

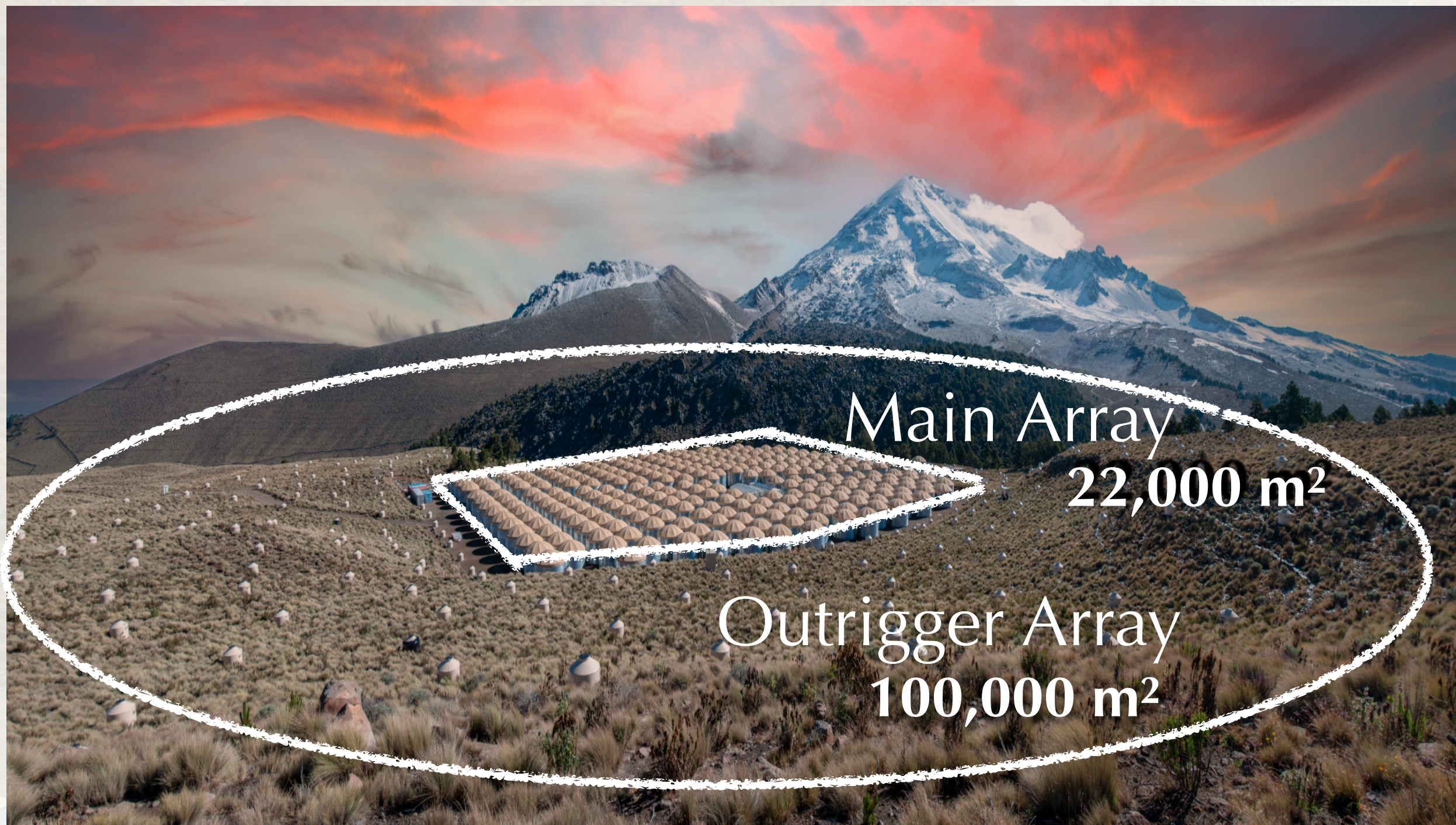
# Comparison Between PeVatron Candidates in the HAWC and LHAASO Data Sets

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For the HAWC collaboration , TEVPA 2022

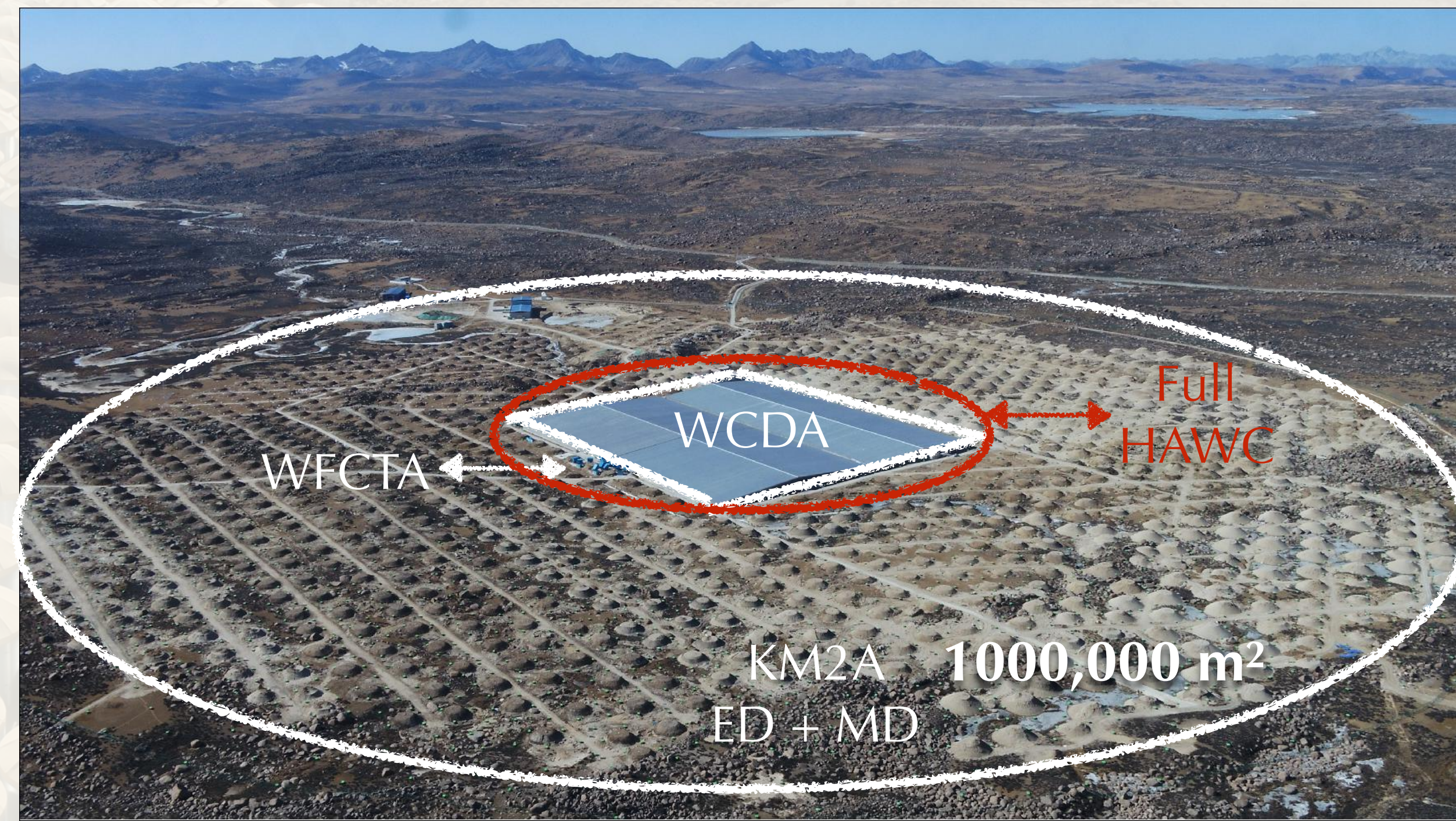


**Michigan  
Technological  
University**



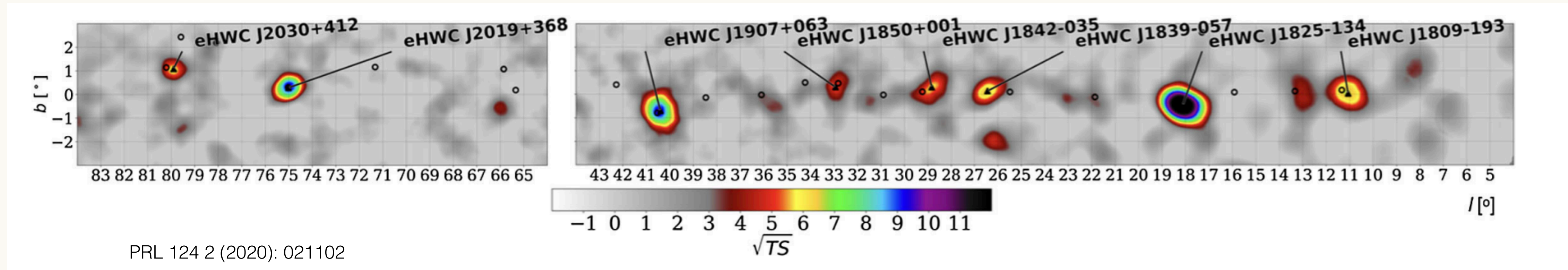


- Puebla, Mexico at 4100 meters a.s.l., 19°N
- 300 WCDs for Main array + 345 WCDs for outrigger
- $\gamma$ -ray 300 GeV - >100 TeV
- >95% duty circle
- Instantaneous field of view: ~ 2sr



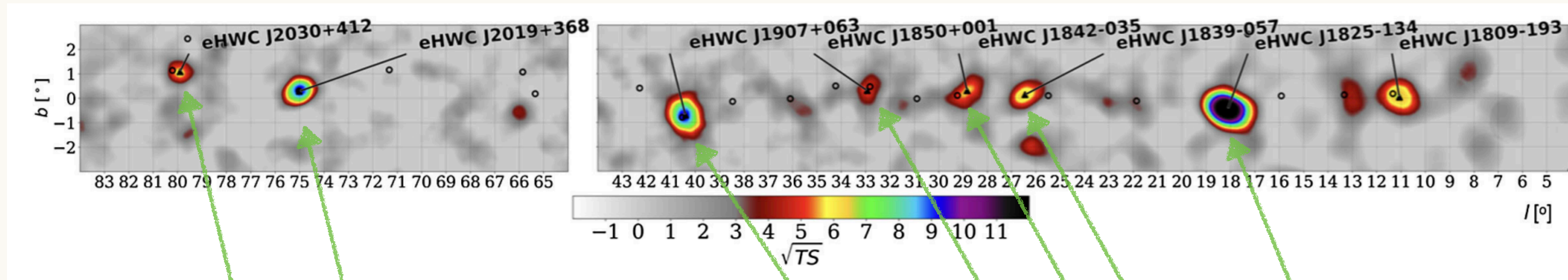
- Daocheng, China at 4410 meters a.s.l., 29° N
- 18 WFCTAs + 3 WCDA ponds + KM2A (5195ED, 1188MD)
- Cosmic-ray and  $\gamma$ -ray sub TeV - 1000 TeV
- >95% duty circle
- Instantaneous field of view: ~ 2sr

# Search PeVatrons In HAWC's FOV

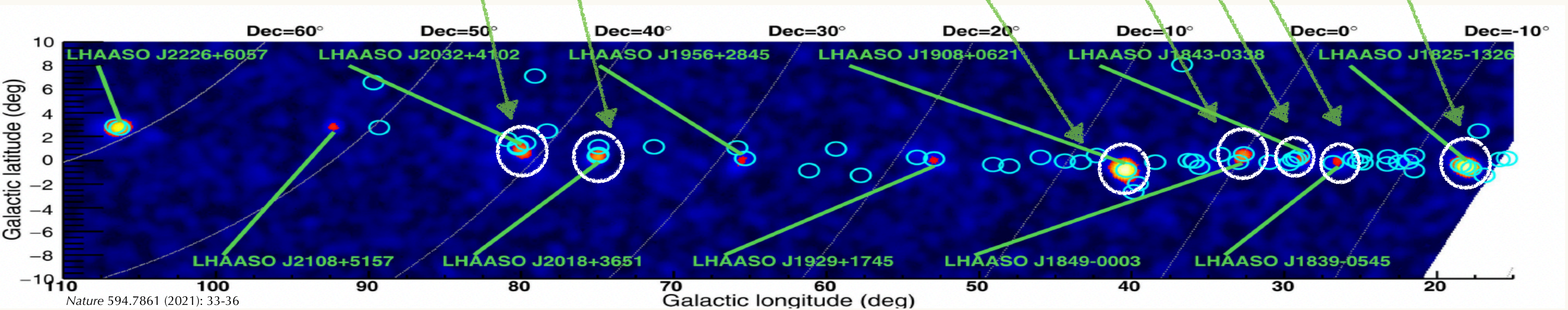


- ~1038.8 days of data
- June 2015 to July 2018
- 9 sources are detected with energy above 56 TeV

# Search PeVatrons In HAWC's FOV



## LHAASO-KM2A

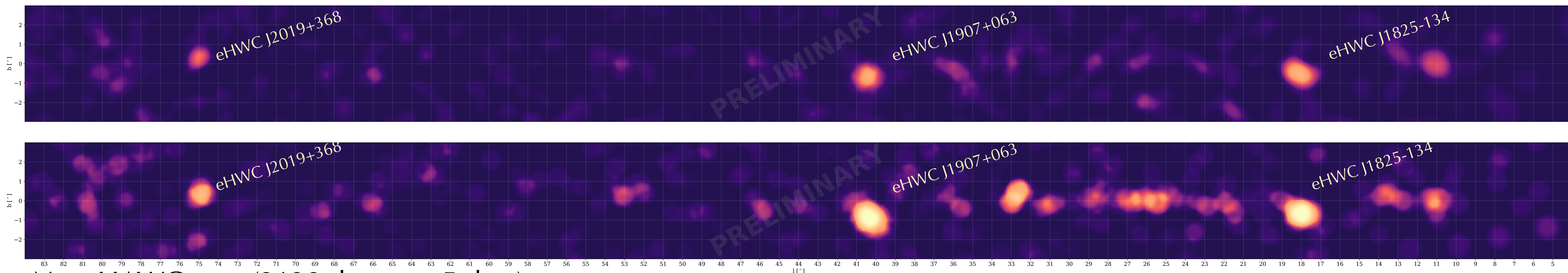


- LHAASO confirm all eHWC sources in the overlap region with energy  $> 100$  TeV

