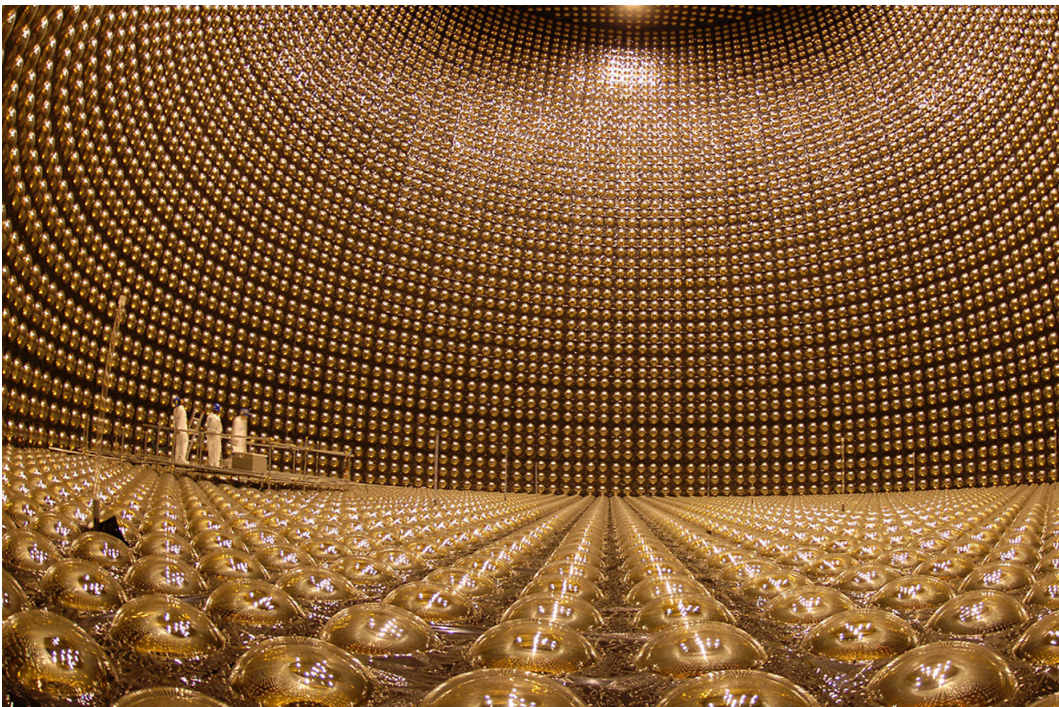
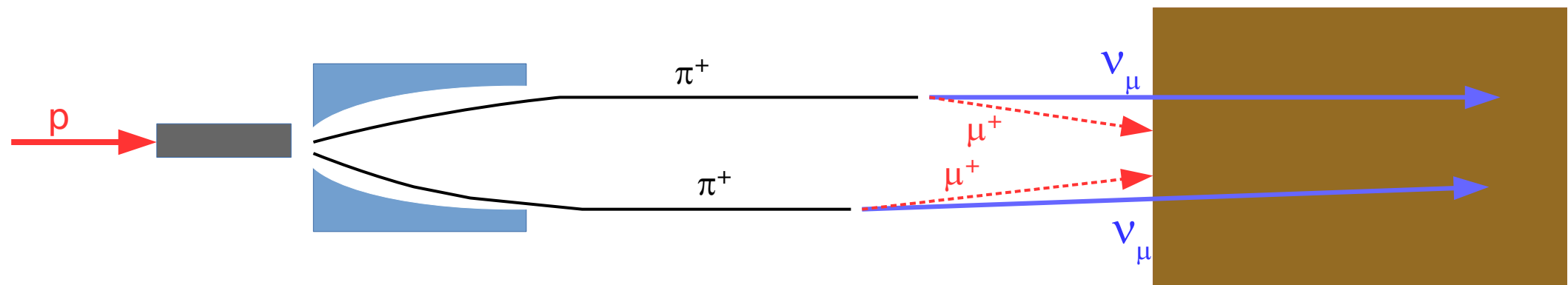


# Current Long Baseline Neutrino Experiments



© Fermi National Accelerator Laboratory



**Scott Oser**

**(UBC/TRIUMF)**

TAUP 2017  
July 26, 2017

# The neutrino 3x3 mixing matrix

Different L/E values pick up different  $\Delta m^2$  pairs, probing different parts of mixing matrix.

$$\underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}$$

Atmospheric  $\nu$ 's:

$$\theta_{23} \approx 46^\circ$$

Maximal mixing! (?)

$$\underbrace{\begin{pmatrix} c_{13} & 0 & e^{i\delta} s_{13} \\ 0 & 1 & 0 \\ -e^{-i\delta} s_{13} & 0 & c_{13} \end{pmatrix}}$$

Short baseline reactor  $\nu$ 's:

$$\theta_{13} \approx 9^\circ$$

Small, quark-like mixing

$$\underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}$$

Solar  $\nu$ 's:

$$\theta_{12} \approx 33^\circ$$

Large, non-maximal mixing

Compare to identical parameterization of CKM matrix ...

$$\theta_{23} \approx 2.4^\circ$$

$$\theta_{13} \approx 0.2^\circ$$

$$\theta_{12} \approx 13^\circ$$

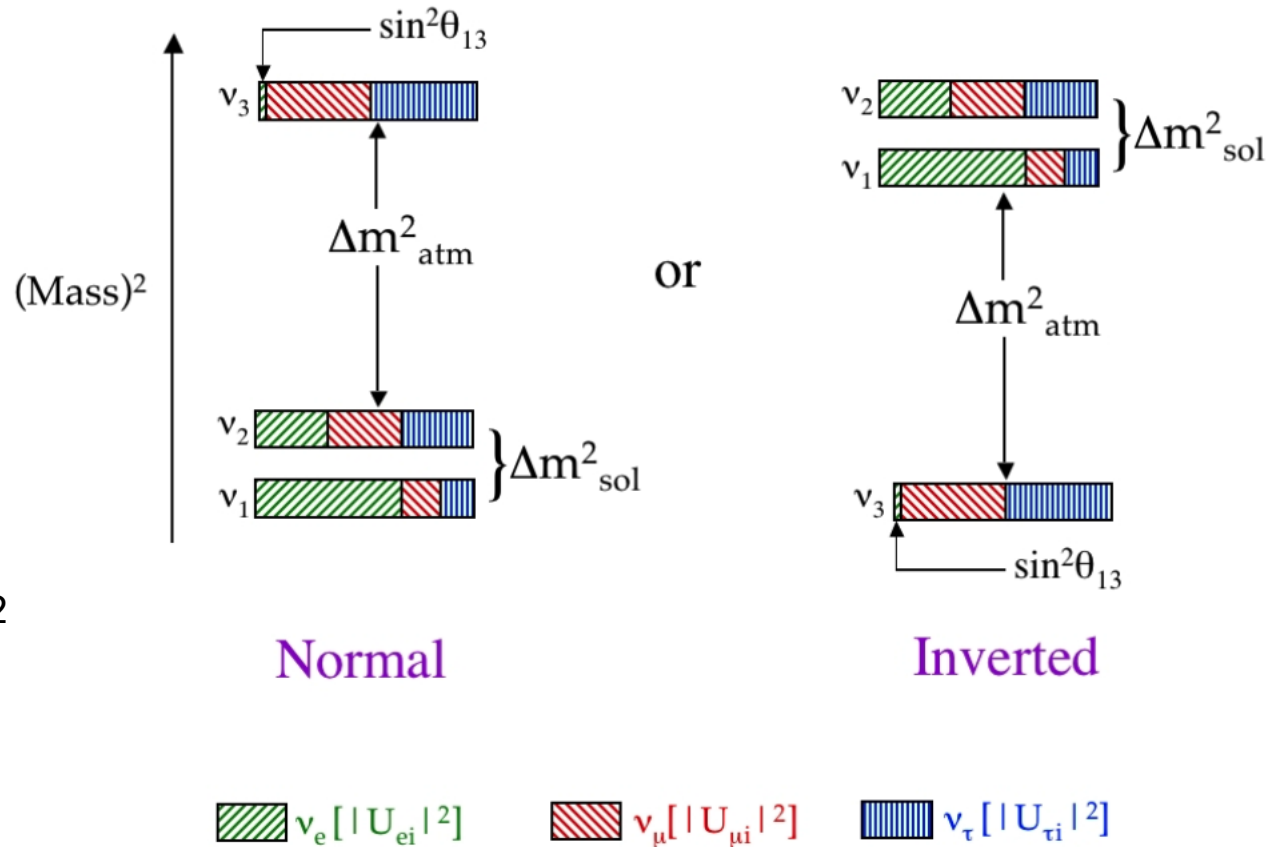
# Mass Hierarchy

Currently unknown:

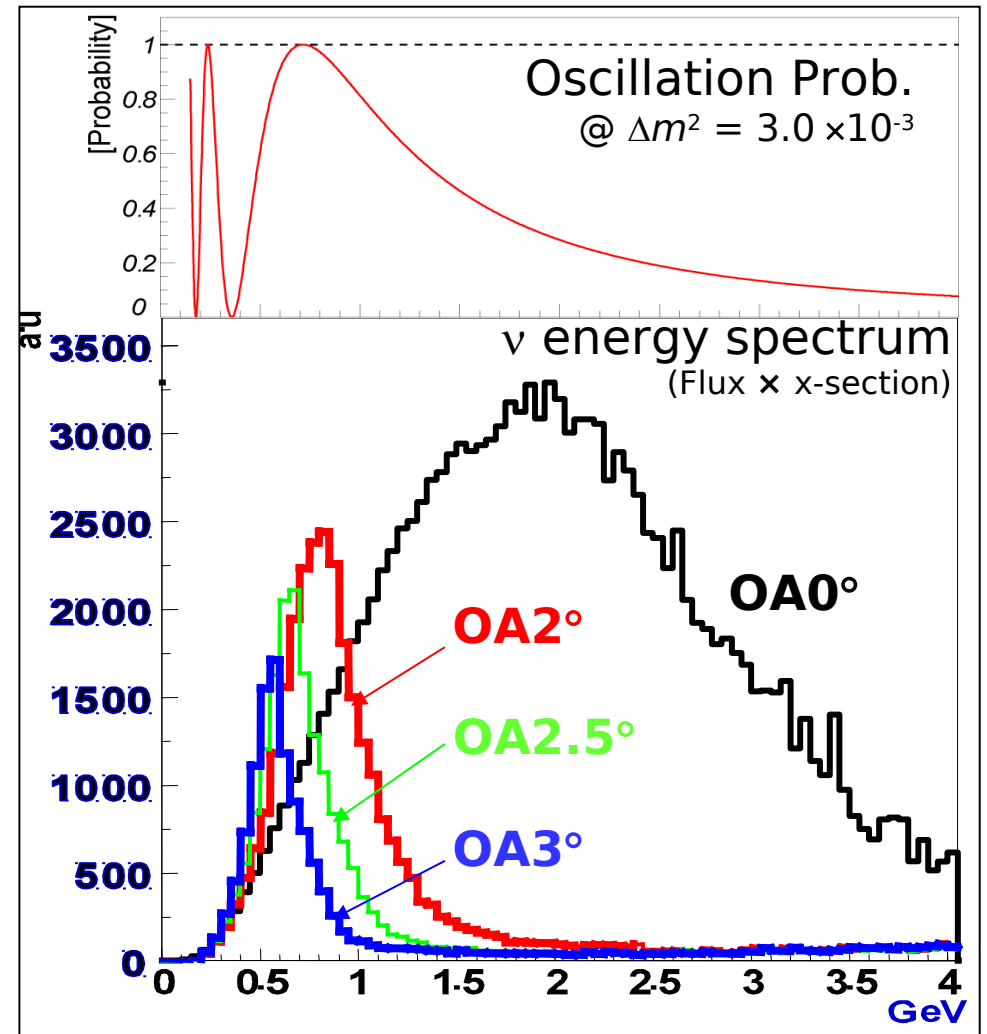
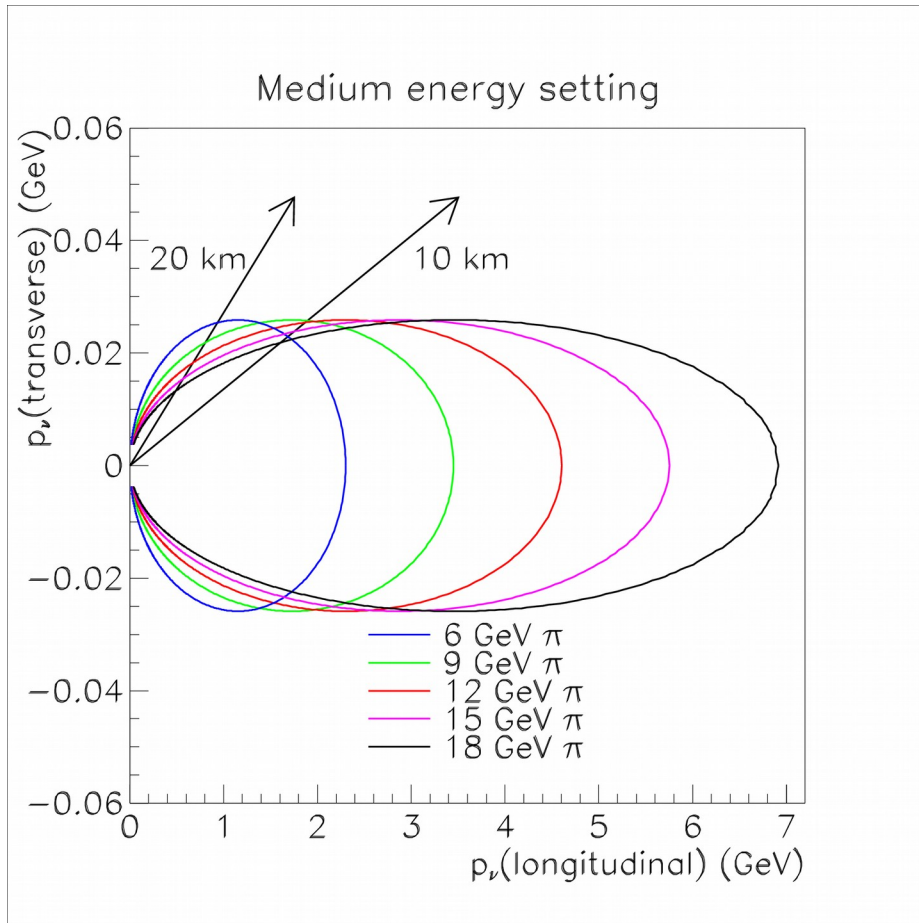
- value of  $\delta_{CP}$
- sign of the mass hierarchy

$$|\Delta m_{32}^2| = 2.4 \times 10^{-3} \text{ eV}^2$$

$$\Delta m_{21}^2 = 7.5 \times 10^{-5} \text{ eV}^2$$



# Off-Axis $\nu_\mu$ Beam



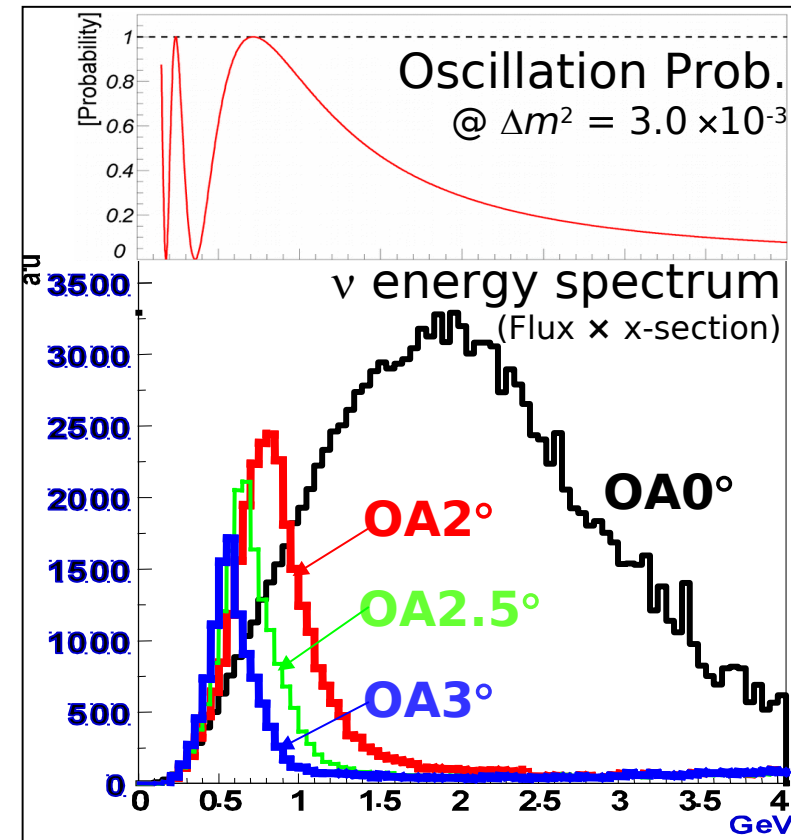
Off-axis beam: more flux near peak oscillation energy, less flux at higher energies where  $\nu_e$  backgrounds are produced.

# LBL signature #1: $\nu_\mu$ disappearance

Starting from  $P_{\nu_a \rightarrow \nu_b}(L, E) = \left| \sum_{j,k} U_{aj}^* U_{bj} U_{ak} U_{bk}^* e^{-i \frac{\Delta m_{jk}^2 L}{2E}} \right|$

$$P(\nu_\mu \rightarrow \nu_\mu) \approx 1 - \cos^4 \theta_{13} \sin^2 2\theta_{23} \sin^2 \left( \frac{\Delta m_{32}^2 L}{4E} \right) - \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \left( \frac{\Delta m_{32}^2 L}{4E} \right) \approx 1 - \sin^2 2\theta_{23} \sin^2 \left( \frac{\Delta m_{32}^2 L}{4E} \right)$$

Sensitive to  $\Delta m_{32}^2$  and  $\theta_{23}$ .  
 Same formula for neutrinos and for antineutrinos, if CPT holds.



# LBL signature #2: $\nu_e$ appearance

$$P(\nu_\mu \rightarrow \nu_e) \sim \boxed{\sin^2 2\theta_{13}} \times \sin^2 \theta_{23} \times \frac{\sin^2[(1-x)\Delta]}{(1-x)^2} \\
 -\alpha \sin 2\theta_{13} \times \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \times \sin \Delta \frac{\sin[x\Delta]}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} \\
 +\alpha \sin 2\theta_{13} \times \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \times \cos \Delta \frac{\sin[x\Delta]}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} \\
 x \equiv 2\sqrt{2}G_F N_e \frac{E_\nu}{\Delta m_{31}^2} \quad \Delta \equiv \frac{\Delta m_{31}^2 L}{4E_\nu}$$

Dominant term corresponds to a  $\sim 5\%$  transition probability at the oscillation maximum

# LBL signature #2: $\nu_e$ appearance

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) \sim & \sin^2 2\theta_{13} \times \sin^2 \theta_{23} \times \frac{\sin^2[(1-x)\Delta]}{(1-x)^2} \\
 & -\alpha \sin 2\theta_{13} \times \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \times \sin \Delta \frac{\sin[x\Delta]}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} \\
 & +\alpha \sin 2\theta_{13} \times \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \times \cos \Delta \frac{\sin[x\Delta]}{x} \frac{\sin[(1-x)\Delta]}{(1-x)} \\
 x \equiv & 2\sqrt{2}G_F N_e \frac{E_\nu}{\Delta m_{31}^2} \quad \Delta \equiv \frac{\Delta m_{31}^2 L}{4E_\nu}
 \end{aligned}$$

Terms containing  $\delta$  are sensitive to CP phase.  
The  $\delta$ 's flip sign for antineutrinos.

The  $x$  parameter (matter effect) also flips sign for antineutrinos. The matter effect is subdominant at T2K due to low beam energy, but larger at NO $\nu$ A.

# CP, T, and CPT for neutrinos

$$P(\nu_a \rightarrow \nu_b) \longleftrightarrow_{\text{CP}} P(\bar{\nu}_a \rightarrow \bar{\nu}_b)$$

$$P(\nu_a \rightarrow \nu_b) \longleftrightarrow_{\text{T}} P(\nu_b \rightarrow \nu_a)$$

$$P(\nu_a \rightarrow \nu_b) \longleftrightarrow_{\text{CPT}} P(\bar{\nu}_b \rightarrow \bar{\nu}_a)$$

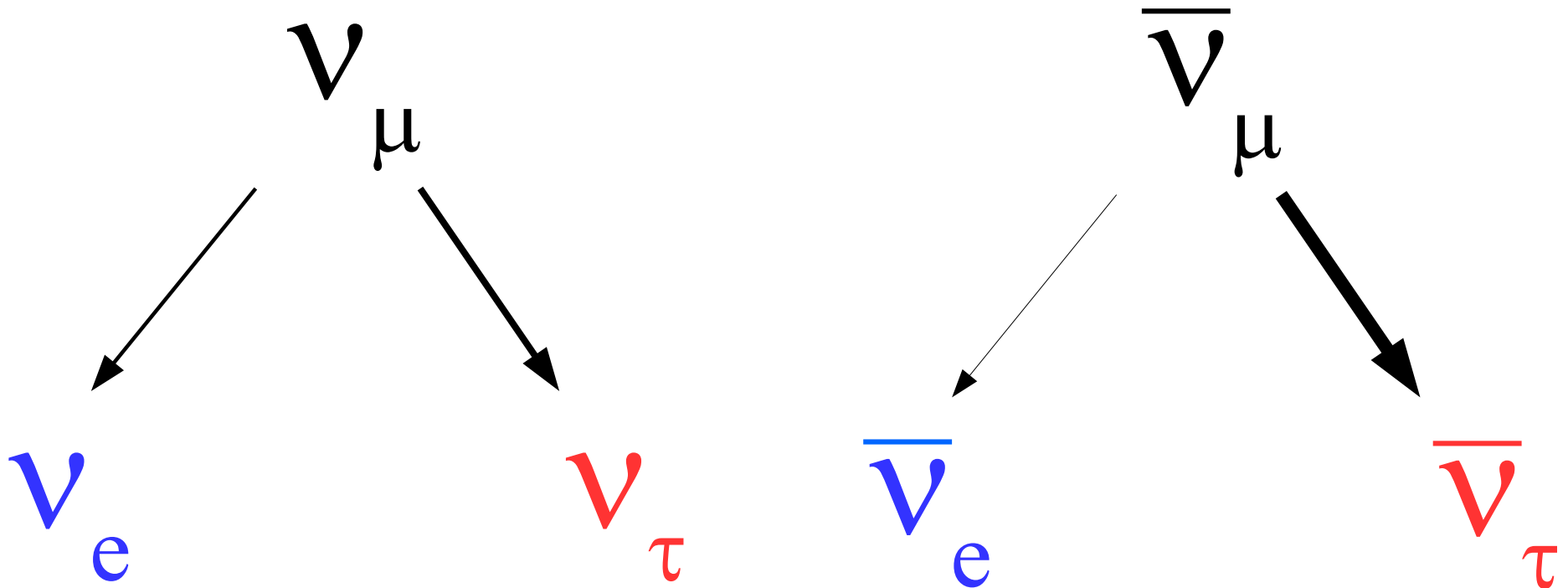
$$P(\nu_a \rightarrow \nu_a) \longleftrightarrow_{\text{CPT}} P(\bar{\nu}_a \rightarrow \bar{\nu}_a)$$

If CPT holds in neutrino sector, neutrino survival probability equals antineutrino survival probability.

