

# Speeding up complex multivariate data analysis in Borexino with **parallel** computing based on Graphics Processing Unit

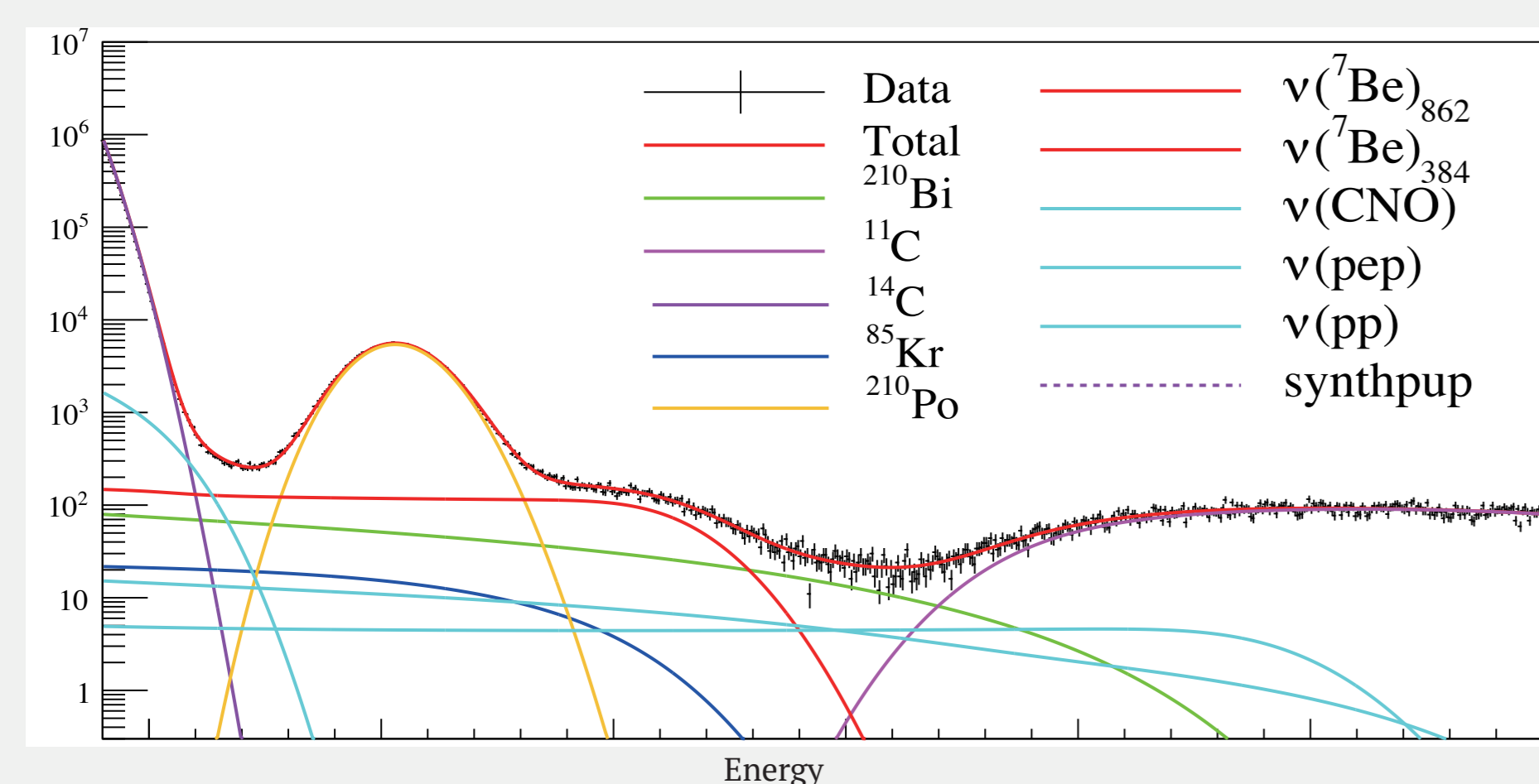
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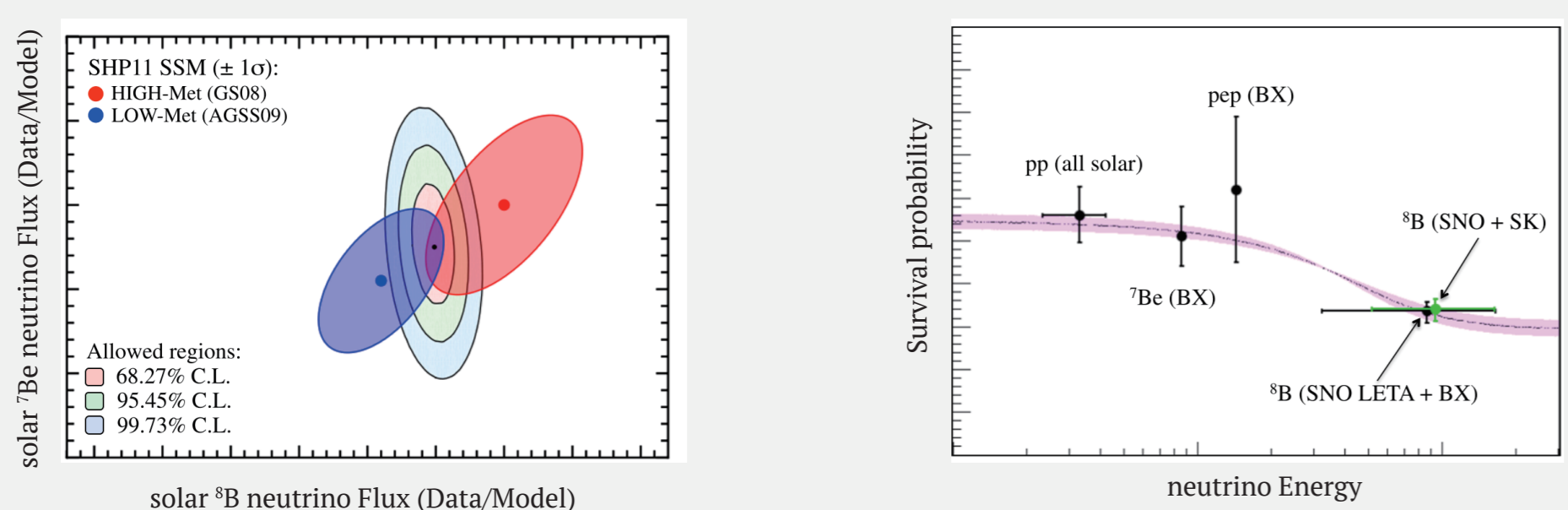