# Pushing to the highest energies ( $>100$ TeV)

High-energy paper (1038 days pass4 data)

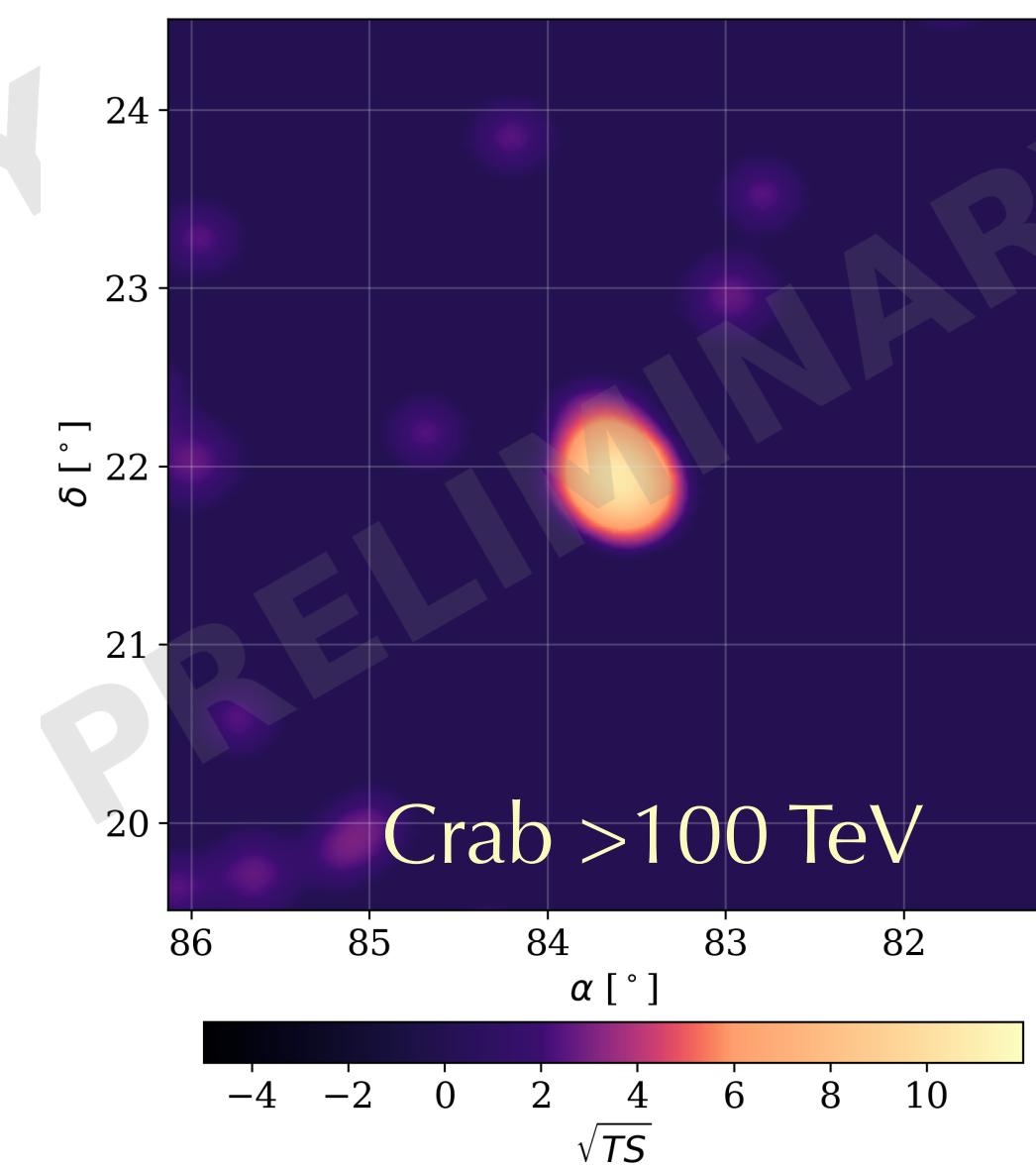
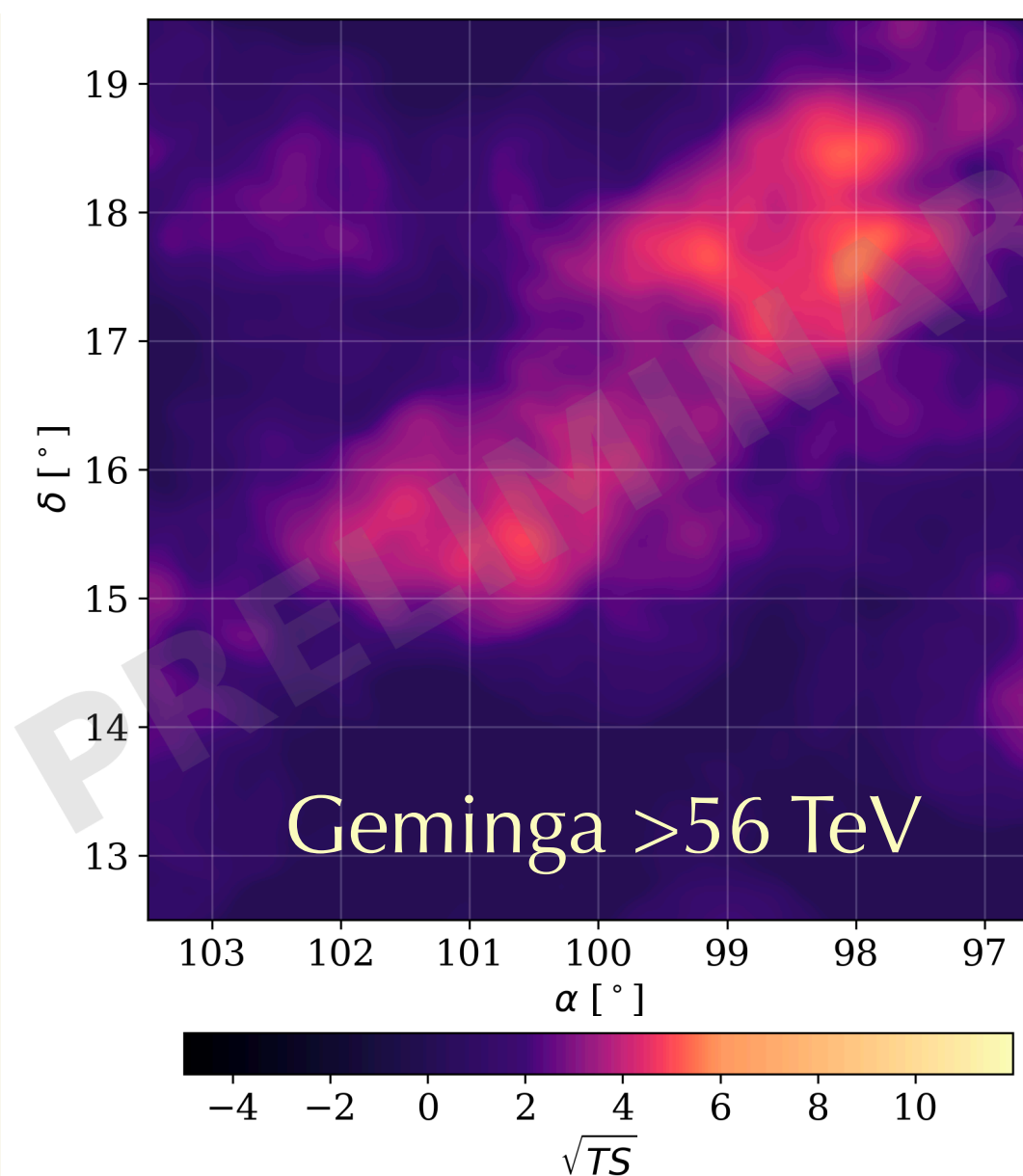
0.5° Extended Source Map



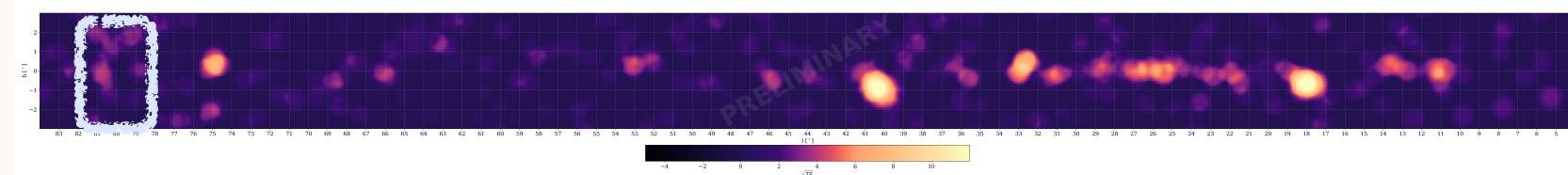
New HAWC map (2139 days pass5 data)

Better angular resolution; Wider FOV; Better background rejection

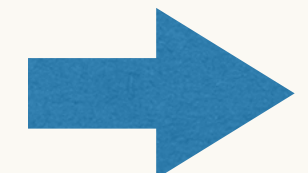
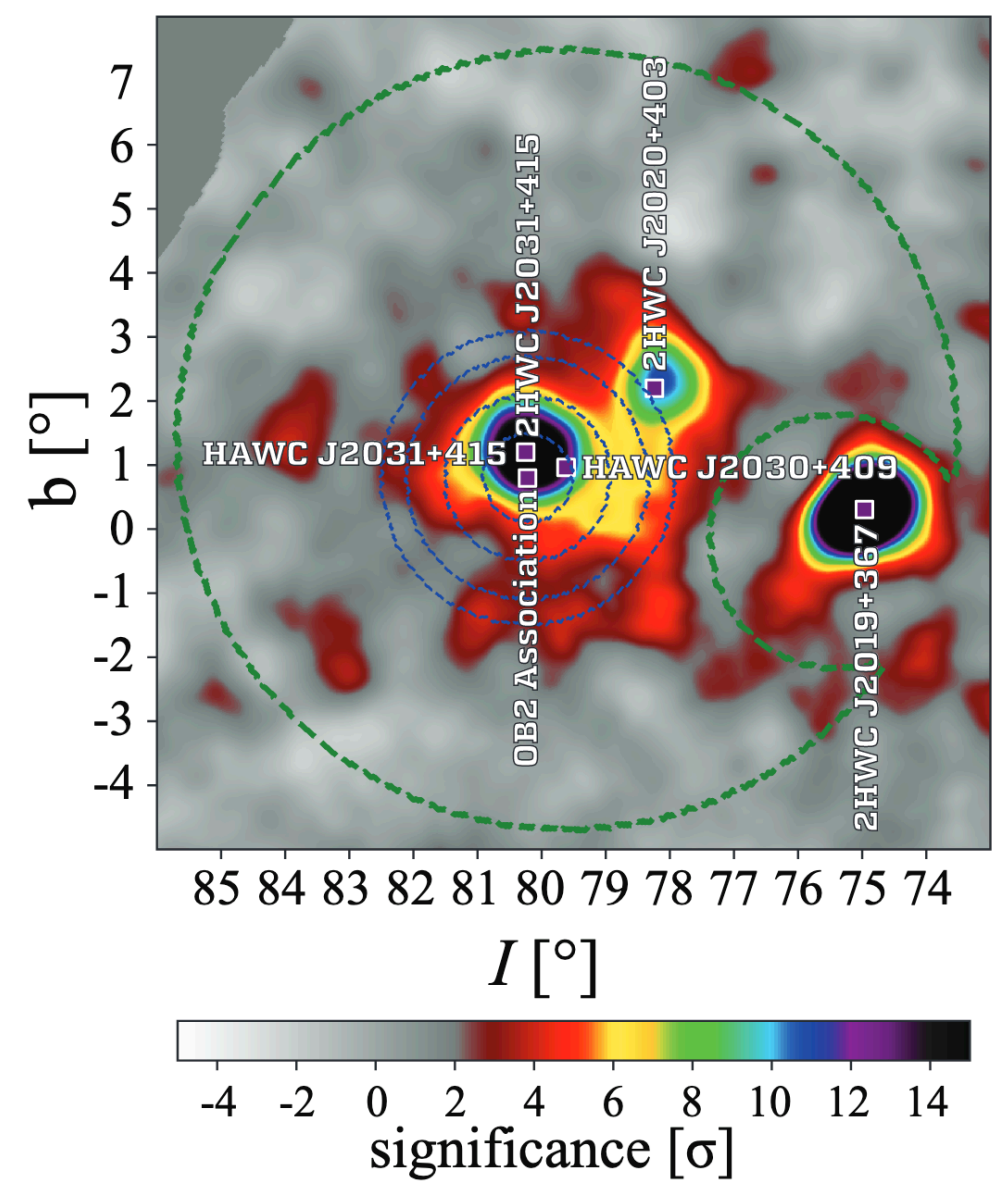
- New Map - **25** sources are identified above **56 TeV**
- New Map - **18** sources are identified above **100 TeV**
- New Map - **4** sources are identified above **177 TeV**
- The majority of UHE gamma-ray sources appear to be associated with PWN, but that detailed morphology studies to clearly distinguish associations are needed



