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Neutrino Oscillation Results from the NOvA Experiment

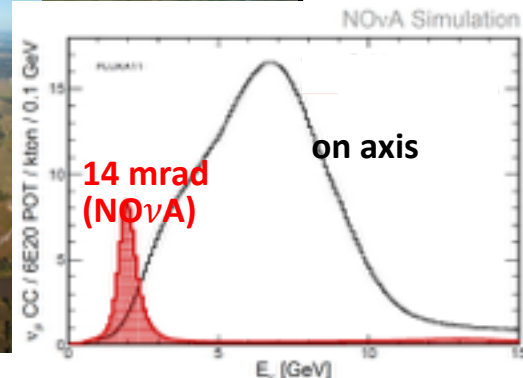
Evan Niner for the NOvA Collaboration

Lake Louise Winter Institute 2016

11 February 2016

NuMI Off-axis ν_e Appearance Experiment

- Two functionally identical detectors 809 km apart and 14 milli-radians off-axis from NuMI beam.
- Narrow-band off-axis muon neutrino beam centered at 2 GeV.
- Neutrino and anti-neutrino beam modes.
- Measure θ_{13} , θ_{23} , mass hierarchy, and δ_{cp}
- First analysis data from February 2014 to May 2015.
- 3.45×10^{20} protons-on-target (POT) accumulated.
 - Full-detector equivalent 2.74×10^{20} POT (7.6% planned exposure)



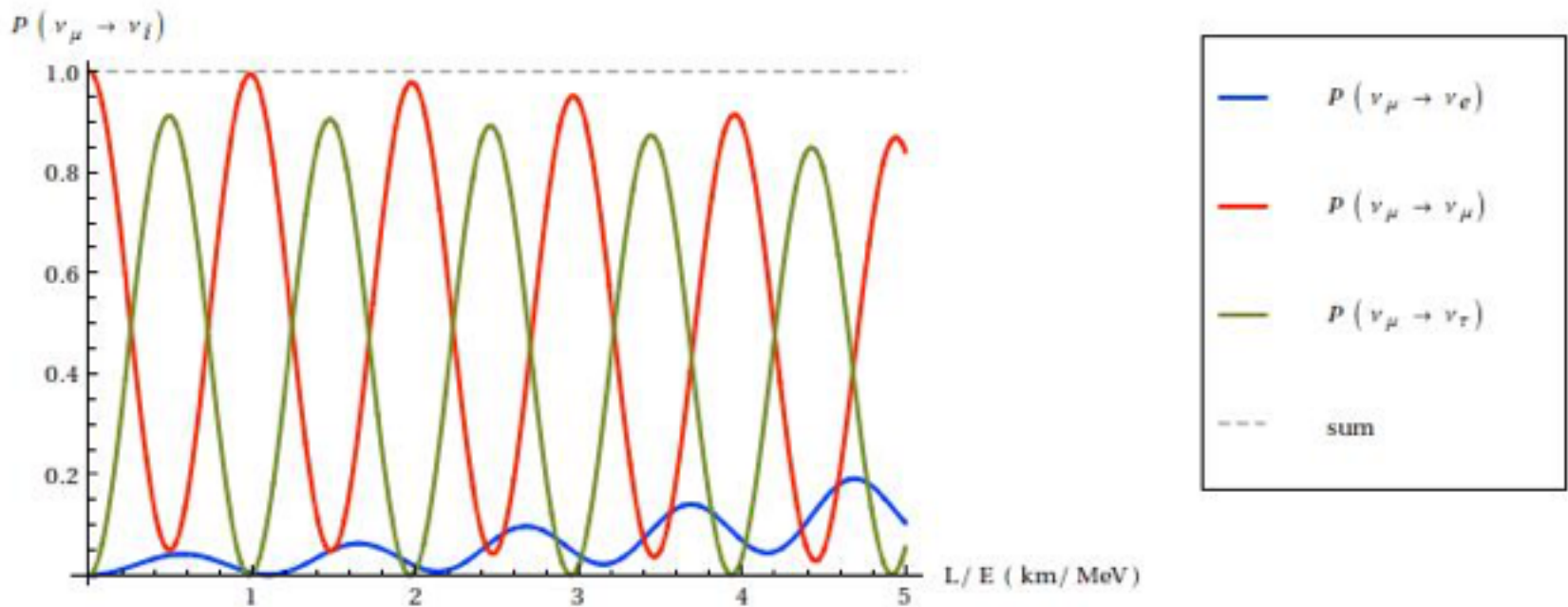


Image from: <http://invisibles.eu/outreach/entry/ceaseless-transformation-three-neutrinos>

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$\nu_\mu \rightarrow \nu_\mu$$

$$\nu_\mu \rightarrow \nu_\tau$$

atmospheric and
long-baseline

$$\nu_e \rightarrow \nu_e$$

$$\nu_\mu \rightarrow \nu_e$$

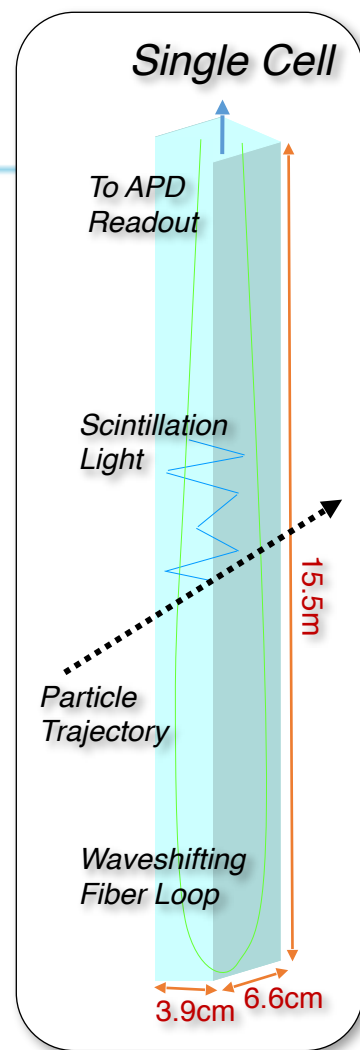
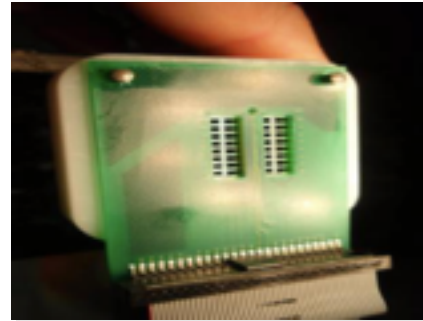
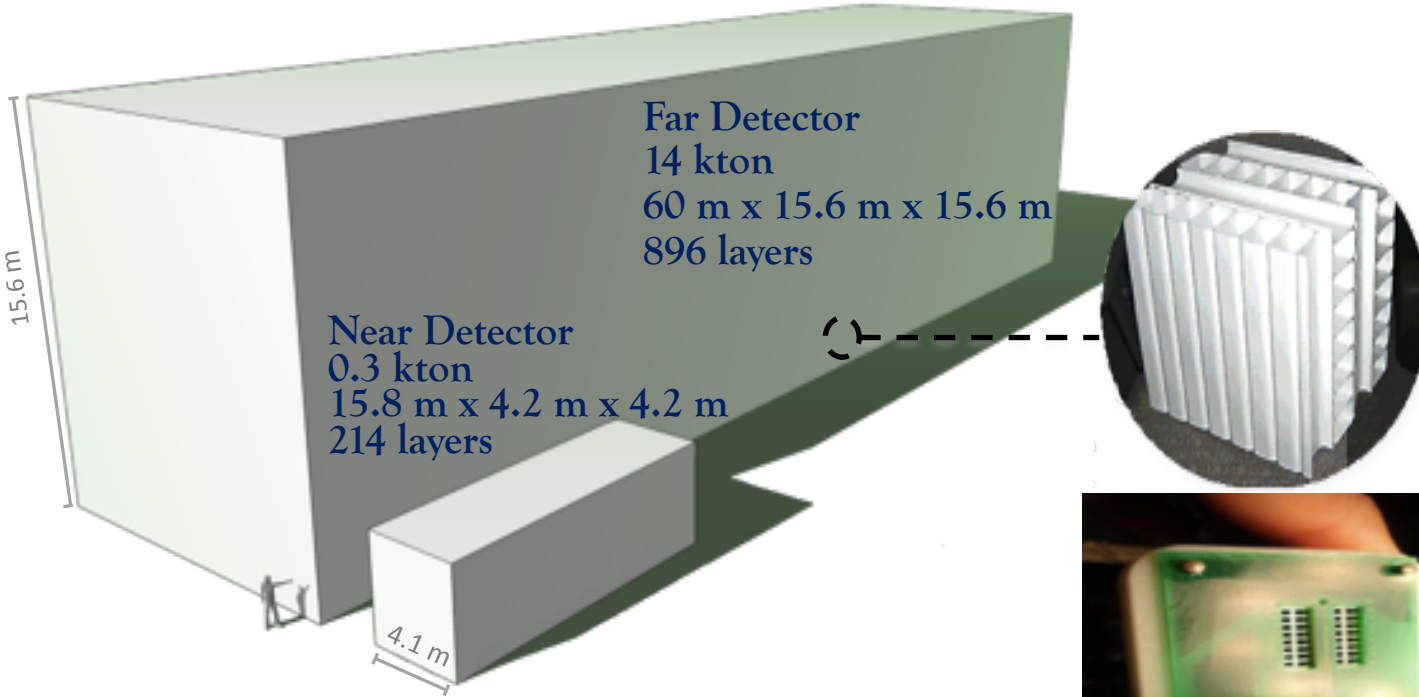
reactor and
long-baseline