## Solar neutrino with Borexino

- Disentangle signal-background with **spectral fitting**

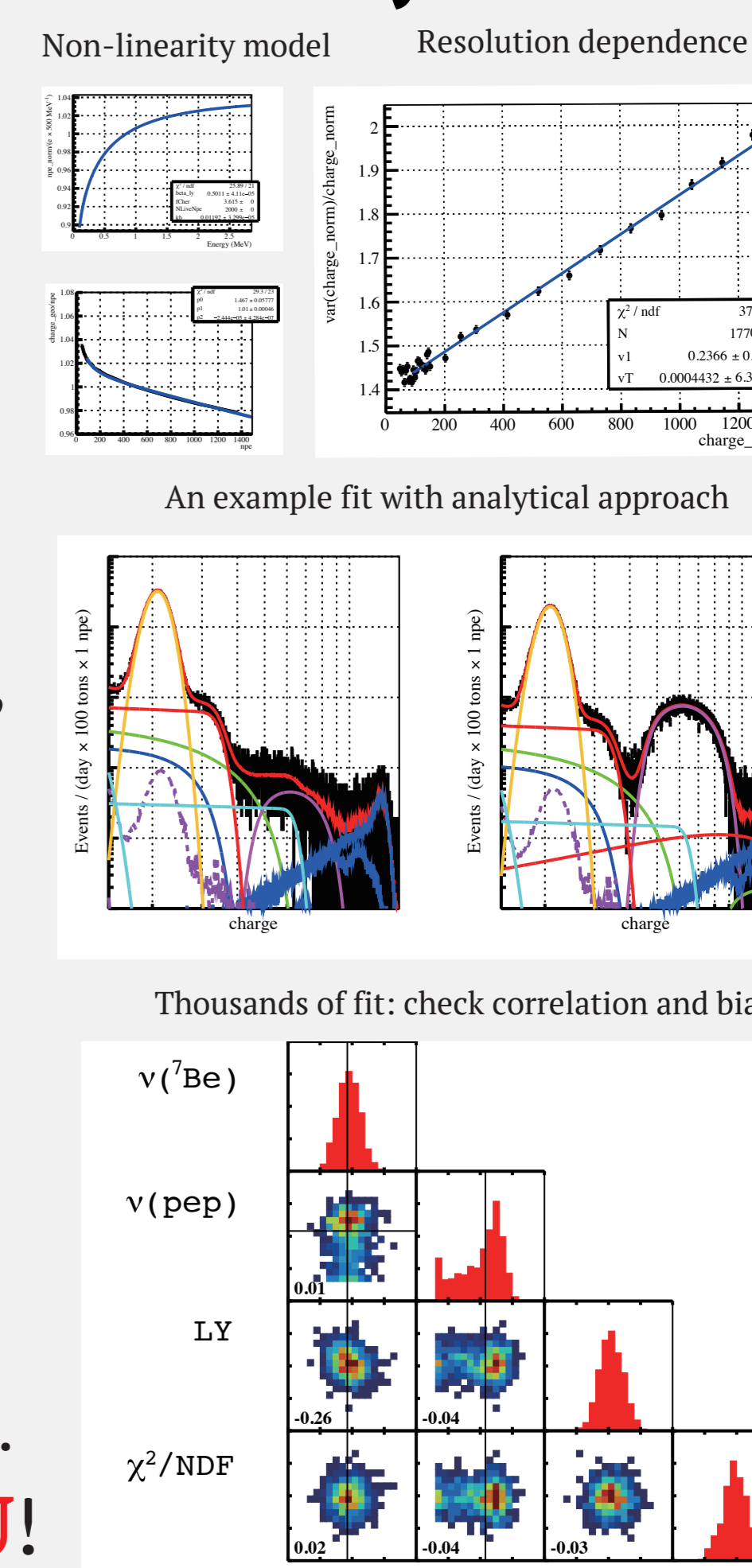


- Study the **neutrino flux** or the **neutrino oscillations**.



## Challenges and Complexity in Borexino spectral analysis

- Various **energy estimators** with advantage in different energy regions:
    - Number of fired PMTs, number of hits**: free of extra PMT charge resolution
    - Sum of Integrated PMT Charge**: more linear, better resolution at high energies
  - Need of analytical approach:
    - Fit the **light yield**. Important for understanding the solar pp neutrino.
    - Fit the **resolution**. For charge variable it can fit the single p.e. charge resolution.
  - Challenge to get the precise analytical detector response model
    - The shape of mono-energetic uniform-in-fiducial-volume events**: scaled poisson, modified gaussian, generalized gamma. Validation of it on Monte Carlo. [1]
