



Canadian Association
of Physicists

Association canadienne
des physiciens et physiciennes

The T2K, Super-Kamiokande, and Hyper-Kamiokande Experiments

CAP CONGRESS, JUNE 8, 2021

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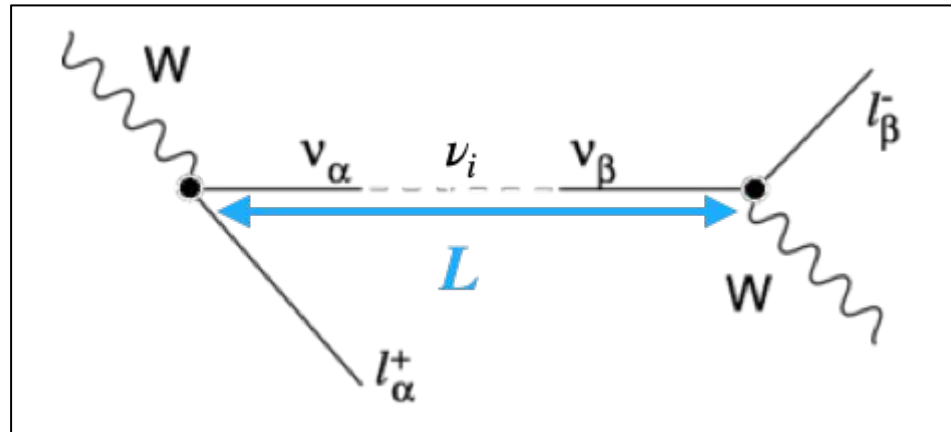


Neutrino Oscillation Formalism

Flavor Eigenstate $|\nu_\alpha\rangle = \sum_i U_{\alpha i}^* |\nu_i\rangle$ Hamiltonian Eigenstate ν_1, ν_2, ν_3

Superposition (Unitary transf.)

Produced as weak/flavour state



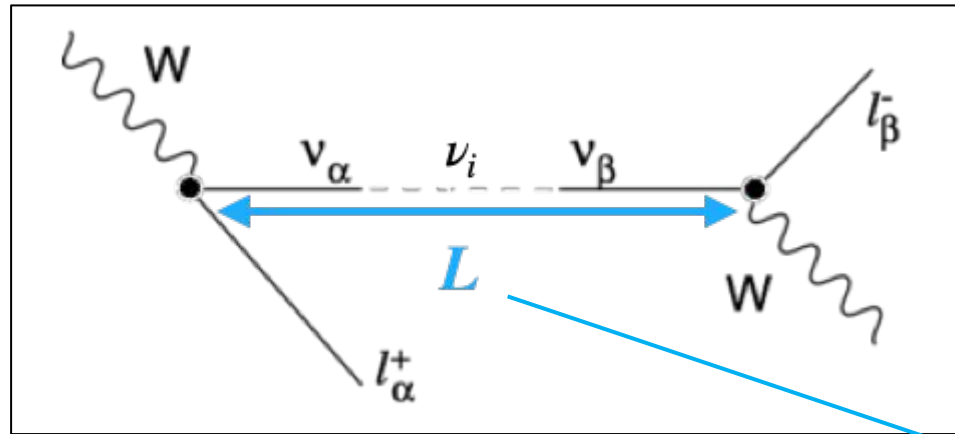
Interact as weak/flavour state

Propagate as mass states with relative phases

Neutrino Oscillation Formalism

Flavor Eigenstate $|\nu_\alpha\rangle = \sum_i U_{\alpha i}^* |\nu_i\rangle$ Hamiltonian Eigenstate ν_1, ν_2, ν_3

Superposition (Unitary transf.)



Mixing matrix (oscillation) parameters

ν mass² differences

Oscillation probability

$$P(\nu_\alpha \rightarrow \nu_\beta) = 4 \sum \text{Re} \left(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^* \right) \sin^2 \left[\frac{\Delta m_{ij}^2 L}{4 E} \right] + \dots$$

ν energy

Neutrino Oscillation Formalism

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$c_{ij} = \cos \theta_{ij}$, and $s_{ij} = \sin \theta_{ij}$.

Flavor
Eigenstate

$$|\nu_\alpha\rangle$$

$$= \sum_i U_{\alpha i}^* |\nu_i\rangle$$

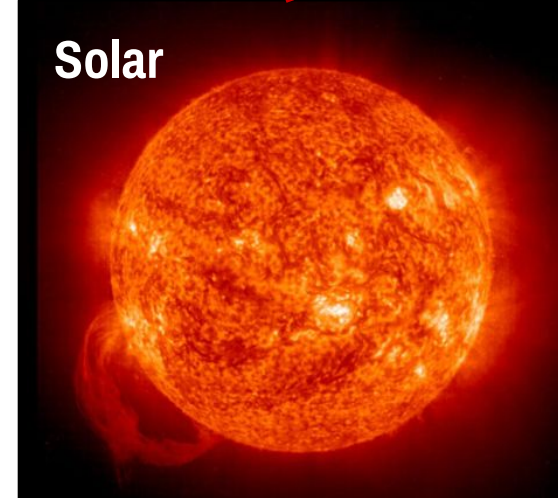
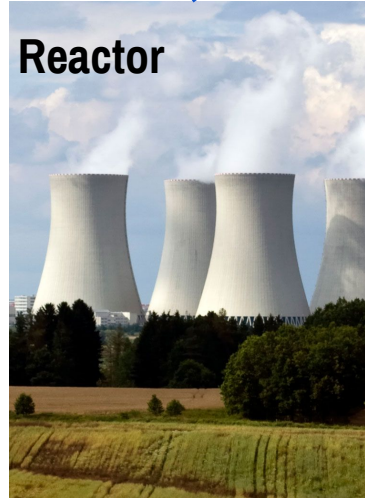
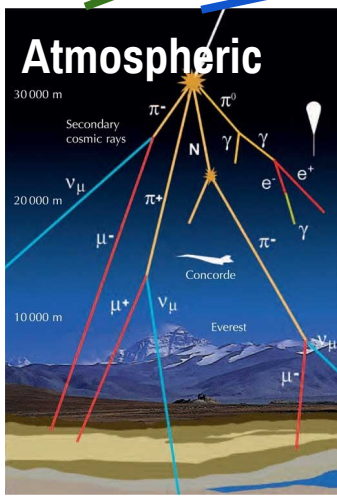
Superposition
(Unitary transf.)

Hamiltonian
Eigenstate
 ν_1, ν_2, ν_3

Neutrino Sources

$c_{ij} = \cos \theta_{ij}$, and $s_{ij} = \sin \theta_{ij}$.

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



Neutrino Knowns and Unknowns

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$c_{ij} = \cos \theta_{ij}$, and $s_{ij} = \sin \theta_{ij}$.

