

CAP Congress 2023 - UNB Fredericton - June 2023

# Light Sub-GeV Dark Matter at Accelerators

Adam Ritz

University of Victoria



Work over several years with B. Batell, A. Berlin, P. deNiverville,  
S. Foroughi, D. McKeen, M. Pospelov, P. Schuster, N. Toro

CAP Congress 2023 - UNB Fredericton - June 2023

# (A Decade of) Light Sub-GeV Dark Matter at Accelerators\*

Adam Ritz

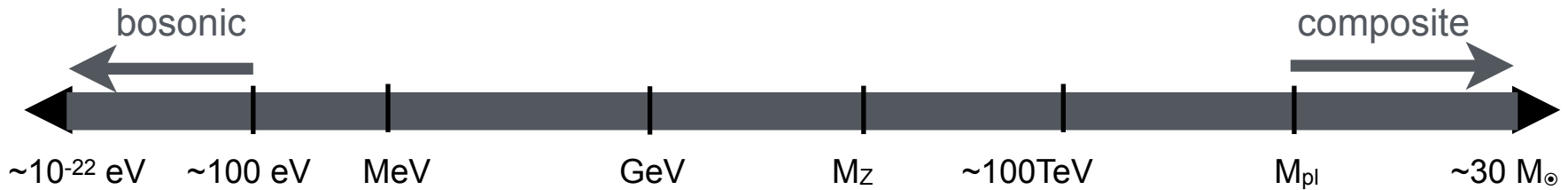
University of Victoria



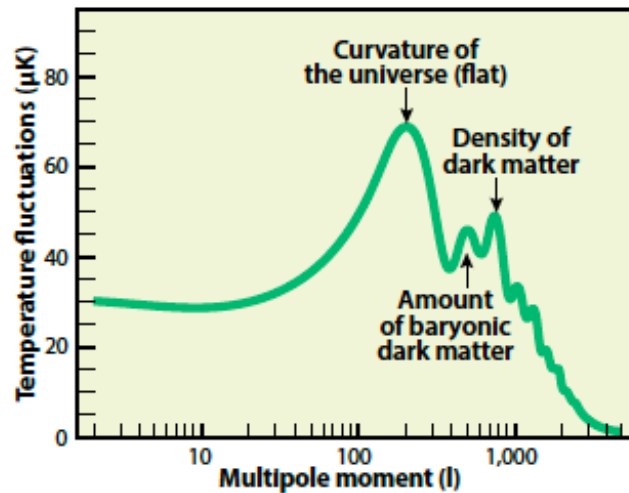
Work over several years with B. Batell, A. Berlin, P. deNiverville,  
S. Foroughi, D. McKeen, M. Pospelov, P. Schuster, N. Toro

\* Disclaimer for historical incompleteness!

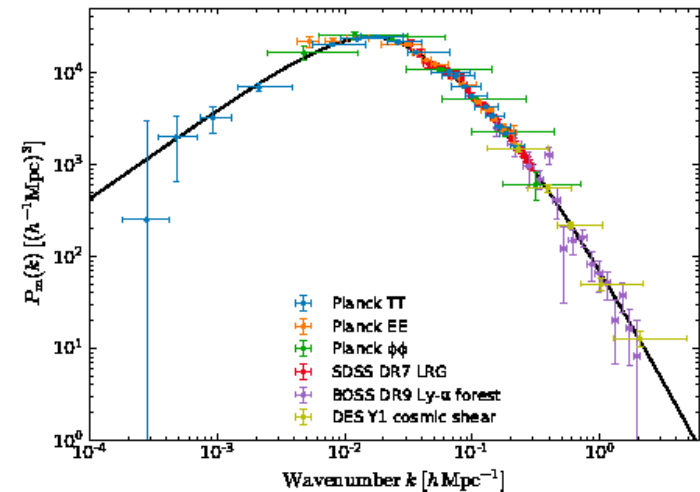
# Cold dark matter landscape



- Gravitational evidence for DM from multiple cosmological & astrophysical scales (CMB, LSS, Lensing, etc)
- Empirical evidence for dark matter (and neutrino mass) arguably points to a dark/hidden sector (but not directly to a specific mass scale)

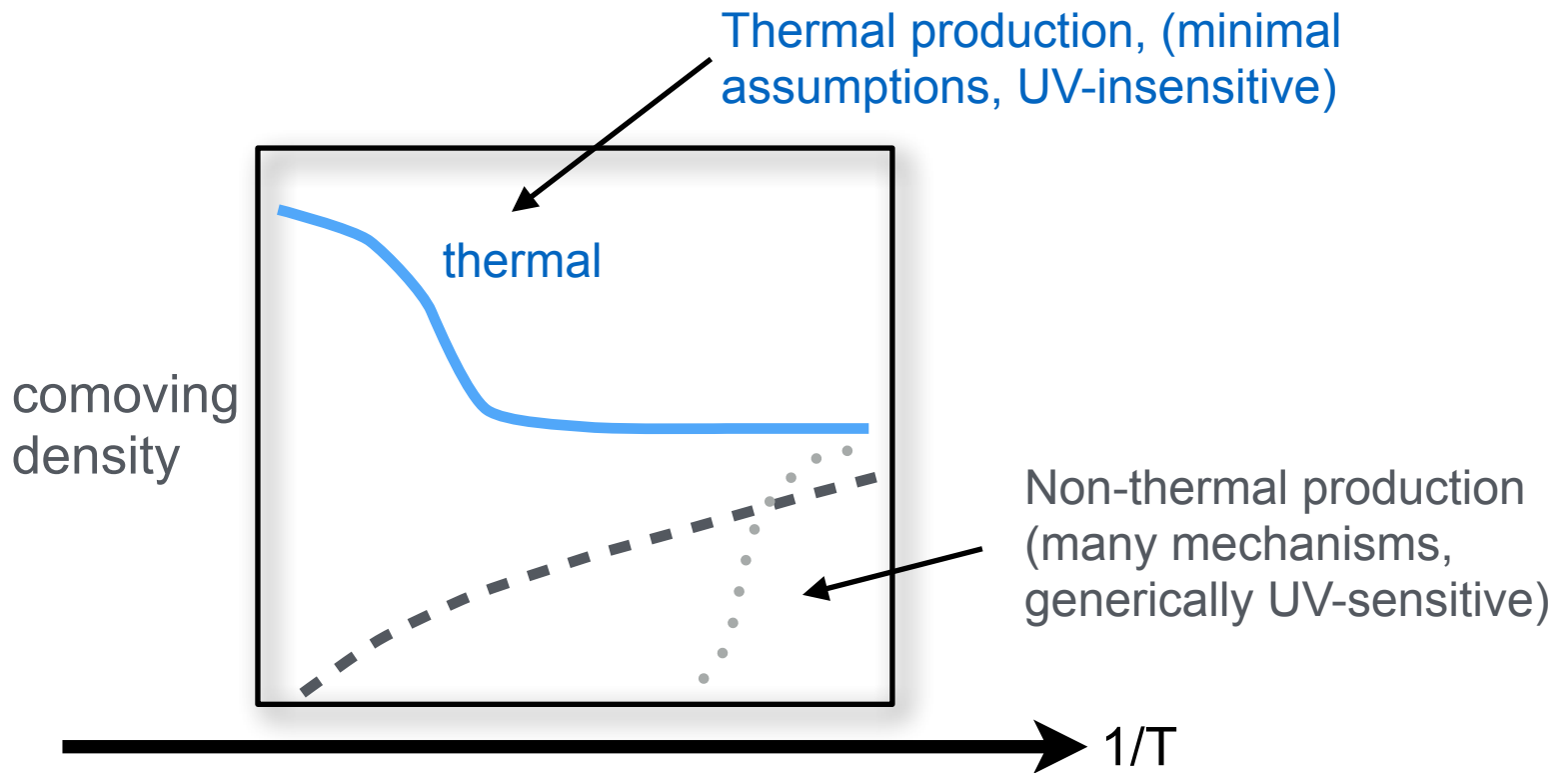
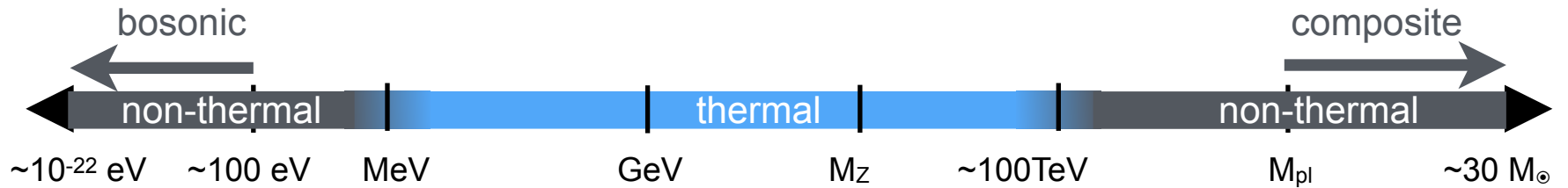


- $\rho_{\text{CDM}} \sim 5\rho_{\text{baryons}}$
- Cold enough...
- Dark enough...
- Structure growth prior to recombination...



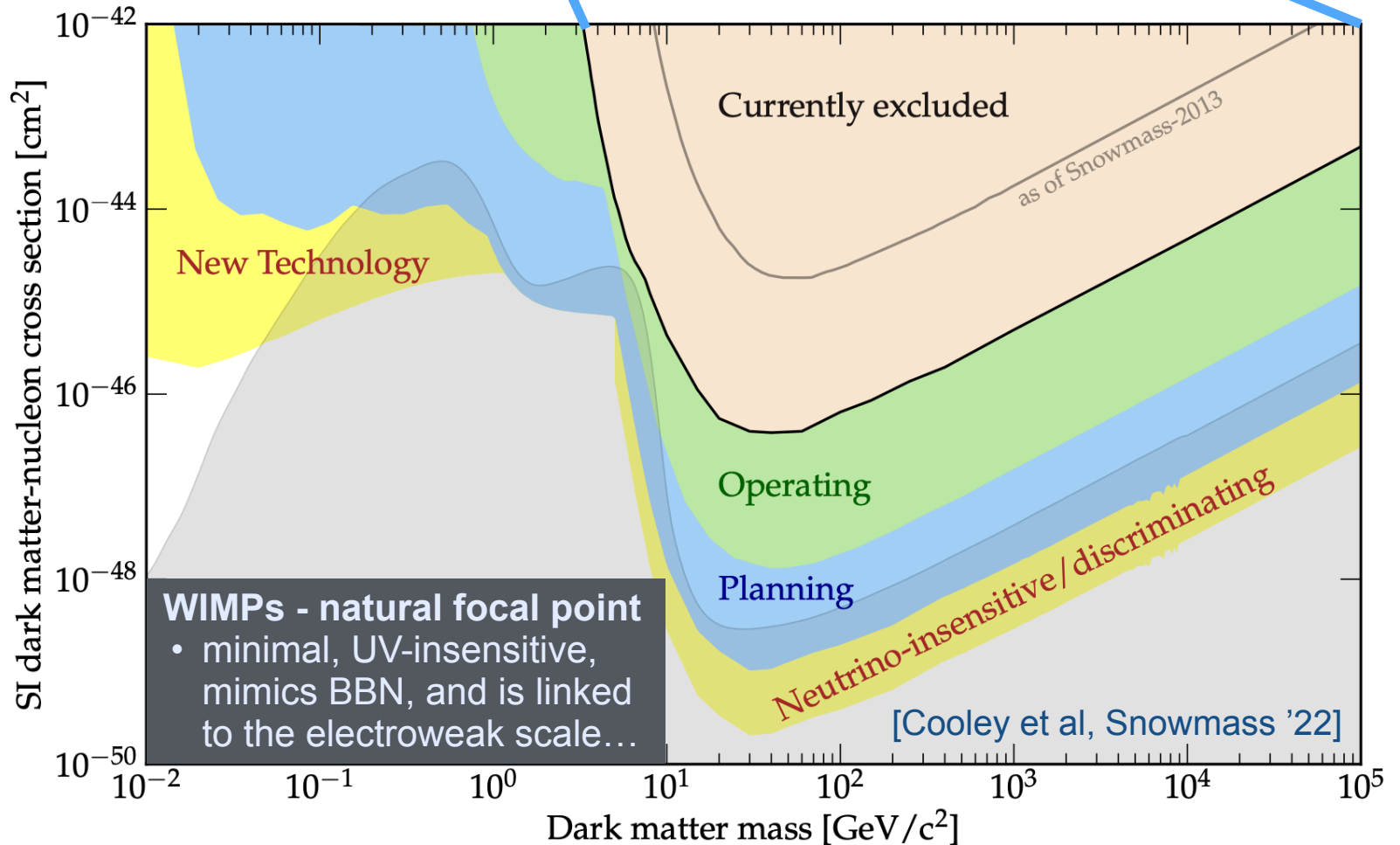
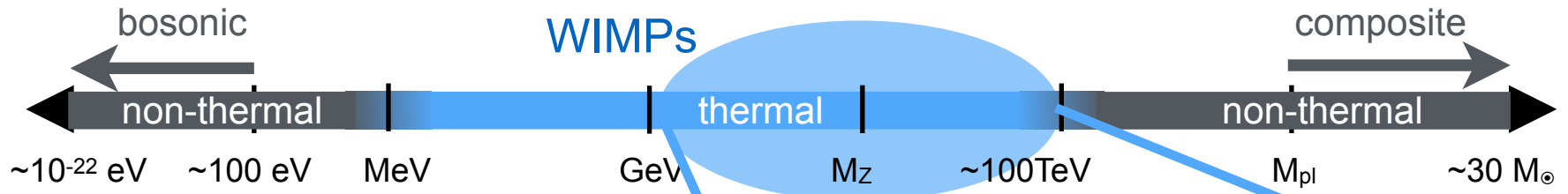
⇒ Empirical evidence for DM (and neutrino mass) allows a vast mass range for new physics, so what are the appropriate theory priors?

