

Recent Solar neutrino Results from Super-Kamiokande

Yuuki Nakano (Kamioka Observatory)

for the Super-Kamiokande collaboration

26th, July 2017 (Wed)

TAUP 2017 @ Sudbury, Canada.



Supported by Grant-in-Aid for
Young Scientists (B) 17K17880



Contents

- **Super-Kamiokande**
 - Detector
 - Physics motivation
 - Recent progress in solar neutrino analysis
- **^8B solar neutrino flux measurement**
 - Observed signal
 - Yearly flux measurement
 - Periodic analysis
- **Spectrum analysis**
- **Summary**

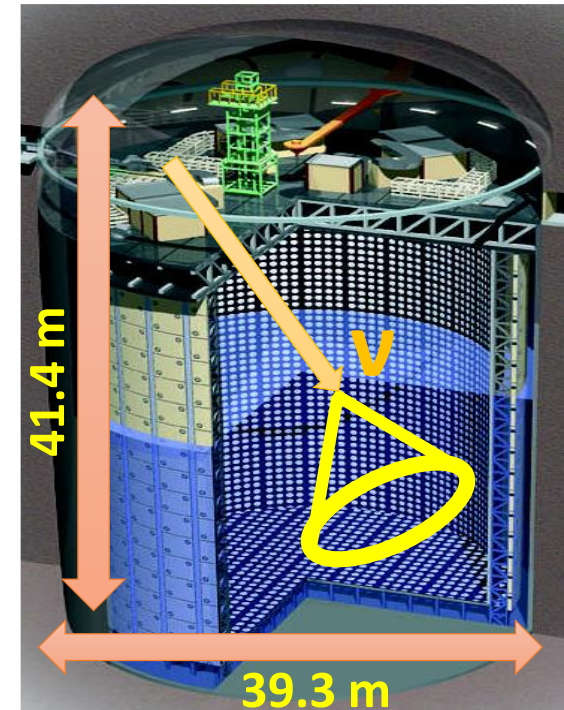
Super-Kamiokande

• Super-Kamiokande detector

- Located at Kamioka, Japan
 - 1000 m under Ikenoyama mountain
 - 2700 m water equivalent
- 50 kton ultra pure water tank
 - More than 11,000 20-inch PMTs for ID
 - 22.5 kton for the fiducial volume
- Water Cherenkov technique
 - Energy, direction, particle ID

• Many physics targets

- Astrophysical neutrino (Solar, Supernova)
- Atmospheric neutrino
- Proton decay
- Long base line neutrino (T2K)
- Dark matter search etc...



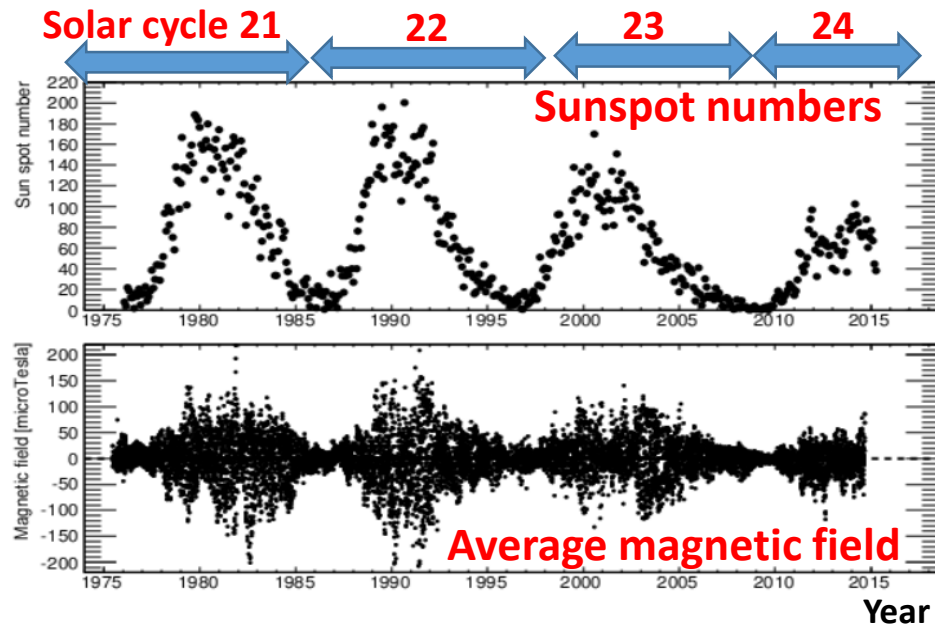
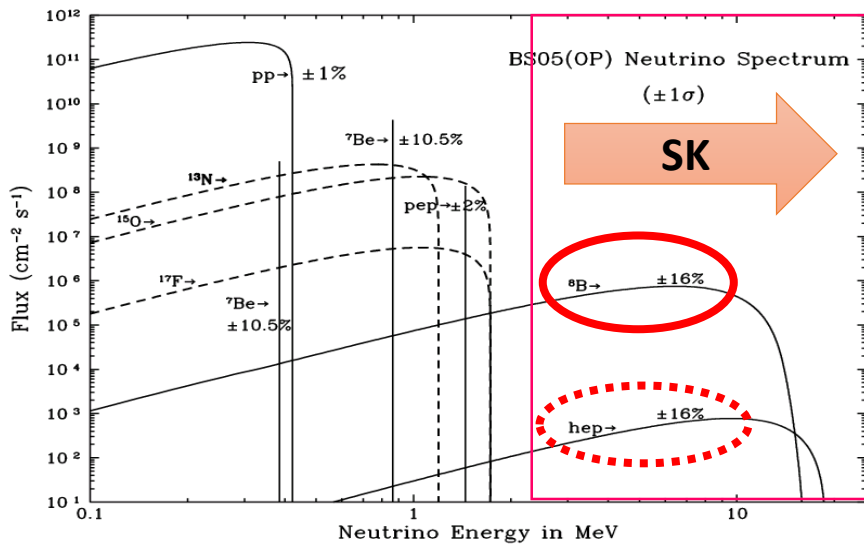
Solar ν
3.5-19.5 MeV

Supernova ν
20~100.0 MeV

Atmospheric ν and proton decay
~100 MeV GeV TeV PeV

Physics motivation (1)

- **Standard Solar Model (SSM)** describes the profile of the Sun well.
 - SSM predicts the solar neutrino fluxes.
 - Super-K has observed ^8B neutrino because of its high energies ($< \sim 14$ MeV).
- **Several periodical solar activities** are observed.
 - Variation of the sunspot number at the surface (~ 11 years).
 - The Sun itself oscillates due to the acoustic waves (~ 5 minutes).
- **Searching for the periodic modulation of ν fluxes** is interesting.



Left figure: *Astrophys. J.* 621 85 (2005).