KNOWNNS

($\sim 1\sigma$ accuracy)

$\Delta m^2 / \text{eV}^2 = 2.48 \times 10^{-3}$	(1.3%)
$\delta m^2 / \text{eV}^2 = 7.34 \times 10^{-3}$	(2.2%)
$\sin^2 \theta_{13} = 0.0225$	(3.0%)
$\sin^2 \theta_{12} = 0.303$	(4.4%)
$\sin^2 \theta_{23} = 0.545$	($\sim 5\%$)

UNKNOWNNS

(>1 σ hints)

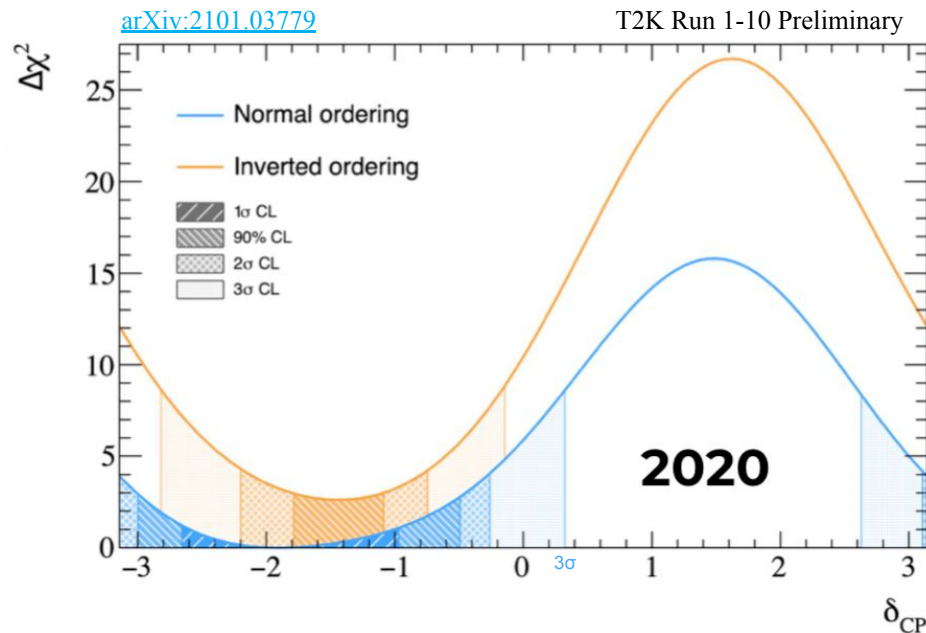
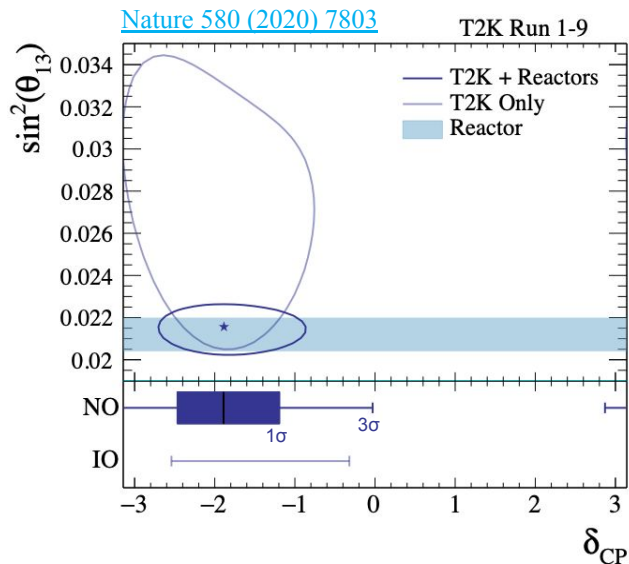
Dirac or Majorana	
Mass ordering	(>3 σ NO)
Absolute mass	(<sub-eV)
Dirac CP phase δ_{cp}	(1.6 σ CPV)
Octant of θ_{23}	

E. Lisi (TAUP2019)

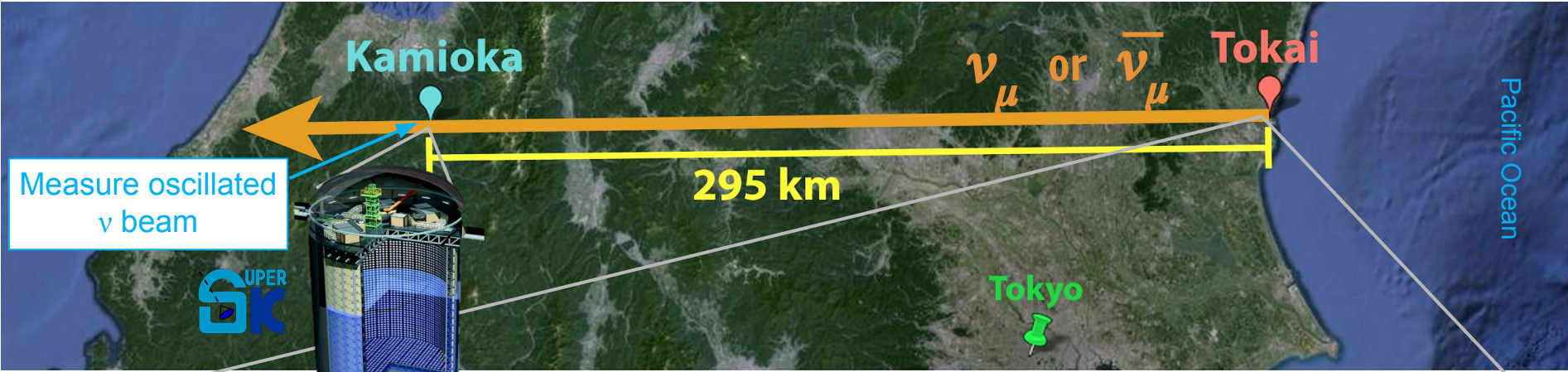
T2K CP Violation Constraints



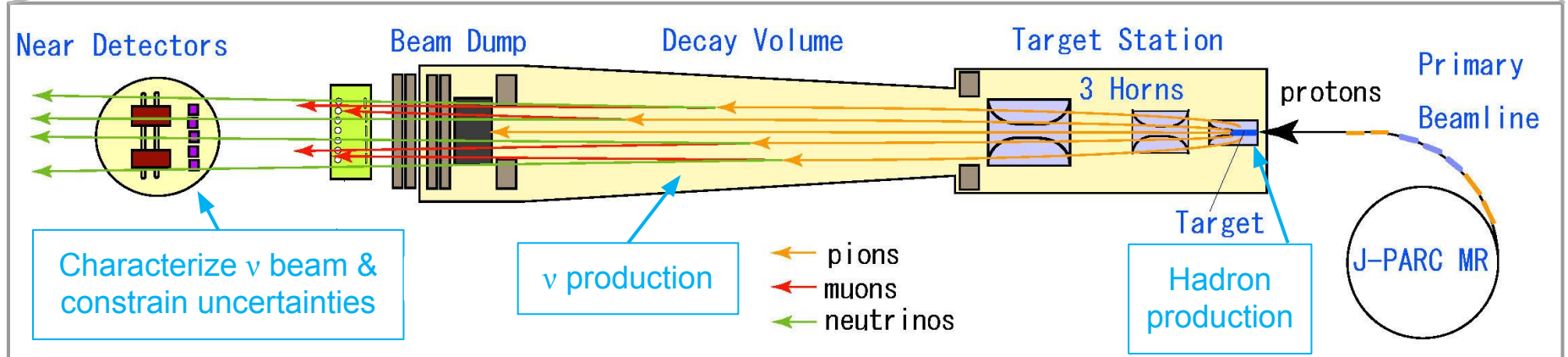
- 2019 analysis:
 - Disfavored $\delta_{CP} = 0$ at 3σ
 - Disfavored IO at 1σ
- 2020 analysis slightly looser constraints



