

Winter Institute 2016,  
Lake Louise, 7-14 Feb 2016

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On behalf of the OPERA Collaboration



# Latest results from OPERA



- $\nu_{\mu} \rightarrow \nu_{\tau}$  appearance
- Sterile  $\nu$  searches



Laboratori Nazionali  
del Gran Sasso, INFN, Italy

# The long way to appearance



- **Disappearance** a “leading” effect: deficit of atmospheric  $\nu_\mu \rightarrow$ 
  - 1998 discovery of  $\nu$ -oscillations by Super-K, MACRO, K2K ...
- **Appearance** on the other hand:

- At the **solar scale**. Reactors and solar  $\nu$ .

$\nu_e \rightarrow \nu_\mu$  "IMPOSSIBLE":  $\mu$  is below threshold

- At the **atmospheric scale**. Atmospheric- $\nu$ , artificial beams.

$\nu_\mu \rightarrow \nu_e$  "RARE"

$\theta_{13}$  suppression ?

$\nu_\mu \rightarrow \nu_\tau$  "DIFFICULT" !

mass suppression, small  $c\tau$

Today's perspective

- Reactors: **no...  $\theta_{13}$  is BIG !**
- Appearance seen by **T2K** and **NOvA** (with few POT)

Confirmed difficult, but event-by-event detection achieved by **OPERA** & **SK** (with a much lower S/B)

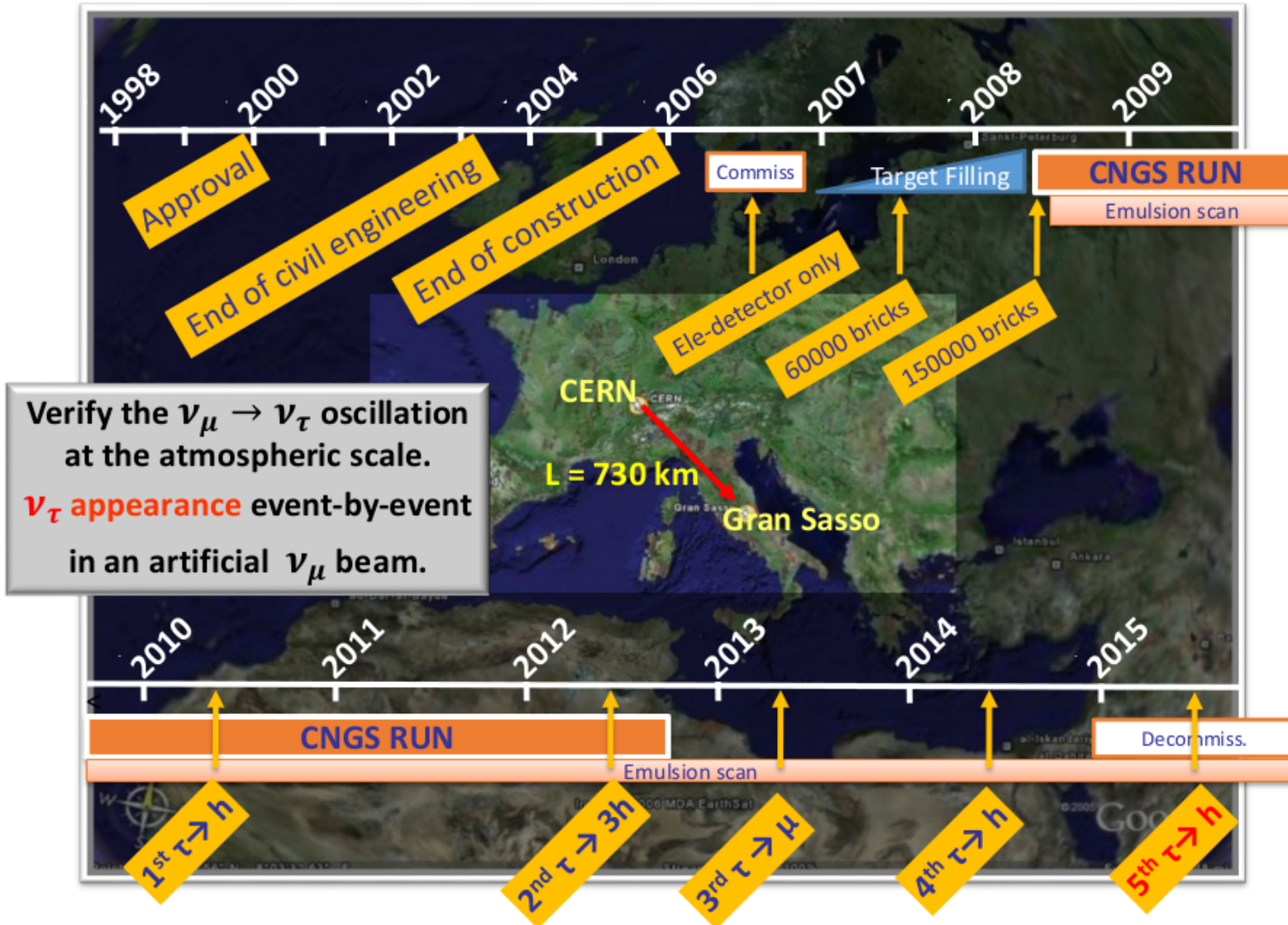


# The OPERA road map



An experimental and technological challenge:

- **Beam:** O(10) more energetic (17 GeV) than any other LBL ( $m_\tau$ ). 732 km baseline
- **Detector:** “fine-grained” O(100) more massive (1.25 kt) than SBL precursors (CHORUS)





# The CNGS beam for $\nu_{\mu} \rightarrow \nu_{\tau}$

$\langle E_{\nu} \rangle$	<b>17 GeV</b>
$L / \langle E_{\nu} \rangle$	<b>43 km/GeV</b>

The oscillation peak for  $L=732$  km at  $\sim 1.5$  GeV (similar to NuMI) but here the goal is to produce  $\tau$  leptons  
 $\rightarrow$  unbalance at higher energies

$$N(\tau) \sim \text{Pr}(\nu_{\mu} \rightarrow \nu_{\tau}) \times \sigma_{\nu(\tau)\text{CC}}(E) \times \text{flux}$$

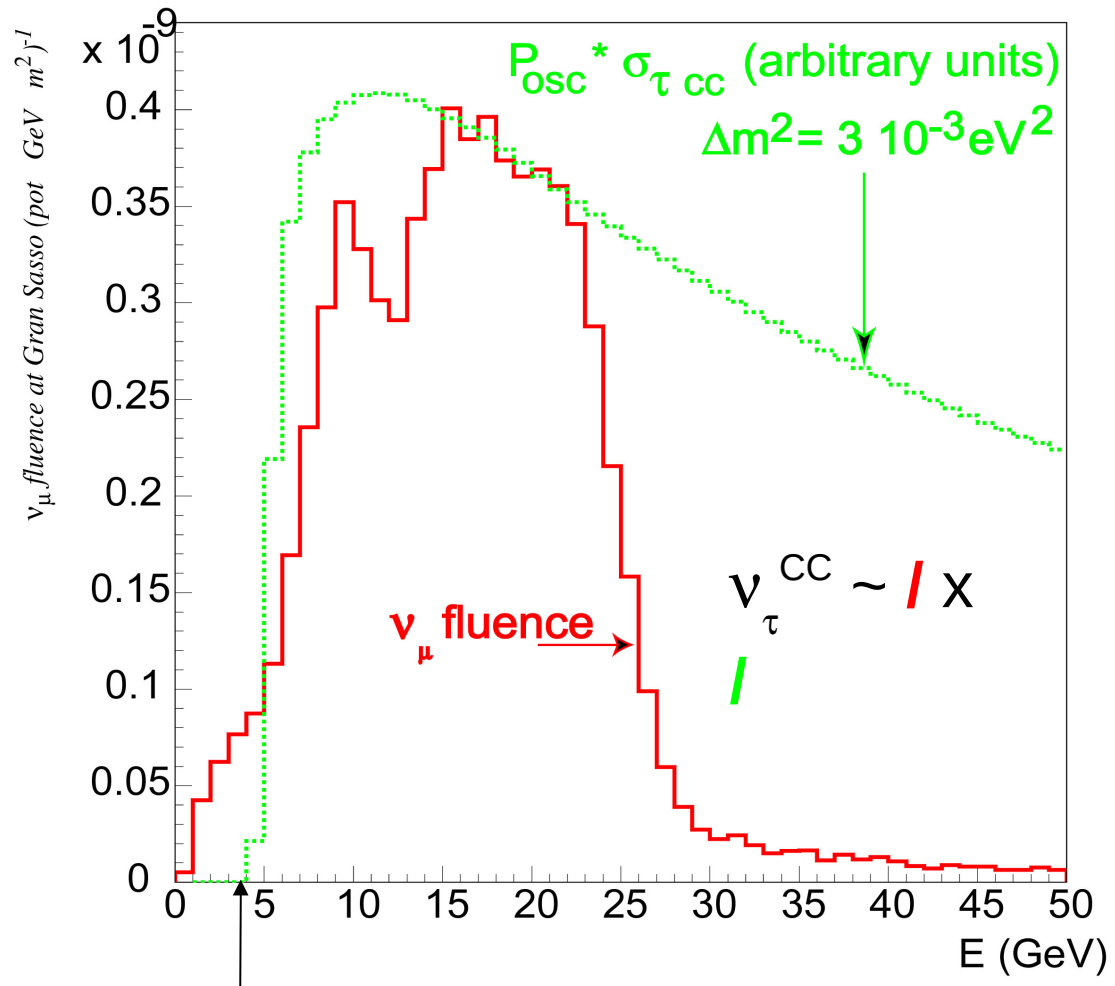
Fluxes:

$(\nu_e + \bar{\nu}_e) / \nu_{\mu}$	<b>0.9 %</b>
$\bar{\nu}_{\mu} / \nu_{\mu}$	<b>2.1 %</b>
$\nu_{\tau}$ prompt (from $D_s$ )	<b>negligible</b>

Interaction rates ( $1.8 \times 10^{20}$  pot):

$\sim 20k \nu_{\mu}$  CC+NC  
 **$66.4 \nu_{\tau}$  CC (not efficiency corrected)**

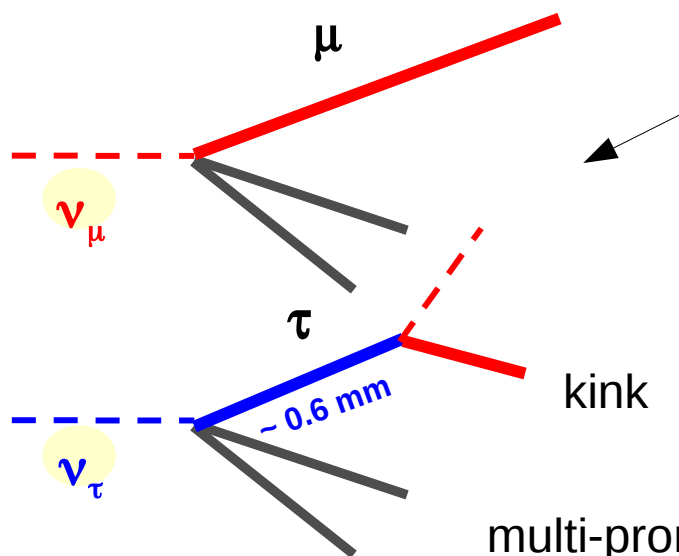
DESIGN:  $4.5 \cdot 10^{19}$  pot/year, 200 days/y per 5 y



**Threshold for  $\tau$  at  $\sim 3.5$  GeV.**  
**Slow rise.**



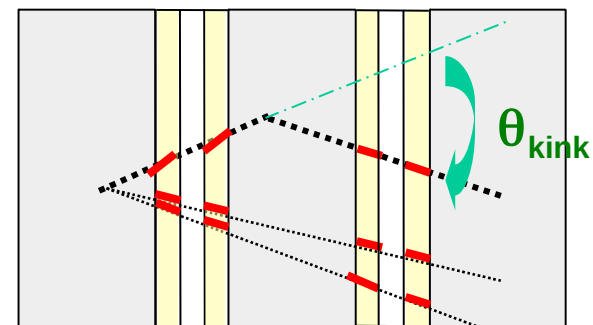
# The $\nu_\tau$ detection challenge



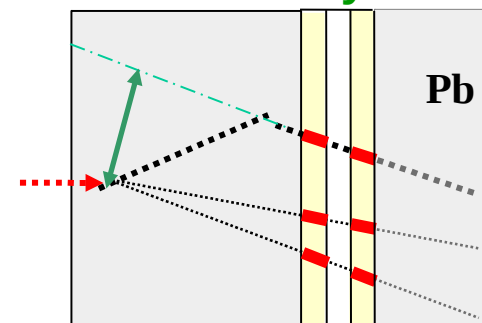
Detect a few  $\nu_\tau^{CC}$  from the bulk of  $\nu_\mu^{CC}$

$\tau^- \rightarrow \mu^- \nu_\tau \nu_\mu$	17 %
$\tau^- \rightarrow e^- \nu_\tau \nu_e$	18 %
$\tau^- \rightarrow h^- \nu_\tau n(\pi^0)$	50 %
$\tau^- \rightarrow \pi^+ \pi^- \pi^- \nu_\tau n(\pi^0)$	14 %

“long” decays: kink



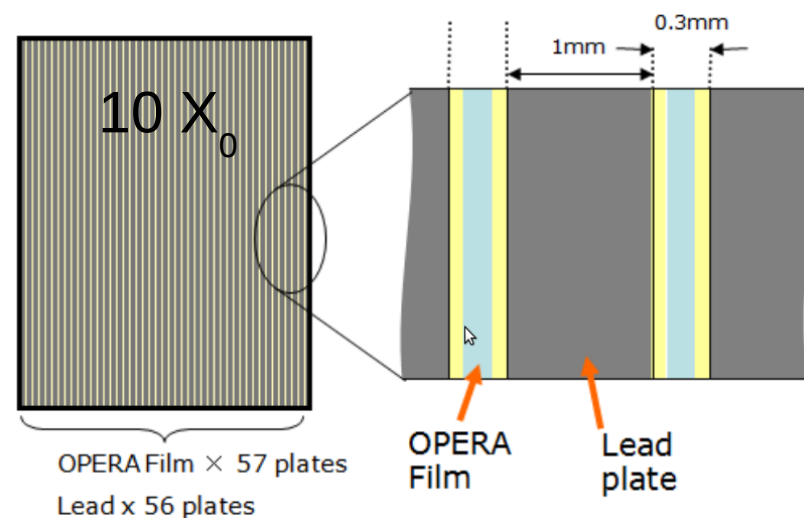
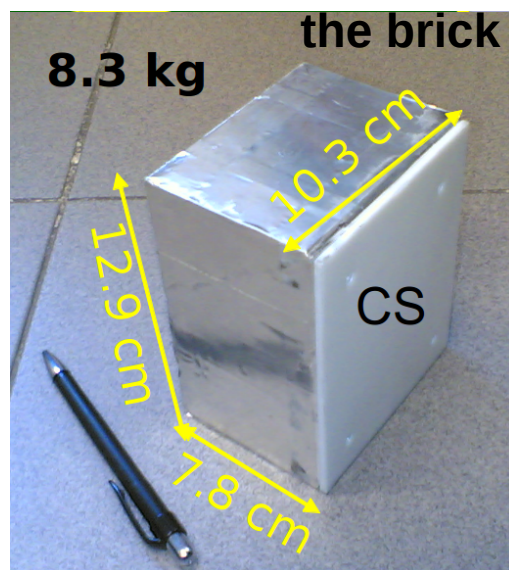
“short” decays: I.P.



Modular detector of “Emulsion Cloud Chambers” (or bricks)

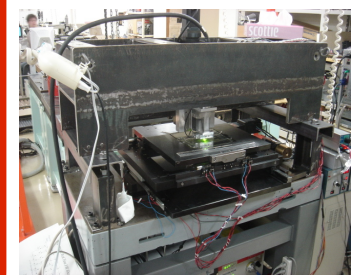
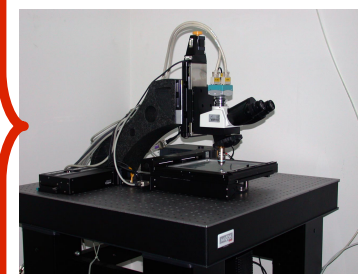
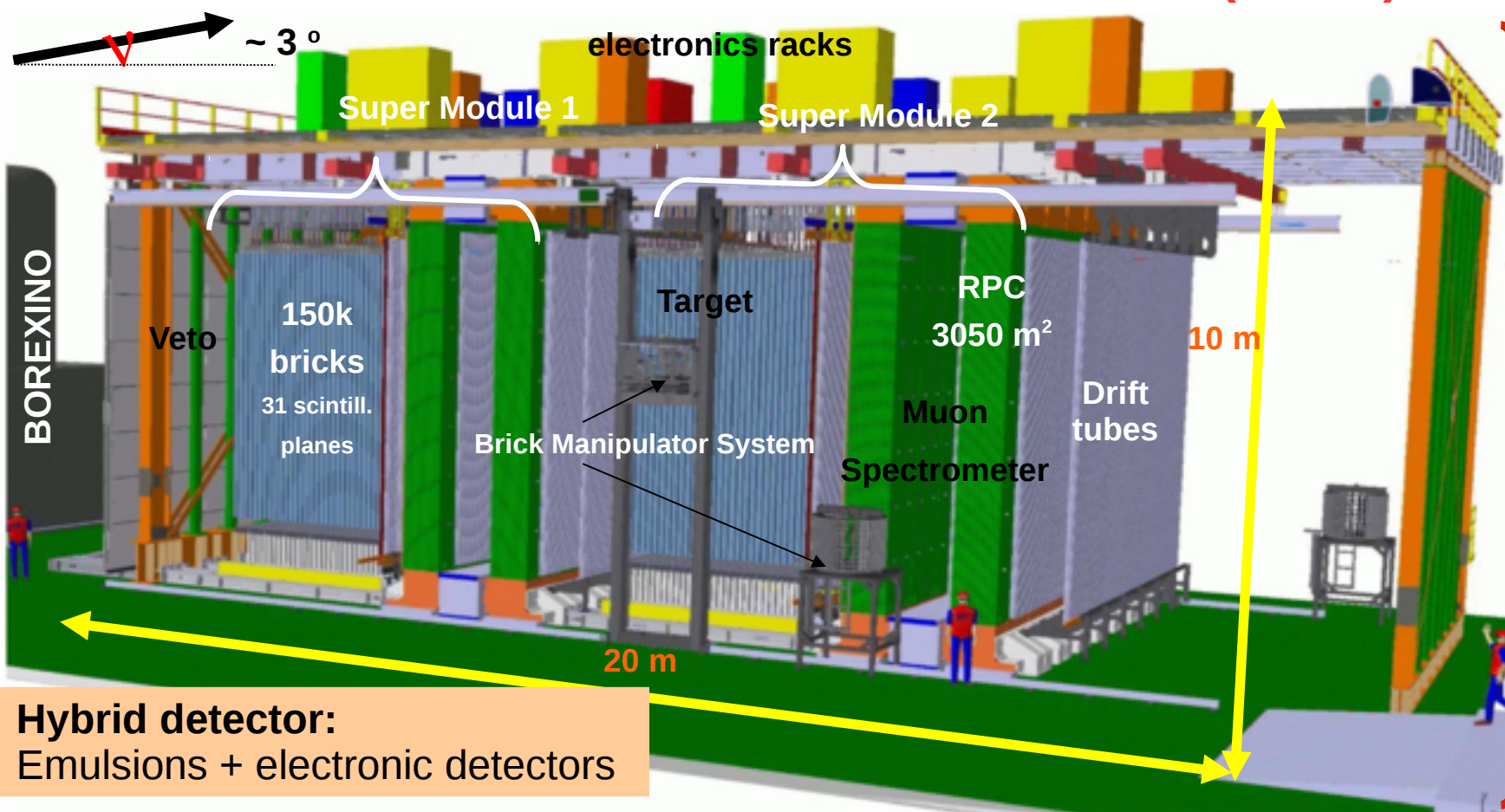
Reconciles the need for:

- **Large mass**
- $N_\tau \propto (\Delta m^2)^2 M_{\text{target}}$
- **Extreme granularity**
- $\sim \mu\text{m}$



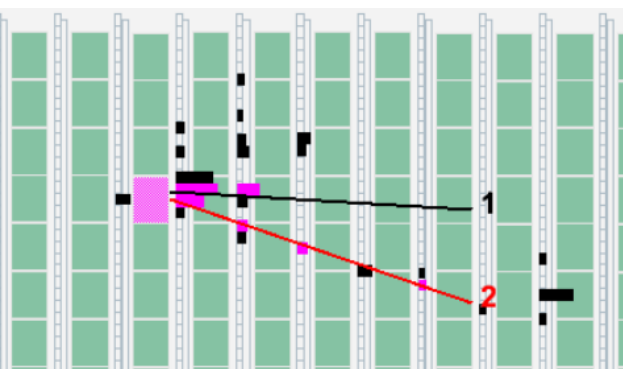
# The OPERA detector

$O(\mu\text{m})$  resolution  
over  $O(100\text{m}^3)$  !!!



