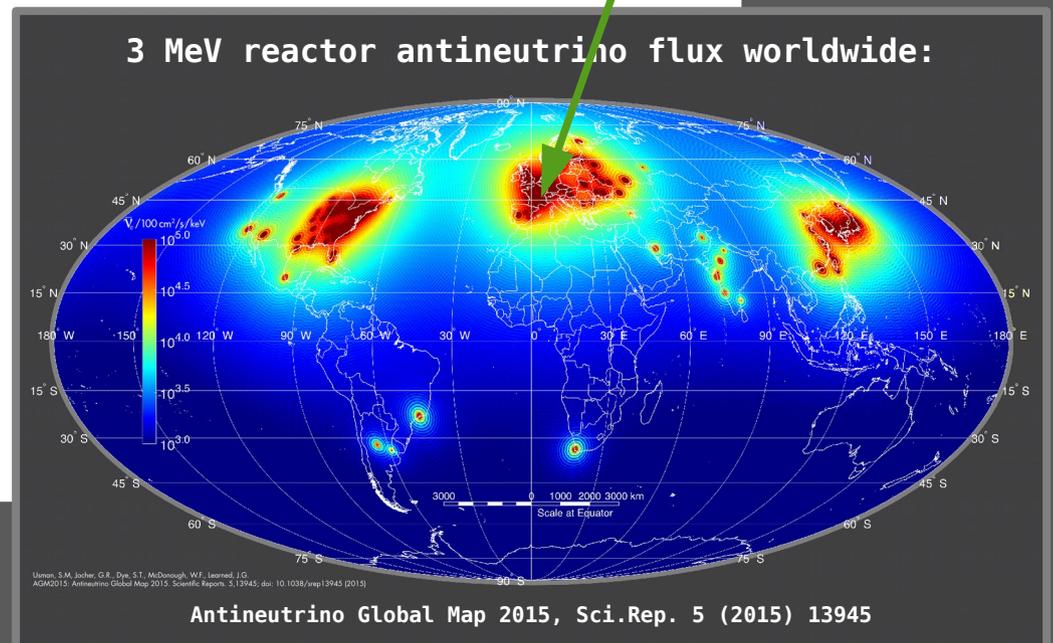


Search for eV Sterile Neutrinos The Stereo Experiment

Stefan Schoppmann
for the Stereo collaboration



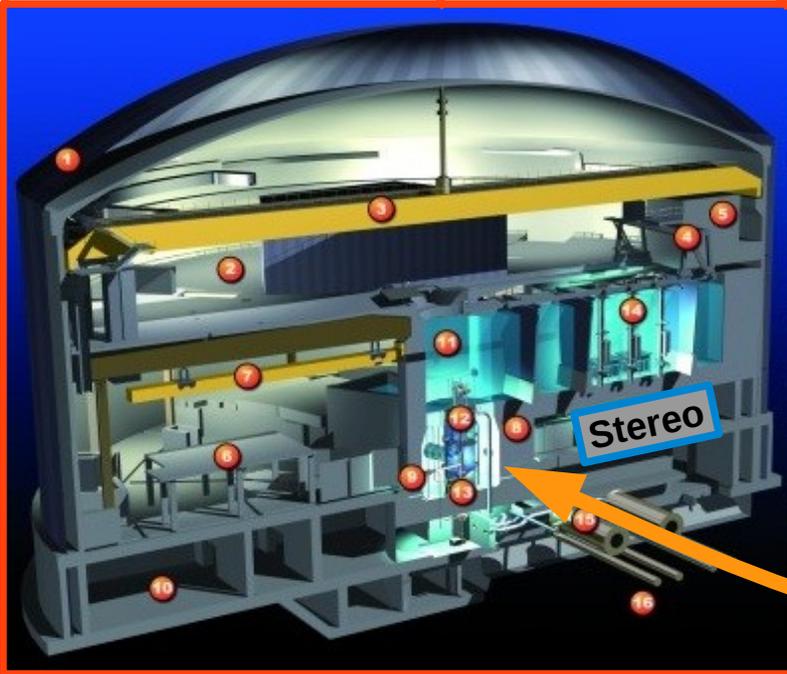
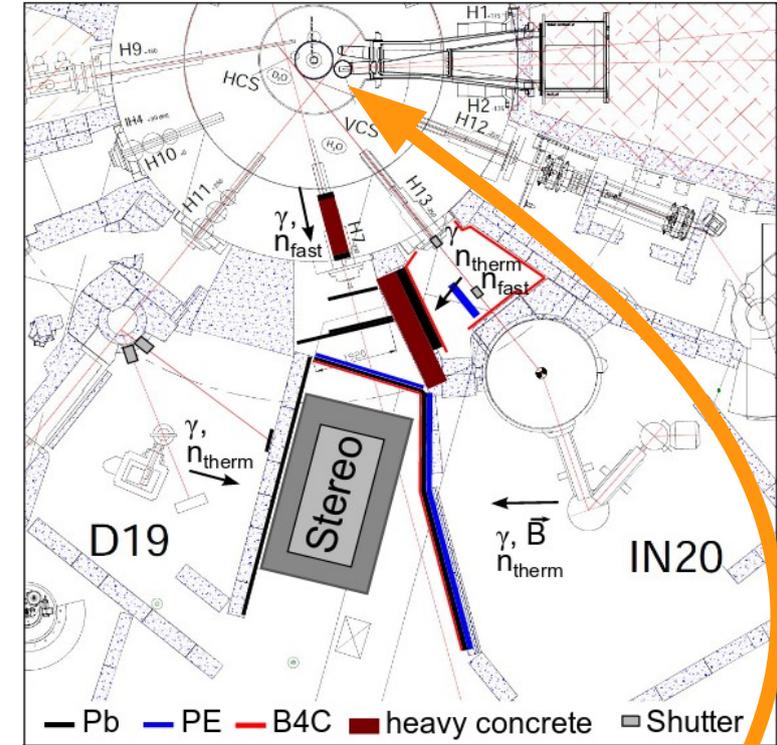
- Experimental Setup
- Calibration and Energy Reconstruction
- Signal and Backgrounds
- Summary/Outlook





Experimental Setup

Experiment Site



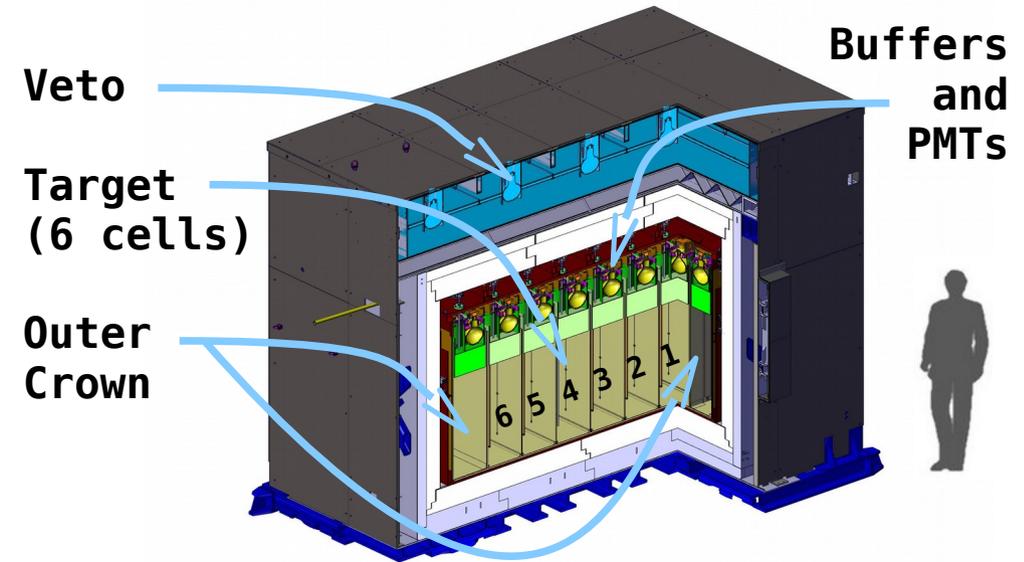
- Neutrino source:
- reactor at ILL Grenoble
 - 58.3 MW_{thermal}
 - Height: 80cm
 - Diameter: 40cm
 - Highly enriched in ^{235}U (93%)

Baseline: 9 - 11 m
Overburden: ~15 m.w.e.