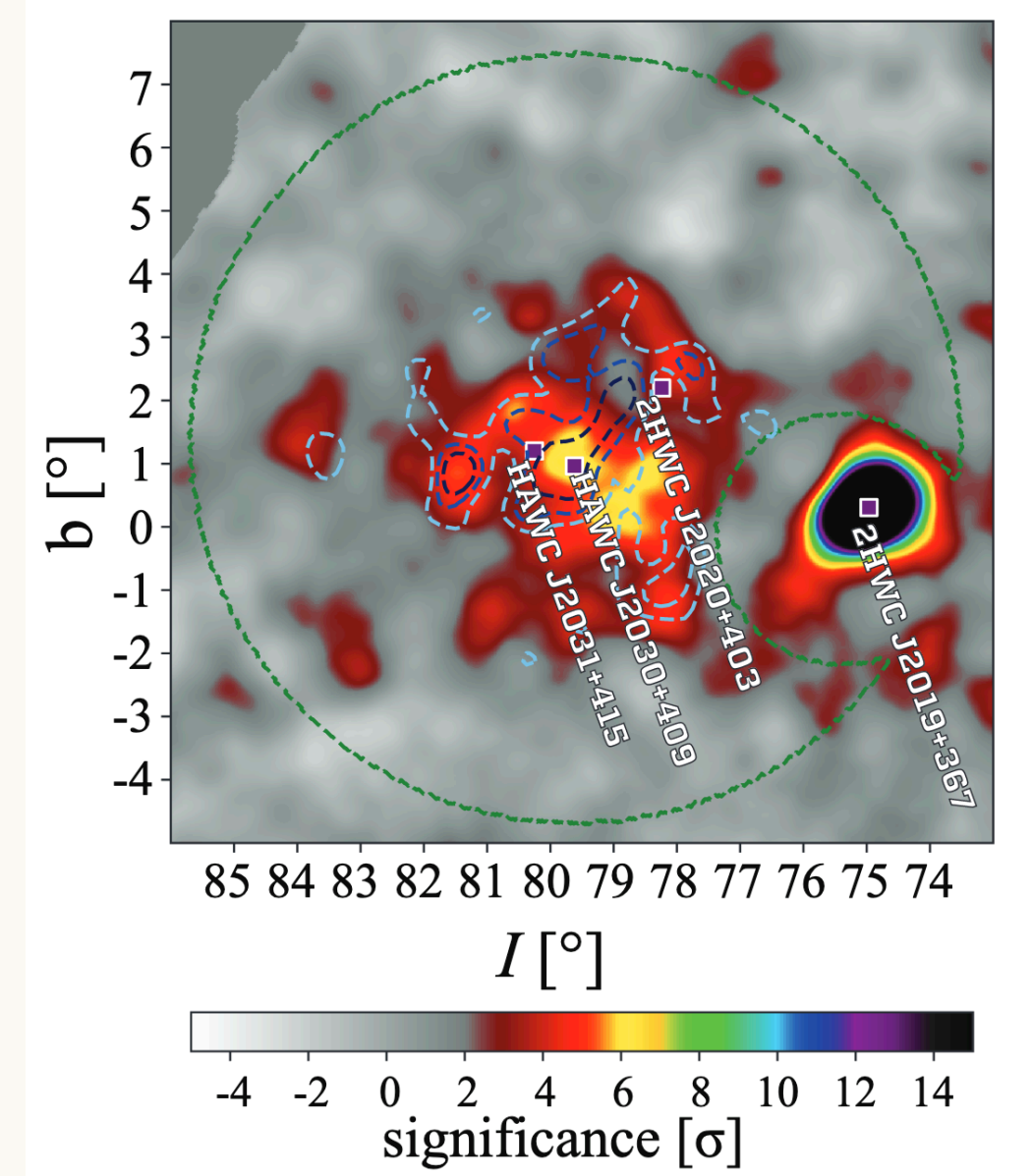
# Cygnus Cocoon



## HAWC Data



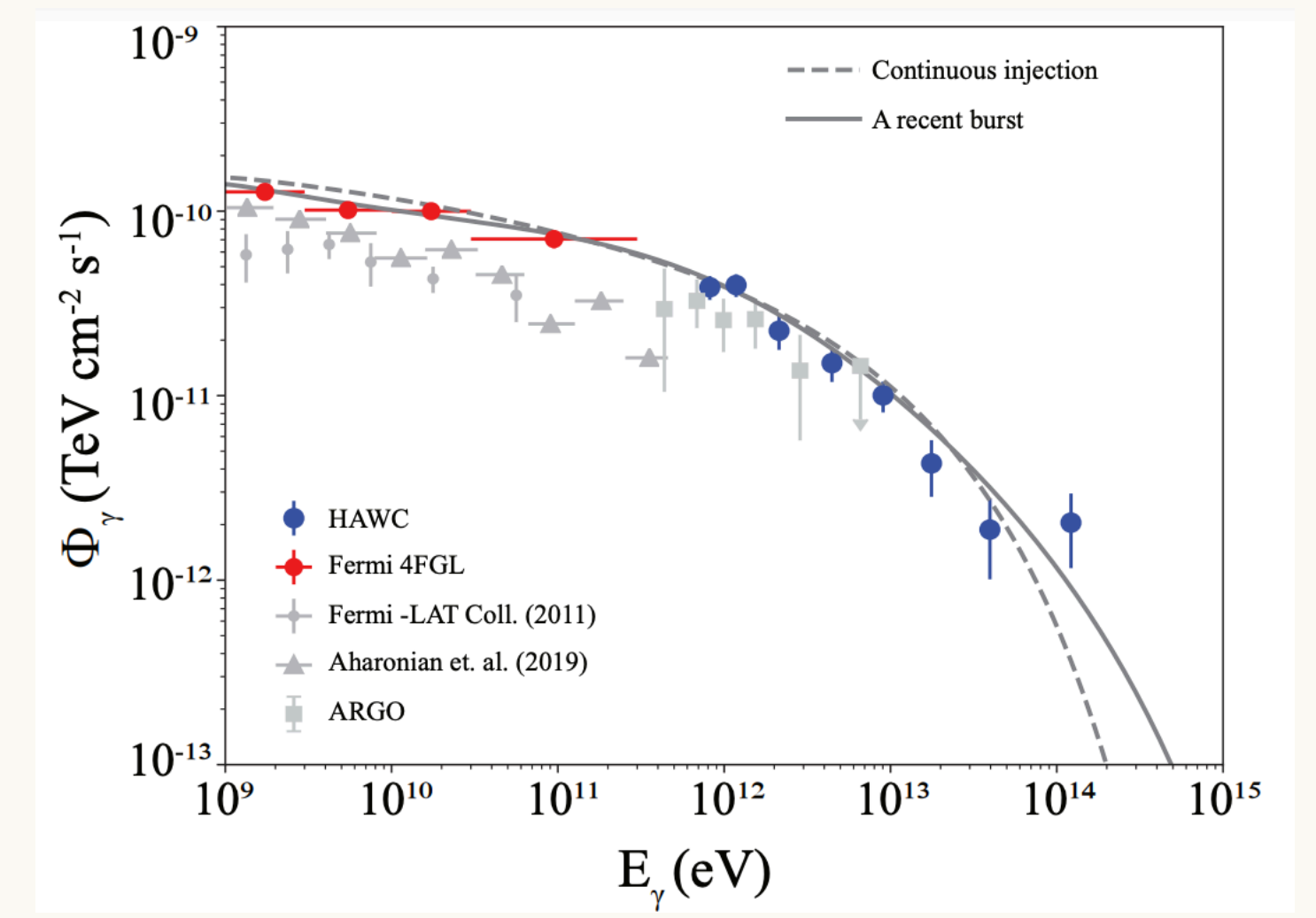
## Subtract PWN & Gamma Cygni



- Resolved three sources from 1343 days of data
  - PWN, Cocoon, Gamma Cygni
- First time revealed star forming region (Cyg OB2) could accelerate particles to PeV
- Cocoon spectrum have good agreement with ARGO and Fermi result
- LHAASO found the 1.4 PeV photon from Cocoon region (Looking forward with morphological study)
- HAWC flux at 100 TeV match with LHAASO measurement if using same morphology to calculate flux

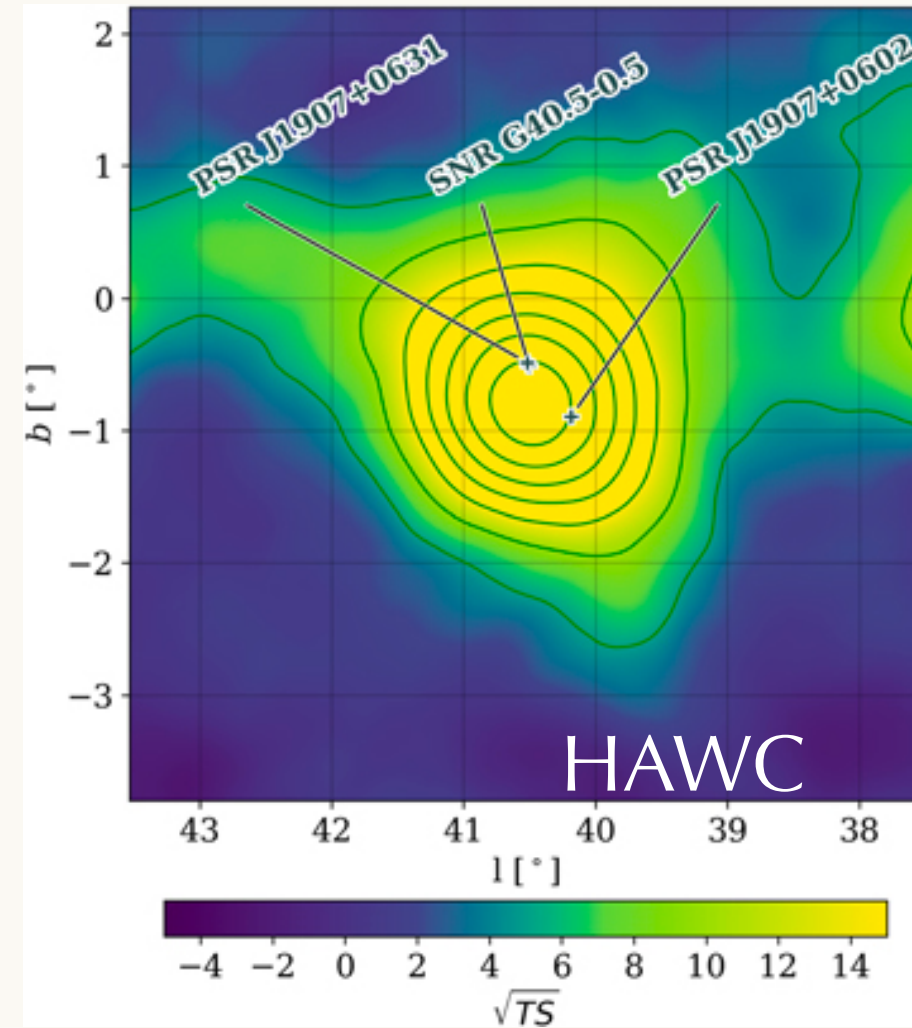
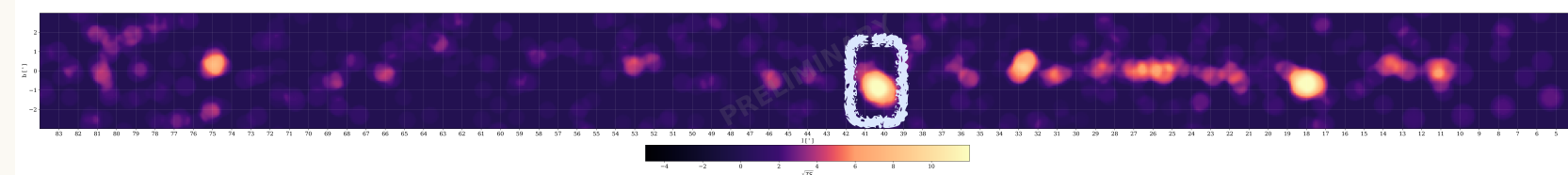
Nat Astron 5, 465–471 (2021)

- Pass 5 analysis agree with three components model
- Results coming soon!

