# Primordial Black Holes Dark Matter in the Context of Extra Dimensions

Based on arxiv:2201.11761 (PRD 105, 103508)

August 8th, 2022

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- If PBHs survive they act as Cold Dark Matter
- PBHs have been studied extensively in 4D
  - But in extra-dimensions black holes behave differently!

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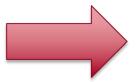
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• When r >> R

$$V(r) = \frac{m_1 m_2}{M_*^{2+n} R^n r}$$



$$M_*^{2+n}R^n = M_{pl}^2$$

#### Extra Dimensional Black Holes

Black holes with
 r<sub>h</sub> << R have
 modified size and
 temperature</li>

$$r_s = \frac{a(N_{\rm ED})}{M_*} \left(\frac{M}{M_*}\right)^{\frac{1}{N_{\rm ED}+1}}$$

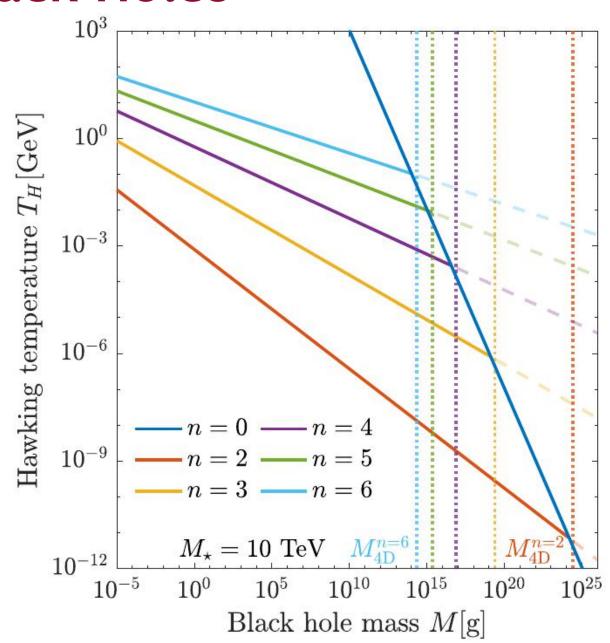
$$T_{BH} = \frac{N_{\rm ED} + 1}{4\pi r_s}$$

See Conley and Wizanksy arxiv:hep-ph/0611091

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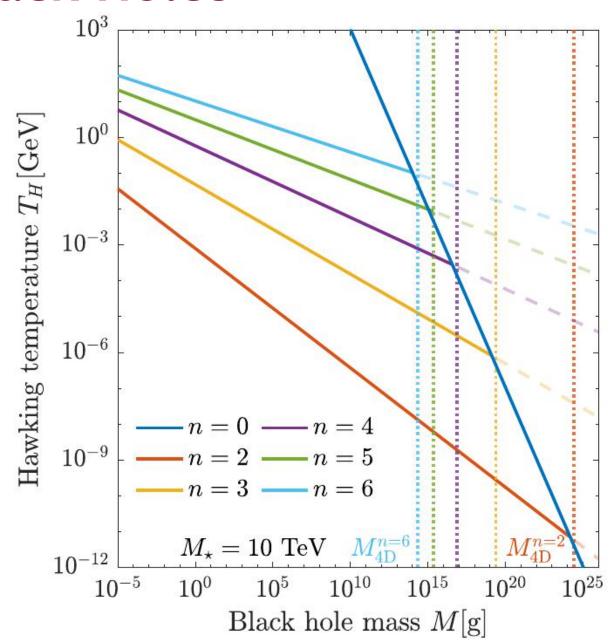


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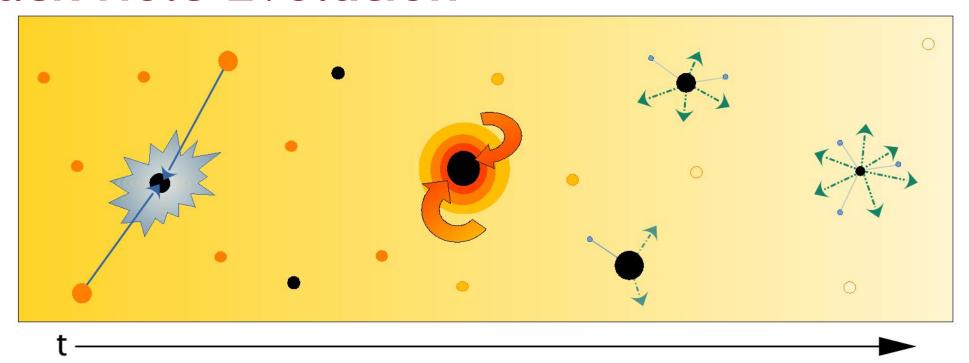
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 Particle collisions at E > M\* can produce microscopic black holes