As a result CP violation is observable only in appearance channels, in which the flavour of the appearing lepton is detected. This is why LBL neutrinos are interesting!



# CP and three-flavour mixing



CP violation affects the mix of flavours that results from the oscillation

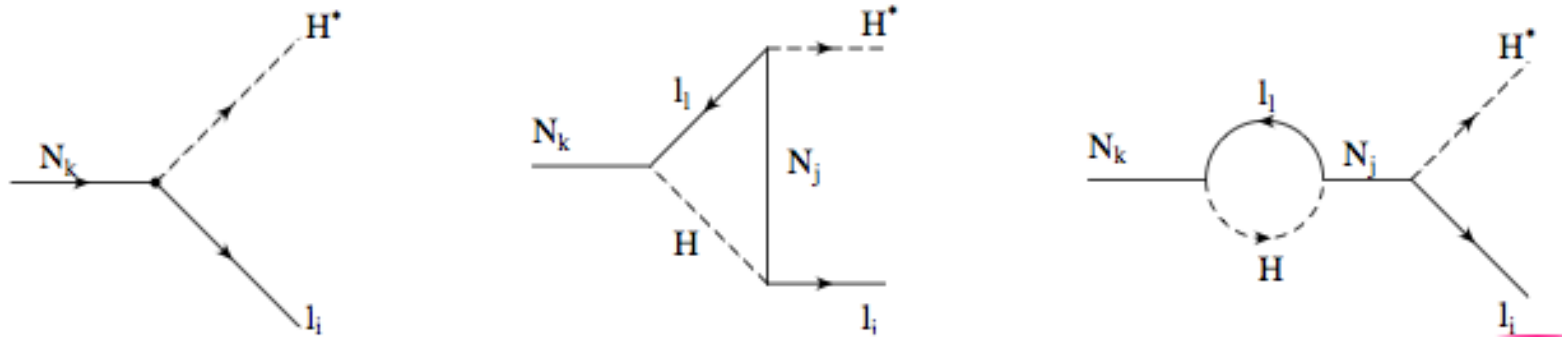
# Leptogenesis

CP violation in quark sector not enough to explain observed matter-antimatter asymmetry in universe.

Neutrino mixing provides another possible source of CPV.

- **Standard Leptogenesis: decays of RH neutrinos (CPV in decay)**

Quantum interference of tree diagram and one-loop diagram



Usual scenario: decay of heavy Majorana neutrinos [Phys.Lett B 174, 45 \(1986\)](#)

Many alternates, eg. leptogenesis with only Dirac  $\nu$ 's [PRL 89:271601 \(2002\)](#)

Relation of  $\delta_{CP}$  to leptogenesis is model-dependent, but observation of leptonic CP violation is an important milestone.

# Matter Effects and $\nu_e$ Appearance

Matter effects modify the oscillation formula. Because the Earth is made of electrons and not heavier leptons, the effective “index of refraction” for  $\nu_e$  is different than that for  $\nu_\mu$ . At the oscillation maximum, the  $\nu_e$  appearance probability changes to:

$$P(\nu_\mu \rightarrow \nu_e) \approx \left( 1 + 2 \frac{E}{E_R} \right) P_{vac}(\nu_\mu \rightarrow \nu_e)$$

where

$$E_R = \frac{\Delta m_{32}^2}{2\sqrt{2}G_F N_e} = \pm 11 \text{ GeV}$$

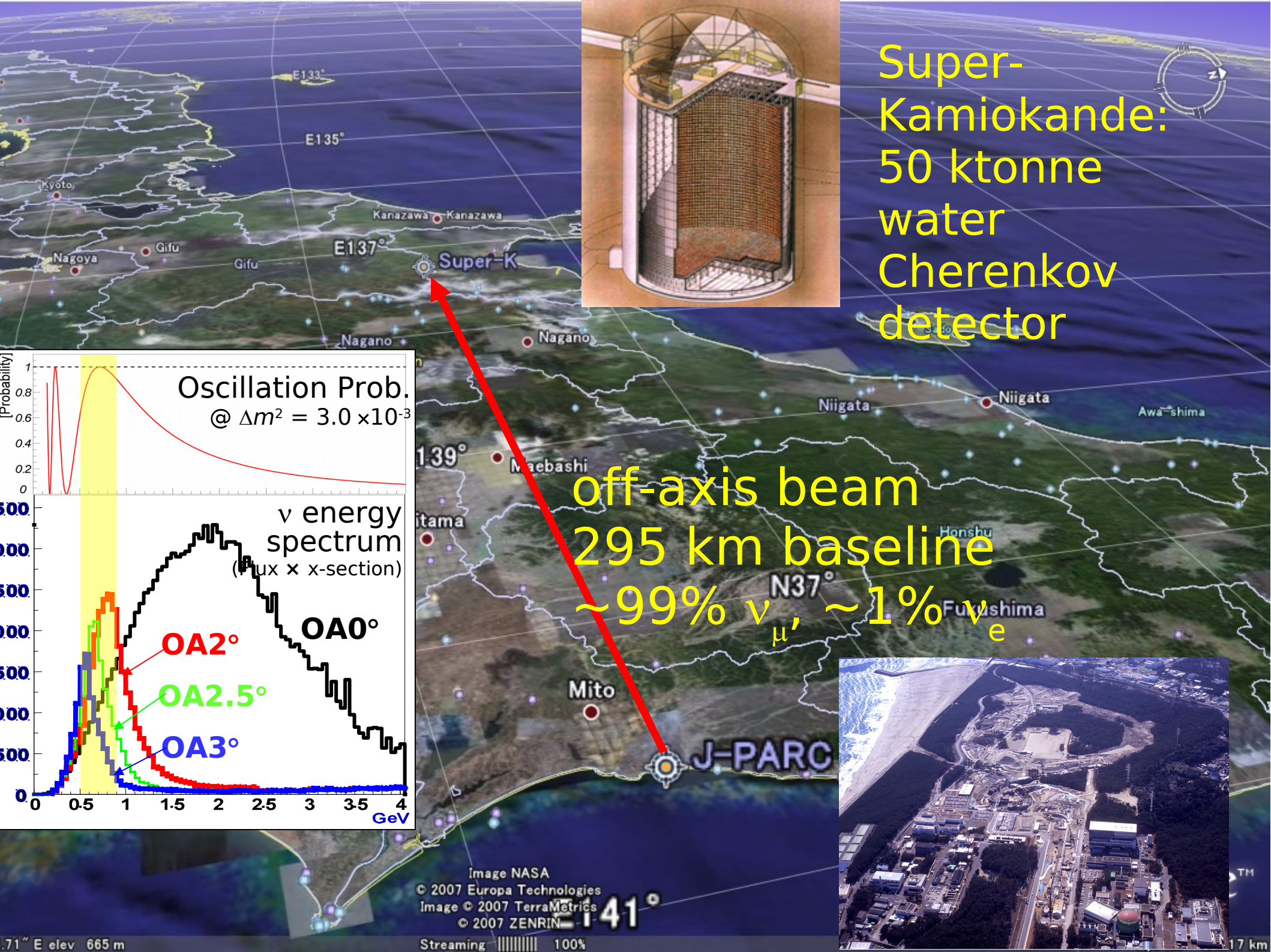
The sign of the matter effect is opposite for neutrinos and anti-neutrinos, and depends on the sign of  $\Delta m^2$  as well.

# T2K

JAPAN PROTON  
ACCELERATOR RESEARCH  
COMPLEX (J-PARC):  
Tokai, Japan  
30 GeV proton synchrotron  
design power: 0.75MW  
(upgradable to >1MW)



Image NASA  
© 2007 Europa Technologies  
Image © 2007 TerraMetrics  
© 2007 ZENRIN



Super-Kamiokande:  
50 ktonne  
water  
Cherenkov  
detector

off-axis beam  
295 km baseline  
 $\sim 99\% \nu_{\mu}$ ,  $\sim 1\% \nu_e$

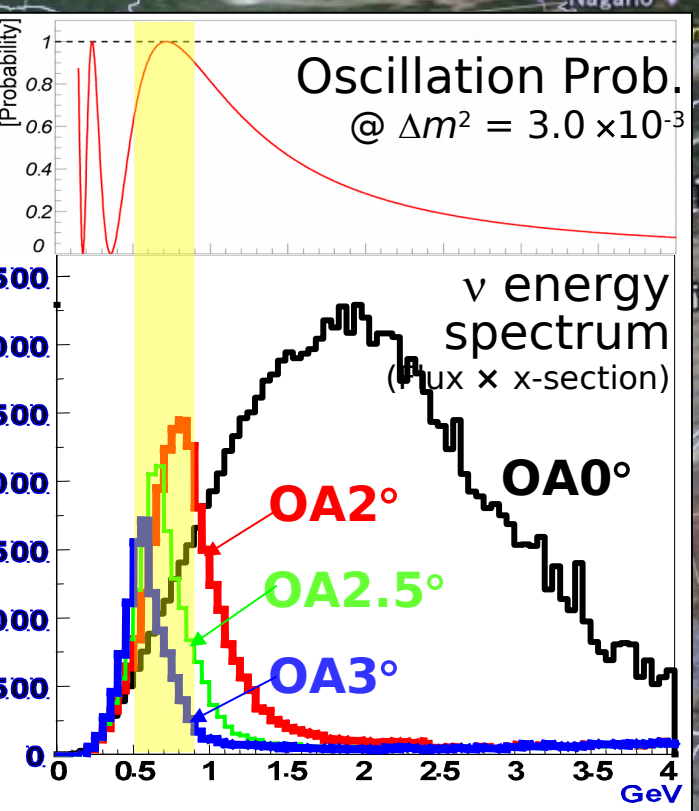


Image NASA  
© 2007 Europa Technologies  
Image © 2007 TerraMetrics  
© 2007 ZENRIN

Streaming 100%



Sophisticated on-axis and off-axis near detectors 280m from proton target

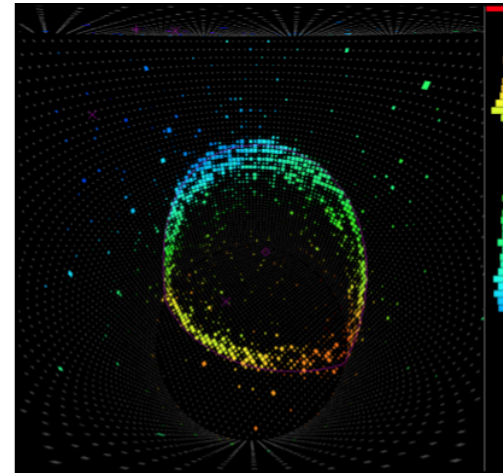
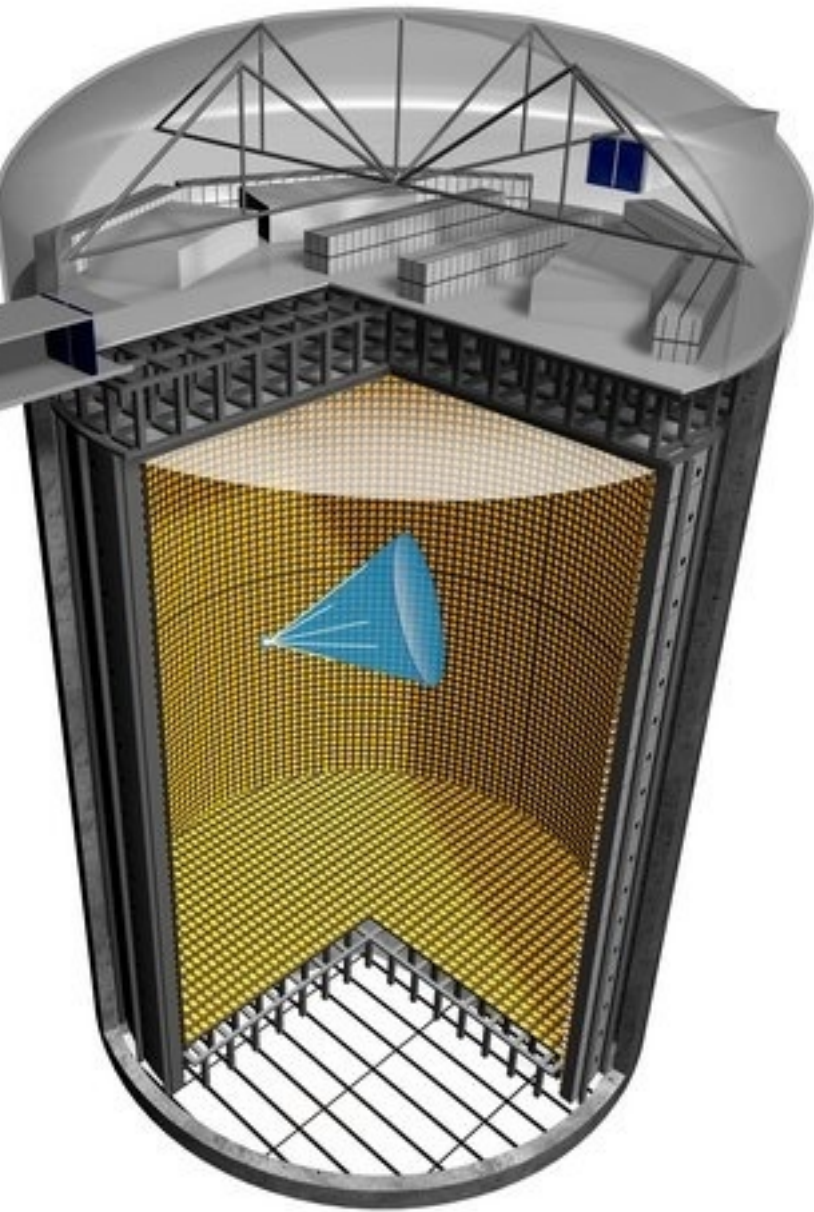
Image NASA  
 © 2007 Europa Technologies  
 Image © 2007 TerraMetrics  
 © 2007 ZENRIN

71" E elev 665 m

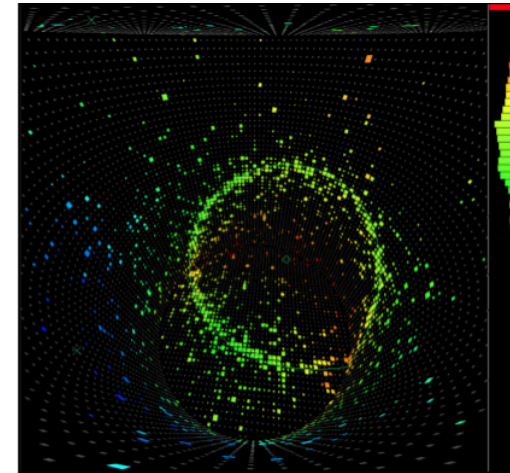
Streaming 100%

17 km

# Super-Kamiokande



muon-like ( $\nu_\mu$ )

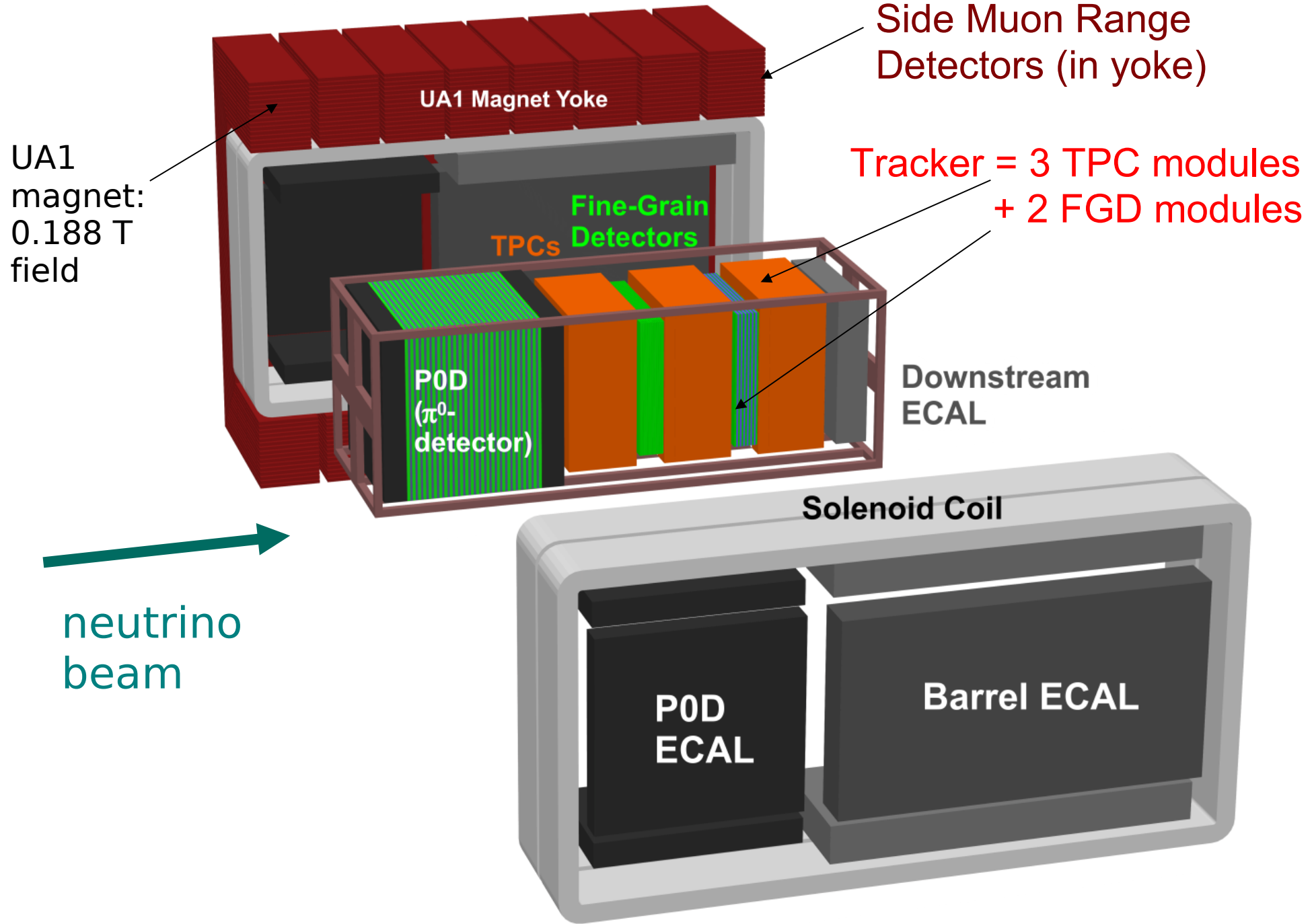


electron-like ( $\nu_e$ )

Water Cherenkov detection.

