

# **PROJECT 8**

# A FREQUENCY-BASED TRITIUM ENDPOINT EXPERIMENT

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#### Goal of the Project 8 Experiment

Measure the absolute neutrino mass  $m_\beta$  in case of inverted ordering, or limit  $m_\beta$  to 0.04 eV in the case of normal ordering with smaller mass.

#### Outline

- Neutrino mass determination by measuring cyclotron radiation frequency.
- 2. Challenges ahead, a **phased approach**.
- 3. Quantum-limited microwave amplification.





Figure by Symmetry Magazine

# **NEUTRINO MASS MEASUREMENT: STATUS**





#### Direct methods:

- Rely on the distributions of the neutrino and electron **kinetic energy** in β-decay processes.
- Current experimental limit from KATRIN is  $0.8\,{\rm eV/c^2}.$  The projected sensitivity is  $0.2\,{\rm eV/c^2}$

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Figure by GERDA



Figure by Symmetry Magazine

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Indirect methods:

- Neutrino Oscillations set a lower bound on  $m_{\beta}$ .
- Neutrinoless double beta decay: Currently not yet observed.
- **Cosmology**, through the signatures of growth and evolution of large scale structures in the cosmic microwave background.
- $\rightarrow$  Indirect is model dependent.

# NEUTRINO MASS MEASUREMENTS: TRITIUM ENDPOINT METHOD



How to **scale down** the experiment while **increasing** the neutrino mass **sensitivity**?

- 1. Use a **source inside** the **detector** volume.
- 2. The use of **cyclotron radiation** emitted by electrons in a magnetic field enables **frequency detection** of microwaves.

The **Project 8 collaboration** employs Cyclotron Radiation Emission Spectroscopy (**CRES**)

# Tritium decays inside a **uniform** $\vec{B}$ -field.



Electron performs **cyclotron motion** with frequency

$$f(B, E_{kin}) = \frac{1}{2\pi} \frac{eB}{m_e + E_{kin}/c^2}$$

PHASE I

- First detection of cyclotron radiation from a single electron. Gaseous <sup>83m</sup>Kr used as a source.
  Phys. Rev. Lett. 114, 1162501 (2015)
- Analysis steps
  - 1. Time series
  - 2. Frequency spectrum each  $\mathcal{O}(0.1\,\mathrm{ms})$ .





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  - 3. 2D spectrogram
  - 4. Track identification
  - 5. Start frequency fitting







#### PHASE I

PHASE II

• First Project 8 **limit on the neutrino mass** using gaseous tritium and a waveguide antenna.

Publication in preparation



Figure by A. Lindman

 First Project 8 limit on the neutrino mass using gaseous tritium and a waveguide antenna.







#### Figure by A. Lindman

Figure by T. Weiss

PHASE III

Atomic tritium is required due to uncertainty in final states of molecular tritium.





Wouter Van De Pontseele

Figure by M. Fertl

Atomic tritium is required due to uncertainty in final states of molecular tritium.

- ⇒ Create a pure atomic tritium source delivering  $\mathcal{O}(10^{18} \text{atoms /m}^3)$  density.
- A **dissociator** to produce atoms from molecules (thermal cracker).
- + Accomodator to cool down the atoms to  ${\approx}8\,{\rm K}$
- Velocity and state based selection of trappable atoms.
- A magnetic neutral atom trap (Halbach of Ioffe array)



# PHASE III

# ATOMIC TRITIUM DEMONSTRATOR

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All steps being developed at Project 8 institutions High-flow hydrogen cracking demonstrated.



Free space CRES detection with an antenna array.

Current R&D focus!

- Understanding the different components of the electron trajectory and its radiation.
- · Commercial ≈1T MRI magnet deployed at University of Washington.
- Magnetic field and efficiency calibration with an electron gun.
- Reconstruction techniques using digital beam forming and matched filters.



# **OPTION 2: MICROWAVE CAVITY**

# **Cavity** CRES detection. Current R&D focus!

- Higher radiation energy collection.
- Allows for lower magnetic fields and tritium density.
- Mode filtering and open ended cavity prototype design in progress.



# Phase IV, putting it all together!

Neutrino mass experiment covering the inverted ordering, probing  $m_{\beta} \approx 40 \text{ meV}$ .



• The signal power of cyclotron radiation per electron in a fixed magnetic field is constant.  $P \sim B^2$ ,  $P \approx 1 \,\text{fW}$  for  $B \approx 1 \,\text{T}$ 

SNR ratio needs to be maximised by minimising the noise.

**QUANTUM AMPLIFIERS & PROJECT 8** 

# JOSEPHSON TRAVELING-WAVE PARAMETRIC AMPLIFIER (JTWPA) [MACKLIN ET AL., 2015]



Chain of phase matched cells with Josephson Junctions, bandwidth  $\mathcal{O}(GHz)$ .





