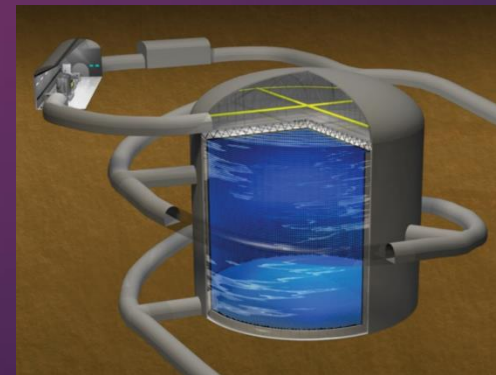
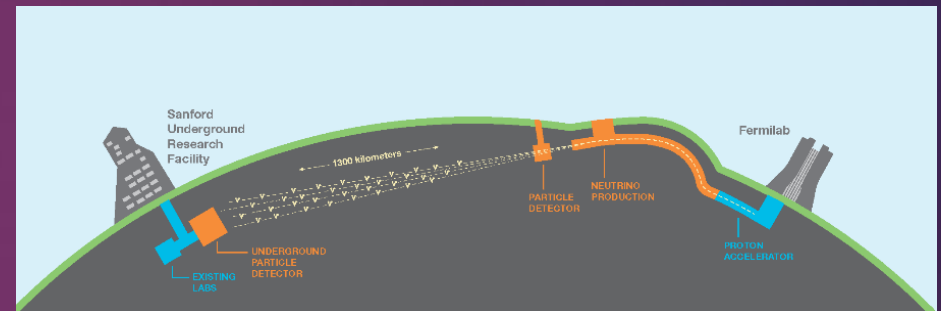


# Long-Baseline Neutrino Experiments: The Future

LISA WHITEHEAD KOERNER  
UNIVERSITY OF HOUSTON  
JULY 26, 2017

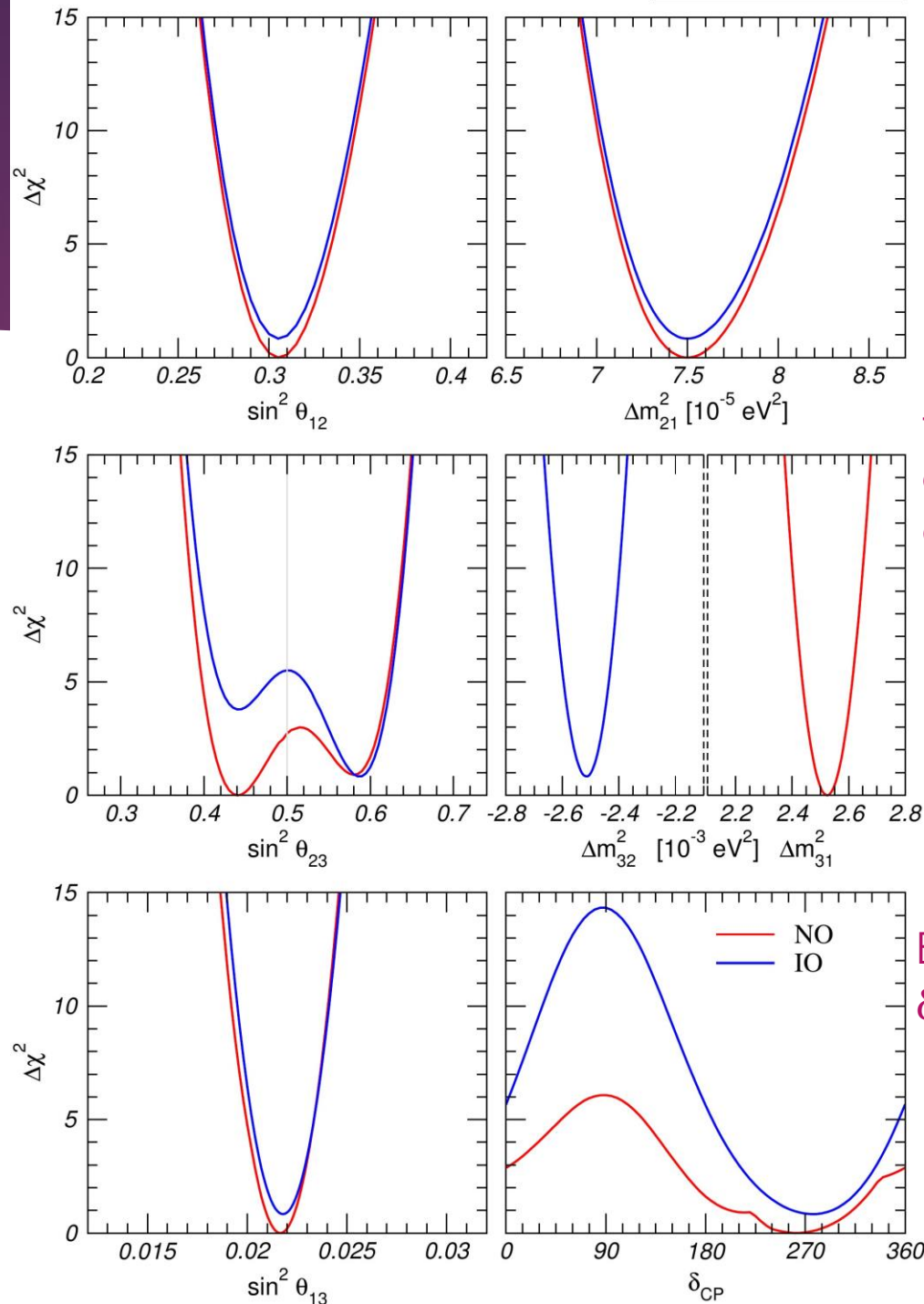


# Introduction

- Some open questions in 3-neutrino mixing:
  - Is CP violated in neutrino oscillations ( $\delta \neq 0, \pi$ )?
  - Is  $\theta_{23}$  maximal? ( $=\pi/4$ )?
  - What is the neutrino mass ordering?
- Or new physics?

<http://www.nu-fit.org>

NuFIT 3.0 (2016)

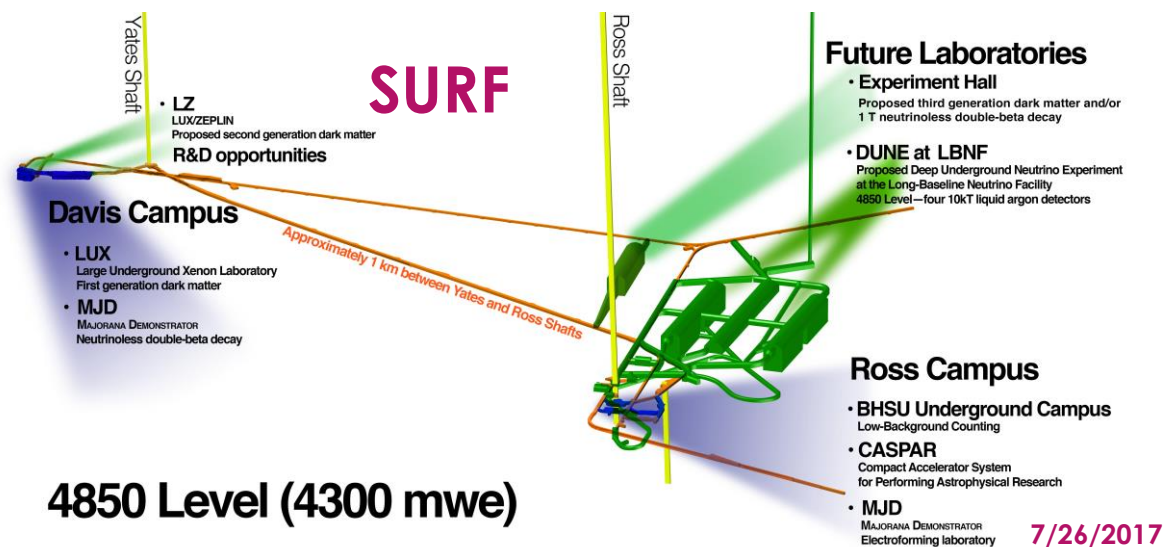
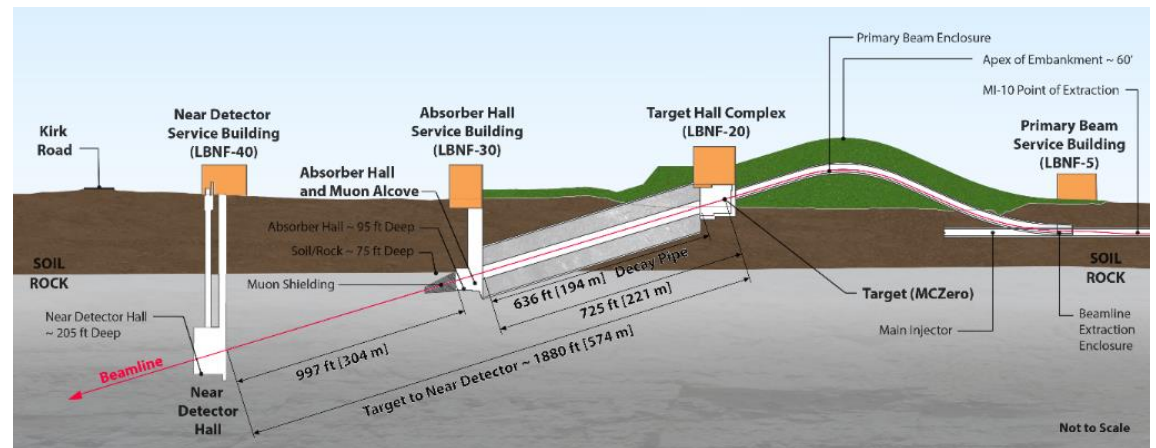


Recent global fit of neutrino oscillation data

Best fit  
 $\delta_{CP} = 261^{+51}_{-59}$   
 ( $\approx -\pi/2$ )

# DUNE (Deep Underground Neutrino Experiment)

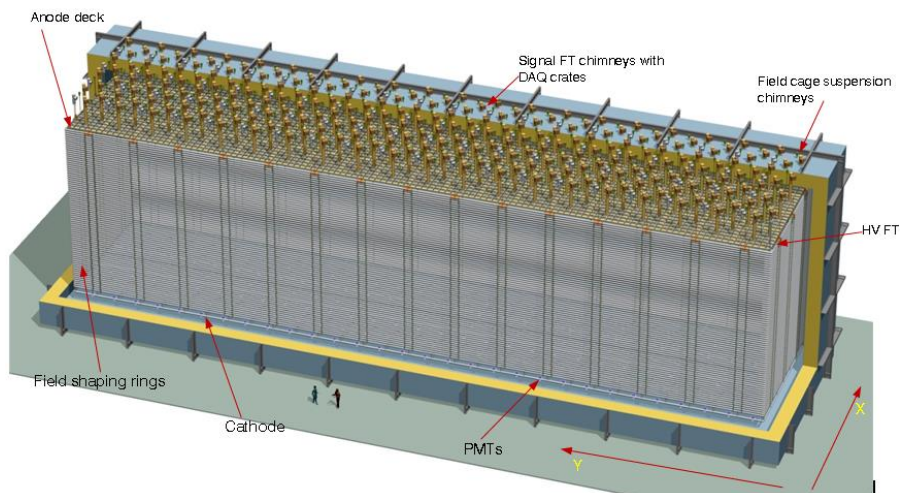
- Muon neutrino beam from Fermilab (LBNF – Long Baseline Neutrino Facility)
  - On-axis broadband beam
  - Beam intensity 1.2 MW, upgradable to 2.4 MW (120 GeV primary protons)
- Far detector at SURF in South Dakota
  - 1300 km baseline
  - 4300 mwe overburden



