

# Current and Future of DarkSide

Dark Matter Experiment

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on behalf of the **DarkSide Collaboration**

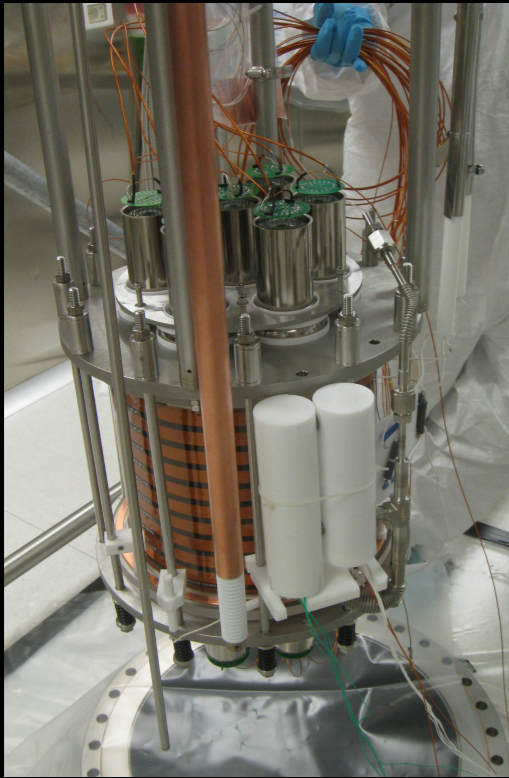
25 Aug. 2015

# DarkSide Program

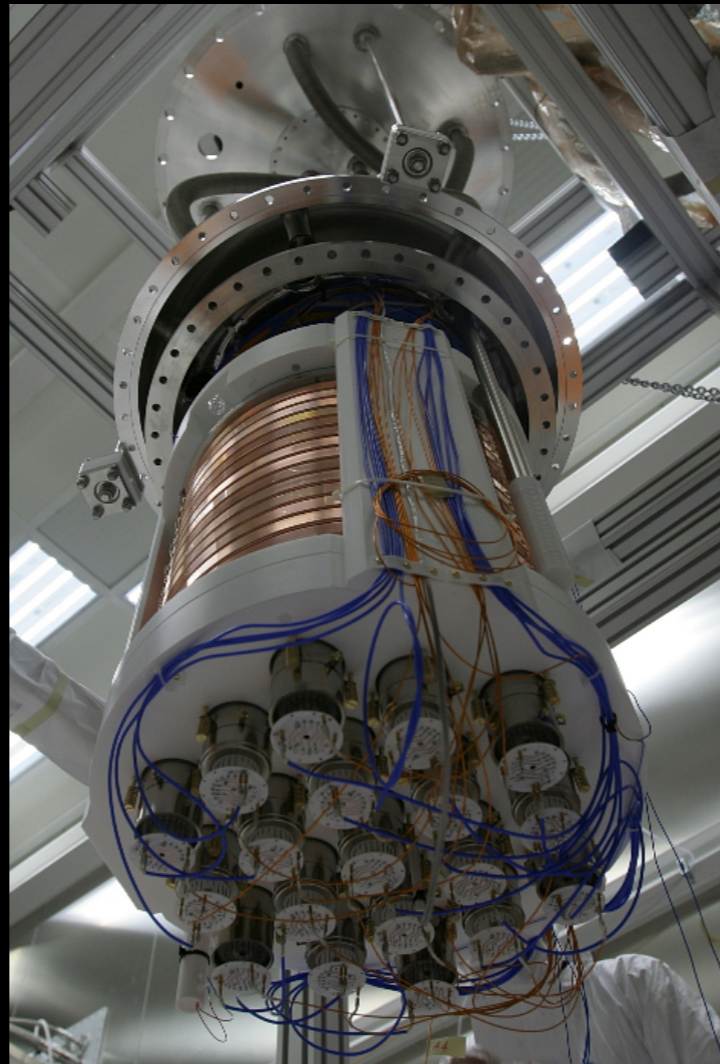
- **Direct detection** search for WIMP dark matter
- Based on a **two-phase argon** time projection chamber (**TPC**)
- Design philosophy based on having very low background levels that can be further reduced through **active suppression**, for **background-free** operation from backgrounds (both from neutrons and  $\beta/\gamma$ 's)

# DarkSide Program

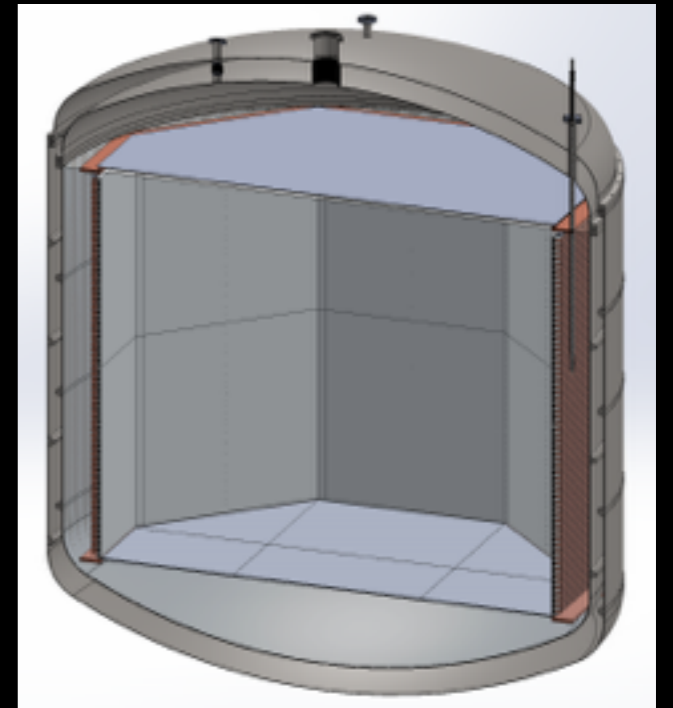
Multi-stage program at Gran Sasso National Laboratory in Italy



DarkSide 10  
Prototype detector



DarkSide 50  
First physics detector  
Commissioned Oct.2013



DarkSide-20k  
30 tonne (20 tonne  
fiducial depleted argon  
detector proposed to  
LNGS for operations  
within **2020**)

+ multiple smaller test setups and prototypes

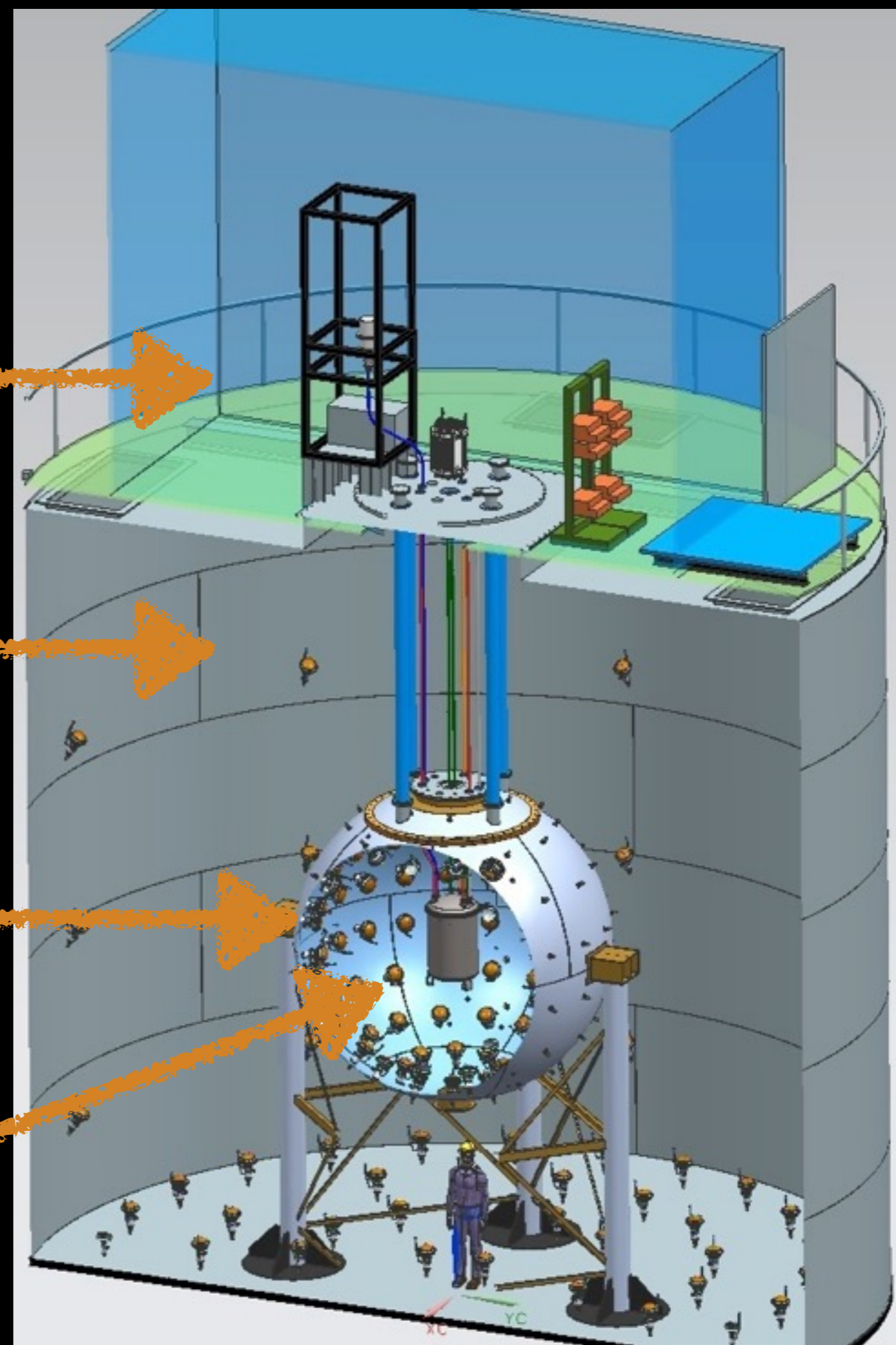
# DarkSide 50

Radon-free Assembly  
**Clean Room**

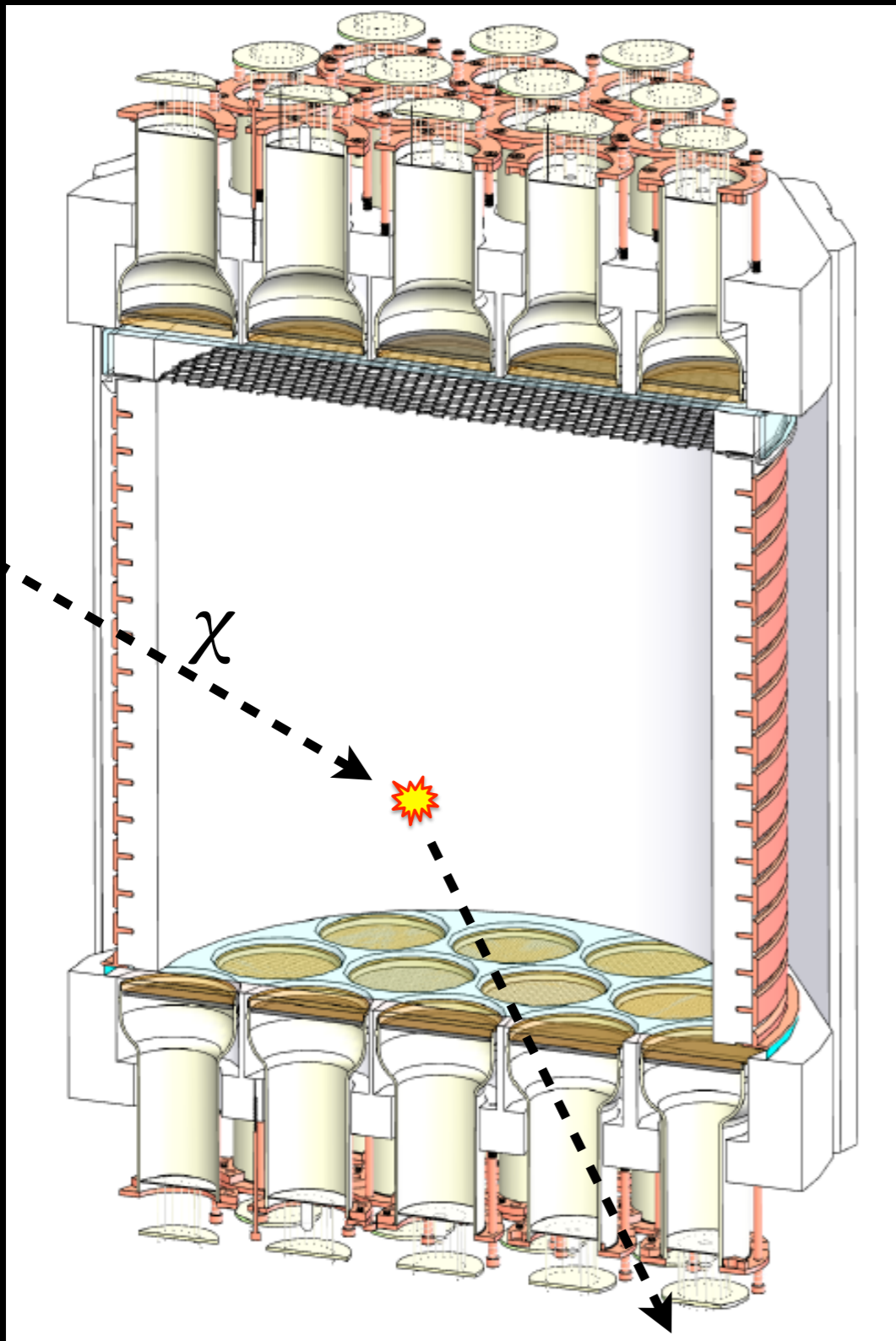
1,000-tonne Water-based  
Cherenkov **Cosmic Ray Veto**

30-tonne Liquid Scintillator  
**Neutron and  $\gamma$ 's Veto**

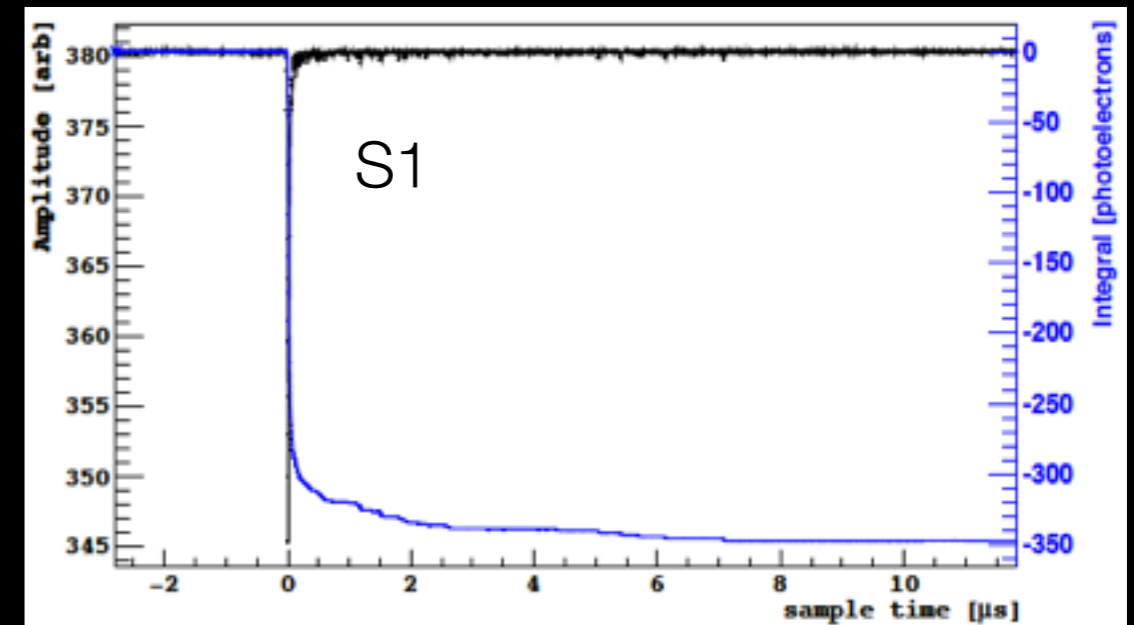
Inner detector **TPC**



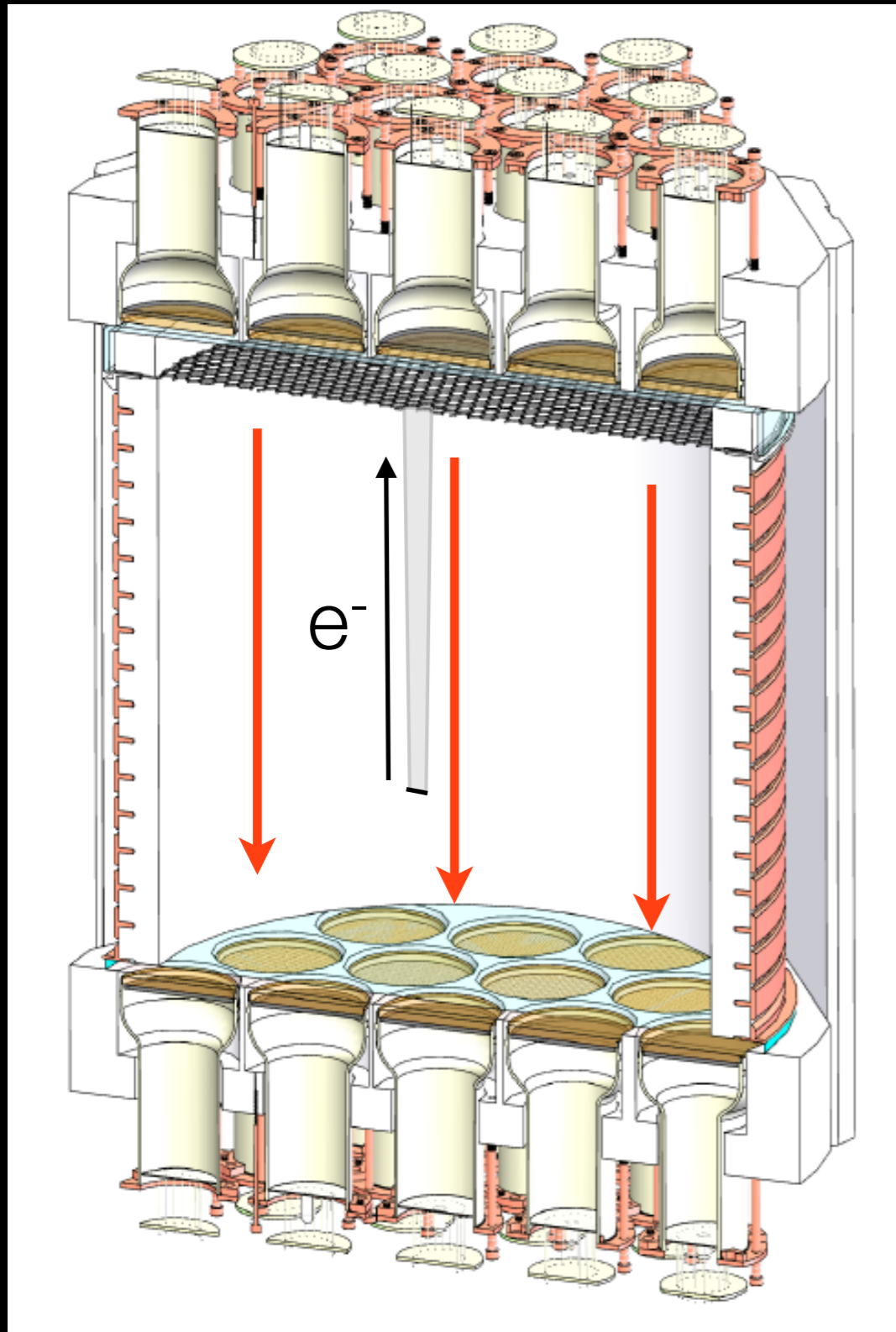
# Two Phase Argon TPC



Nuclear Recoil excites and ionizes the liquid argon, producing scintillation light (S1) that is detected by the photomultipliers



