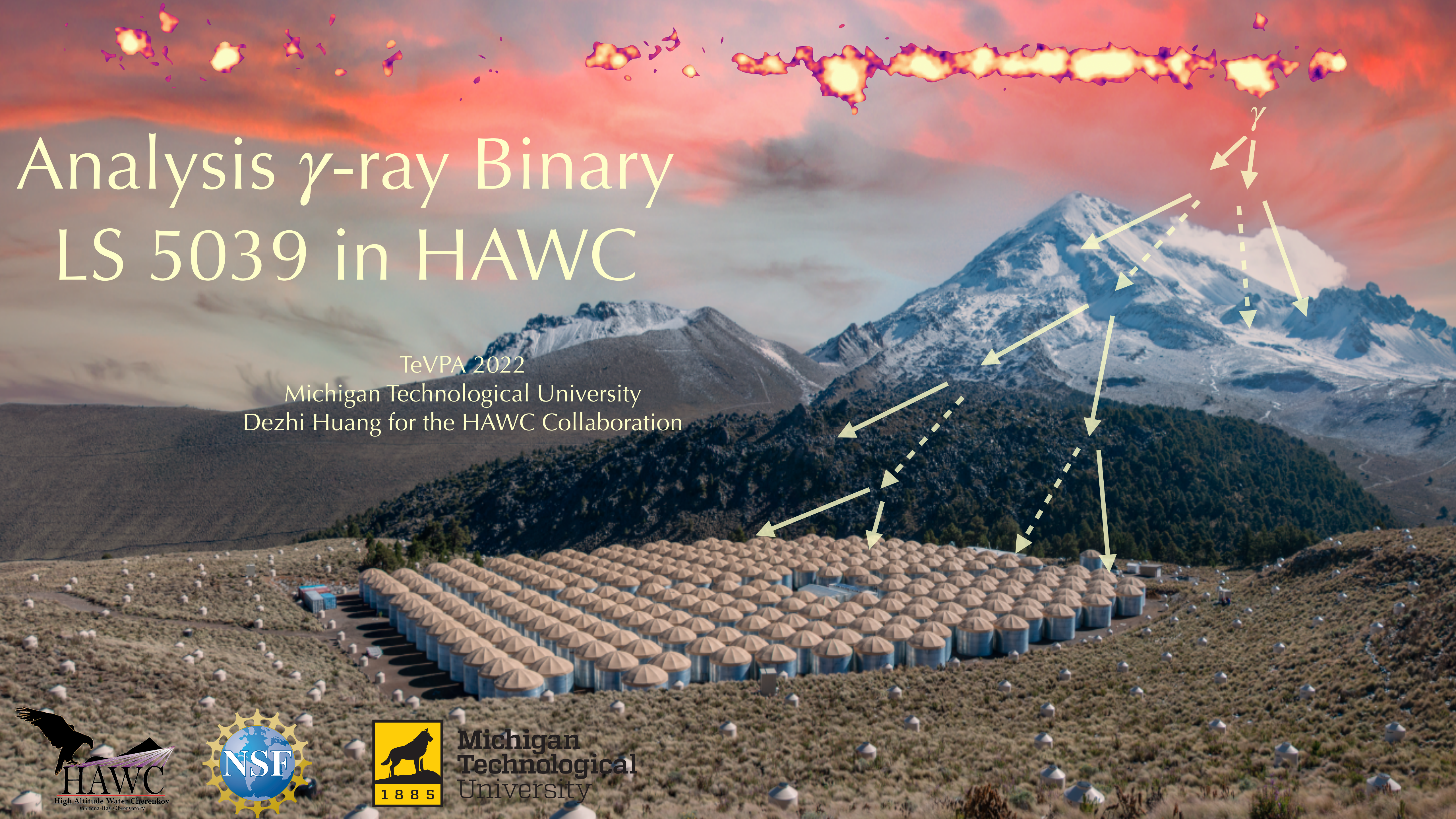


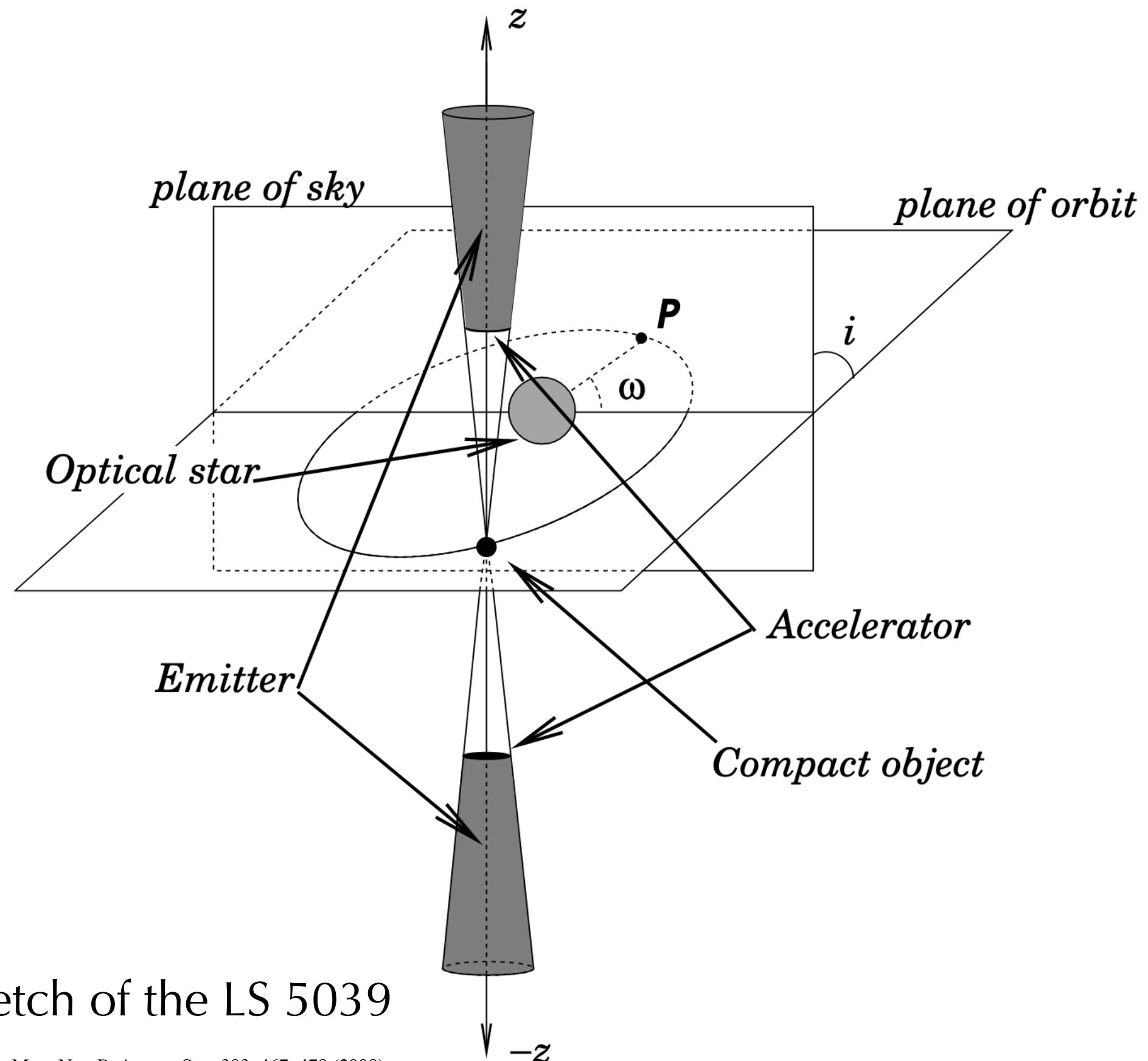
# Analysis $\gamma$ -ray Binary LS 5039 in HAWC

TeVPA 2022  
Michigan Technological University  
Dezhi Huang for the HAWC Collaboration



Michigan  
Technological  
University

# High-mass $\gamma$ -ray Binary LS 5039



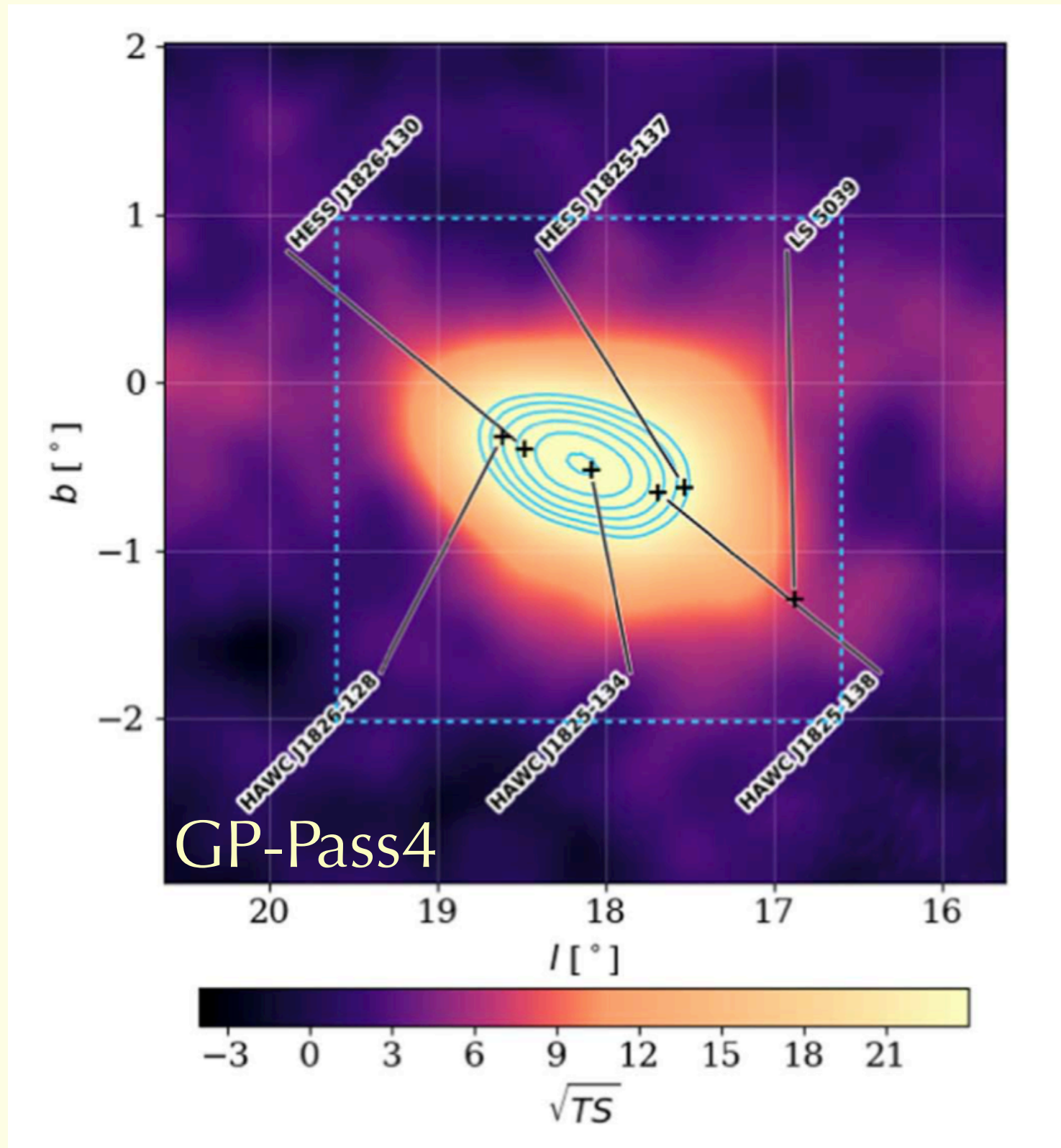
Sketch of the LS 5039

Credits: Mon. Not. R. Astron. Soc. 383, 467–478 (2008)

