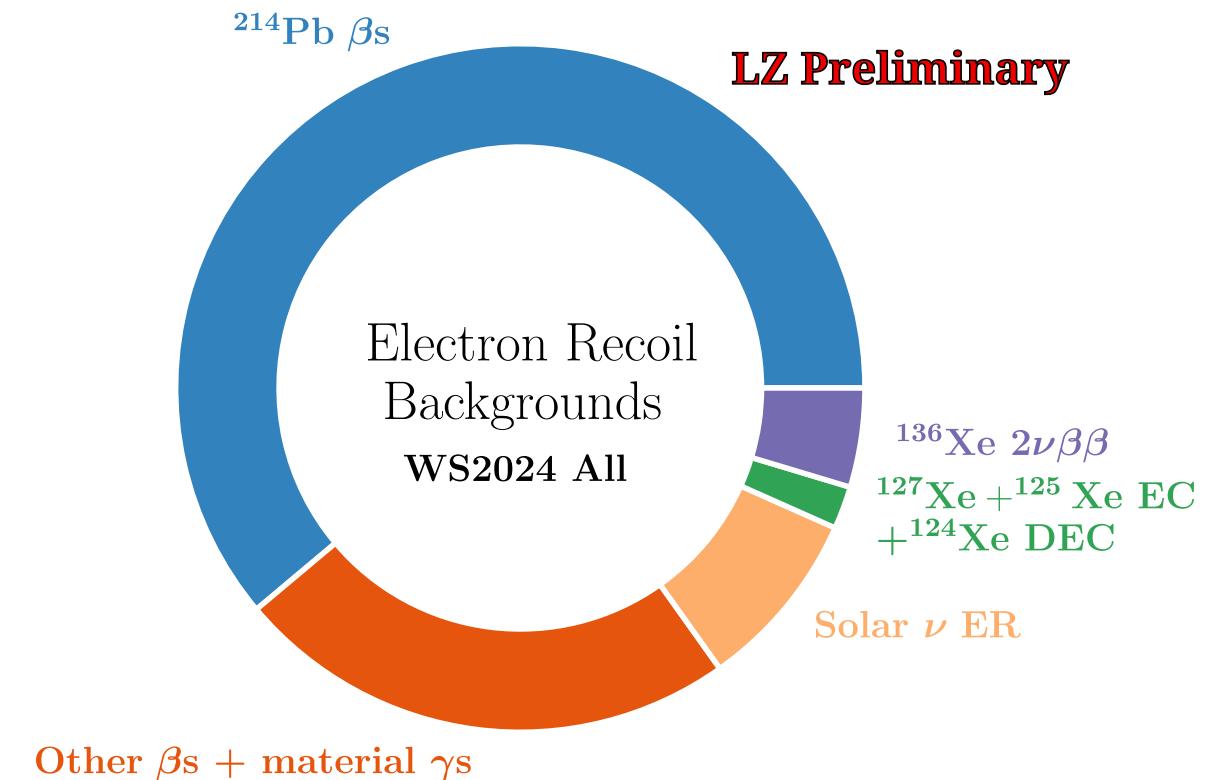
- Inject fake events (salt) into the data stream
- Generated by pairing S1 and S2 pulses from calibration
 - Embed measured waveforms back into the data stream
- Events sampled as follows:
 - Unknown rate below LZ's WS2022 result
 - Recoil spectra of a WIMP of unknown mass
 - Additional contribution for high mass WIMP searches with flat NR spectrum
- Allows us to understand the ROI whilst minimising bias



WS2024 Background Model

- Understand backgrounds from in-situ measurement of sidebands and assays
- Expect 1207 ER background events in WS2024:
 - ^{214}Pb β -decay is dominant at 60%
 - Double electron capture (DEC)
 - Solar neutrinos
- Expect 0.18 NR CEvNS events
 - Excluded by region of interest for dedicated search
- Neutrons from spontaneous fission in detector components and (α, n) reactions
- Accidental backgrounds from isolated S1 and S2 pulses

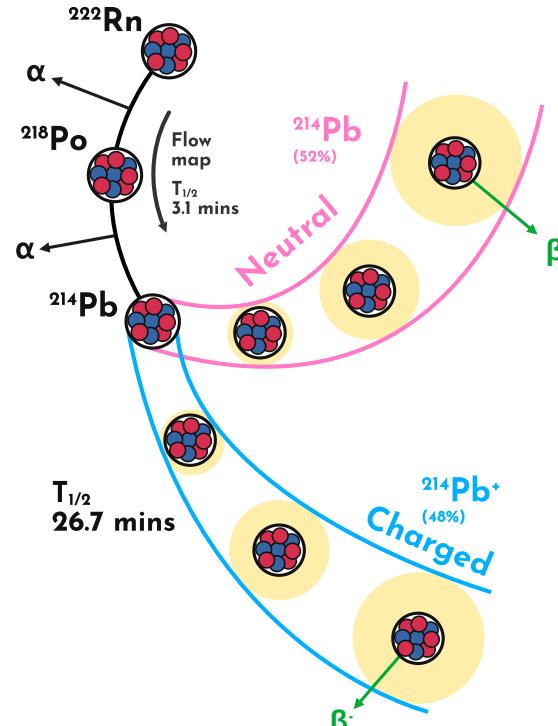
WS2022 background model Phys. Rev. D 108, 012010



Radon tagging

See Simran's talk 5:00 PM

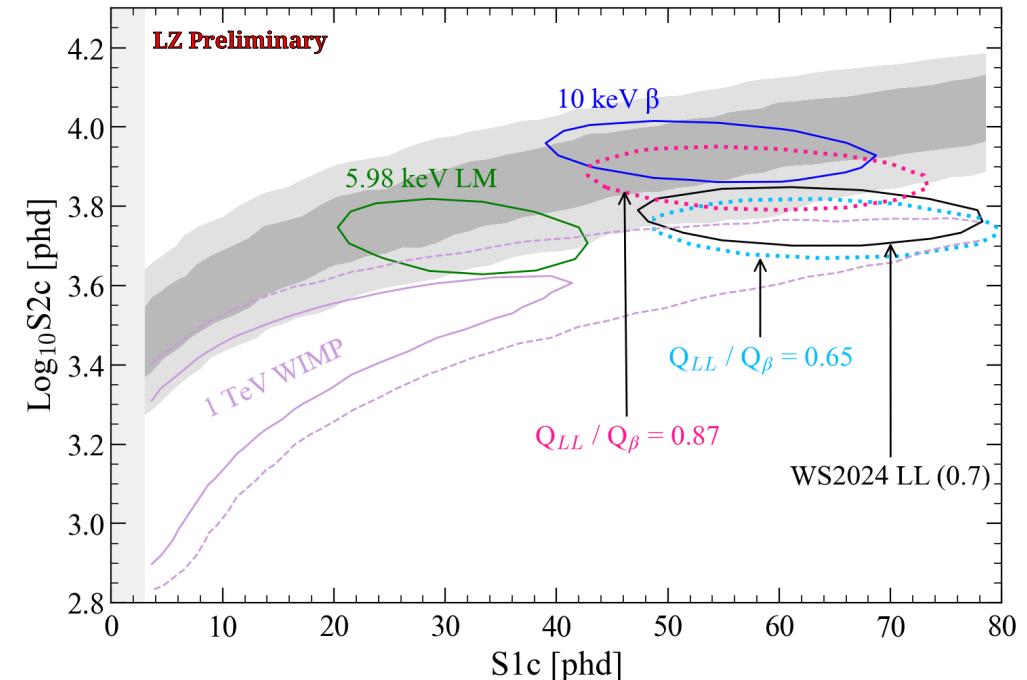
- Predict future ^{214}Pb decays
 - Observe ^{218}Po α -decay
 - Tag interactions around Xe streamlines
 - Track for 81 minutes ($\sim 3 \times ^{214}\text{Pb } T_{1/2}$)
 - Tag incorporated into statistical analysis
 - **~60% tagging efficiency**