# Detecting WIMPs

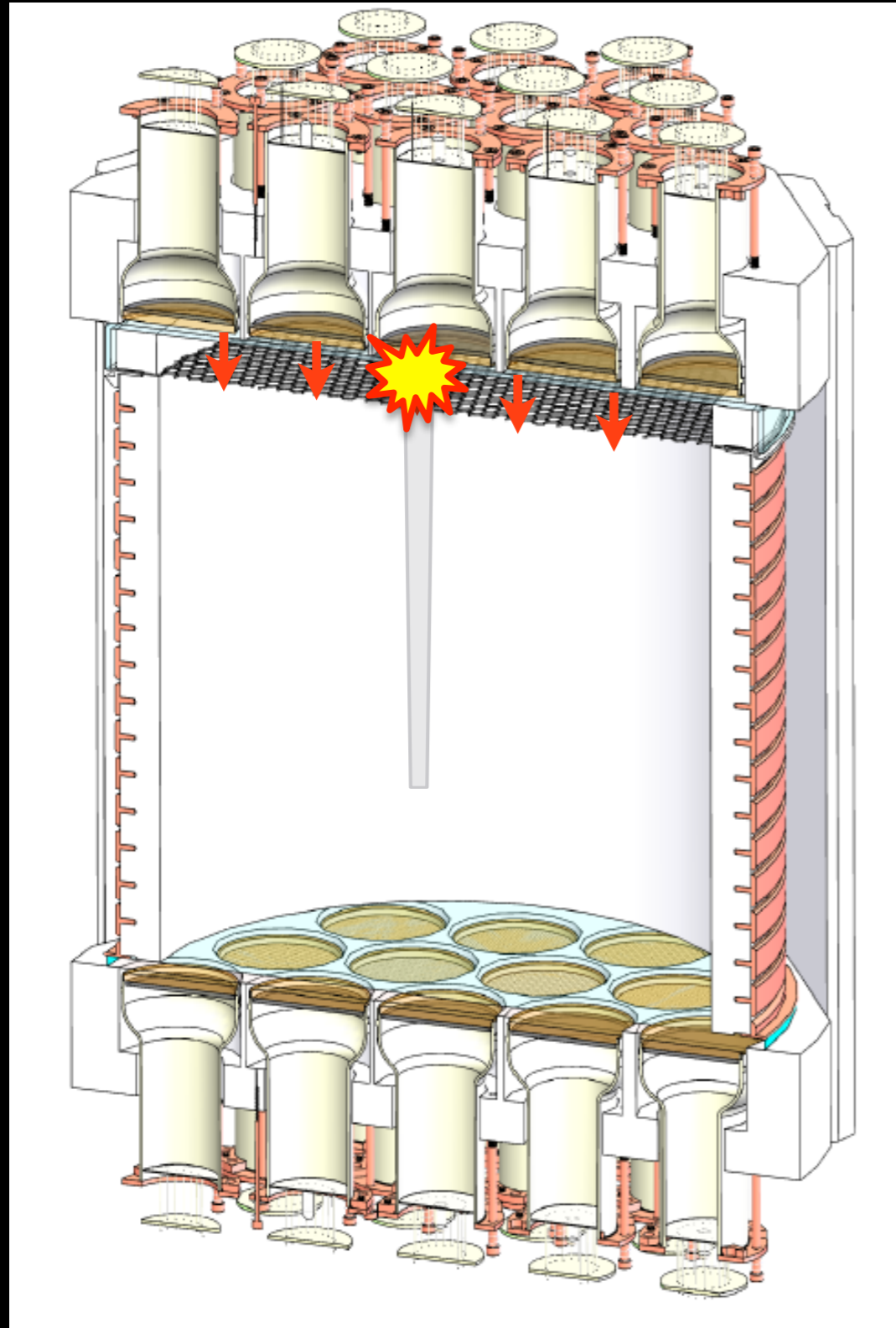


The ionized electrons that survive recombination are drifted towards the liquid-gas interface by the electric field.

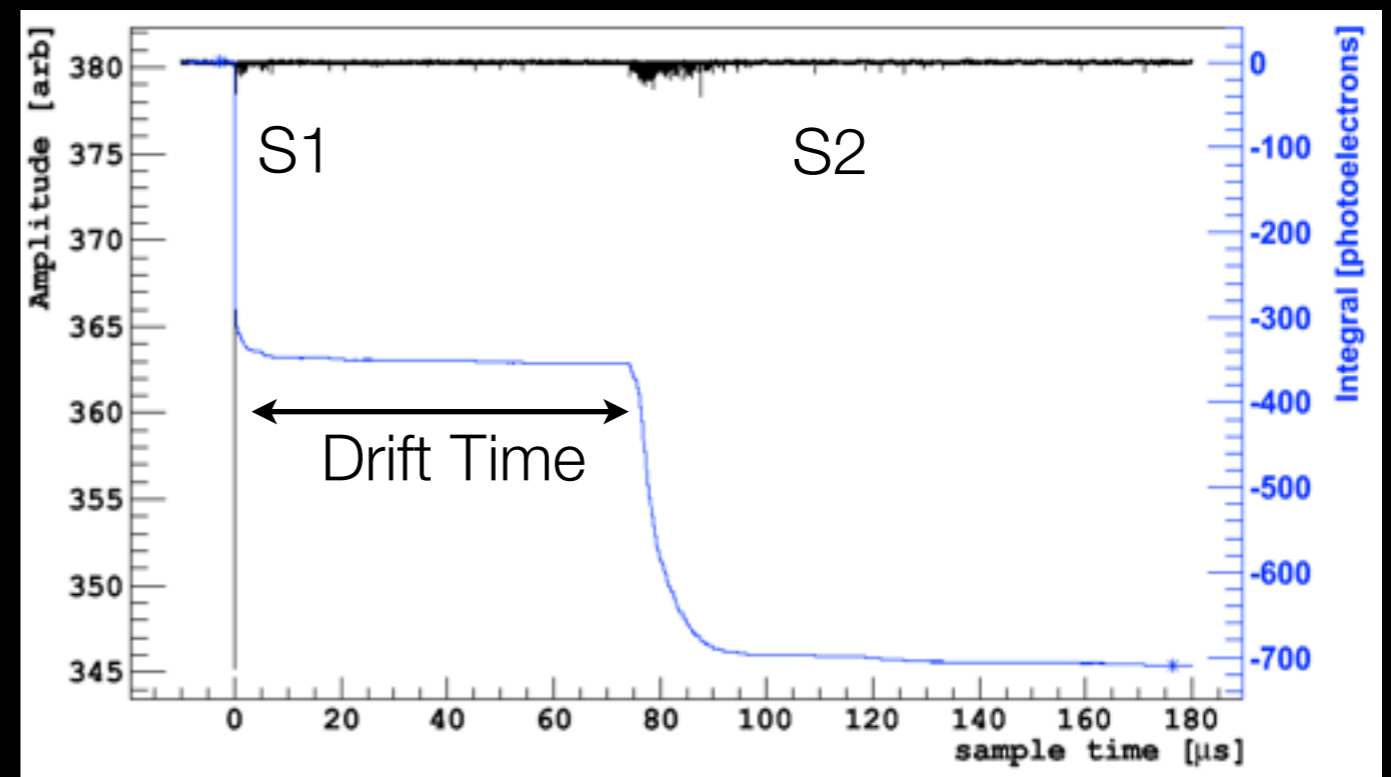
Electron drift lifetime  $> 5$  ms,  
compared to max. drift time of  $\sim 375 \mu\text{s}$

Electron drift speed =  $0.93 \pm 0.01$  mm/ $\mu\text{s}$

# Detecting WIMPs



The electrons are extracted into the gas region, where they induce electroluminescence (S2)



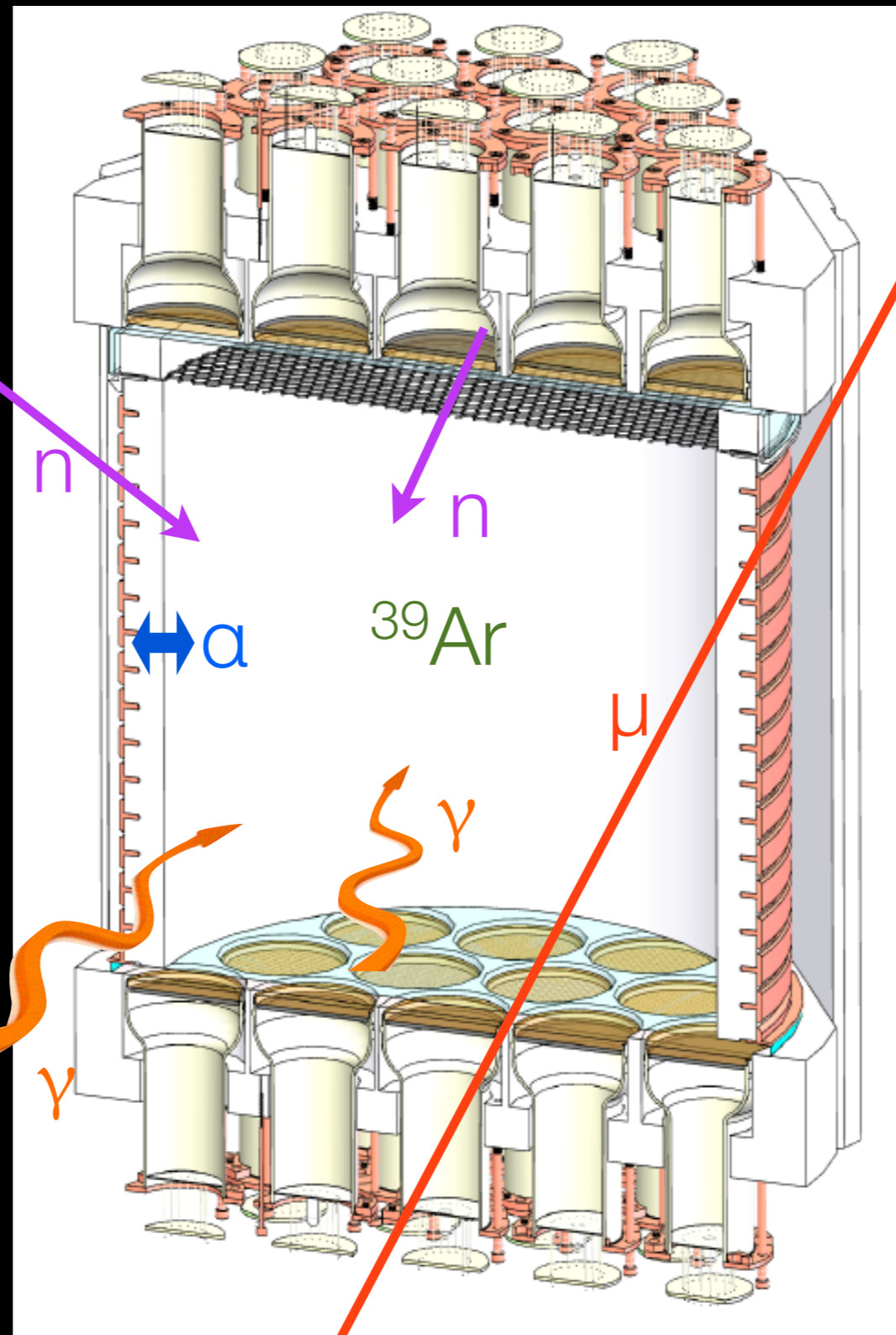
The time between the S1 and S2 signals gives the vertical position

# Backgrounds

[30-200] keVr

## ELECTRON RECOILS

## NUCLEAR RECOILS



$^{39}\text{Ar}$   
 $\sim 9 \times 10^4$  evt/kg/day

$\mu$   
 $\sim 30$  evt/m<sup>2</sup>/day

Radiogenic  $n$   
 $\sim 6 \times 10^{-4}$  evt/kg/day

$\gamma$   
 $\sim 1 \times 10^2$  evt/kg/day

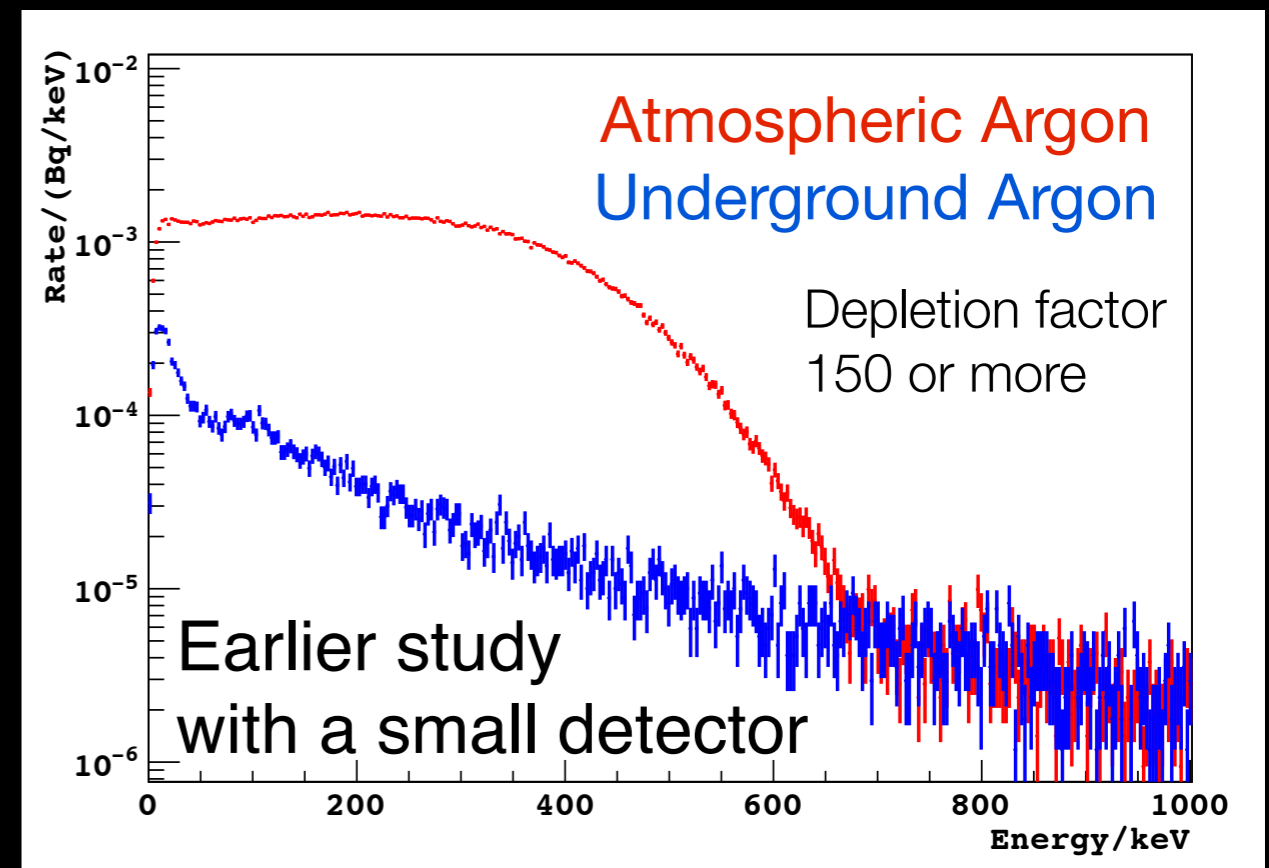
$\alpha$   
 $\sim 10$  evt/m<sup>2</sup>/day