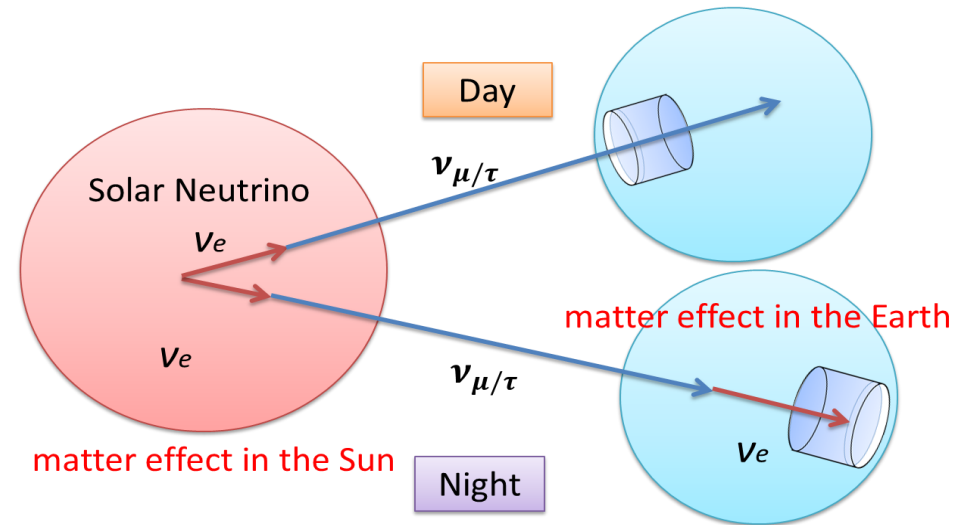
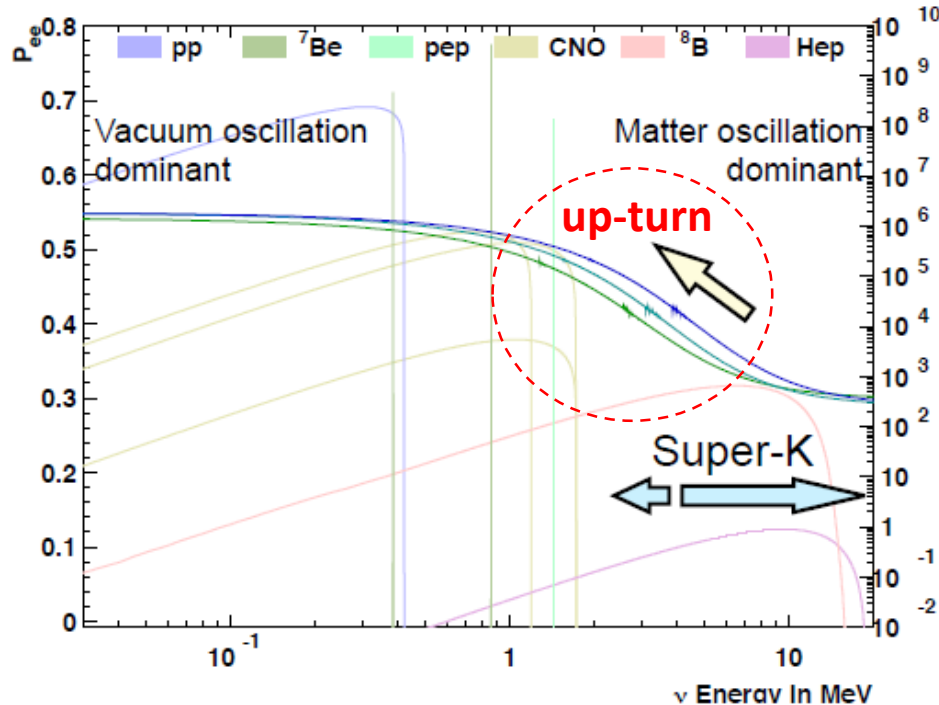
Sunspot number is taken from <http://solarscience.msfc.nasa.gov/greenwch/spotnum.txt>

Average magnetic field is taken from: http://wso.stanford.edu/meand/MF_timeseries.txt

Physics motivation (2)

- Neutrino oscillation due to the matter effect (MSW effect)
 - Spectrum “**up-turn**” expected by MSW effect in the Sun.
 - MSW effect leads to a resonant conversion of the higher energy ν .
 - Day/Night flux asymmetry due to terrestrial matter effect.
 - Regeneration of electron- ν is expected when ν pass through the Earth.

$$i \frac{d}{dt} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} -\frac{\Delta m^2}{4E} \cos 2\theta + \sqrt{2} G_F N_e & \frac{\Delta m^2}{4E} \sin 2\theta \\ \frac{\Delta m^2}{4E} \sin 2\theta & \frac{\Delta m^2}{4E} \cos 2\theta \end{pmatrix} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix}$$

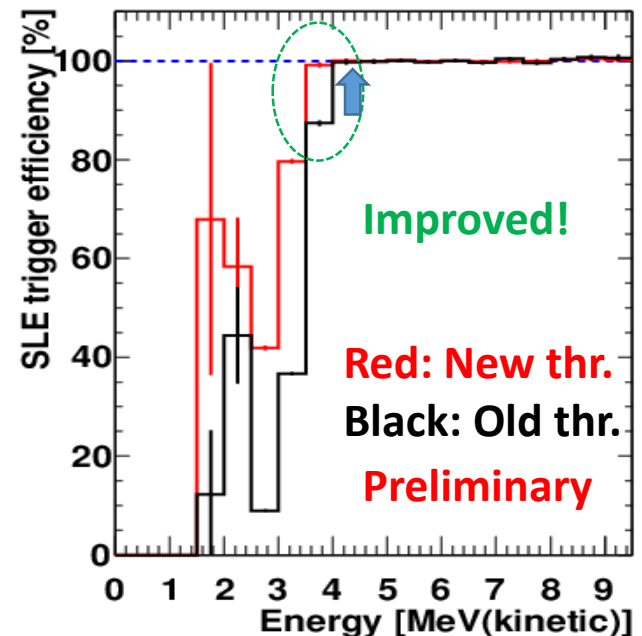


$$A_{\text{DN}} = \frac{\Psi_{\text{day}} - \Psi_{\text{night}}}{(\Psi_{\text{day}} + \Psi_{\text{night}})/2}$$

Super-K Day/Night result
Phys. Rev. Lett. 112 (2014) 091805.

Solar neutrino analysis in Super-K

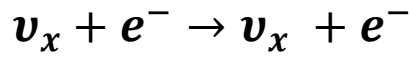
- Last year, we submitted a paper about the solar neutrino analysis results (using SK-IV 1664 days data sample).
 - **Phys. Rev. D 94, 052010 (2016), arXiv: 1606.07538**
- Updated results are presented.
 - ^8B solar neutrino flux with SK-IV 2365 days sample (5200 days).
 - Energy spectrum with SK-IV 2645 days sample (5480 days).
- Recent progress of solar analysis.
 - Yearly ^8B flux plot to see any correlation with solar activity.
 - Periodic modulation of solar ν flux.
 - Lower trigger threshold in May 2015.
 - Detection efficiency in 3.5-4.0 MeV_{kin.}
~84% → ~99%.



Observed ^8B solar neutrino signal

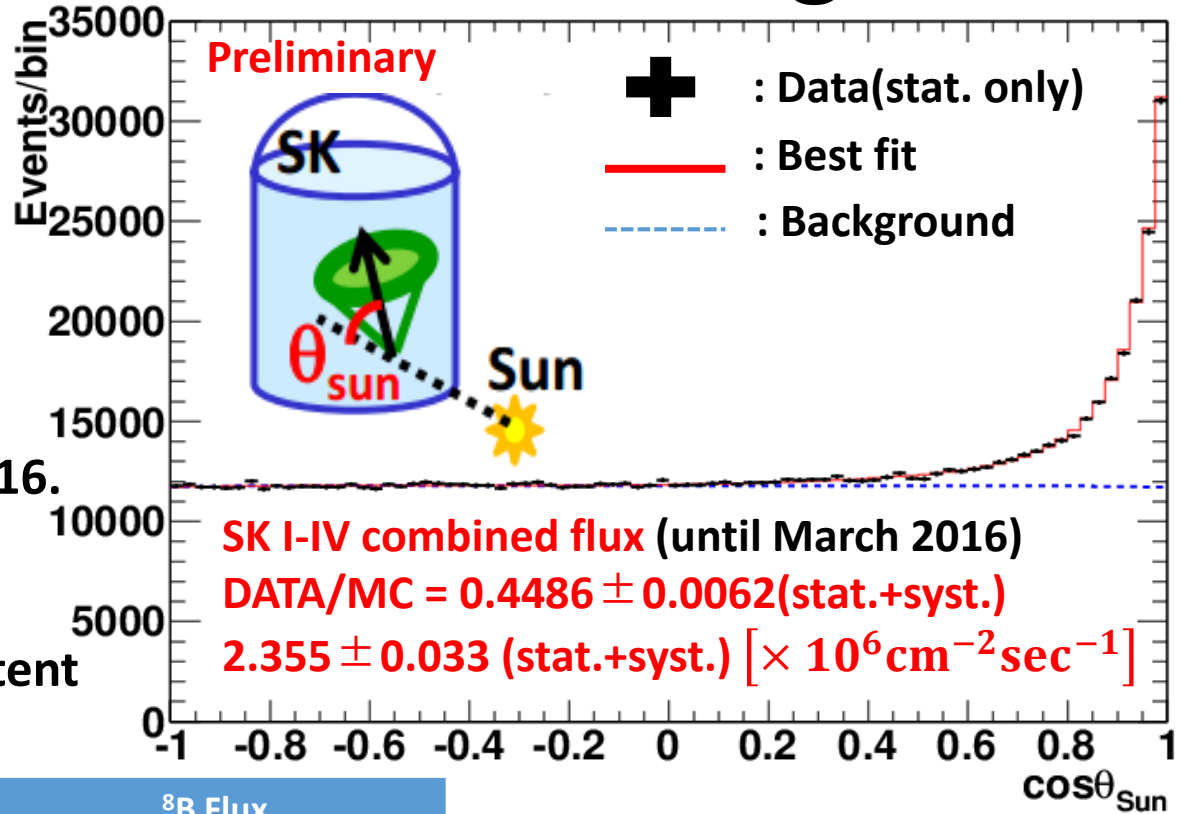
◆ ^8B neutrino measurement

Cherenkov light generated by recoil electron scattered with ν .