Electron Captures (EC)

See Olivia's talk 4:45 PM

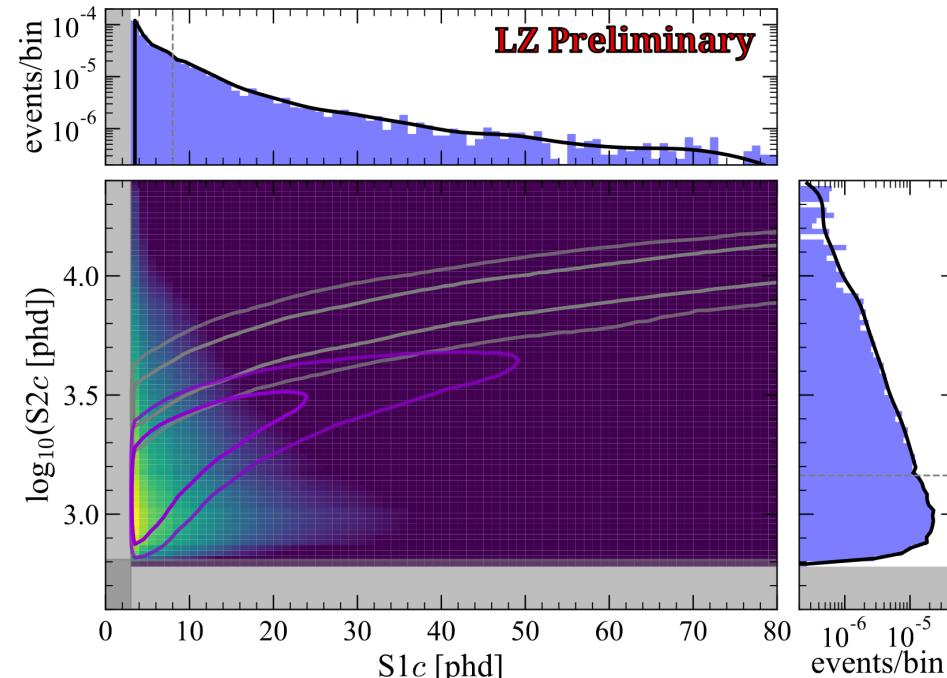
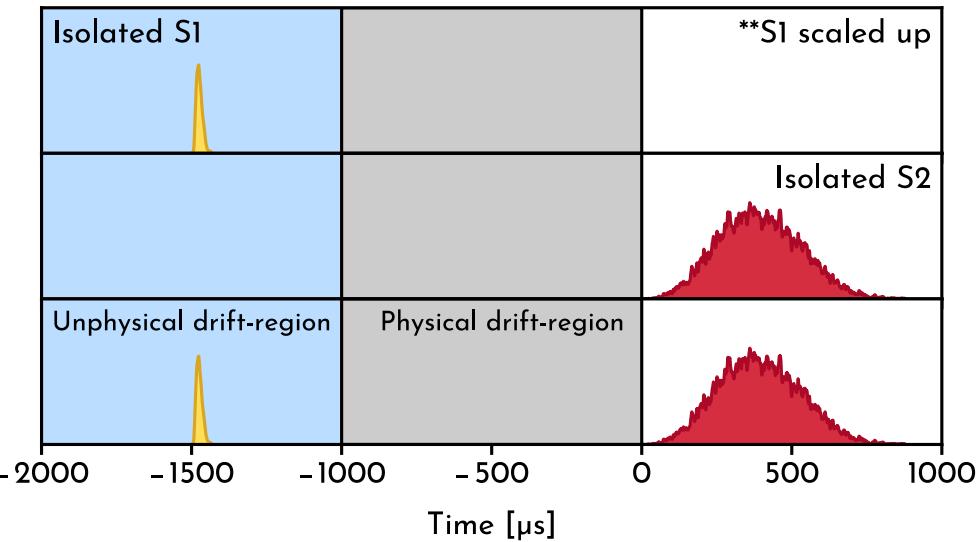
- Background in LZ (5.2 keV L-shell):
 - Single EC: $^{125}/^{127}\text{Xe}$ - from neutron activation
 - Double EC: ^{124}Xe - $T_{1/2} \sim 10^{22} \text{ years!}$ [1,2]
- EC suffers from charge suppression [3]
 - **Looks more NR like than normal**



- [1] XENONnT measurement Nature 568, 532–535 (2019)
[2] LZ measurement J. Phys. G: Nucl. Part. Phys. 52 015103
[3] XELDA measurement Phys. Rev. D 104, 112001 (2021)