Building a Neutrino Beam (in Japan)



Japan Proton Accelerator Research Complex & Neutrino Beamline



T2K Beam Exposure and History

POT Total: 3.64×10^{21} (ν -mode : $\bar{\nu}$ -mode $\sim 6 : 5$)

— Total Accumulated POT for Physics
— ν -Mode Accumulated POT for Physics
— $\bar{\nu}$ -Mode Accumulated POT for Physics
• ν -Mode Beam Power
• $\bar{\nu}$ -Mode Beam Power

2020 Result

[arXiv:2101.03779](https://arxiv.org/abs/2101.03779)

Constraint

[Nature 580, 339–344](https://doi.org/10.1038/s41586-020-2000-4)

Search for CPV

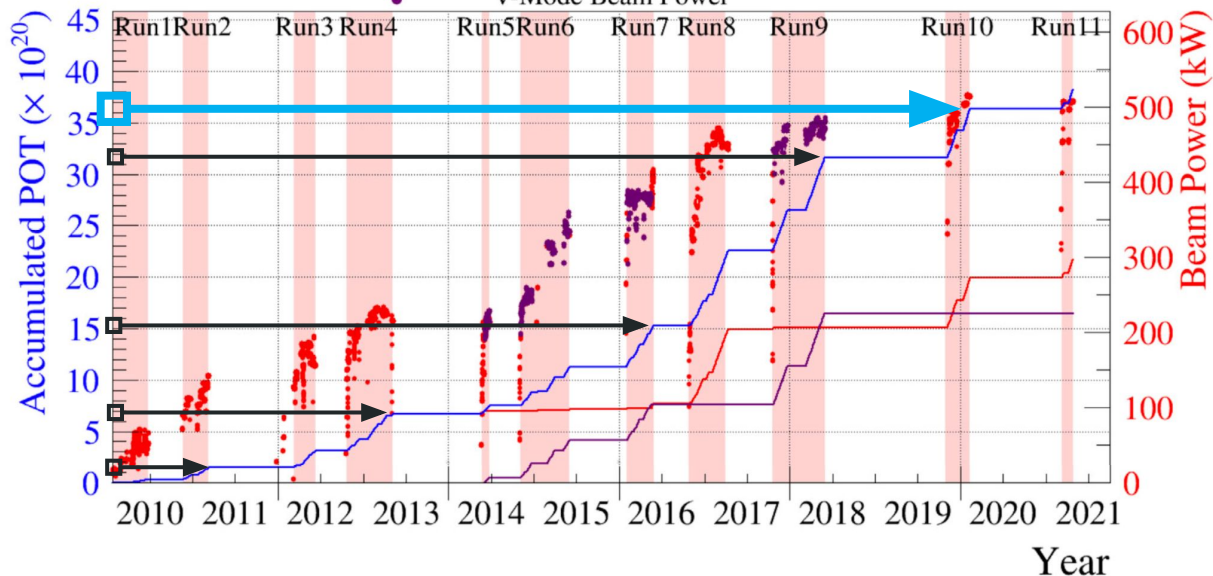
[PRL 121, 171802](https://doi.org/10.1126/science.1211718)

Observation of ν_e appearance

[PRL 112, 061802](https://doi.org/10.1126/science.1120618)

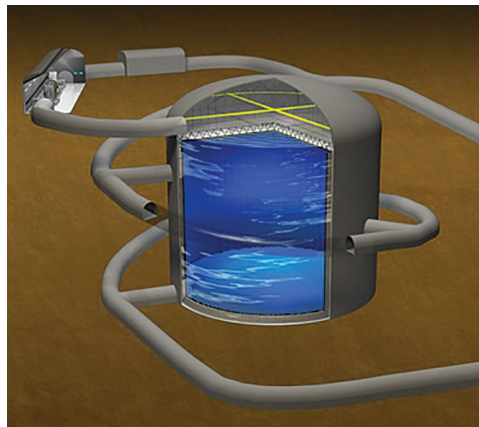
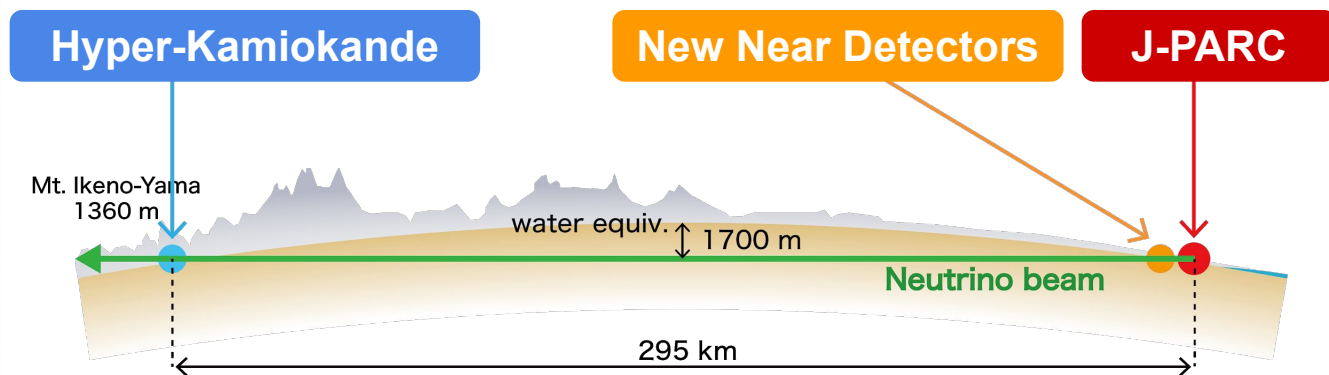
Indication of ν_e appearance

[PRL 107, 041801](https://doi.org/10.1126/science.1107041)



- **515 kW operation achieved in 2020**
- **33% increase of ν -mode data since previous analysis**
- Collected Run 11 data in 2021 with SK-Gd