    - Reliable and flexible non-linearity and resolution model**.
    - Temporal dependence of detector response**.
  - Describe **the pile-up with 14C effect**, critical effect for solar pp neutrinos.
  - Describe the spikes for the **normalized integer variable**.
  - Various studies need to be done to get robust fitting procedure and fit result:
    - Validation with correlation plot: fit of toyMC spectra, reproduce the injected rate
    - Scan of parameters: Varying fit range, binning, all fixed parameter etc.
    - Systematics: Inject deformation of detector response and see how stable is the fit.
- Common challenges for all experiments! Need to be fast! Need GPU!**



## Parallel computing

This project is based on C++11, ROOT, cuda and GooFit

Graphic Processing Unit: thousands of cores, data parallelization

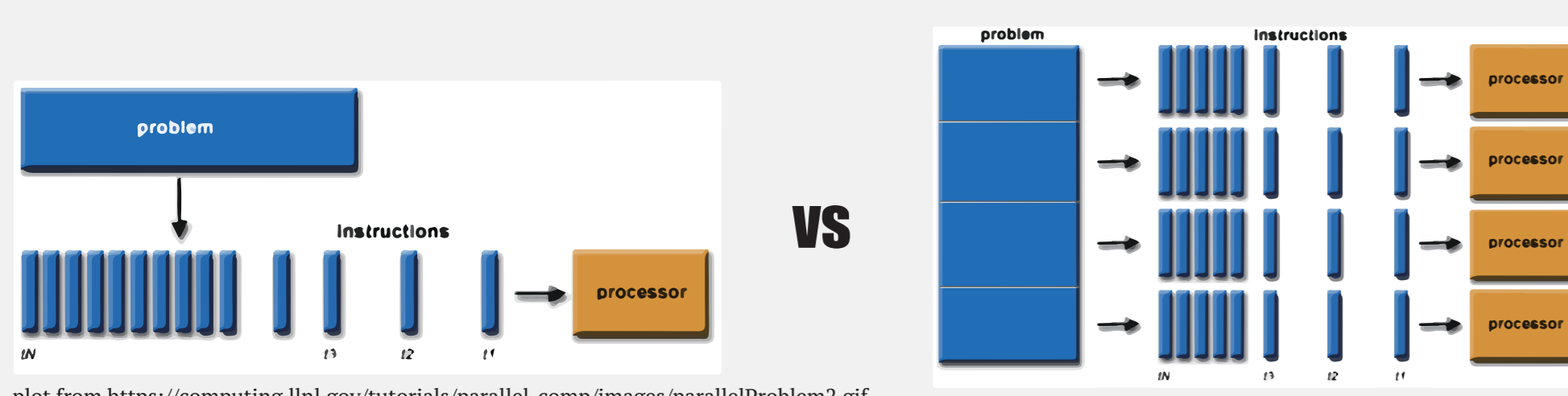
CPU: large cache, instruction parallelization