See Conley and Wizanksy arxiv:hep-ph/0611091



#### **Black Hole Evolution**

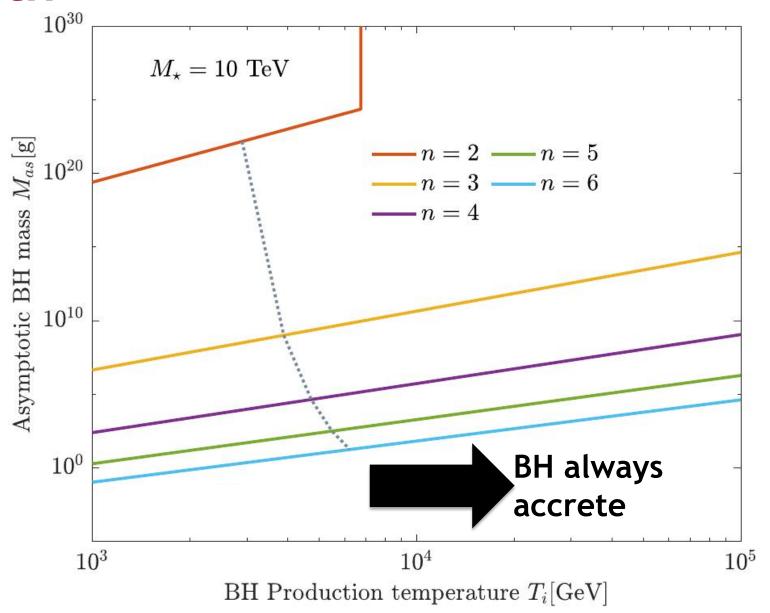


Black hole mass evolves due to accretion and evaporation

$$\frac{dM}{dt} = \left(-\alpha + \beta \frac{T^4}{T_H^4}\right) T_H^2$$

#### **Black Hole Growth**

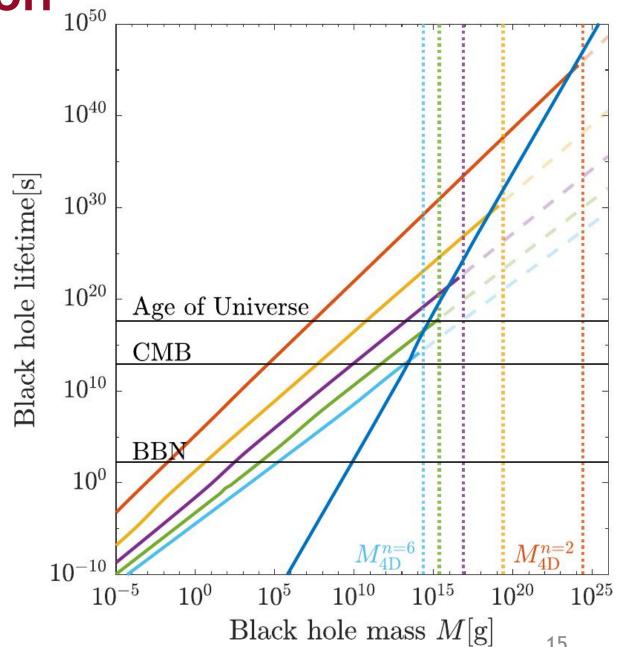
Black holes in extra-dimensions initially **grow** via accretion depending if they are created when the universe is sufficiently hot



# **Black Hole Evaporation**

 After accretion stops due to the universe cooling, black holes evaporate via Hawking radation

 Lifetime depends on initial mass and number of dimensions



# **Evaporation Spectrum**

 Black holes evaporate to all sufficiently light particles with a grey-body spectrum

$$dP = \sigma_i \frac{\omega}{\exp(\omega/T_H) \mp 1} \frac{d^3p}{(2\pi)^3}$$

# **Evaporation Spectrum**

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- Secondary photons and electrons produced from cascades of unstable evaporation products
  - PPPC4DMID used for secondary spectrum
    Cirelli et al. arxiv:1012.4515

#### **Constraining Large Extra-Dimensions**

- Astrophysical constraints on PBHs come from a variety of sources:
  - Big Bang Nucleosynthesis (BBN)
  - CMB angular power spectrum
  - Galactic centre photon flux
  - Isotropic x-ray and gamma-ray flux

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# **Isotropic Light Constraints**

- PBH evaporation post-recombination produces an isotropic X-ray and gamma ray background
- There are two seperate components:
  - Extragalactic evaporation
  - Isotropic galactic evaporation

# **Extragalactic Evolution**

- Extragalactic photon spectrum tracked from recombination to today
- Each redshift step updated using:

$$\frac{d\Phi_{\gamma,\text{EBL}}}{dE_i} = \frac{V_{i-1}}{V_i} \frac{dE_{i-1}}{dE_i} \frac{d\Phi_{\gamma,\text{EBL}}}{dE_{i-1}} e^{-\tau} + \frac{d\Phi_{\gamma,\text{comp}}}{dE_i dz_i} \Delta z + \frac{d\Phi_{\gamma,\text{inj}}}{dE_i dz_i} \Delta z$$

Universe Expansion

Attenuation

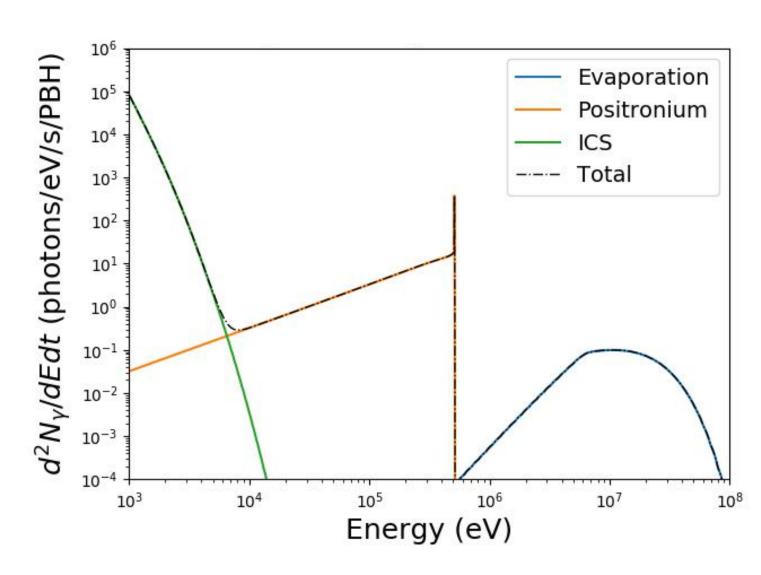
Scattering

Injected Photons

#### **Extragalactic Photon Production**

Extragalactic Photons produced via 3 mechanisms:

