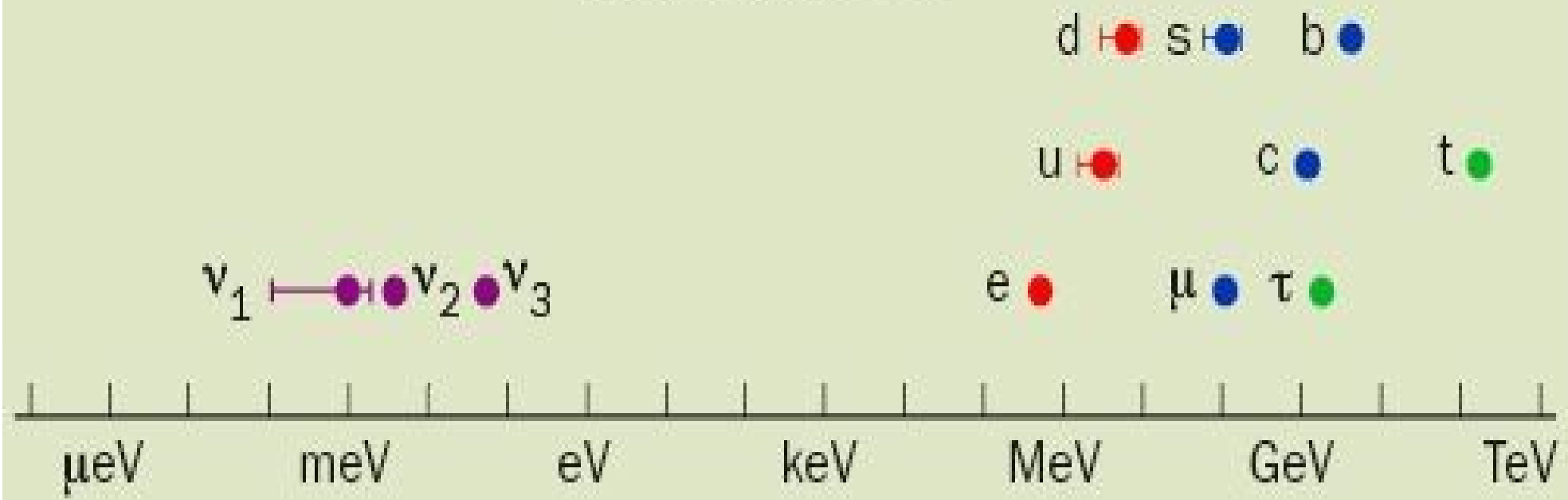


# Lecture 2

*In which the origin of mass is considered and  
unsuccessfully measured*

# The mystery of neutrino mass

## fermion masses



Why are neutrino masses so small?

# $\nu$ Mass in the Standard Model

Dirac Lagrangian mass term for fermions contains a mass term with a Dirac mass,  $m_D$

$$L_\nu = \bar{\psi} (i \gamma_\mu \partial^\mu - m_D) \psi \Rightarrow L_{mass} = m_D \bar{\psi} \psi$$

Can rewrite mass term in terms of chiral states

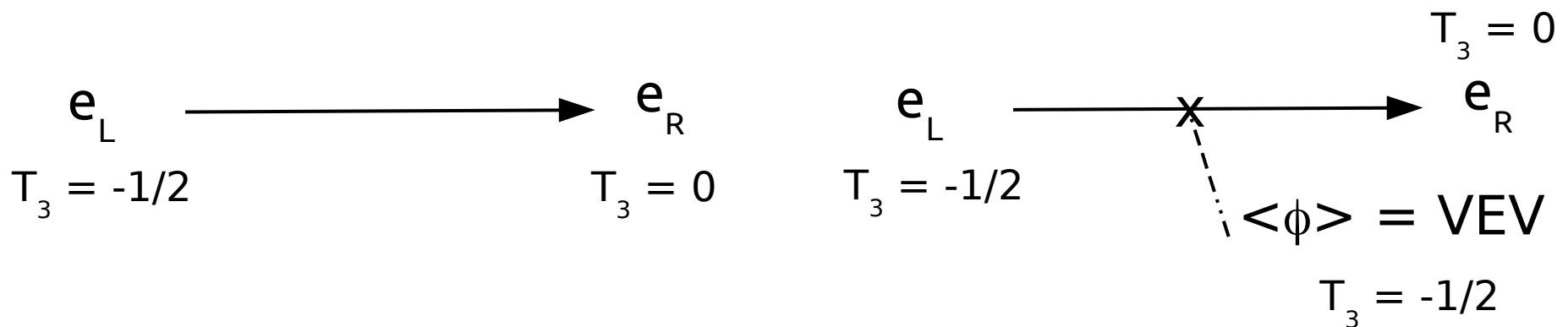
$$L_{mass} = m_D \bar{\psi} \psi = m_D (\bar{\psi}_L + \bar{\psi}_R) (\psi_L + \psi_R) = m_D (\bar{\psi}_L \psi_R + \bar{\psi}_R \psi_L)$$

Mass term is the only place that the L- and R- chiral sectors of the SM meet.

Unfortunately, as it stands, such a term does *not* preserve gauge invariance. You need the Higgs mechanism to fix this.

# Higgs mechanism

$$L_{mass} = m_D (\bar{\psi}_L \psi_R + \bar{\psi}_R \psi_L) \rightarrow Y_\psi \langle \phi \rangle (\bar{\psi}_L \psi_R + \bar{\psi}_R \psi_L)$$

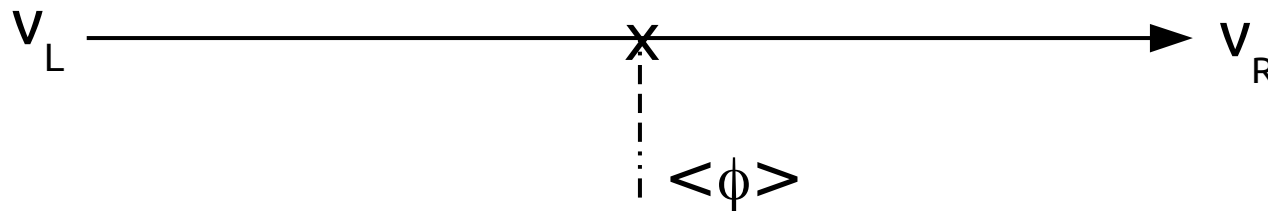


**Dirac mass:**  $m_D = Y_\psi \langle \phi \rangle$   $\langle \phi \rangle = 246 \text{ GeV}$

- ▶ Higgs mechanism provides a means to give mass to fermions
- ▶ Preserves gauge invariance of the mass term
- ▶ Does not predict the mass, however. Still need to measure the Yukawa coupling.

# Neutrino Dirac Mass

$$L_{mass} = Y_\nu \langle \phi \rangle (\bar{\nu}_L \nu_R + \bar{\nu}_R \nu_L)$$



- ▶ Addition of a sterile right-handed neutrino state to the SM which is, in principle, undetectable (apart from flavour oscillations)
- ▶ Tiny  $m_\nu$  implies tiny Yukawa coupling :  $Y_\nu < 10^{-13}$ 
  - ▶ Smallness of neutrino mass is not addressed by this mechanism

# Majorana Neutrinos

Mass terms need a R-chiral field. Neutrinos only have L-chiral field.

Can one build a R-chiral field only from the L-chiral field?

Yes : Ettore Majorana showed  $\nu_L^C = C \bar{\nu}_L^T$  is right-handed

C = charge conjugation operator

Can form a *Majorana* neutrino :  $\nu = \nu_L + \nu_L^C$

This is self-conjugate :  $\nu = \nu^C$  : **particle is identical to the antiparticle**

The neutrino is the only fundamental fermion with potential to be Majorana.

We can also now write down a mass term for Majorana neutrinos

$$L_{Maj} = \frac{1}{2} m_L (\bar{\nu}^C \nu + \bar{\nu} \nu^C) = \frac{1}{2} m_L (\bar{\nu}_L^C \nu_L + \bar{\nu}_L \nu_L^C)$$

We are now coupling neutrinos and antineutrinos, leading to a process which **violates lepton number by 2**

# Damn

The left-handed Majorana mass term also violates gauge invariance.

$$\overline{\nu_L^C} \nu_L \quad \nu_L \xrightarrow{\quad \times \quad} \nu_L^C$$

$Y = -1 \quad \xrightarrow{\Delta Y = +2} \quad Y = +1$

To maintain gauge invariance this has to couple to a Higgs-y thing with  $Y = +2$  and  $T_3 = 1$  - that is a Higgs weak triplet with hypercharge +2.

No such field exists in the Standard Model (although you do get them if you expand the Higgs sector to include both a scalar doublet and triplet)

We are forced then to consider the existence of an independent right-handed U(1) singlet Majorana neutrino field :  $N = N_R^C + N_R$

The existence of neutrino mass implies physics beyond the Standard Model, either from a right-handed state needed for the Dirac mass mechanism, or a Higgs triplet, or a new mass mechanism.



# The general mass term

Suppose : once upon a time there were 2 Majorana neutrinos.  
An almost massless one, and a very heavy one. The mass term looks like

$$L_{mass} = m \overline{\nu}_m \nu_m + M \overline{N}_m N_m = \begin{pmatrix} \overline{\nu}_m & \overline{N}_m \end{pmatrix} \begin{pmatrix} m & 0 \\ 0 & M \end{pmatrix} \begin{pmatrix} \nu_m \\ N_m \end{pmatrix} \quad \begin{array}{l} \text{Written in the mass basis} \\ \text{States of definite mass} \end{array}$$

We have, potentially, 4 separate chiral fields to play with :

$$\nu_L, \nu_L^C, N_R, N_R^C$$

If we're resigned to having right-handed fields anyway we can write down 4 different mass terms

$$\left. \begin{array}{l} L_L^M = m_L \overline{\nu}_L^C \nu_L \\ L_R^M = m_R \overline{N}_R^C N_R \\ L_L^D = m_D \overline{N}_R^C \nu_L \\ L_R^D = m_D \overline{\nu}_L^C N_R^C \end{array} \right\} \begin{array}{l} \text{Two Majorana mass terms} \\ \text{Two Dirac mass terms} \end{array}$$



# The general mass term

The most general mass term combines all of these

$$L_{mass} = L_L^D + L_R^D + L_L^M + L_R^M$$

$$L_{mass} = \begin{pmatrix} \overline{\nu}_L^C & \overline{N}_R \end{pmatrix} \begin{pmatrix} 0 & m_D \\ m_D & m_R \end{pmatrix} \begin{pmatrix} \nu_L \\ N_R^C \end{pmatrix}$$

I've set  $m_L = 0$  because of the gauge issue.

Mass eigenstates are mixes of the chiral eigenstates

Physical masses are the eigenvalues of the diagonalised mass matrix ( $m_1, m_2$ ).

$$\begin{pmatrix} m & 0 \\ 0 & M \end{pmatrix} = Z^{-1} \begin{pmatrix} 0 & m_D \\ m_D & m_R \end{pmatrix} Z \quad m, M = \frac{1}{2} \left[ m_R \pm \sqrt{m_R^2 + 4 m_D^2} \right]$$

# See-Saw mechanism

$$m, M = \frac{1}{2} \left[ m_R \pm \sqrt{m_R^2 + 4 m_D^2} \right]$$

- ▶ M is the mass of a right-handed (singlet) neutral fermion
- ▶ Suppose that this is around the GUT scale :  $\Lambda$

$$M \sim m_R \sim \Lambda \qquad m \sim \frac{m_D^2}{m_R} \sim \frac{\langle \Phi \rangle^2}{\Lambda}$$

right-handed  
heavy neutral lepton



“our” neutrino

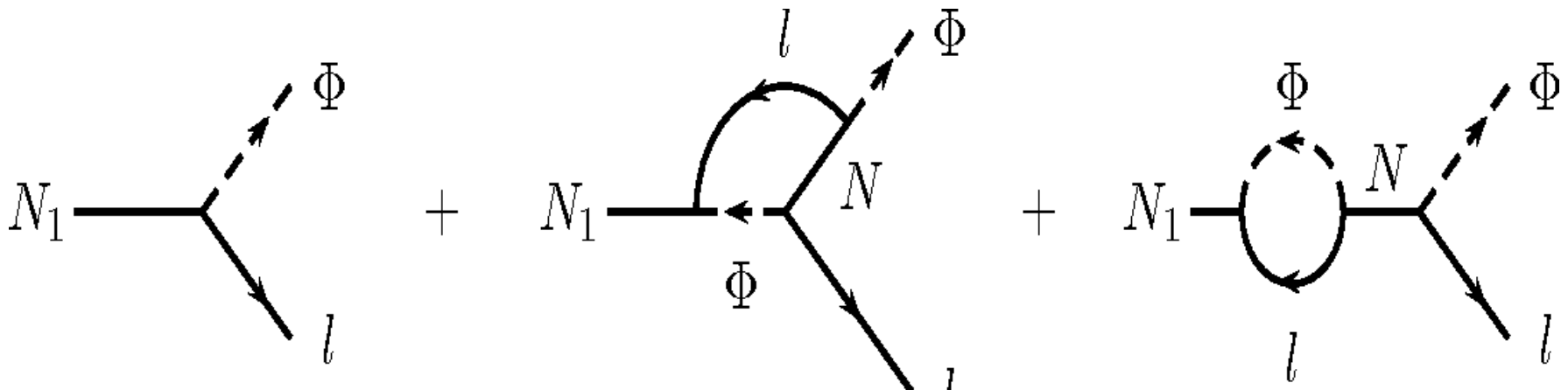
- ▶ Mass of “our” neutrino suppressed by the GUT scale
- ▶ Currently our only “natural” way to explain why the neutrino mass is so much smaller than other Dirac particles

# Leptogenesis

Seesaw mechanism requires a GUT scale heavy Majorana neutrino partner.