Hybrid detector:  
Emulsions + electronic detectors

## “Brick-finding”



## + several ancillary facilities “off-site”:

- Emulsions “refreshing”
- Assembly/disassembly of bricks
- Labelling and X ray marking
- Automatised development
- Scanning of CS doublets
- Scanning of bricks

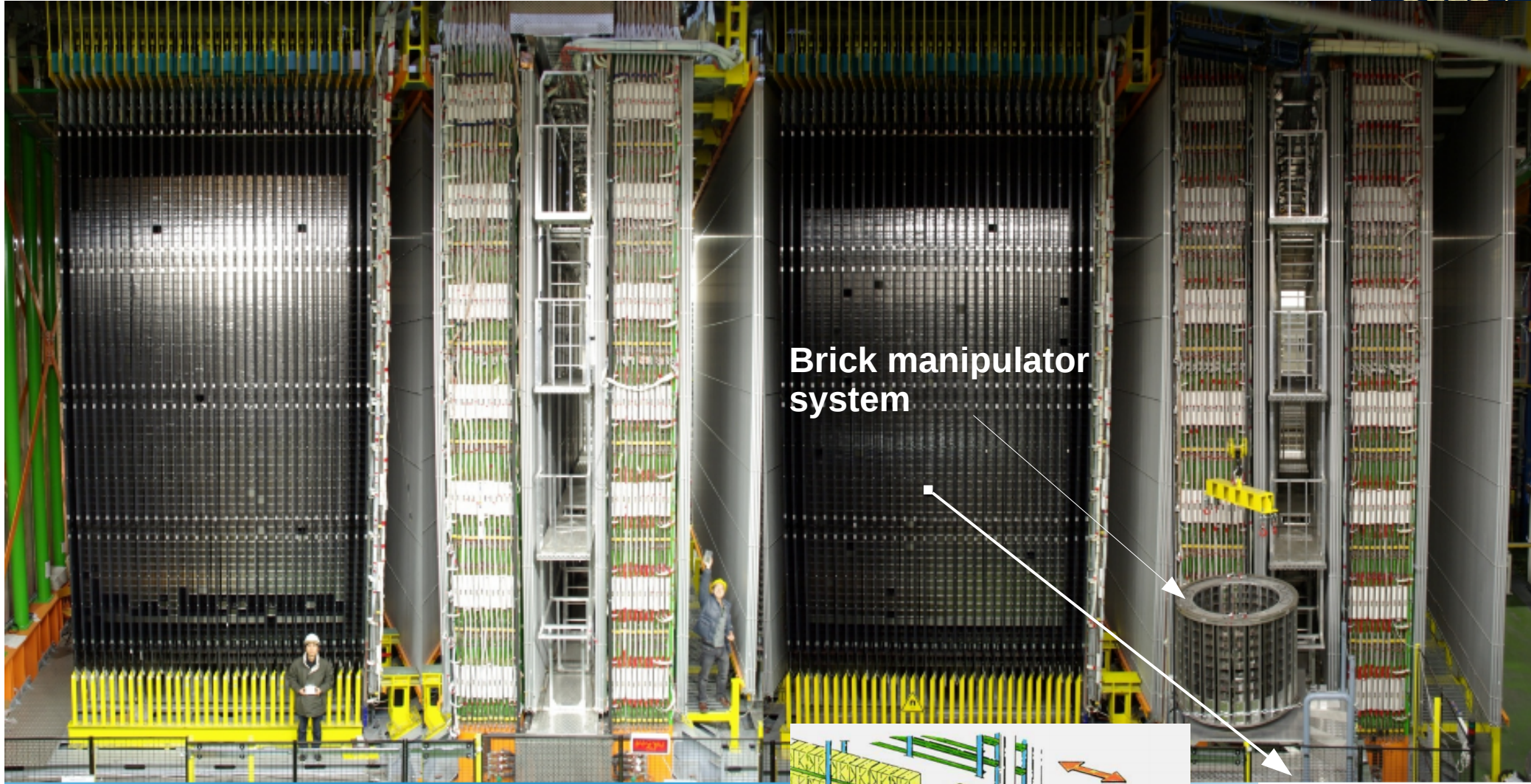
JP & Gran Sasso  
Gran Sasso  
Gran Sasso  
Gran Sasso  
Gran Sasso  
EU & JP



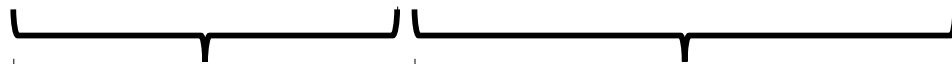


← Super Module 1 →

← Super Module 2 →



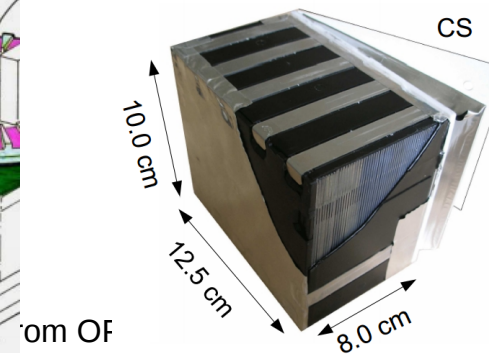
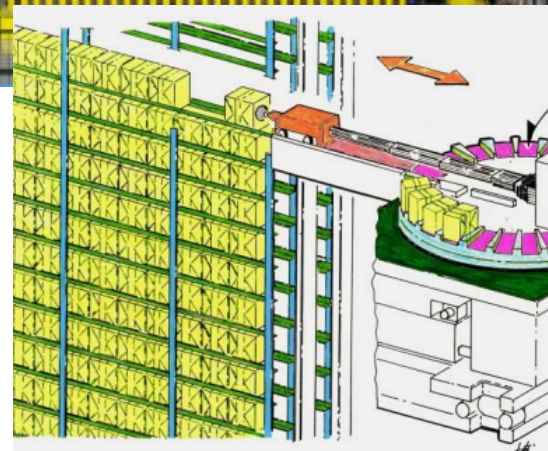
Brick manipulator system



Target area

μ spectrometer

~ 150.000 bricks in total.  
1.25 kt mass



10m OF



Super Module 1

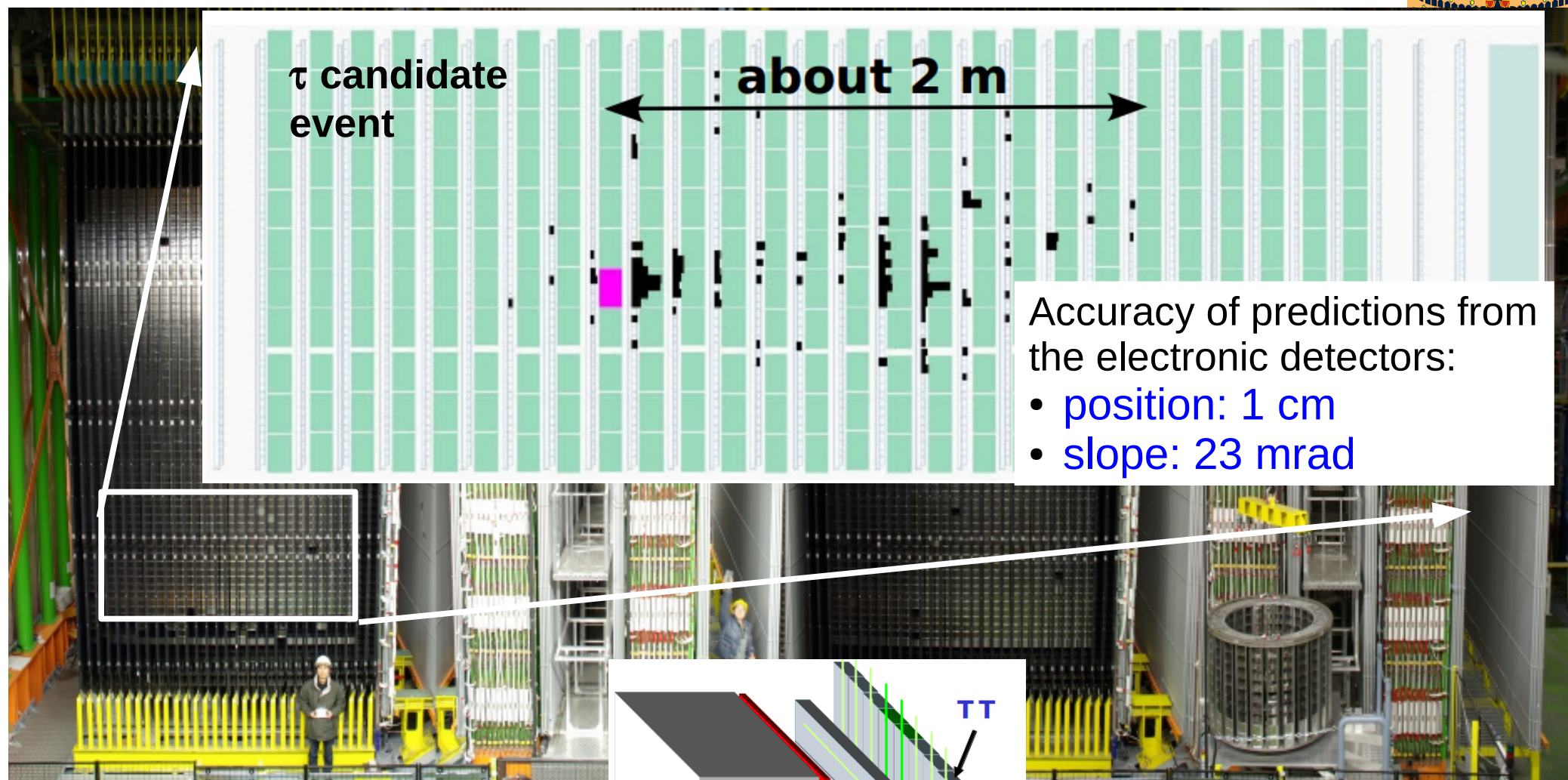
Super Module 2

$\tau$  candidate event

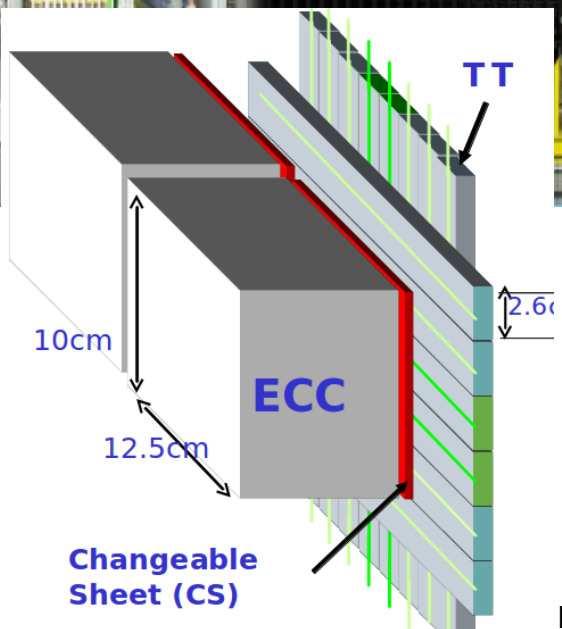
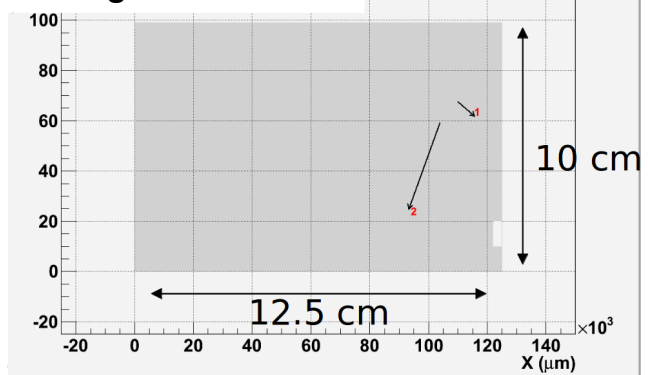
about 2 m

Accuracy of predictions from the electronic detectors:

- position: 1 cm
- slope: 23 mrad



CS: confirmation, triggering of the scanning in the brick:



Location efficiency:

- CC: 74 %
- NC: 48 %



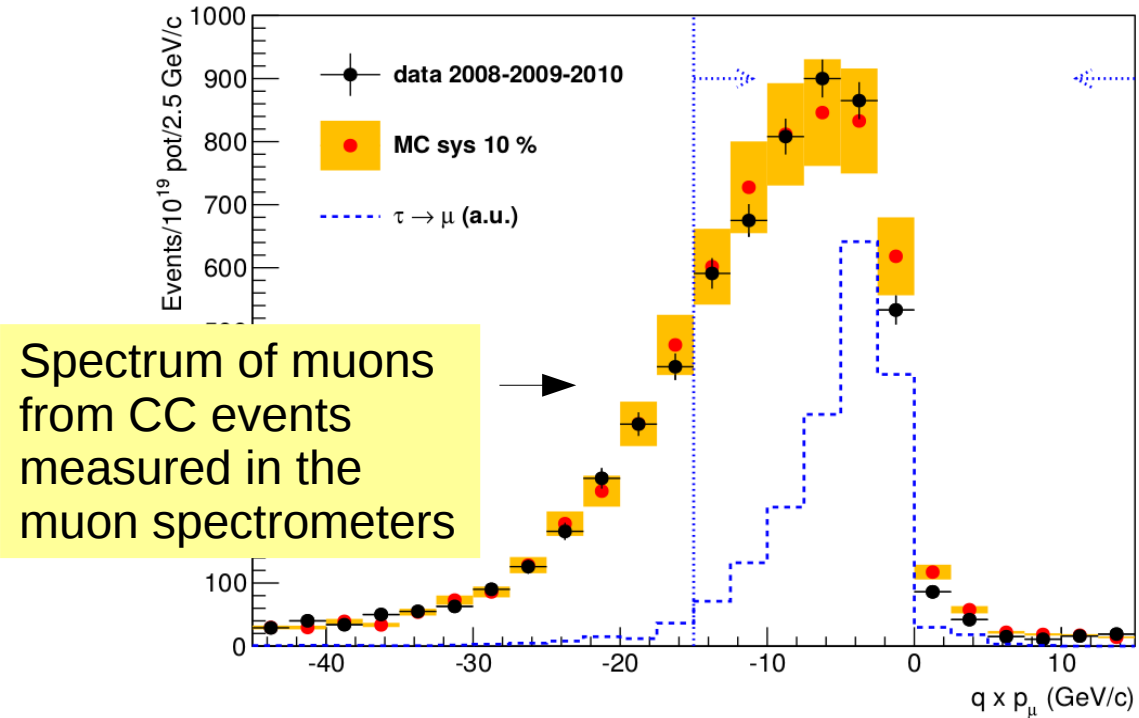
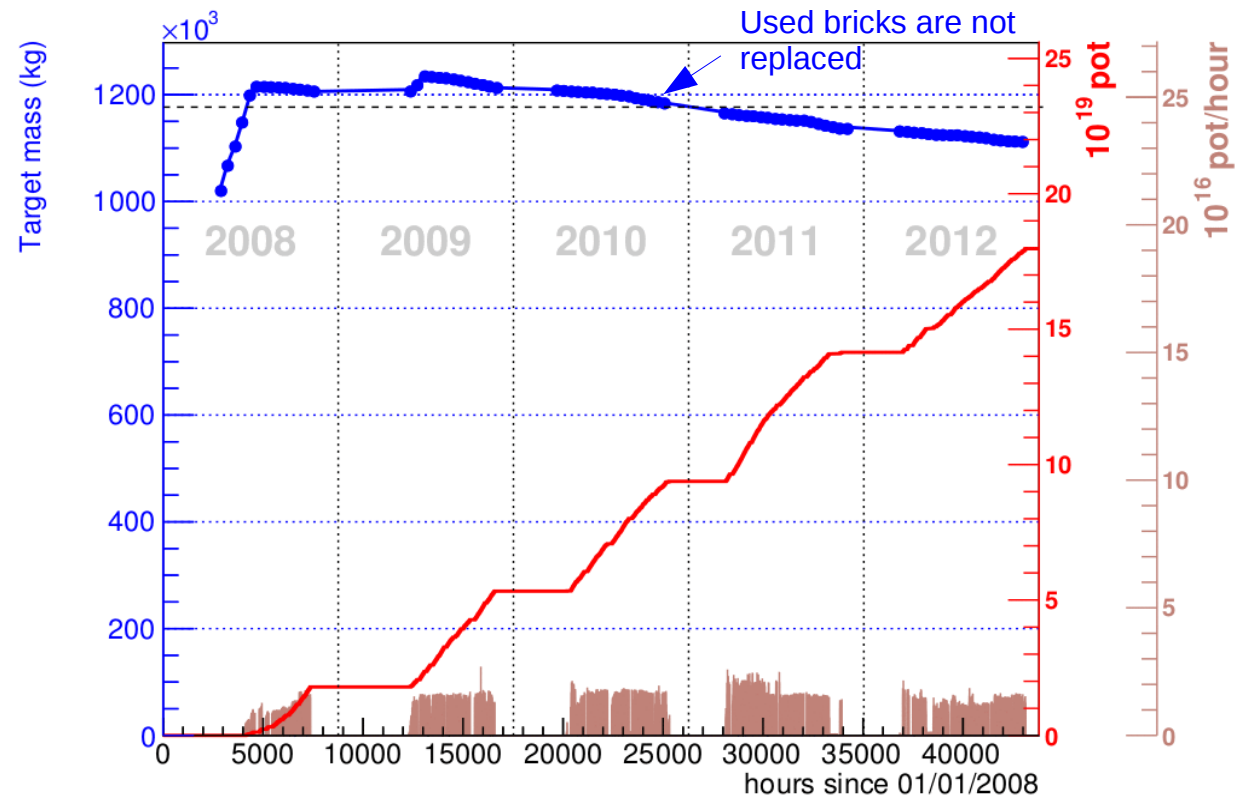
# Collected samples

The 5 year long CNGS run has ended in 2012.

$1.8 \times 10^{20}$  p.o.t. collected  
80% of the design ( $2.25 \times 10^{20}$ )

19505 neutrino interactions in the emulsion targets.

