

SUPERCDCMS SNOLAB

A SNAPSHOT OF THE DARK MATTER EXPERIMENT



Presented at the Lake Louise Winter Institute

19th – 24th February, 2023

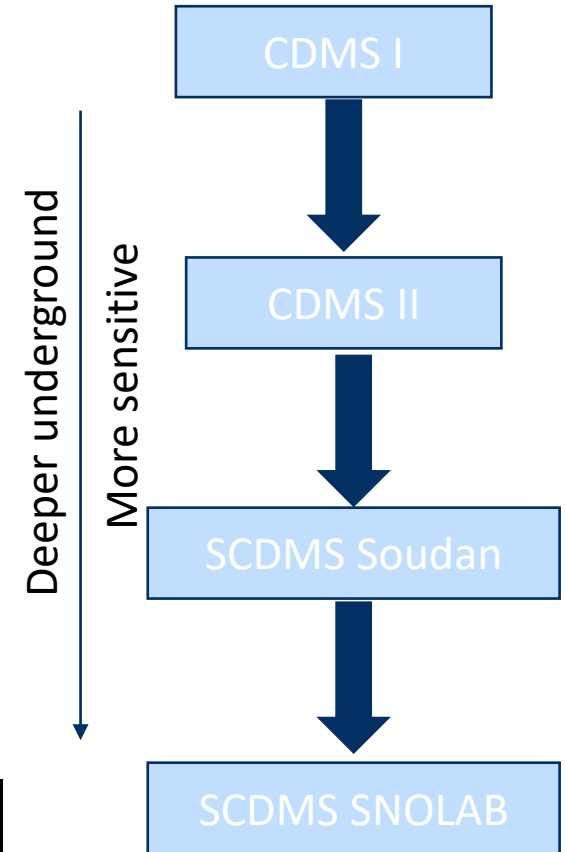
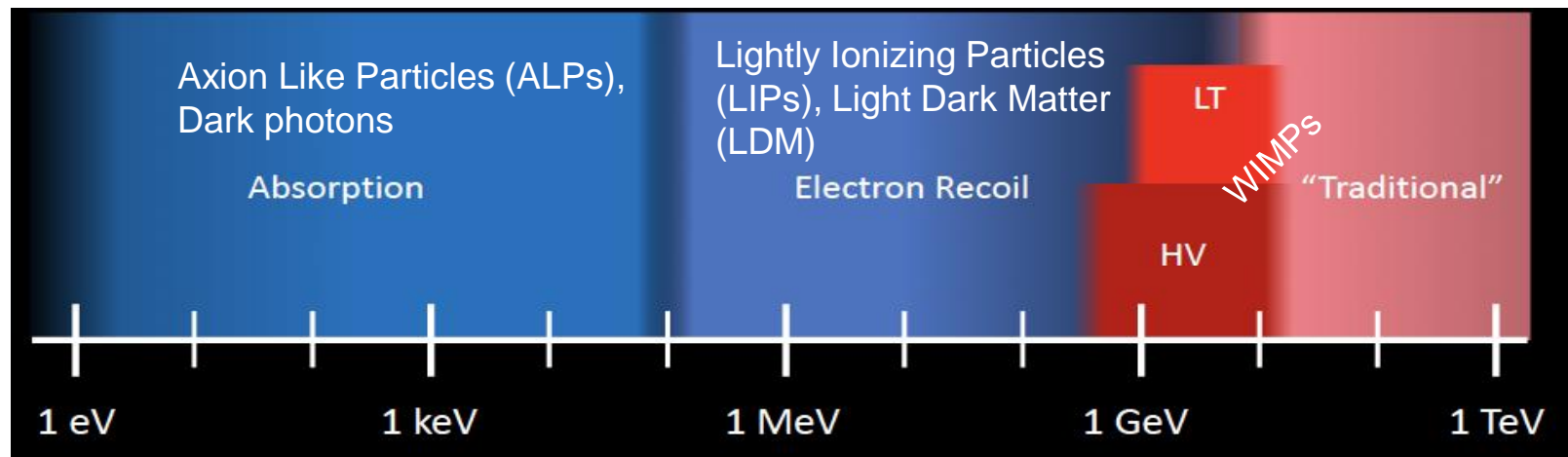
Aditi Pradeep
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University of British Columbia

on behalf of the SuperCDMS collaboration

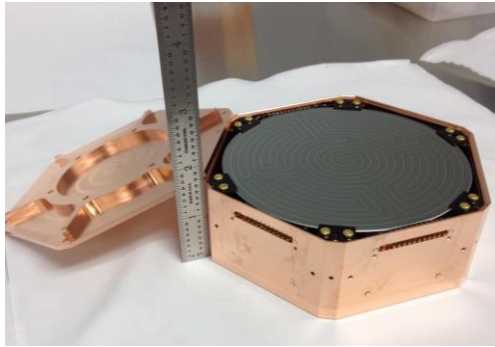
OVERVIEW

- ✓ Many theories motivate DM searches at lower energy scales (< 10 GeV)
 - ✓ Asymmetric DM that could explain matter antimatter asymmetry
 - ✓ Hidden sector DM with weak couplings to SM matter
 - ✓ Axions and Axion-Like Particles