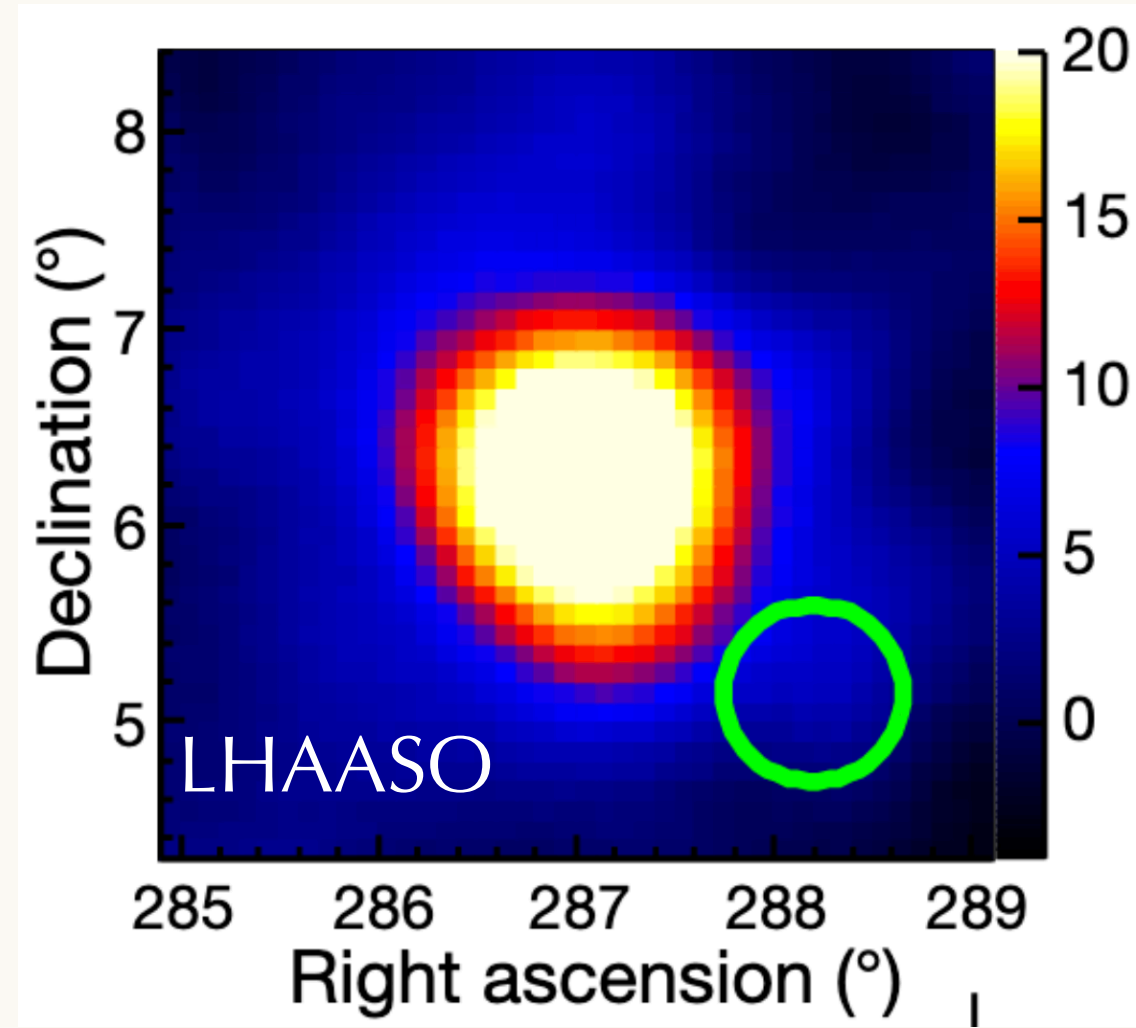




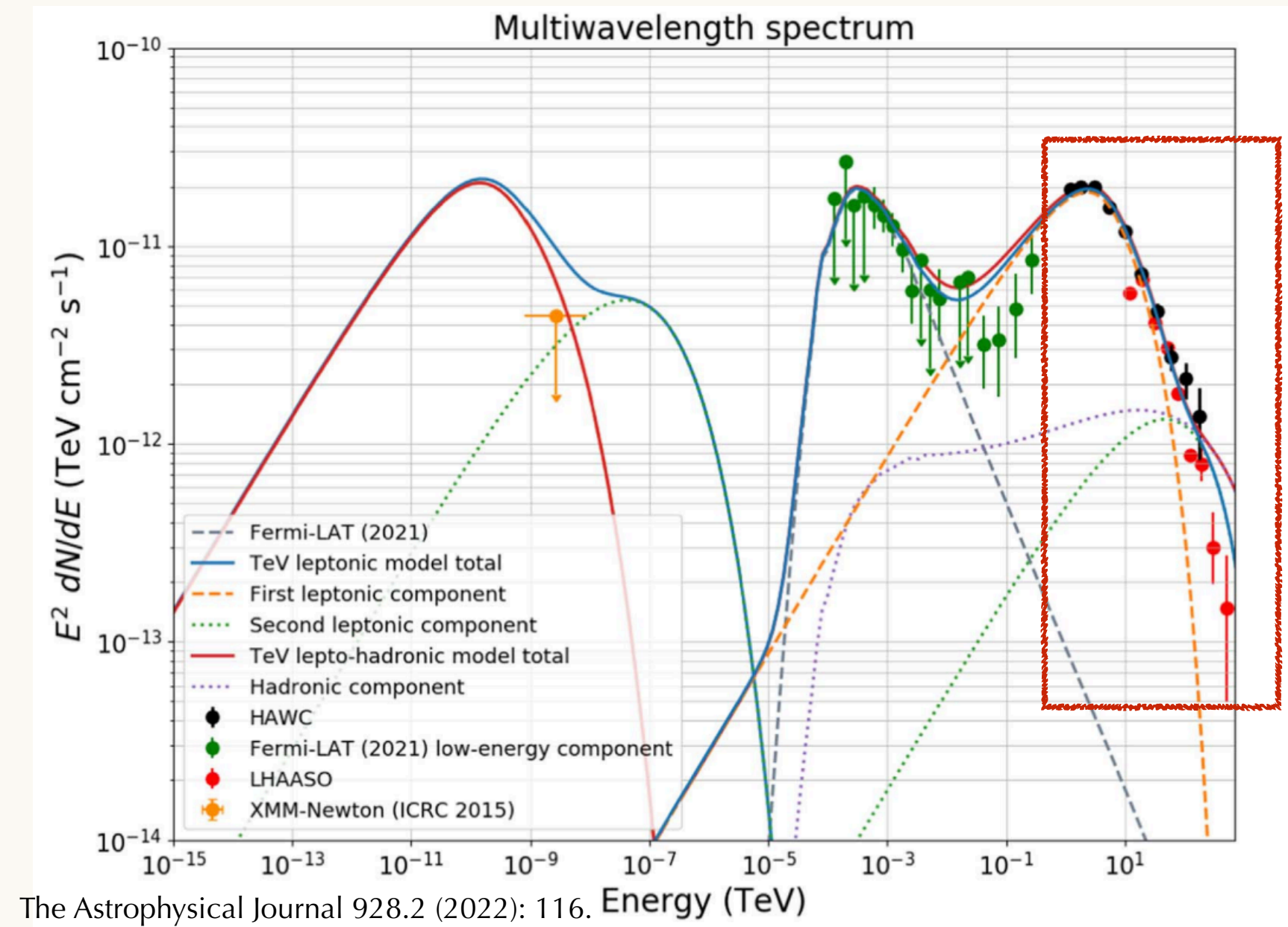
# eHWC J1907+63



The Astrophysical Journal 928.2 (2022): 116.



Nature 594.7861 (2021): 33-36

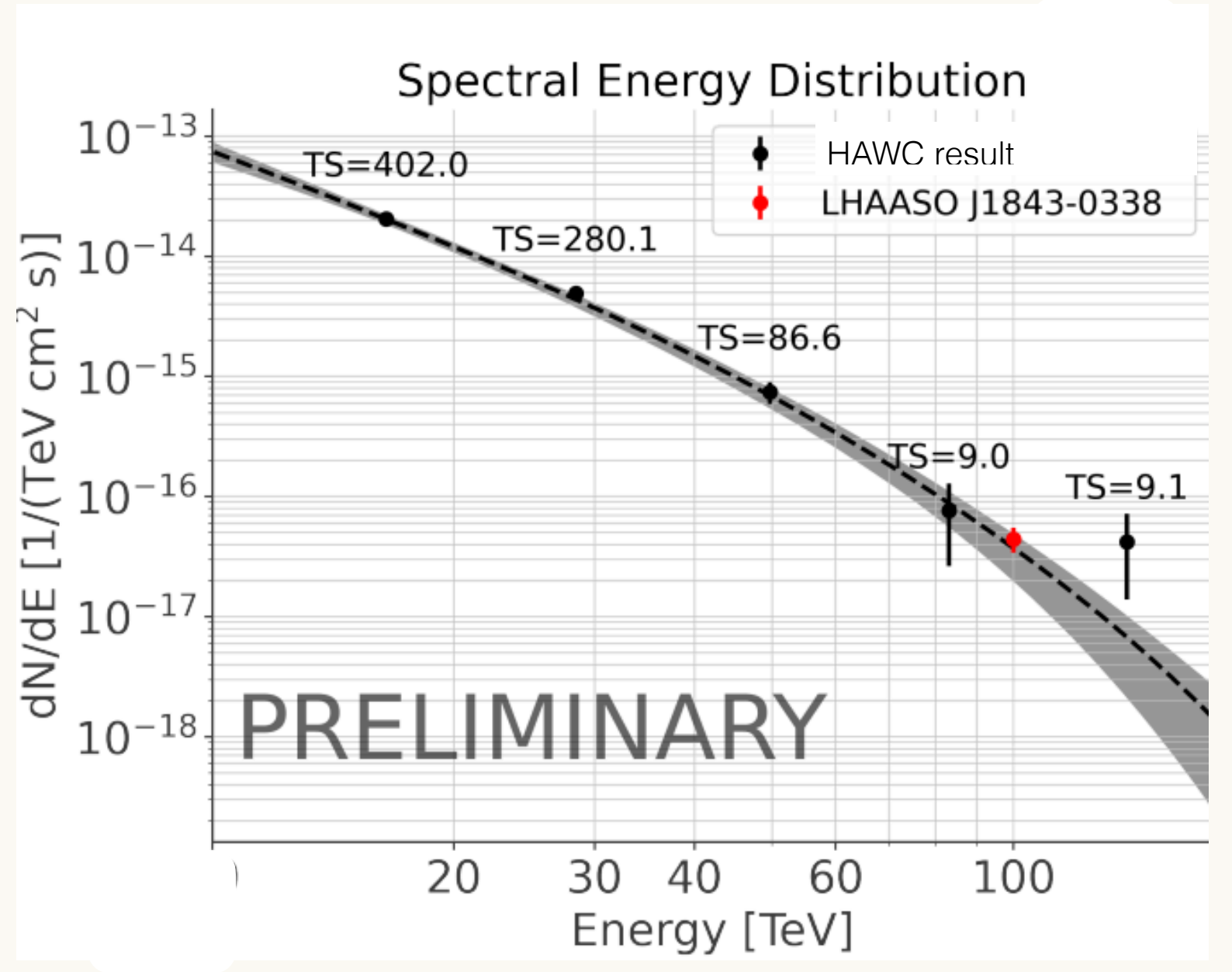
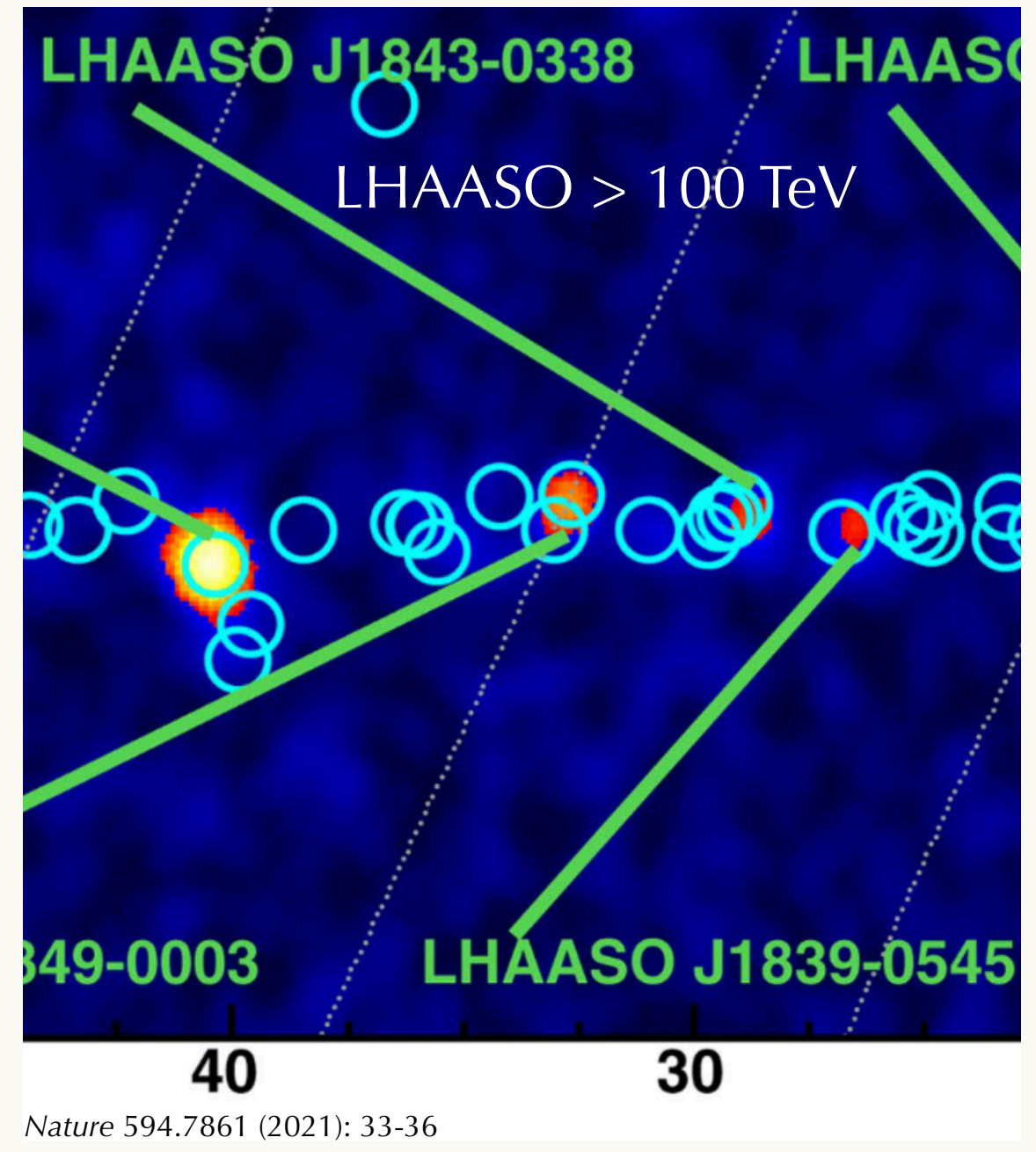
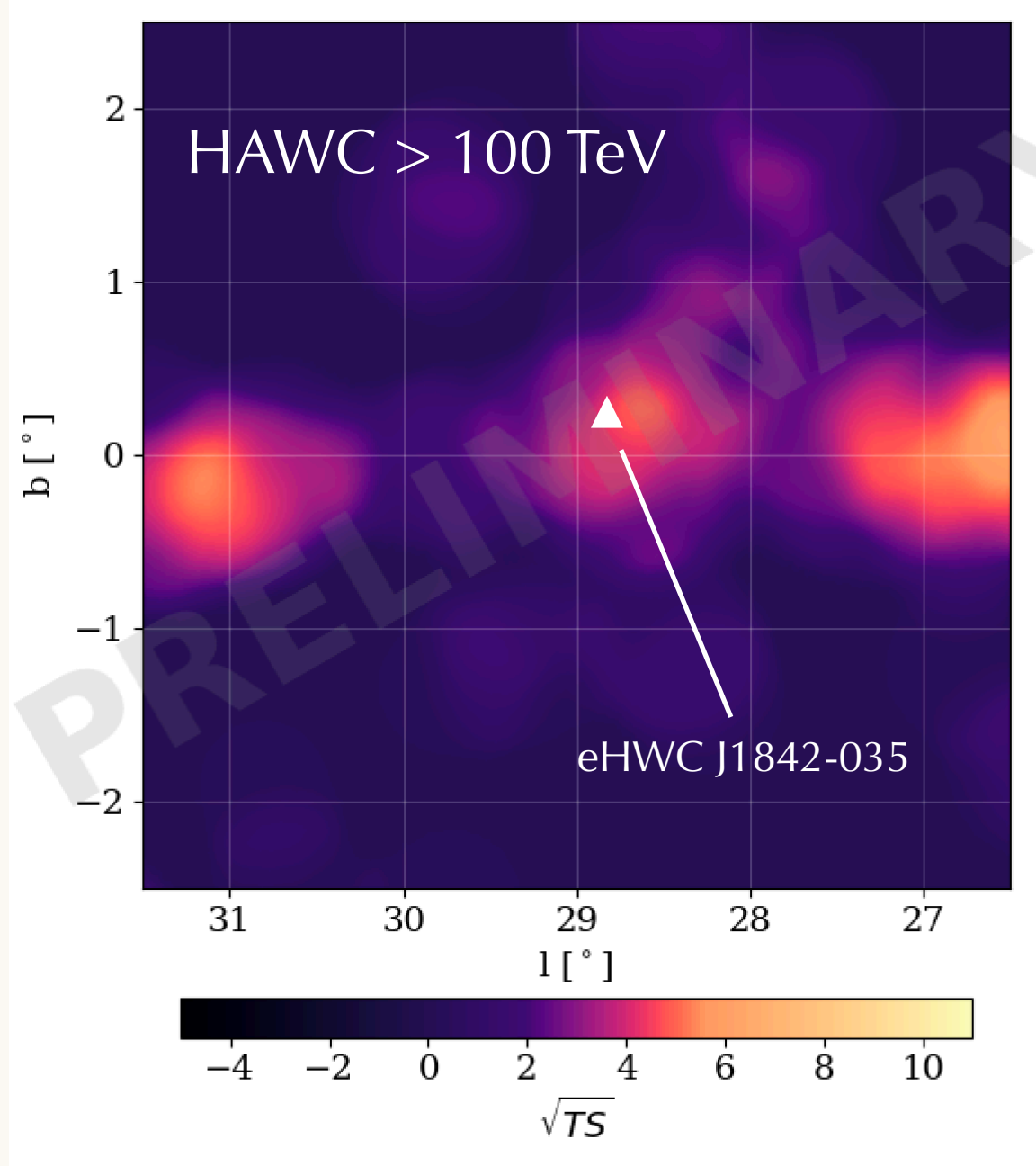
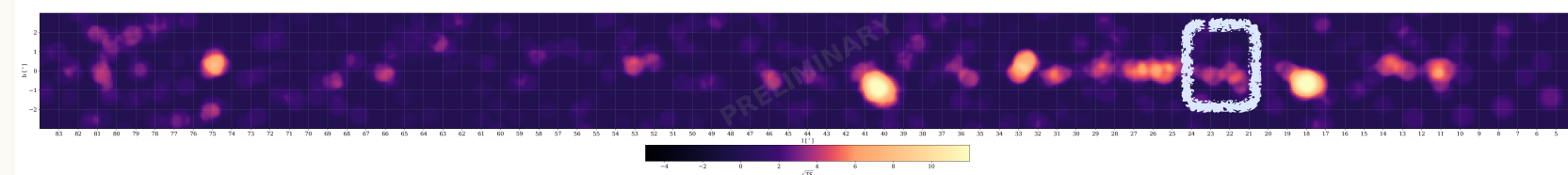