$$\nu_e \rightarrow \nu_e$$

$$\nu_e \rightarrow \nu_\mu + \nu_\tau$$

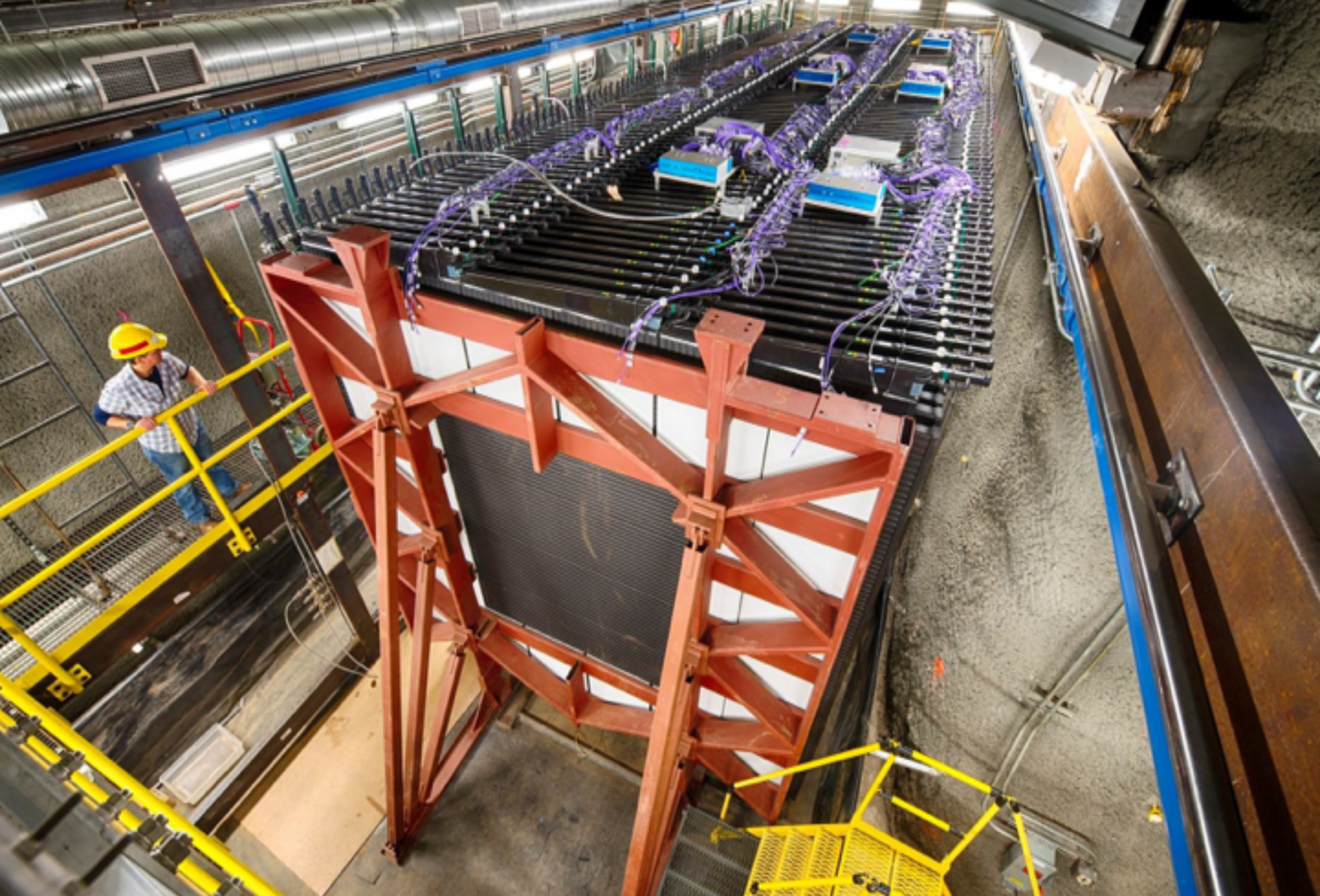
solar and
reactor

NOvA Design



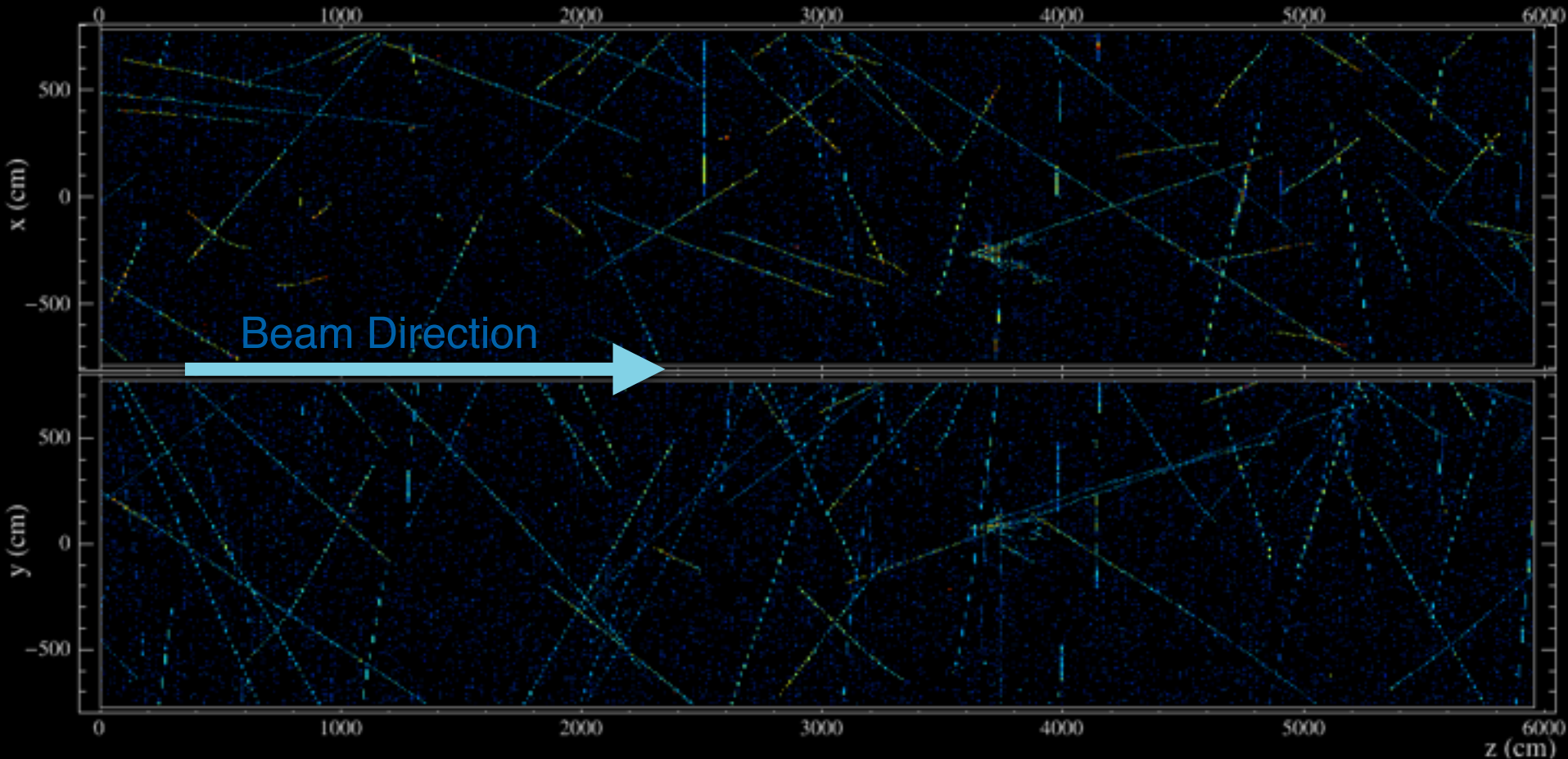
- Functionally identical detectors made of planes of extruded plastic PVC cells.
- Each cell filled with liquid scintillator, light collected in a wavelength shifting fiber, coupled to APD.
- Low-z design enhances electron/photon separation.
- Near Detector has 10 steel planes at downstream end to range out muons.





Far Detector 550 μ s Readout Window

Cell hits colored by charge deposition



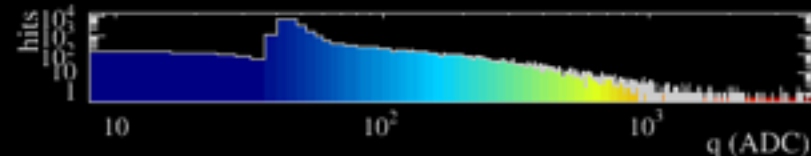
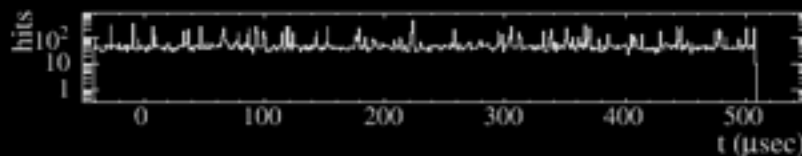
NOvA - FNAL E929

Run: 18620 / 13

Event: 178402 / --

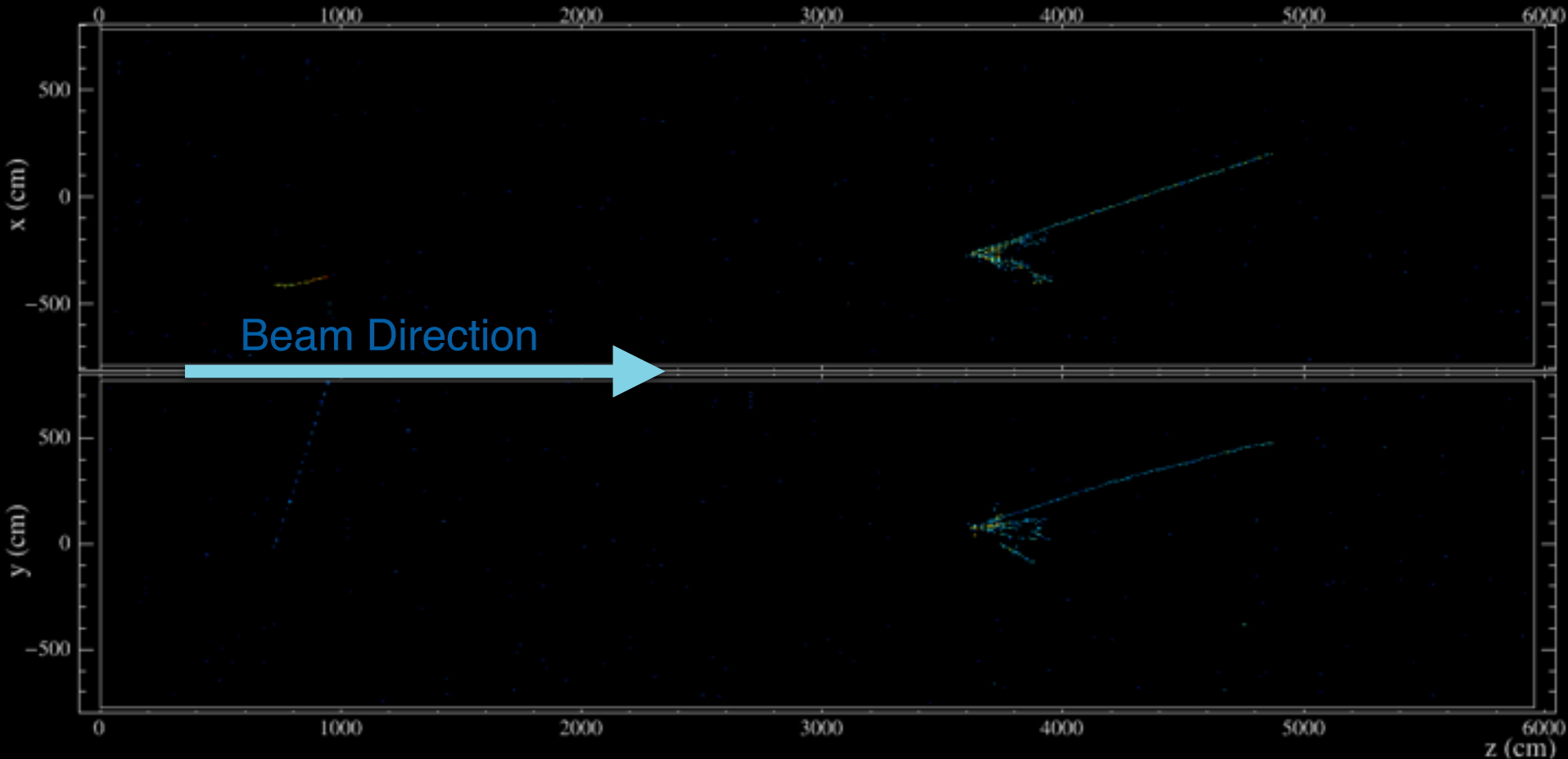
UTC Fri Jan 9, 2015

00:13:53.087341608



Far Detector 10 μ s NuMI Beam Window

Cell hits colored by charge deposition



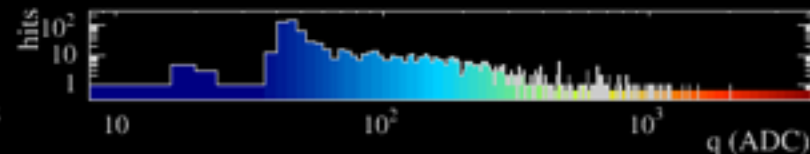
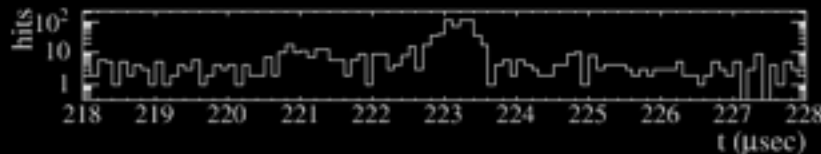
NOvA - FNAL E929

Run: 18620 / 13

Event: 178402 / --

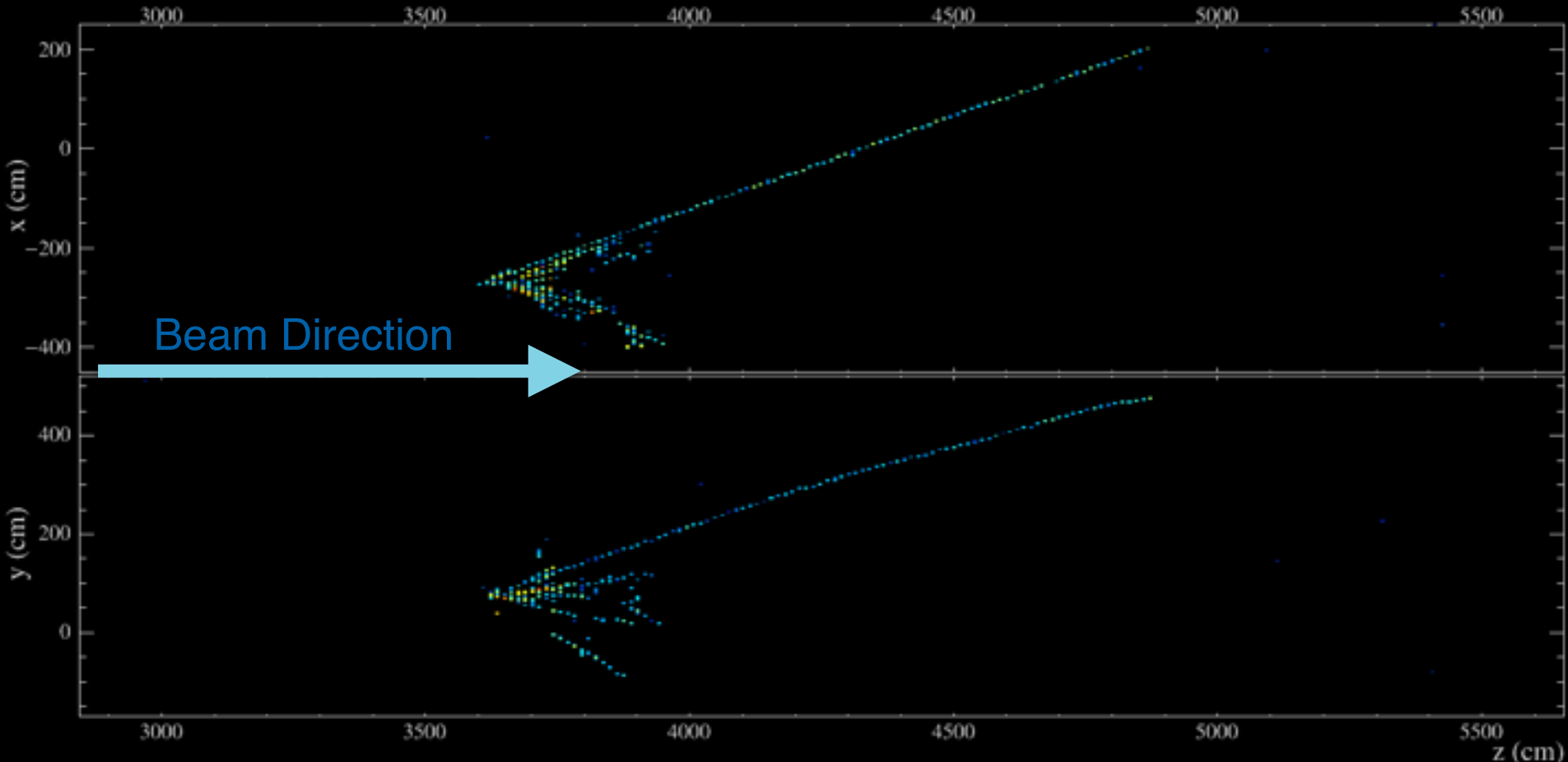
UTC Fri Jan 9, 2015

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Far Detector Neutrino Interaction