GPU: thousands of cores, data parallelization



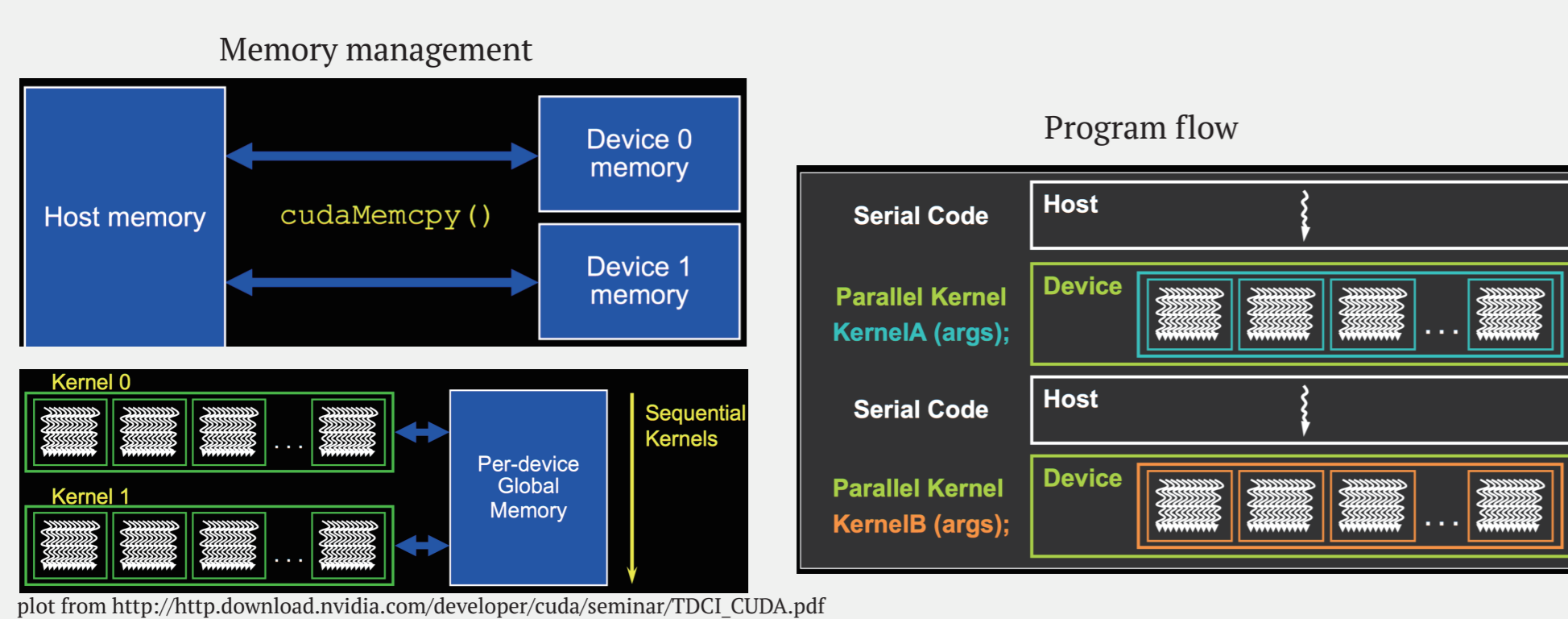
plot from [https://www.ogf.org/OGF25/materials/1605/CUDA\\_Programming.pdf](https://www.ogf.org/OGF25/materials/1605/CUDA_Programming.pdf)