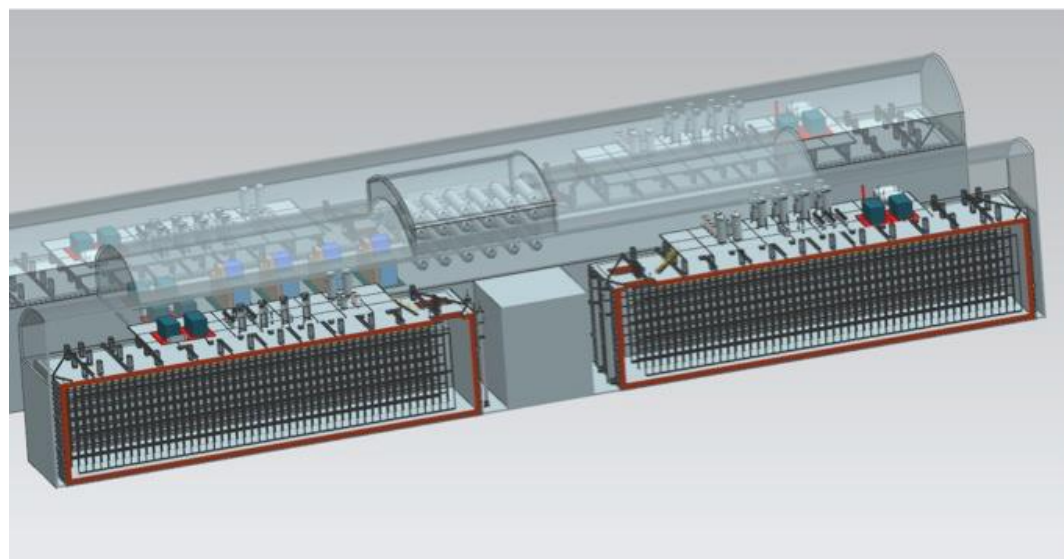
# DUNE Far Detector

**40 kt liquid argon (LAr) TPC**  
**4 x 10 kt modules**  
 (Modules not necessarily identical)

**Dual-phase TPC**  
 (single module with amplification in gas phase)



L.W. Koerner, University of Houston



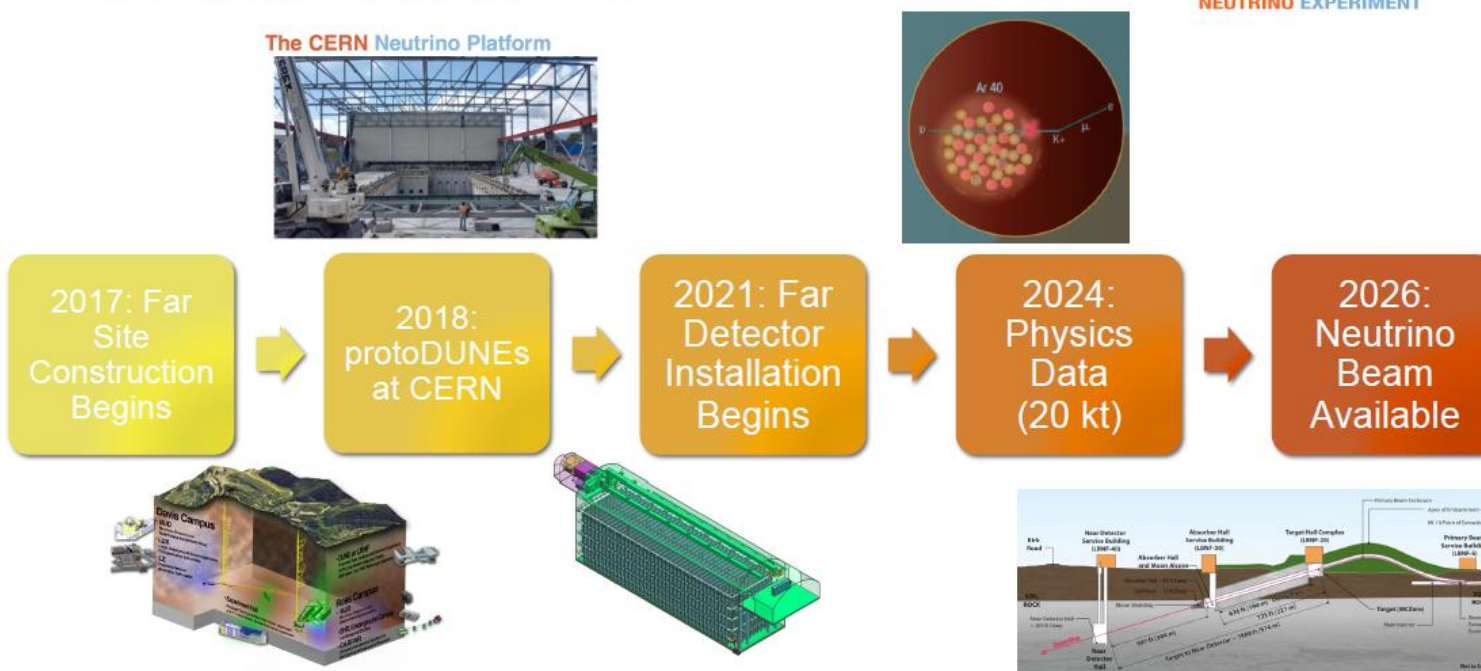
**Single-phase TPC**

**Suspended anode (APA) and cathode (CPA) assemblies – 3.6 m spacing**

7/26/2017

# DUNE Status and Timeline

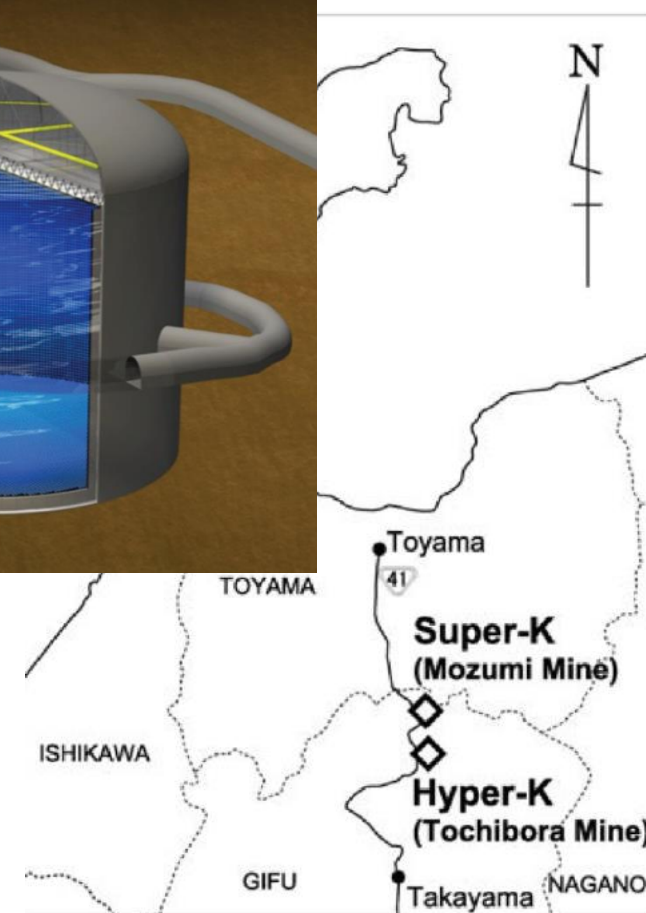
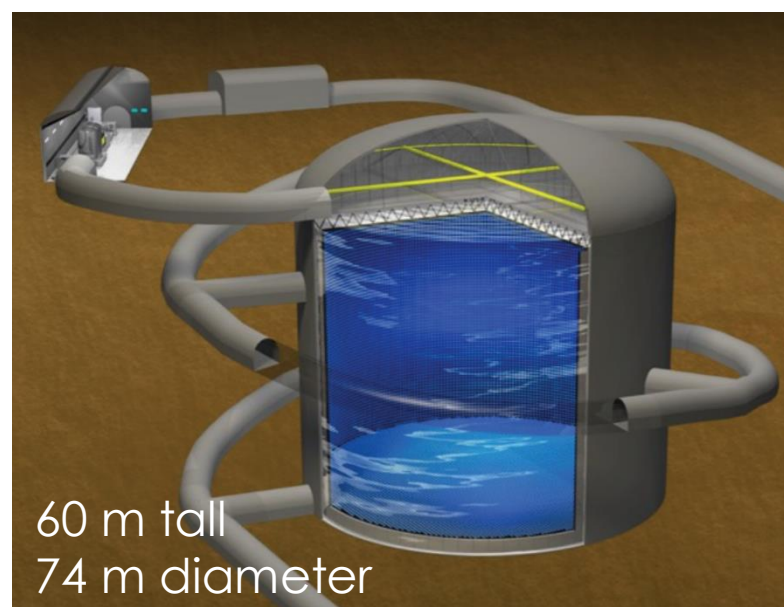
## DUNE Timeline



- ▶ International collaboration
  - ▶ Began in 2015
  - ▶ Nearly 1000 collaborators from 30 countries
- ▶ Far site ground breaking ceremony July 21!