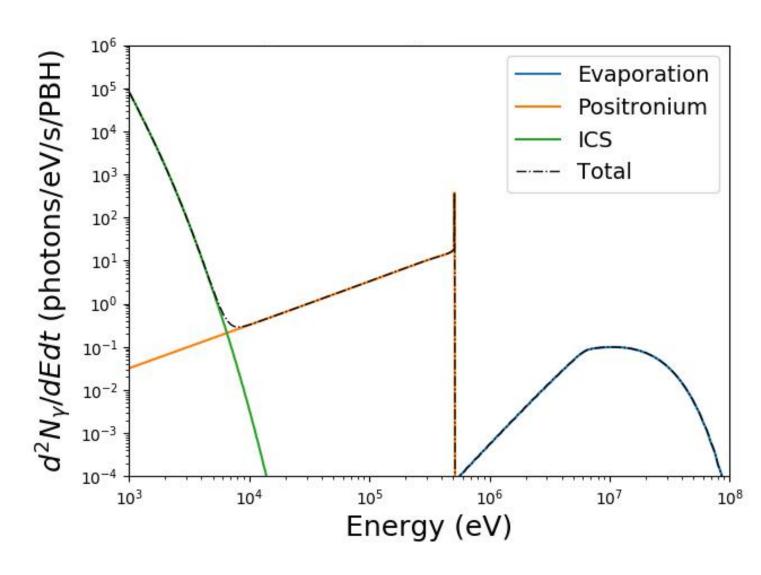
Evaporation



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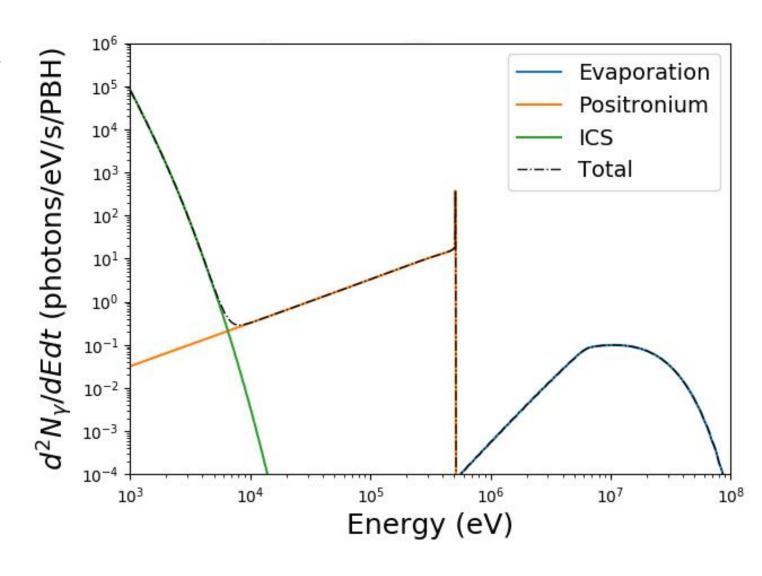
- Evaporation
- Positronium annihilation



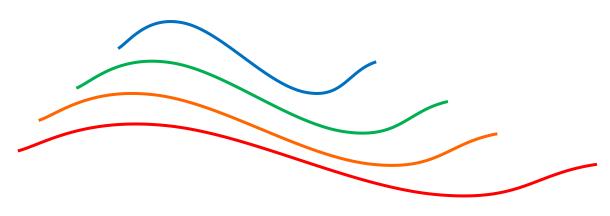
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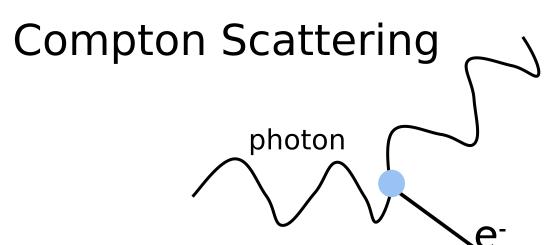
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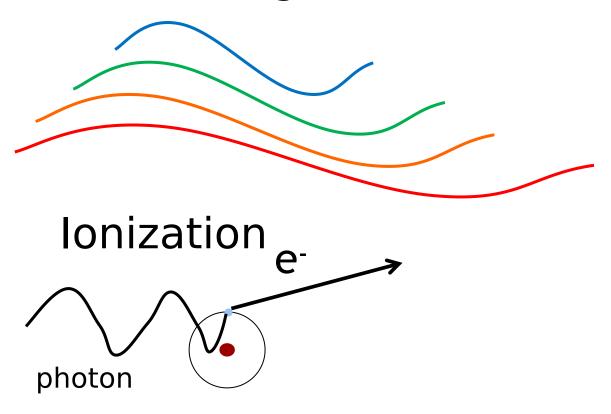
- Evaporation
- Positronium annihilation
- Inverse Compton scattering (ICS)