In GUT theories, B-L (baryon # - lepton #) is a global U(1) symmetry and is absolutely conserved

Suppose there is direct CP violation in the heavy neutrino decay? This generates a violation of L.



$$\Gamma(N_i \rightarrow l_i + \overline{H^0}) \neq \Gamma(N_i \rightarrow \bar{l}_i + H^0)$$

# Leptogenesis

If  $L$  is violated then, to keep  $B-L$  conserved, one needs to violate  $B$  as well.

Generation of baryon asymmetry from lepton asymmetry  
(via non perturbative *sphaleron* transitions 🙄 )

Could neutrino mass help explain CP violation in the baryons? In other words, could neutrinos help explain why there is more matter than antimatter in the universe?

This idea requires

- ⦿ the neutrino to be massive
- ⦿ the neutrino must be Majorana
- ⦿ a GUT scale heavy neutral lepton must exist

# Leptogenesis

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*(Attempts at) mass measurements*

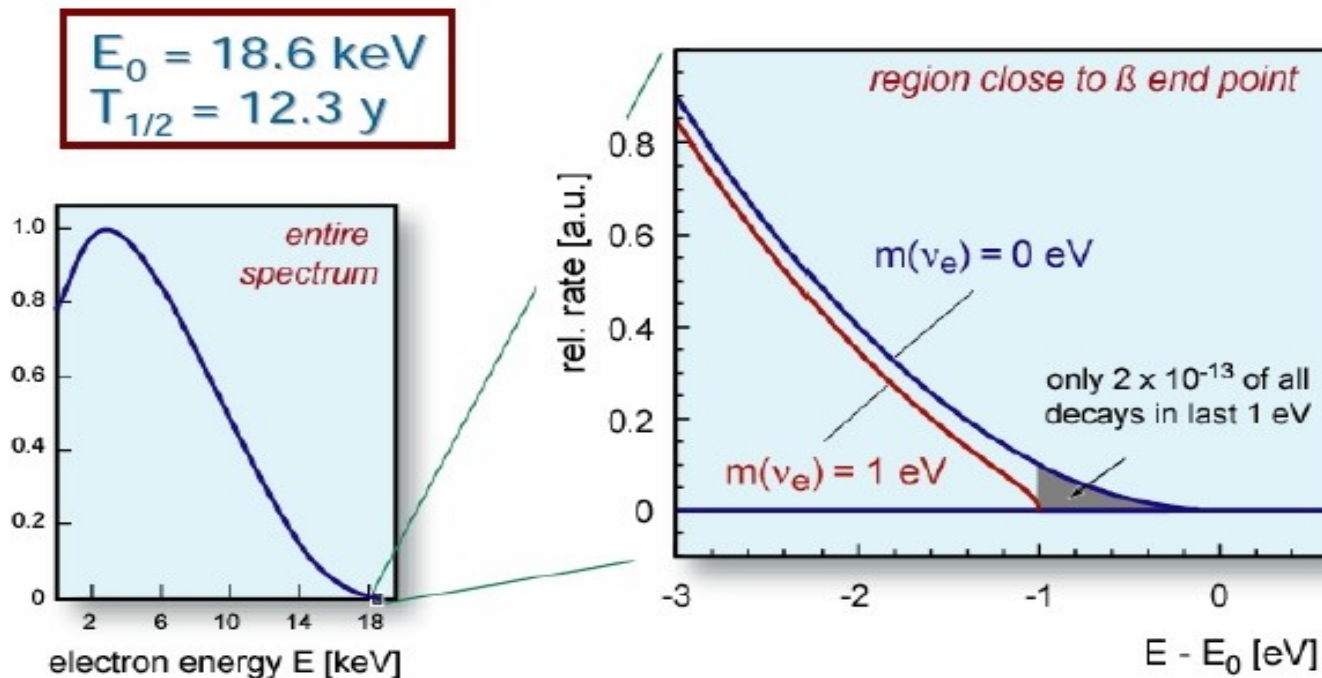


# $\nu_e$ mass

Measurement of  $\nu_e$  mass from kinematics of  $\beta$  decay.

$$\frac{d\Gamma_i}{dE} = C p(E + m_e)(E_0 - E) \sqrt{(E_0 - E)^2 - m_\nu^2} F(E) \theta(E_0 - E - m_\nu)$$

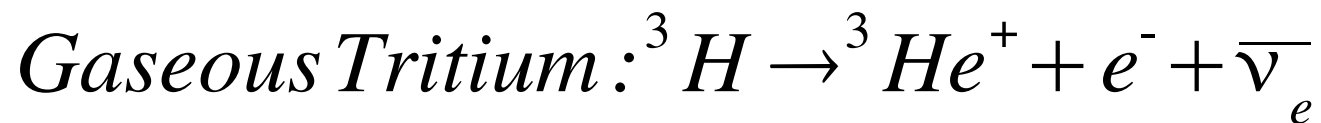
Observable is  $m_\nu^2$





# Requirements

- ▶ # electrons close to the endpoint should be large
- ▶ Good (and well-understood) electron energy resolution
- ▶ No (or minimal) electron energy loss within the source
- ▶ Minimal atomic and nuclear final state effects, of excited transitions



Endpoint is at 18574 eV

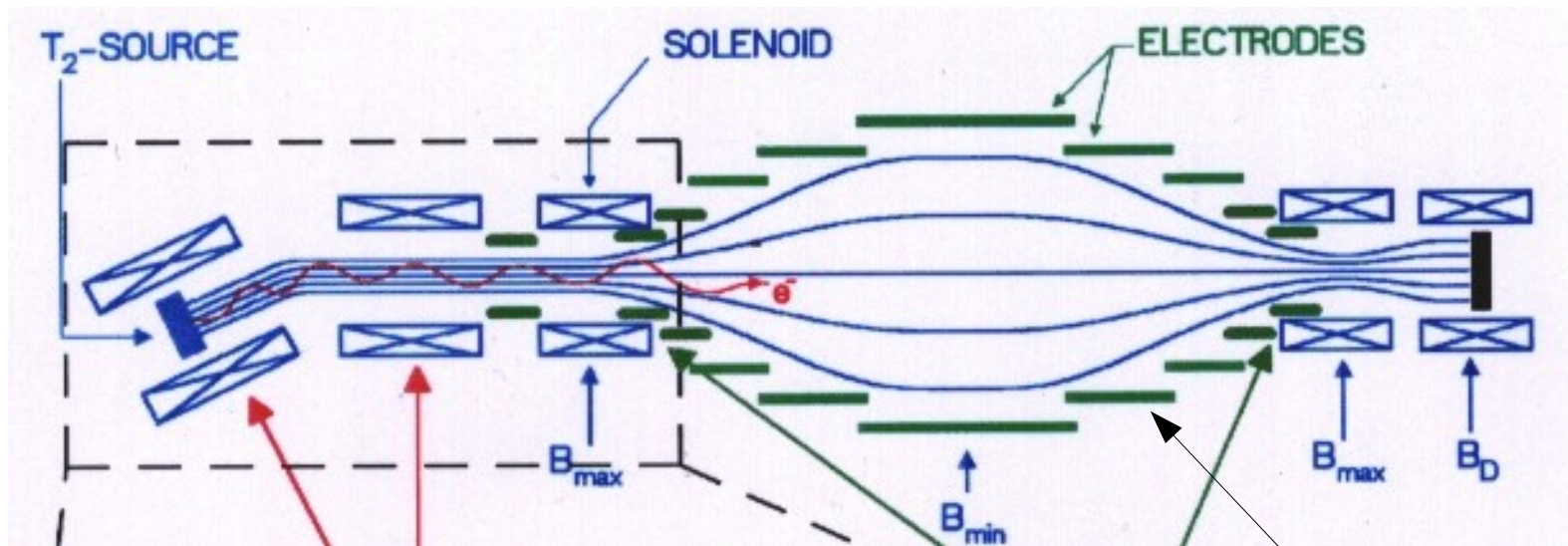
No molecular excitation above 18547 eV

Still only  $10^{-9}$  electrons in this region

Gaseous so you can have a very large source

# Mainz Experiment

The current standard for tritium beta decay experiments



- $2\pi$  acceptance
- High energy resolution

$$\frac{\Delta E}{E} \sim 0.03\%$$

Electrostatic  
MAC-E Filter

# Present Status



Troitsk

windowless gaseous  $T_2$  source

analysis 1994 to 1999, 2001

$$m_{\nu}^2 = -2.3 \pm 2.5 \pm 2.0 \text{ eV}^2$$

$$m_{\nu} \leq 2.2 \text{ eV (95\% CL.)}$$

Mainz

quench condensed solid  $T_2$  source

analysis 1998/99, 2001/02

$$m_{\nu}^2 = -1.2 \pm 2.2 \pm 2.1 \text{ eV}^2$$

$$m_{\nu} \leq 2.2 \text{ eV (95\% CL.)}$$

Both experiments have reached the intrinsic limit of their sensitivity.

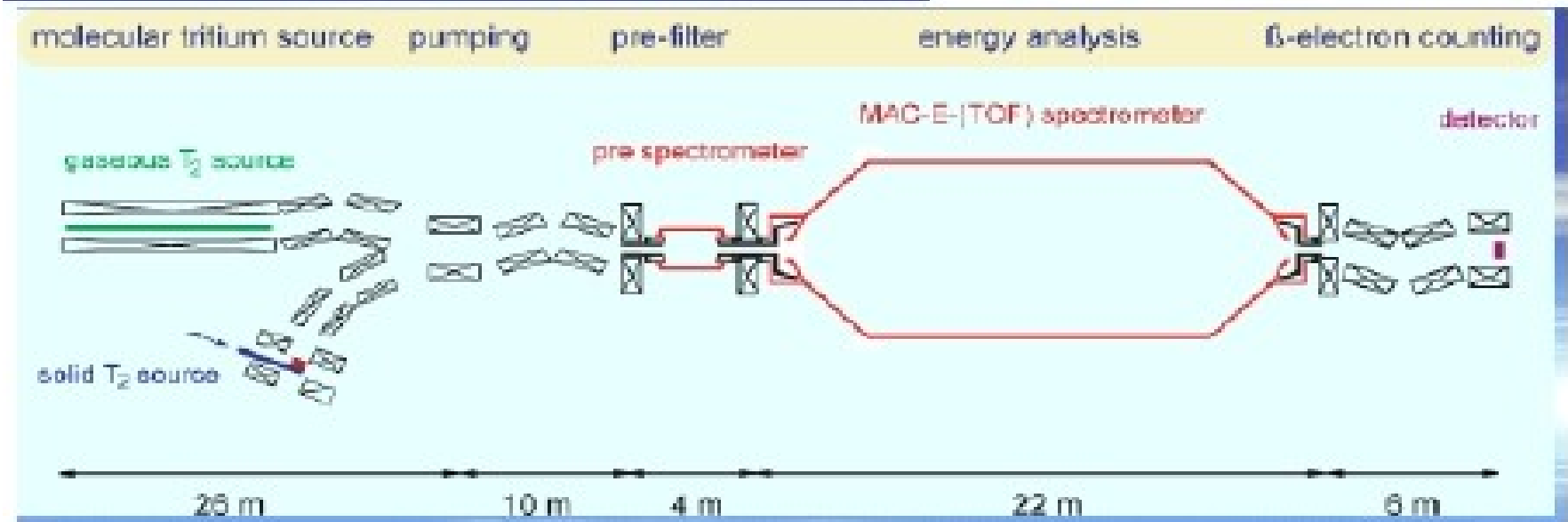
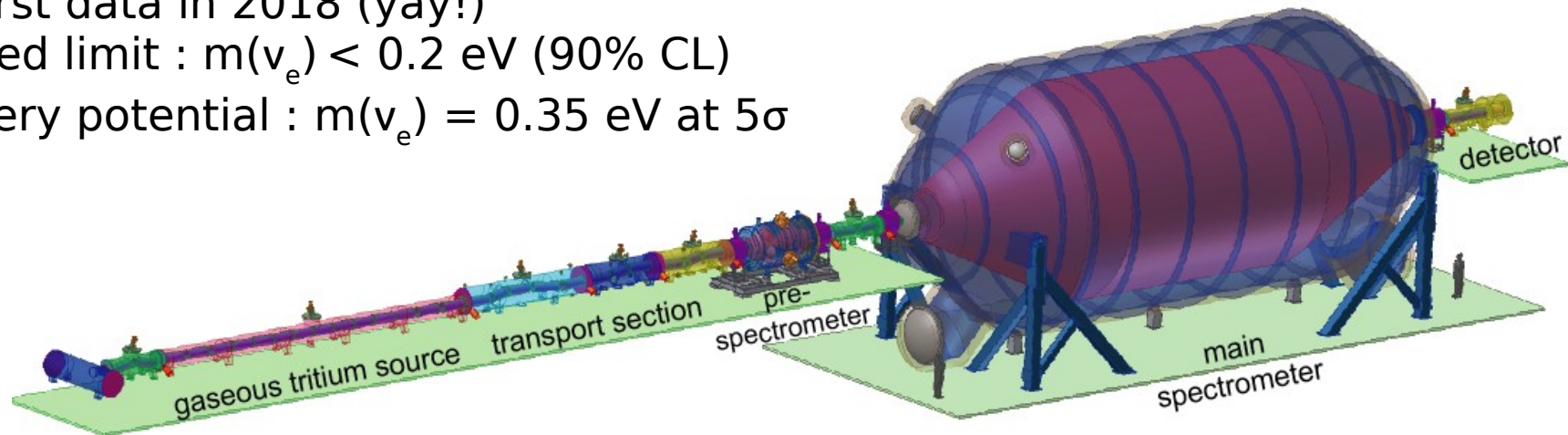


# KATRIN

Took first data in 2018 (yay!)

