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Dark matter indirect detection limits including complete annihilation patterns

Céline **Armand**, Björn **Herrmann**

UNIVERSITÉ DE GENÈVE
FACULTÉ DES SCIENCES
Département d'astronomie

FNSNF
FONDS NATIONAL SUISSE
SCHWEIZERISCHER NATIONALFONDS
FONDO NAZIONALE SVIZZERO
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Enigmass
The enigma of mass

DARK MATTER

85% of the total matter of our Universe

Relic density observed experimentally by Planck:

$$\Omega_\chi h^2 \simeq 0.1200 \pm 0.0012$$

Ref: Ade et al. 2016, Astrophys. 594, A13

Its identification would **reveal new Physics**

Proving its existence and nature would **improve our understanding** of the Universe

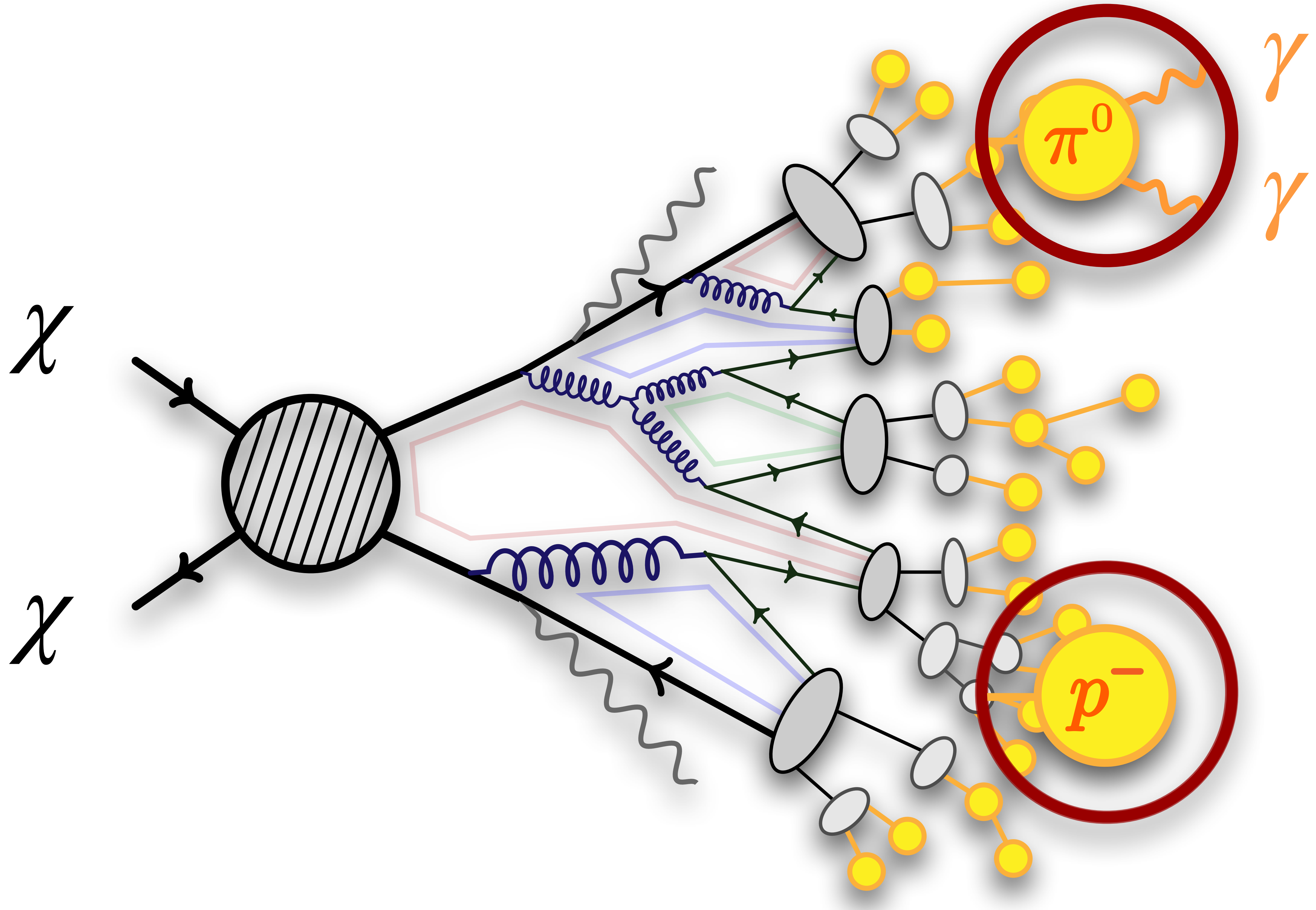
GOALS

Study of the impact of a more complete particle model

New prediction of DM upper limits with CTA mockdata of Sculptor

- **Previously:** use of individual annihilation channels
- **This work:** Collaboration with a theoretician to include a more complex and more complete model

INDIRECT **SEARCHES**



Dark Matter (DM) annihilation → Standard Model particles (bosons, quarks, leptons) → Final state products such as γ rays

INDIRECT **SEARCHES**

INDIRECT SEARCHES

Expected γ-ray flux from DM annihilation

Astrophysical
J factor

$$\frac{d\Phi\left(\langle\sigma v\rangle, J\right)}{dE} = \frac{1}{4\pi}\frac{\langle\sigma v\rangle}{2m_\chi^2}\sum_f \mathrm{BR}_f\frac{\mathrm{d}N_f}{\mathrm{d}E} \times \int_{\Delta\Omega}\int_{\mathrm{los}} \rho_{\mathrm{DM}}^2 ds d\Omega$$

Particle Physics
factor

where

- **<σv>** = annihilation cross-section
- **m_χ** = DM particle mass
- **BR_f** = branching ratio
- **dN_f/dE** = differential spectrum
- **ρ_{DM}** = DM density

STATISTICAL ANALYSIS

LOG-LIKELIHOOD RATIO TEST STATISTICS

Constrained minimization

$$\Lambda = -2 \ln \frac{\mathscr{L}_{H_0}}{\mathscr{L}_{H_1}} = -2 \ln \frac{\mathscr{L}(\langle \sigma v \rangle_0 \,|\, \hat{\hat{N}}_B, \hat{\hat{J}})}{\mathscr{L}(\langle \hat{\sigma v} \rangle, \hat{N}_B, \hat{J})}$$