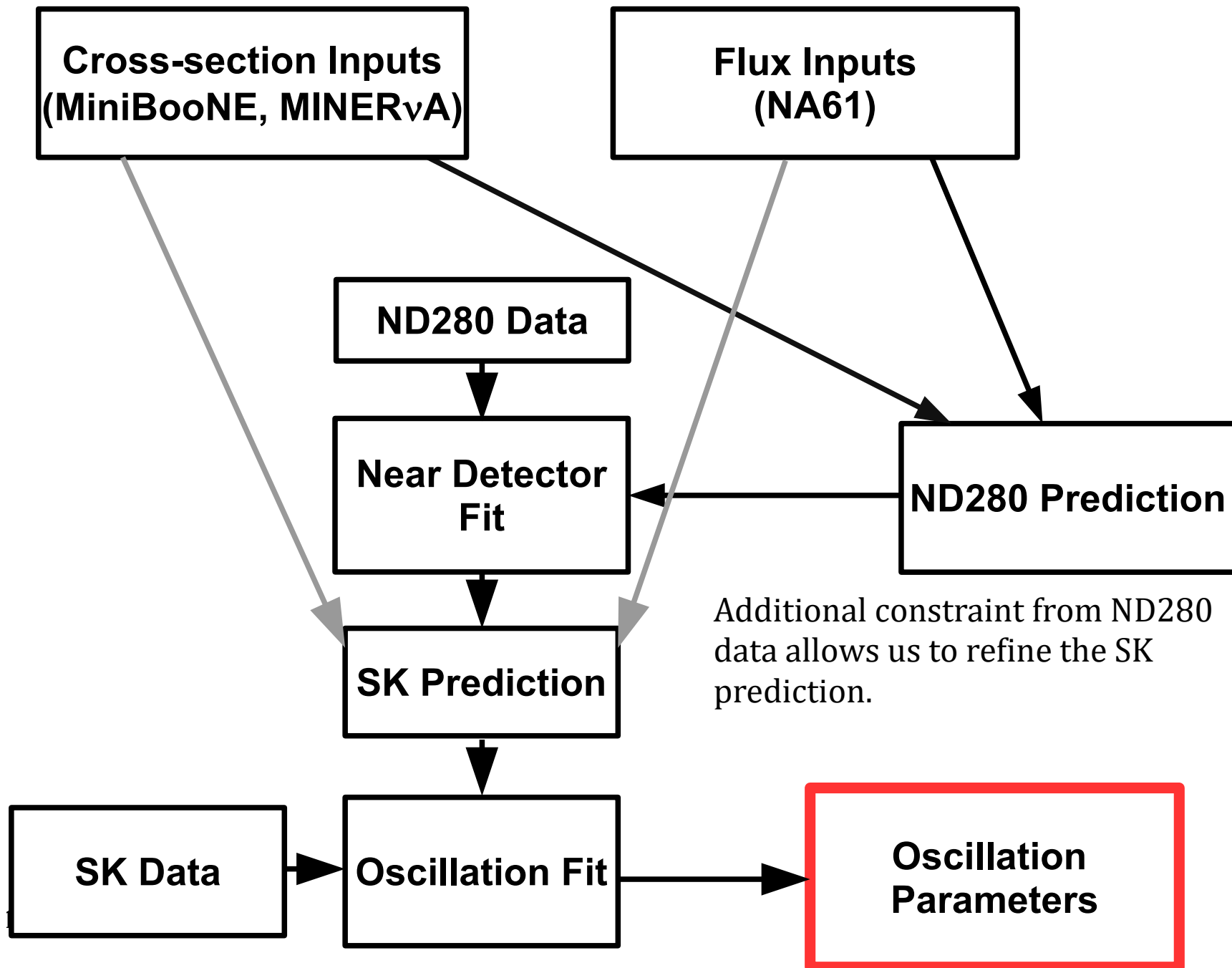
Primary signal channel is CCQE single-ring events. Reconstruct  $\nu$  energy in CCQE hypothesis from lepton kinematics.

# Off Axis Near Detector

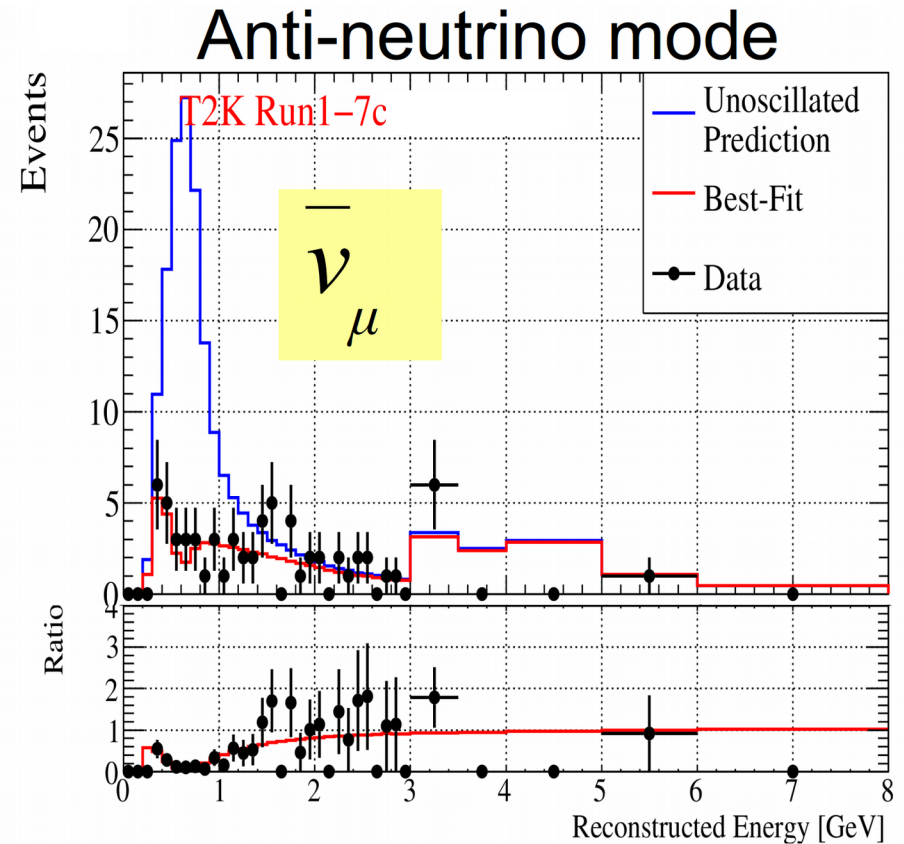
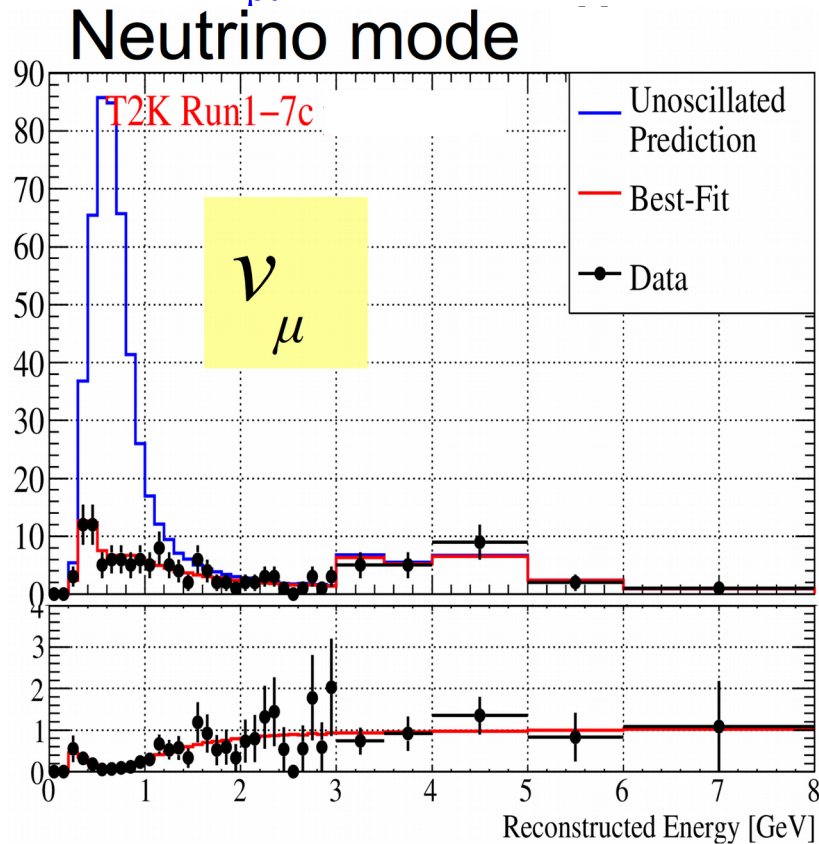




# T2K Oscillation Analysis Structure



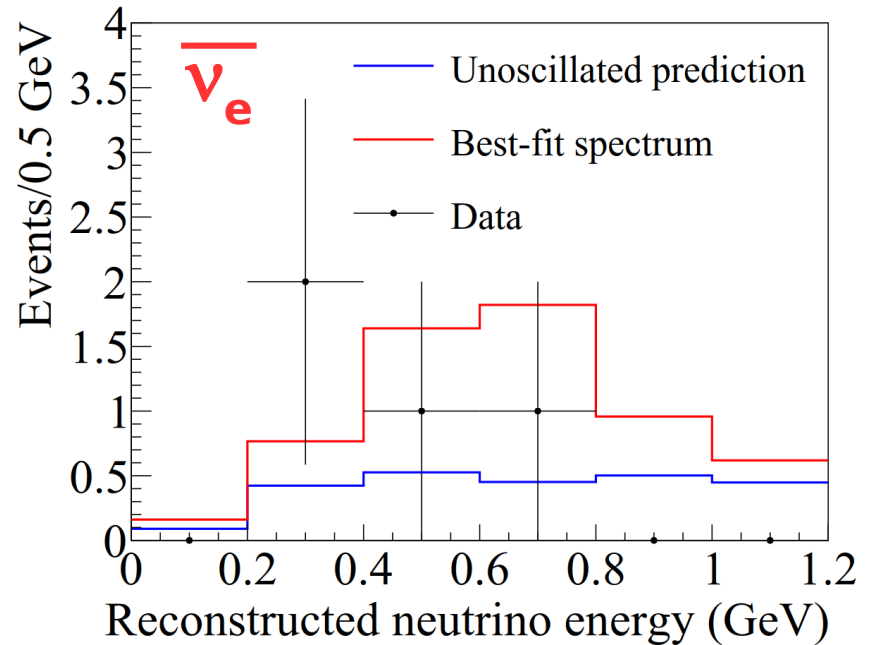
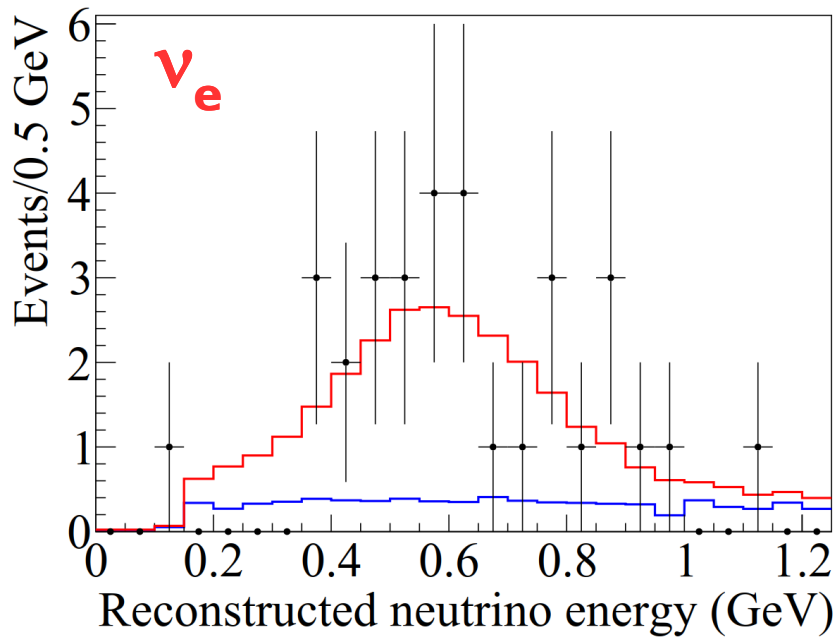
# $\nu_\mu$ disappearance results



Beam mode	Sample	Exp. Not Osc	Exp. $\delta_{CP} = 0$ (NH)	Observed
neutrino	$\mu$ -like	521.8	135.5	135
antineutrino	$\mu$ -like	184.8	64.1	66

$7.5 \times 10^{20}$  protons on target of  $\nu$  data  
 +  $7.5 \times 10^{20}$  protons on target of anti- $\nu$  data

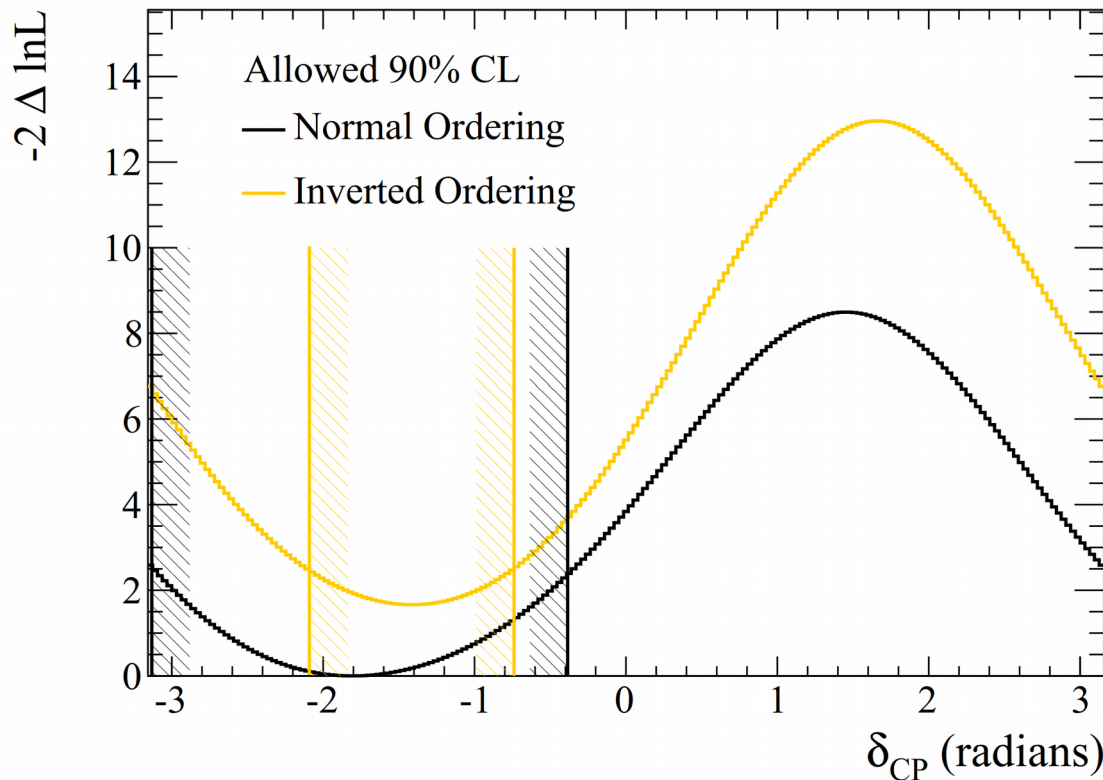
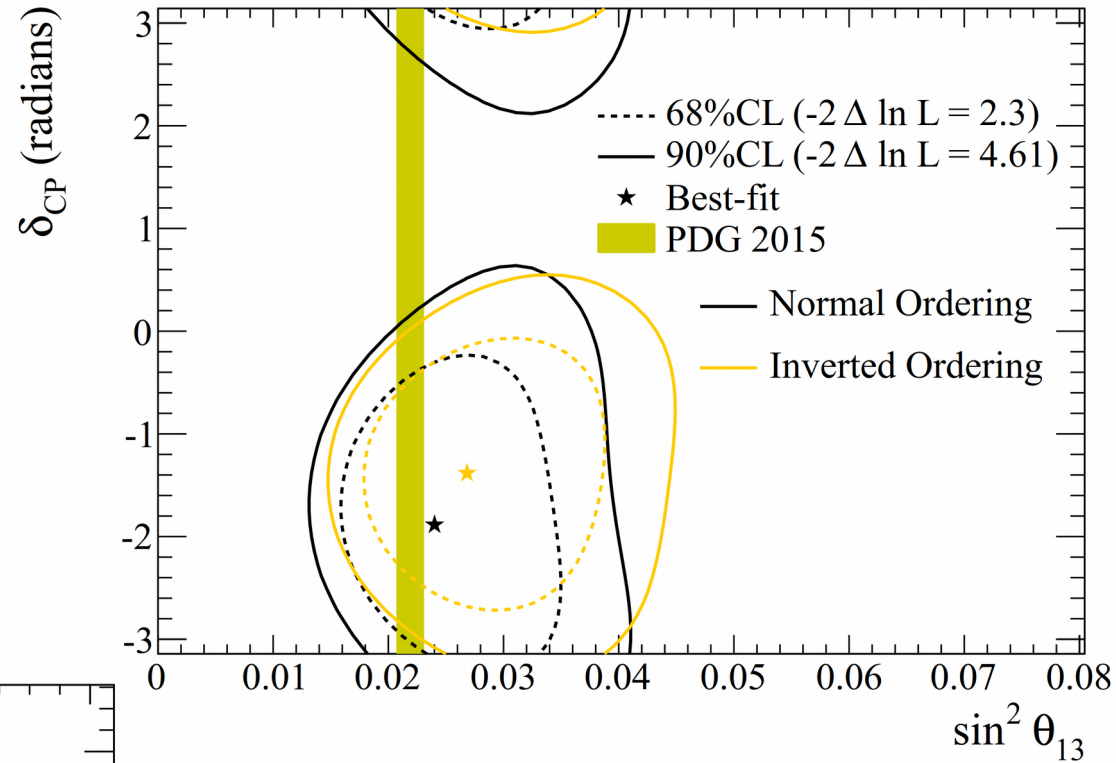
# $\nu_e$ appearance results



Normal	$\delta_{CP} = -\pi/2$	$\delta_{CP} = 0$	$\delta_{CP} = \pi/2$	$\delta_{CP} = \pi$	Observed
$\nu_e$	28.7	24.2	19.6	24.1	32
$\bar{\nu}_e$	6.0	6.9	7.7	6.8	4
Inverted	$\delta_{CP} = -\pi/2$	$\delta_{CP} = 0$	$\delta_{CP} = \pi/2$	$\delta_{CP} = \pi$	Observed
$\nu_e$	25.4	21.3	17.1	21.3	32
$\bar{\nu}_e$	6.5	7.4	8.4	7.4	4

# $\delta_{CP}$ contours

The best-fit CP phase is close to  $-\pi/2$ : maximal CP effect. Formally speaking  $\delta_{CP}=0$  is excluded at 90% CL.

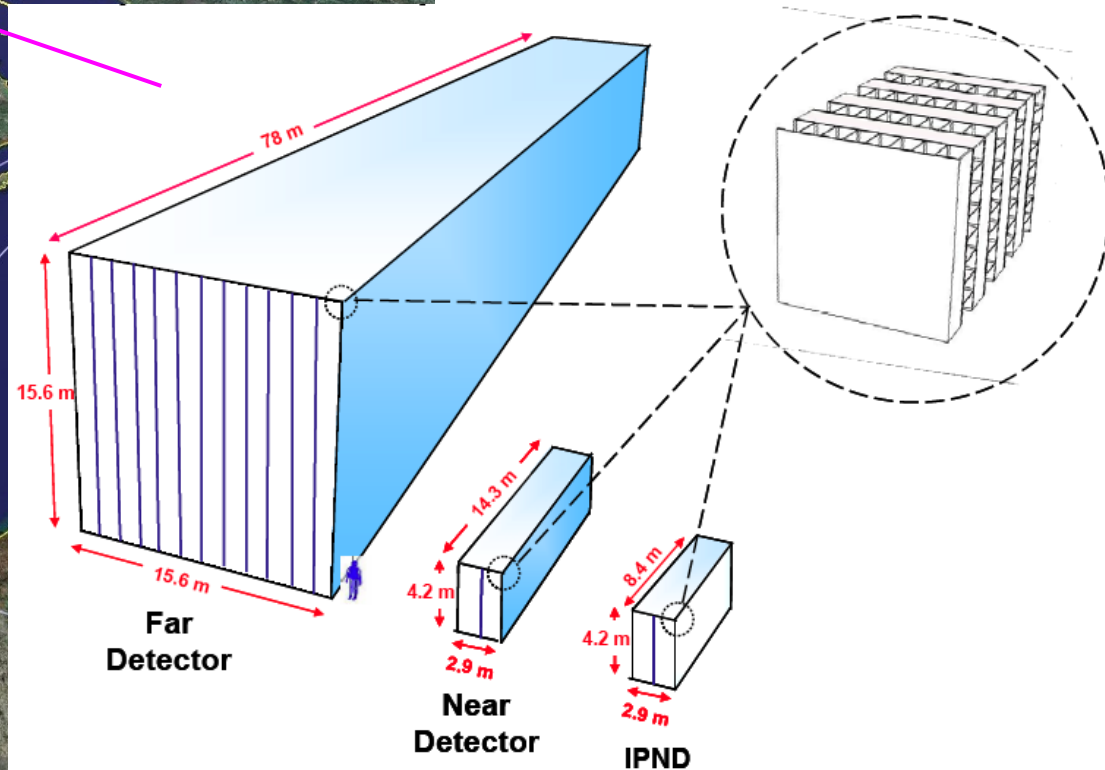
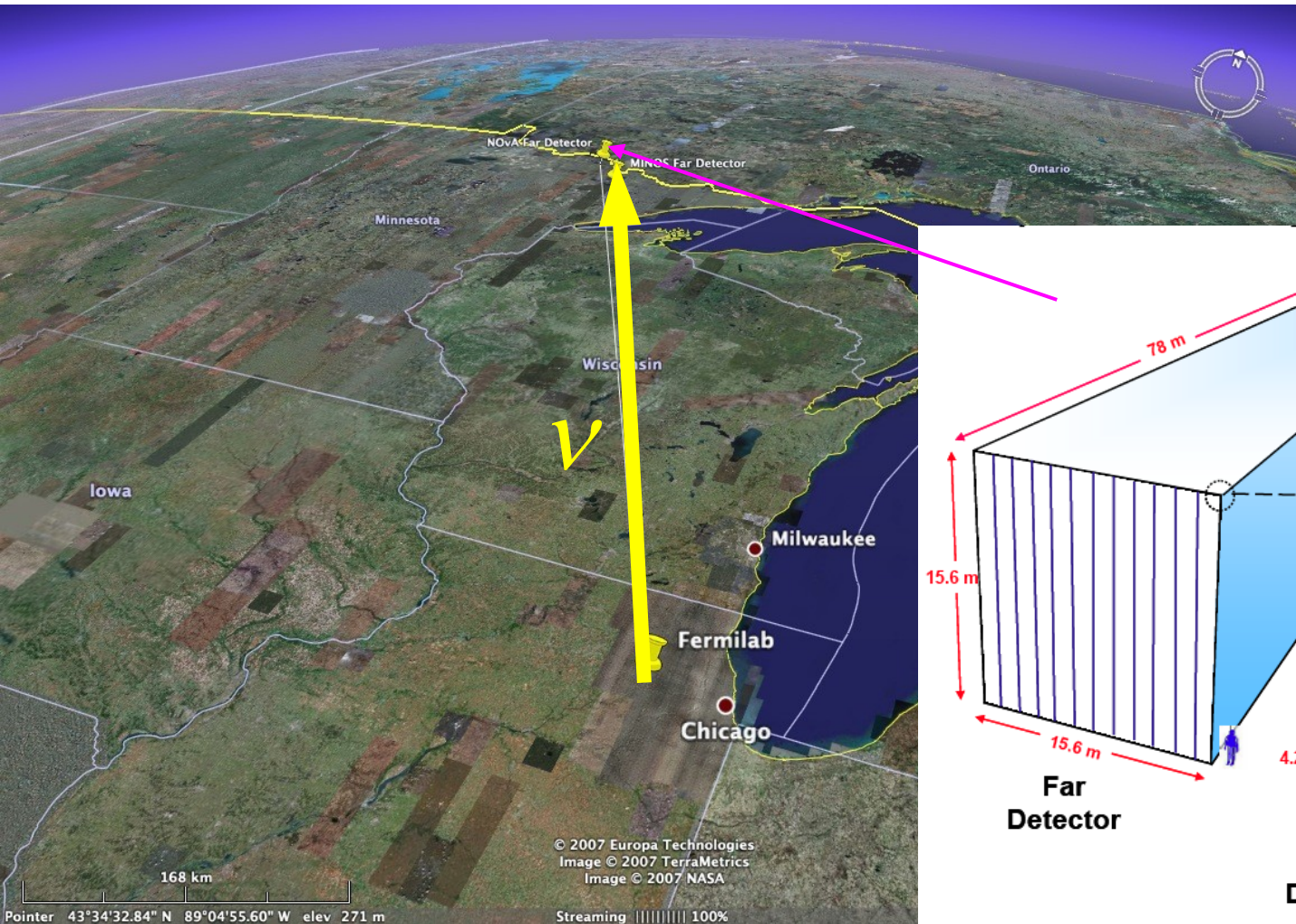


Caveat: the result is primarily driven by very high  $\nu_e$  appearance rate at T2K, beyond expectations of model. In other words, limit is better than expected sensitivity.