Overview: Next Generation Experiment

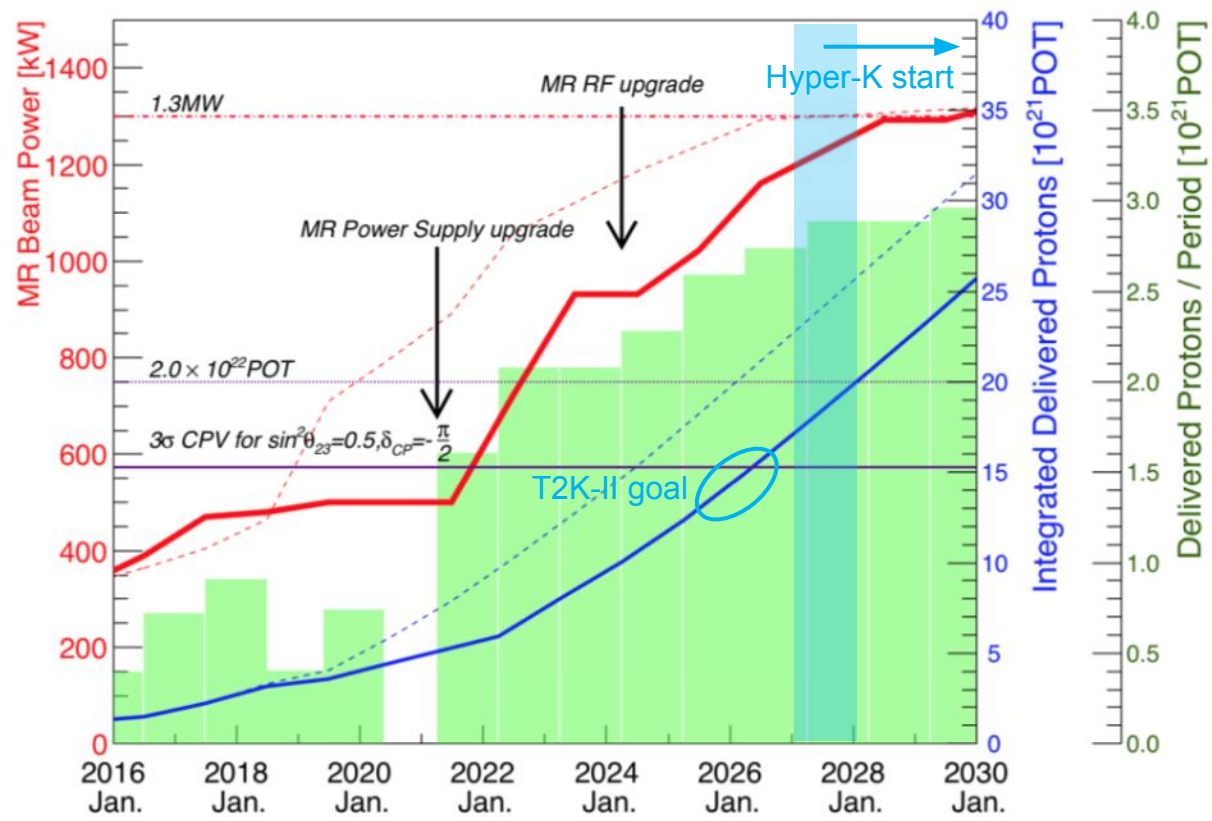


- Bigger and more sensitive than ever
 - Fiducial mass **8x** Super-K
 - J-PARC beam **2.5x** more powerful
 - Neutrino rates **20x** T2K
- Precise systematic understanding becomes critical to the % level
 - New near detectors and photon detectors
 - New calibration and event reconstruction techniques
 - New supporting external data

Beam Line Upgrades Towards T2K-II and Hyper-K

- Increase beam power from ~500 kW to 1.3 MW
- Many upgrades to neutrino beamline components
 - Target, beam monitors, etc.
- Increase horn current from 250 → 320 kA
 - 10% more neutrinos and reduced wrong-sign background

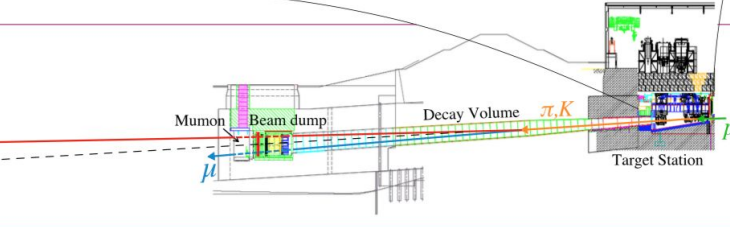
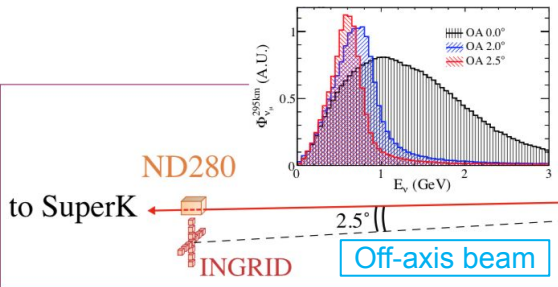
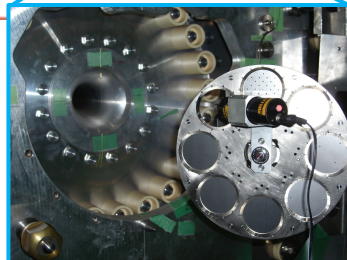
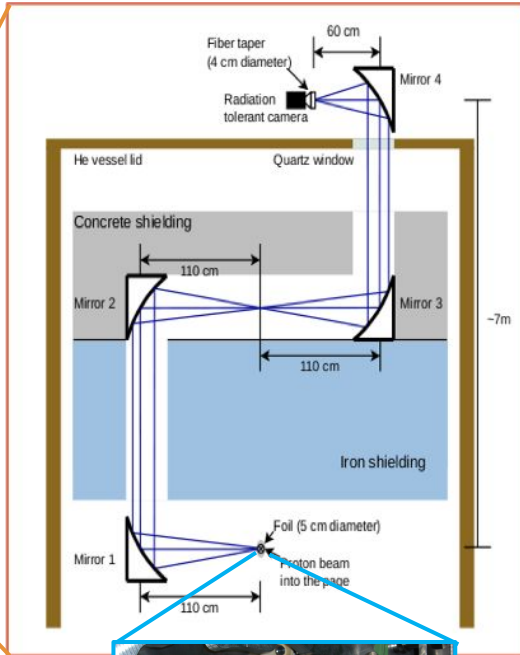
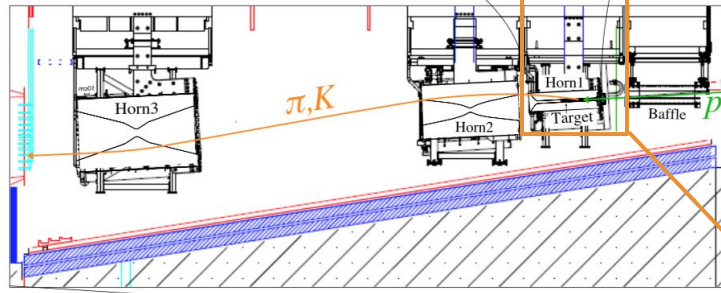
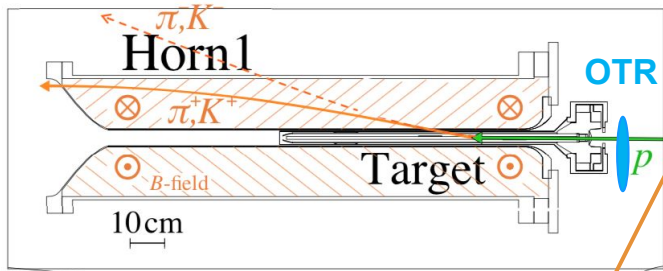
T2K-II Target POT (Protons-On-Target)



