

The π -axion and π -axiverse of Dark QCD

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Talk Based on:
Alexander, Gilmer, Manton, EM '23. arXiv:2304.11176
Maleknejad & EM PRD '22

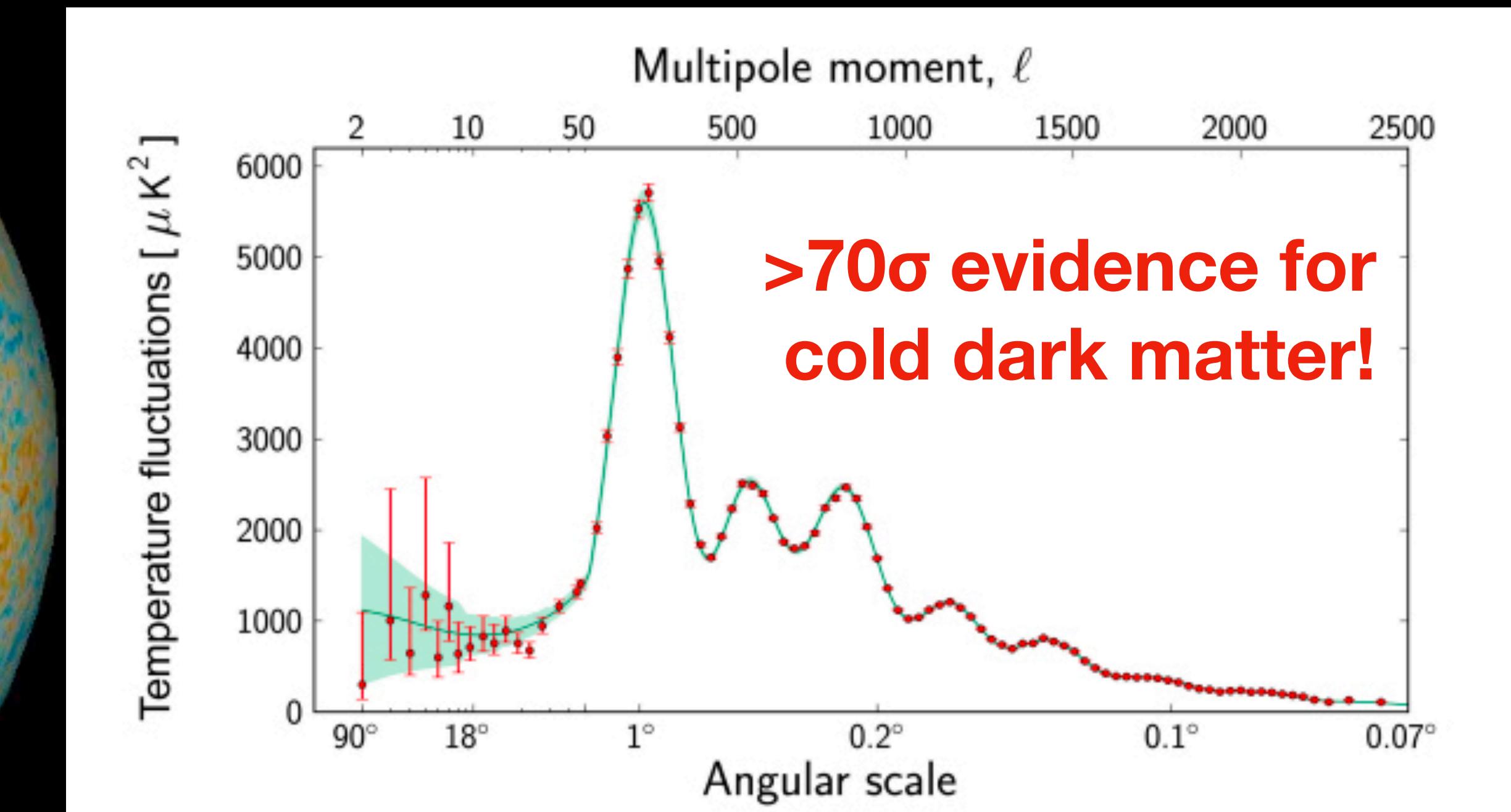
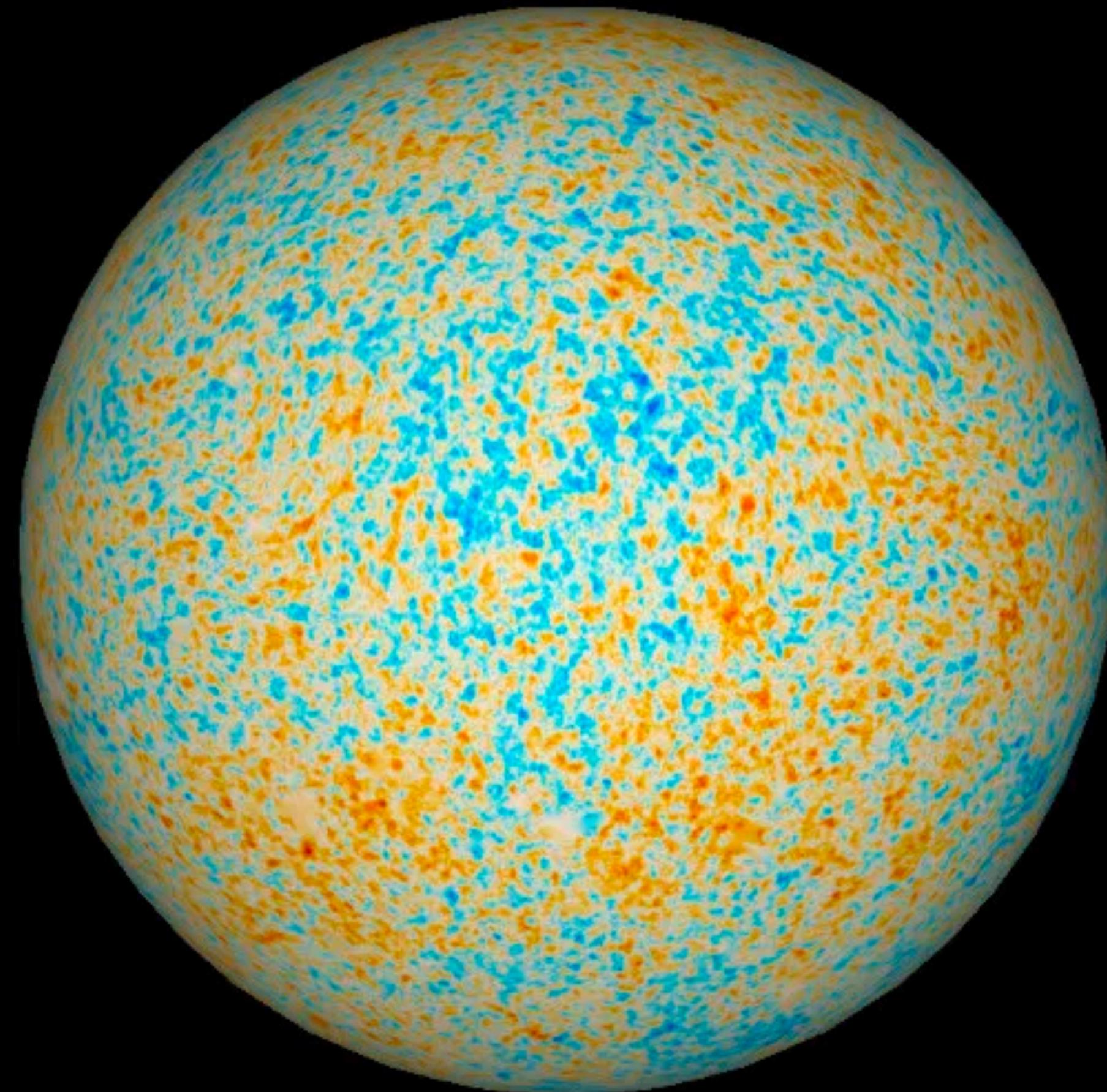
Outline