100 GeV,  $10^{-45}$  cm<sup>2</sup> WIMP Rate  $\sim 10^{-4}$  evt/kg/day



# $^{39}\text{Ar}$

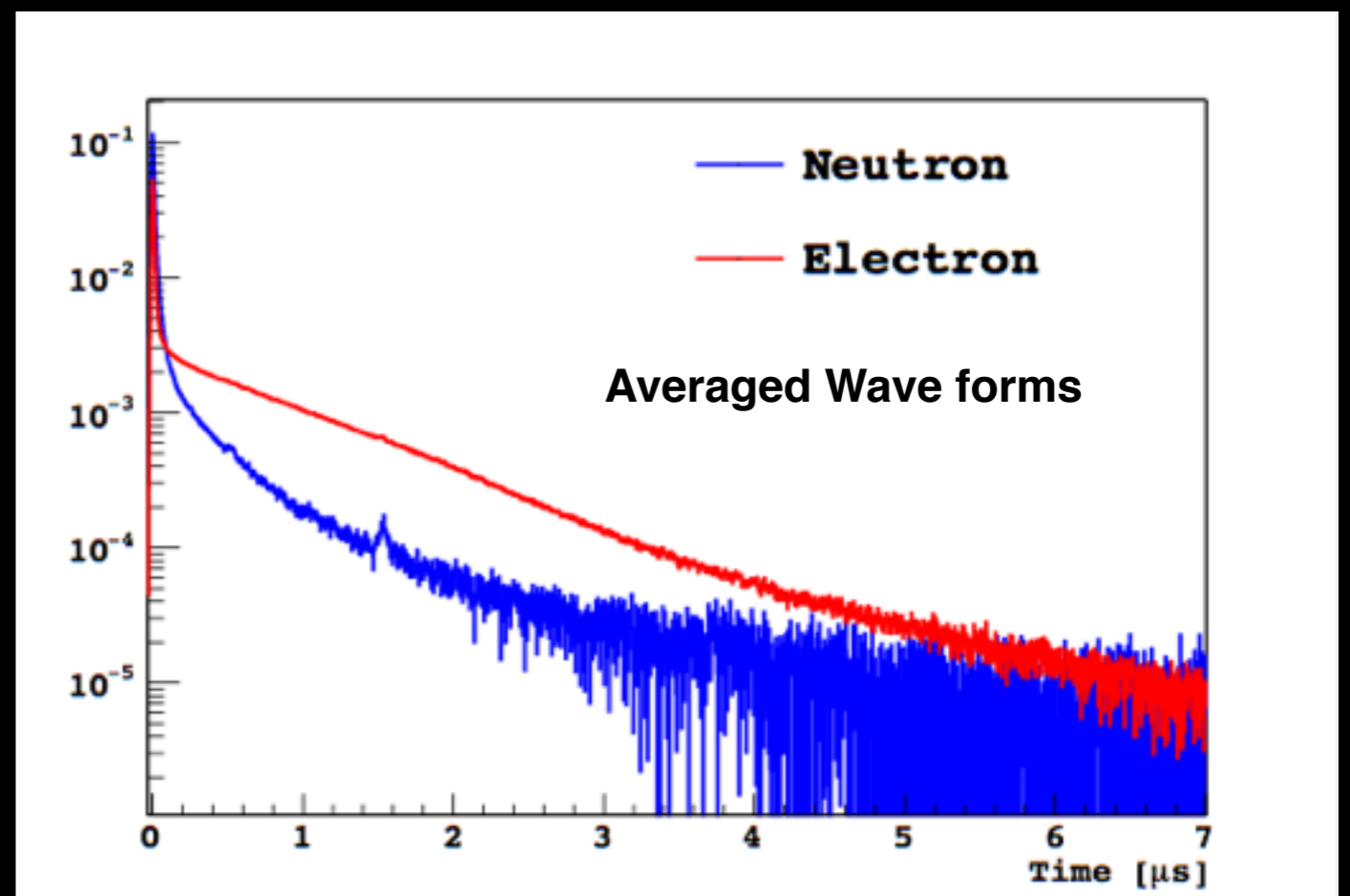
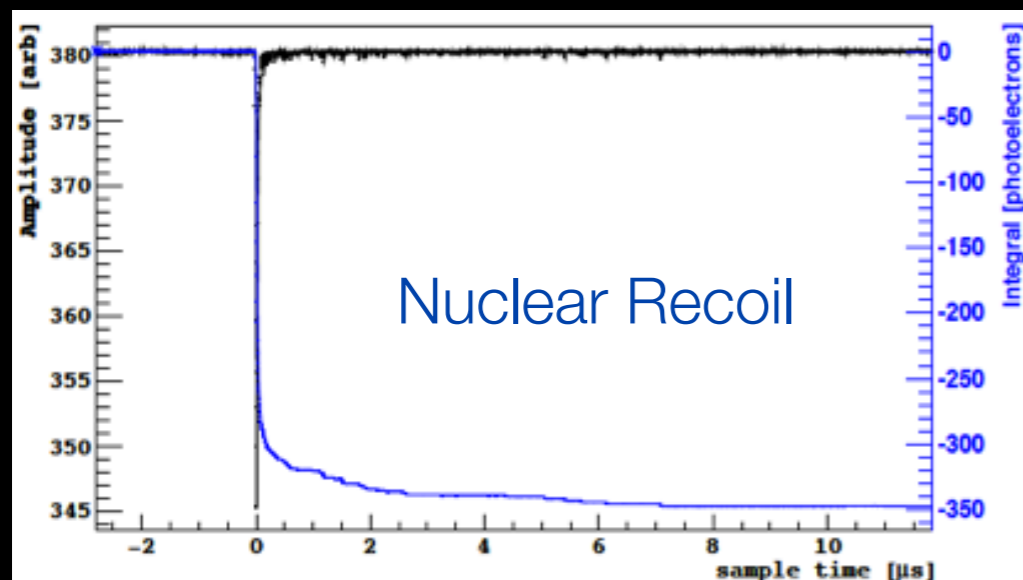
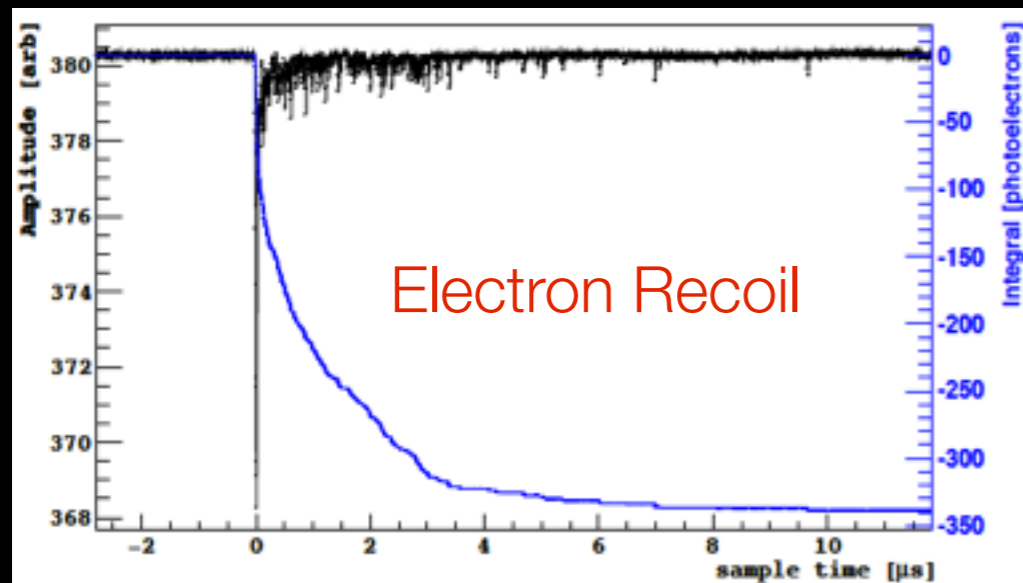
- Intrinsic  $^{39}\text{Ar}$  radioactivity in **atmospheric argon** is the primary background for argon-based detectors
- $^{39}\text{Ar}$  activity sets the dark matter detection threshold at low energies (where pulse shape discrimination is ineffective)
- $^{39}\text{Ar}$  is a **cosmogenic isotope**, and the activity in argon from **underground sources** can be significantly lower compared to **atmospheric argon**
- Recently DarkSide deployed underground argon. **Update will be at the end of this talk.**



# Pulse Shape Discrimination

Electron and nuclear recoils produce different excitation densities in the argon, leading to different ratios of singlet and triplet excitation states

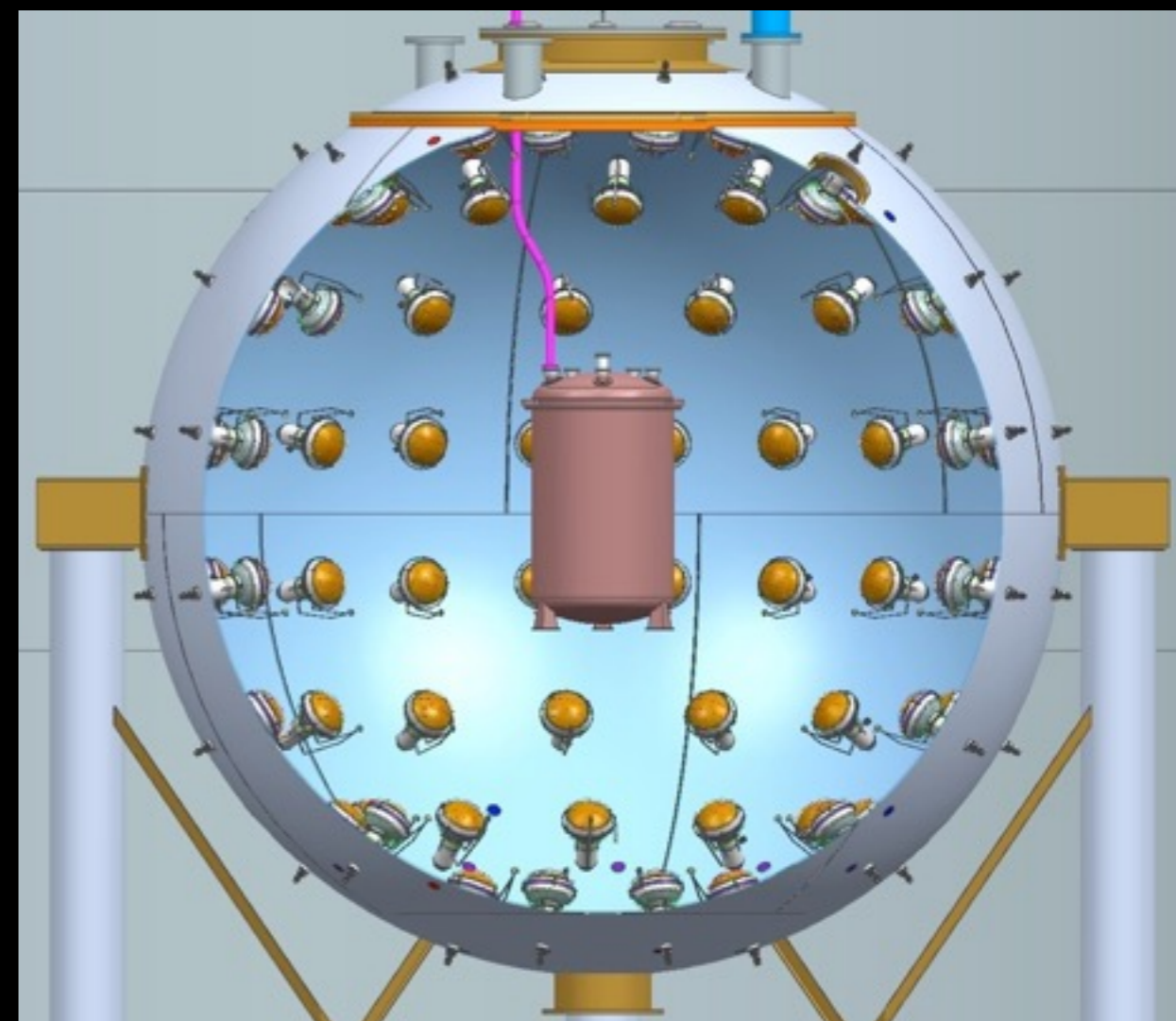
$\tau_{\text{singlet}} \sim 7 \text{ ns}$   
 $\tau_{\text{triplet}} \sim 1500 \text{ ns}$



# Liquid Scintillator Veto

Liquid scintillator allows coincident veto of **neutrons (and  $\gamma$ 's)** in the TPC and provides *in situ* measurement of the neutron background rate

- 4 m diameter sphere containing PC + TMB scintillator
- Instrumented with 110 8" PMTs

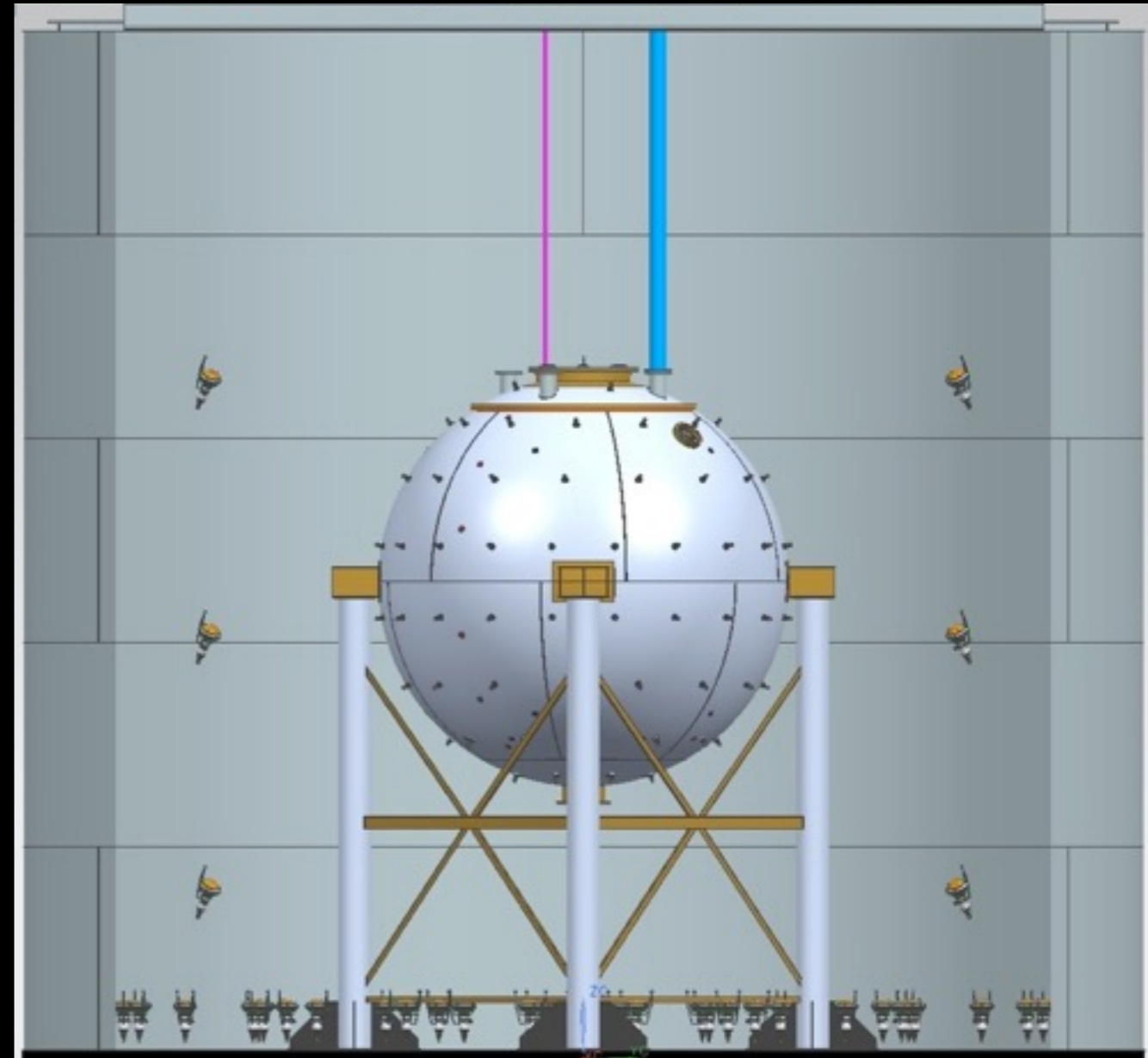


Odd time structure:  $^{14}\text{C}$  content is too high ( $\sim 98\%$  efficiency) to achieve design efficiency ( $\sim 99.5\%$ ) after the first fill.

The TMB was replaced with new low  $^{14}\text{C}$  TMB (Jan. 2015).  $^{14}\text{C}$  activity decreased from **150 kBq** to **0.3 kBq**.

# External Water tank

- 80 PMTs within water tank (11 m diameter x 10 m height)
- Acts as a **muon and cosmogenic veto** (~ 99% efficiency)
- Provides **passive  $\gamma$ 's and neutron shielding**



# Radon-Free Clean Rooms

Radon daughters plate out on surfaces of the detector causing dangerous alpha-induced nuclear recoils.

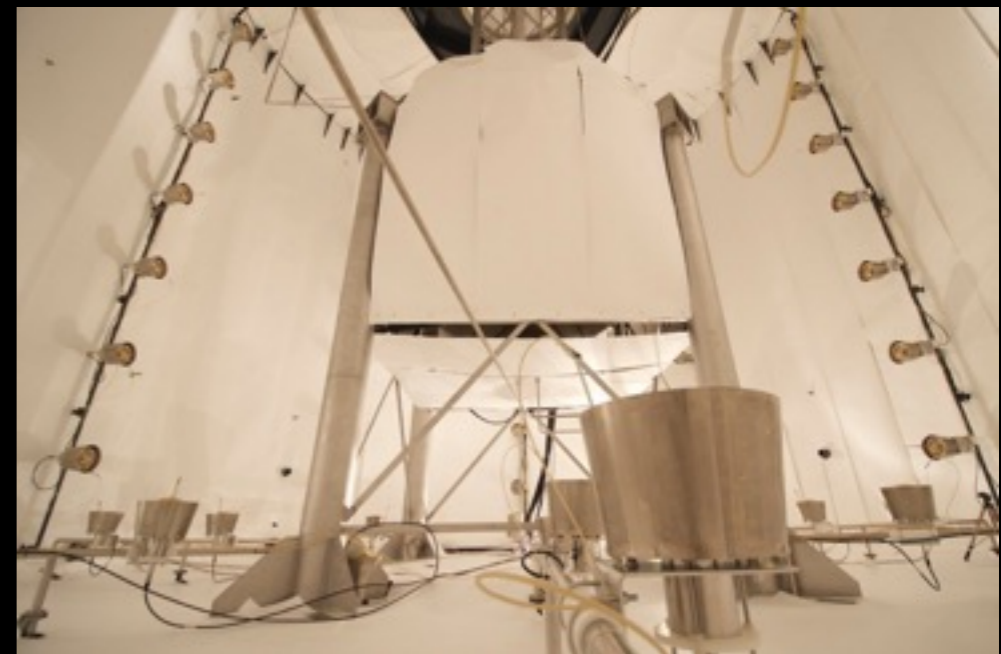
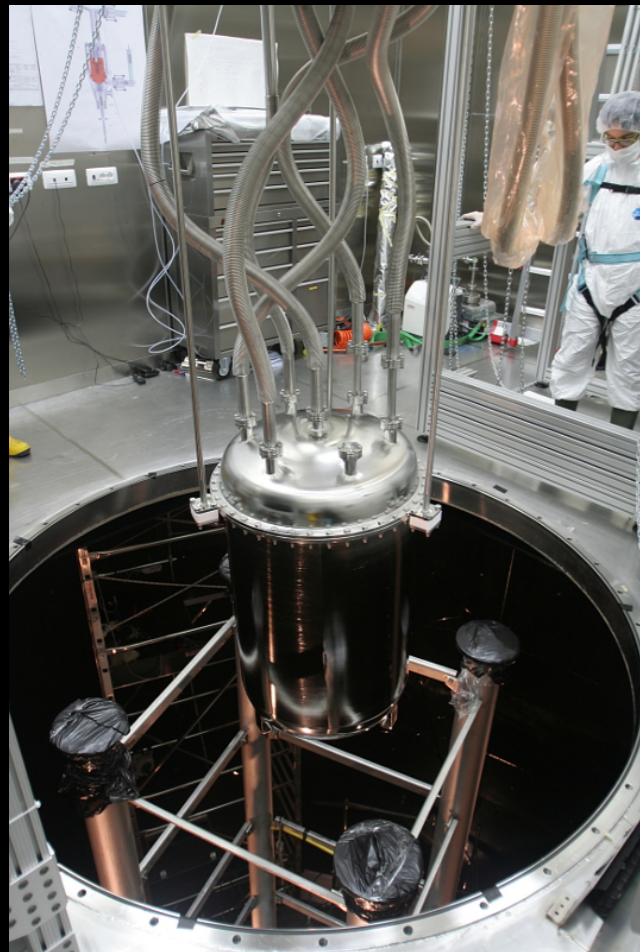
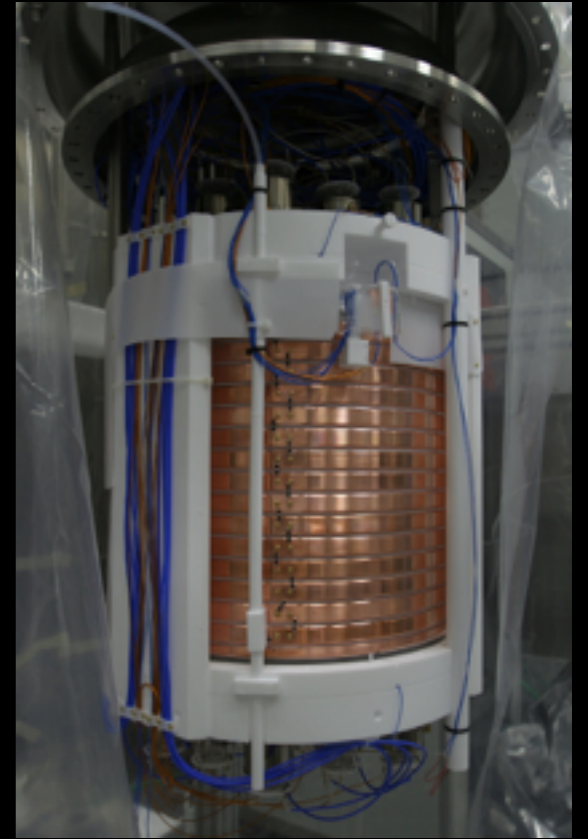
**Final preparation, cleaning, evaporation and assembly** of all inner detector parts was carried out in radon-free clean rooms.