# NO<sub>v</sub>A

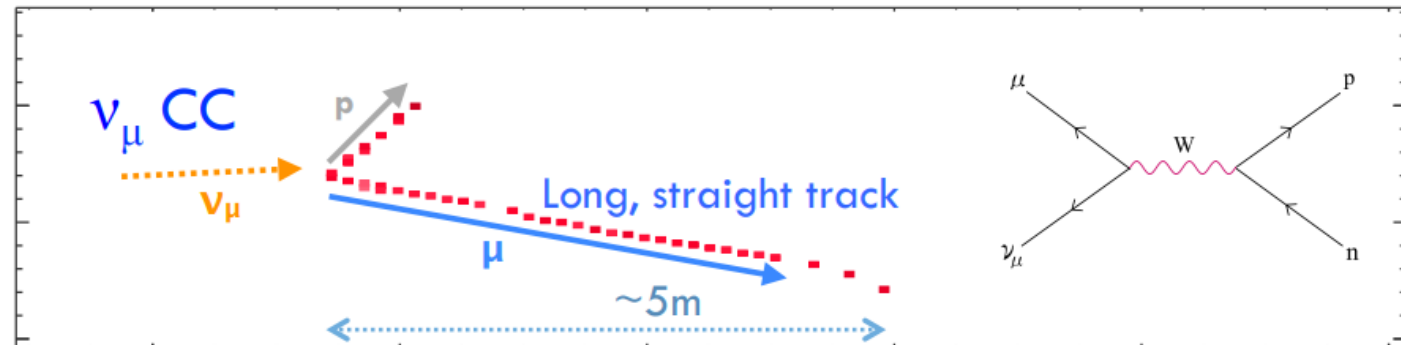
*FNAL to Ash River: 810km as the neutrino flies*

- 14 kt liquid scintillator tracker
- 290t near detector
- 0.8 deg off axis
- ~2 GeV beam, up to 700 kW

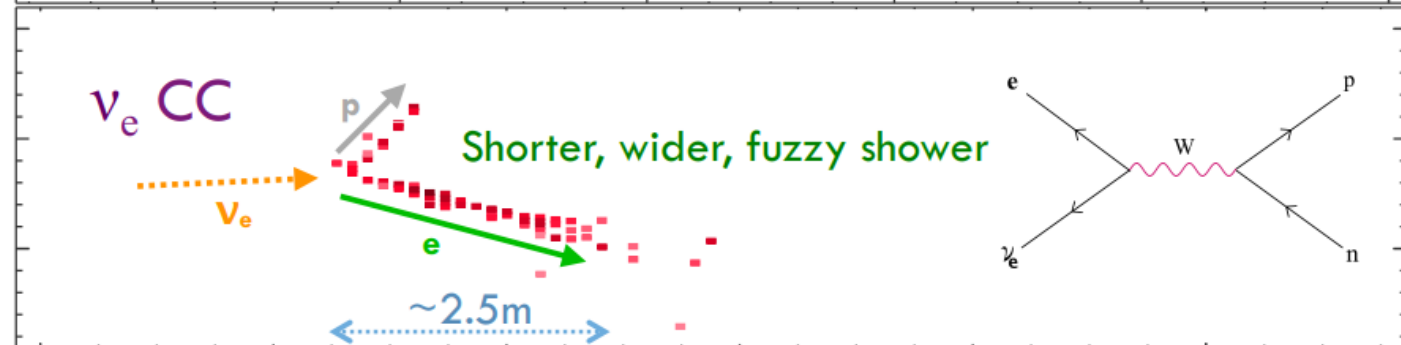


# $\text{NO}\nu\text{A}$ vital characteristics

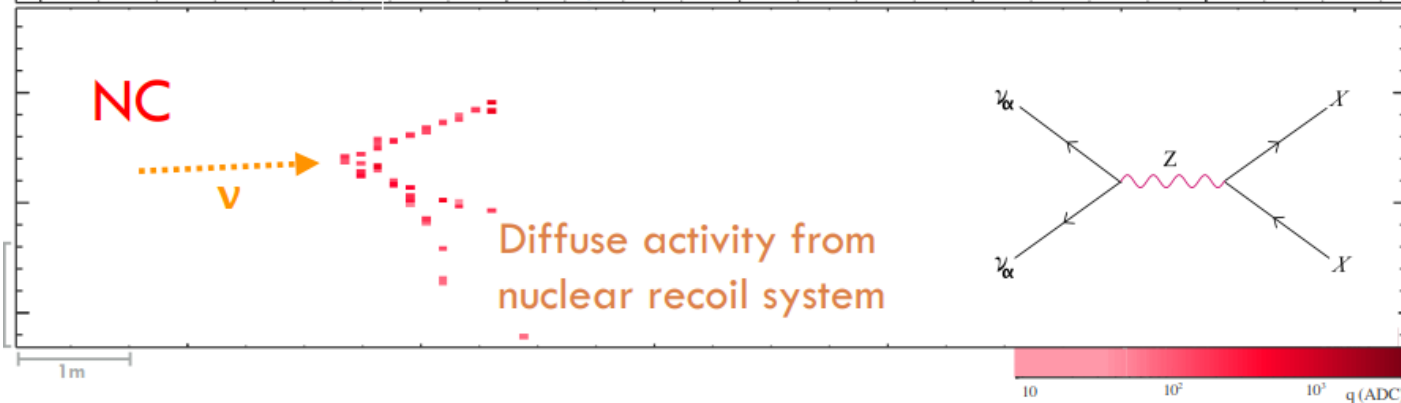
Flavour ID with  
NC sensitivity



Reconstruction  
informed by  
computer vision,  
machine learning

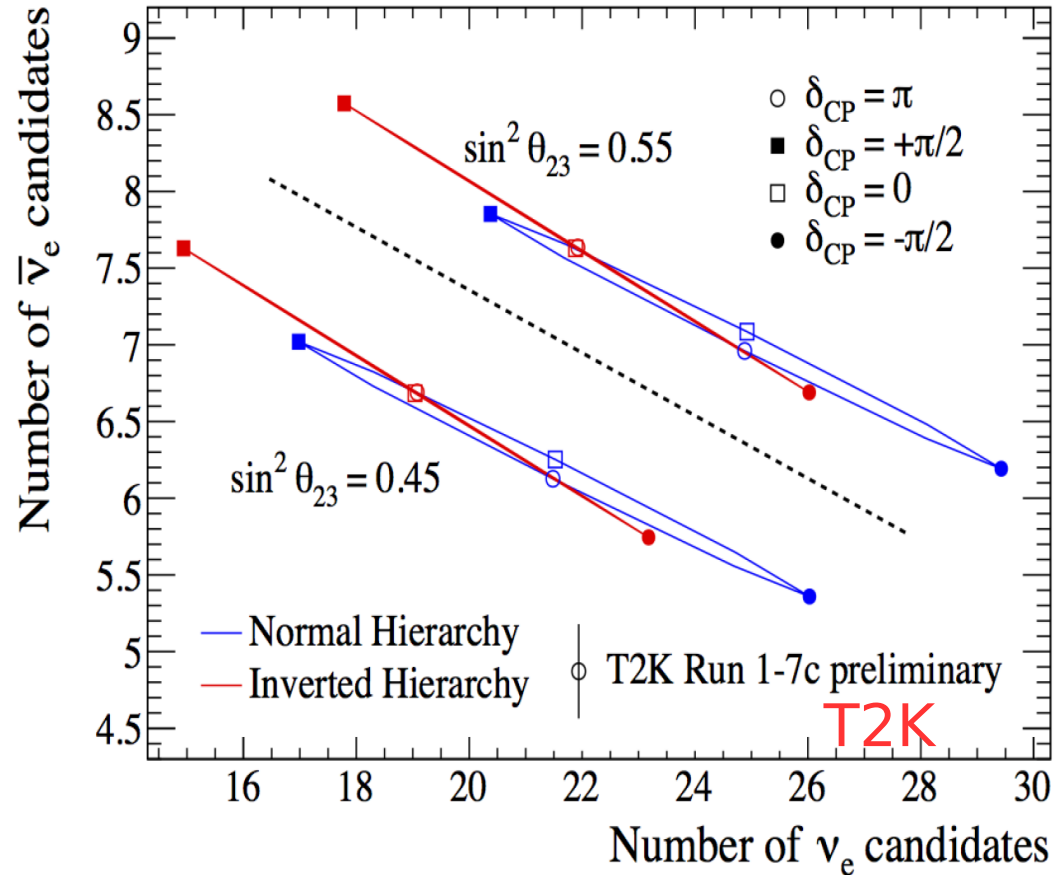
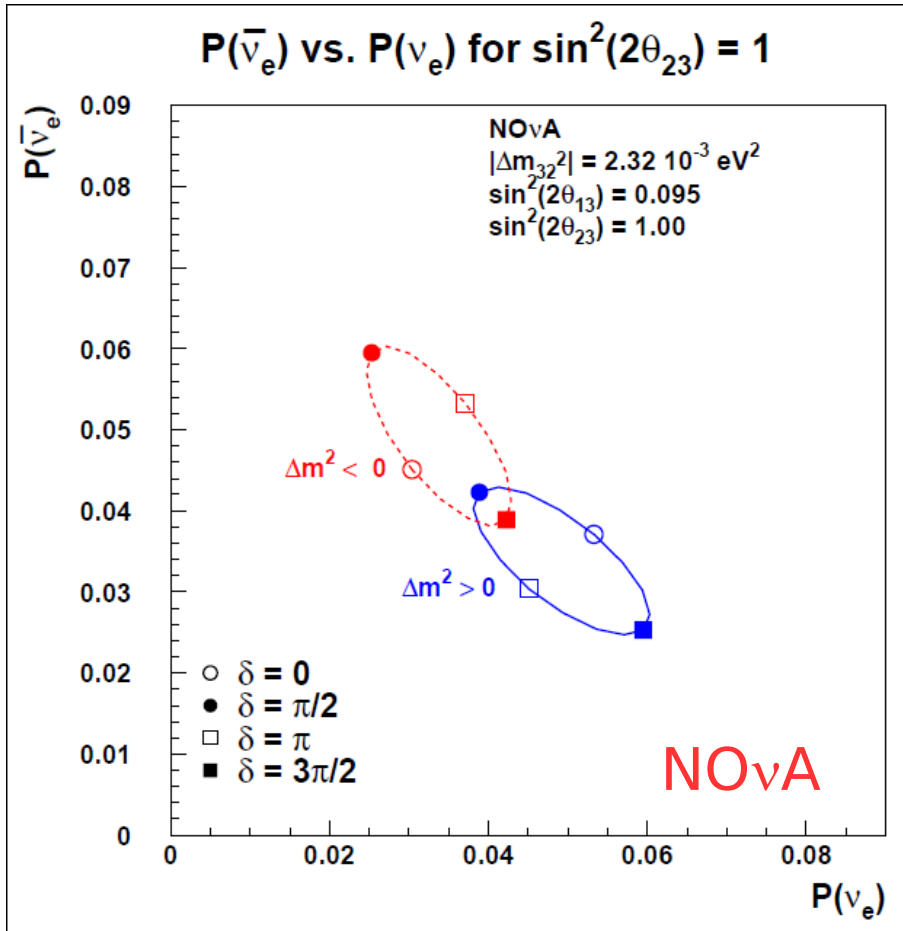


Beam energy of  
 $\sim 2\text{GeV}$  more  
sensitive than



T2K to hadronic production. Energy estimation is more calorimetric than T2K, where hadrons usually are below detection threshold. (But hadrons are harder to model.)

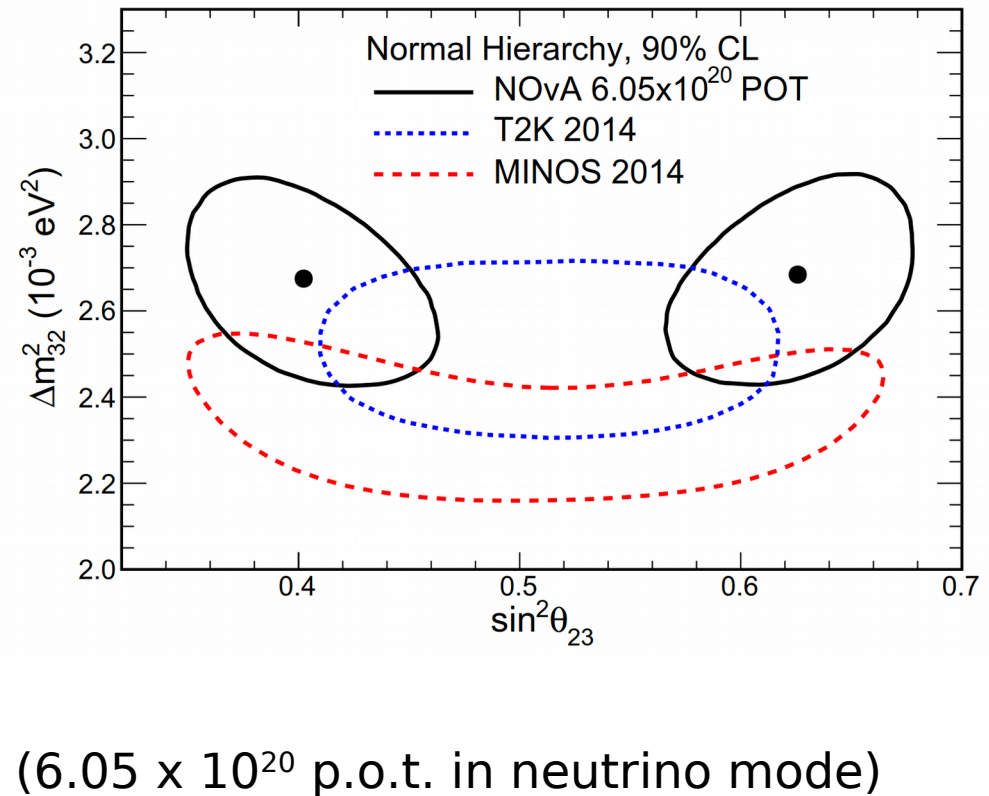
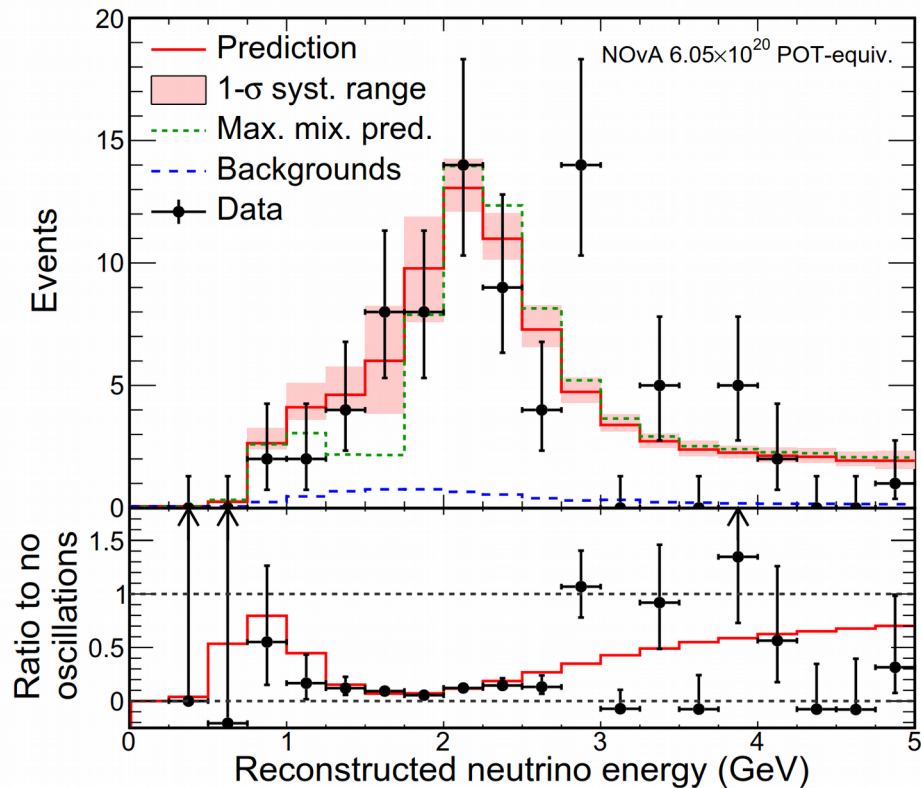
# NO<sub>v</sub>A matter effect sensitivity



$$P(\nu_\mu \rightarrow \nu_e) \approx \left(1 + 2 \frac{E}{E_R}\right) P_{vac}(\nu_\mu \rightarrow \nu_e)$$

Higher energy gives NO<sub>v</sub>A better sensitivity to matter effects than T2K

# NO $\nu$ A Oscillation Results



33  $\nu_e$  events on background of  $8.2 \pm 0.8$   
Maximal mixing disfavoured at  $2.6\sigma$

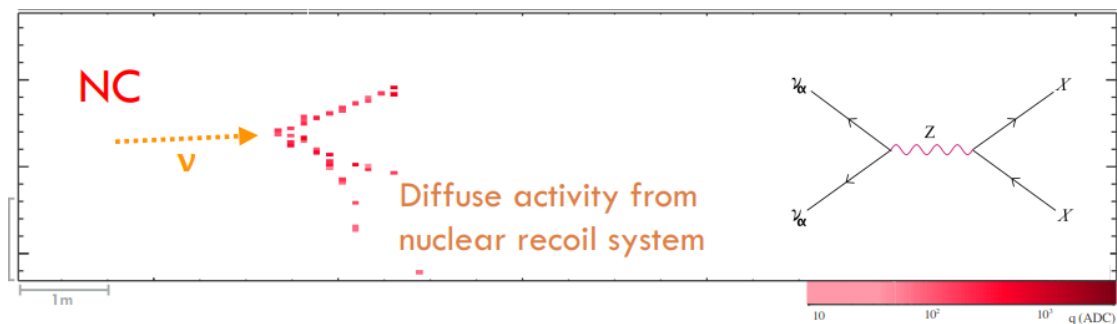
Phys. Rev. Lett. 118, 231801 (2017) arXiv: 1703.03328

Phys. Rev. Lett. 118, 151802 (2017), arXiv: 1701.05891



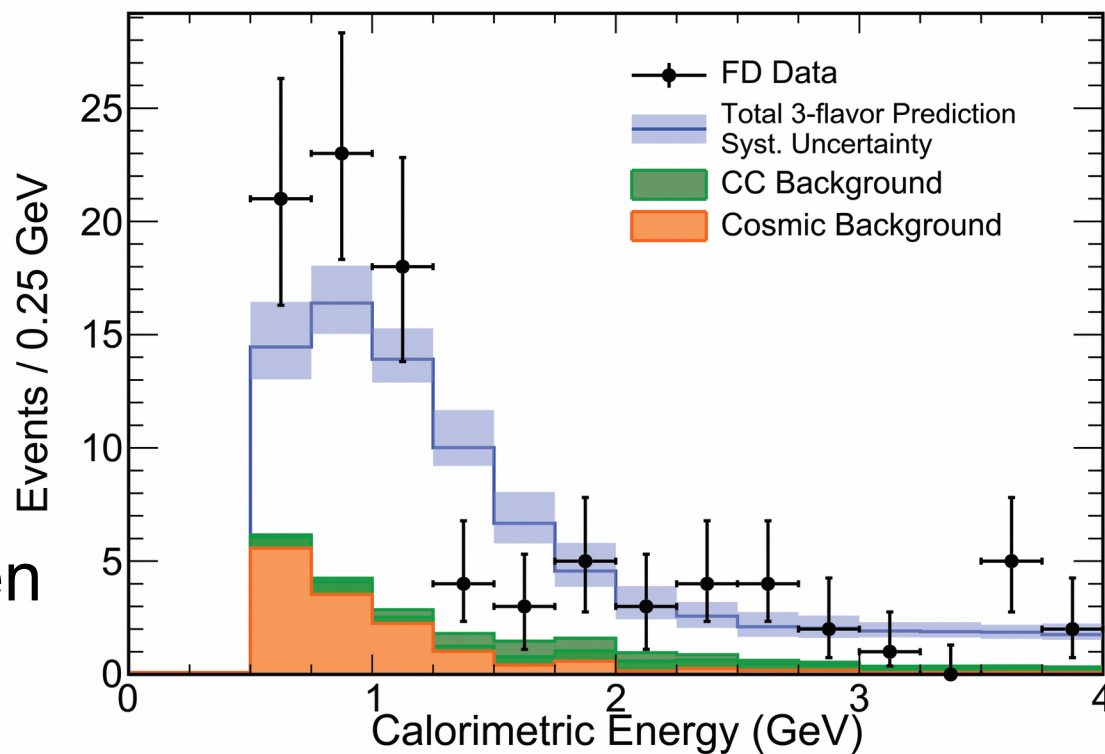
# NO $\nu$ A sterile neutrinos results

Neutral current sensitivity is unique capacity for NO $\nu$ A. A convolutional neural network is used to distinguish events with no charged lepton.