Year	Days	p.o.t. ( $10^{19}$ )	$\nu$ interactions
2008	123	1.74	1698
2009	155	3.53	3693
2010	187	4.09	4248
2011	243	4.75	5131
2012	257	3.86	3923
<b>tot</b>	<b>965</b>	<b>17.97</b>	<b>19505</b>

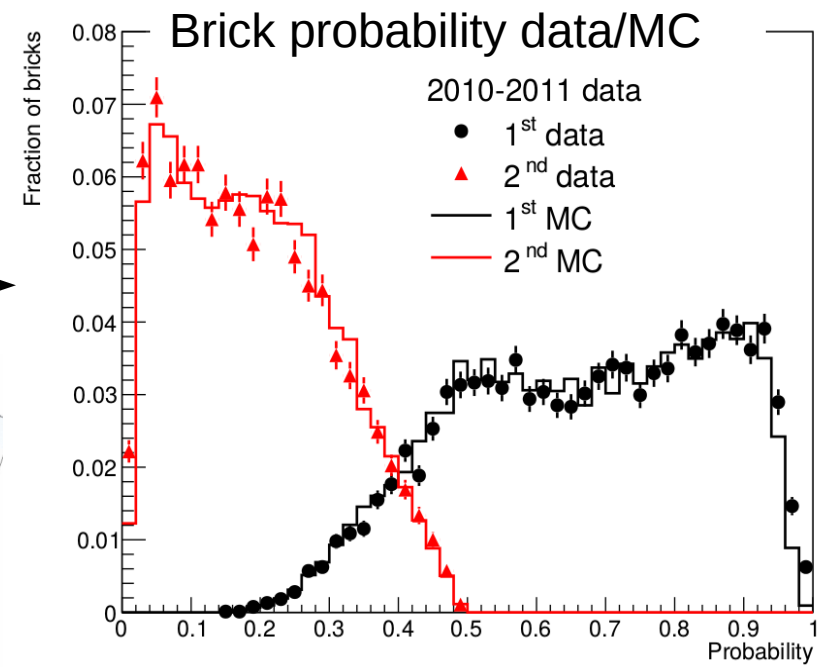
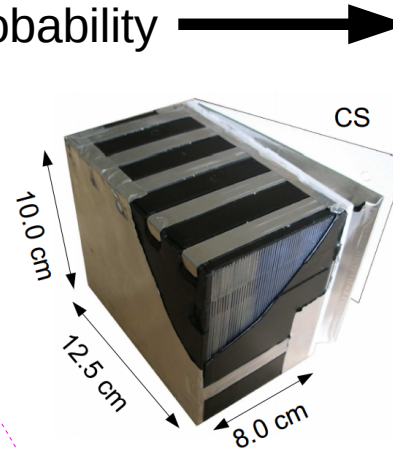


# Progress in the analysis of the emulsion films

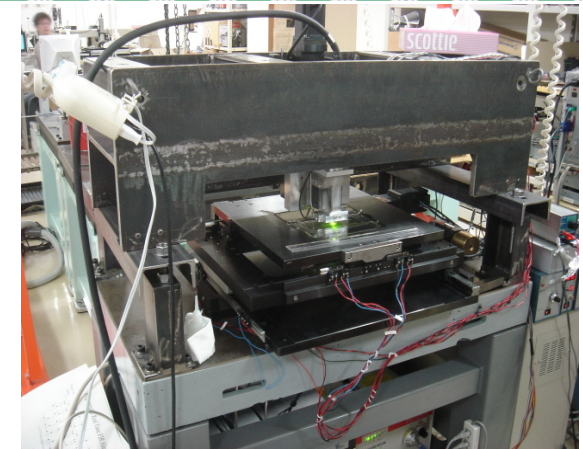
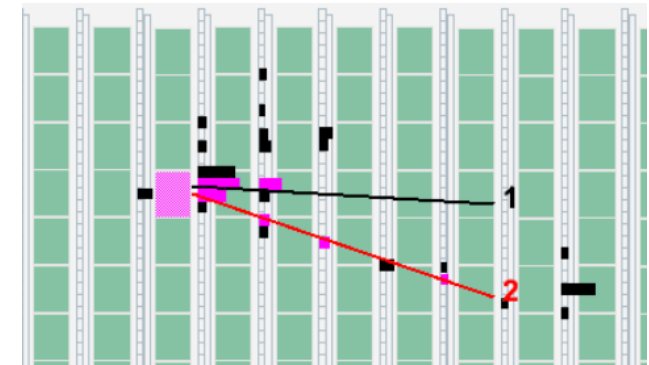
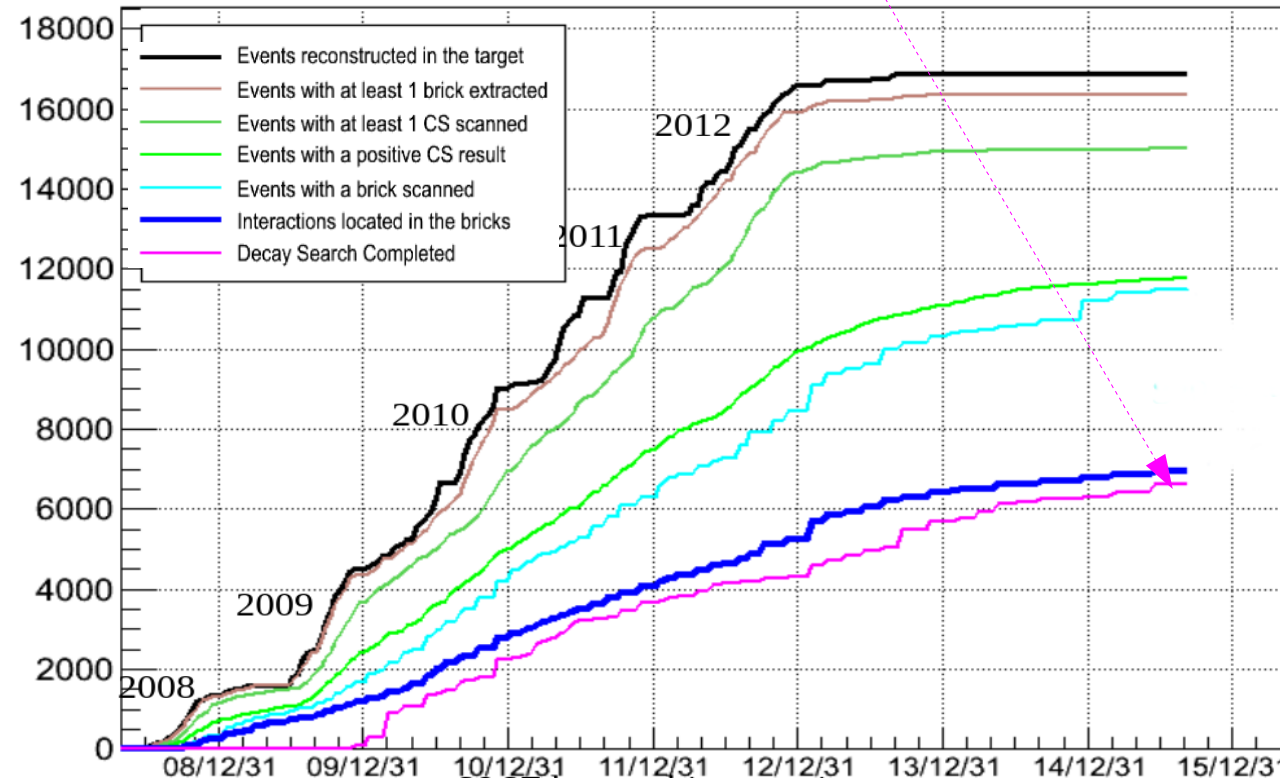
Bricks are ordered according to their probability of containing the interaction vertex

Analysis status:

- 2008-2009 completed up to 4<sup>th</sup> brick
- 2010-2012 completed up to 2<sup>nd</sup> bricks

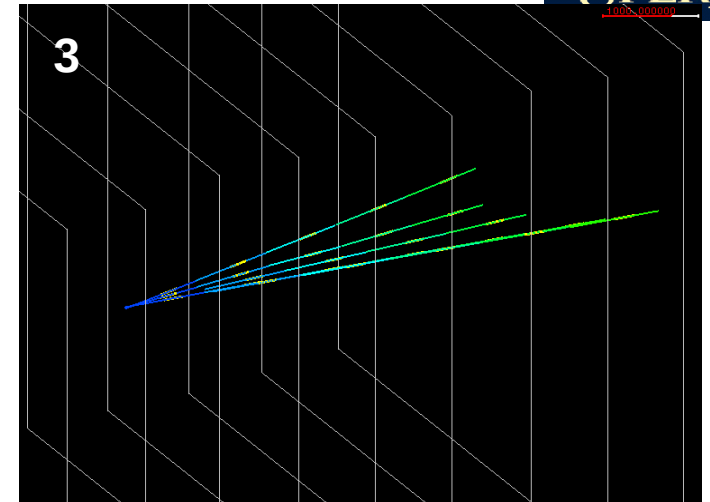
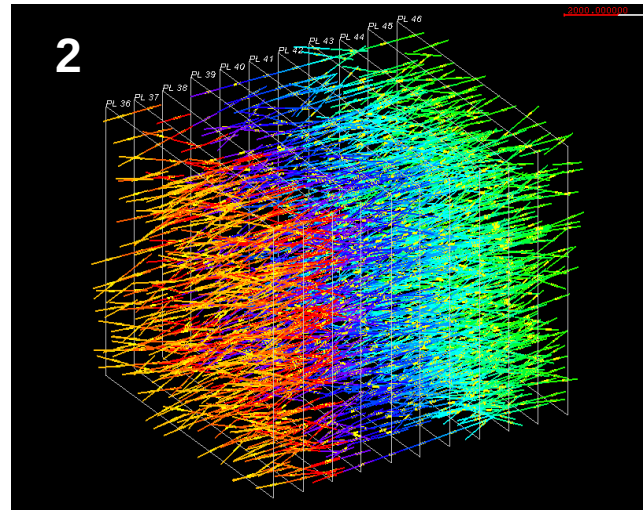
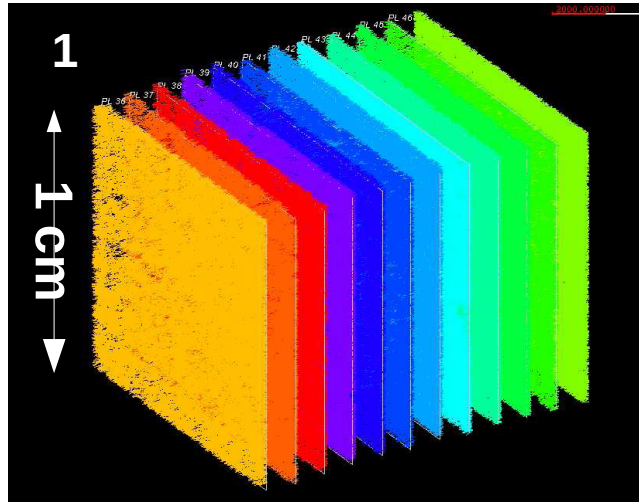


> 6600 fully analysed bricks

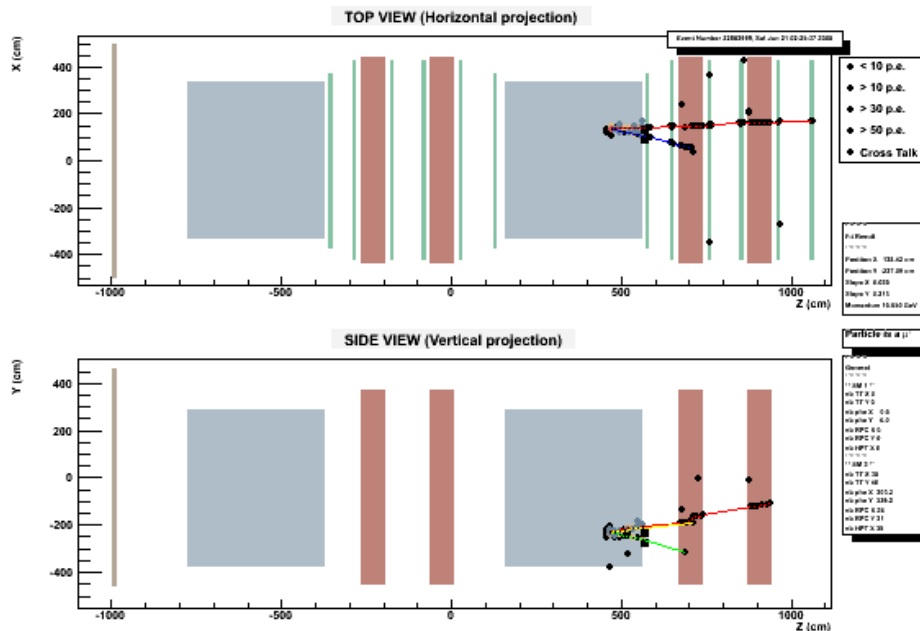




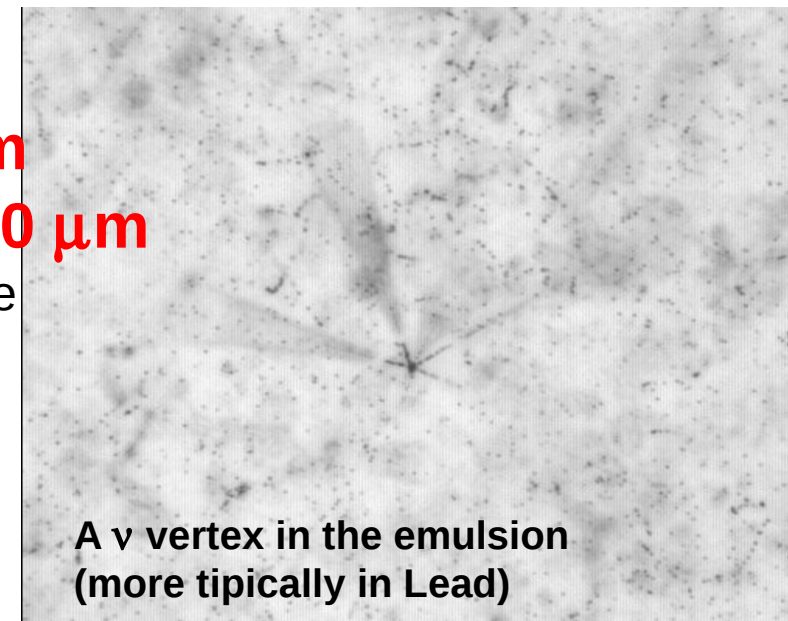
# Vertex hunting in the brick



- 0) tracks tagged in the **CS films** are **followed upstream** until a **stopping point** is found
- 1) base-tracks in the 12 films of the a **volume centered in the stopping point** are reconstructed
- 2) cosmic ray tracks (from a dedicated exposure) are used for the fine **alignment** of films
- 3) passing-through tracks discarded → **vertexing algorithm**



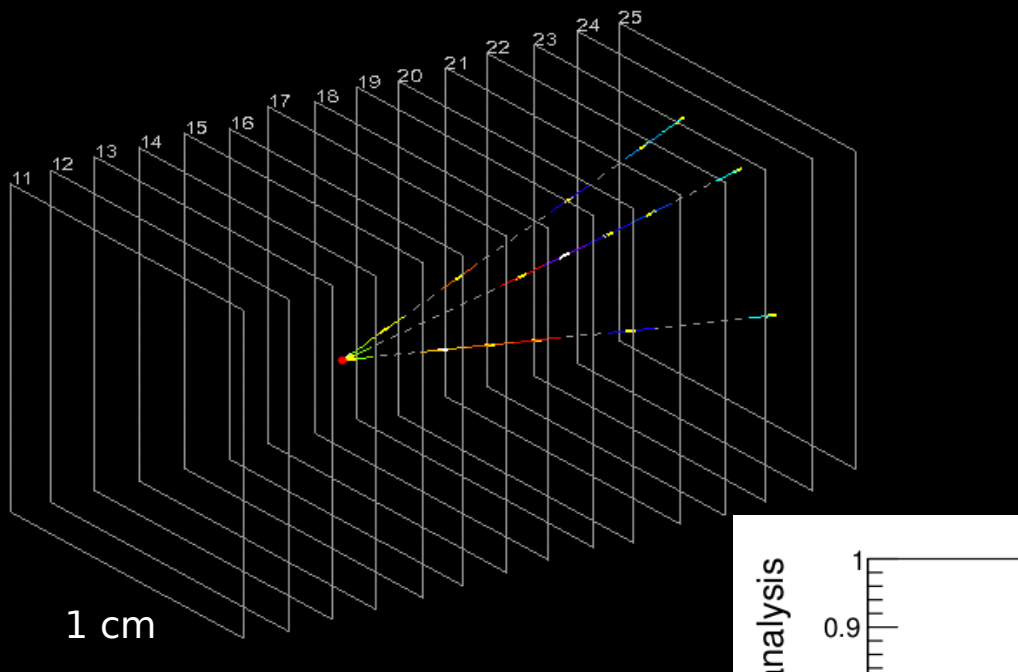
**Leap from  
20 m → 100  $\mu\text{m}$**   
(essential role  
of CS films)



**A v vertex in the emulsion  
(more typically in Lead)**

# Location efficiency

JHEP 11 (2013) 036

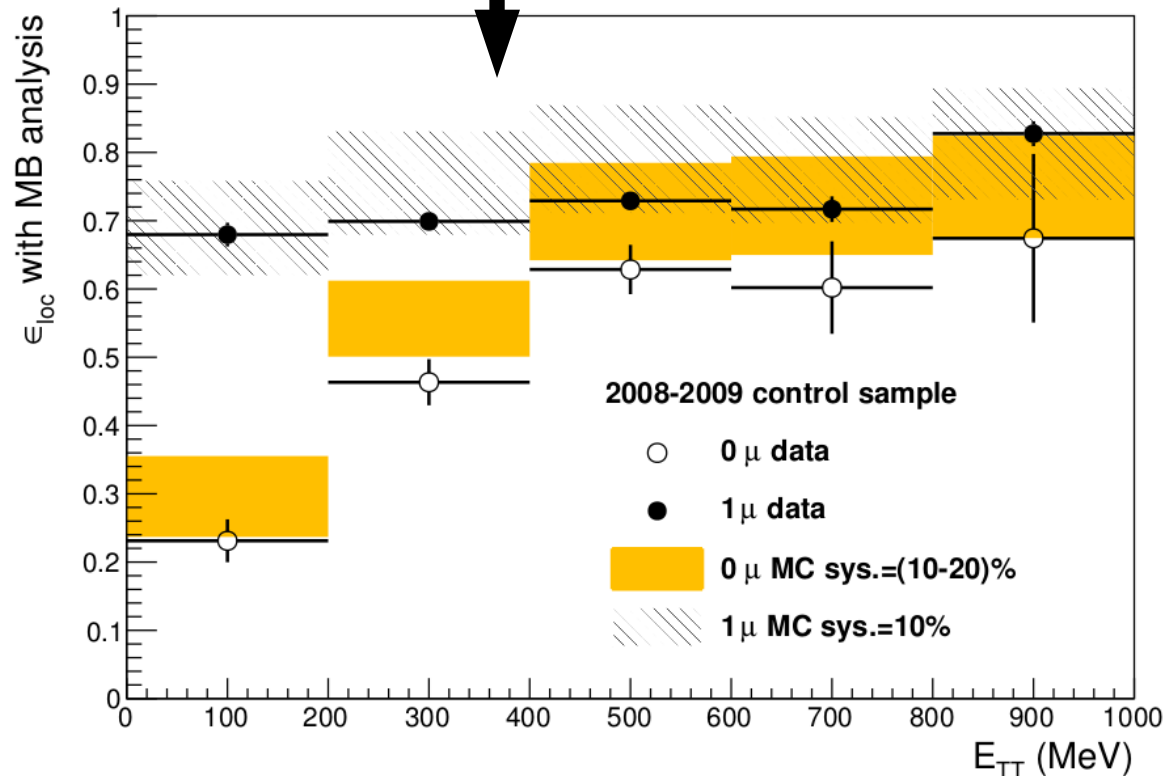


$0\mu$  and  $1\mu$  samples

Data-Monte Carlo comparison of the location efficiency as a function of the visible energy in the target scintillators

Hybrid detector:  
a complex simulation!  
Reasonable agreement.

The prediction for the  $\tau$  signal and backgrounds relies on the efficiencies as measured in real data





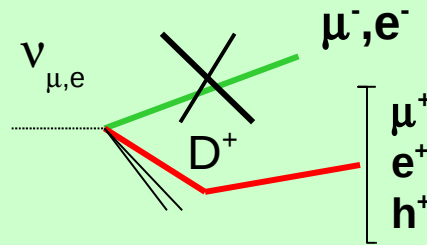
# $\nu_{\mu} \rightarrow \nu_{\tau}$ background characterisation

Monte Carlo simulation benchmarked on control samples.