**Typical radon in air ~ 30 Bq/m<sup>3</sup>**

**Cleanroom radon levels < 5 mBq/m<sup>3</sup>**

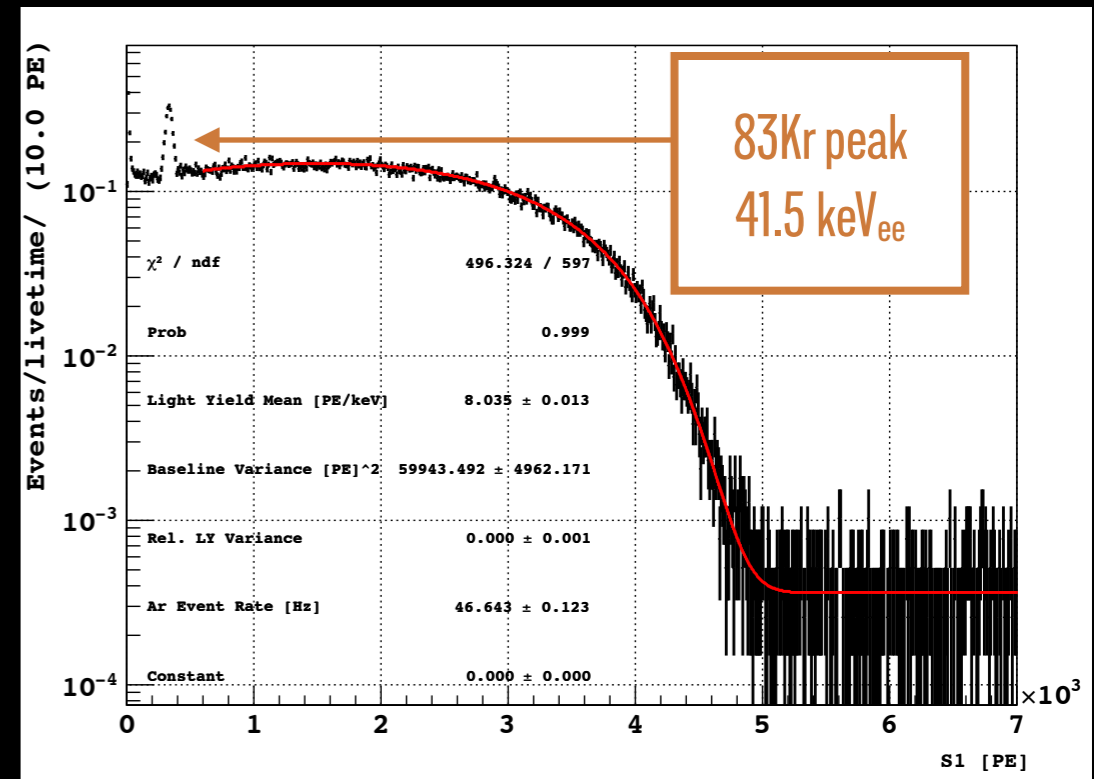
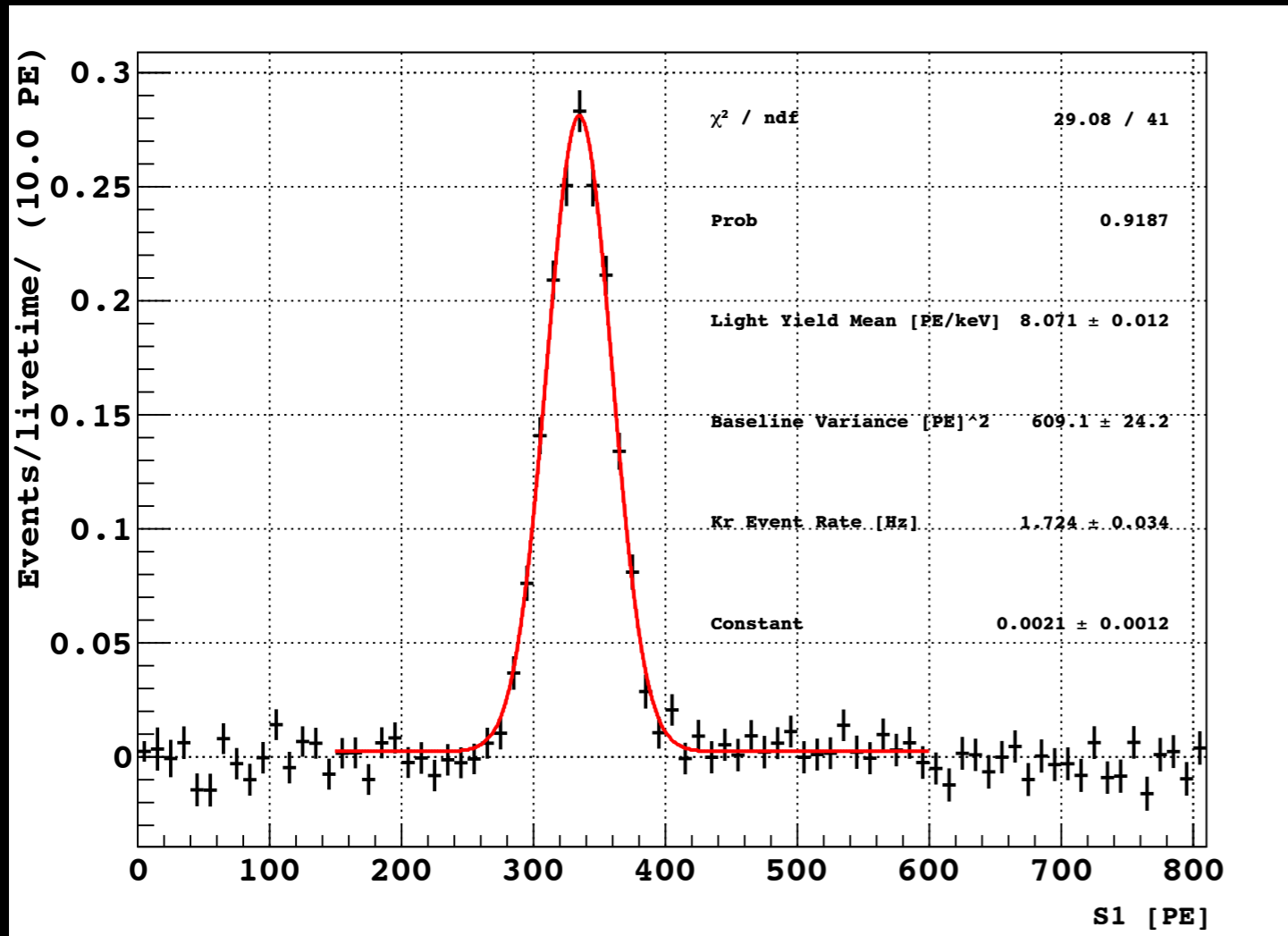
# DS50 Commissioning



# TPC Calibration

# TPC: ER calibration @ null field

Subtracted BG



$^{83m}\text{Kr}$  Half-life = 1.83 hours

$^{83m}\text{Kr}$  gas deployed into detector (41.5 keV<sub>ee</sub>)

Fits to  $^{39}\text{Ar}$  and  $^{83m}\text{Kr}$  spectrum indicate  
**AVERAGE LIGHT YIELD:  $7.9 \pm 0.4$  PE/keV<sub>ee</sub>**

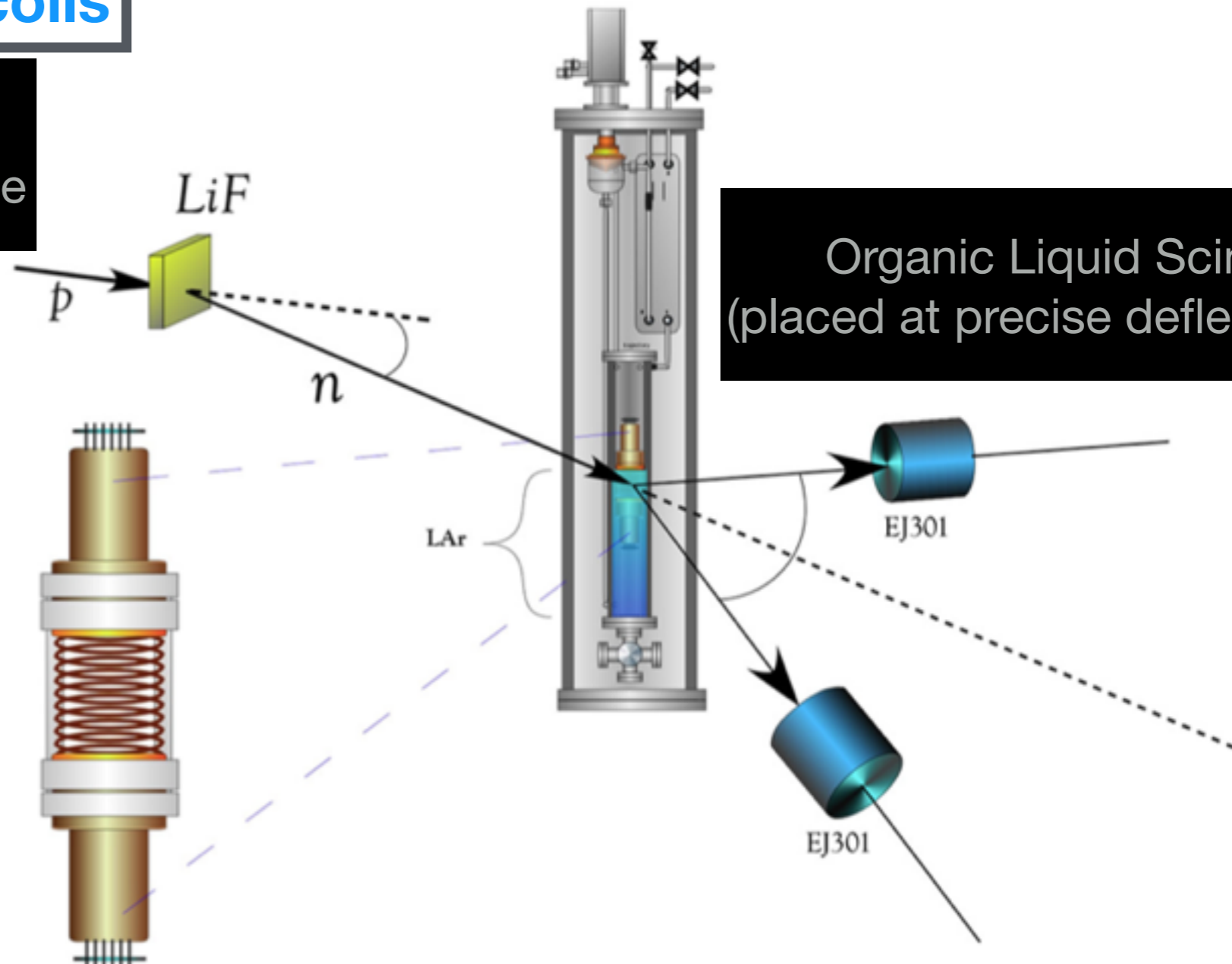


# SCENE

## (Scintillation Efficiency of Nuclear Recoils in Noble Elements)

### For Nuclear Recoils

Proton Beam at  
University of Notre Dame



Organic Liquid Scintillators  
(placed at precise deflection angles)

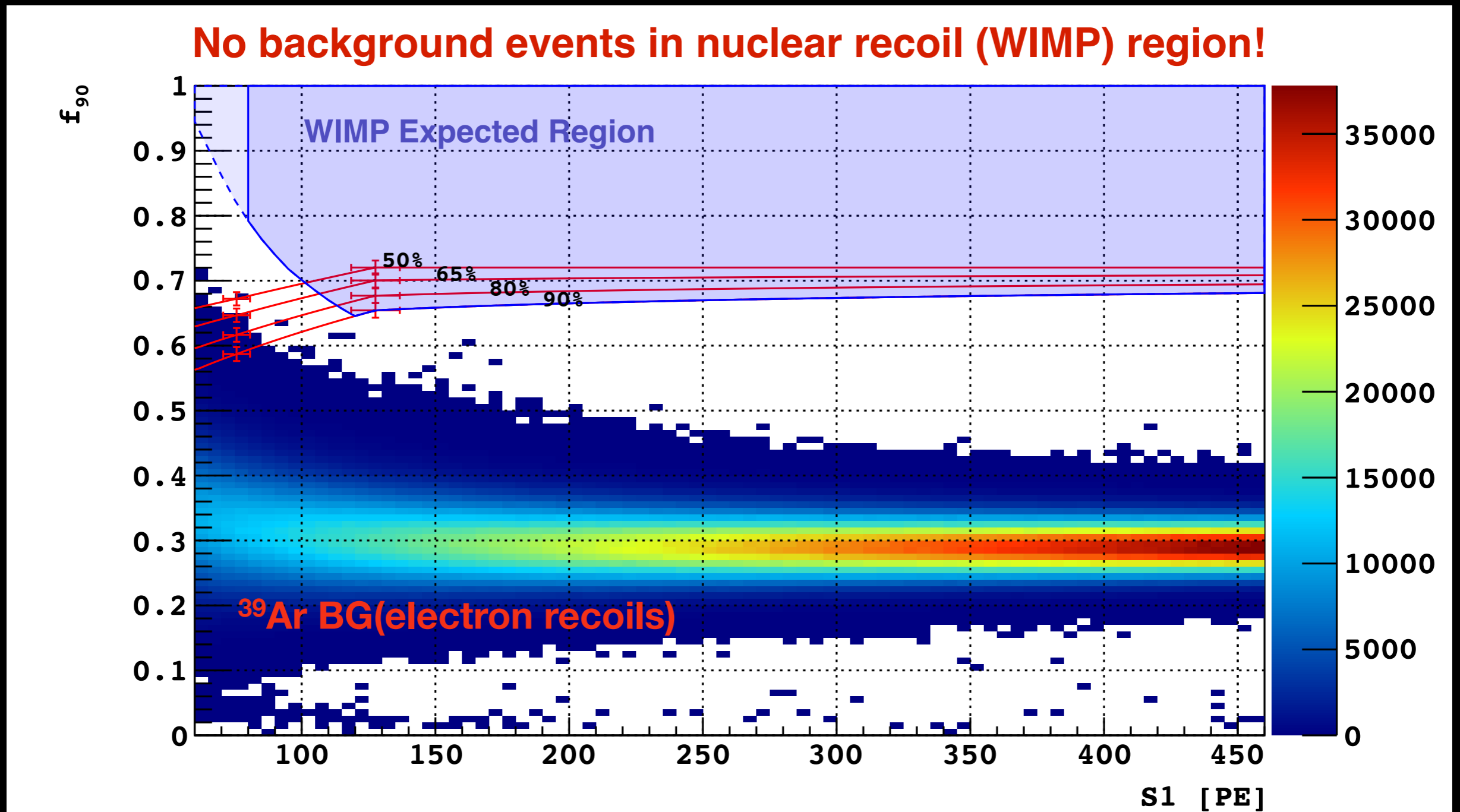
Liquid Argon TPC  
(based on DarkSide  
design)

${}^7Li(p, n){}^7Be$  reaction produces low energy monoenergetic neutrons  
TOF measurement between target, LAr and organic scintillators allows  
clean identification of elastic neutron interactions of known energy