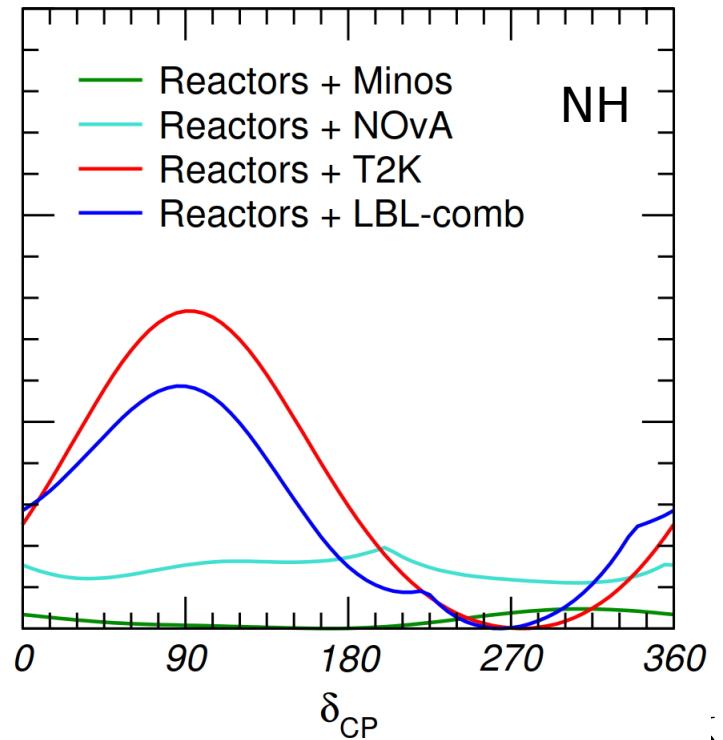
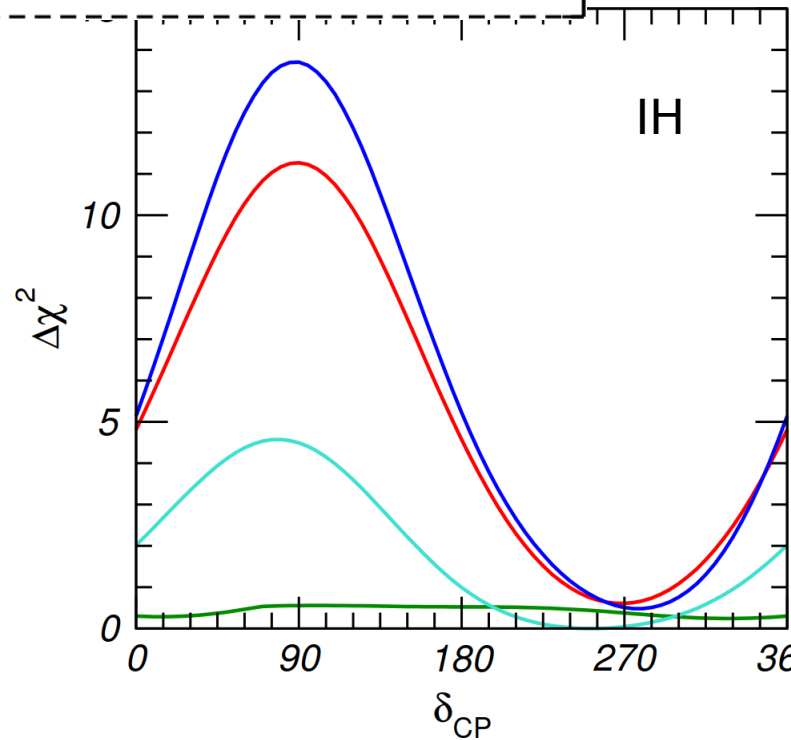
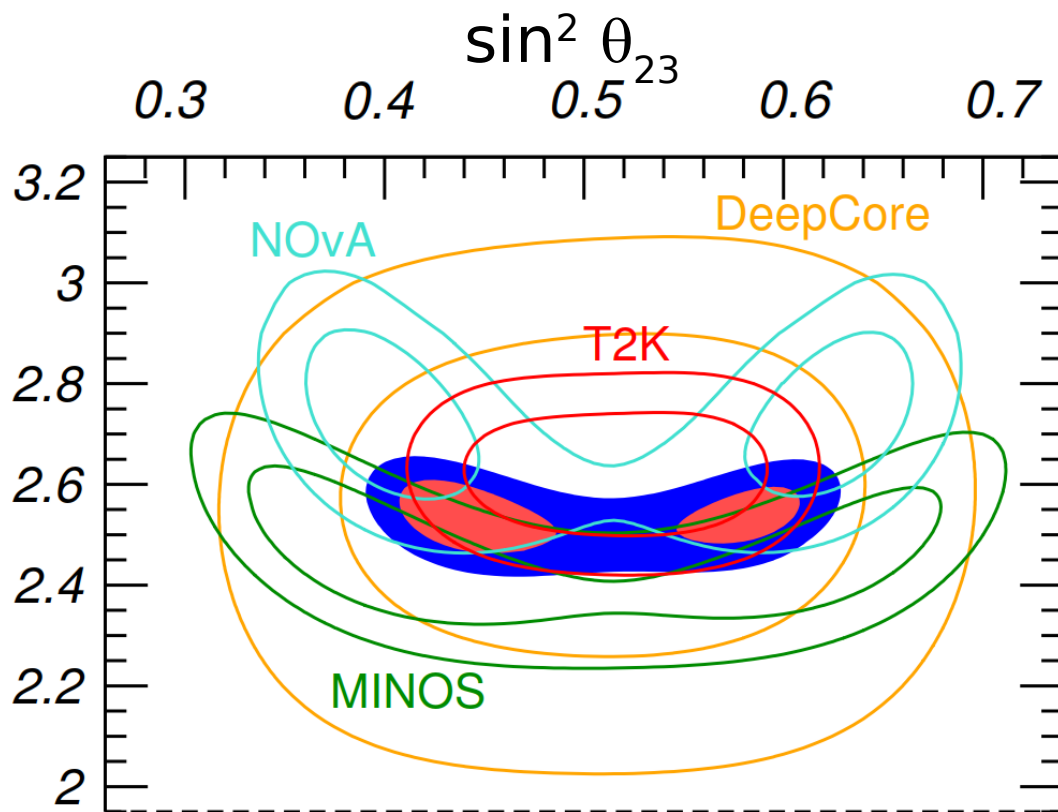
If neutrinos mix to a sterile state, then the NC rate may be less than expected.

Recent NO $\nu$ A arXiv submission 1706.04592 saw 95 events where  $83.5 \pm 9.7(\text{stat}) \pm 9.4(\text{syst})$  were predicted assuming mixing only occurs between active neutrino species.

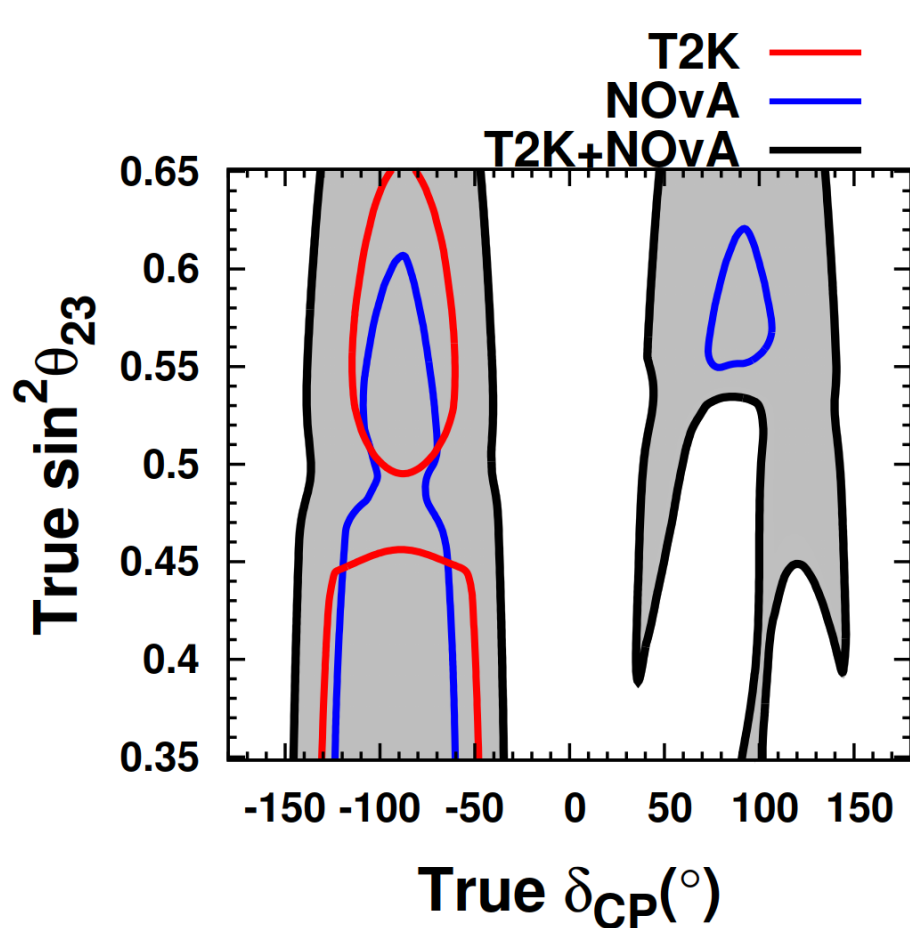


# Global oscillation parameter fits

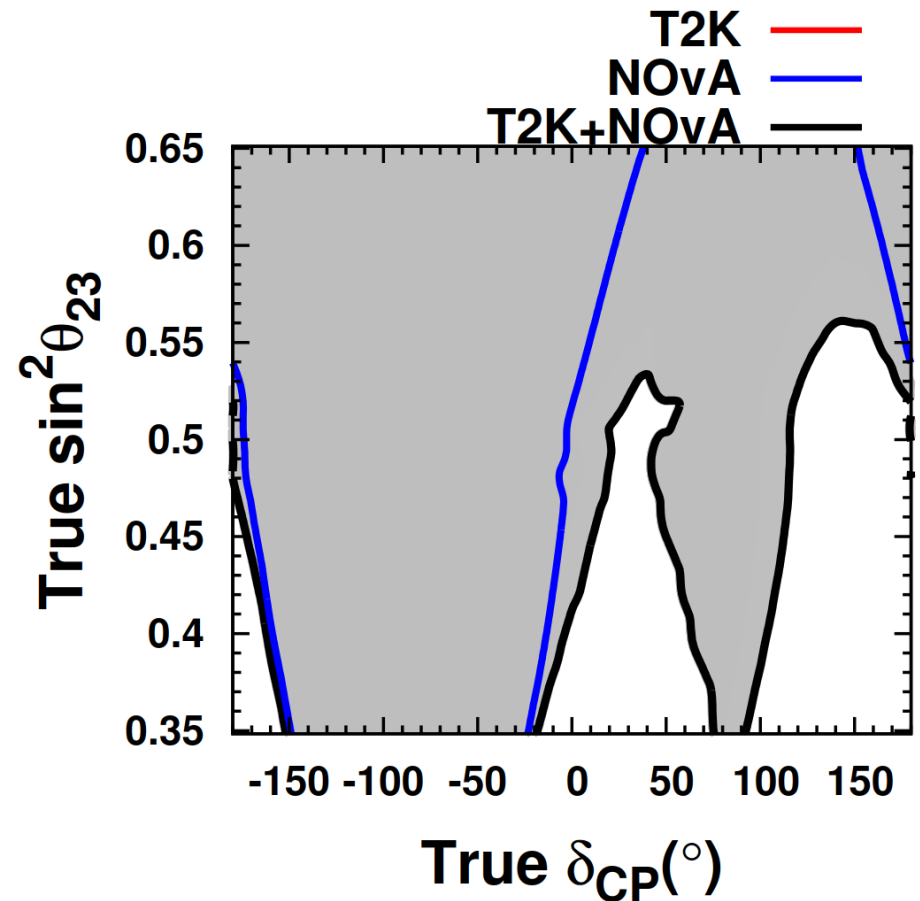
([www.nu-fit.org](http://www.nu-fit.org))



# T2K+NOvA combined sensitivity

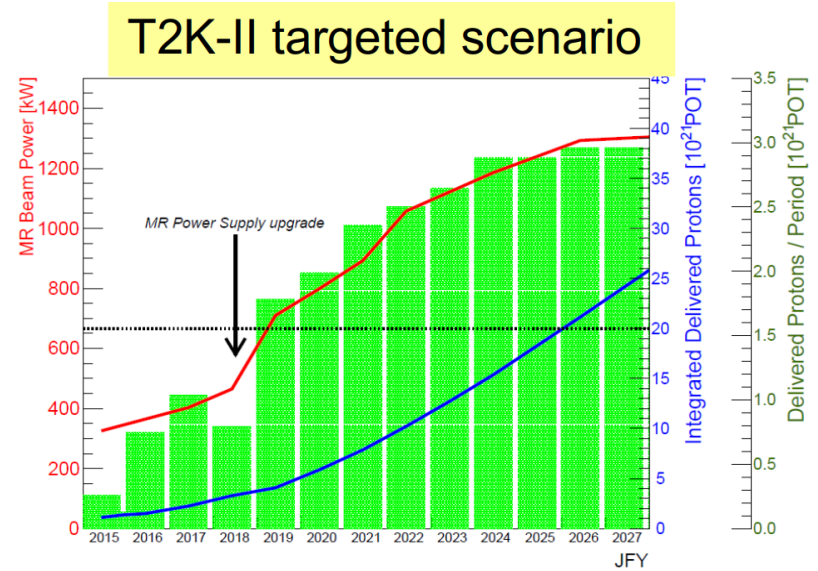
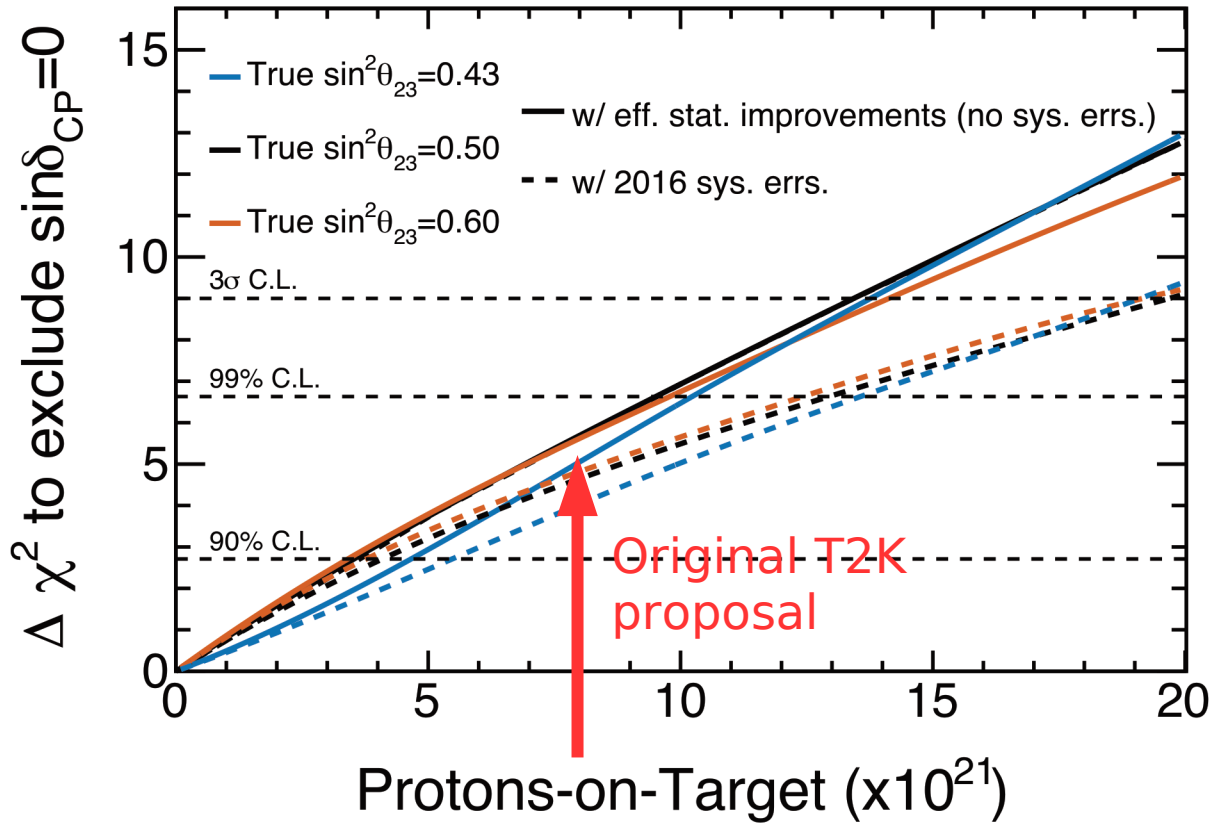


Regions for which  $\delta_{CP}=0$  is rejected at 90% CL (for normal hierarchy)



Regions for which the wrong mass hierarchy is rejected at 90% CL (for normal hierarchy)

# T2K-II



Proposal to continue running T2K until  $\sim 2025$ , with beam power and near detector upgrades, with an aim to achieve  $3\sigma$  sensitivity to non-zero  $\delta_{CP}$ .

# Conclusions

Long baseline neutrino experiments with flavour sensitivity are the only window we have on CP violation in the neutrino sector.

T2K and NO $\nu$ A results are consistent with each other and PMNS paradigm. First limits on CP violation, although still statistically weak, favour maximal CP effect.

LBL experiments dominate  $\Delta m^2_{32}$  and  $\theta_{23}$  determination.

See parallel session talks by Mark Scott (T2K—Tuesday, 17:15, Neutrino 4), Nicoletta Mauri (OPERA—Tuesday, 17:00, Neutrino 4) and Kirk Bays (NO $\nu$ A, Wednesday, 13:00, Neutrino 5) for details.

# Backup Slides

# CP Violation and $\nu_e$ Appearance

CP symmetry requires  $P(\nu_\mu \rightarrow \nu_e) = P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$

For  $\nu_e$  appearance at  $\Delta m_{32}^2$ :

$$A_{CP} = \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)} \simeq \frac{\Delta m_{12}^2 L}{4 E_\nu} \frac{\sin 2\theta_{12}}{\sin \theta_{13}} \sin \delta_{CP}$$

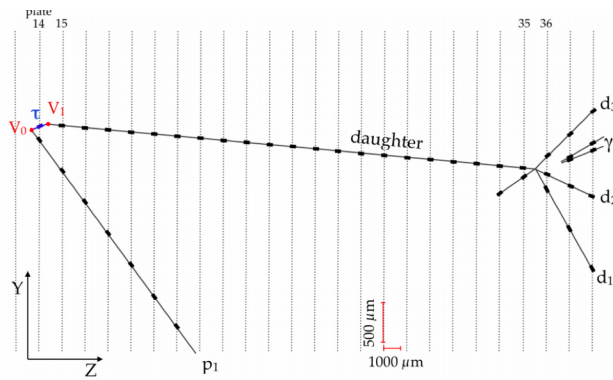
This may be a big asymmetry!

## SO WHAT?

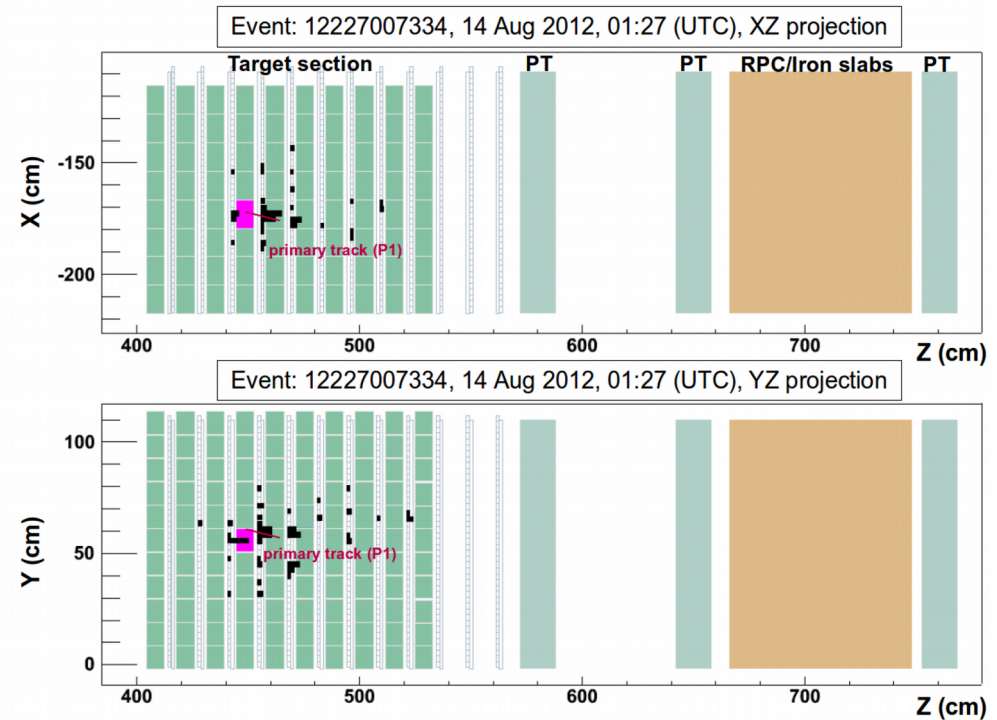
- Our universe is made of matter but not anti-matter.
- A cosmological asymmetry requires CP violation.
- Regular quark CP violation not enough---is this the missing piece?