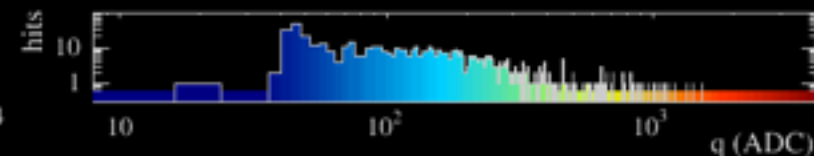
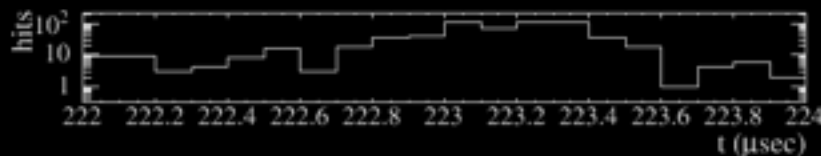
Cell hits colored by charge deposition



NOvA - FNAL E929

Run: 18620 / 13
Event: 178402 / --

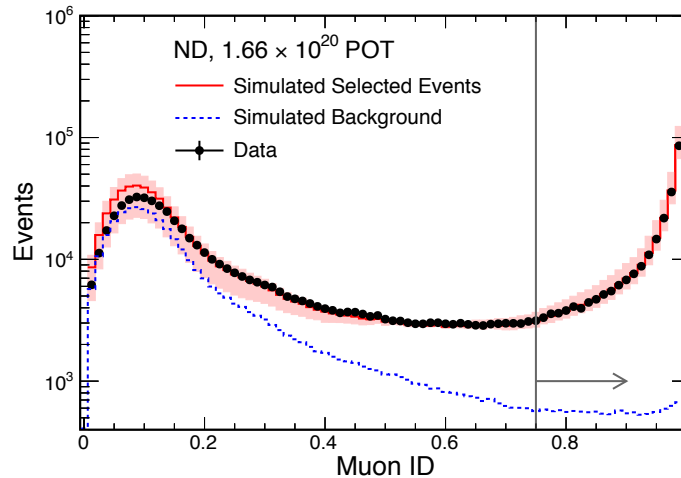
UTC Fri Jan 9, 2015
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Muon Neutrino Disappearance Analysis

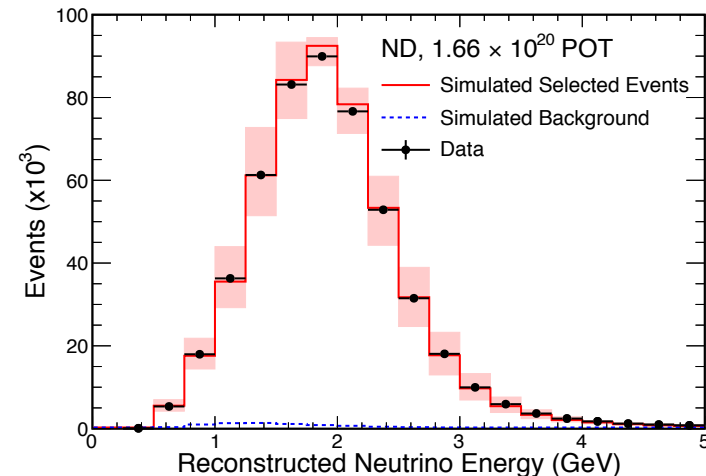
Muon Neutrino Disappearance Analysis

- Use ND selected ν_μ 's to extrapolate FD prediction, reduces systematics.
- Cosmic ray backgrounds rejected by factor 10^7 .
- 99% percent reduction of neutral current events.



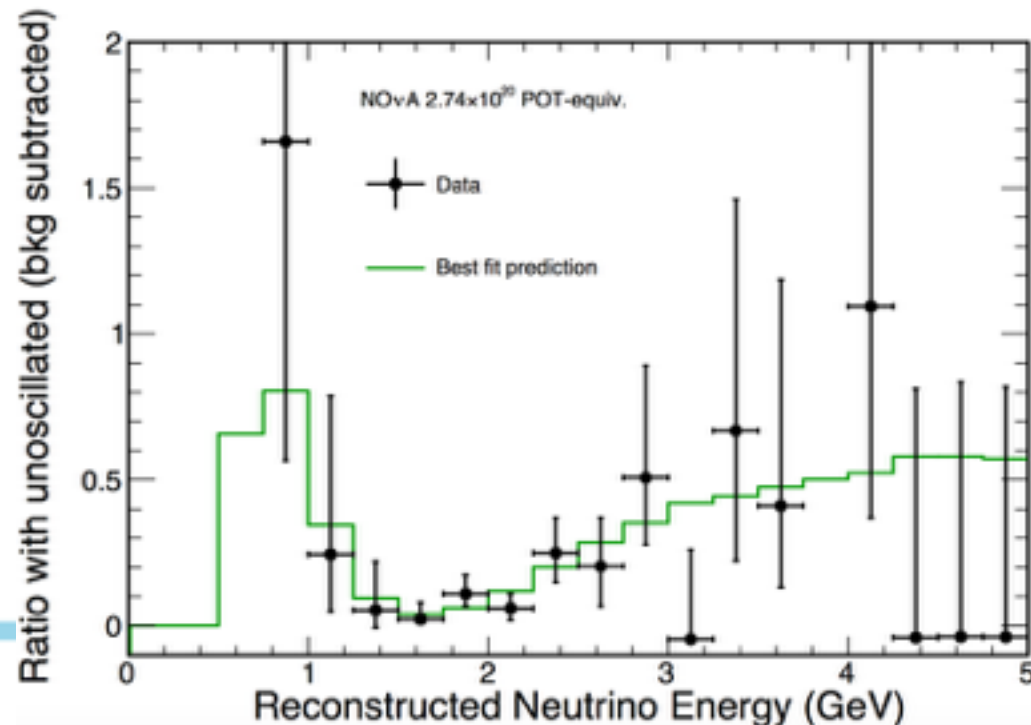
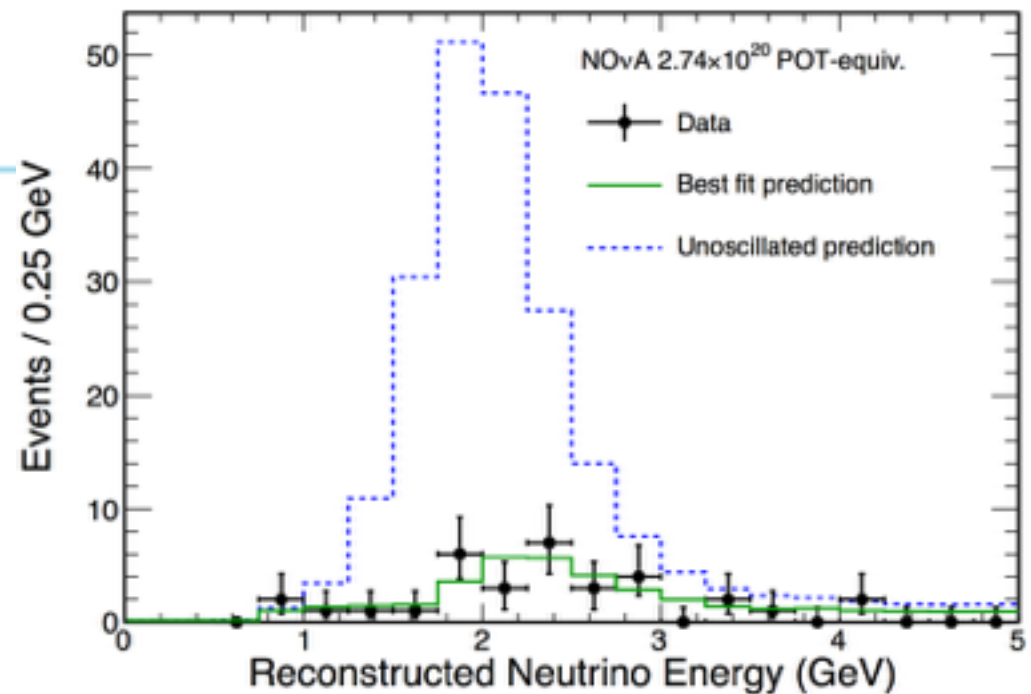
kNN based on track length, dE/dx , multiple-scattering, fraction of overlapping hadronic energy.

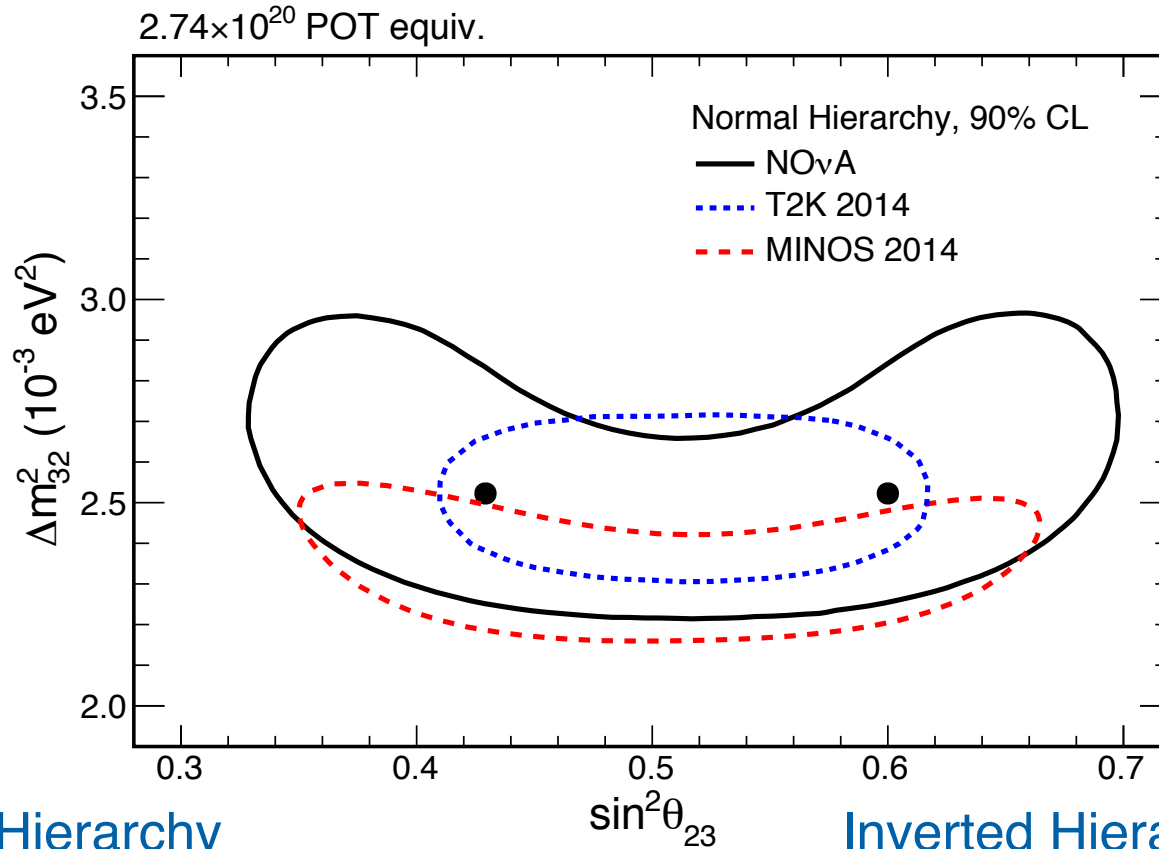
Energy spectrum in Near Detector with 14% hadronic energy correction, equivalent to 6% neutrino energy shift.



ν_μ Spectrum

- 211.8 ± 12.5 (syst.) events predicted in the absence of oscillations.
- 33 candidate events between 0 and 5 GeV observed.