# The First Physics Result from DS-50

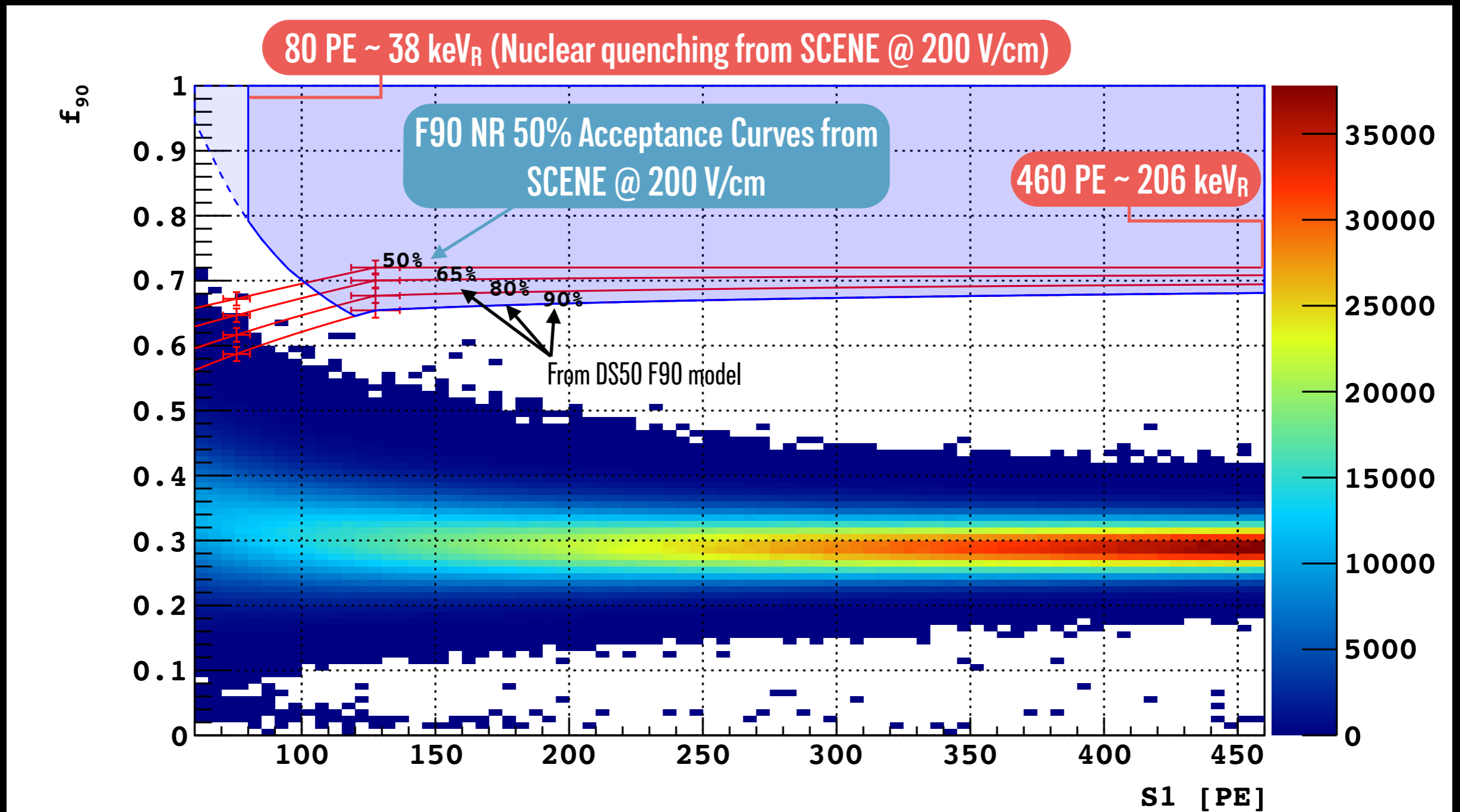
# Background-free exposure of $1422 \pm 67$ kg·day

[Phys. Lett. B 743 \(2015\) 456](#)



Selected only single-hit interactions in the TPC fiducial volume (36.9 kg) with no energy deposition in the veto

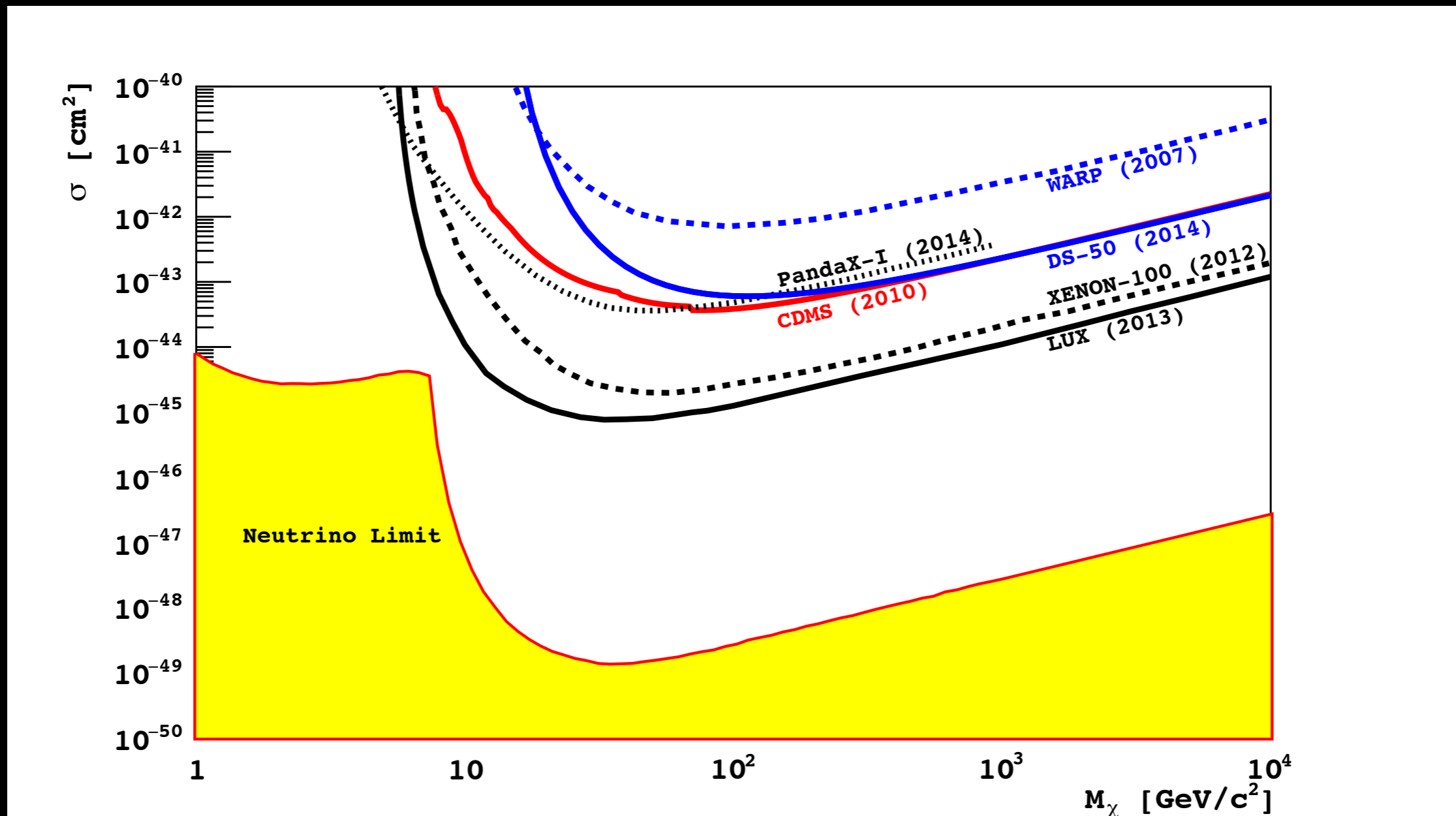
# Background-free exposure of $1422 \pm 67$ kg·day



Selected only single-hit interactions in the TPC fiducial volume (36.9 kg) with no energy deposition in the veto

# Dark Matter exclusion plot

This is the most sensitive dark matter search performed with an **argon** target. The WIMP-nucleon spin-independent cross section is  $6.1 \times 10^{-44}$  cm<sup>2</sup> for a WIMP mass of 100 GeV/c<sup>2</sup>.

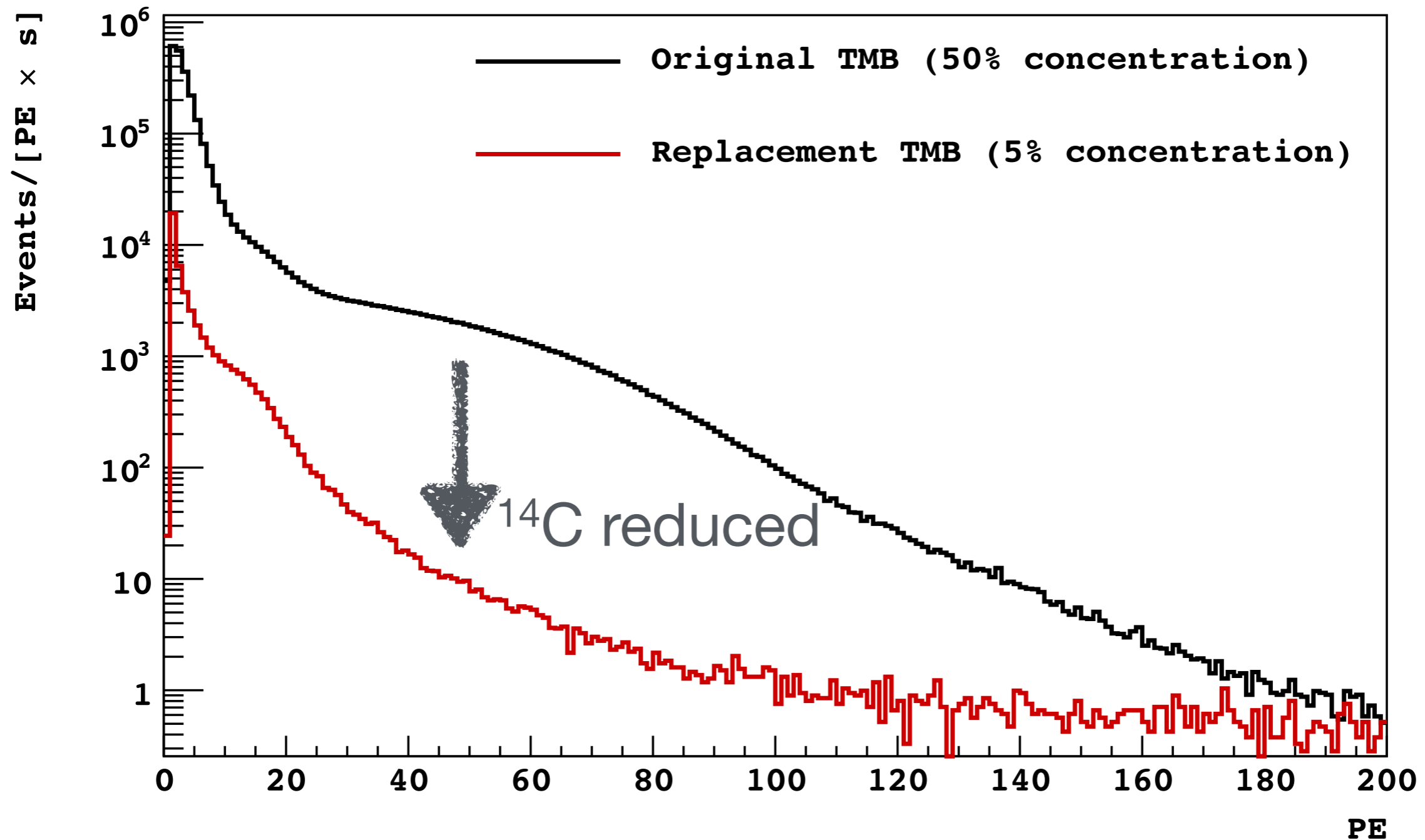


# DS50 Timeline

- **Oct. 2013:** LArTPC, Neutron Veto and Muon Veto commissioned.
  - TPC filled with **atmospheric argon** (AAr).
- **Up to June 2014:** data taken with high  $^{14}\text{C}$  content in LSV.
  - **47.1 live days** (1422 kg · day fiducial) for the first physics result.
  - TMB ( $^{14}\text{C}$ ) was removed to reduce the  $^{14}\text{C}$  rate.
- **Oct. to Dec. 2014:** Calibration of TPC w/ radioactive sources.
- **Jan. 2015:** Add radiopure TMB at 5% concentration.
- **Mar. to Apr. 2015:** Fill with **UAr** and re-commissioning the detector.
- **Apr. to Aug. 2015:** Accumulate data with **UAr** for **dark matter search**.

# DS-50 Current Status

# Status of Liquid Scintillator Veto



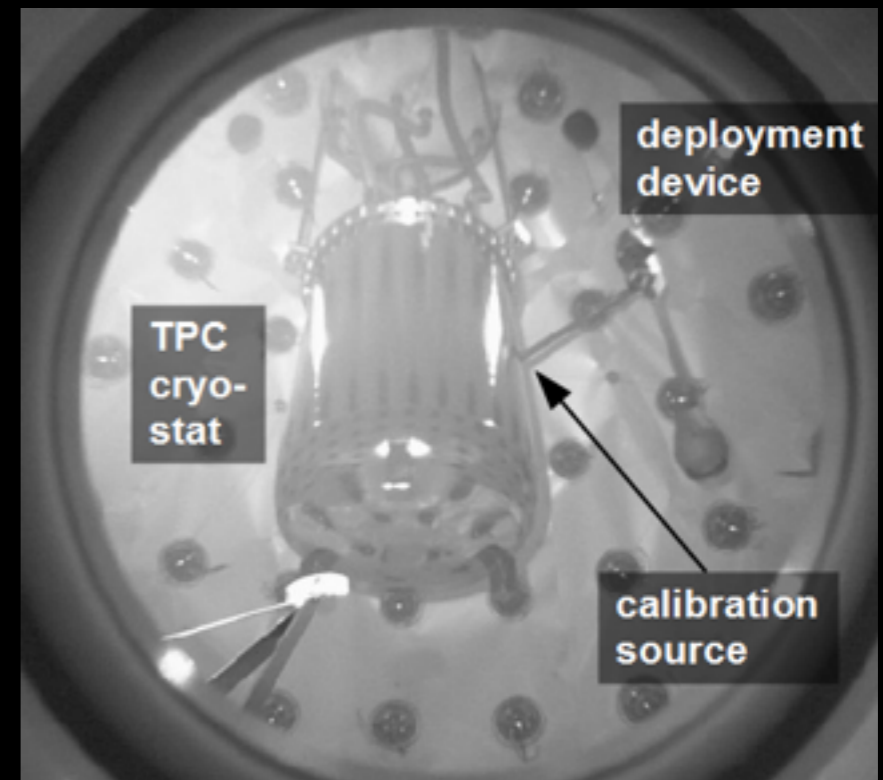
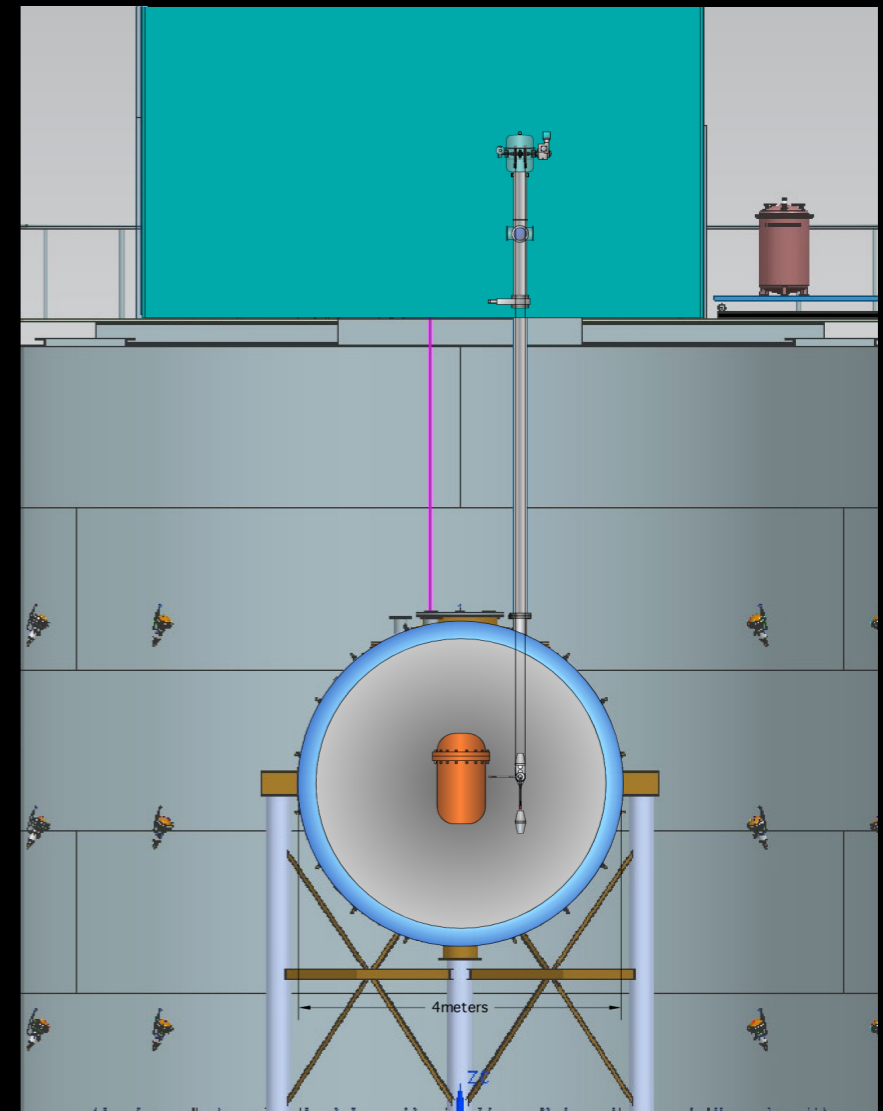
$^{14}\text{C}$  activity decreased from 150 kBq to 0.3 kBq.