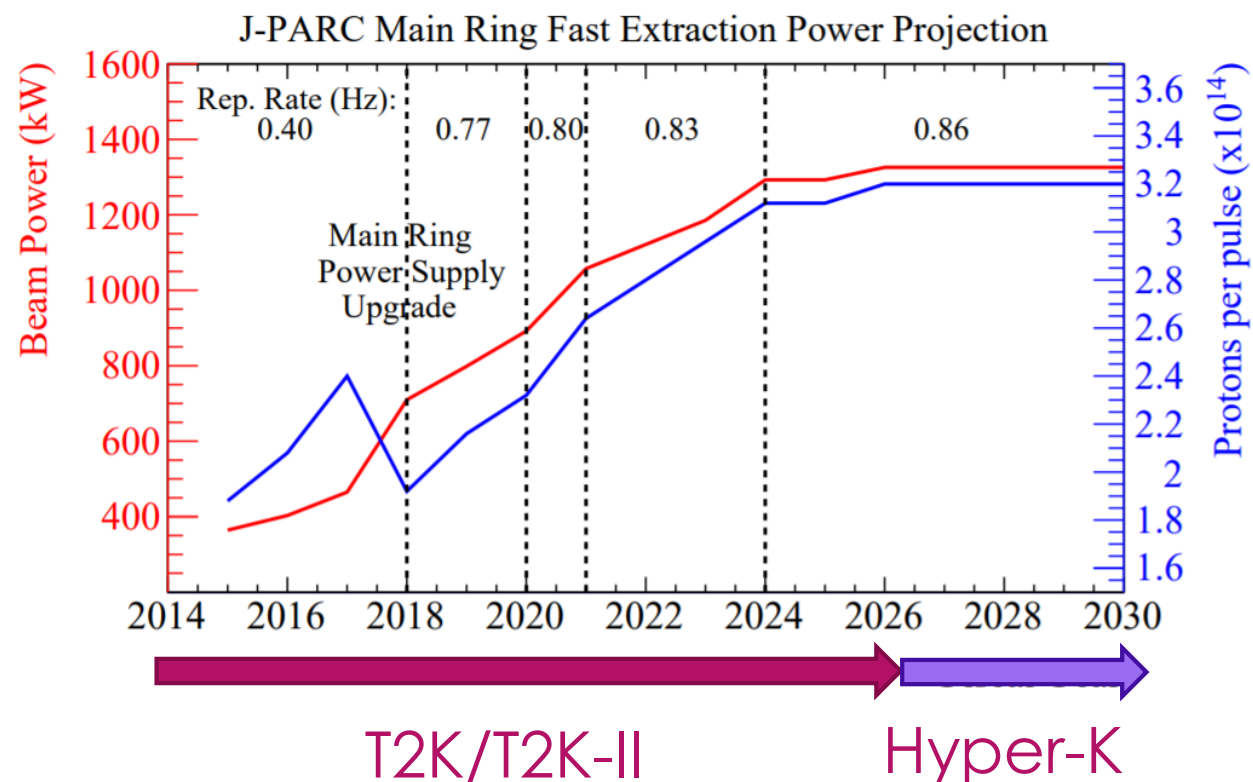
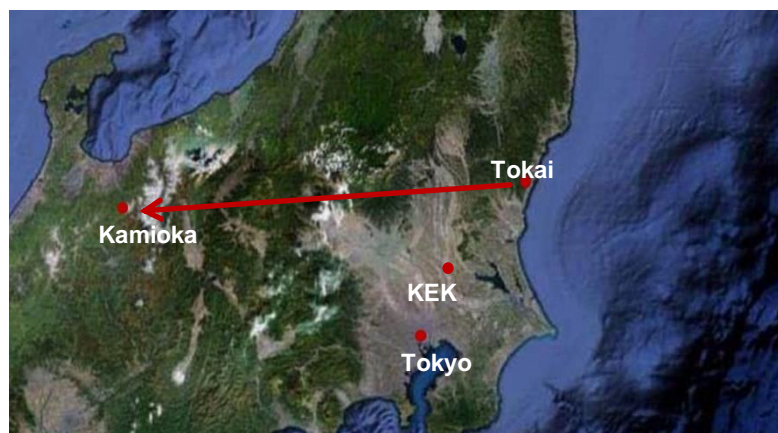
# Hyper-Kamiokande Detector

- ▶ Water Cherenkov detector
  - ▶ 260 kton ultra pure water (Fiducial mass 187 kton)
  - ▶ New 50 cm photo sensors with improved single photon detection efficiency (2x Super-K PMTs)
  - ▶ 40% photocathode coverage
  - ▶ 650 m (1750 mwe) depth
- ▶ Aiming for a quick start with one tank
  - ▶ Second tank under consideration (time, design, location...)

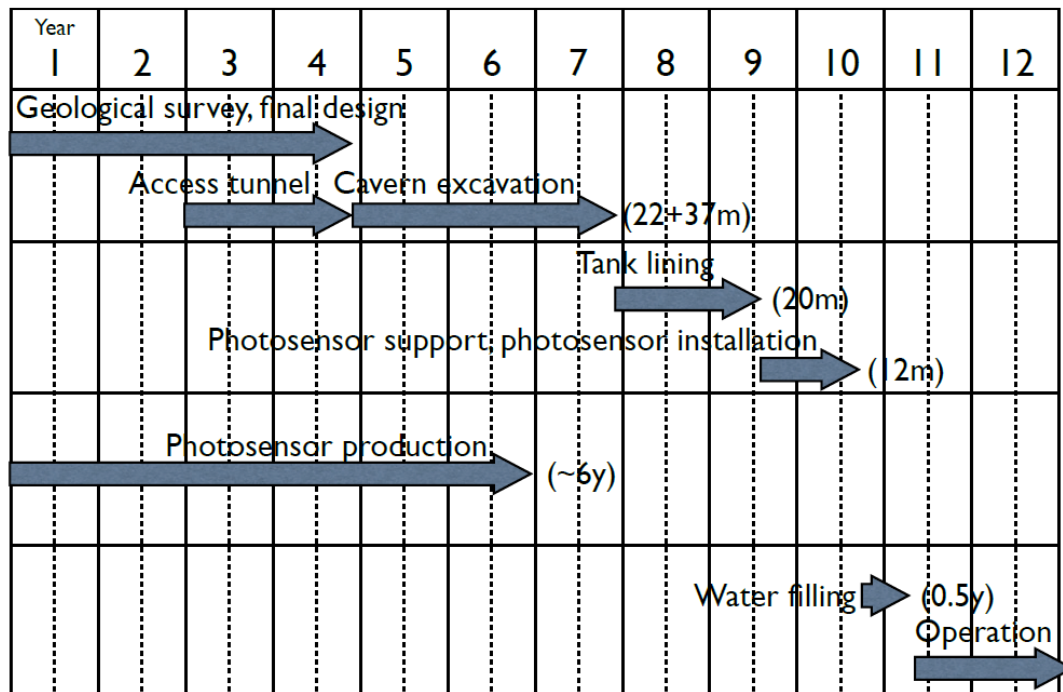


# Hyper-K Long-Baseline Program

- ▶ Muon neutrino beam from J-PARC (currently being used for T2K)
  - ▶ Same 295 km baseline,  $2.5^\circ$  off-axis angle
  - ▶ Beam power of 1-1.3 MW achievable after upgrades



# Hyper-K Status and Timeline



- ▶ International Collaboration
  - ▶ Began in 2015
  - ▶ As of April 2017, 300 members from 15 countries
- ▶ Just last week, a draft of the MEXT (funding agency) Roadmap for Large Projects was released and includes Hyper-K as an important component
- ▶ Budget request to start construction in JFY2018
  - ▶ Aim to begin operation in 2026



# Sensitivity Assumptions

## DUNE

- ▶ Staging: Begin with 20 kton, 1.07 MW beam; 40 kton in year 4, 2.14 MW in year 7
- ▶ Neutrino: Antineutrino = 1:1
- ▶  $\theta_{23}$  from global fit (non-maximal)

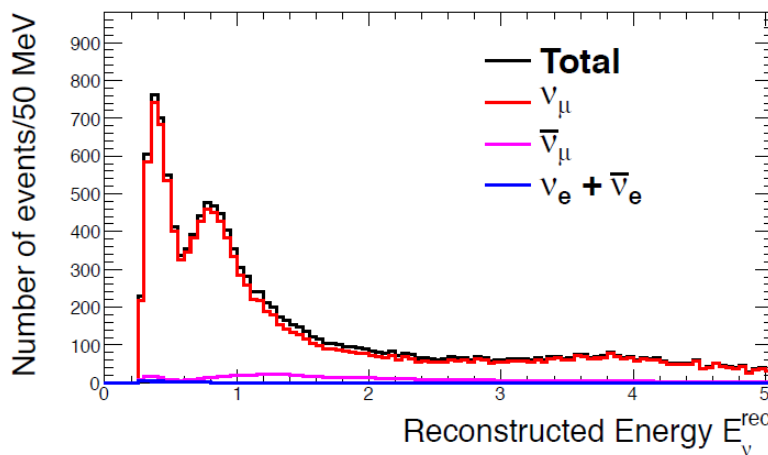
## Hyper-K

- ▶ Staging: Begin with single 187 kton fiducial tank and 1.3 MW beam; second tank in year 7
- ▶ Neutrino: Antineutrino = 1:3
- ▶  $\theta_{23}$  maximal

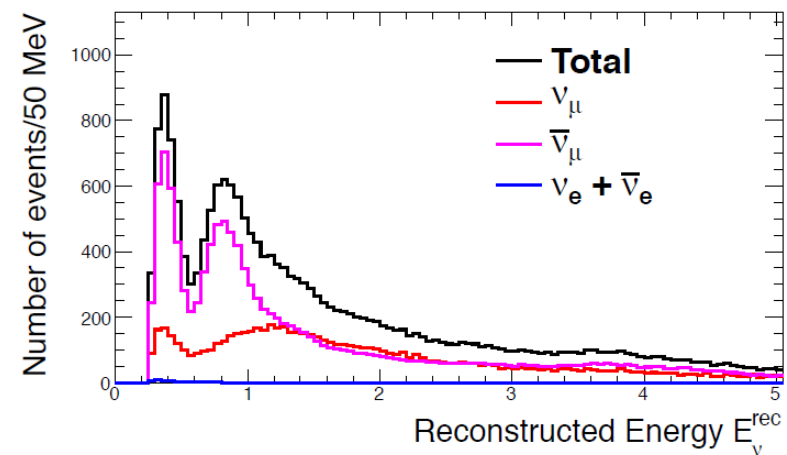
# Disappearance Spectra

Hyper-K  
(10 yrs  
total)

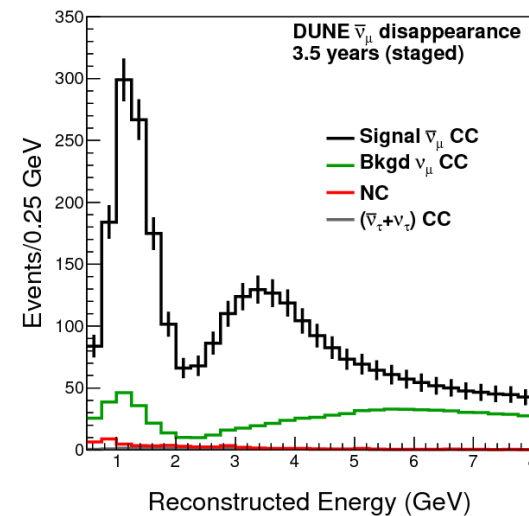
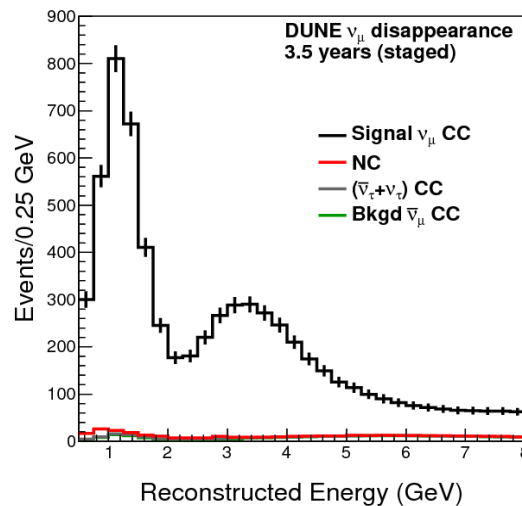
Disappearance  $\nu$  mode