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1. Introduction: Axions past present and future
2. The π -axiverse
3. π -axion DM
4. Experimental Arenas
5. Summary: π -axions vs axions

Introduction

The case for Dark Matter



Axions circa late 1970's

The Strong CP Problem

[PQWW,
DFSZ,
KSVZ]

$$\mathcal{L} = \frac{g^2}{32\pi^2} \theta G\tilde{G} \quad \theta < 10^{-10}$$

$$\mathcal{L}(\Phi \equiv \phi e^{ia/f_a}) = \frac{1}{2} |\partial_\mu \Phi|^2 - \lambda(|\Phi|^2 - f_a^2)^2 - \Lambda^4 \cos(\theta + \frac{a}{f_a}) + \frac{a}{f_a} G\tilde{G}$$

Around the same time: Choi, Kim
Composite (dynamical, colored) axion:

Solve strong CP with:
Dark quarks charged under both dark ("axi-") color and SM color
New heavy quarks; other colored particles

Axion Dark Matter

$$\Phi = \phi e^{ia/f_a}$$

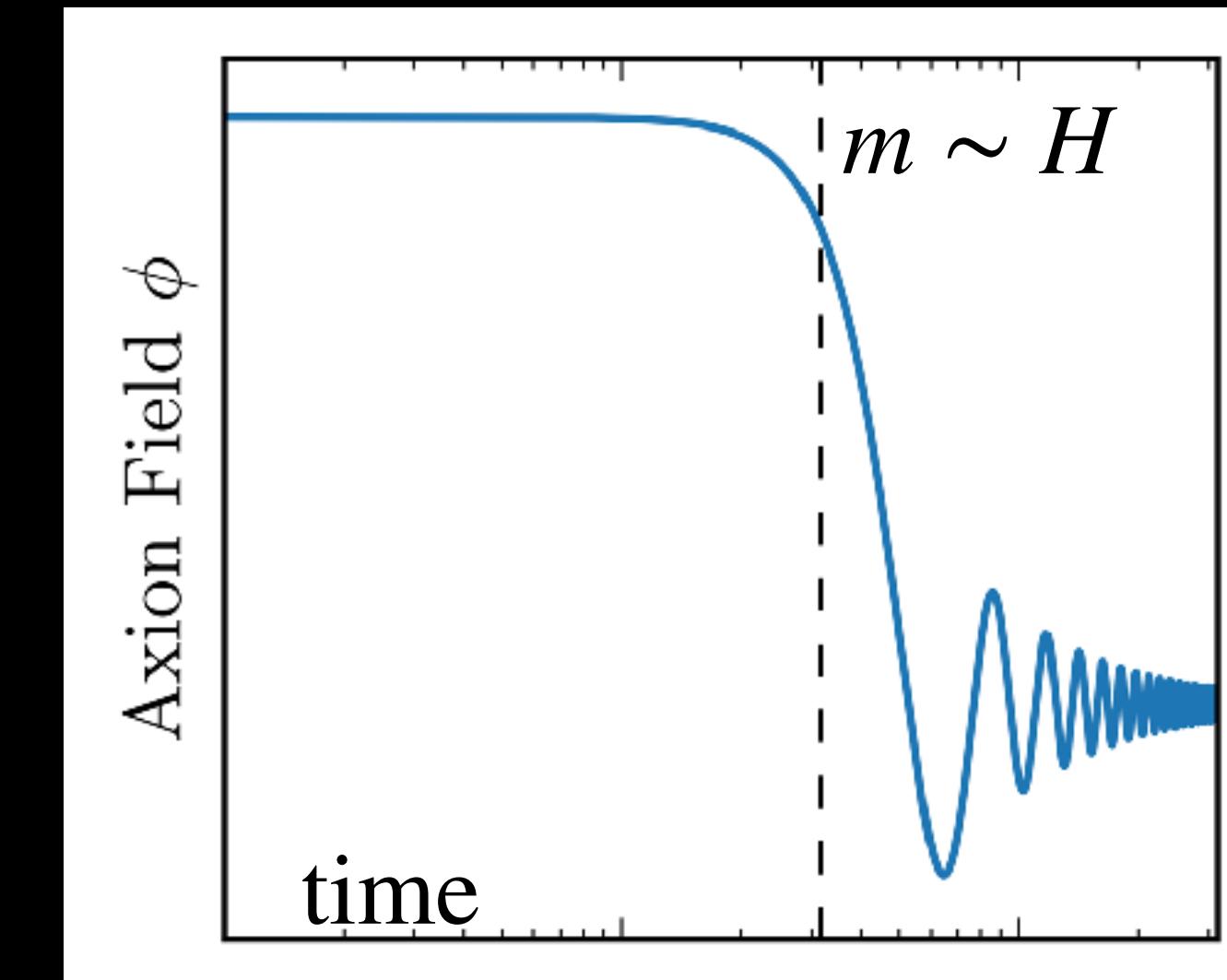
$$\mathcal{L} = |\partial_\mu \Phi|^2 - \lambda(|\Phi|^2 - f^2)^2 - \Lambda^4 \cos\left(\frac{a}{f_a}\right)$$

ALP: axion-like particle = Goldstone Boson of Spontaneous breaking of global U(1) symmetry

$$\ddot{a} + 3H\dot{a} + m^2 a = 0 \quad m \ll \text{eV}$$

See e.g. Marsh '15
For a review

Ω_{DM} set by
initial “misalignment” ϕ_i



Lots of neat particle physics:

$$\mathcal{L}_{\text{int}} = \frac{a}{f} \epsilon_{\mu\nu\rho\sigma} F^{\mu\nu} F^{\rho\sigma} = \frac{a}{f} E \cdot B$$

Axions (ALPs) from String Theory: The String Axiverse

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Arvanitaki et al '09
Goodsell, Ringwald '12

Gauge invariance in 10 dimensions: $B_{\mu\nu} \rightarrow B_{\mu\nu} + \partial_{[\mu} f_{\nu]}$

Shift symmetry of scalar in 4 dimensions: $b \rightarrow b + c$

Many axions! In type IIB:

Fundamental Axion

4-form axions

2-form axions

$$C_0 \quad \vartheta = \int_{\Sigma^4} C_4 \quad c = \int_{\Sigma^2} C_2 \quad b = \int_{\Sigma^2} B_2$$

Enter the π -axiverse

Pions: textbook physics – and a lot like axions!

