



Recent results from the ANTARES Deep Sea Neutrino Telescope

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on behalf of the ANTARES Collaboration

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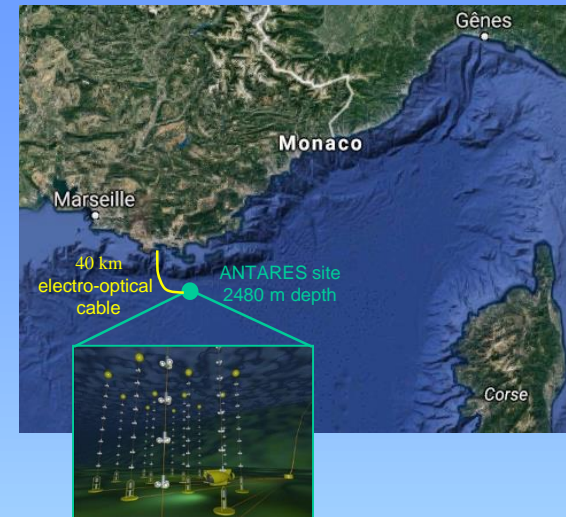
The ANTARES Neutrino Telescope

📖 NIM A 656 (2011) 11

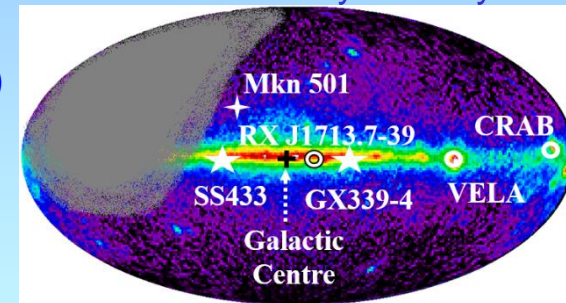
- **Largest underwater neutrino telescope** operating for 10 years now (complete in 2008)
- 12 line detector with 885 10" PMTs installed by 2500 m depth off the coast of Provence (France)
- **O(11000) neutrinos detected** with $E_\nu > 10 \text{ GeV}$
- **Excellent view of Galactic Centre region** with high angular resolution (0.3° - 0.4° median)
→ **interesting constraint of possible galactic component** of the IceCube HE signal
- **Real time** data processing → generation of alerts (~5s) for multi-messenger searches

➤ Science scope of ANTARES:

- Neutrino astrophysics, search for HE CR origin
- Multi-messenger observations
- Indirect searches for Dark Matter
- Atmospheric neutrinos (oscillations, sterile neutrinos)
- Exotic searches (magnetic monopoles, nuclearites)
- Earth & Sea Sciences, environmental studies



ANTARES sky visibility

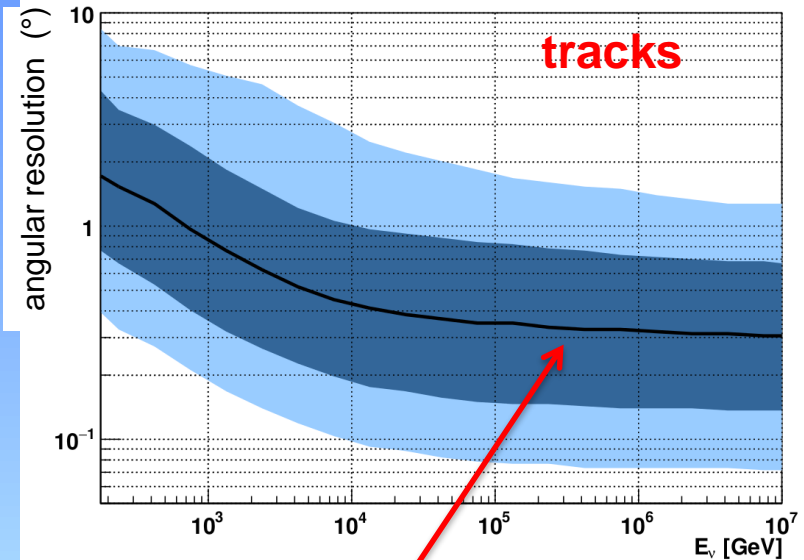


} This talk

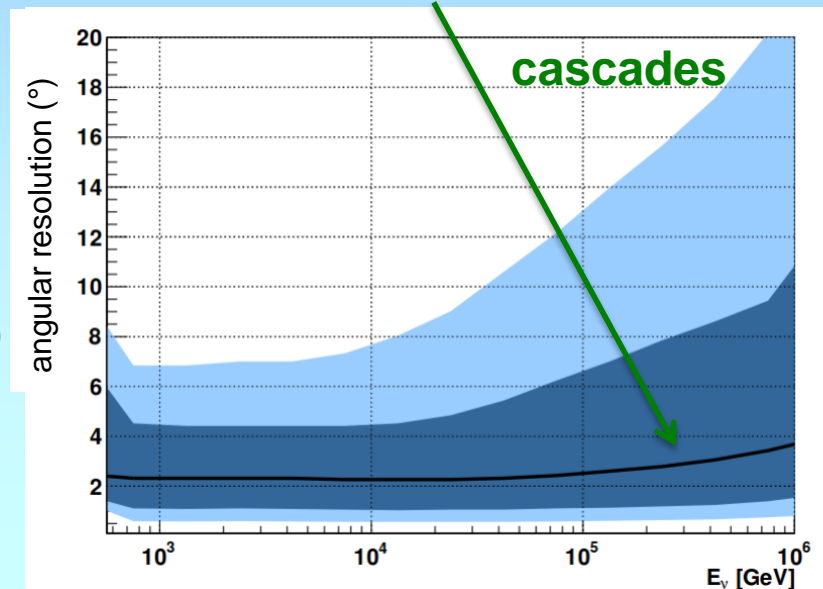


Reconstruction performances

- Upgoing **track events** (ν_μ CC)
 - Angular resolution $< 0.4^\circ$ for $E_\nu > 10$ TeV
 - Energy resolution : **factor 3**
 - **90% purity** of neutrinos
 - Large detection volume from μ range
→ ideal for neutrino astronomy
→ but large atmospheric μ bkg
-
- Upgoing **cascade events** (ν_e/ν_τ CC, NC)
 - Angular resolution $< 3^\circ$
 - Energy resolution for ν_e CC $< 10\%$
 - Contained events (small detection volume)
→ almost no atmospheric bkg

