## Is dark energy dependent on cosmological curvature?

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Within present observational uncertainties, the time evolution of dark energy discovered by DESI (2025) is consistent with a simple model in which the dark energy density maintains a direct dependence on the  $|\Omega - 1|$  measure of spatial curvature. This, together with Bousso's (2002) conjecture that the holographic bound of the universe saturates at the observer's apparent horizon, and Gibbons & Hawking's (1977) postulate that the cosmological event horizon in de Sitter space is fully physically equivalent to an inverted event horizon of a Schwarzschild black hole of identical surface area, predicts with surprising accuracy both the strength and change in dark energy. A notable feature of this model is that  $\rho_{DE}$  remains small as long as space remains nearly flat, and  $\rho_{DE}$  acts to re-flatten space if  $\Omega$  ever deviates significantly from unity.

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