Disappearance  $\bar{\nu}$  mode



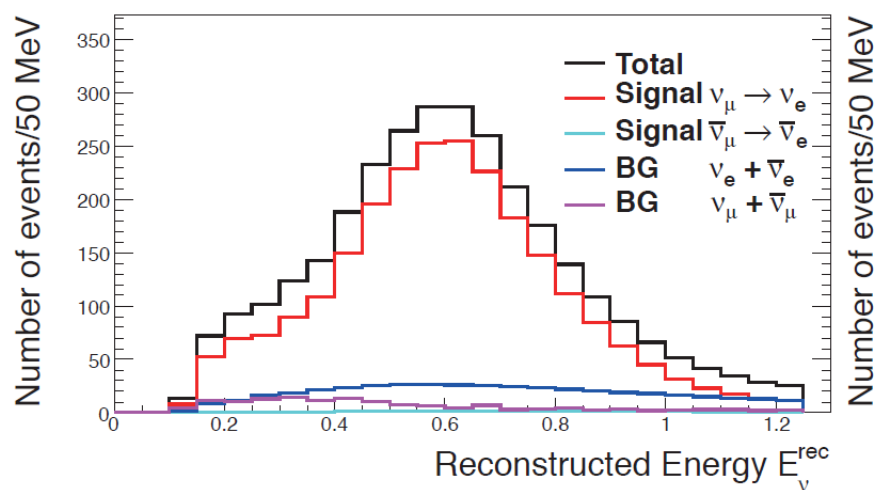
DUNE  
(7 yrs total)



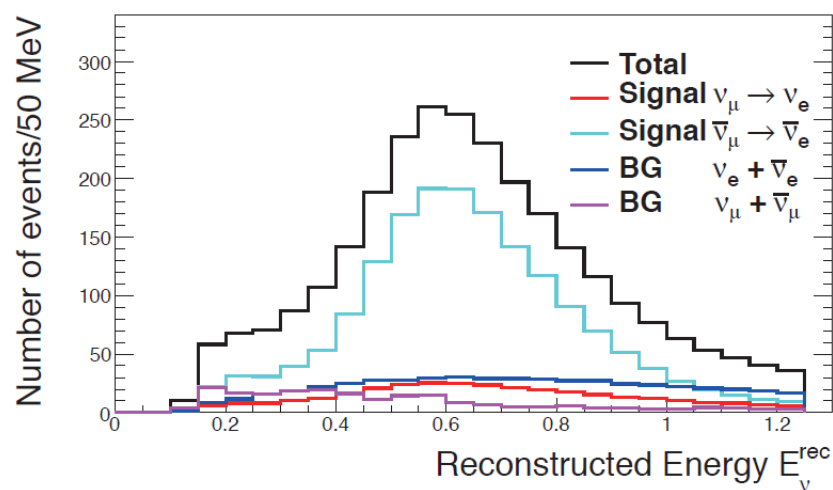
# Appearance Spectra

Hyper-K  
(10 yrs total)

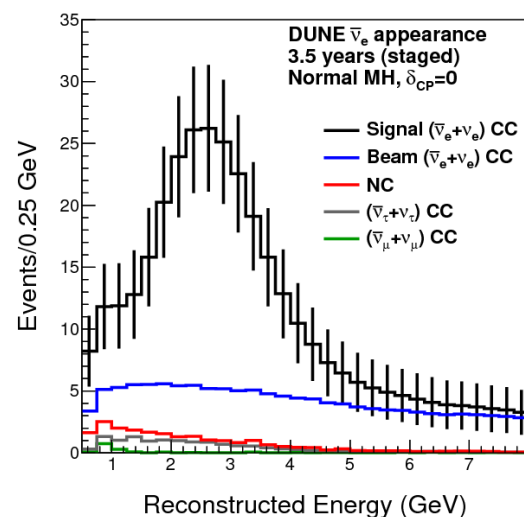
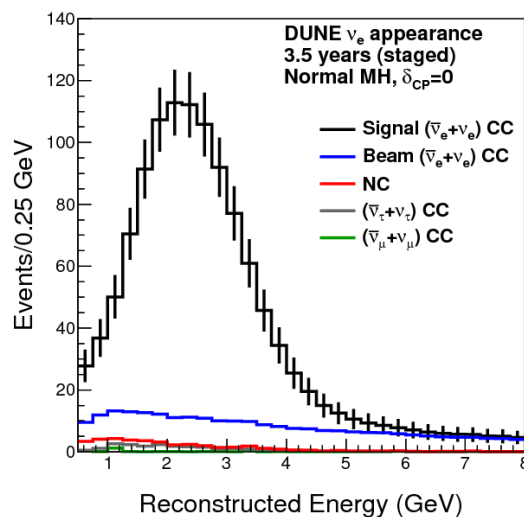
Appearance  $\nu$  mode



Appearance  $\bar{\nu}$  mode

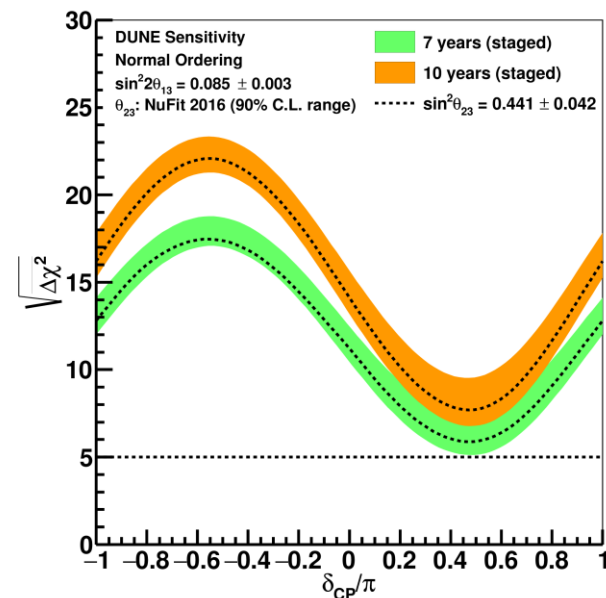


DUNE  
(7 yrs total)



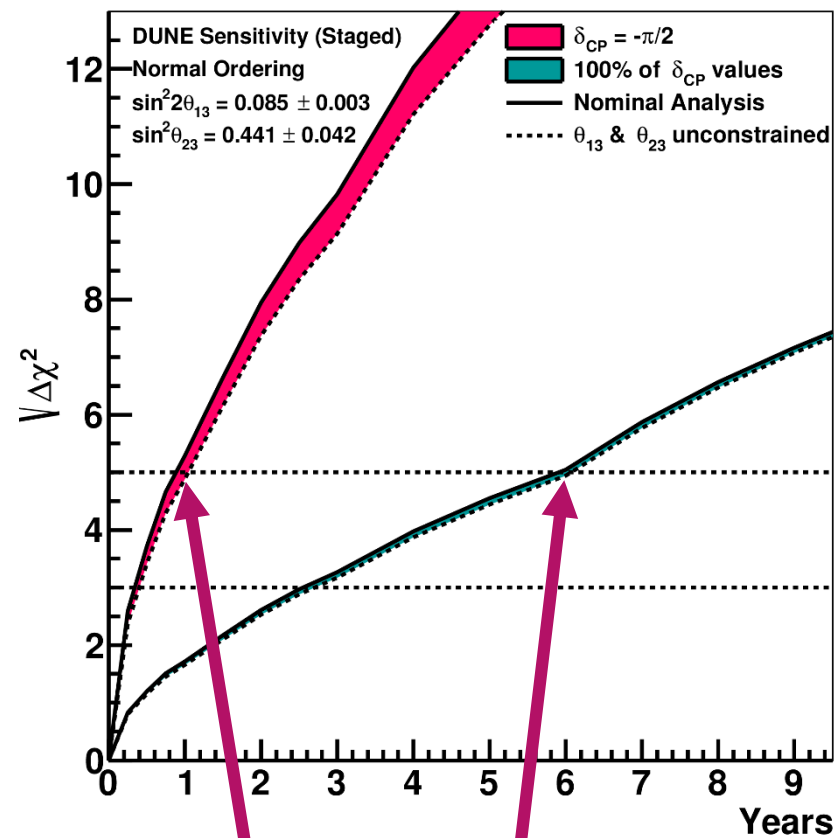
# Mass Hierarchy

DUNE Mass Hierarchy Sensitivity



Width of band indicates variation in sensitivity for  $\theta_{23}$  values in the NuFit 2016 90% C.L. range

DUNE MH Sensitivity



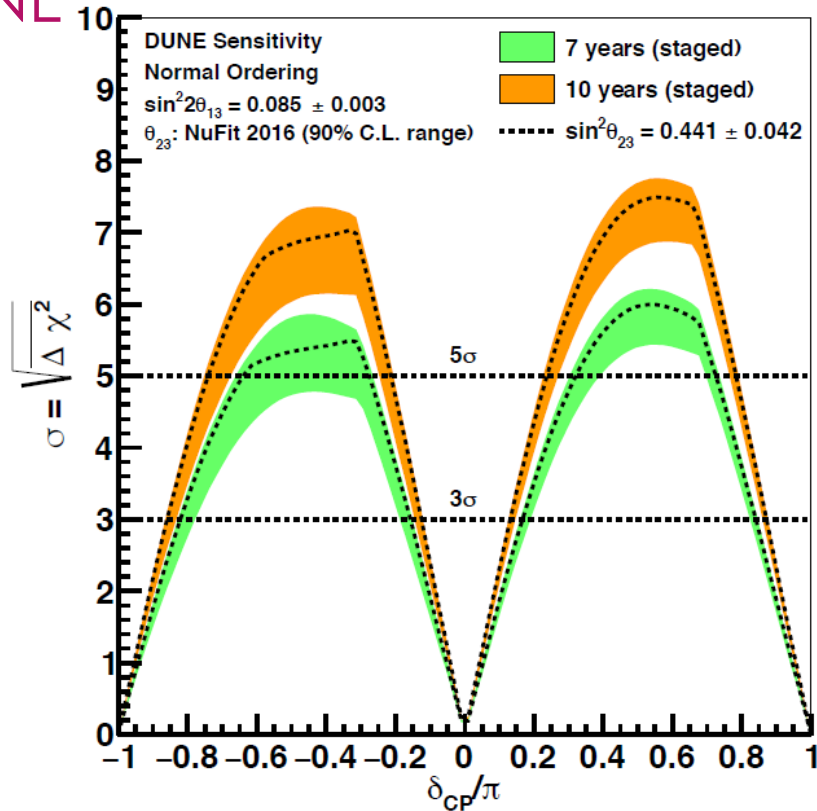
$\sim 1$  yr

$\sim 6$  yrs

- ▶ DUNE should be able to make a relatively quick determination of the mass hierarchy
- ▶ Hyper-K is less sensitive due to the shorter baseline
  - ▶ Combined analysis with beam and atmospheric neutrinos leads to  $>3\sigma$  determination in about 5 years with one tank Hyper-K