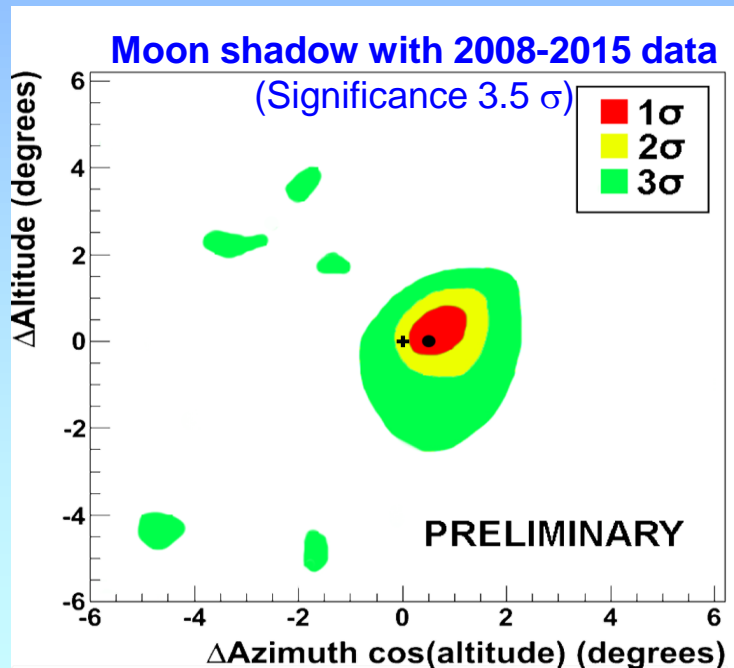
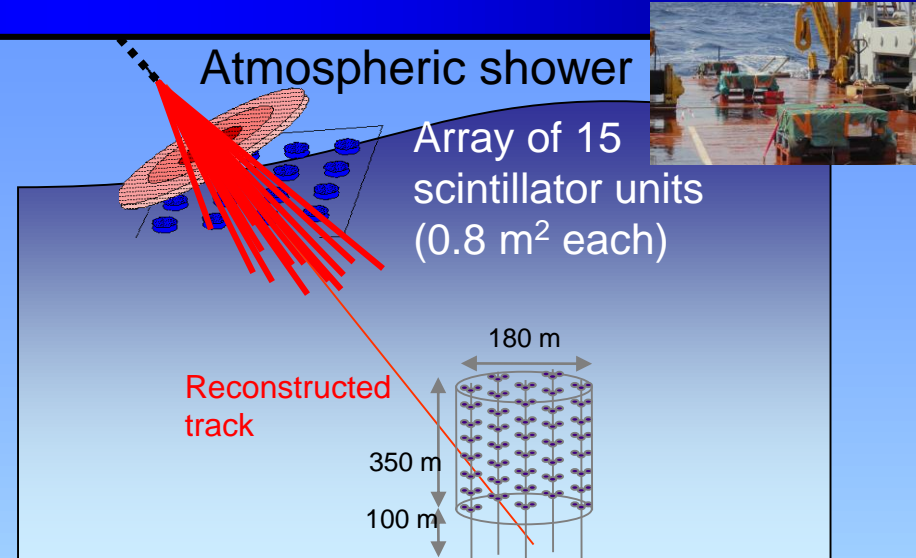
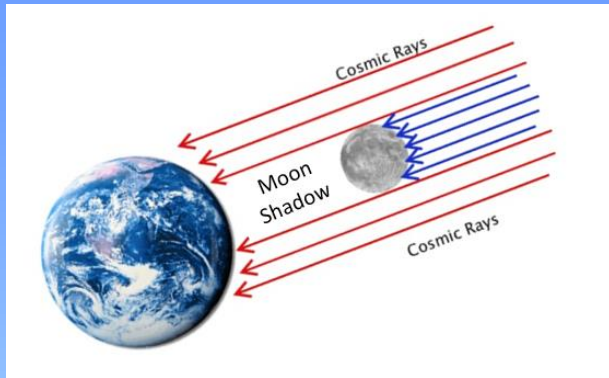


Median angular resolution vs E_ν



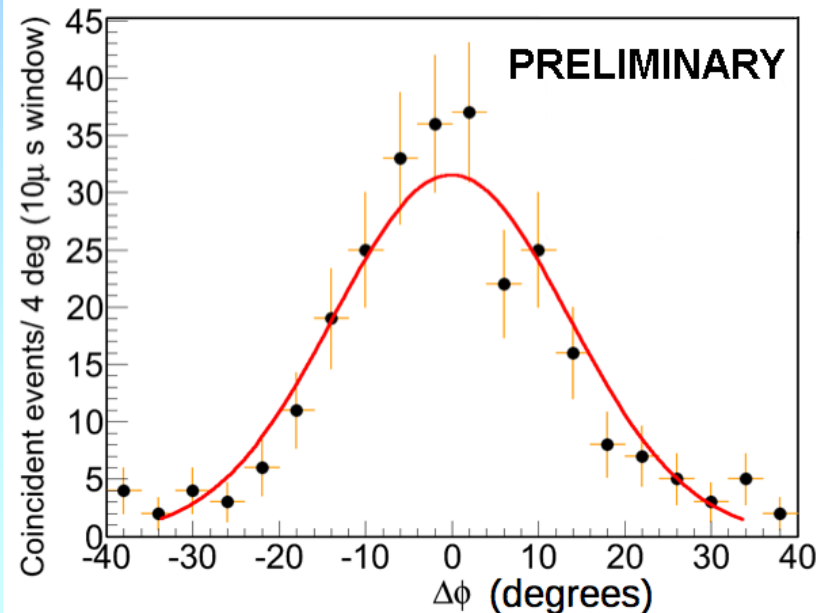


Check of Detector Absolute Pointing



Angular resolution for down-going muons :
 $0.73^\circ \pm 0.15^\circ$

Coincidence with Surface Array detector
(2 sea compans)



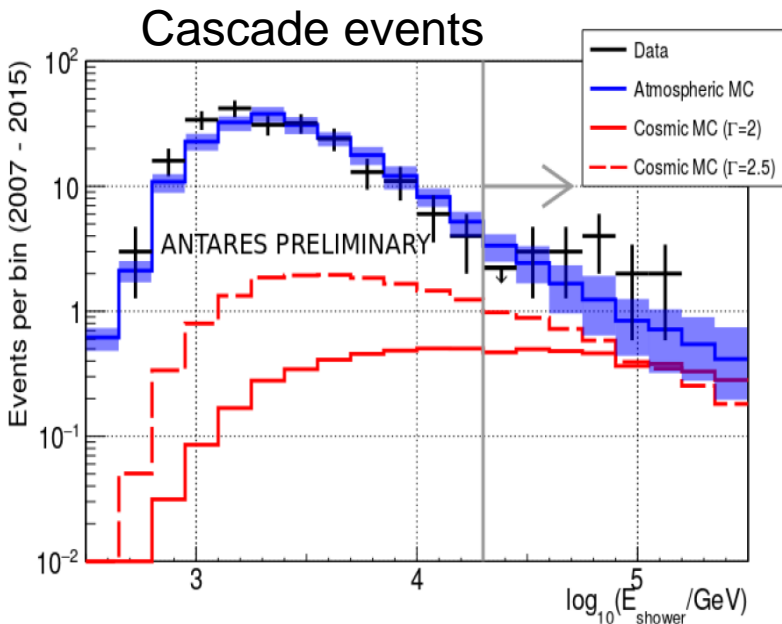
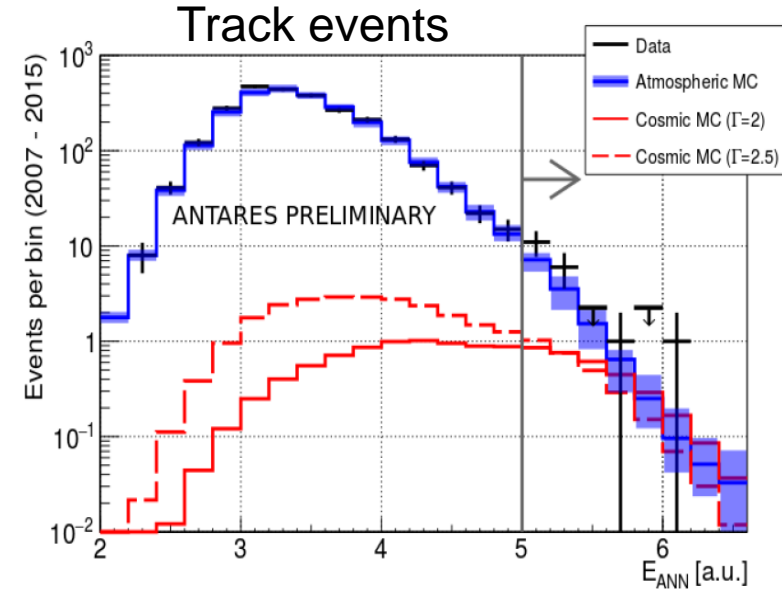