Proton Beam Monitoring: Optical Transition Radiation



- Crucial proton beam monitoring and ν beam constraints
- New OTR installation in spring 2022 for T2K-II era and beyond
 - Improving calibration systems
 - New simulations
 - Stress testing new foils

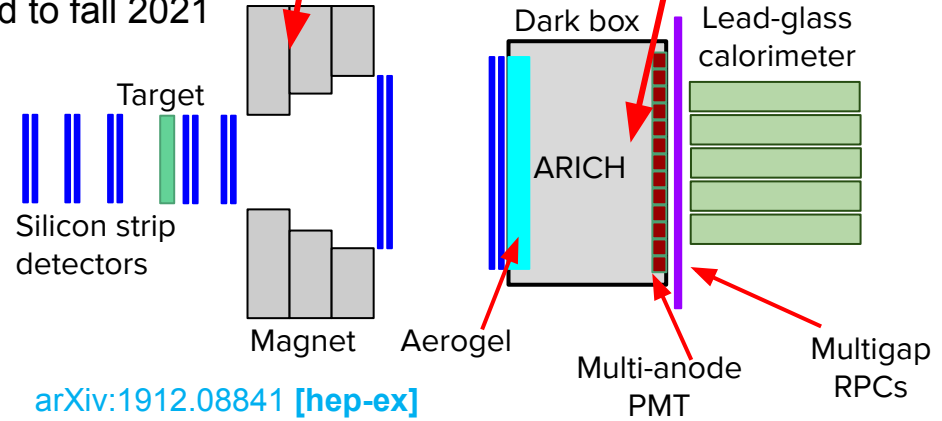
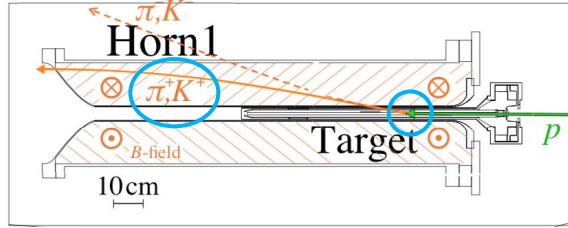
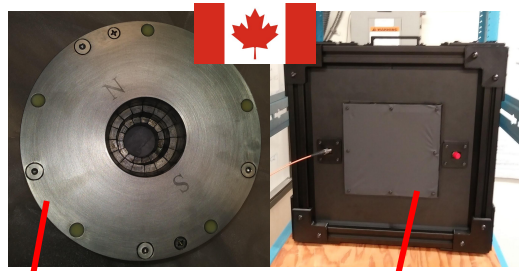


Hadron Production for Neutrino Flux Modeling: EMPHATIC

- Experiment to Measure the Production of Hadrons At a Testbeam In Chicagoland
- Constraints on beam and atmospheric ν flux predictions
 - For T2K, SK, HK, NO ν A, DUNE

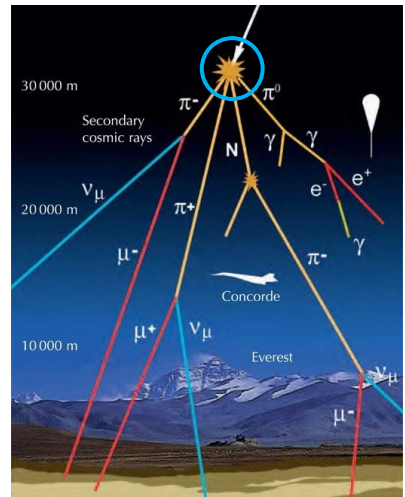
π and K elastic and QE interactions (< 10 GeV/c)
Important systematic uncertainty

- At Fermilab Test Beam Facility
 - **2018**: Pilot run, paper finished collaboration review
 - **2020**: Phase I (limited acceptance 150 mrad) \rightarrow postponed to fall 2021
 - **2022**: Phase II, full acceptance 400 mrad



arXiv:1912.08841 [hep-ex]

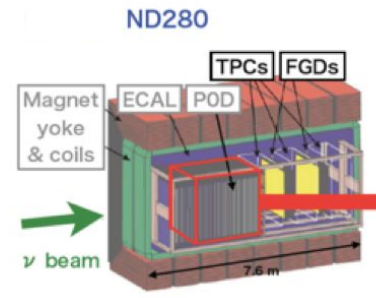
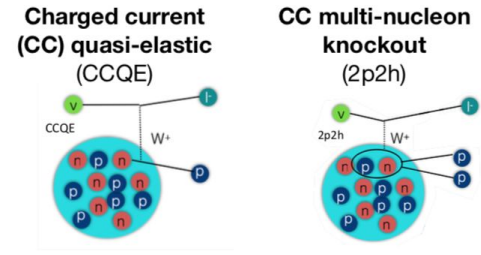
● See poster by [Bruno Ferrazzi](#) (Regina)



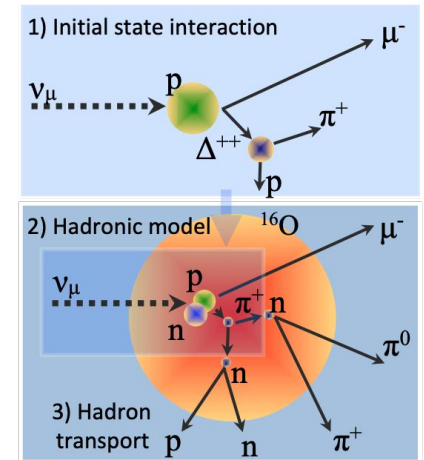
T2K-ND280 ν Cross-Section Measurements



- ν -nucleus interactions: a theoretical challenge that must be met by rich experimental measurements
- Many recent T2K xsec papers, e.g.:
 - CC0 π on C+O ([PRD 101. 112004 \(2020\)](#))
 - $\nu_{\mu}/\bar{\nu}_{\mu}$ CC0 π ([PRD 101. 112001 \(2020\)](#))
 - $\nu_e/\bar{\nu}_e$ CC - First $\bar{\nu}_e$ in 43 years! ([JHEP 2020. 114 \(2020\)](#))
 - CC1 π +p - Nuclear effects ([arXiv:2102.03346 \[hep-ex\]](#))
- See also talk by [Mitchell Yu \(York\)](#):
 - $\nu_{\mu}/\bar{\nu}_{\mu}$ CC coherent π production



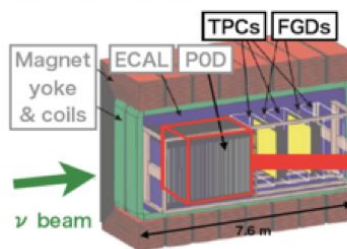
- ND280 off-axis detector
- Active **scintillator** + passive **water** targets
 - **Tracking** with time projection chambers
 - **Magnetized** for charge & momentum measurement