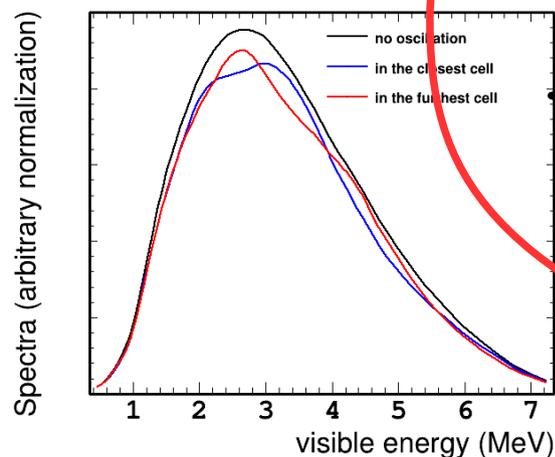
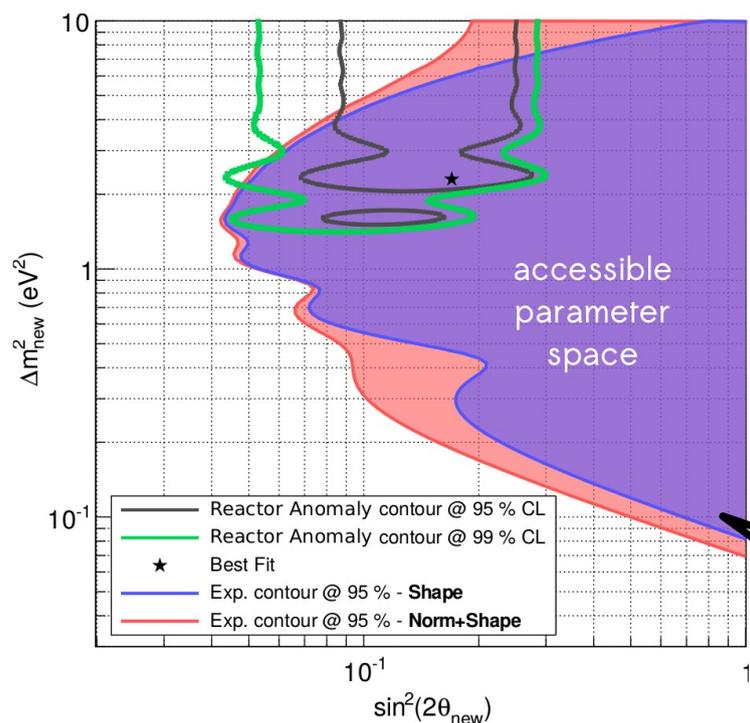
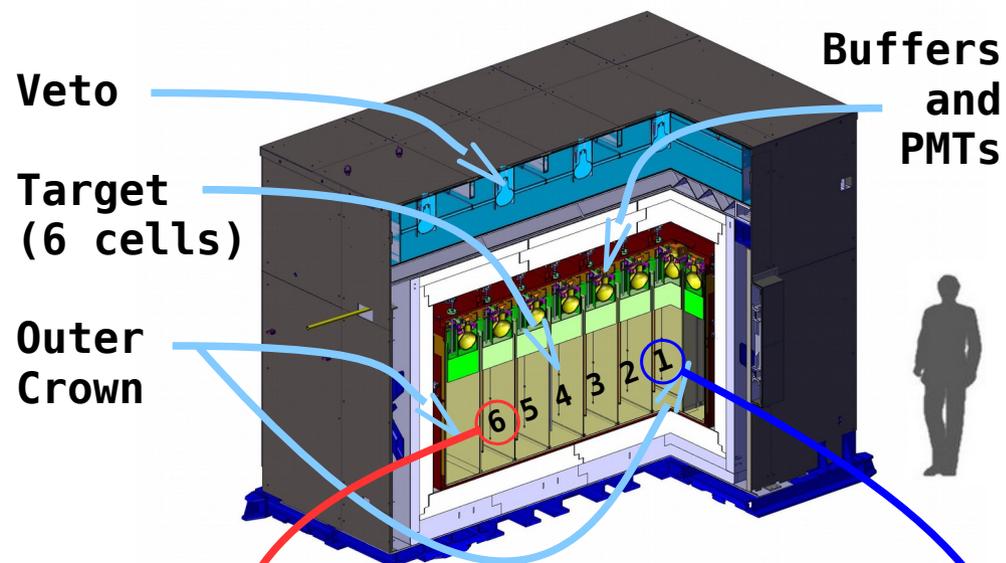
Detector/Measurement Idea

- Target segmented in 6 cells
→ 1800 l of Gd-loaded liquid scintillator
- Surrounding outer crown to capture gammas
- 48 PMTs of 8 inch diameter
- Layers of acrylic and oil as buffer
- Water Cherenkov veto on top
- About 90 tons of shielding material
→ lead, polyethylene, B_4C , iron

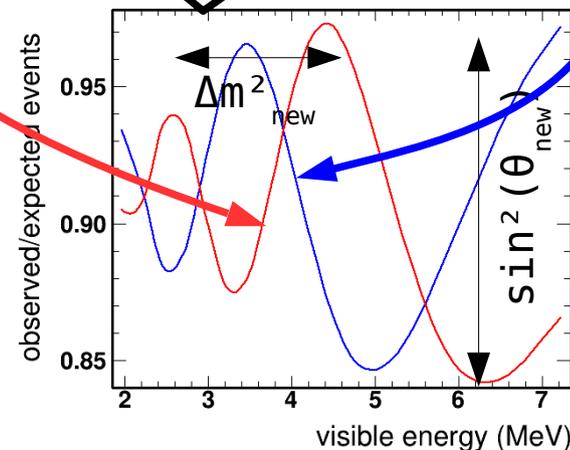


Detector/Measurement Idea

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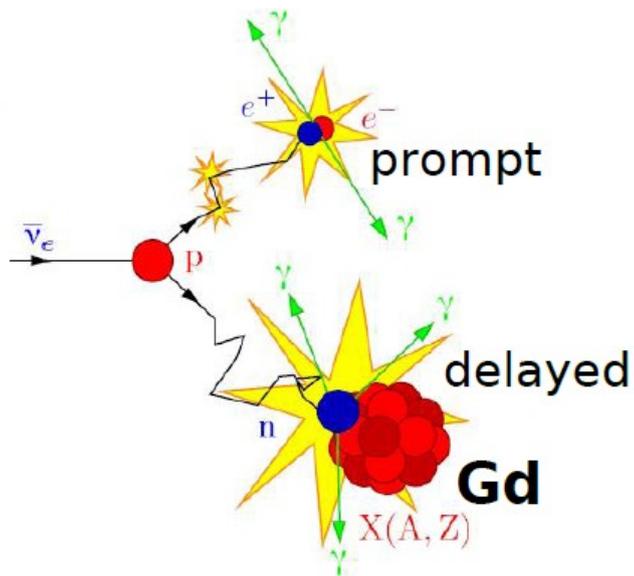
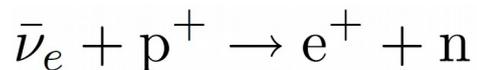


Visible shape effects in energy

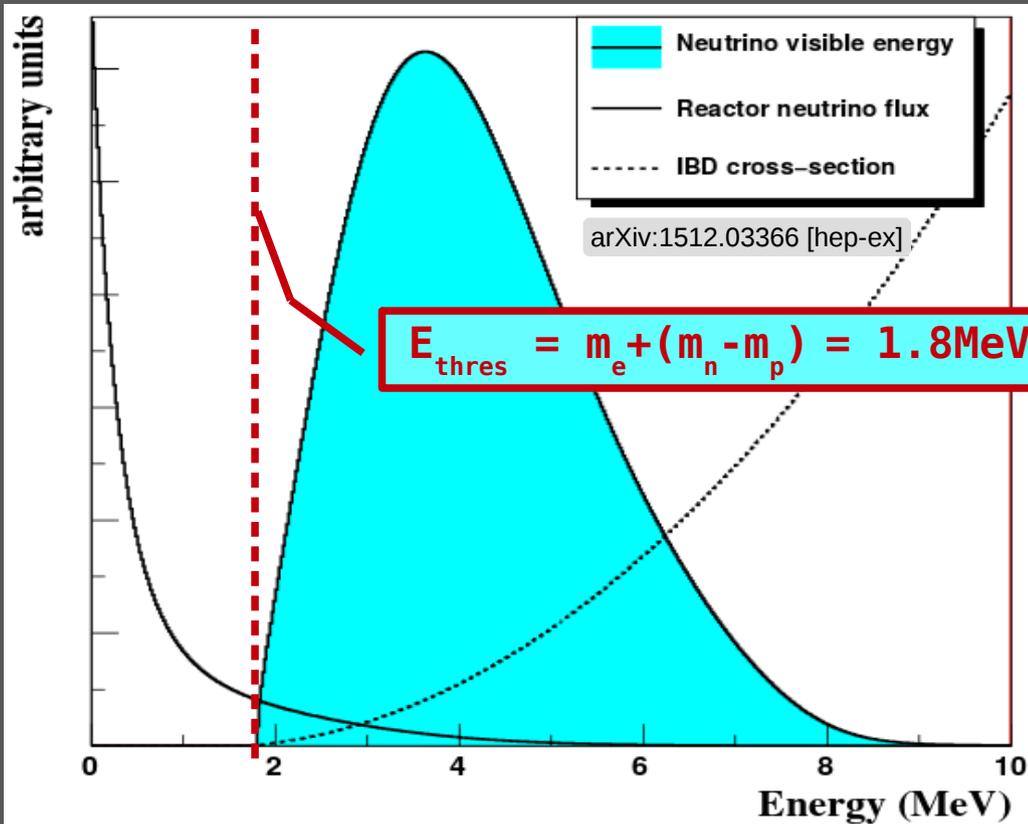


sensitivity after 1 year

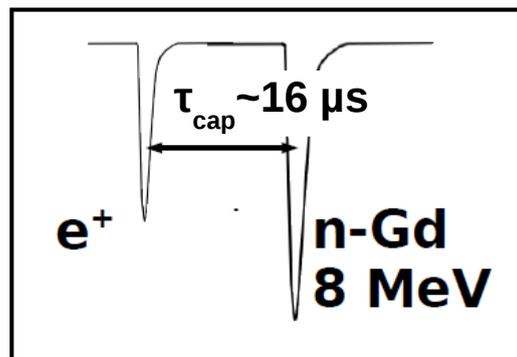
Inverse beta-decay (IBD):