Search for HE Neutrino Excess (aka 'Diffuse Flux')

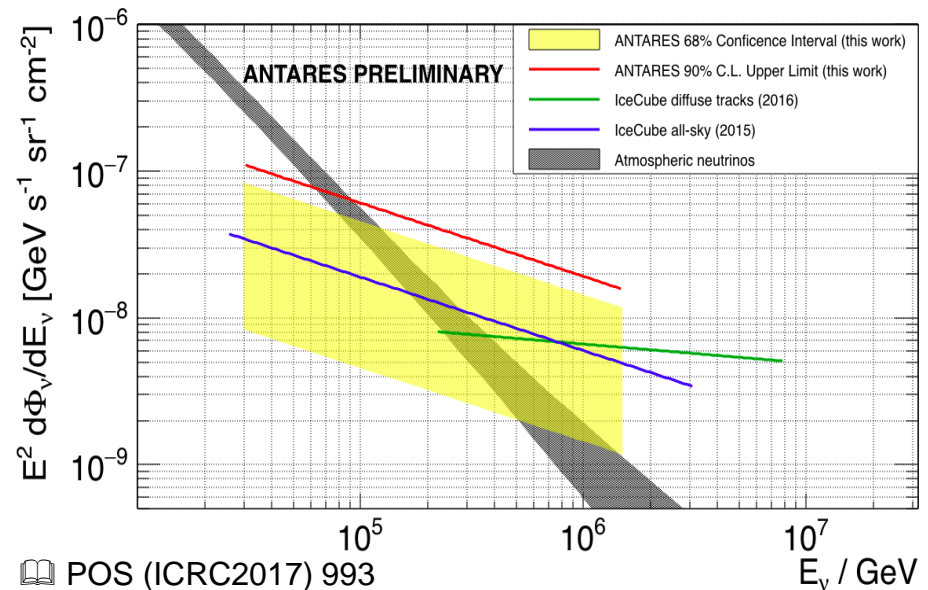
All sky/all flavour neutrino search with 2007-2015 data (2450 livedays)

- Total events above energy thresholds :
- Expected : 24 ± 7 bkg, ~ 8 from IceCube flux
 - Observed : 33

→ Small excess not significant but compatible with IceCube flux

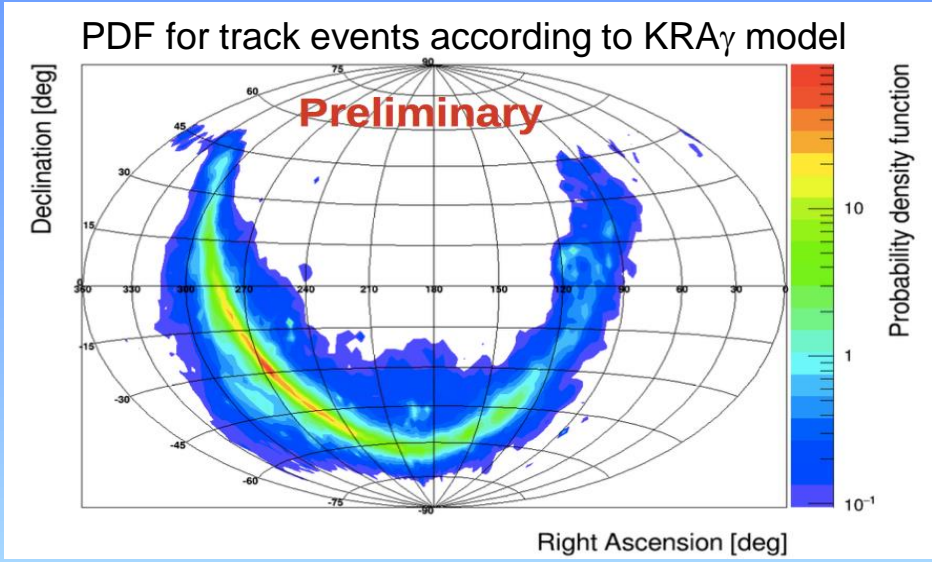


Combined upper limit assuming $E^{-2.5}$ spectrum





Study of Galactic Ridge



- Guaranteed galactic neutrinos from CR interactions with matter
- Does it contribute to IC flux?
- Test 'KRA γ ' model -> spectrum and morphology reproduce Fermi & Milagro data
- Analysis uses full model morphology and spectrum

New all-flavour analysis :

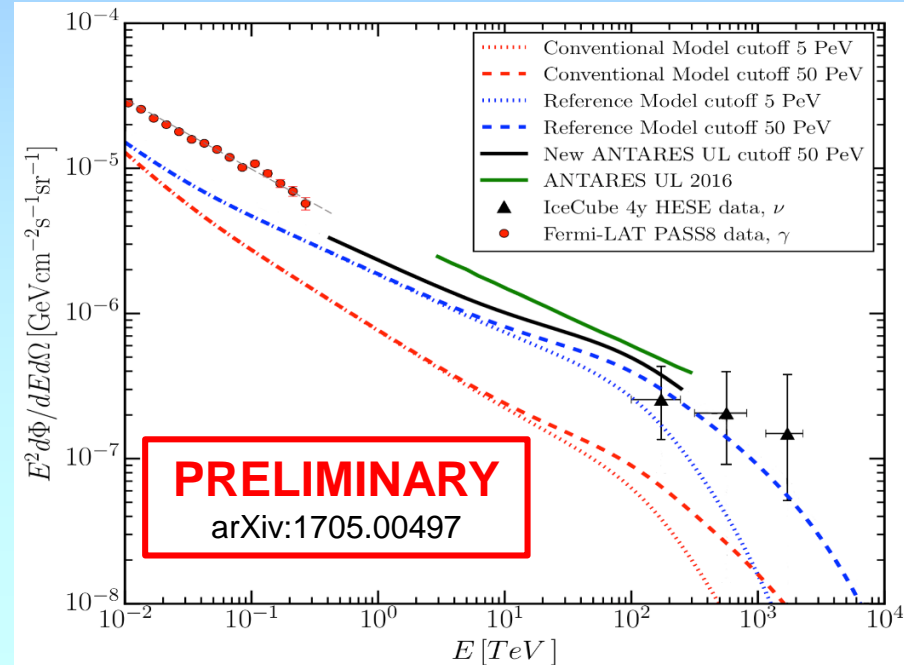
2007-2015 data, track & cascade events
Likelihood approach probing KRA_γ model