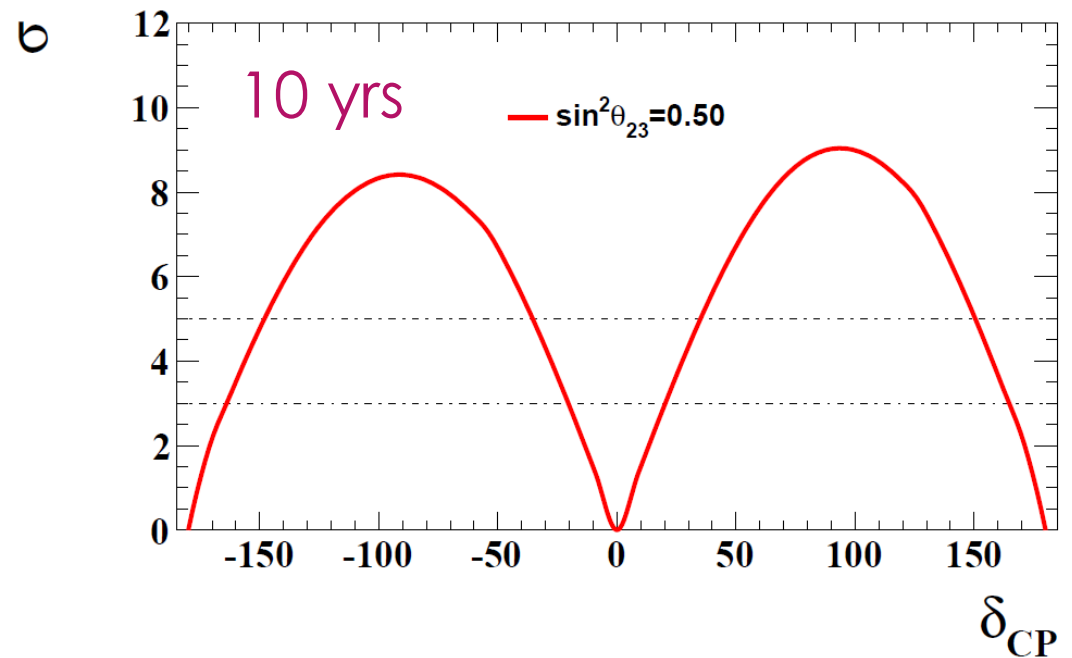
# CP Violation

DUNE



Hyper-K

## Normal mass hierarchy

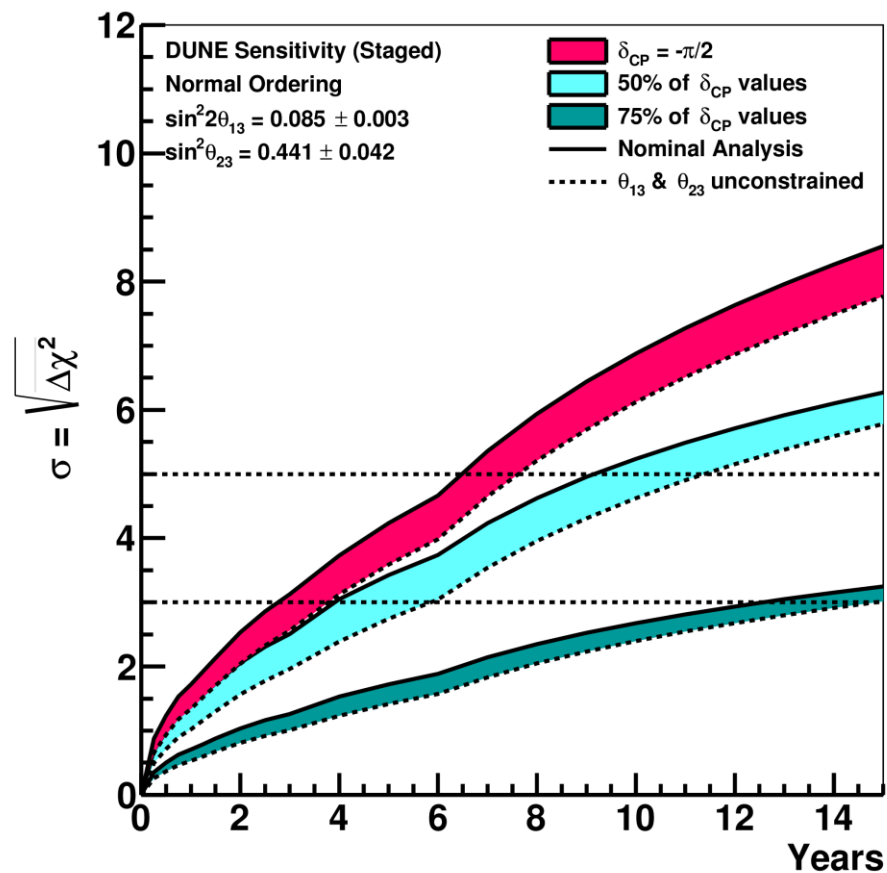


Width of band indicates variation in sensitivity for  $\theta_{23}$  values in the NuFit 2016 90% C.L. range

