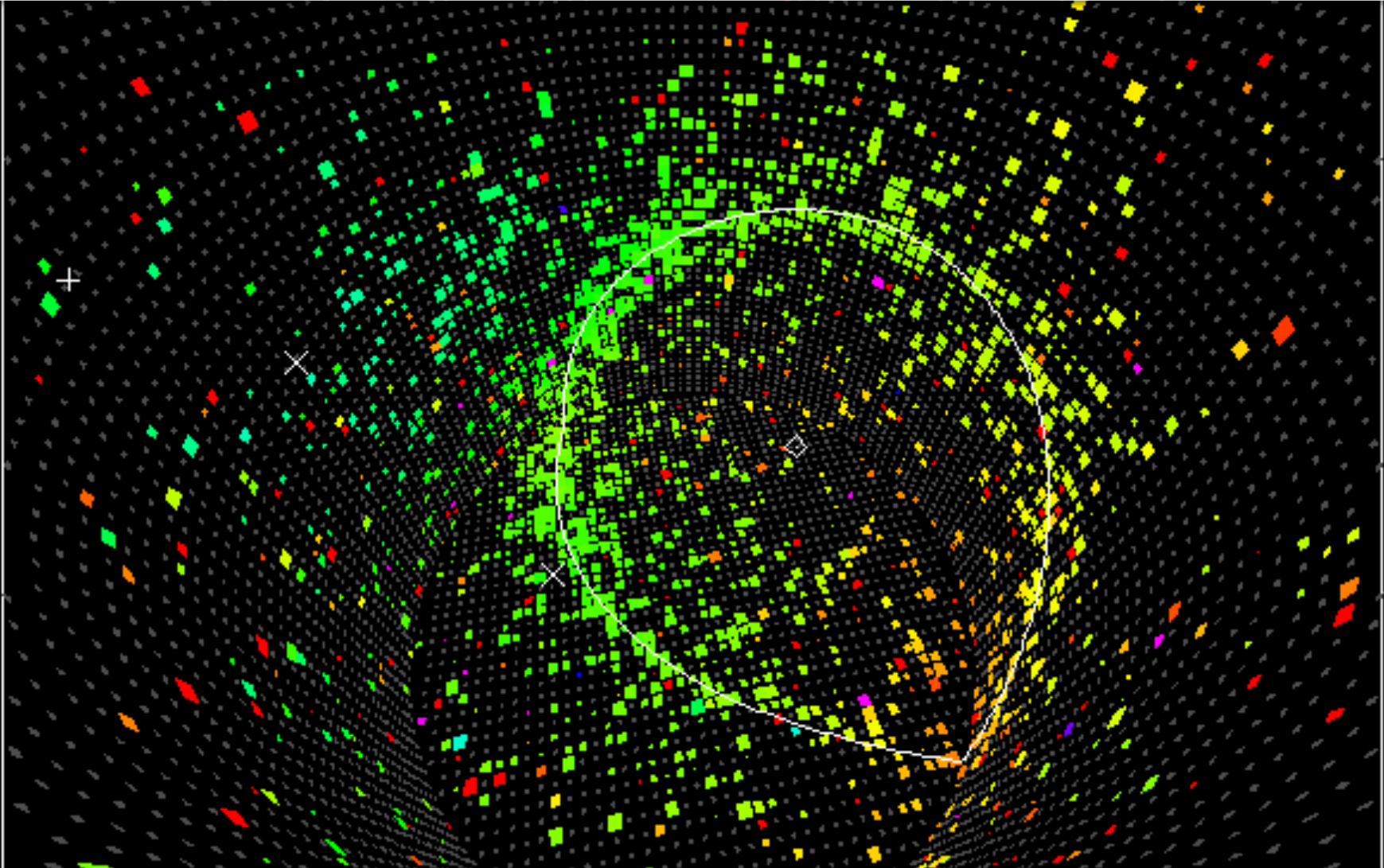
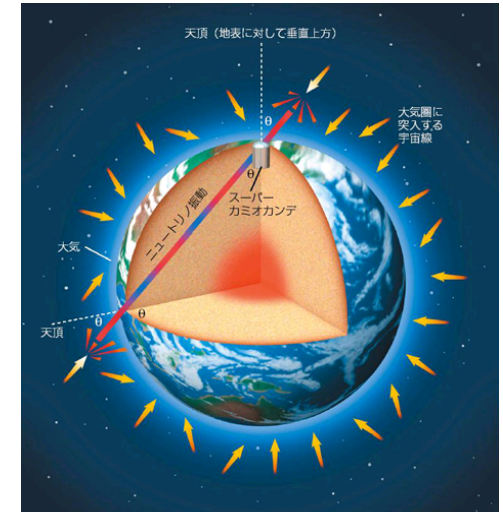
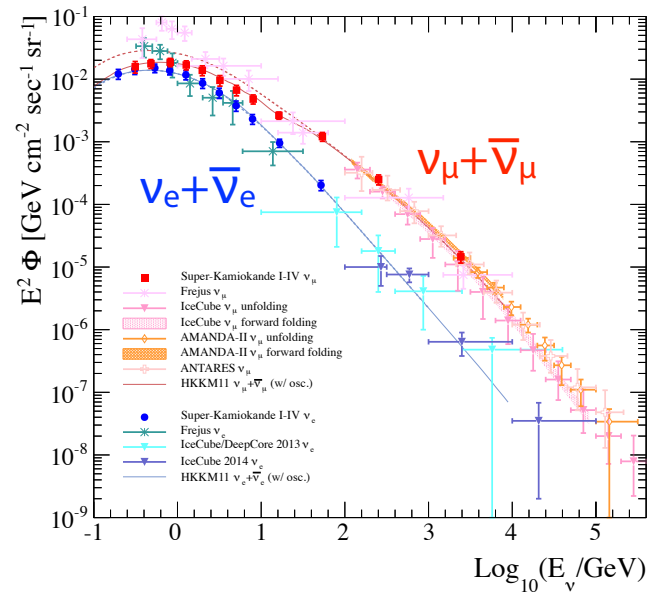
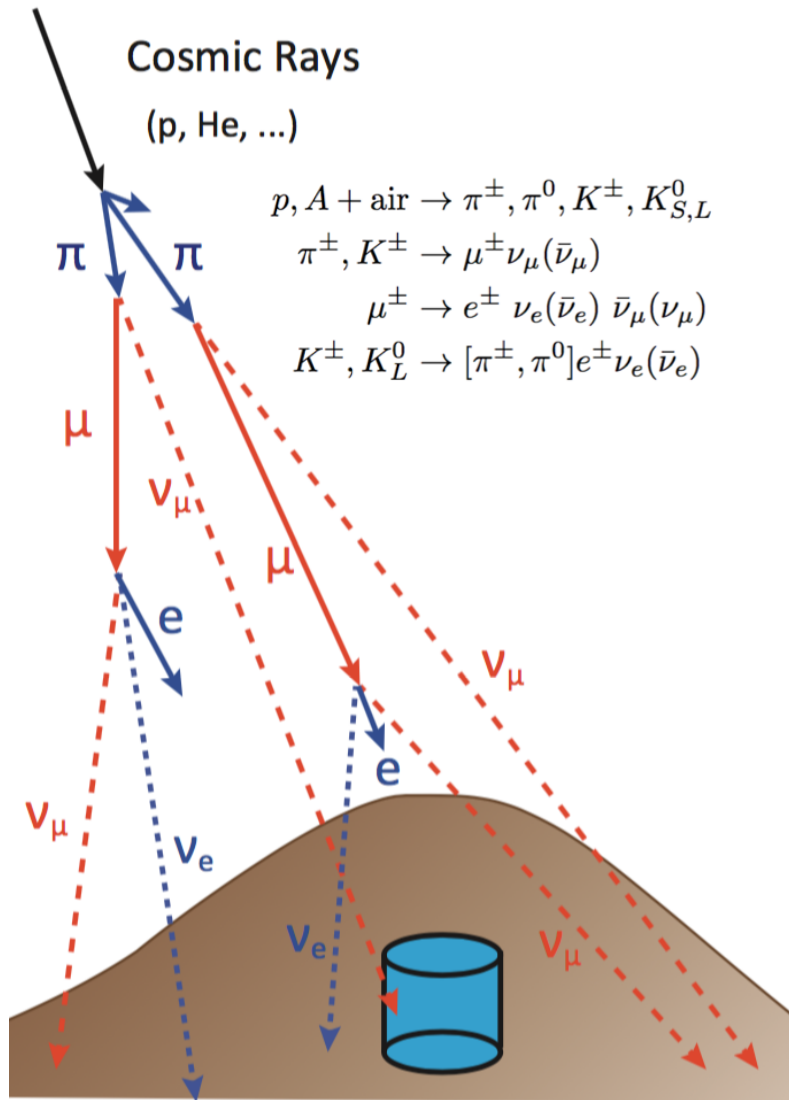