Limit at 1.25 x KRA_γ model

At most 19% of HESE IC dataset can originate from CR diffusion

↔ 5.6 out of the 30 HESE with $E_\nu > 60\text{TeV}$

Combination with IceCube ongoing

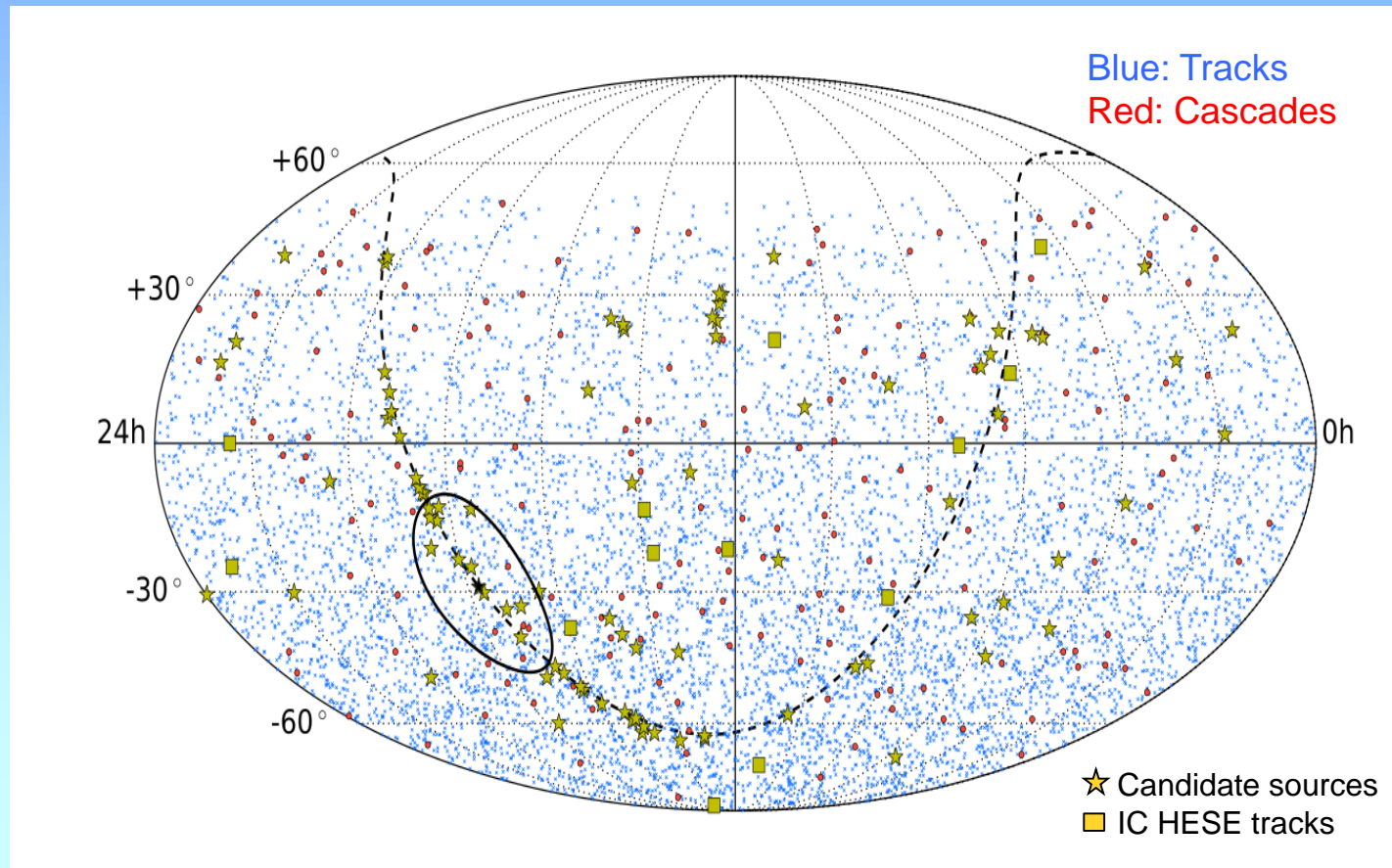




All flavour point source search with ANTARES

- 2007-2015 data set (2424 livedays) :
7629 tracks, 180 cascades
- Search for 106 candidate sources (Pulsars, SNRs,...)
and 13 IceCube tracks from the HESE sample

