Long-Baseline Oscillations at DUNE

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Big Questions in Neutrino Oscillation Physics

$$U_{\rm PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & e^{-i\delta_{\rm CP}}s_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{\rm CP}}s_{13} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

- What is the neutrino mass ordering? Is it normal ordering (NO) or inverted ordering (IO)?
- Is there CP violation? What is the value of δ_{CP} ?
- Are there symmetries in neutrino mixing (PMNS matrix)? What is the octant of θ_{23} ?





DUNE Physics Goals

DUNE has a **rich** physics program which includes:

- 1. Make precise measurements of the oscillation parameters θ_{23} , θ_{13} and Δm_{32}^2
- 2. Resolve the neutrino mass hierarchy, i.e. whether $m_3^2 > m_2^2$ or $m_3^2 < m_2^2$
- 3. Determine the octant of θ_{23}
- 4. Determine whether CP is violated in neutrinos and make a measurement of δ_{CP}
- 5. Search for **T** appearance
- 6. Check the unitarity of the PMNS matrix
- 7. Search for nucleon decay
- 8. Be ready to detect low-energy neutrinos from a supernova
- 9. Search for Beyond Standard Model physics, e.g. sterile neutrinos, heavy neutral leptons etc .



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- This talk will focus on (1-4)
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Long-Baseline Neutrino Oscillation

Measure neutrino and antineutrino oscillation as a function of L/E

Does the three-flavor model describe the data?

- If yes: measure the mixing angles, mass splittings, and CP phase
- If no: characterize the new physics

Need for a global program: different energies, matter effects, systematic uncertainties, etc.



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Deep Underground Neutrino Experiment (DUNE)





Neutrino Beam

- LBNF beamline will produce world-leading intensity
 - Phase 1: 1.2 MW
 - Phase 2: Upgrade to \rightarrow **2 MW**
- On-axis beam —> broad range of energies
 - Covers 1st & 2nd oscillation maxima







Far Detector (FD)

• Liquid argon provides **precise reconstruction** of lepton and hadronic energy over **a broad energy range**

• Will consist of **4 x 10kt fiducial mass LArTPC** modules







Near Detector (ND)

Main purpose: enable prediction of Far Detector reconstructed spectra

Movable detector system: LArTPC with muon spectrometer (PRISM)

Same target, same technology \rightarrow inform predictions of reconstructed E_v in Far Detector

ND-LAR: liquid argon TPC with muon spectrometer (**TMS**) 50t detector with pixel readout to reduce pileup

- Characterize beam flux and neutrino interactions
- SAND: System for on-Axis Neutrino Detection
 - Inner tracker & ECAL in magnetic field serve as beam monitor





ND – PRISM



Precision Reaction-Independent Spectrum Measurement (PRISM)

Off-axis data in different neutrino fluxes constrains energy dependence of neutrino cross sections, independent of interaction model





Current Canadian contributions



DAQ development at ProtoDUNE

- **ProtoDUNE**: FD prototypes at CERN
- Readout system tests & developments
- ProtoDUNE second run with SPS hadron beam





ND-LAr prototyping with 2x2 modules

- On cosmics and with Fermilab NuMI beamline
- Reusing MINERvA detector planes for "fast tracker"

ND-LAr event reconstruction





Latest developments

Far detector cavern excavation complete!



ProtoDUNE run 2 event display



ND-LAr prototype event display



DUNE's Sensitivities

Long-baseline analysis

 $N(\text{Observables}) = \int \frac{\text{Flux}(E_{\nu}, \text{time}) \times \text{Interaction prob}(E_{\nu}, \text{final state})}{\times \text{Detector Efficiency}(\text{final state}) \times \text{Osc}(E_{\nu})}$

- Measure event rates → product of oscillations and flux/interaction/detector models
- Near detector has lots of events and assumed to have **no oscillations** → **constrain the systematics**
- Far detector has oscillations \rightarrow apply systematic constraints \rightarrow infer oscillation parameter values

CP Sensitivity

- After 10 years exposure there is significant CP violation (δ_{CP} ≠ 0, π) discovery potential across true values of δ_{CP} and for both hierarchies
- DUNE can establish CPV over 75% of δ_{CP} values at >3 σ

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Mass Ordering Sensitivity

 Obtain a definitive answer for the mass hierarchy within 7 years, regardless of the values of the other oscillation parameters

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Resolution to oscillation parameters

Ultimate precision 6-16° in δ_{CP}

World-leading precision (for long-baseline experiment) in θ_{13} and $\Delta m^2 \rightarrow$ comparisons with reactor measurements are sensitive to new physics

Beyond the 3-flavour paradigm

See Sergio's talk next for more!