# OPERA

Beam from CERN to Gran Sasso, looking primarily for  $\nu_\tau$  appearance using emulsion technology.



Five  $\nu_\tau$  candidates seen on a background of  $0.25 \pm 0.05$ , & a claimed significance of  $5.1\sigma$ .



Phys. Rev. Lett. 115, 121802 (2015)



# Flavour Oscillation

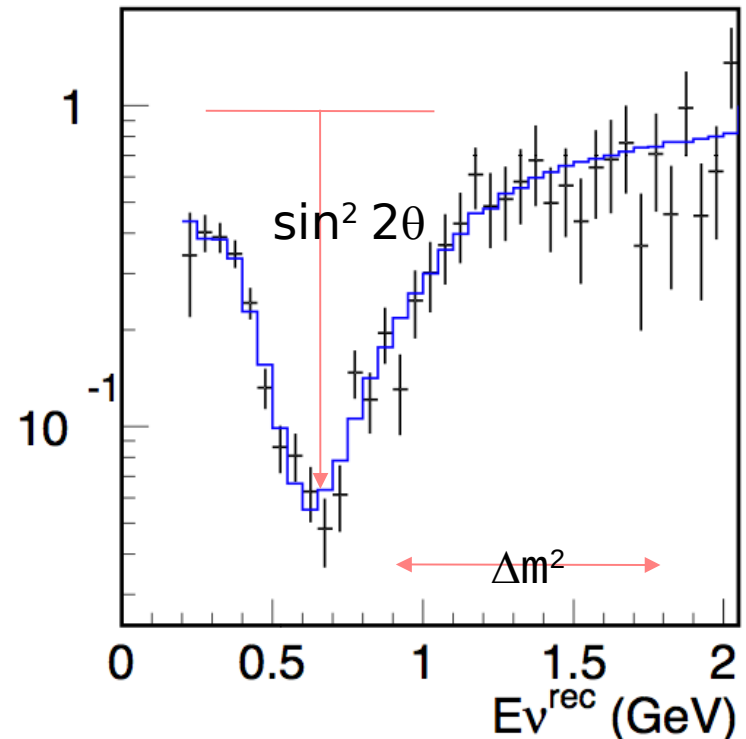
Because a flavour eigenstate produced by a weak interaction is a mix of mass eigenstates which, if  $m_1 \neq m_2$ , propagate with different kinematics, oscillation can occur.

$$|\nu(t=0)\rangle = |\nu_e\rangle = \cos\theta |\nu_1\rangle + \sin\theta |\nu_2\rangle$$

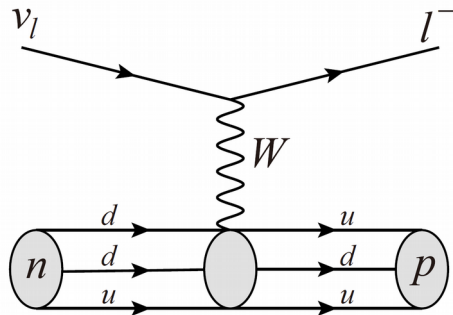
$$|\nu(t)\rangle = e^{i\sqrt{p^2+m_1^2}t} \cos\theta |\nu_1\rangle + e^{i\sqrt{p^2+m_2^2}t} \sin\theta |\nu_2\rangle$$

$$Prob(\nu_e \rightarrow \nu_e) = 1 - \sin^2(2\theta) \sin^2\left(\frac{1.27 \Delta m^2 L}{E}\right)$$

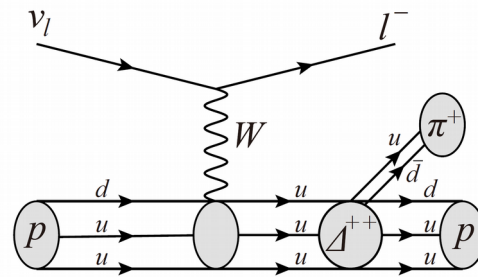
Units: [L] = km; [E] = GeV;  
 $\Delta m^2 = [\text{eV}^2]$



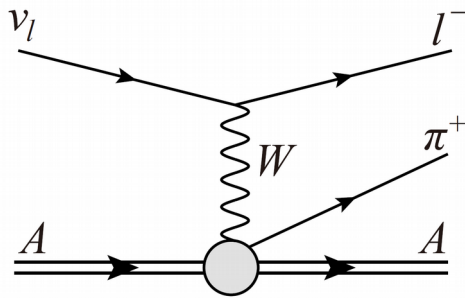
# Neutrino interactions at $\sim 1$ GeV



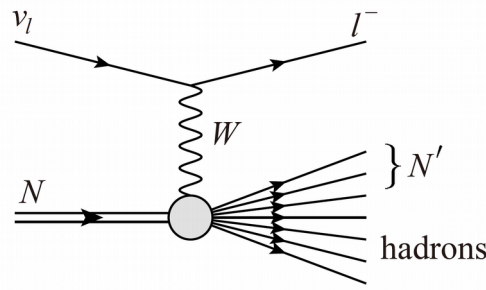
Charged-current quasielastic



Resonant Pion Production  
(CC 1 $\pi$ )

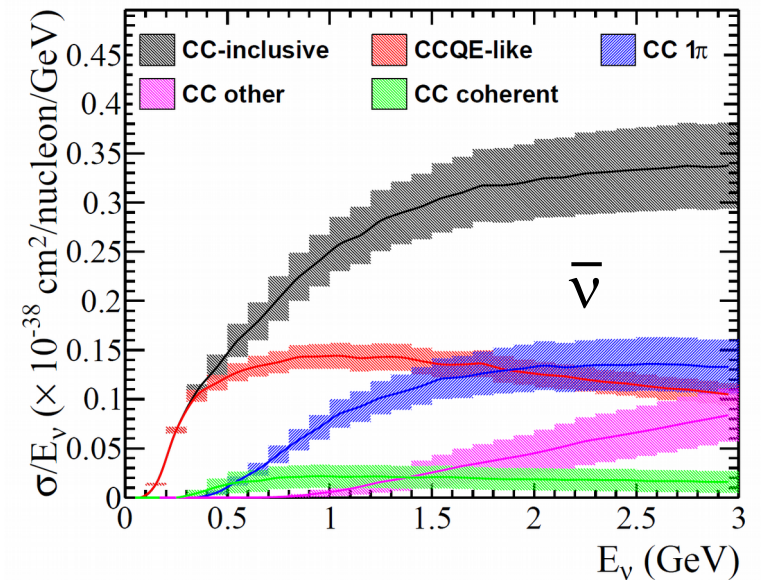
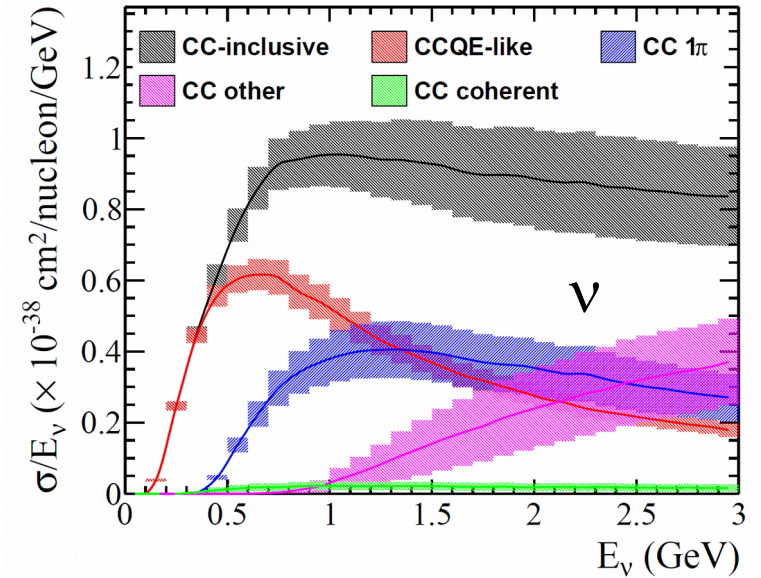


Coherent



Deep Inelastic Scattering

T2K looks for final-state leptons from charged-current interactions  
 Cross-sections not well measured – need to use near detector to normalize.



# Neutrino-Nucleon Interactions

The basic neutrino-nucleon interaction model is the dipole form factor model

- For CCQE interactions, this depends on a single physical parameter, the axial mass  $M_A$ .

CCQE interactions are particularly useful as the energy depends only on the outgoing lepton kinematics  $p_\mu, \theta_\mu$

- This is the main signal for T2K

$$E_{\text{reco}} = \frac{m_p^2 - (m_n - E_b)^2 - m_\mu^2 + 2(m_n - E_b)E_\mu}{2(m_n - E_b - E_\mu + p_\mu \cos\theta_\mu)}$$

# Nuclear Effects

Dipole form factor model is sufficient for interactions with nucleons – not nuclear targets. In a nucleus there is binding energy, and Fermi motion of nucleons (no longer at rest). A simplistic nuclear model is used:

## Relativistic Fermi Gas

Simple model of nuclear effects for CCQE interactions

- Nucleus is modeled as a Fermi gas of non-interacting neutrons and protons

Uses two nucleus-dependent parameters

- $E_B$  : the nucleon binding energy
- $p_F$  : the Fermi momentum
- Different for each nucleus

## Random Phase Approximation

Correction to the RFG model

- Includes first-order nucleon-nucleon correlations not found in the RFG model

Models long range correlations between nucleons at low energies

- Not strongly nucleus dependent

# Modelling CCQE interactions

Axial form factor for nucleon-neutrino interactions modelled by dipole parametrization:

$$F_A(Q^2) = g_A \left( 1 + \frac{Q^2}{M_A^2} \right)^{-2}$$

Measurements of the axial mass from different experiments are all over the map!

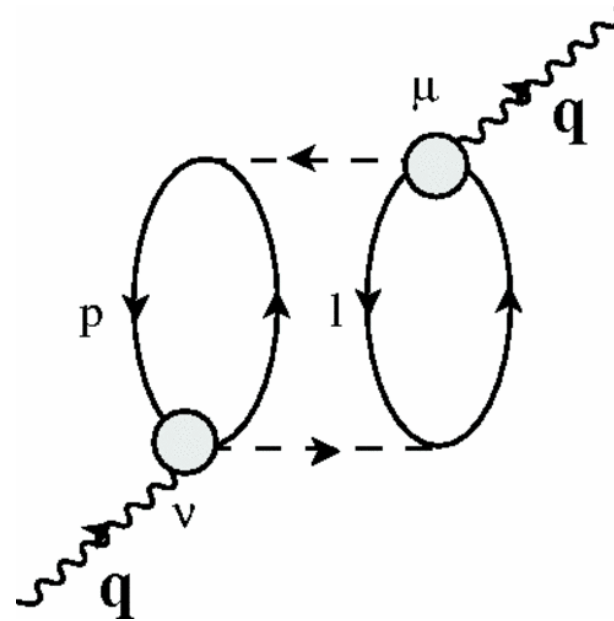
World average:	$1.012 \pm 0.031 \pm 0.060$ GeV
Deuterium experiments:	$0.99 \pm 0.04$ GeV
MiniBooNE (carbon):	$1.35 \pm 0.17$ GeV
K2K (carbon):	$1.20 \pm 0.12$ GeV

It looks like other effects confound the free nucleon form factor.

# Multi-nucleon effects

Other multinucleon correlations are not covered in the RFG + RPA model

- These interactions can produce multiple protons or neutrons in the final state – difficult to identify separately from CCQE
- Irreducible experimental background

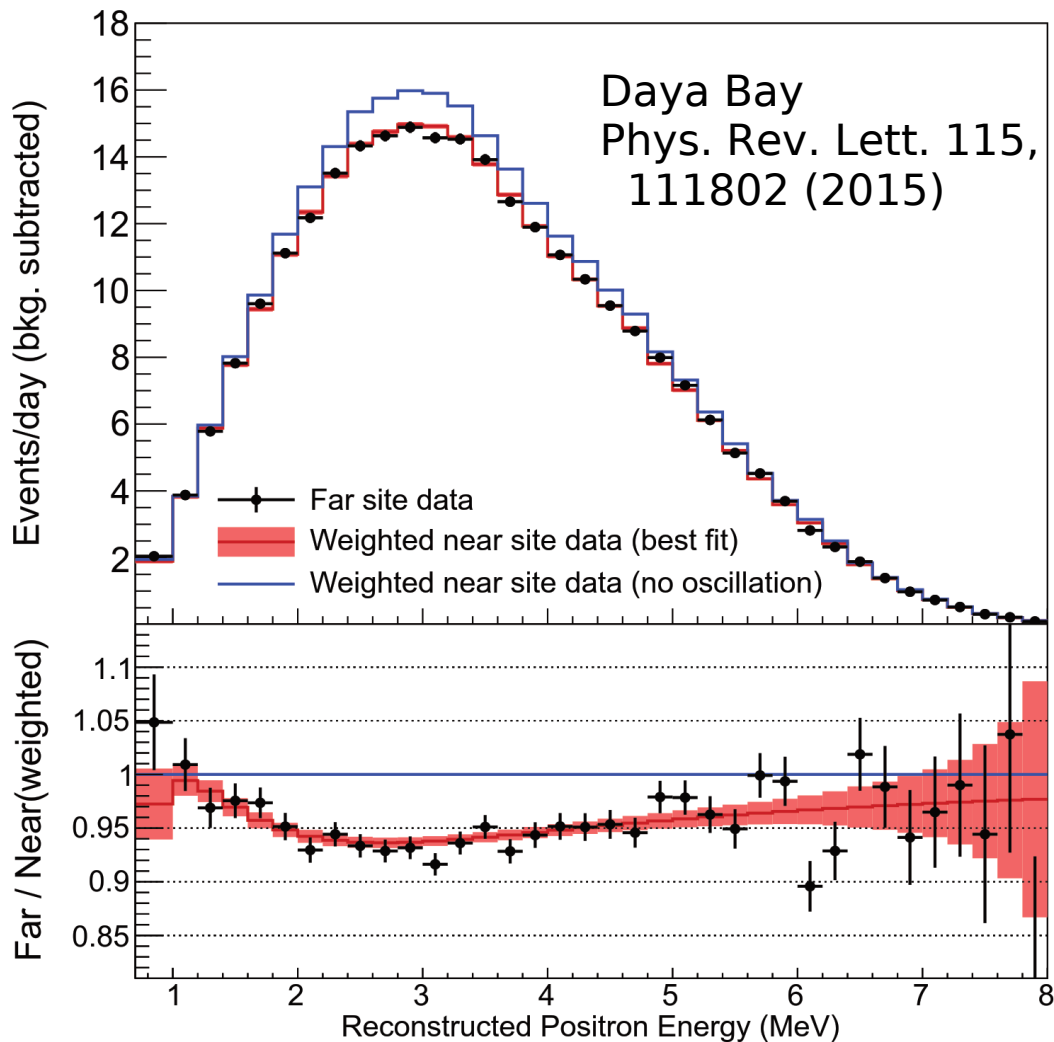


T2K models 2p-2h interactions, where two particle – hole pairs are propagated through the nucleus

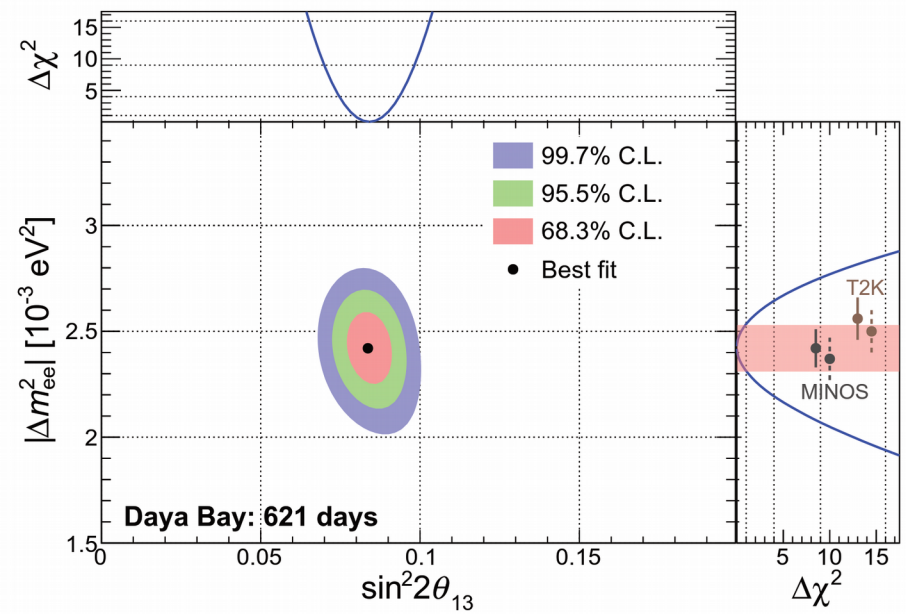
- Can produce two nucleons in the final state
- If not modeled, can have a significant effect on axial mass measurement: originally introduced to solve tensions between the axial mass measured with MiniBooNE and global averages
- T2K uses the Nieves 2p-2h model in the NEUT generator

# Reactor neutrinos & $\theta_{13}$

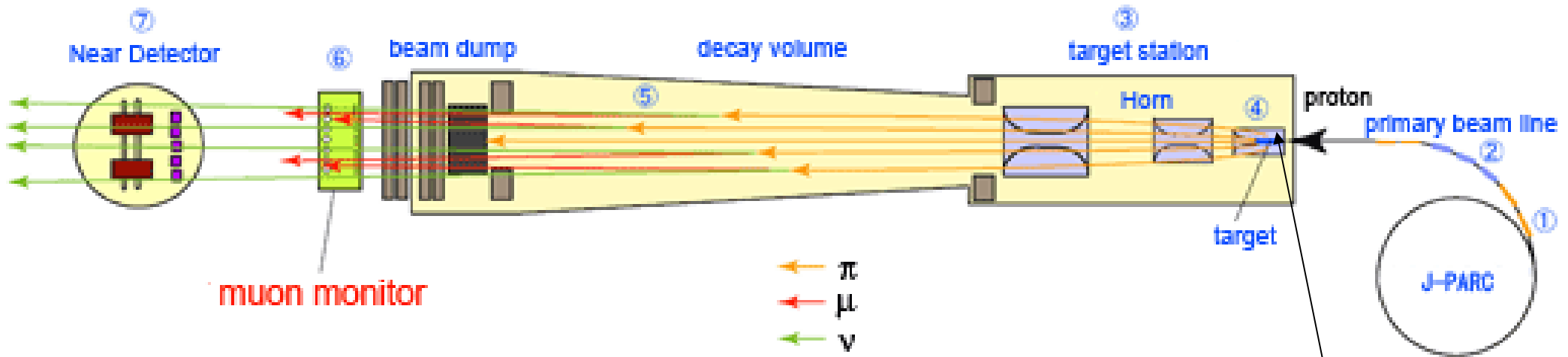
$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{13} \sin^2 \left( \frac{1.27 \Delta m_{31}^2 L}{E} \right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left( \frac{1.27 \Delta m_{21}^2 L}{E} \right)$$



Daya Bay, RENO,  
Double CHOOZ look for  
disappearance of  
reactor neutrinos at  
 $\sim 1\text{km}$  baseline.



# How To Make A Neutrino Beam

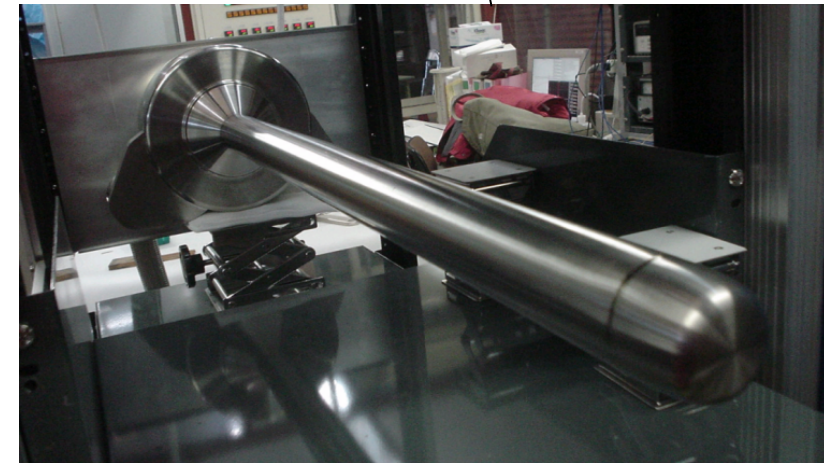


30 GeV protons hit graphite target

3 magnetic horns focus  $\pi^+$ , defocus  $\pi^-$ .