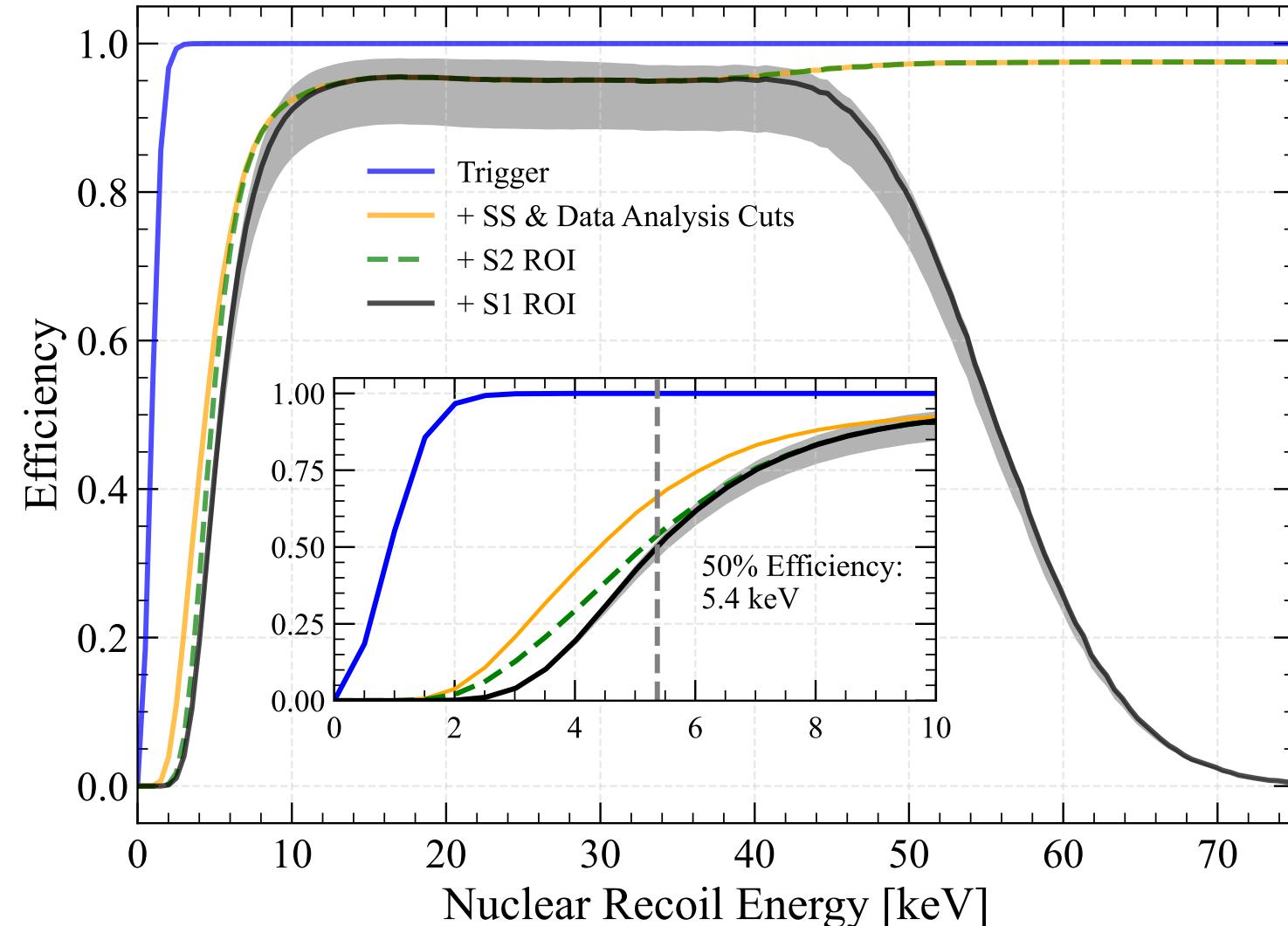
Accidental background

- Pile-up of unrelated S1-like and S2-like pulses
 - Looks like a single scatter and can mimic a WIMP
- Fraction of these have an unphysical drift time
 - Population to calculate rate with physical drift-time
- Model as product of isolated S1-like and S2-like pulses
- Distribution peaks in the low energy NR region
- Analysis cuts specifically tested on and tuned for this background
 - 99.5% rejection efficiency
 - Expect 2.8 ± 0.6 events in WS2024



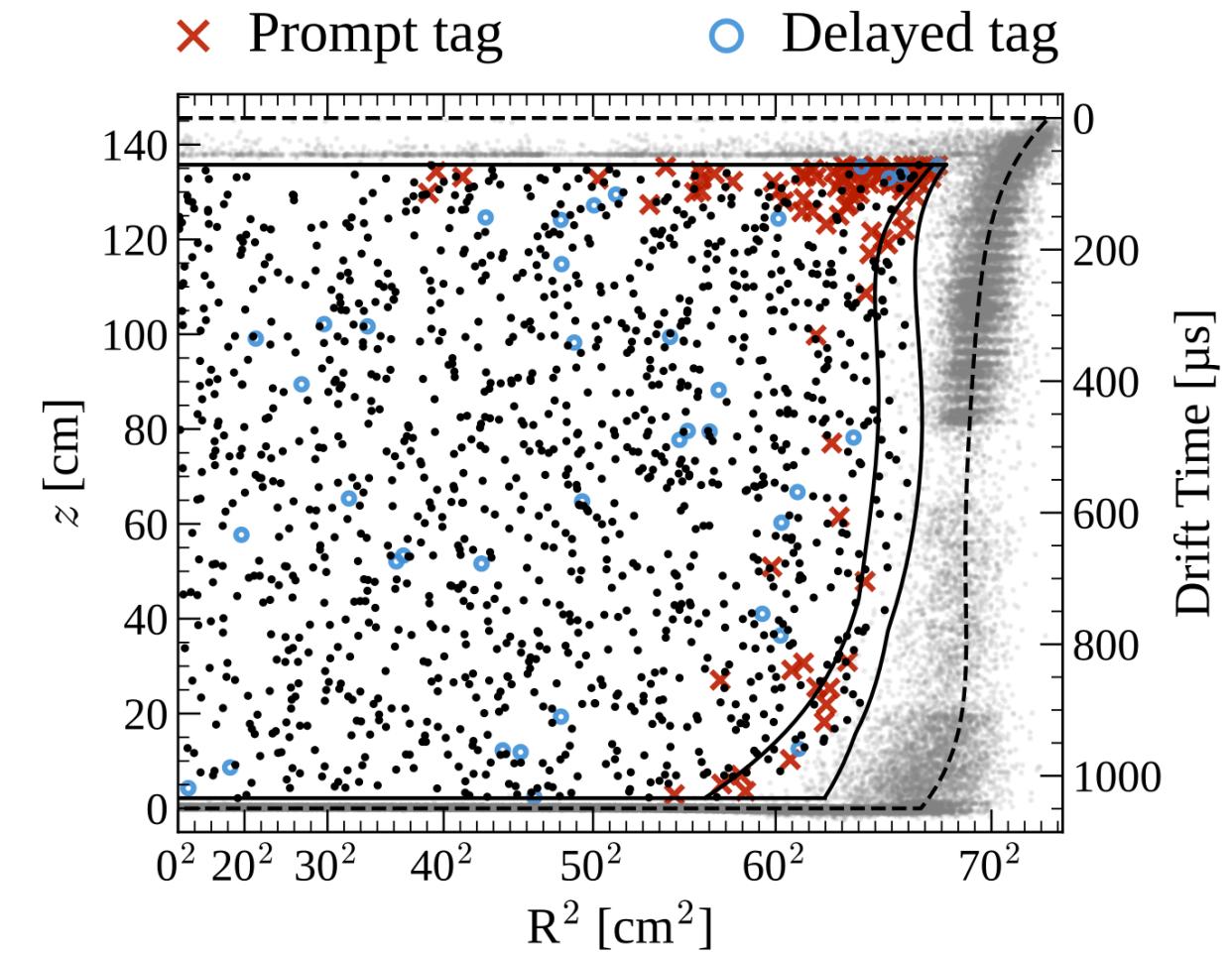
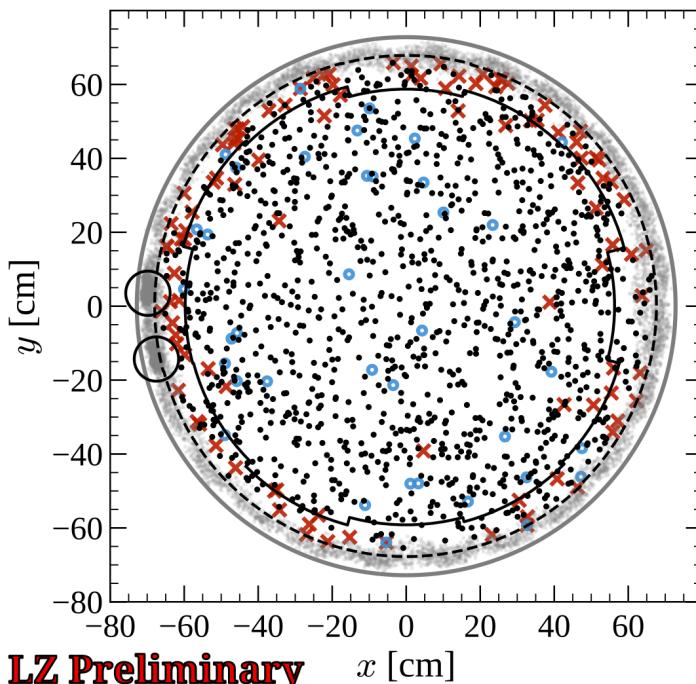
Signal Acceptance