Expected limit :  $m(\nu_e) < 0.2 \text{ eV}$  (90% CL)

Discovery potential :  $m(\nu_e) = 0.35 \text{ eV}$  at  $5\sigma$











# Katrin on the move

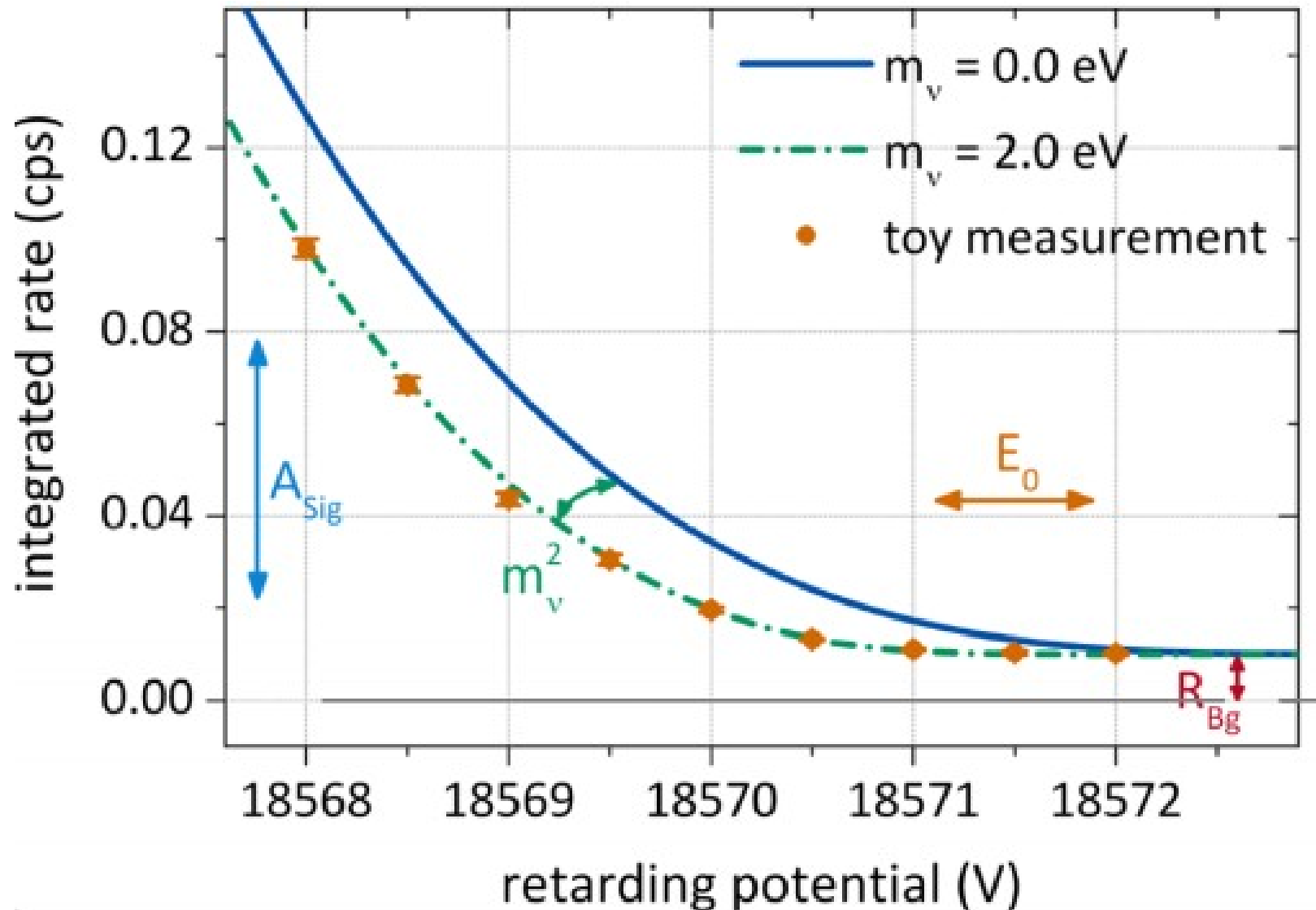




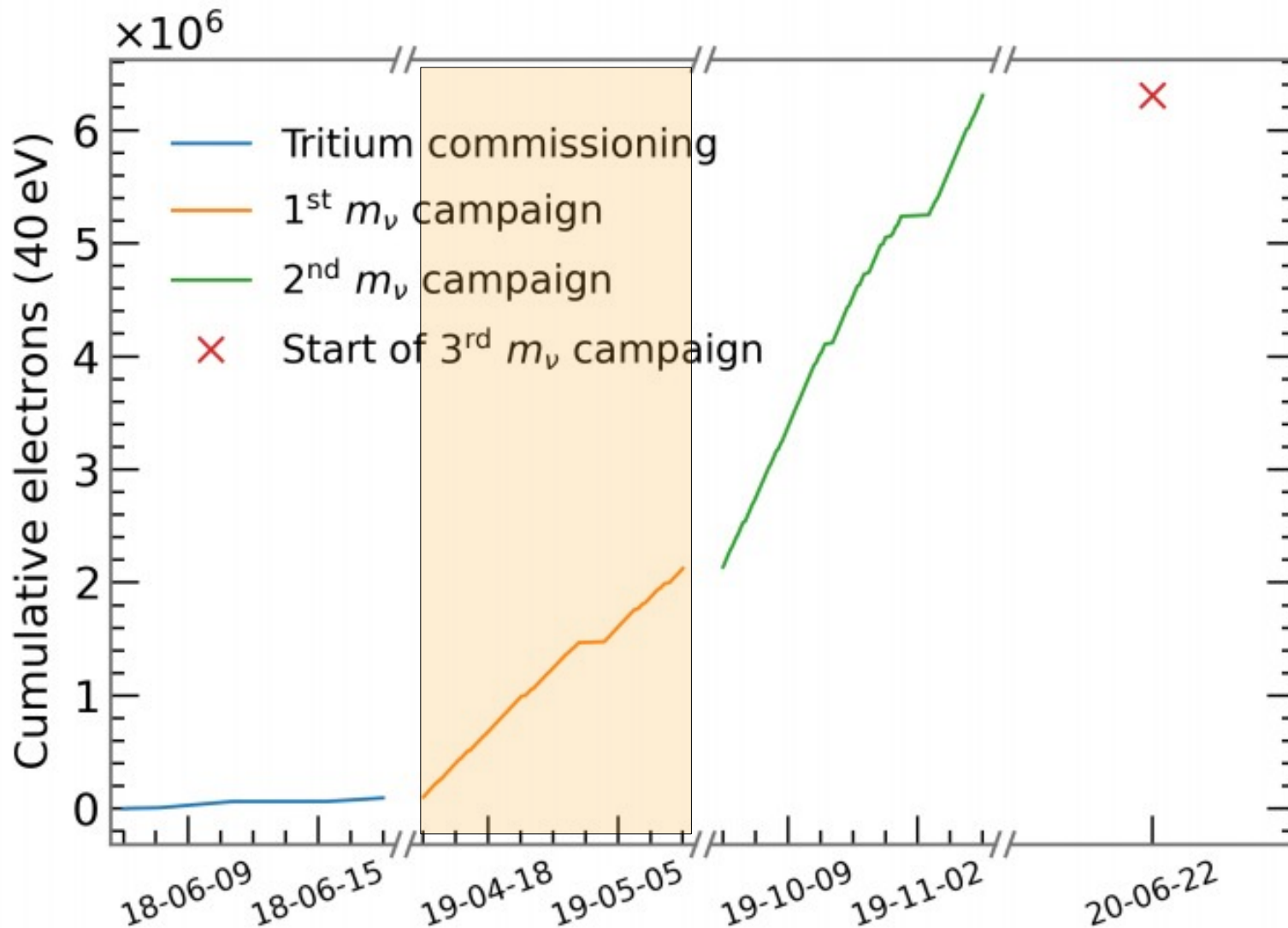
# KATRIN on the move



# Katrin data

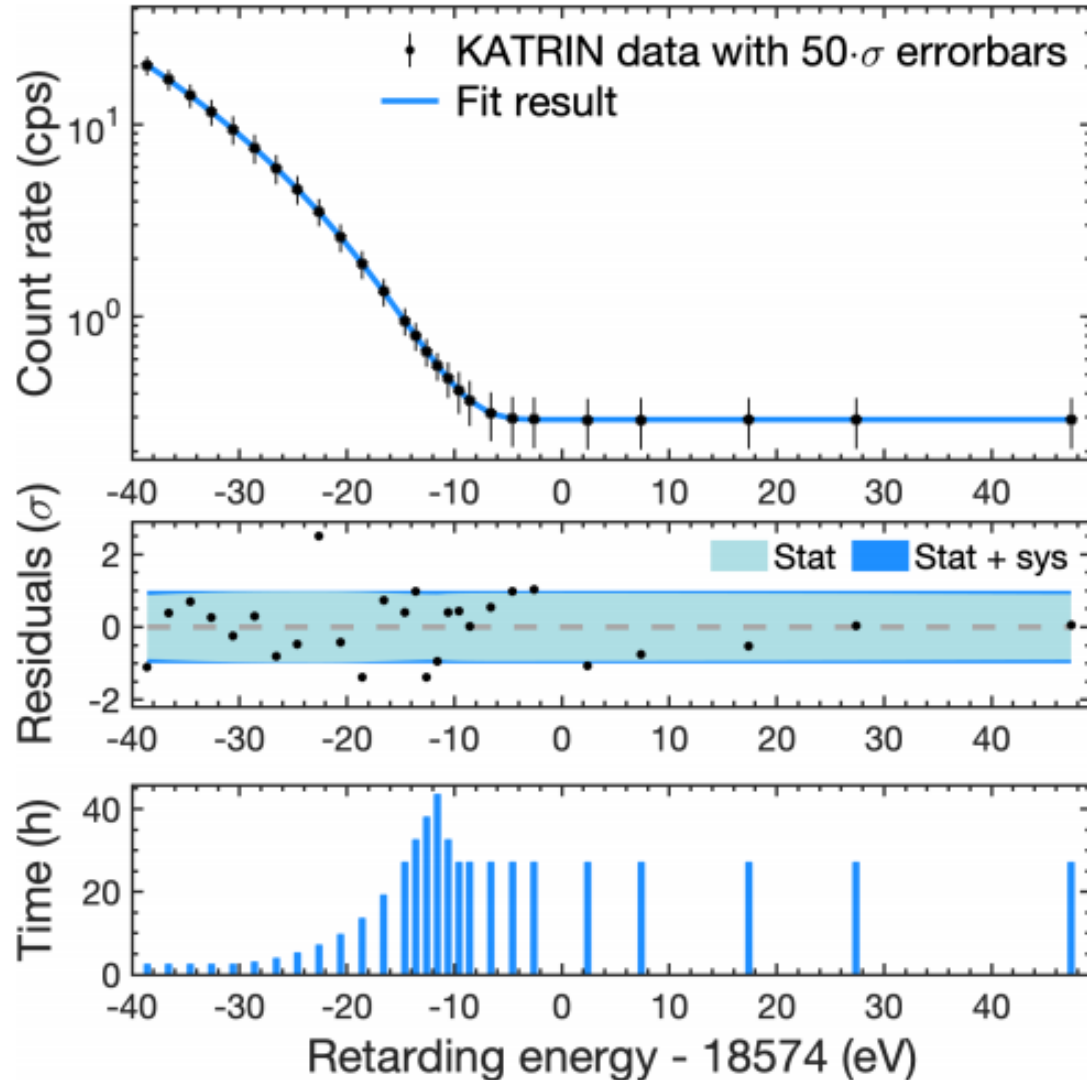


# KATRIN Data-taking





# First KATRIN result



▶ 2 million electrons

▶ Fit for electron neutrino mass:

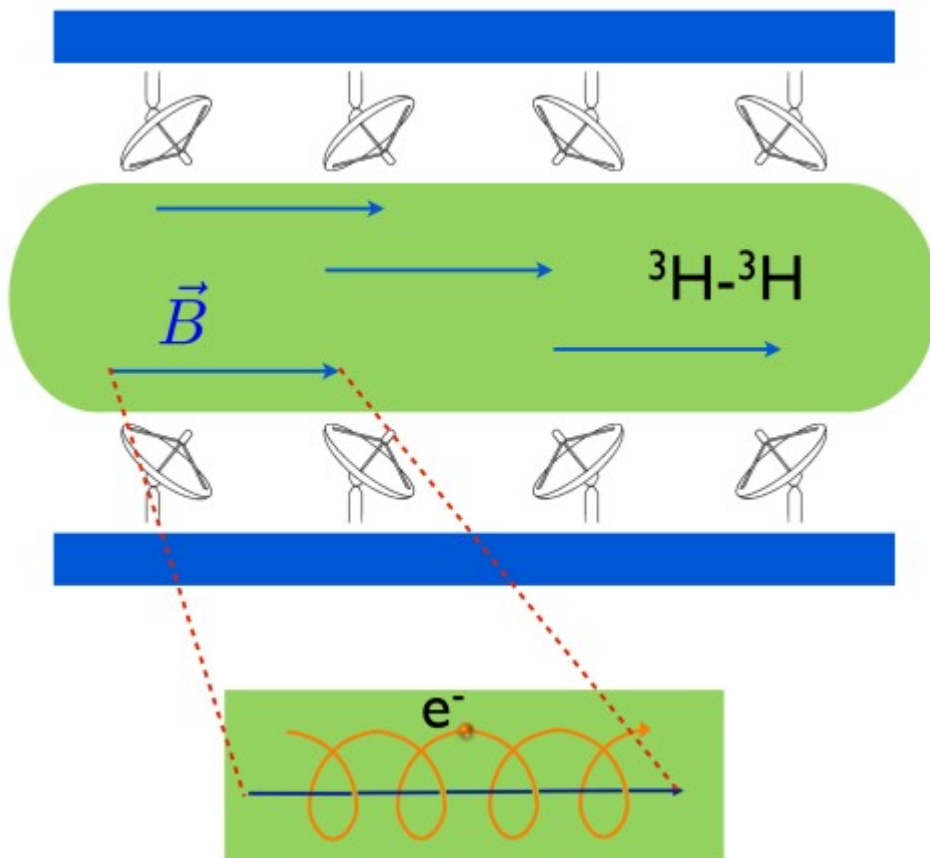
$$m_\nu^2 = (-1.0 \pm 1.0) \text{ eV}^2$$

▶ Upper limit

$$m_\nu < 0.9 \text{ eV @ } 90\% \text{ CL}$$

# Project 8

## Project 8



- ▶ Tritium beta decay in a magnetic field.
- ▶ Electron from beta decay spirals around the field lines
- ▶ Emits cyclotron radiation at a particular frequency

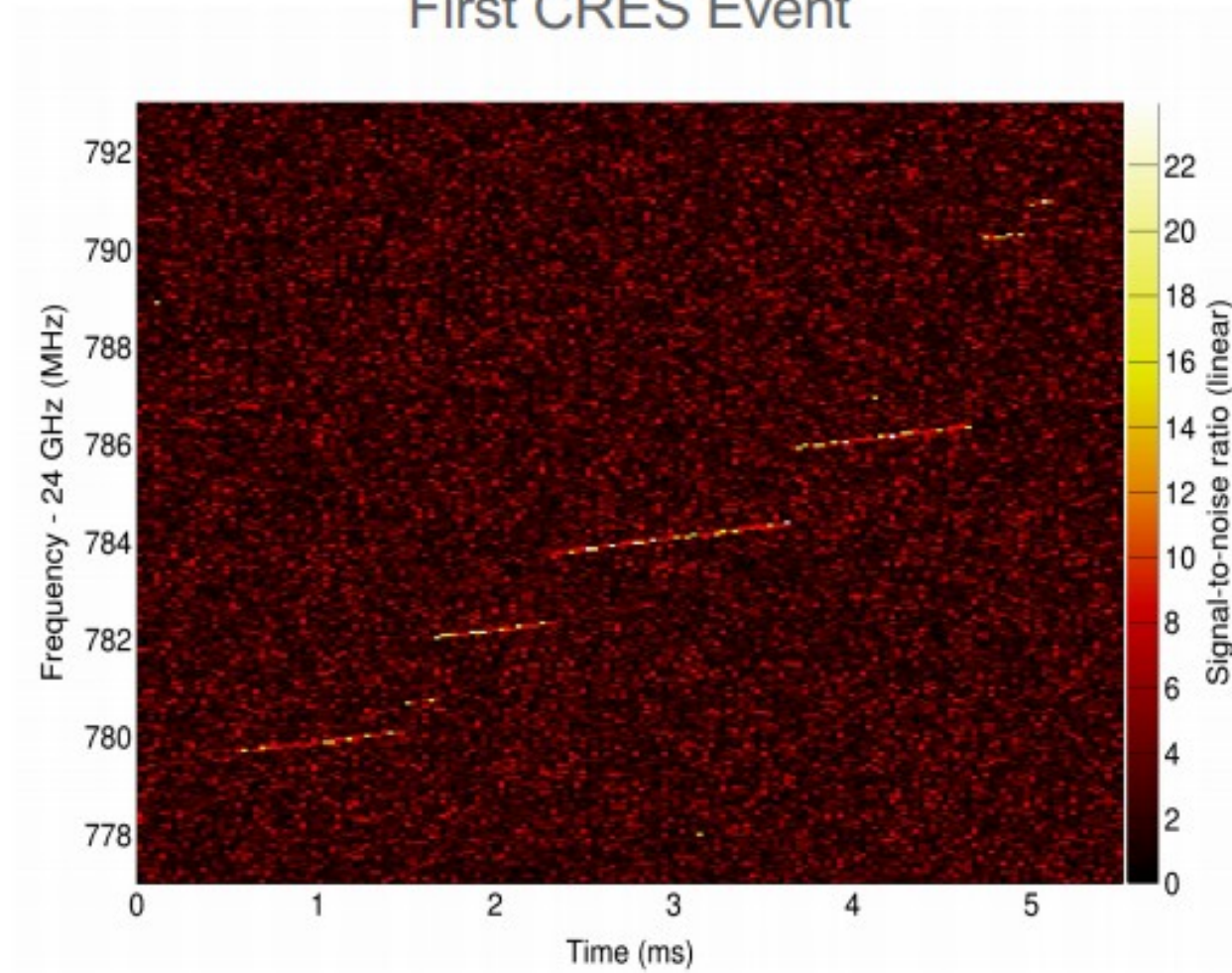
$$\omega = \frac{\omega_c}{E + m_e}$$

- ▶ Measures electron energy from the frequency of the cyclotron radiation!

# Project 8

## Project 8 Demonstrator

### First CRES Event

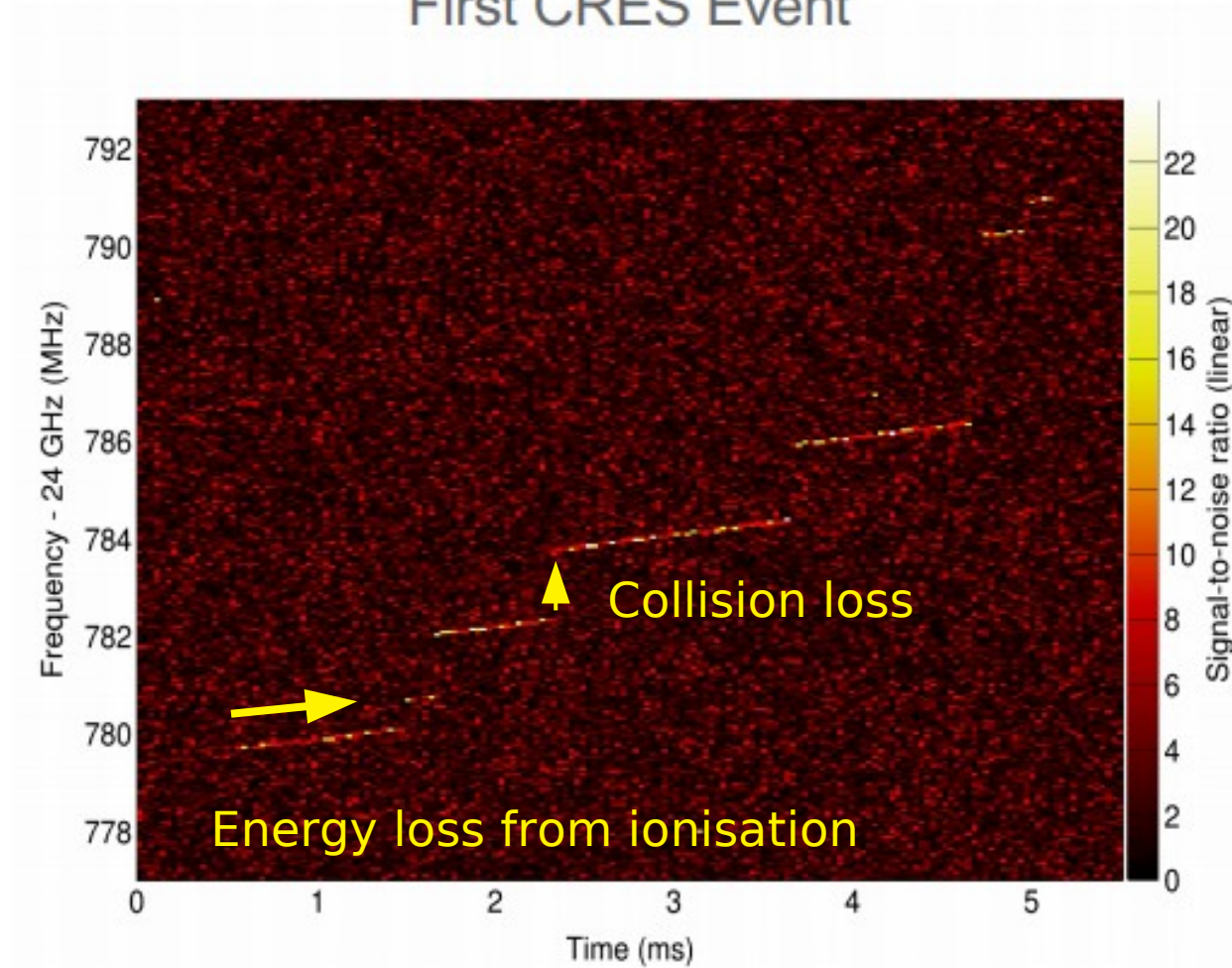




# Project 8

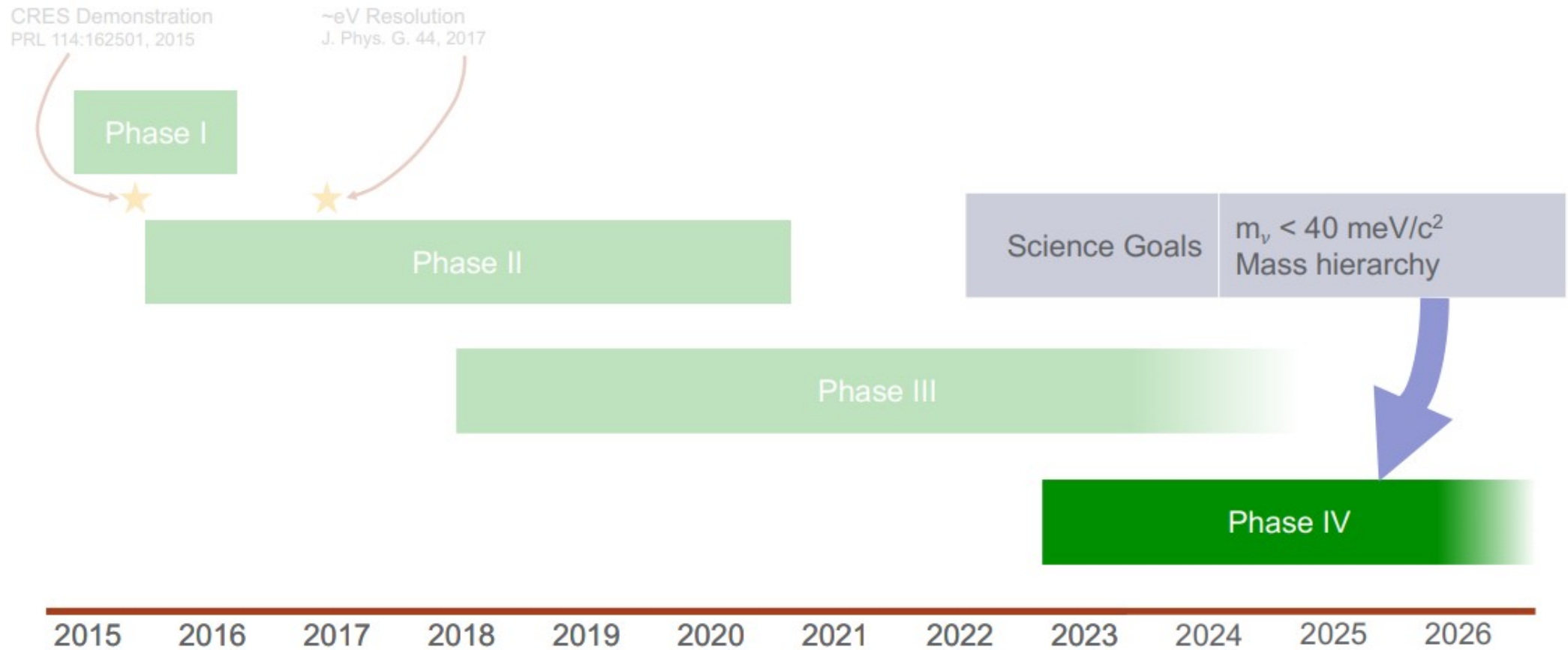
## Project 8 Demonstrator

### First CRES Event



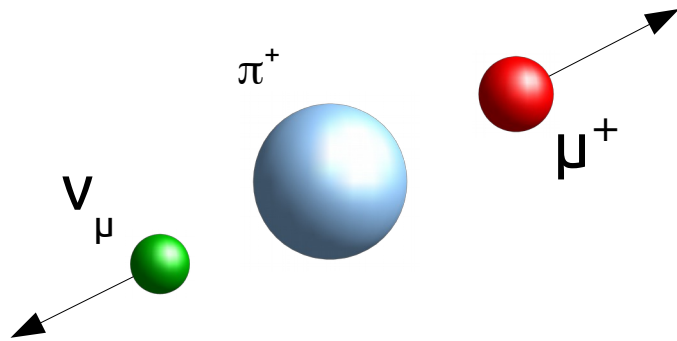


# Project 8



# $\nu_\mu$ mass

Easiest way is to use pion decay at rest



$$m_{\nu_\mu}^2 = m_\pi^2 + m_\mu^2 - 2 m_\pi \sqrt{p_\mu^2 + m_\mu^2}$$

$$m_\pi = 139.57037 \pm 0.00021 \text{ MeV}$$

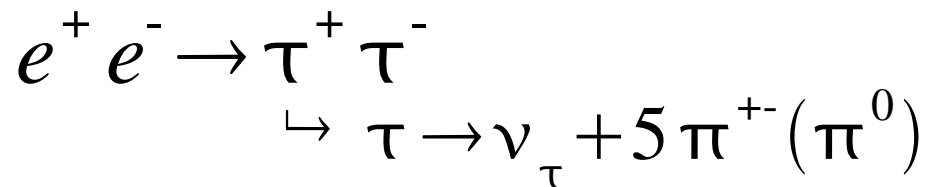
$$m_\mu = 105.658389 \pm 0.000034 \text{ MeV}$$