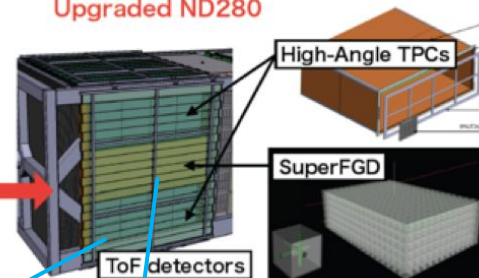
Future of the Near Detector Suite



Current ND280

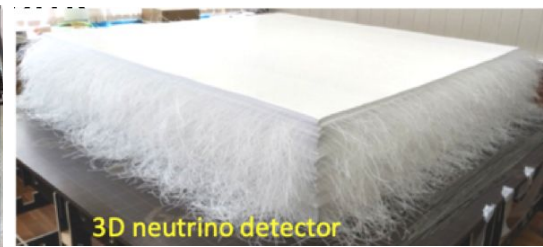
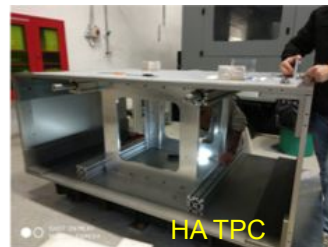


Upgraded ND280



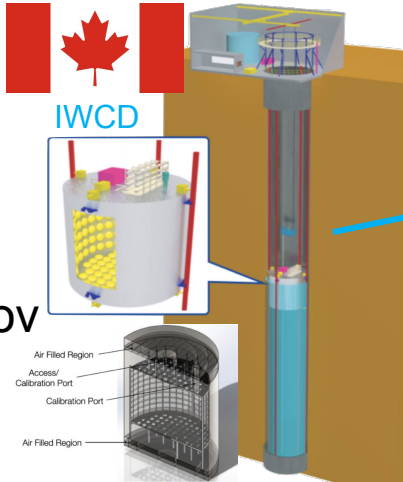
- ND280 upgrade in 2022

- Increase phase space coverage, similar to SK
- Lower proton energy threshold and neutron detection capability



Future of the Near Detector Suite

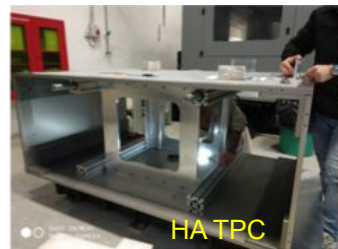
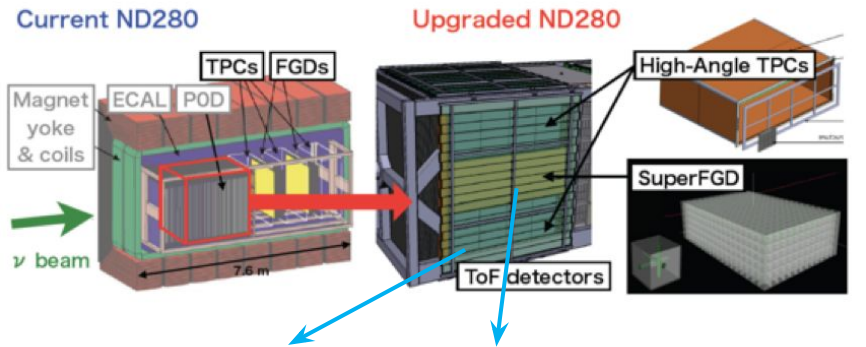
- Canada-led, novel off-axis spanning Intermediate Water Cherenkov Detector for Hyper-K



- Handle on far detector observables' dependence on neutrino energy
- Precise cross-section measurements on water
- See [Mark Hartz's \(TRIUMF\) talk at TIPP](#)

- ND280 upgrade in 2022

- Increase phase space coverage, similar to SK
- Lower proton energy threshold and neutron detection capability



IWCD & Hyper-K Photosensor Development

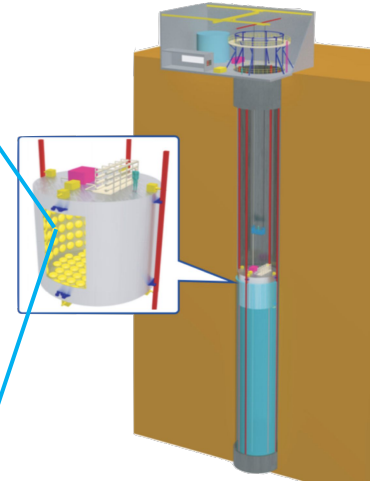
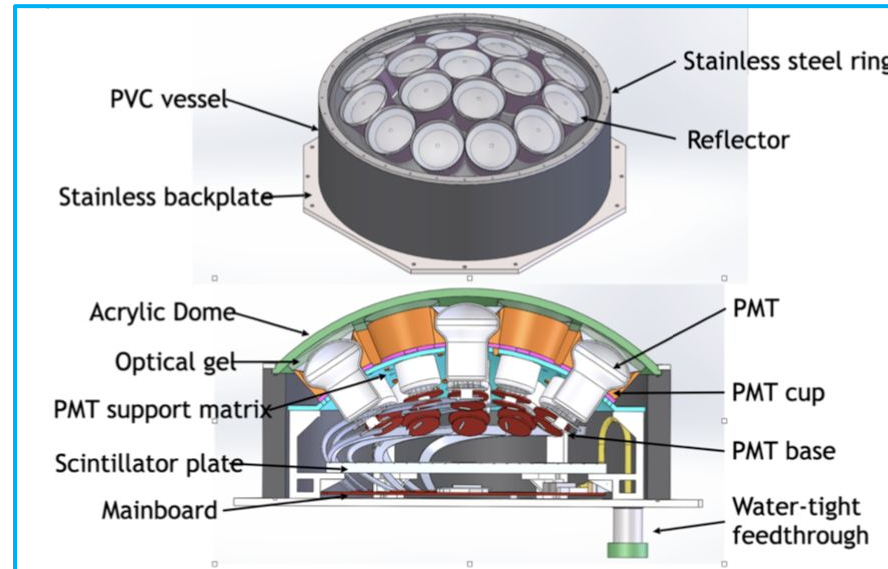


- Multi-PMT:
19 x 3" diameter
PMTs in a water-
tight vessel with
HV and electronics

- Canadian
contribution
to the IWCD

- ~250 out of 480 mPMTs
- Aiming for 400 from CFI 2023 for Hyper-K far detector contribution