- Region of interest (ROI):
 - $S1_c = (3, 80)$ phd
 - $S2_c = (645, 10^{4.5})$ phd
 - Excludes ${}^8\text{B}$ for dedicated analysis
- Multiple event and pulse level cuts
 - FV, ROI, single scatter
 - Veto anti-coincidence
 - Delayed neutron capture
 - Prompt γ -ray interactions
 - $S1$ & $S2$ based cuts targeting accidentals
- Cuts developed using data outside ROI
- 50% efficiency at 5.4 keV



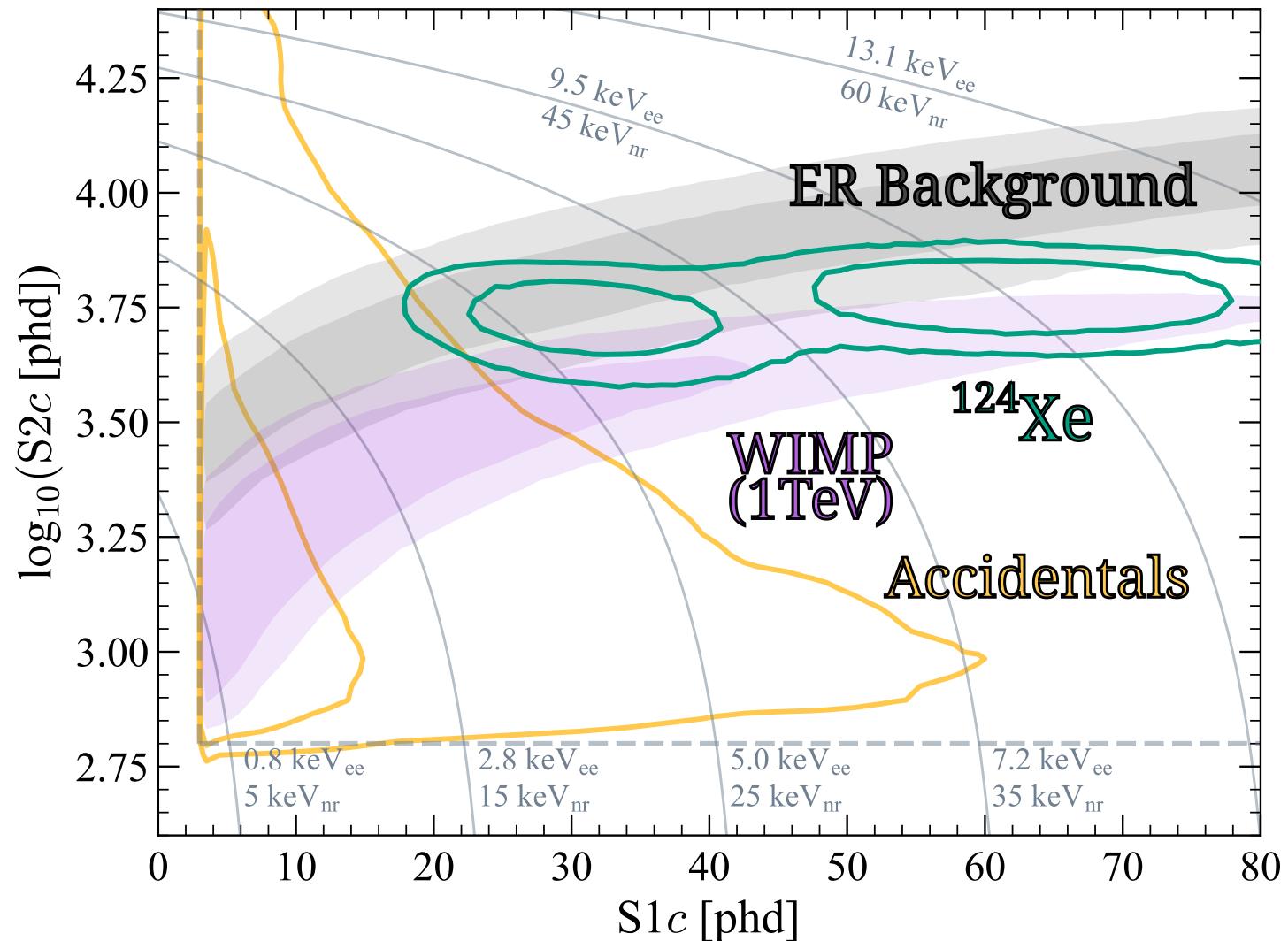
Event positions, fiducial volume (FV), and vetos

- Majority of backgrounds are peripheral
 - Self-shielding prevents infiltration
- Reject the majority of backgrounds with a fiducial cut
 - Azimuthal dependence added for WS2024
 - Full range shown by two solid lines in $z-R^2$ plane
 - Defined to admit **<0.01 wall background events**
- Fiducial mass of **5.5 ± 0.2 tonnes**