Goldstone Bosons of Chiral Symmetry Breaking:

$$SU(N_f)_L \times SU(N_f)_R \rightarrow SU(N_f) \quad \Phi = \phi e^{\frac{2i\pi^{a(x)}\tau_a}{F_\pi}}$$

Pion Potential: $V(\pi) \propto \text{Tr}[M\Phi + M^\dagger\Phi^\dagger] \rightarrow \cos\left(\frac{\pi}{F_\pi}\right) \quad (M_{ij} = m_\pi \delta_{ij})$

Energy scales: $m_\pi^2 \sim \Lambda m_q$, $F_\pi \sim \Lambda$

ALP-DM-like cosmological evolution:

$$m_{\pi_0} < \text{eV}, F_\pi \gtrsim 10^{10} \text{GeV}$$

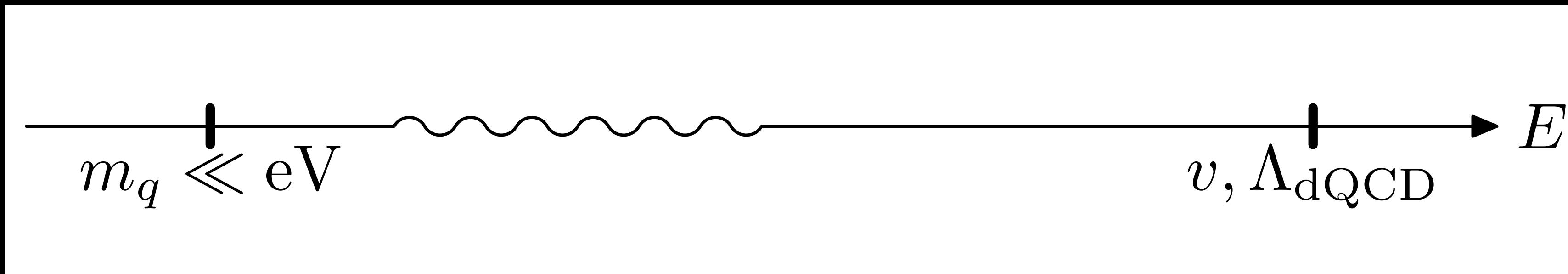
“ π -axions”

$$\Rightarrow m_q < 10^{-19} \text{eV}, \Lambda \gtrsim 10^{10} \text{GeV}$$

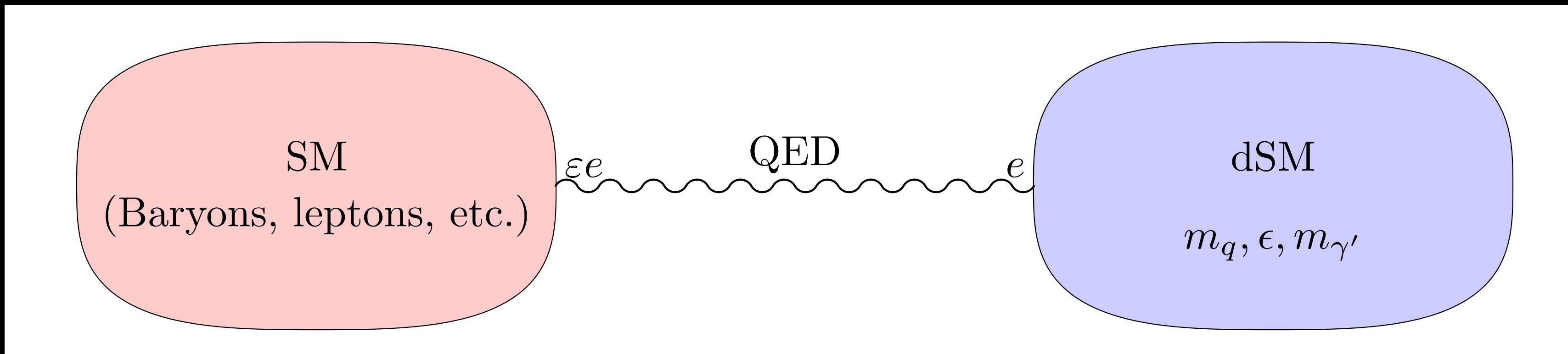
The Dark Standard Model

Key points

1. Two energy scales



2. SM portal: photon kinetic mixing (millicharges)



Dark Pions

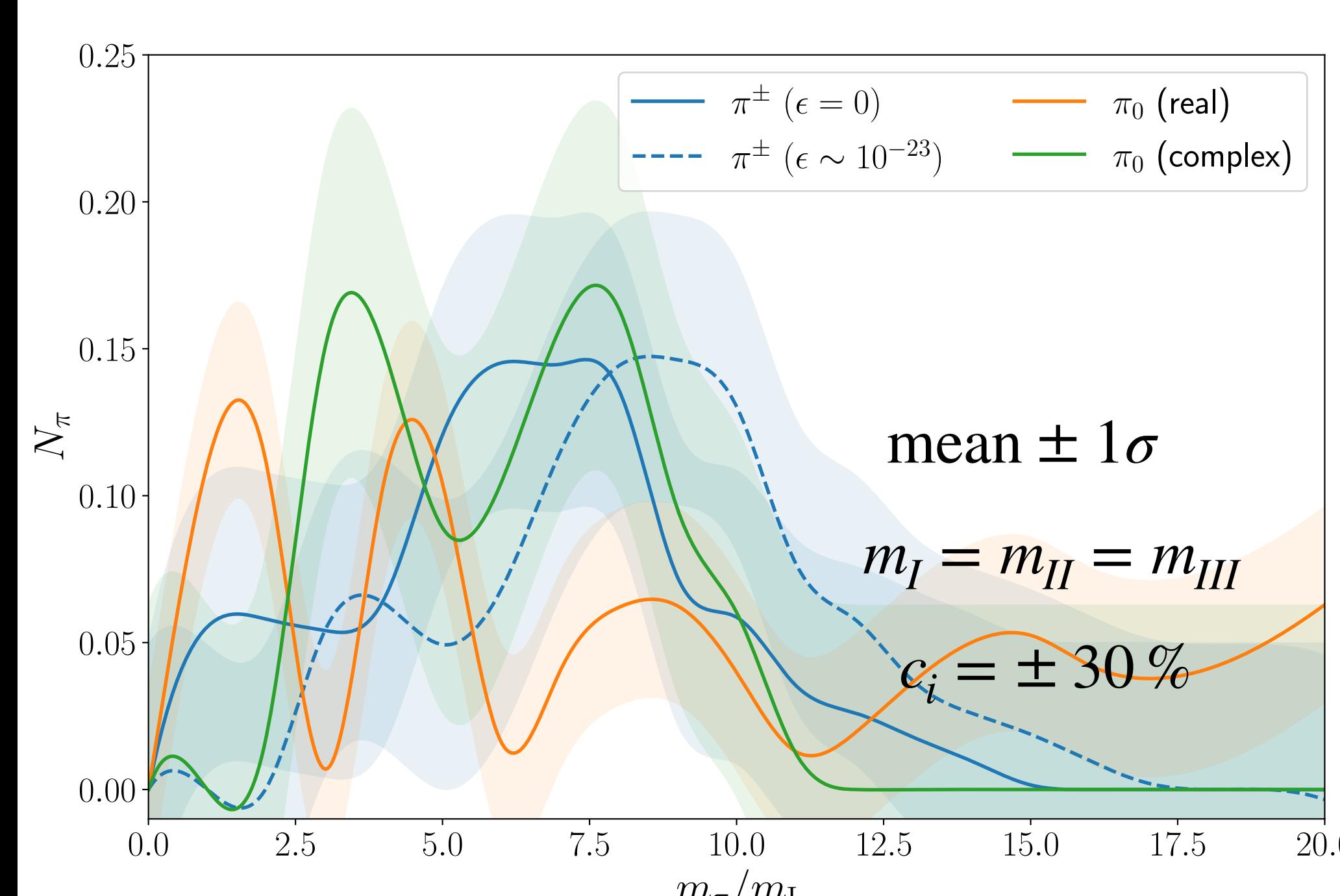
$$m_\pi^2 \simeq \frac{\langle q\bar{q} \rangle}{F_\pi^2} \sum_i m_{q_i}$$