G. Mention, PhD thesis (2005)
Université Lyon I



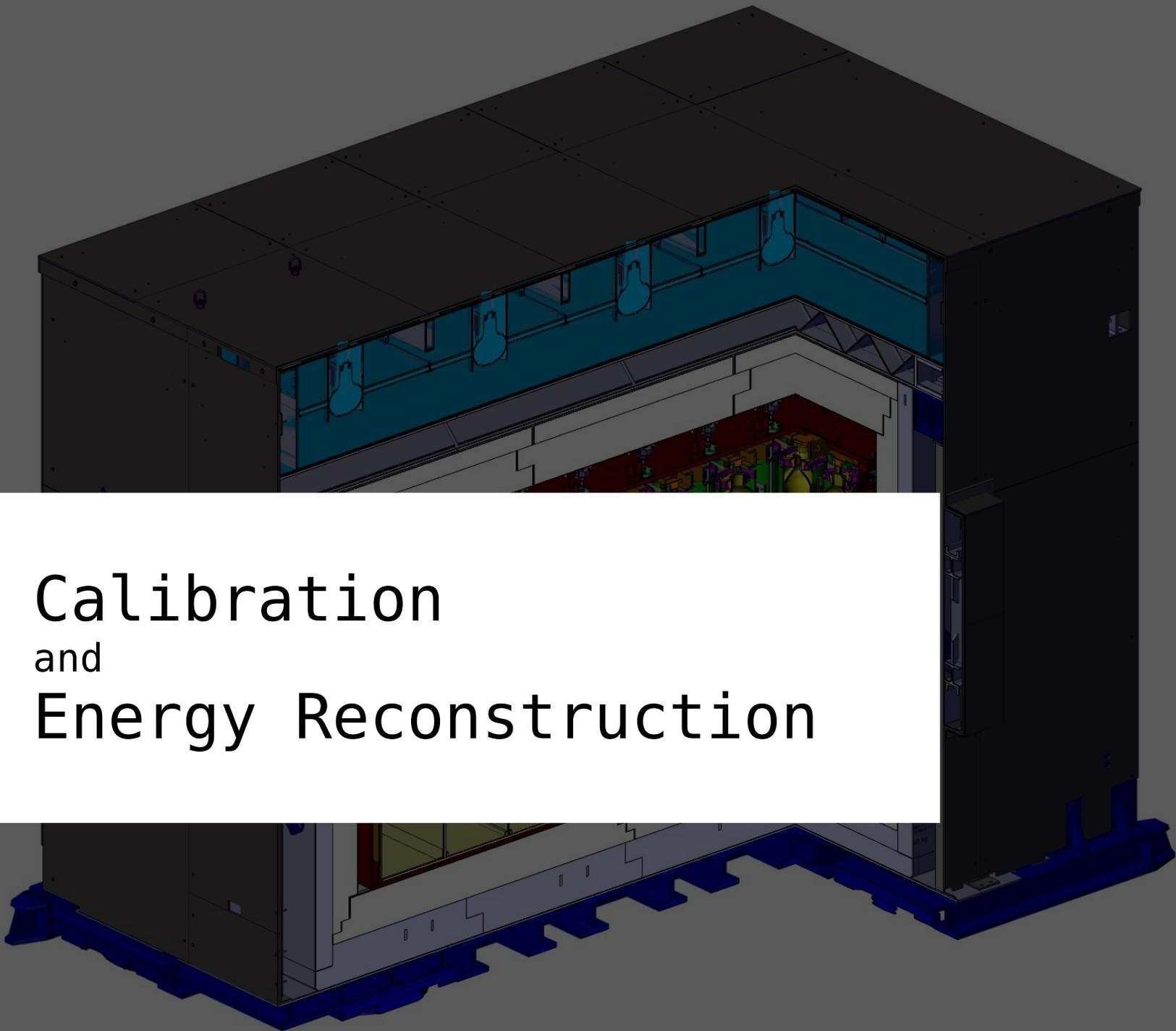
Delay due to:
thermalisation
of neutron
before capture
+
time constant
of the capture
process



Visible prompt energy =
neutrino energy + corrections

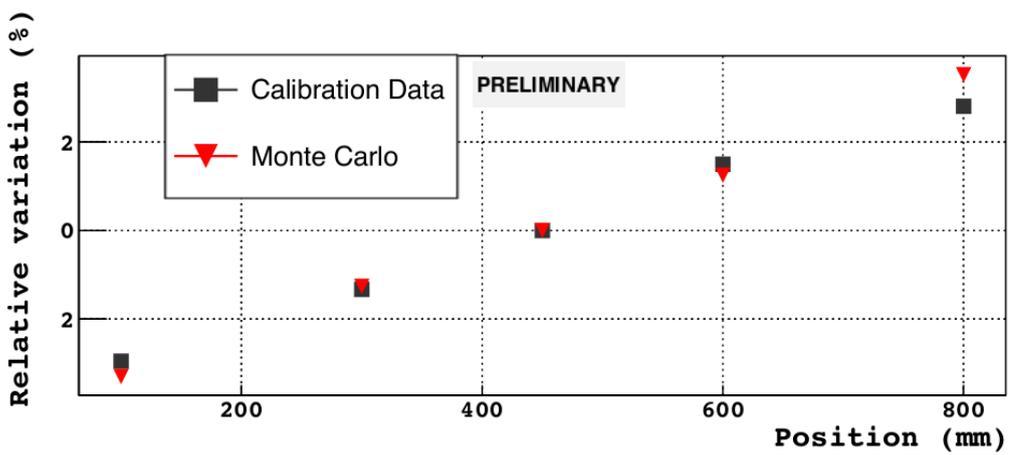
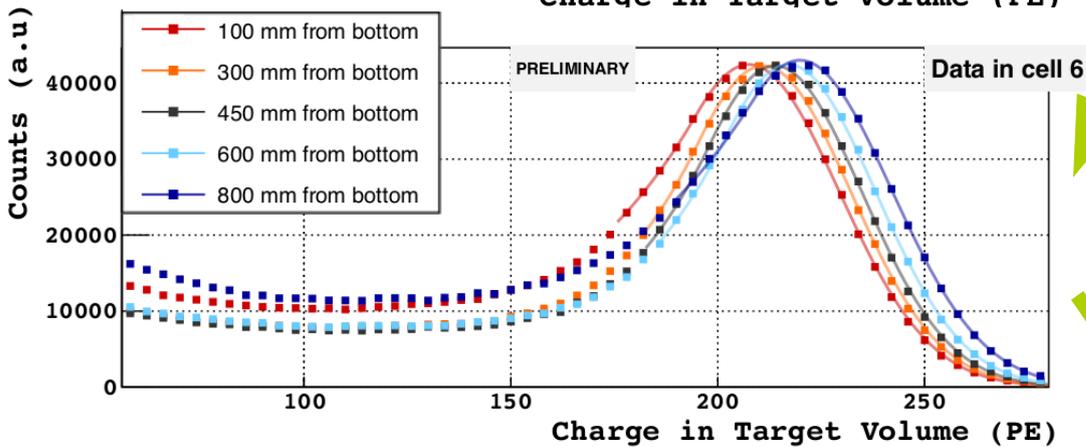
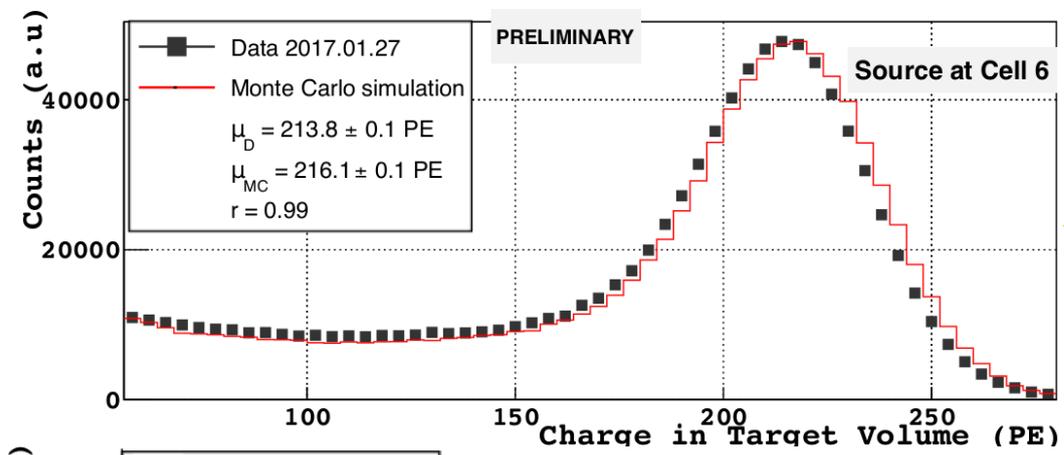
$$E_{\bar{\nu}_e} = E_{e^+} + (m_n - m_p) + \mathcal{O}(E_{\bar{\nu}_e}/m_n)$$