- Optical type O star has the mass  $\sim 22.9 (+ 3.4, -2.9) M_{\odot}$
- The compact object  $3.7 (+ 1.3, -1.0) M_{\odot}$  Pulsar? Black Hole?
- The orbit period is  $\sim 3.9$  days
- Located  $\sim 1.5^{\circ}$  from galactic plane

# LS 5039 at HAWC

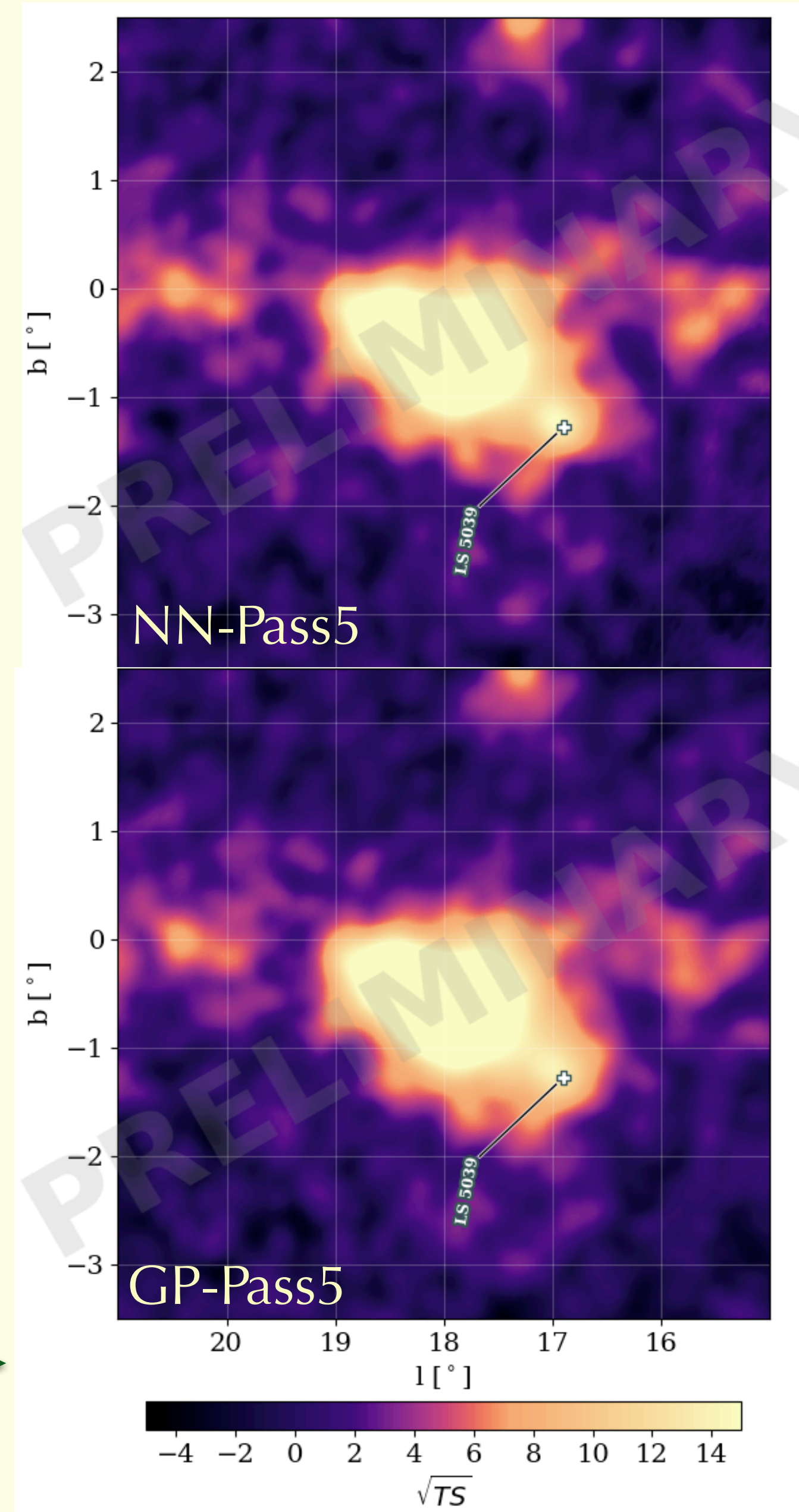
- We are unable to resolve LS 5039 in HAWC Pass4 maps



Pass4 Pass5

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Improvement of angular resolution



Two independent energy estimators

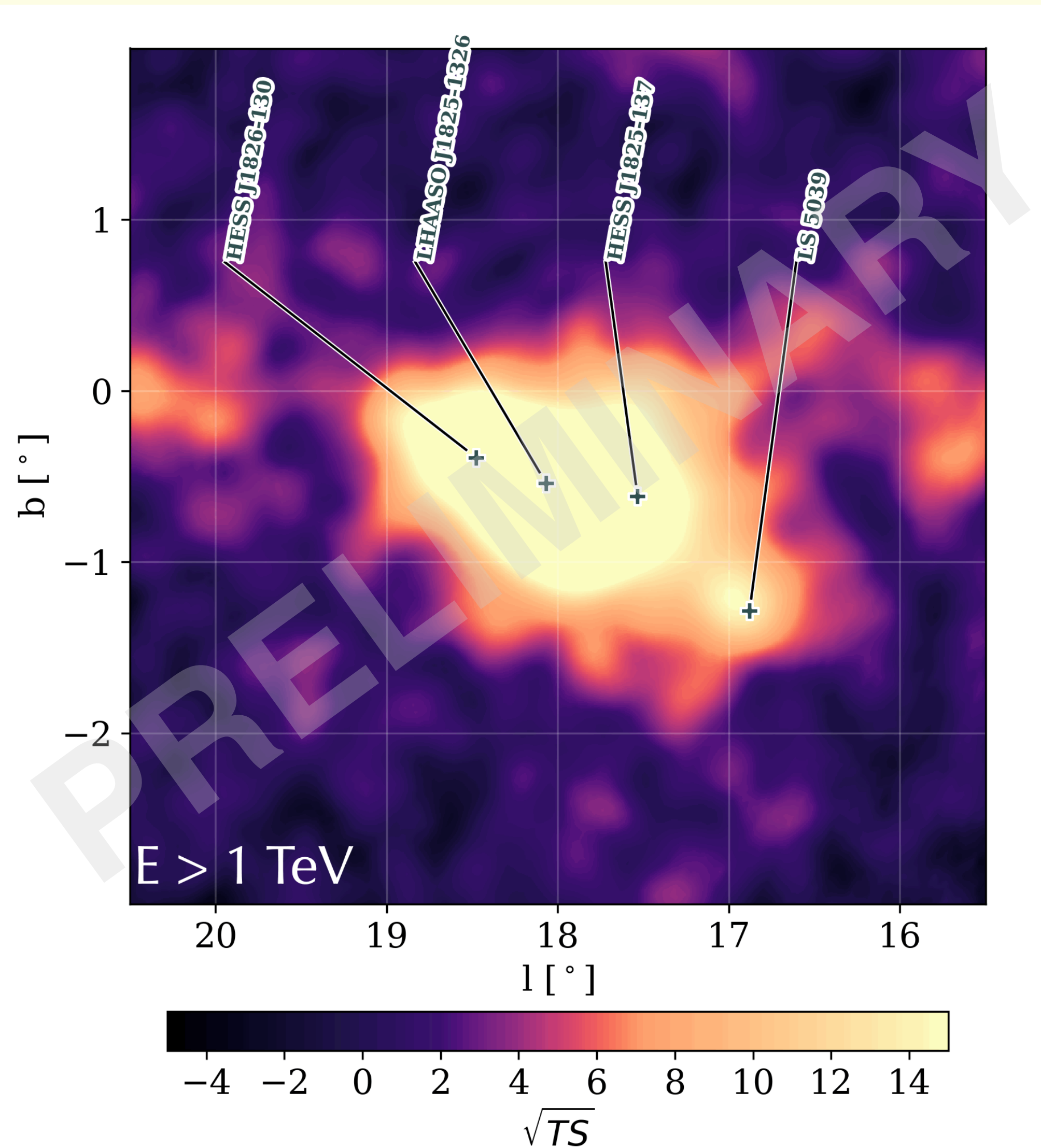
Neural network(NN)

- Using artificial neural network to estimate the energy of primary gamma-rays

Ground Parameter(GP)

- Using charge density at certain distance from shower axis to estimate the energy of primary gamma-rays