# MEASUREMENT SETUP FOR JTWPAS AT MIT

- **Dilution refrigerator** setup using network analyser and spectrum analysers for readout.
- through line for insertion loss measurements.
- Open Transmission line **qubit** for **power calibration** and noise measurements.
- Proto-type JTWPAs in the 4 GHz to 7 GHz and O(23 GHz) range, fabricated by Lincoln Laboratories.

Cryogenic testing of antenna and cavity designs using generated waveforms starting. Wouter Van De Pontseele



# **MEASUREMENT SETUP FOR JTWPAS AT MIT**





# CONCLUSIONS

#### Project 8: A Frequency-based Neutrino Mass Measurement

- Measure electron kinetic energy from tritium decays to obtain the neutrino mass.
- · In practice: A high-precision cryogenic microwave frequency experiment.
- Phase I&II completed

<sup>83m</sup>Kr calibration and **molecular tritium endpoint** measurement **using CRES**.

#### Phase III: Paving the path towards phase IV

- Demonstration of Scalibility: Explore CRES in a large source volume with free-space antenna or cavity-based detection. Measure the tritium endpoint with a  $m_{\beta}$  sensitivity of  $\mathcal{O}(5 \text{ eV})$ .
- Atomic Tritium Demonstrator A pure atomic tritium source at the optimum Phase IV density of  $\mathcal{O}(10^{18} \text{atoms /m}^3)$ .

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# PROJECT 8: THE CHALLENGES OF SCALING UP

- An accurate neutrino mass measurement relies on **high statistics** in the tail of the tritium endpoint.
- Scattering time and atomic tritium source/trap restrictions constrain the density.
- $\rightarrow$  Phase IV will need a high triggering efficiency in combination with a large volume.
  - The **power of cyclotron radiation** per electron is

 $P \sim B^2$ ,  $P \approx 1 \, \text{fW}$  for  $B \approx 1 \, \text{T}$ 

In a constant field, the signal power is constant

• The proposed free space multi-antenna readout will have a lower coverage and lower received power per channel.

ightarrow The signal to noise power ratio needs to be maximised by minimising the noise power.

TIME LINE











# QUANTUM AMPLIFIERS: INTRODUCTION

# Driven by Quantum Computing

- Superconducting qubit signals are weak microwaves.
- First stage amplifier limits performance.
- Amplifier **bandwidth** enables **multiplexing** of multiple qubits.

# Similarities with Project 8

- Cyclotron emission is in the microwave region.
- Trigger efficiency ultimately depends on the noise performance of the first stage amplifier.
- Interest in multiplexing of antenna channels.
- · Bandwidth requirements driven by calibration sources.



# PARAMETRIC AMPLIFICATION [FASOLO ET AL., 2020]



# Principle: A non-linearity and a pump wave

- Swing process: amplification by changing the centre of mass (**pump**) with a certain frequency and phase.
- Amplification of a signal by exchanging pump power into signal and idler. Exploits Josephson junction as a non-linear circuit element.
- Parametric amplification requires the pump ( $\omega_p$ ), signal ( $\omega_s$ ),and idler ( $\omega_i$ ) frequencies to satisfy **energy** conservation,

$$2\omega_p = \omega_{\rm s} + \omega_i$$

and momentum conservation (phase matching),

$$\Delta k = 2k_p - k_s - k_i = 0$$

# JOSEPHSON PARAMETRIC AMPLIFIERS [ELO ET AL., 2019]

- Single cell with Josephson Junction, small footprint.
- Circular device, input and output over the same line.
- Small amplification bandwidth of *O*(100 MHz).
- Demonstrated with central frequency from 0.6 GHz to 7 GHz.
- $\cdot\,$  Loops are sensitive to magnetic fields.



# WHY QUANTUM AMPLIFIERS: THE NOISE TEMPERATURE

Quantum amplifiers have a noise temperature an order of magnitude below the current best off-shelf cryogenic amplifiers.

Current R&D to demonstrate JTWPA's in the frequency range and magnetic fields proposed by Project 8. Noise Temperatures for different amplifiers



# TWPA CHARACTERISATION: GAIN AS A FUNCTION OF PUMP POWER/FREQUENCY



Transmission network analyser measurement, compared to a JTWPA without pump.



# TWPA CHARACTERISATION: NOISE AS A FUNCTION OF PUMP POWER/FREQUENCY



Spectrum analyser measurement of the noise floor, compared to a JTWPA without pump.



# TWPA CHARACTERISATION: SIGNAL TO NOISE RATIO IMPROVEMENT



#### JTWPA OPTIMAL PUMP SETTINGS COMPARED









### JTWPA OPTIMAL PUMP SETTINGS COMPARED

In practice, optimal operating point chosen to limit the spurious peaks.

Gain above 20 dB with a Signal-to-Noise Ratio Improvement (SNRI) above 15 dB in the 3 GHz bandwidth.