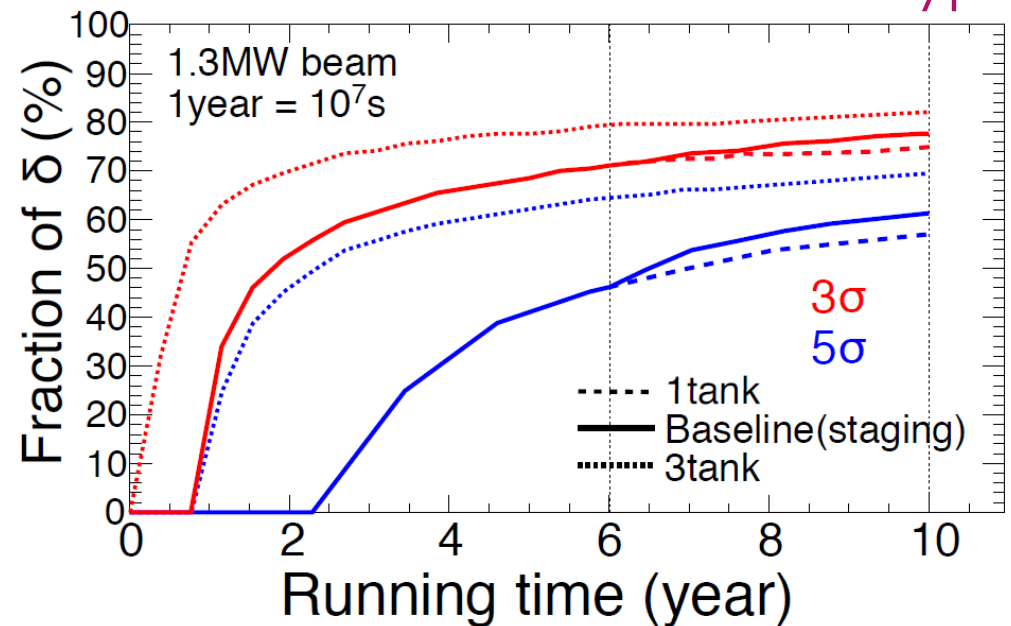
# CP Violation

DUNE

CP Violation Sensitivity



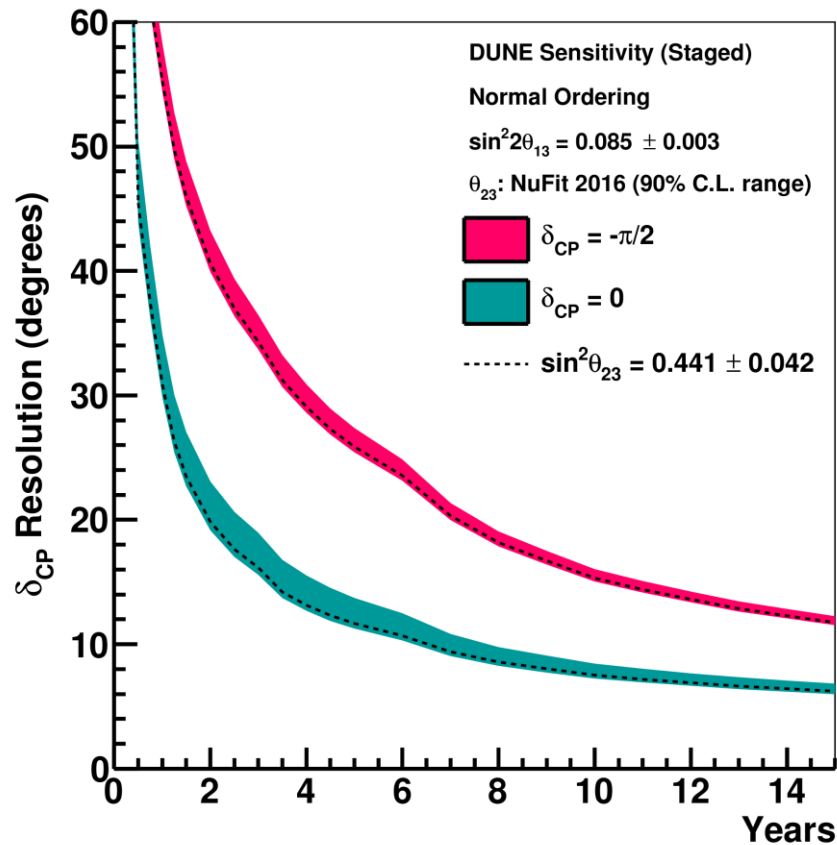
Hyper-K



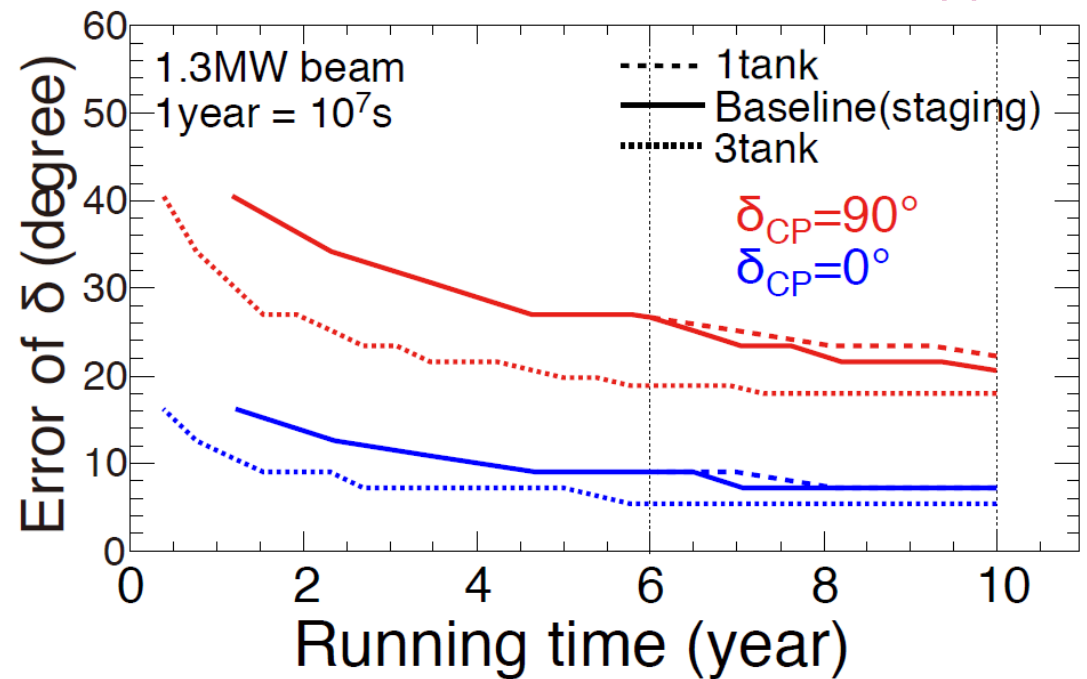
- ▶ Ultimately sensitivities of 5 $\sigma$  with 50% CP coverage and 3 $\sigma$  75% CP coverage

# CP Phase

DUNE

 $\delta_{CP}$  Resolution

Hyper-K

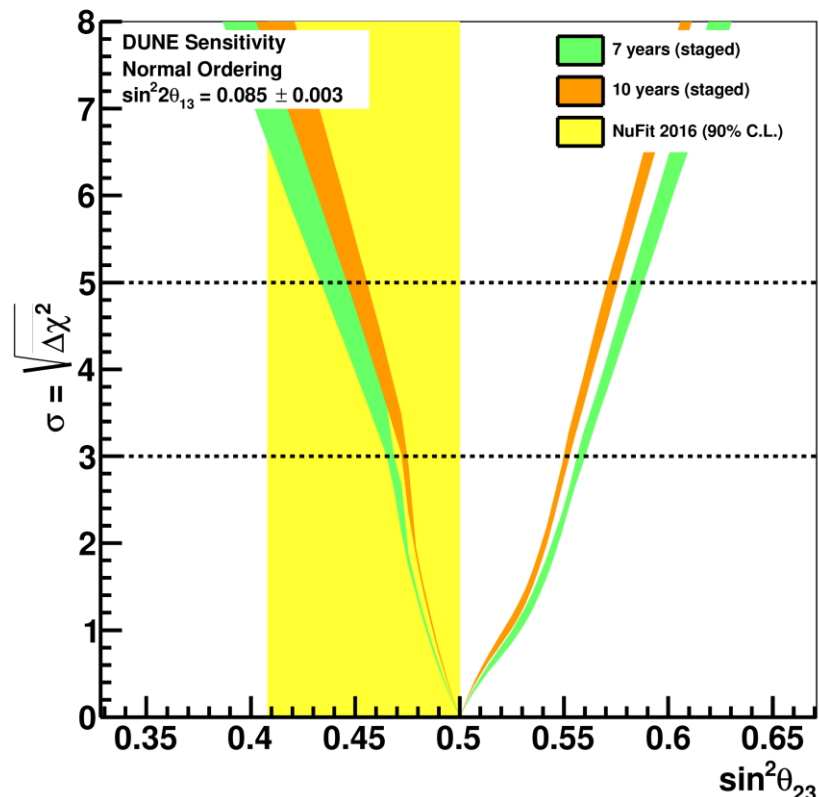


- Resolution on measurement of  $\delta_{CP}$  of  $\sim 10$ - $20^\circ$

# Octant

DUNE

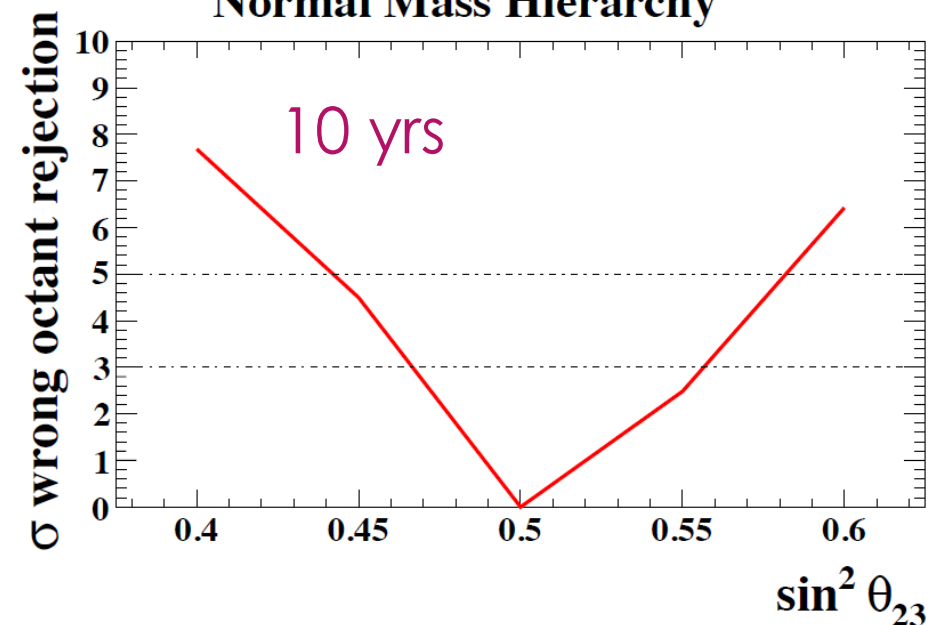
Octant Sensitivity



Width of band indicates variation in sensitivity for different  $\delta_{CP}$  values

Hyper-K

Normal Mass Hierarchy



- ▶ Potential to reject maximal mixing at  $3\sigma$  or  $5\sigma$  in the range of the current global best fit
- ▶ Enhanced sensitivity with combined beam and atmospheric neutrino data

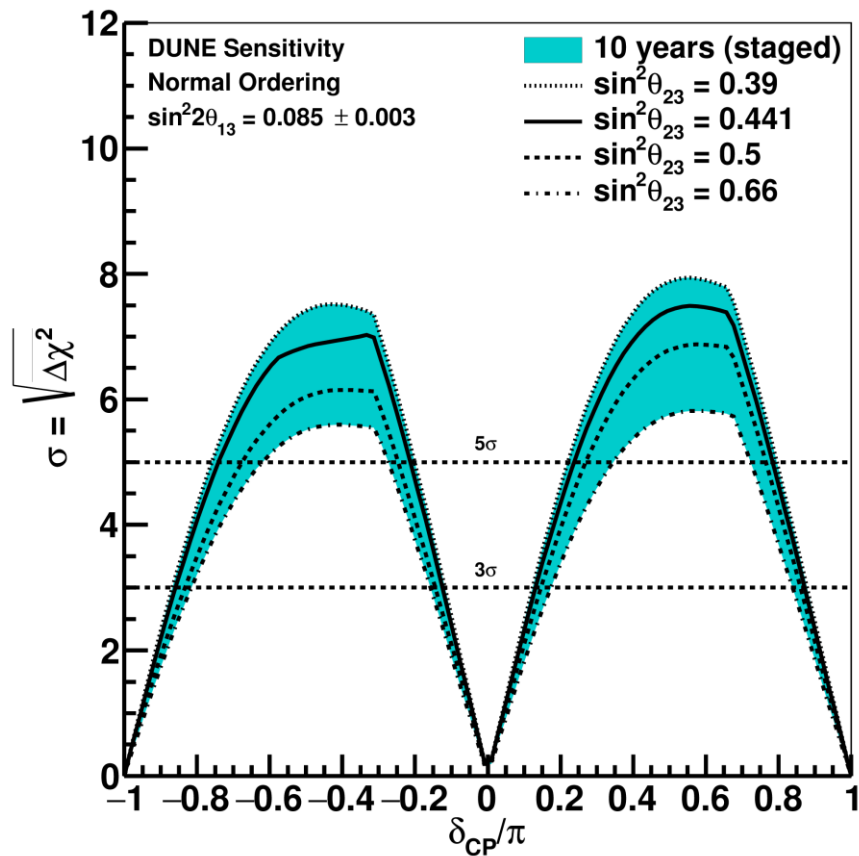
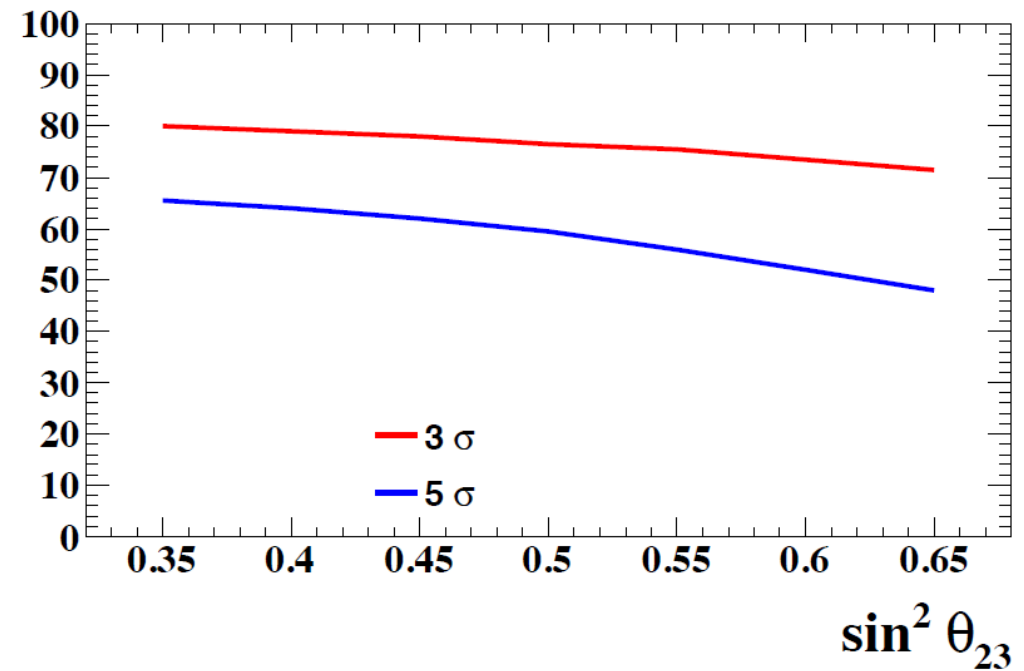