Broad range of L/E at ND and FD \rightarrow search for non-SM oscillations Sterile neutrino mixing can be probed by comparing 3-flavour and 3+1flavour neutrino disappearance and appearance probabilities

Summary

- DUNE will enable an exciting physics program and aims to make precise measurements of the oscillation parameters:
 - Definitively measure the MO regardless of other oscillation parameters
 - Sensitivity to **CPV** and θ_{23} octant

Thanks for your time!!!

References

- DUNE TDR II <u>arXiv:2002.03005</u>
- Long-baseline neutrino oscillation physics potential of the DUNE experiment Eur. Phys. J. C 80, 978 (2020)

Inside ProtoDUNE!

Backup slides

DUNE & Hyper-K Complementarity

- Very long baseline \rightarrow large matter effect
- On-axis location & broad energy spectrum
- Reconstruct ν energy over broad range \rightarrow LArTPC technology
- Highly-capable near detectors to constrain systematic uncertainties

- Shorter baseline \rightarrow small matter effect
- Off-axis location & narrow energy spectrum
- Lower energy and mostly quasi-elastic interactions → very large water Cherenkov detector
- Highly-capable near detectors to constrain systematic uncertainties

Together we'll have resolved parameter degeneracy, better precision, better parameter sensitivity irrespective of true δ_{CP} value, and increased sensitivity to new physics.

Why both DUNE & Hyper-K?

Effect on matter on neutrino oscillations complicates some measurements Matter does not have same effect on neutrino and anti-neutrino oscillations – complicates CPV measurement Possible strategies:

Small oscillations length (~300 km) = insignificant matter effects Off axis beam gives high flux at oscillation maximum, narrow energy range

Large oscillations length (~1000 km) = significant matter effects

On axis beam gives wide range of neutrino energy – differentiate CPV effects from matter effects through energy dependence

Why both DUNE & Hyper-K?

- Count v events in neutrino mode (x axis), and count v in antineutrino mode (y axis)
- Ellipses represent effect of δ_{CP} . Matter effect splits the NH and IH ellipses about y=x, sin² θ_{23} moves ellipses along y=x

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ProtoDUNE

Prototypes of 2 DUNE far detector (FD) modules, located at CERN

Two LArTPC designs:

- Horizontal drift (HD) technology
- Vertical drift (VD) technology
- ProtoDUNE Horizontal drift is an 800t active mass TPC, making it the largest LArTPC constructed.
- ProtoDUNE successfully operated in 2018 and is preparing for its second run now

Figure: CERN Neutrino Platform.

Figure: CERN. LHC P.A.2 ALICH SPS B. LHC P.A.1 ATLAS

HD

Segmented Pixelated ND LArTPC

- 130t of active Argon
- ~20 v events (energy from ~50 events)/ 10 µs spill (with a 1.2 MW beam)
 (FD sees 3.4 v events/hr)

- **Necessary to reduce pileup**
- Optical: the detector is sliced up int 7x5 modules to contain prompt scintillation light
- Charge: pixelated charge readout

Beyond the Engineering Drawings

- Have built four 14% scale modules for Near Detector, tested their performance with cosmic rays @ Bern
- Mapping out pixel response: 4mmx4mm pixels, checking our ability to simulate this new detection technique

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Neutrino Test of Near Detector Strategy

- Moving beyond tests of cosmic rays
- Want to test strategy in operating neutrino beam
- Fermilab is already home to most intense neutrino beam in the world
- Also reusing MINERvA | detector planes for "fast tracker"
- ~4 days of NuMI antineutrino data was taken in July 2024
- Plan to take data in Spring 2024!
- 4 modules
- 2.4 instrumented tons of Aron

Exposure

Staging scenario is FD3 after 3 years, FD4 after 5 years

Mass ordering and CP sensitivity – best case

For best-case oscillation scenarios, DUNE has

- >5 σ mass ordering sensitivity in 1 year
- $>3\sigma$ CPV sensitivity in 3.5 years

Mass ordering and CP sensitivity – otherwise

For worst-case oscillation scenarios, DUNE has $>5\sigma$ mass ordering sensitivity in 3 years In long term, DUNE can establish CPV over 75% of δ_{CP} values at >3 σ