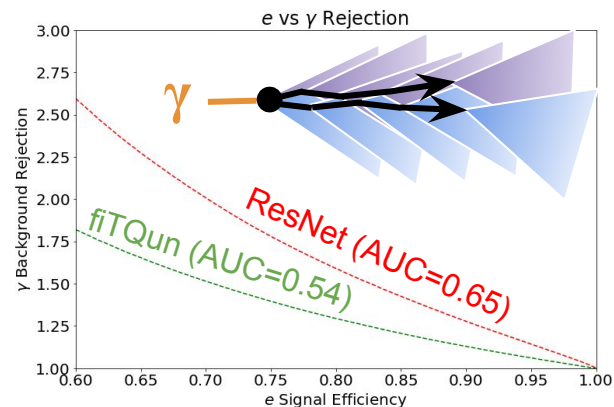
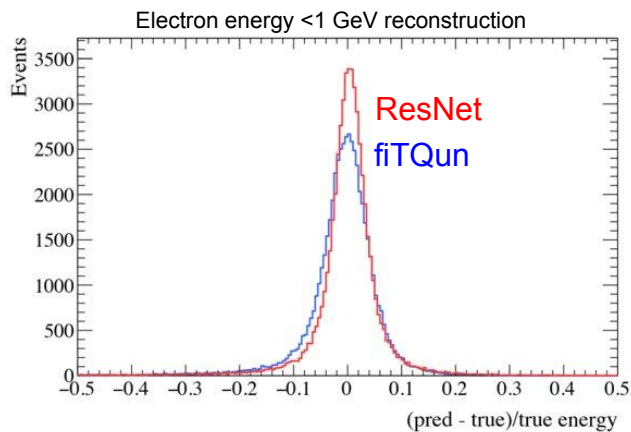
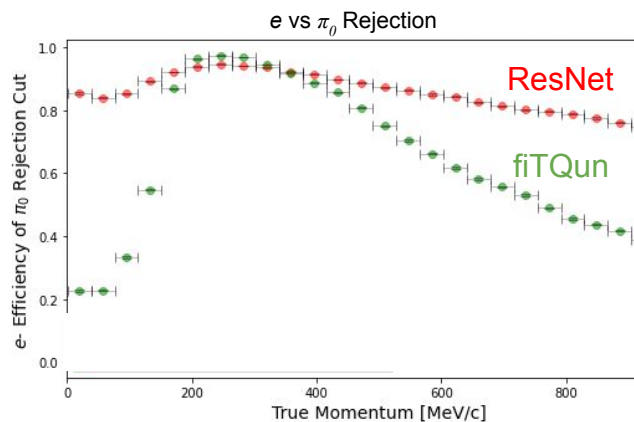
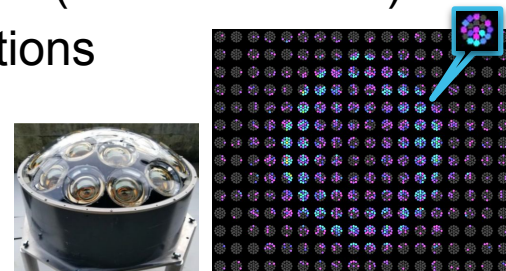
- See talk by [Luan Koerich \(Regina\)](#)



Machine Learning Event Reconstruction



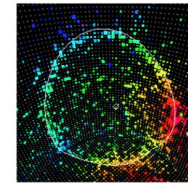
- Canada leading the Water Cherenkov Machine Learning (WatChMaL) consortium
 - Towards a unified platform and knowledge base across many such detectors
- Improved particle classification and regression/reconstruction (in IWCD below)
- Massive processing speed-up enables multitudes of simulations for detector design and systematics studies
- See talk by [Nick Prouse \(TRIUMF\)](#)



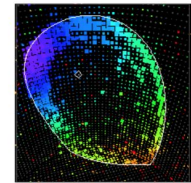
Super-Kamiokande - 25 Years of ν and Astrophysics



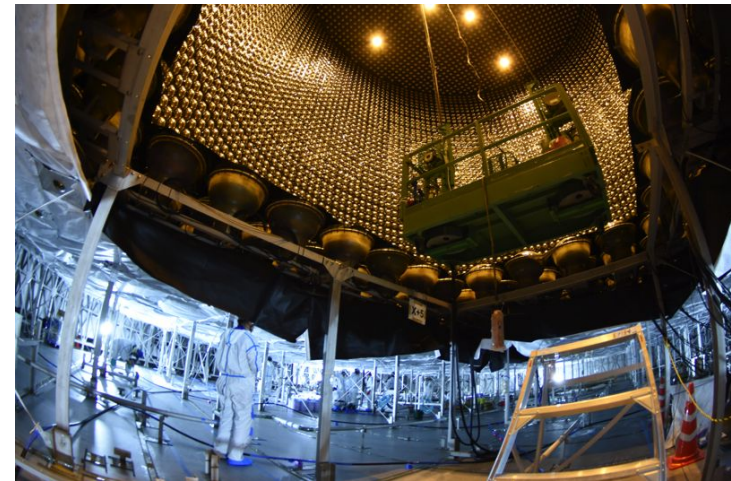
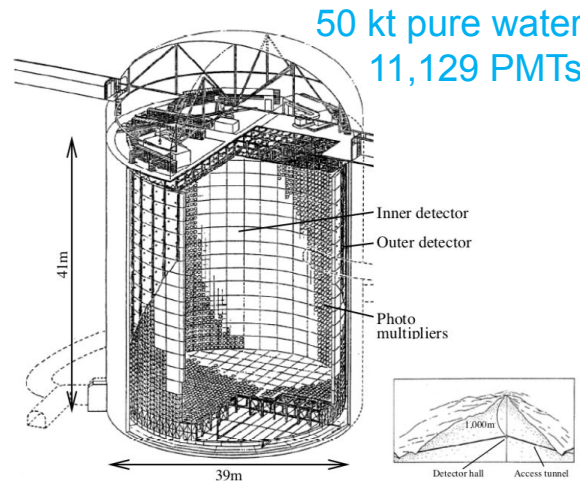
- 1998: discovery of atmospheric ν flavor transformation
- 2001: discovery of solar ν flavor transformation with SNO
- 2004: confirmation of atmospheric ν oscillation by K2K
- 2012: first evidence for τ appearance
- 2013: first direct indication of ν osc. matter effects
- Ongoing searches for nucleon decay, DM, supernovae...