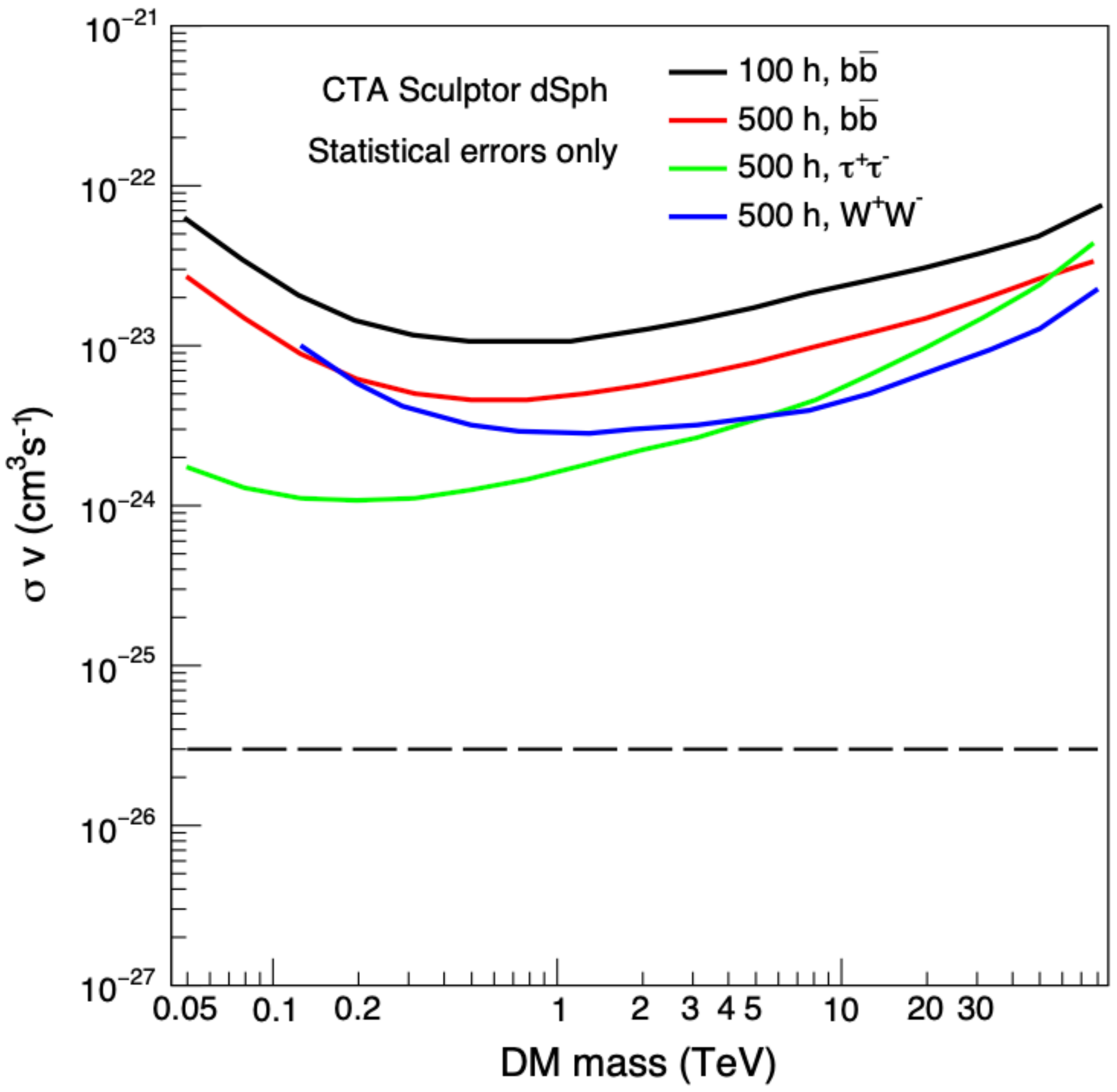
Global minimization

Ref: Cowan et al, 2010
Eur.Phys.J.C71:1554,2011

<σv>	Parameter of interest
N_B, J	Nuisance parameters

Λ	2.71 at 95% Confidence Level

UPPER LIMITS



- Each annihilation channel **treated independently**
- Corresponding to a **branching ratio of 100%**
- **Simplest** model possible where all DM particles annihilate through the same channel

WHAT IF

We change the particle physics model?

SINGLET SCALAR DARK MATTER

Standard model extended by an additional scalar field (DM)

$$V_{\text{scalar}} \supset 2\lambda_H v^2 h^2 + \frac{1}{2}\mu_S^2 S^2 + \frac{1}{4}\lambda_{SH} v^2 S^2 + \frac{1}{4}\lambda_{SH} v S^2 h + \lambda_{SH} S^2 h^2$$

DM mass

$$m_S^2 = \mu_S^2 + \frac{1}{2}\lambda_{SH} v^2$$

DM – Higgs interaction

("Higgs portal")

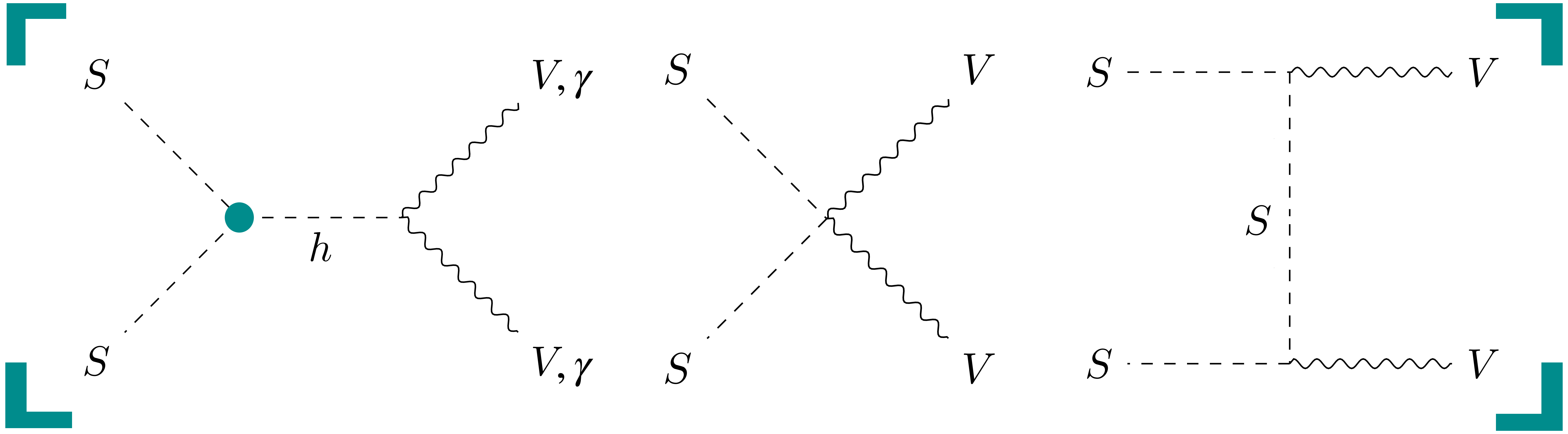
Phenomenology governed by

m_S (DM mass)