Normal Hierarchy

$$\Delta m_{32}^2 = (2.52^{+0.20}_{-0.18}) \times 10^{-3} \text{ eV}^2$$

$$\sin^2(\theta_{23}) = [0.38, 0.65]$$

(68% CL)

Degenerate best fit points at 0.43 and 0.60

Inverted Hierarchy

$$\Delta m_{32}^2 = (-2.52 \pm 0.19) \times 10^{-3} \text{ eV}^2$$

$$\sin^2(\theta_{23}) = [0.37, 0.64]$$

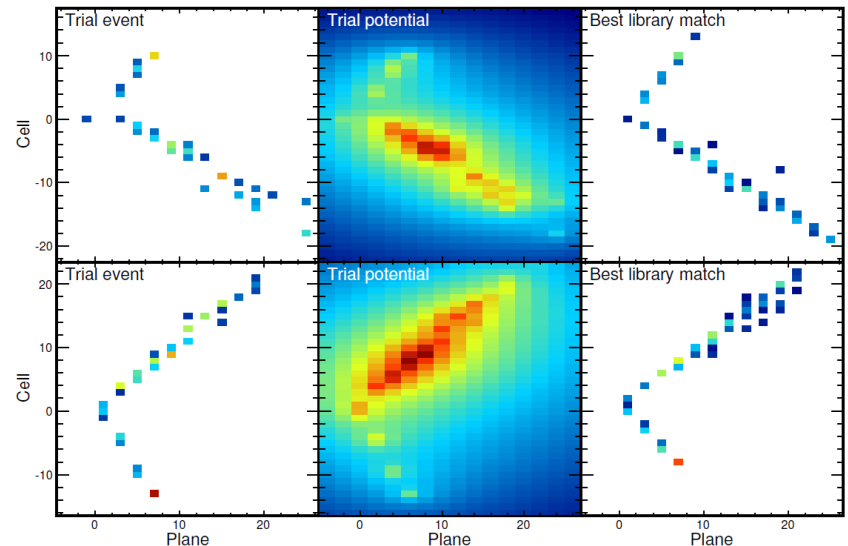
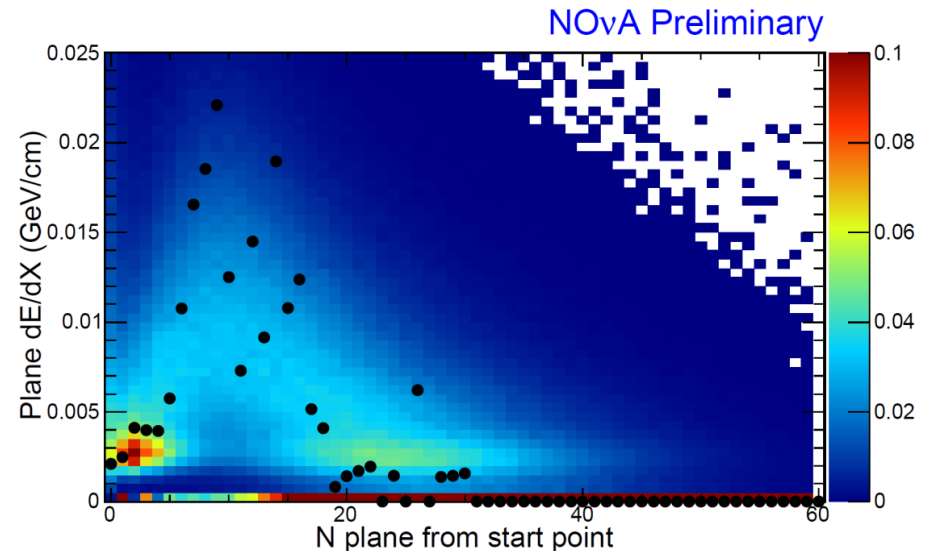
(68% CL)

Degenerate best fit points at 0.44 and 0.59

Electron Neutrino Appearance Analysis

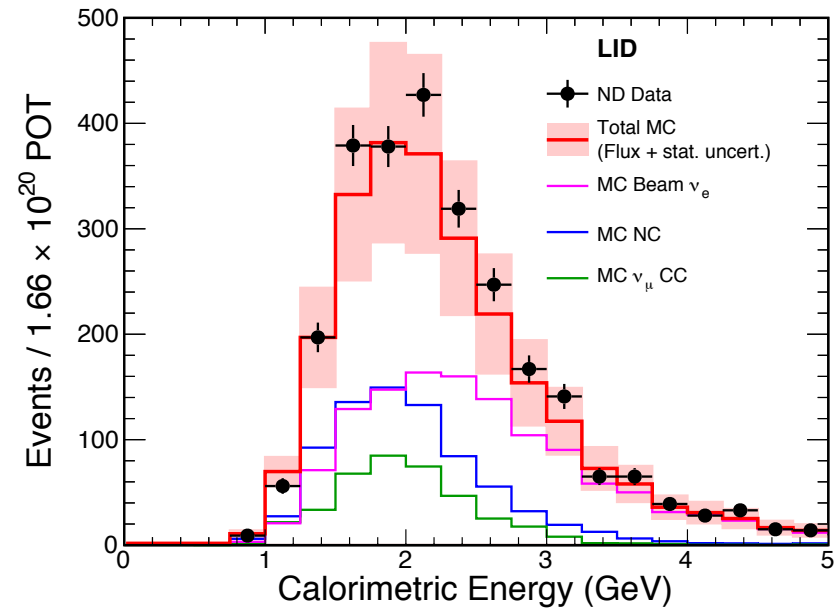
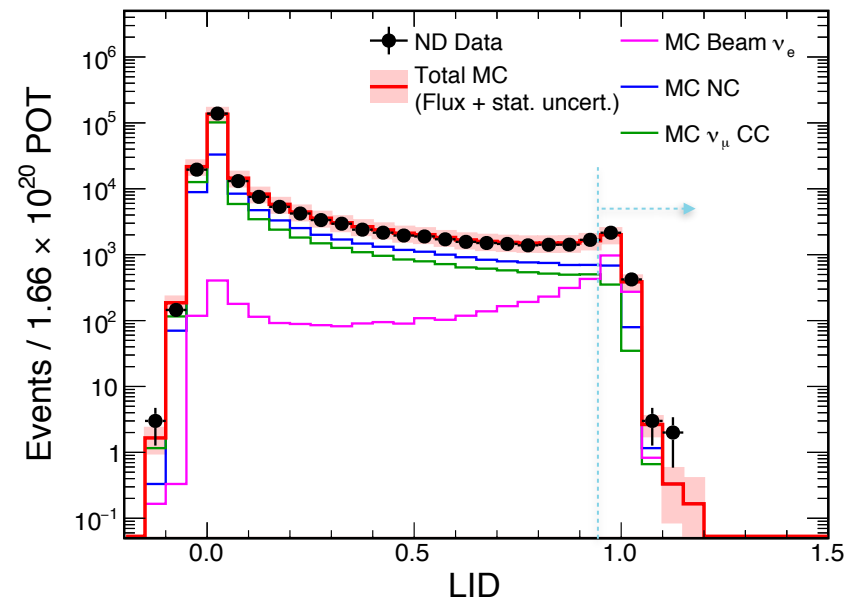
ν_e Identification

- Likelihood Identifier (LID)
 - Compare longitudinal and transverse dE/dx in leading shower to template histograms for $e/p/n/\mu/\pi^\pm/\pi^0/\gamma$.
 - Build neural net from these inputs and reconstructed quantities.
- Library Event Matching (LEM)
 - Compares input event to simulated event library.
 - Properties from most similar events fed into decision tree.
- 62% event overlap between selectors.
- LID chosen before unblinding as primary selector.



ν_e Appearance Analysis Strategy

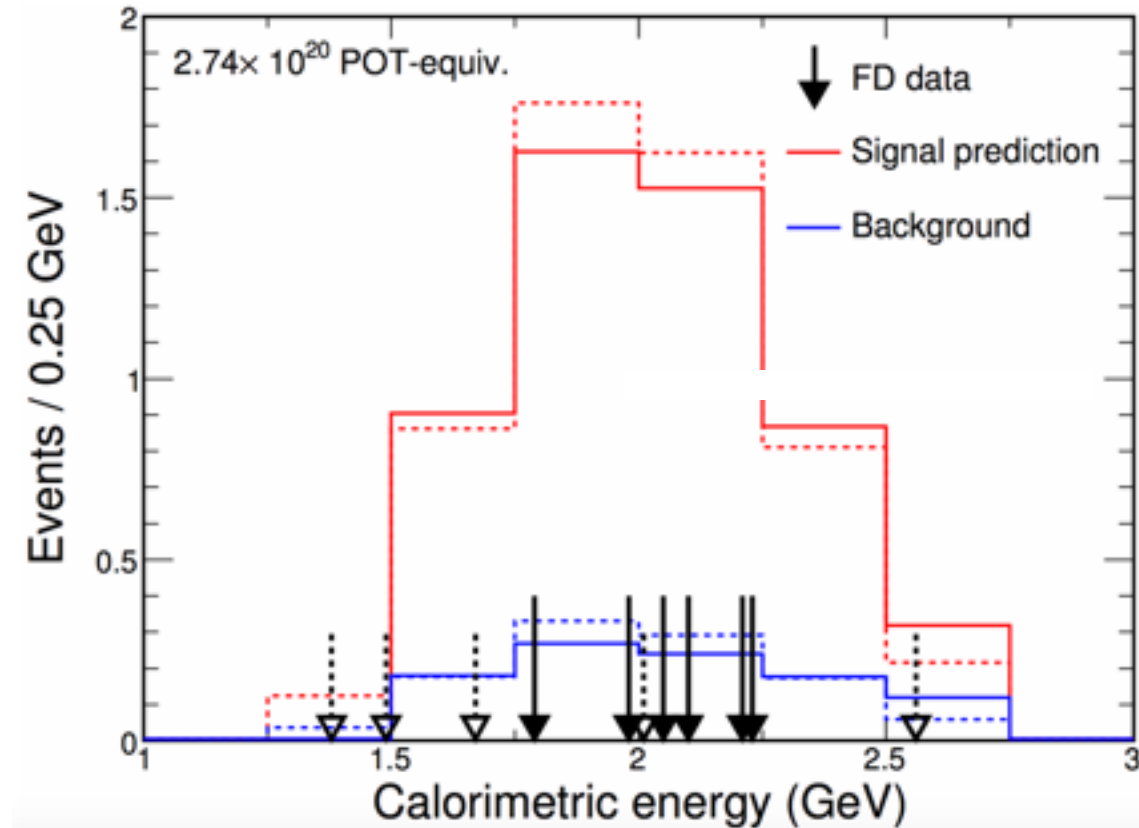
- FD background prediction extrapolated from ND
 - ND selects $\sim 7\%$ more background in data relative to simulation.
- Combination of containment, topology and event classifier achieve cosmic rejection factor $>10^8$. Effective FD fiducial volume of 10 kT.
- “Cut and count” analysis between 1.5 and 2.7 GeV in FD for the primary selector.



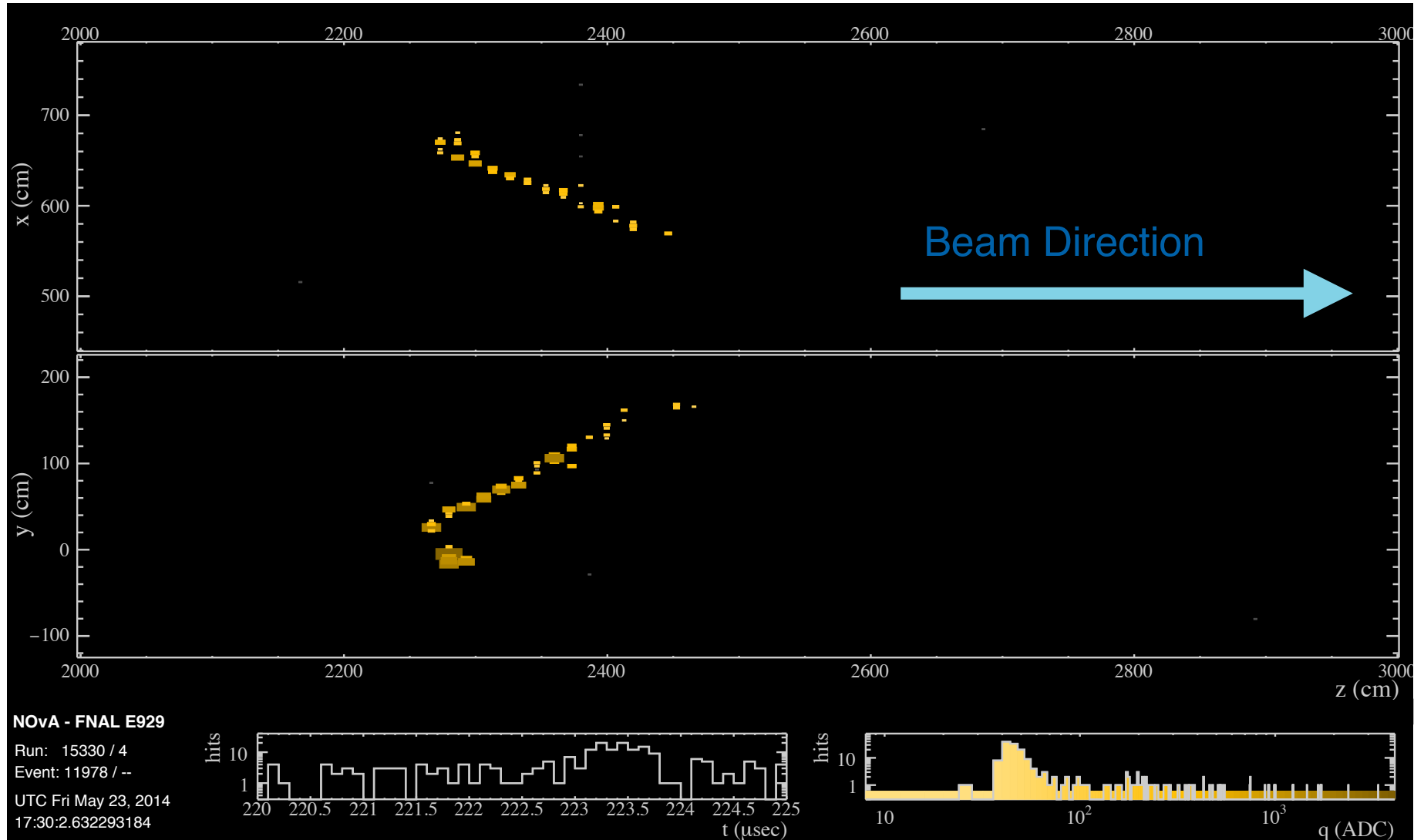
ND energy spectrum with LID > 0.95

ν_e Appearance Analysis Results

- LID observed 6 events on a background prediction of 0.99 ± 0.11 (syst), 3.3σ excess.
- LEM observed 11 events on a background of 1.07 ± 0.14 (syst), 5.3σ excess.
- All LID events in LEM set.
- 7.8% probability of this overlap configuration or one less likely.