Atmospheric Neutrino and Proton Decay Results in Super-Kamiokande

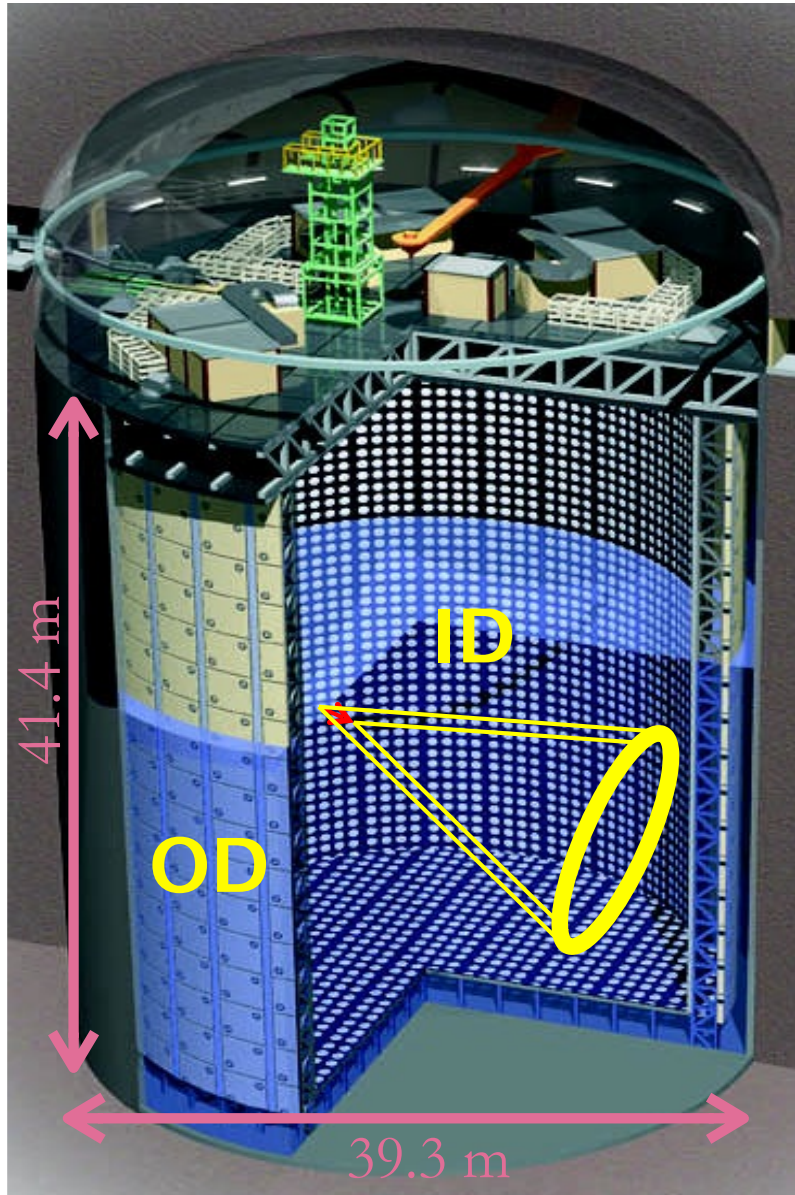


Atmospheric Neutrinos



- Decay products of secondaries by cosmic ray interactions with atmosphere. ($\nu_\mu : \nu_e \sim 2 : 1$ at GeV energy)
- **Power-law like energy spectrum.** Affected by cutoff due to geomagnetic field below several GeV.
- Path length: distributed in **O(10)km ~ 13,000km** depending on zenith angle direction
- Neutrino oscillation driven by Δm^2_{32} below O(10) GeV

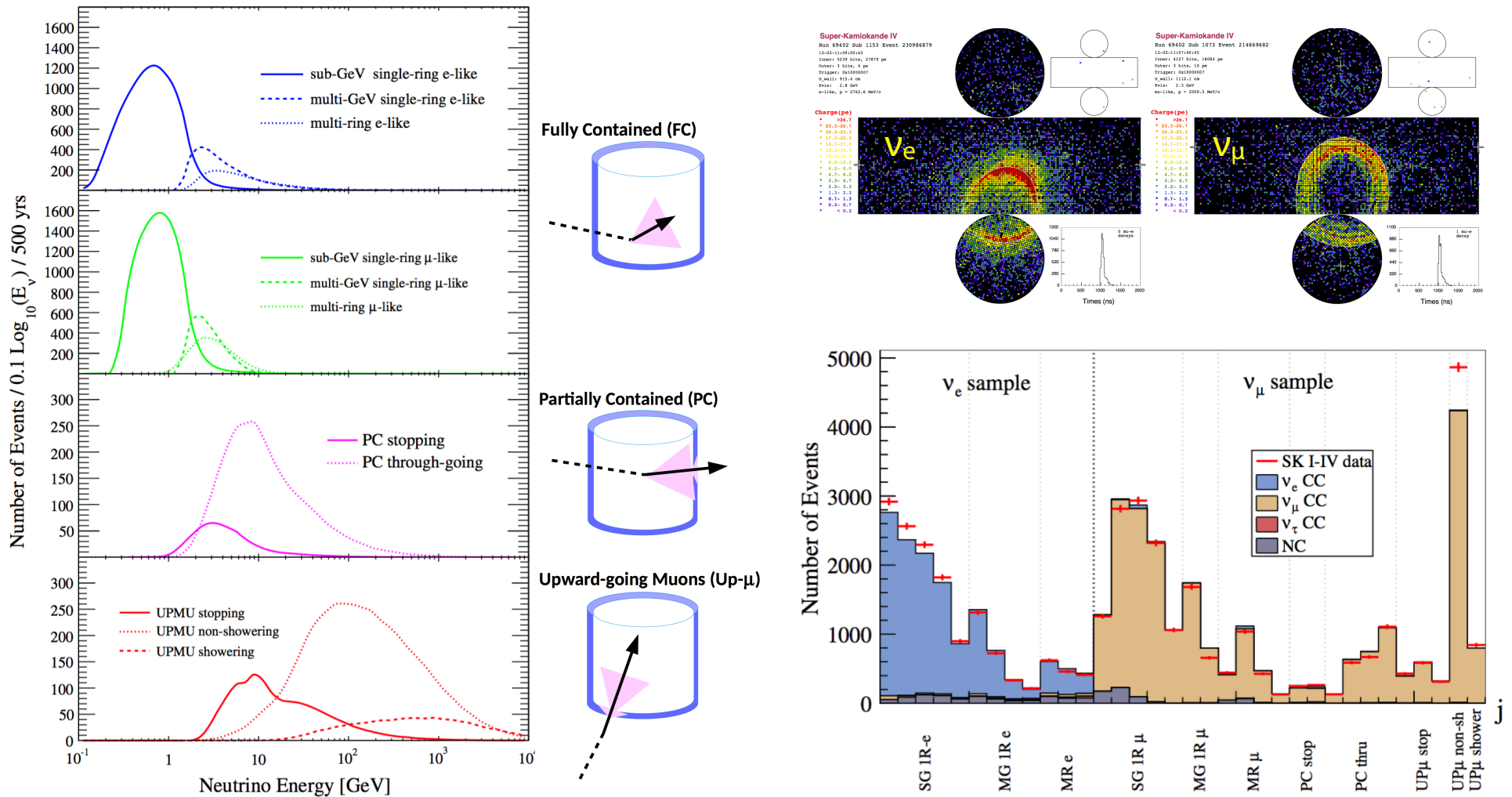
Super-Kamiokande Detector



- Water Cherenkov imaging detector
- **1000 m underground** in Kamioka mine
- **50 kton volume** (fiducial 22.5 kton)
- **11129 20" PMTs** in inner detector (ID) for Cherenkov ring imaging
- 1885 8" PMTs for outer detector (OD)

Phase	Period	# of PMTs
SK-I	1996.4 ~ 2001.7	11146 (40%)
SK-II	2002.10 ~ 2005.10	5182 (20%)
SK-III	2006.7 ~ 2008.8	11129 (40%)
SK-IV	2008.9 ~	