The Astrophysical Journal 928.2 (2022): 116.

HAWC & LHAASO flux points from 1 TeV to 440 TeV

|                                | HAWC   | LHAASO   |
|--------------------------------|--|--|
| <b>Location</b>                | R.A. 287.05° Dec. 6.39°  | R.A. 287.05° Dec. 6.35°  |
| <b>Maximum measured energy</b> | >200 TeV   | 440 TeV  |
| <b>Origin of TeV emission</b>  | Suggest leptonic in nature. Powered by PSR J1907+0602<br><br>one-population leptonic, two-population leptonic, and lepton-hadronic allowed | Both hadronic and leptonic origin are tested in LHAASO data<br><br>No preference in current data |

# eHWC J1842-035

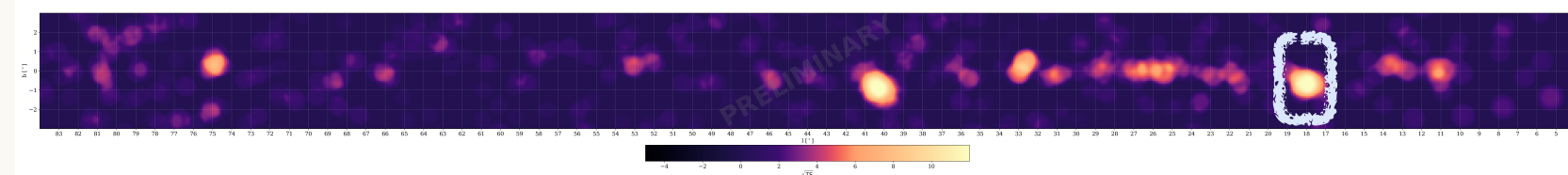


|                                | HAWC                     | LHAASO                   |
|--------------------------------|--------------------------|--------------------------|
| <b>Location</b>                | R.A. 280.72° Dec. -3.51° | R.A. 280.75° Dec. -3.65° |
| <b>Maximum measured energy</b> | >56 TeV                  | 260 TeV                  |
| <b>Morphology</b>              | 0.39° extension          | 0.3° extension template  |

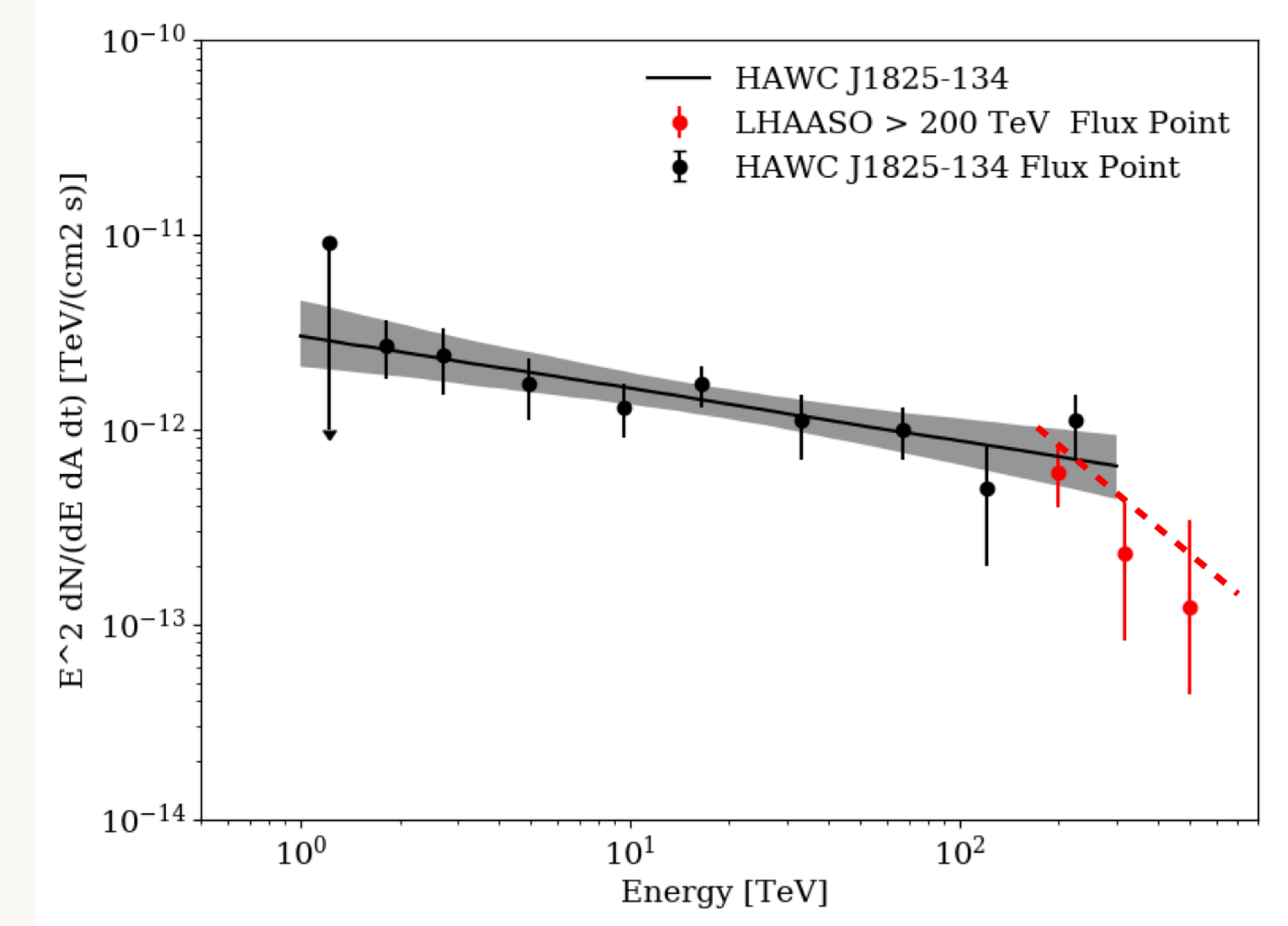
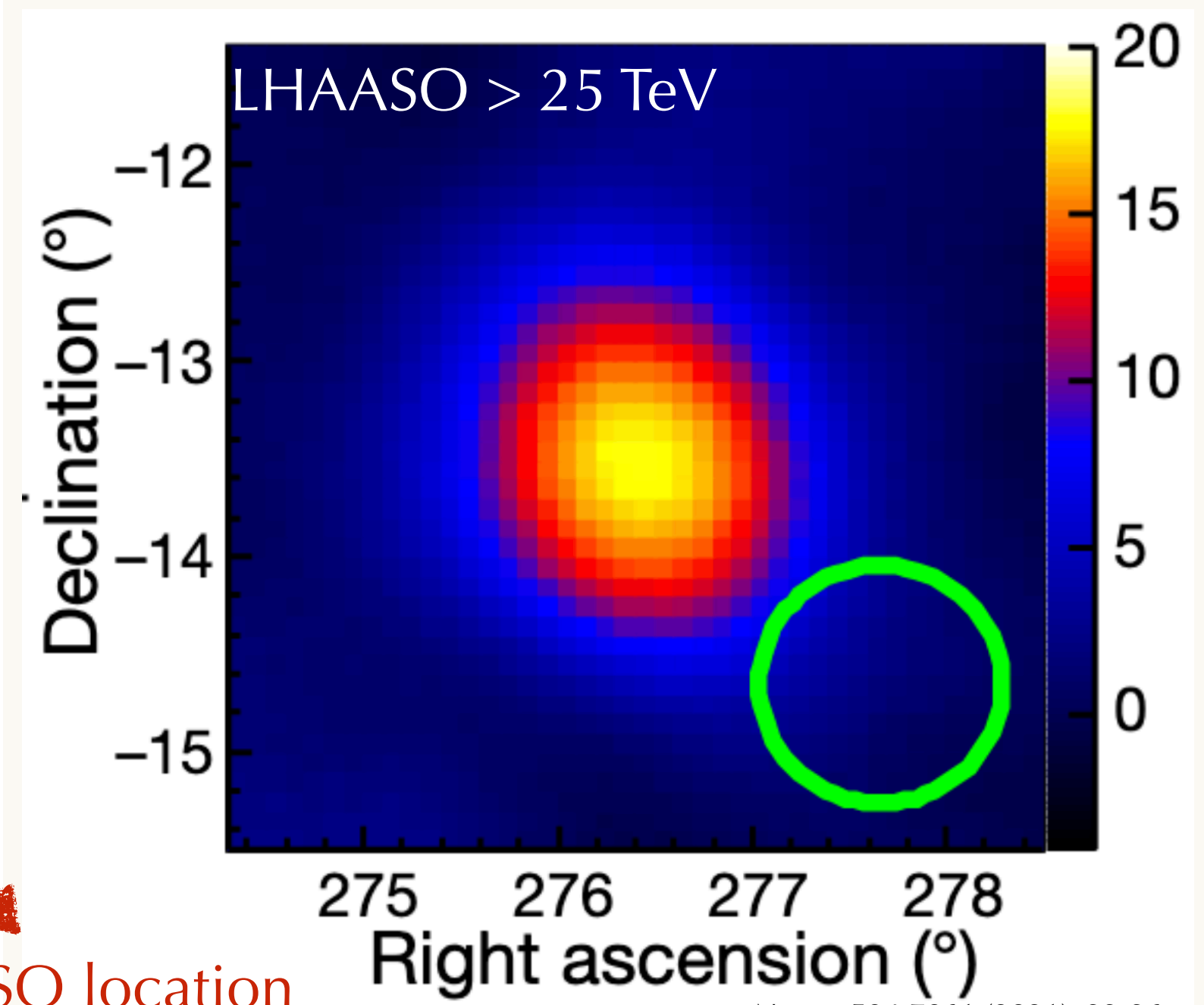
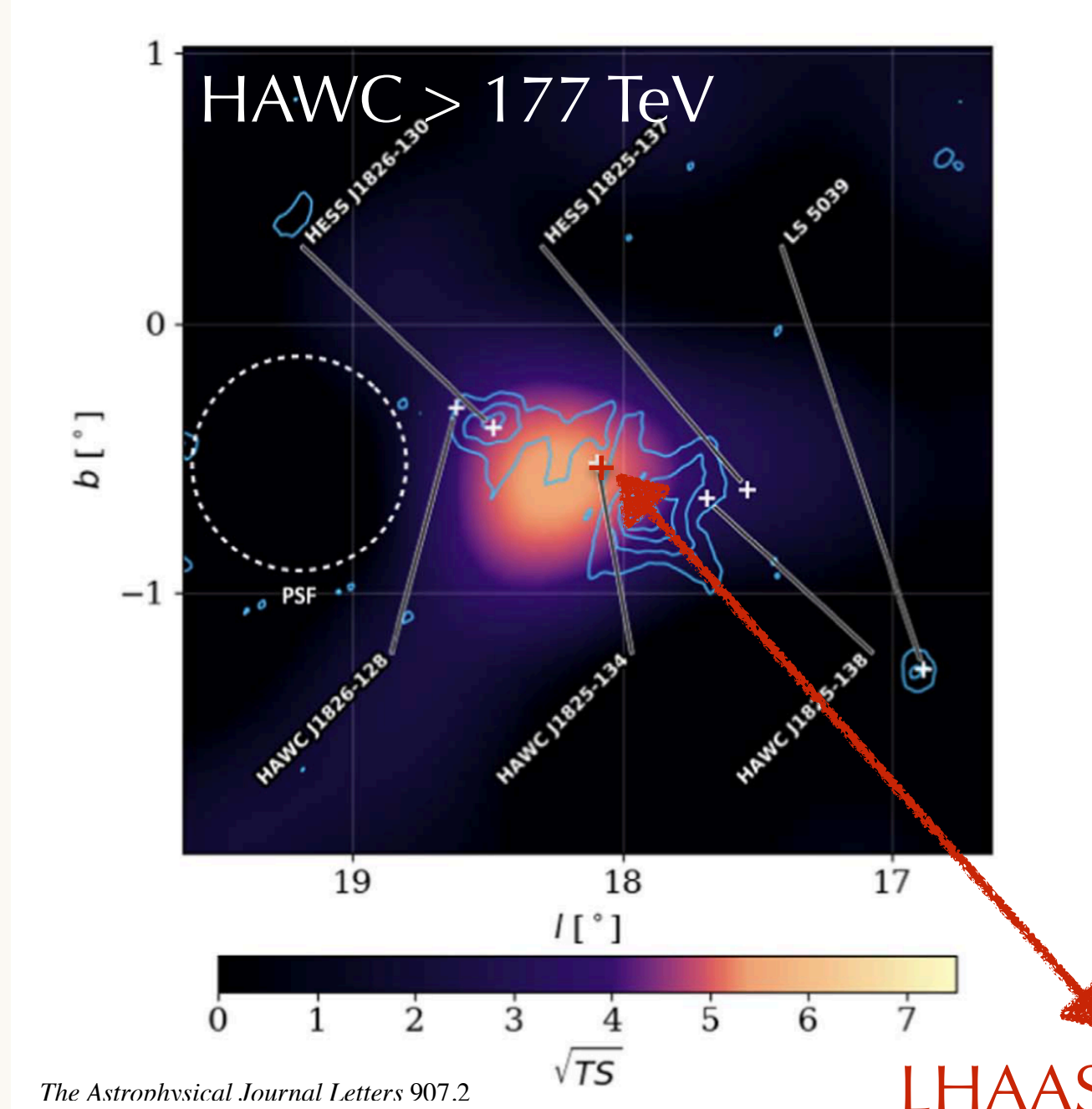
- Multi-source analysis reveal 3 components in Pass5 data
- The extended component spectrum shows great agreement with LHAASO flux point at 100 TeV