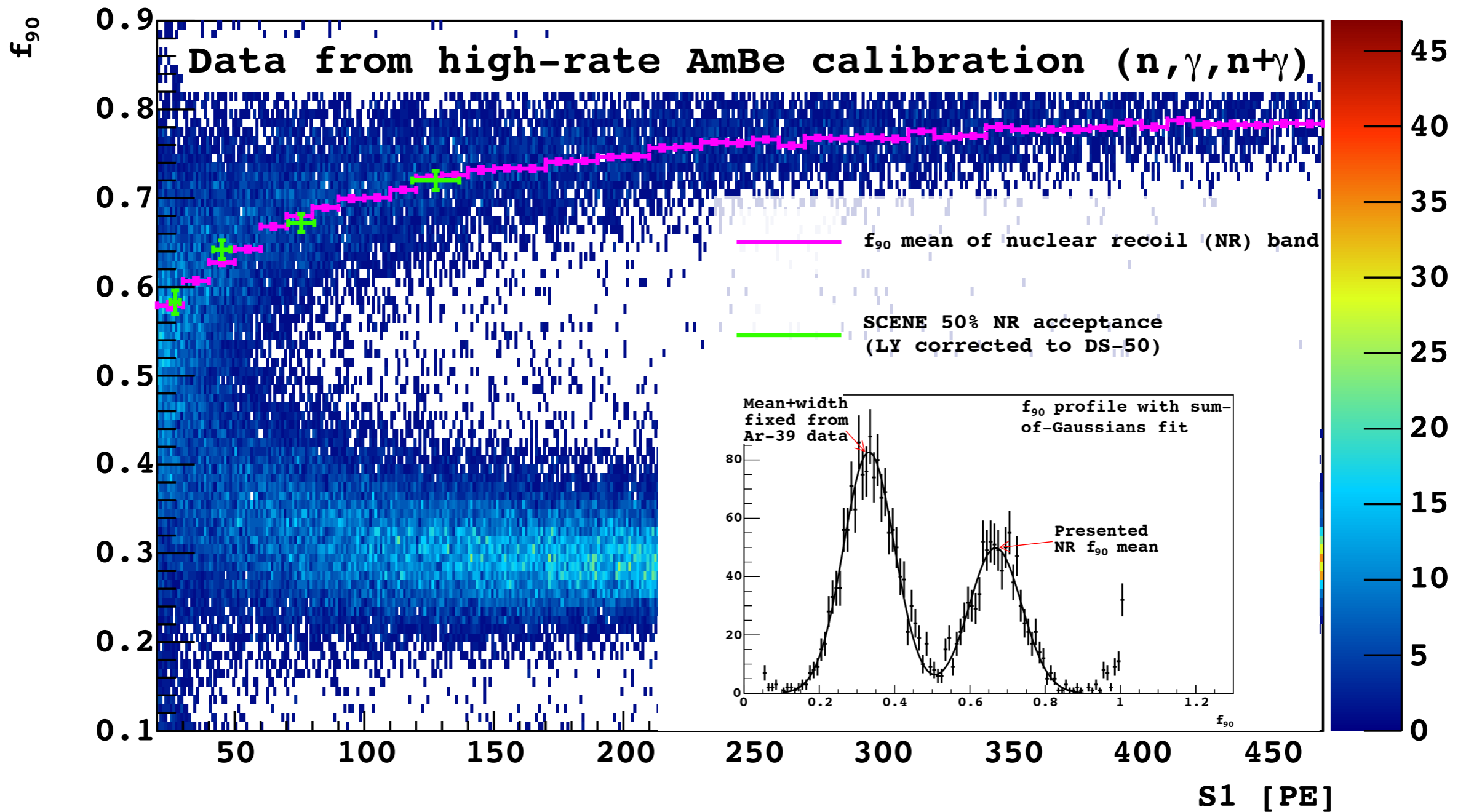
# CALIS - CALibration Insertion System

Calibrate both **TPC** and **Neutron veto**

- **Gamma sources:**  $^{57}\text{Co}$  (122 keV),  $^{133}\text{Ba}$  (356 keV),  $^{137}\text{Cs}$  (663 keV)
- **Neutron source:** AmBe w/ and w/o collimator
- **Different drift fields:** null, 100 V/cm, 150 V/cm, 200 V/cm



# NR from AmBe source



NR band matches with the points extrapolated from SCENE.

# Underground Ar



Plant at Colorado

**1. Extraction at Colorado** (CO<sub>2</sub> Well)  
Extract a crude argon gas mixture (Ar, N<sub>2</sub>, and He)

**2. Purification at Fermilab**  
Separate Ar from He and N<sub>2</sub>



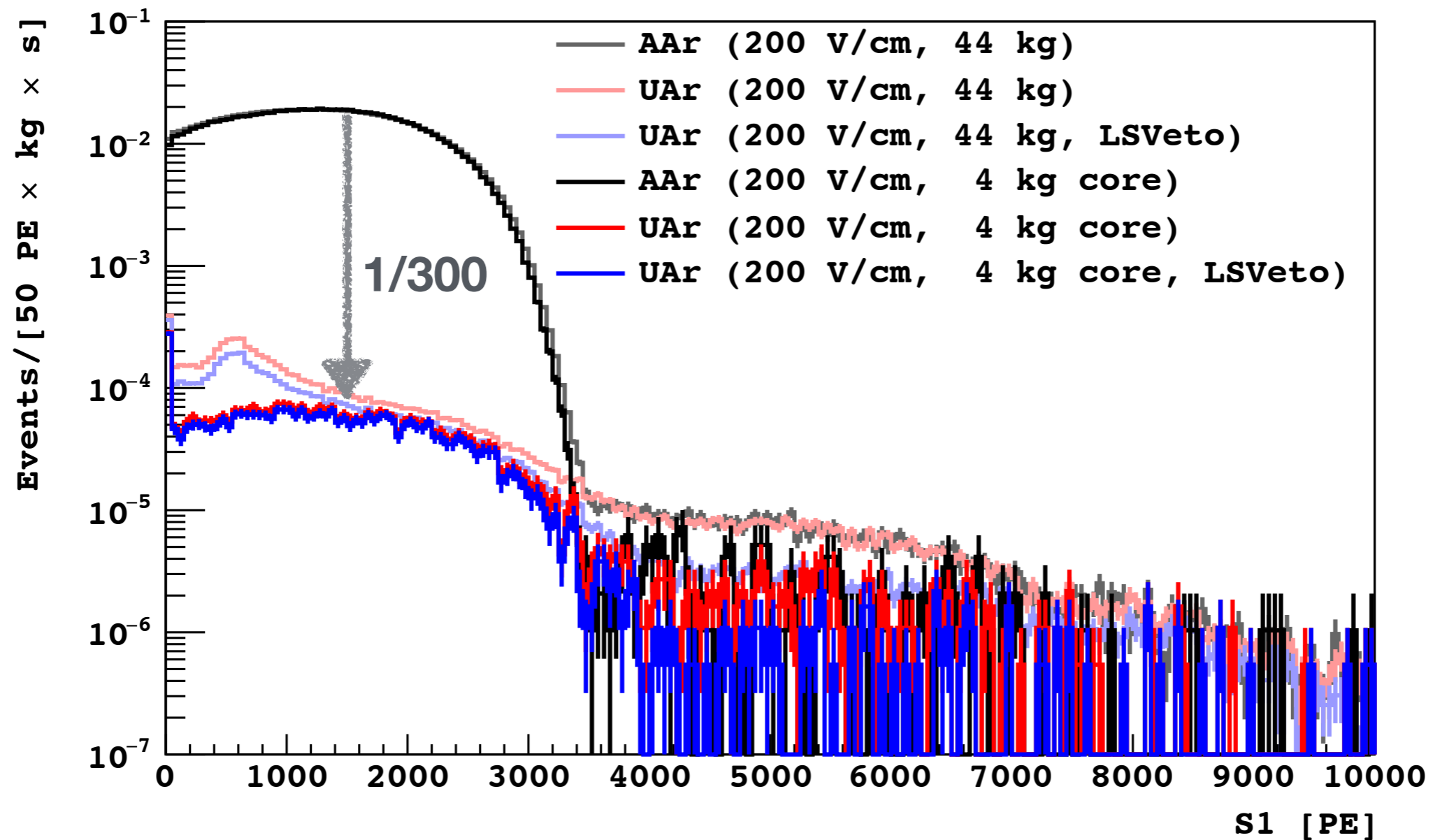
Distillation Column at Fermilab



UAr bottles at LNGS

**3. Arrived at LNGS**  
Ready to fill into DS-50

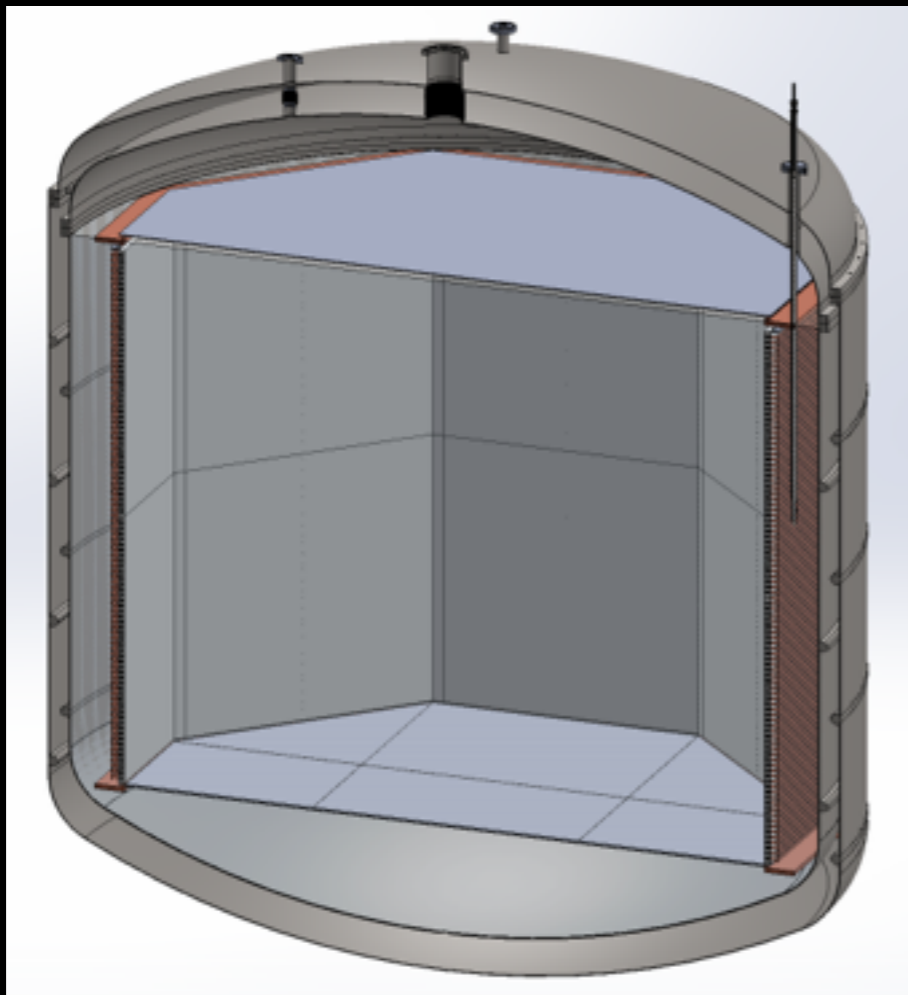
# Underground Ar



Concentration of  $^{39}\text{Ar}$  in **UAr** is at least 300 times lower than in **AAr**.

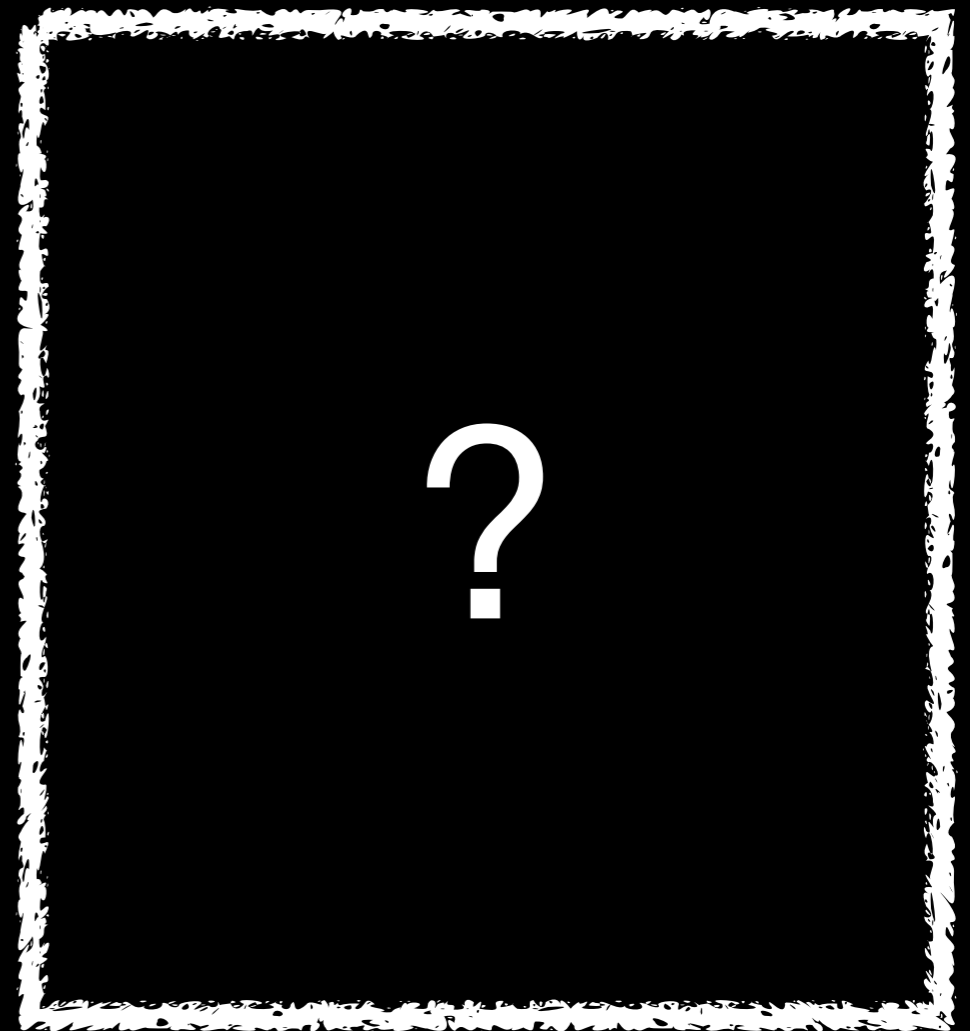
Low level of  $^{39}\text{Ar}$  allows extension of DarkSide program to **ton-scale** detector.

# Future Detectors



**DS-20k**

30 tonne (20 tonne fiducial) detector



**ARGONIE**

300 tonne (200 tonne fiducial) detector

# Requirements for DS-20k

## Neutron Background:

- **Cosmogenic:** Veto system
- **Radiogenic:** radiopure SiPM & ultra-clean Titanium (TPC cryostat)

## $\beta/\gamma$ background:

- **$^{39}\text{Ar}$ :** Underground Argon (Urania Project) & Depleted Argon (Aria Project)
- **$\gamma$ :** SiPM & ultra-clean Titanium

# Further Depletion of Ar

## **Urania** (Underground Argon):

- Expansion of the argon extraction plant in Cortez, CO, to reach capacity of **100 kg/day** of Underground Argon

## **Aria** (UAr Purification):

- Very tall column in the Seruci mine in Sardinia, Italy, for high-volume chemical and isotopic purification of Underground Argon

Experiment	$\sigma$ [cm <sup>2</sup> ] @1 TeV/c <sup>2</sup>	$\sigma$ [cm <sup>2</sup> ] @10 TeV/c <sup>2</sup>
LUX [10k kg×day Xe]	$1.1 \times 10^{-44}$	$1.2 \times 10^{-43}$
XENON [7.6k kg×day Xe]	$1.9 \times 10^{-44}$	$1.9 \times 10^{-43}$
DS-50 [1.4k kg×day Ar]	$2.3 \times 10^{-43}$	$2.1 \times 10^{-42}$
ArDM [1.5 tonne×yr Ar]	$8 \times 10^{-45}$	$7 \times 10^{-44}$
DEAP-3600 [3.0 tonne×yr Ar]	$5 \times 10^{-46}$	$5 \times 10^{-45}$
XENON-1ton [2.7 tonne×yr Xe]	$3 \times 10^{-46}$	$3 \times 10^{-45}$
LZ [15 tonne×yr Xe]	$5 \times 10^{-47}$	$5 \times 10^{-46}$
DS-20k [100 tonne×yr]	$9 \times 10^{-48}$	$9 \times 10^{-47}$
1 Neutrino Event [400 tonne×yr Ar or 300 tonne×yr Xe]	$2 \times 10^{-48}$	$2 \times 10^{-47}$
ARGO [1,000 tonne×yr]	$9 \times 10^{-49}$	$9 \times 10^{-48}$



# DarkSide-20k and Argo Lol Signatories

D. Franco, A Tonazzo - [APC Paris](#)  
D. Alton - [Augustana College](#)  
A. Kubankin - [Belgorod National Research University](#)  
K. Keeter, B. Mount - [Black Hills State University](#)  
L. Romero, R. Santorelli - [CIEMAT](#)  
S. Horikawa, K. Nikolics, C. Regenfus,  
A. Rubbia - [ETH Zürich](#)  
S. Pordes - [Fermilab](#)  
A. Gola, C. Piemonte - [FBK & TIFPA](#)  
S. Davini - [GSSI](#)  
E. Hungerford, A. Renshaw - [University of Houston](#)  
M. Guan, J. Liu, Y. Ma, C. Yang, W. Zhong - [IHEP Beijing](#)  
N. Canci, F. Gabriele, G. Bonfini, A. Razeto, N. Rossi,  
F. Villante - [LNGS](#)  
C. Jollet, A. Meregaglia - [IPHC Strasbourg](#)  
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K. Fomenko, A. Sotnikov, O. Smirnov - [JINR](#)  
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E. Unzhakov - [PNPI Saint Peterburg](#)  
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H. O. Back - [PNNL](#)  
M. Ghioni, A. Gulnatti, L. Pellegrini, I. Rech, A. Tosi,