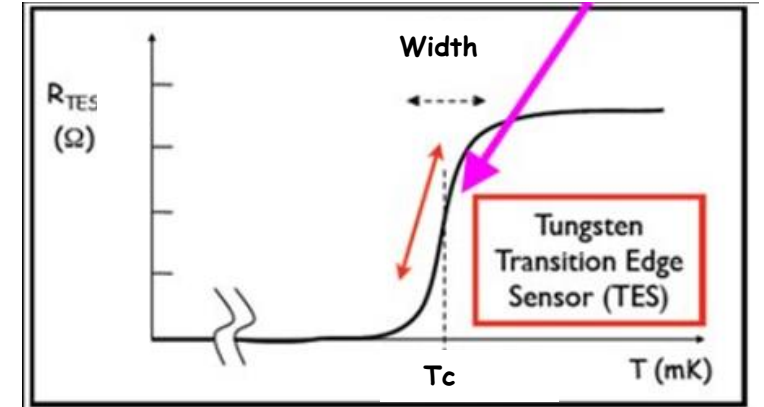
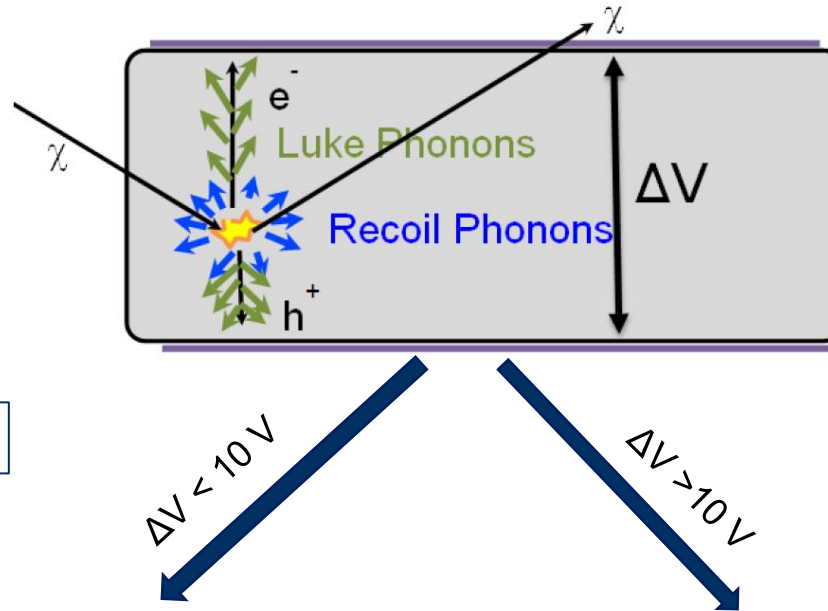
SuperCDMS SNOLAB searches for dark matter across a spectrum of candidates using different detection channels



HOW DO WE FIND DARK MATTER?



Prototype iZIP detector

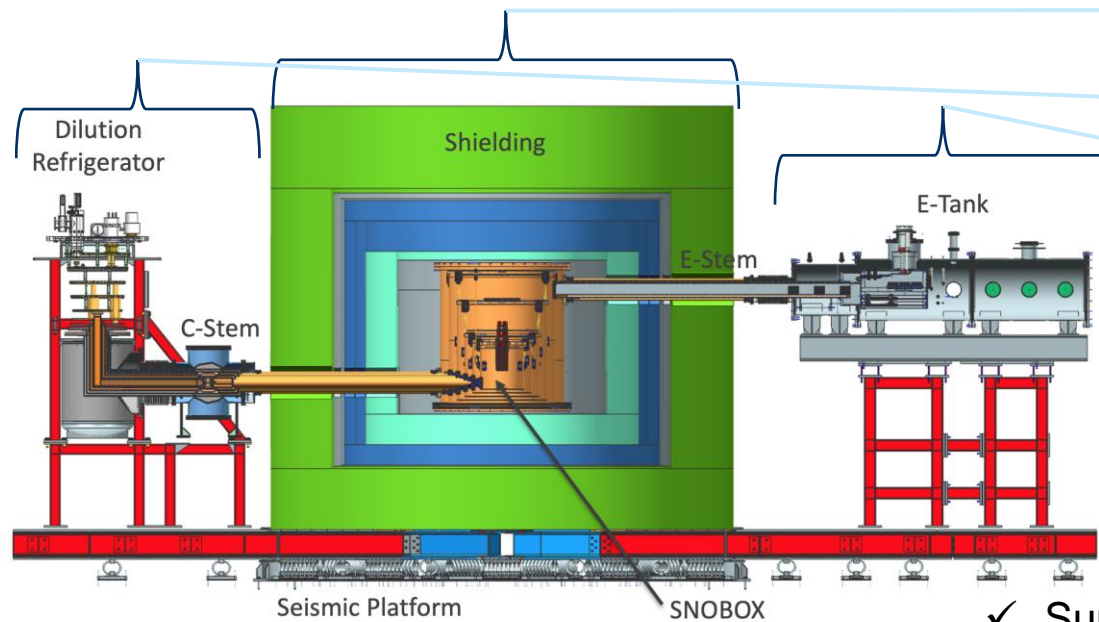


Phonon signal collected via Tungsten TES sensors; charge signals collected at electrodes

- ✓ Recoil/prompt phonons dominate for nuclear recoils
- ✓ Readout 12 phonon + 4 charge channels
- ✓ Low background detectors (event by event ER/NR discrimination) : **iZIP**
- ✓ Threshold of 133 eV_{pt} (231 eV_{pt})* in Si (Ge)
- ✓ Resolution of 19 eV_{pt} (33 eV_{pt})* in Si (Ge)