λ_{SH} (DM coupling)

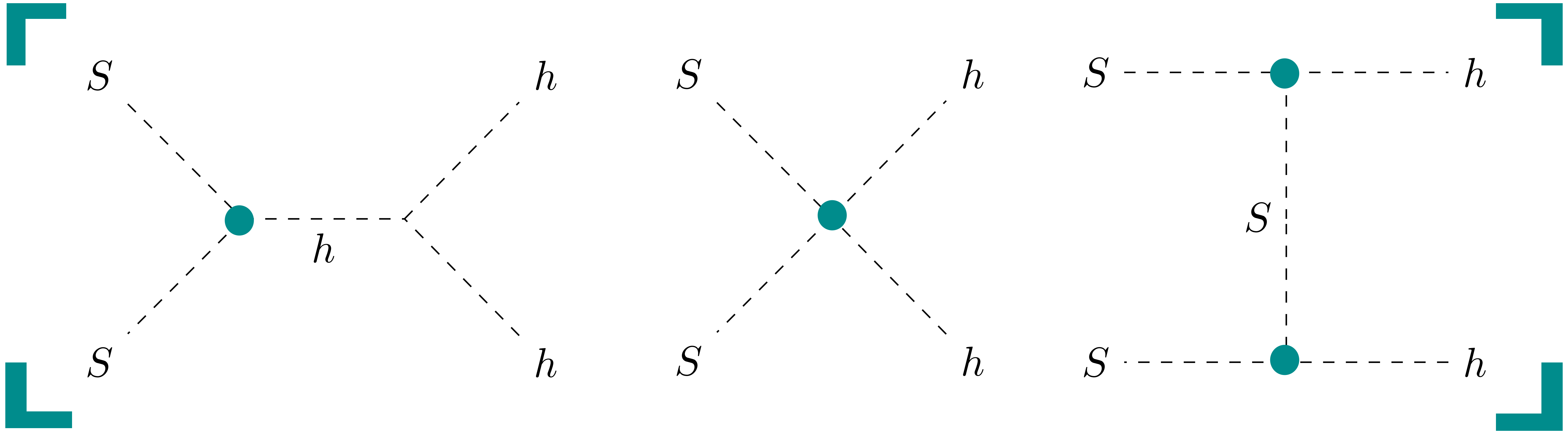
SINGLET SCALAR DARK MATTER

Possible dark matter annihilation channels (DM relic density + indirect detection)