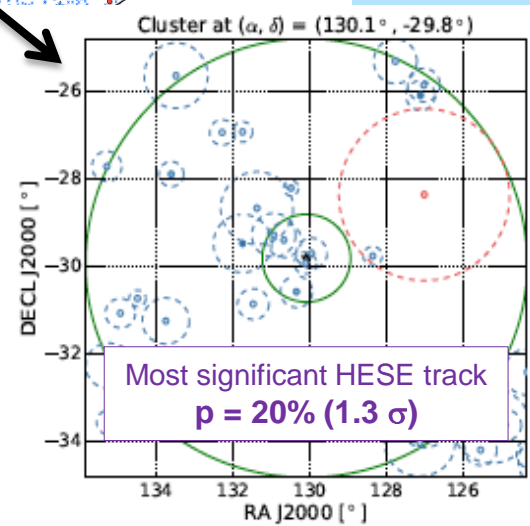
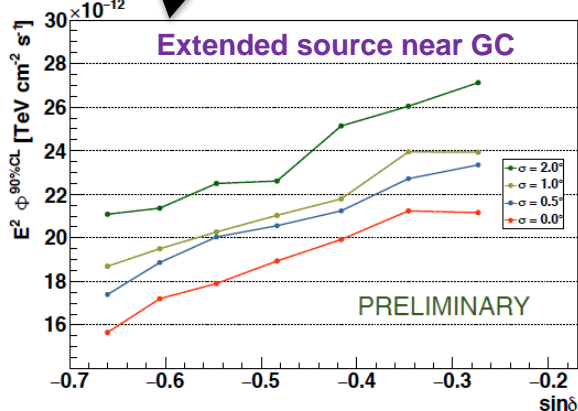
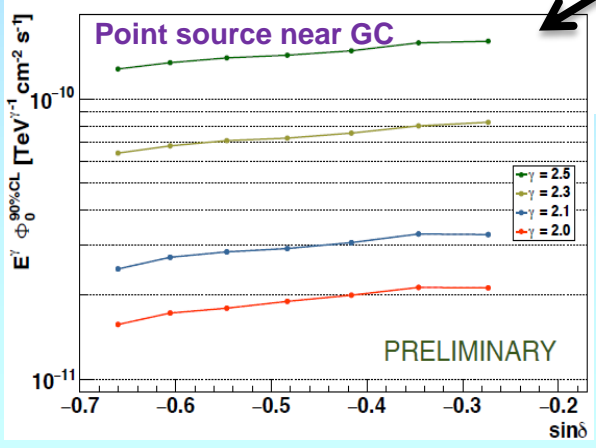
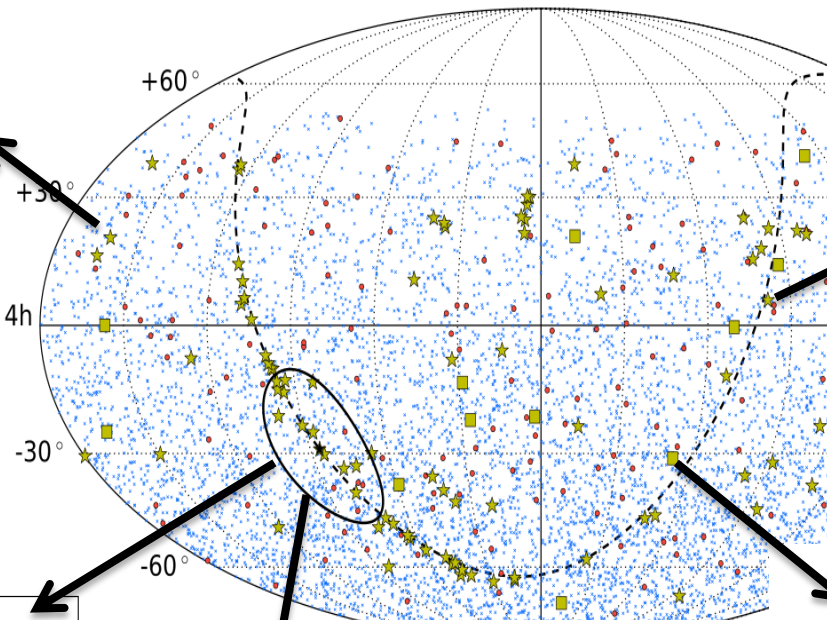
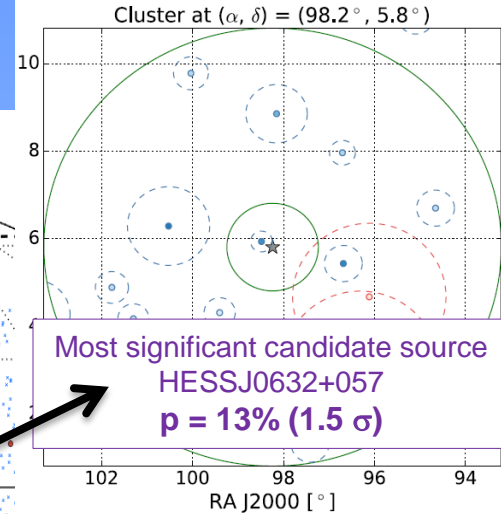
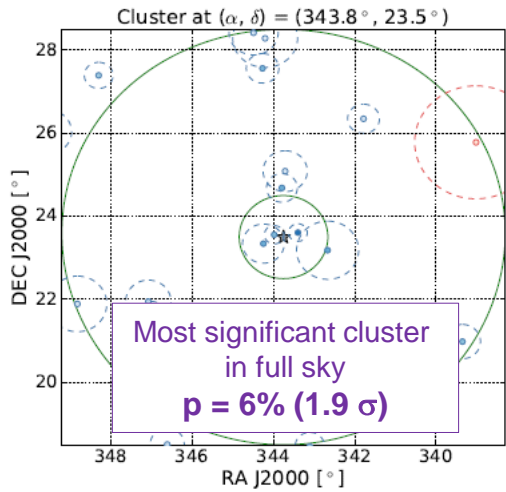
arXiv:1706.01857





All flavour point source search with ANTARES

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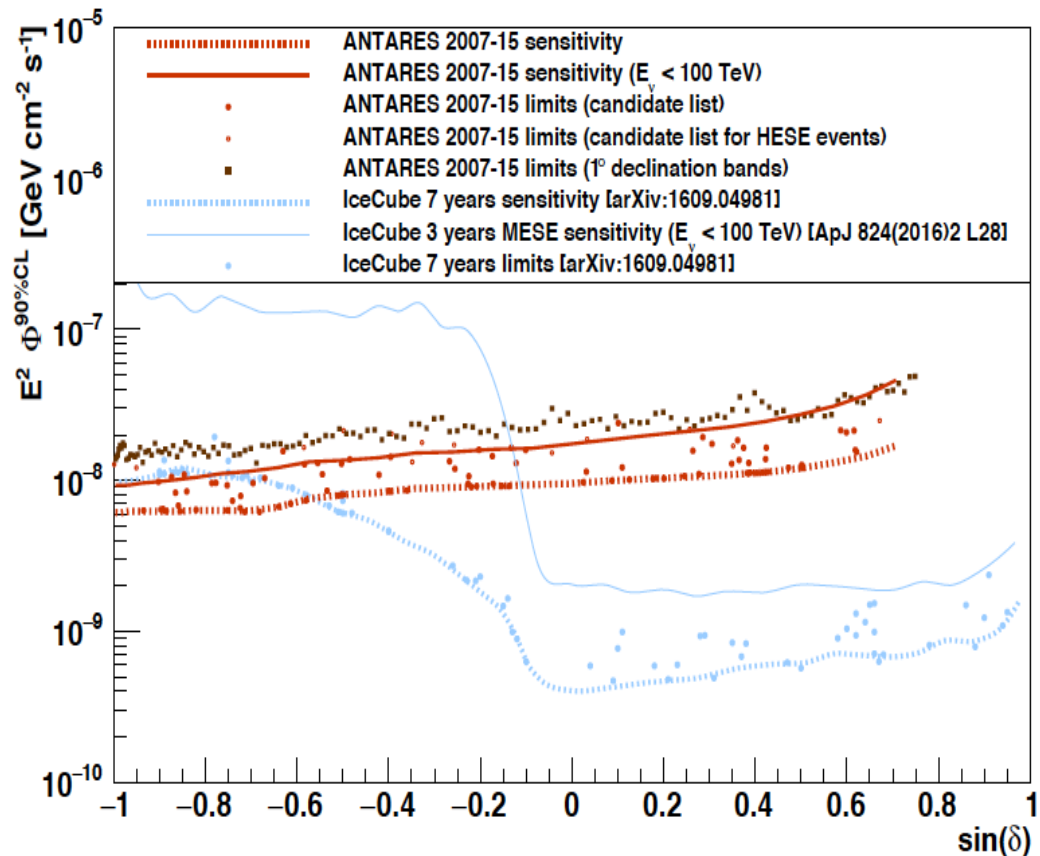
All flavour point source search with ANTARES