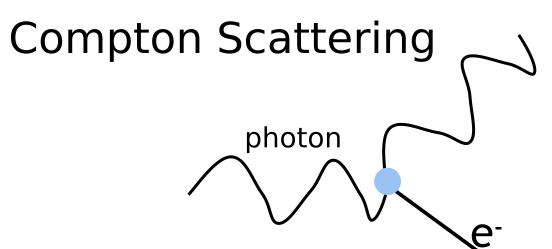


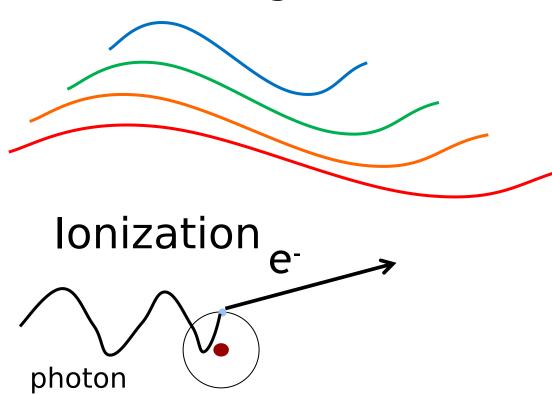


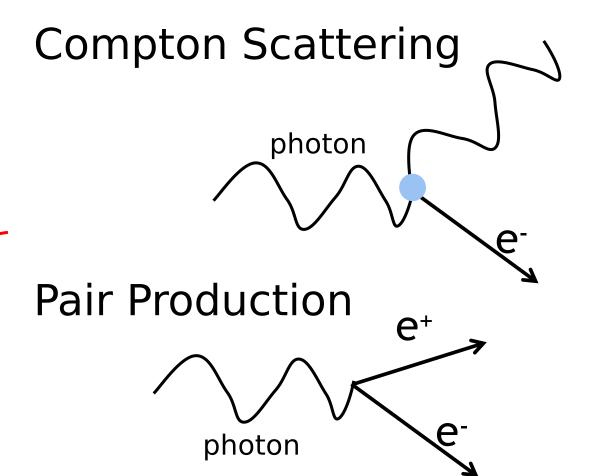


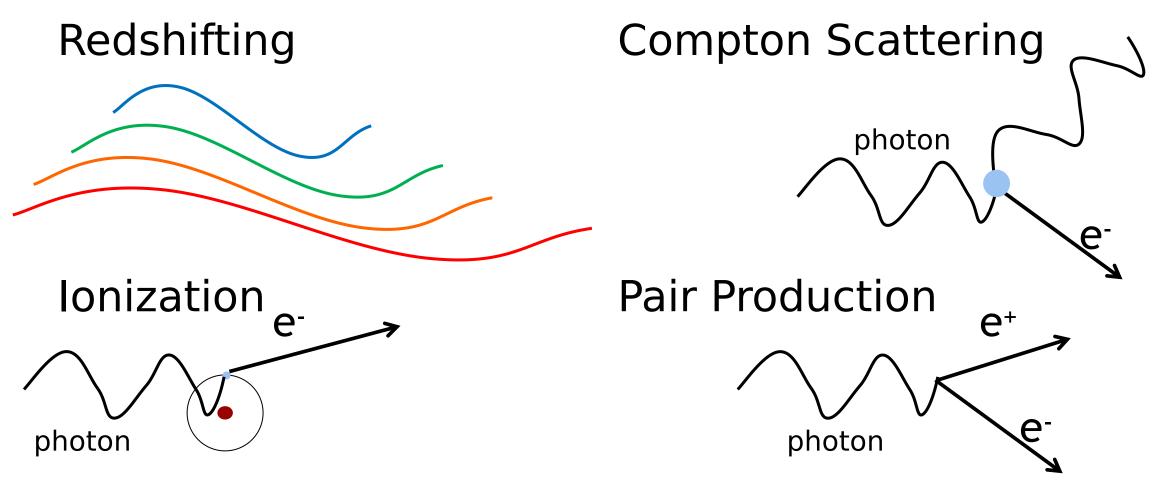












For the observed energy range (1 keV - 100 MeV) attenuation is important for high redshift (z > 100)

# Galactic Isotropic Signal

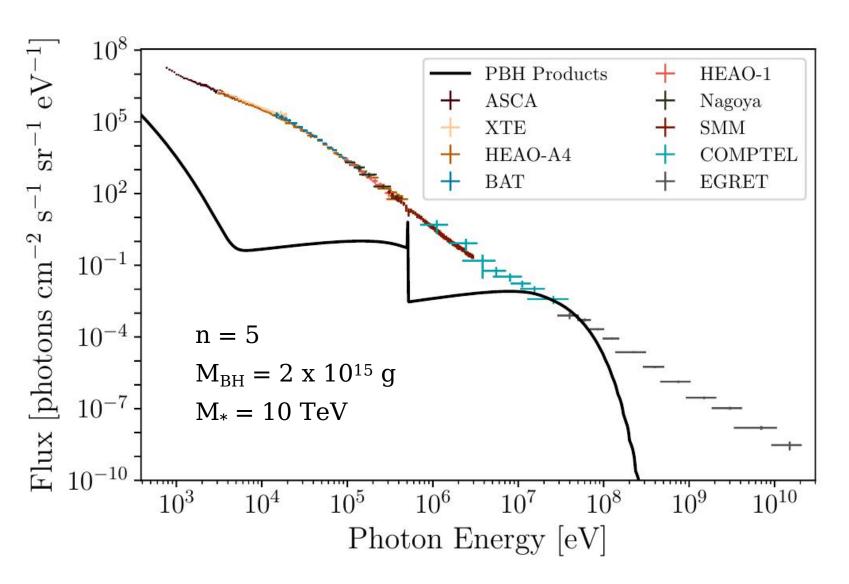
- Galactic photon flux is highly anisotropic but non-zero in all directions
- Isotropic component can be determined with

$$\frac{d\Phi_{\gamma,\text{gal}}}{dE} = \frac{f_{\bullet,0}}{4\pi M} \frac{d^2 N_{\gamma}}{dE dt} (E, M) \mathcal{D}_{\text{min}}$$