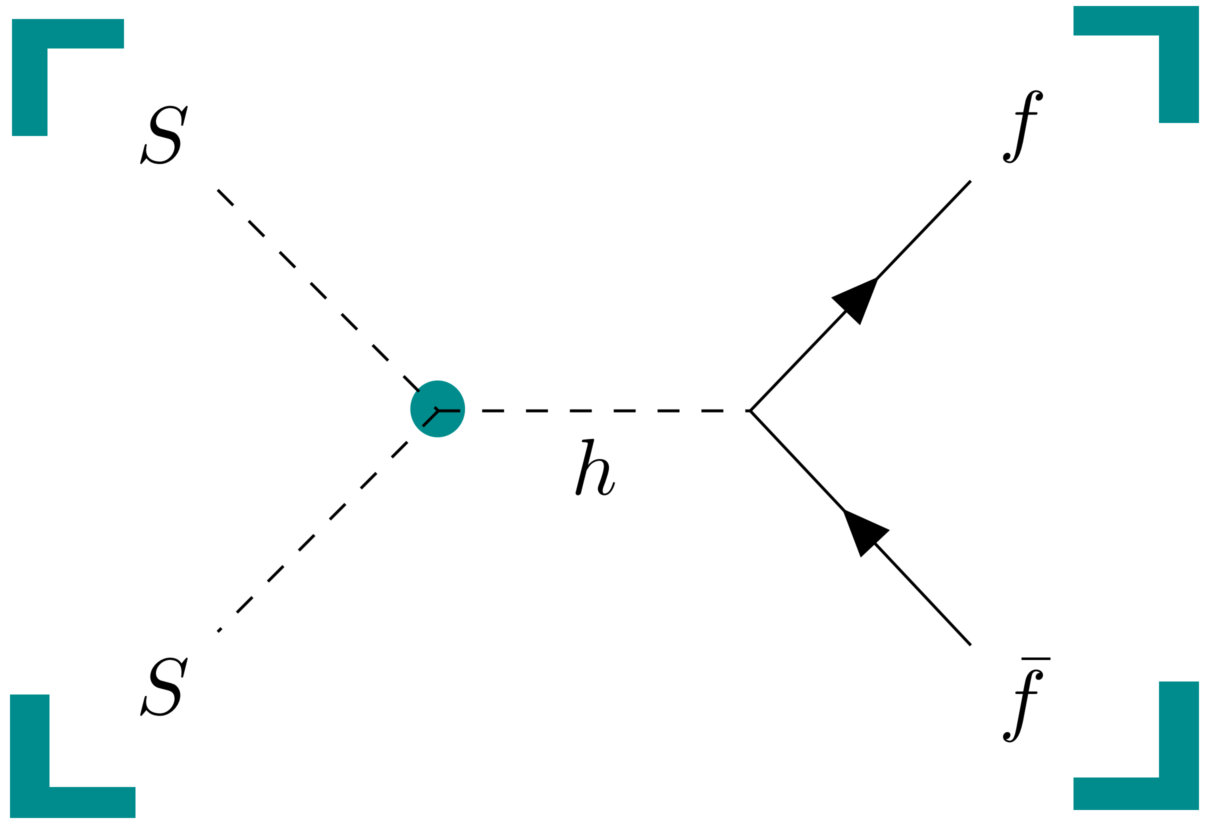


Gauge boson final states

$V = Z^0, W^{\pm}$



Higgs boson final states



Quark or lepton final states

$f = u, d, c, s, b, t, e, \mu, \tau$

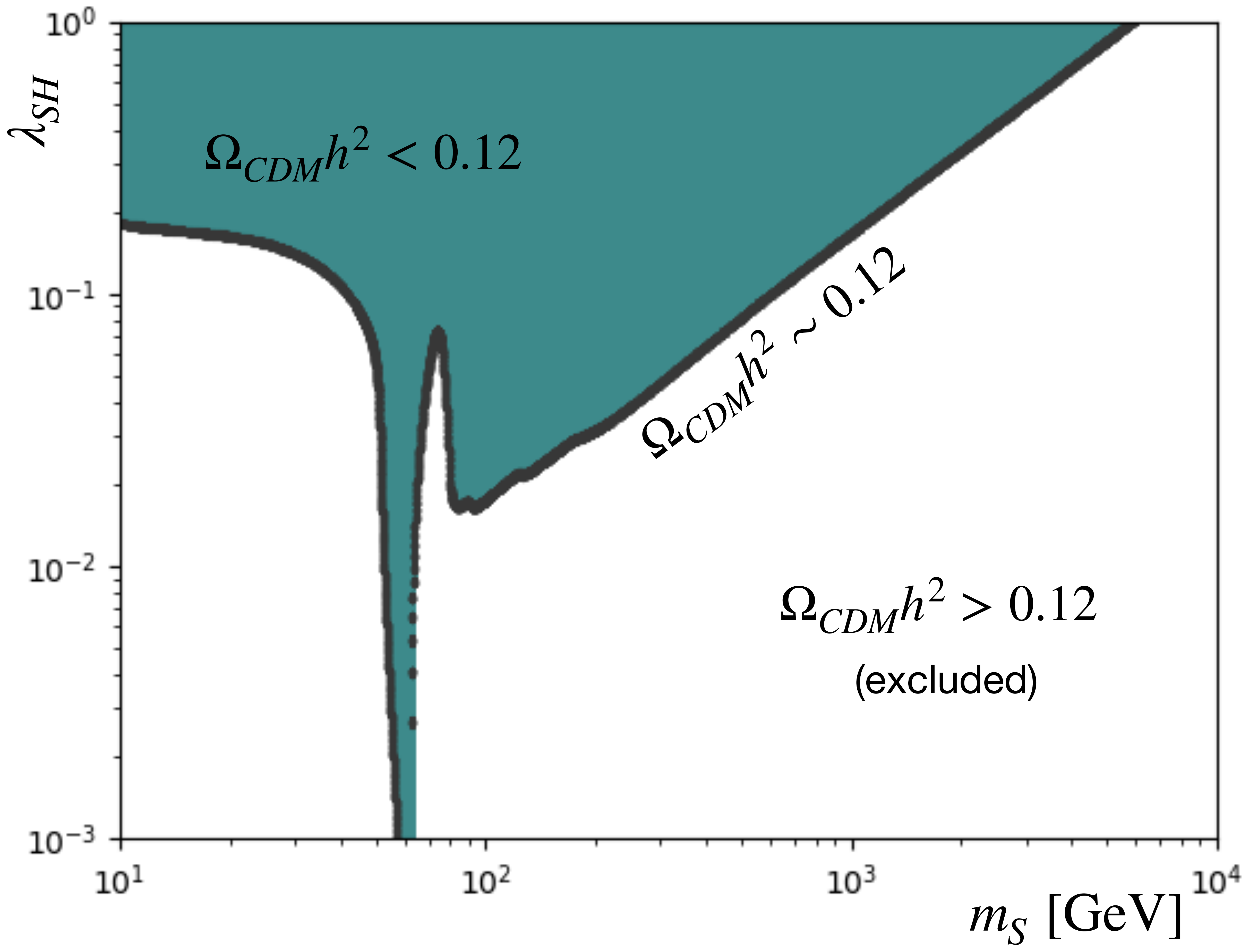
SINGLET SCALAR DARK MATTER

SINGLET SCALAR DARK MATTER

DM coupling vs DM mass

Relic density and branching ratio grid computed using **micrOMEGAs**