A total of **84k** solar neutrinos were observed until March 2016. (89k events until March 2017)

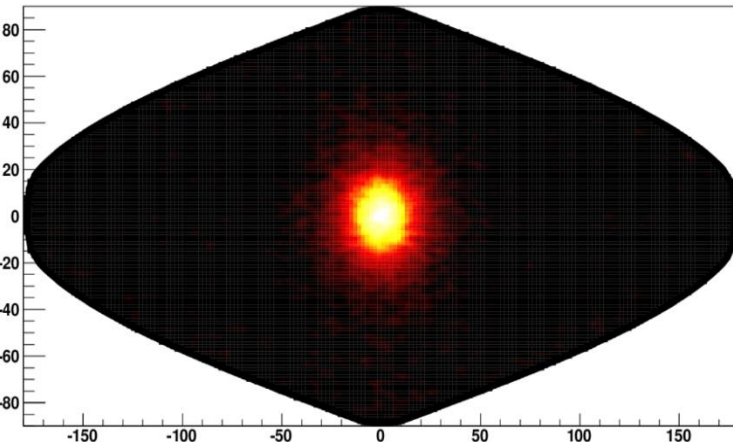
Measured ^8B fluxes are consistent within uncertainties.



SK phase	Energy threshold [MeV(kin)]	Live time [day]	^8B Flux [$\times 10^6/\text{cm}^2/\text{sec}$]
SK I	4.5-19.5	1496	$2.38 \pm 0.02 \pm 0.08$
SK II	6.5-19.5	791	$2.41 \pm 0.05^{+0.16}_{-0.15}$
SK III	4.0-19.5	548	$2.40 \pm 0.04 \pm 0.05$
SK IV	3.5-19.5	2365 2645	$2.32 \pm 0.02 \pm 0.04$ Under preparation

MC: $5.25 \times 10^6/\text{cm}^2/\text{sec}$

(SNO: NC current, Phys. Rev. C88 (2013) 022501.)



^8B solar neutrino yearly flux

◆ Solar activity cycle

Sun spot numbers are strongly correlated with the solar activity cycle (~11 years).

SK has observed ^8B solar neutrino for ~20 years (More than 1.5 cycle).
Data taken until March 2016 is used.

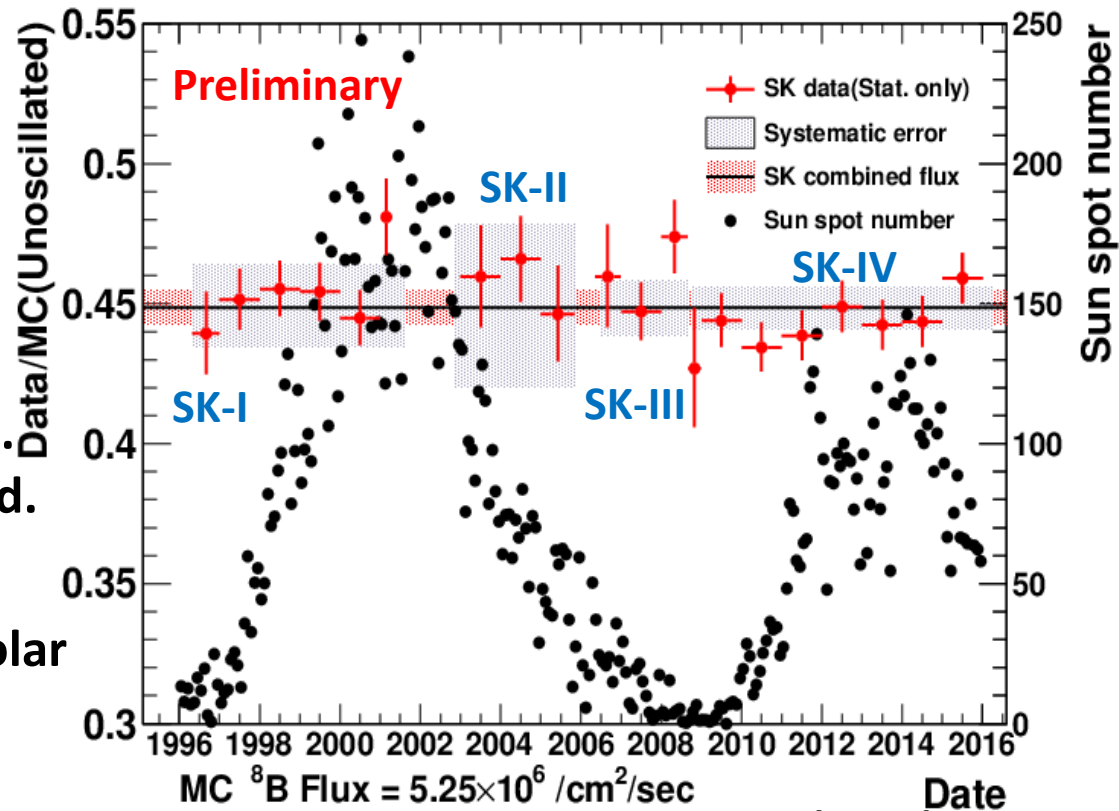
◆ ^8B flux vs. sun spot

No correlation with the 11 years solar activity is observed.

$$\chi^2 = 15.52/19(\text{D. O. F})$$

$$\text{Prob.} = 68.9\% \quad \text{Preliminary}$$

Super-K solar rate measurements are fully consistent with a constant solar neutrino flux emitted by the Sun.



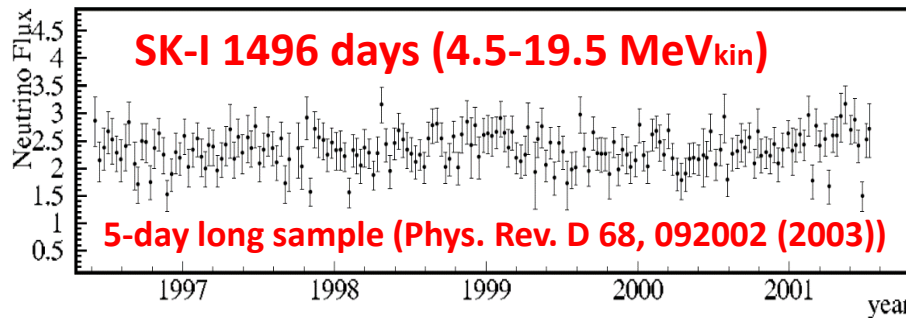
SK I-IV combined flux (until March 2016)
DATA/MC = 0.4486 ± 0.0062 (stat.+syst.)
 2.355 ± 0.033 (stat.+syst.) [$\times 10^6 \text{ cm}^{-2} \text{ sec}^{-1}$]

Sun spot number: <http://www.sidc.be/silso/datafiles>

Source: WDC-SILSO, Royal Observatory of Belgium, Brussels.

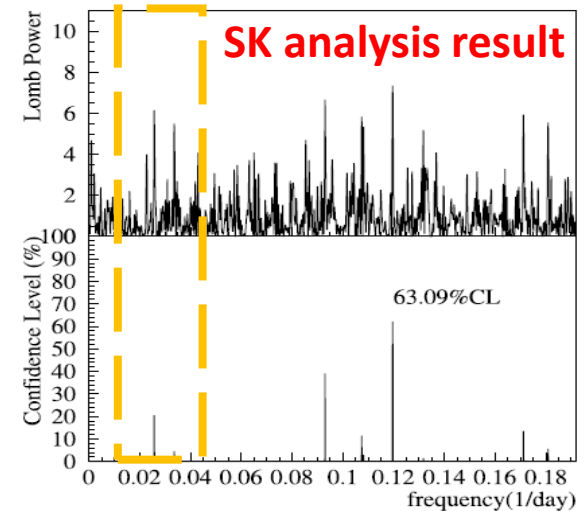
Periodic modulation analysis

- SK collaboration reported the time variation of 5-day long sample of the observed ^8B ν flux (Phys. Rev. D 68, 092002 (2003)).
- SK performed a periodic analysis using Lomb-Scargle (LS) method.