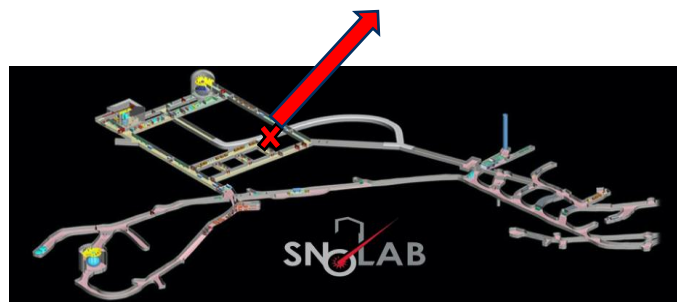
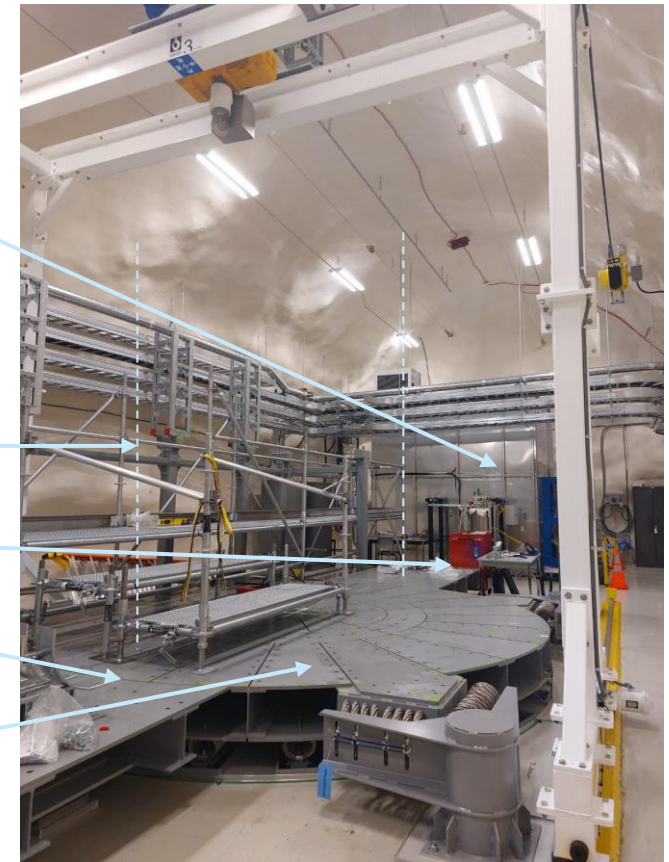
- ✓ Luke phonons dominate
- ✓ Readout 12 phonon channels
- ✓ Low threshold detectors: **HV**
- ✓ Threshold of 91 eV_{pt} (238 eV_{pt})* in Si (Ge)
- ✓ Resolution of 13 eV_{pt} (34 eV_{pt})* in Si (Ge)

SUPERCDCMS @ SNOLAB



Low Radon
cleanroom

Seismic
platform



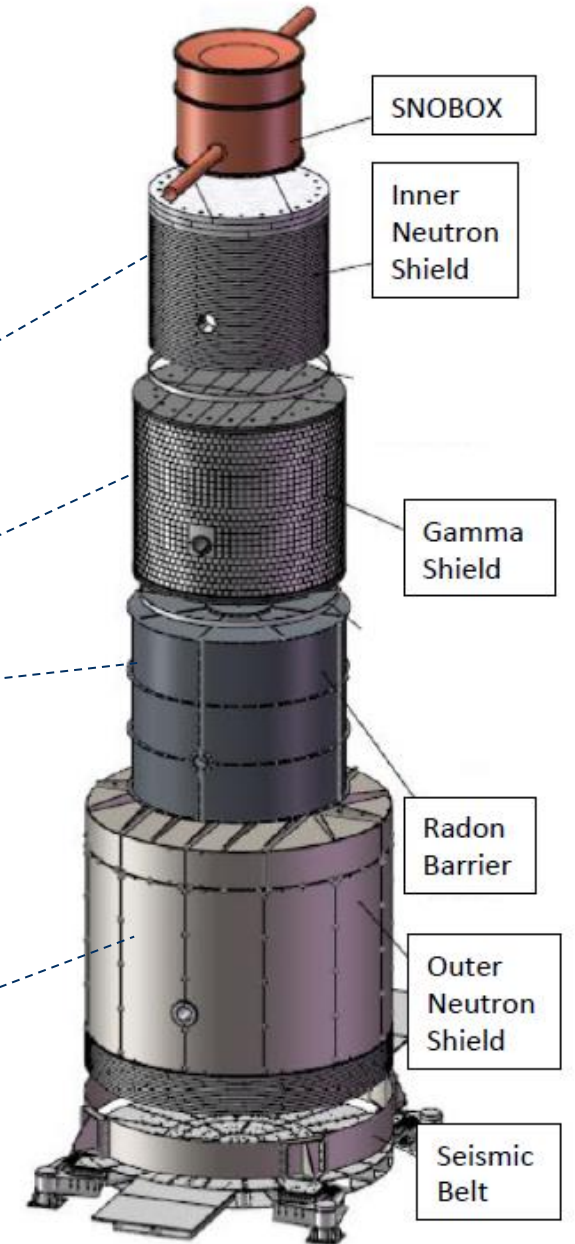
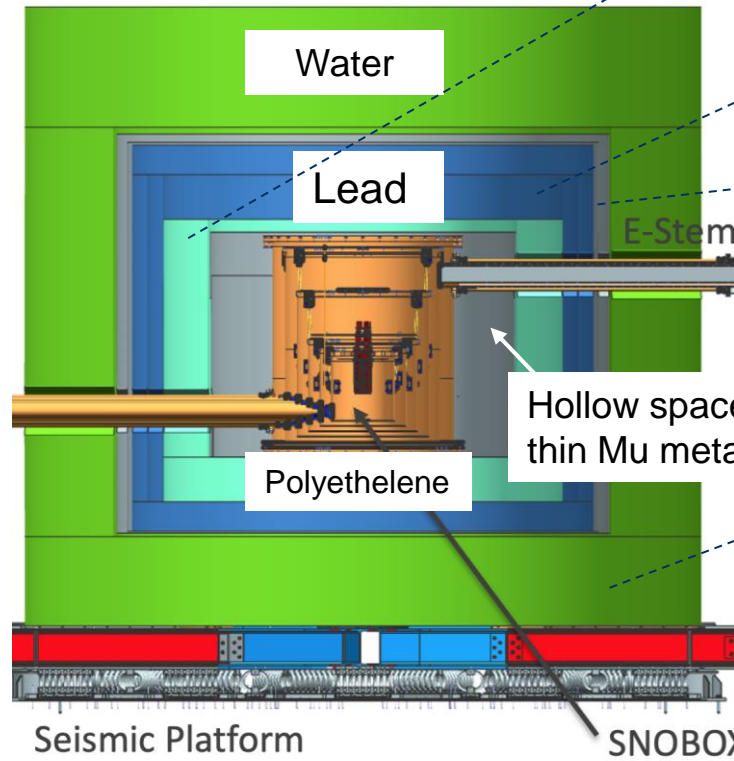
- ✓ SuperCDMS SNOLAB is under construction in SNOLAB, the deepest underground clean lab in North America
- ✓ SNOLAB provides 2 km natural rock overburden
- ✓ Commissioning to begin this year!

