

Energy response and position reconstruction in the DEAP-3600 dark matter experiment

TAUP 2017, Sudbury

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on behalf of the DEAP-3600 collaboration

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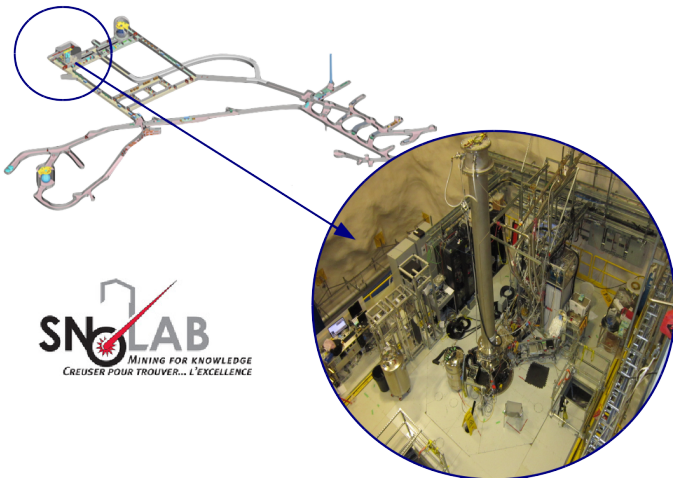
Laurentian University

24/07/2017





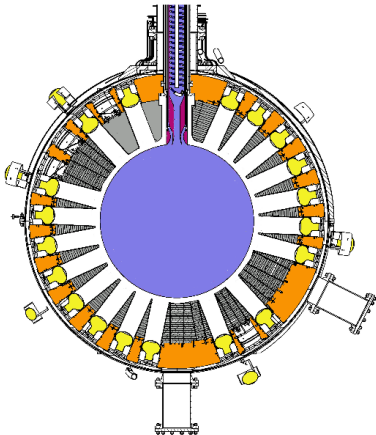
Located at SNOLAB, 2 km underground at 6000 mwe:



SNOLAB
MINING FOR KNOWLEDGE
CREUSER POUR TROUVER... L'EXCELLENCE



The experiment:



- Pixelated detector
- capable to hold 3600 kg LAr target material, currently filled to 3260 kg
- 255 PMTs to measure energy and position of events in the LAr
- AV coated with wavelength shifter TPB
- Detection of WIMPs via nuclear recoils with a target sensitivity to WIMP-nucleon cross section 10^{-46} cm^2 at WIMP masses of 100 GeV

Running stable since November
2016



Radioactive Calibration Sources

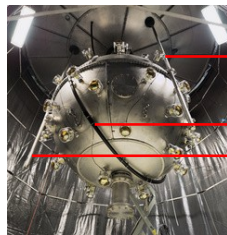
^{39}Ar

Internal source

- β^- emitter with $Q = 565$ keV
- From cosmic ray interaction on ^{40}Ar
- Isotropically distributed in LAr

^{22}Na

External source



Cal E

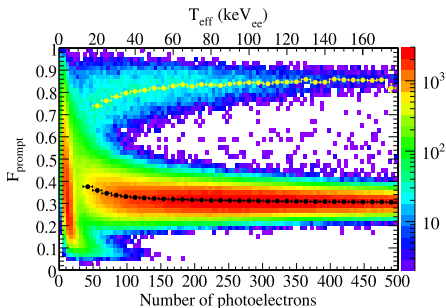
Cal F

Cal A



Discriminating the ^{39}Ar signal using PSD:

$$f_{\text{prompt}} = \frac{q_{\text{prompt}}}{q_{\text{event}}}$$



- Ar Dimer states with different life times:

- Singlet τ 6 ns -predominantly nuclear recoils
- Triplet τ 1500 ns -predominantly electromagnetic events