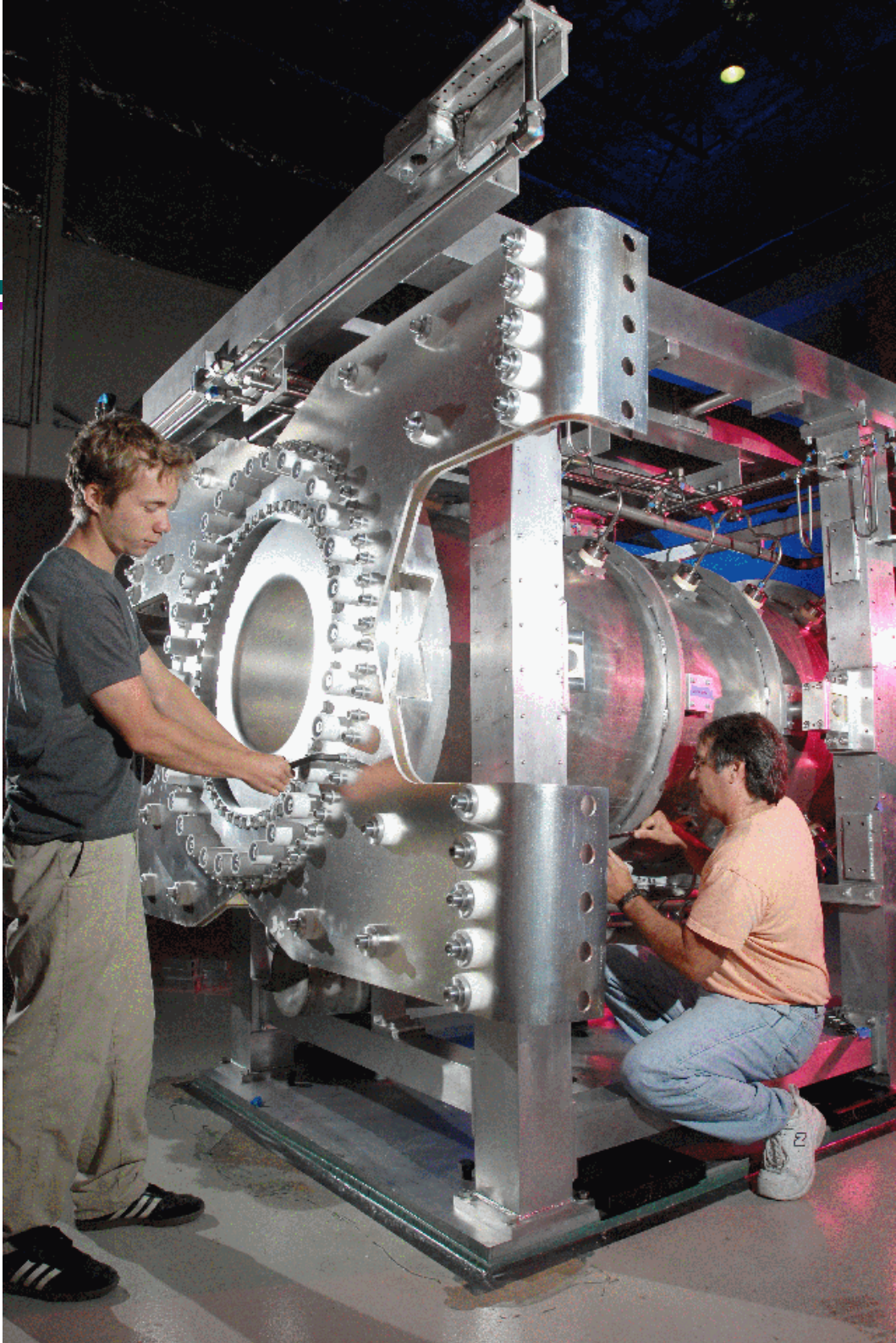
$\pi^+ \rightarrow \mu^+ + \nu_\mu$  in 110m long decay pipe

$\mu$  monitor at far end of beam dump:  
fluence:  $10^8 \mu/\text{cm}^2/\text{spill}$  at full power



T2K's 90cm graphite target

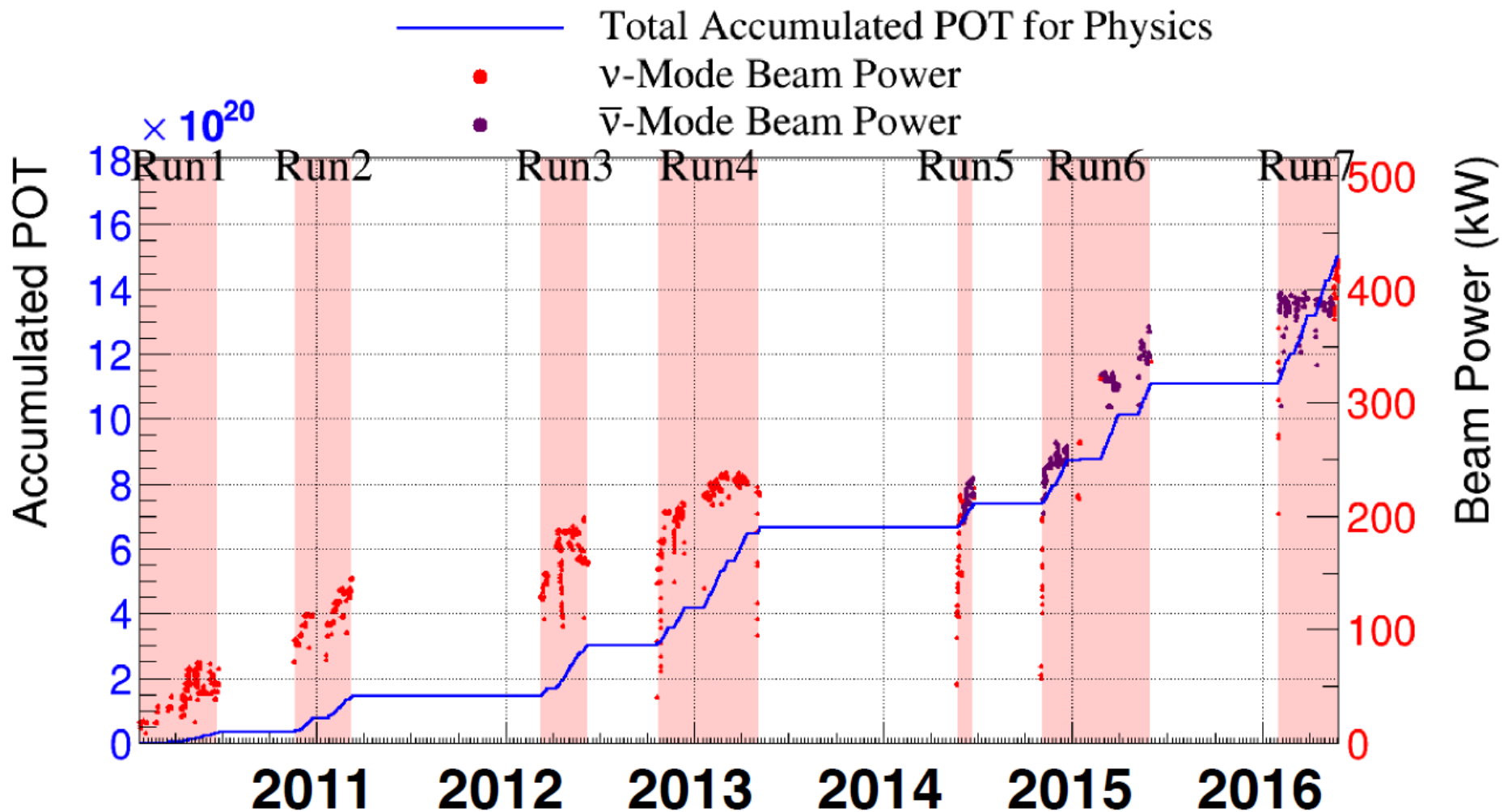




↑  
Inside the decay  
volume

← The 2<sup>nd</sup> focusing  
horn

# T2K data collection



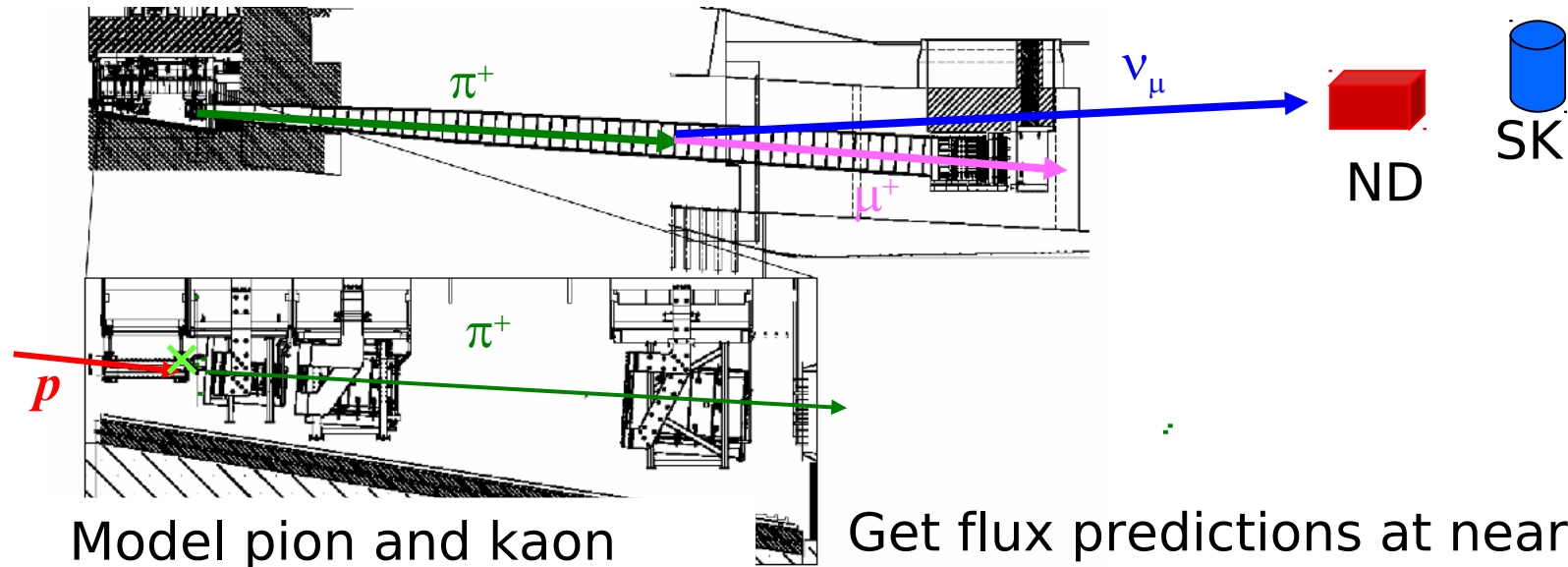
27 May 2016

POT total:  $1.510 \times 10^{21}$

ν-mode POT:  $7.57 \times 10^{20}$  (50.14%)

ν̄-mode POT:  $7.53 \times 10^{20}$  (49.86%)

# T2K: Flux prediction (Beam MC)

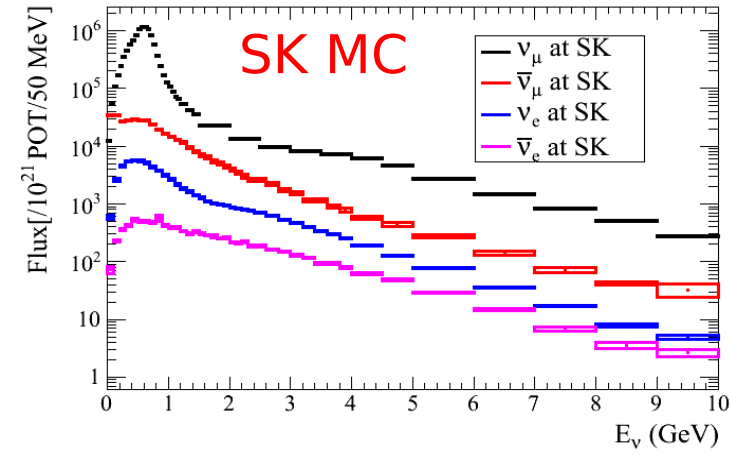
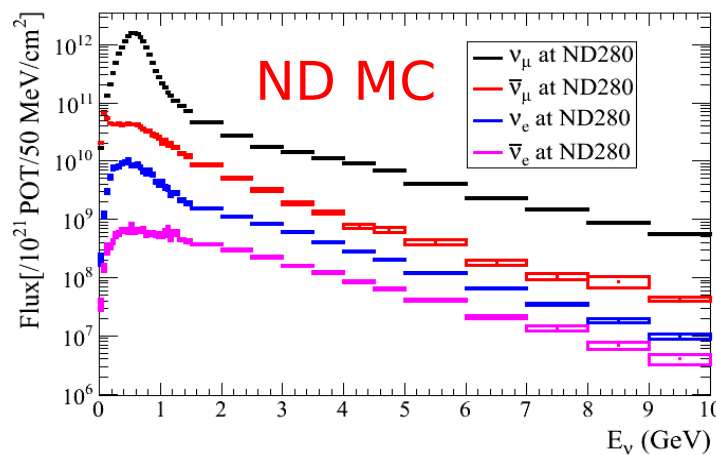


Simulate hadron production on target using FLUKA simulation

Model pion and kaon propagation and decay through horns and beamline

Get flux predictions at near detector and SK

Particle production cross sections tuned to external data from NA61 and others.



# Backgrounds to $\nu_e$ Appearance

Intrinsic beam  $\nu_e$ :

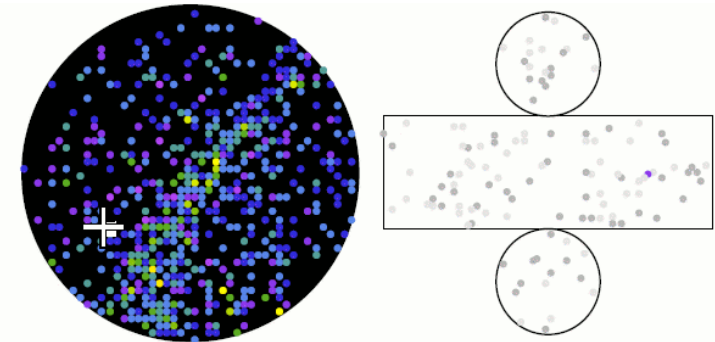
- reduce with E cut
- measure at ND

$\pi^0$  production, with one  $\gamma$  from event not detected at Super-K:

- better ID algorithms
- measure at ND
- measure  $\pi^0$  in SK

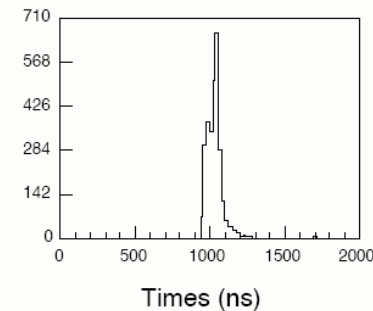
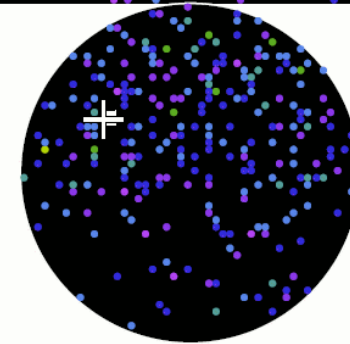
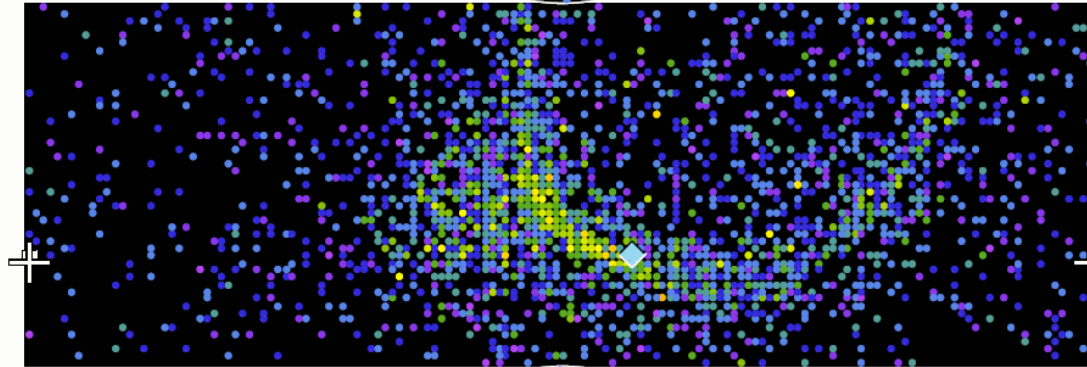
The intrinsic beam events are a more significant background.

*Do you see the 2<sup>nd</sup> ring?*

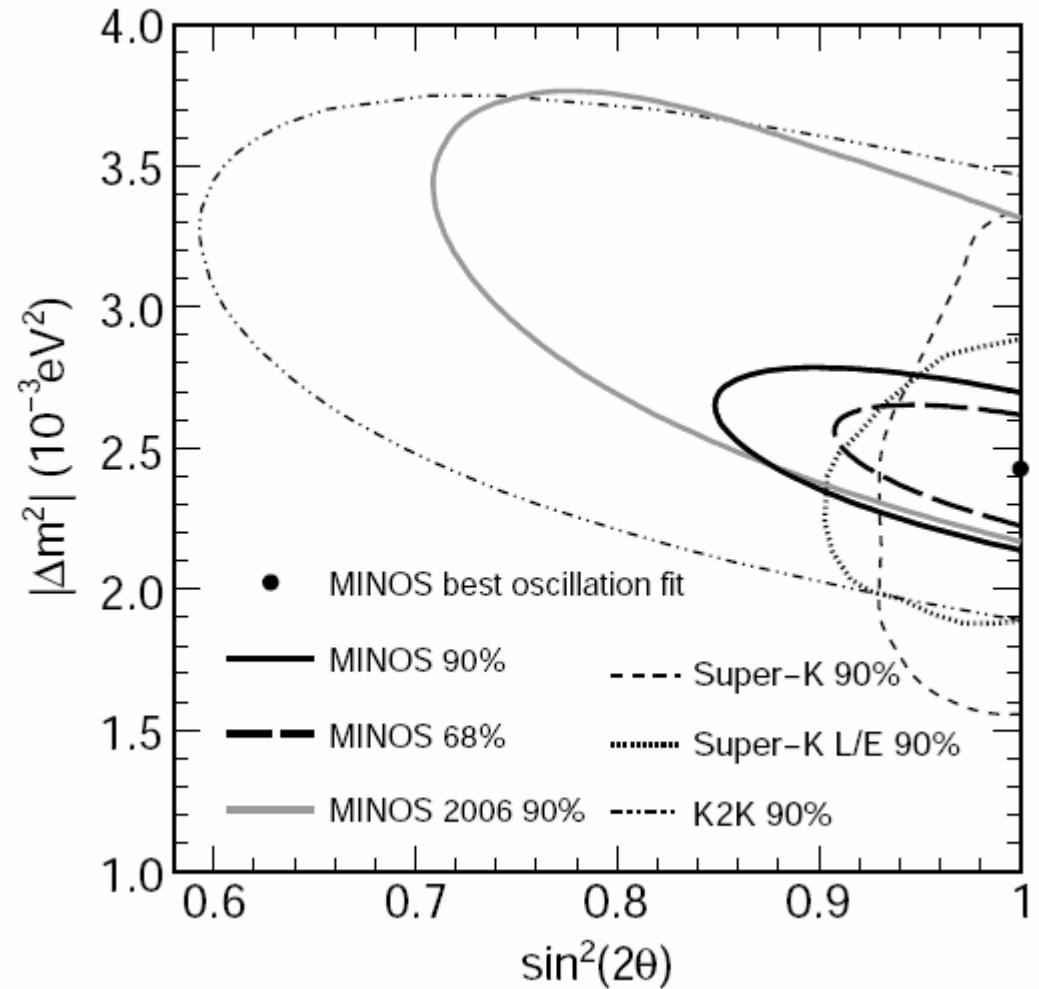
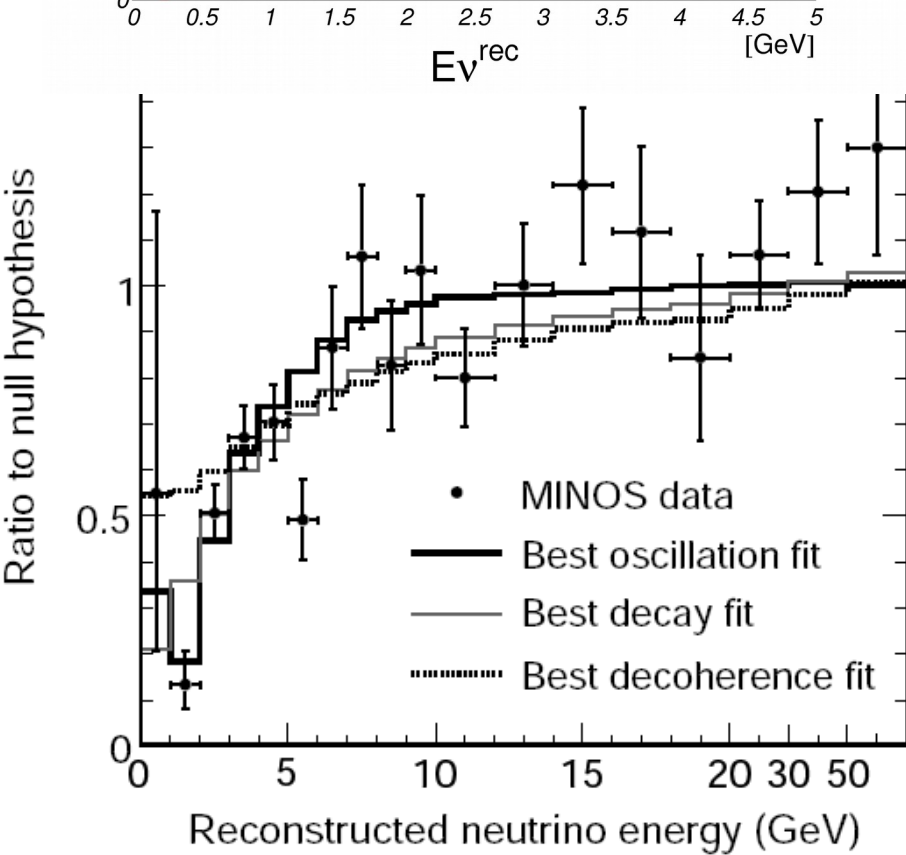
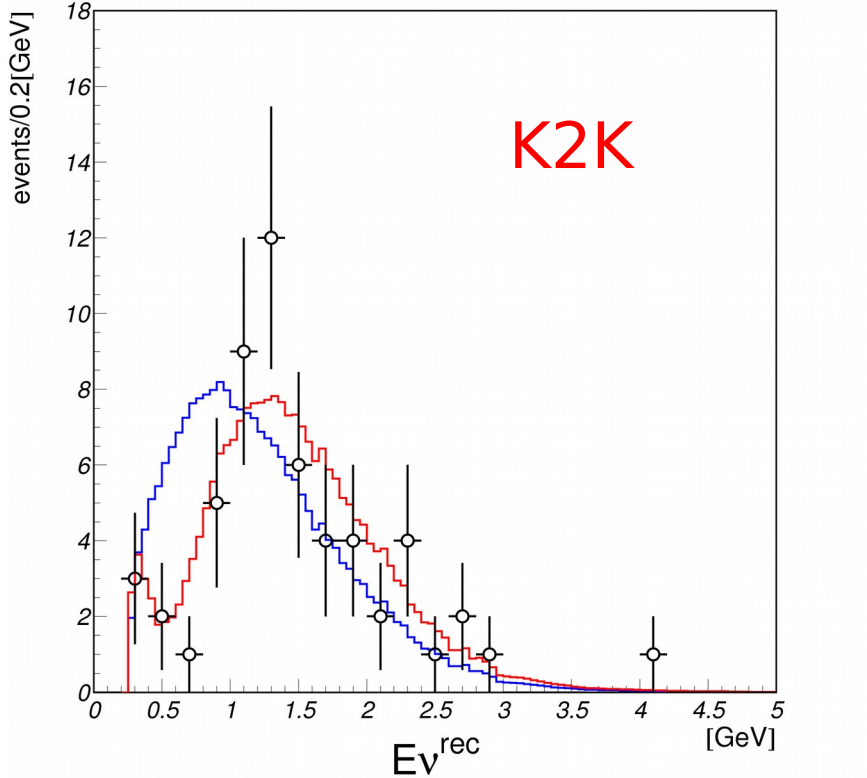