SHIELDING

- ✓ Several layers of nested shielding
- ✓ Inner layers attenuate by-products from outer layers: “Self shielding”
- ✓ Mu metal provides magnetic shielding

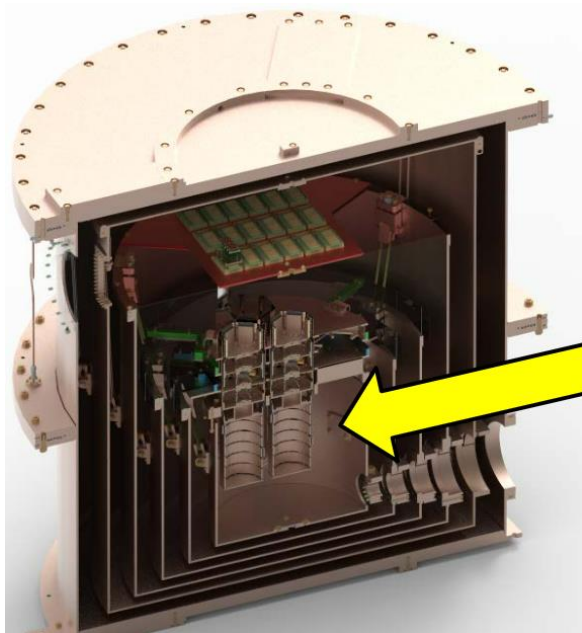


Completely assembled Lead shield



DILUTION FRIDGE & SNOBOX

- ✓ Commercial dilution fridge that cools to temperatures of order ~ 10 mK
- ✓ SNOBOX contains 7 nested Cu cans for staged cooling
- ✓ Innermost can houses detector towers!



SCDMS SNOLAB SNOBOX



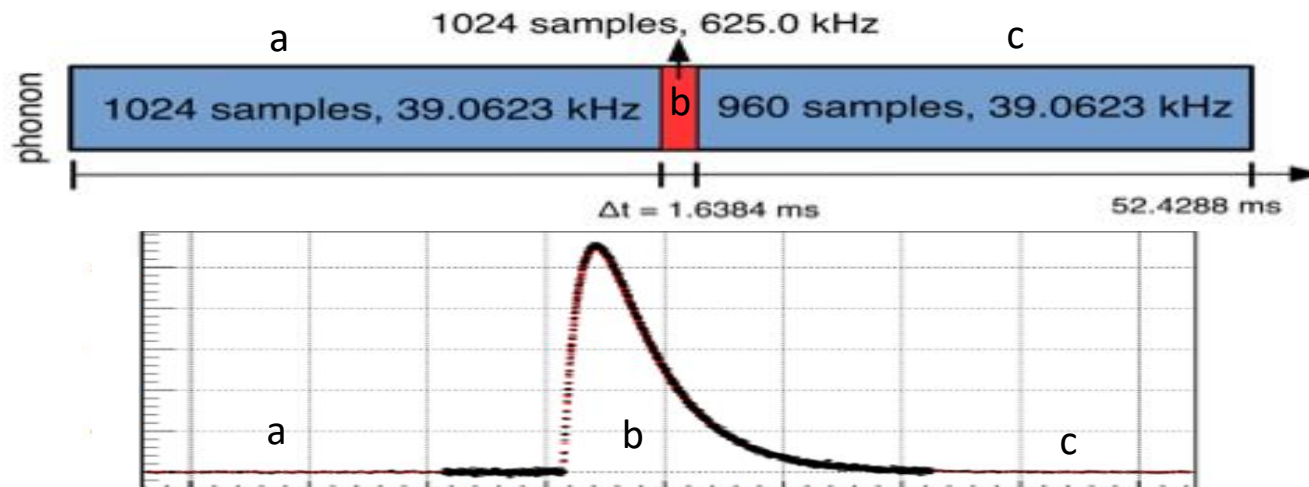
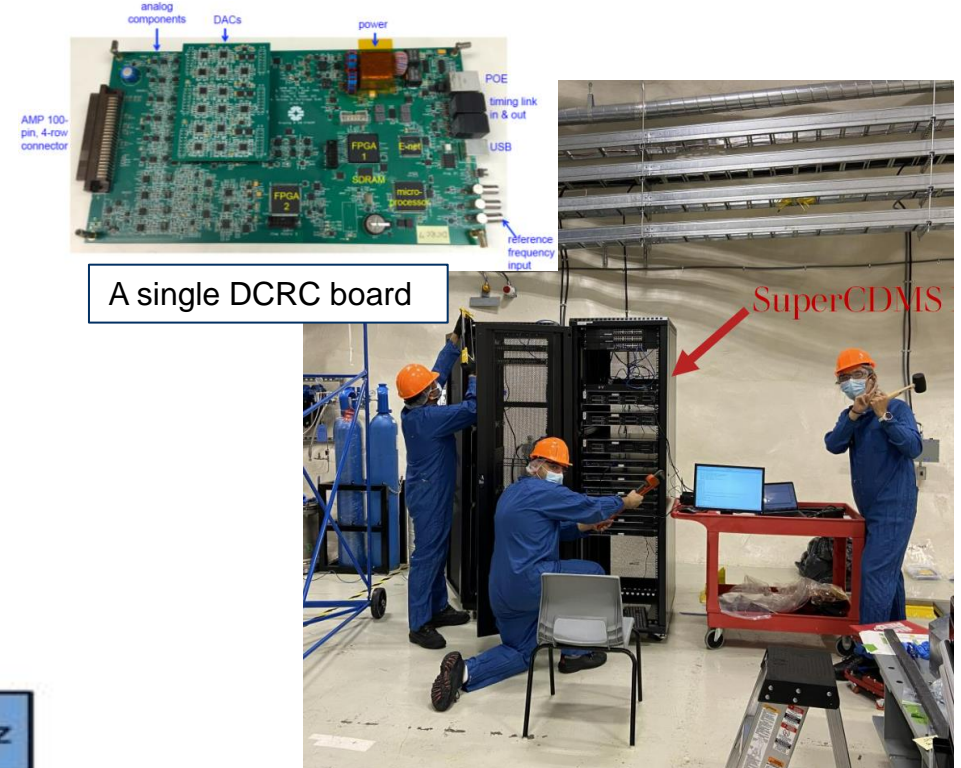
Inside the SCDMS fridge