In order of decreasing relevance

## CC with charm production

(all channels) IF the primary lepton is not identified and the daughter charge is not (or incorrectly) measured



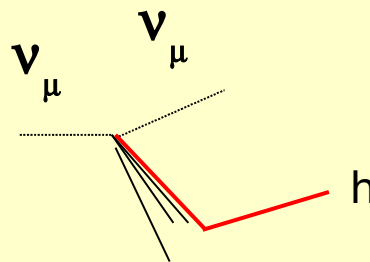
MC tuned on CHORUS data (cross section and fragmentation functions), validated with measured OPERA charm events.

*Eur. Phys.J. C74 (2014) 2986*

Reduced by "track follow down", procedure and large angle scanning

## Hadronic interactions

Background for  $\tau \rightarrow h$



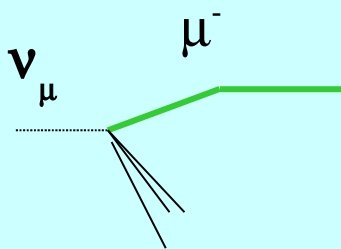
FLUKA + pion test beam data

Reduced by large angle scanning and nuclear fragment search

*PTEP9 (2014) 093C01*

## Large angle muon scattering

Background for  $\tau \rightarrow \mu$



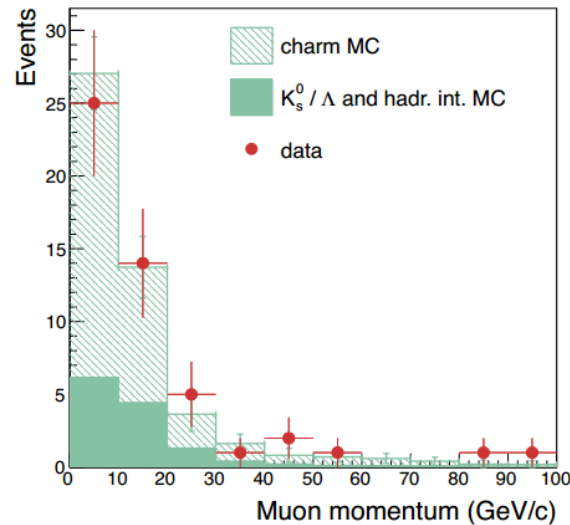
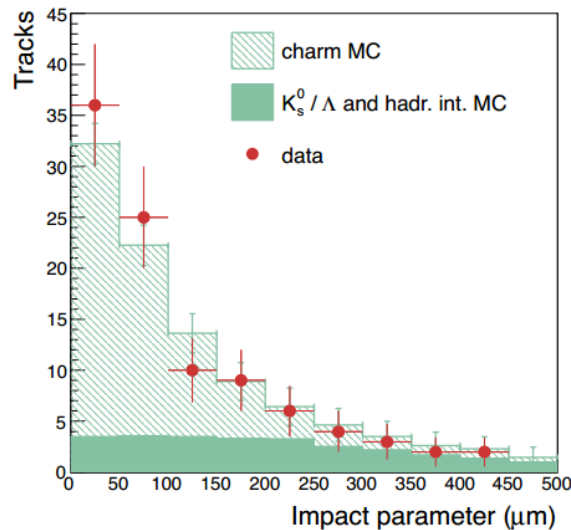
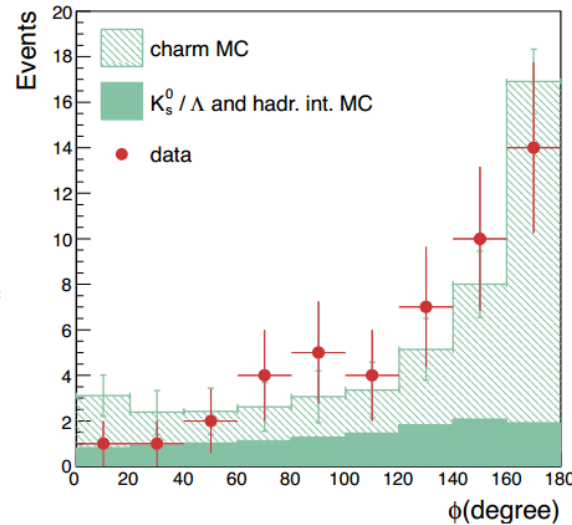
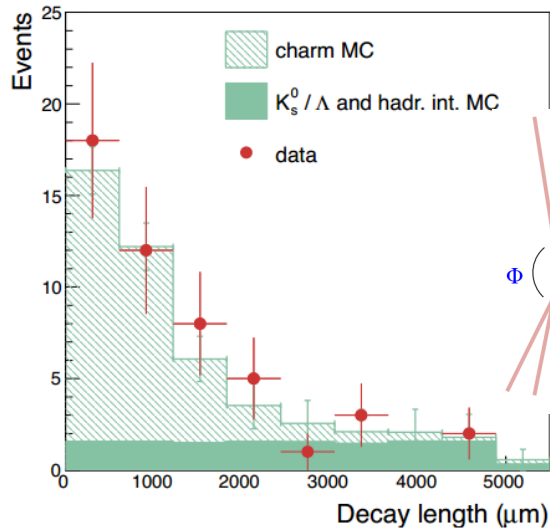
Improved knowledge bringing this contribution to a **negligible level**: GEANT4 simulation benchmarked on real data from the literature.

*IEEE Trans. Nucl. Sci. 62,5, 2216-2225*

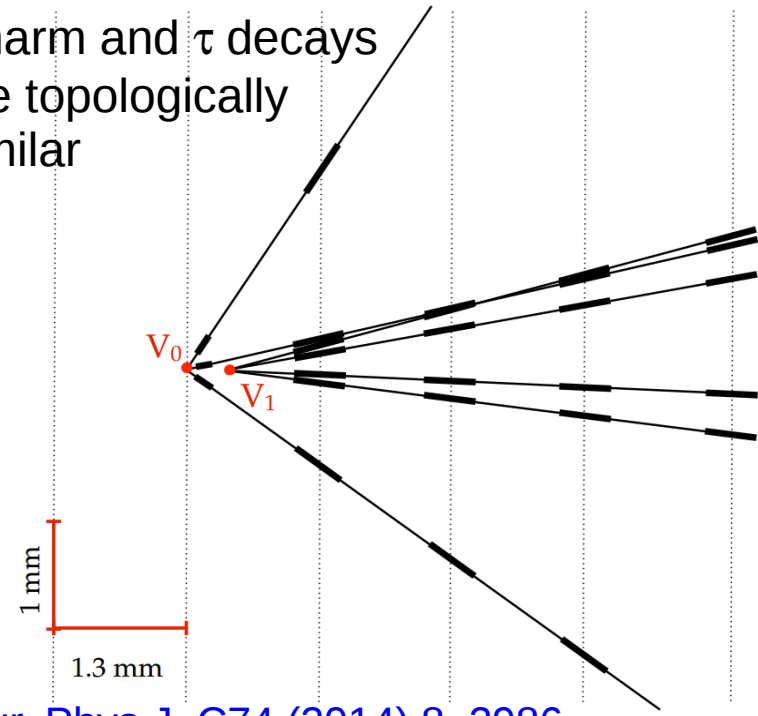
# Validation with the CNGS charm events sample

Test for: reconstruction efficiencies, description of kinematical variables, charm background.

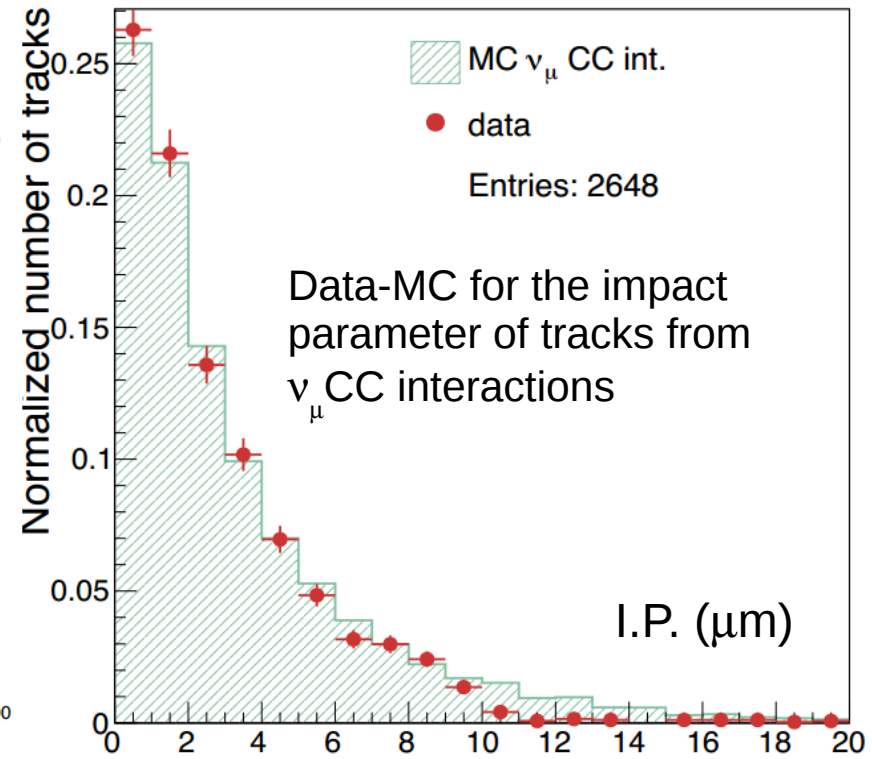
**54 ± 4 expected ↔ 50 observed**



Charm and  $\tau$  decays are topologically similar



Eur. Phys.J. C74 (2014) 8, 2986



# OPERA $\nu_\tau$ appearance



Channel	Total background	Expected signal	Observed
$\tau \rightarrow 1h$	$0.04 \pm 0.01$	$0.52 \pm 0.10$	3
$\tau \rightarrow 3h$	$0.17 \pm 0.03$	$0.73 \pm 0.14$	1
$\tau \rightarrow \mu$	$0.004 \pm 0.001$	$0.61 \pm 0.12$	1
$\tau \rightarrow e$	$0.03 \pm 0.01$	$0.78 \pm 0.16$	0
Total	$0.25 \pm 0.05$	$2.64 \pm 0.53$	5

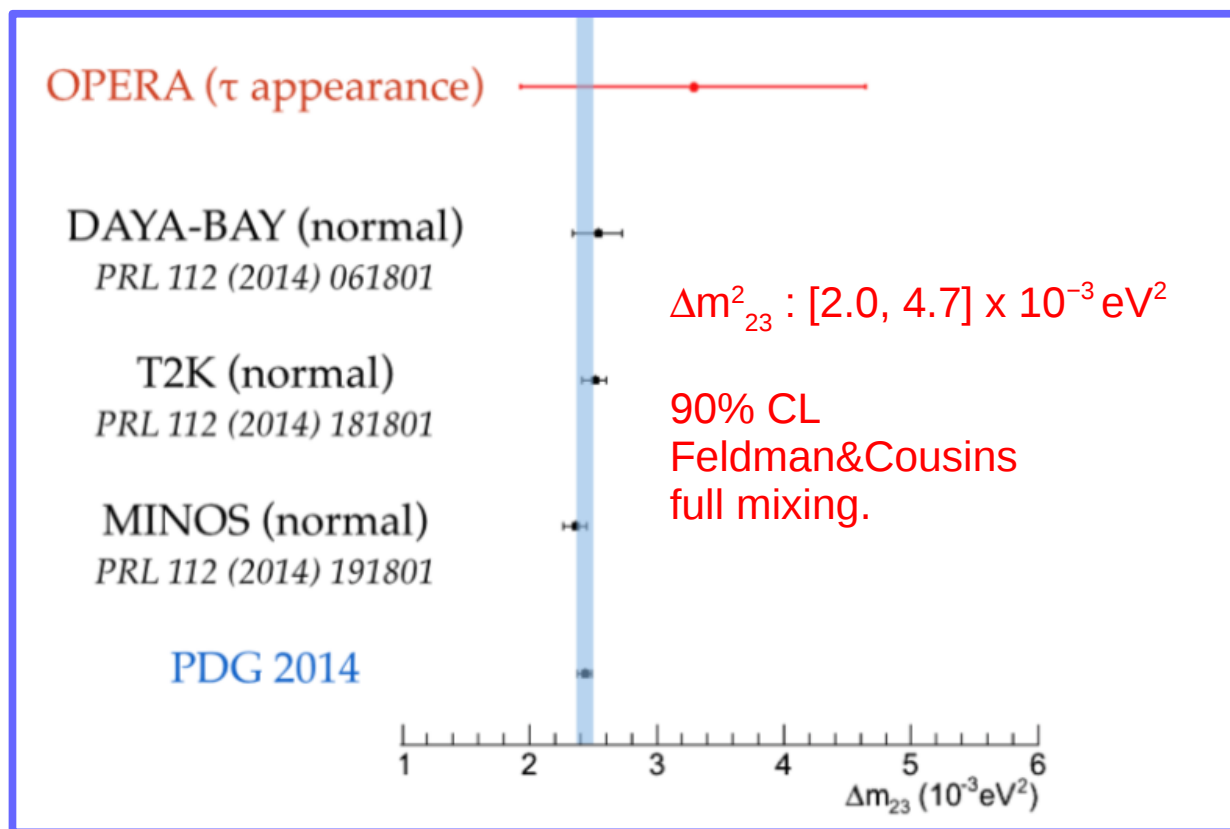
Signal/background  $\sim 10$

$$N_\tau \propto (\Delta m_{32}^2)^2 L^2 \int \phi(E) \epsilon(E) \frac{\sigma(E)}{E^2} dE$$

5 candidates (1 leptonic+4 hadronic) fulfilling the kinematic selection defined in the experiment proposal

→  
5.1  $\sigma$  exclusion of the background-only hypothesis

[PRL 115 (2015) 12, 121802]







# $\nu_\mu \rightarrow \nu_\tau$ : sterile?

How is the appearance probability is modified by one possible extra (sterile) state (3+1 scheme) ?

$$\begin{bmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{bmatrix}$$

$$P(\nu_\mu \rightarrow \nu_\tau) = 4 |U_{\mu3}|^2 |U_{\tau3}|^2 \sin^2 \frac{\Delta_{31}}{2} \quad \text{standard}$$

$$+ 4 |U_{\mu4}|^2 |U_{\tau4}|^2 \sin^2 \frac{\Delta_{41}}{2} \quad \text{exotic}$$

$$\begin{aligned} &+ 2 \Re [U_{\mu4}^* U_{\tau4} U_{\mu3} U_{\tau3}^*] \sin \Delta_{31} \sin \Delta_{41} \\ &- 4 \Im [U_{\mu4}^* U_{\tau4} U_{\mu3} U_{\tau3}^*] \sin^2 \frac{\Delta_{31}}{2} \sin \Delta_{41} \\ &+ 8 \Re [U_{\mu4}^* U_{\tau4} U_{\mu3} U_{\tau3}^*] \sin^2 \frac{\Delta_{31}}{2} \sin \frac{\Delta_{41}}{2} \\ &+ 4 \Im [U_{\mu4}^* U_{\tau4} U_{\mu3} U_{\tau3}^*] \sin \Delta_{31} \sin \frac{\Delta_{41}}{2} \end{aligned}$$

[JHEP 06 (2015) 069]

interference

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{2 E}$$

A rich structure:

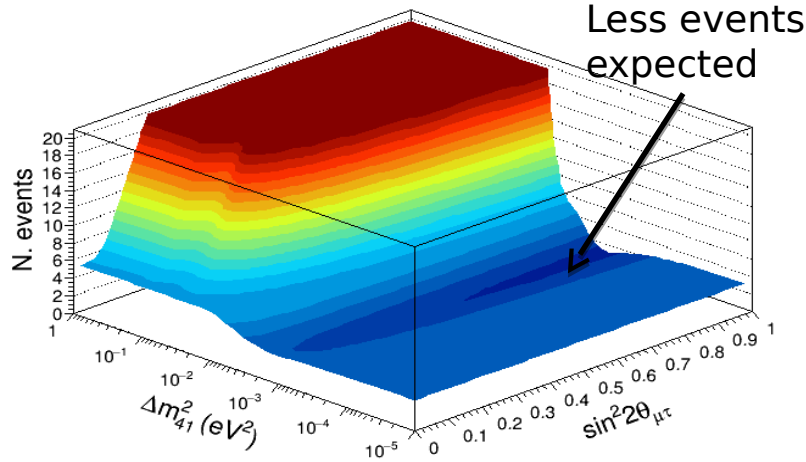
- can result in an **increase or decrease** of expected number of  $\nu_\tau$  events
- **sensitive to new CPV phases** + mass hierarchy (unlike short-baseline exp.)

# $\nu_\mu \rightarrow \nu_\tau$ : effect of a sterile ?

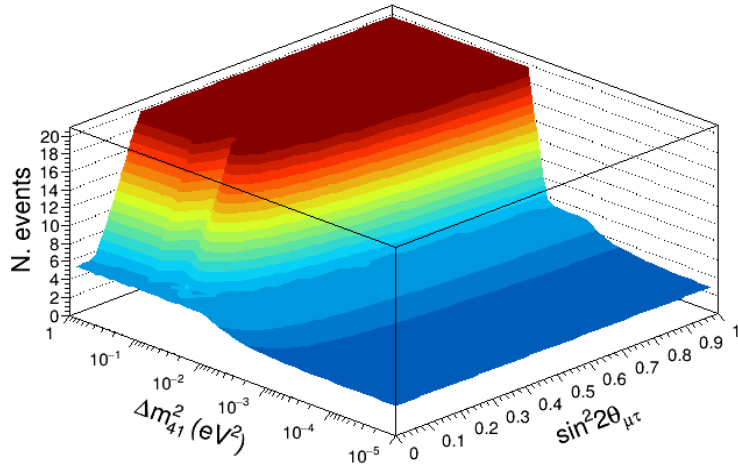
$\nu_\tau$  rate only analysis

Expected  $\nu_\tau$  events

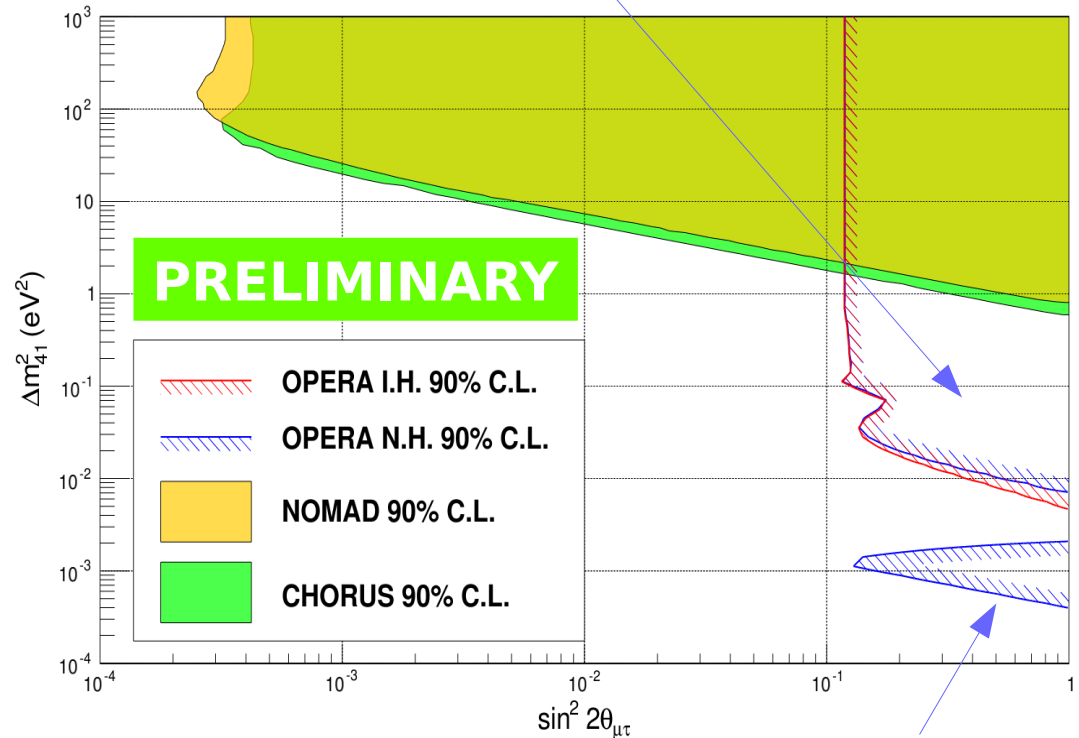
Normal hierarchy



Inverted hierarchy



Long baseline, high-E:  
a wide range of  $\Delta m_{41}^2$  from  $10^{-3}$  to 1 excluded  
for  $\sin^2 2\theta_{\mu\tau} > O(10^{-1})$