$$E_{vis}^{prompt} = E_{e^+} + m_{e^-}$$



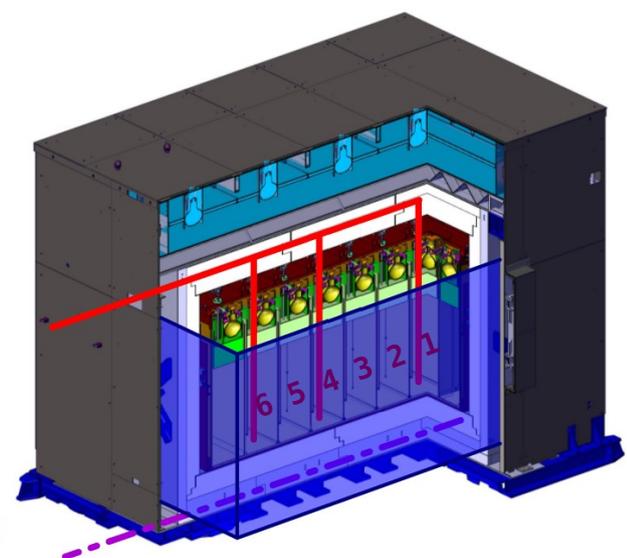
Calibration and Energy Reconstruction

Homogeneity



- three calibration systems
- various $\gamma+n$ sources

- internal calibration (cell 1,4,6)
- external calibration (2D, inside shielding)
- underneath calibration

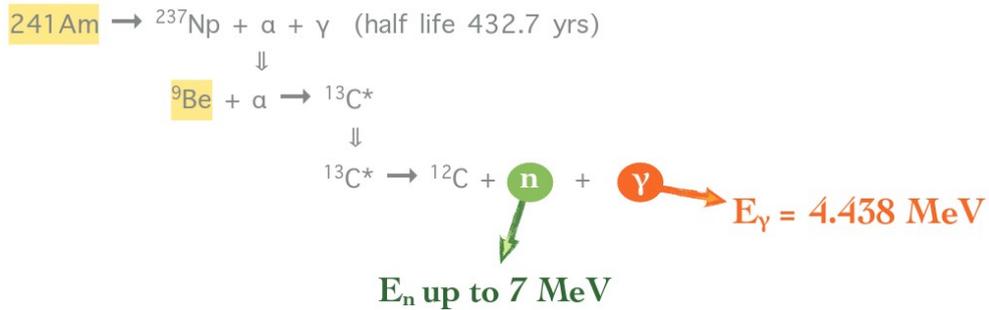


Mn-54 source at different heights in one cell shows:

- low z-dependence
- excellent MC to data agreement at current stage of analysis

Detection Efficiency

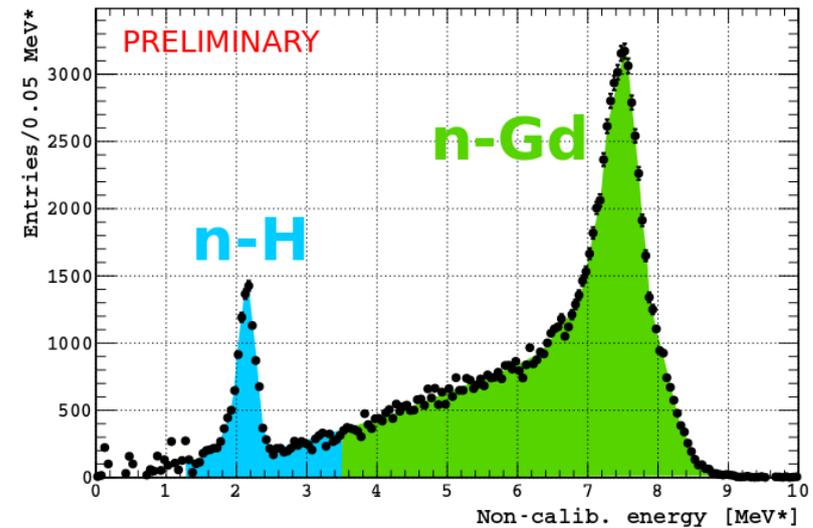
Using $\gamma+n$ events of an AmBe source



→ study delayed event energy

→ study delay time

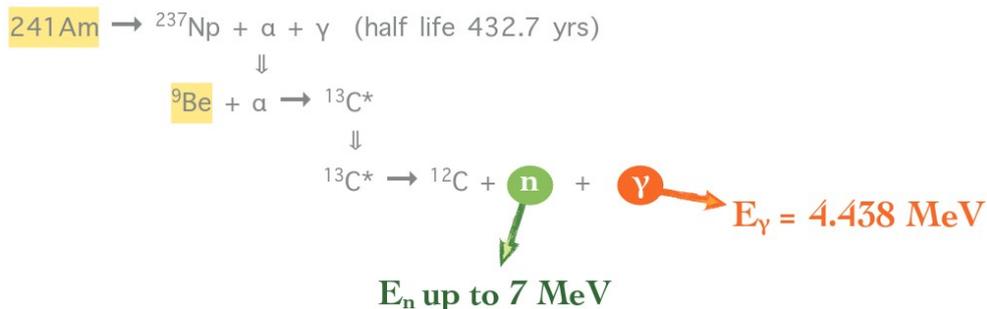
Source at centre of cell 6



⇒ n-Gd capture fraction ~86%

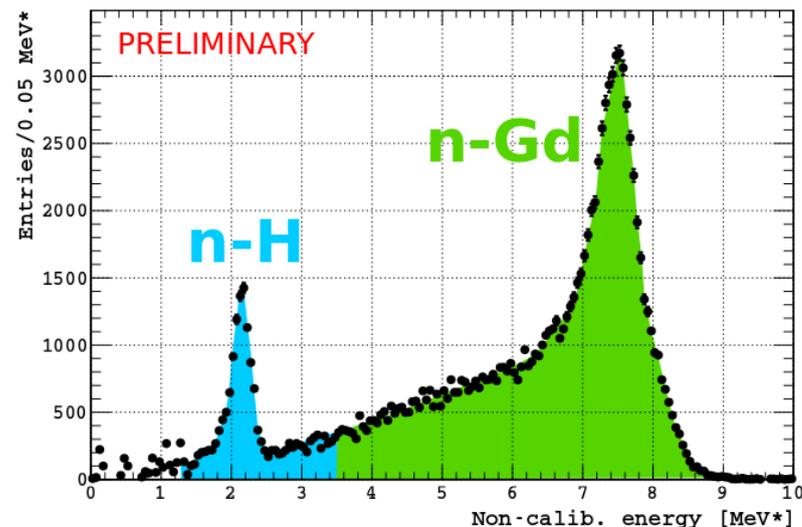
Detection Efficiency

Using $\gamma+n$ events of an AmBe source



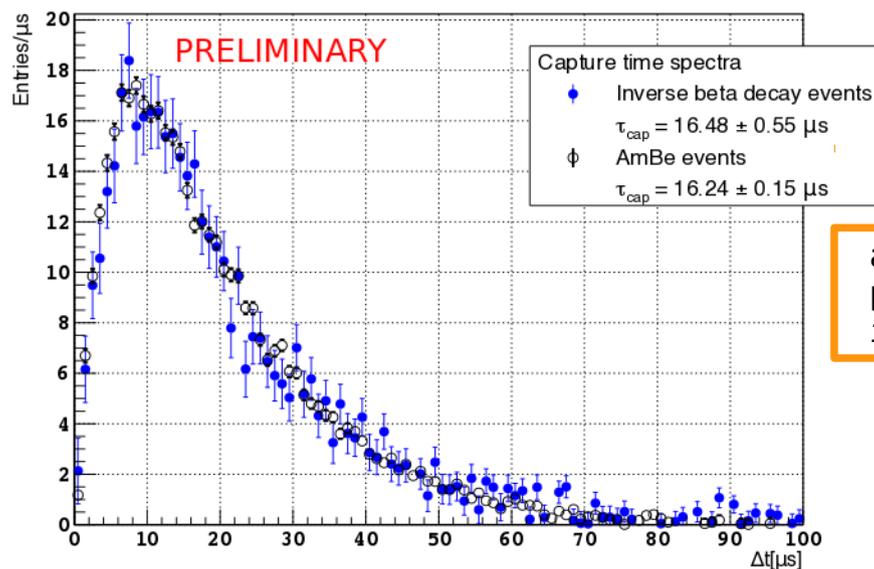
- study delayed event energy
- study delay time