The fridge has arrived at SNOLAB and will be installed soon. SCDMS gas handling system (blue box on the left with buttons) also currently at SNOLAB.

DATA ACQUISITION & PROCESSING

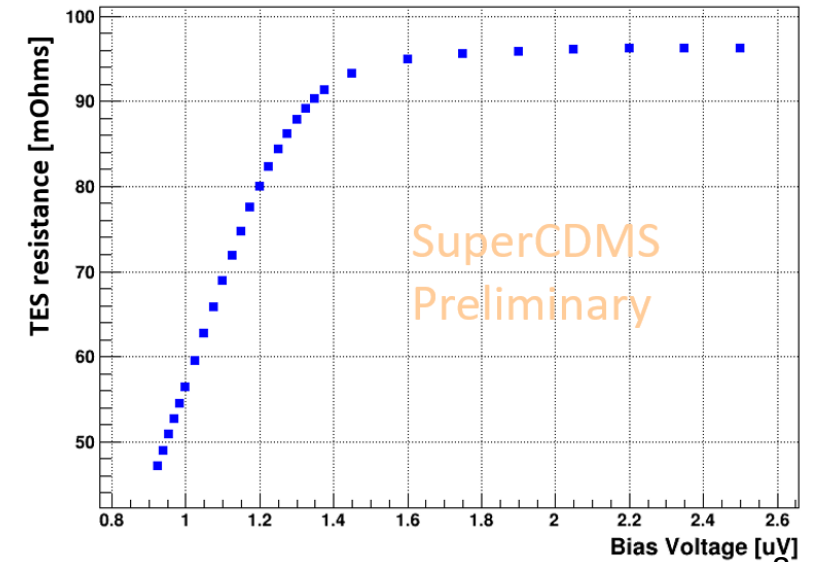
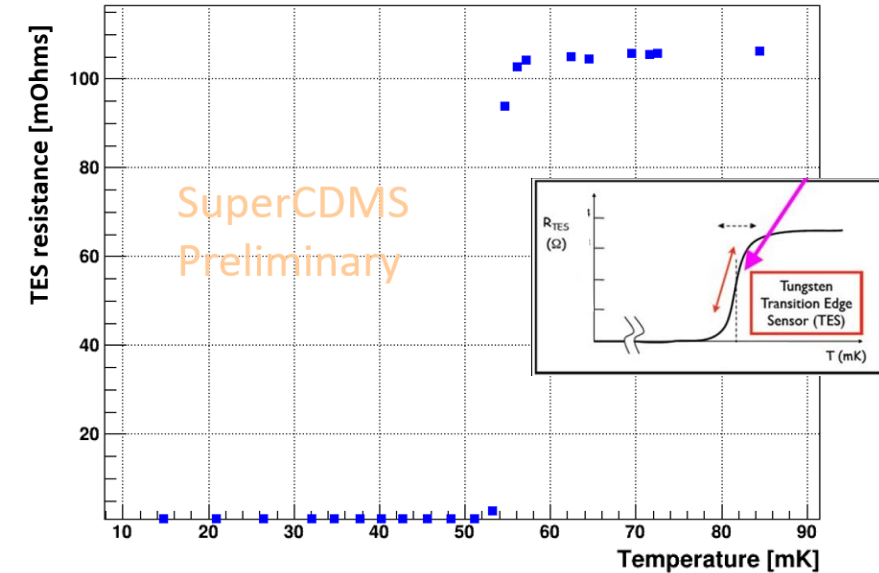
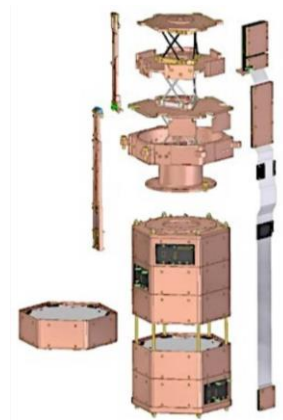
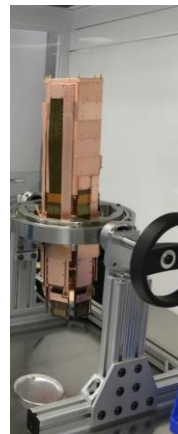
- ✓ E-Tank houses DAQ electronics
- ✓ Detector Control and Readout Cards (DCRCs) read out data from detectors (throughput of ~5 MB/sec)
- ✓ DAQ is limited by bandwidth of electronics; data will be **hybrid sampled**
- ✓ Advanced processing algorithms developed and deployed
- ✓ Massive computing effort: 100 TB/yr of anticipated raw data