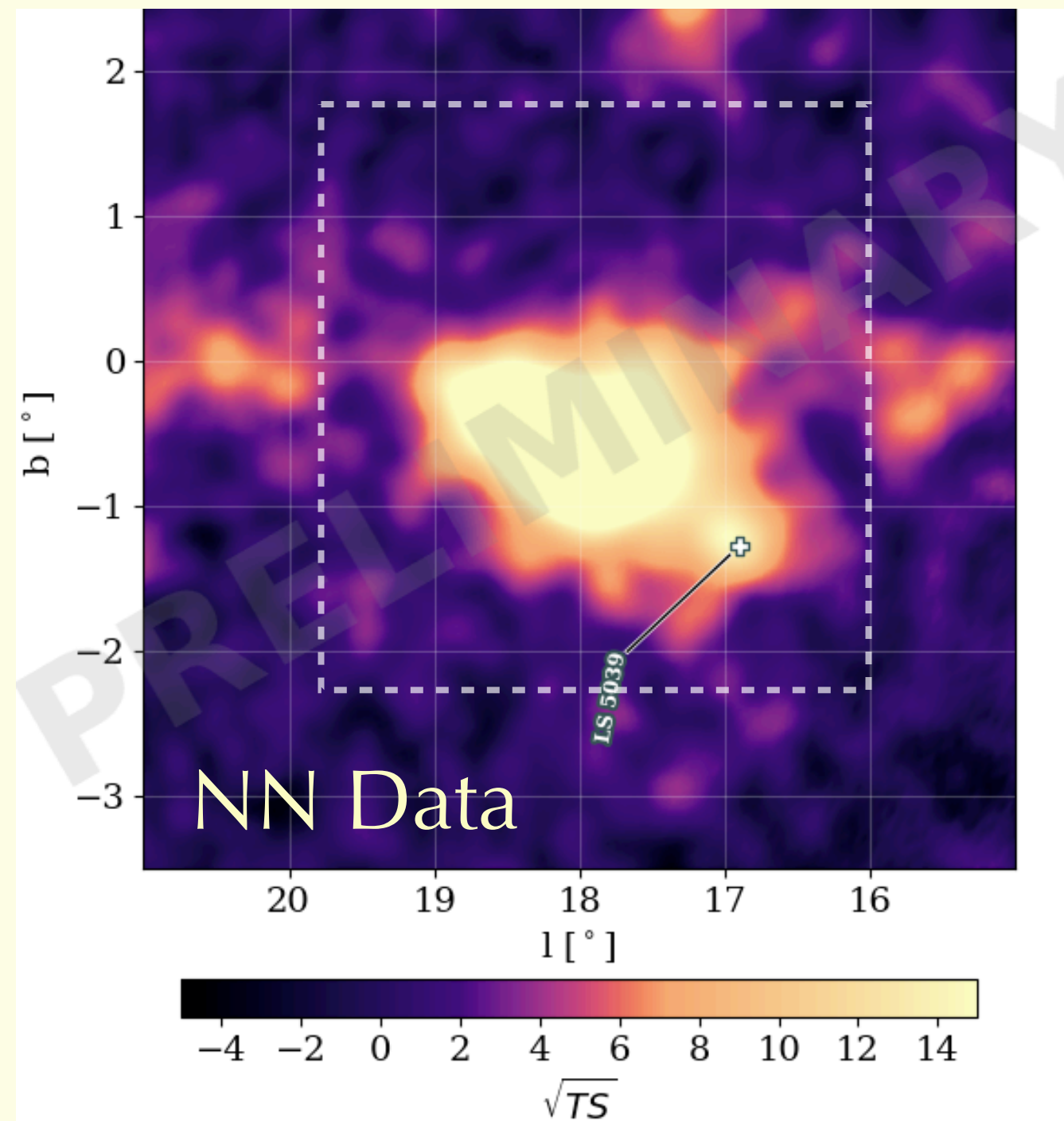
# LS 5039 at HAWC



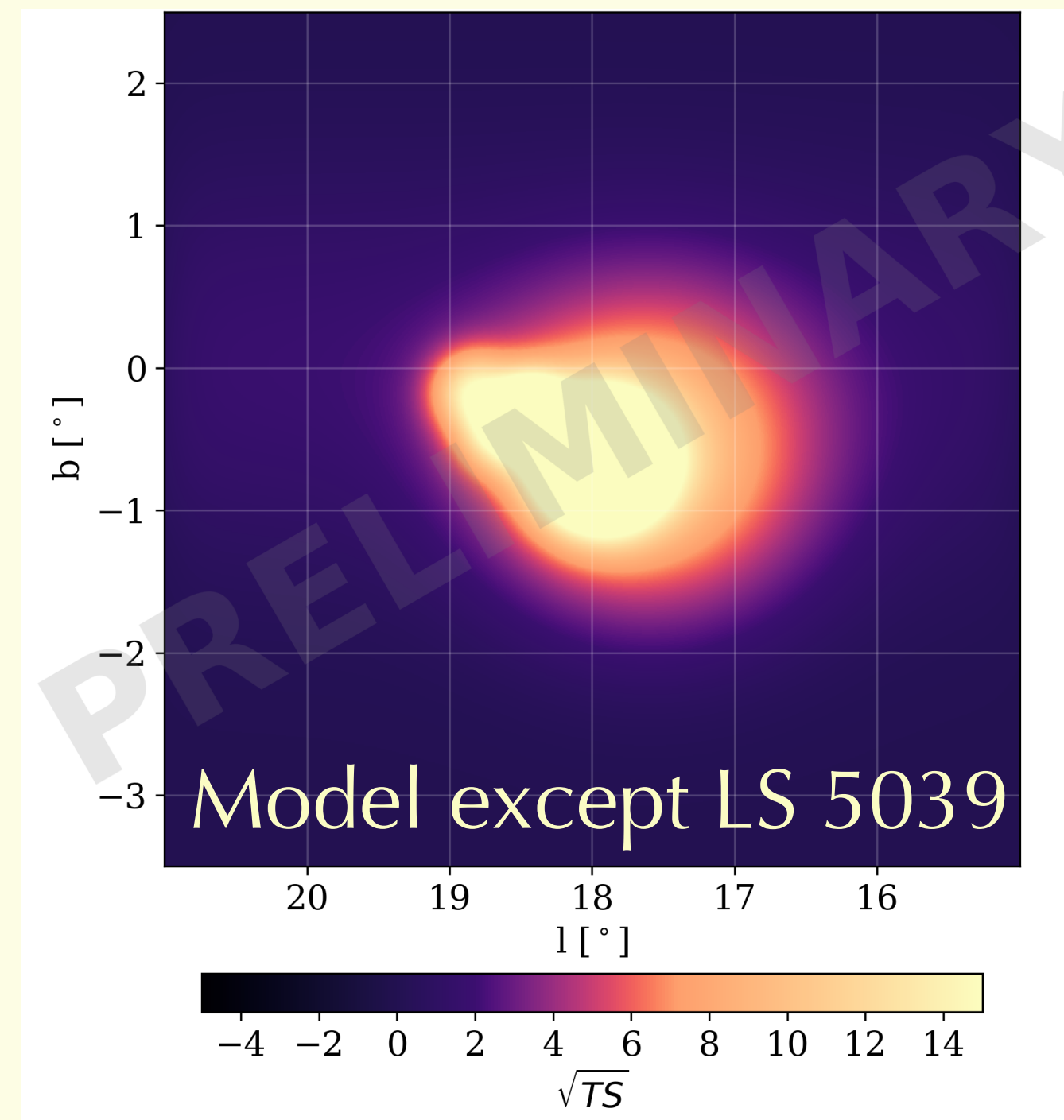
- Hotspot found in the LS 5039 location
- Hotspot start separate from J1825 region when we move to high energy
- To isolate the LS 5039 from this complex region:
  1. Diffuse background emissions
  2. J1825 contamination

# Modeling the Region

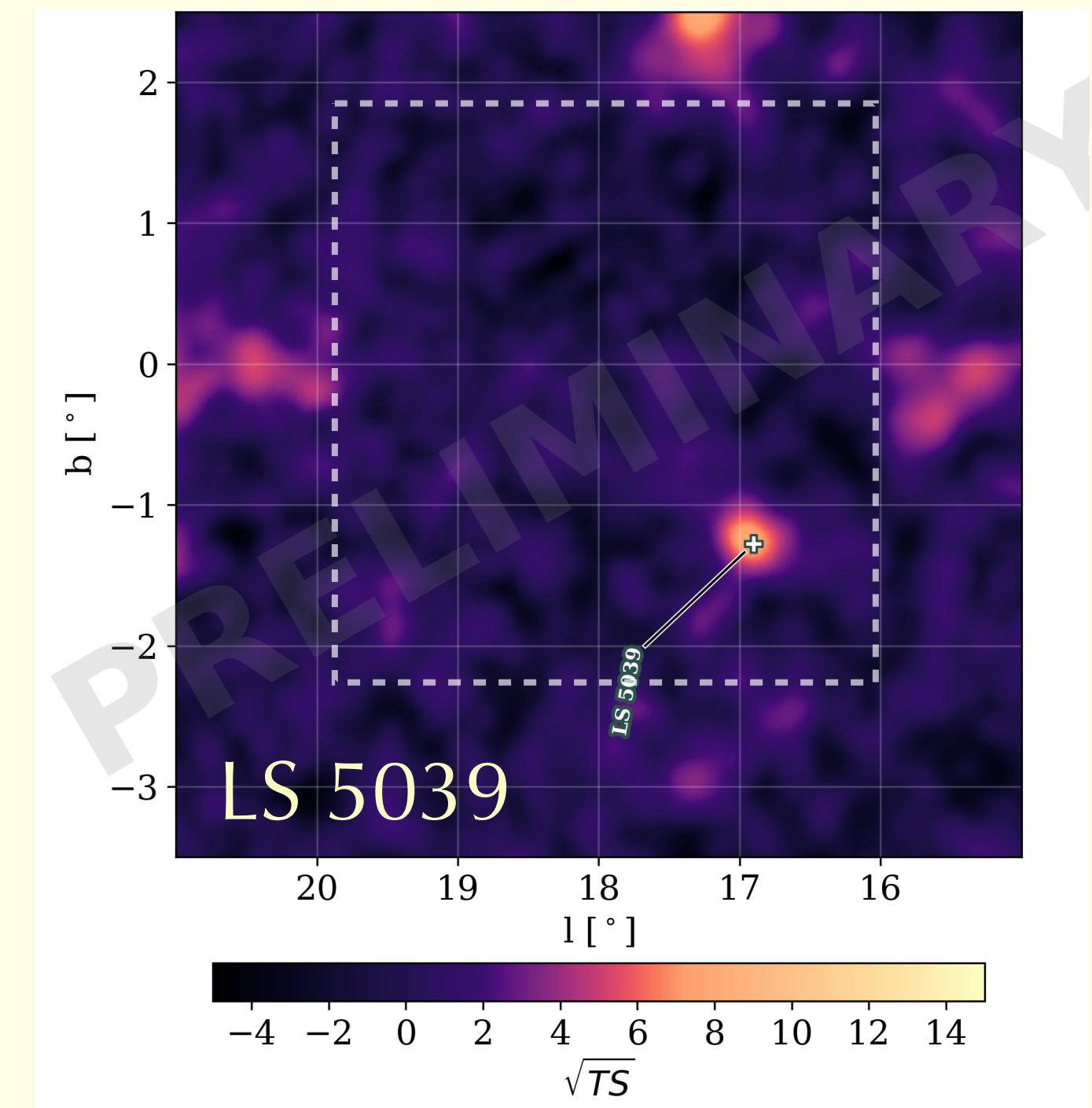
- Simultaneously likelihood fit performed inside the region of interest
- Model includes diffuse background emission and all background sources
- Same analysis repeated using GP map

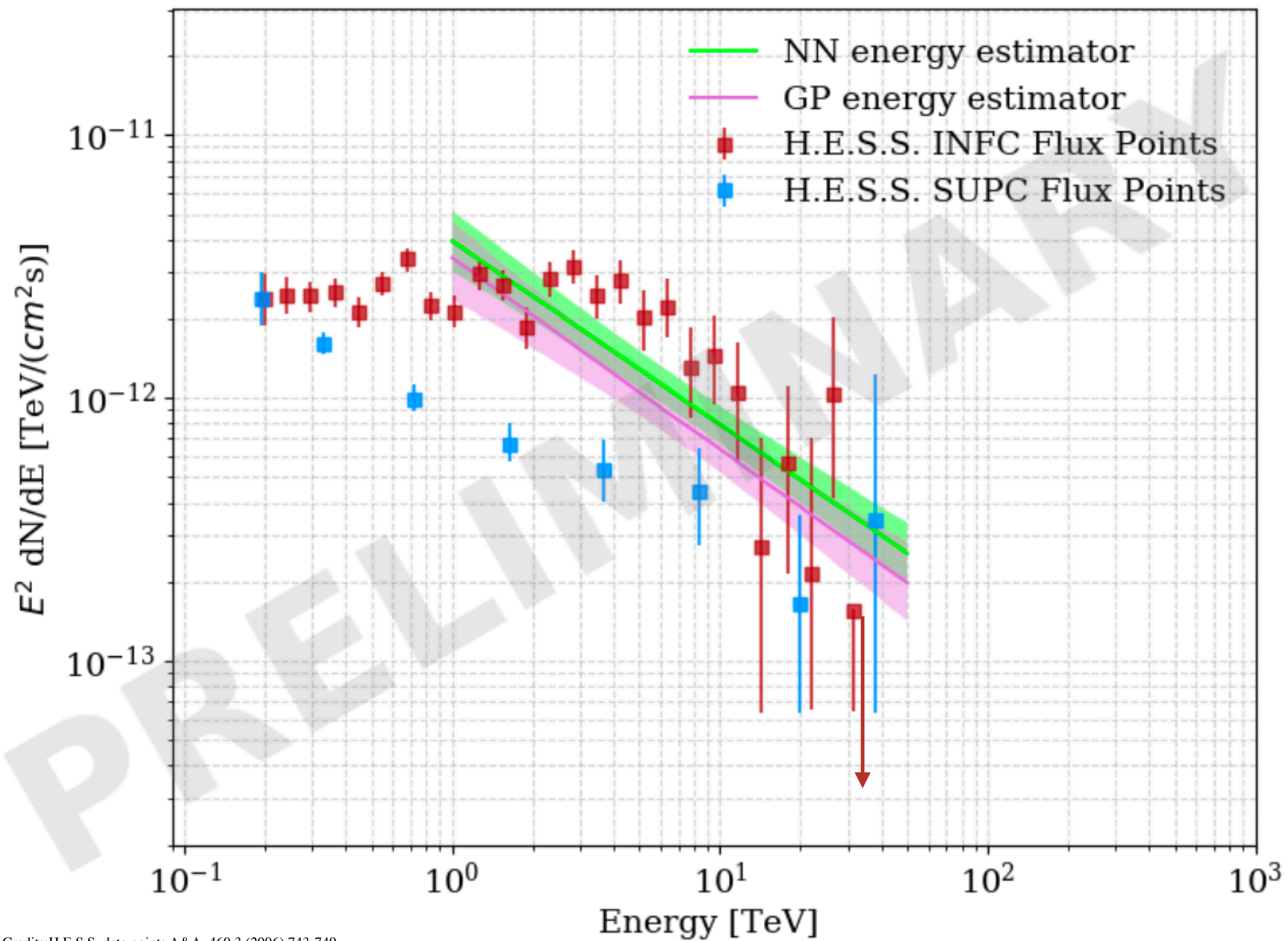


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Credits: H.E.S.S. data points [A&A, 460 3 \(2006\) 743-749](#)

- Pure powerlaw is preferred in both of HAWC maps
- HAWC spectrum is located in between of H.E.S.S. INFC and SUPC
- Can we distinguish high and low states HAWC ?