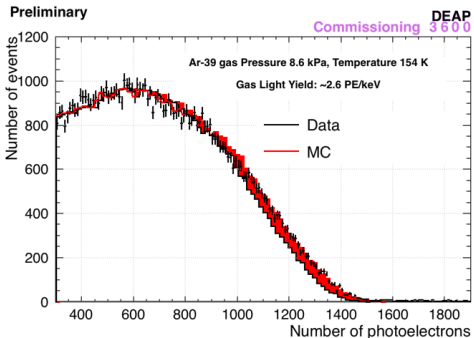
→ Percentage of light signal in prompt light as indication of singlet state population

DEAP-1 calibration data

Astroparticle Physics 85 (2016) 1-23



Understanding the energy response using ^{39}Ar , Gas phase calibration:

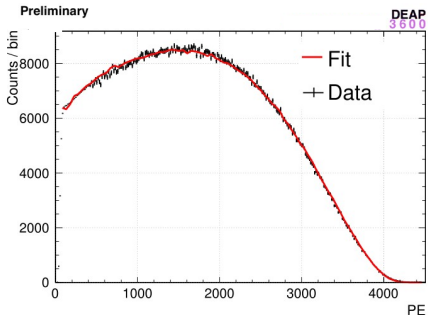


- Light yield uniformly scaled to match the simulation to data

Cool down phase, before fill



Understanding the energy response using ^{39}Ar , LAr phase calibration:

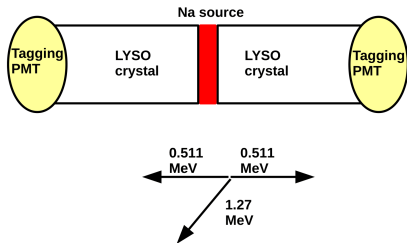


- Light yield uniformly scaled to match the simulation to data

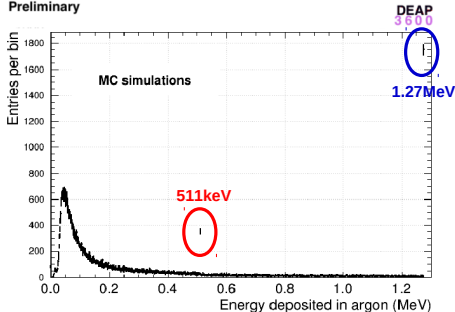
First fill data



External ^{22}Na source allows tagged monoenergetic gamma rays:



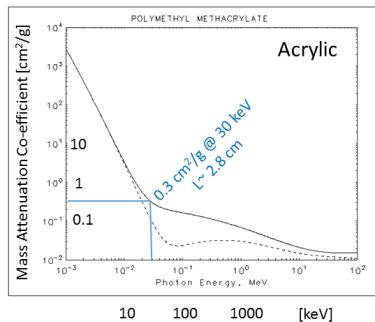
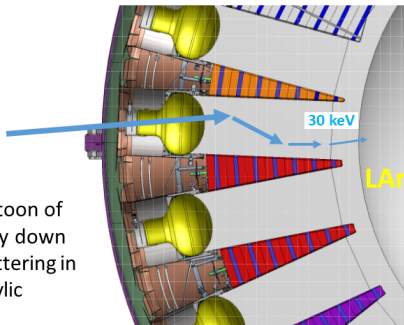
Preliminary





The low energy feature

Cartoon of γ -ray down scattering in acrylic

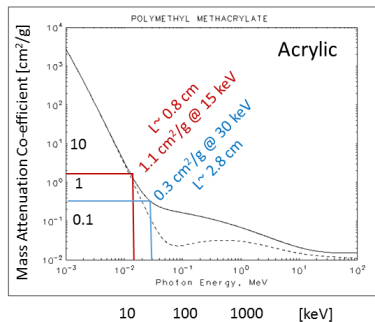
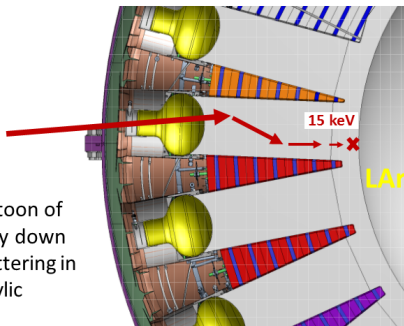