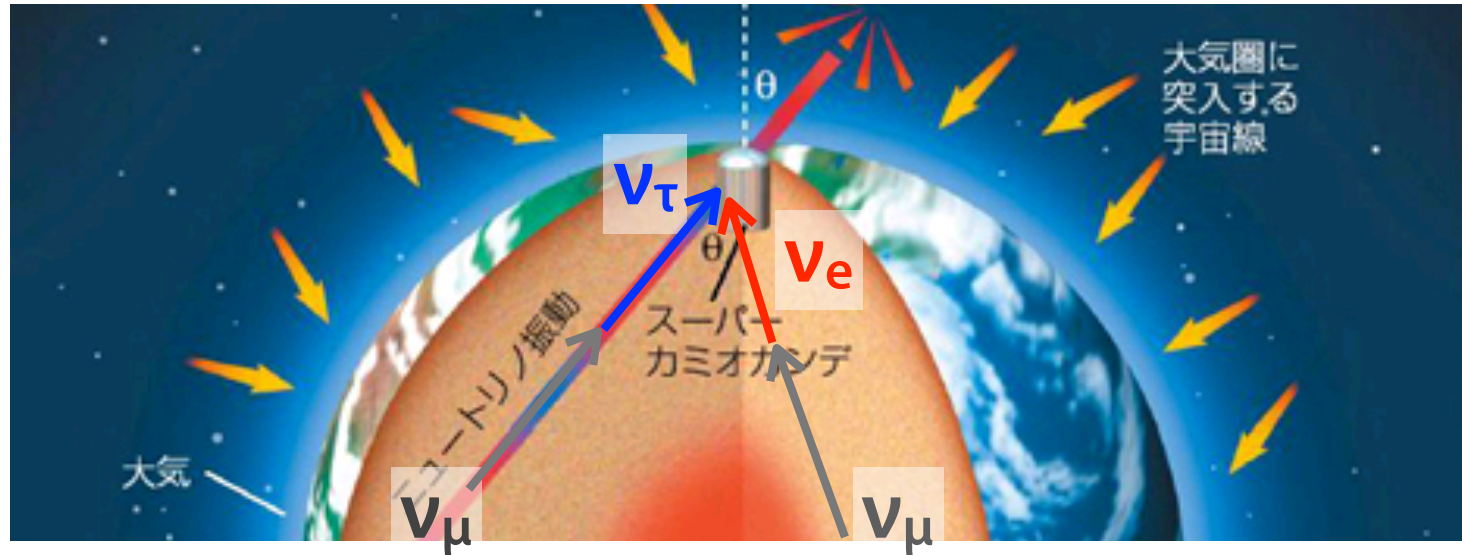
Super-K Atmospheric Event Sample



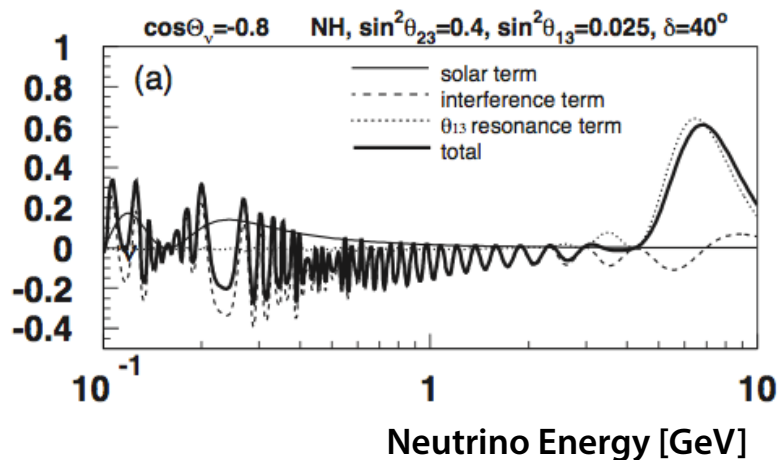
- Cover wide energy range **from 100 MeV up to 10 TeV** with three different event topologies

- High purity of ν_μ and ν_e CC sample with excellent particle ID performance

Atmospheric Oscillation Physics



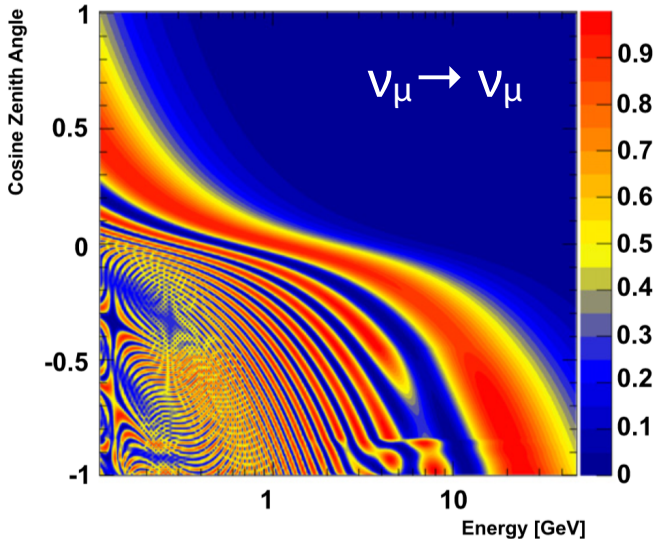
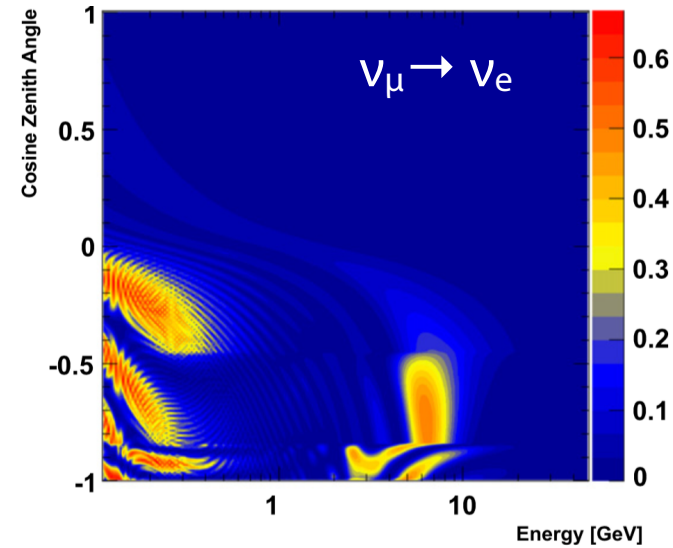
ν_e flux change due to sub-dominant oscillation:



- Many opportunities to test three flavor mixing:
 - ν_μ disappearance by $\nu_\mu \rightarrow \nu_\tau$ (Δm^2_{32} , θ_{23})
 - Sub-dominant oscillation in ν_e sample: **mass hierarchy (sign of Δm^2_{32}), δ_{CP} , θ_{23} octant**
 - **ν_τ appearance**
- Exotic mode (sterile, NSI, ..)

Matter Effect and Mass Hierarchy

Normal hierarchy ($\Delta m_{32}^2 > 0$)



- Neutrino is affected by additional potential due to forward scattering with electrons (**matter effect**)

$$i \frac{d\nu(t)}{dt} = H_0 \nu(t) \quad H_0 \rightarrow H_0 + \frac{1}{2E} \begin{pmatrix} A & 0 \\ 0 & 0 \end{pmatrix}$$

$$A = \pm 2\sqrt{2}G_F E_\nu n_e$$

- Effective mixing angle in matter:

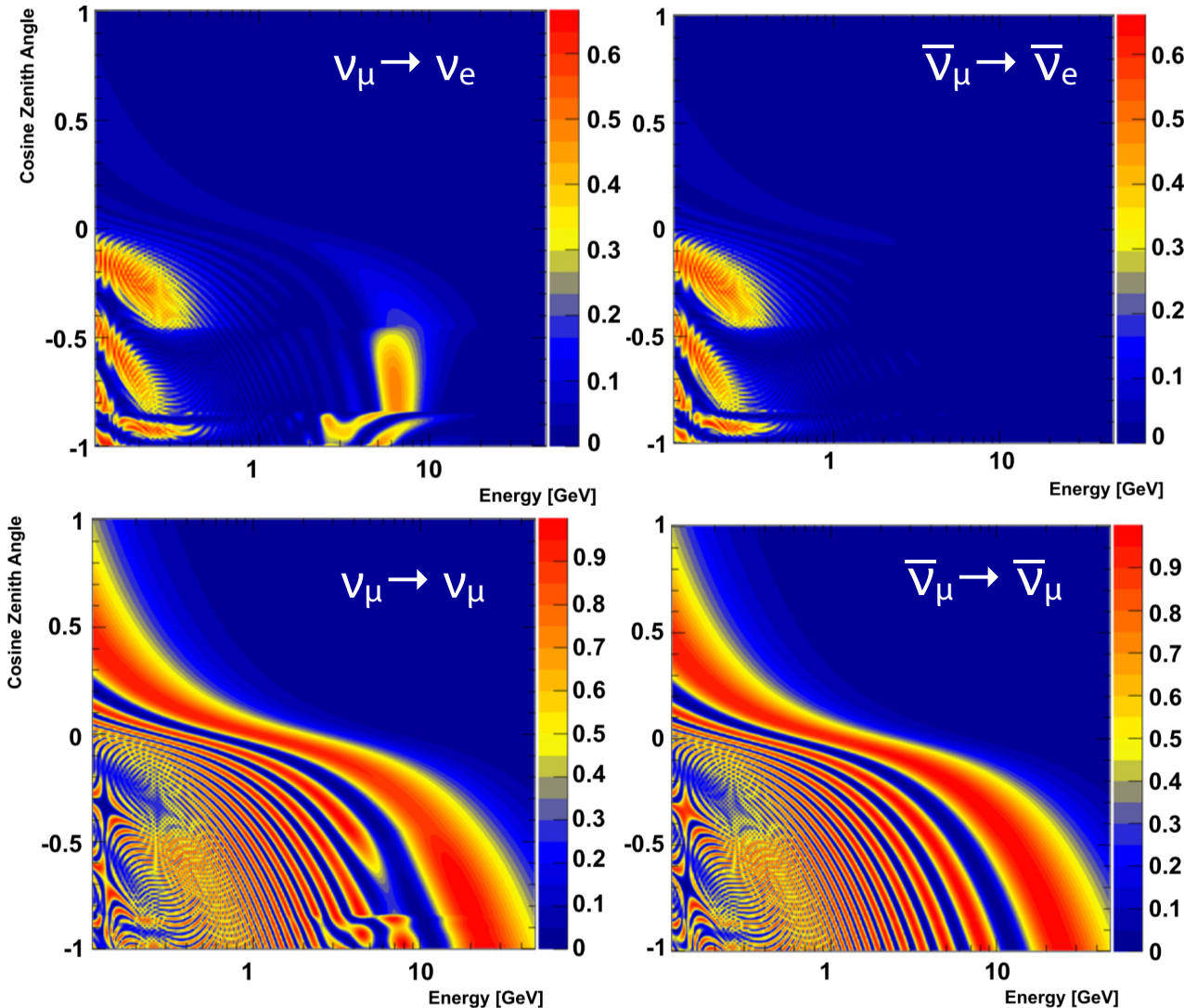
$$\sin 2\theta_{13}^M = \frac{\sin 2\theta_{13}}{\sqrt{\left(\frac{A}{\Delta m_{32}^2} - \cos 2\theta_{13}\right)^2 + \sin^2 2\theta_{13}}}$$