Parallel computing: divide into small tasks and solved simultaneously



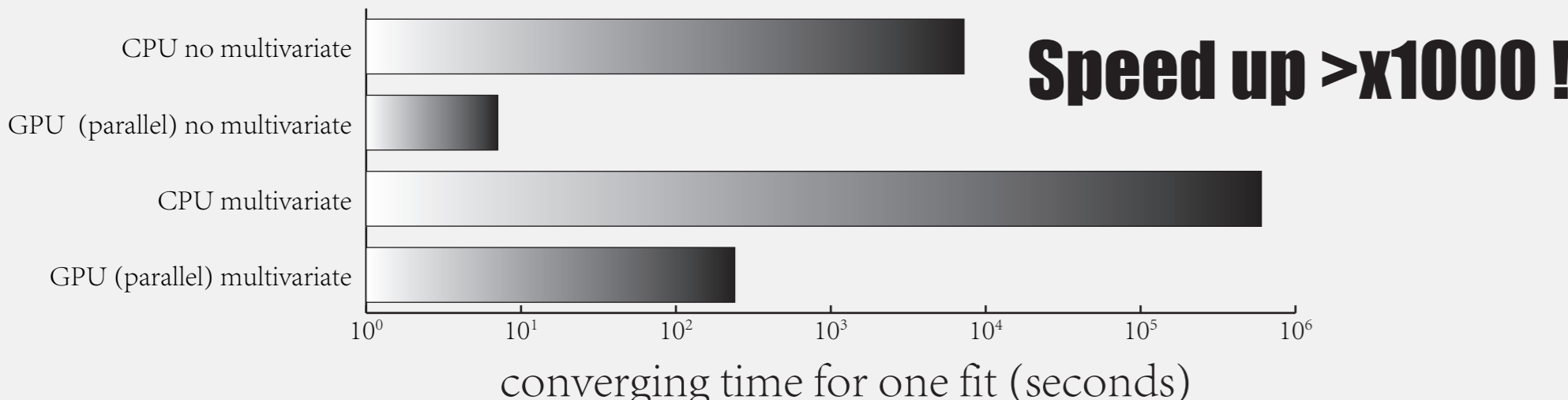
plot from [https://computing.llnwd.net/tutorials/parallel\\_comp/images/parallelProblem2.gif](https://computing.llnwd.net/tutorials/parallel_comp/images/parallelProblem2.gif)

Scheme of Graphic Processing Unit based parallelization



plot from [http://http.download.nvidia.com/developer/cuda/seminar/TDCI\\_CUDA.pdf](http://http.download.nvidia.com/developer/cuda/seminar/TDCI_CUDA.pdf)

## Performance and physics results



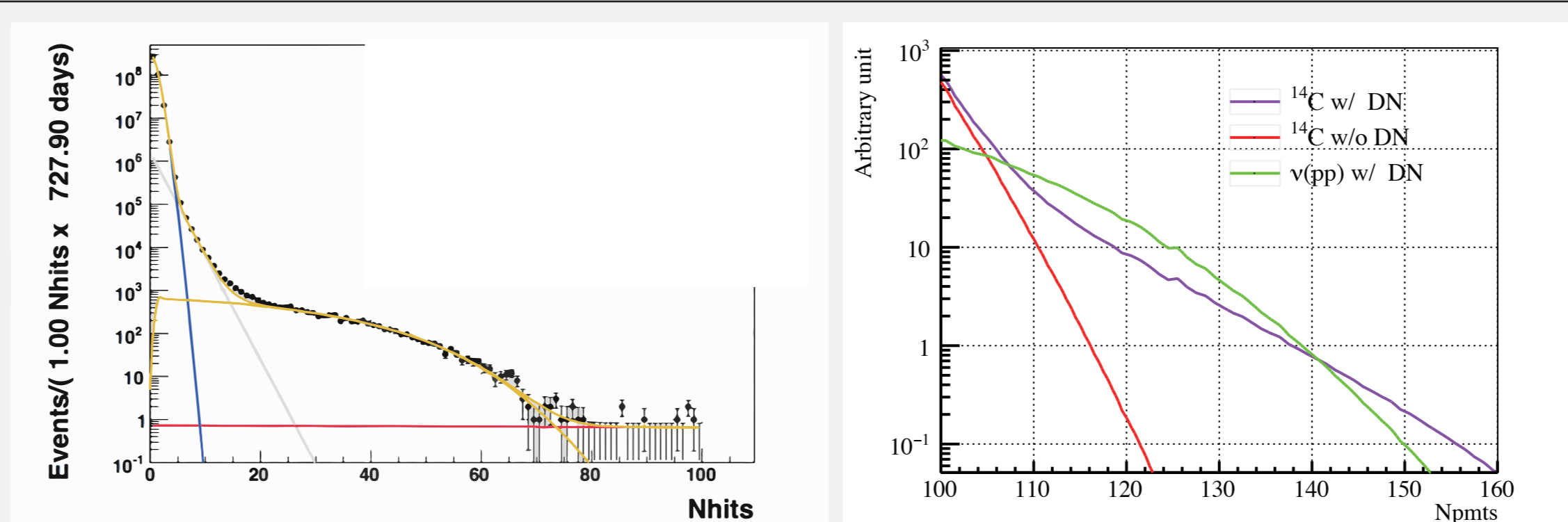
The physics result got with this tool

- The global spectral-fit of solar neutrino rate with analytical approach
- The systematic uncertainty, which is got with tens of thousands of fits
- The upper limit for the neutrino magnetic moment with phase-II data