$$\mathcal{D}_{\min} = \int_{R_{\circ}}^{\infty} dr \rho_{DM}(r)$$

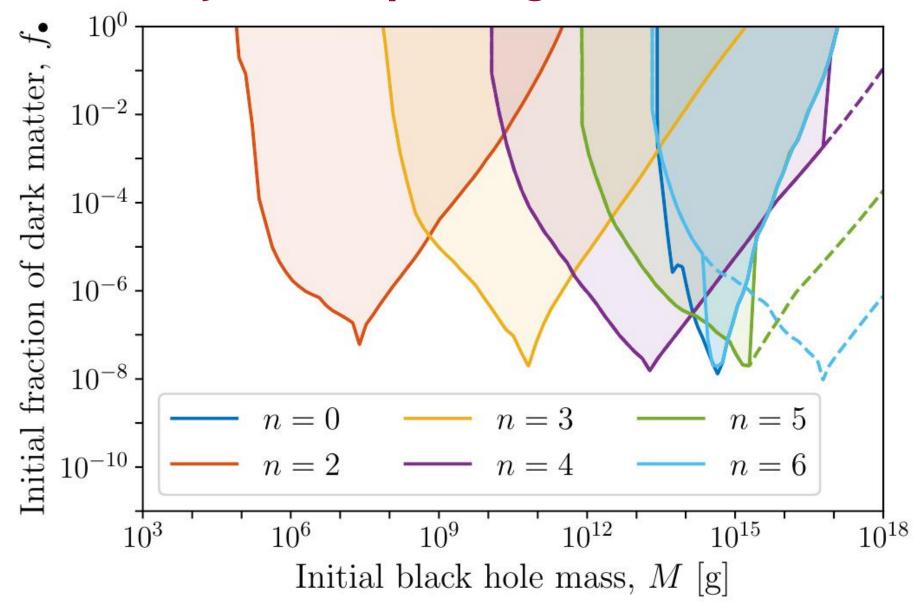
See: Iguaz et al. arxiv:2104.03145

#### **Observed Flux**

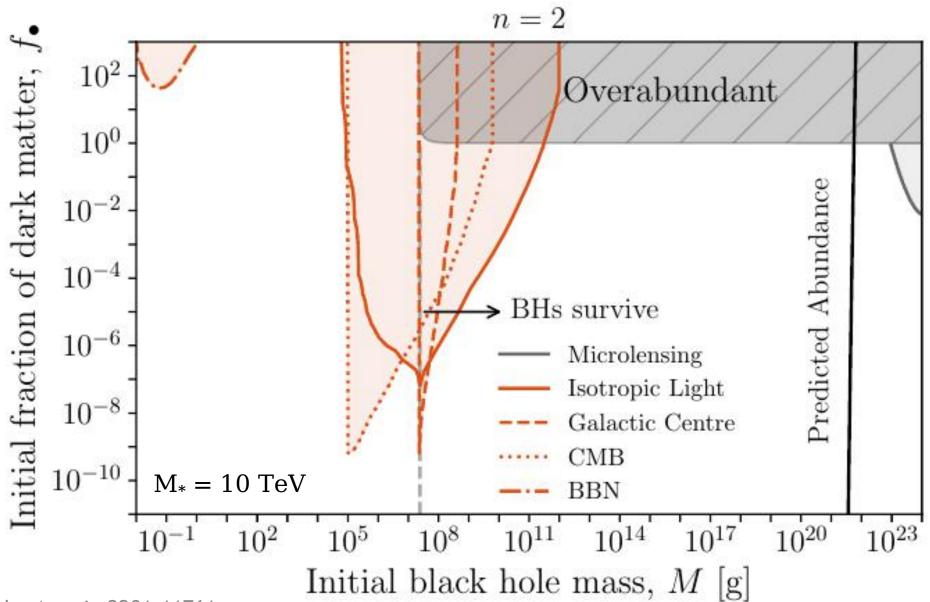


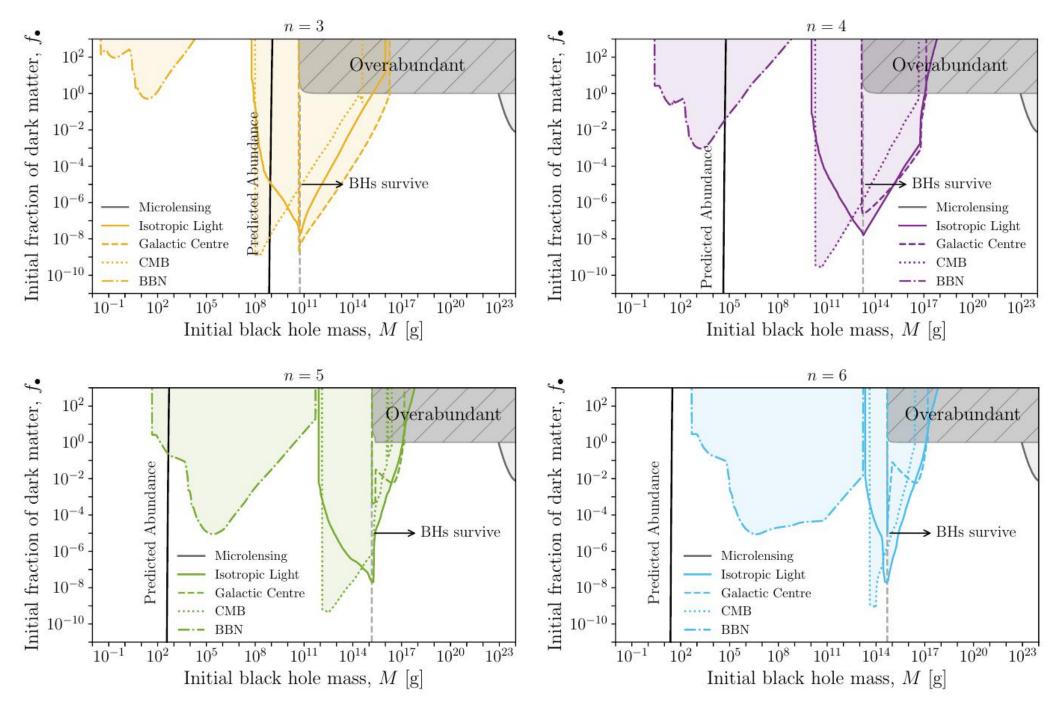
Calculated flux is compared to various x-ray and gammaray telescopes as compiled by Ajello et al. (arxiv:0808.3377)

# Preliminary Isotropic Light Constraints



#### **Combined Constraints**





#### Conclusions

Theories of Large Extra Dimensions predict the existence of Primordial Black Holes

- Astrophysical observables of LED PBHs is very different from regular PBHs
- Primordial Black Holes with two extra dimensions might comprise all of dark matter!
- Check out the paper for more! (arxiv:2201.11761)

# Questions?

# Extra Slides

### The Hierarchy Problem

- Two scales exist in the Standard Model
  - Electroweak scale (~10³ GeV)
  - Planck/Quantum Gravity scale (~10<sup>18</sup> GeV)
- Higgs boson mass is set by the electroweak scale but without fine tuning, quantum corrections would be expected to increase Higgs mass to the Planck

#### Extra Dimensional Black Holes

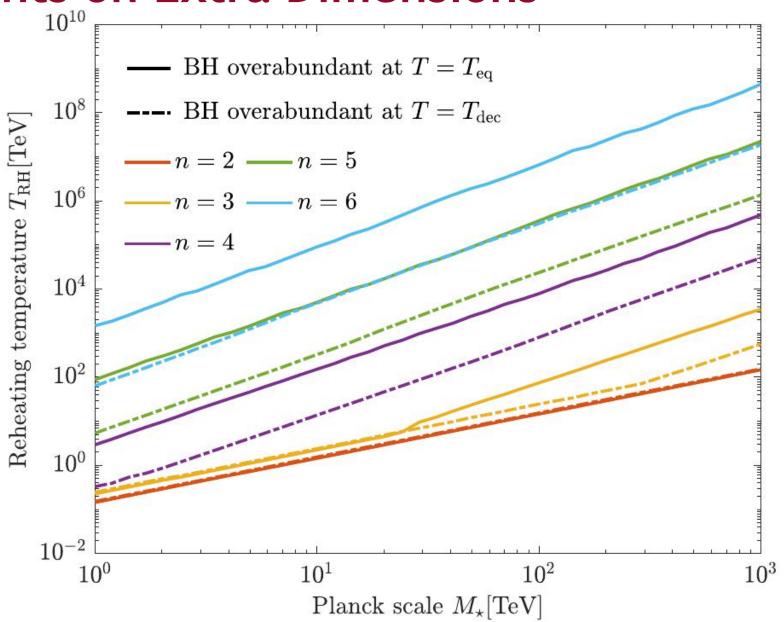
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$$r_h = \frac{a_n}{M_{\star}} \left(\frac{M}{M_{\star}}\right)^{1/(n+1)}$$