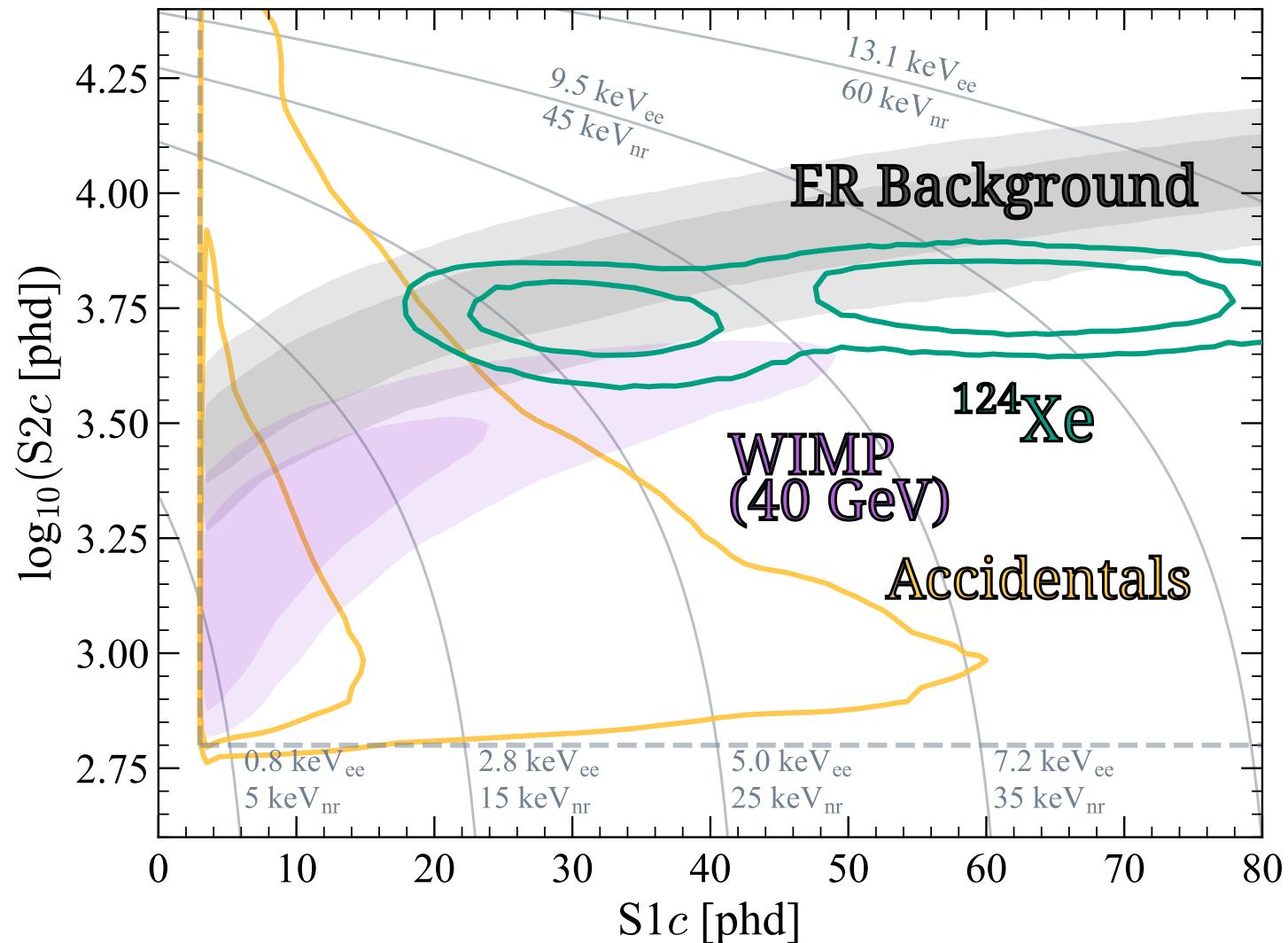
Analysis of WS2024

- Statistical analysis in $\log_{10}(S_{2c})$ - S_{1c} space
 - $S_{1c} = (3, 80)$ phd
 - $S_{2c} = (645, 10^{4.5})$ phd
 - Excludes 8B for dedicated analysis
- Generate templates of each fit component using simulations
 - In-situ measurements & assays provide rate priors
 - Find the best fit of each component for several WIMP masses
- WIMP template (PDF) has a longer tail for larger masses
 - They all peak at low energies



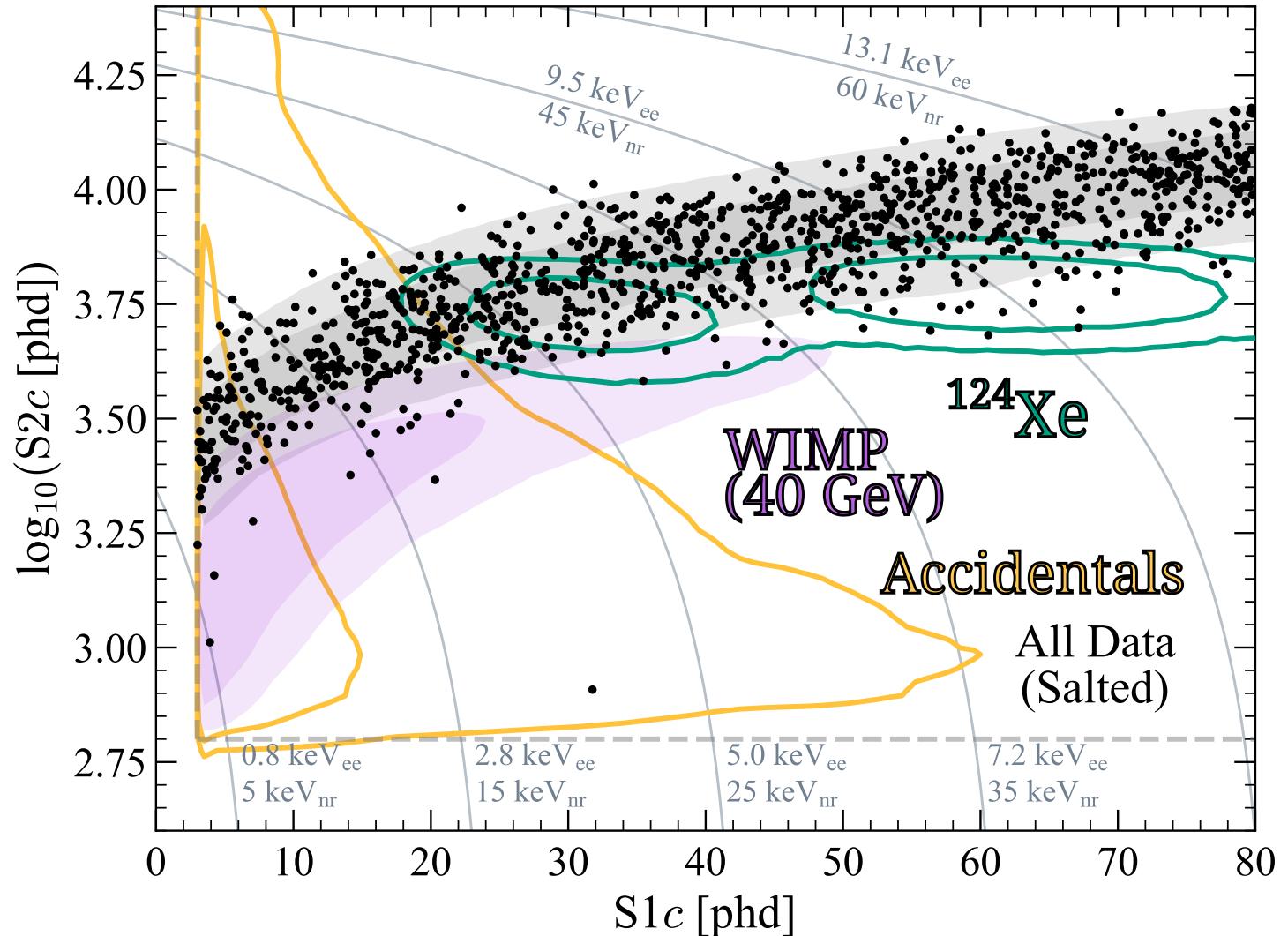
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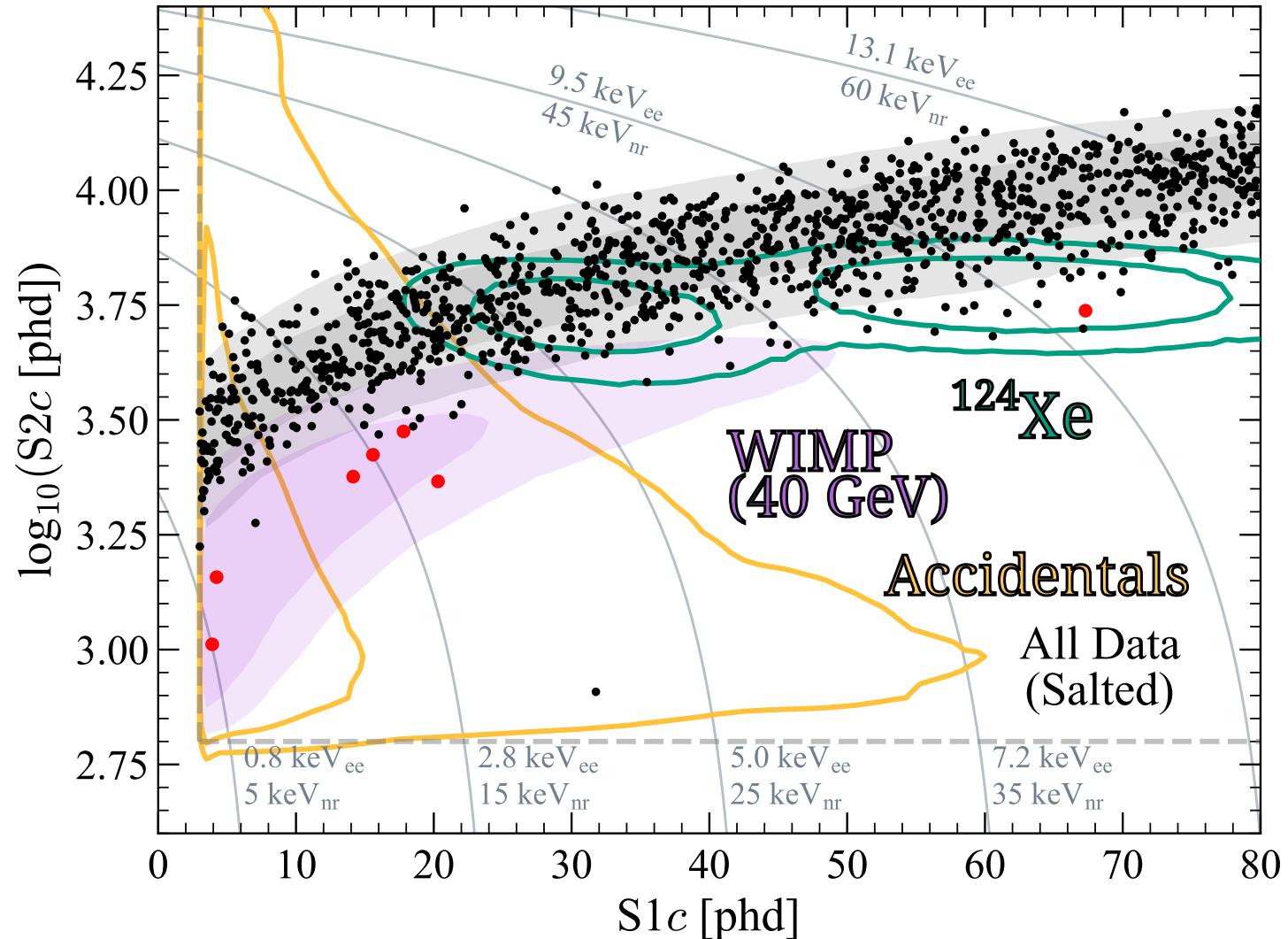
Salted WS2024 result