## Physics motivated Statistical tool for Borexino analysis

### How to fit pp neutrino? Account for Pile-up from 14C

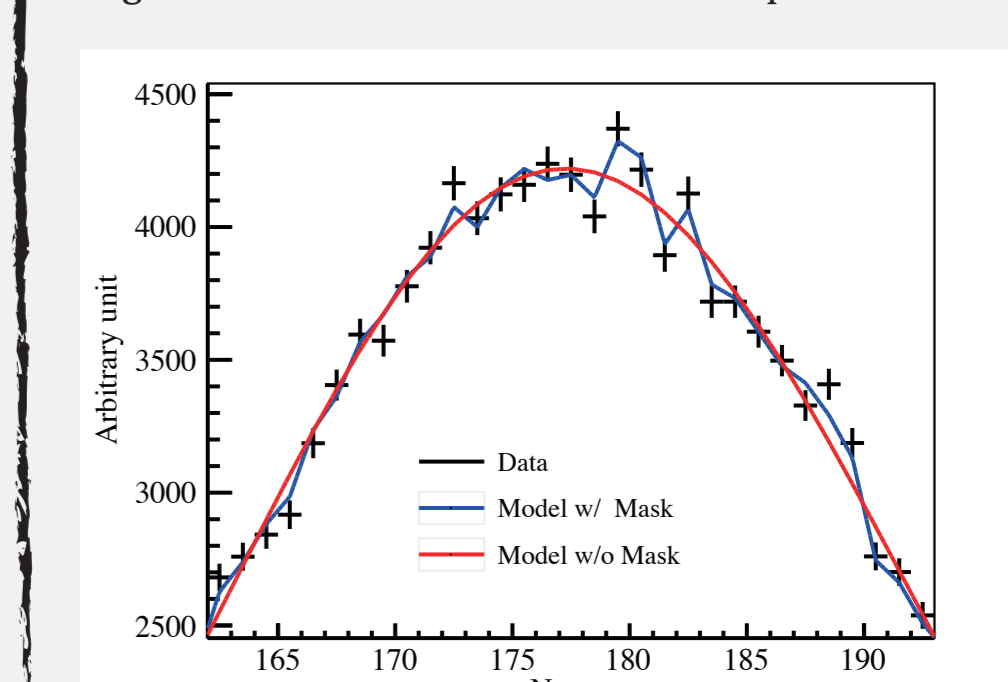
- The major background for pp neutrino is <sup>14</sup>C. The **pile-up effect modifies its shape** in region of interest of pp neutrino drastically.
- <sup>14</sup>C will **pile-up with everything**: <sup>14</sup>C, <sup>210</sup>Po etc.
- Take **spectrum with random trigger**, which is the spectrum to be piled on <sup>14</sup>C
- Convolve the analytical 14C shape with such spectrum** to account for the pile-up effect



Spectrum acquired with random trigger Primary bkg. for v(pp): <sup>14</sup>C-<sup>14</sup>C pile up