# Analysis Procedure

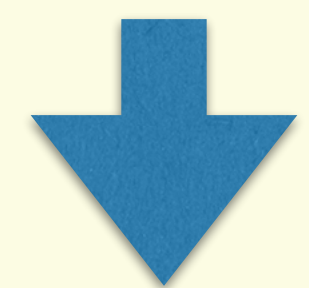
## HAWC Data Using NN Energy Estimator

Inferior conjunction (INFC)

High State  $0.45 < \phi \leq 0.9$

Superior conjunction (SUPC)

Low State  $\phi \leq 0.45$  or  $\phi > 0.9$



Separate HAWC data

High State Data

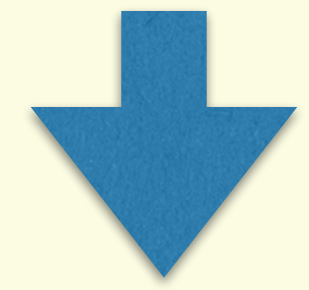
Low State Data



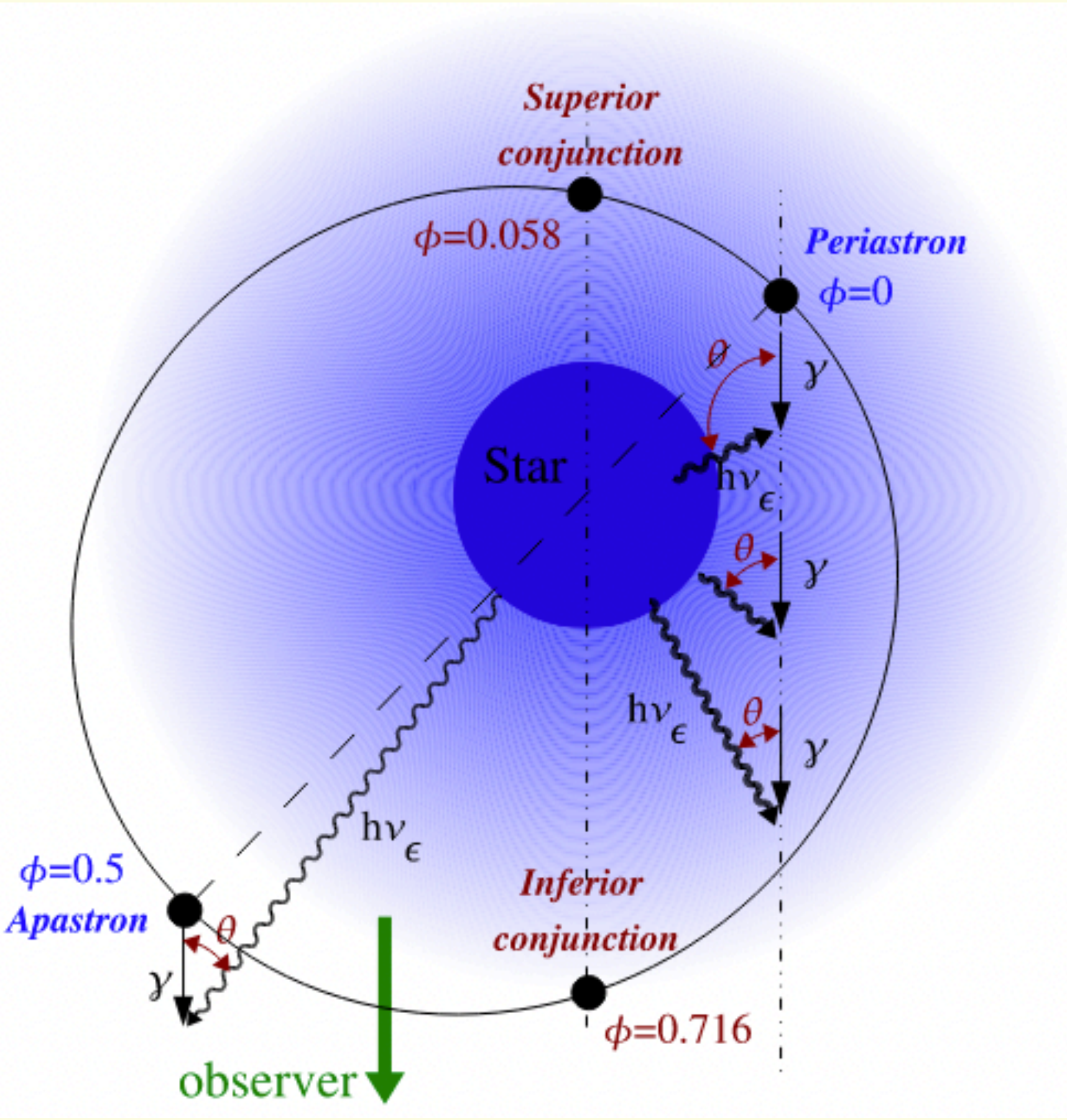
Map Maker

High State Map 817 Days

Low State Map 1003 Days

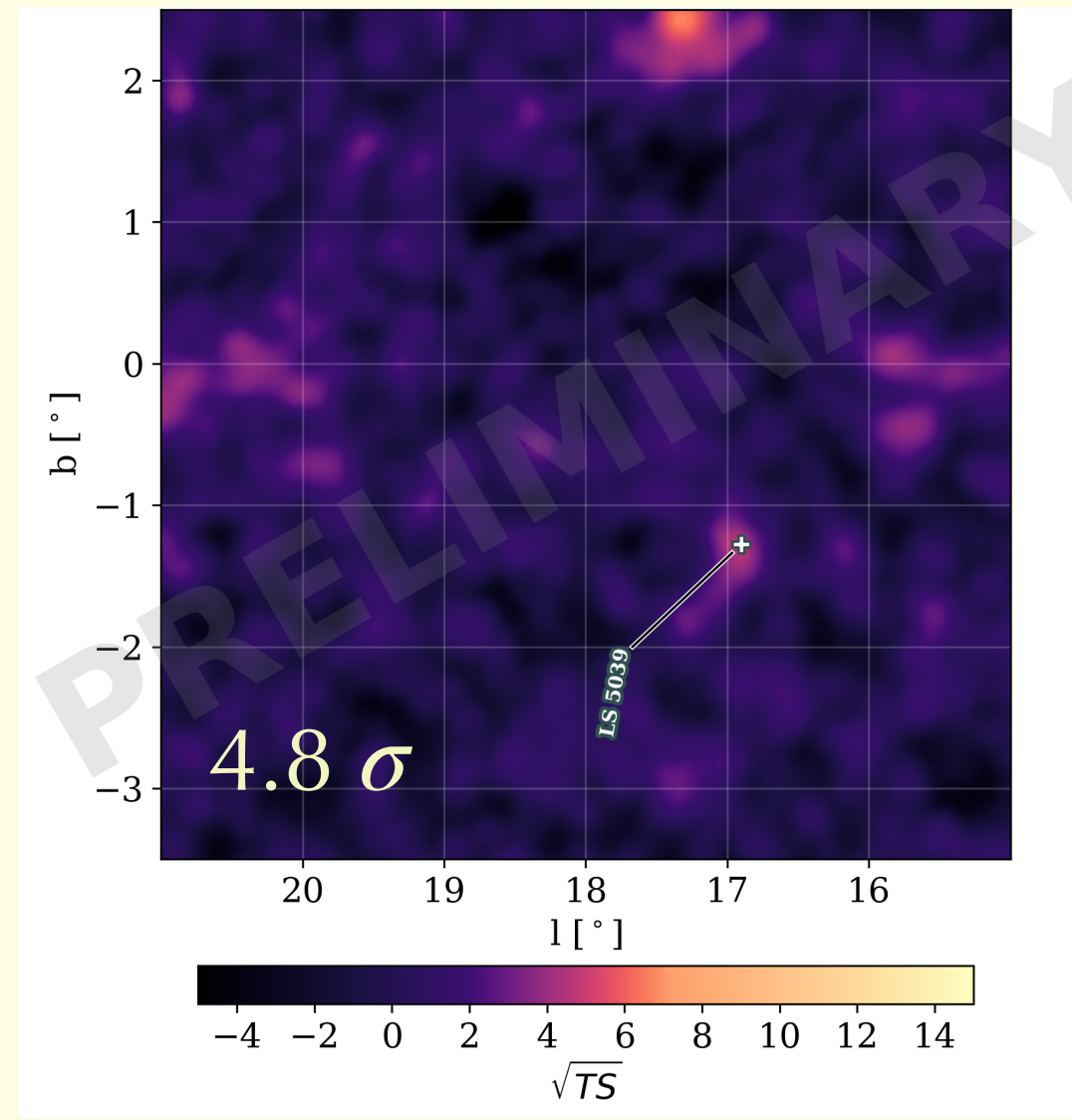
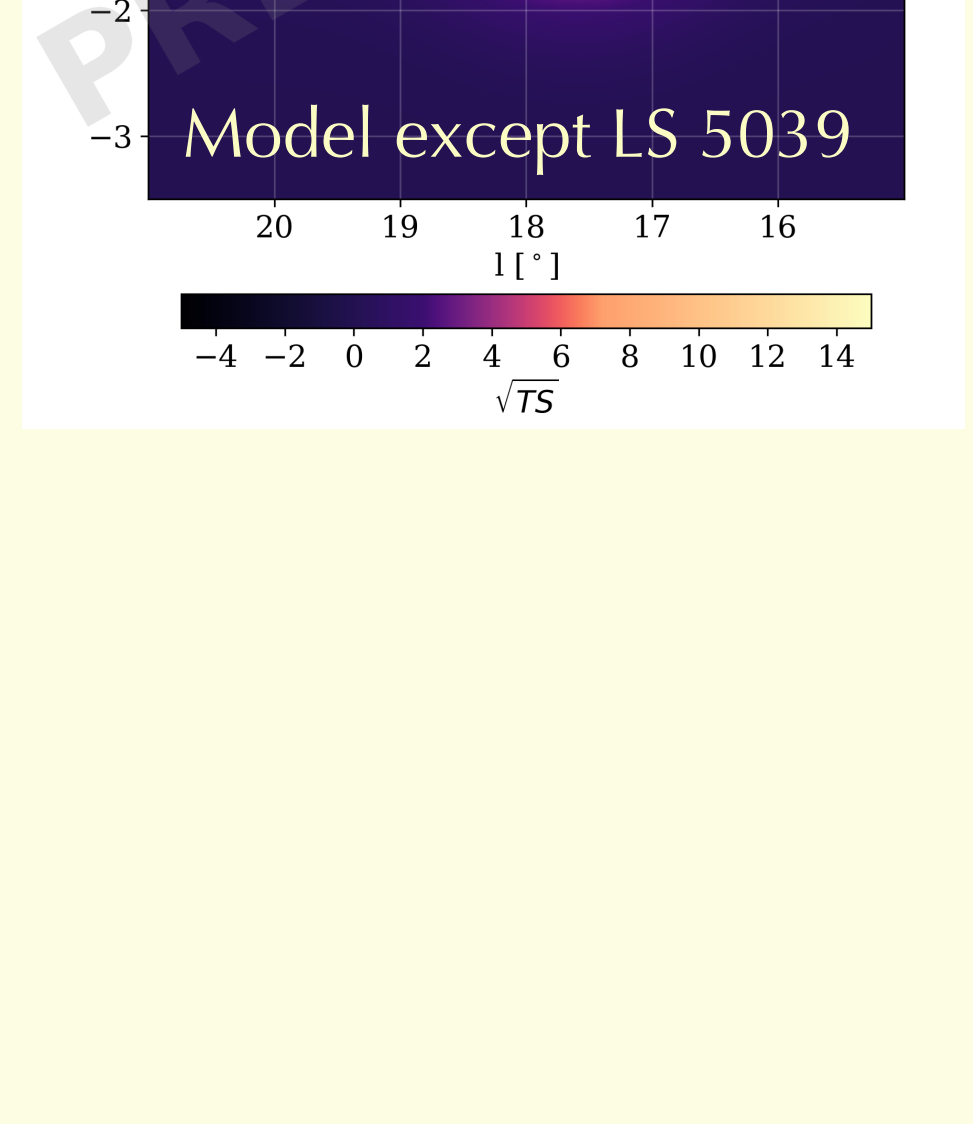
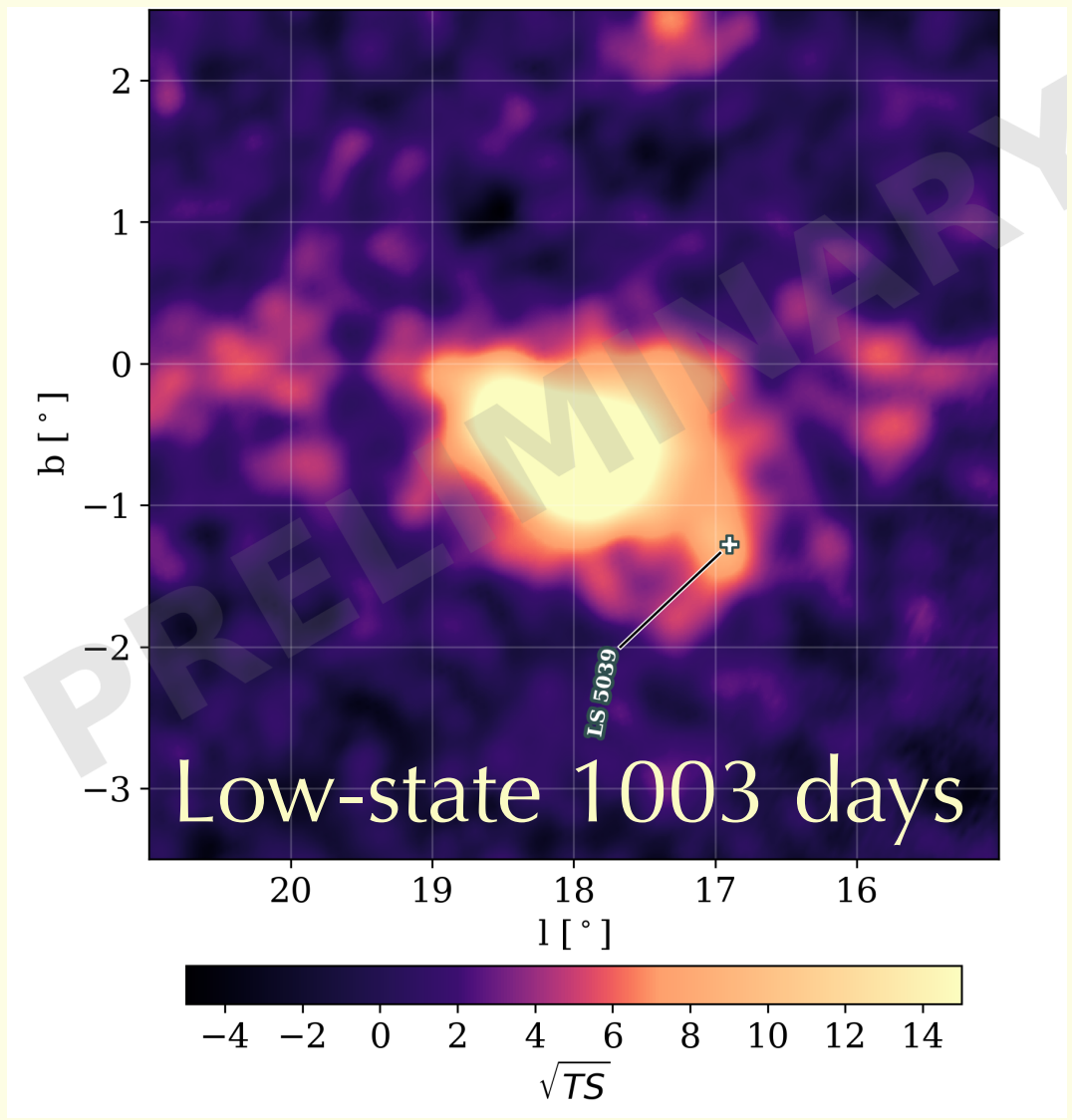
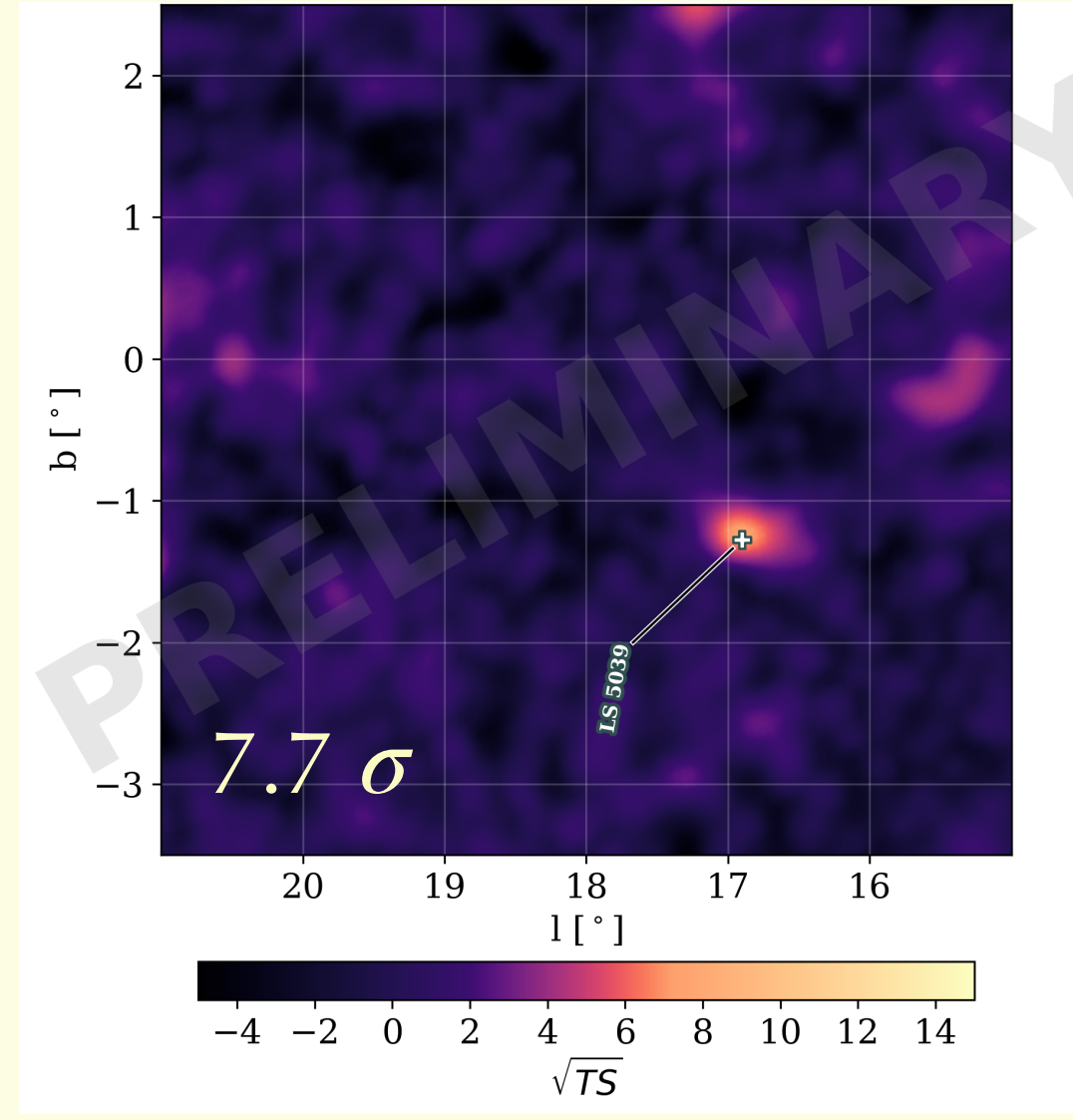
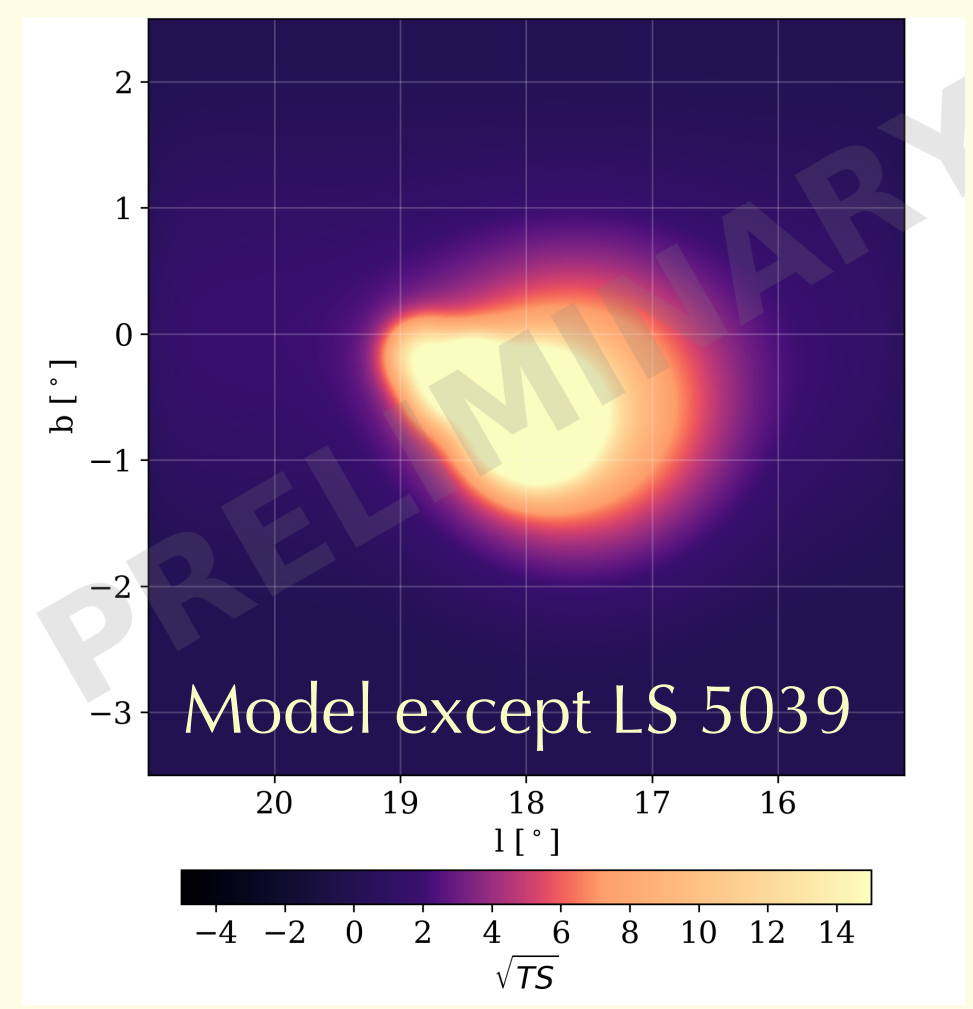
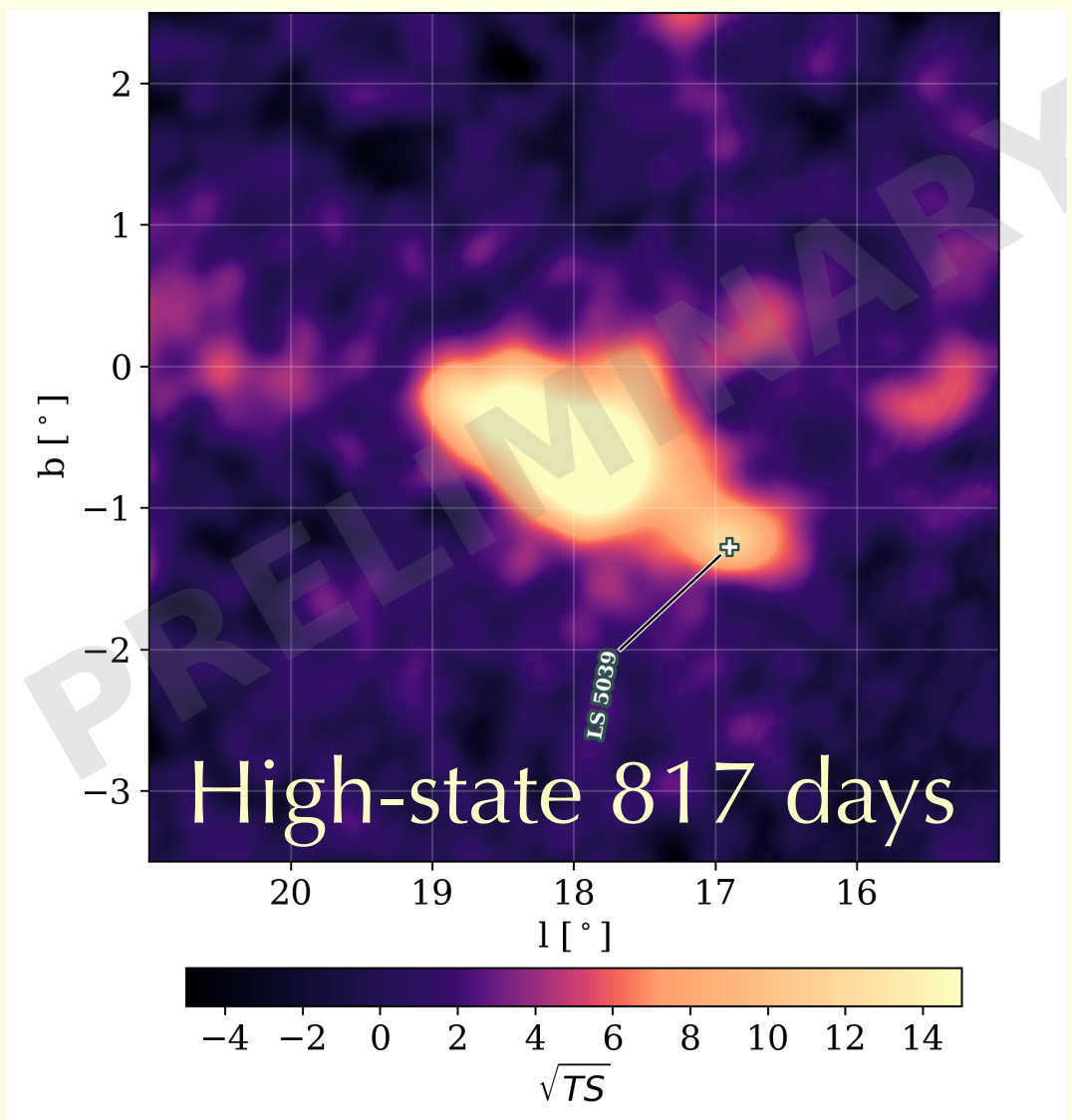