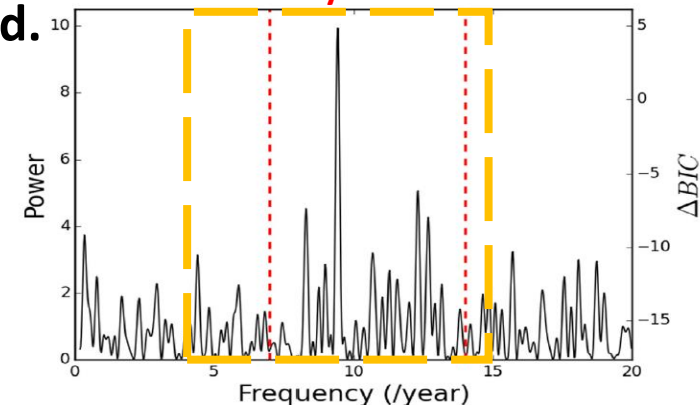


LS method

No clear periodic signal 5-15 year⁻¹.



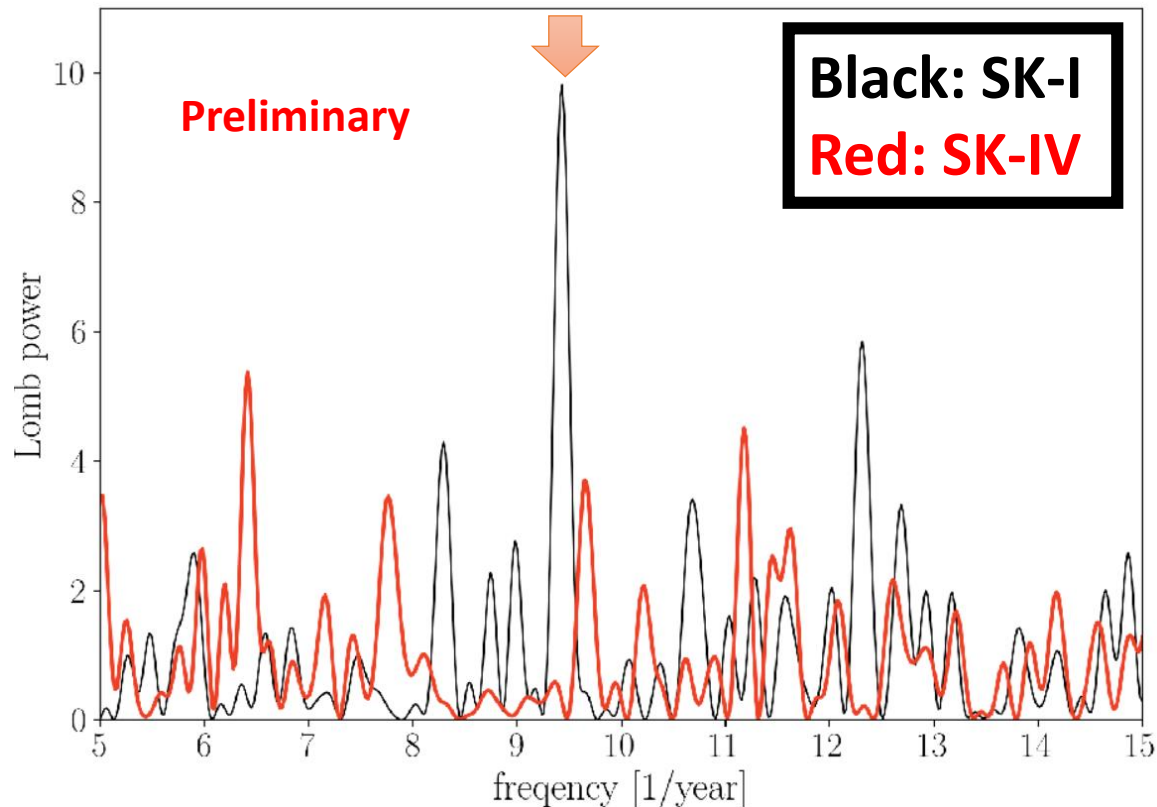
Several researchers found a peak at around 9.42 year⁻¹



- Several papers reported that a maximum peak is observed at around 9.42 year⁻¹.
 - Cf.) Astropart. Phys. 82, 86-92 (2016).
 - Generalized Lomb-Scargle (GLS) method is used.
- SK has reanalyzed SK-I data with GLS method provided by astroML.
- SK-IV data is also analyzed with GLS.

Periodic modulation results

- Using the Generalized LS method, both SK-I and SK-IV are analyzed.
- 5-day long sample is made from SK-I data and SK-IV data.
 - SK-I: 1496 days data (4.5-19.5 MeV_{kin}), Phys. Rev. D 68, 092002 (2003).
 - SK-IV: 1664 days data (4.5-19.5 MeV_{kin}), Phys. Rev. D 94, 052010 (2016).
- Search region 5-15 year⁻¹.
- **Maximum peak at around 9.42 year⁻¹ is not found in SK-IV.**



Global oscillation analysis input

◆SK

- SK-I 1496 days, Spectrum : 4.5-19.5MeV(kin.) + D/N : $E_{kin} \geq 4.5\text{MeV}$
- SK-II 791 days, Spectrum : 6.5-19.5MeV(kin.) + D/N : $E_{kin} \geq 7.0\text{MeV}$
- SK-III 548 days, Spectrum : 4.0-19.5MeV(kin.) + D/N : $E_{kin} \geq 4.5\text{MeV}$
- SK-IV 2645 days, Spectrum : 3.5-19.5MeV(kin.) + D/N (1664 days) : $E_{kin} \geq 4.5\text{MeV}$
Spectrum (until March 2017): Updated from Phys. Rev. D 94, 052010 (2016).
Day/Night flux: Updated from PRL 112 (2014) 091805.

◆SNO

- Parameterized analysis (c_0, c_1, c_2, a_0, a_1) of all SNO phased published in Phys. Rev. C88 (2013) 025501.

The same method is applied to both SK and SNO with a_0 and a_1 to LMA expectation.

◆Radiochemical (Ga, Cl)

- Ga rate 66.1 ± 3.1 SNU (All Ga global), Phys. Rev. C80 (2009) 015807.
- Cl rate 2.56 ± 0.23 SNU, Astrophys. J. 496 (1988) 505.

◆Borexino

- ${}^7\text{Be}$ flux, Phys. Rev. Lett. 107 (2011) 141302. **Does NOT include Borexino pp 2014. Nature 512 (2014) 383.**

◆KamLAND reactor

- 3-flavor analysis , Phys. Rev. D88 (2013) 033001.

◆ ${}^8\text{B}$ spectrum

- Winter 2006, Phys. Rev. C73 (2006) 025503.

SK I-IV combined recoil electron spectrum

◆ Spectrum shape

SK searches for the “**up-turn**” in its recoil electron energy spectrum.