- Statistical analysis in $\log_{10}(S_{2c})$ - S_{1c} space
 - $S_{1c} = (3, 80)$ phd
 - $S_{2c} = (645, 10^{4.5})$ phd
 - Excludes 8B for dedicated analysis
- Injected artificial events (salt) from calibration for bias mitigation
- Exposure of 220 days \times 5.5 tonnes:
 - 3.3 tonne-years



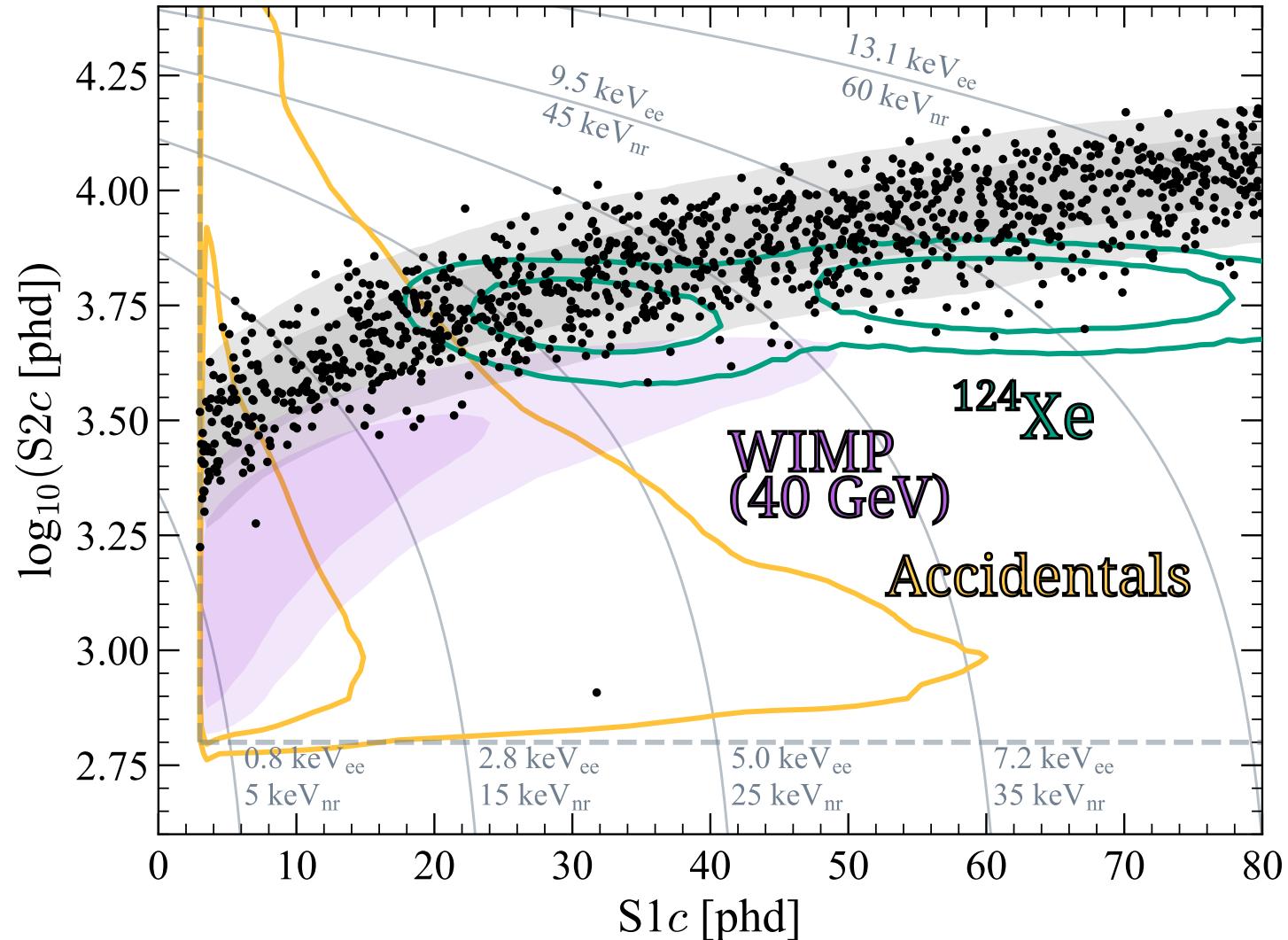
Unsalted WS2024 result

- Statistical analysis in $\log_{10}(S_{2c})$ - S_{1c} space
 - $S_{1c} = (3, 80)$ phd
 - $S_{2c} = (645, 10^{4.5})$ phd
 - Excludes 8B for dedicated analysis
- Injected artificial events (salt) from calibration for bias mitigation
- Exposure of 220 days \times 5.5 tonnes:
 - 3.3 tonne years
- 8 salt events injected
 - 1 was removed by cuts
 - This is consistent with the signal efficiency