# Cold dark matter landscape

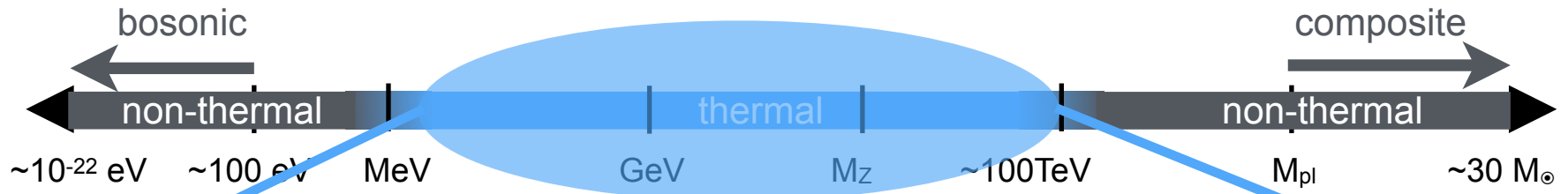


DM thermal history in the early universe

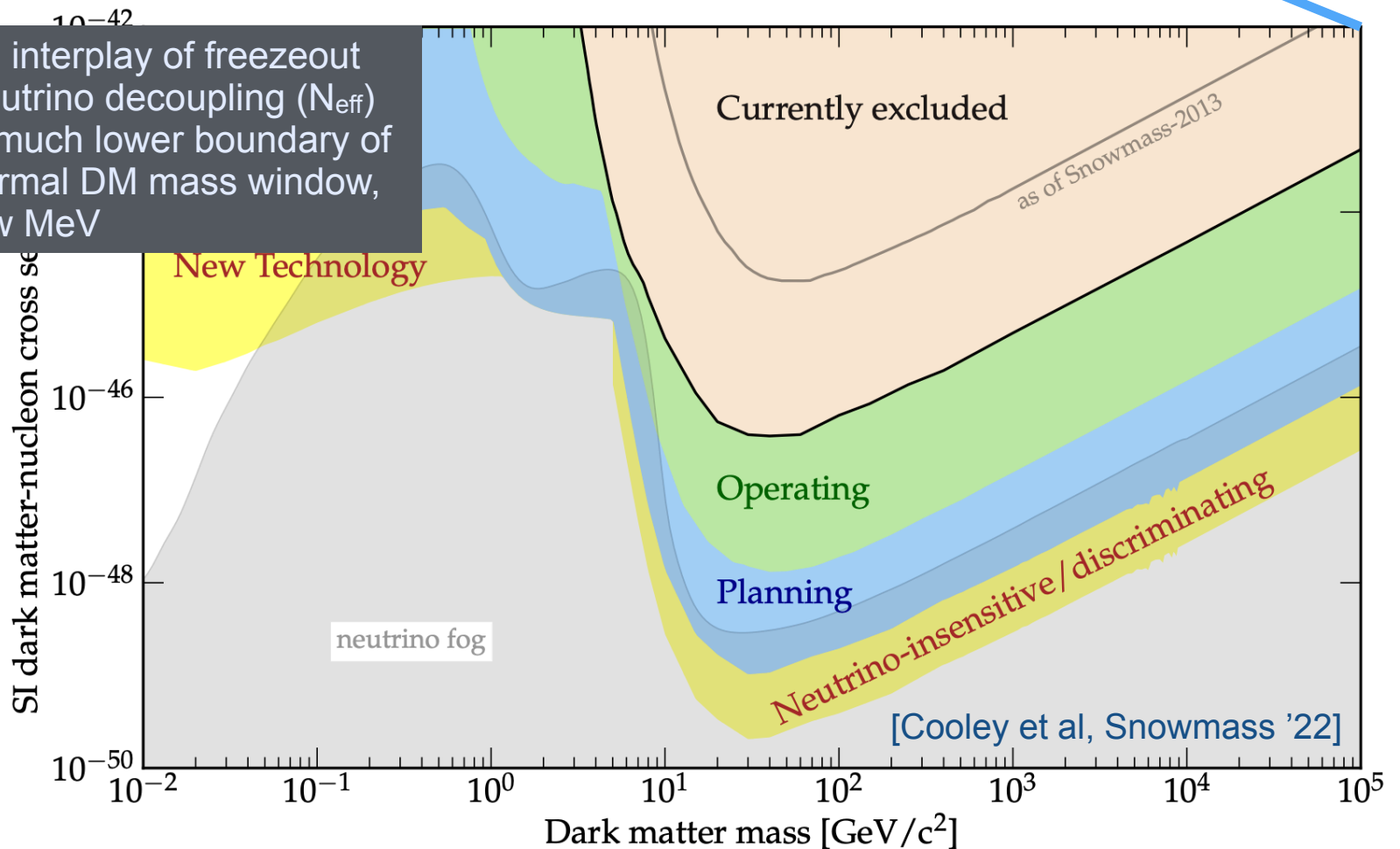
# Cold dark matter landscape



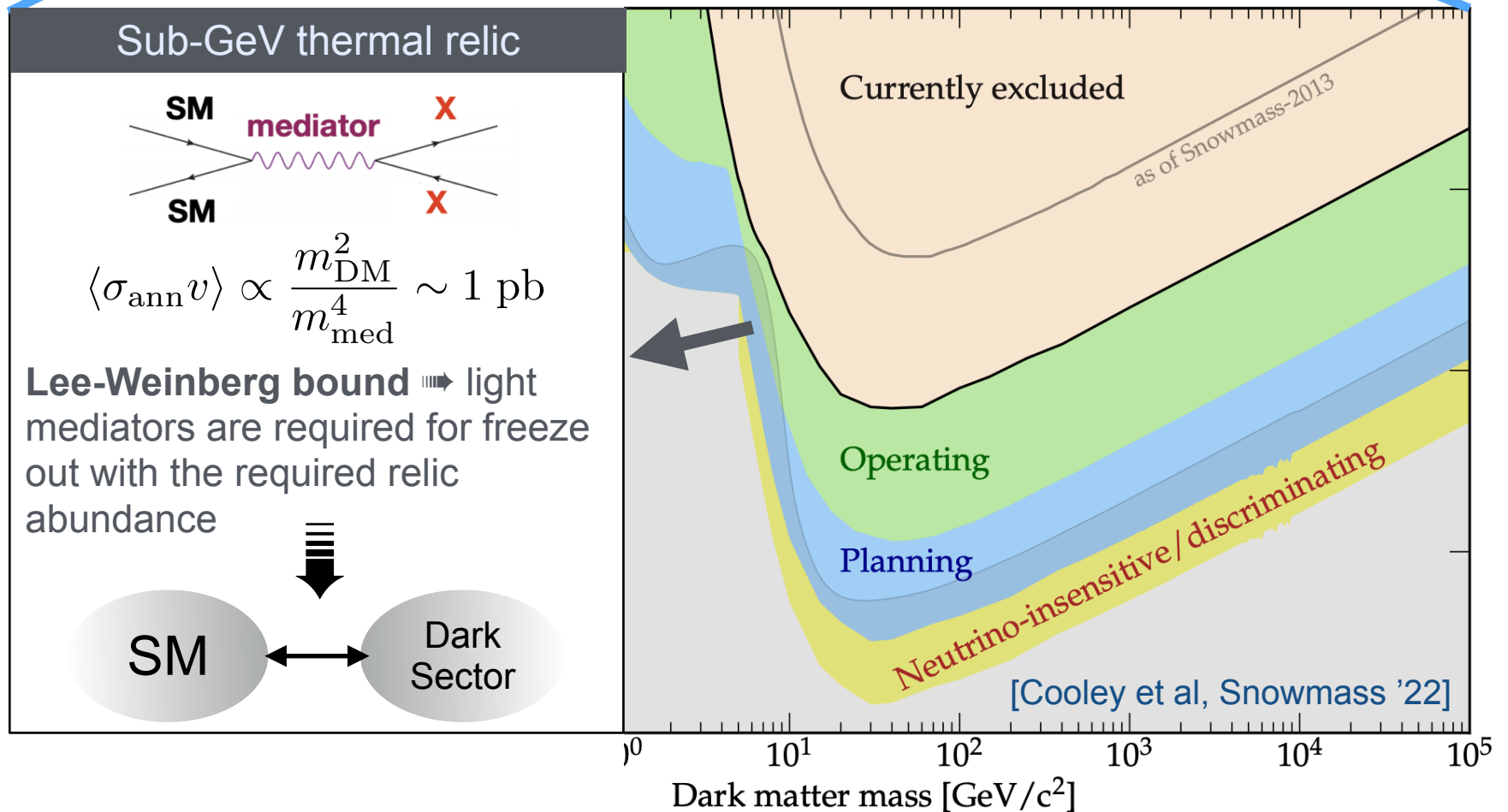
# Cold dark matter landscape



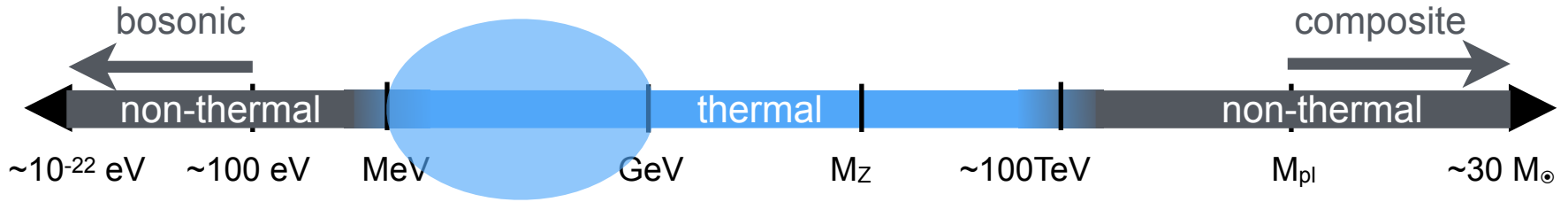
But the interplay of freezeout with neutrino decoupling ( $N_{\text{eff}}$ ) sets a much lower boundary of the thermal DM mass window, at a few MeV



# Cold dark matter landscape



# Cold dark matter landscape

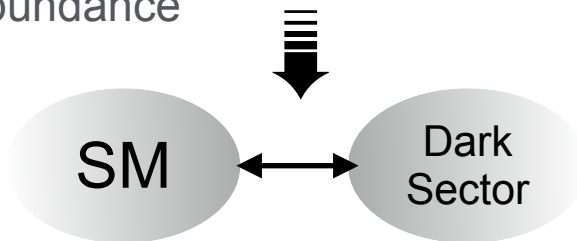


## Sub-GeV thermal relic



$$\langle \sigma_{\text{ann}} v \rangle \propto \frac{m_{\text{DM}}^2}{m_{\text{med}}^4} \sim 1 \text{ pb}$$

**Lee-Weinberg bound**  $\Rightarrow$  light mediators are required for freeze out with the required relic abundance



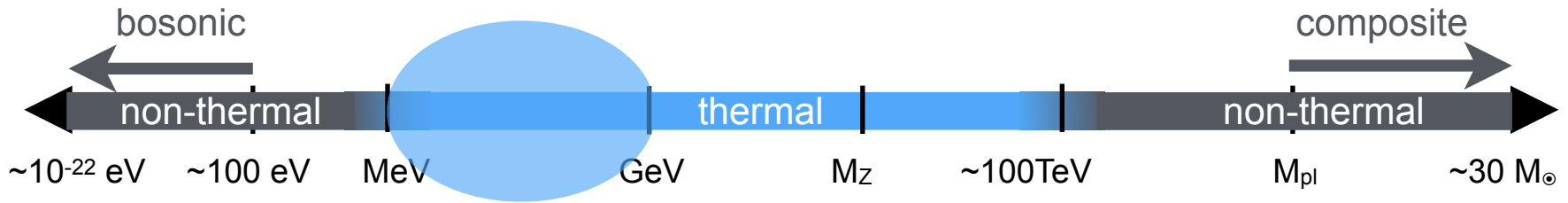
## EFT for mediators to a (neutral) dark sector

There are just three UV-complete relevant or marginal “portals” to a SM-neutral dark sector, unsuppressed by a (possibly large) new physics scale  $\Lambda$

$$\begin{aligned} \mathcal{L} &= \sum_{n=k+l-4} \frac{c_n}{\Lambda^n} \mathcal{O}_k^{(\text{SM})} \mathcal{O}_l^{(\text{med})} = \mathcal{L}_{\text{portals}} + \mathcal{O}\left(\frac{1}{\Lambda}\right) \\ &= -\frac{\epsilon}{2} B^{\mu\nu} \underbrace{A'_{\mu\nu}}_{\text{Vector portal}} - H^\dagger H \underbrace{(AS + \lambda S^2)}_{\text{Higgs portal}} - Y_N^{ij} \bar{L}_i H \underbrace{N_j}_{\text{Neutrino portal}} \end{aligned}$$



# Cold dark matter landscape

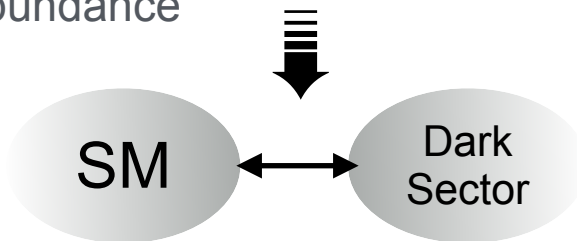


## Sub-GeV thermal relic



$$\langle \sigma_{\text{ann}} v \rangle \propto \frac{m_{\text{DM}}^2}{m_{\text{med}}^4} \sim 1 \text{ pb}$$

**Lee-Weinberg bound**  $\Rightarrow$  light mediators are required for freeze out with the required relic abundance



## EFT for mediators to a (neutral) dark sector

There are just three UV-complete relevant or marginal “portals” to a SM-neutral dark sector, unsuppressed by a (possibly large) new physics scale  $\Lambda$

$$\mathcal{L} = \sum_{n=k+l-4} \frac{c_n}{\Lambda^n} \mathcal{O}_k^{(\text{SM})} \mathcal{O}_l^{(\text{med})} = \mathcal{L}_{\text{portals}} + \mathcal{O}\left(\frac{1}{\Lambda}\right)$$

$$= -\frac{\epsilon}{2} B^{\mu\nu} \underbrace{A'_{\mu\nu}}_{\text{Vector portal}} - H^\dagger H \underbrace{(AS + \lambda S^2)}_{\text{Higgs portal}} - Y_N^{ij} \bar{L}_i H \underbrace{N_j}_{\text{Neutrino portal}}$$

Focus for this talk - thermal DM (scalar, fermion in dark sector) charged under U(1)' vector portal

# Accelerator-based strategy for light DM

To test this scenario, return to an old hypothesis - is CDM more like the CvB (abundant, but with KE too low for direct detection recoil thresholds)?