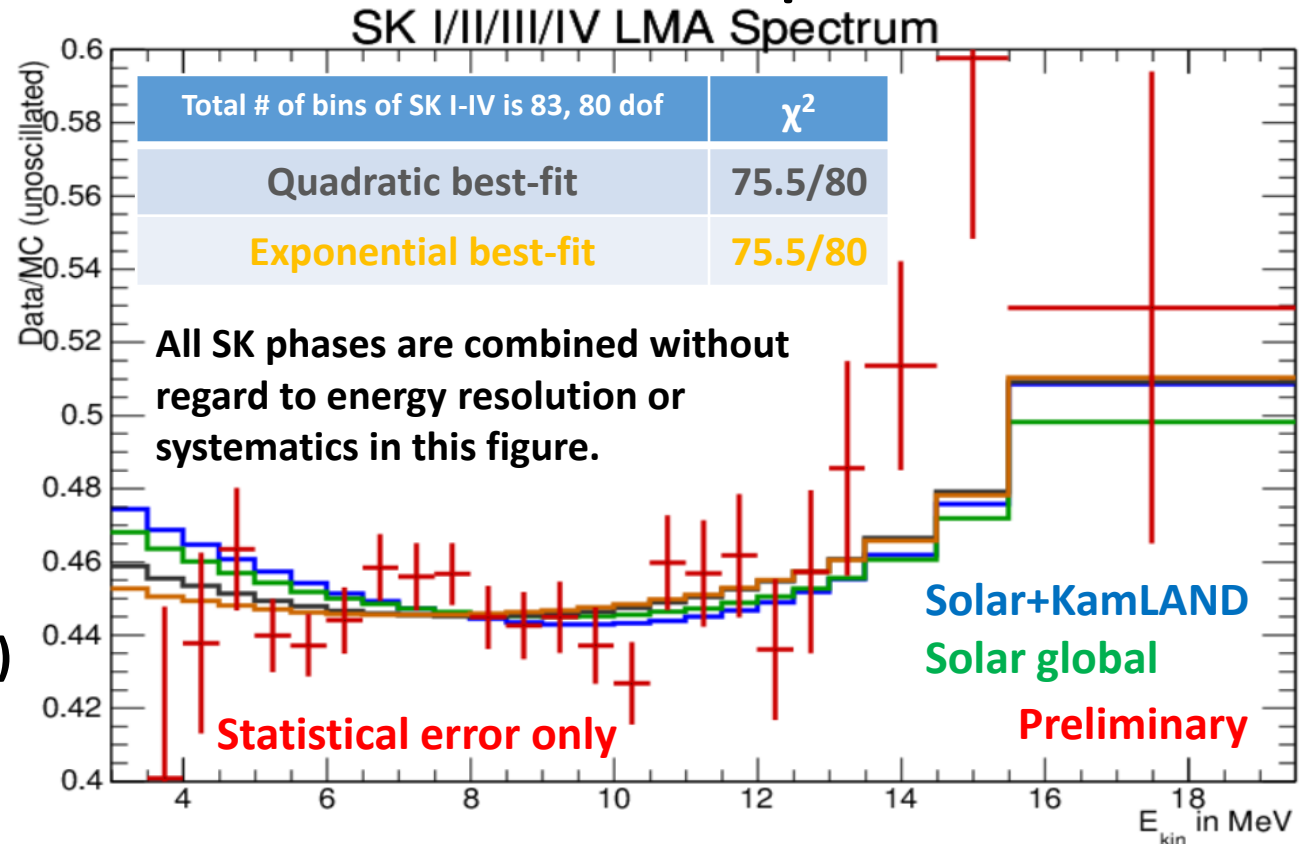
SK-I: 1496 days ($E \geq 4.5$)

SK-II: 791 days ($E \geq 6.5$)

SK-III: 548 days ($E \geq 4.0$)

SK-IV: 2645 days ($E \geq 3.5$)

Total: 5480 days sample



MC: $5.25 \times 10^6/\text{cm}^2/\text{sec}$ (SNO: NC current, Phys. Rev. C88 (2013) 025501.)

The SK recoil electron spectrum is consistent within $\sim 1\sigma$ with the MSW up-turn for the solar global best fit parameters, and marginally consistent within $\sim 2\sigma$ with the MSW up-turn for the solar+KamLAND best fit parameters.

Summary

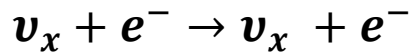
- Super-Kamiokande has taken solar neutrinos for more 20 years.
- Precise solar neutrino flux measurement is performed.
 - SK-IV result: $2.32 \pm 0.02(\text{stat.}) \pm 0.04(\text{syst.}) [\times 10^6/\text{cm}^2/\text{sec}]$.
- SK solar neutrino flux measurements agree across all phase.
 - Combined result: $2.355 \pm 0.033 (\text{stat.}+\text{syst.}) [\times 10^6/\text{cm}^2/\text{sec}]$.
 - Fully consistent with a constant solar neutrino flux emitted by the Sun.
 - **No correlation with the solar activity cycle is seen.**
- Preliminary periodic modulation analysis is performed.
 - Maximum peak at around 9.43 year^{-1} is **not found** in SK-IV.
- Recoil electron energy spectrum is **consistent with the MSW up-turn.**
 - $\sim 1\sigma$ for the solar global best-fit parameters
 - $\sim 2\sigma$ for the solar+KamLAND best-fit parameters.

Back up

Typical low energy event in SK

◆ How to detect

Elastic scattering(ES) reaction is used for solar neutrinos



◆ Reconstruction

- Timing information → Vertex position
- Ring pattern → Direction
- Number of hit PMTs → Energy (~6hits/MeV)

◆ Resolutions

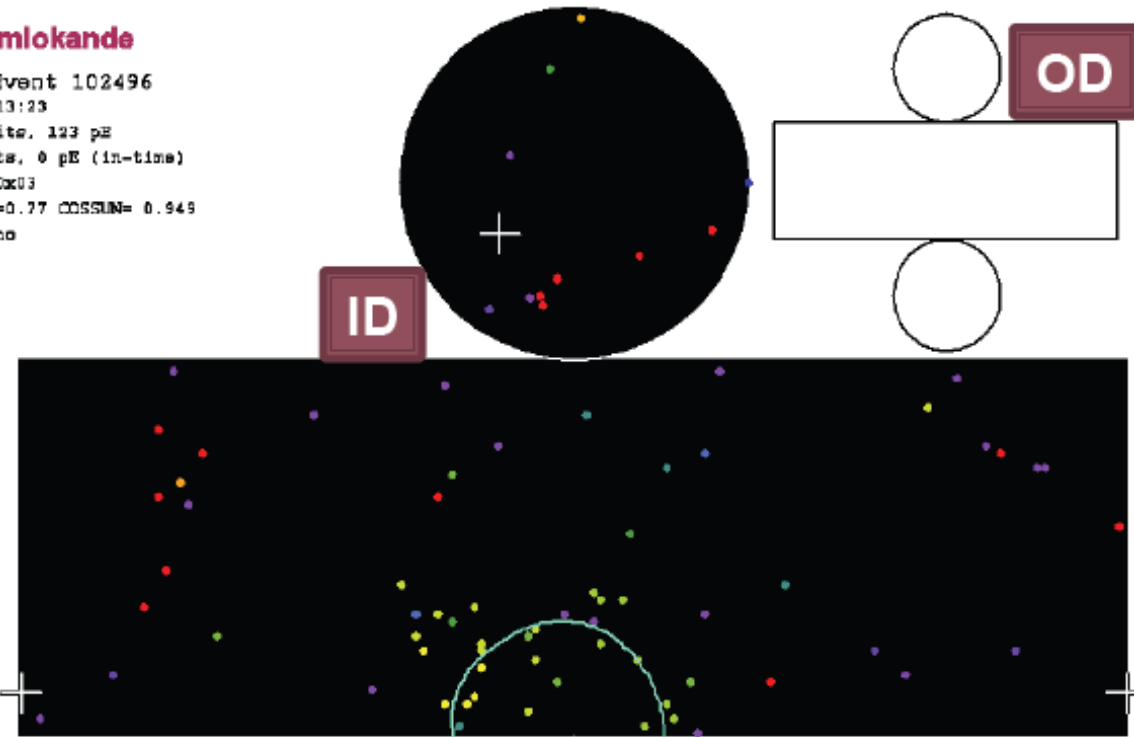
Energy : 14 % Vertex : 55 cm Direction : 23°
 (for E = 9.5 MeV(kin.) electron)

Super-Kamioke

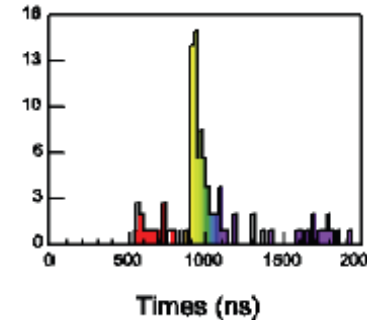
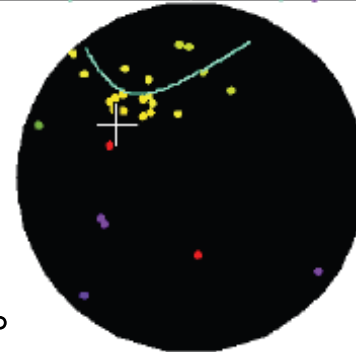
Run 1742 Event 102496
 96-05-31:07:13:23
 Inner: 193 hits, 123 pE
 Outer: -1 hits, 0 pE (in-time)
 Trigger ID: 0x03
 E= 9.086 GeV=0.77 COSθ= 0.949
 Solar Neutrino

Time(ns)

- < 815
- 815- 835
- 835- 855
- 855- 875
- 875- 895
- 895- 915
- 915- 935
- 935- 955
- 955- 975
- 975- 995
- 995-1015
- 1015-1035
- 1035-1055
- 1055-1075
- 1075-1095
- >1095