$$p_\mu = 29.792 \pm 0.00011 \text{ MeV}$$

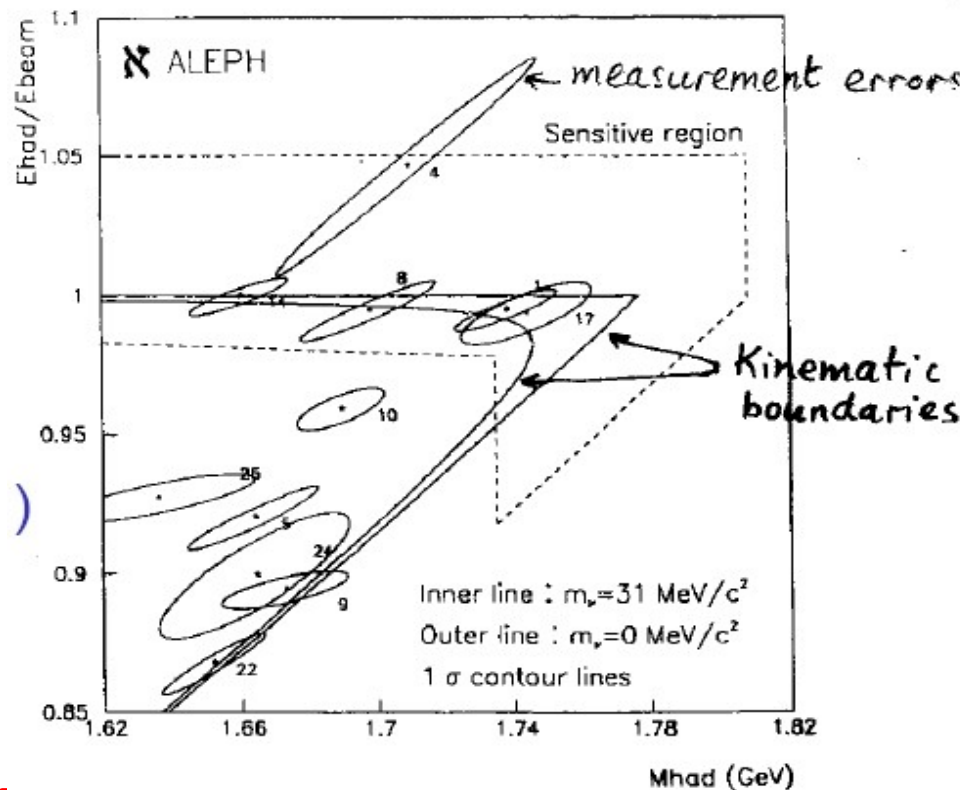
$$m_{\nu}^2 = (-0.016 \pm 0.023) \text{ MeV}^2$$

$$m_{\nu} < 190 \text{ keV} \text{ (90\% CL)}$$

# $\nu_\tau$ mass



$$E_\tau = \frac{\sqrt{s}}{2}$$



$$m_\tau < 19.2 \text{ MeV} (95\% \text{ CL})$$



# Cosmology

- Density fluctuations are affected by neutrino mass in the early universe
- Highly model dependent
- WMAP, 2dF, ACBAR, CBI, PLANCK, BOSS, BAO, SDSS

$$\sum m_{\nu_i} < (0.14 - 0.33) \text{ eV}$$

(rather model dependent)



$m_{\nu} = 0 \text{ eV}$

$m_{\nu} = 1 \text{ eV}$

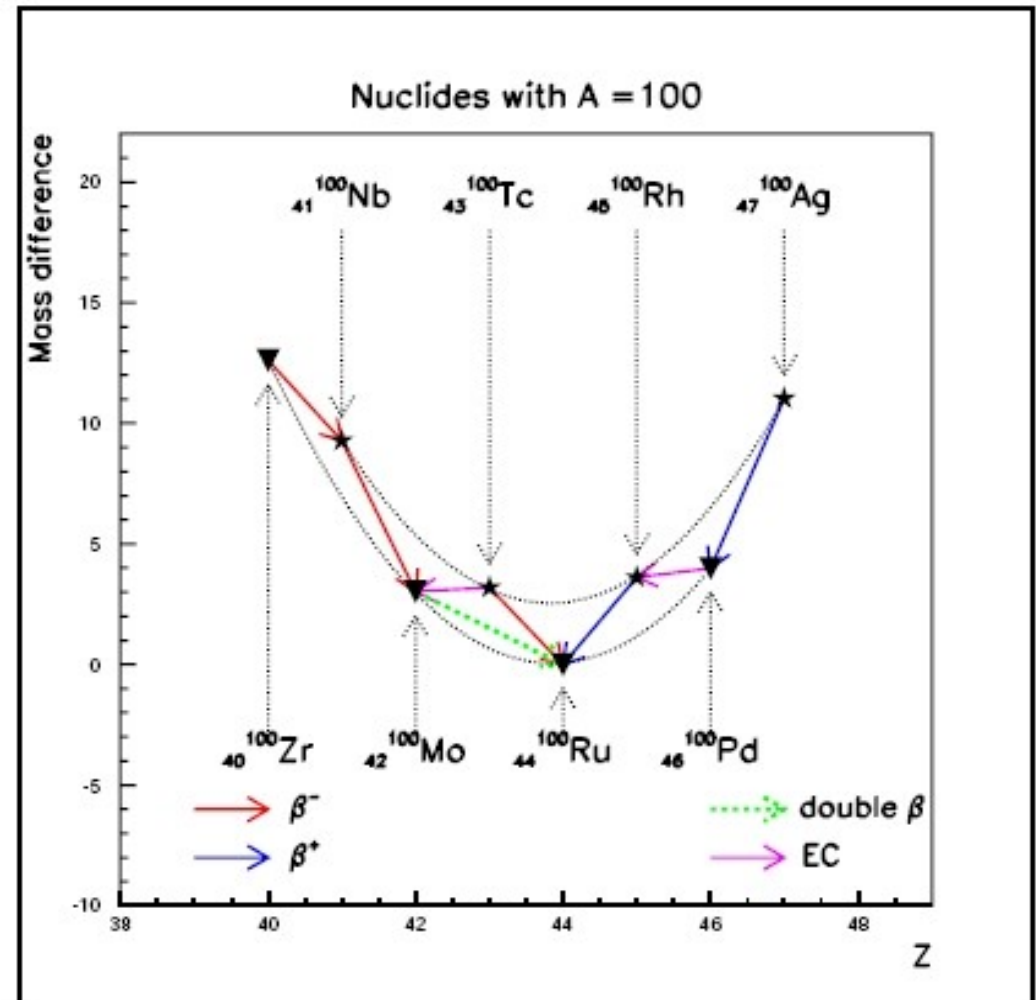
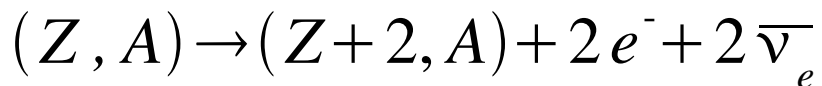
$m_{\nu} = 7 \text{ eV}$

$m_{\nu} = 4 \text{ eV}$

# $2\nu\beta\beta$ Decay

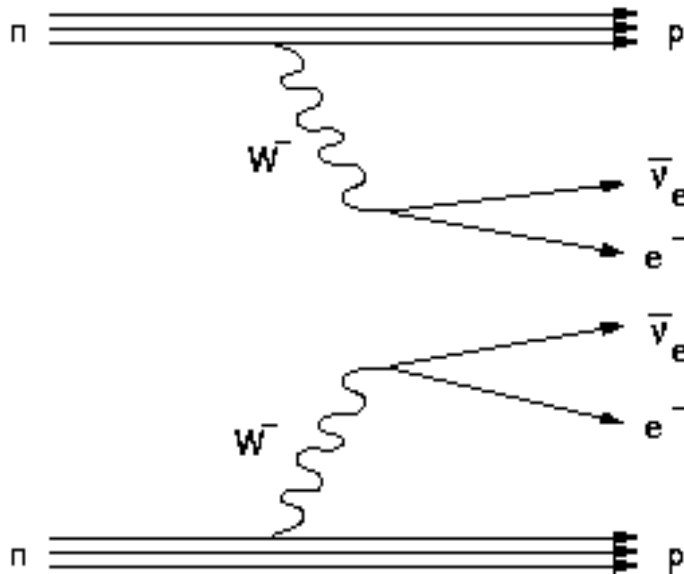
*Neutrinoless double beta decay* is considered a **golden** channel for the measurement of neutrino mass.

In some nuclei  $\beta$  decay is forbidden but double beta decay is not





# $2\nu\beta\beta$ Decay



$$\left[ T_{1/2}^{2\nu} \right]^{-1} = G^{2\nu}(Q, Z) |M^{2\nu}|^2$$

Calculable  
phase space

Nuclear  
matrix element

- Second order process in perturbation theory
- Severe test for nuclear matrix element calculation
- Nuclear structure effects cause variations in the nuclear matrix elements of factors of 10

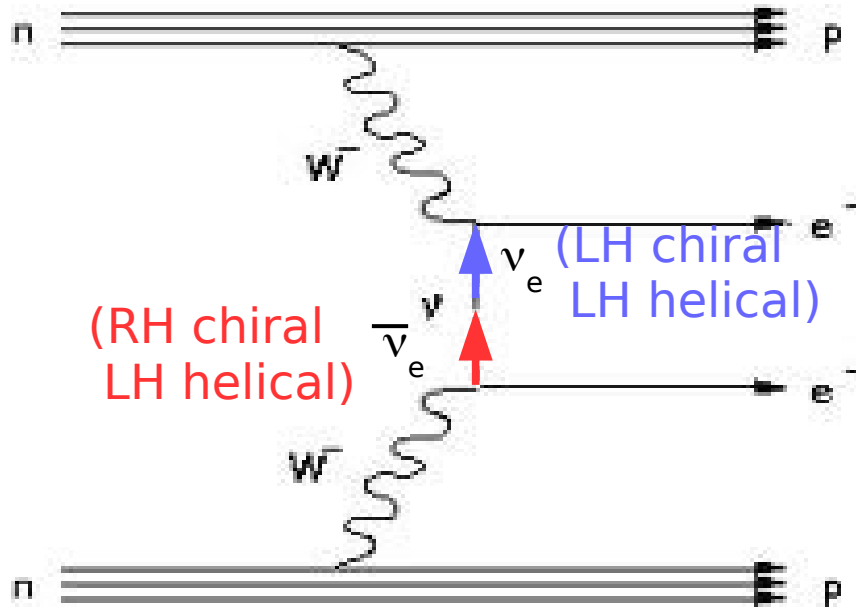
# $2\nu\beta\beta$ Decay

| $2\nu\beta\beta$ mode                                       | Half life ( $\times 10^{24}$ years) |
|---|-------------------------------------|
| ${}^{48}_{20}\text{Ca} \rightarrow {}^{48}_{22}\text{Ti}$   | 4.1                                 |
| ${}^{76}_{32}\text{Ge} \rightarrow {}^{76}_{34}\text{Se}$   | 40.9                                |
| ${}^{82}_{34}\text{Se} \rightarrow {}^{82}_{36}\text{Kr}$   | 9.3                                 |
| ${}^{96}_{40}\text{Zr} \rightarrow {}^{96}_{42}\text{Mo}$   | 4.4                                 |
| ${}^{100}_{42}\text{Mo} \rightarrow {}^{100}_{44}\text{Ru}$ | 5.7                                 |
| ${}^{110}_{46}\text{Pd} \rightarrow {}^{110}_{48}\text{Cd}$ | 18.6                                |
| ${}^{116}_{48}\text{Cd} \rightarrow {}^{116}_{50}\text{Sn}$ | 5.3                                 |
| ${}^{124}_{50}\text{Sn} \rightarrow {}^{124}_{52}\text{Te}$ | 9.5                                 |
| ${}^{130}_{52}\text{Te} \rightarrow {}^{130}_{54}\text{Xe}$ | 5.9                                 |
| ${}^{136}_{54}\text{Xe} \rightarrow {}^{136}_{56}\text{Ba}$ | 5.5                                 |
| ${}^{150}_{60}\text{Nd} \rightarrow {}^{150}_{62}\text{Sm}$ | 1.2                                 |