⇒ Muon neutrinos were instead discovered at BNL in a fixed target experiment, via production and (weak) scattering of a *relativistic* beam

OBSERVATION OF HIGH-ENERGY NEUTRINO REACTIONS AND THE EXISTENCE OF TWO KINDS OF NEUTRINOS\*

G. Danby, J-M. Gaillard, K. Goulianos, L. M. Lederman, N. Mistry, M. Schwartz,<sup>†</sup> and J. Steinberger<sup>†</sup>

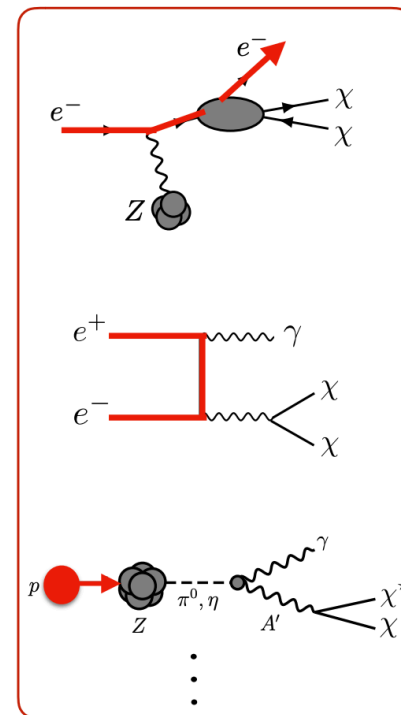
Columbia University, New York, New York and Brookhaven National Laboratory, Upton, New York  
(Received June 15, 1962)

⇒ **accelerator-based search strategy**

Low DM mass, and low dimension portal couplings



High-luminosity medium energy colliders (B-factories) and e-beam and p-beam fixed target experiments



[Krnjaic et al, Snowmass '22]

# Accelerator-based strategy for light DM

To test this scenario, return to an old hypothesis - is CDM more like the CNB (abundant, but with KE too low for direct detection recoil thresholds)?

⇒ Muon neutrinos were instead discovered at BNL in a fixed target experiment, via production and (weak) scattering of a *relativistic* beam

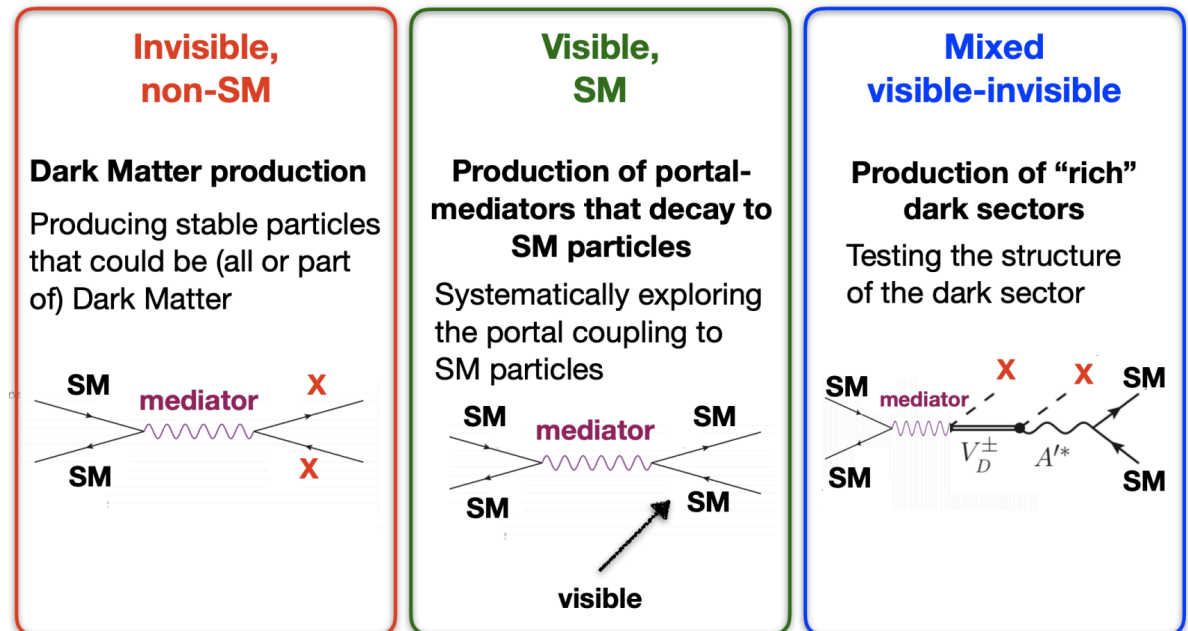
OBSERVATION OF HIGH-ENERGY NEUTRINO REACTIONS AND THE EXISTENCE OF TWO KINDS OF NEUTRINOS\*

G. Danby, J-M. Gaillard, K. Goulianos, L. M. Lederman, N. Mistry, M. Schwartz,<sup>†</sup> and J. Steinberger<sup>†</sup>

Columbia University, New York, New York and Brookhaven National Laboratory, Upton, New York  
(Received June 15, 1962)

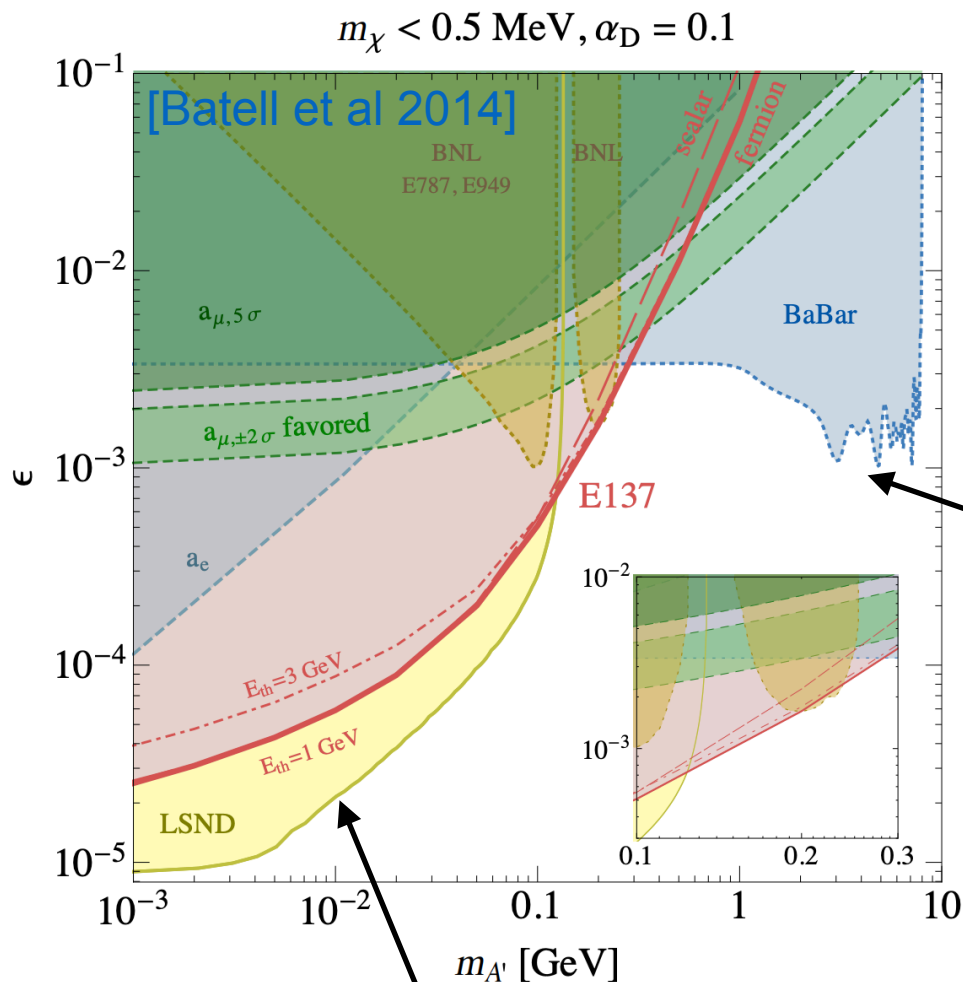
⇒ accelerator-based search strategy

⇒ probe the full kinematics of thermal freezeout of dark sector DM + mediator



[Gori, Williams, Snowmass RF6 '22]

# Status a decade ago (vector portal DM)



Initial efforts to recast existing data and analyses by theorists

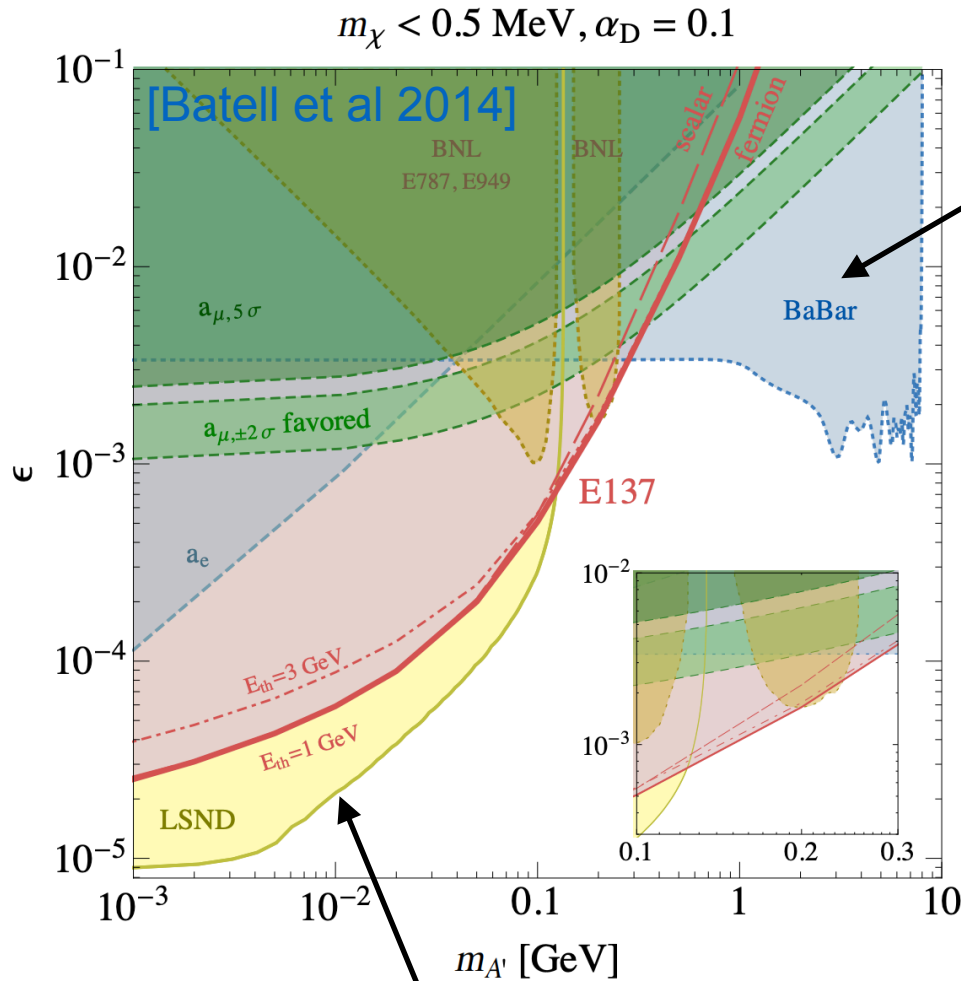
Recast of BaBar monophoton (missing E) data [Essig et al 2013]

Signal rate  $\sim \epsilon^2$

Recast of LSND e-scattering data [deNiverville, Pospelov & AR 2011]

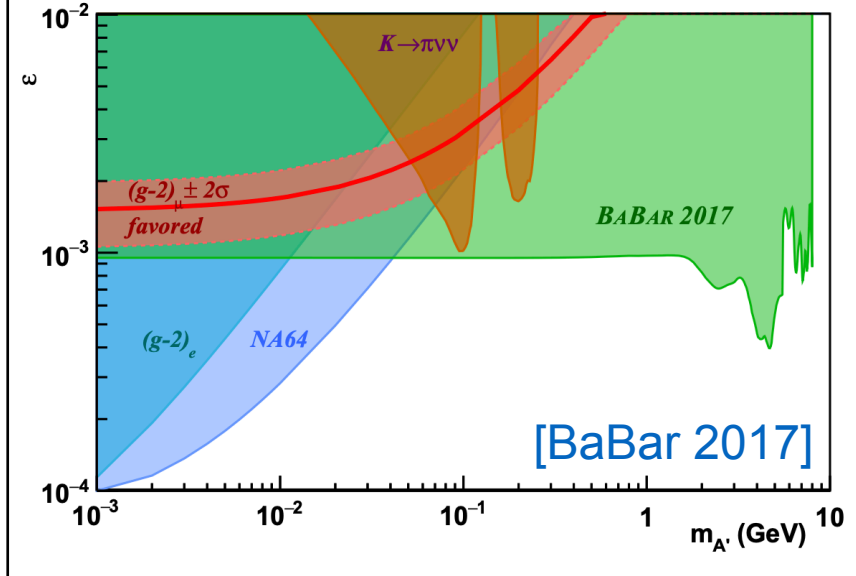
Signal rate  $\sim \epsilon^4$

# Status a decade ago (vector portal DM)



Recast of BaBar monophoton (missing E) data [Essig et al 2013]

Analysis of the full dataset by BaBar improved this limit in 2017



Recast of LSND e-scattering data [deNiverville, Pospelov & AR 2011]

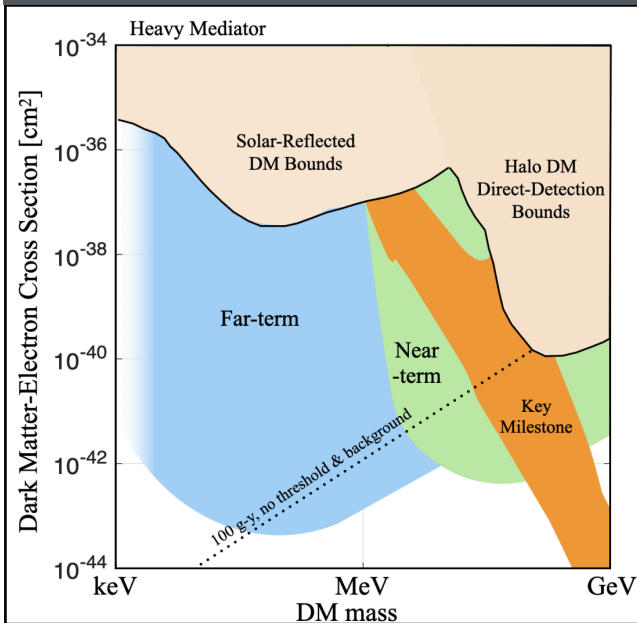
# “New” experiments...