- At resonance region in multi-GeV:

$$A \sim \Delta m_{32}^2 \cos 2\theta_{13} \quad \rightarrow \quad \theta_{13}^M \gg \theta_{13}$$

Matter Effect and Mass Hierarchy

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- At resonance region in multi-GeV:

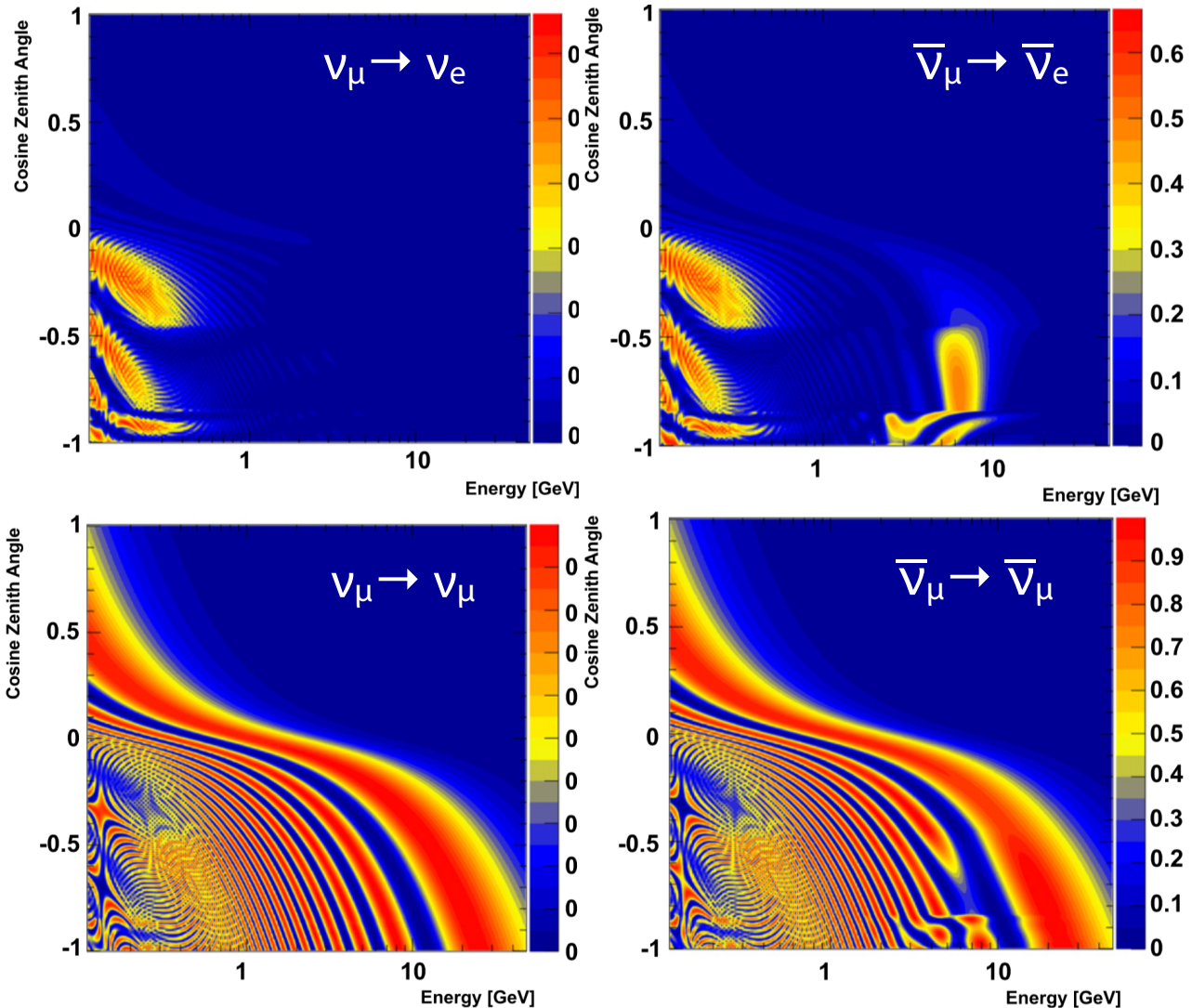
$$A \sim \Delta m_{32}^2 \cos 2\theta_{13} \quad \rightarrow \quad \theta_{13}^M \gg \theta_{13}$$

- Presence of resonance depends:

- $\nu / \bar{\nu}$ ($A \rightarrow -A$)

Matter Effect and Mass Hierarchy

Inverted hierarchy ($\Delta m_{32}^2 < 0$)



- Neutrino is affected by additional potential due to forward scattering with electrons (**matter effect**)

$$i \frac{d\nu(t)}{dt} = H_0 \nu(t) \quad H_0 \rightarrow H_0 + \frac{1}{2E} \begin{pmatrix} A & 0 \\ 0 & 0 \end{pmatrix}$$

$$A = \pm 2\sqrt{2}G_F E_\nu n_e$$

- Effective mixing angle in matter:

$$\sin 2\theta_{13}^M = \frac{\sin 2\theta_{13}}{\sqrt{\left(\frac{A}{\Delta m_{32}^2} - \cos 2\theta_{13}\right)^2 + \sin^2 2\theta_{13}}}$$

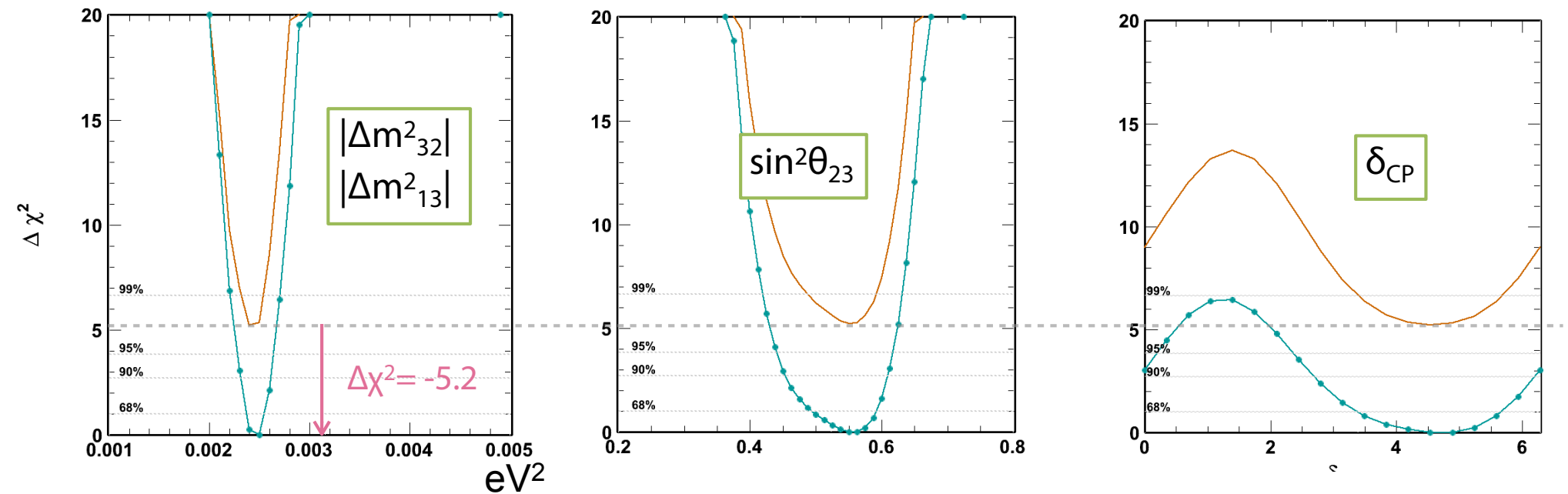
- At resonance region in multi-GeV:

$$A \sim \Delta m_{32}^2 \cos 2\theta_{13} \quad \rightarrow \quad \theta_{13}^M \gg \theta_{13}$$

- Presence of resonance depends:

- $\nu / \bar{\nu}$ ($A \rightarrow -A$)
- **Mass hierarchy** ($\Delta m_{32}^2 \rightarrow -\Delta m_{32}^2$)

Three Flavor Fit (w/ reactor and T2K constraints)