- 2007-2015 data set (2424 livedays) :
7629 tracks, 180 cascades
- Search for 106 candidate sources (Pulsars, SNRs,...)
and 13 IceCube tracks from the HESE sample

arXiv:1706.01857

No significant excess found

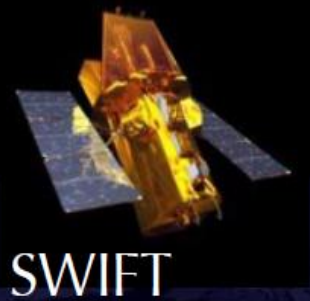
- Best limits for part of Southern Hemisphere for $E_\nu < 100$ TeV
- Relevant energy range for Galactic sources
- Results to be combined with latest IceCube search





ANTARES Multi-Messenger Alerts

Send **real-time alerts** (5 s delay, precision 0.3° - 0.4°) for neutrinos pairs, single neutrino close to direction local galaxies, single VHE neutrinos

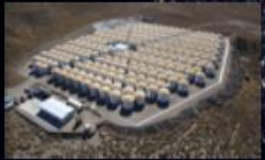


SWIFT

256 alerts sent to robotic optical telescopes, 13 to Swift and a few to MWA & HESS
→ no significant transient associated to neutrinos so far



MASTER



HAWC



ANTARES



SVOM/GWAC



TAROT



HESS



MWA



ZADKO

📖 Astropart. Phys. 35 (2012) 530

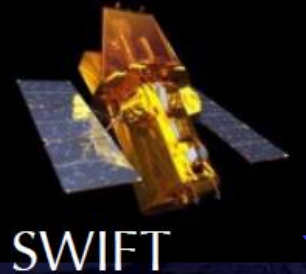
📖 Astrophys. J. 820 (2016) 24

📖 JCAP 02 (2016) 062

📖 arXiv:1705.08258



ANTARES Multi-Messengers Analyses



SWIFT



Maxi



Fermi



Ligo

Milagro



HAWK

TA



Virgo

ANTARES



HESS



Auger



MWA



Parkes



Utmost



IceCube



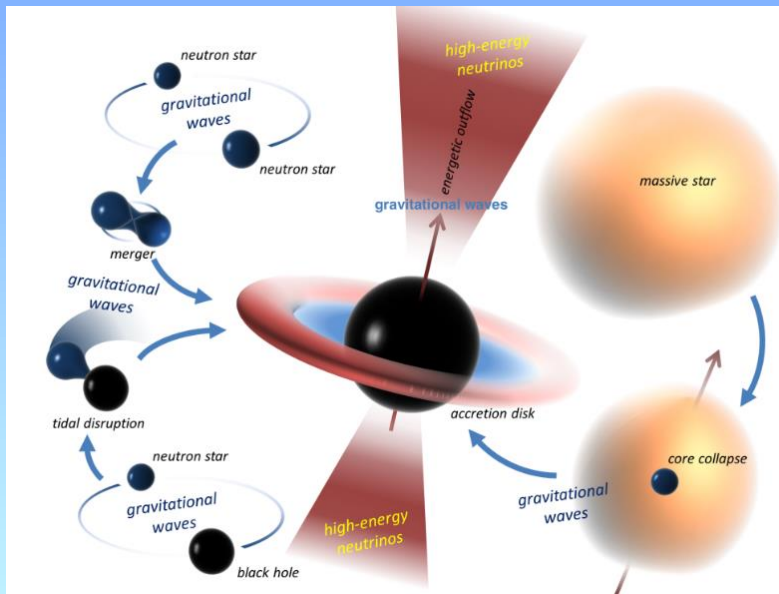
Multi-Messengers / Time Correlation Analyses

Object	Messenger	Telescope	Publication
Flaring Blazars	Gamma	FERMI/LAT	Astrop. Ph. 36 (2012) 204 JCAP 12 (2015) 014 POS (ICRC2017) 970
Flaring X-ray binaries	X & Gamma-rays	Swift, MAXI RXTE/ASM, Fermi/LAT	JCAP 04 (2017) 019
Flares from Mrk 421 and Mrk 501	Gamma	HAWK	POS (ICRC2017) 946
Gamma Ray Bursts	Gamma	Swift, GCN	JCAP 03 (2013) 006 A&A 559 (2013) A9 MNRAS 469 (2017) 906 EPJC (2017) 77:20
IceCube Events	Neutrino	IceCube	A&A Lett. 576 (2015) L8 arXiv:1706.01857
UHECR	CR	Auger, TA	ApJ 774 (2013) 19 POS (ICRC2017) 990
Galactic plane	Model fit	Fermi, Milagro	PLB 760 (2016) 143 arXiv:1705.00497
Fast Radio Bursts	Radio	SUPERB@Parkes, UTMOST, ASKAP	MNRAS (2017) 469 (4):4465
Fermi Bubbles	Gamma	Fermi	EPJC (2014) 74:2701 POS (ICRC2017) 1001
BH/NS mergers	Gravitational waves	Ligo/Virgo (+Icecube)	JCAP 06 (2013) 006 PRD 93 (2016) 122010 PRD 96 (2017) 022005



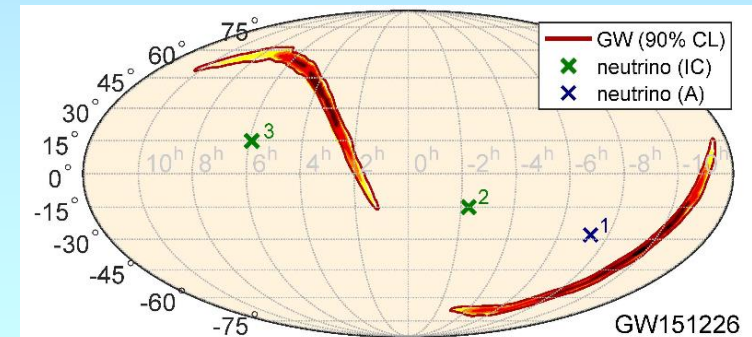
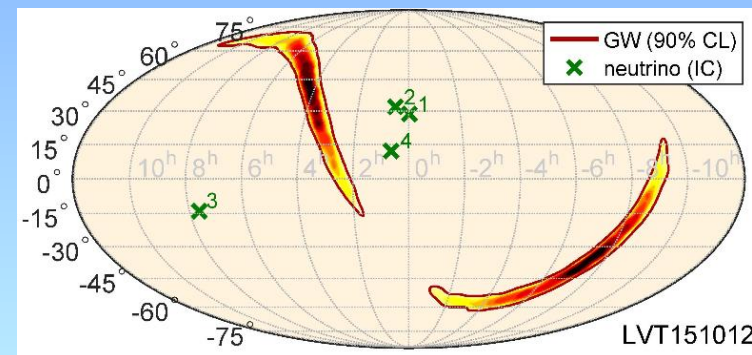
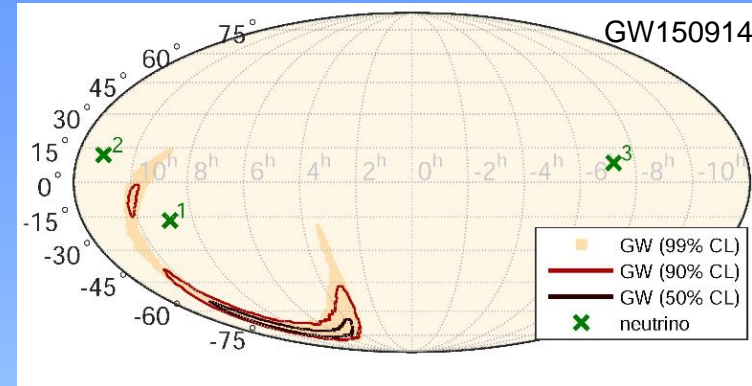
Search for Coincidences with Gravitational Waves