Ref: Bélanger, Pukhov et al. 2003 – 2022



SINGLET SCALAR DARK MATTER

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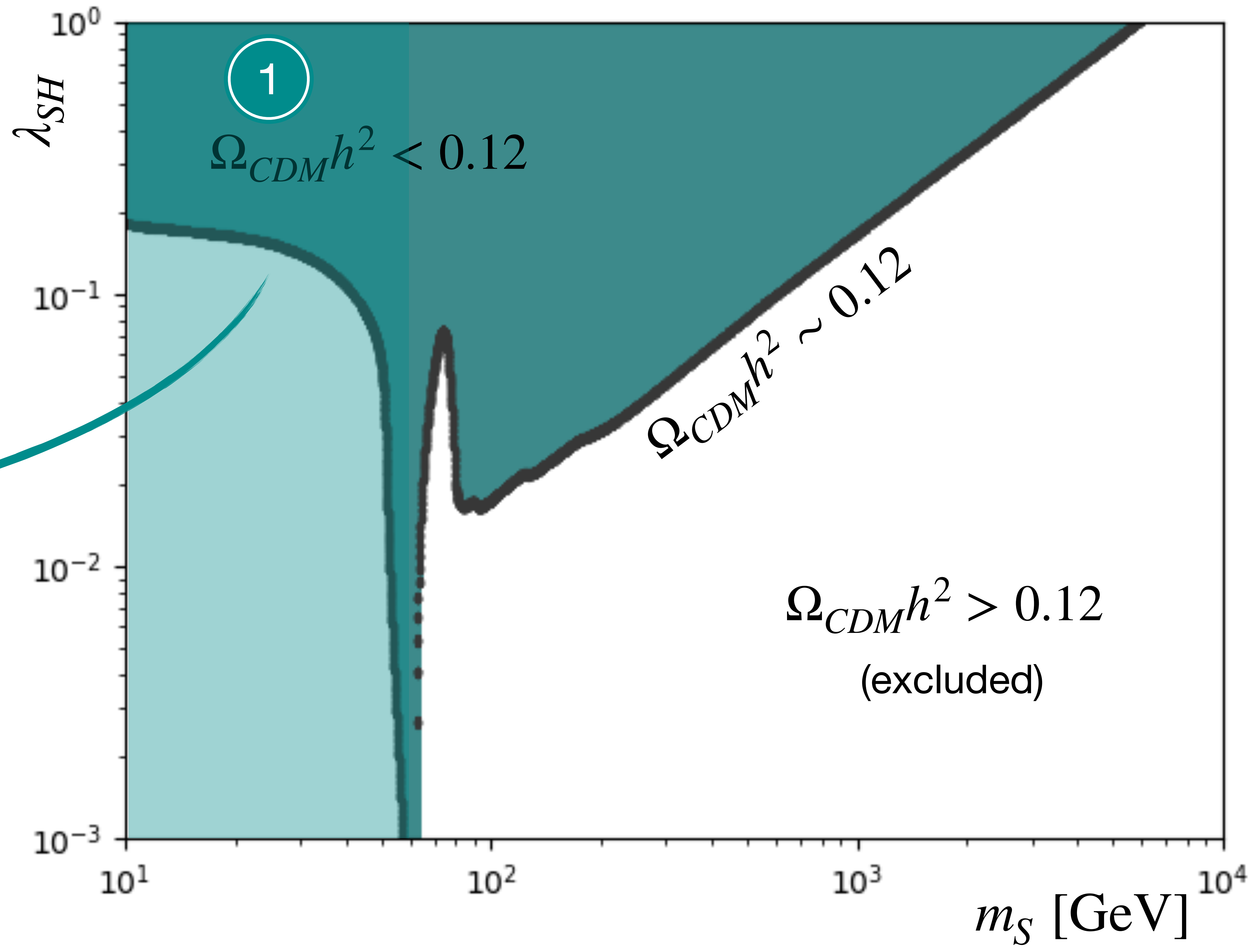
$SS \to b\bar{b} \quad \sim 90\,\%$

$SS \to \tau^+\tau^- \quad \sim 10\,\%$



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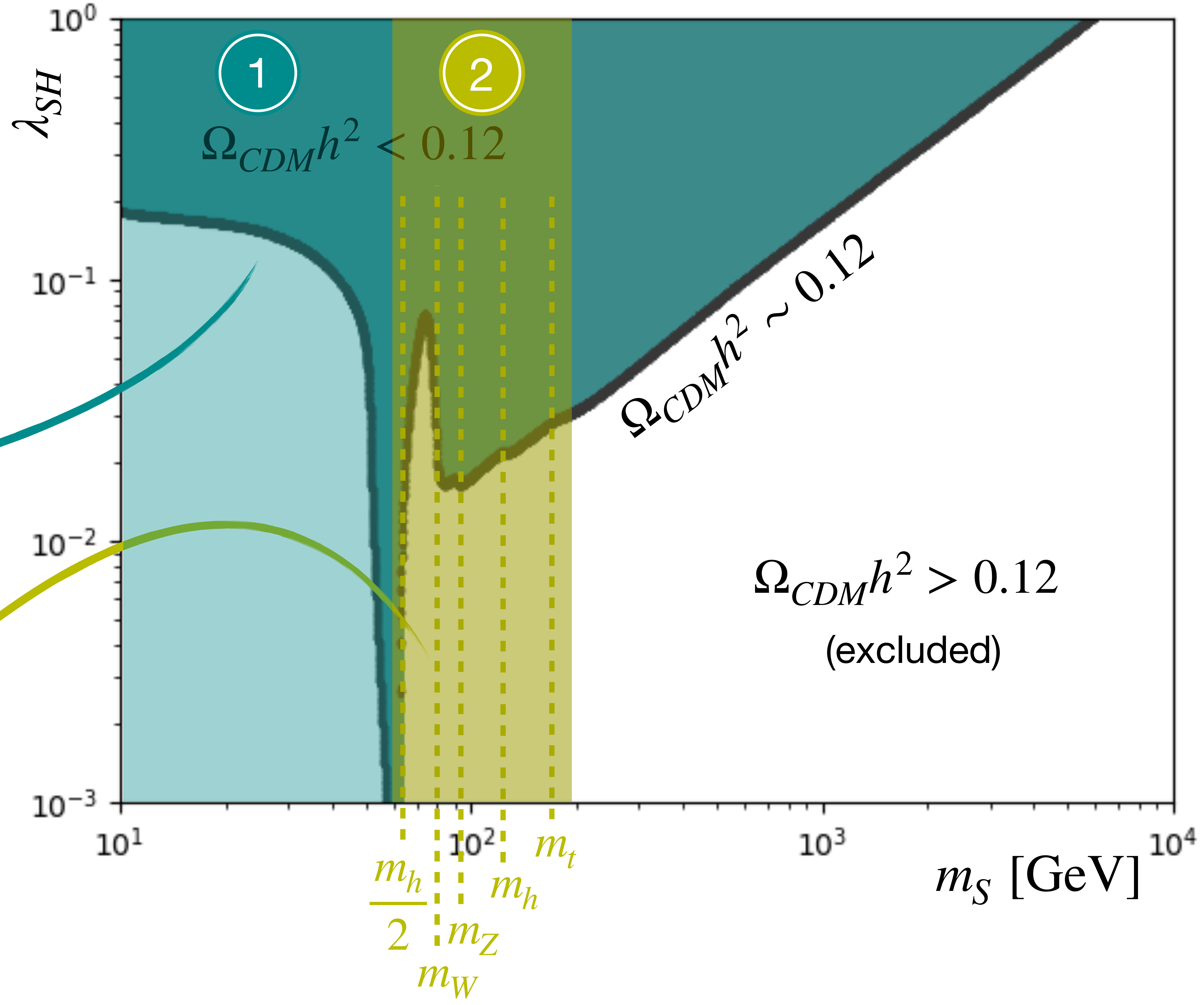
Region of frequent change of the dominant annihilation channel