All 36 DCRCs for SCDMS SNOLAB and other DAQ hardware installed at SNOLAB

TOWER TESTING: SLAC

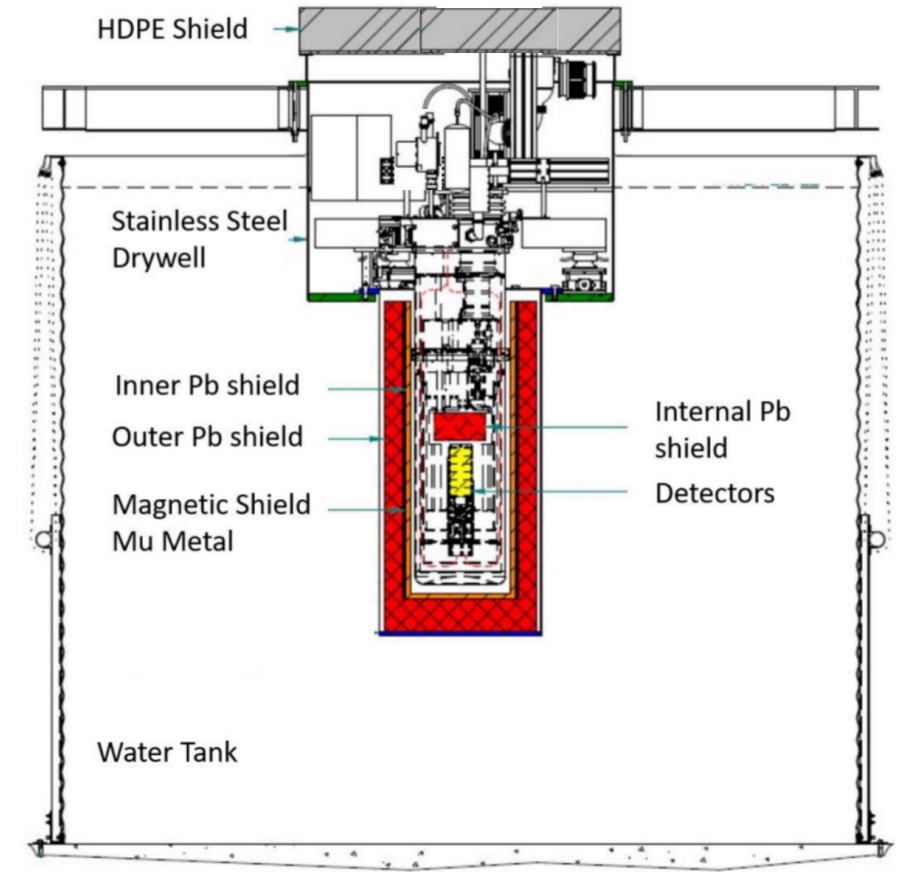
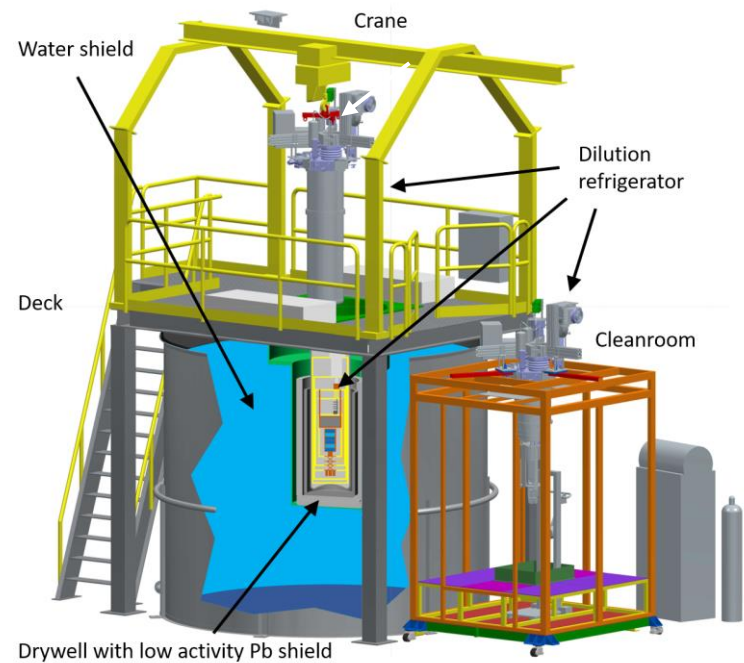
- ✓ 4 detector towers: 2 iZIP (T1 & T4) + 2 HV (T2 & T3) towers
- ✓ Detector composition: [6 Ge] (T1) + [4 Ge, 2 Si] (T2, T3, T4)
- ✓ **All towers were tested successfully for basic functionality at SLAC**
- ✓ **Crystal properties measured (T_c , resistances) indicate optimal sensitivity & thresholds**
- ✓ Limitations: surface testing with single layer of Cu shielding- high background & event rates; source-less testing, i.e., no energy scale estimate (yet!)
- ✓ More testing at CUTE!