- ▶ Only occur in 36 known sources
- ▶ Rarest natural radioactive decay
- ▶ extremely long half-lives



# Neutrinoless $\beta\beta$ Decay



## Requirements

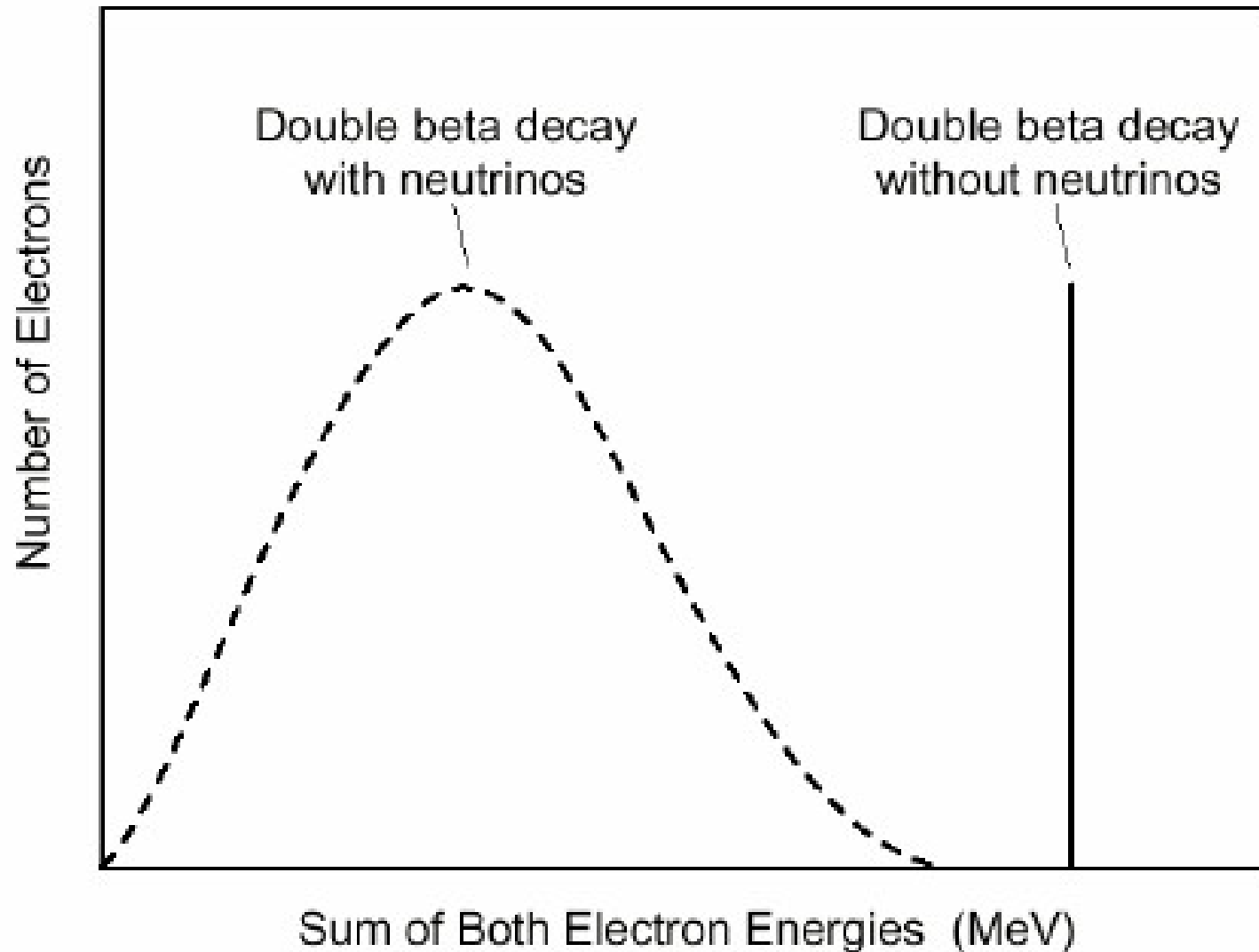
- Neutrino must have mass
- Neutrino is Majorana
- Violation of lepton number conservation

$$|\nu_L\rangle = |\nu_{h=-1}\rangle + \frac{m}{E} |\nu_{h=+1}\rangle$$

$\uparrow$  helicity states  $\uparrow$

$$\Gamma_{0\nu} = G_{0\nu} |M_{0\nu}|^2 \left| \sum_i |U_{ei}|^2 m_i \right|^2 \Rightarrow T_{1/2} \sim 10^{27} \text{ years}$$

# $0\nu\beta\beta$ signal

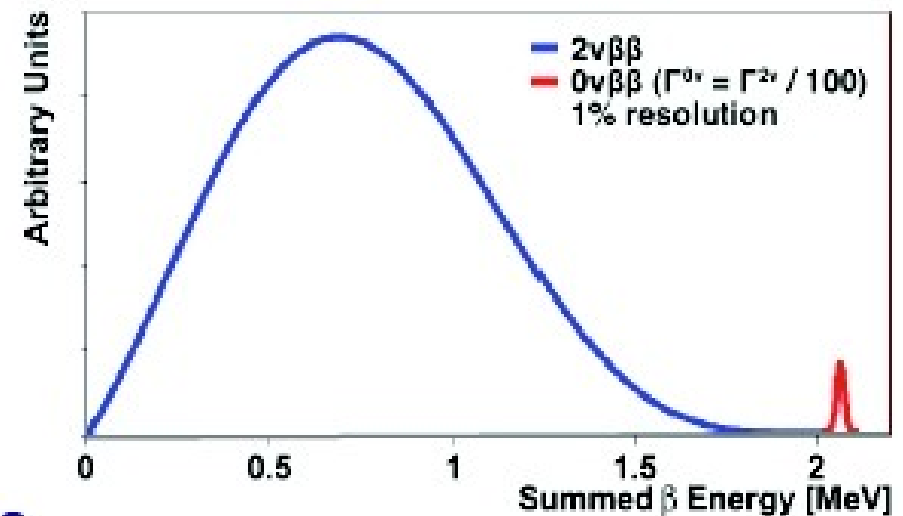


# Experimental Requirements

*Extremely* slow decay rates  
( $0\nu\beta\beta T_{1/2} \sim 10^{26} - 10^{27}$  years)

Best case,  
0 background !

$$\propto \text{Source Mass} \cdot \text{time}_{\text{exp}}$$



## Requires

Large, highly efficient source mass

- detector as source

Best possible energy resolution

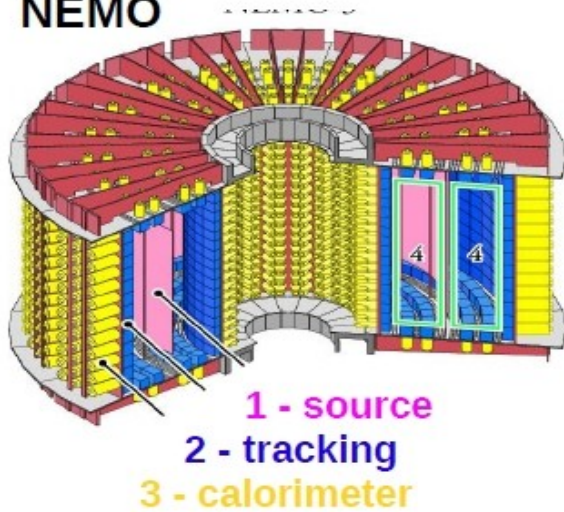
- minimize  $0\nu\beta\beta$  peak ROI to maximize S/B
- separate from  $0\nu\beta\beta$  from irreducible  $2\nu\beta\beta$  ( $\sim T_{1/2} \sim 10^{19} - 10^{21}$  years)

Extremely low (near-zero) backgrounds in the  $0\nu\beta\beta$  peak region

- requires ultra-clean radiopure materials
- the ability to discriminate signal from background

# Types of experiments

## NEMO



### 1. the source is inserted as thin foil inside a tracking detector

- $2e^-$  are detected separately
  - different channels of  $0\nu\text{DBD}$  can be distinguished
- **particle identification**
  - background suppression
- **poor energy resolution**
  - important  $2\nu\text{DBD}$  background (limitation on isotope choice)



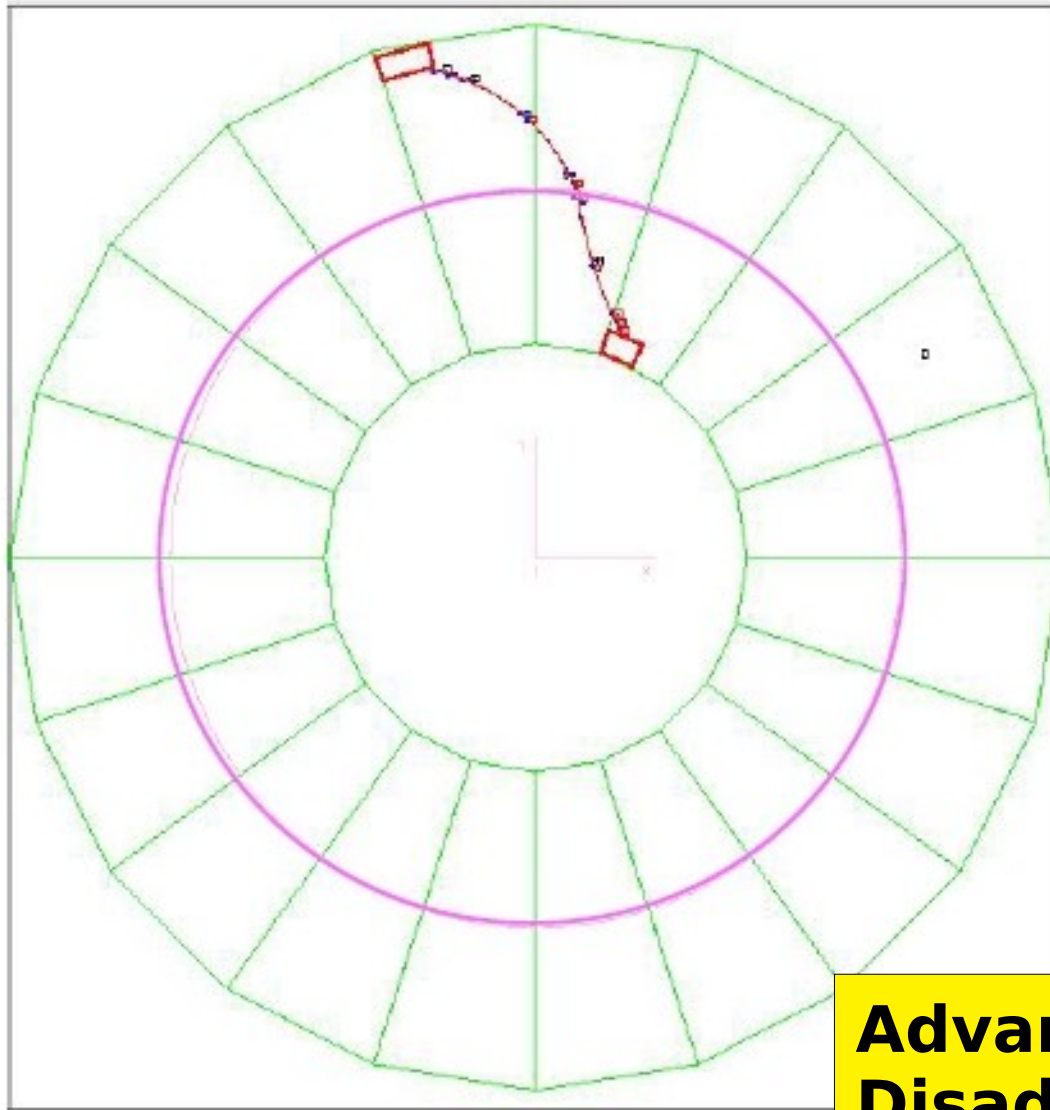
### 2. the detector is itself the source

- **solid state detectors**
  - several candidates, high resolution  
no info on kinematic  
techniques for background suppression
- **gaseous detectors for Xe**