$$\Delta m_{\pi_i^\pm}^2 \simeq \begin{cases} \alpha'_e F_\pi^2 & , \quad m_{\gamma'} < F_\pi \\ \alpha_e \epsilon^2 F_\pi^2 & , \quad m_{\gamma'} > F_\pi, \end{cases}$$

Spectrum of π -axions in Dark QCD				
#	π -axion	quark content	mass squared ($m_{\pi_i}^2$)	charge [ϵ]
5	Real Neutral:			
	π_3	$u\bar{u} - d\bar{d}$	$(c_1 + c_2)m_I F_\pi$	0
	π_8	$u\bar{u} + d\bar{d} - 2s\bar{s}$	$((c_1 + c_2)m_I + c_3 m_{II}) F_\pi$	0
	π_{29}	$c\bar{c} - b\bar{b}$	$(c_4 m_{II} + c_5 m_{III}) F_\pi$	0
	π_{34}	$c\bar{c} + b\bar{b} - 2t\bar{t}$	$(c_4 m_{II} + (c_5 + c_6) m_{III}) F_\pi$	0
	π_{35}	$-u\bar{u} - d\bar{d} - s\bar{s} + c\bar{c} + b\bar{b} + t\bar{t}$	$((c_1 + c_2)m_I + (c_3 + c_4)m_{II} + (c_5 + c_6)m_{III}) F_\pi$	0
6	Complex Neutral:			
	$\pi_6 \pm i\pi_7$	$d\bar{s}/ds$	$(c_2 m_I + c_3 m_{II}) F_\pi$	0
	$\pi_9 \pm i\pi_{10}$	$u\bar{c}/\bar{u}c$	$(c_1 m_I + c_4 m_{II}) F_\pi$	0
	$\pi_{17} \pm i\pi_{18}$	$d\bar{b}/\bar{d}b$	$(c_2 m_I + c_5 m_{III}) F_\pi$	0
	$\pi_{19} \pm i\pi_{20}$	$s\bar{b}/\bar{s}b$	$(c_3 m_{II} + c_5 m_{III}) F_\pi$	0
	$\pi_{21} \pm i\pi_{22}$	$u\bar{t}/\bar{u}t$	$(c_1 m_I + c_6 m_{III}) F_\pi$	0
	$\pi_{30} \pm i\pi_{31}$	$c\bar{t}/\bar{c}t$	$(c_4 m_{II} + c_6 m_{III}) F_\pi$	0
9	Charged:			
	$\pi_1 \pm i\pi_2$	$ud/\bar{u}\bar{d}$	$(c_1 + c_2)m_I F_\pi + 2\xi_1(e\epsilon F_\pi)^2$	± 1
	$\pi_4 \pm i\pi_5$	$u\bar{s}/\bar{u}s$	$(c_1 m_I + c_3 m_{II}) F_\pi + 2\xi_2(e\epsilon F_\pi)^2$	± 1
	$\pi_{15} \pm i\pi_{16}$	$u\bar{b}/\bar{u}b$	$(c_1 m_I + c_5 m_{III}) F_\pi + 2\xi_3(e\epsilon F_\pi)^2$	± 1
	$\pi_{11} \pm i\pi_{12}$	$d\bar{c}/\bar{d}c$	$(c_2 m_I + c_4 m_{III}) F_\pi + 2\xi_4(e\epsilon F_\pi)^2$	∓ 1
	$\pi_{23} \pm i\pi_{24}$	$d\bar{t}/\bar{d}t$	$(c_2 m_I + c_6 m_{III}) F_\pi + 2\xi_5(e\epsilon F_\pi)^2$	∓ 1
	$\pi_{13} \pm i\pi_{14}$	$s\bar{c}/\bar{s}c$	$(c_3 + c_4)m_{II} F_\pi + 2\xi_6(e\epsilon F_\pi)^2$	∓ 1
	$\pi_{25} \pm i\pi_{26}$	$s\bar{t}/\bar{s}t$	$(c_3 m_{II} + c_6 m_{III}) F_\pi + 2\xi_7(e\epsilon F_\pi)^2$	∓ 1
	$\pi_{27} \pm i\pi_{28}$	$c\bar{b}/\bar{c}b$	$(c_4 m_{II} + c_5 m_{III}) F_\pi + 2\xi_8(e\epsilon F_\pi)^2$	± 1
	$\pi_{32} \pm i\pi_{33}$	$b\bar{t}/\bar{b}t$	$(c_5 + c_6)m_{III} F_\pi + 2\xi_9(e\epsilon F_\pi)^2$	∓ 1

Quark masses: $c_1 m_I$, $c_2 m_I$, $c_3 m_{II}$, $c_4 m_{II}$, $c_5 m_{III}$, $c_6 m_{III}$

Statistical Distribution of masses



Photon Portal to the SM:

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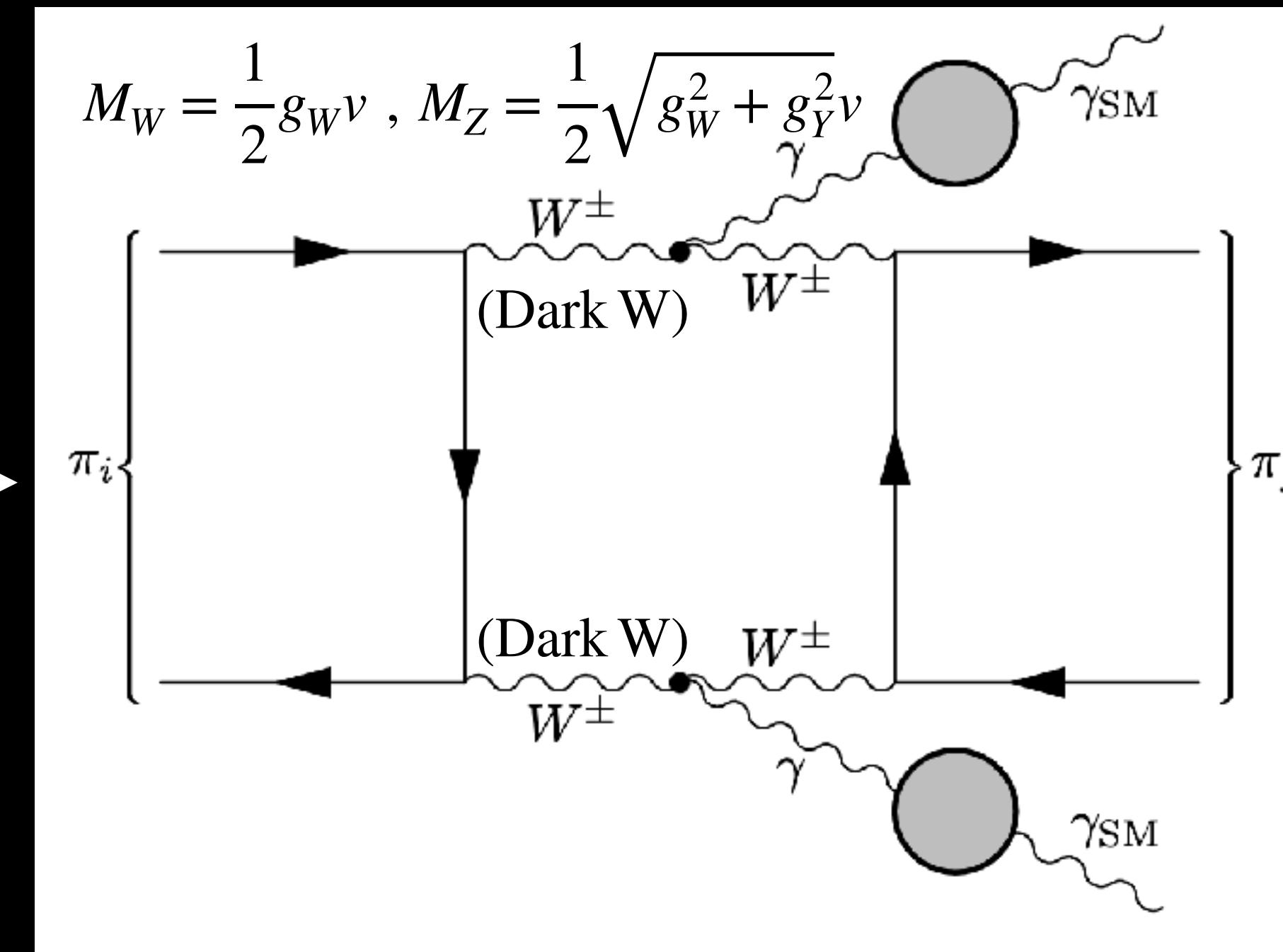
$\epsilon F_{\mu\nu} \tilde{F}^{\mu\nu} \Rightarrow$ millicharges, π -axion–photon couplings

1. Neutral scalar pion: $\mathcal{L} = \lambda_1 \frac{\epsilon^2}{F_\pi} \pi^0 F \tilde{F}$

2. Charged pion: $\mathcal{L} \sim \epsilon^2 \pi^+ \pi^- A_\mu A^\mu$

3. Flavor violating:

$$\mathcal{L} = \lambda \frac{\epsilon^2}{M^2} \pi_i \pi_j^* F_{\mu\nu} F^{\mu\nu} + h.c.$$



π -axion Lifetime:

$$\tau_{\pi^0} \sim H_0^{-1} \left(\frac{F_\pi}{\text{TeV}} \right)^2 \left(\frac{0.3\text{eV}}{m_{\pi^0}} \right)^3 \frac{1}{\epsilon^4}$$

π -axion DM

Misalignment Relic Density

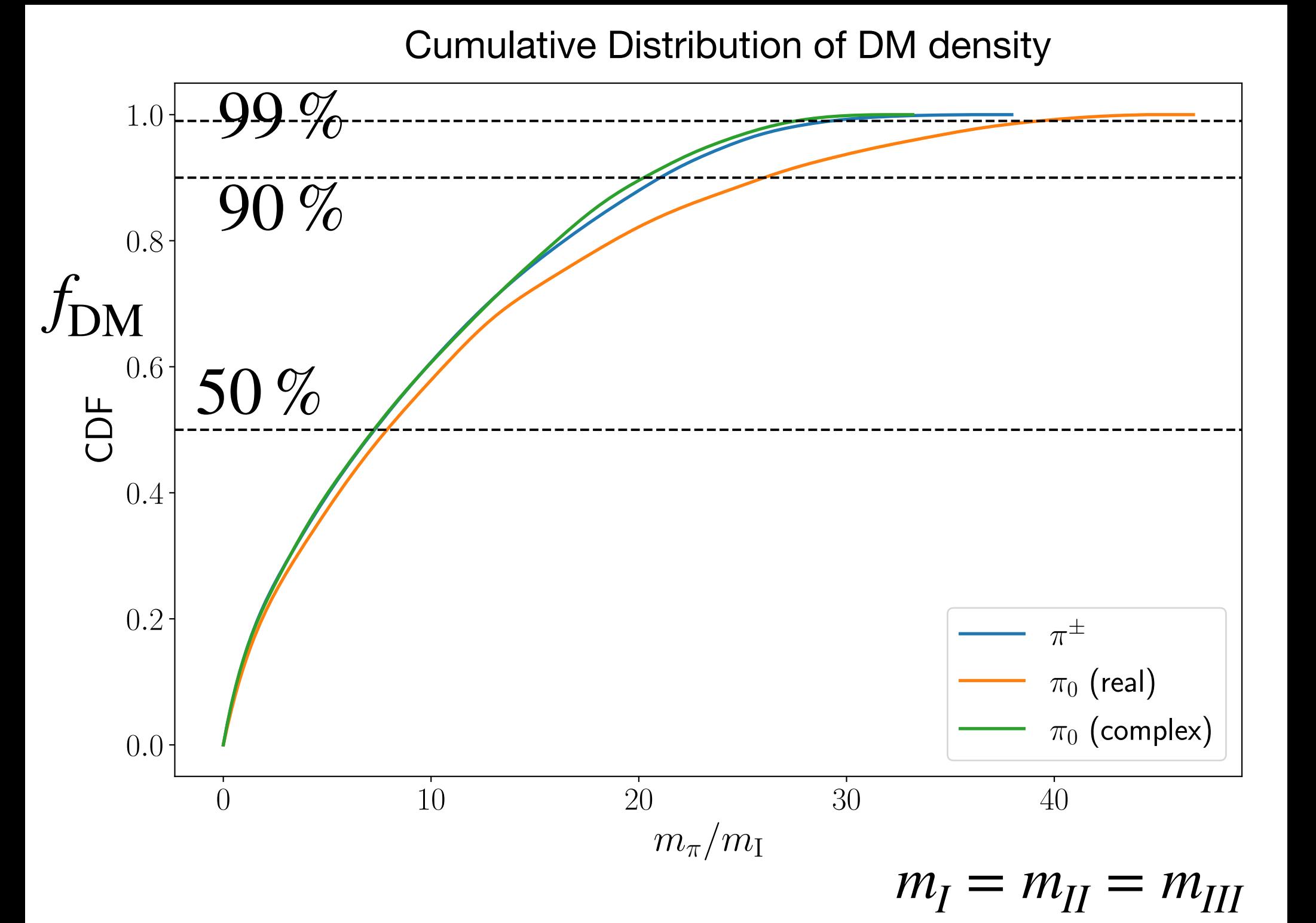
Scenario:

- Symmetry breaking before inflation
- CMB Isocurvature: $H_{\text{inf}} \lesssim 10^{10}(F_\pi/M_P) \text{ GeV}$

$$\text{Cosmological stability: } \tau_{\pi^0} \sim H_0^{-1} \left(\frac{F_\pi}{\text{TeV}} \right)^2 \left(\frac{0.3 \text{eV}}{m_{\pi^0}} \right)^3 \frac{1}{\epsilon^4}$$

- Heavy Dark Photon: $m_{\gamma'} > F_\pi$
- Late-time relic density produced by misalignment:

$$\Omega_{\pi_i} = \frac{1}{6} (9\Omega_r)^{\frac{3}{4}} \frac{F_\pi^2}{M_p^2} \sqrt{\frac{m_i}{H_0}} \theta_i^2$$

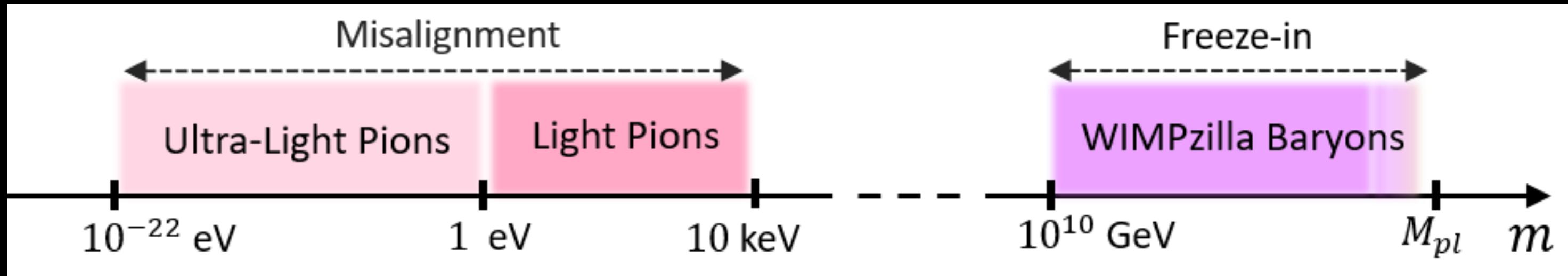