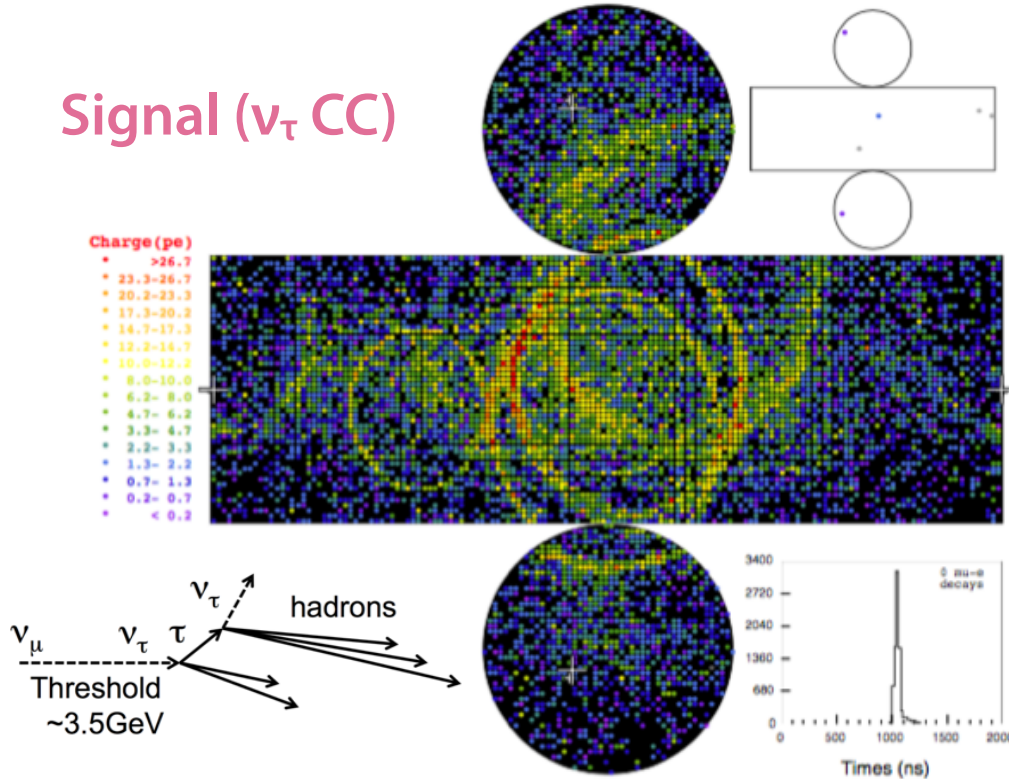
- Perform full parameter fit with additional constraints from reactor (θ_{13}) and T2K public data (Δm^2_{32} and θ_{23})
- Best-fit at NH, $\delta_{CP} \sim 3\pi/2$, $\sin^2\theta_{23} = 0.55$
- **Normal hierarchy is slightly preferred** ($\Delta\chi^2 = \chi^2_{NH} - \chi^2_{IH} = -5.2$)
- p-value of Inverted hypothesis is 0.024 and 0.001 for $\sin^2\theta_{23} = 0.6$ and 0.4, respectively

Best-fit parameters:

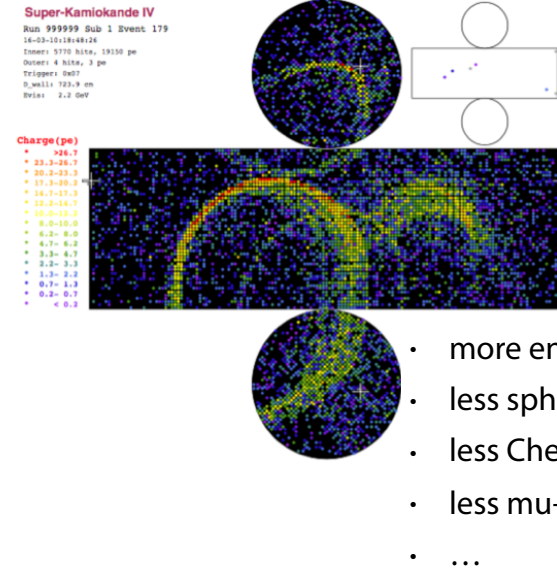
	δ_{CP}	$\sin^2\theta_{23}$	$ \Delta m^2_{32} $ (eV ²)
Inverted	4.538	0.55	2.5×10^{-3}
Normal	4.887	0.55	2.4×10^{-3}

Tau Appearance Analysis

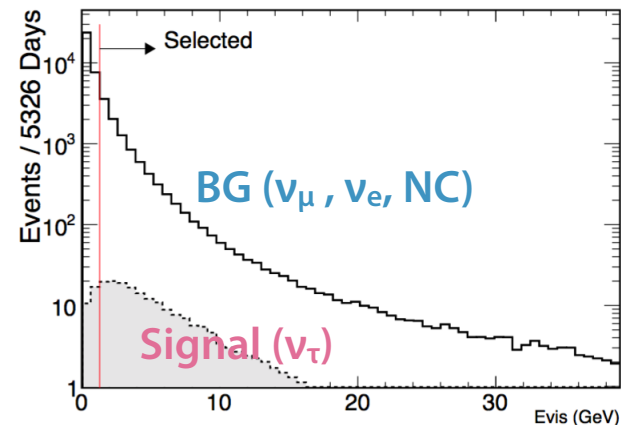
Signal (ν_τ CC)



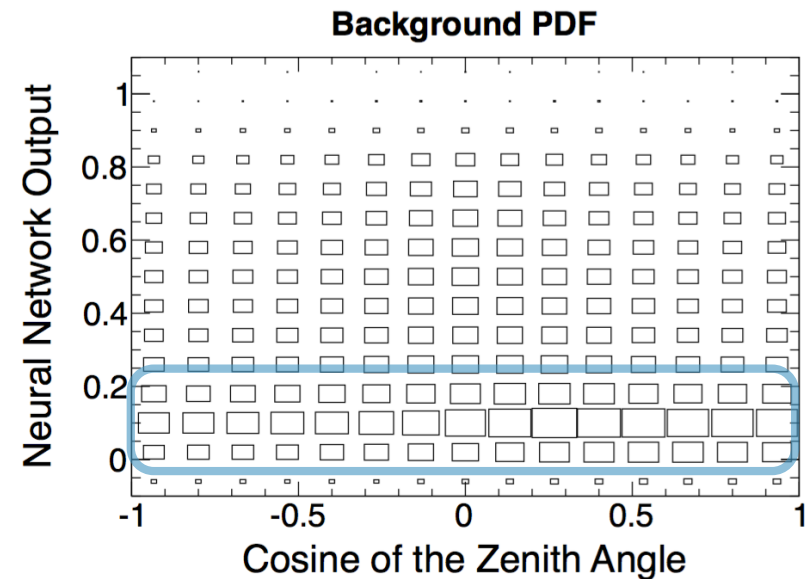
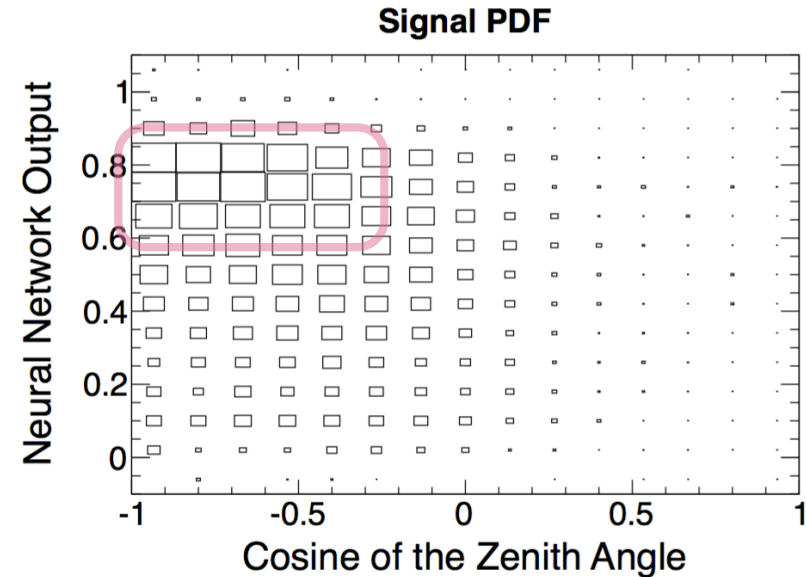
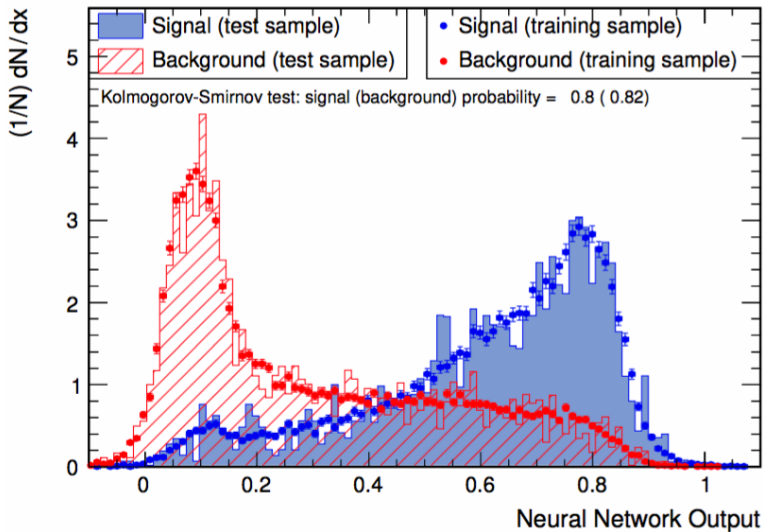
Background (ν_μ CC)



- Detection of tau appearance induced by $\nu_\mu \rightarrow \nu_\tau$ is critical for verifying three-flavor mixing scheme
- Detection is challenging: **low signal rate** (~ 1 event / kton year) with **huge backgrounds**
- Search for **hadronic modes of tau decay** (branching ratio: 65%)