# eHWC J1825-134



Point source spectrum



The Astrophysical Journal Letters 907.2

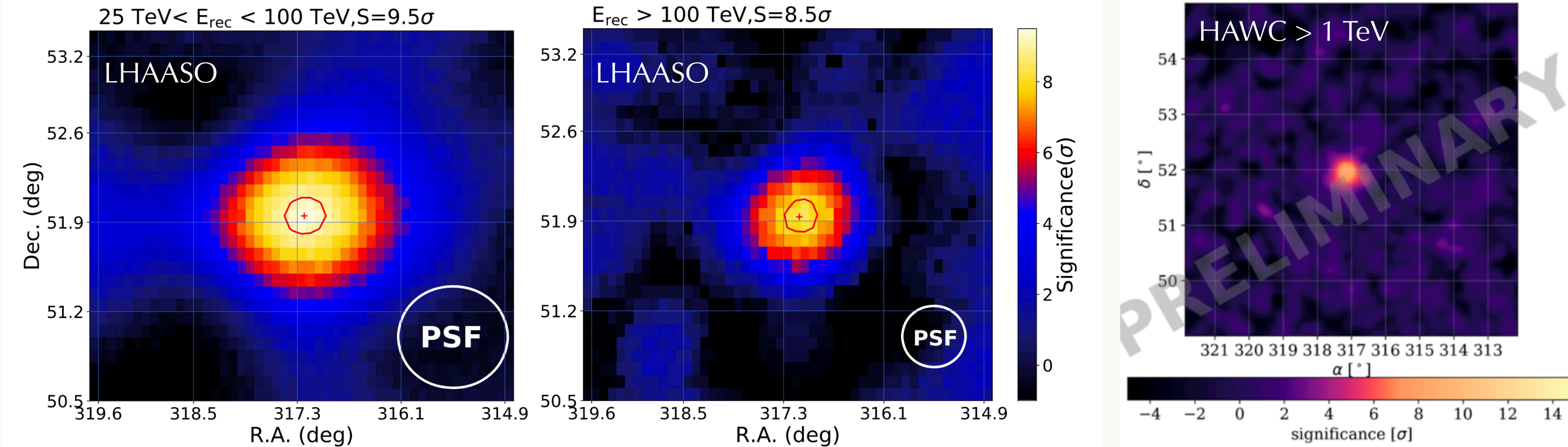
Nature 594.7861 (2021): 33-36

LHAASO location

|                                | HAWC (point source)   | LHAASO                    |
|--------------------------------|---|---------------------------|
| <b>Location</b>                | R.A. 276.44° Dec. -13.42°   | R.A. 275.45° Dec. -13.45° |
| <b>Morphology</b>              | 2 extended sources + 1 point source                                 | 0.3 ° extension template  |
| <b>Maximum measured energy</b> | $>200$ TeV  | 420 TeV                   |
| <b>Origin of TeV emission</b>  | Proton accelerated by SFR<br>Electron accelerated by PSR J1826-1334 |                           |

- Two extended HAWC sources are cutoff around 25 TeV
- Assume LHAASO  $> 200$  TeV flux associate with HAWC point source
- LHAASO spectrum shows good continuity with HAWC

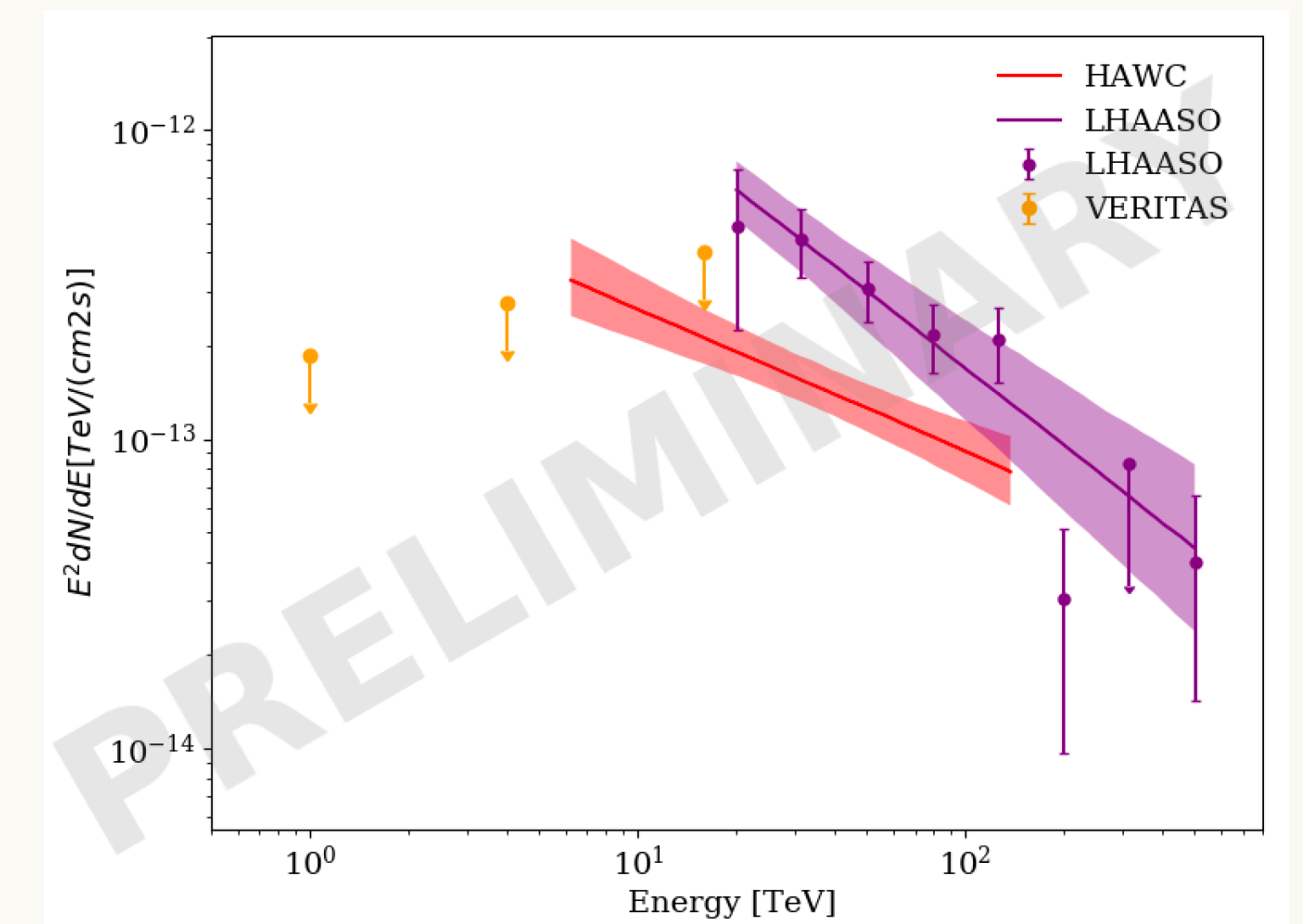
# LHAASO J2108+5157



- HAWC detected LHAASO J2108+5157 in new Pass5 data
- Joint analysis with VERITAS is undergoing.
- See more on Wednesday 15:50 Bangale. P, Galactic Source

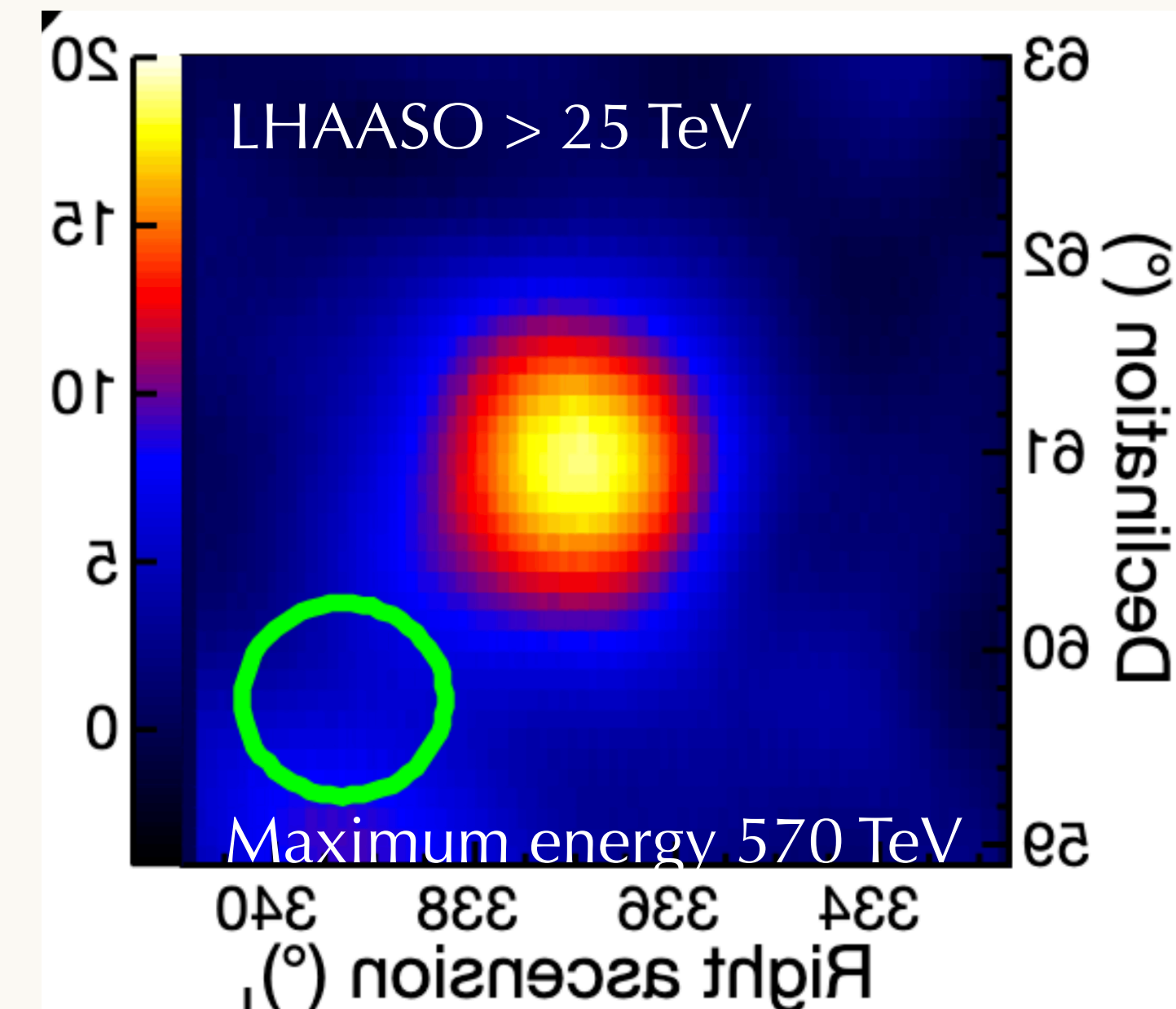
The Astrophysical Journal Letters, 919:L22 (9pp)

|                                | HAWC                     | LHAASO                   |
|--------------------------------|--------------------------|--------------------------|
| <b>Location</b>                | R.A. 317.14° Dec. 51.93° | R.A. 371.22° Dec. 51.95° |
| <b>Maximum measured energy</b> | >137 TeV                 | 430 TeV                  |
| <b>Morphology</b>              | Point-like               | Point-like               |