Charge (pe)

- >26.7
- 23.3-26.7
- 20.2-23.3
- 17.3-20.2
- 14.7-17.3
- 12.2-14.7
- 10.0-12.2
- 8.0-10.0
- 6.2- 8.0
- 4.7- 6.2
- 3.3- 4.7
- 2.2- 3.3
- 1.3- 2.2
- 0.7- 1.3
- 0.2- 0.7
- < 0.2

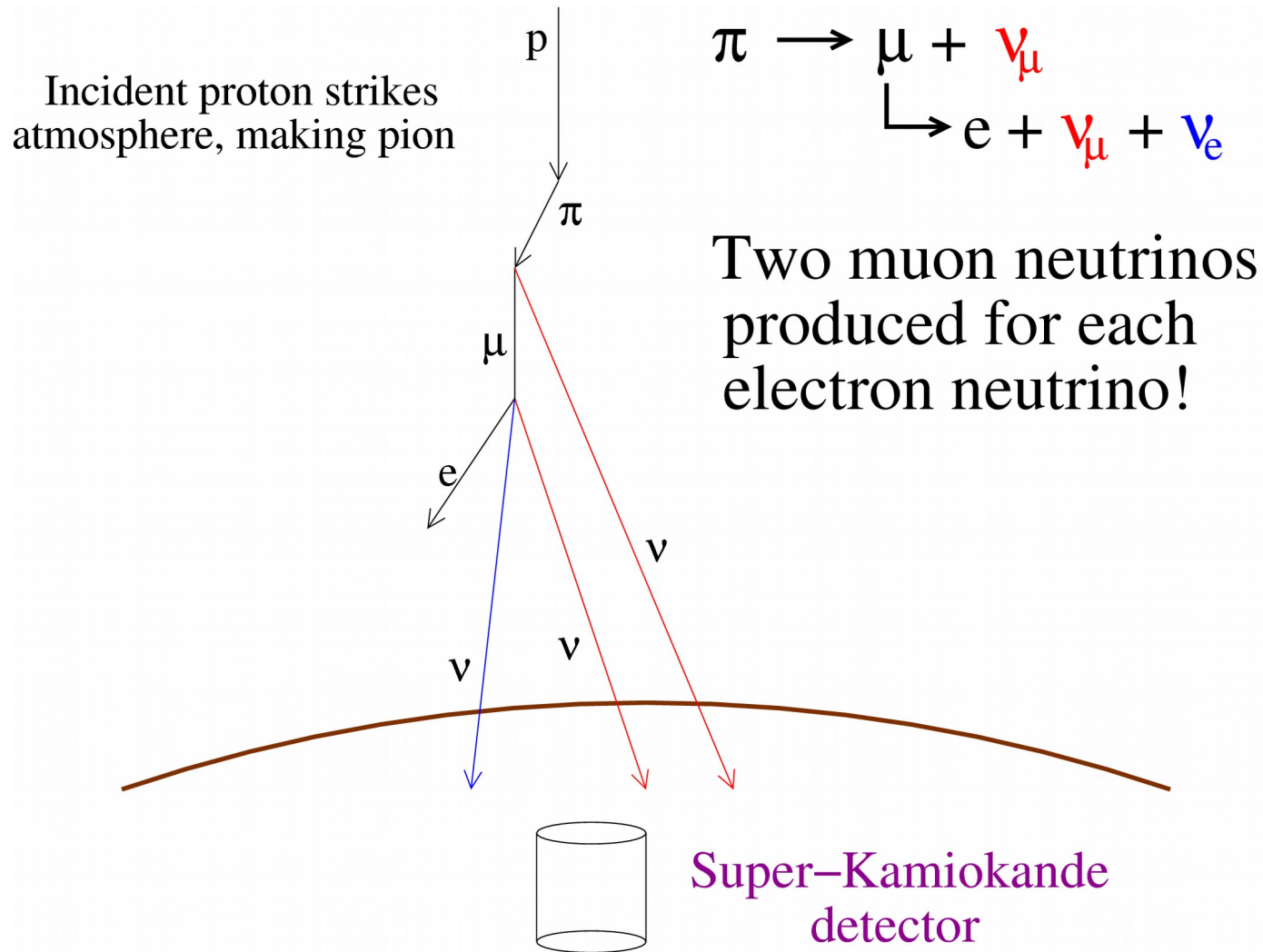


# K2K & MINOS

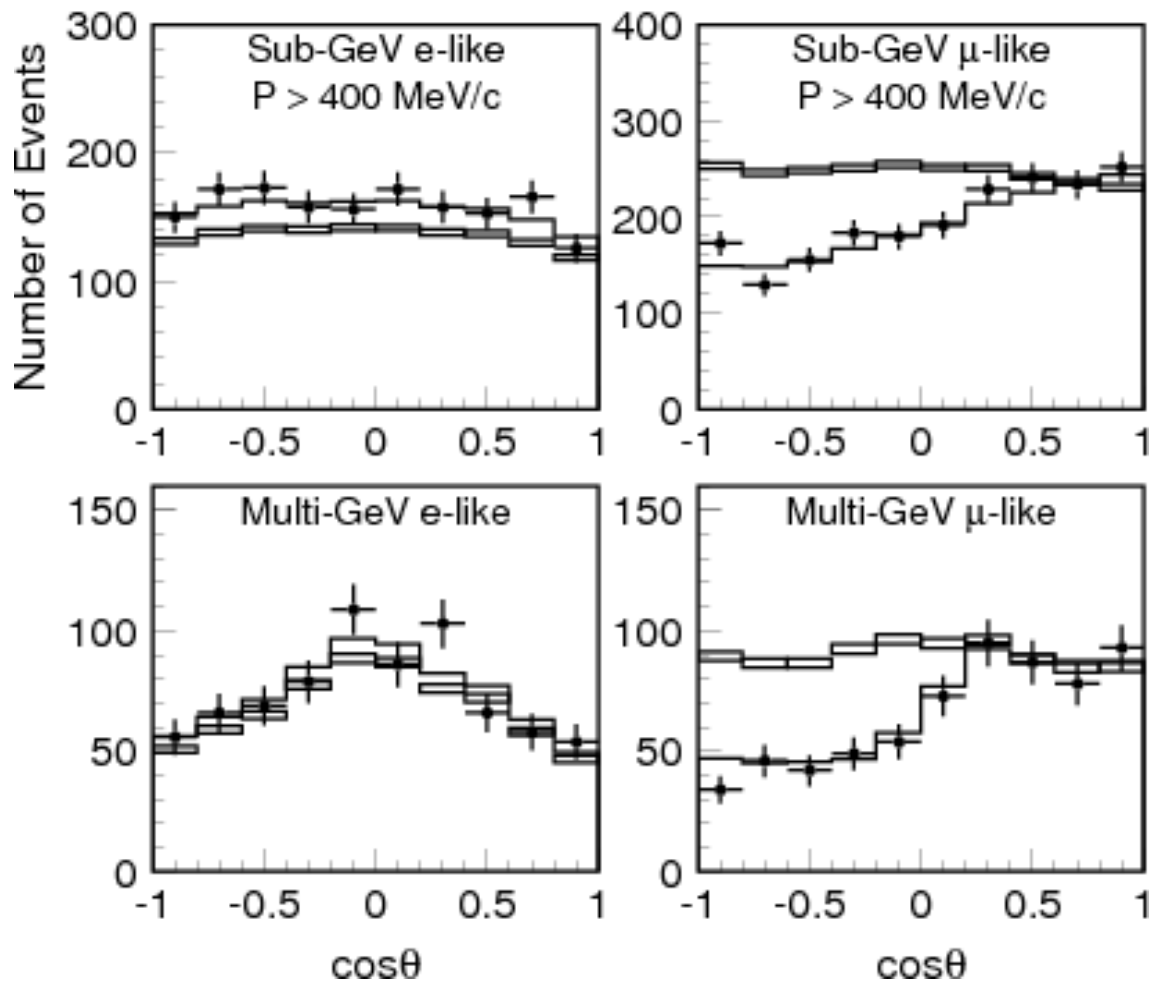


Consistency between atmospheric and long-baseline  $\nu$  oscillation results.

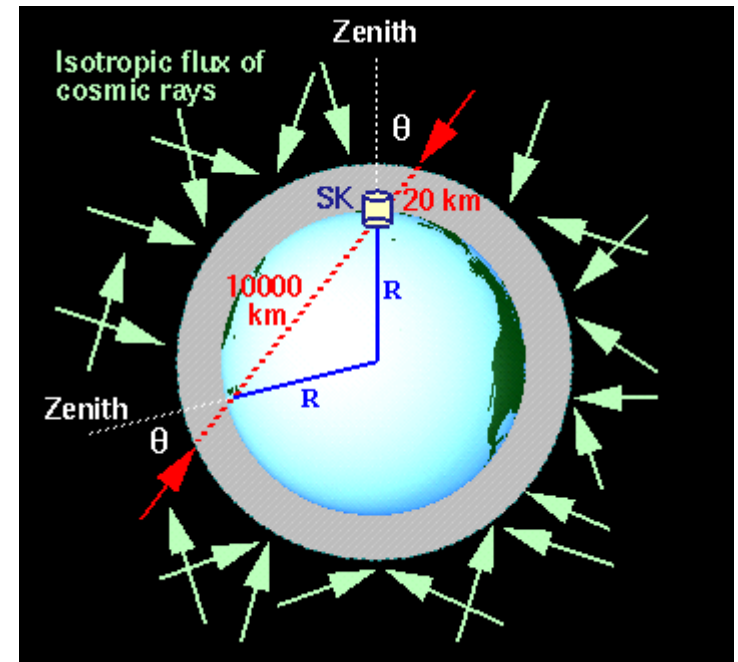
# Atmospheric Neutrinos



# Super-K atmospheric $\nu$ results



PRL 93:101801, 2004  
PRD 71:112005, 2005

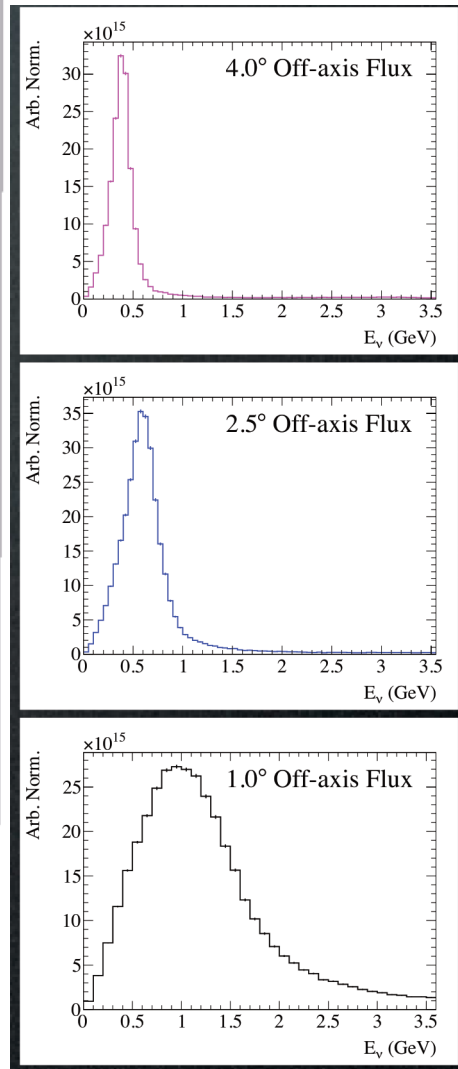
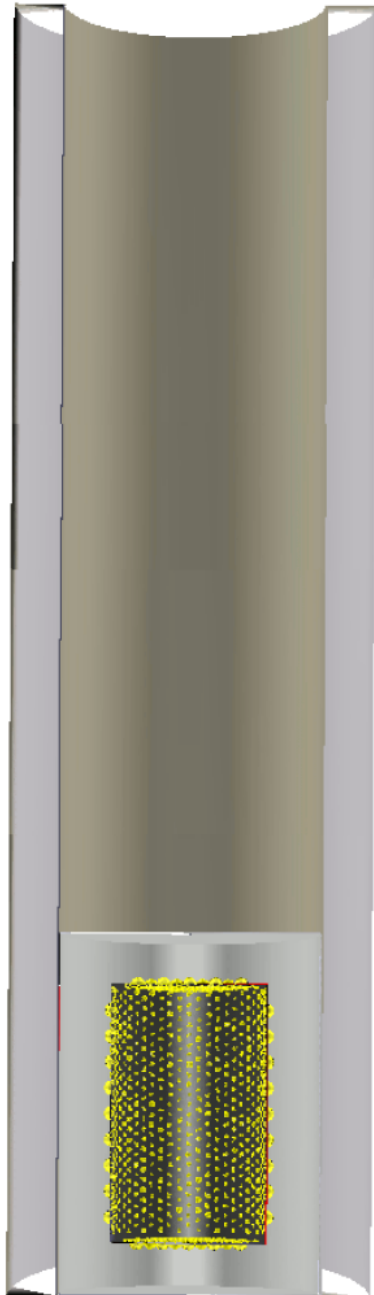


Deficit of upward-going  $\nu_\mu$  relative to downward-going.

No deficit for  $\nu_e$ .

Seems like  $\nu_\mu \rightarrow \nu_\tau$

# NuPRISM concept



New concept to exploit the variation in neutrino energy with off-axis angle: tall water Cherenkov near detector spanning range of off-axis angles.

Data taken at different angles can directly predict neutrino interactions with arbitrary neutrino fluxes, including effects from oscillation.

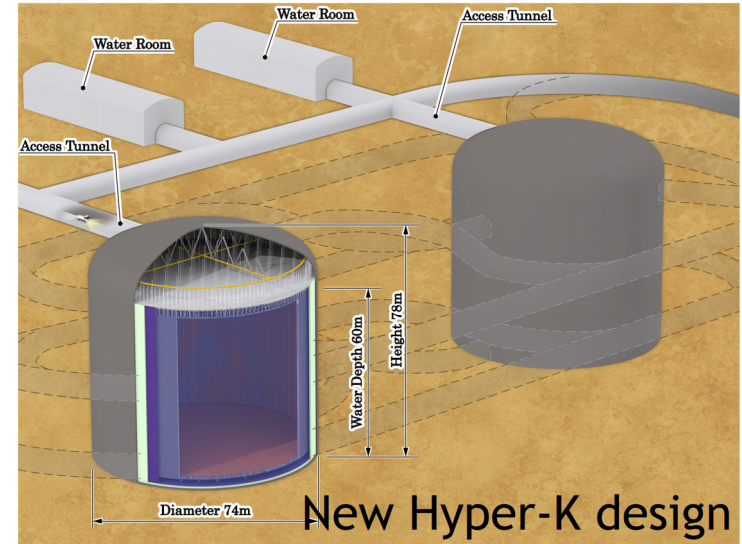
Main proposal to Canadian government under review.



# Hyper-K

New technical design report:  
187kton fiducial volume tank with  
enhanced photosensor coverage.

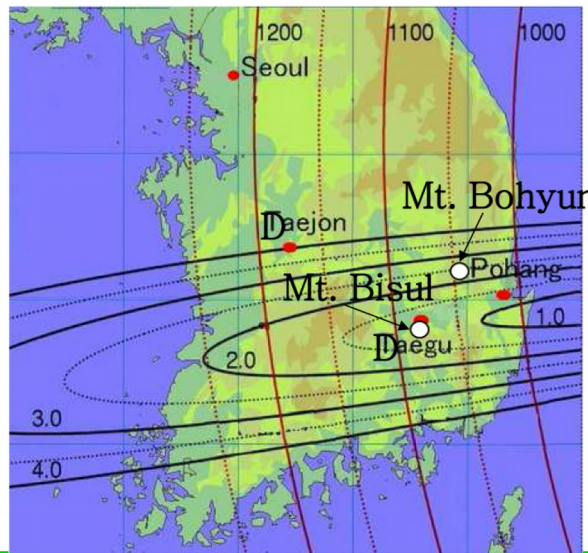
Proposal to locate a second  
detector in Korea at the second  
oscillation maximum.



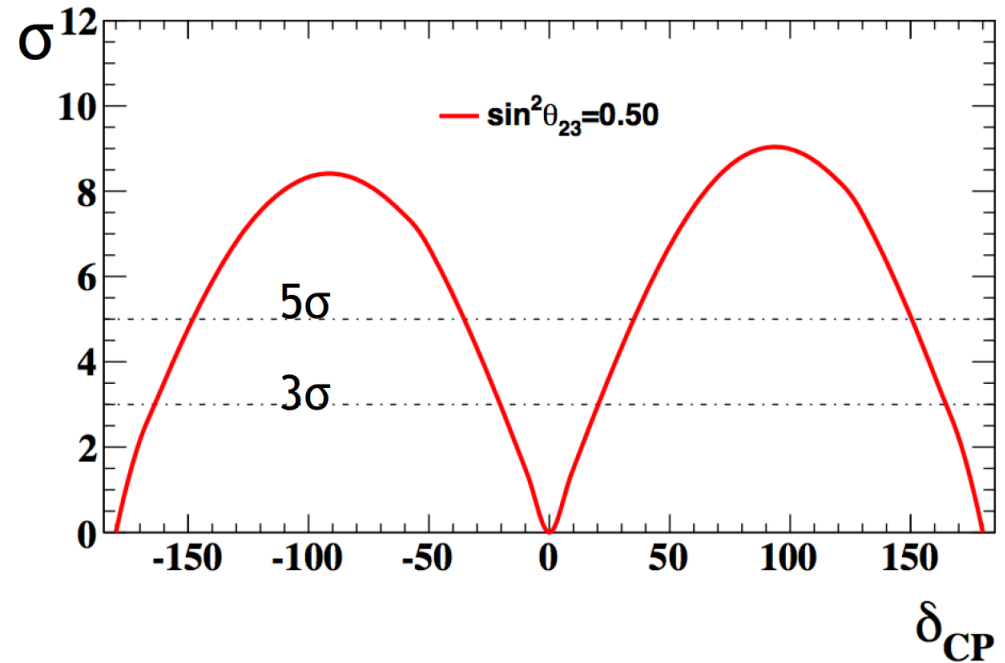
New Hyper-K design

High sensitivity for  $\delta_{CP}$

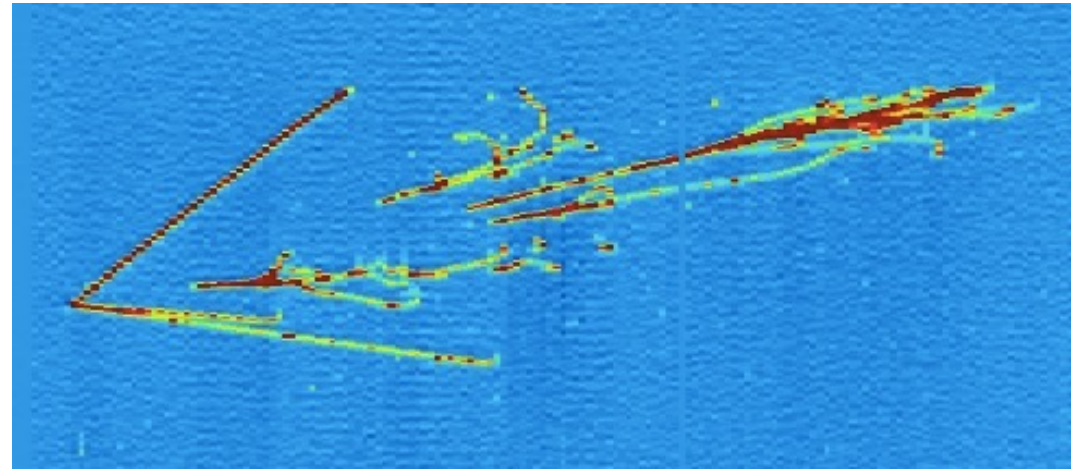
arXiv: 1502.05199



arXiv: 1611.06118



# DUNE



Very long baseline experiment aiming at CP violation using liquid argon TPCs as the detector technology.

Higher energy and baseline than Hyper-K.