Unsalted WS2024 result

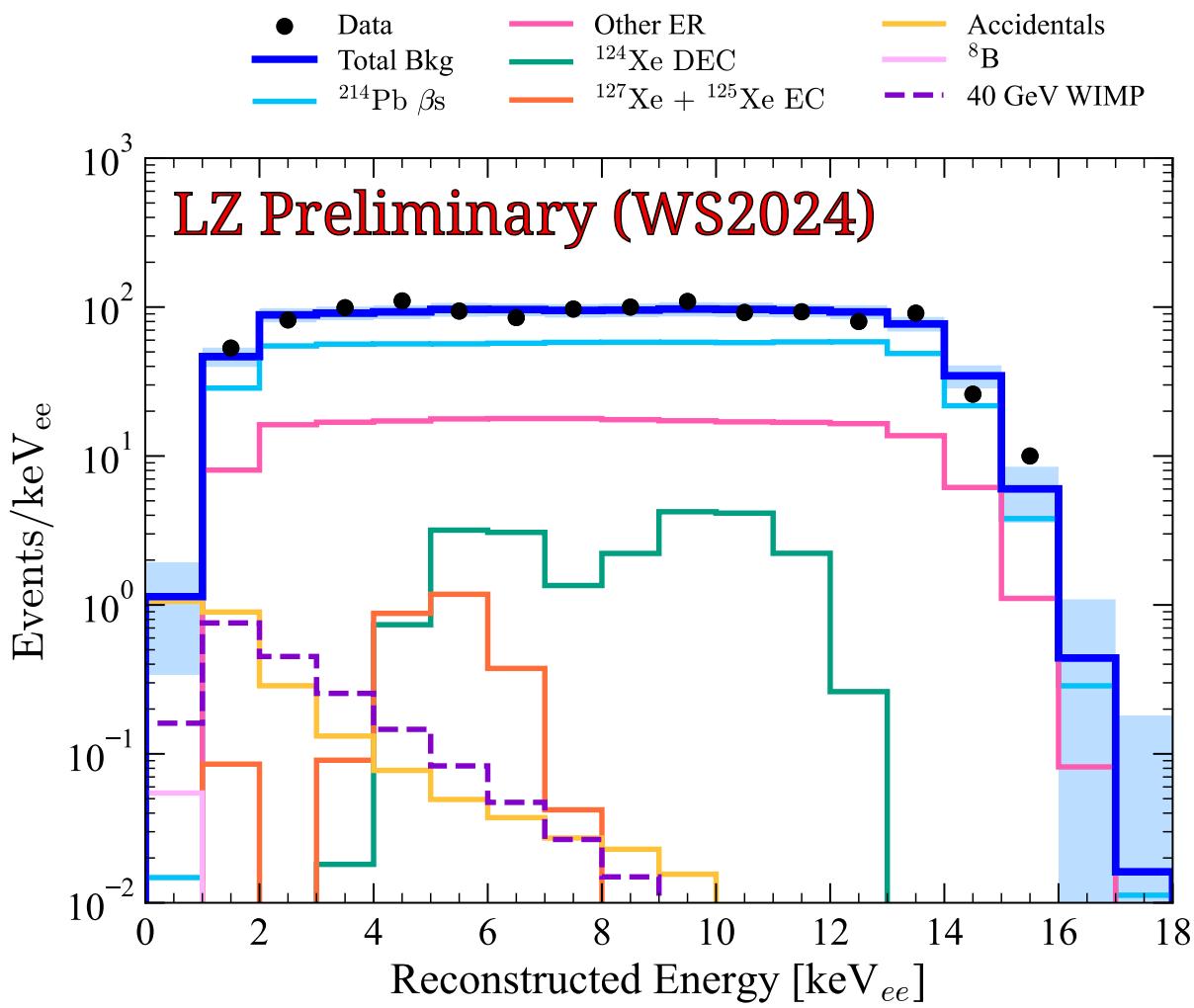
- Statistical analysis in $\log_{10}(S2_c)$ - $S1_c$ space
 - $S1_c = (3, 80)$ phd
 - $S2_c = (645, 10^{4.5})$ phd
 - Excludes 8B for dedicated analysis
- Injected artificial events (salt) from calibration for bias mitigation
- Exposure of 220 days \times 5.5 tonnes:
 - 3.3 tonne years
- 8 salt events injected
 - 1 was removed by cuts
 - This is consistent with the signal efficiency
- 1220 events remain after un-salting
- **No changes to analysis post un-salting**



Fit results in WS2024

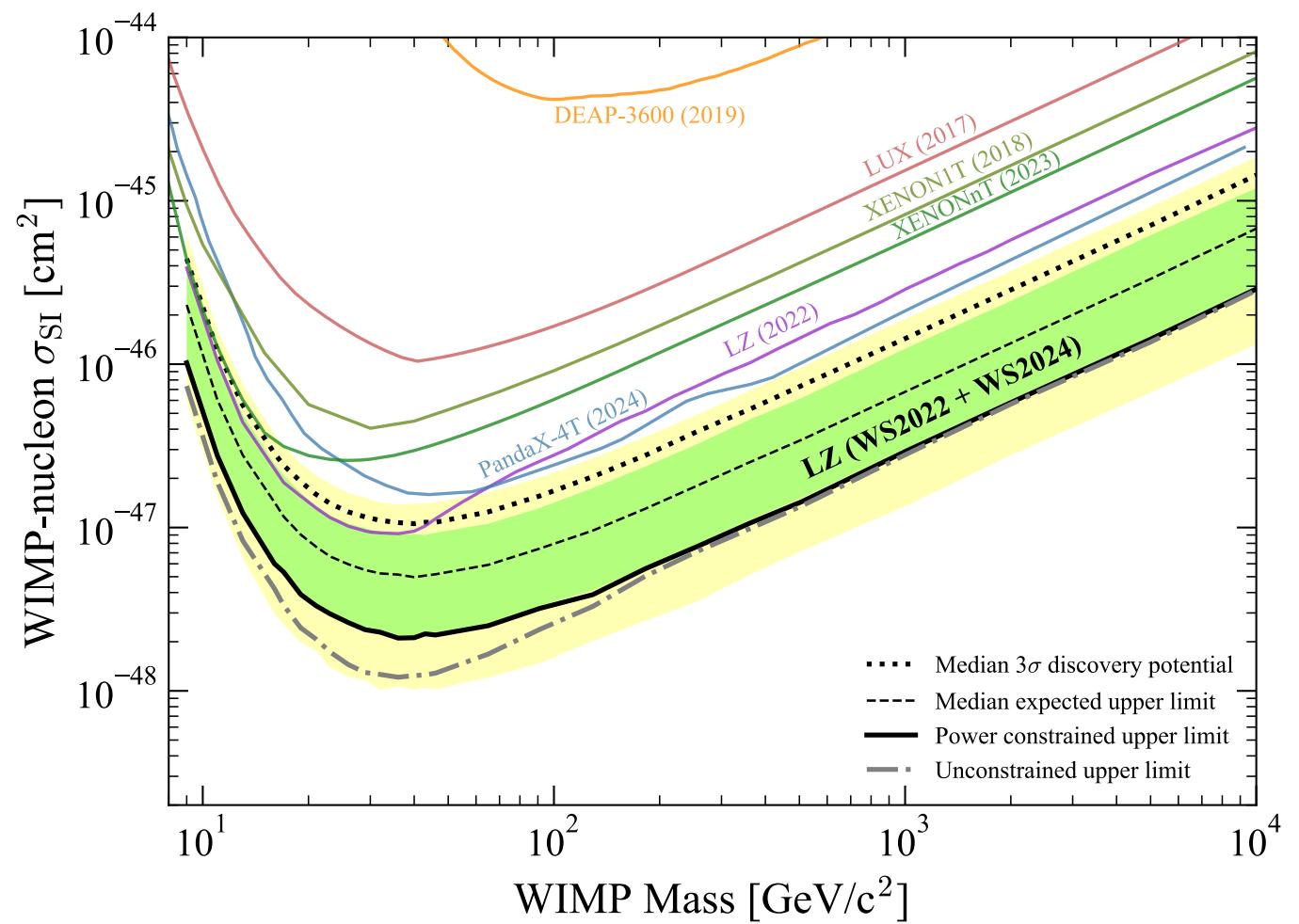
- Best fit of zero WIMPs at all masses ($9 \text{ GeV} \rightarrow 100 \text{ TeV}$)
- Good agreement with background in all studied spaces