Plot and data from NIST.gov
X-ray mass attenuation coefficients



The low energy feature

Cartoon of γ -ray down scattering in acrylic



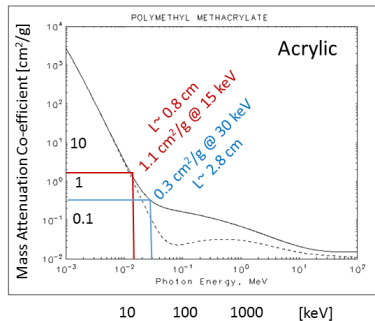
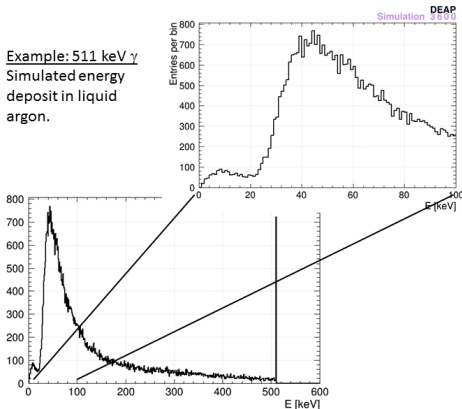
Plot and data from NIST.gov
X-ray mass attenuation coefficients



The low energy feature

Both the Rising and Falling Edge in Distribution Energy Deposit Arise from Electromagnetic Physics

Example: 511 keV γ
Simulated energy
deposit in liquid
argon.

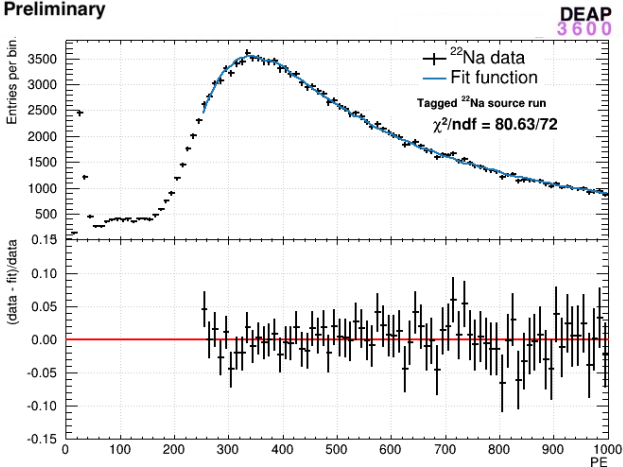


Plot and data from NIST.gov
X-ray mass attenuation coefficients



Fits on the ^{22}Na spectrum: Fit on low energy feature

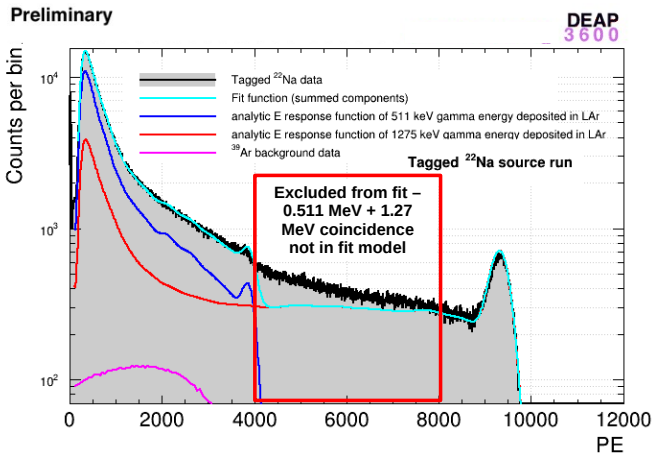
Preliminary

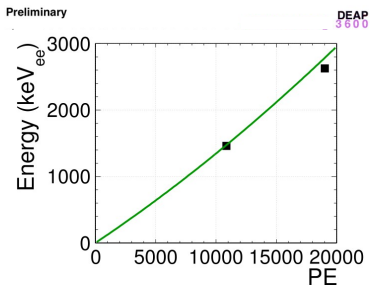




Fits on the ^{22}Na spectrum:

Fit on full spectrum: consistency check only



Combining ^{39}Ar and ^{22}Na :

Saturation effects at high energies not yet accounted for

WIMP ROI: 80 – 240 PE

$$c_0 + c_1\text{PE} + c_2\text{PE}^2$$

Preliminary light yield:

$$LY = 7.36_{-0.52}^{+0.61}(\text{fit syst.}) \pm 0.22(\text{SPE syst.})\text{PE}/\text{keV}_{ee} @80 \text{ PE}$$



Measurement of event position:

Two main approaches possible:

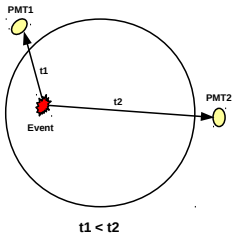
- Time-based
- Charge-based



Measurement of event position:

Two main approaches possible:

- Time-based



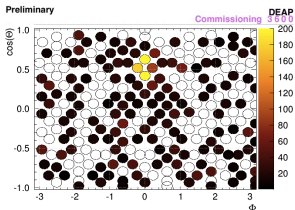
- Finite speed of light
- PMT hit time proportional to source distance from PMT
- Absolute vertex resolution uniform across volume
- Dependent on scintillator response times, PMT transit time, DAQ quality



Measurement of event position:

Two main approaches possible:

- Charge-based



- Charge patterns of the PMTs
- Point-like source: closer PMTs expected to have more photon hits and charges
- Pattern detector dependent
- Vertex resolution improved towards the edge of the detector



Measurement of event position:

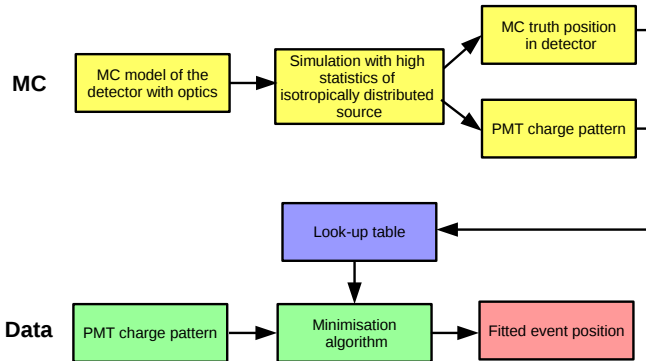
Two main approaches possible:

- Time-based
- Charge-based

DEAP-3600 small enough for charge-based vertex reconstruction to deliver the better position resolution



How it is done:

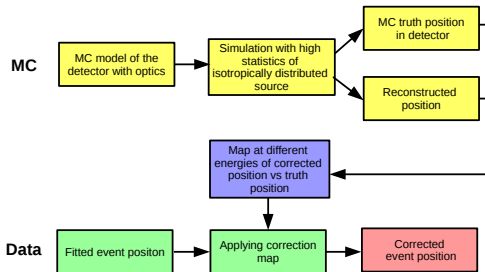


Work in progress!