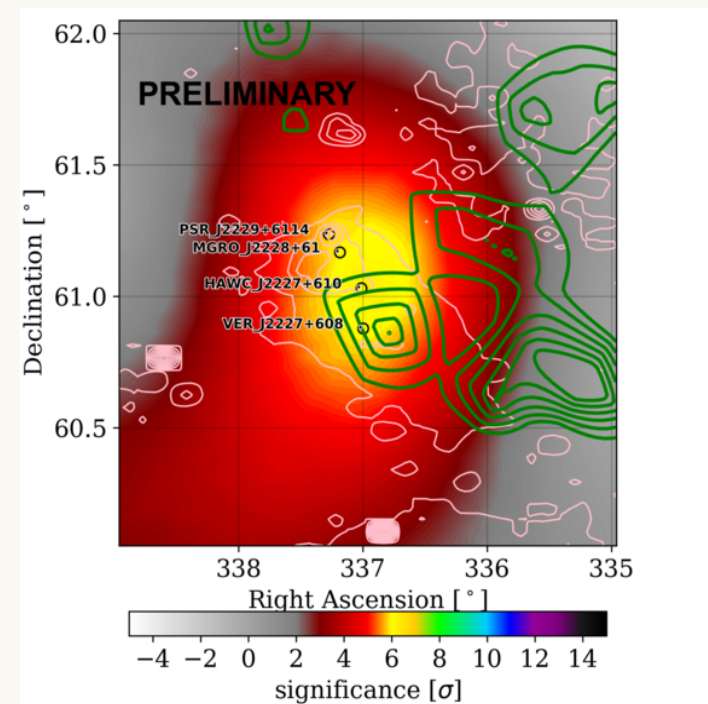




# HAWC J2227+ 610 (Boomerang region)



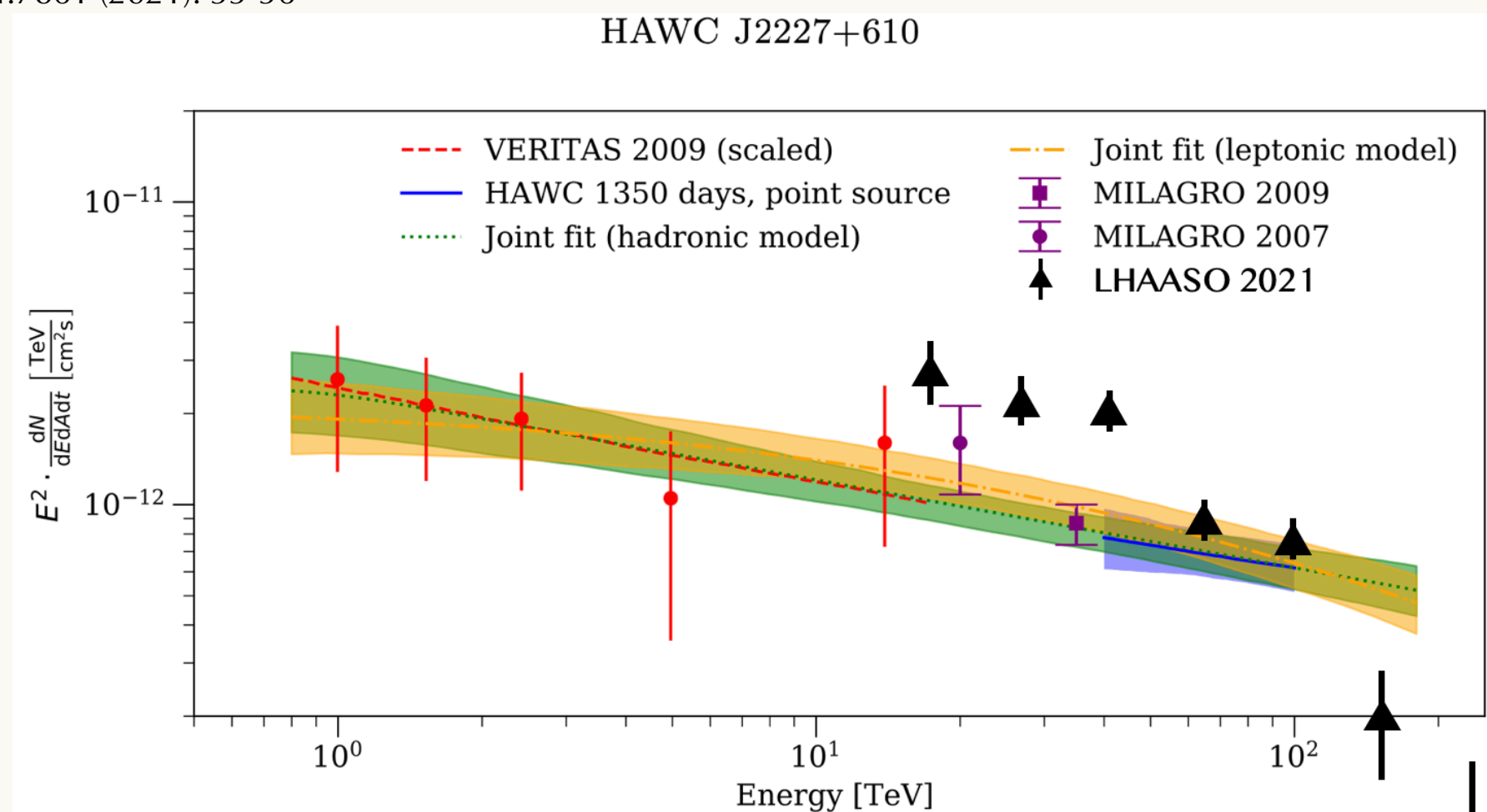
HAWC 1347 days



Maximum energy 180 TeV

- SNR G106.3+2.7 is a comet-shaped radio source
- PSR J2229+6114, seen in radio, X-rays, and gamma rays and its pulsar wind nebula (PWN), Boomerang Nebula is contained in the remnant
- Pass 4 analysis
- HAWC morphology and location are consistent with VERITAS
- The joint VERITAS-HAWC spectrum is well fit by a power law from 900 GeV to 180 TeV
- If hadronic, the cutoff energy in the underlying proton spectrum is constrained to be above 800 TeV

Nature 594.7861 (2021): 33-36

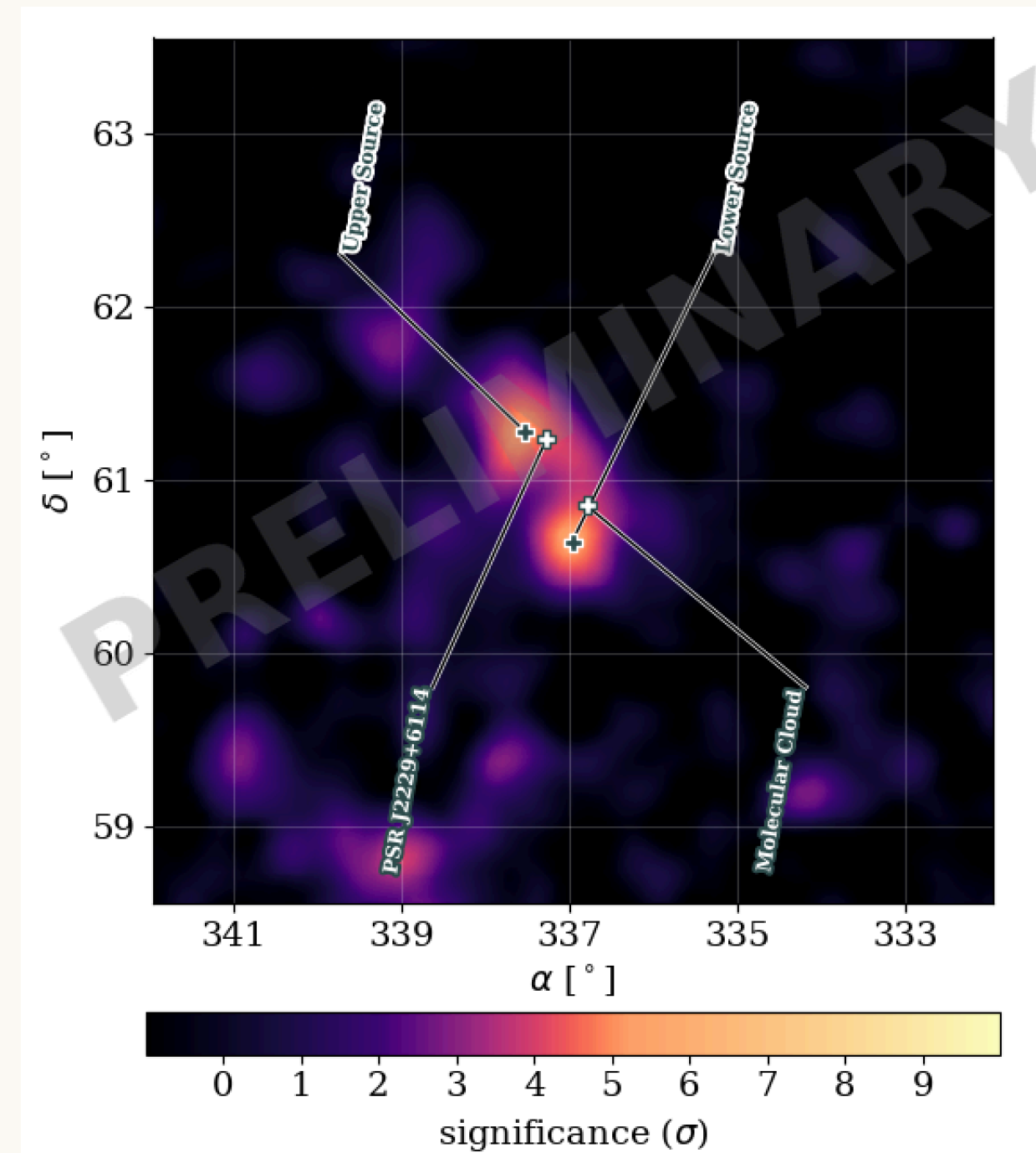


A. Albert et al 2020 ApJL 896 L29

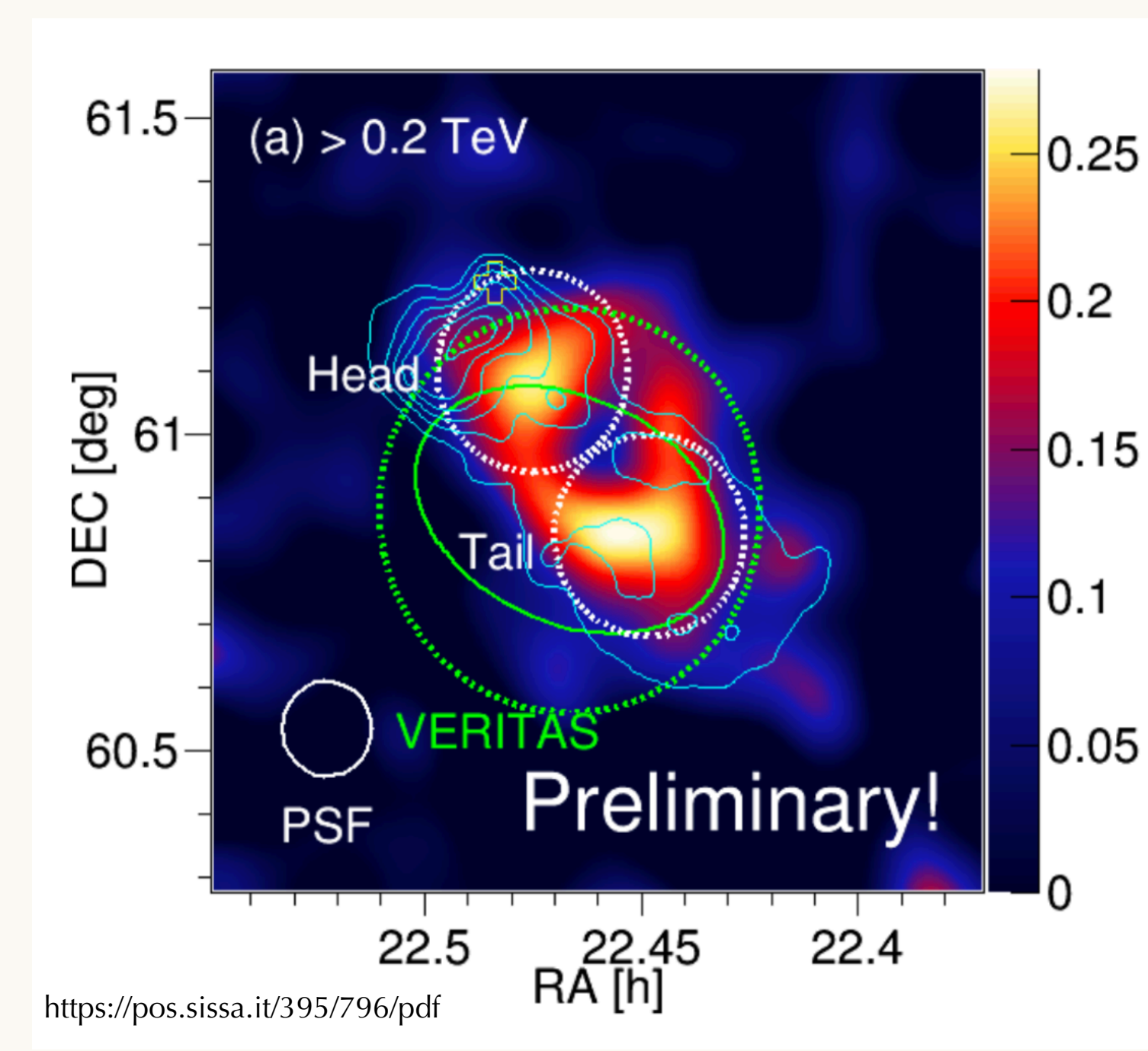
# HAWC J2227+ 610 (Boomerang region)