Source at centre of cell 6



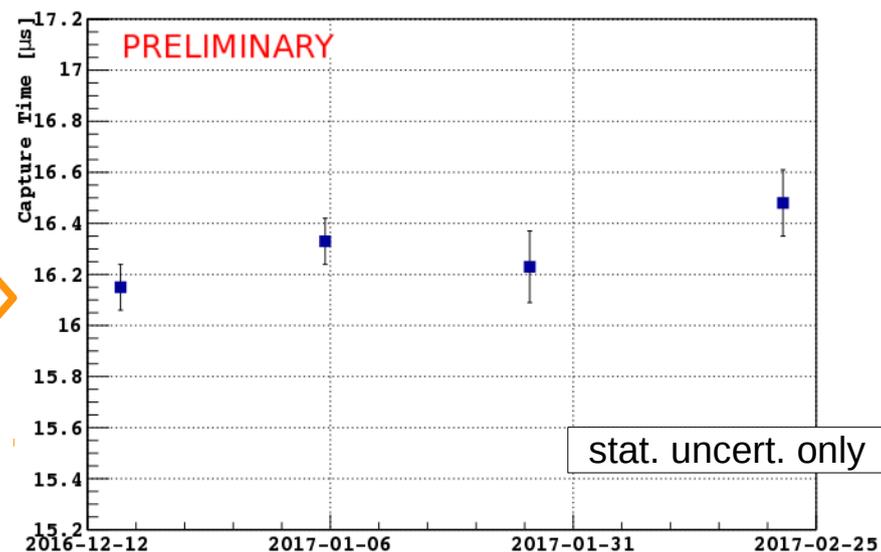
⇒ n-Gd capture fraction ~86%

Neutron capture time comparison



at many points in time

Capture time stability



Energy Reconstruction

- Collected charge in one target cell given by
- light produced in that cell and
 - light cross-talking from neighbouring cells j :

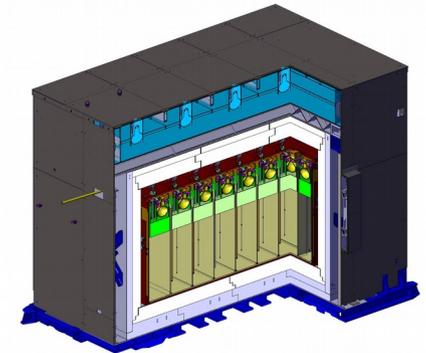
$$Q_i = \alpha_i^{\text{geom}} \sum_{j=\text{cells}} E_j^{\text{dep}} \times f_j \times L_{j \rightarrow i}$$

photon acceptance cell i

deposited energy in cell j

photons per MeV

light crosstalk between cells



→ Conversion factors obtained simultaneously and in-situ from calibration data

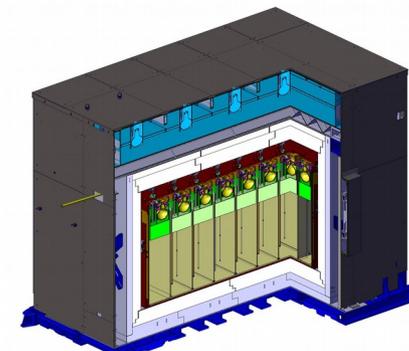
$$Q_i = \sum_{j=\text{Cells}} E_j^{\text{dep}} C_j L_{j \rightarrow i}$$

Energy Reconstruction

- Collected charge in one target cell given by
- light produced in that cell and
 - light cross-talking from neighbouring cells j :

$$Q_i = \alpha_i^{\text{geom}} \sum_{j=\text{cells}} E_j^{\text{dep}} \times f_j \times L_{j \rightarrow i}$$

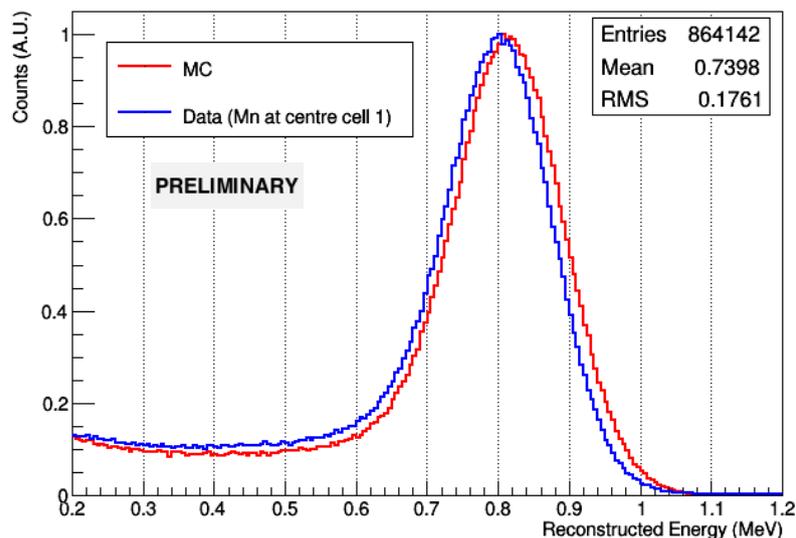
α_i^{geom} : photon acceptance cell i
 E_j^{dep} : deposited energy in cell j
 f_j : photons per MeV
 $L_{j \rightarrow i}$: light crosstalk between cells



→ Conversion factors obtained simultaneously and in-situ from calibration data

$$Q_i = \sum_{j=\text{Cells}} E_j^{\text{dep}} C_j L_{j \rightarrow i}$$

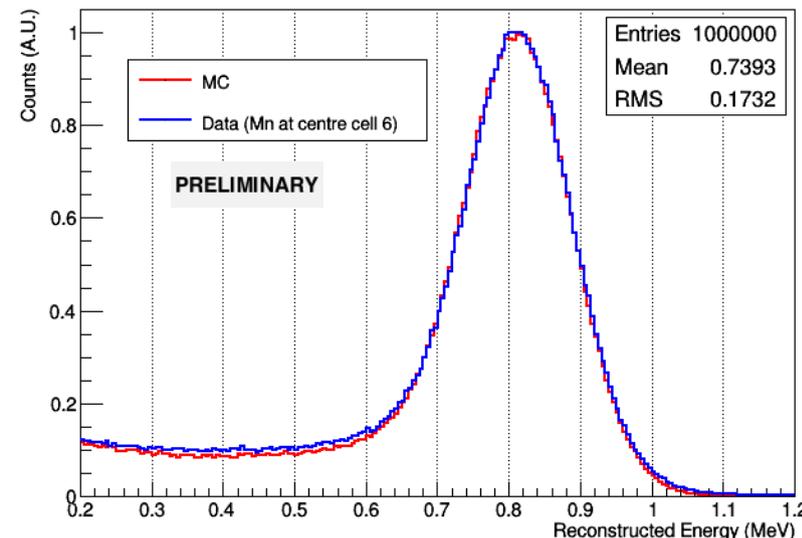
Cell 1



Mn-54 source at centre of different cells:

→ Data-MC agreement goal: 2%
 ⇒ already achieved

Cell 6



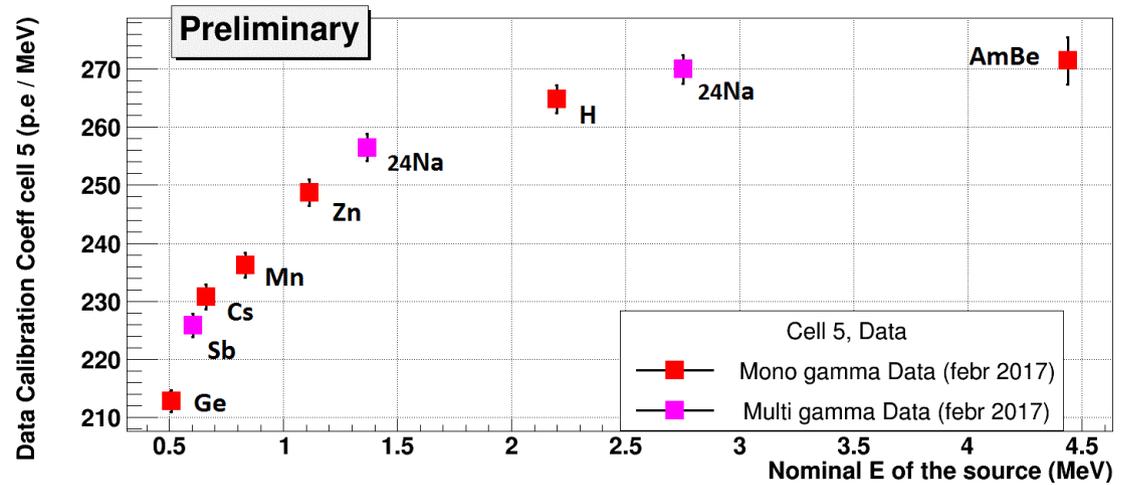
Energy Reconstruction



Energy response function

- quenching curve
- gamma sources
- single gamma lines

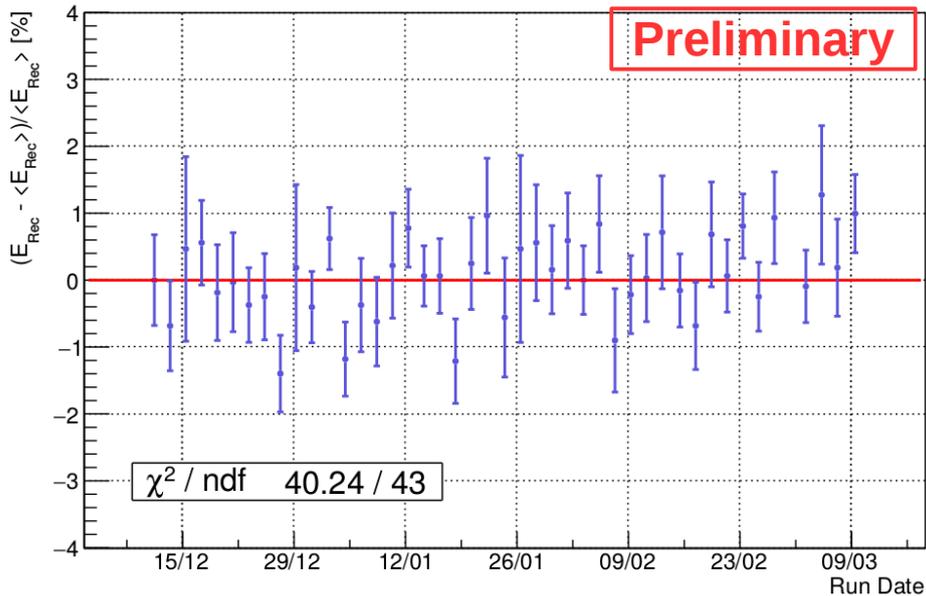
- ⇒ expected behaviour of liquid scintillator ionisation quenching (Birk's law)



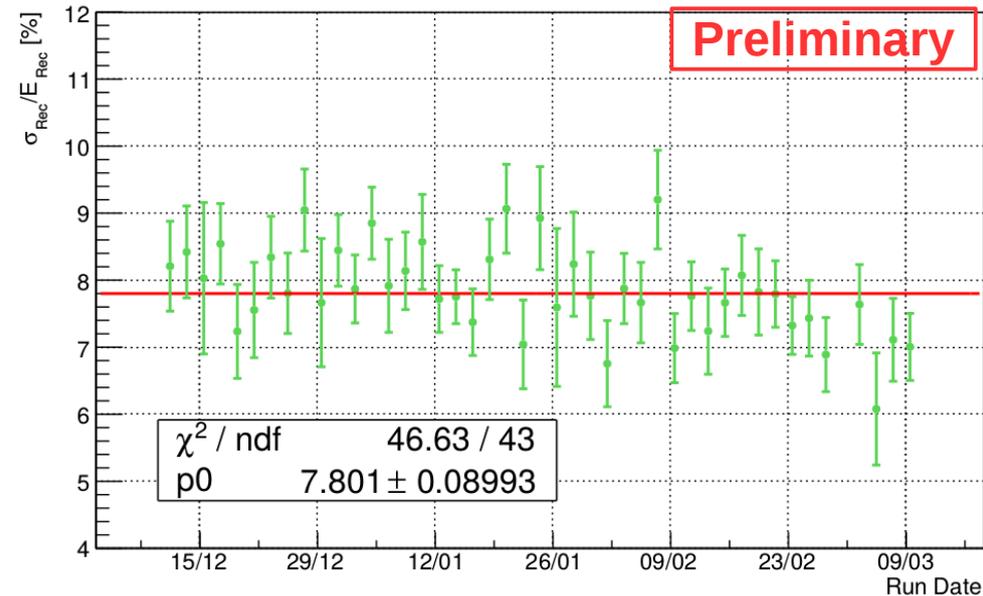
Energy stability from n-H captures of spallation neutrons

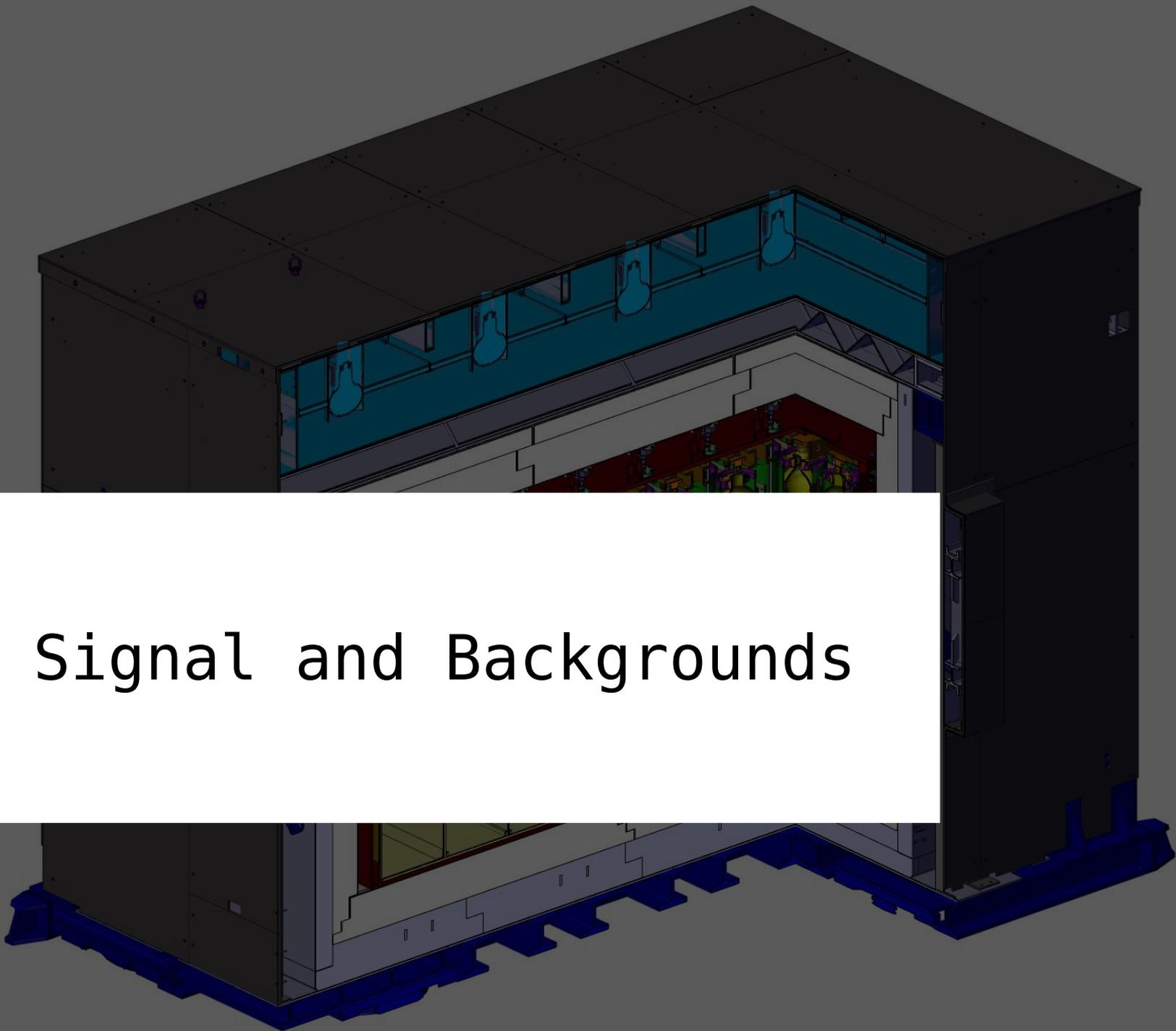
⇒ in good agreement with stable behaviour

Evolution of Target Energy Peak of H Capture



Evolution of Target Energy Width of H Capture





Signal and Backgrounds



© NSF/J. Yang

Muon induced

Examples:

- fast neutrons
- stopping muons

Counter measures:

- water pool
- active veto
- pulse shape discr.
- PMT cut



Courtesy of J. Haser

Reactor induced

Examples:

- neutrons
- γ -radiation via neutron capture

Counter measures:

- shielding
- pulse shape discr.
- event topology



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Radioactive decays

Examples:

- thorium/uranium
- radon/argon
- potassium

Counter measures:

- shielding
- event topology
- precise measurement



© NSF/J. Yang

Muon induced

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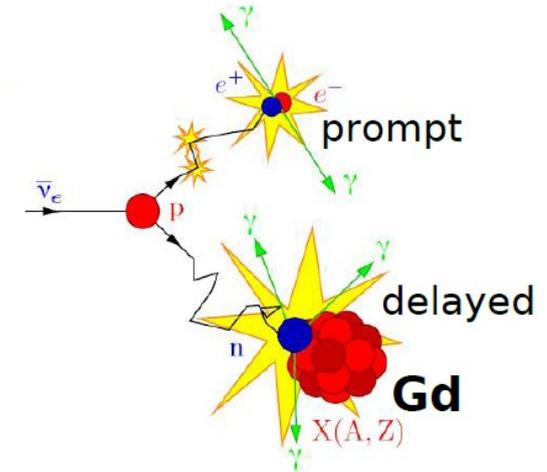
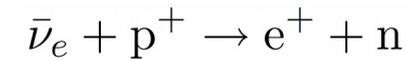
Neutrino Candidate Search

Signal selection (proposal cuts):

	Cut 1	Cut 2	Cut 3
E_{prompt}	$1.5 \text{ MeV} < E_{\text{tot}} < 8 \text{ MeV}$	$E_{\text{GC}} < 1.1 \text{ MeV}$	2.5σ PSD
E_{delayed}	$5 \text{ MeV} < E_{\text{tot}} < 10 \text{ MeV}$	$E_{\text{TG}} > 1 \text{ MeV}$	
T_{delay}	$< 70 \mu\text{s}$		

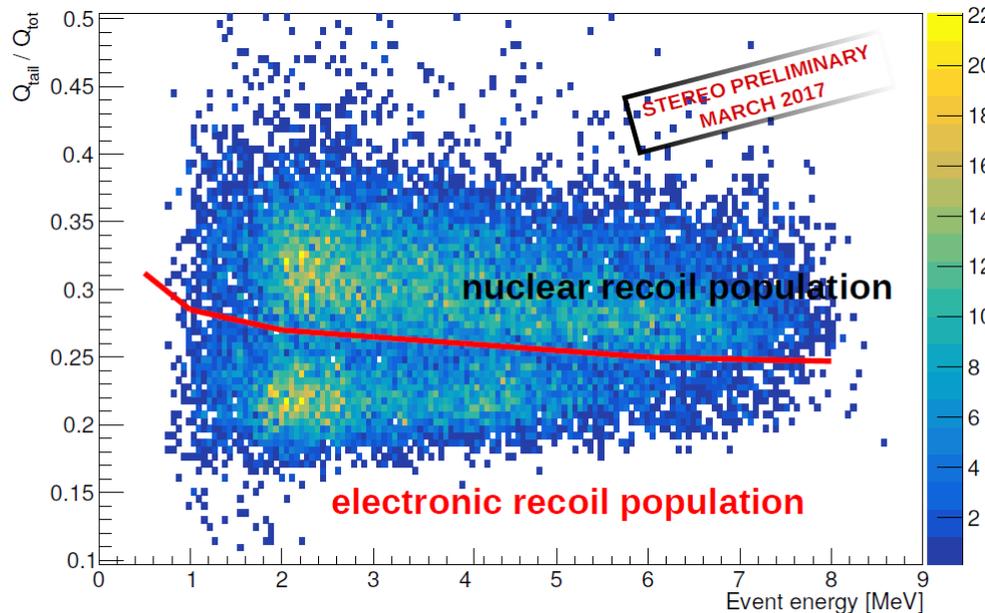
+ after muon veto
+ additional cleaning cuts

Inverse beta-decay (IBD):

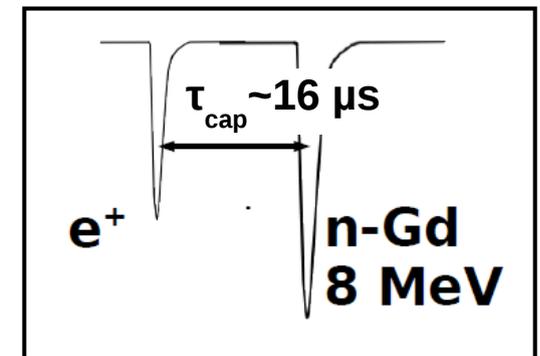


G. Mention, PhD thesis (2005)
Université Lyon 1

Background rejection by Pulse Shape Discrimination (PSD)



IBD seen by detector:



Neutrino Candidate Rates

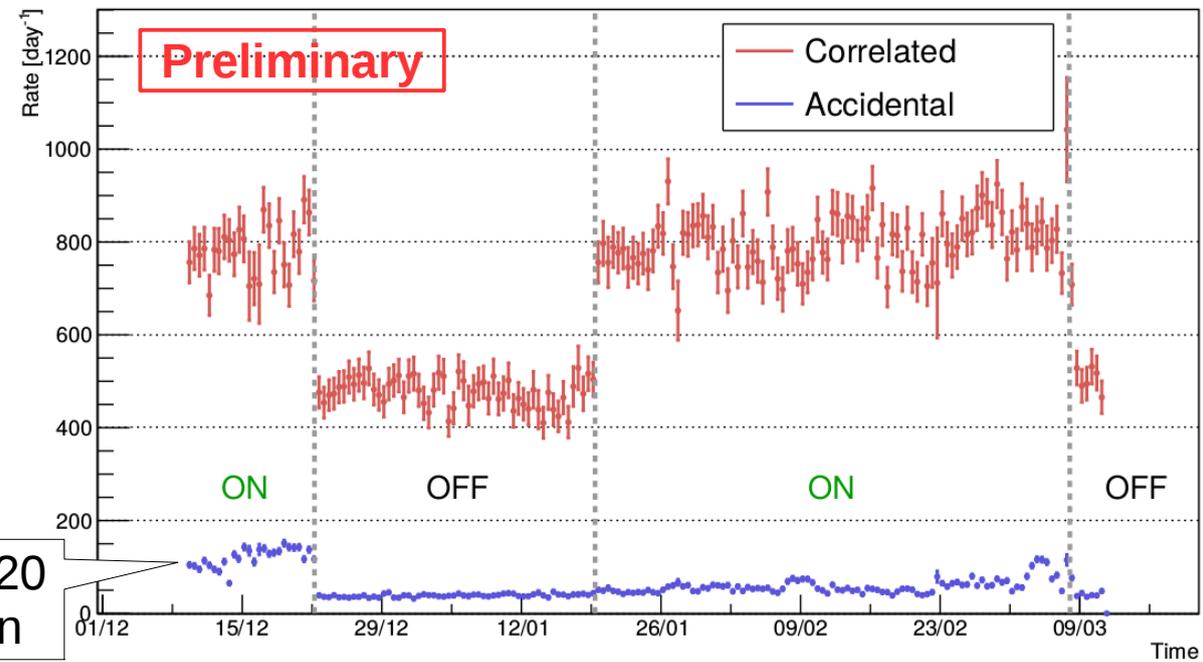


Neutrino candidate rate by comparing rates at reactor-on (75 days) vs. reactor-off (28 days):

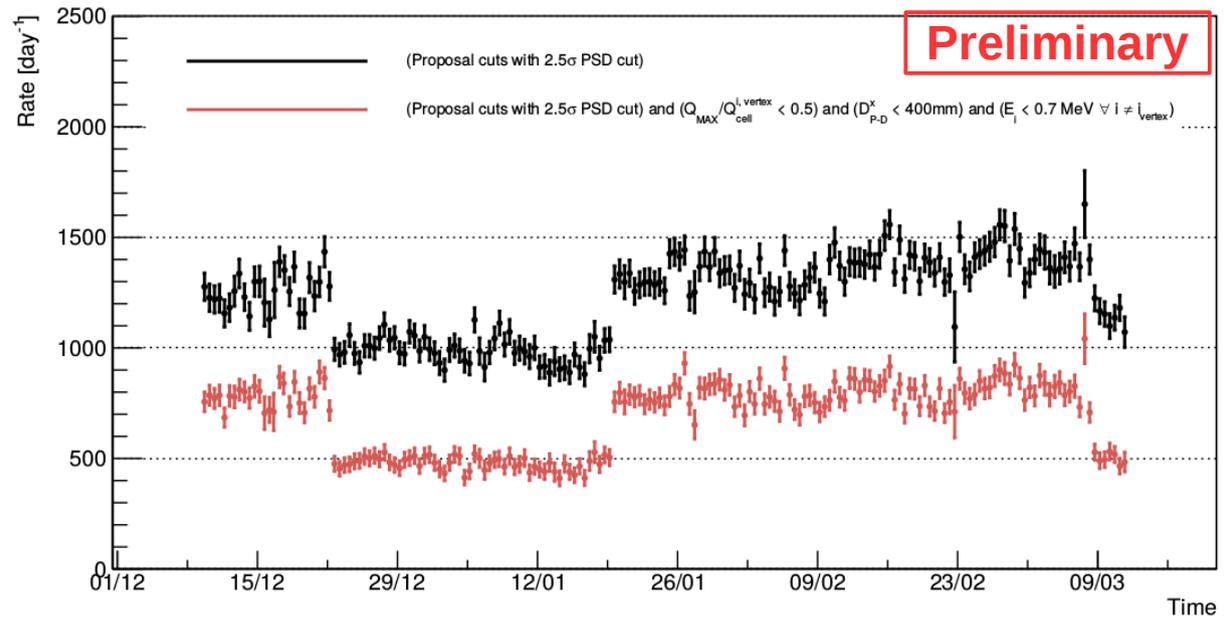
~300/day
(with current proposal cuts)