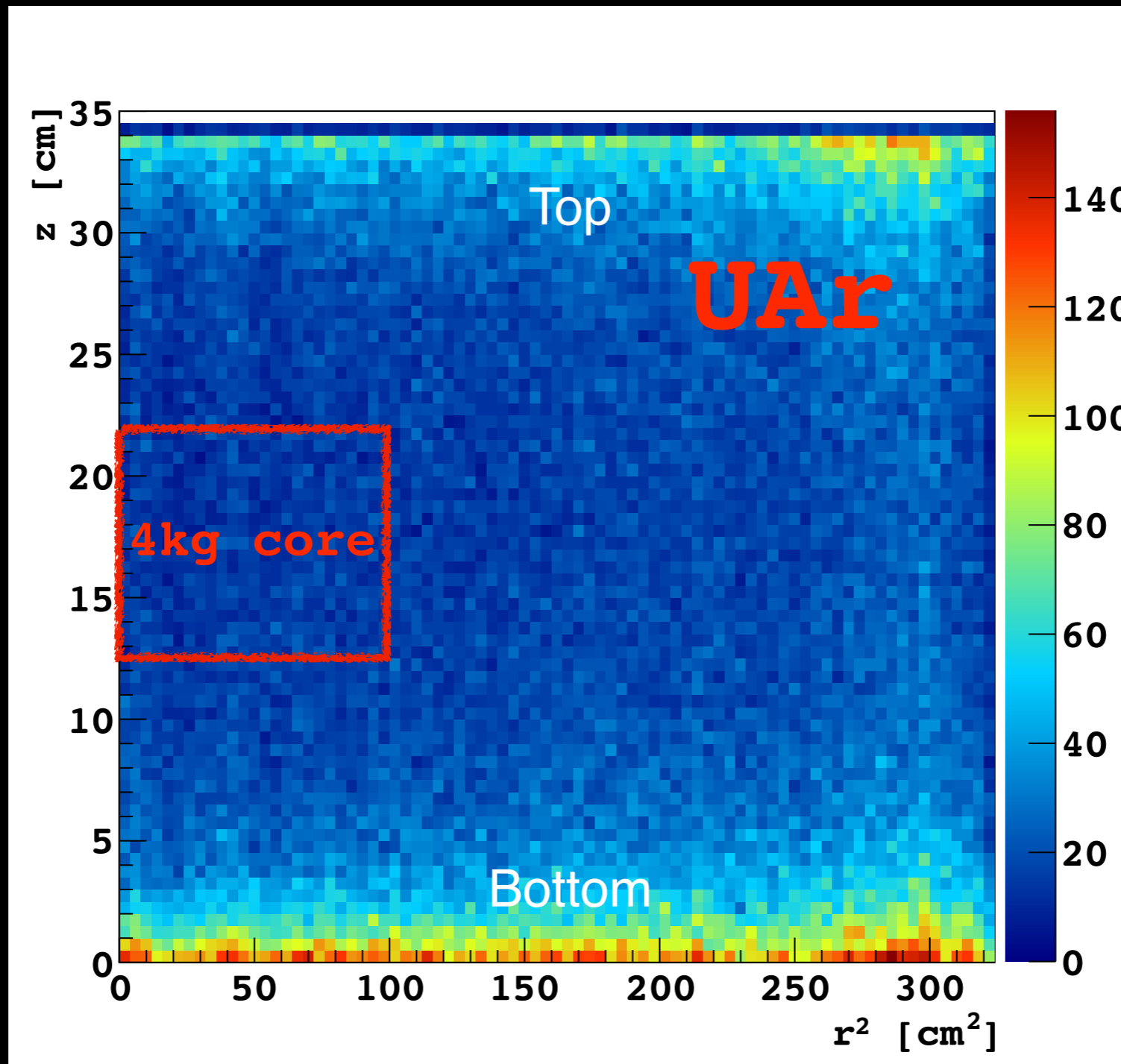
F. Zappa - [Politecnico di Milano](#)  
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I. Zilcov - [SINP MSU Moscow](#)  
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E. Pantic - [UCDavis](#)  
Y. Suvorov, H. Wang - [UCLA](#)  
A. Pocar - [UMass Amherst](#)  
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A. Devoto, M. Lissia, M. Mascia,  
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M. Pallavicini, G. Testera,  
S. Zavatarelli - [Università & INFN Genova](#)  
D. D'Angelo, G. Ranucci - [Università & INFN Milano](#)  
F. Ortica, A. Romani - [Università & INFN Perugia](#)  
S. Catalanotti, A. Cocco, G. Covone, G. Fiorillo,  
B. Rossi - [Università Federico II & INFN Napoli](#)  
C. Dionisi, S. Giagu, M. Rescigno - [Università La Sapienza & INFN Roma](#)  
S. Bussino, S. Mari - [Università & INFN Roma 3](#)  
J. Maricic, R. Milincic, B. Reinhold - [University of Hawaii](#)  
P. Cavalcante - [Virginia Tech](#)

# Summary

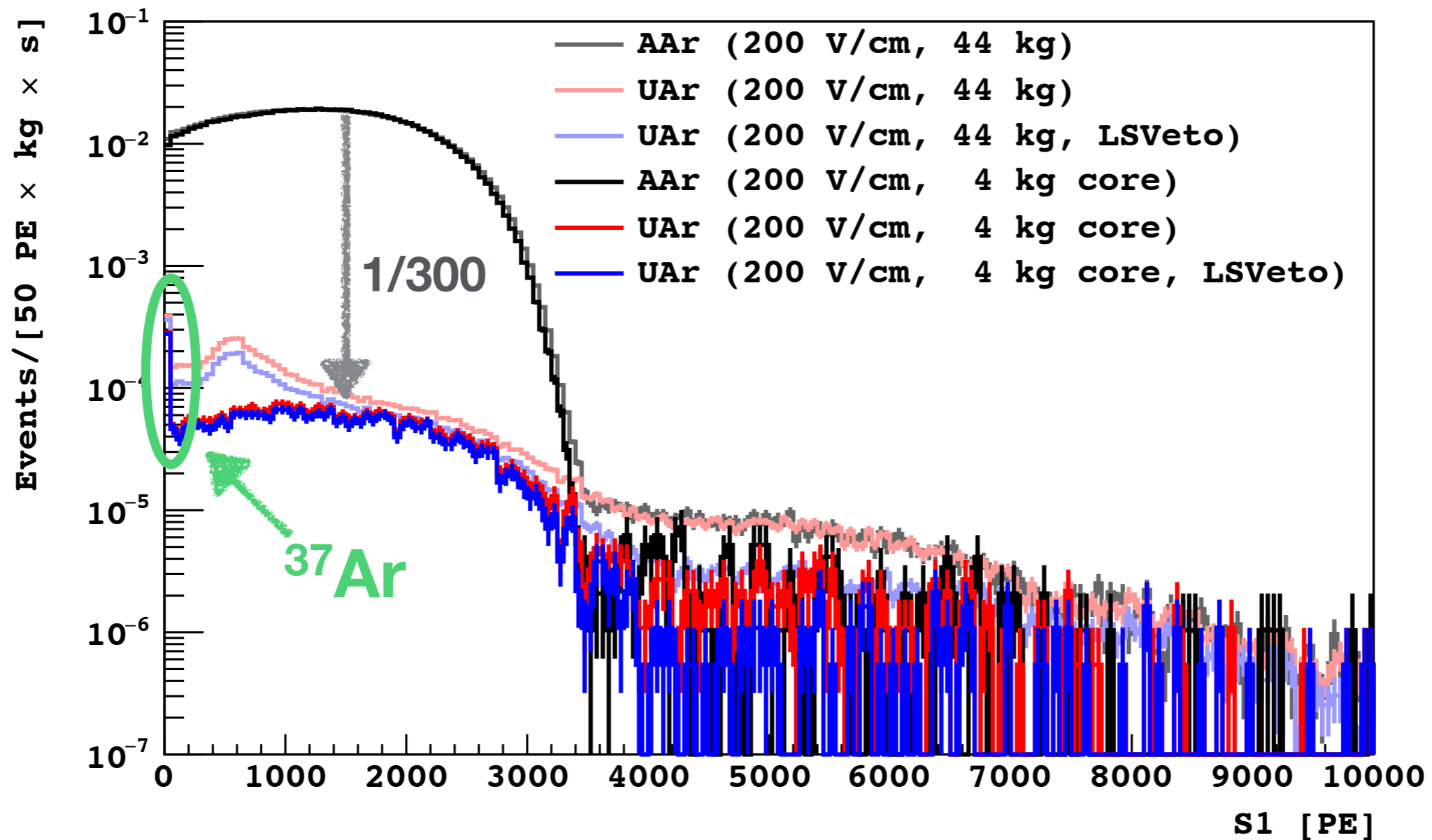
- **Background free**  
 $^{39}\text{Ar}$  BG from **47.1 live days** (1422 kg · day fiducial) of AAr corresponds to that expected in **38.7 years of UAr** DS-50 run (»planning physics run time, 3 years).
- Concentration of  $^{39}\text{Ar}$  in **UAr** is at least **300** times lower than in **AAr**.
- With the BG-free exposure of 1422 kg · day fiducial, DarkSide demonstrates  $^{39}\text{Ar}$  BG rejection at level of **1 tonne·year** with UAr.
- Future detectors are planned and Letter of Intent was submitted to LNGS April 27 2015.

THE END

# Event Position in TPC



# $^{37}\text{Ar}$

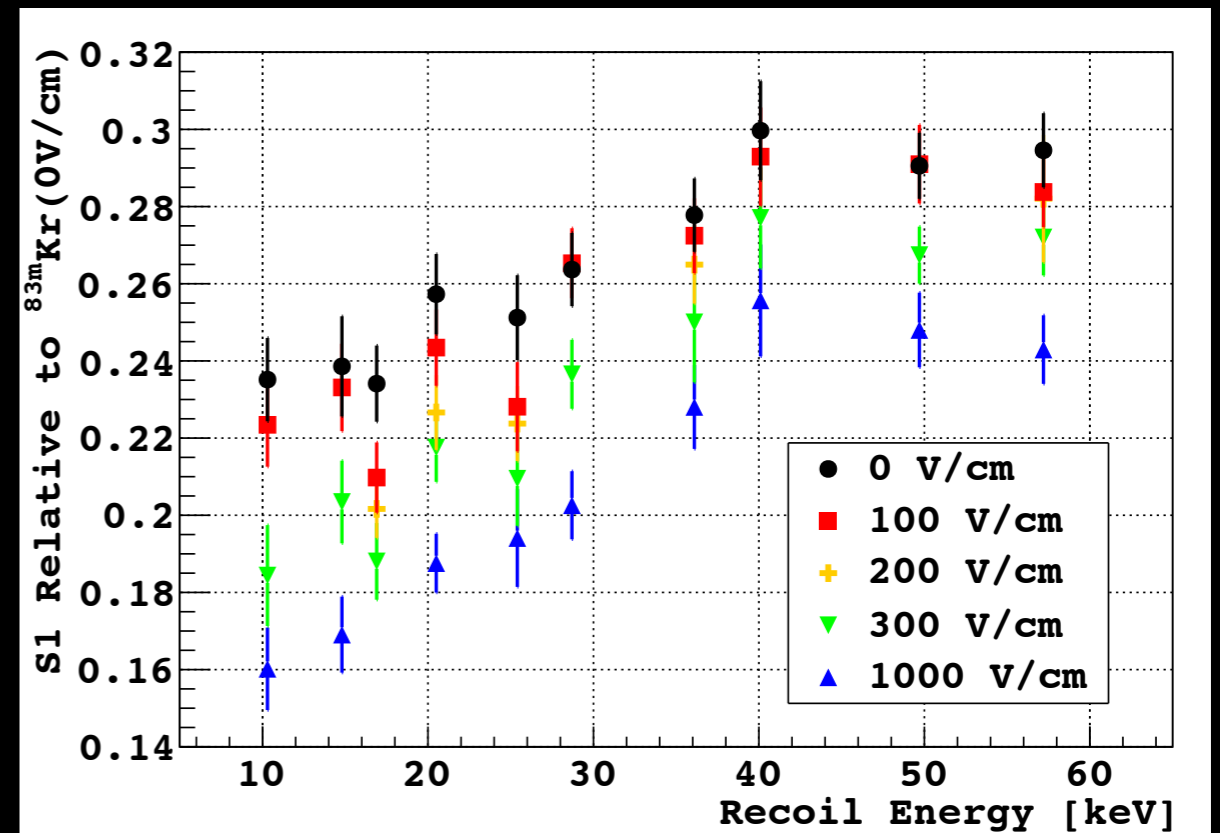
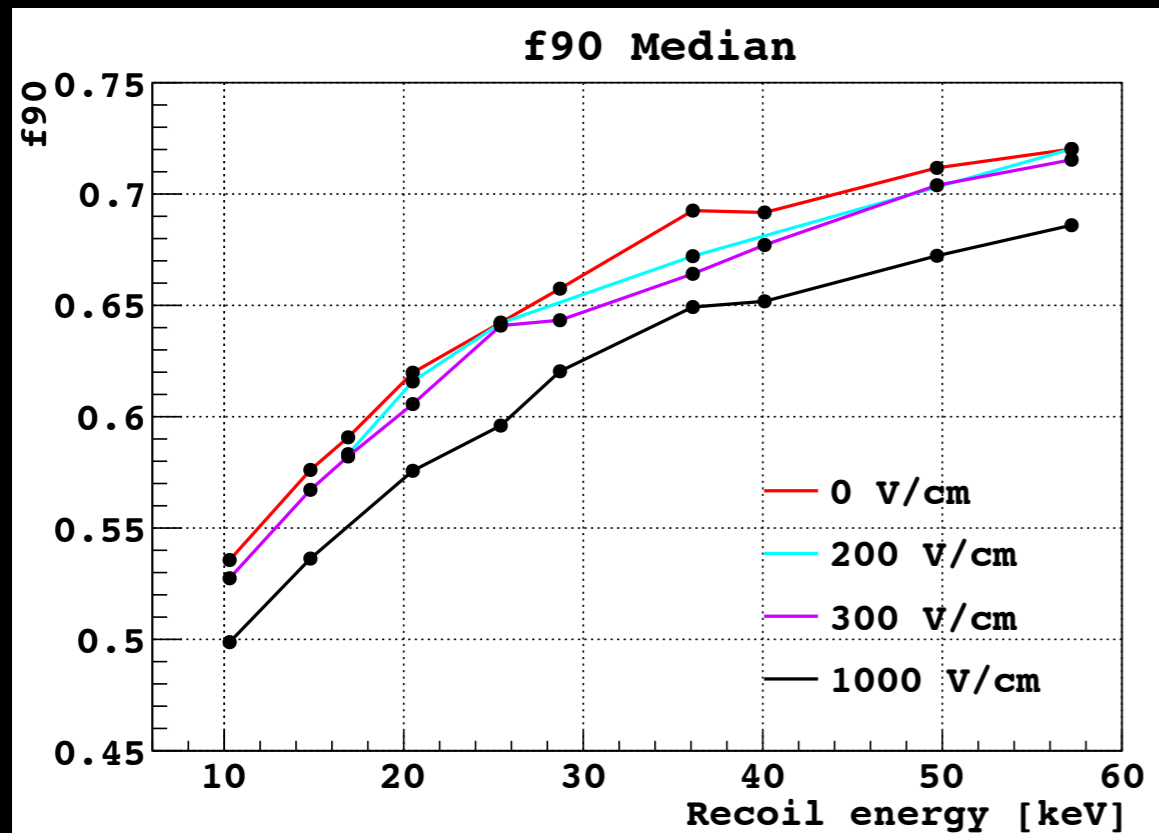


$^{37}\text{Ar}$  is activated by cosmic rays.

Only 35 day half life

Provide low energy (~2-3 keV) calibration point.

# Data from SCENE and the plots

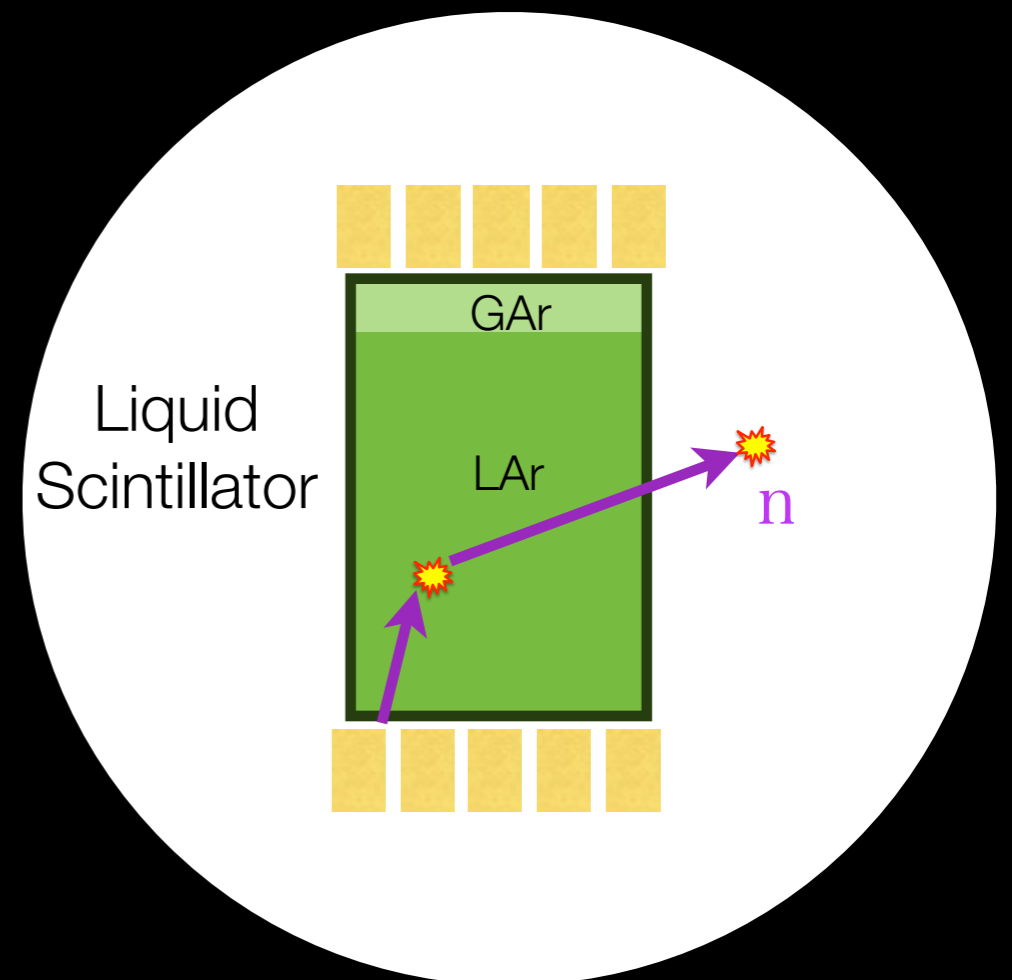
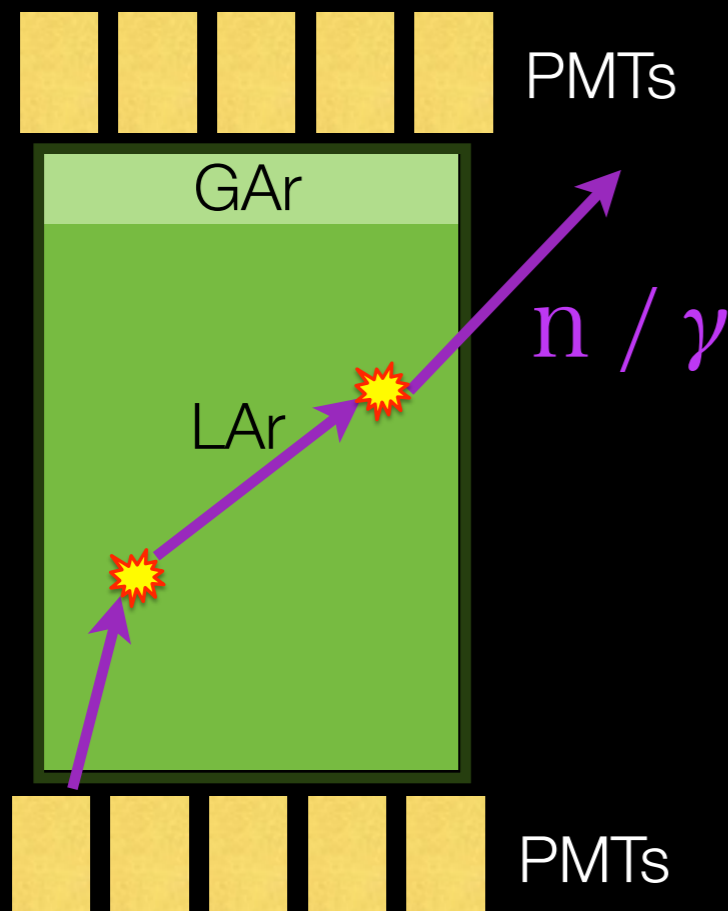


- Left : the median of the f90 distribution for nuclear recoils as a function of energy as measured in the SCENE experiment
- Right: the quenching factor for nuclear recoils as measured by the SCENE experiment

Kr calibration data is used for cross calibration of light yield between DS50 and SCENE.