New annihilation channels open, Higgs resonance at $m_h/2$



SINGLET SCALAR DARK MATTER

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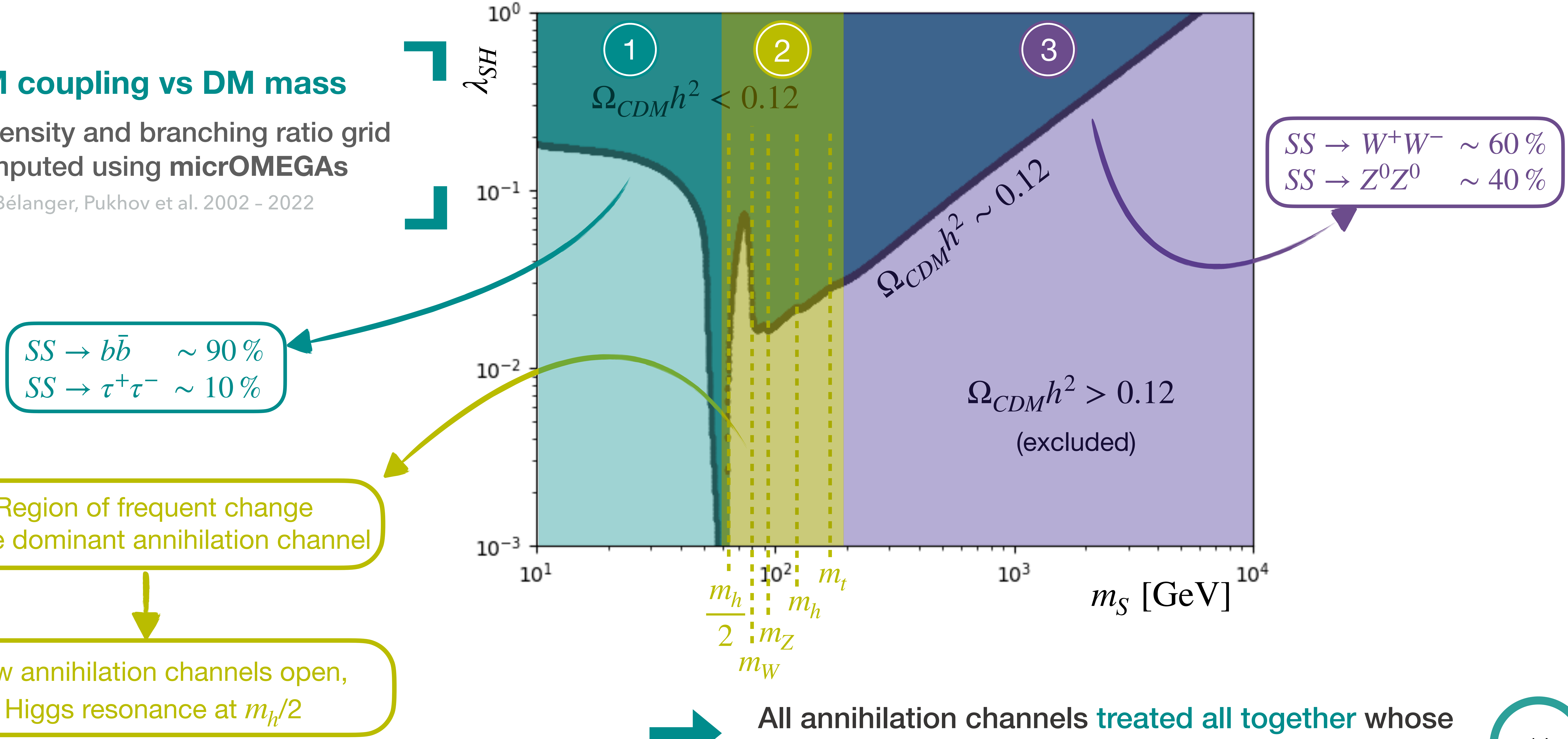


All annihilation channels treated all together whose branching ratio varies with respect to the DM mass

Even in such a simple setup, the "100% hypothesis" is not justified…

More complex models invoke an even richer phenomenology…

TARGET SOURCE

Dwarf galaxy selected for the CTA dark matter program

South Hemisphere
l = 287.62°, b = -83.16°

J factor
Log10 $J = 18.3 \pm 0.3$

Ref: Bonnivard et al, 2015 ApJ 808 L3

Sculptor

Mock data prepared with Gammapy 0.18.2

Simulated events for 500h of observation at 20° zenith angle

NEW UPPER LIMITS

NEW UPPER LIMITS

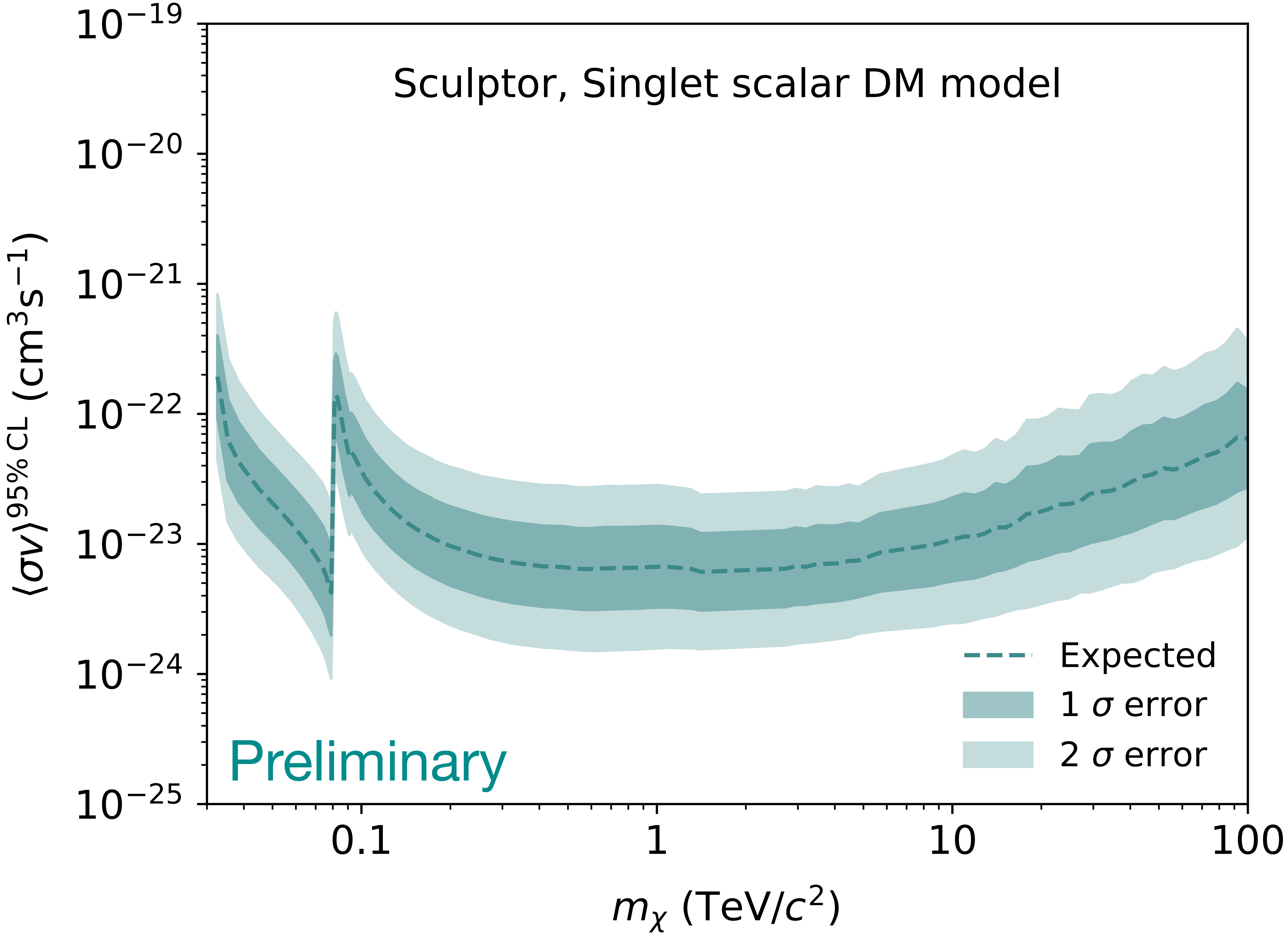
Computation of the predicted DM cross section
VS
DM particle mass