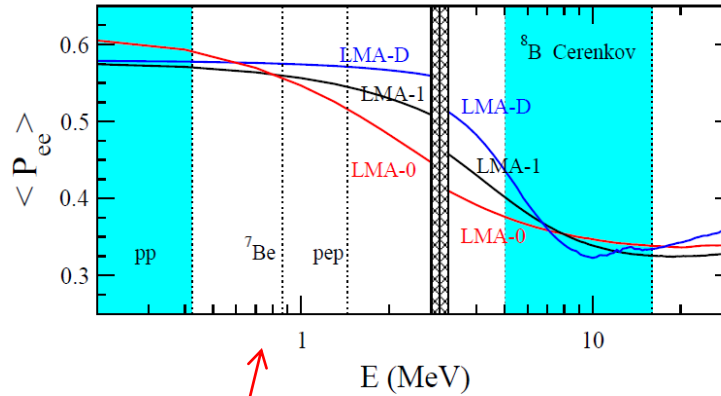
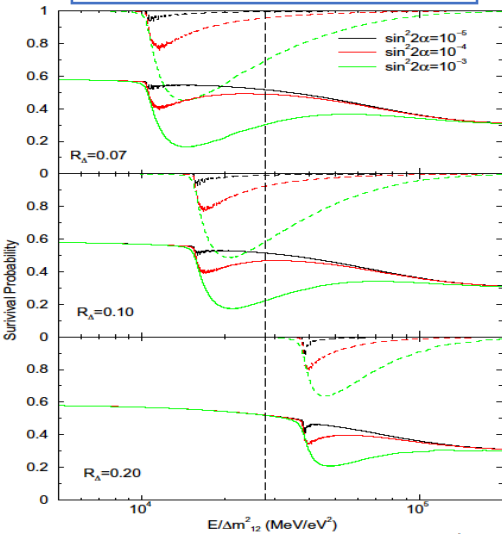


$E_{total} = 9.1 \text{ MeV}$
 $\cos\theta_{sun} = 0.95$



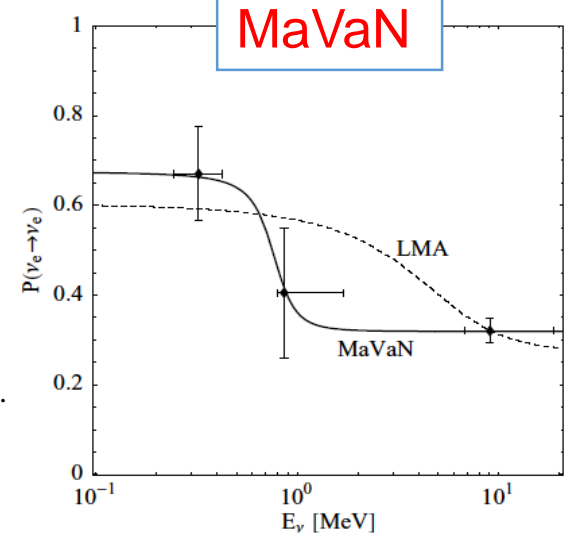
Non-standard models to predict flat spectrum p.16

Sterile neutrino



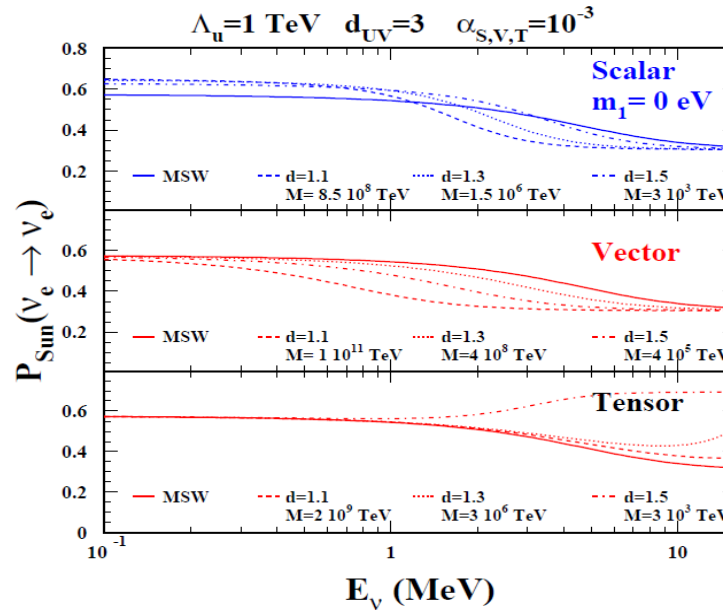
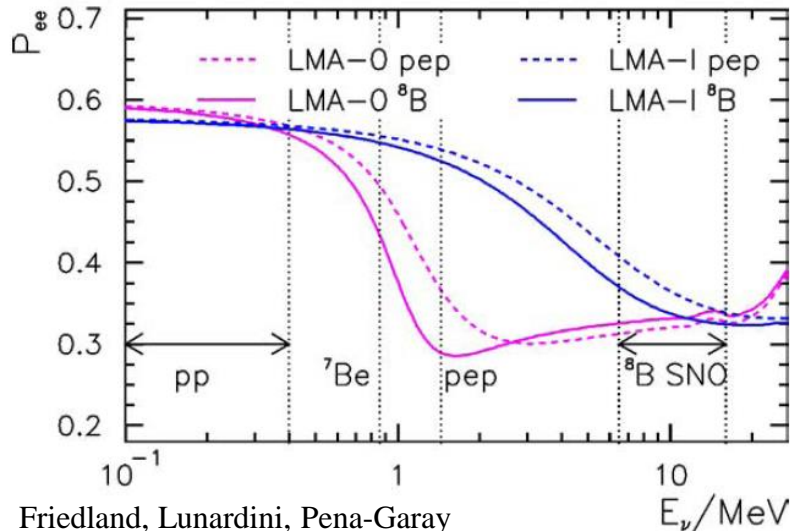
Miranda, Tortola and Valle, JHEP 0610:008,2006. (hep-ph/0406280)

Non standard Interaction



Barger, Huber and Marfatia, Phys.Rev.Lett.95:211802,2005 (hep-ph/0502196)

$R_\Delta \equiv \frac{\Delta m_{01}^2}{\Delta m_{21}^2}$ Holanda and Smirnov, Phys.Rev.D69(2004)113002. (hep-ph/0307266)



Gonzalez-Garcia, Holanda, Zukanovich, Funchal, JCAP 0806:019,2008. (hep-ph/0803.1180)

Unparticle

Friedland, Lunardini, Pena-Garay PLB594(2004)347(hep-ph/0402266)

Tight Fiducial volume

3.5-4.0 MeV

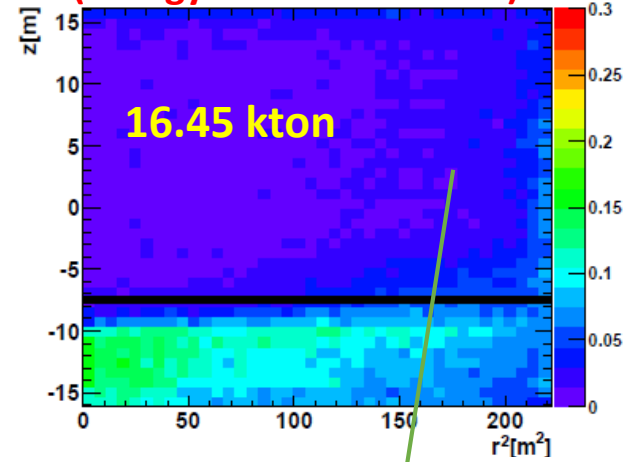
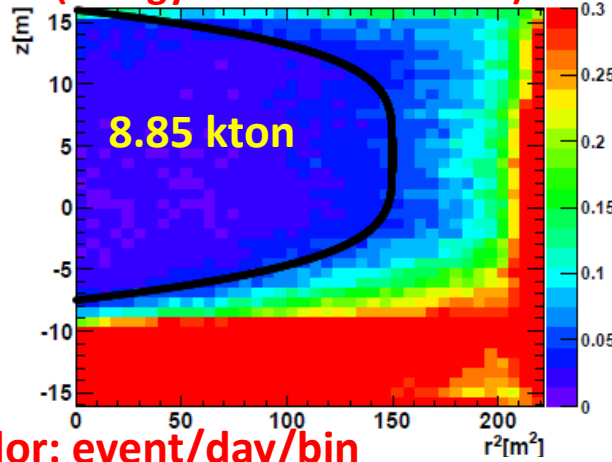
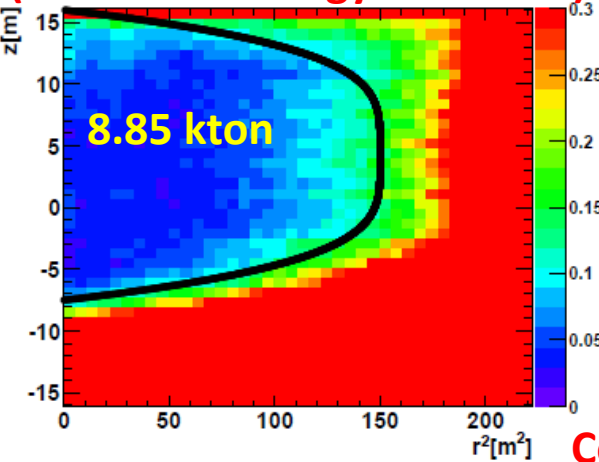
4.0-4.5 MeV