Analysis



Credits: H.E.S.S. *A&A*, 460 3 (2006) 743-749

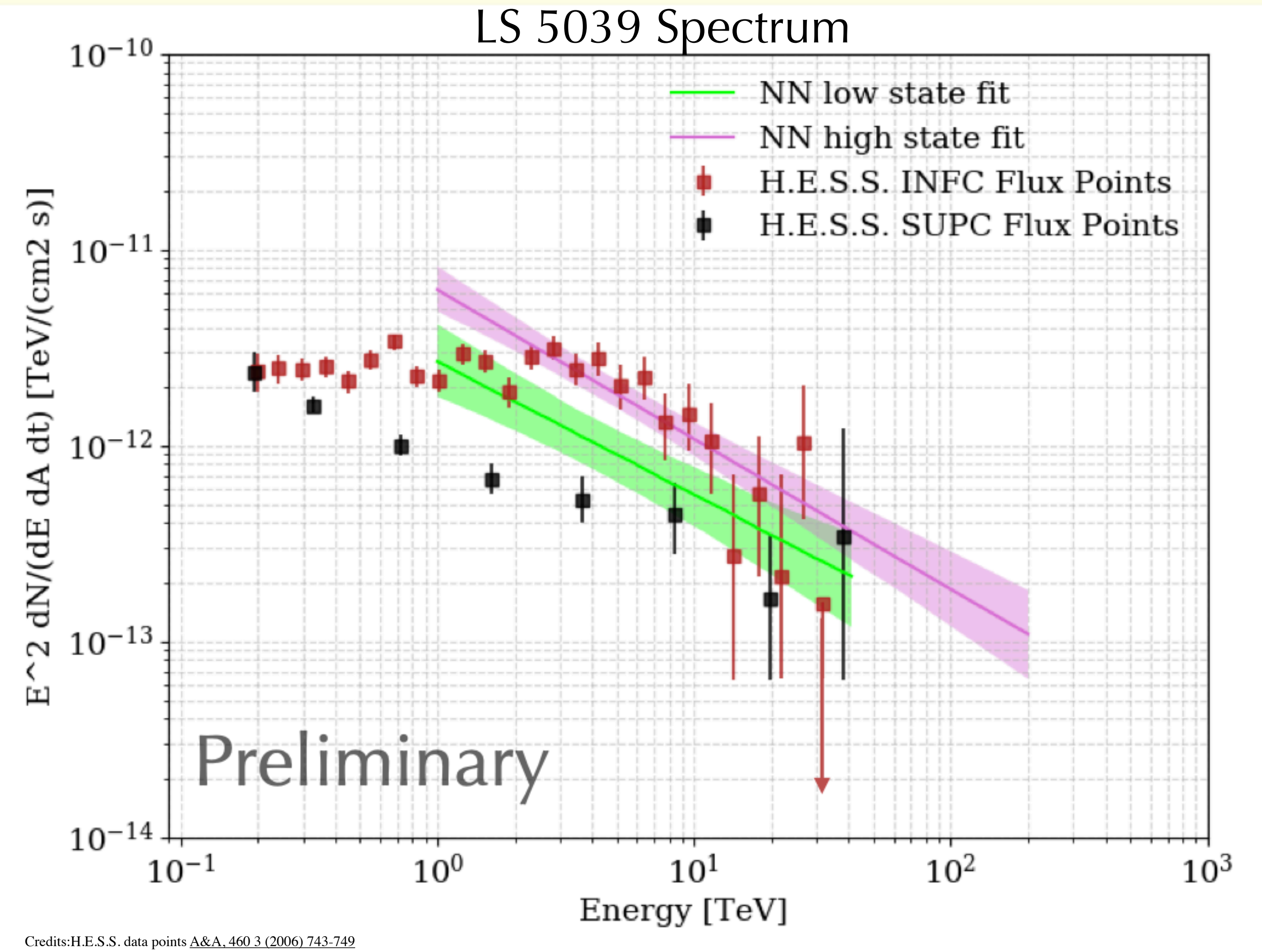
# Study the Flux Modulation



- High state accumulate data at  $\sim 0.27\sigma$  per day

- Low state accumulate data at  $\sim 0.15\sigma$  per day

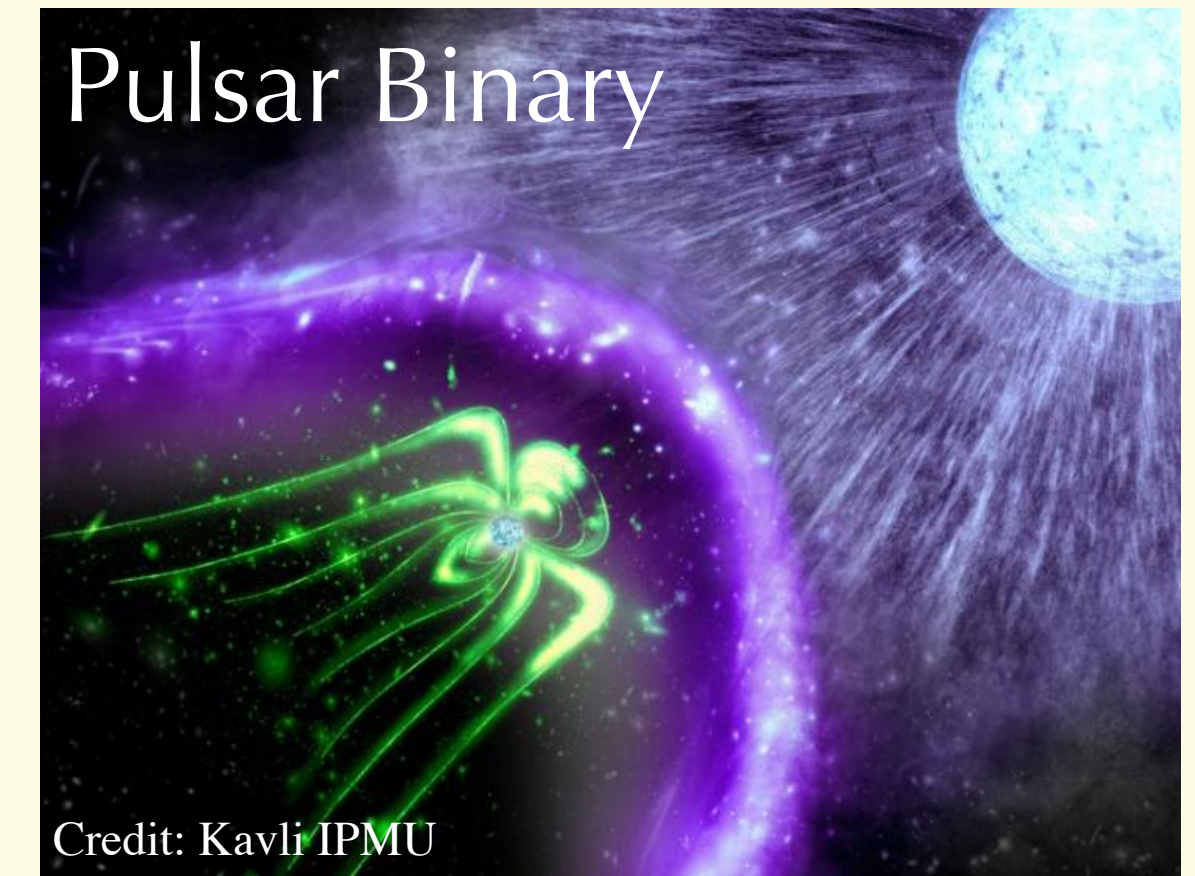
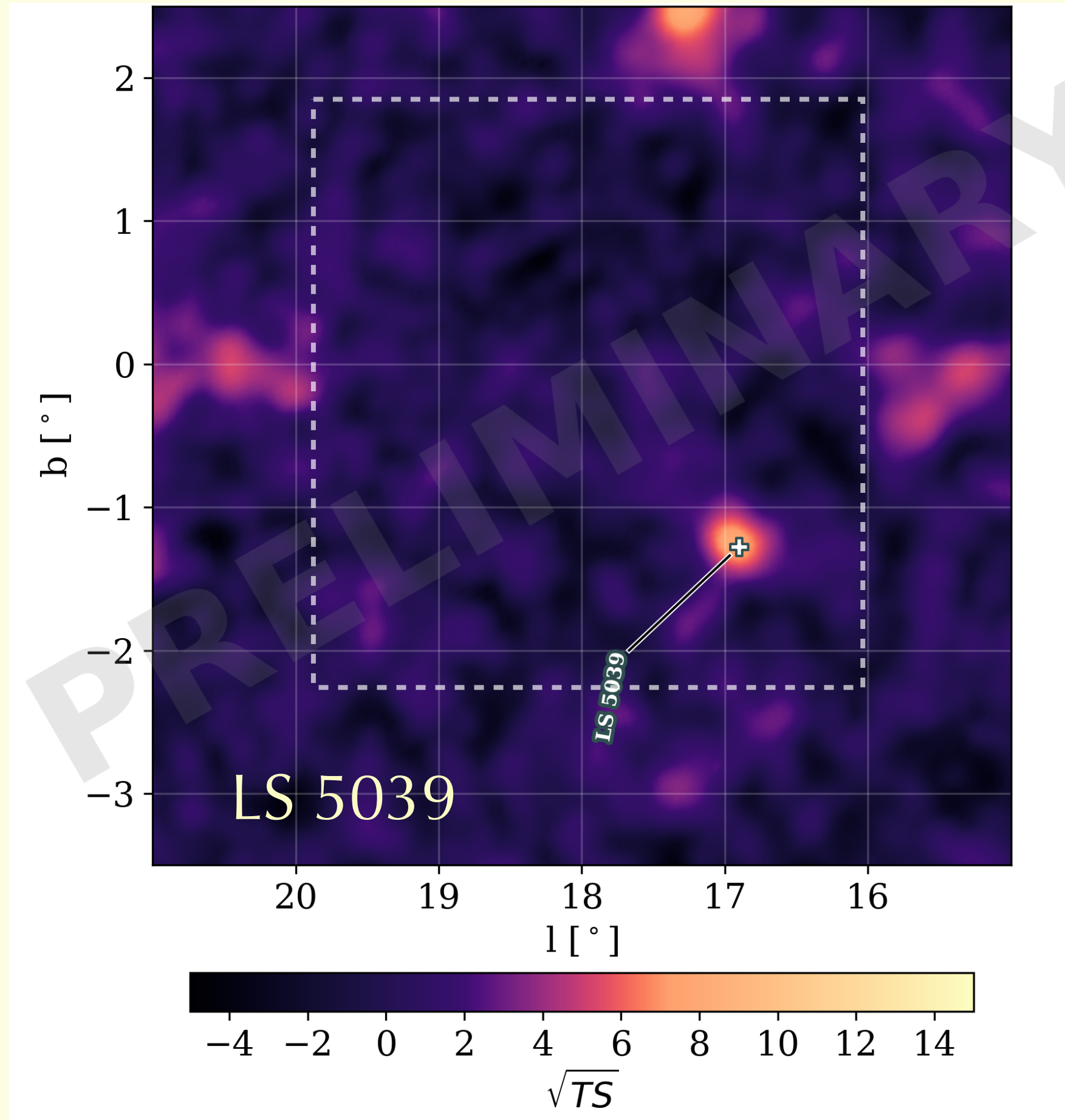




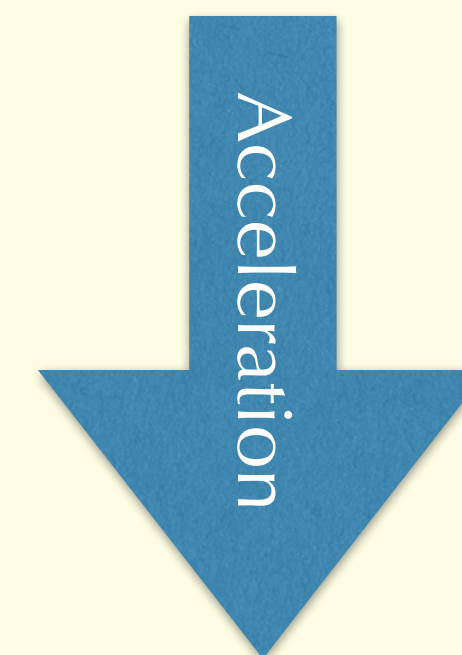
- HAWC see flux modulation at LS5039
- High state flux have a factor of two compare to low state flux
- High and low state have similar powerlaw index
- No cutoff found in both low state and High state maps yet