A broad target mass range (MeV-GeV) provides an incentive to explore new technologies, and low-cost synergies with existing (e.g. neutrino) experiments

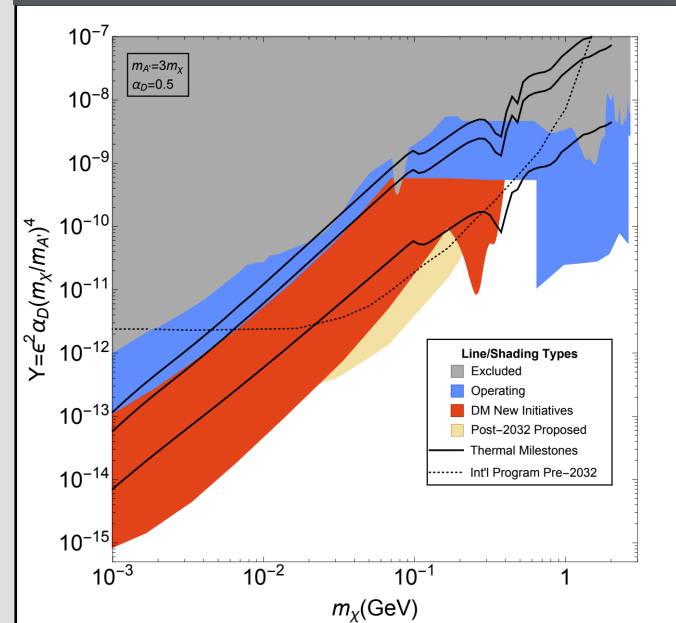
Complementary kinematics

Focus of this talk

## Direct detection via (non-rel) electron scattering

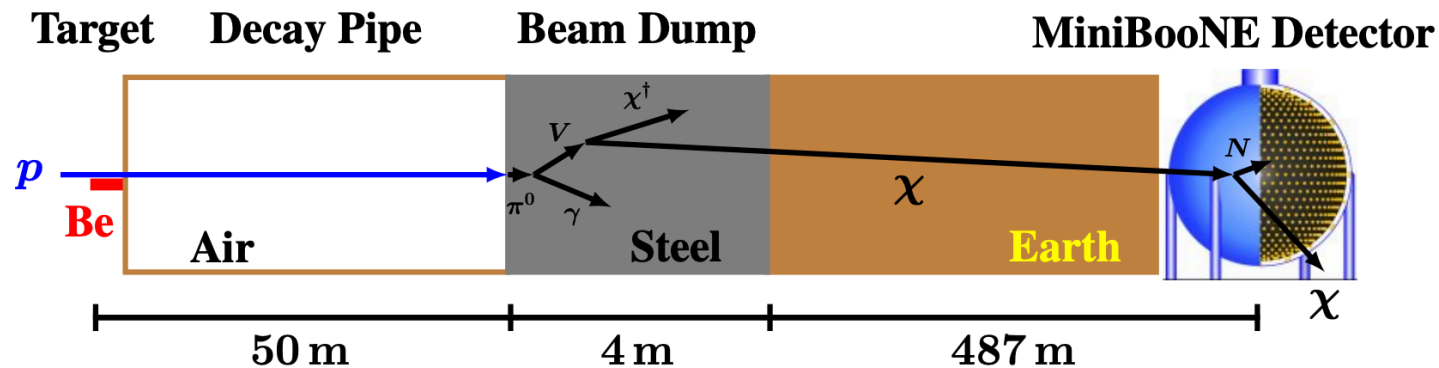


## Accelerator-based missing E/mtm or (rel) scattering



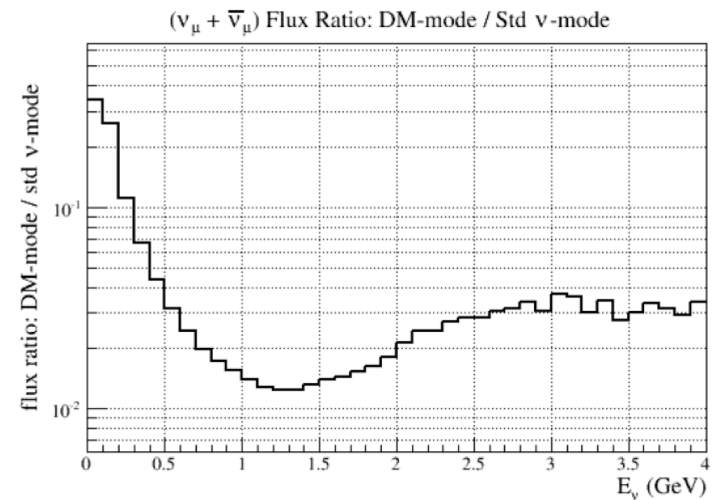
# Neutrino beams & scattering - MiniBooNE

**Proposal:** Synergy with the neutrino short- and long-baseline program, using the (near) detector as a dark matter detector, looking for recoil, but now for a relativistic beam [Batell, Pospelov & AR 2009, deNiverville, Pospelov & AR 2011, Dharmapalan et al 2012]



Neutrino “background” can be reduced significantly:

- p-beam off-target (factor  $\sim 70$  reduction)
- Timing (10ns delay for DM, pulsed beams)
- Recoil energy cuts (forward e-scattering)



# MiniBooNE proposal to FNAL PAC

## A Proposal to Search for Dark Matter with MiniBooNE

Submitted to the FNAL PAC Dec 16, 2013

### The MiniBooNE Collaboration

R. Dharmapalan, & I. Stancu  
*University of Alabama, Tuscaloosa, AL 35487*

R. A. Johnson, & D.A. Wickremasinghe  
*University of Cincinnati, Cincinnati, OH 45221*

R. Carr, G. Karagiorgi, & M. H. Shaevitz  
*Columbia University, New York, NY 10027*

B.C. Brown, F.G. Garcia, R. Ford, T. Kobilarcik, W. Marsh,  
C. D. Moore, D. Perevalov, C. Polly, & W. Wester  
*Fermi National Accelerator Laboratory, Batavia, IL 60510*

J. Grange, & H. Ray  
*University of Florida, Gainesville, FL 32611*

R. Cooper, R. Tayloe, & R. Thornton  
*Indiana University, Bloomington, IN 47405*

G. T. Garvey, W. Huelsnitz, W. Ketchum, Q. Liu, W. C. Louis, G. B. Mills,  
J. Mirabal, Z. Pavlovic, C. Taylor, R. Van de Water, & D. H. White  
*Los Alamos National Laboratory, Los Alamos, NM 87545*

B. P. Roe  
*University of Michigan, Ann Arbor, MI 48106*

A. A. Aguilar-Arevalo & I. L. de Icaza Astiz  
*Instituto de Ciencias Nucleares, Universidad Nacional Autónoma de México, D.F. México*

Stefan Edelstein  
*Saint Mary's University of Minnesota, Winona, MN 55987*

T. Katori  
*Queen Mary University of London, London, E1 4NS, UK*

C. Mariani  
*Virginia Tech, Blacksburg, VA, 24061*

### The Theory Collaboration

B. Batell  
*University of Chicago, Chicago, IL, 60637*

P. deNiverville, M. Pospelov, & A. Ritz  
*University of Victoria, Victoria, BC, V8P 5C2*

D. McKeen  
*University of Washington, Seattle, WA, 98195*

[Dharmapalan et al 2012]

PAC presentation  
in Jan 2014

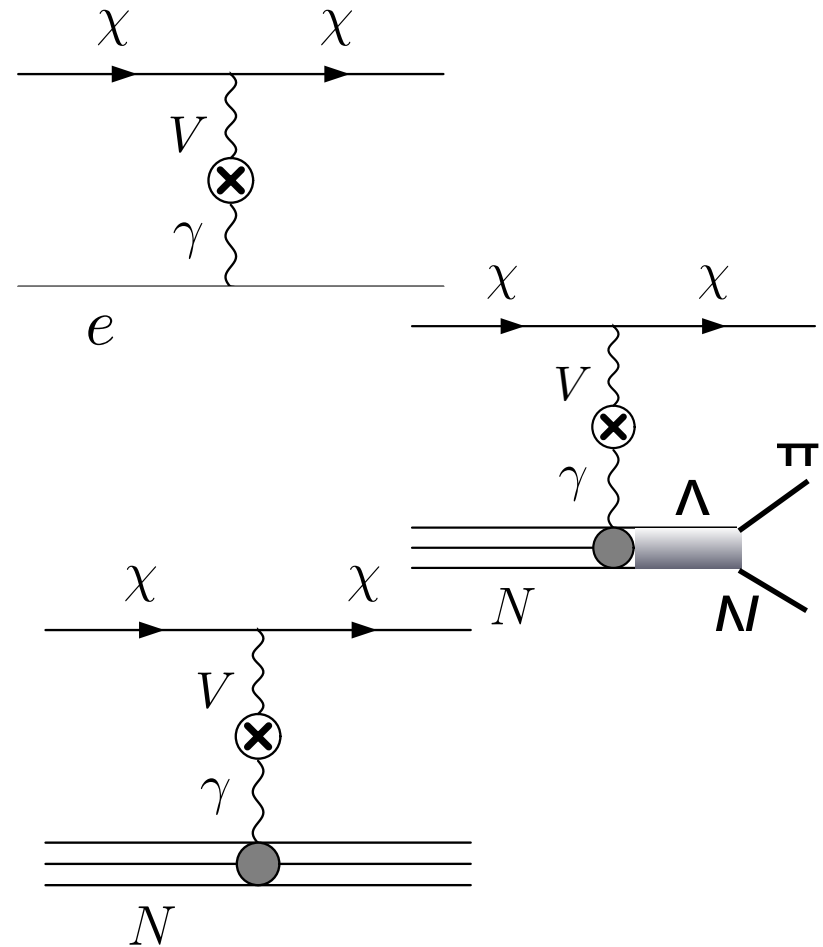
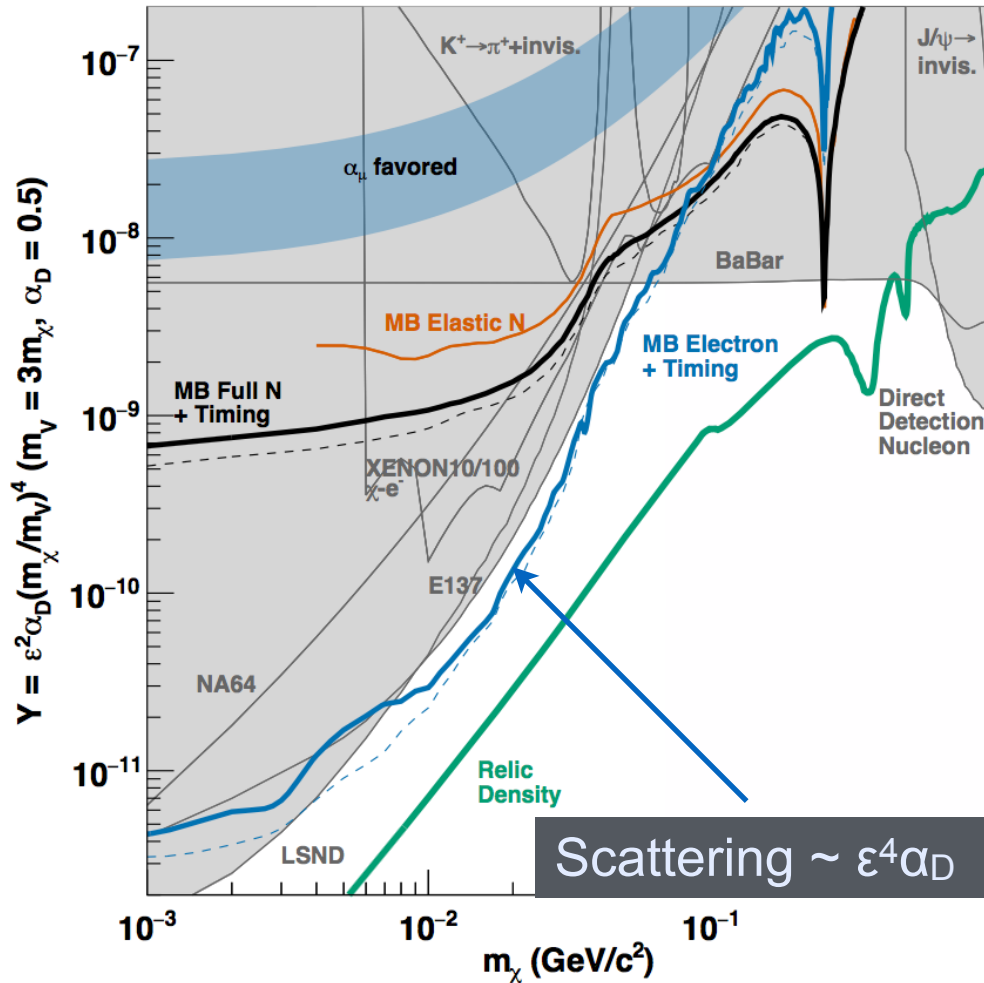
Request  $2 \times 10^{20}$   
POT in 2014 with  
beam off-target