DETECTOR TESTING: CUTE

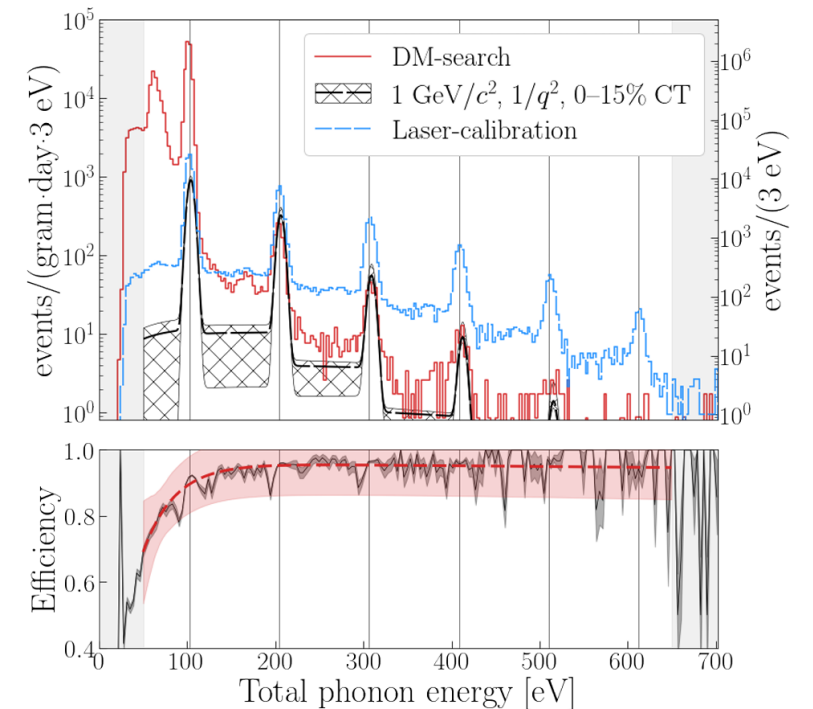
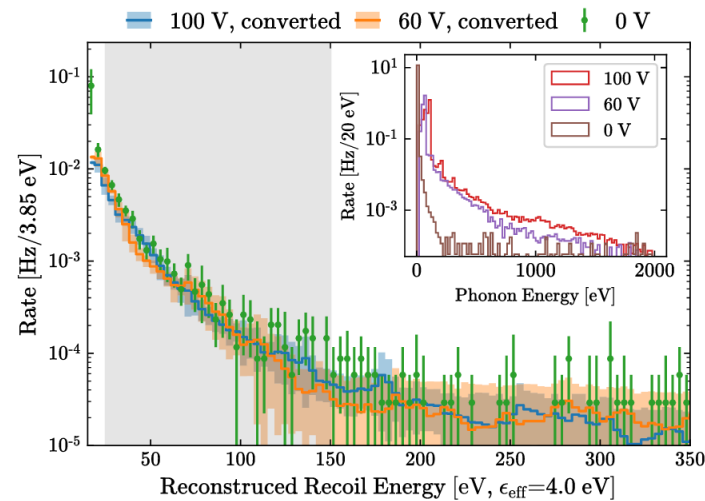
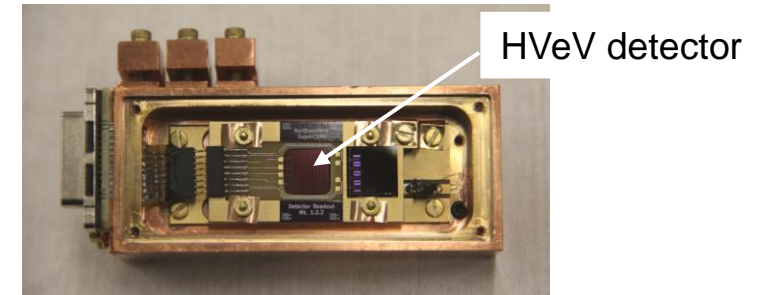
- ✓ Low bkgd testing of SCDMS detectors for early science
- ✓ Purging out potential software & hardware issues
- ✓ Several prototype and RnD devices already tested
- ✓ Tower testing & calibration to begin in a couple of months