$$\sin 2\theta_{\mu\tau} = 2|U_{\mu 4}||U_{\tau 4}|$$

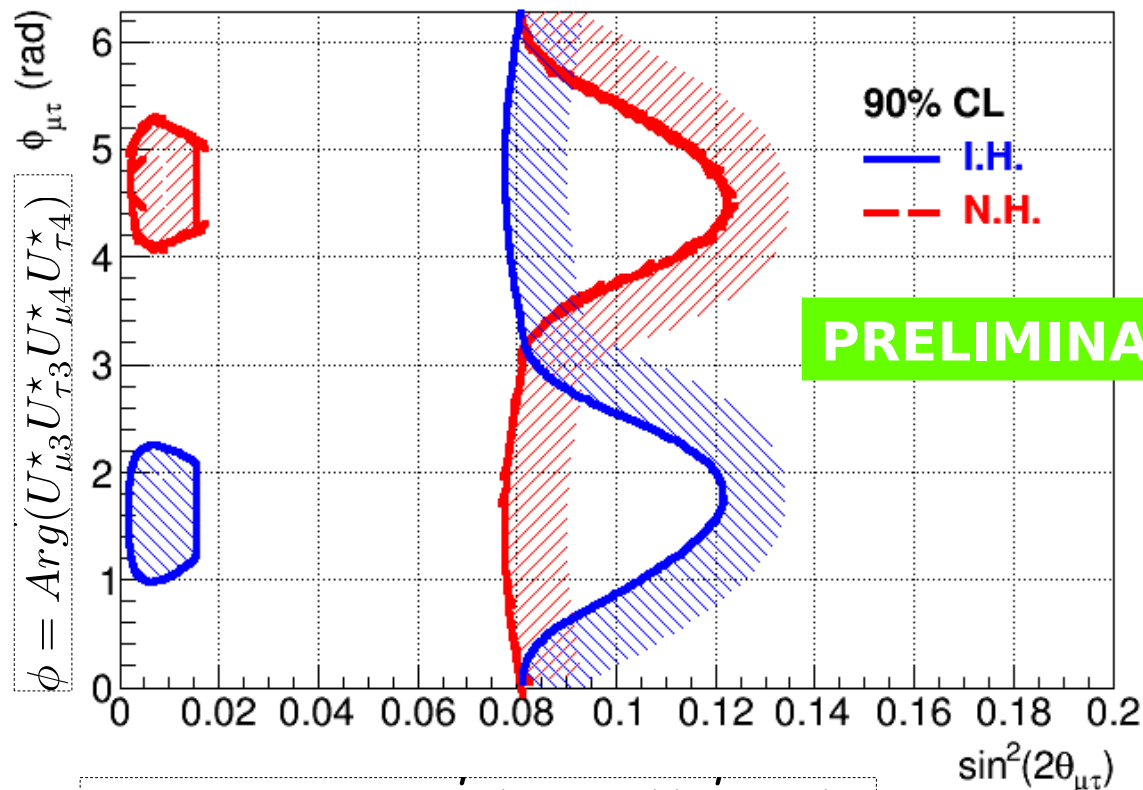
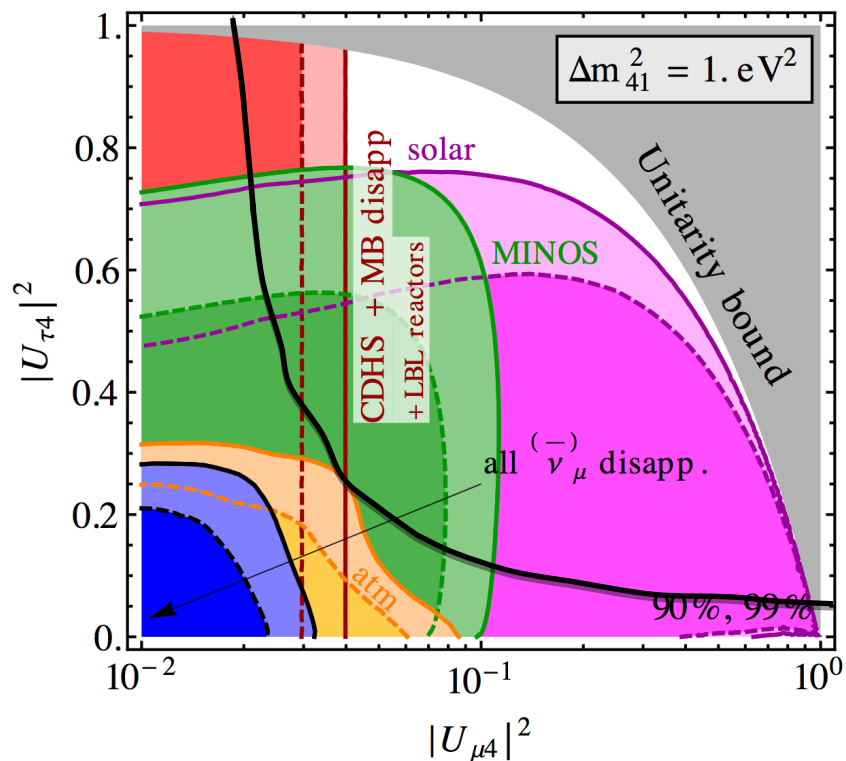
With normal hierarchy a sterile neutrino close enough to  $m_3$  would suppress  $\nu_\tau$  appearance well below the 5 observed events

# $\nu_\mu \rightarrow \nu_\tau$ : effect of a sterile ?

- For  $|\Delta m_{41}^2| > 1 \text{ eV}^2$
- Dependence on the new CP violating new phase  $\phi$ 
  - Peculiar of long-baseline

$$\phi = \text{Arg}(U_{\mu 3}^* U_{\tau 3}^* U_{\mu 4}^* U_{\tau 4}^*)$$

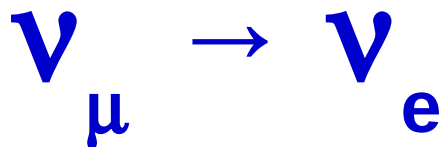
## OPERA



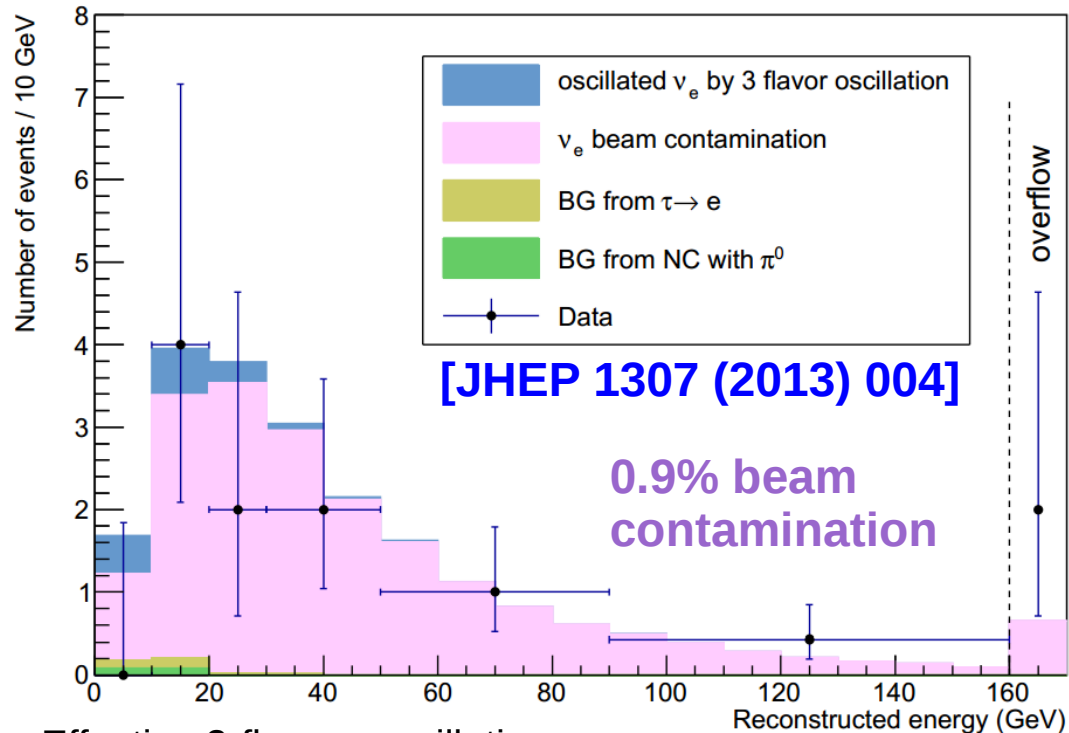
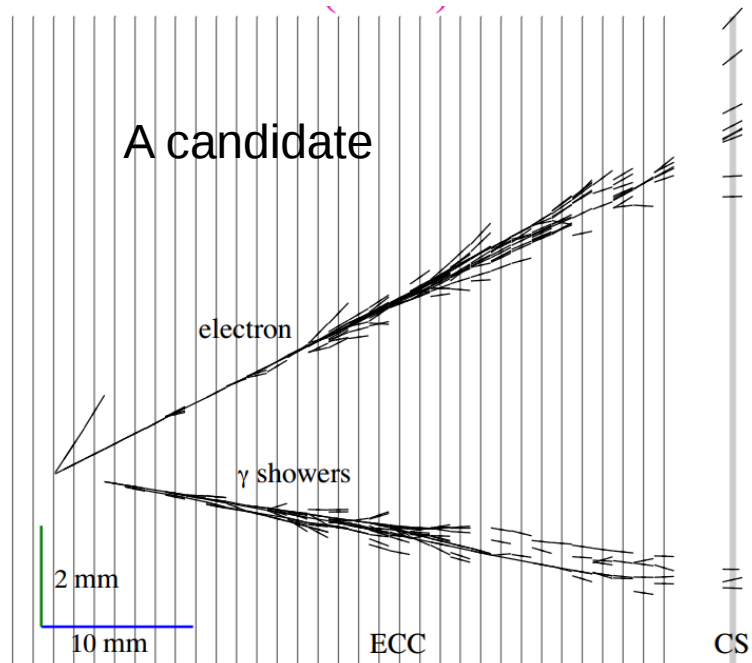
$$\sin 2\theta_{\mu\tau} = 2|U_{\mu 4}||U_{\tau 4}|$$

- $\sin^2 2\theta_{\mu\tau} < 0.119$  at 90% CL integrating over  $\phi$
- in the  $|U_{\tau 4}|^2$  vs  $|U_{\mu 4}|^2$  plane: complementarity with disappearance experiments (MINOS) is visible

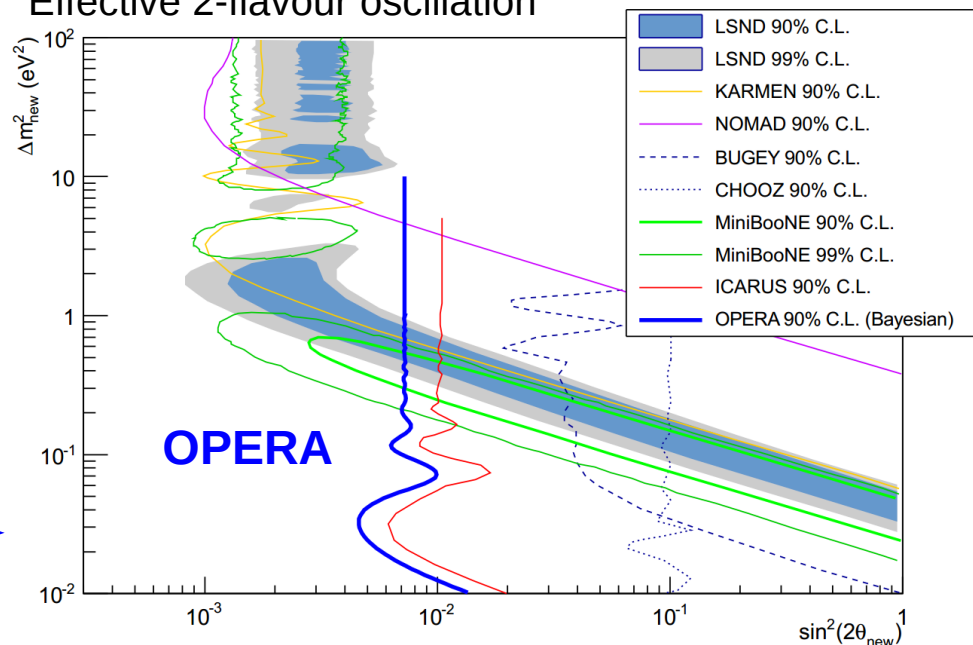




505  $0_{\mu}$  interactions (2008-09 runs)



Effective 2-flavour oscillation



		$E < 20$ GeV
$\nu_e$ candidates	19	4
Background	$19.8 \pm 2.8$ (sys.)	4.6

$$\sin^2 2\theta_{\text{NEW}} < 7.2 \times 10^{-3} \text{ (90\% CL)} \rightarrow$$

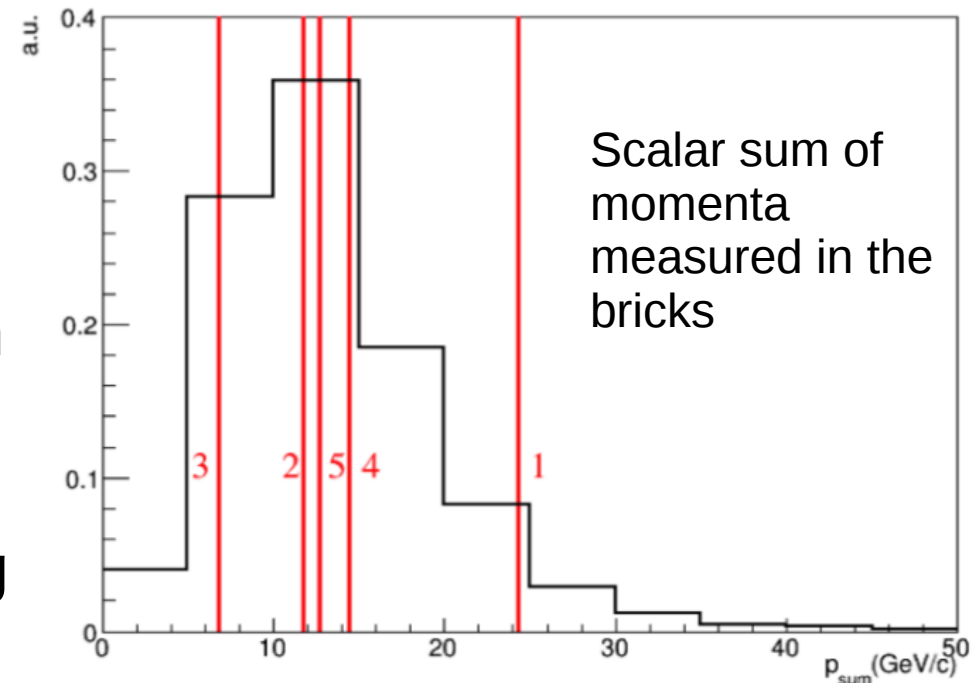
Final analysis with the full data sample close to completion ( $\times 2.5$  increase in sample)  
 Will include rigorous 3+1 treatment – A. Palazzo PRD 91, 91301R (2015)

# Conclusions

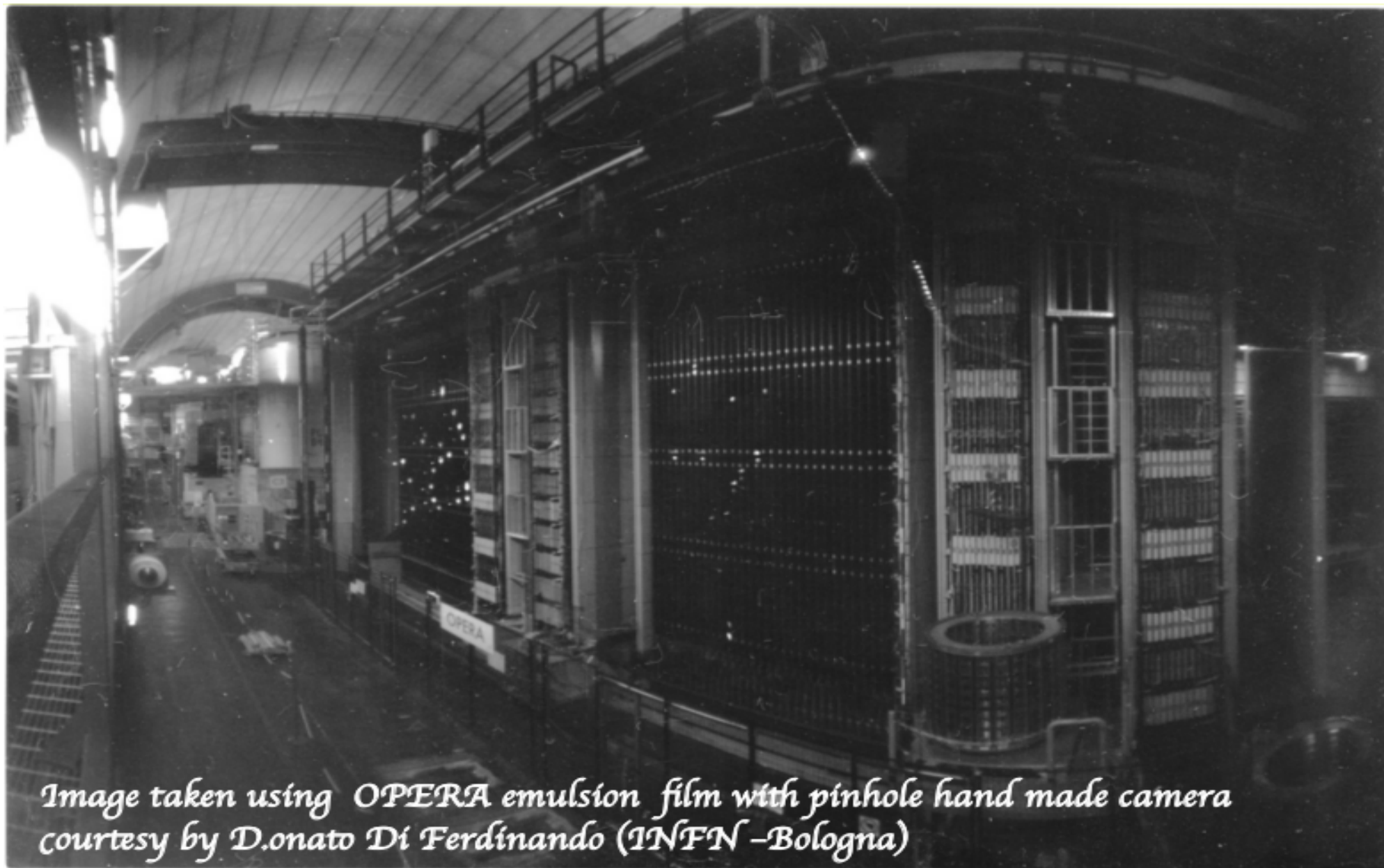
- 5  $\nu_\tau$  candidates with a 0.25 event background
  - No oscillation hypothesis excluded at  $5.1 \sigma$
  - Search for anomalies in  $\nu_\mu \rightarrow \nu_e$  and  $\nu_\mu \rightarrow \nu_\tau$  at a unique L/E.
  - First limits on  $|U_{\mu 4}|^2 |U_{\tau 4}|^2$  from direct measurement of  $\nu_\tau$
- discovery of  $\nu_\tau$  appearance**