✓ - beam dump run approved initially for 12 months,  
but ultimately extended until MicroBooNE switched on



# MiniBooNE results

Scattering signatures mimic neutral current neutrino interactions



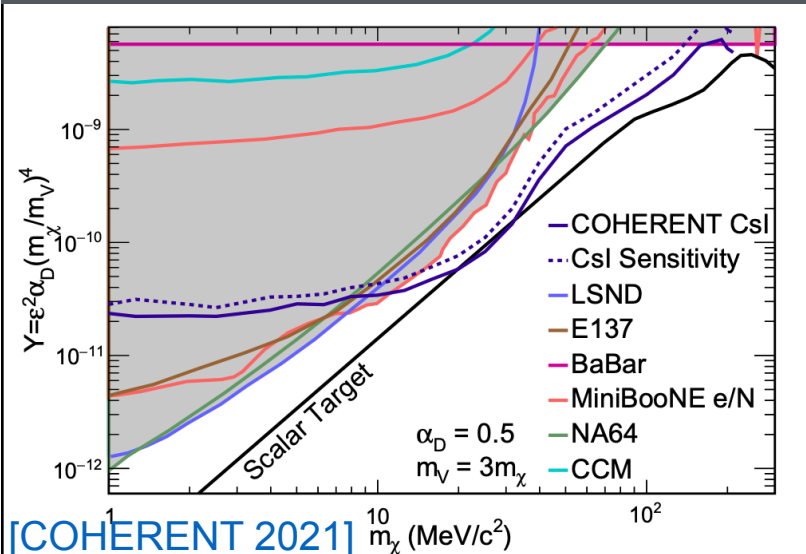
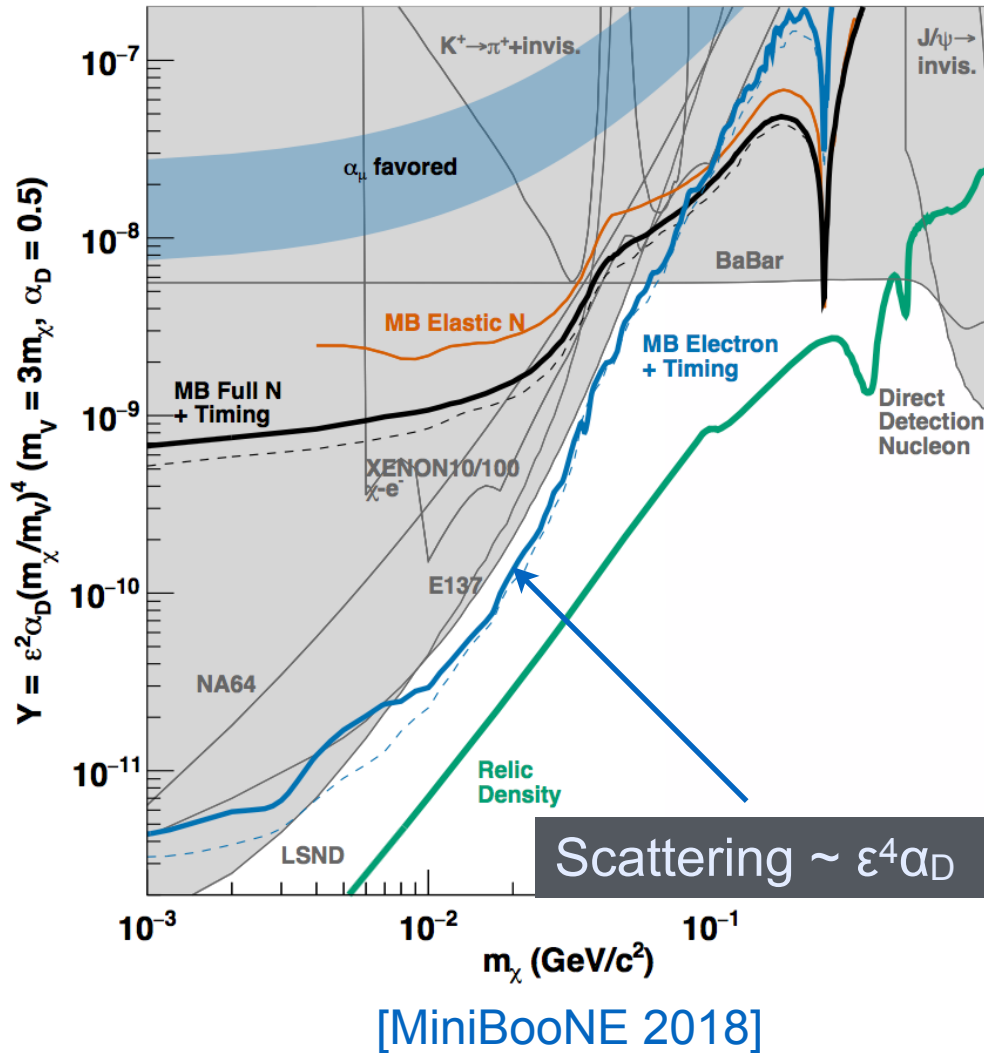
[MiniBooNE 2018]

# MiniBooNE (& COHERENT) results

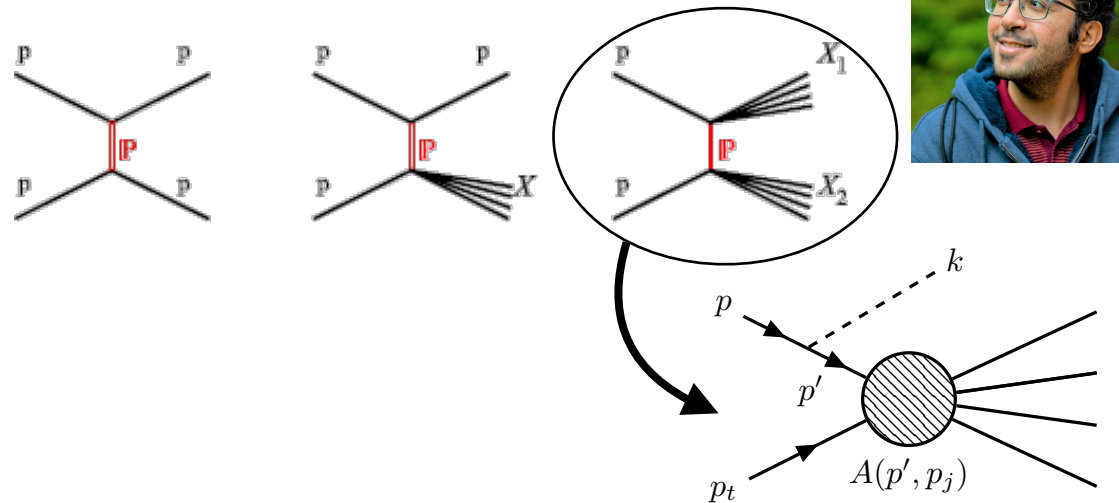
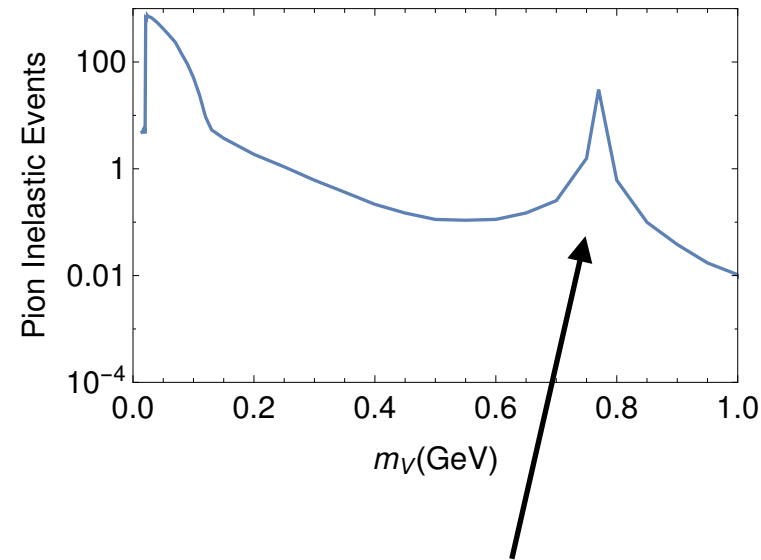
Scattering signatures mimic neutral current neutrino interactions

Small improvement in sensitivity, but this was a full experimental analysis, not a theoretical recast (as for the LSND limit), and pioneered a number of tools for background reduction

Improved higher-mass sensitivity using similar techniques in 2021 from COHERENT CsI (at SNS)

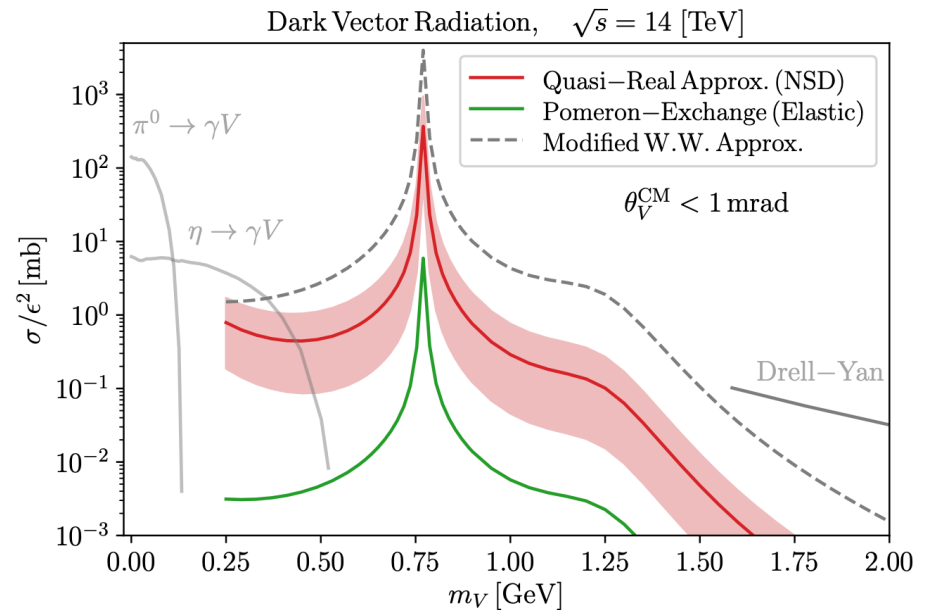


# Neutrino Beams - production rate



Revisiting  $A'$  production in proton bremsstrahlung, including the  $\rho/\omega$  resonance region