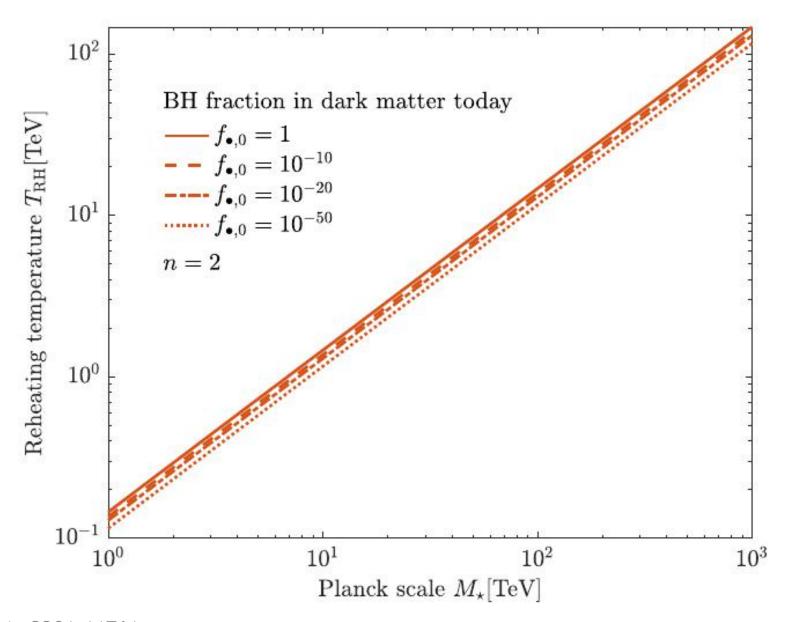
$$T_H = \frac{n+1}{4\pi r_h}$$

See Conley and Wizanksy arxiv:hep-ph/0611091

#### **Constraints on Extra Dimensions**



#### **Predicted PBH Abundance**



#### **Detailed BBN Effects**

- There are four main mechanism for PBHs to affect BBN
  - Increased universe expansion rate causing earlier neutron freeze-out
  - Hadrons and mesons converting protons to neutrons after freeze-out
  - Energetic mesons dissociating Helium nuclei
  - Energetic photons dissociating Helium nuclei

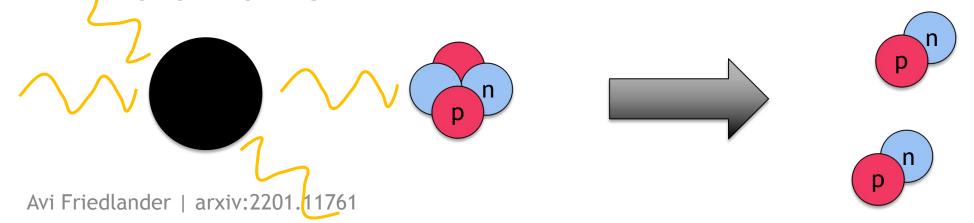
# **Calculating BBN Constraints**

- The effect of decaying dark matter on BBN has been studied in detail (Kawasaki et al. arxiv:1709.01211)
- PBHs can be mapped onto decaying dark matter models ensuring key properties match
  - Dark matter/PBH density
  - Average injected fermion energy
  - Average time of injected energy (Kieth et al. arxiv:2006.03608)

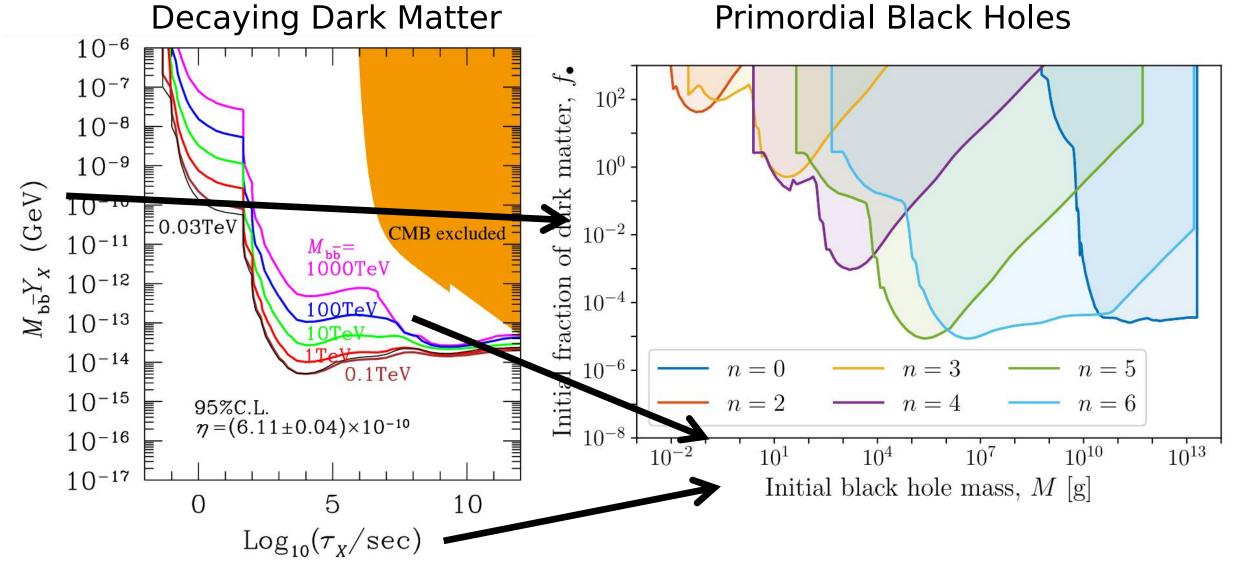
# PBH and Big Bang Nucleosynthesis

When the universe cools to ~1 MeV light nuclei
 form

- PBHs can change the expansion rate during BBN
- Evaporation products can dissociate primordial elements



### **Preliminary BBN Results**



Kawasaki et al. arxiv:1709.01211

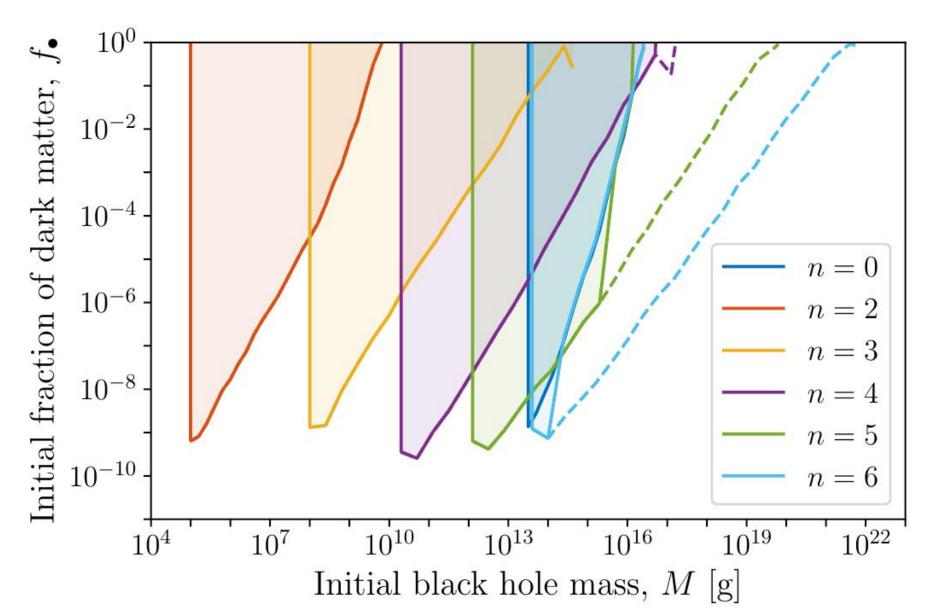
See Kieth et al. arxiv:2006.03608 for recasting details