## Outlook

- Emulsion analysis ~ complete
- Results  $\nu_\mu \rightarrow \nu_e$  with full statistics soon
- Ongoing characterisation of events failing the cut-based appearance analysis (**marginal events**) still having a significant  $\nu_\tau$  purity



# OPERA taking a "selfie"... Thank you!



*Image taken using OPERA emulsion film with pinhole hand made camera courtesy by D.onato Di Ferdinando (INFN -Bologna)*

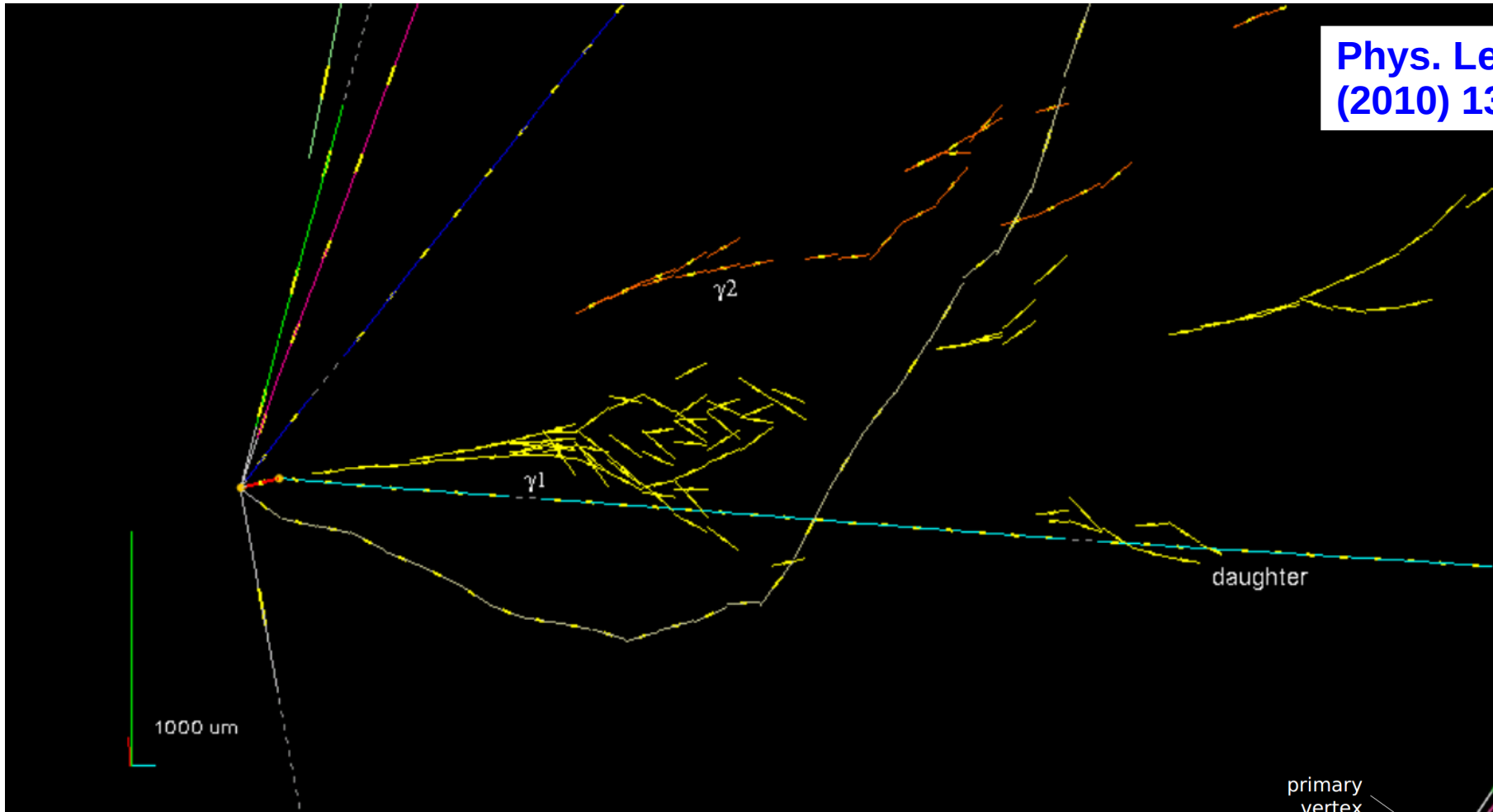


# Institutions

<p><b>Belgium</b> IIHE-ULB Brussels</p> 	<p><b>Italy</b> Bari Bologna Frascati L'Aquila, LNGS Naples Padova Rome Salerno</p> 	<p><b>Russia</b> INR RAS Moscow LPI RAS Moscow ITEP Moscow SINP MSU Moscow JINR Dubna</p> 
<p><b>Croatia</b> IRB Zagreb</p> 	<p><b>Japan</b> Aichi Toho Kobe Nagoyn Nihon</p> 	<p><b>Switzerland</b> Bern</p> 
<p><b>France</b> LAPP Annecy IPHC Strasbourg</p> 	<p><b>Korea</b> Jinju</p> 	<p><b>Turkey</b> METU, Ankara</p> 
<p><b>Germany</b> Hamburg</p> 	<p><b>Israel</b> Technion Haifa</p> 	

# The 1<sup>st</sup> candidate ( $\tau \rightarrow 1h$ )

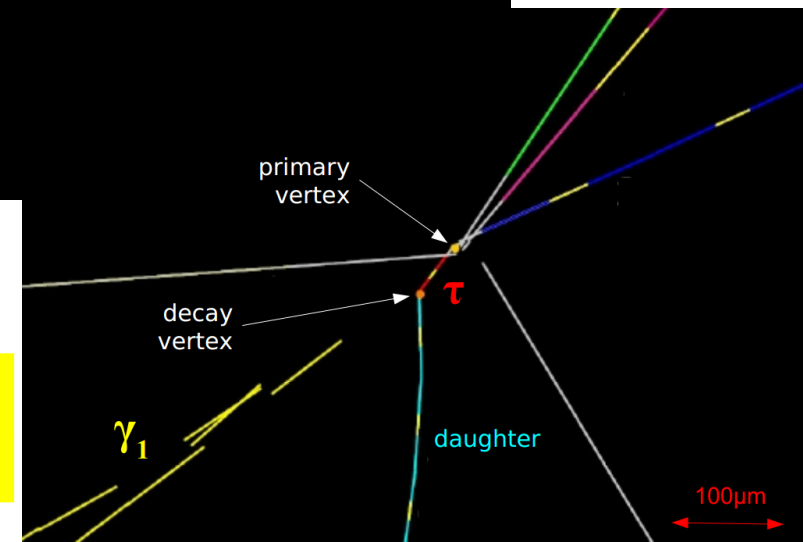
Phys. Lett. B691  
(2010) 138



$$\tau^- \rightarrow \rho^- + \nu_\tau \quad (\text{B.R.} \sim 25\%)$$

$$\rho^- \rightarrow \pi^0 + \pi^- \quad 640^{+125}_{-80} \text{ (stat.) } ^{+100}_{-90} \text{ (sys.) MeV/c}^2$$

$$\pi^0 \rightarrow \gamma\gamma \quad 120 \pm 20 \text{ (stat.) } \pm 35 \text{ (sys.) MeV/c}^2$$

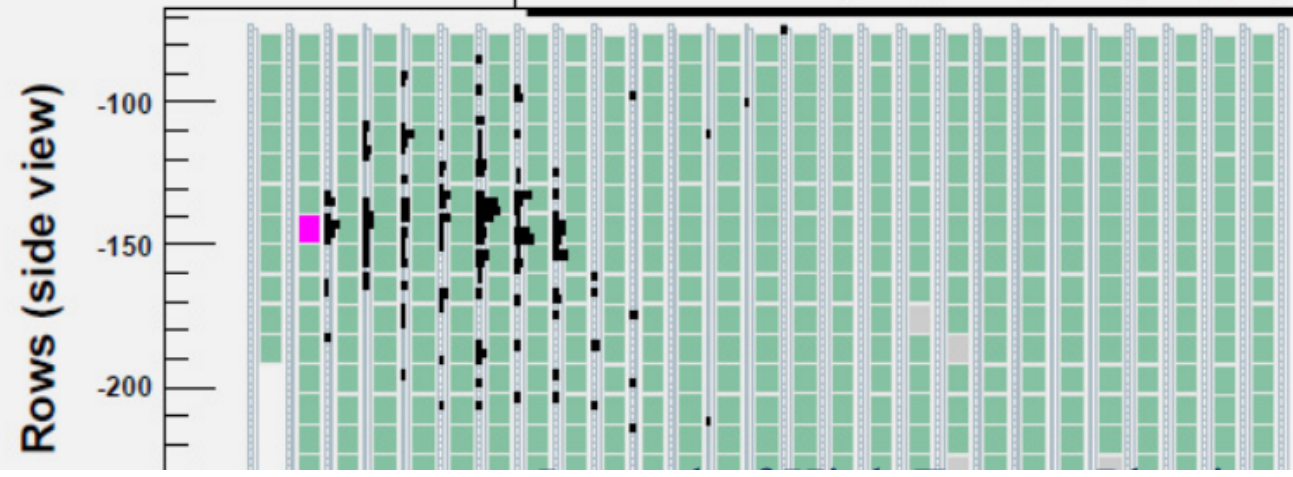




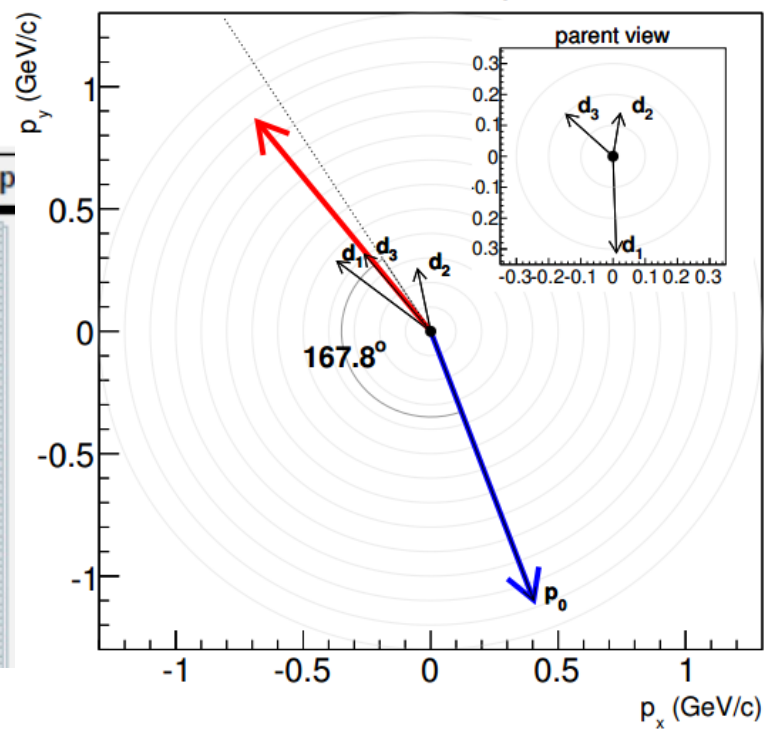


# The 2<sup>nd</sup> candidate ( $\tau \rightarrow 3h$ )

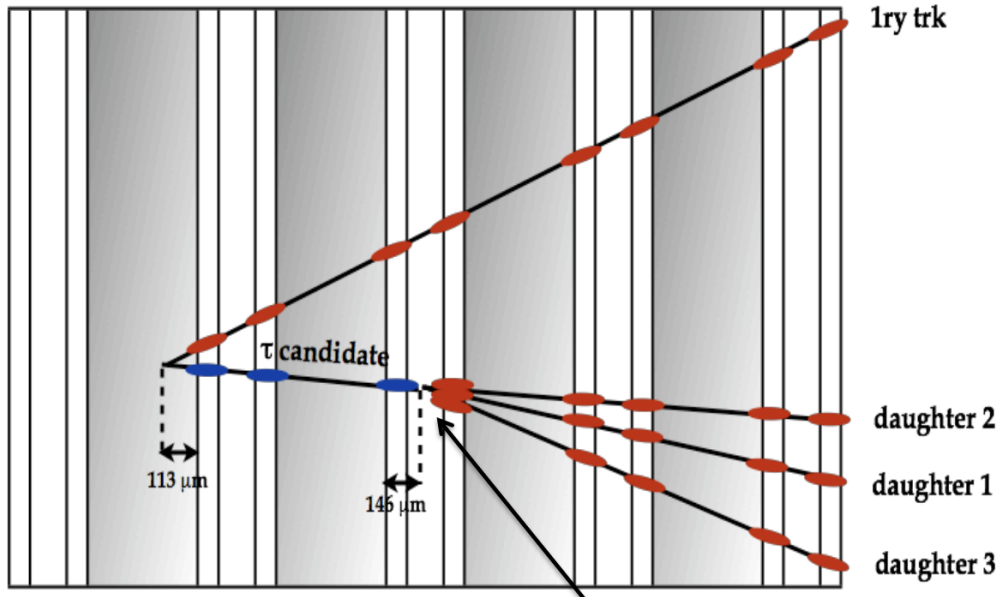
Event: 11113019758, 23 Apr 2011, 07:15 (UTC), YZ p



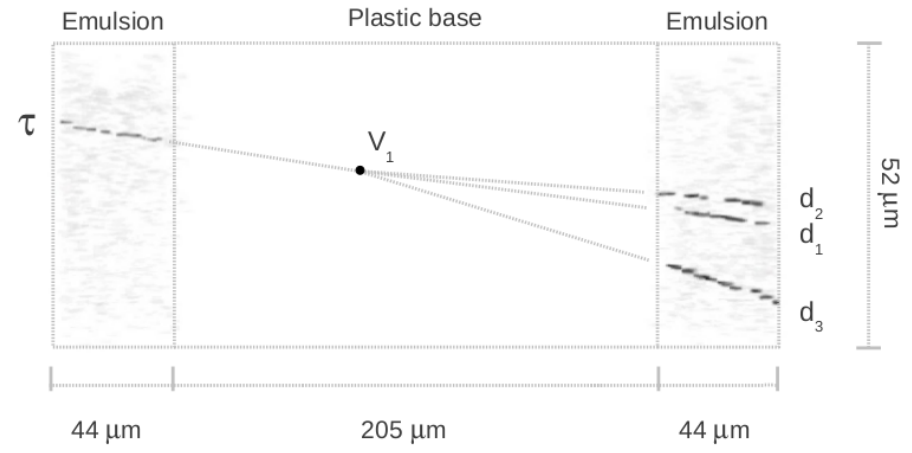
CNGS transverse-plane view



300  $\mu\text{m}$  1 mm 300  $\mu\text{m}$



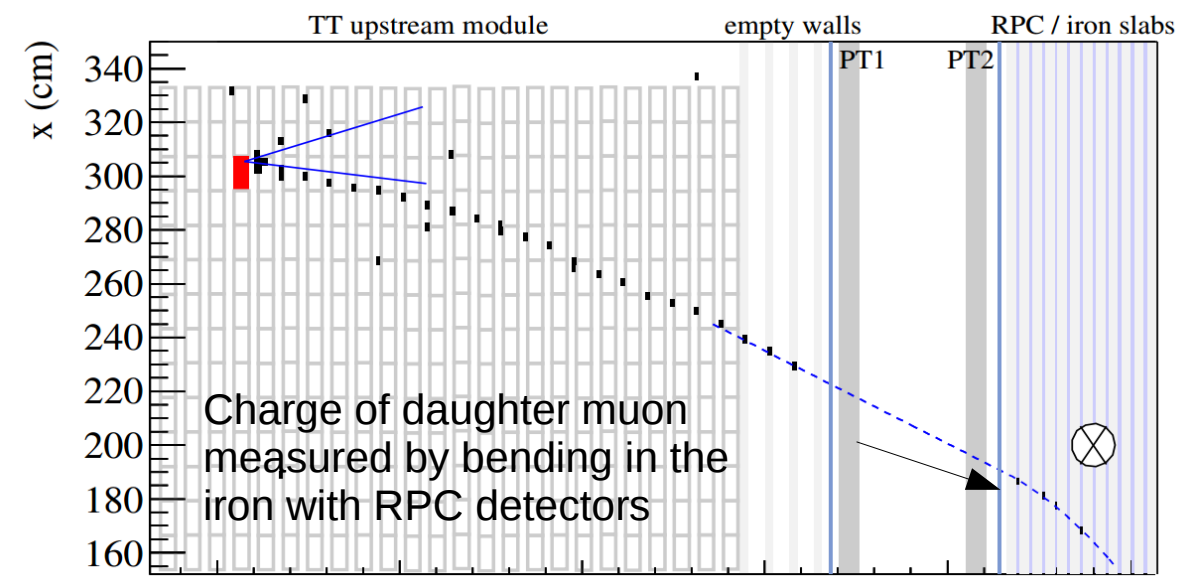
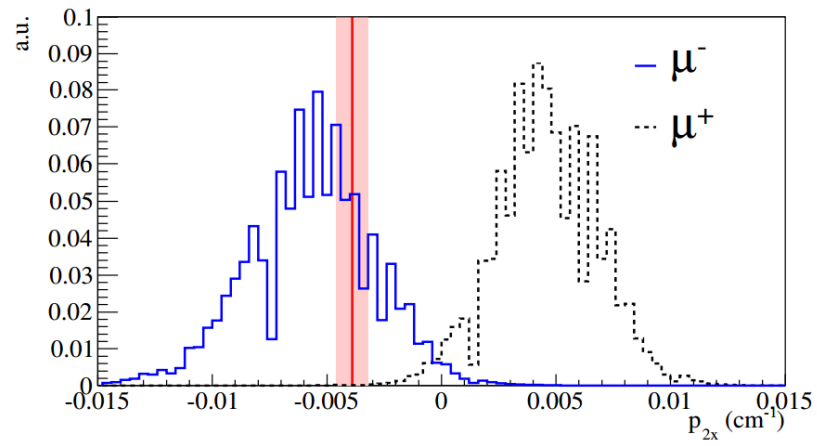
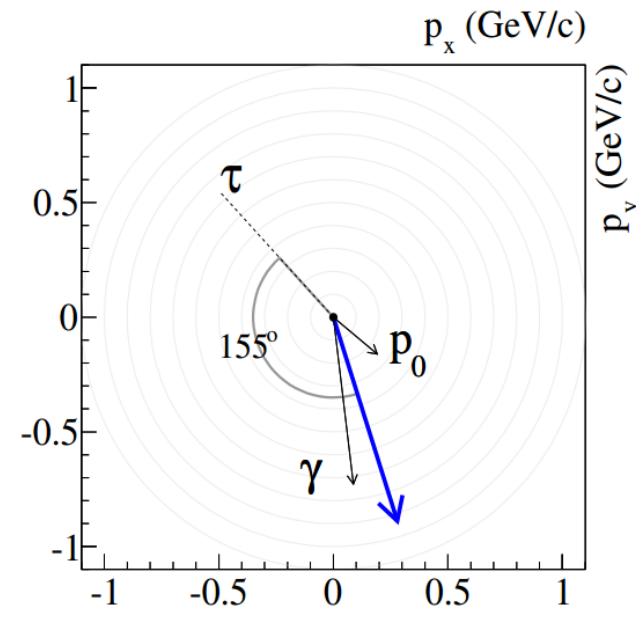
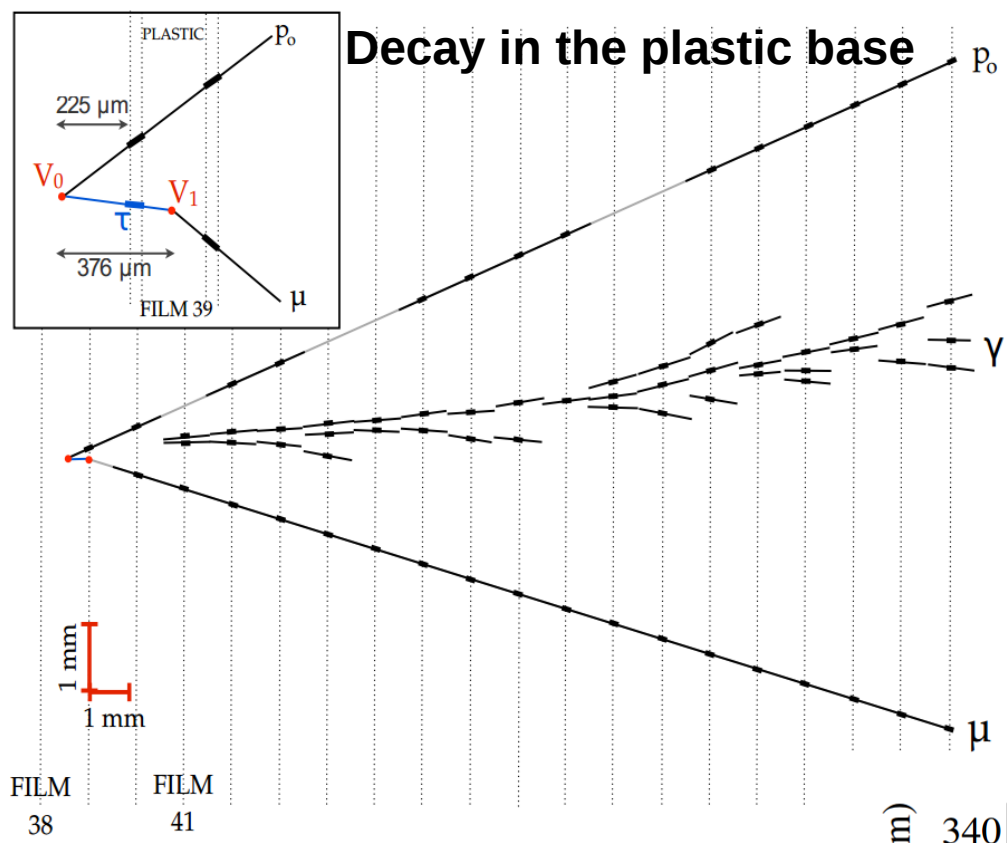
Decay in the plastic base