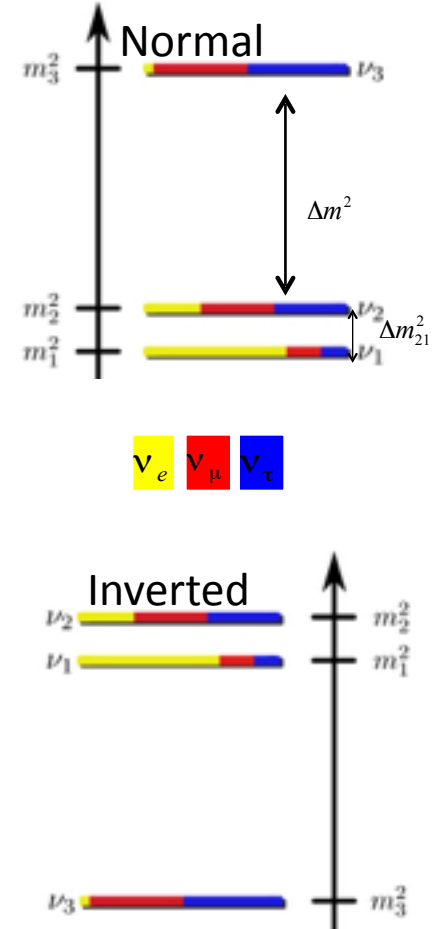
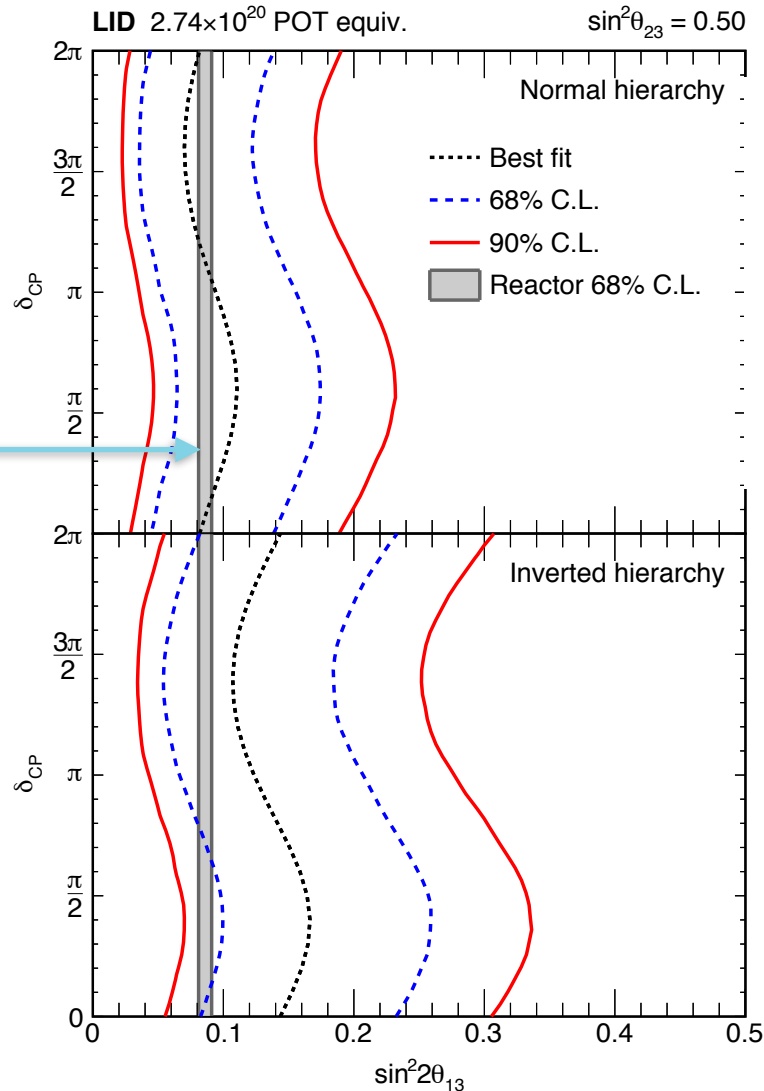


Far Detector ν_e CC candidate



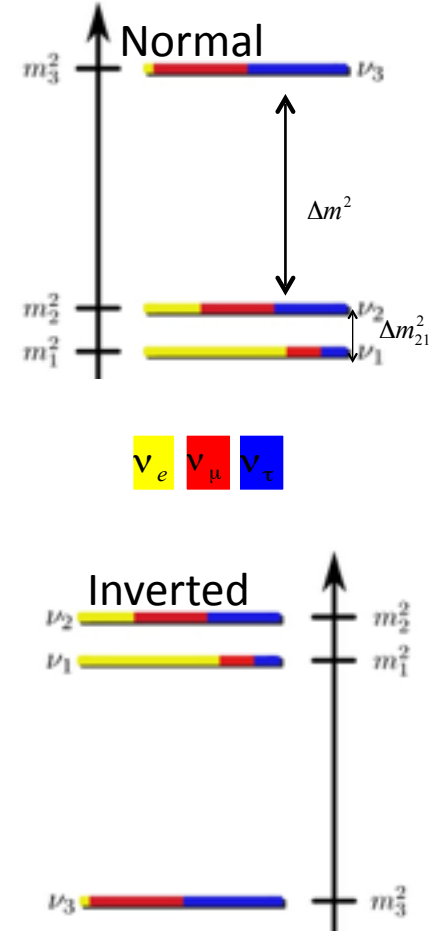
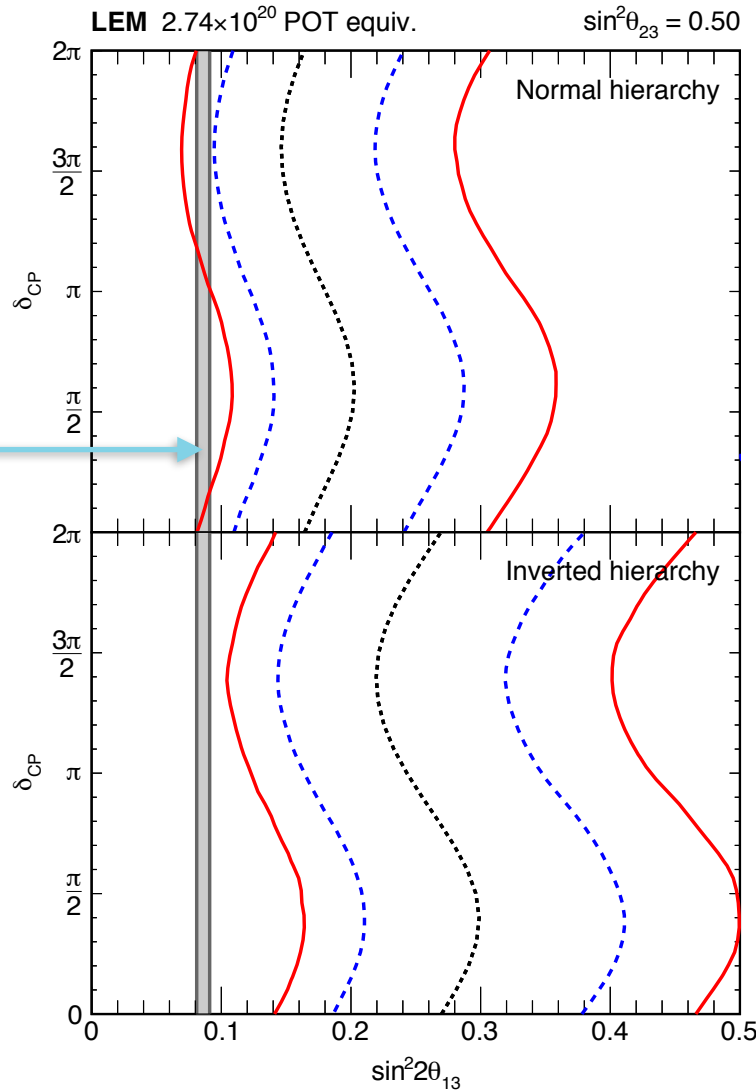
ν_e Appearance Analysis Results

Global best fit from reactor experiments

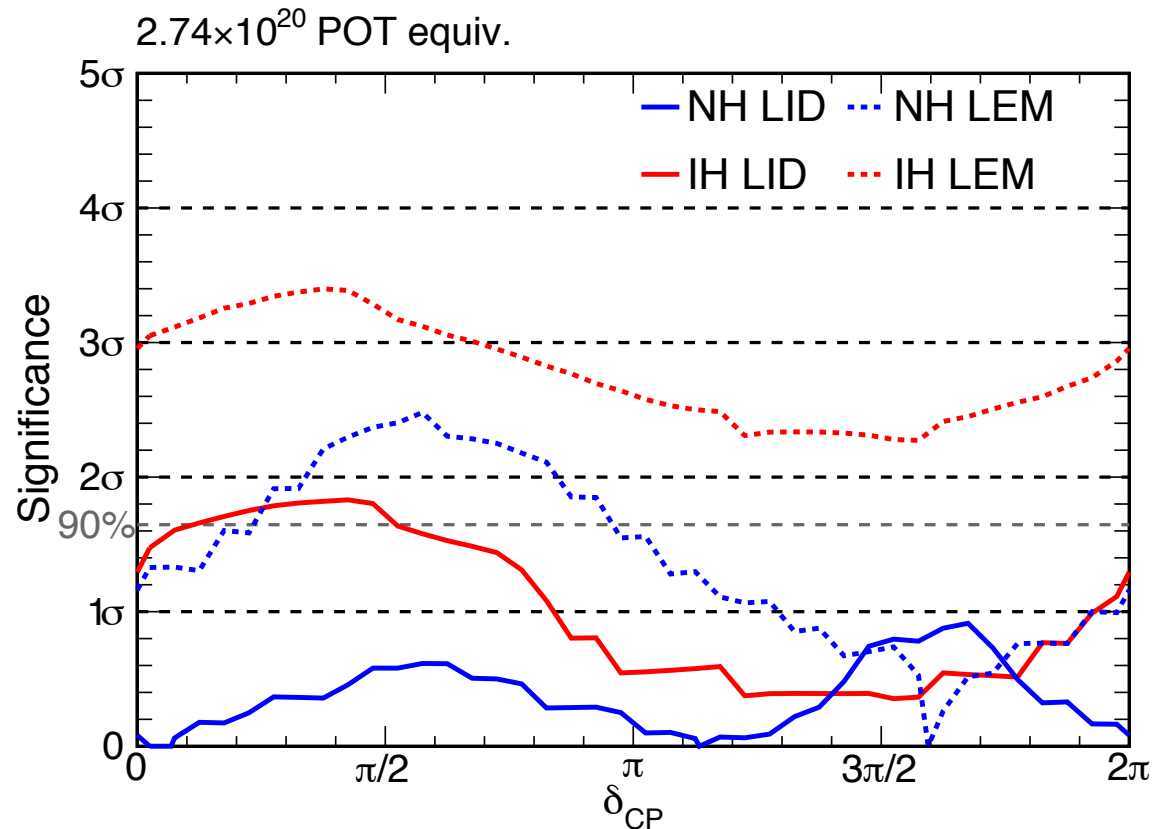


ν_e Appearance Analysis Results

Global best fit from reactor experiments



- Apply global reactor constraint
 - $\sin^2\theta_{13}=0.086\pm 0.05$
- Marginalize over θ_{23} .
- Both selectors weakly prefer normal mass hierarchy and $\pi < \delta_{CP} < 2\pi$.
- This preference is consistent with T2K (arXiv:1502.01550)



Summary

- First oscillation results from NOvA program with 7.6% of planned exposure.
- ν_μ disappearance results consistent with MINOS and T2K.
- ν_e appearance results hint at normal mass hierarchy and $\pi < \delta_{cp} < 2\pi$, consistent with T2K.
- Near Detector cross section studies in progress with ν_e CC and coherent π^0 results shown at NuINT, publications in progress.
- Expect to produce second analysis with double the statistics in 2016.
- Stay Tuned!