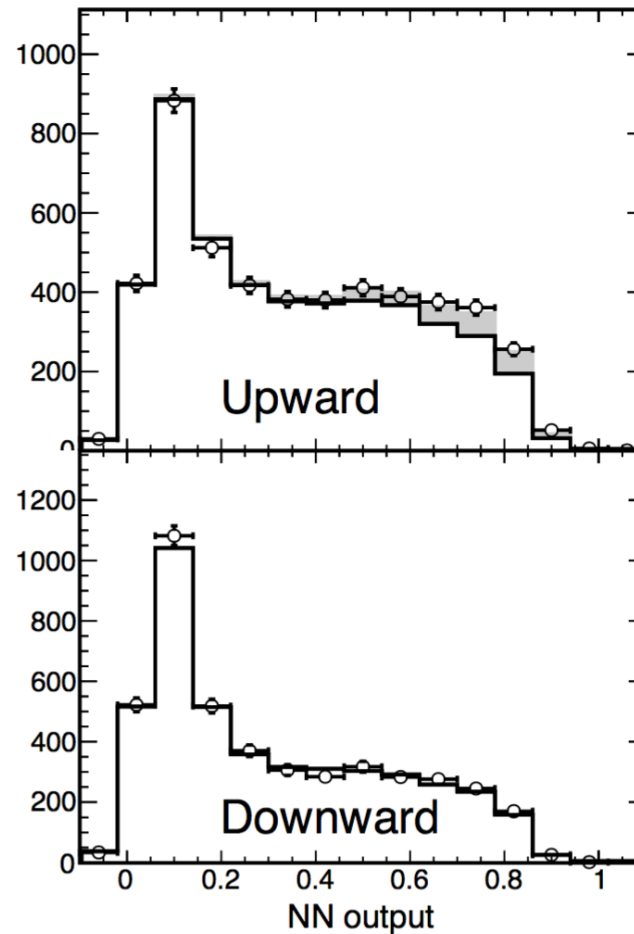
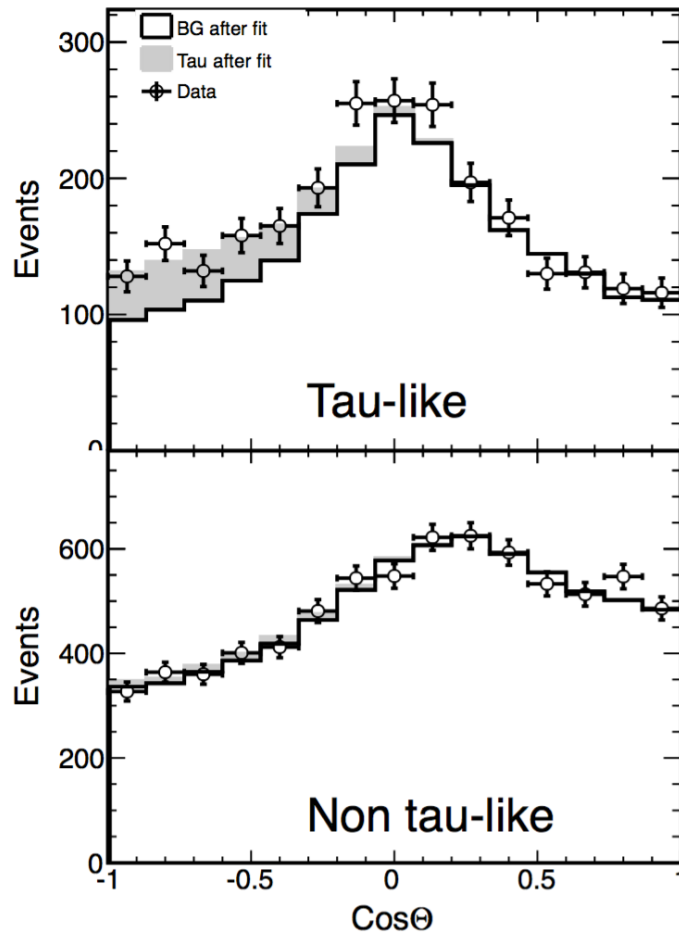
Tau Signal Discrimination



- Employ **neural network (NN)** technique to discriminate tau signal from background
 - Signal eff. 76%, 26% of background remains by NN>0.5 cut
- Tau events have higher NN output and enhanced in **upward direction**
- Perform 2-dim. fit with signal scale parameter:

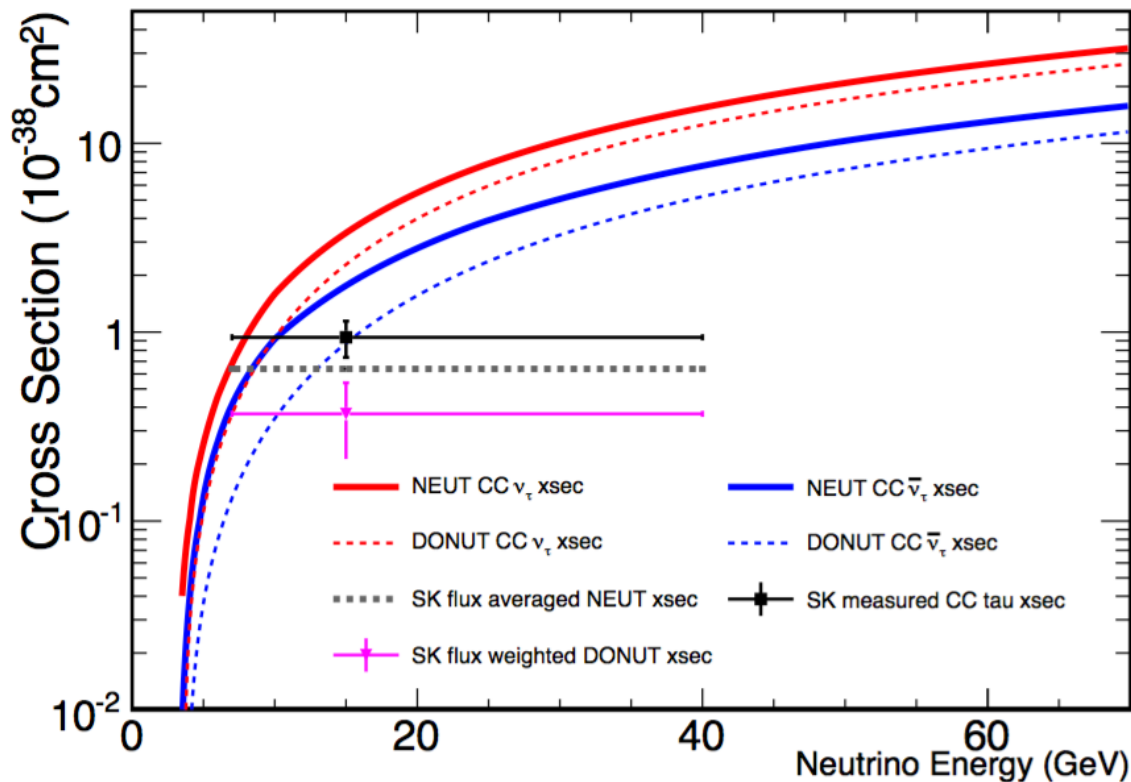
$$Data = PDF_{BG} + \alpha \times PDF_{\tau} + \sum \epsilon_i \times PDF_i$$

Tau Appearance Result



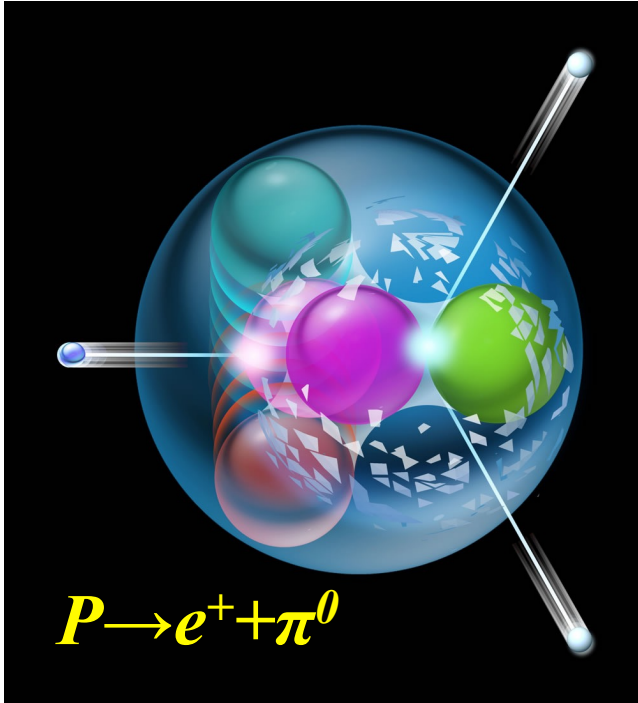
- Data: SK-I~IV 5,326 days
- Fitted tau normalization:
 $\alpha = 1.47 \pm 0.32$ (stat+syst)
- Observed events:
 338.1 ± 72.7 events
(exp'd: 224.5)
- Excluding no tau appearance hypothesis with 4.6σ
(exp'd 3.3σ)
- Still dominated by statistical uncertainty