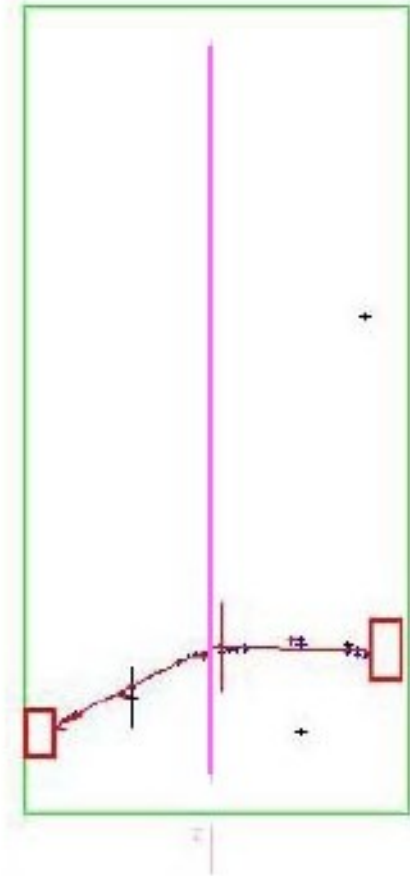




Typical  $\beta\beta 2\nu$  event observed from  $^{100}\text{Mo}$



Top view

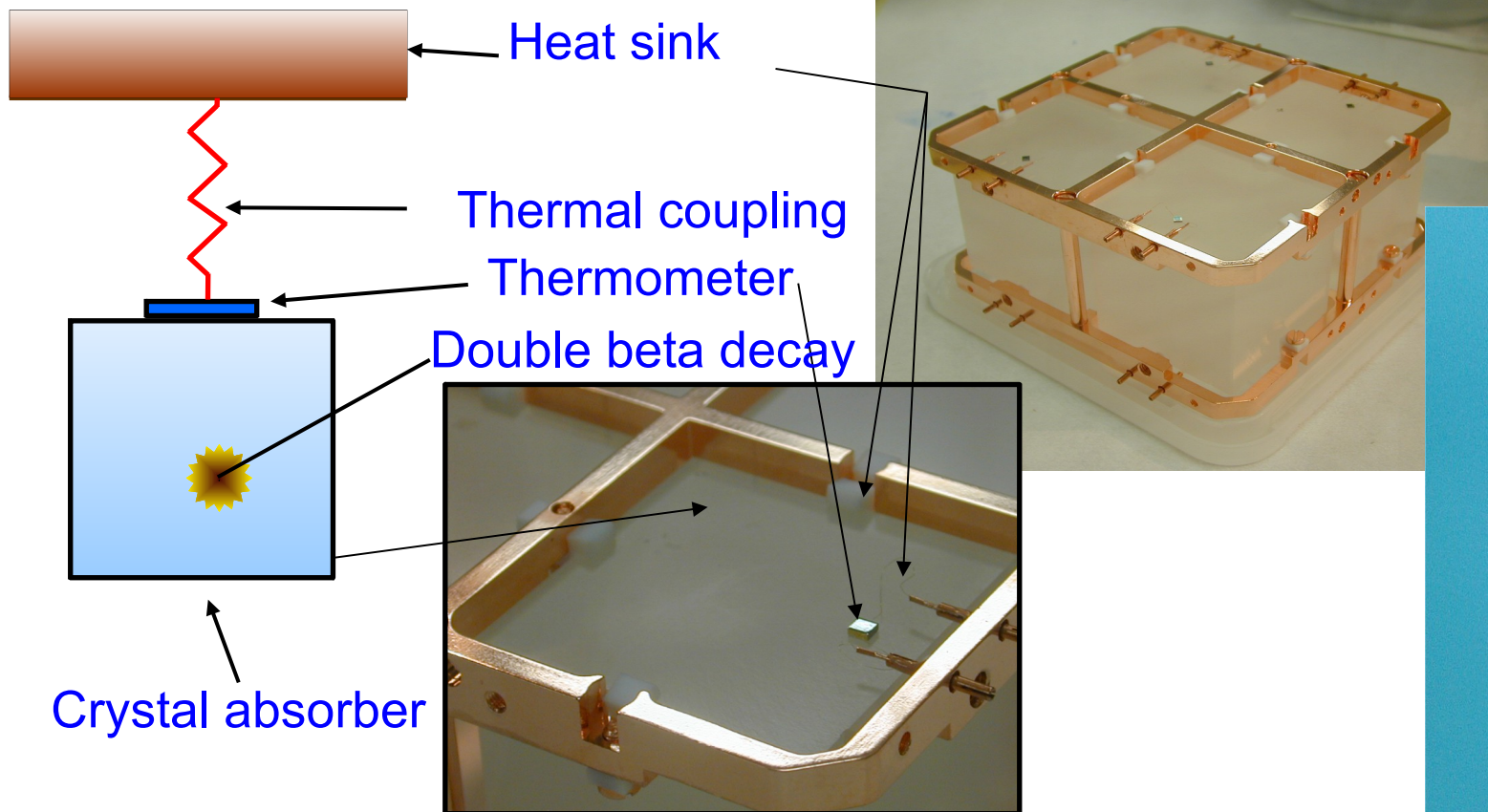


Side view

**Advantage** : electron tracking  
**Disadvantage** : limited source material and relatively poor energy resolution



# Bolometry : Cuore



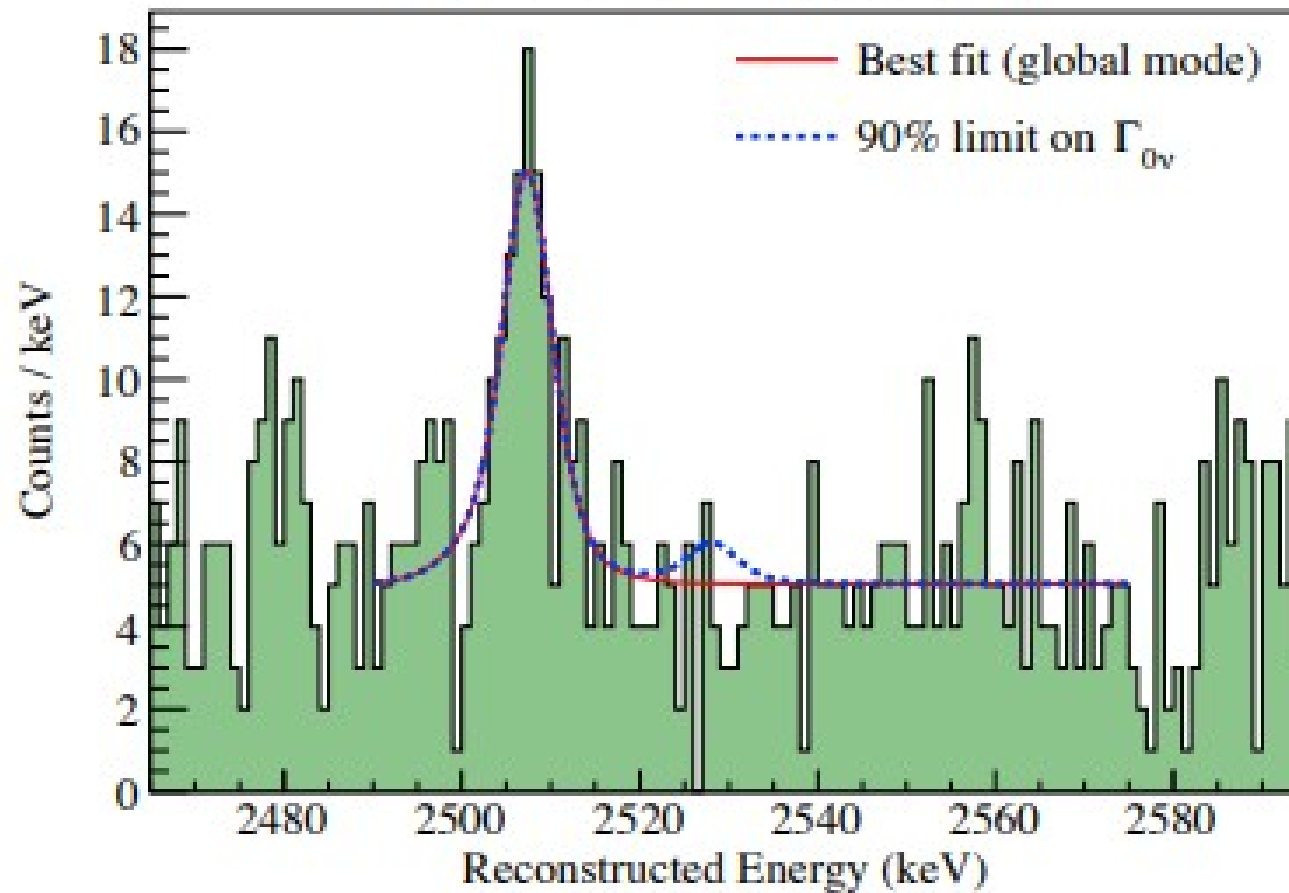
**example: 750 g of  $\text{TeO}_2$  @ 10 mK**

$$C \sim T^3 \text{ (Debye)} \Rightarrow C \sim 2 \times 10^{-9} \text{ J/K}$$

$$1 \text{ MeV } \gamma\text{-ray} \Rightarrow \Delta T \sim 80 \mu\text{K}$$

$$\Rightarrow \Delta U \sim 10 \text{ eV}$$

# Cuore Results

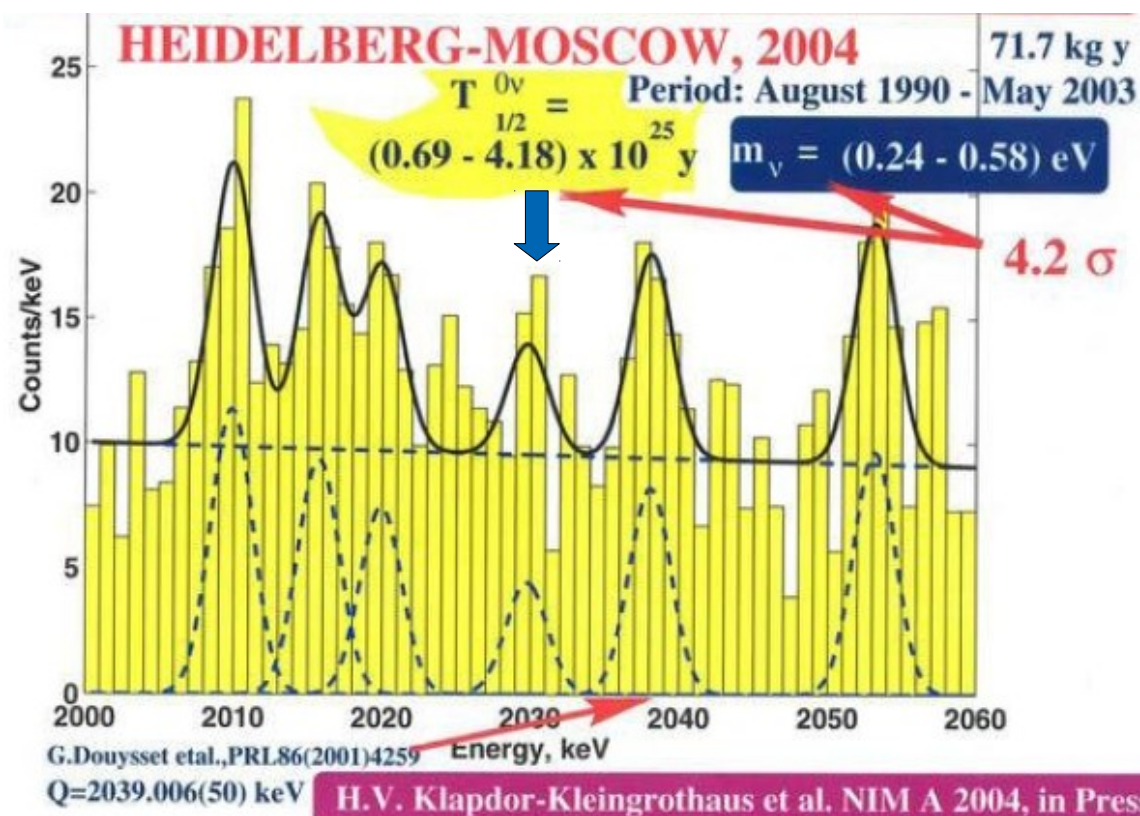
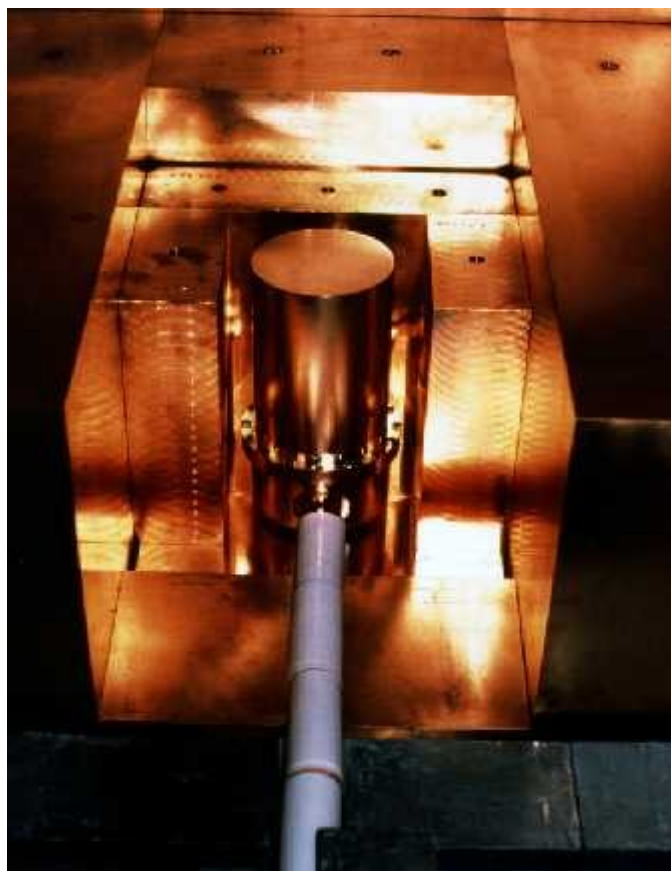


$$T_{1/2}^{0\nu} > 3.2 \times 10^{25} \text{ years} \Rightarrow \langle m_{\nu} \rangle < 0.76 - 3.5 \text{ eV}$$