4.5-5.0 MeV

(SK-IV new energy threshold)

(energy threshold in SK-III)

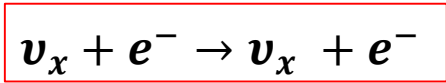
(energy threshold in SK-I)



Color: event/day/bin

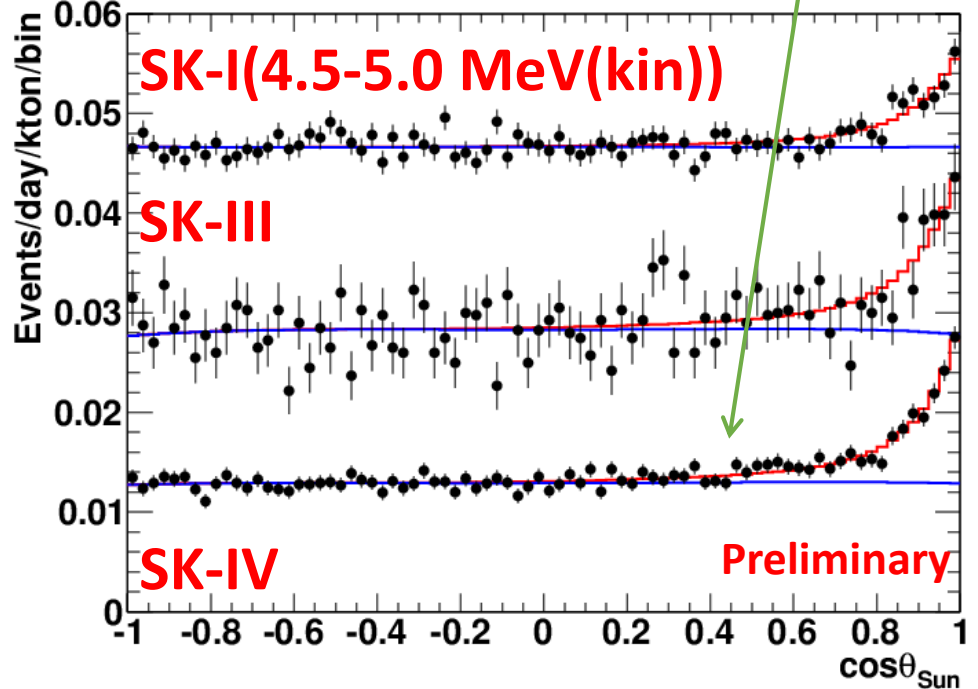
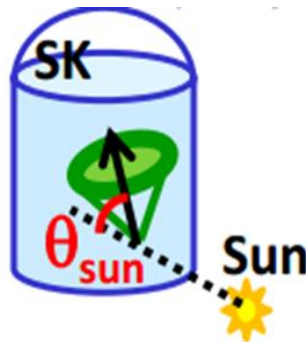
◆ solar ν observation

Elastic scattering

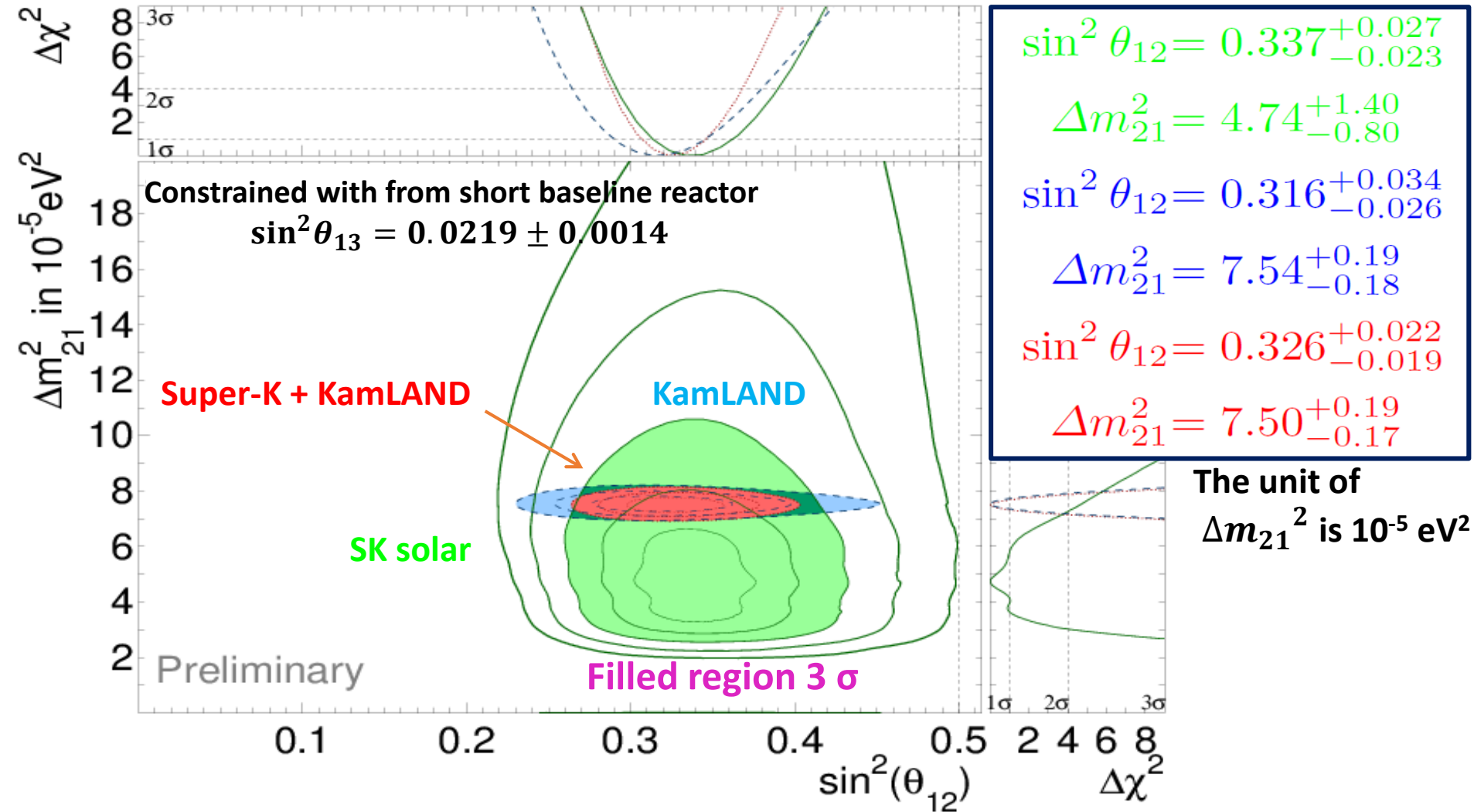


Solar neutrino signals are seen around $\cos\Theta_{\text{sun}} = 1$.