### Spikes from integer

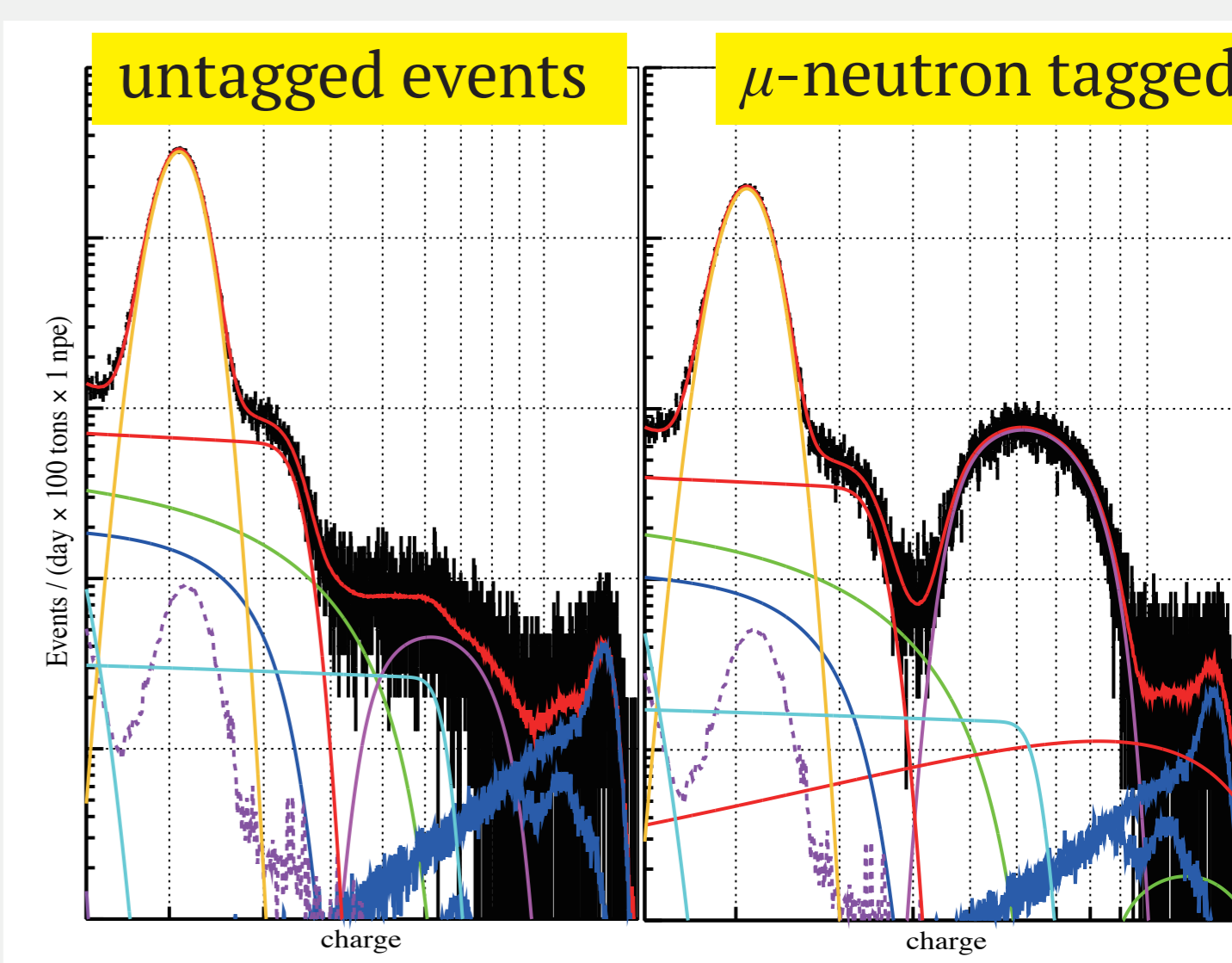
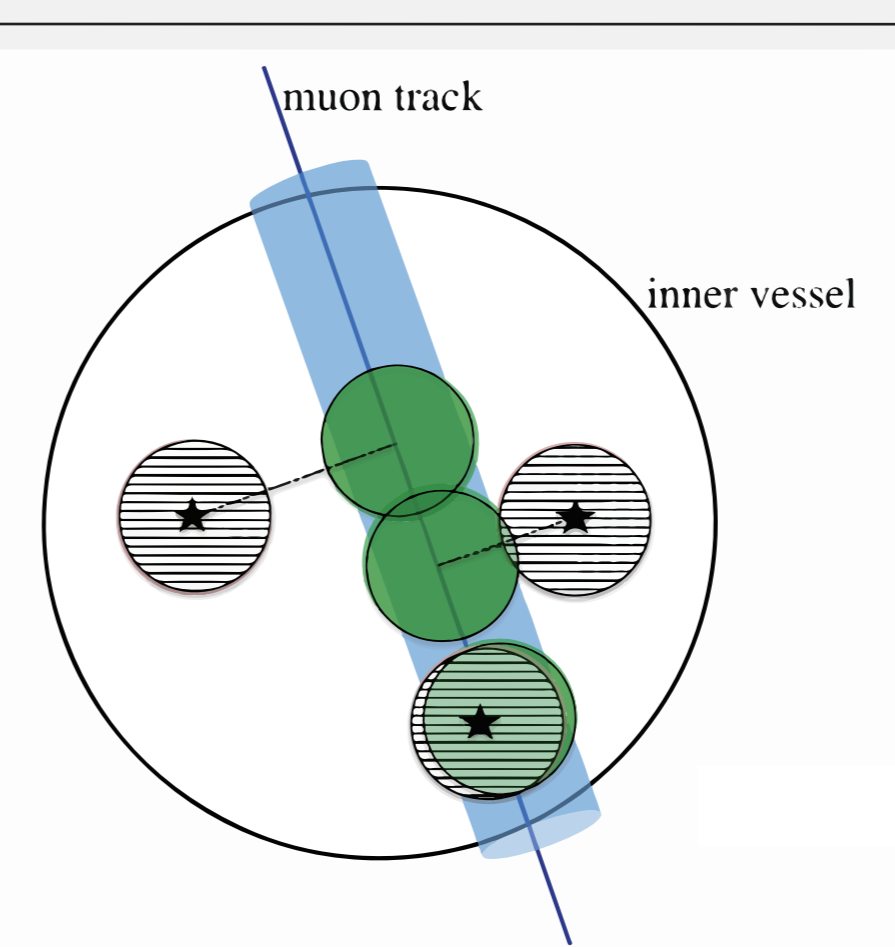
Use number of fired pmts as energy estimator, which is integer. Need correction to account for spikes due to binning.