- Neutrino follow-up relevant in case of baryonic ejecta (mostly for BH/NS or NS/NS merging)
- Coincident neutrino signal would allow pinpointing the source ($< 1^\circ$)



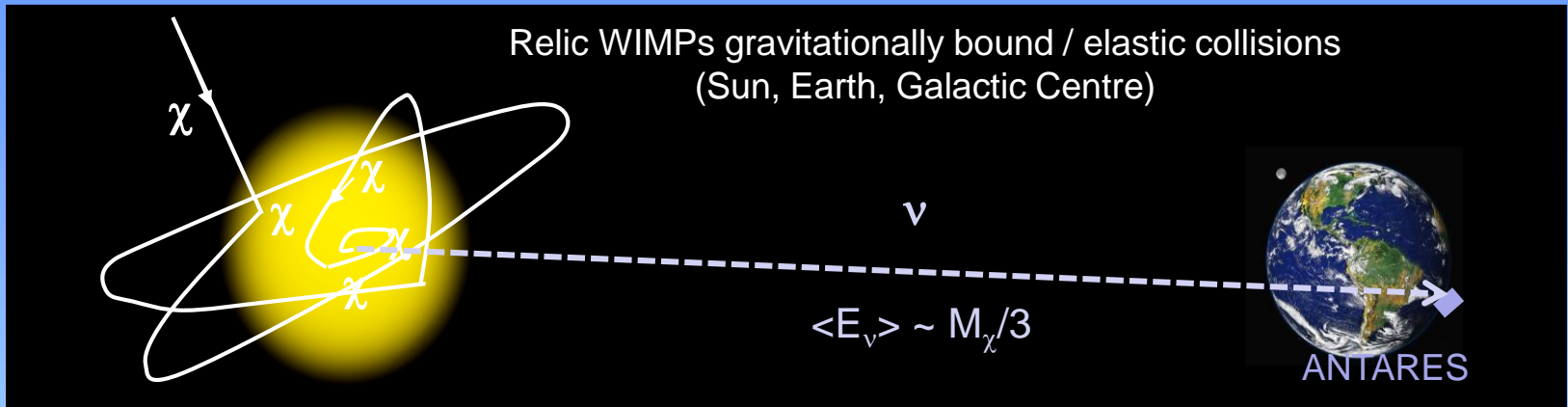
- Combined analysis Ligo/Virgo, IceCube and ANTARES (dominates for $E_\nu < 100$ TeV)
- **No counterpart observed so far (in ± 500 s)**
 → **Total energy radiated in neutrinos $< 10\%$**

Now real time follow-up of ongoing science run





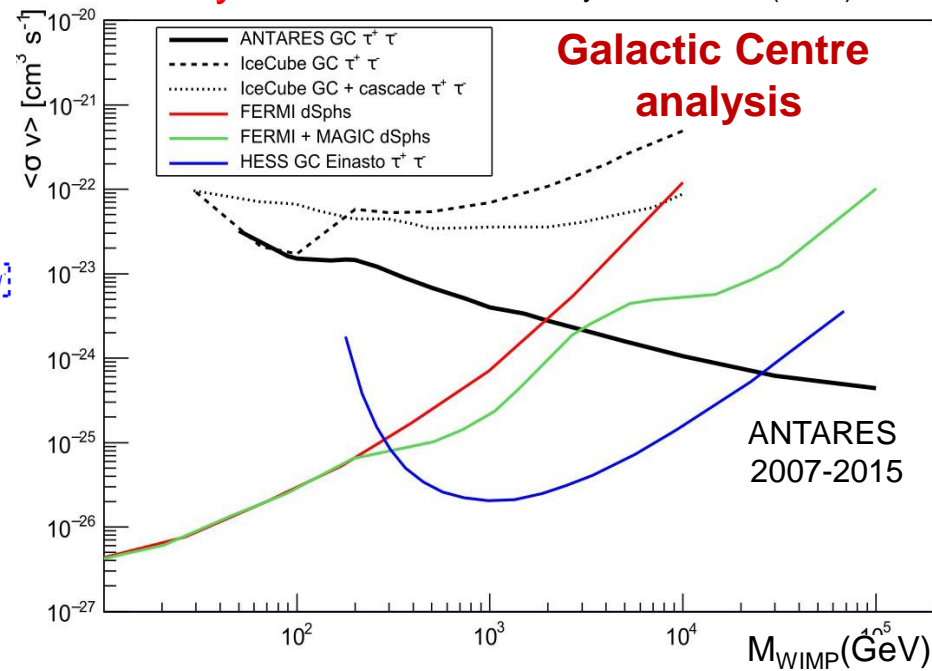
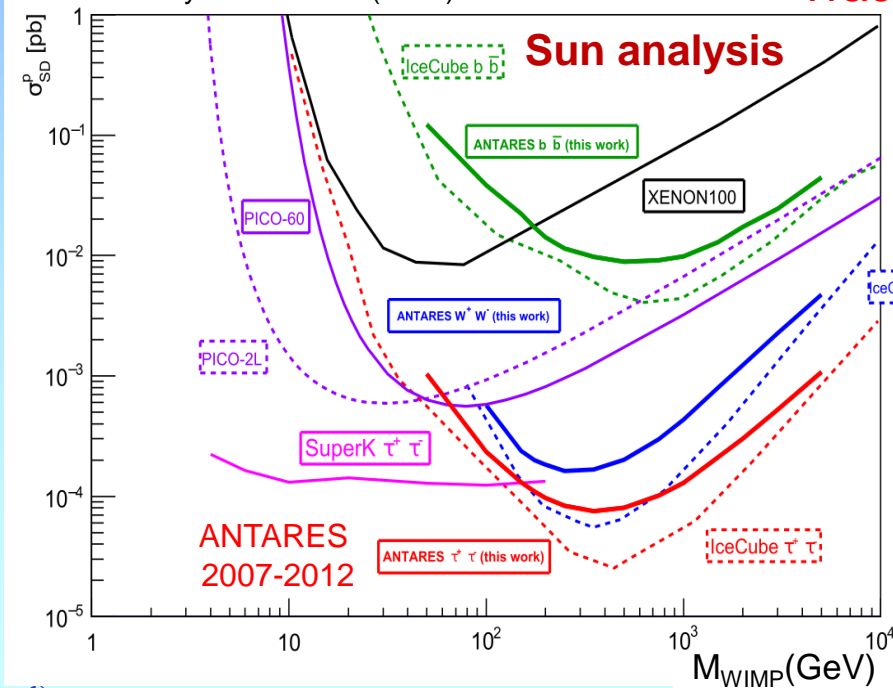
Indirect Searches for Dark Matter