Background event → Other direction ($\cos\Theta_{\text{sun}} < 0$)

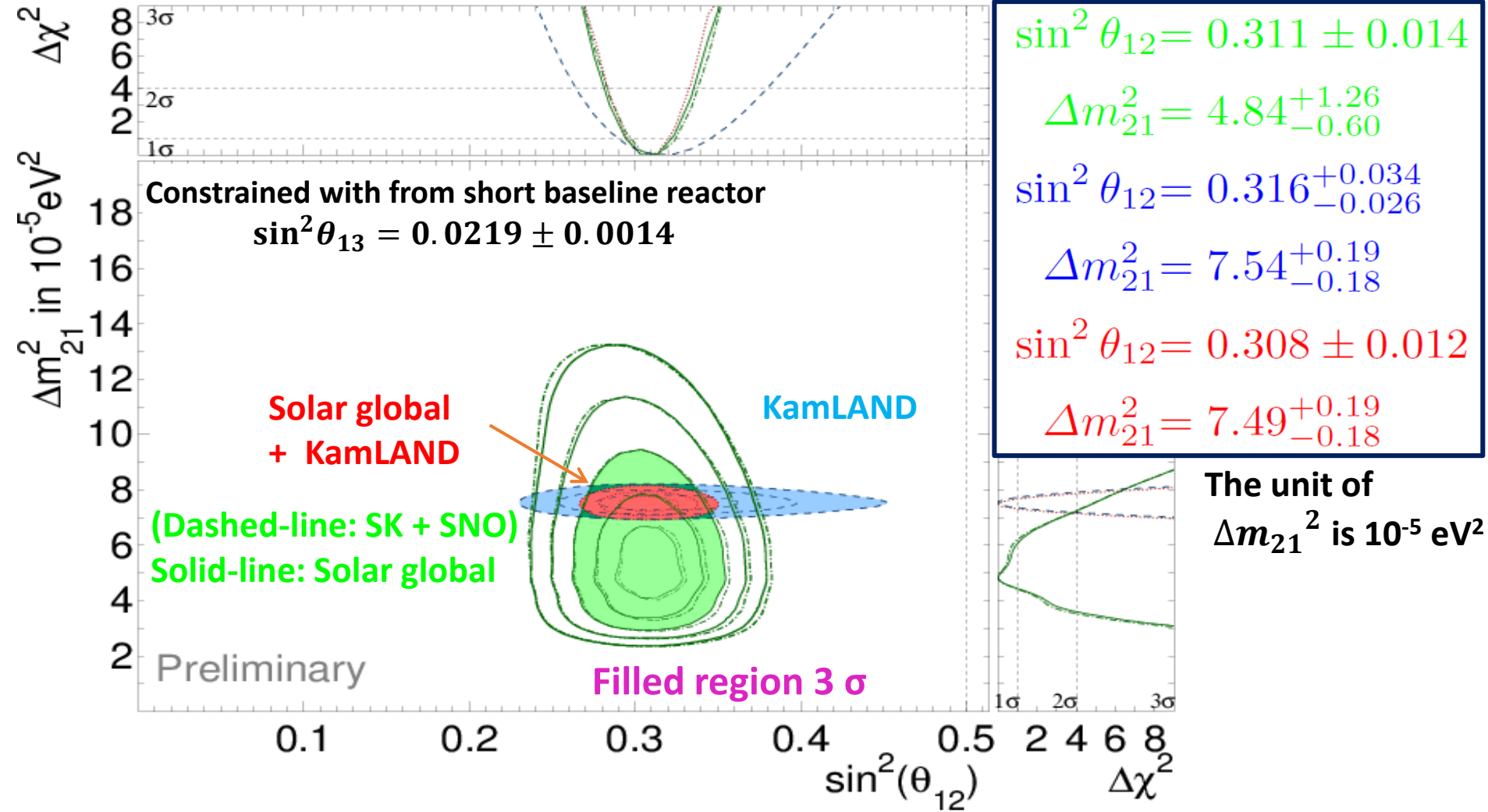


Super-K vs. KamLAND



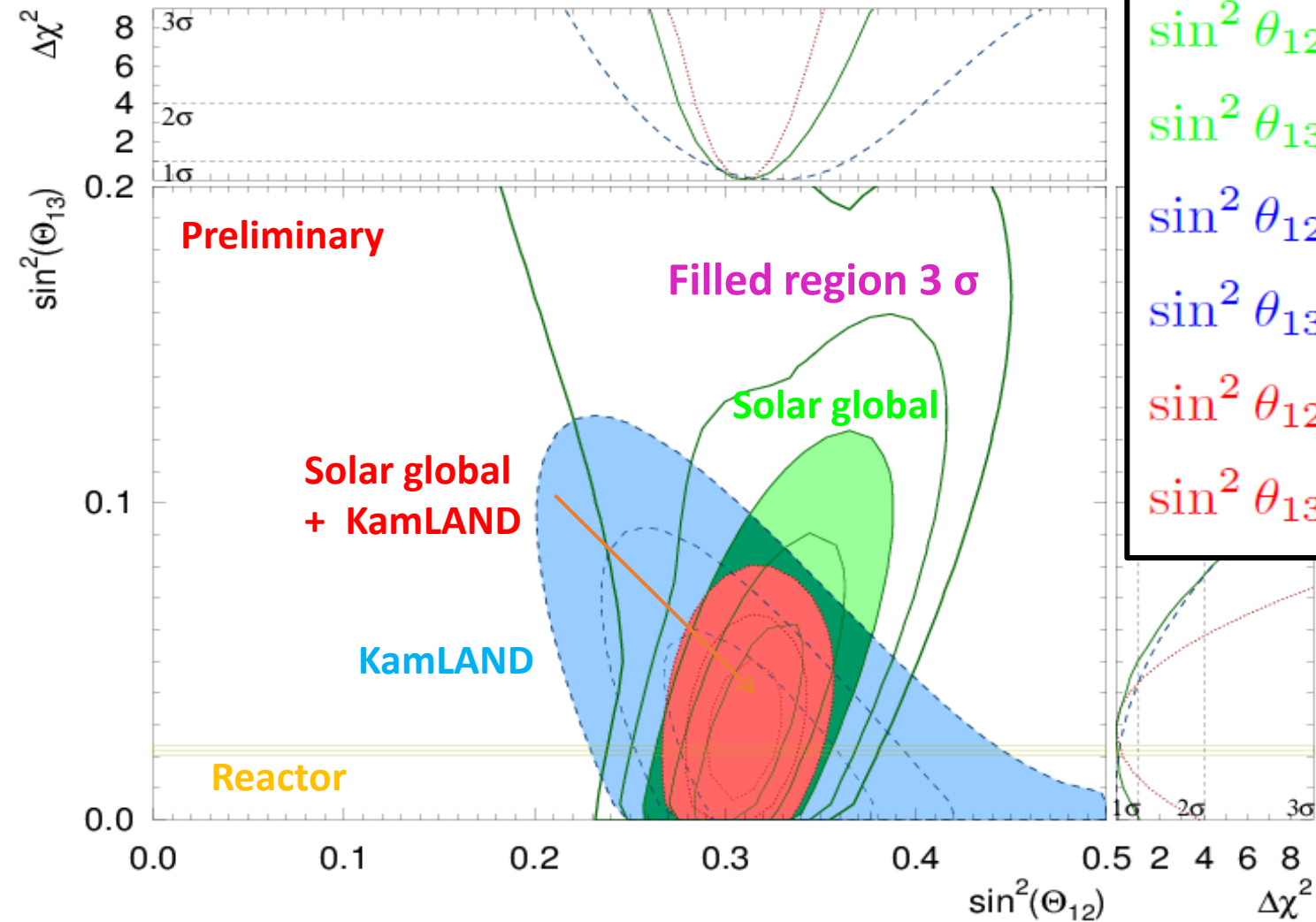
SK result uniquely selects the Large Mixing Angle MSW region by more than 3σ .
 SK significantly contributes to the measurement of the solar angle.

Solar global vs. KamLAND



The SK spectrum and D/N data favor a lower m_{21}^2 value than KamLAND's by more than 2σ and mostly determine this parameter in the solar neutrino oscillation fit.

3-flavor oscillation analysis



$$\sin^2 \theta_{12} = 0.311^{+0.022}_{-0.017}$$

$$\sin^2 \theta_{13} = 0.027^{+0.025}_{-0.027}$$

$$\sin^2 \theta_{12} = 0.316^{+0.034}_{-0.026}$$

$$\sin^2 \theta_{13} = 0.010^{+0.033}_{-0.034}$$

$$\sin^2 \theta_{12} = 0.310^{+0.014}_{-0.013}$$

$$\sin^2 \theta_{13} = 0.029^{+0.014}_{-0.015}$$

2σ level non-zero θ_{13} is obtained from Solar global + KamLAND.
Consistent with the reactor results.