Red vs Blue: w/ vs w/o correction

### reject <sup>11</sup>C with μ-neutron tagging: Complementary

- cosmogenic 11C** is the major background for pep/CNO neutrinos
- The production of **cosmogenic 11C** is usually **associated with one or more spallation neutrons**.
- If a physics event is associated with a muon track and a tagged neutron, it is tagged to be more likely a cosmogenic <sup>11</sup>C.
- Rather than cut these events, we build another spectrum and sum up the likelihood of both. **Sensitivity on solar pp and 7Be neutrino species is improved**.
- For analytical approach, the enhanced <sup>11</sup>C **improves the knowledge on the resolution model**.
- Massive tests have been performed and this method is validated.

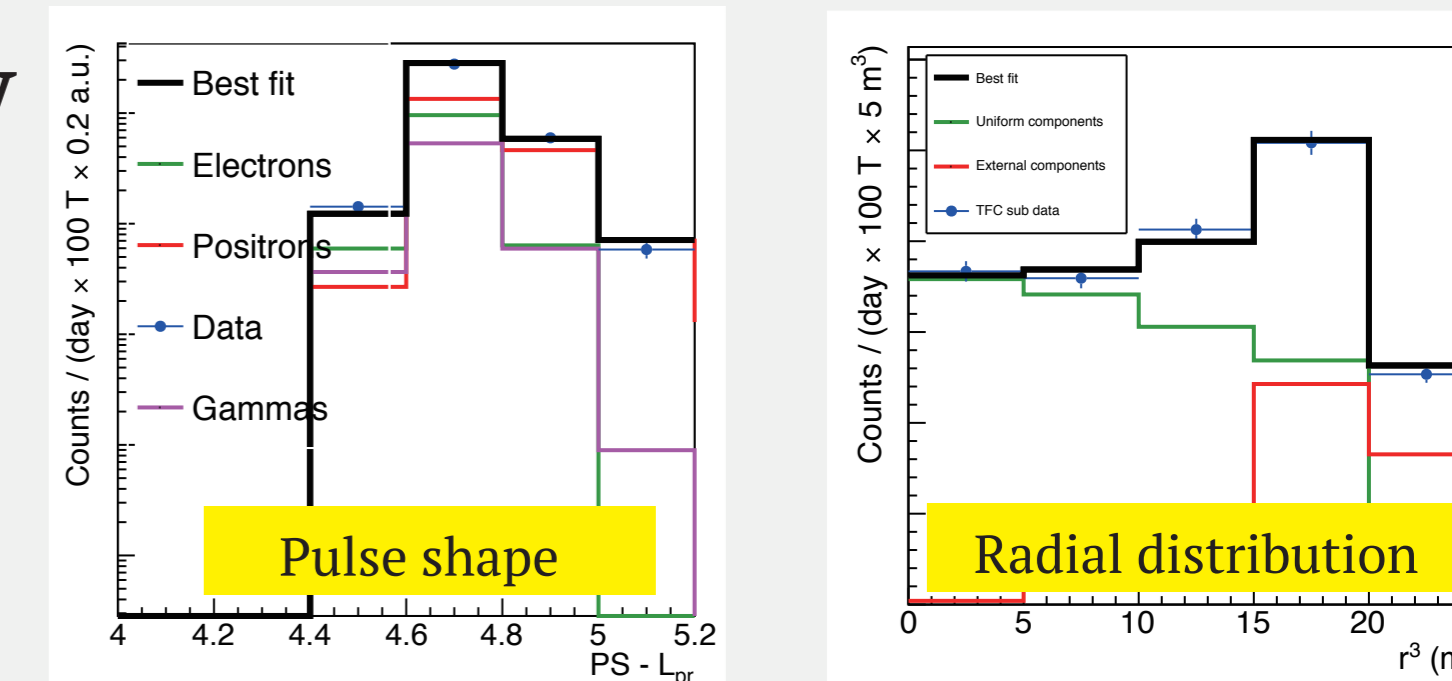


### Further improvement of pep and CNO sensitivity

**β<sup>+</sup>/β Pulse Shape Discrimination** to suppress <sup>11</sup>C  
**Radial Distribution** to suppress external γ

Integrate PSD/Radial distribution into the likelihood:

**Multivariate fit**



And... **Year-by-year simultaneous fitting**, constrain neutrino to be constant while background different each year, North/South splitting, constrain light yield, resolution etc. follow **linear trend over time** etc.

[1] Oleg Smirnov. NIM, 595, 2008 (410-418) [2] Bellini, G. et al. PRD, 89, 112007, 2014 [3] New Borexino paper