#### **Effect on the CMB**

- Evaporation products during and after recombination scatter CMB photons changing the observed power spectrum
- We modify ExoCLASS to determine the effect from LED PBHs

Constraints are from performing a MCMC

#### **CMB Constraints**



#### **Galactic Centre Photons**

$$\frac{d\Phi_{\gamma}}{dEd\Omega} = \frac{1}{4\pi} \frac{dN}{dEdt} \frac{f_{\bullet,0}}{M} \frac{1}{\Delta\Omega} \mathcal{D}(\Omega)$$

$$\mathcal{D}(\Omega) \equiv \int_{l.o.s.\Delta\Omega} \rho_{DM}(\vec{x}) d\Omega dx$$

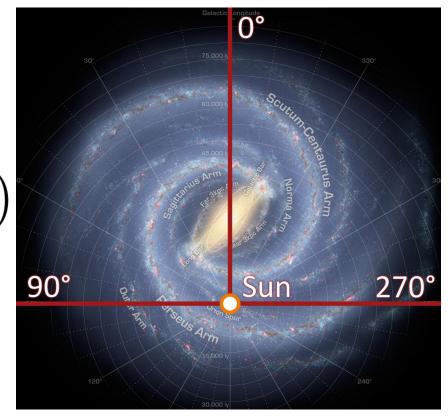


Image from NASA

 Compared to previously processed 6-years of INTEGRAL/SPI data (Bouchet et al. arxiv:1107.0200)

#### **Galactic Centre Positrons**

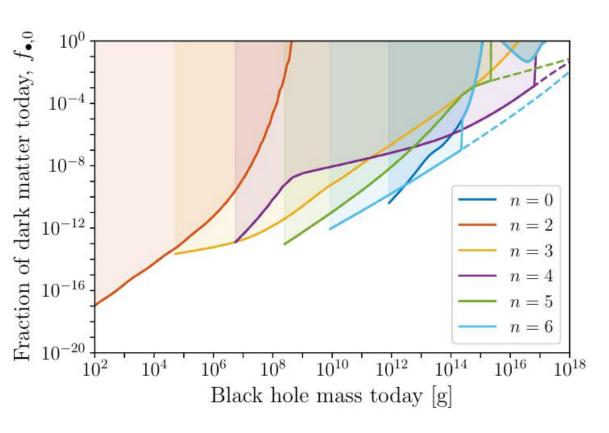
 If PBHs evaporate to positrons, they will annihilate to produce an additional 511 keV signal

$$\frac{d\Phi_{511}}{d\Omega} = 2(1 - 0.75f_P) \frac{dN_{e^+}}{dt} \frac{1}{4\pi} \frac{1}{M} \frac{1}{\Delta\Omega} \mathcal{D}(\Omega)$$

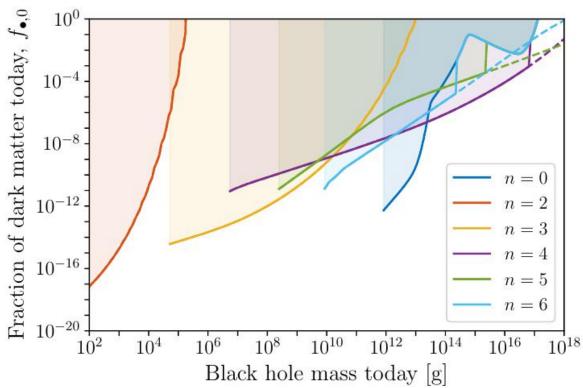
 Compared to INTEGRAL/SPI 511 keV line data (Siegert et al. arxiv:1906.00498)