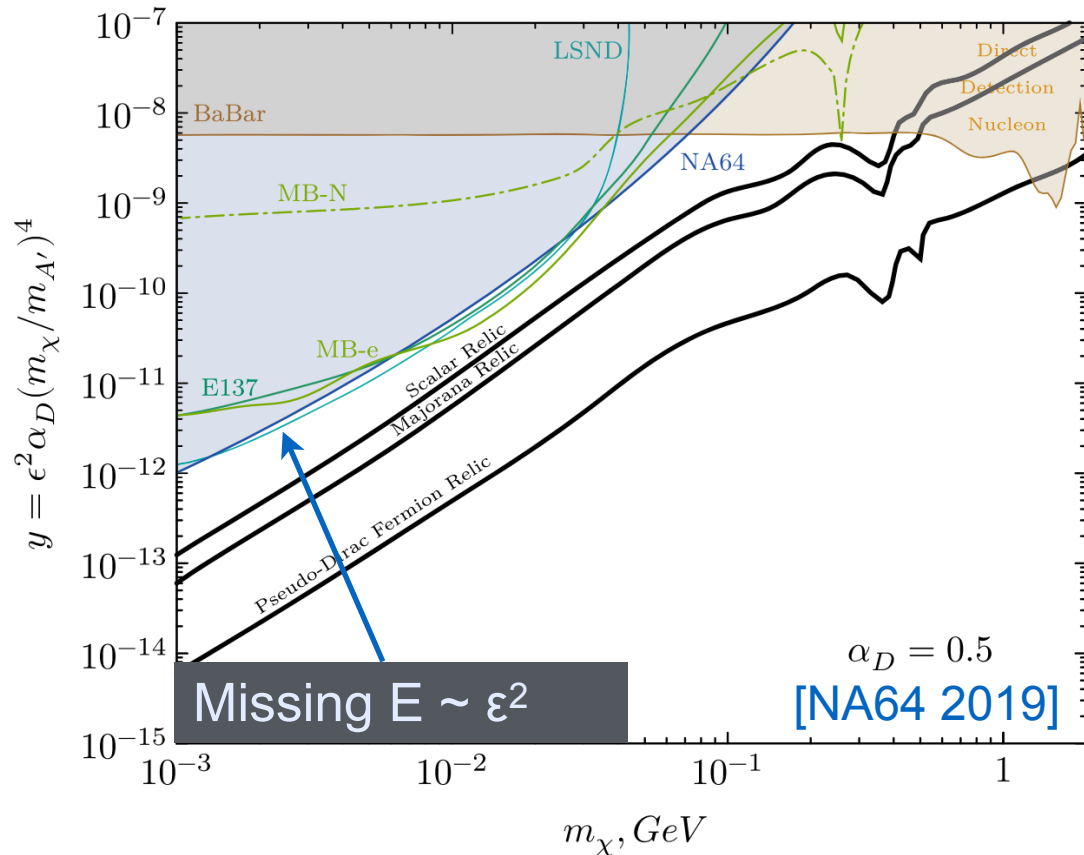
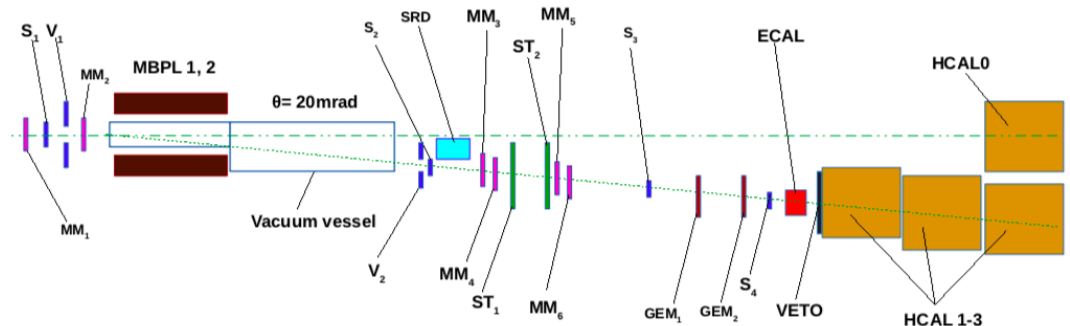
[Foroughi-Abari, AR '21 + work in progress]



# Missing mass - NA64

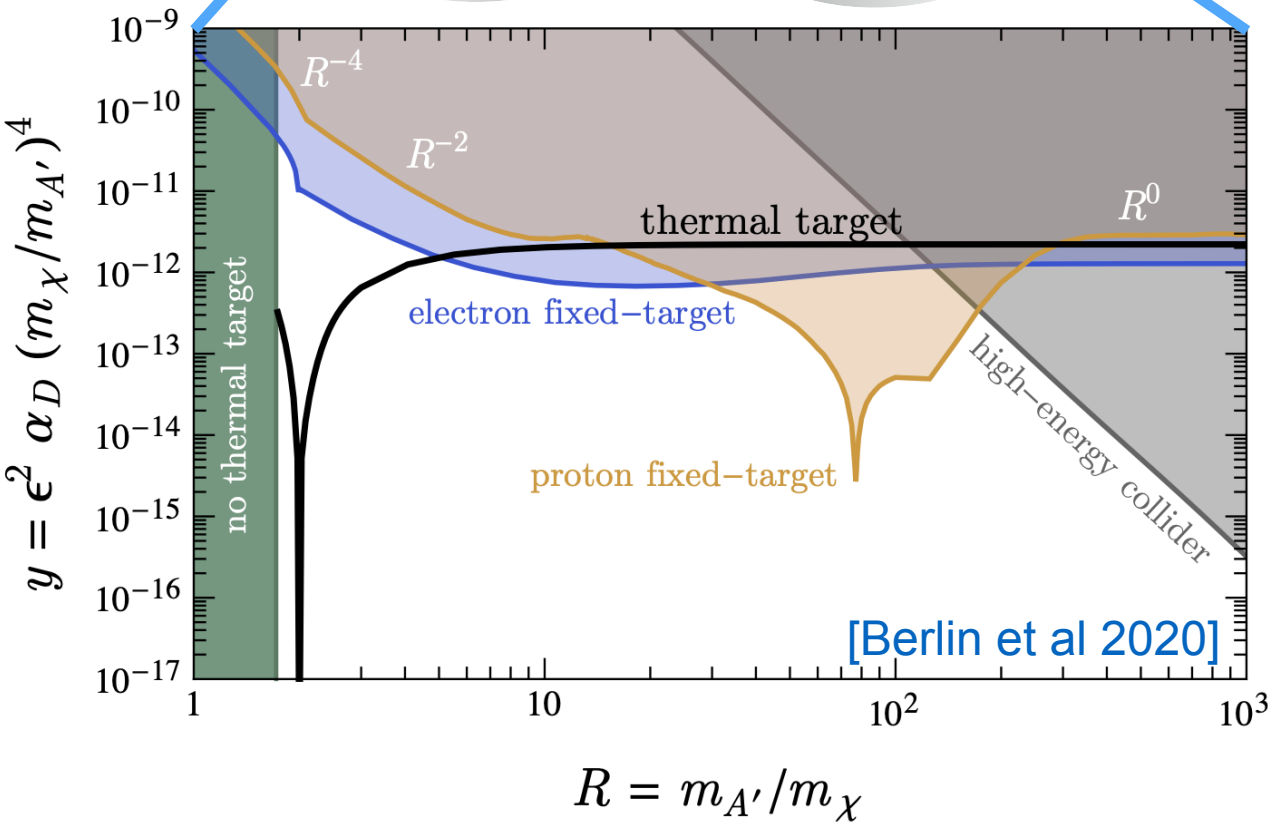
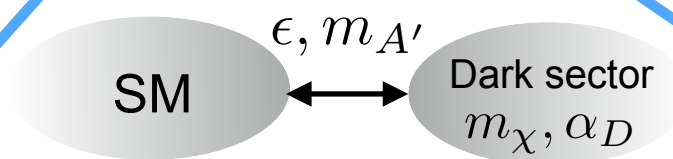
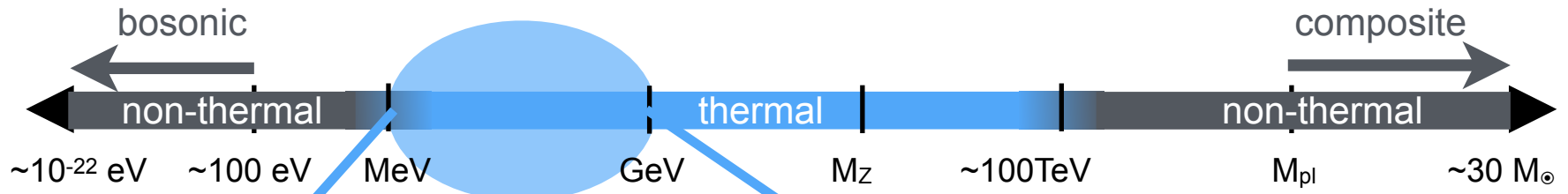
The signal rate for production and missing mass is  $\sim \epsilon^2$

Proposal for NA64 approved for CERN north area with operations starting in 2016



Improvement on E137, MB-e and LSND limits, relying purely on the  $A'$ -electron coupling

# Sub-GeV thermal DM landscape today

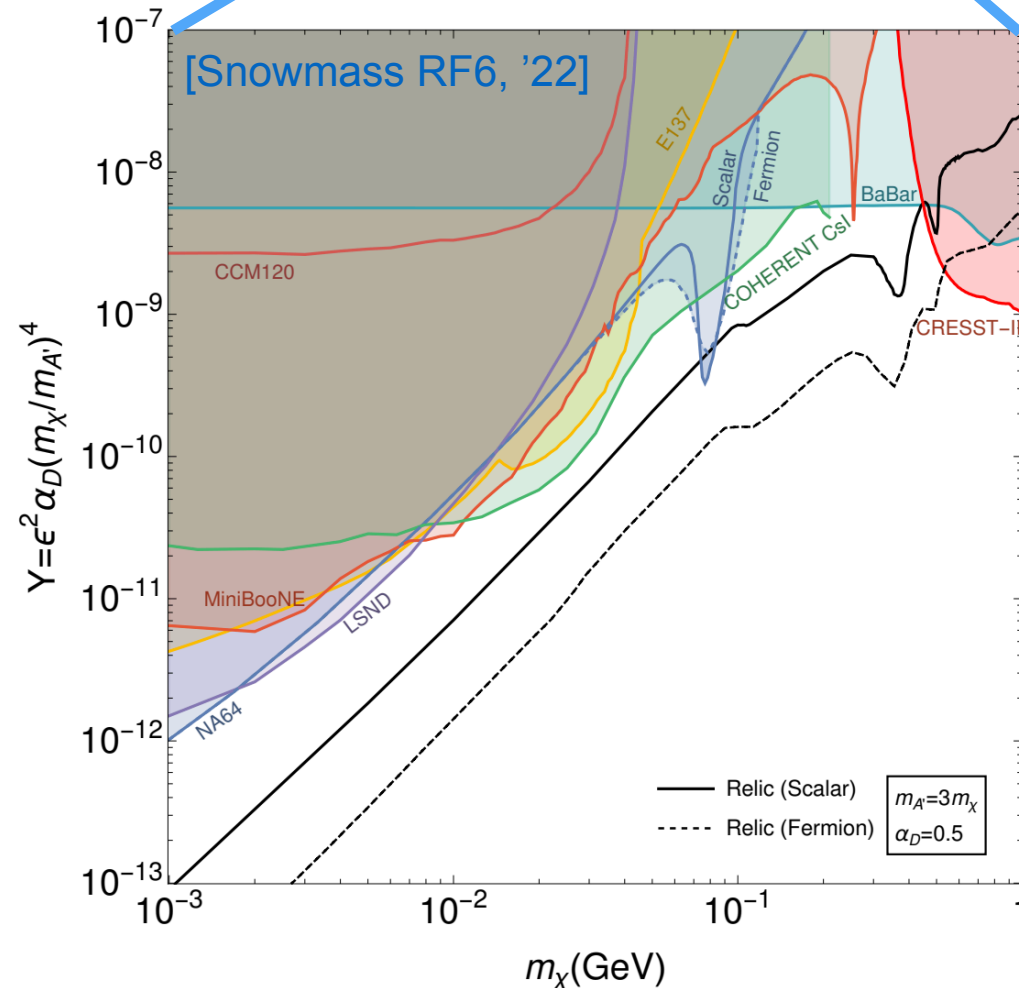
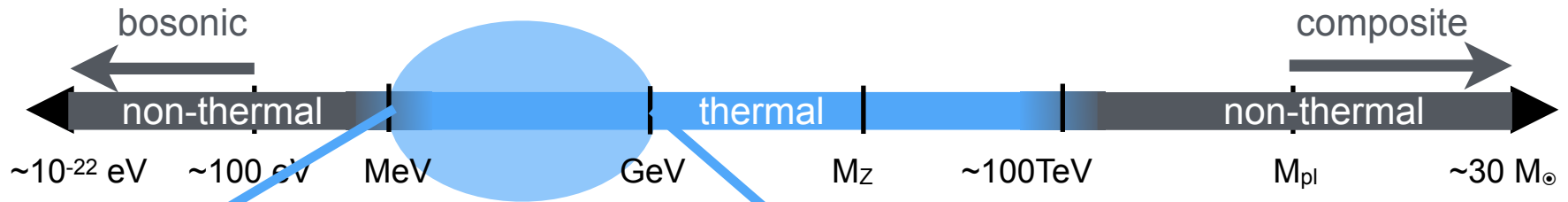


**Invisible,  
non-SM**

**Dark Matter production**  
Producing stable particles that could be (all or part of) Dark Matter

Feynman diagram showing SM particles interacting via a mediator to produce two X particles.