Phys.Lett. B 759 (2016) 69

Track channel only

Phys. Let. B 769 (2017) 249



Also Secluded DM from the Sun : JCAP 05 (2016) 016
and DM from the centre of the Earth : Physics of the Dark Universe 16 (2017) 41

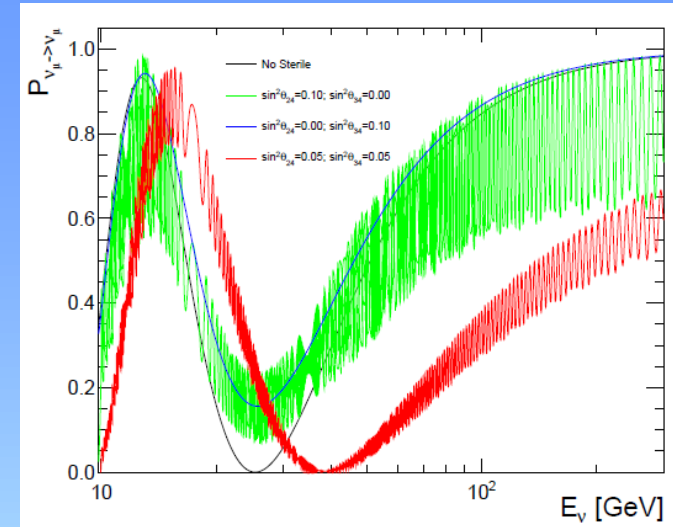
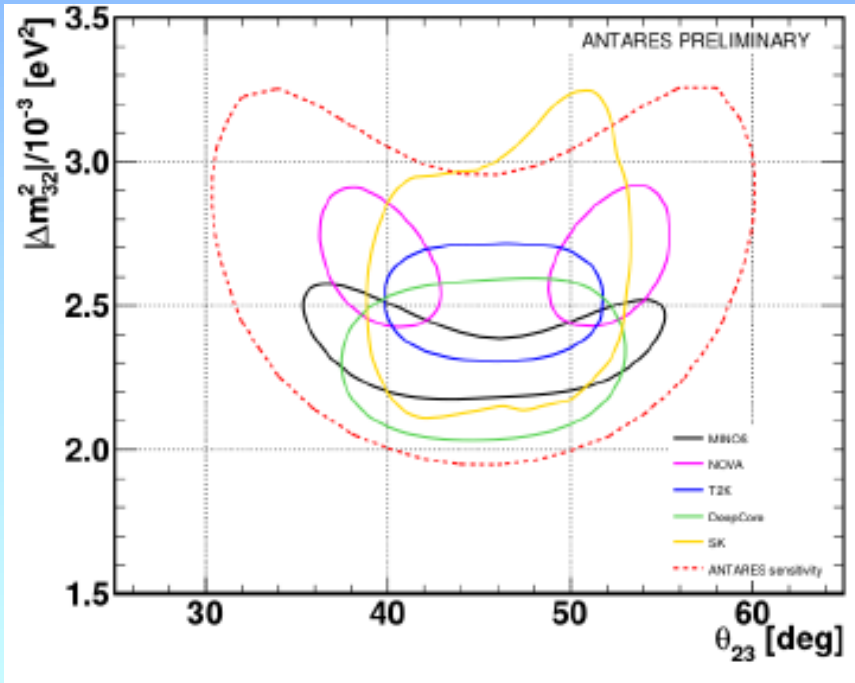


Oscillations of atmospheric neutrinos

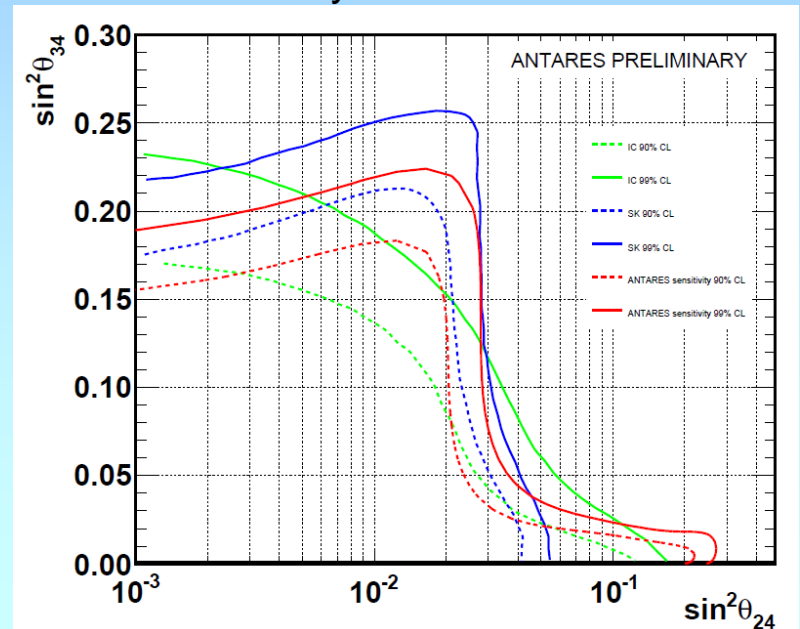
New analysis on-going with dedicated energy reconstruction optimized for low energy events

POS (ICRC2017) 1026

Sensitivity to neutrino mixing parameters



Sensitivity to sterile neutrino



Data unblinding & results after Summer



Summary & Perspectives

- **ANTARES**, the largest Neutrino Telescope in the Northern Hemisphere, is celebrating **10 years of continuous data taking**
- Cascades events included in analyses with up to **3° angular resolution** → all flavour neutrino searches
- **Strongest limits for galactic point sources** in Southern Hemisphere with $E_\nu < 100$ TeV
- Interesting constraints on **origin of IceCube signal**
- **Small excess of HE neutrinos**, of magnitude expected from IceCube cosmic flux
- **Many multi-messenger analyses** looking for transient sources (Radio, Optical, X/ γ -Rays, TeV gammas, GW)
- **Competitive results in Dark Matter searches**

Demonstration of great potential of Deep Sea Neutrino Telescopes
→ **To be exploited with KM3NeT**

(cf talks of P. Sapienza on Tue. 25th & J. Hofestädt on Thu. 27th)