- In new HAWC data, HAWC resolves two sources
- MAGIC sees two sources
- Head Region (Upper Source)
  - Contains PWN and PSR
  - IC scattering in the PWN
- Tail Region (Lower Source)
  - Molecular cloud nearby
  - Both pion decay and IC scattering are plausible

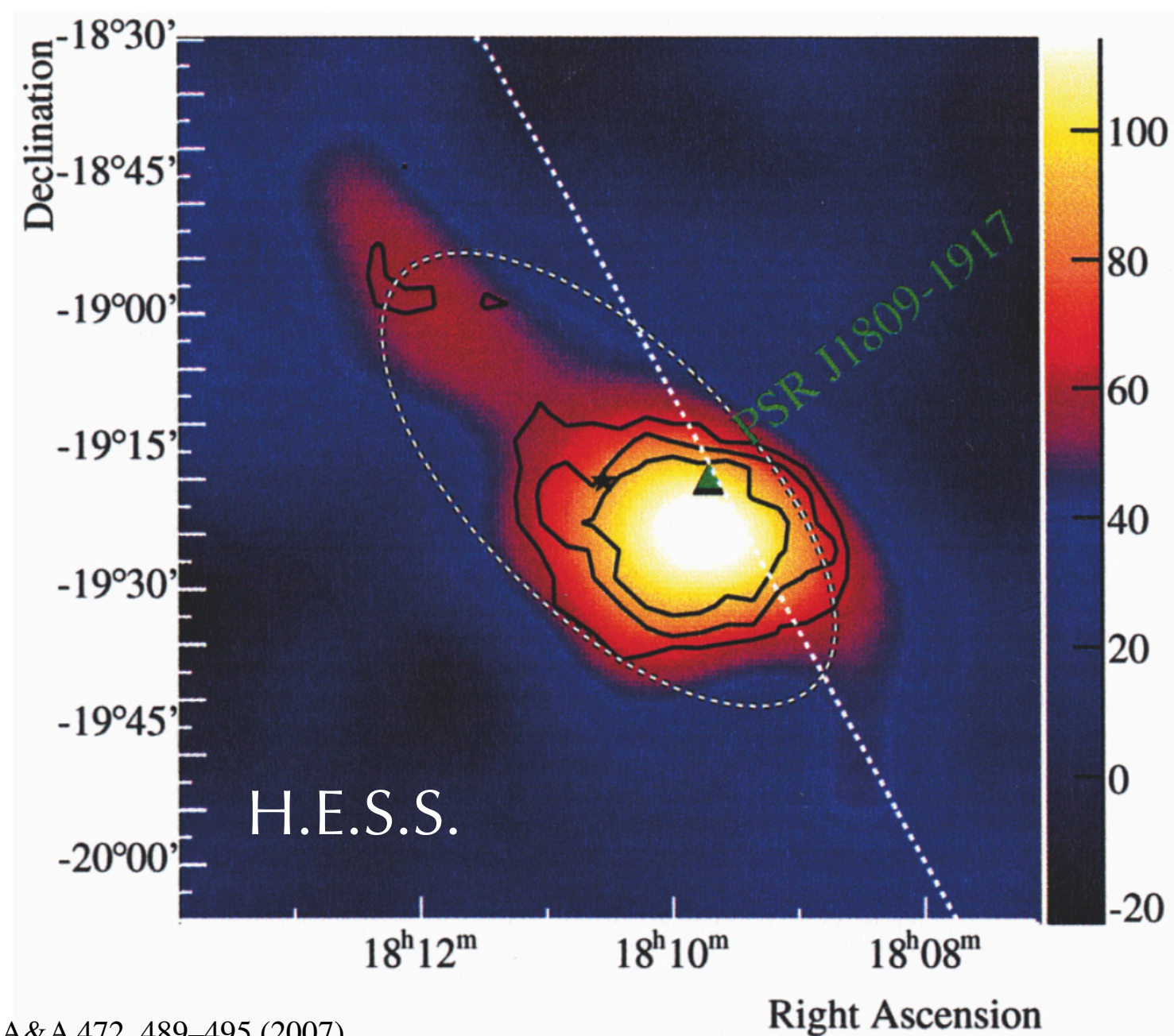
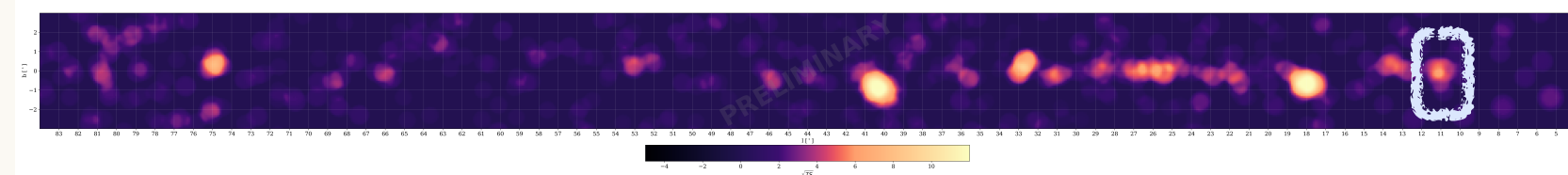
HAWC Pass5 Data



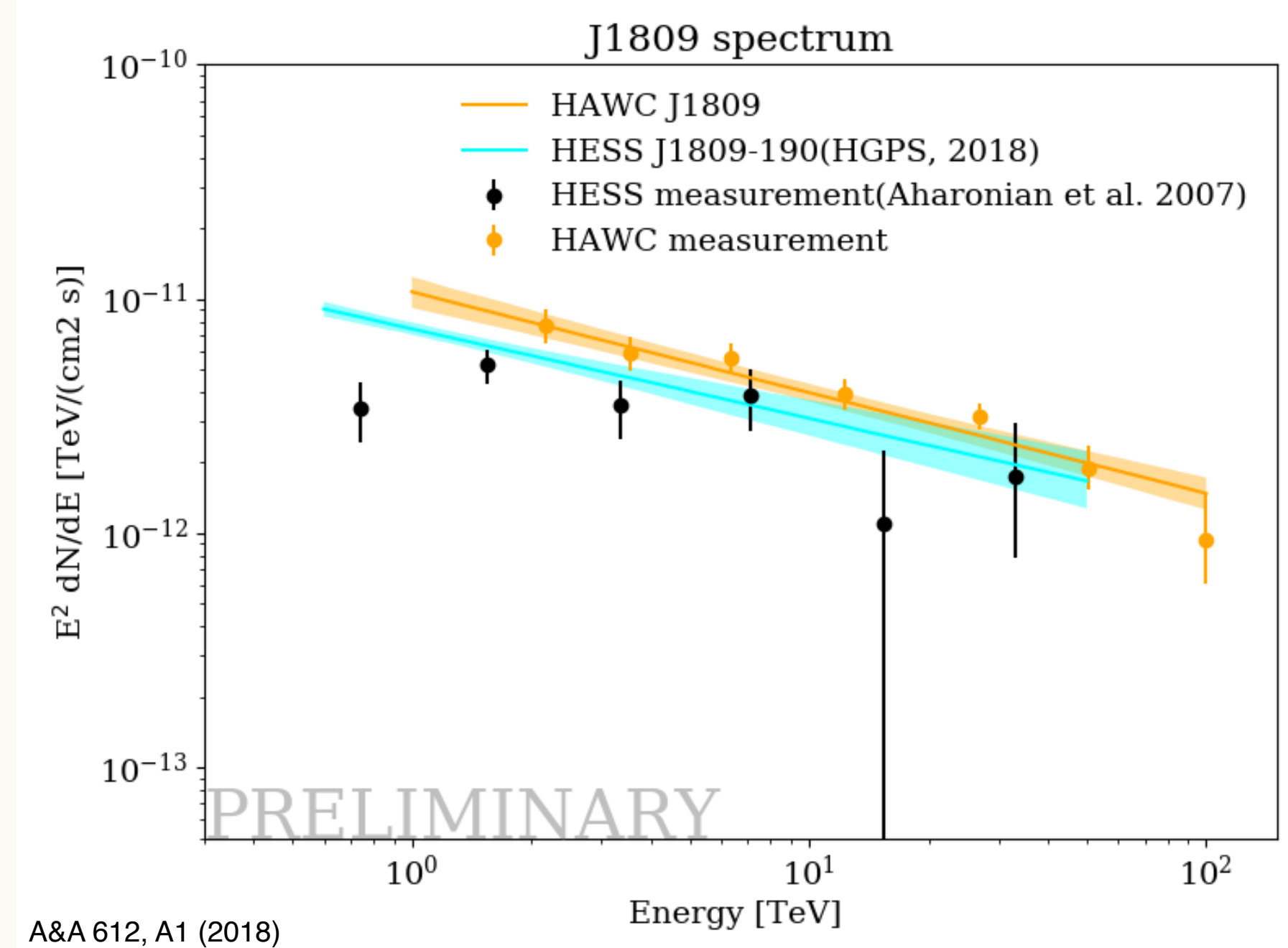
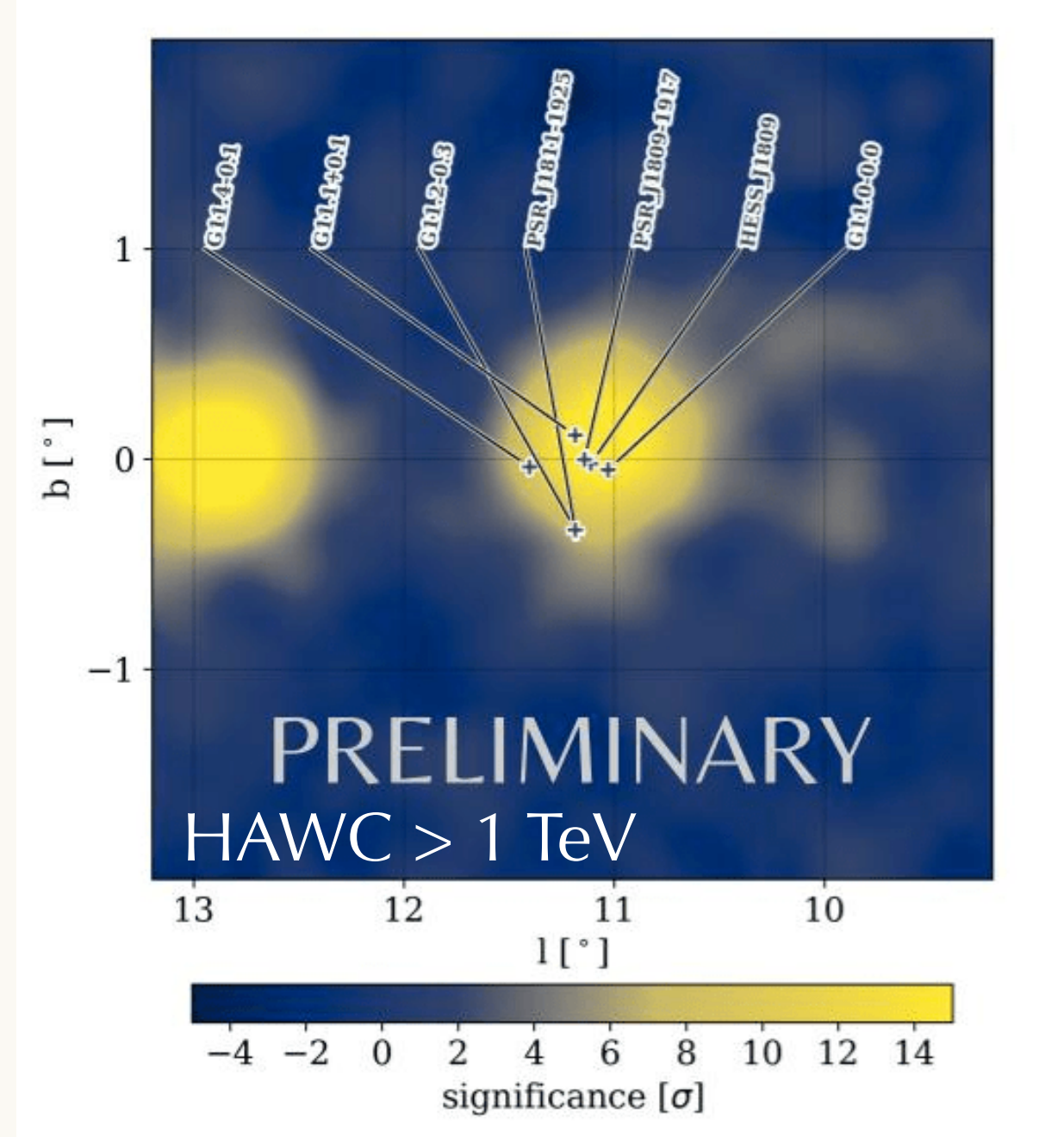
MAGIC



# HESS J1809-1917



A&A 472, 489–495 (2007)



- Not detected by LHAASO's 12 UHE sources, at the edge of their FOV
- PSR J1809-1917 is a young (age = 51 kyr) pulsar,  $E = 1.8 \times 10^{36} \text{ erg s}^{-1}$
- Several SNRs are in the region, G11.0-0.0 is spatially coincident with the peak of the H.E.S.S. emission
- Possibly Hadronic PeVatron
- Preliminary analysis shows good agreement with H.E.S.S., SED extended to 100 TeV without cutoff



# Conclusion and Outlook

- eHWC sources are confirmed by LHAASO with higher energies
- More PeVatron candidates found in HAWC pass5 dataset
- Improved angular reconstruction allows to associate  $\gamma$ -ray source with accelerator
- More sensitivity results using outriggers are coming!