# Sub-GeV thermal DM landscape today

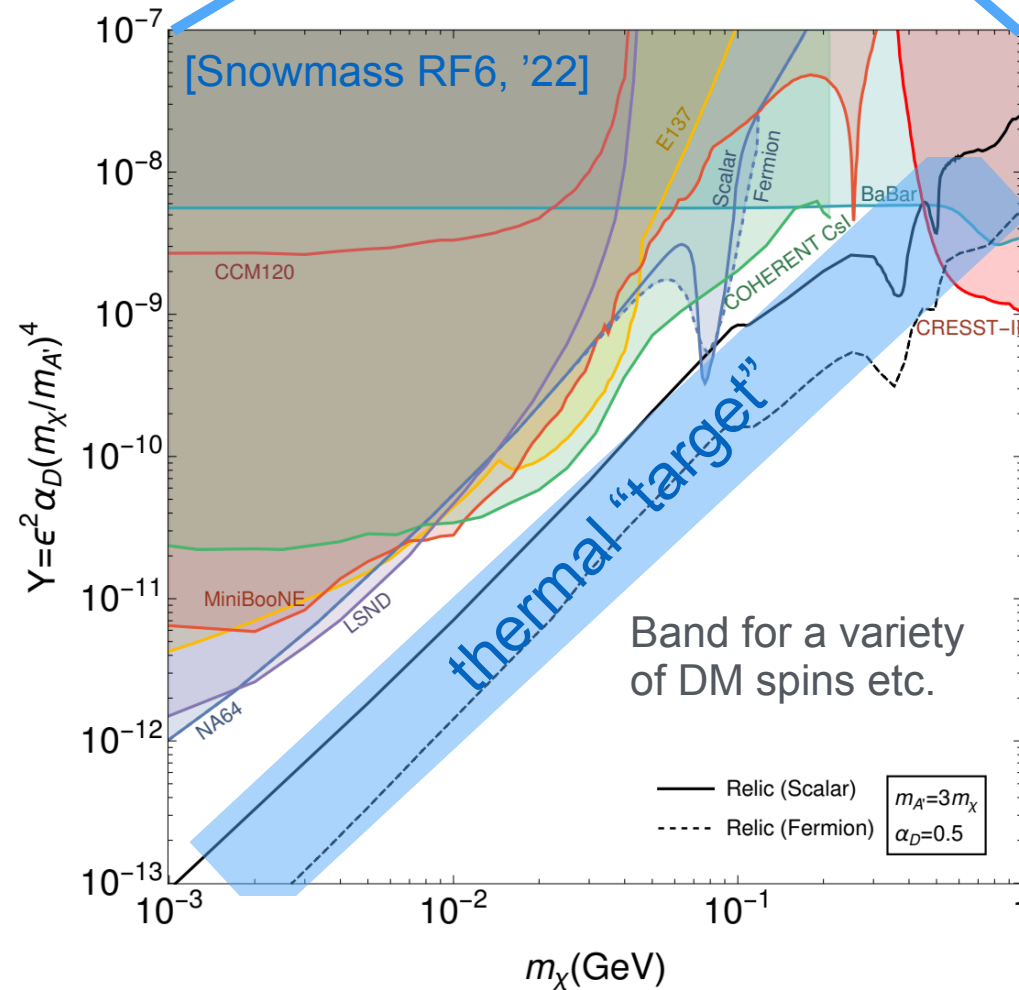
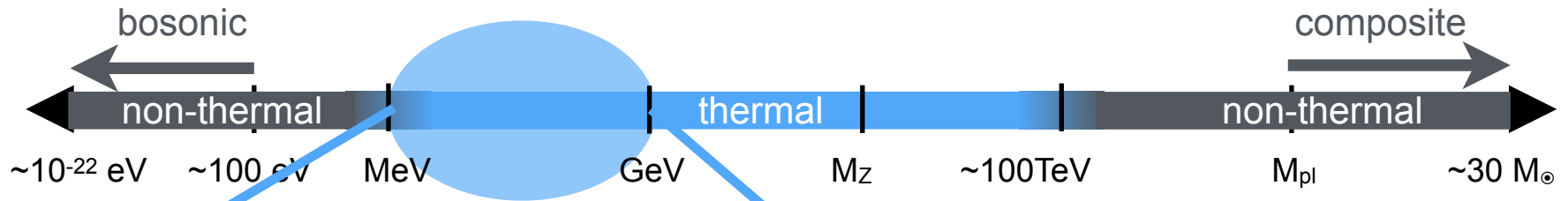


**Invisible, non-SM**

**Dark Matter production**  
Producing stable particles that could be (all or part of) Dark Matter

Feynman diagram showing the production of two dark matter particles (X) from two Standard Model (SM) particles via a mediator. The mediator is represented by a wavy line connecting the two interaction vertices.

# Sub-GeV thermal DM landscape today

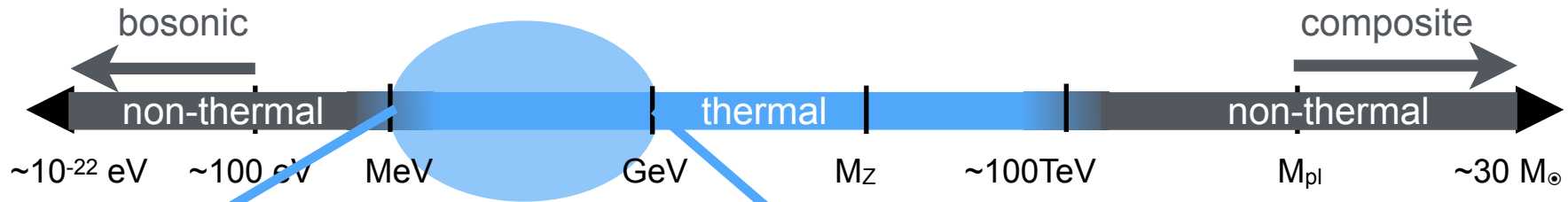


**Invisible,  
non-SM**

**Dark Matter production**  
Producing stable particles that could be (all or part of) Dark Matter

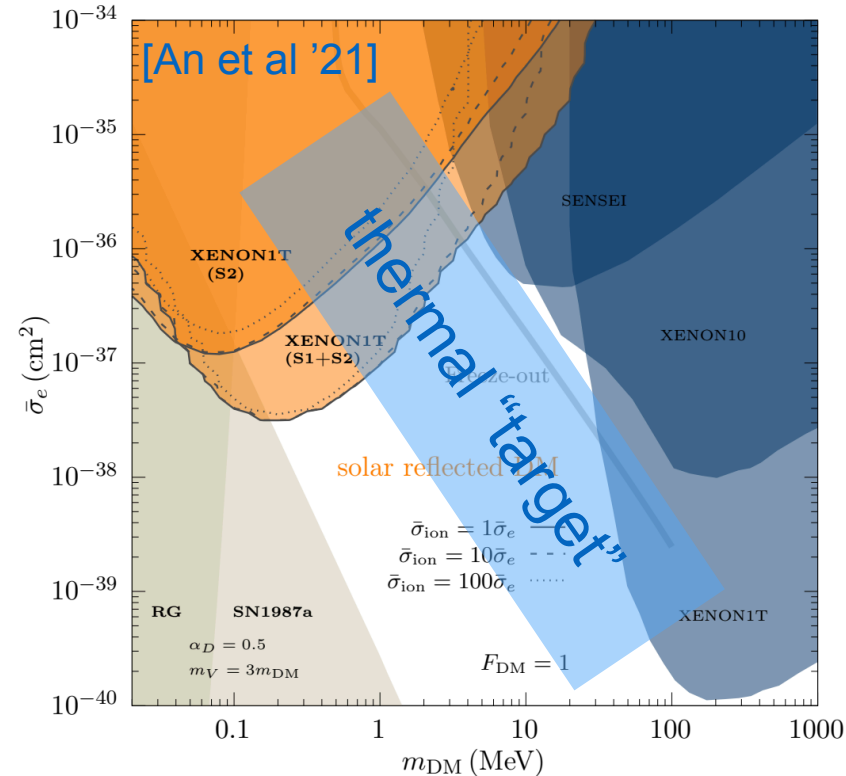
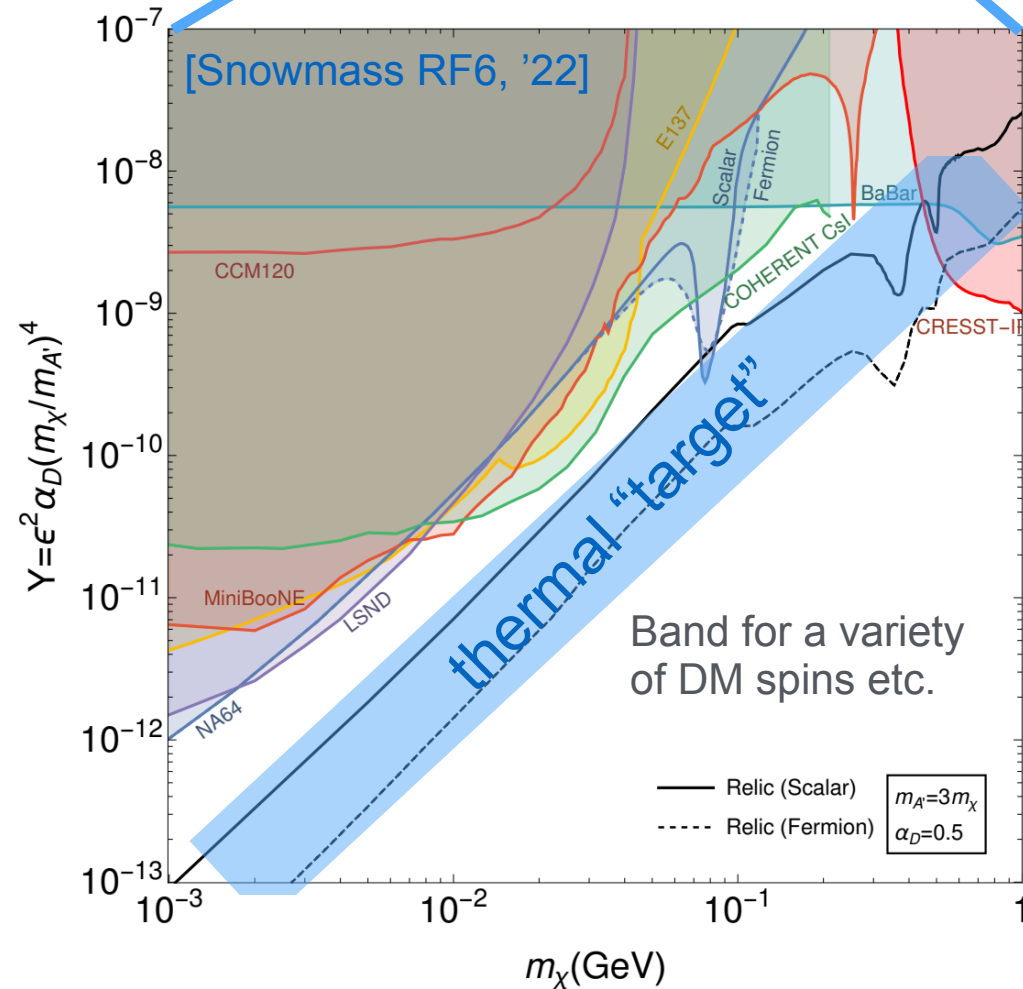
$SM + SM \rightarrow \text{mediator} \rightarrow X + X$

# Sub-GeV thermal DM landscape today



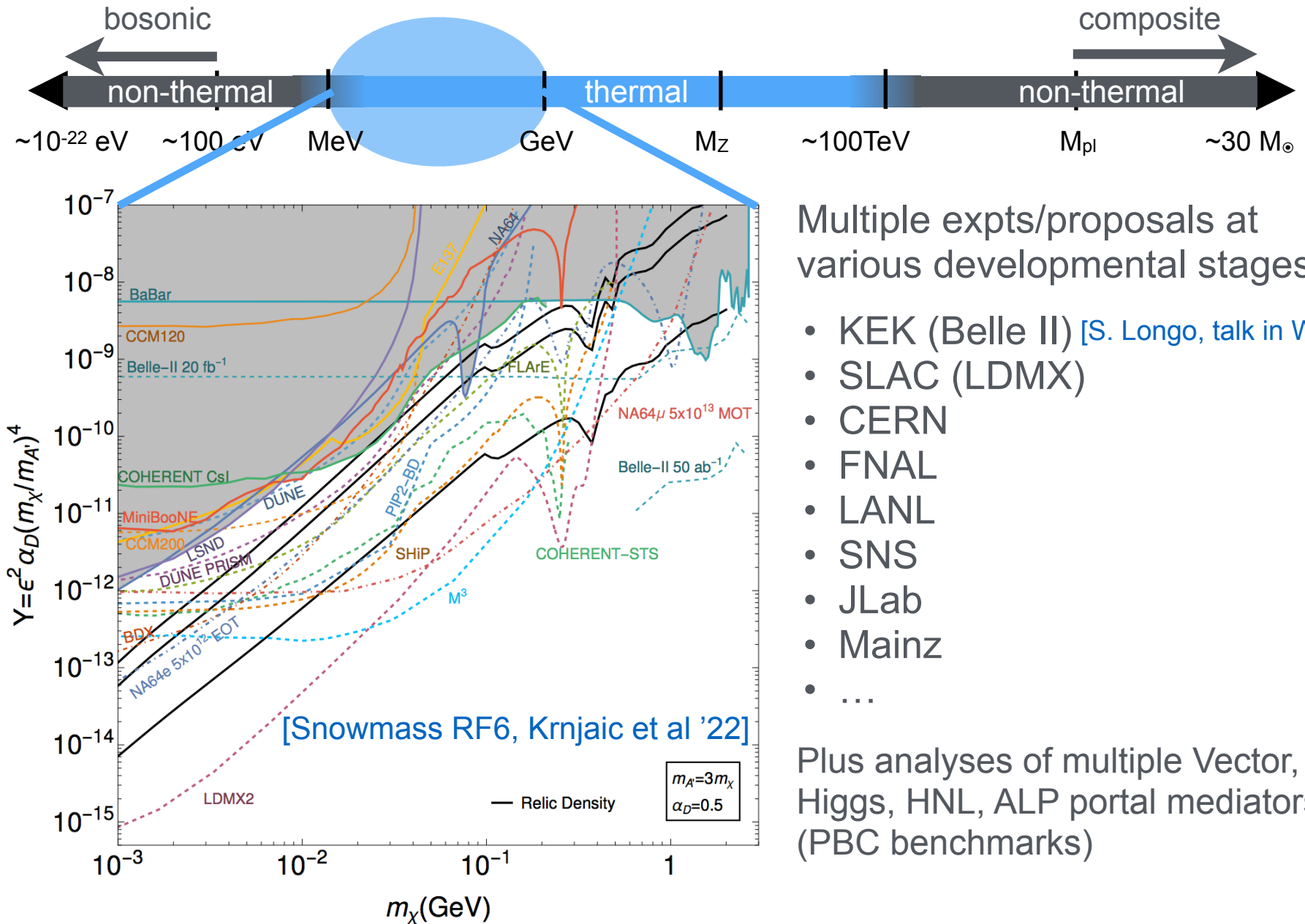
Complementary thermal target for next-gen low mass direct detection, (and improved sensitivity to  $N_{\text{eff}}$ )

[Snowmass WP Essig et al, '22]





# Sub-GeV thermal DM landscape today



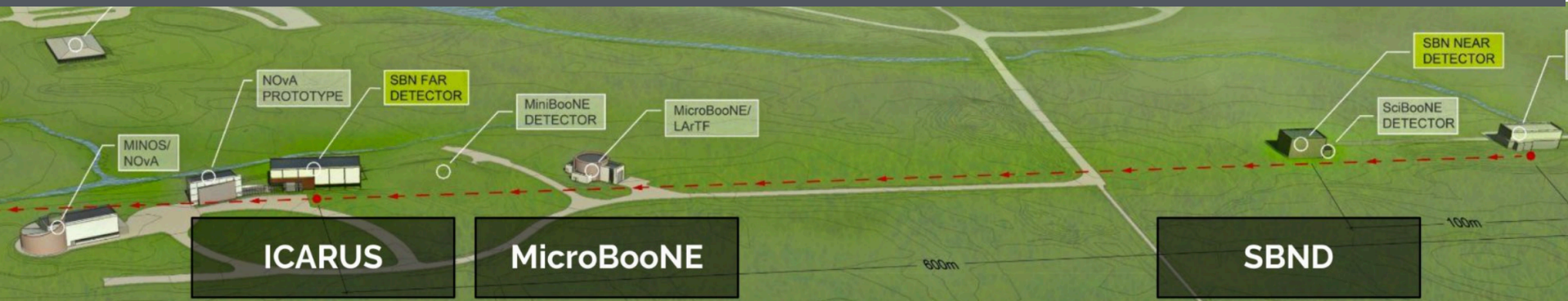
Multiple expts/proposals at various developmental stages

- KEK (Belle II) [S. Longo, talk in W1-1]
- SLAC (LDMX)
- CERN
- FNAL
- LANL
- SNS
- JLab
- Mainz
- ...