IN20
on

Rate evolution for event categories (atmospheric pressure corrected)



Evolution of the IBD candidates rate in the Stereo detector



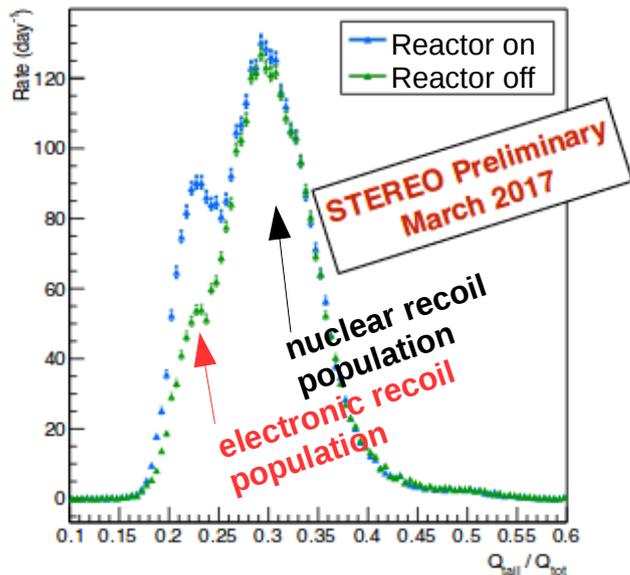
Neutrino Candidate Rates



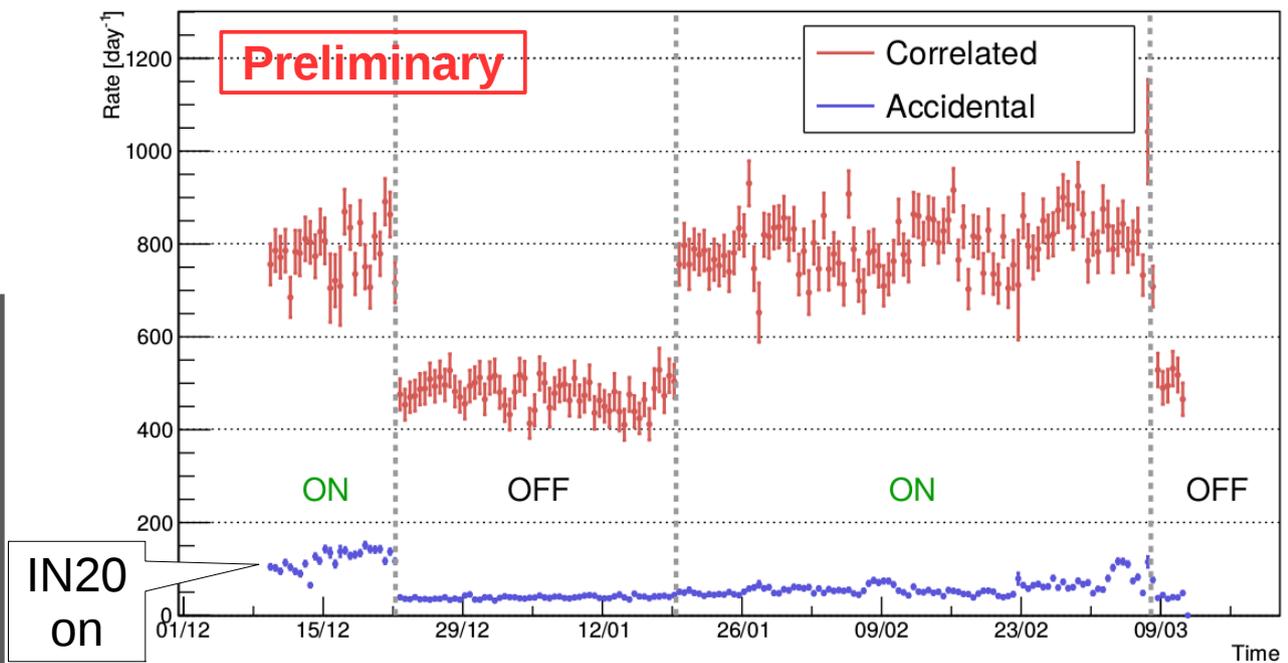
Neutrino candidate rate by comparing rates at reactor-on (75 days) vs. reactor-off (28 days):

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(with current proposal cuts)

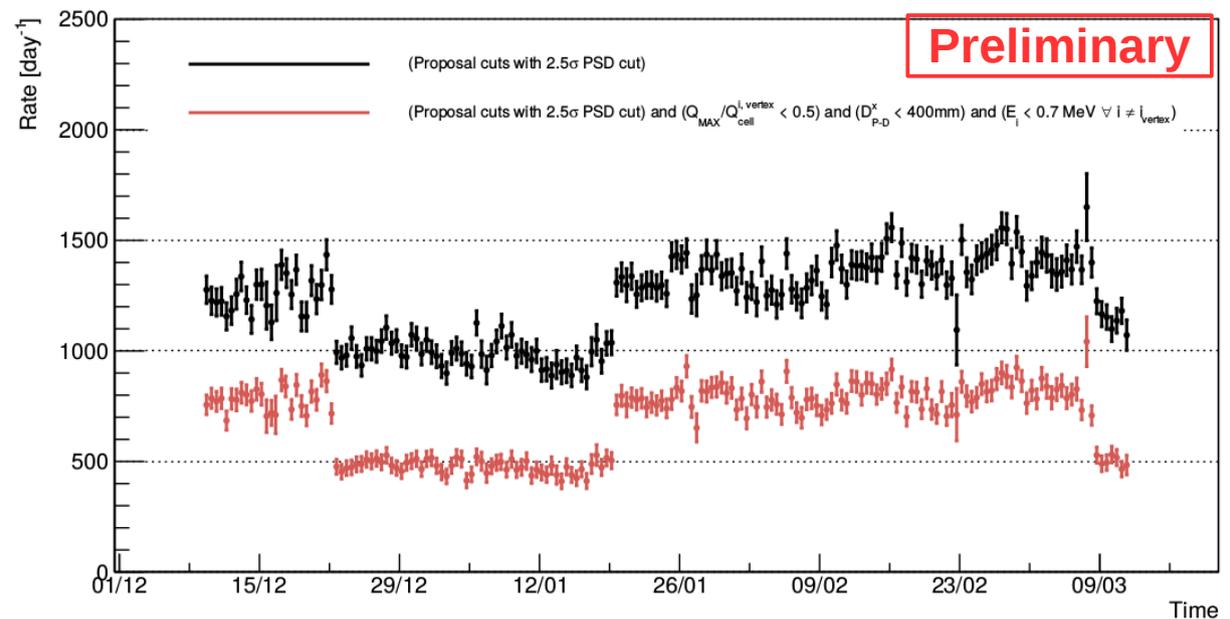
Comparison reactor-on vs. reactor-off via PSD
 → additional contribution only in signal region
 → shielding against reactor-induced background sufficient



Rate evolution for event categories (atmospheric pressure corrected)



Evolution of the IBD candidates rate in the Stereo detector





Summary/Outlook

Summary

- first data taking phase completed (75 days reactor on + 28 days reactor off)
- detector shows well-understood/stable response
- first signal and background selections yield promising results

Outlook

- cut optimisation
- finalisation of energy reconstruction
 - for spectral shape oscillation analysis
- studies of backgrounds + systematics
- re-installation of detector (currently maintenance at reactor site)
- further data taking
 - 2017: + 60 days reactor on
 - 2018: + ~150 days reactor on

The Stereo Collaboration



Supported by:





Appendix

The Stereo Timeline



Spring-Autumn 2016:

Mounting and installation of the STEREO detector and its shielding

9th November 2016:

ASN approval

⇒ detector filled and commissioned (2 weeks)

Winter 2016-2017:

Data taking of

75 days reactor ON +

28 days reactor OFF

3rd March 2017:

Reactor shutdown

(up to October 2017)

⇒ End of the first data acquisition phase

March-August 2017 :

STEREO retracted for

reactor works, used for

detector maintenance



Shielding types:

- 6 tons of borated polyethylene
- 65 tons of lead
- B_4C sheets all around the detector
- Magnetic shielding (soft iron + μ -metal)
- Casemate reinforcements (lead + borated polyethylene)