Thank You



ν_e Appearance Oscillation Probability

$$P(\nu_\mu \rightarrow \nu_e) \approx \left| \sqrt{P_{atm}} e^{-i(\Delta_{32} + \delta)} + \sqrt{P_{sol}} \right|^2$$

$$= P_{atm} + P_{sol} + 2\sqrt{P_{atm}P_{sol}} (\cos \Delta_{32} \cos \delta \mp \sin \Delta_{32} \sin \delta)$$

$$\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \frac{\sin(\Delta_{31} - aL)}{\Delta_{31} - aL} \Delta_{31}$$

Depends on relative sign of “a” and Δ_{31}

$$\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \frac{\sin(aL)}{aL} \Delta_{21}$$

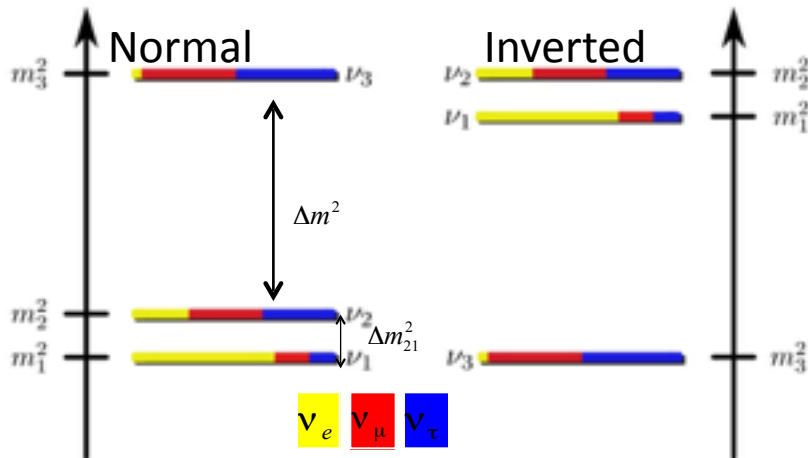
$aL=0.08$ for $L=295\text{km}$ T2K baseline
 $aL=0.23$ for $L=810\text{km}$ NOvA baseline

$$a = \frac{G_F N_e}{\sqrt{2}} \approx \frac{1}{3500\text{km}}$$

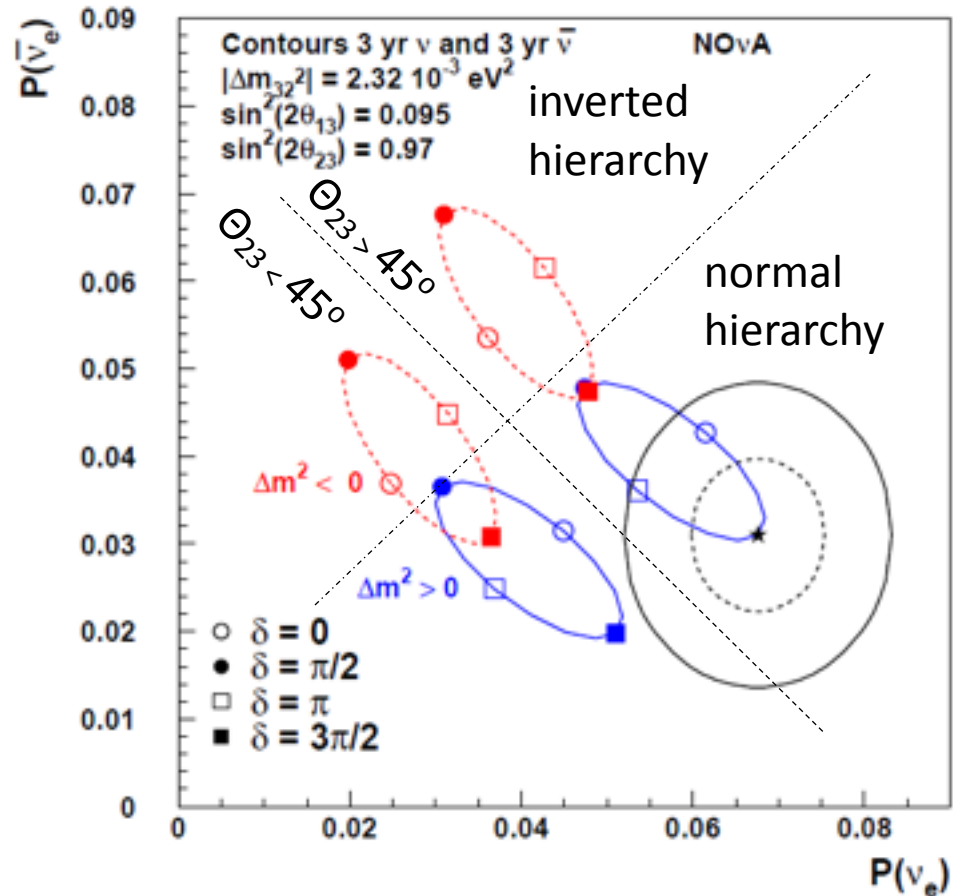
Oscillation probability is sensitive to: **mass ordering**, **CP violating phase**, and θ_{23} **octant**.

Relation of Oscillation Parameters in NOvA

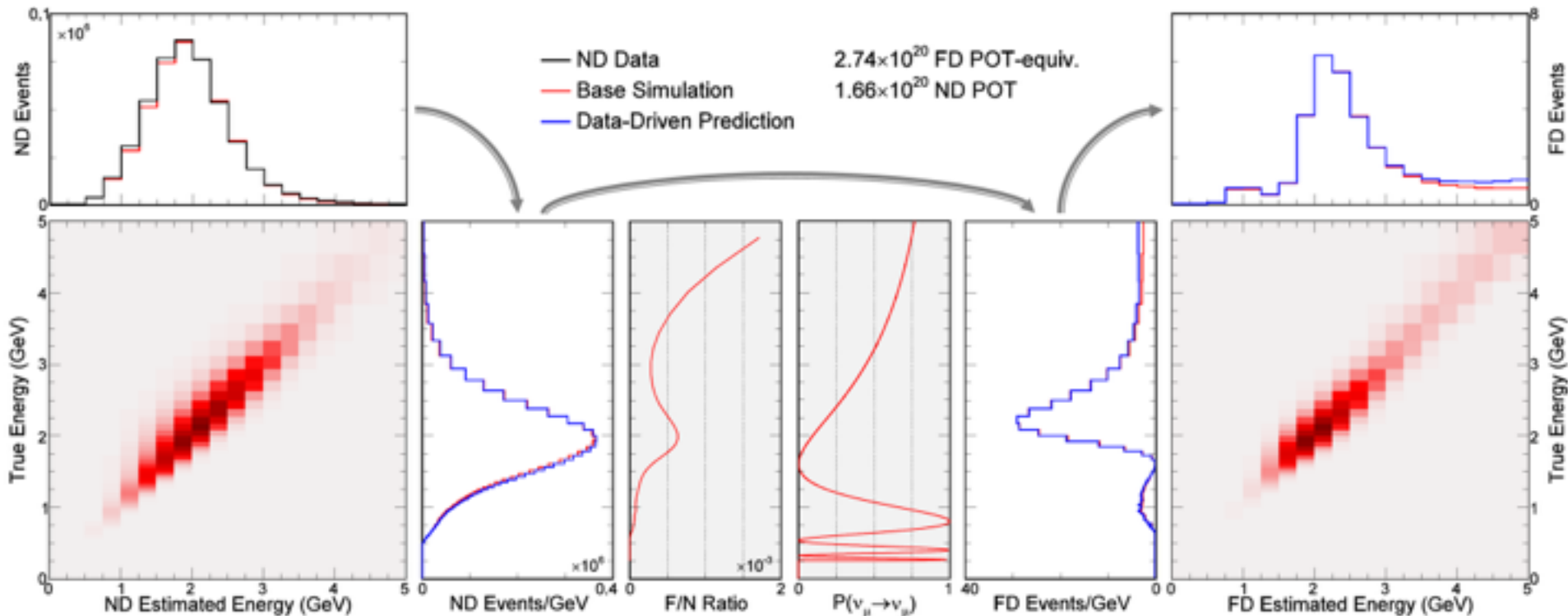
- NOvA makes a measurement of the oscillation probabilities:
 - $P(\nu_\mu \rightarrow \nu_e)$
 - $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$
- The measured probabilities depend on the mass hierarchy, θ_{23} octant, and δ_{cp} .

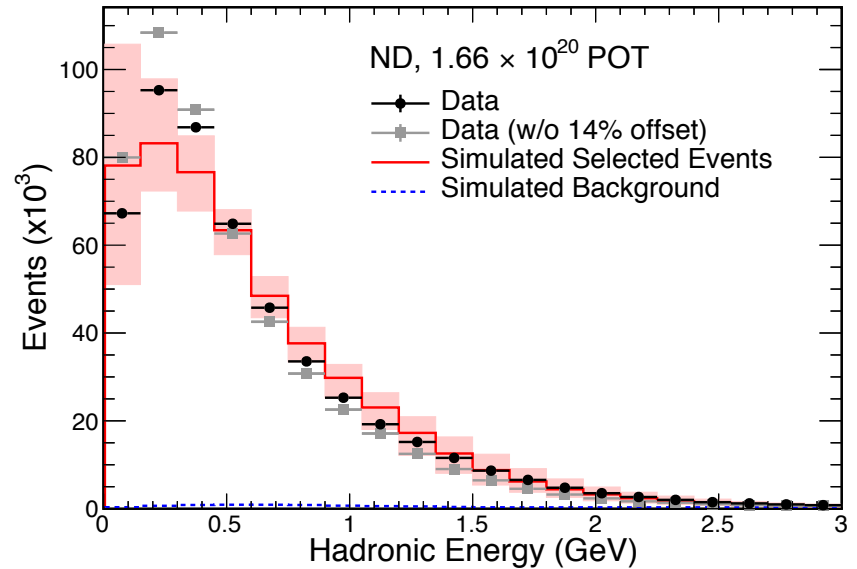
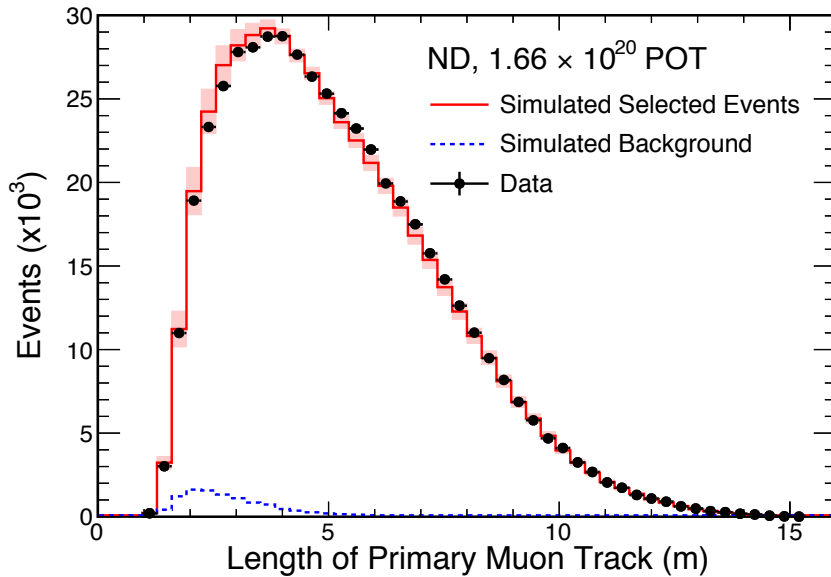