#### **Galactic Centre Constraints**

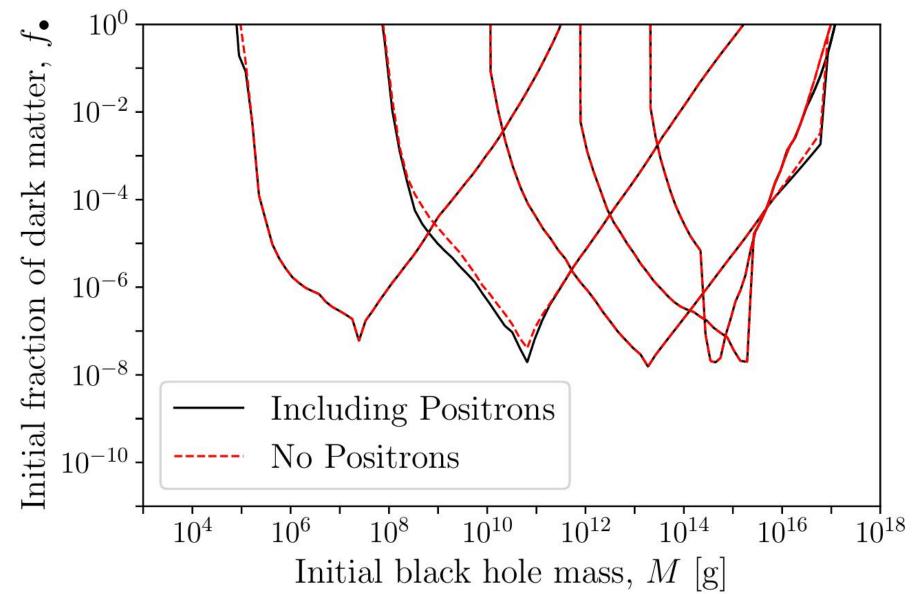
#### Gamma ray continuum



#### **Positron Annihilation**

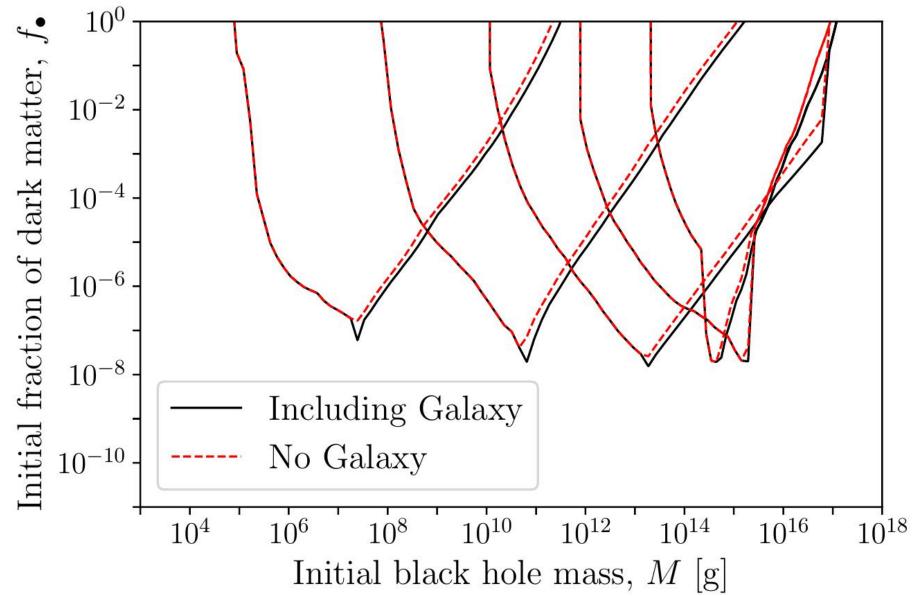


#### Impact of Positronium on Isotropic Photon Limits

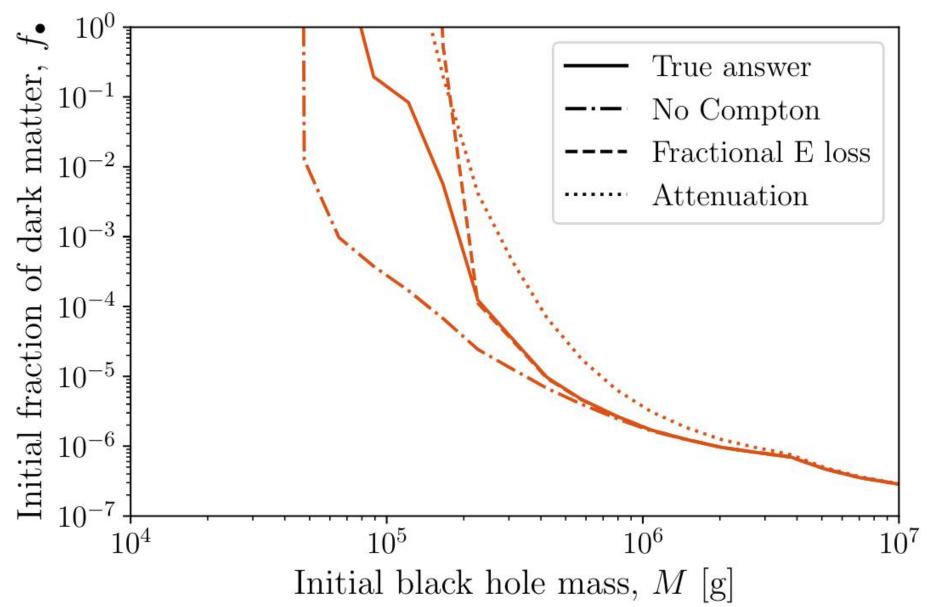


Avi Friedlande

### Impact of Milky Way on Isotropic Photon Limits



# **Compton Scattering Approximations**



### **Inverse Compton Scattering**

 High energy electrons and positrons upscatter CMB photons

$$\frac{d^2 N_{\gamma, \text{ics}}}{dE dt}(E, M_{BH}) = 2 \int_0^\infty dE_e \frac{d^2 N_{e^-}}{dE dt}(E_e, M_{BH}) \frac{d\tilde{N}_{\gamma, \text{ics}}}{dE}(E, E_e, T_{CMB})$$

Photons produced by ICS

Electron spectrum from evaporating PBH

ICS photon spectrum per electron calculated by DarkHistory (arxiv:1904.09296)

#### **Positronium**

- Positrons annihilate into photons via formation of positronium
  - Each annihilation produces spectrum

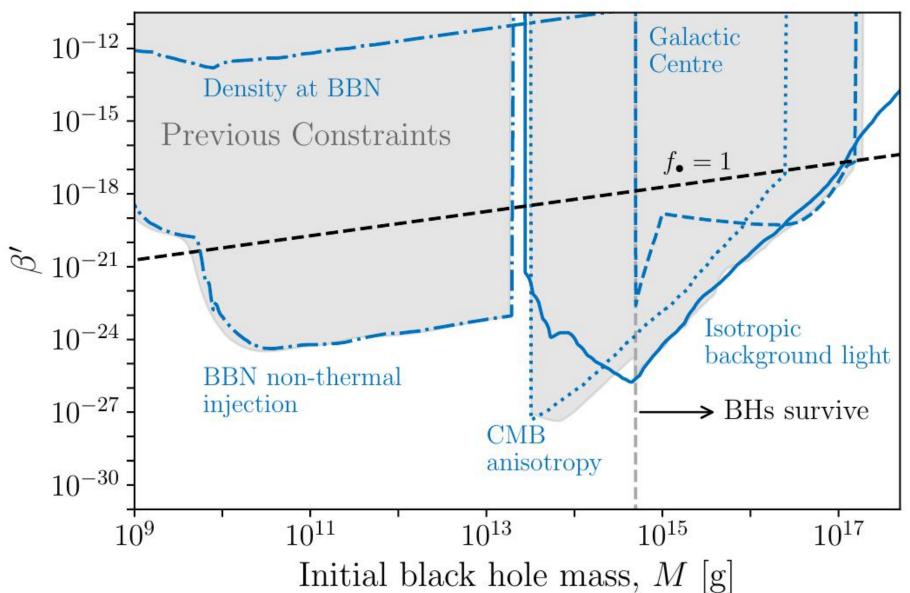
$$\frac{d\tilde{N}_{\gamma}^{\text{ann}}}{dE}(E) = \frac{1}{2}\delta(E - m_e) + \frac{3}{4}\frac{dN_{\gamma}^{\text{ann}}}{dE}\Big|_{\text{triplet}}$$

- All positrons assumed to immediately annihilate

See: Iguaz et al.

arxiv: 2104.03145

#### **4D PBH Constraints**



#### **4D Galactic Constraints**