CUTE facility



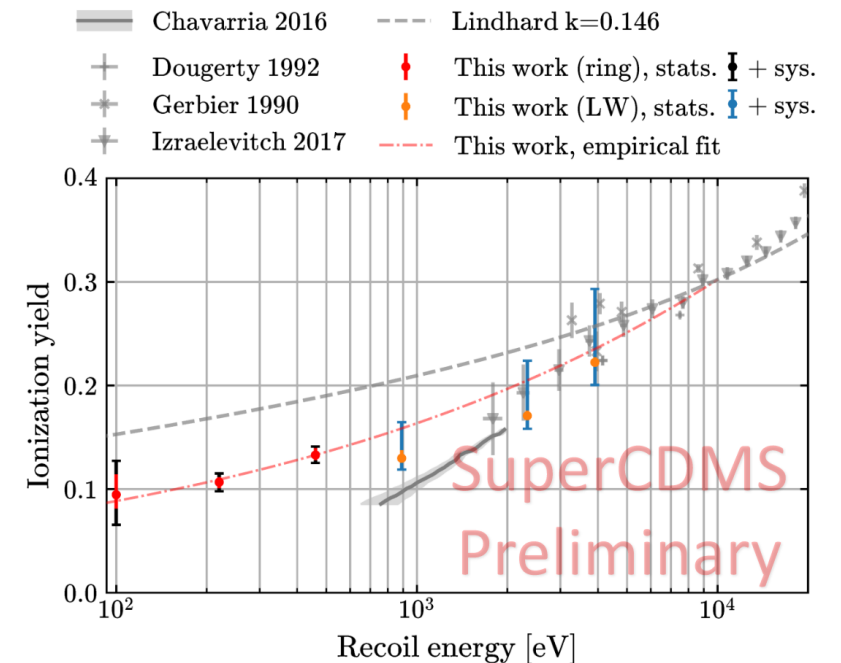
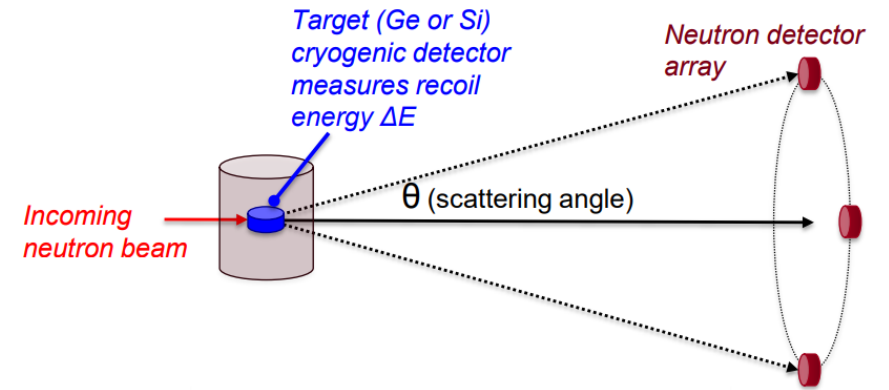
DETECTOR TESTING: NEXUS

- ✓ Testing gram-scale, low threshold devices
- ✓ HVeV devices have single e^-/h^+ pair resolution ($\sim 3\text{eV}$)
- ✓ Observed low energy excess events similar to other experiments (but consistent with background)
- ✓ Relevant publications: Phys. Rev. D 105, 112006 (2022), Phys. Rev. D 102, 091101(R)



IONIZATION YIELD: IMPACT

- ✓ Ionization yield is instrumental to background discrimination in SCDMS
- ✓ Total phonon energy, $E_T = E_R \left(1 + Y(E_R) \frac{eV_b}{\epsilon_{eh}} \right)$, where, E_R : recoil energy, Y : yield, V_b : applied bias, ϵ_{eh} : average e^-/h^+ pair production energy
- ✓ Yield described by Lindhard theory. Deviations observed below 5 keV_{NR}
- ✓ IMPACT is a neutron scattering experiment to measure Y at low recoil energies
- ✓ **Look out for upcoming publication for Si ionization yield!**

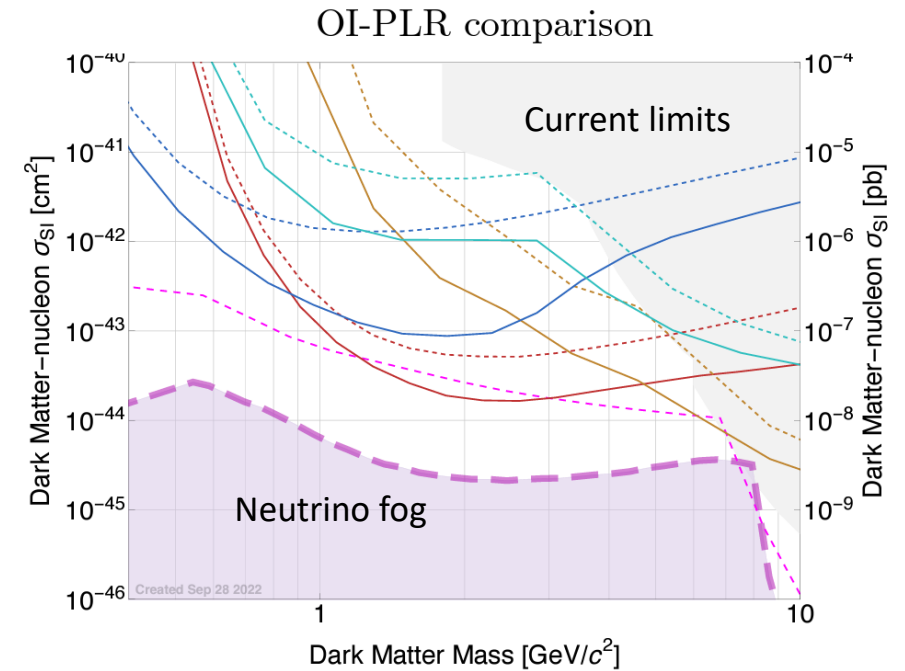


SCDMS SNOLAB SENSITIVITY

- ✓ SNOMASS paper: [arXiv:2203.08463](https://arxiv.org/abs/2203.08463)
- ✓ Current and future sensitivity of SCDMS SNOLAB experiment for NRDM, ERDM (ALPs, dark photons, LDMs)
- ✓ Several detector upgrade & bkgd scenarios considered

Other upcoming papers:

- ❖ EFT analysis of CDMSLite Run 2 data : [arXiv:2205.11683](https://arxiv.org/abs/2205.11683)
- ❖ Inelastic DM scattering analysis with CDMSLite Run 3



Comparison of expected NRDM sensitivity for SCDMS SNOLAB between OI (dashed) and Profile likelihood methods (solid)

Legend: Ge HV; Si HV; Ge iZIP; Si iZIP; “single neutrino” sensitivity, where one neutrino event can be expected on average;

We
Are
SUPERCDMS



BACKUP SLIDES

DILUTION FRIDGE SCHEMATIC