# High Energy Emission from LS 5039

## HAWC Observation



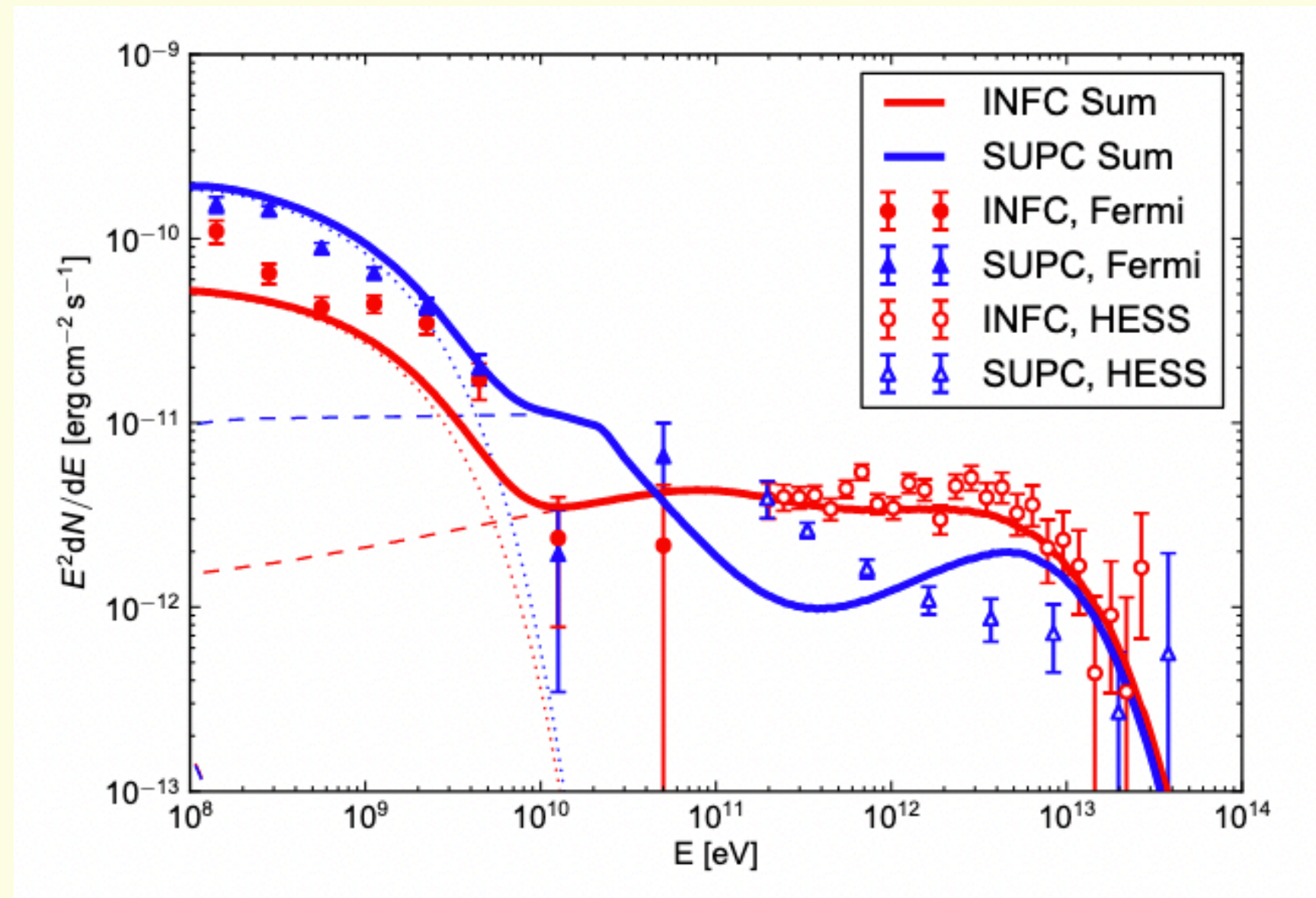
## Powerful Outflows



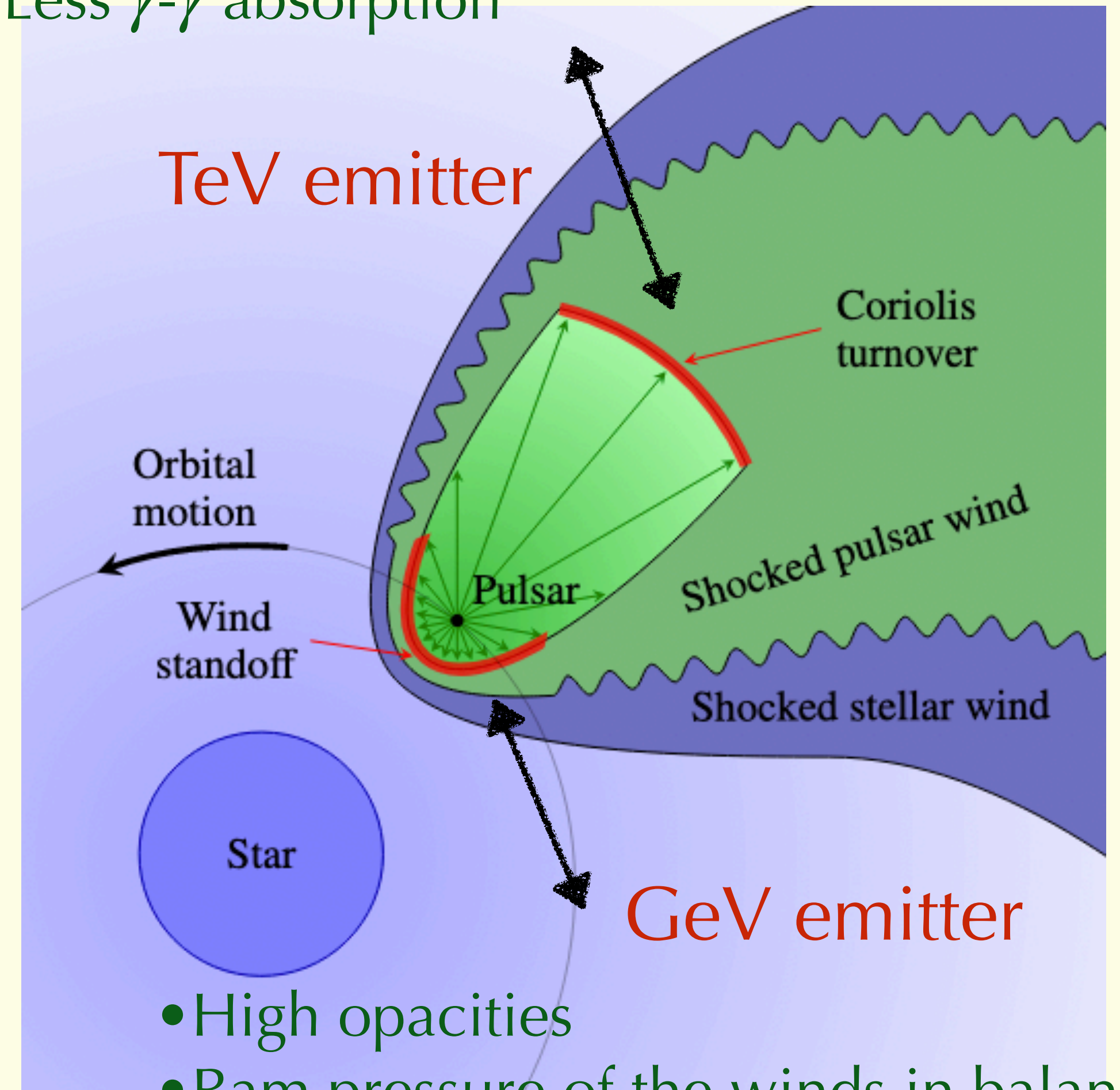
Jet formation through matter accretion

Interaction between pulsar wind and star wind

# Mirrored Modulation



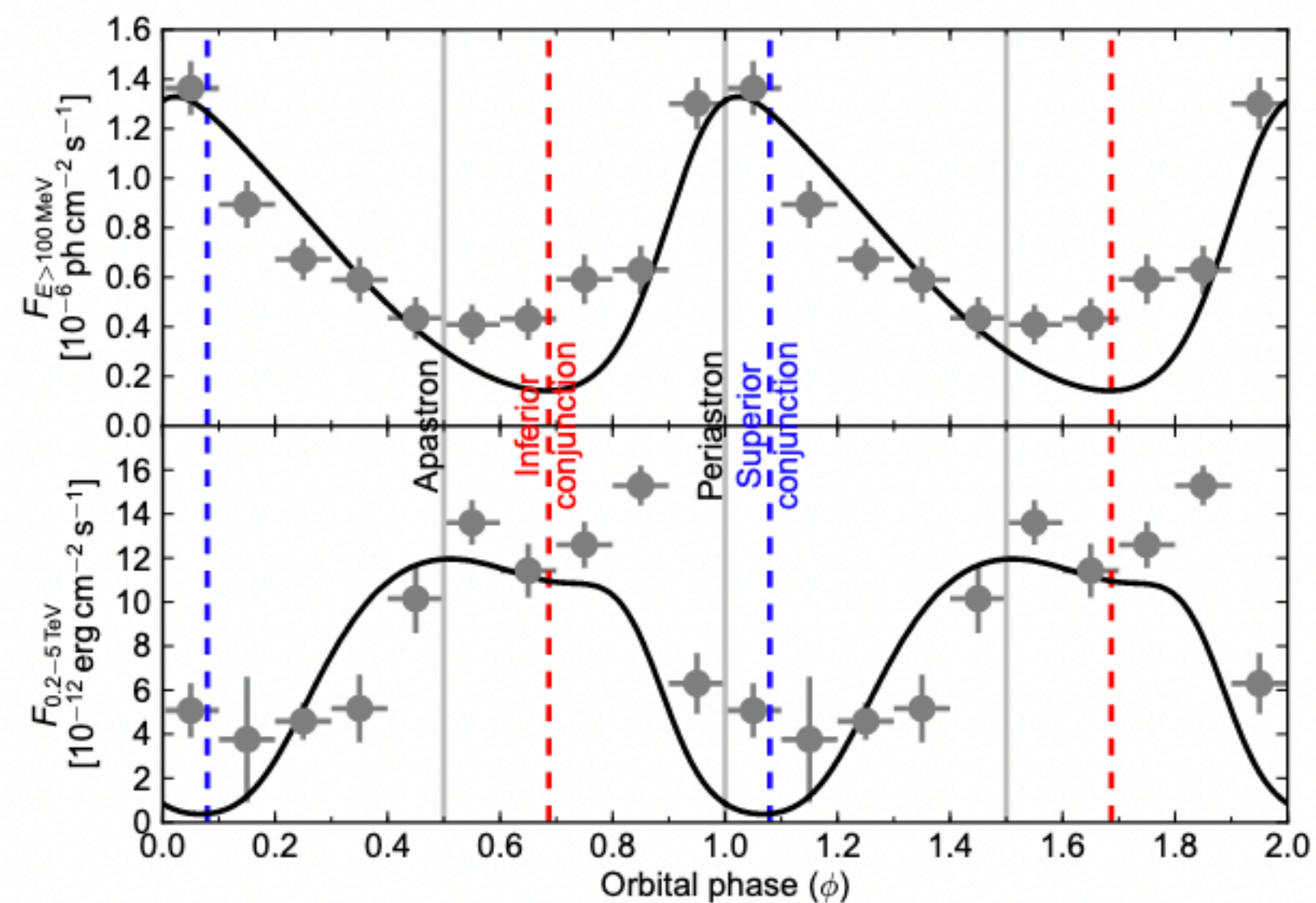
- Ram pressure of the stellar wind owing to Coriolis forces
- Less  $\gamma$ - $\gamma$  absorption



- High opacities
- Ram pressure of the winds in balance

Fermi

100 MeV to 30 GeV

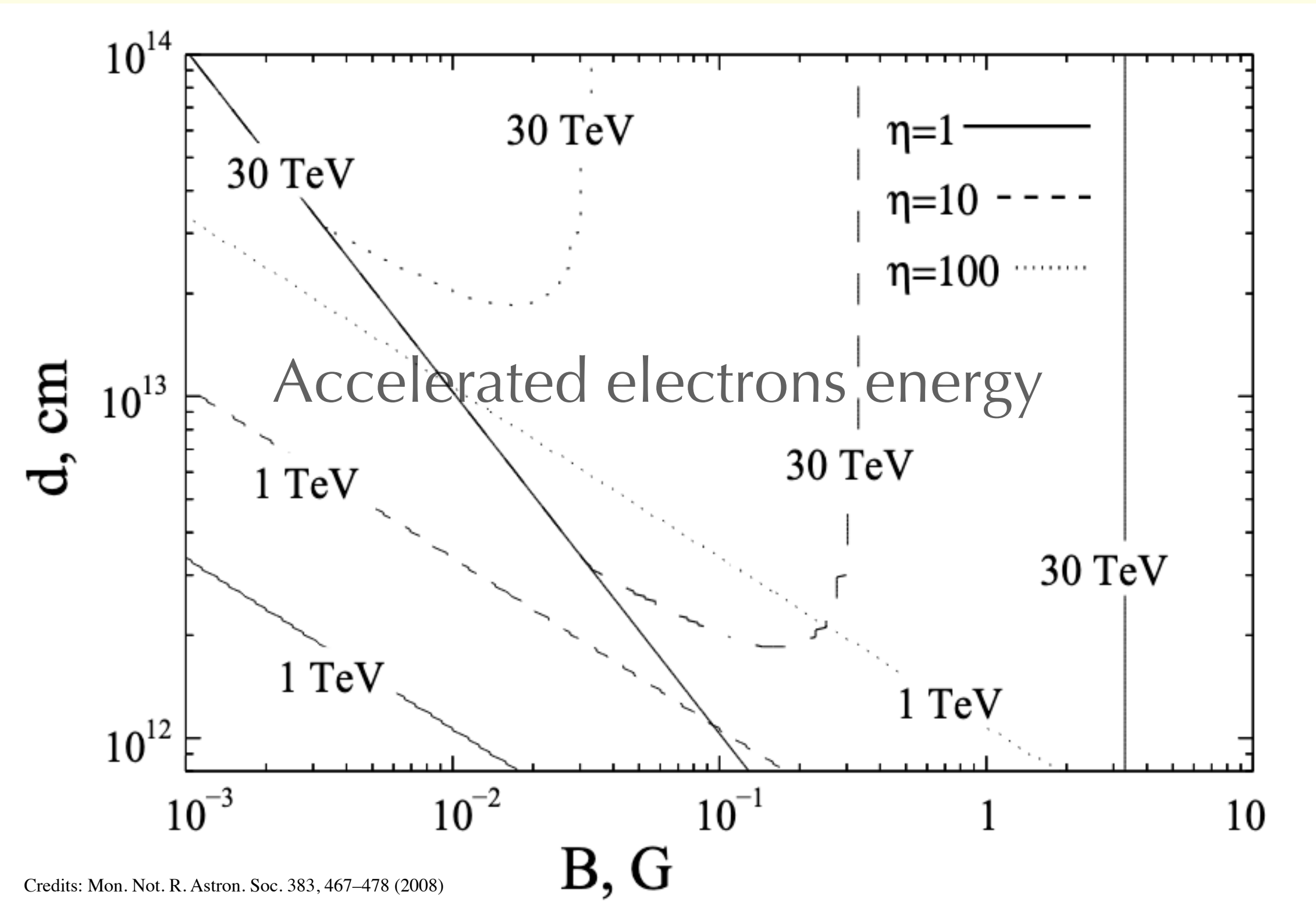


H.E.S.S.

200 GeV to 5 TeV

# High Energy Emission from LS 5039

Maximum  $\gamma$ -ray energy from HAWC can constrain



**How efficient the accelerator is?**

Very efficiently with  $\eta$  even  $< 10$

**Where is  $\gamma$ -ray produced?**

Not be located deep inside the binary

**What is the magnetic field?**

$B < 0.1$  G? Cannot explain if X-ray emission modulated as TeV emission

# Conclusion & Outlook

- HAWC able to disentangle the LS 5039 from complex J1825 region with improved angular resolution
- HAWC see flux modulation at LS 5039 respect to different orbital phase
- Spectrum analysis in both high and low states prefer pure powerlaw and similar index without cutoff
- Detailed studies at highest energy will help us understand the nature of LS 5039
- More results will come in our up coming publication

## Acknowledgments:



Thanks for your attention!  
Question are welcome!

Backup