Fiducialisation and de-biasing using ^{39}Ar :

- Isotropic ^{39}Ar distribution
- Map true radius to reconstructed radius
- Account for energy dependence



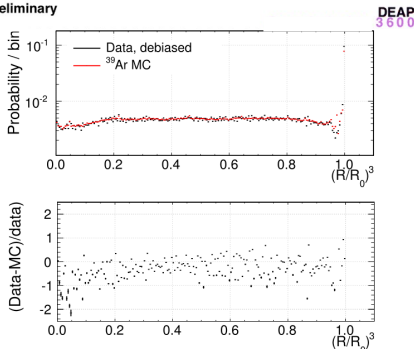


Fiducialisation and de-biasing using ^{39}Ar :

- Fiducial mass from activity of de-biased ^{39}Ar decay spectrum after applying fiducial cuts

consistent with 2222 kg of LAr

Preliminary



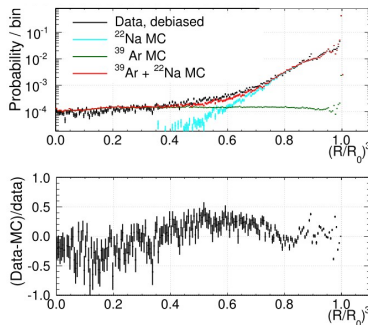
Work in progress!



^{22}Na studies to understand surface backgrounds:

- ^{22}Na low energy feature at low energies near ROI helps determine fiducial cut parameter

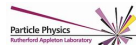
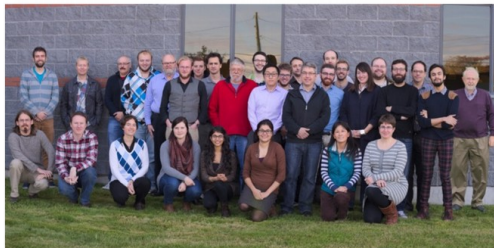
Preliminary

DEAP
3600

Work in progress



The Deap-3600 collaboration:



The speakers operational support was provided by NSERC



Back Up



DEAP-3600 calibration program:

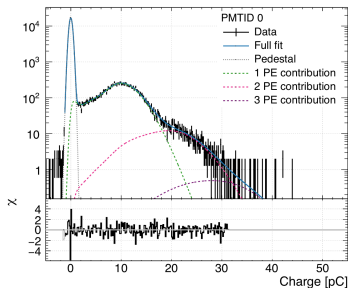
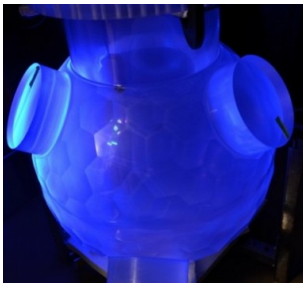
Calibration Source	Calibration goal	Notes
Laserball	Optical (PMT) calibration	vacuum runs only
LED Light Injection	Optical (PMT) calibration, monitoring	used in all run phases
^{22}Na	Energy and position reconstruction, gamma response	Argon phase
AmBe	Energy calibration, gamma and neutron response	Argon phase
^{39}Ar	Intrinsic, energy and position reconstruction	Argon phase

- Argon phase: gas phase (GAr), partial fill phase, liquid argon phase (LAr)
- LED Light Injection system with fibres installed on PMTs
- External calibration sources: ^{22}Na (1 MBq) and AmBe (74 MBq)
- Intrinsic calibration source: ^{39}Ar (expected 1.01 Bq/kg)



Single Photon counting:

Ideal measurement: single photon counting correcting for PMT effects
 De-excitation photons (128 nm) →
 TPB (420 nm) → Photoelectron
 cascades in PMTs

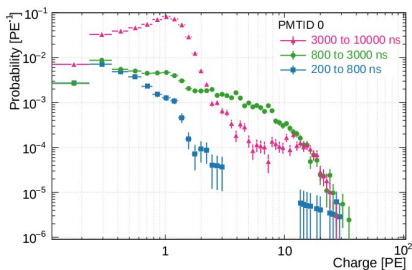


arXiv:1705.10183

- Translation of PMT pulses to number of photoelectrons observed using charge division (qPE)



Correction of different effects necessary:



Effects to correct on PE estimator:

● PMT effects:

- After-pulse (AP): caused by back-scatter of electrons on PMT dynodes
- Saturation of PMTs
- Dark noise

● Other effects:

- Pile-up of two or more events in same event window

arXiv:1705.10183