# The 3<sup>rd</sup> candidate ( $\tau \rightarrow \mu$ )

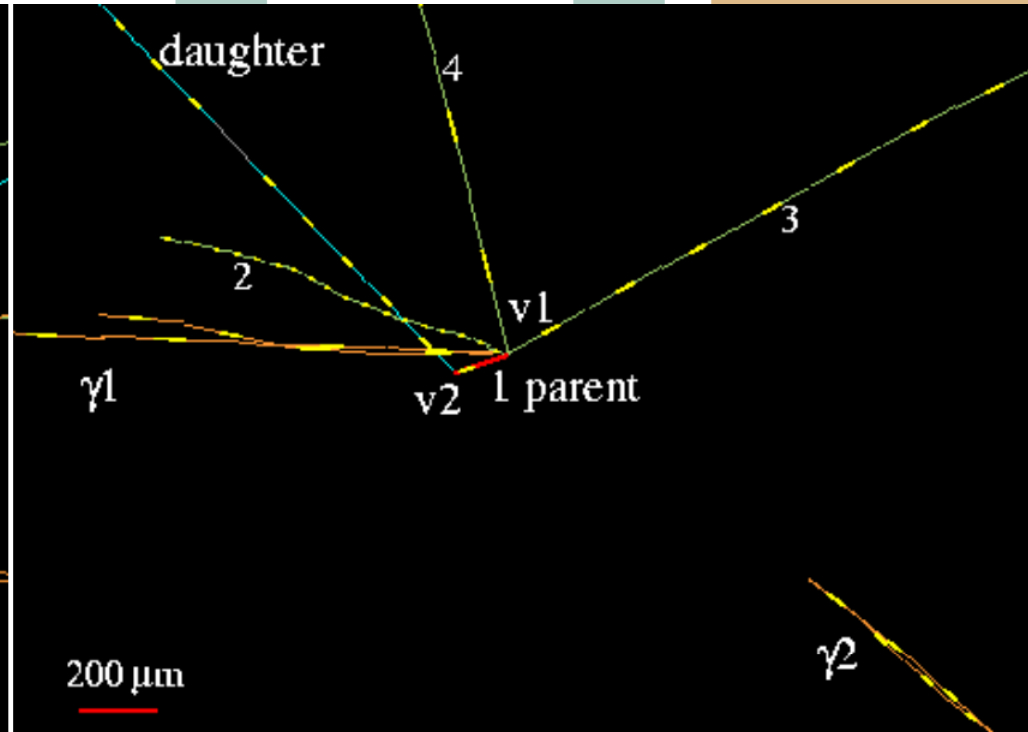
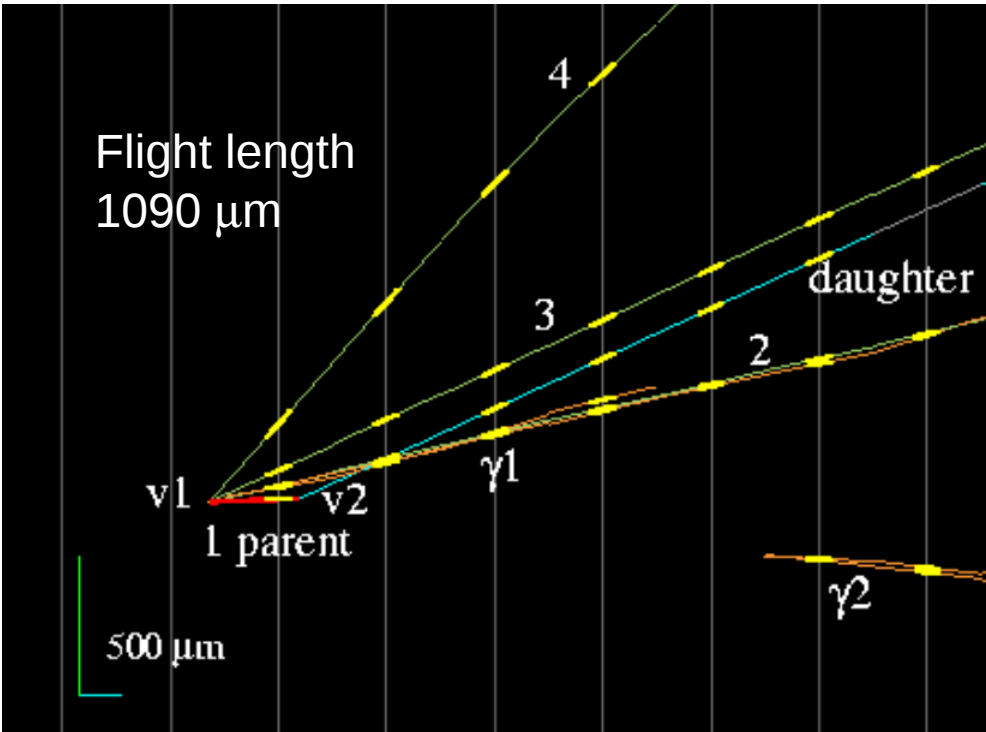
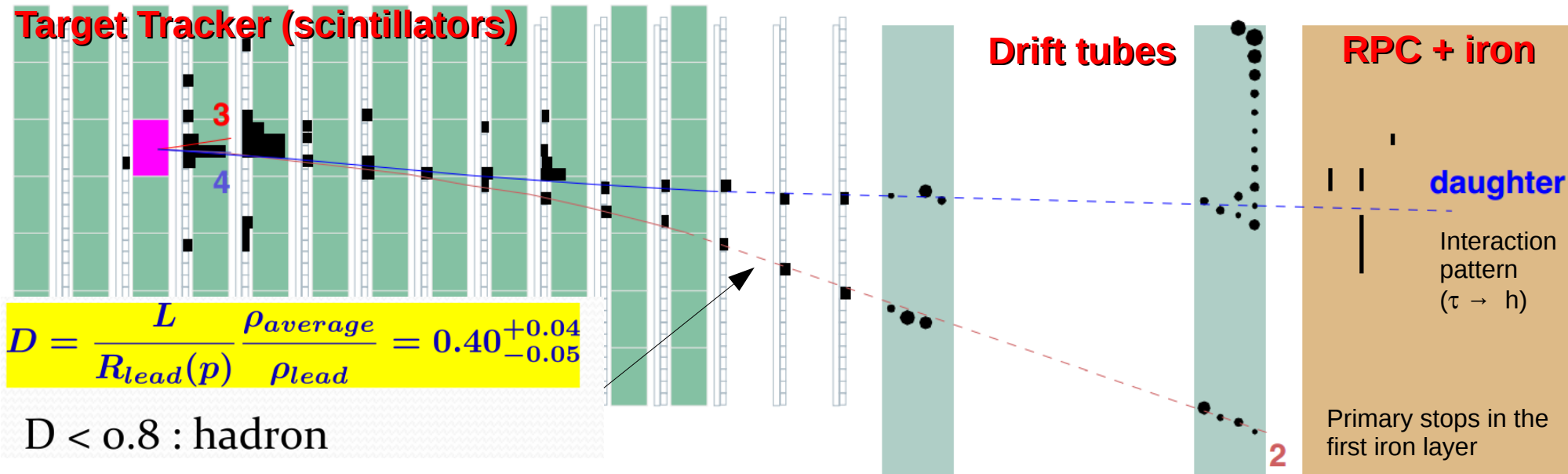


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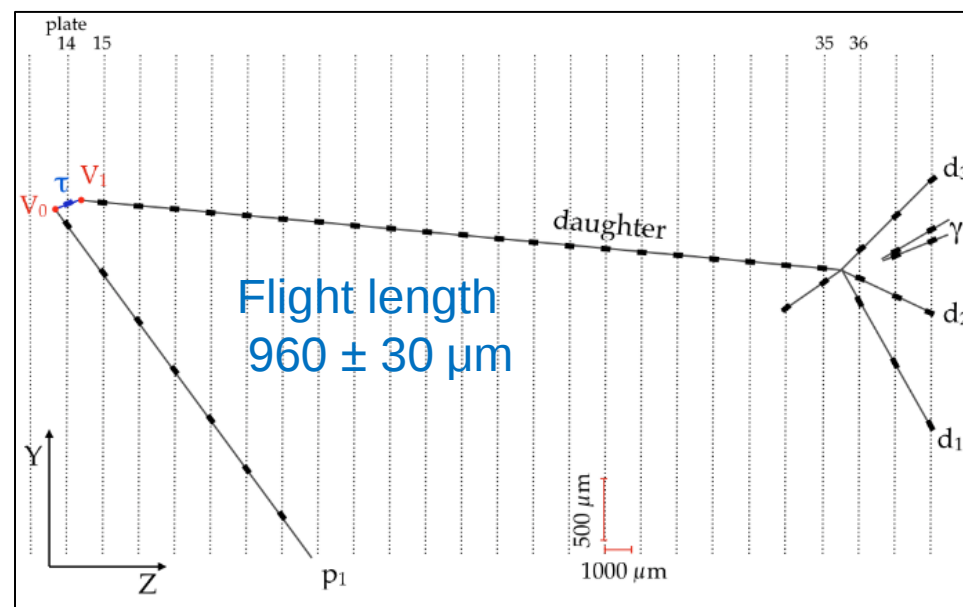
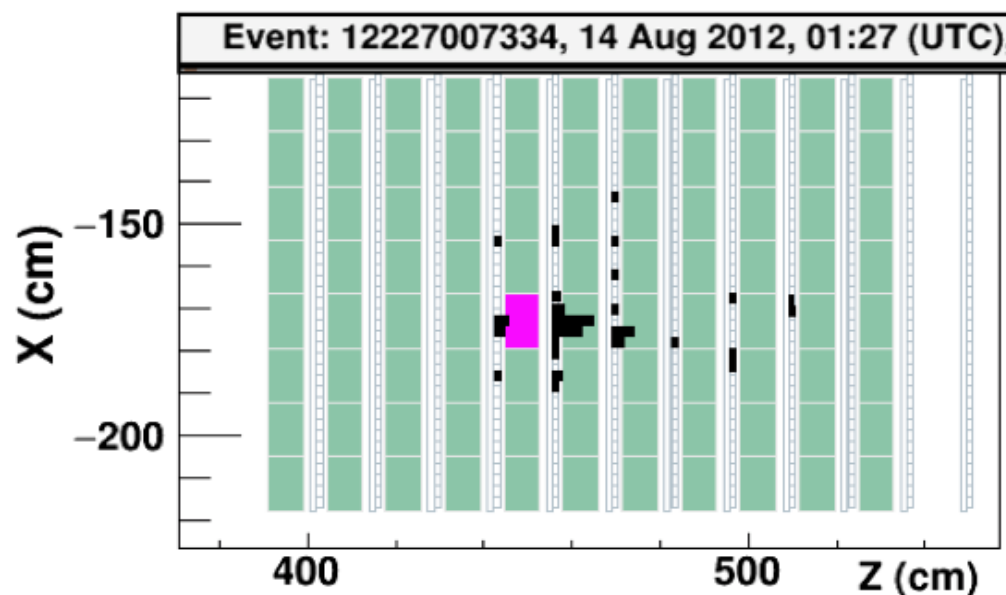


# The 4<sup>th</sup> candidate ( $\tau \rightarrow 1h$ )



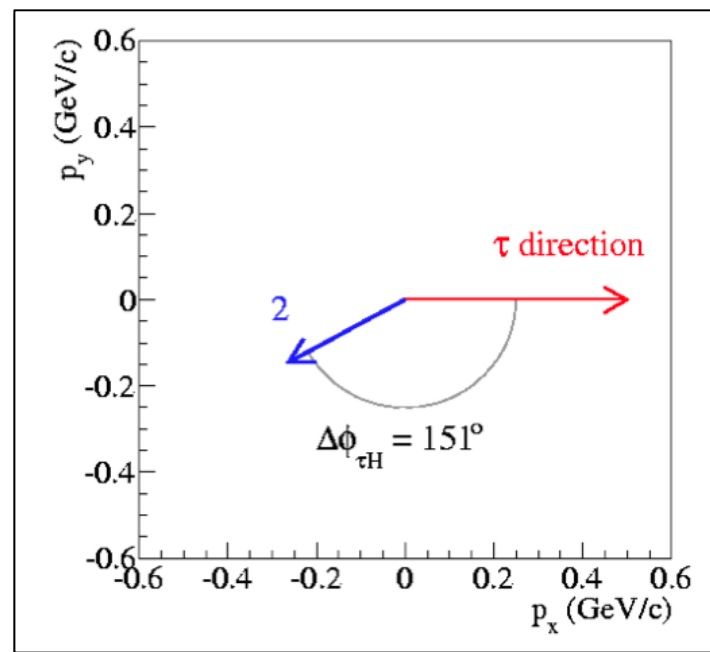


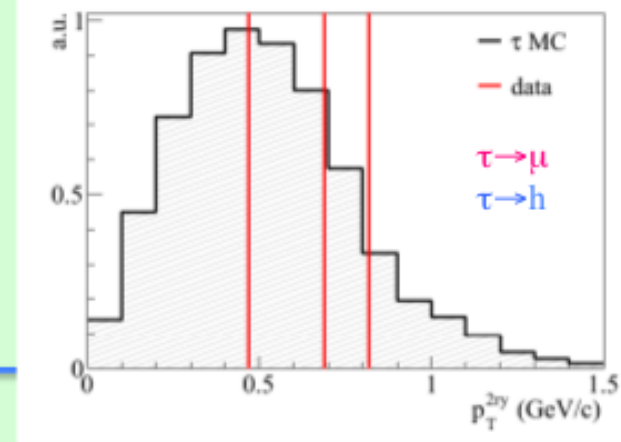
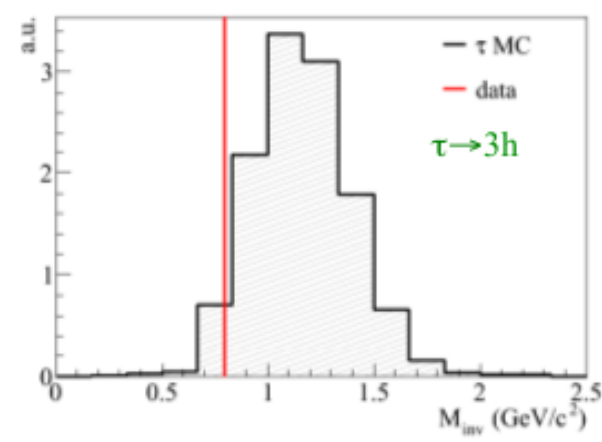
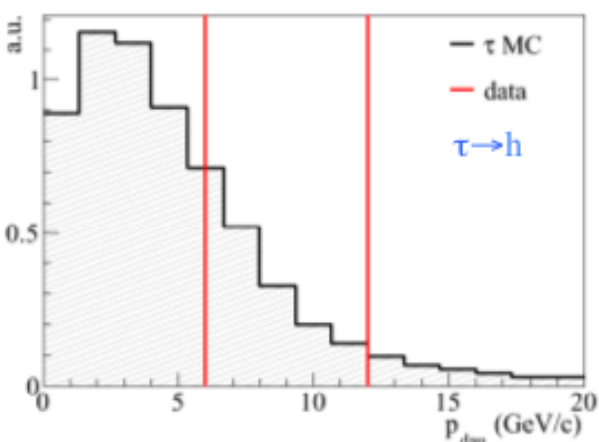
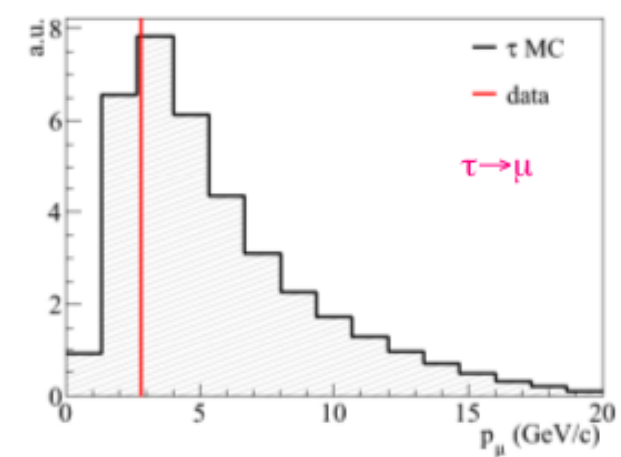
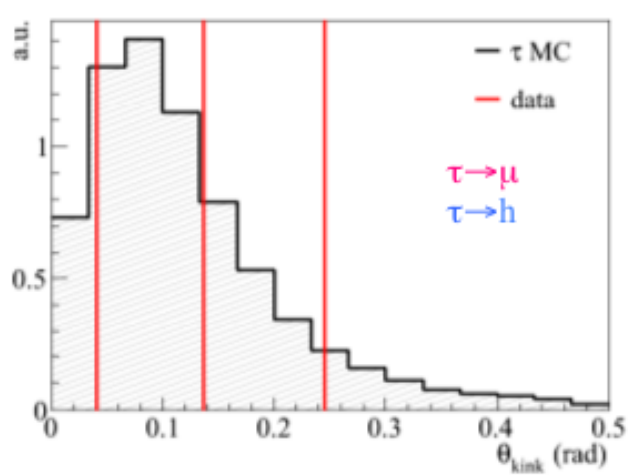
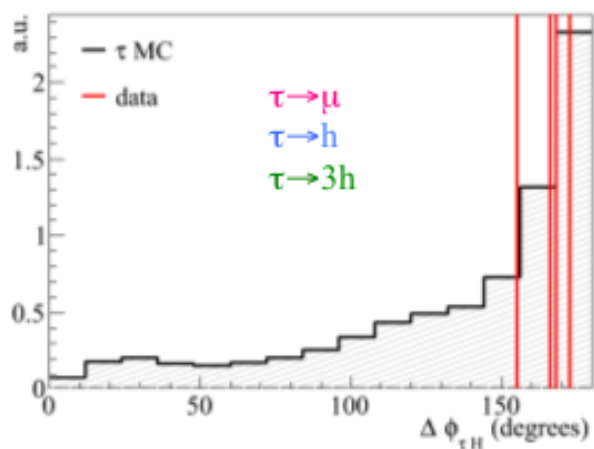
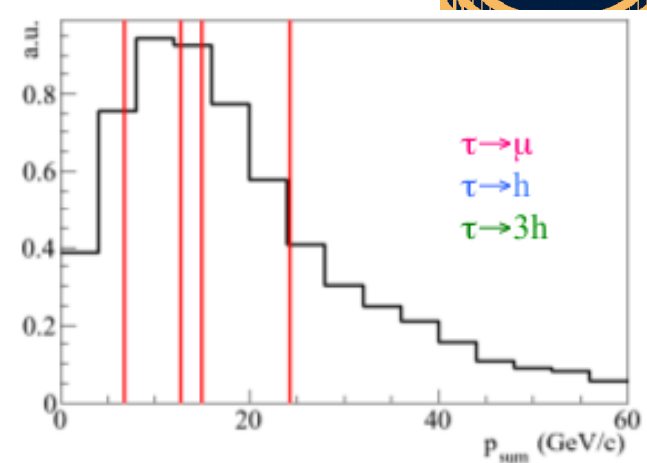
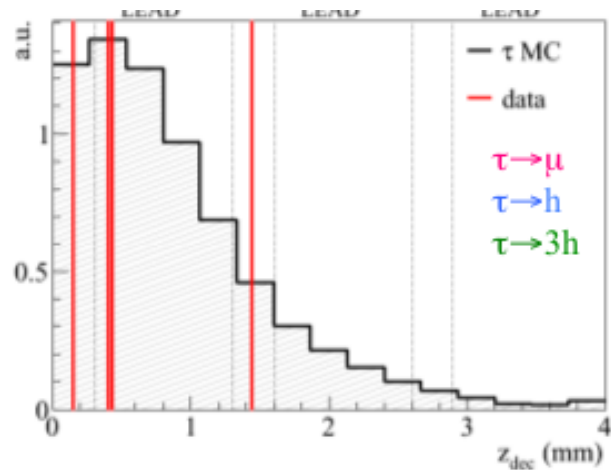
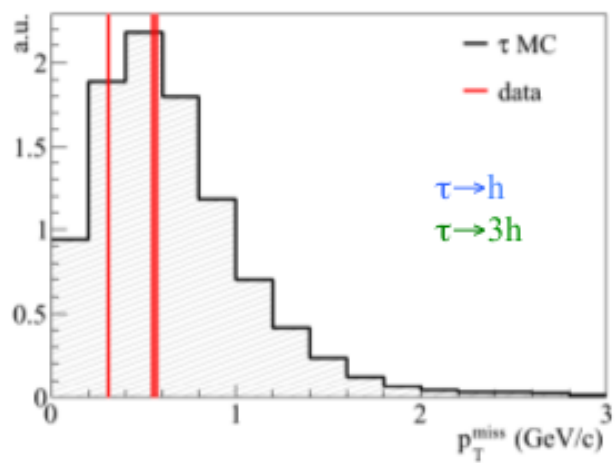
# The 5<sup>th</sup> candidate ( $\tau \rightarrow 1h$ ) [Phys. Rev. Lett. 115 (2015) 12, 121802]