Tau Neutrino Cross Section

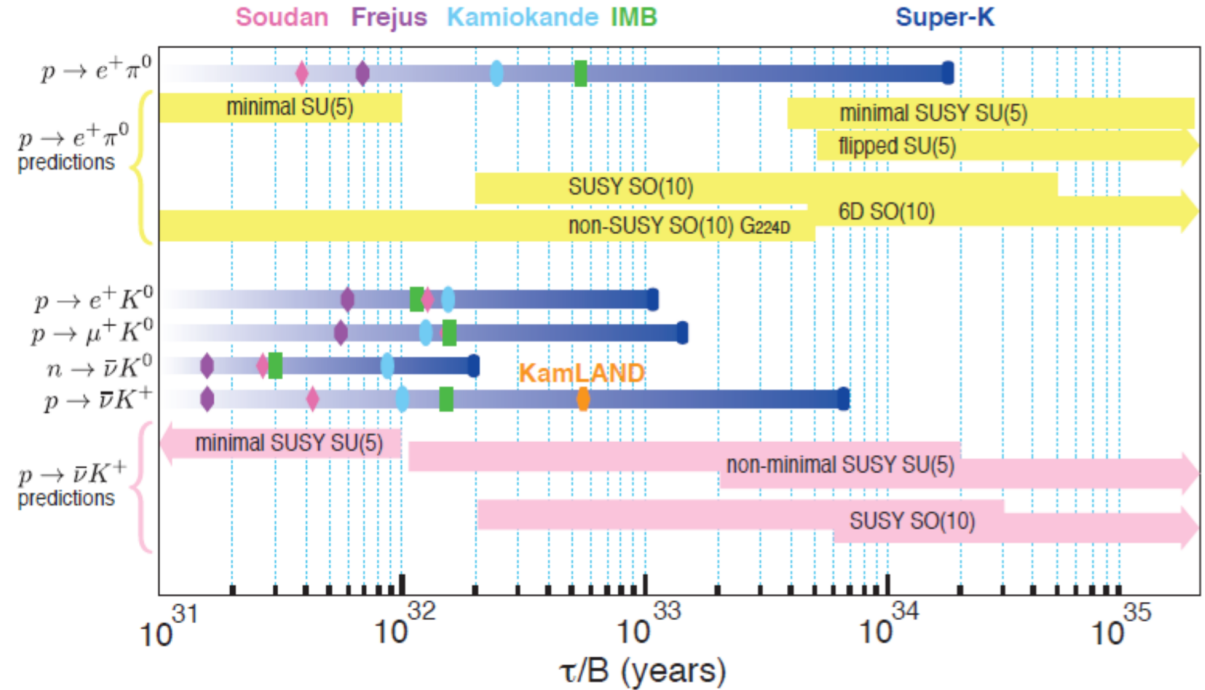


- Large sample of CC ν_τ sample offers the opportunity to **measure CC ν_τ cross section**
- Sensitive energy: 3.5 ~ 70 GeV
- Flux averaged cross section ($\times 10^{-38} \text{cm}^2$):
 - measured: 0.94 ± 0.20**
 - theory: 0.64**
- **Consistent with SM prediction** within 1.5 sigma
- Larger than scaled σ measured by DONUT at 111 GeV

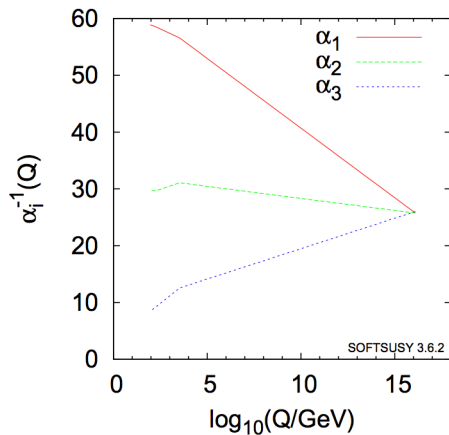
Proton Decay



Ed Kearns (Snowmass 2013)



MSSM: $m_0=M_{1/2}=2$ TeV, $A_0=0$, $\tan\beta=30$



- Proton decay is predicted by **GUTs (Grand Unified Theory)**
 - Provide the method for baryon asymmetry Universe
 - Open direct path to "Beyond SM" if detected
- **Many GUT predictions** — SU(5), SO(10), SUSY GUT
- Major decay modes: $P \rightarrow e^+ + \pi^0$, $P \rightarrow \bar{\nu} + K^+$

Proton Decay Measurement in Super-K

- World leading experiment in proton decay
large mass, high efficiency, various modes
- Categories:
 - decay to anti-lepton + meson
 - decay to $\bar{\nu} + K$ (Updated)**
 - others (di-nucleon decay, $n-\bar{n}$, ..)

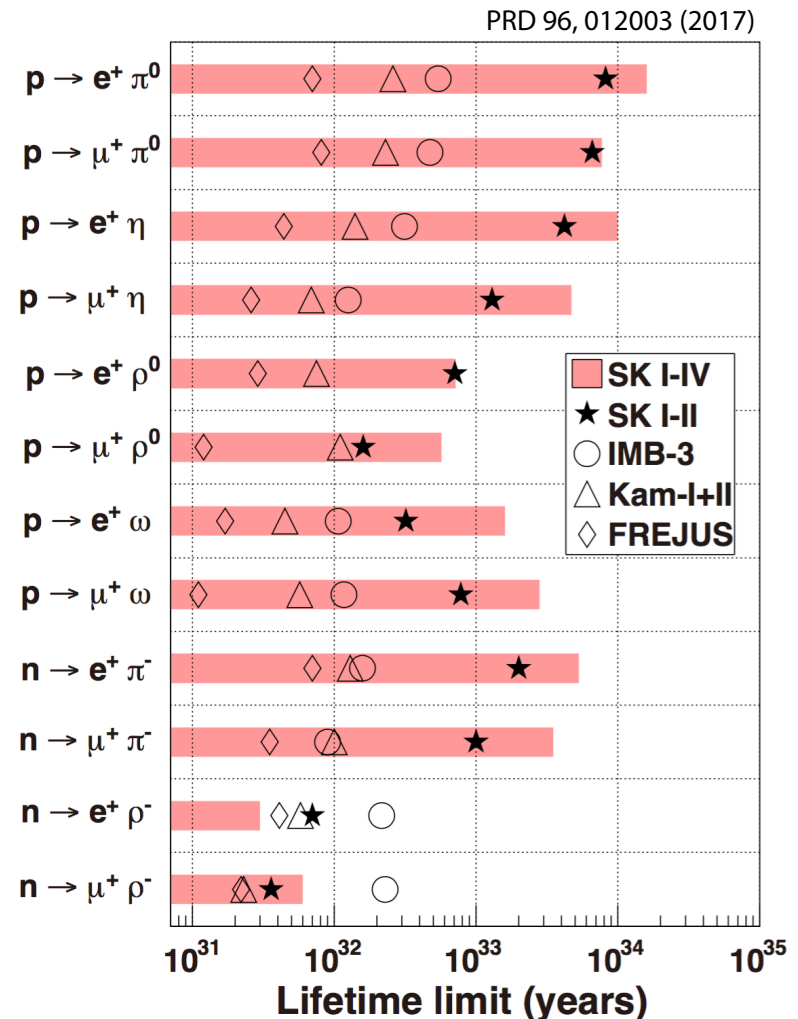
(S. Mine, NNN16)

Decay mode	$ \Delta(B-L) $	Lifetime lower limit at 90% CL (years)	Paper
$p \rightarrow e^+ \pi^0$	0	1.6×10^{34}	arXiv:1610.03597 (submitted to PRD) (*)
$p \rightarrow \nu K^+$	$0(\bar{\nu}), 2(\nu)$	6.6×10^{33}	PRD 90, 072005 (2014) ←
$p \rightarrow \mu^+ \pi^0$	0	7.7×10^{33}	arXiv:1610.03597 (submitted to PRD) (*)
$p \rightarrow (e^+, \mu^+)(\eta, \rho, \omega),$ $n \rightarrow (e^+, \mu^+)(\pi, \rho)$	0	$(0.03-10) \times 10^{33}$	will submit to PRD (**)
$p \rightarrow \mu^+ K^0$	0	1.6×10^{33}	PRD 86, 012006 (2012)
$\bar{n} \rightarrow \nu \pi^0, \bar{p} \rightarrow \nu \pi^+$	0	$1.1 \times 10^{33}, 3.9 \times 10^{32}$	PRL 113, 121802 (2014)
$p \rightarrow (e^+, \mu^+) \nu \nu$	$0(\bar{\nu}\nu),$ $2(\nu\nu, \bar{\nu}\bar{\nu})$	$1.7/2.2 \times 10^{32}$	PRL 113, 101801 (2014)
$p \rightarrow (e^+, \mu^+) X$?	$7.9/4.1 \times 10^{32}$	PRL 115, 121803 (2015)
$n \rightarrow \nu \gamma$	$0(\bar{\nu}), 2(\nu)$	5.5×10^{32}	PRL 115, 121803 (2015)
$pp \rightarrow K^+ K^+$	2	1.7×10^{32}	PRL 112, 131803 (2014)
$pp \rightarrow \pi^+ \pi^+, pn \rightarrow \pi^+ \pi^0,$ $nn \rightarrow \pi^0 \pi^0$	2	$7.2 \times 10^{31}, 1.7 \times 10^{32},$ 4.0×10^{32}	PRD 91, 072009 (2015)
$np \rightarrow (e^+, \mu^+, \tau^+) \nu$	$0(\bar{\nu}), 2(\nu)$	$(0.22-5.5) \times 10^{32}$	PRL 115, 121803 (2015)
$n-\bar{n}$ oscillation	2	1.9×10^{32}	PRD 91, 072006 (2015)

(*) published in PRD 95, 012004 (2017)