1 and 2 σ Contours for Starred Point



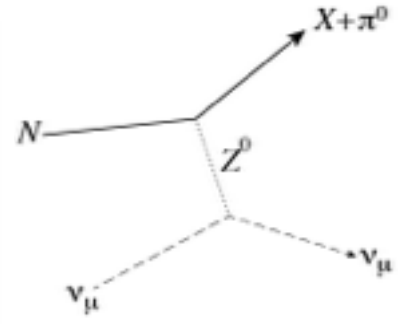
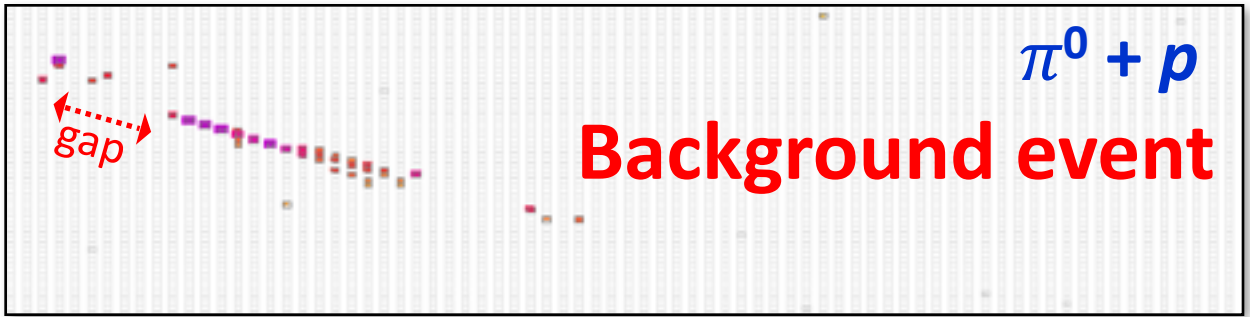
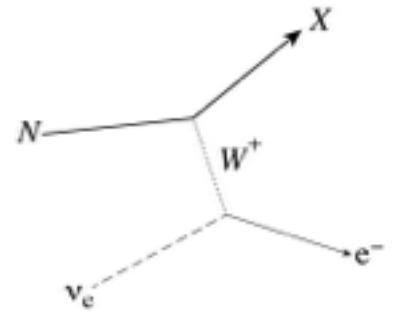
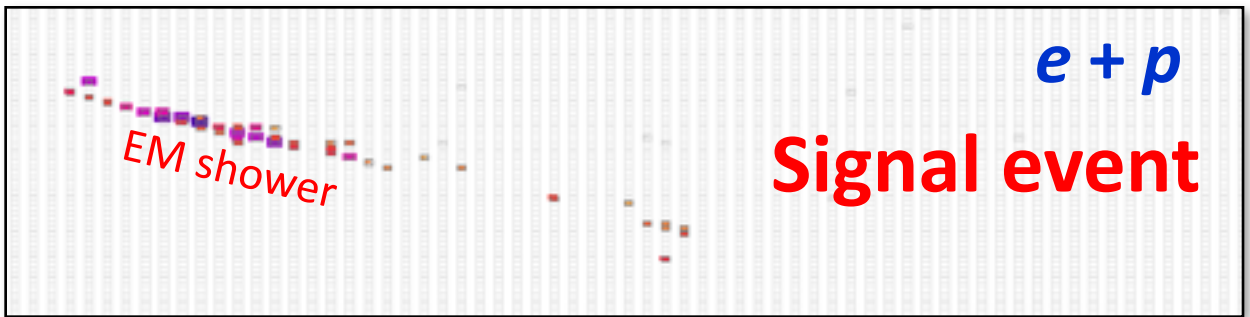
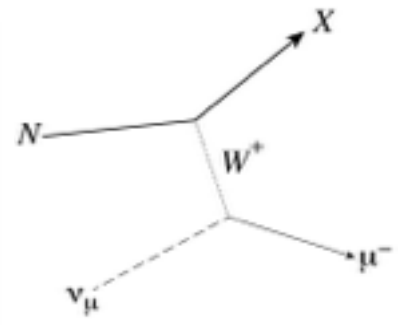
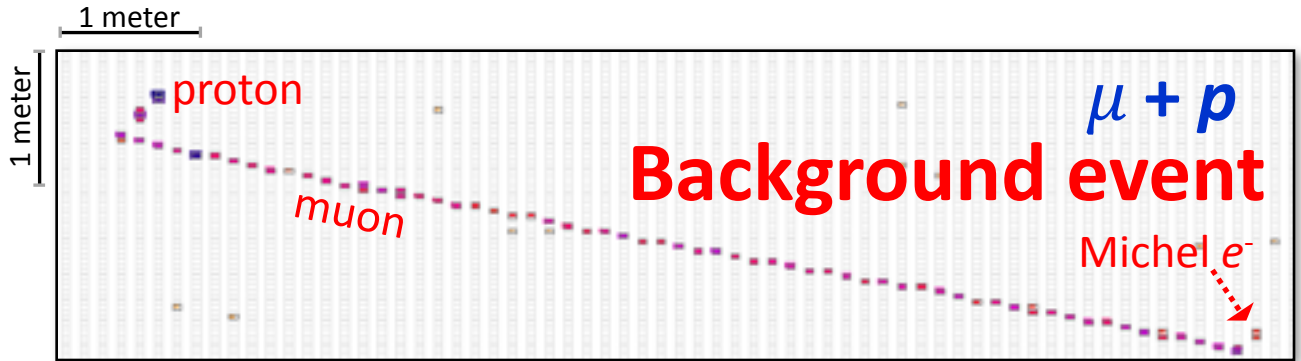
- (1) Estimate the underlying **true energy distribution** of selected ND events
 - (2) Multiply by expected **Far/Near event ratio** and $\nu_\mu \rightarrow \nu_\mu$ **oscillation probability** as a function of true energy
 - (3) Convert FD true energy distribution into **predicted FD reco energy distribution**
- Systematic uncertainties** assessed by **varying all MC-based steps**





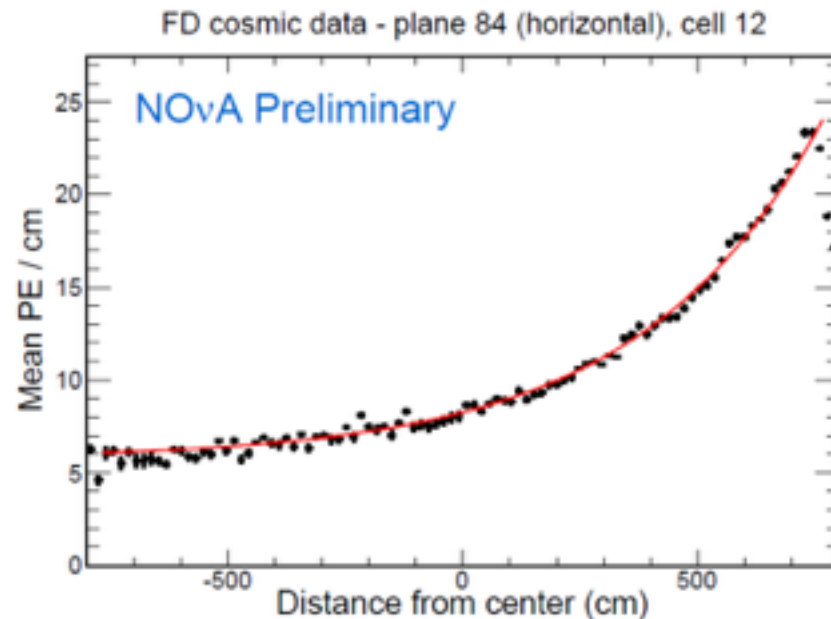
$$E_{\nu} = E_{\mu} + E_{\text{hadrons}}$$

- Muon variables in agreement
- Best fit to hadronic energy prefers 14% increase in data

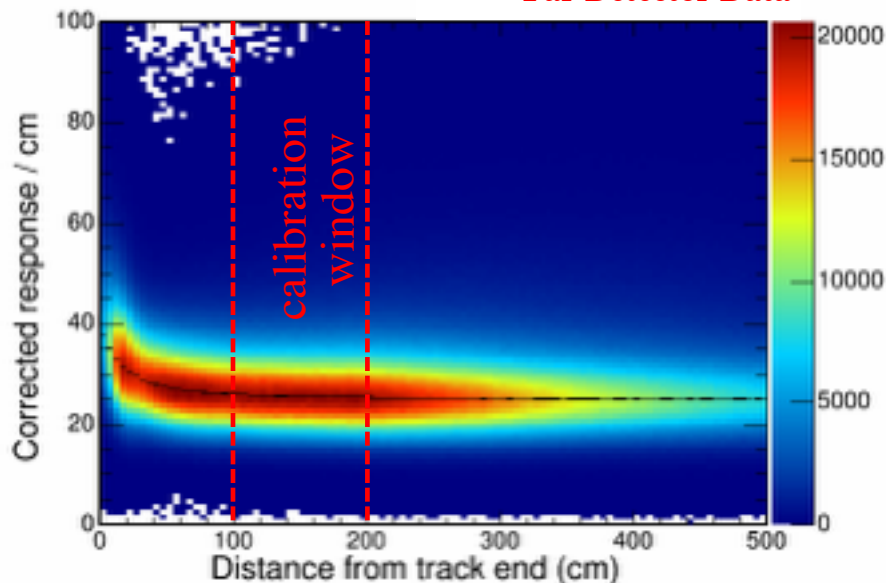


1 radiation length = 38cm (6 cell depths, 10 cell widths)

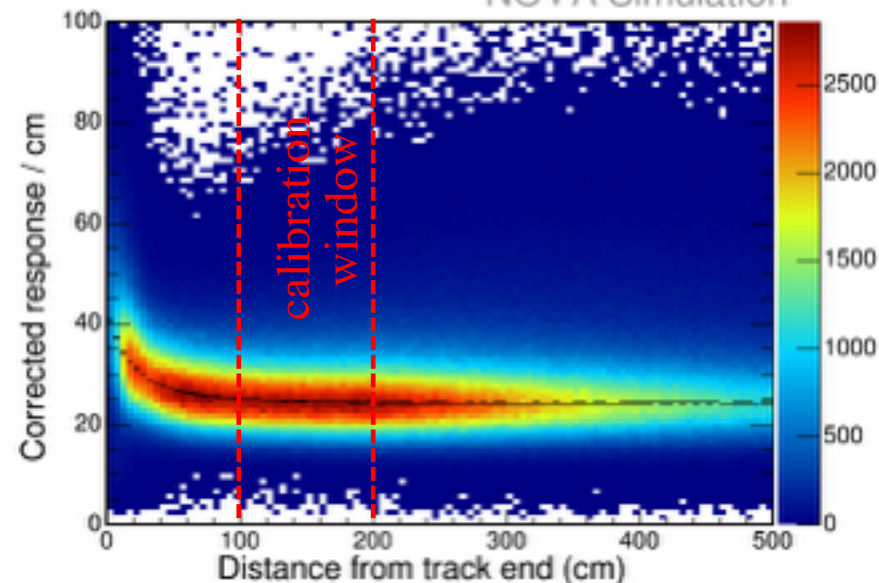
- Correct for attenuation in each cell using through-going cosmic muons (right)
- Set absolute energy scale using stopping muons in data and tuned Monte Carlo (below)