1. **Expected** limits - Sample of **300 Poisson realizations** of the simulated background events

2. **Mean expected limits**
   Mean of the derived <σv> distribution

   Statistical uncertainty bands
   Standard deviation at 1 and 2σ

RESULTS



Predicted upper limit and uncertainties

Assuming a singlet scalar DM model

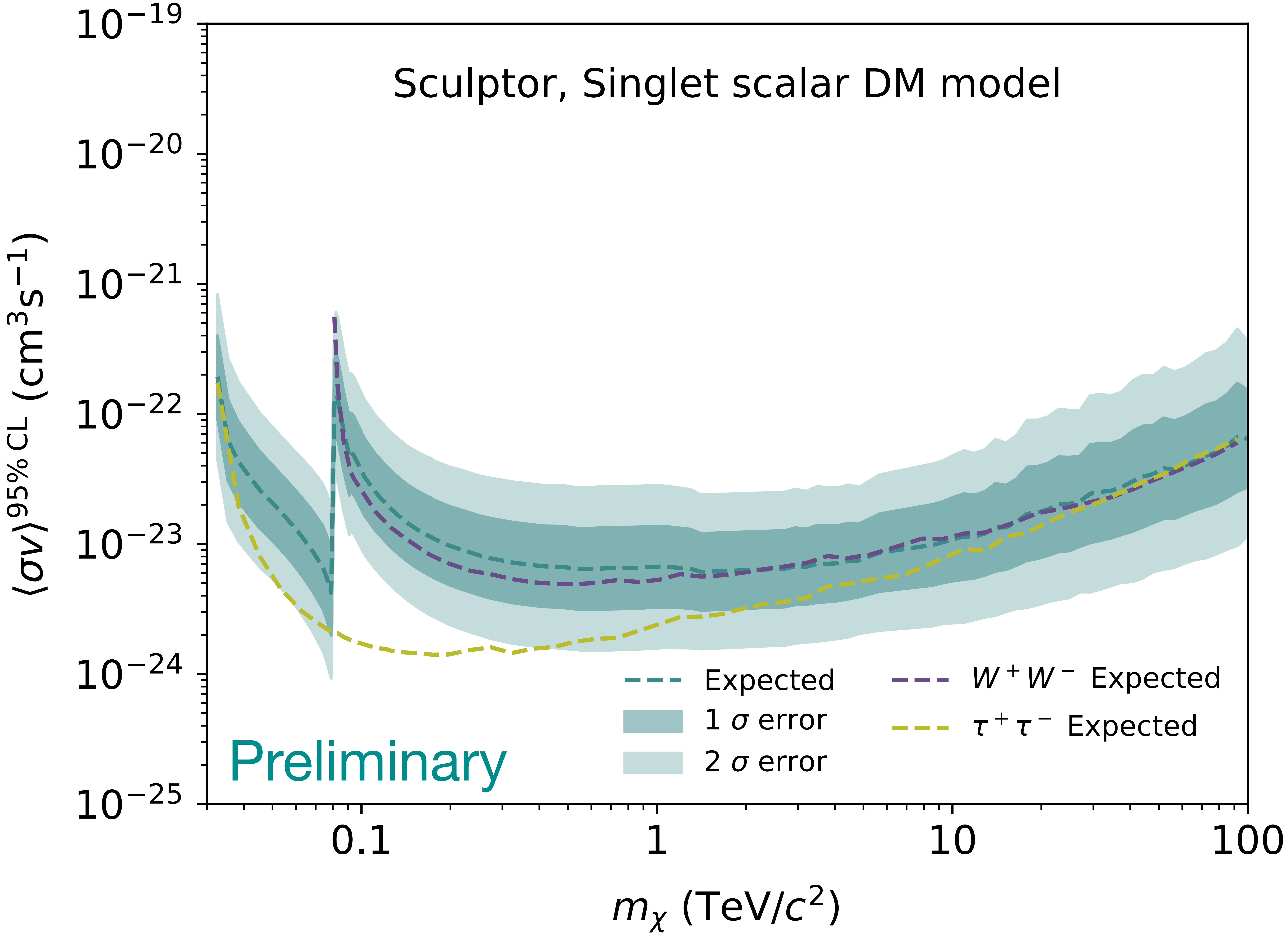
Inflection point

Due to the Higgs resonance

Sudden increase

Due to the opening of the WW channel

COMPARISONS



SINGLET SCALAR MODEL VS 100% W^+W^-

More conservative limit with the singlet scalar DM model

Below the W mass

No upper limit for 100% WW since the WW channel **does not exist**

~0.1-1 TeV

New contributions in addition to WW and ZZ

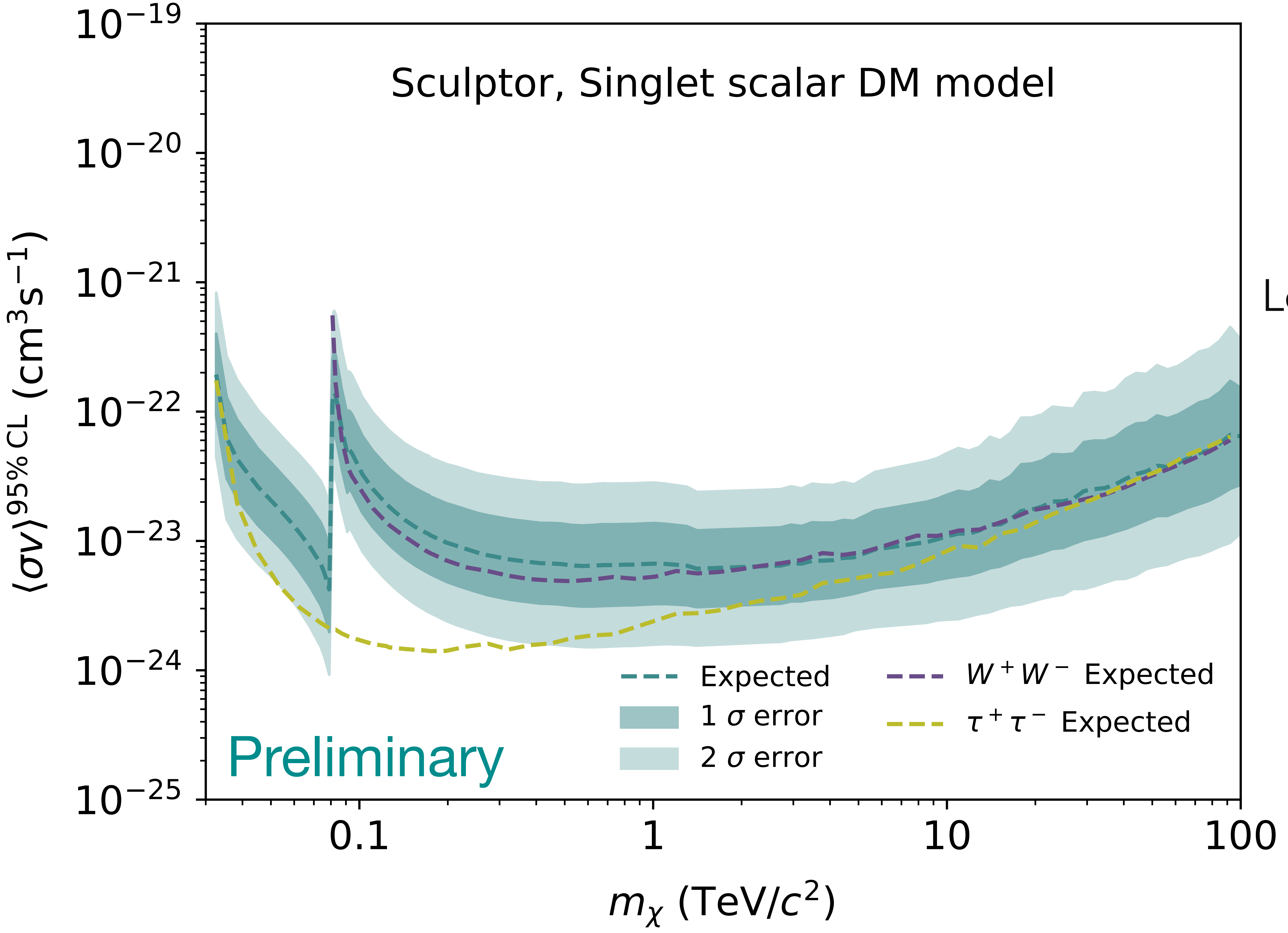
Slight difference between 100% WW and singlet scalar DM model

Above 1 TeV

~60% WW - 40% ZZ

Limits similar to the 100% WW case since WW and ZZ have **similar cross sections**

COMPARISONS



SINGLET SCALAR MODEL
VS
100% $\tau^+\tau^-$

100% $\tau^+\tau^-$ produces more γ rays
Leads to **more constraining** upper limits

However, in the singlet scalar model, **this $\tau^+\tau^-$ channel is never dominant**

100% $\tau^+\tau^-$ = over estimation of the contribution

CONCLUSION & PERSPECTIVES

- Use of a more complex and more complete particle physics model