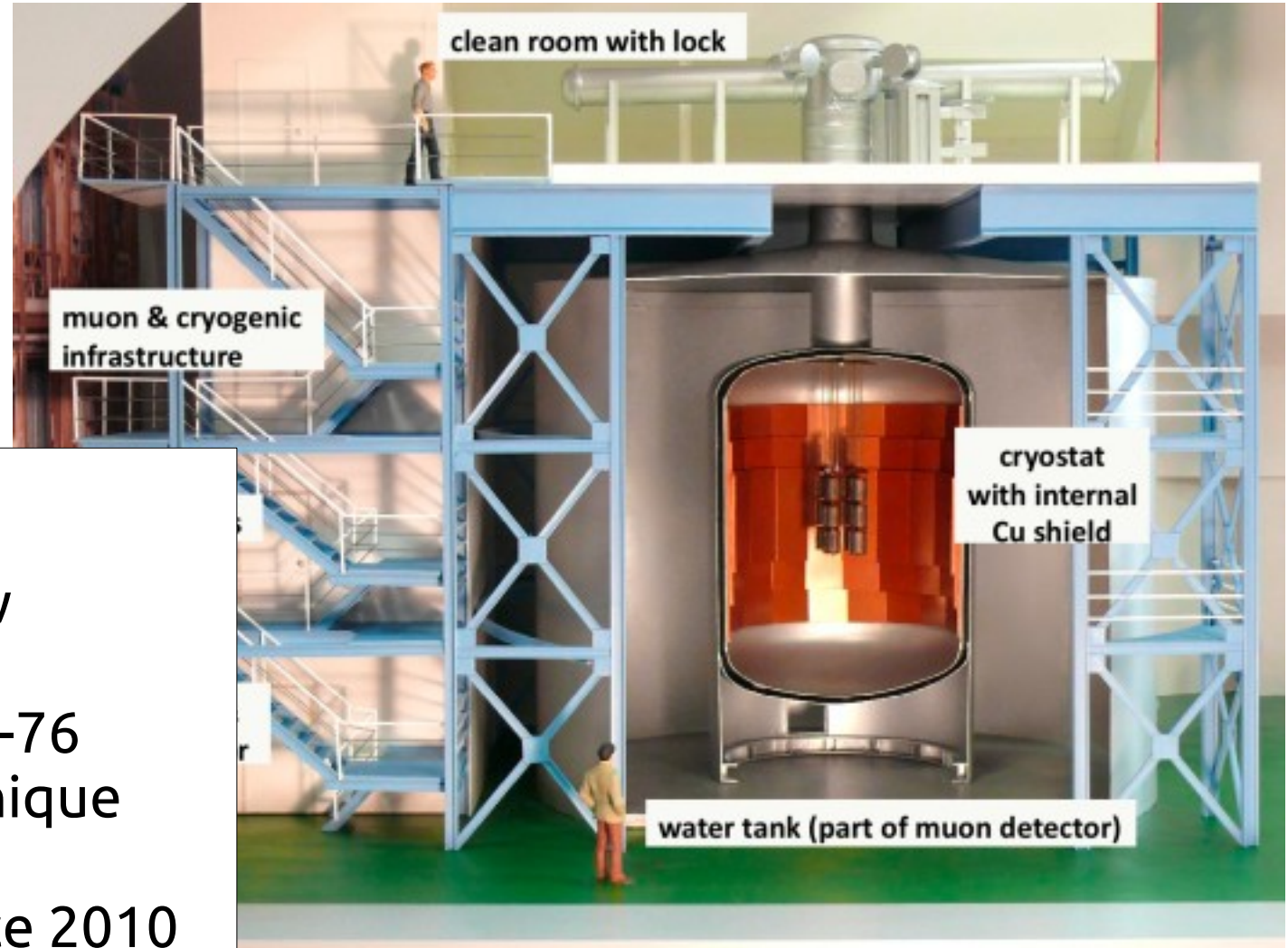


# Heidelberg-Moscow (HdM)

11 kg of Ge enriched to 86% of  $^{76}\text{Ge}$  in the form of 5 Ge diodes surrounded by Cu,Pb,Bn shielding  
 $0\nu\beta\beta$  electrons detected by Ge detectors themselves  
Sum of electron energy is measured



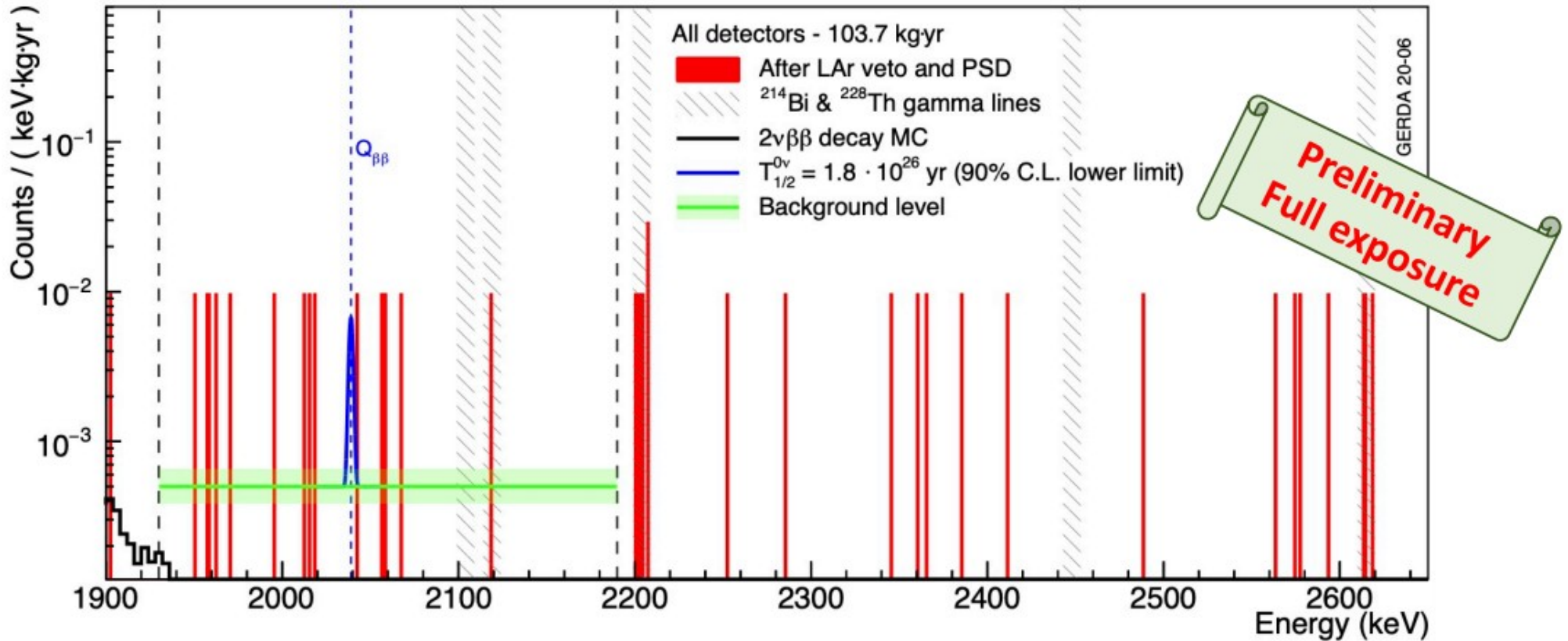
# GERDA



- ▶ Designed to test Heidelberg-Moscow
- ▶ Uses the same Ge-76 isotope and technique
- ▶ Been running since 2010



# GERDA



$$T_{1/2} > 1.4 \times 10^{26} \text{ yr @ 90\% CL}$$

$$m(\nu_e) < 260 \text{ meV @ 90\% CL}$$

Inconsistent with HdM, but  
not definitive (yet)

# Future Program

| Collaboration          | Isotope                              | Technique   | mass (0νββ isotope) | Status       |
|------------------------|--------------------------------------|---|---------------------|--------------|
| CANDLES-III            | <sup>48</sup> Ca                     | 305 kg CaF <sub>2</sub> crystals in liquid scintillator   | 0.3 kg              | Operating    |
| CANDLES-IV             | <sup>48</sup> Ca                     | CaF <sub>2</sub> scintillating bolometers                 | TBD                 | R&D          |
| GERDA                  | <sup>76</sup> Ge                     | Point contact Ge in active LAr                            | 44 kg               | Complete     |
| MAJORANA DEMONSTRATOR  | <sup>76</sup> Ge                     | Point contact Ge in Lead                                  | 30 kg               | Operating    |
| LEGEND 200             | <sup>76</sup> Ge                     | Point contact Ge in active LAr                            | 200 kg              | Construction |
| LEGEND 1000            | <sup>76</sup> Ge                     | Point contact Ge in active LAr                            | 1 tonne             | R&D          |
| SuperNEMO Demonstrator | <sup>82</sup> Se                     | Foils with tracking                                       | 7 kg                | Construction |
| SELENA                 | <sup>82</sup> Se                     | Se CCDs   | <1 kg               | R&D          |
| NνDEx                  | <sup>82</sup> Se                     | SeF <sub>6</sub> high pressure gas TPC                    | 50 kg               | R&D          |
| ZICOS                  | <sup>96</sup> Zr                     | 10% <sup>nat</sup> Zr in liquid scintillator              | 45 kg               | R&D          |
| AMoRE-I                | <sup>100</sup> Mo                    | <sup>40</sup> CaMoO <sub>4</sub> scintillating bolometers | 6 kg                | Construction |
| AMoRE-II               | <sup>100</sup> Mo                    | Li <sub>2</sub> MoO <sub>4</sub> scintillating bolometers | 100 kg              | Construction |
| CUPID                  | <sup>100</sup> Mo                    | Li <sub>2</sub> MoO <sub>4</sub> scintillating bolometers | 250 kg              | R&D          |
| COBRA                  | <sup>116</sup> Cd/ <sup>130</sup> Te | CdZnTe detectors  | 10 kg               | Operating    |
| CUORE                  | <sup>130</sup> Te                    | TeO <sub>2</sub> Bolometer                                | 206 kg              | Operating    |
| SNO+                   | <sup>130</sup> Te                    | 0.5% <sup>nat</sup> Te in liquid scintillator             | 1300 kg             | Construction |
| SNO+ Phase II          | <sup>130</sup> Te                    | 2.5% <sup>nat</sup> Te in liquid scintillator             | 8 tonnes            | R&D          |
| Theia-Te               | <sup>130</sup> Te                    | 5% <sup>nat</sup> Te in liquid scintillator               | 31 tonnes           | R&D          |
| KamLAND-Zen 400        | <sup>136</sup> Xe                    | 2.7% in liquid scintillator                               | 370 kg              | Complete     |
| KamLAND-Zen 800        | <sup>136</sup> Xe                    | 2.7% in liquid scintillator                               | 750 kg              | Operating    |
| KamLAND2-Zen           | <sup>136</sup> Xe                    | 2.7% in liquid scintillator                               | ~tonne              | R&D          |
| EXO-200                | <sup>136</sup> Xe                    | Xe liquid TPC   | 160 kg              | Complete     |
| nEXO                   | <sup>136</sup> Xe                    | Xe liquid TPC   | 5 tonnes            | R&D          |
| NEXT-WHITE             | <sup>136</sup> Xe                    | High pressure GXe TPC                                     | ~5 kg               | Operating    |
| NEXT-100               | <sup>136</sup> Xe                    | High pressure GXe TPC                                     | 100 kg              | Construction |
| PandaX                 | <sup>136</sup> Xe                    | High pressure GXe TPC                                     | ~tonne              | R&D          |
| AXEL                   | <sup>136</sup> Xe                    | High pressure GXe TPC                                     | ~tonne              | R&D          |
| DARWIN                 | <sup>136</sup> Xe                    | <sup>nat</sup> Xe liquid TPC                              | 3.5 tonnes          | R&D          |
| LZ                     | <sup>136</sup> Xe                    | <sup>nat</sup> Xe liquid TPC                              |                     | R&D          |
| Theia-Xe               | <sup>136</sup> Xe                    | 3% in liquid scintillator                                 | 50 tonnes           | R&D          |

R&D

Construction

Operating

Complete

# Direct mass measurements

|                         |  |  |
|-------------------------|--|--|
| • Tritium $\beta$ decay | $\left(\sum_i  U_{ei} ^2 m_i^2\right)^{\frac{1}{2}}$ | $< 0.9 \text{ eV}$   |
| • $0\nu 2\beta$ decay   | $\left \sum_i U_{ei}^2 m_i\right $                   | $< 0.3 \text{ eV}$<br>$\langle m_{\beta\beta} \rangle = 440 \text{ meV}$ from HM |
| • Cosmology             | $\sum_i m_i < 0.15 \text{ eV}$                       | Model dependent  |
| • Pion decay            | $m_{\nu\mu} < 190 \text{ keV}$                       | Fairly pointless   |
| • Tau decay             | $m_{\nu\tau} < 18.2 \text{ MeV}$                     | Entirely pointless   |

# Question

Is there an experimental way of directly showing that the neutrino is a Dirac particle? What about an indirect approach?