$$m_I = m_{II} = m_{III}$$

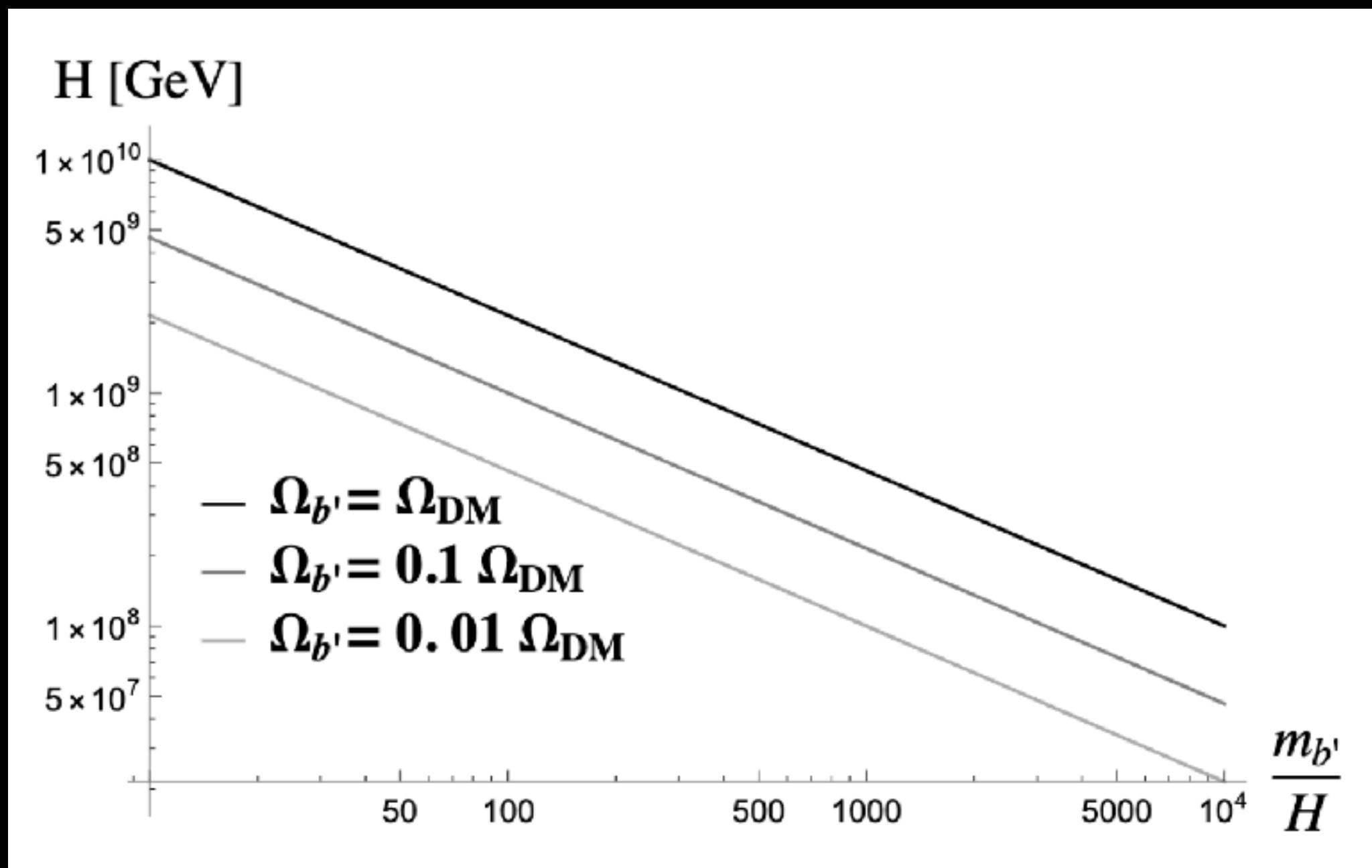
The WIMPzilla Connection: Dark Baryons

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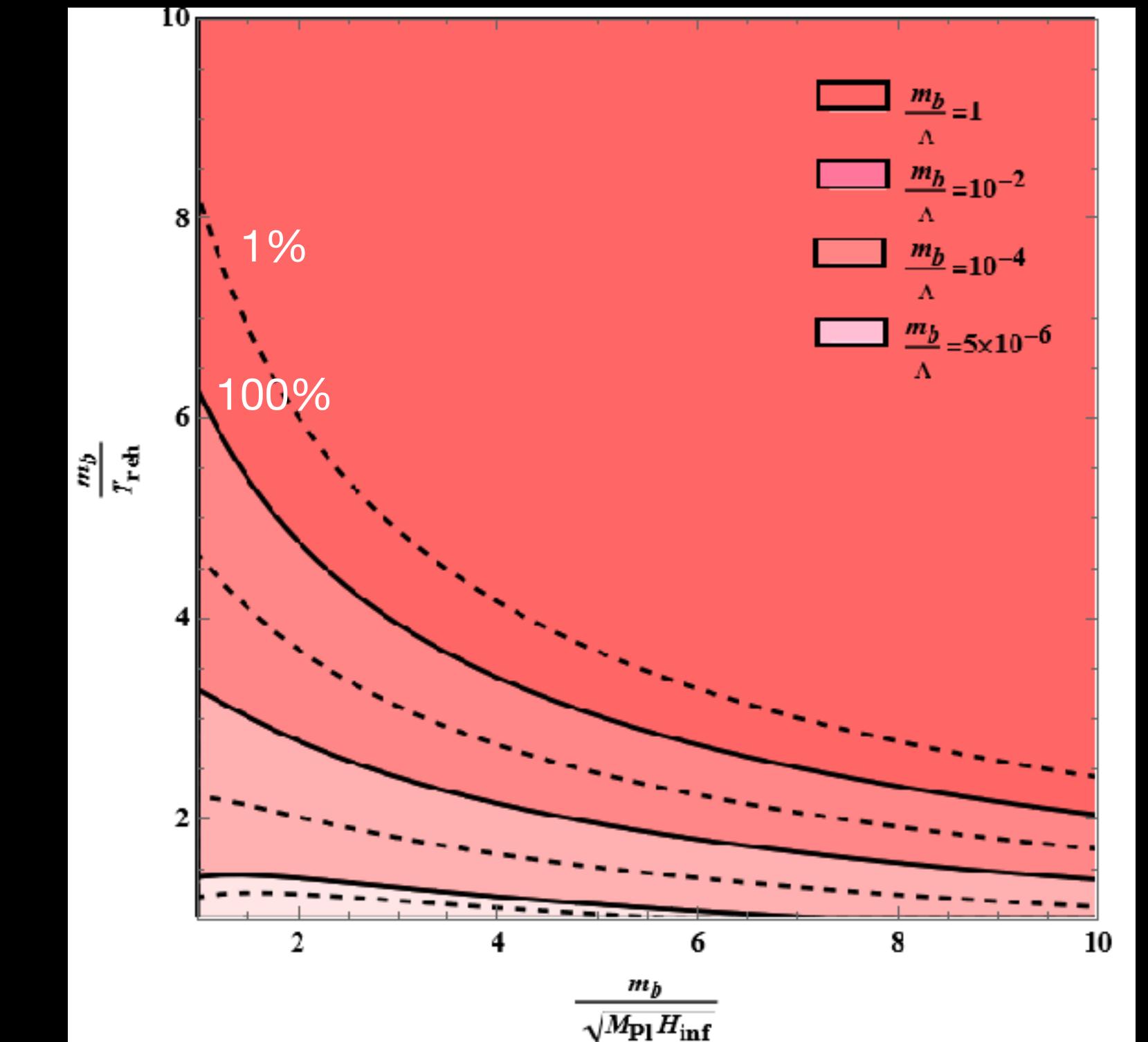
Maleknejad, EM '22



Gravitational Production
from Inflation:



Freeze-In Production:
Higgs-, QED-, or inflaton- portal



Experimental Arenas

Parity-odd portal: $\mathcal{L} \sim g_{\pi\gamma\gamma} \pi F \tilde{F}$

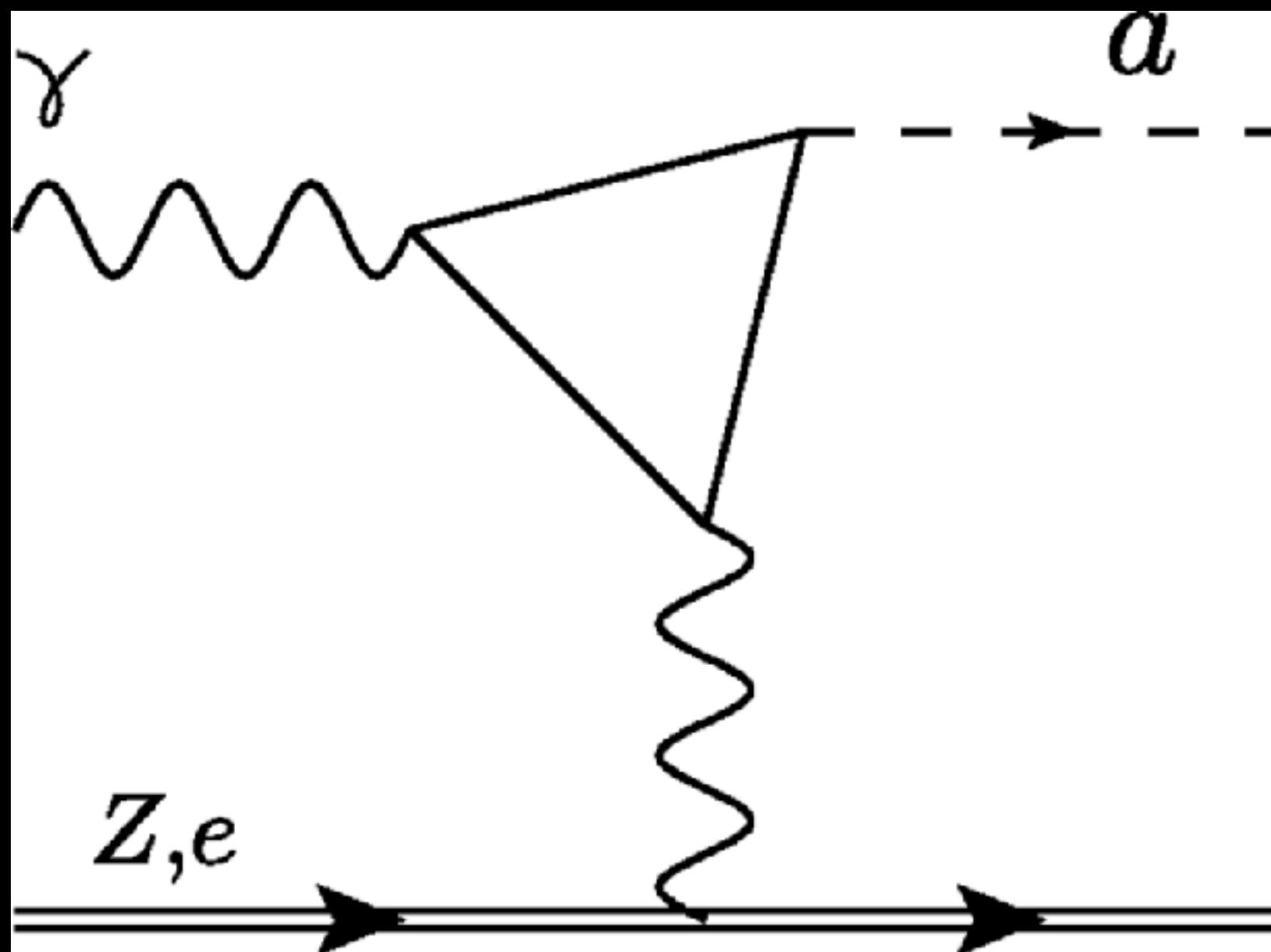
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Recent review:

Adams et al Axion DM Snowmass

Experiments such as ADMX

- Axion-photon couplings in strong B-field background



Neutral pseudo-scalar pions have coupling:

$$g_{\pi\gamma\gamma} \sim c \frac{\alpha_e \varepsilon^2}{F_\pi}$$

Five neutral π -axions

\Rightarrow Multiple distinct resonances

But ε —suppressed
relative to conventional axion

Note:

$$\tau_{\pi^0} \sim H_0^{-1} \left(\frac{F_\pi}{\text{TeV}} \right)^2 \left(\frac{0.3\text{eV}}{m_{\pi^0}} \right)^3 \frac{1}{\varepsilon^4} \Rightarrow \text{Can have } \varepsilon = \mathcal{O}(1) !$$

$$m_\pm \sim \varepsilon F_\pi$$

Parity-even portal

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Experiments such as **Atomic Clocks** **Arvanitaki, Huang, Tilburg '15**

$$\mathcal{L} = \lambda \frac{\varepsilon^2}{M^2} |\pi_i|^2 F_{\mu\nu} F^{\mu\nu} + h.c.$$

$$\pi \sim \pi_0 \sin(mt + \delta) \Rightarrow \alpha_e(t) = \alpha_e \left(1 + \frac{2\lambda e^2}{\Lambda^2} \varepsilon^2 \sum_i |\pi_{i,0}|^2 \sin^2(m_i t + \delta_i) \right)$$

- Multiple (light) fields \Rightarrow multiple incoherent contributions
- Benchmark needed for detection:

$$\frac{\rho_{\text{DM}}^i \varepsilon^2}{M^2 m_{\pi_i}^2} \gtrsim 10^{-15}$$

π -axion Star Explosions (Via Parametric Resonance)

Amin, Mou '20

Amin, Mou, Saffin 21

Du et al. '23 ("Axion Star Explosions")

Chung-Jukko et al. '23

$$A''_{\pm} + (k^2 + B_{\pm}(t)k + C(t)) A_{\pm} = 0$$

$$\begin{aligned} B_{\pm}(t) &= \frac{\lambda_2}{\Lambda_2^2} \varepsilon^2 \sum_{i,j}^9 \pi_{i,0}^c \pi_{j,0}^c (2 \cos(\theta_i - \theta_j)) \left[m_i \cos \varphi_i(t) \sin \varphi_j(t) + m_j \sin \varphi_i(t) \cos \varphi_j(t) \right] \\ &\quad + \frac{\lambda_4}{\Lambda_4^2} \varepsilon^2 \left\{ \sum_{i,j}^6 \pi_{i,0}^c \pi_{j,0}^c (2 \cos(\theta_i - \theta_j)) + \sum_{i,j}^5 \pi_{i,0}^r \pi_{j,0}^r \right\} \left[m_i \cos \varphi_i(t) \sin \varphi_j(t) + m_j \sin \varphi_i(t) \cos \varphi_j(t) \right] \\ &\quad \pm \frac{\lambda_3}{F_{\pi}} \varepsilon^2 \sum_i^5 \pi_{i,0}^r m_i \cos \varphi_i(t), \\ C(t) &= \lambda_1 \varepsilon^2 e^2 \sum_{i,j}^9 \pi_{i,0}^c \pi_{j,0}^c \cos(\theta_i - \theta_j) \sin \varphi_i(t) \sin \varphi_j(t) \end{aligned}$$

Even just the simple charged-pion coupling, $|\pi^{\pm}|^2 A_{\mu} A^{\mu}$
can dramatically enhance parameter resonance

Jaeckel, Schenk '21

Summary

π -Axion

Axion

confining gauge theory, Chiral symmetry breaking	complex scalar, spontaneously broken global U(1) symmetry
Many pions, mass splitting due to charges	1 axion per complex scalar
WIMPZilla: dark baryons	WIMPZilla: radial field
Real and complex neutral, and charged	Real neutral
Other degrees of freedom: dark Electroweak , glueballs ...	Other degrees of freedom: Model dependent
Interactions: Parity-odd and parity-even	Parity-odd couplings

$$(e.g. |\pi^\pm|^2 A_\mu A^\mu)$$

Q: Smoking guns of π -axions?

A:

- Combination of parity-odd and parity even couplings, each benchmarked to the millicharge and decay constant
- Heavy dark baryons with mass $\sim \pi$ -axion decay constant
- The spectrum of π -axions: tightly packed discretum (e.g. 5 neutral pi-axions); combination of real scalars,

Take-Home Messages

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- SM Physics can provide an ‘axiverse’: the π -axiverse of a dark SM
 - Dark SM gets you the millicharges, flavor violation, etc.
- DM candidate: π -axion DM
 - DM can be mixture of real scalar, neutral complex scalars, and millicharged complex scalars
 - Natural theory expectation is tightly packed discretum of masses
- Detection: mixture of parity-odd and parity-even signals

Lots to do!

1. Constraints Plot! Statistical distribution of predictions
2. Millicharge signatures
3. Imprint on CMB and Large Scale Structure

Thanks!

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