- Takes into account the full phenomenology with all annihilation channels at once
- Change of dominant annihilation channel(s) along with the DM particle mass
- Affects the predicted upper limits
- Feature can be expected in any particle physics model
- Derivation of a predicted upper limit and its 1σ and 2σ uncertainty bands over the full energy range of CTA
- Could be used as well on the future data of CTA
- Paper in preparation

Thanks for your attention

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Enigmass
The enigma of mass

STATISTICAL ANALYSIS

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Total likelihood

$$\mathcal{L}(\langle\sigma v\rangle, N_B, J) = \prod_{i=1} \mathcal{L}_{P_i}(\langle\sigma v\rangle, N_{B_i}, J \,|\, N_{\mathrm{ON_i}}, N_{\mathrm{OFF_i}}, \alpha)\mathcal{L}^J(J \,|\, \bar{J}, \sigma_J)$$

Poisson likelihood

Log-normal likelihood

Poisson likelihood for each energy bin

$$\mathcal{L}_i^P = \frac{(N_{S_i} + N_{B_i})^{N_{\mathrm{ON}_i}}}{N_{\mathrm{ON}_i}!} e^{-(N_{S_i}+N_{B_i})} \cdot \frac{(\alpha N_{B_i})^{N_{\mathrm{OFF}_i}}}{N_{\mathrm{OFF}_i}!} e^{-\alpha N_{B_i}}$$

ON REGION

OFF REGION

Log-normal likelihood to model the uncertainties of the J factor

$$\mathcal{L}^J = \frac{1}{\ln(10)\sqrt{2\pi}\sigma_J J} \exp - \frac{(\log_{10} J - \log_{10} \bar{J})^2}{2\sigma_J^2}$$

COMPARISONS