## Kinematical variables

Parameter	Measured value	Selection Criteria
$\Delta\phi_{\tau H}$ ( $^\circ$ )	$151 \pm 1$	$> 90$
$p_T^{miss}$ (GeV/c)	$0.3 \pm 0.1$	$< 1$
$\theta_{kink}$ (mrad)	$90 \pm 2$	$> 20$
$z_{dec}$ ( $\mu m$ )	$634 \pm 30$	[44, 2600]
$p_T^{2ry}$ (GeV/c)	$11_{-4}^{+14}$	$> 2$
$p_T^{2ry}$ (GeV/c)	$1.0_{-0.4}^{+1.2}$	$> 0.6$ (no $\gamma$ attached)





# $\nu_\tau$ kinematic analysis

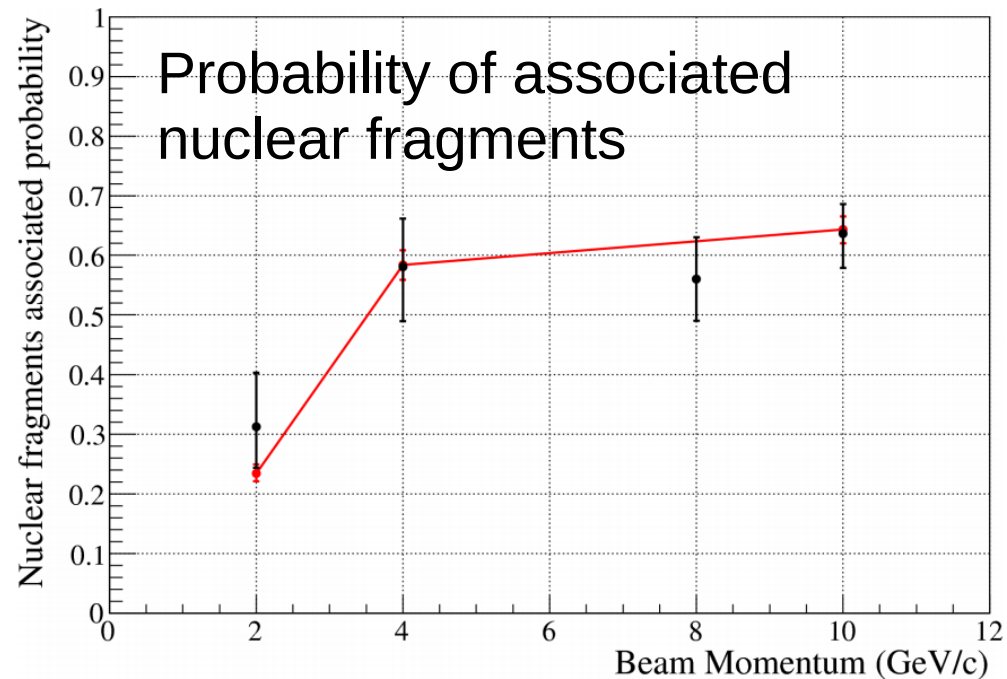
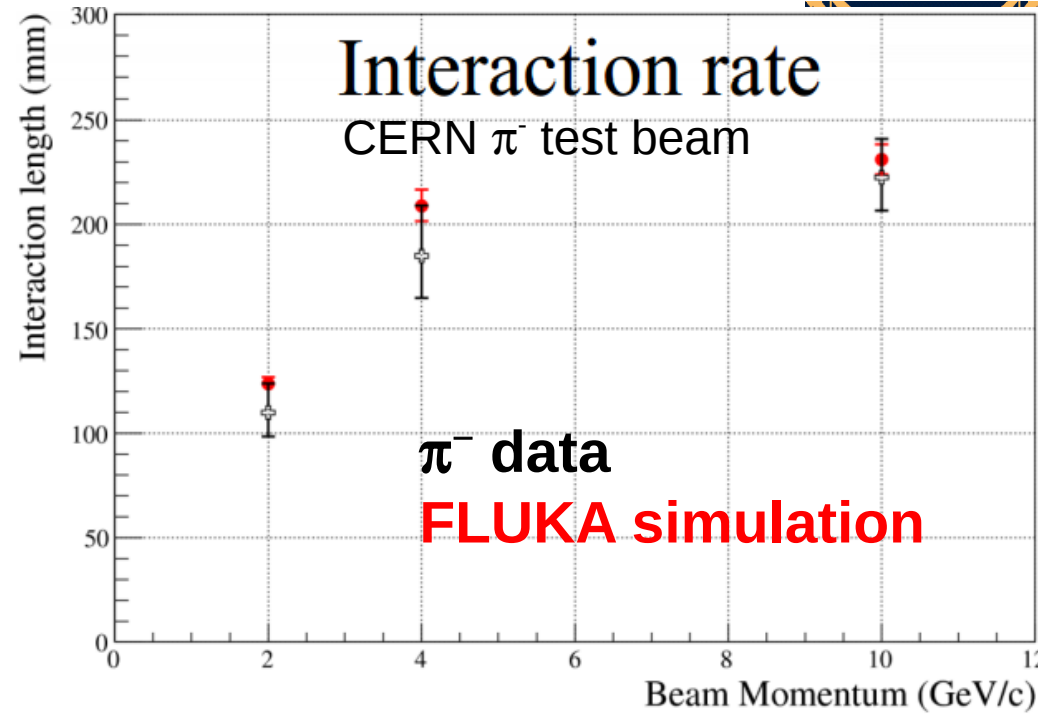
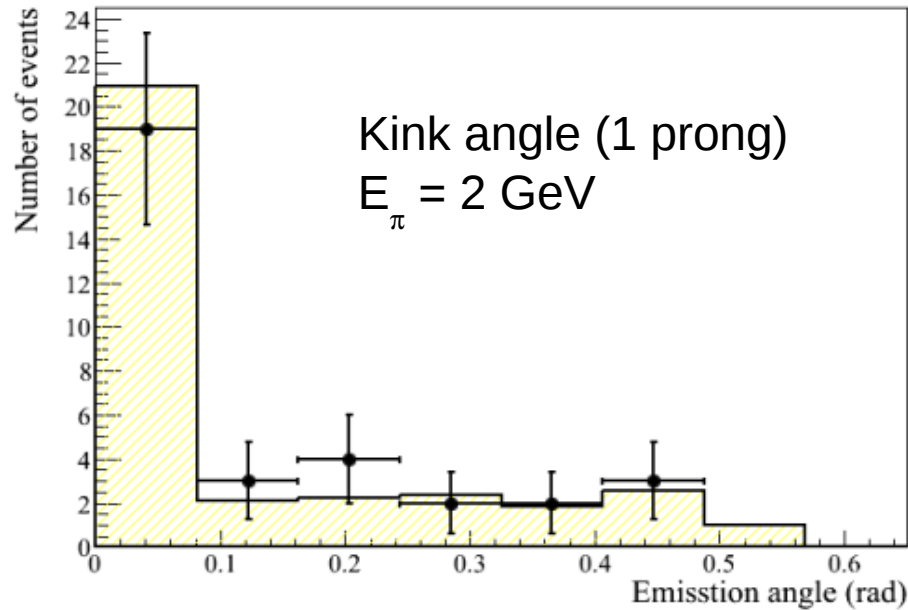
**Candidate events  
have to fulfil  
kinematical cuts:** →

defined in the experiment  
proposal to enhance the  
S/B ratio

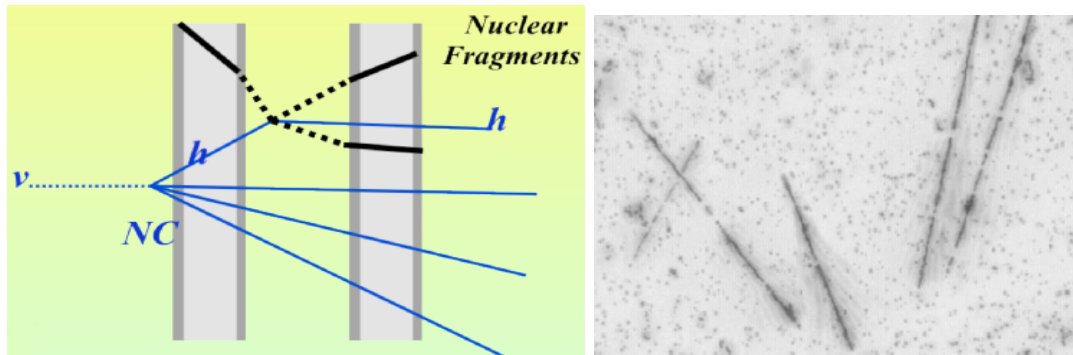
variable	$\tau \rightarrow 1h$	$\tau \rightarrow 3h$	$\tau \rightarrow \mu$	$\tau \rightarrow e$
lepton-tag		No $\mu$ or $e$ at the primary vertex		
$z_{dec}$ ( $\mu\text{m}$ )	[44, 2600]	< 2600	[44, 2600]	< 2600
$p_T^{miss}$ (GeV/c)	< 1*	< 1*	/	/
$\phi_{lH}$ (rad)	> $\pi/2^*$	> $\pi/2^*$	/	/
$p_T^{2ry}$ (GeV/c)	> 0.6(0.3)*	/	> 0.25	> 0.1
$p^{2ry}$ (GeV/c)	> 2	> 3	> 1 and < 15	> 1 and < 15
$\theta_{kink}$ (mrad)	> 20	< 500	> 20	> 20
$m, m_{min}$ (GeV/c <sup>2</sup> )	/	> 0.5 and < 2	/	/



# Hadronic background: $\pi$ test beams

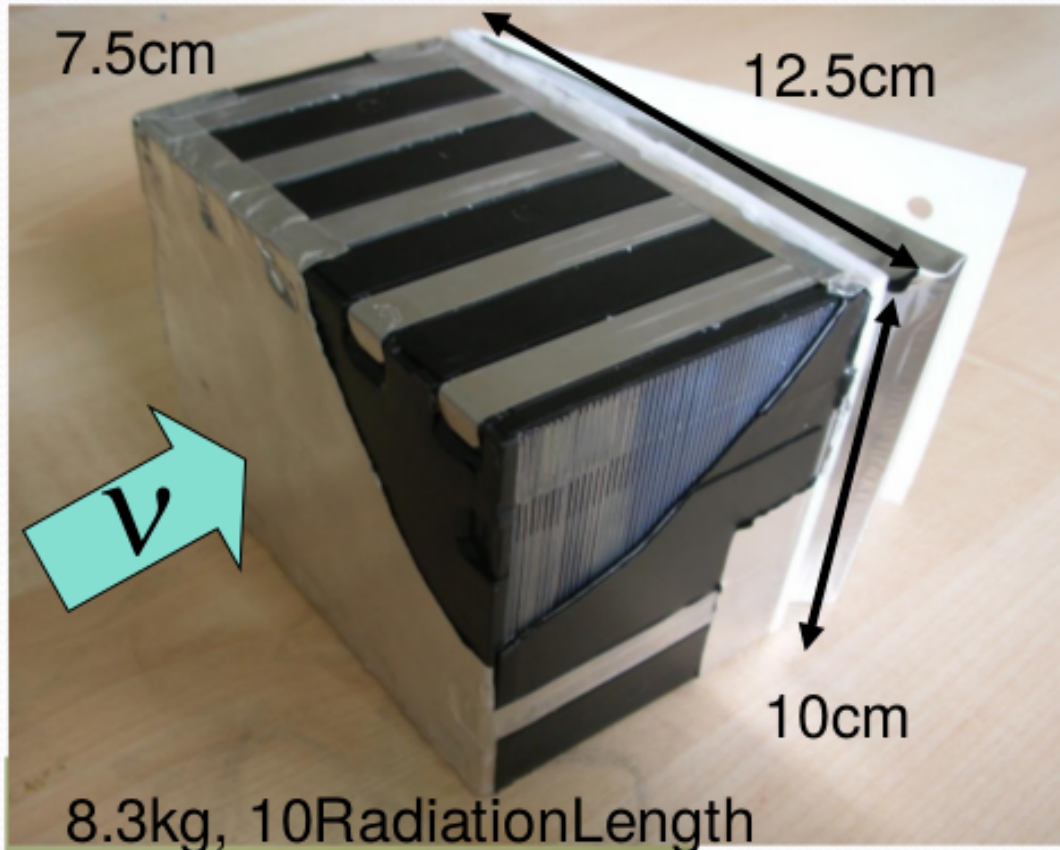


**Nuclear fragments:** a smoking gun for the occurrence of an  $\pi$  interaction instead of a decay.



Hadronic background rate per located event:  $\tau \rightarrow (3)h = (1.5)3.09 \times 10^{-5}$

# The ECC



- **ECC properties**

- 56 of 1mm thick lead plates interleaved with 57 emulsion films.
- 8.3kg / brick
- 10 radiation length

- **150,000 ECC bricks**

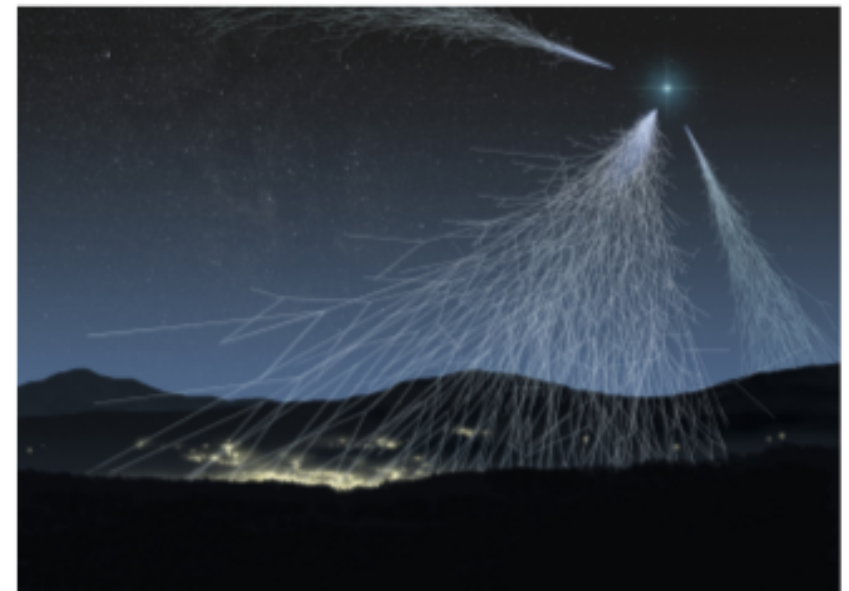
- 1.25 ktons
- 9 million films

- **Capability**

- Micrometric accuracy vertex analysis
- Kinematical analysis
  - Momentum measurement by MCS.
  - EM energy measurement

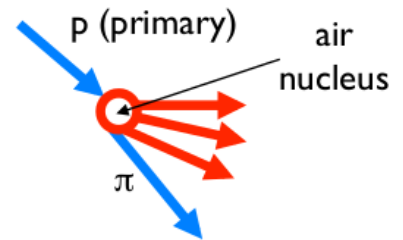
# The atmospheric muon charge ratio

- The atmospheric muon charge ratio  $R_\mu \equiv N_{\mu^+}/N_{\mu^-}$  is being studied and measured since many decades
  - Depends on the **chemical composition** and energy spectrum of the primary cosmic rays
  - Depends on the **hadronic interaction features**
  - At high energy, depends on the **prompt component**
- Possibility to check HE hadronic interaction models ( $E > 1 \text{ TeV}$ ) in the **fragmentation region** (phase space complementary to collider's one)
- Atmospheric muons are kinematically related to atmospheric neutrinos (same sources)  $\rightarrow R_\mu$  provides a benchmark for **atmospheric  $\nu$  flux computations** (e.g. background for neutrino telescopes)





# Cosmic rays: $R = N_{\mu^+} / N_{\mu^-}$



- Highest-E region reached!
- opposite magnet polarities runs → lower systematics
- Strong reduction of the charge ratio for multiple muon events

1 $\mu$	$1.377 \pm 0.006$
Multi- $\mu$	$1.098 \pm 0.023$

- Results compatible with a simple  $\pi$ -K model
- No significant contribution of the prompt component up to  $E_{\mu} \cos \theta^* \sim 10$  TeV
- Validity of Feynman scaling in the fragmentation region up to  $E_{\mu} \sim 20$  TeV ( $E_N \sim 200$  TeV)

$$\phi_{\mu^{\pm}} \propto \frac{a_{\pi} f_{\pi^{\pm}}}{1 + b_{\pi} \mathcal{E}_{\mu} \cos \theta / \epsilon_{\pi}} + R_{K\pi} \frac{a_K f_{K^{\pm}}}{1 + b_K \mathcal{E}_{\mu} \cos \theta / \epsilon_K}$$