Component	Prior	Best fit
^{214}Pb β -decays	743 ± 88	733 ± 34
^{85}Kr & ^{39}Ar & detector γ -rays	162 ± 22	161 ± 21
Solar ν ERs	102 ± 6	102 ± 6
^{212}Pb + ^{218}Po β -decays	62.7 ± 7.5	63.7 ± 7.4
^3H + ^{14}C β -decays	58.3 ± 3.3	59.7 ± 3.3
^{136}Xe $2\nu\beta\beta$ decays	55.6 ± 8.3	55.8 ± 8.2
^{124}Xe DEC	19.4 ± 3.9	21.4 ± 3.6
^{127}Xe + ^{125}Xe EC	3.2 ± 0.6	2.7 ± 0.6
Atm. ν CEvNS	0.12 ± 0.02	0.12 ± 0.02
^8B + hep ν CEvNS	0.06 ± 0.01	0.06 ± 0.01
Det. Neutrons		$0.0^{+0.2}$
Accidentals	2.8 ± 0.6	2.6 ± 0.6
Total	1210 ± 91	1203 ± 42



Combined 2024 & 2022 Spin-independant Result

- Apply a profile likelihood ratio (PLR) analysis to search for WIMPs
 - Analysis power constrained at the -1σ level
- Included the WS2022 likelihood in the PLR
 - No changes to the WS2022 analysis or dataset
- Total exposure of 4.2 ± 0.1 tonne-years
- Peak sensitivity: $2.1 \times 10^{-48} \text{ cm}^2 @ 36 \text{ GeV}/c^2$
- Factor of 4 improvement in sensitivity into new parameter space



Conclusions

- World leading limit to WIMP dark matter
- Radon tag reduces main ER background by 60%
- First observation of charge suppression in DEC of ^{124}Xe
- LZ continuing onwards towards 1000 days of exposure
 - Multiple other areas of interest (^8B CEvNS, $0\nu2\beta$, etc.)

More from LZ

WIMP search 2024 arXiv:2410.17036

WIMP search 2022 Phys. Rev. Lett. 131, 041002

WS2022 backgrounds Phys. Rev. D 108, 012010

Low energy ER searches in WS2022 Phys. Rev. D 108, 072006

MIMP dark matter search in WS2022 Phys. Rev. D 109, 112010

WIMP-nucleon EFT search in WS2022 Phys. Rev. D 109, 092003

2ν DEC in Xe-124 J. Phys. G: Nucl. Part. Phys. 52 015103



LZ Collaboration

- Black Hills State University
- Brookhaven National Laboratory
- Brown University
- Center for Underground Physics
- Edinburgh University
- Fermi National Accelerator Lab.
- Imperial College London
- King's College London
- Lawrence Berkeley National Lab.
- Lawrence Livermore National Lab.
- LIP Coimbra
- Northwestern University
- Pennsylvania State University
- Royal Holloway University of London
- SLAC National Accelerator Lab.
- South Dakota School of Mines & Tech
- South Dakota Science & Technology Authority
- STFC Rutherford Appleton Lab.
- Texas A&M University
- University of Albany, SUNY
- University of Alabama
- University of Bristol
- University College London
- University of California Berkeley
- University of California Davis
- University of California Los Angeles
- University of California Santa Barbara
- University of Liverpool
- University of Maryland
- University of Massachusetts, Amherst
- University of Michigan
- University of Oxford
- University of Rochester
- University of Sheffield
- University of Sydney
- University of Texas at Austin
- University of Wisconsin, Madison
- University of Zürich

38 Institutions, 250 scientists, engineers, and technical staff



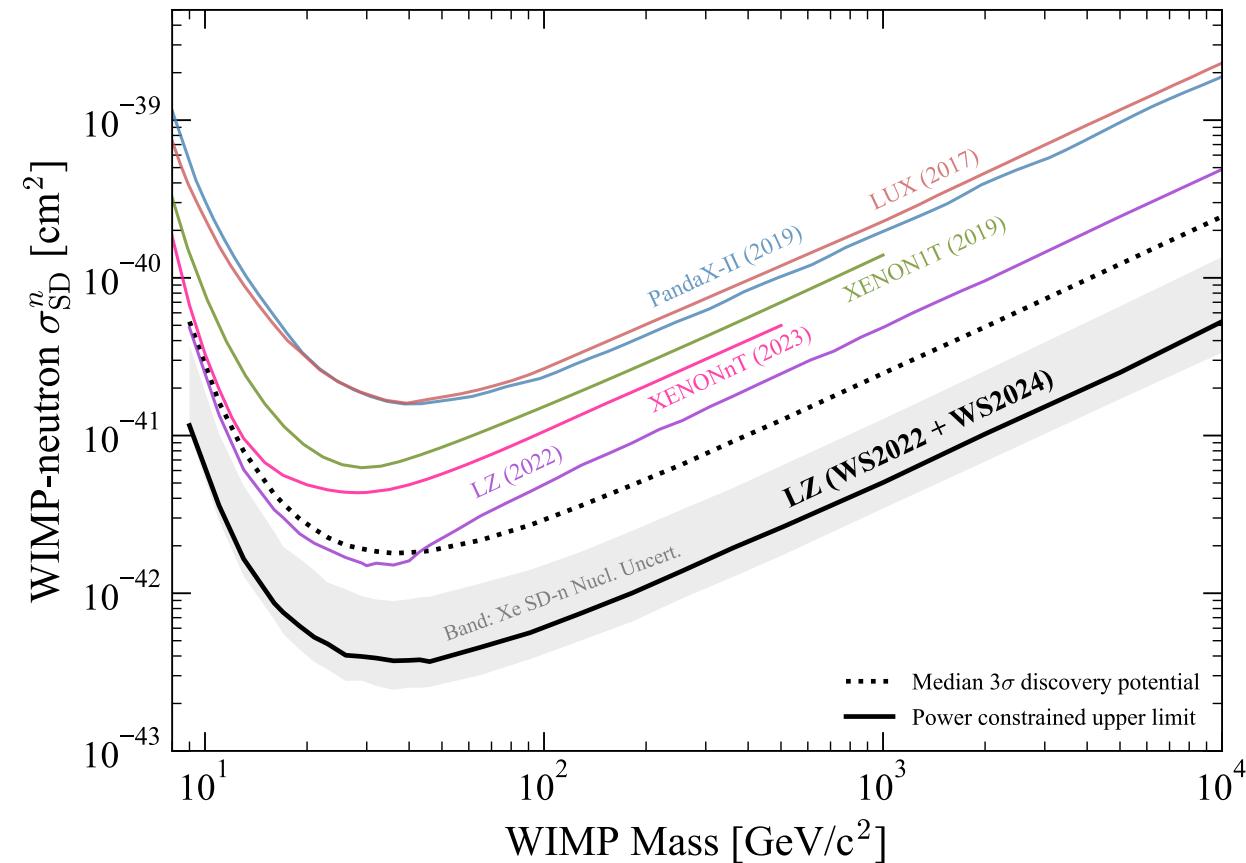
LZ Collaboration Meeting

June 2024 at Brown University

Backup

Combined 2024 & 2022 Spin-dependant Results

Neutron



Proton