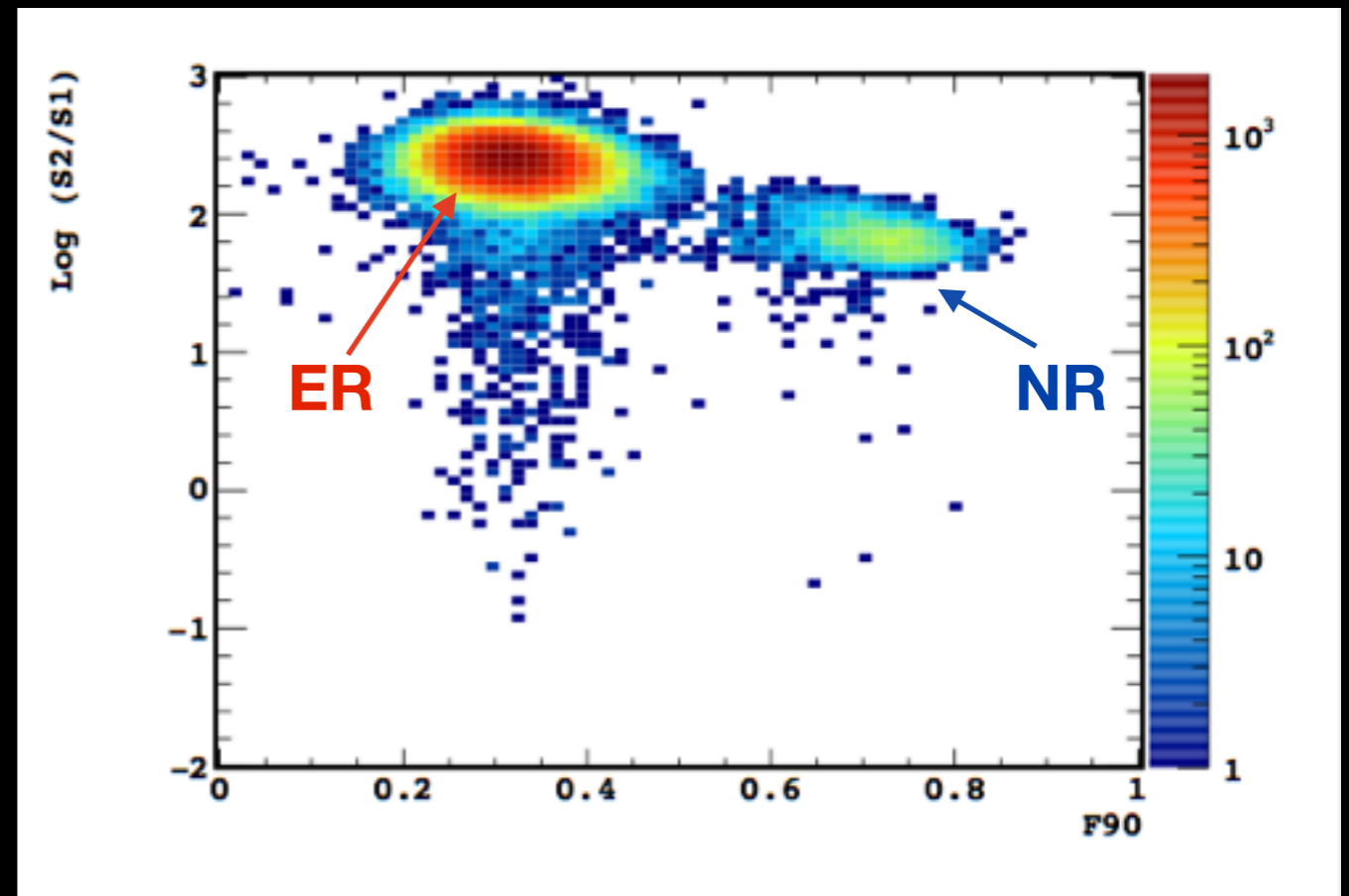
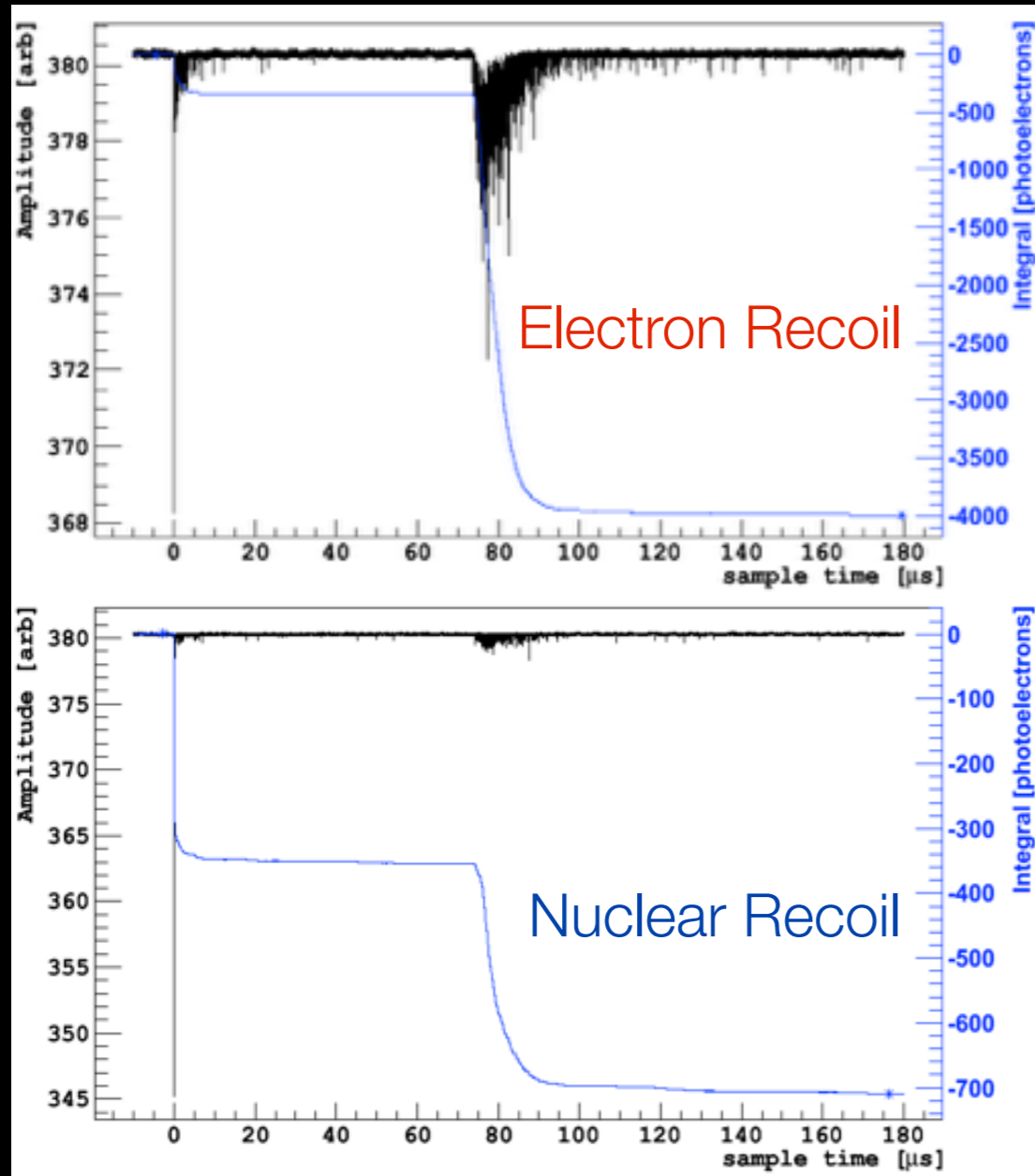
# Multiple Interactions

Expected WIMP signal	Background Rejection Technique	Backgrounds Removed
Single Interaction	Multiple S2 Cut in TPC Liquid Scintillator Veto	Neutrons, Gamma rays



# S2/S1

Electron and nuclear recoils produce different ionization densities that lead to different fractions of electrons that survive recombination



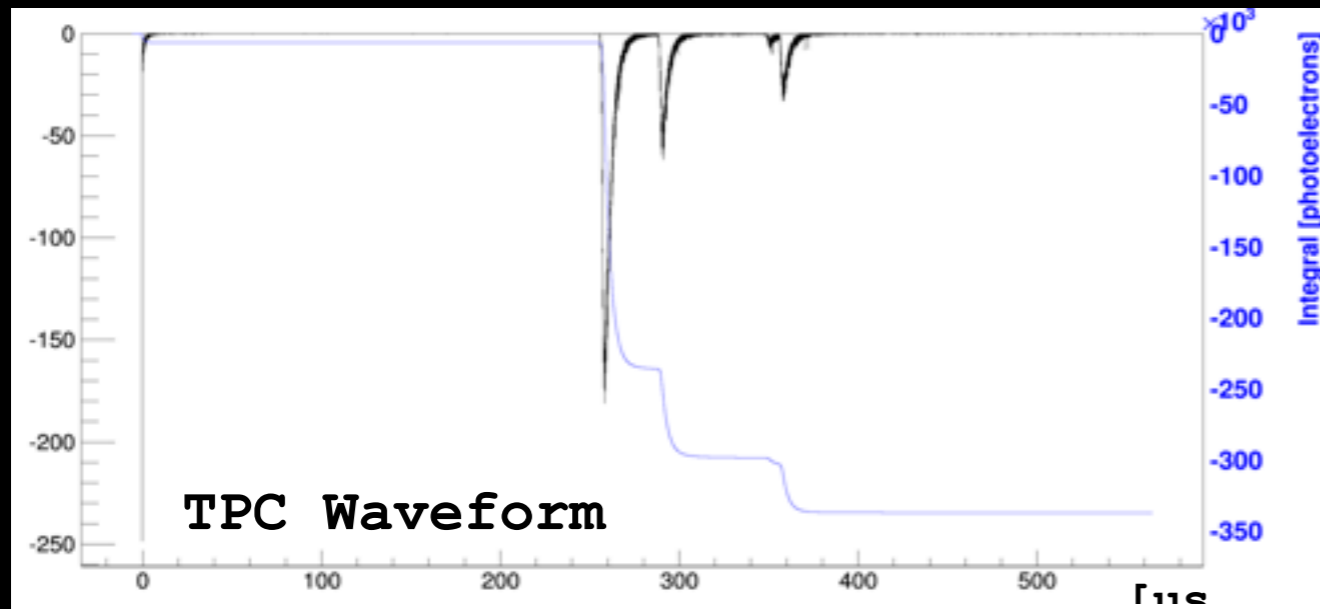
Am-Be Source

Ratio of ionization and scintillation signal (S2/S1) can be used to distinguish between the two populations

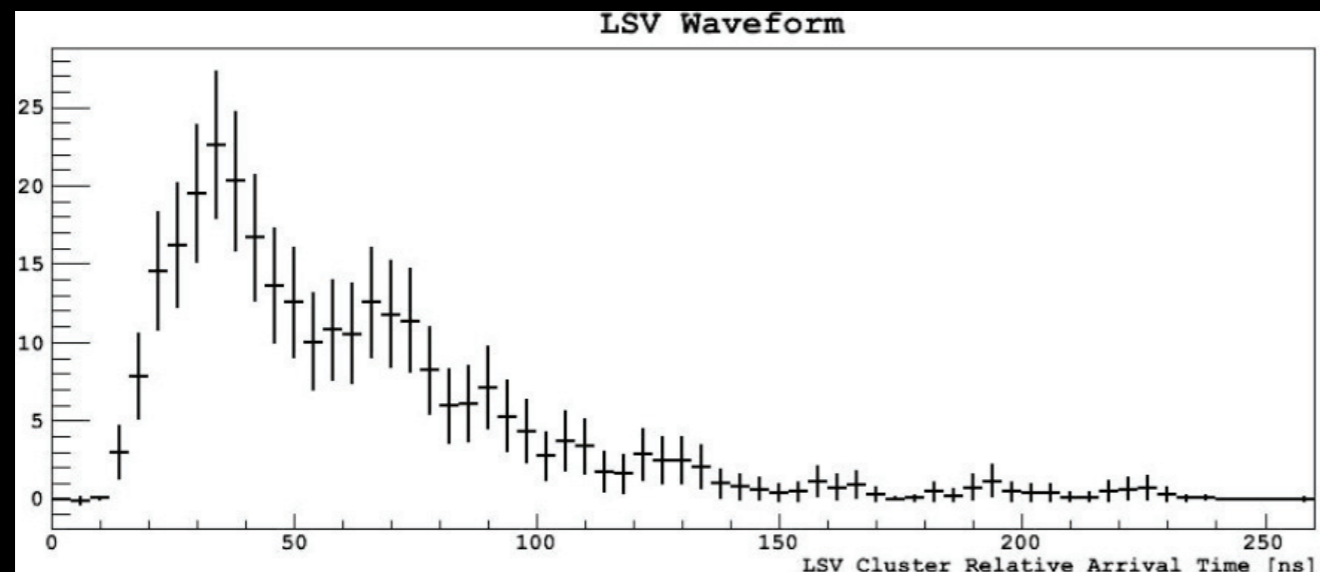


# Neutron Veto Commissioning

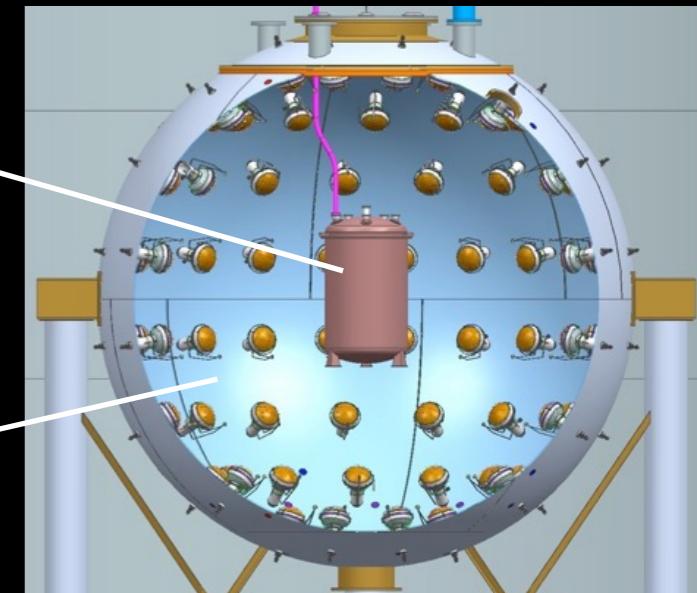
Coincident event in TPC and Neutron Veto



Electron recoil event with multiple S2 signals in TPC



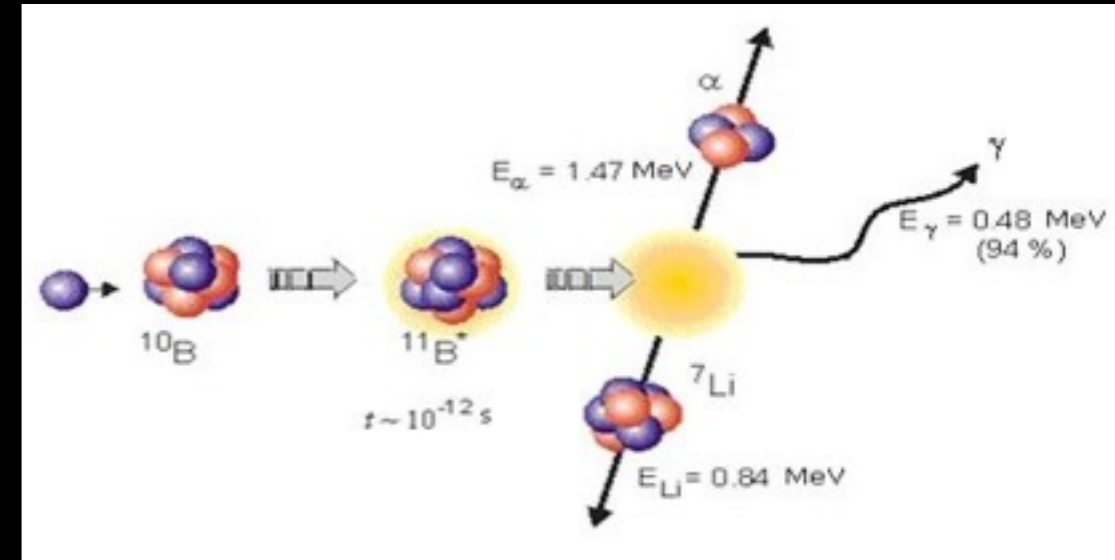
Coincident signal in liquid scintillator veto



Light Yield: liquid scintillator VETO LY of about 0.5 PE/keV<sub>ee</sub>, satisfactory for VETO requirements.

# Borated Liquid Scintillator

- High neutron capture cross section on boron allows for compact veto size
- Capture results in 1.47 MeV  $\alpha$  particle - detected with high efficiency
- Short capture time ( $2.3 \mu\text{s}$ ) reduces dead time loss



	Veto Efficiency (MC)
Radiogenic Neutrons	$> 99\%^*$
Cosmogenic Neutrons	$> 95\%$

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