Plus analyses of multiple Vector, Higgs, HNL, ALP portal mediators (PBC benchmarks)

# Fermilab & CERN

Synergistic with new FNAL short-baseline neutrino program



Multiple proposals to broadly probe LLPs (PBC benchmarks) at CERN

**CODEX-b @ LHCb IP**  
**MOEDAL/MAPP@LHCb IP**

**LHCb**

**ATLAS**

**CMS**

**MATHUSLA @ CMS IP**

**NA64++(e) @ EHN1**  
**HIKE/SHADOWS @ ECN3**  
**SHIP @ ECN3**

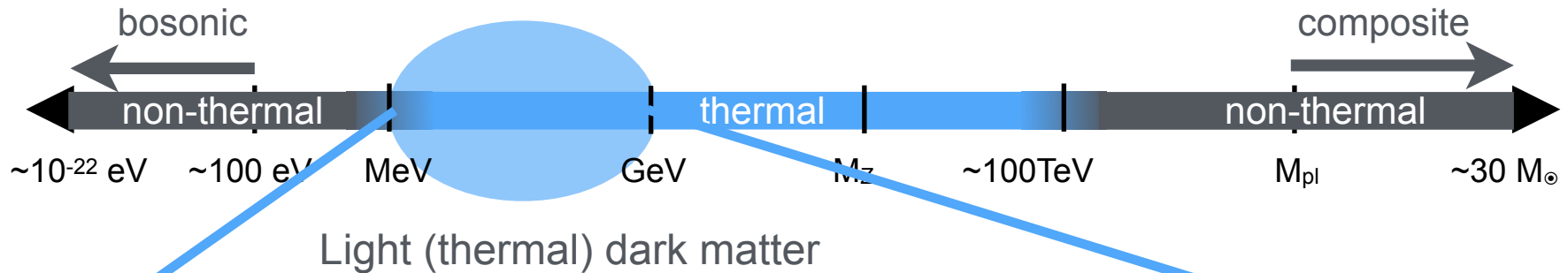
**FASER @ ATLAS IP**  
**ANUBIS @ ATLAS shaft**  
**Forward Physics Facility @ ATLAS IP**

**LHC**

**FASER ready to take data**

[Lanfranchi '21]

# Summary



High-luminosity accelerators have the kinematics to test facets of thermal freezeout in MeV-GeV DM models, a complementary probe to direct detection (N- or e-scattering)

- The effort to broadly search for light DM coupled to the Standard Model via the renormalizable mediator portals has progressed a long way over the past decade.

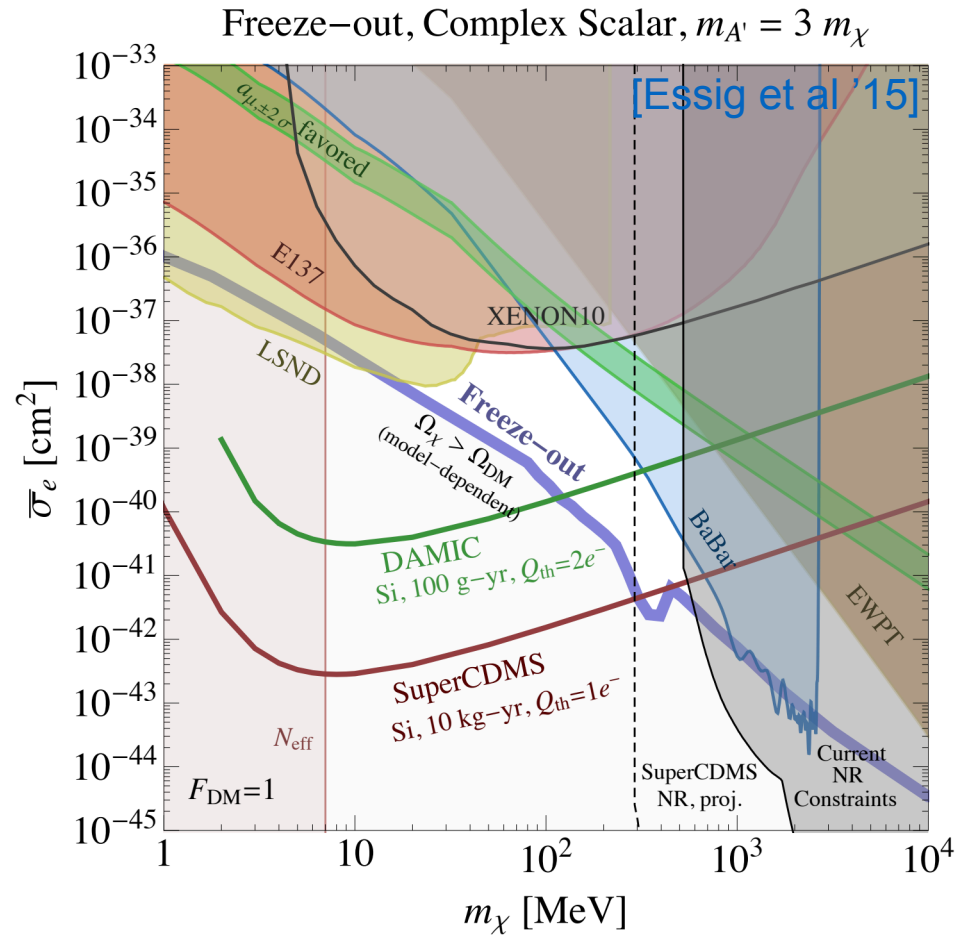
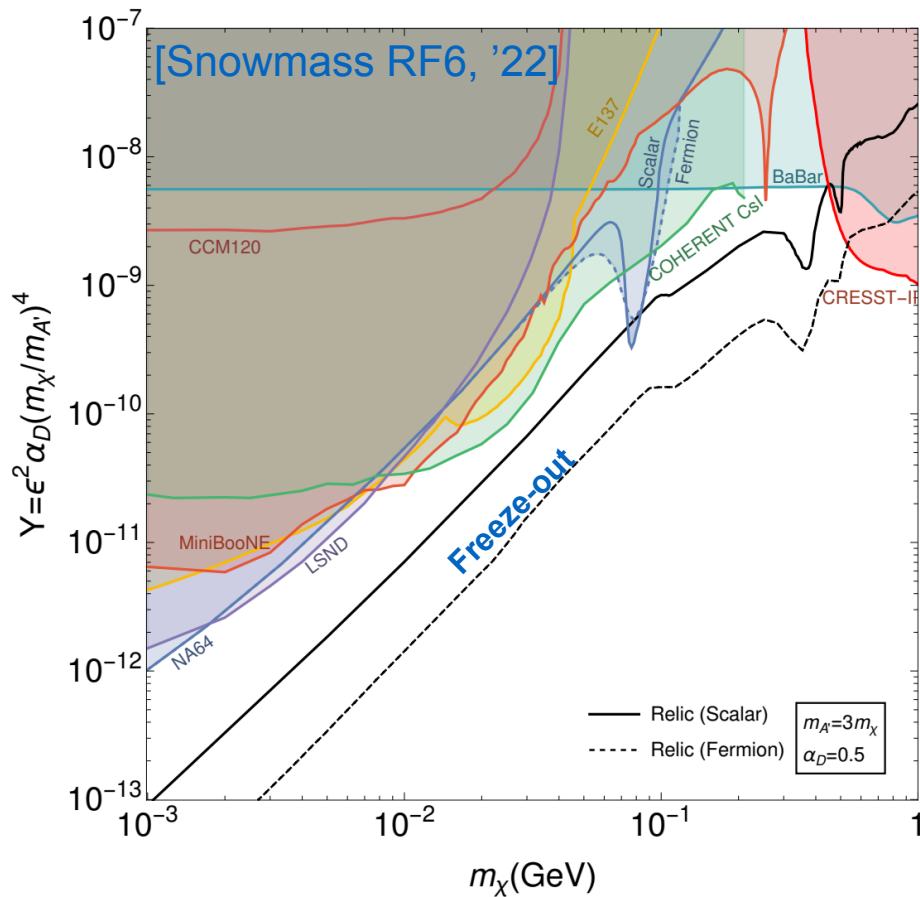
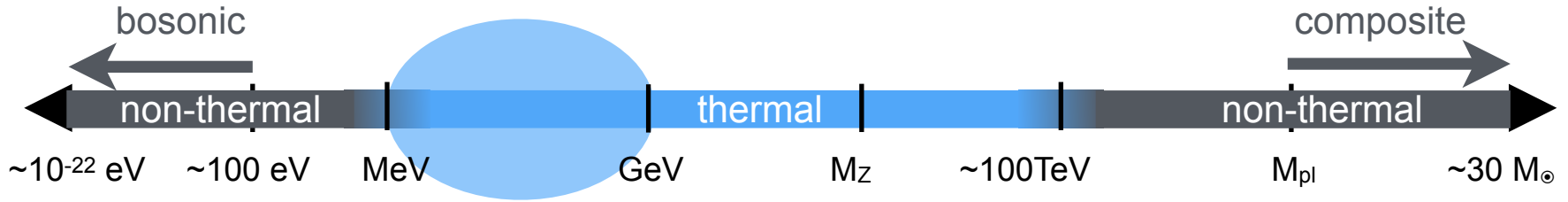
**Natural interplay with long and short baseline neutrino program, which can be pursued further with next-gen facilities**

**Dedicated plans/proposals at multiple labs (CERN, Fermilab, SLAC, LANSCE, KEK, Mainz, JLab) over the coming decade**

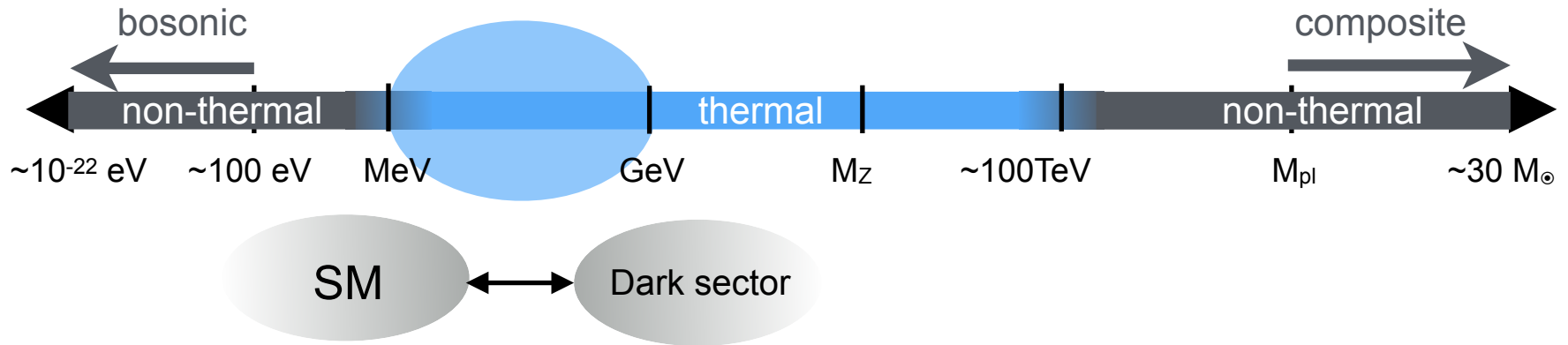
**+ Many complementary efforts on low-mass direct detection, e.g. via e-scattering**

# Backup Slides

# Sub-GeV thermal DM - $Y$ vs $\sigma_e$



# Dark sectors - PBC benchmarks



Broader more systematic framework for light dark sectors and LLPs, covering all low dimension portal mediators, and *focusing on visible decay signatures*

$$\begin{aligned}
 \mathcal{L} &= \sum_{n=k+l-4} \frac{c_n}{\Lambda^n} \mathcal{O}_k^{(\text{SM})} \mathcal{O}_l^{(\text{med})} && \text{CERN PBC benchmark cases} \\
 &= -\frac{\epsilon}{2} B^{\mu\nu} A'_{\mu\nu} - H^\dagger H (AS + \lambda S^2) - Y_N^{ij} \bar{L}_i H N_j \\
 &\quad \text{BC 1-3} \qquad \qquad \qquad \text{BC 4-5} \qquad \qquad \qquad \text{BC 6-8} \\
 &\quad + \frac{1}{f_a} \left( \text{tr}(G\tilde{G}) + c_F F\tilde{F} + c_\psi \partial_\mu j_{A\psi}^\mu \right) a + \mathcal{O}(\text{dim} \geq 5) \\
 &\qquad \qquad \qquad \qquad \qquad \qquad \qquad \qquad \qquad \qquad \qquad \text{BC 9-11}
 \end{aligned}$$