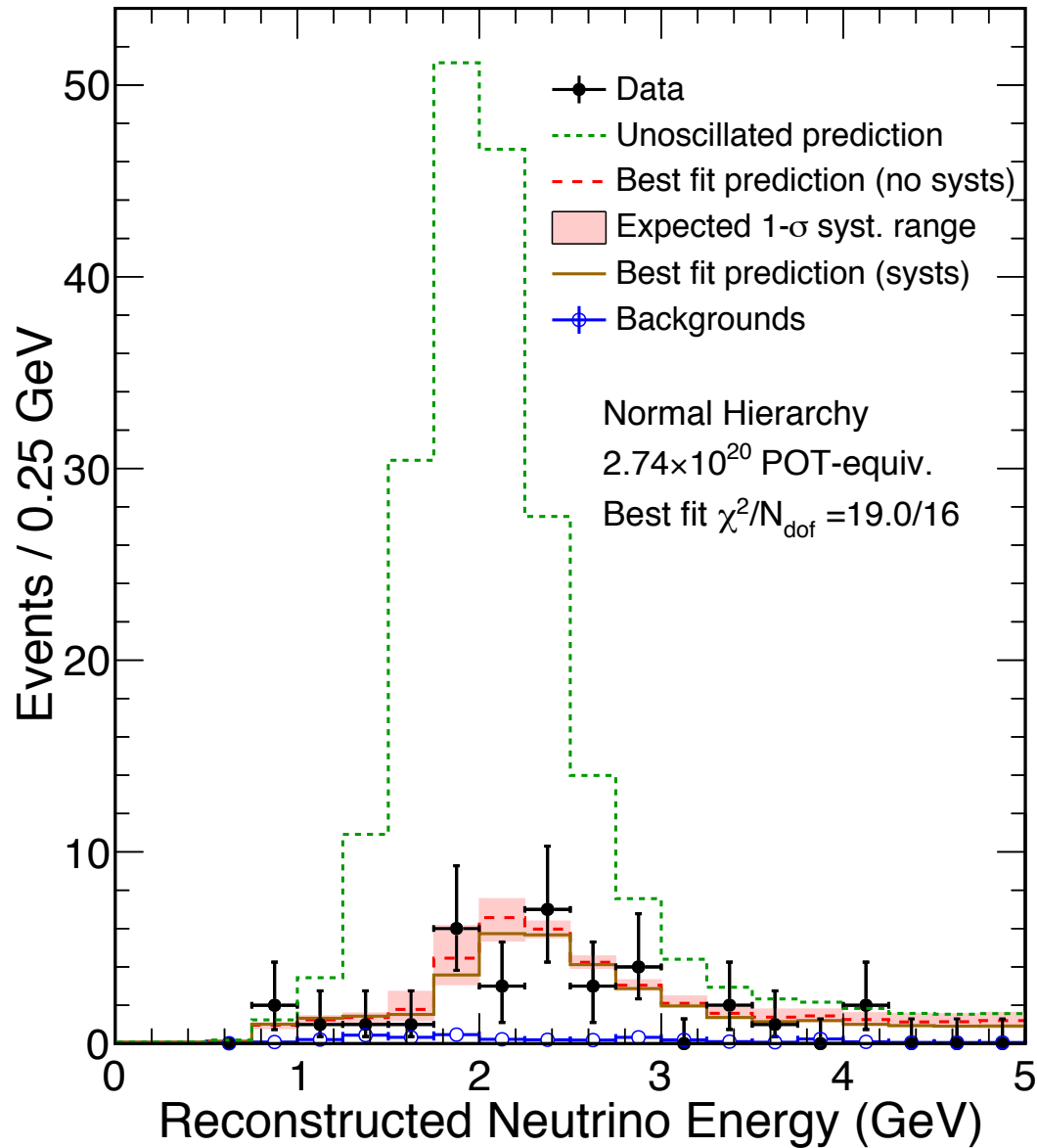
Far Detector Data

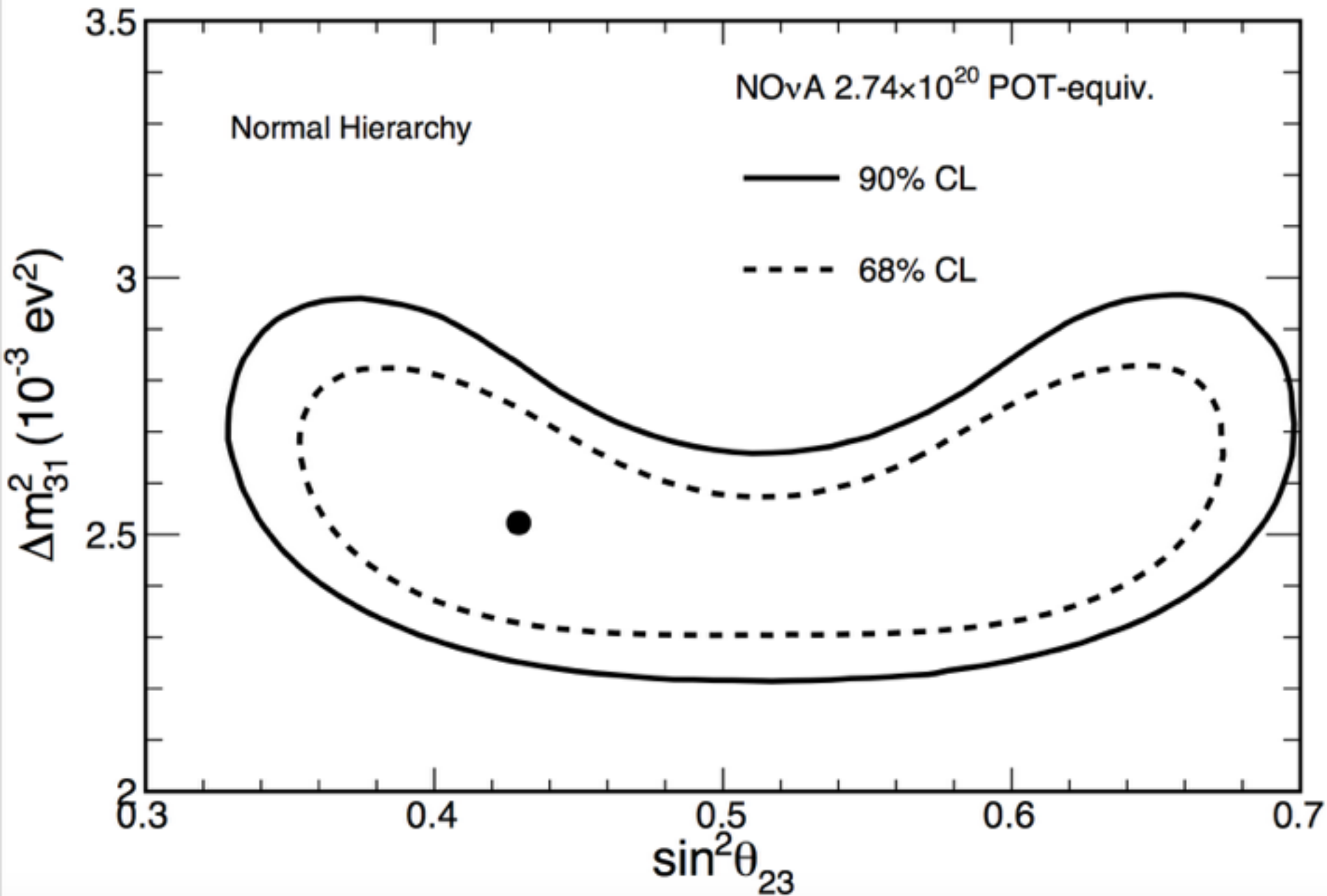


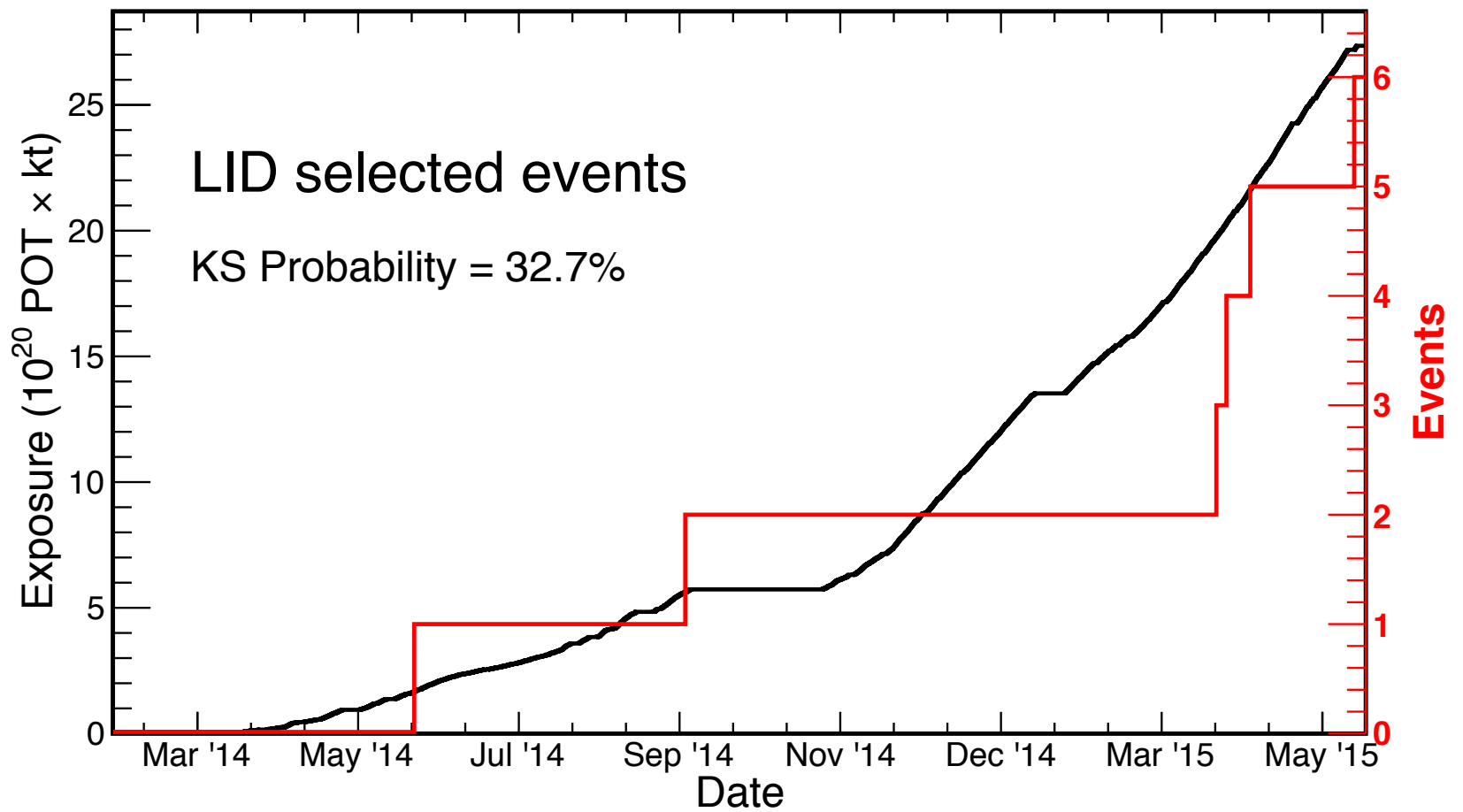
NOvA Simulation

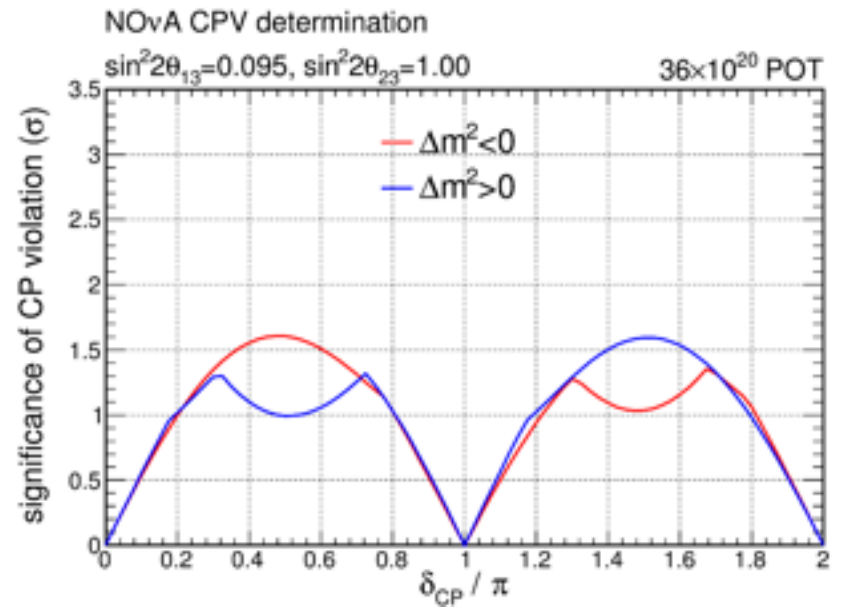
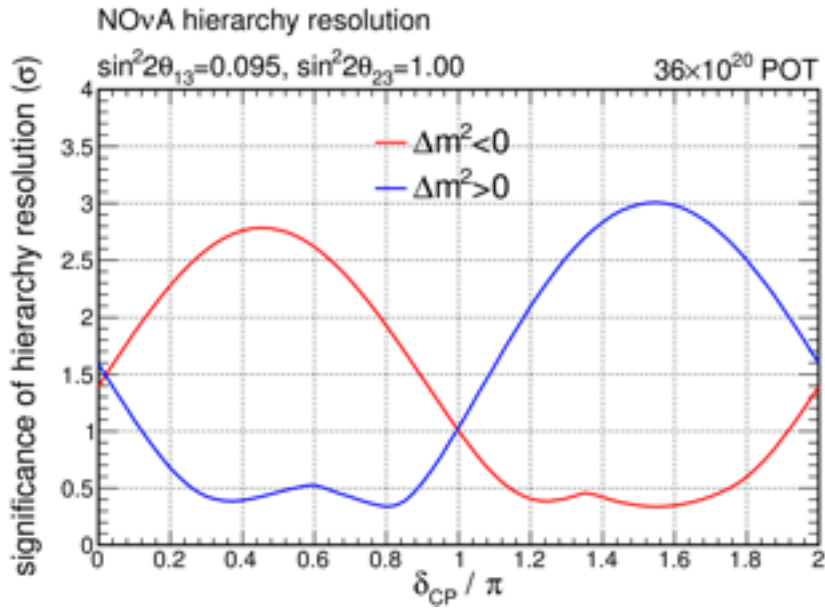


2.74×10²⁰ POT equiv.

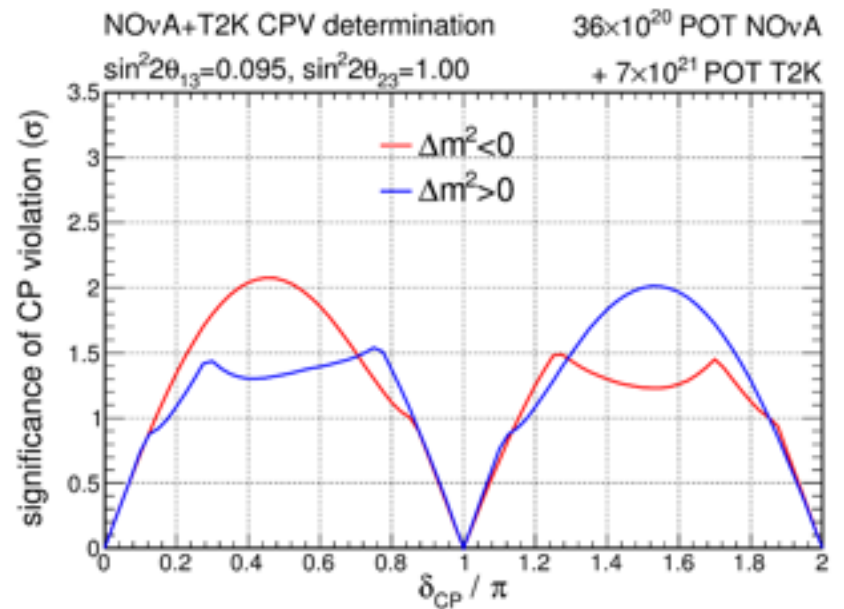
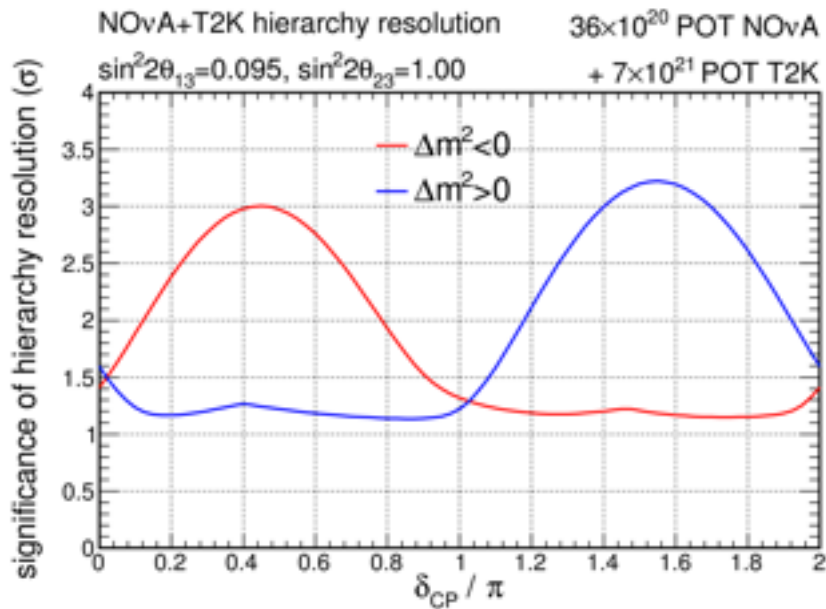








Double-click to edit



Double-click to edit