# Effect of $\theta_{23}$ on CP

DUNE

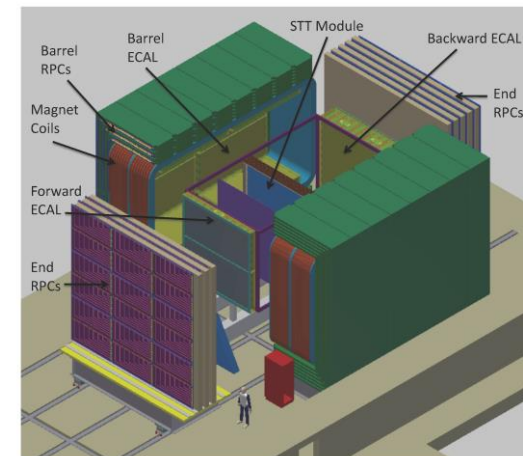
Hyper-K

CP Violation Sensitivity

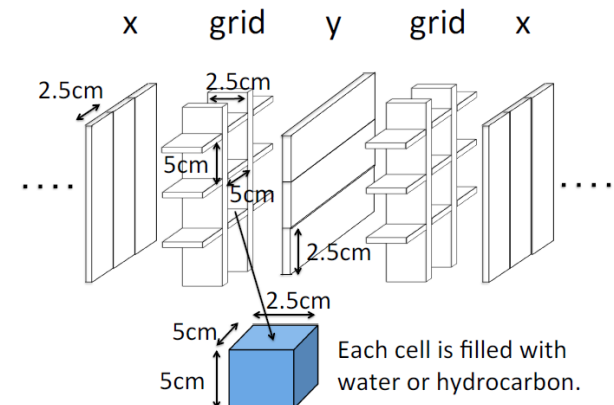
Fraction of  $\delta$  (%)

# How do we get there?

- ▶ Near detectors
  - ▶ DUNE near neutrino detector design is on-going
    - ▶ Fine-grained tracker with multiple targets including argon gas
    - ▶ LAr-TPC + fine-grained tracker
  - ▶ Hyper-K near neutrino detector development
    - ▶ Possible upgrades to T2K ND280
    - ▶ WAGASCI
    - ▶ Intermediate (1-2 km baseline) water Cherenkov detector
- ▶ Both will have muon monitors

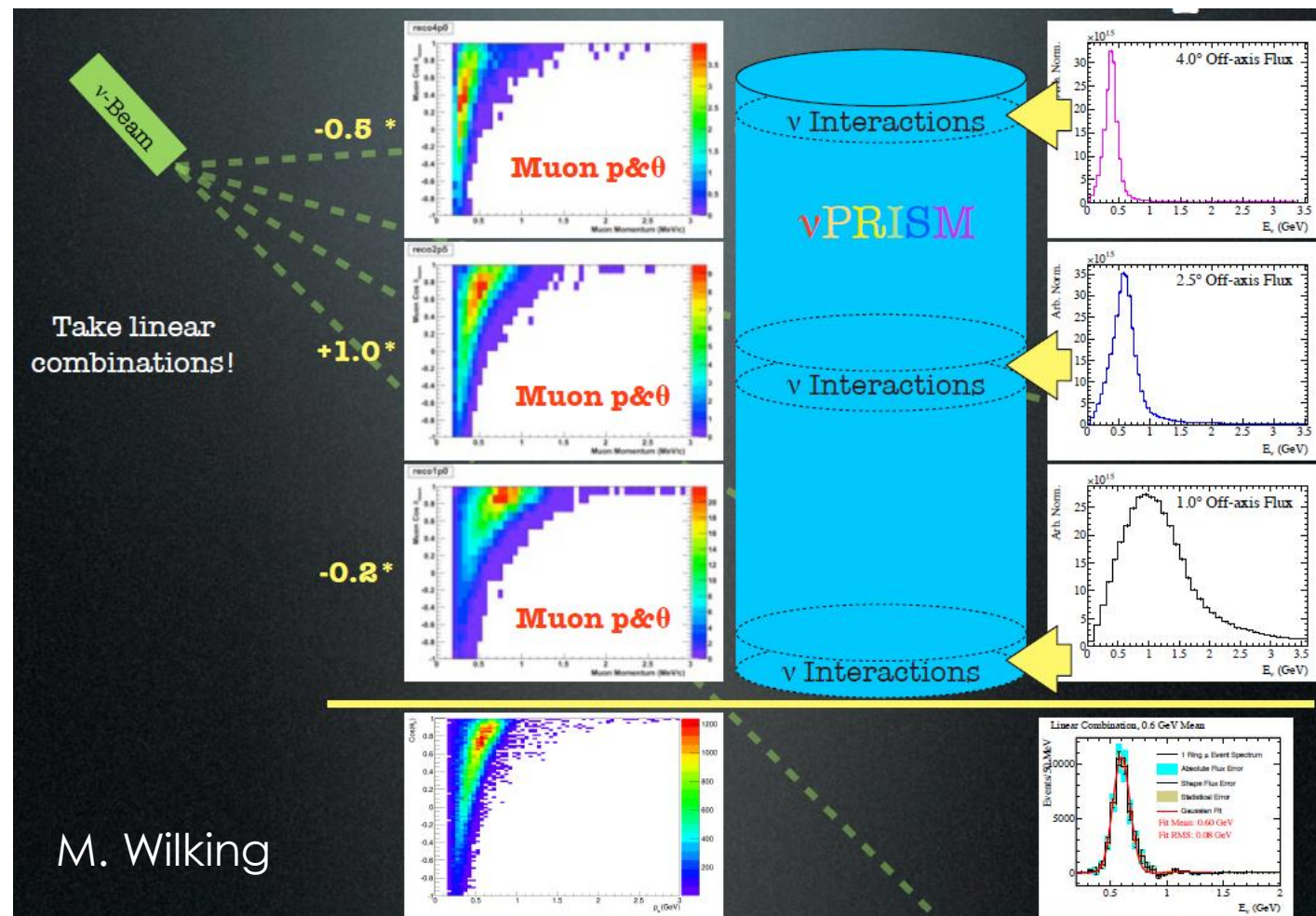


DUNE  
CDR  
Design



WAGASCI

# NuPRISM Concept



- ▶ Status: NuPRISM and TITUS collaborations have merged to form E61 collaboration
- ▶ An intermediate distance water Cherenkov detector with Gd loading
- ▶ 50 m tall, 8m diameter, movable instrumented portion
- ▶ In the proposal stage with J-PARC PAC
- ▶ “Phase 0” date 2021

B. Jamieson, NuInt2017

# How do we get there?

- ▶ Program of supporting measurements
  - ▶ Hadron production measurements: NA61/SHINE considering possibilities for running after the CERN 2019-2020 shutdown
  - ▶ Neutrino Interaction data: ArgoNeuT, MINERvA, NOvA ND, T2K ND280, MicroBooNE, SBND, ICARUS, NuPRISM, ...
  - ▶ Improved models for neutrino interaction generators (GENIE, NuWro, NEUT, ...)
  - ▶ Detector prototypes and calibration measurements (LArLAT, Mini-CAPTAIN, DUNE 35t, ProtoDUNE, ANNIE, ...)

# Supporting Measurements

## ANNIE

- ▶ 26 ton Gd-loaded water Cherenkov detector located 100m from the Fermilab Booster Neutrino Beam (BNB) target
- ▶ Will measure final state neutron multiplicity as a function of topology and kinematics in the 0.5 - 3 GeV range
- ▶ First application of LAPPDs (Large Area Picosecond Photo Detectors) in particle physics
- ▶ Physics data in 2018

M. O'Flaherty, NuInt2017

## CAPTAIN

- ▶ CAPTAIN is a liquid argon TPC (LArTPC) designed to make measurements relevant for the DUNE experiment
- ▶ Mini-CAPTAIN (400 kg instrumented LAr) taking data with Los Alamos neutron beam now
- ▶ Goal is to measure neutron-argon cross section up to 800 MeV

## LArIAT

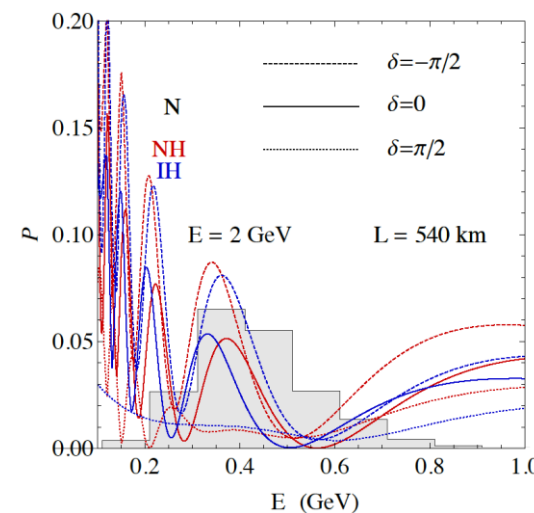
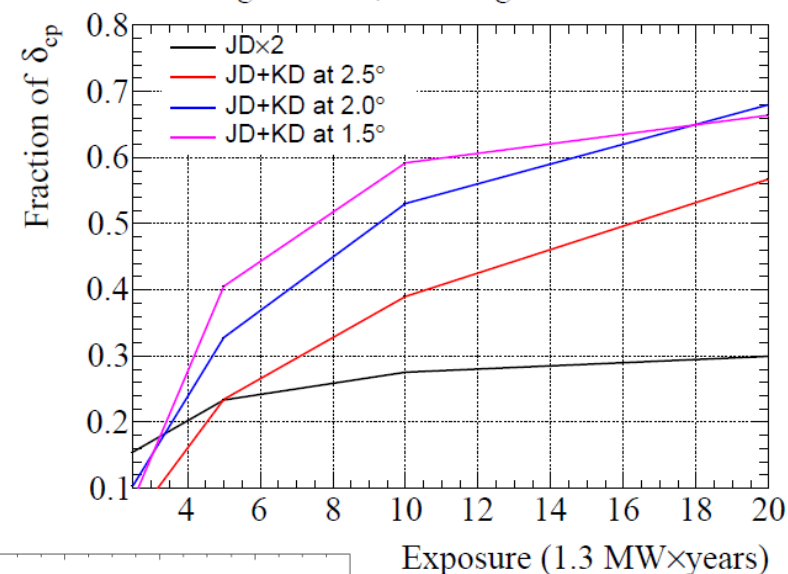
- ▶ Designed to characterize LArTPC performance and measure charged particle interactions in an energy range relevant for current and future LArTPC neutrino experiments
- ▶ 0.25 tons of LAr at Fermilab's Test Beam Facility
- ▶ Just completing its 3<sup>rd</sup> physics run