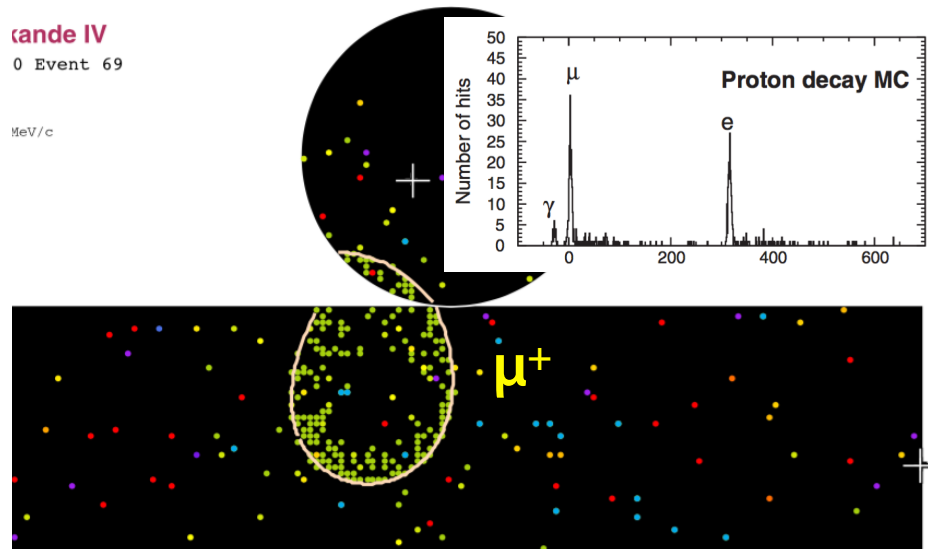
(**) published in PRD 96, 012003 (2017)

Limits on decay modes of anti-lepton + meson

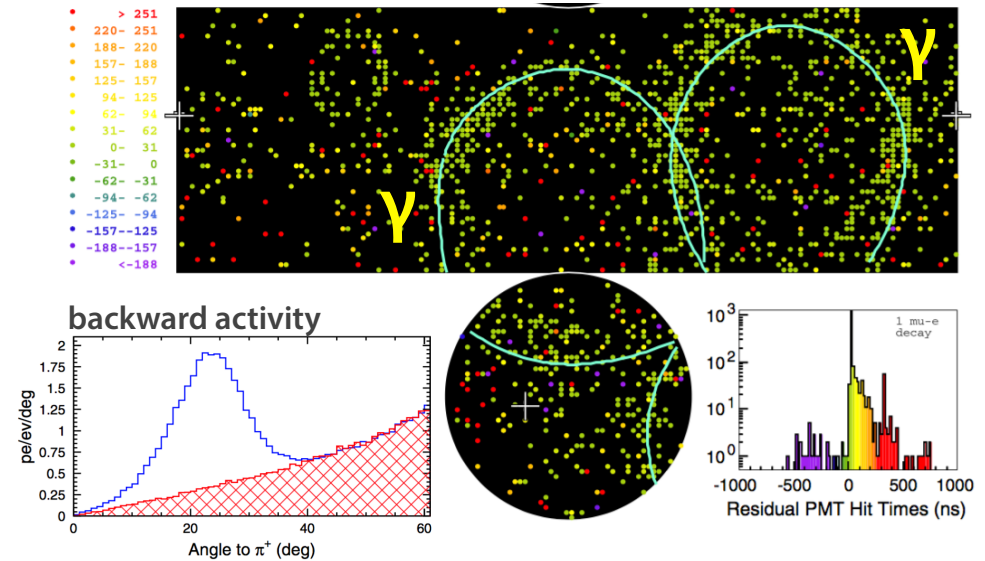


Search for $P \rightarrow \bar{\nu} + K^+$ Decay

(A) $K^+ \rightarrow \nu_\mu + \mu^+$ (BR: 64%)



(B) $K^+ \rightarrow \pi^0 + \pi^+$ (BR: 21%)

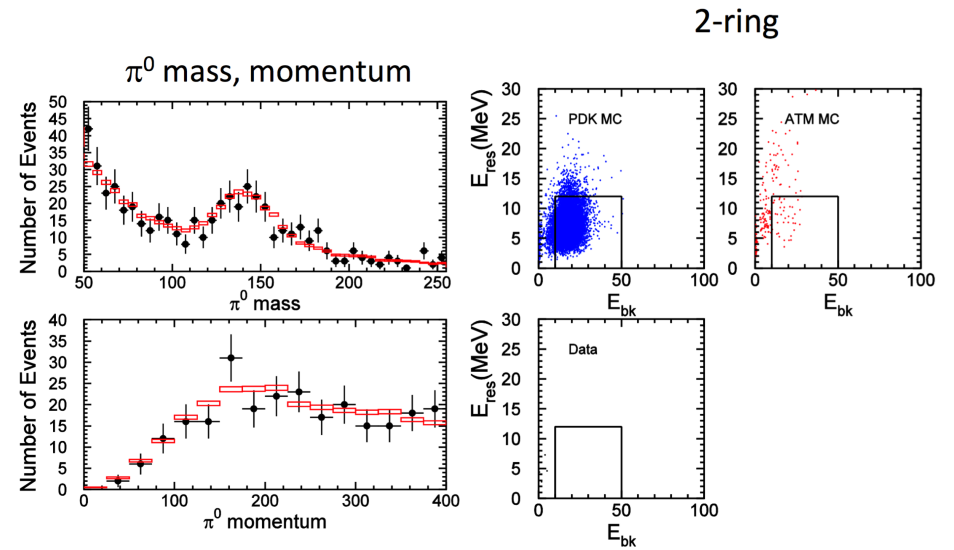
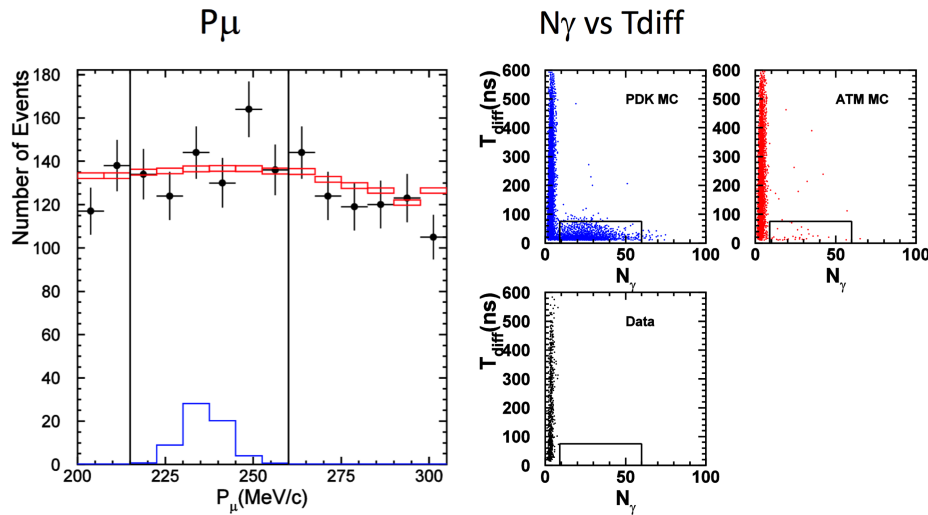


- **Single mono-energetic muon** ($P_\mu=236\text{MeV}/c$) from K^+ decay with following μ -e decay
- Require **prompt 6 MeV gamma** from excited oxygen nuclei
- Search for $\pi^0 \rightarrow 2\gamma$ decay ($P_{\pi^0}=205\text{MeV}/c$) event with **faint π^+ activity** in backward direction

Search for $P \rightarrow \bar{\nu} + K^+$ Decay

(A) $K^+ \rightarrow \nu_\mu + \mu^+$

(B) $K^+ \rightarrow \pi^0 + \pi^+$



	SK1			SK2			SK3			SK4		
	Eff (%)	BG (ev)	Obs (ev)	Eff (%)	BG (ev)	Obs (ev)	Eff (%)	BG (ev)	Obs (ev)	Eff (%)	BG (ev)	Obs (ev)
Pr. γ	7.9 ± 0.1	0.078	0	6.5 ± 0.1	0.082	0	7.5 ± 0.1	0.018	0	9.4 ± 0.1	0.112	0
$\pi^+\pi^0$	7.8 ± 0.1	0.21	0	6.5 ± 0.1	0.19	0	8.3 ± 0.1	0.07	0	9.6 ± 0.1	0.13	0

- No candidate events are observed for both modes in 349 kton-year exposure
- Lifetime limit: $>8.0 \times 10^{33}$ years (90% C.L.)

Summary

Atmospheric Neutrino:

- Various features of atmospheric neutrinos allow us to test **three flavor mixing scheme**.
- **Resonance oscillation by matter effect** in multi-GeV is sensitive to mass hierarchy. According to oscillation fit to data with reactor and T2K constraints, **normal hierarchy is slightly preferred**.
- Tau appearance has **improved to 4.6 sigma**. Measured larger ν_τ cross section than prediction though still **consistent with theory**.

Proton Decay:

- Unique method to probe GUT theory. Super-K has been contributing to explore the possibility of this new “beyond SM” physics.
- The analysis of **$P \rightarrow \bar{\nu} + K^+$ has been updated**. There were no candidate events observed from the searches of two K^+ decay modes. Lower lifetime limit has been **improved to 8.0×10^{33} yrs**.

END