J. Asaadi, NuInt2017

# Other future possibilities

- ▶ T2HKK
  - ▶ Second Hyper-K tank placed in Korea, at a 1100-1300 km baseline and 1-3 degree off-axis angle from J-PARC
  - ▶ Improves sensitivity to leptonic CP violation, neutrino mass ordering as well as nonstandard neutrino interactions.
- ▶ European Spallation Source Neutrino Super Beam
  - ▶ 5 MW, 2 GeV proton beam by 2023
  - ▶ Coverage of 2<sup>nd</sup> oscillation maximum
  - ▶ Design Study is needed

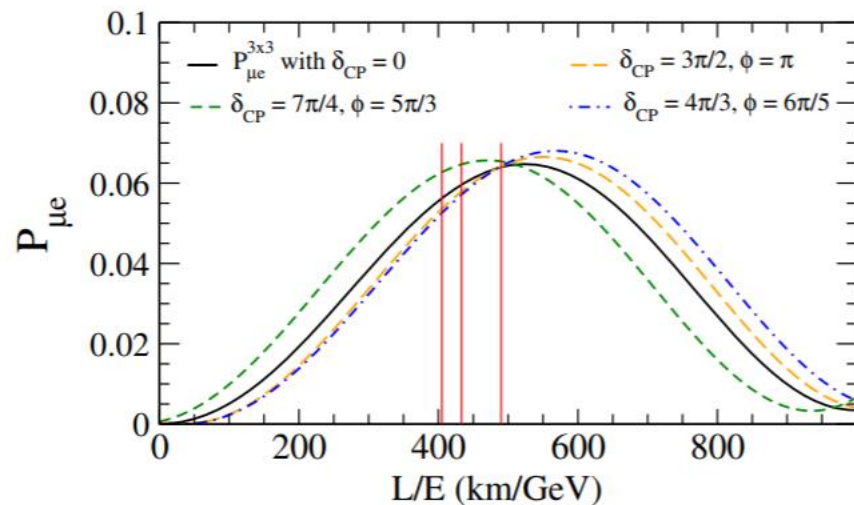
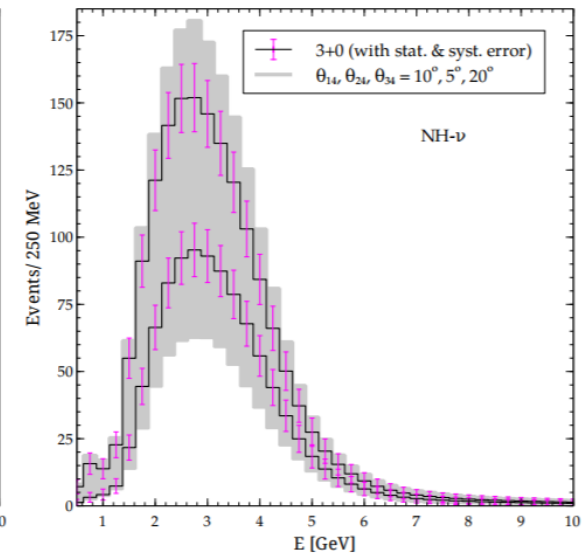
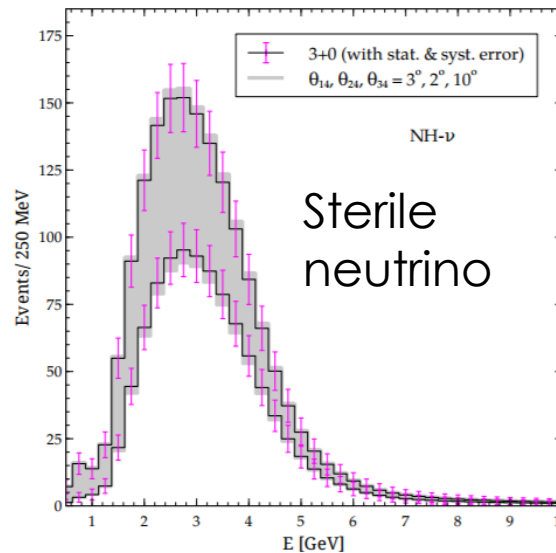
5 $\sigma$  Significance, Ordering Unknown



M. Dracos, WIN 2017

# What about new physics?

- ▶ Future LBL experiments will explore beyond-the-standard model topics such as
  - ▶ Sterile neutrinos
  - ▶ Non-standard interactions
  - ▶ Lorentz and CPT-violation
  - ▶ And more...
- ▶ And new physics could introduce ambiguities of the measurements of CP violation



PRL 117, 061804 (2016)

Non-unitary mixing matrix

JHEP11(2016)122

# Related talks

- ▶ CAPTAIN: Current Neutron and Future Stopped Pion Neutrino Measurements (me)
- ▶ Hyper-Kamiokande (H. Tanaka)
- ▶ The 2nd Hyper-Kamiokande detector in Korea (Seon-Hee Seo)
- ▶ The DUNE Far Detector (E. Kemp)
- ▶ The Hyper-K near detector programme (J. Wilson)
- ▶ Sensitivity of the DUNE Experiment to CP Violation (me)
- ▶ Measurements of the Neutrino Flux Using the DUNE Near Detector (J. Sinclair)
- ▶ Results from ANNIE Phase 1 and Plans for Phase 2 (R. Svoboda)
- ▶ (Apologies if I missed any!)



# Summary

- ▶ Future long-baseline neutrino experiments will address the remaining questions about 3-neutrino mixing, including the neutrino mass ordering, CP violation, and maximal mixing
  - ▶ DUNE: Far site construction and detector prototypes are underway this year
  - ▶ Japan: Program of beam, detector, and analysis upgrades to go from T2K era to Hyper-K era
  - ▶ Different baselines, neutrino energy range, and detector technology make the measurements complementary
- ▶ Program of supporting measurements underway to support every aspect of these experiments
- ▶ Stay tuned!



Check out related talks in  
the parallel sessions!

# Backup

# Neutrino Mixing

3x3 mixing between flavor states and mass states

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U_{PMNS} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Three mixing angles and one phase that affect neutrino oscillations

T2K (arXiv:1701.00432)

$$\sin^2 \theta_{23} = 0.532^{+0.046}_{-0.068}$$

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix}$$

$$\sin^2 \theta_{23} = 0.404^{+0.030}_{-0.022}$$

$$\sin^2 \theta_{23} = 0.624^{+0.022}_{-0.030}$$

NOvA (arXiv:1701.05891)

$$U_{PMNS} =$$

$$\begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$\sin^2 2\theta_{13} = 0.0841 \pm 0.0033$$

Daya Bay  
(arXiv:1610.04802)

$$\tan^2 \theta_{12} = 0.443^{+0.030}_{-0.025}$$

Solar + KamLAND  
(PRC 88, 025501 2013)

$$\delta = ?$$

# Neutrino Oscillations

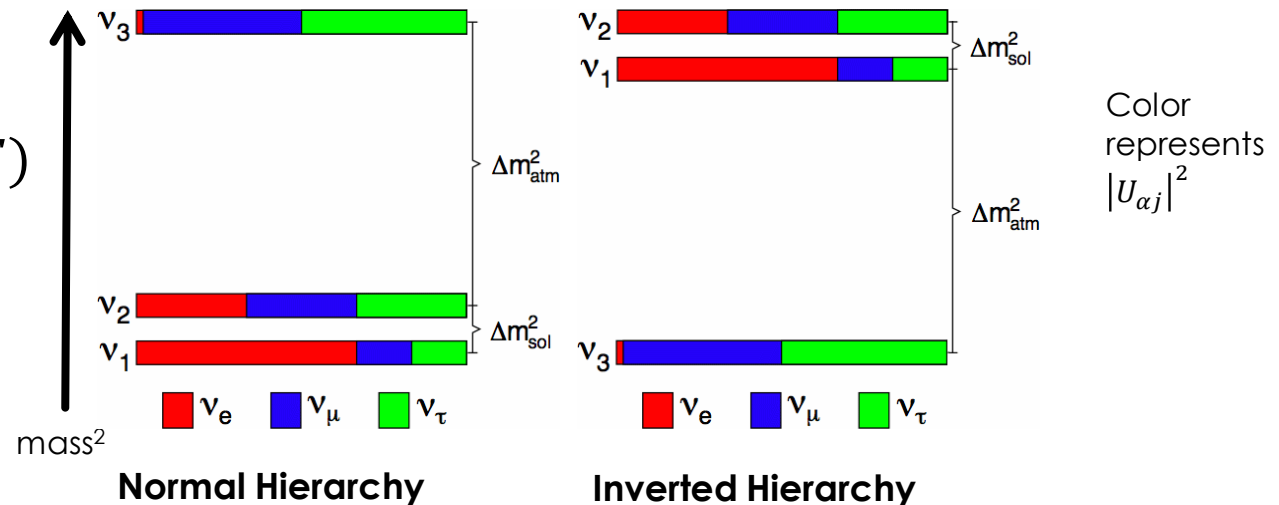
Probability of flavor change  
(2-flavor approximation):

$$P_{\alpha \rightarrow \beta} = \sin^2 2\theta \sin^2(1.27 \Delta m^2 L/E)$$

$L$  = baseline

$E$  = neutrino energy

$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$



$$\Delta m_{21}^2 = 7.46_{-0.19}^{+0.20} \times 10^{-5} \text{eV}^2$$

Solar + KamLAND

PRC 88, 025501 2013

$$|\Delta m_{32}^2| = 2.45_{-0.08}^{+0.08} \times 10^{-3} \text{eV}^2$$

Daya Bay (arXiv:1610.04802)

$$|\Delta m_{32}^2| = 2.545_{-0.084}^{+0.081} \times 10^{-3} \text{eV}^2$$

T2K (arXiv:1701.00432)

(Assuming NH)

# Muon Neutrino to Electron Neutrino Oscillations in Matter

## Oscillations in matter with constant density:

H. Nunokawa, S. J. Parke, and J. W. Valle, Prog.Part.Nucl.Phys. 60, 338 (2008), arXiv:0710.0554 [hep-ph].

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) \approx & \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2(\Delta_{31} - aL)}{(\Delta_{31} - aL)^2} \Delta_{31}^2 \quad (\theta_{13} \text{ term}) \\
 & + \sin 2\theta_{23} \sin 2\theta_{13} \sin 2\theta_{12} \frac{\sin(\Delta_{31} - aL)}{(\Delta_{31} - aL)} \Delta_{31} \frac{\sin(aL)}{(aL)} \Delta_{21} \cos(\Delta_{31} + \delta_{CP}) \quad (\text{Interference term}) \\
 & + \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(aL)}{(aL)^2} \Delta_{21}^2 \quad (\text{solar term})
 \end{aligned}$$

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{4E_\nu}$$

For antineutrinos,  $a \rightarrow -a$  and  $\delta \rightarrow -\delta$

$$a = \frac{G_F N_e}{\sqrt{2}}$$

Matter effect: electron neutrinos have CC interactions with electrons in the Earth, while other flavors don't

$$\cos(\Delta_{31} + \delta_{CP}) = \cos \Delta_{31} \cos \delta_{CP} - \sin \Delta_{31} \sin \delta_{CP}$$

CP-violation