ν_e -like



ν_μ -like

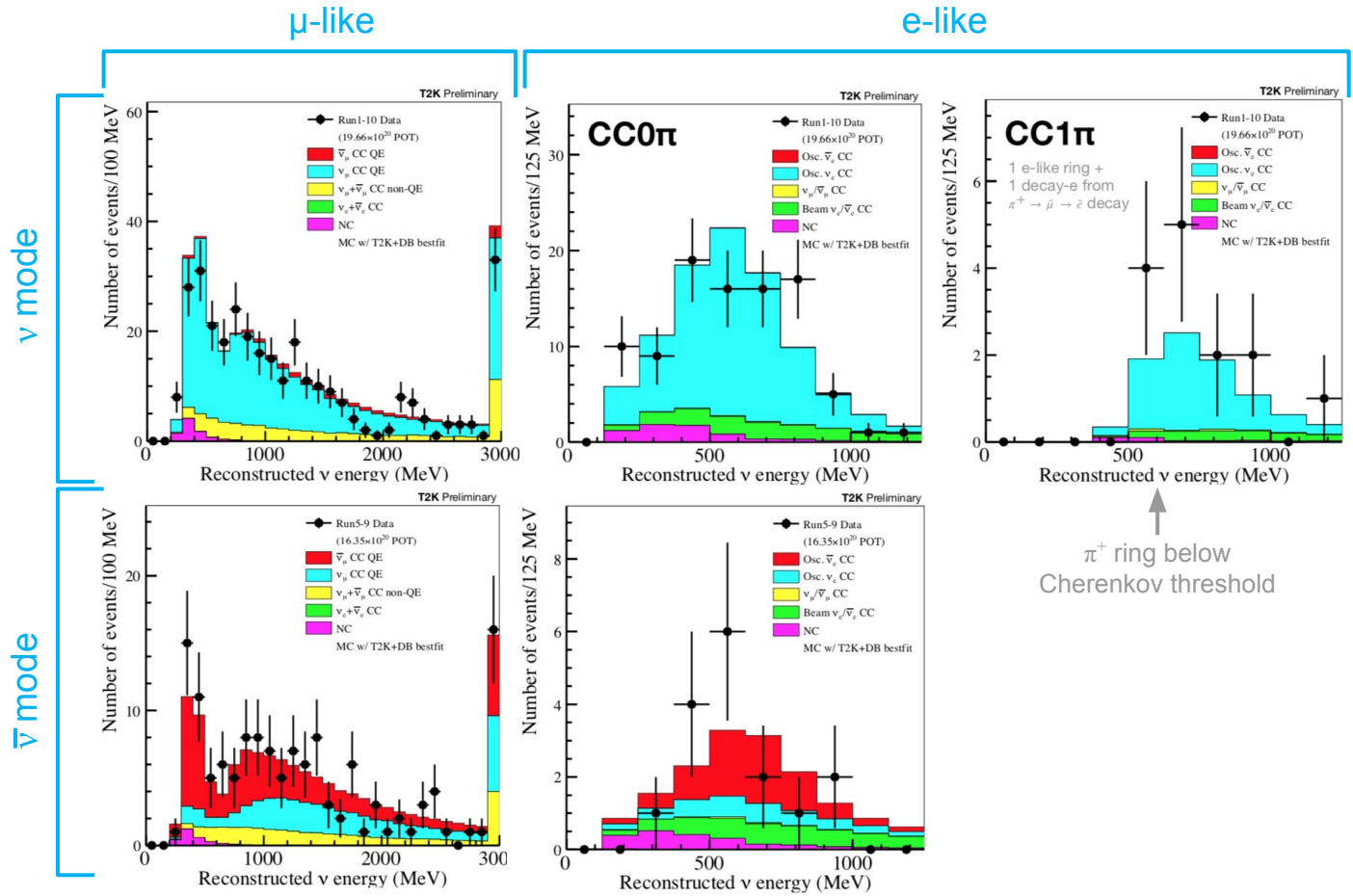


T2K-SK Single-Ring Datasets for CPV Analysis

- Updated Super-K datasets used for 2020 CPV analysis

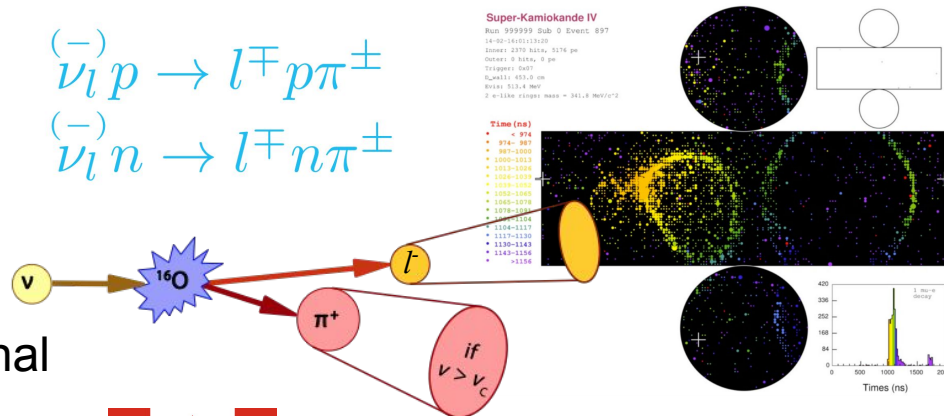
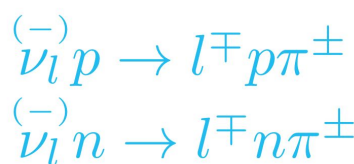
- ν_μ disappearance
- ν_e appearance
- Only one visible Cherenkov ring

- Event reconstruction (fitQun) and detector systematic analyses developed in Canada 🇨🇦

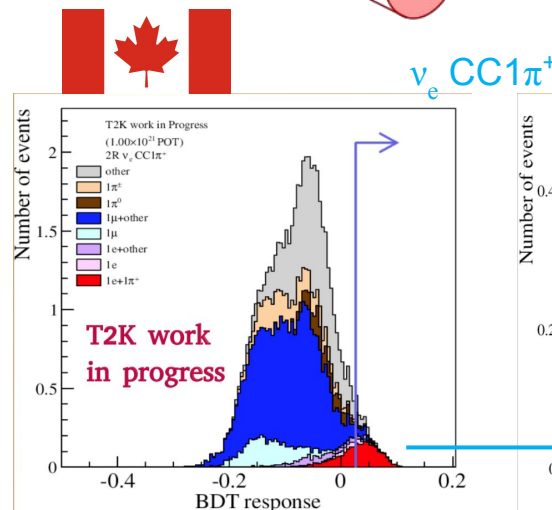
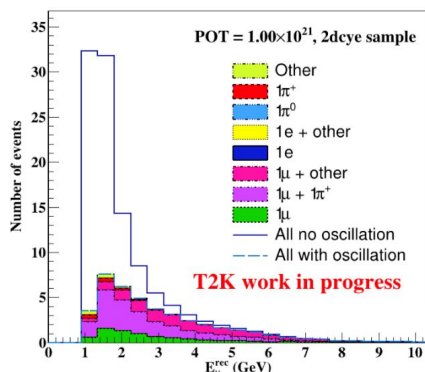
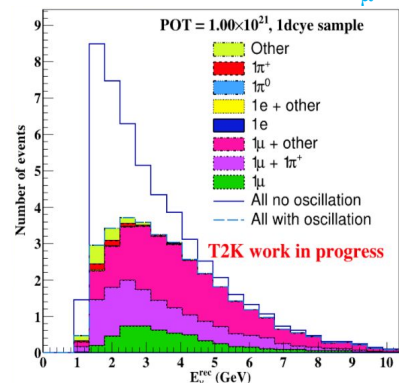


T2K-SK Multi-Ring Datasets for Future Analyses

- Second dominant interaction channel: **resonant 1π production**
- Expected to improve oscillation parameter measurements
 - E.g. $\sim 12\%$ increase in ν_e signal statistics
- New BDT pushing the limits of traditional likelihood reconstruction algorithm



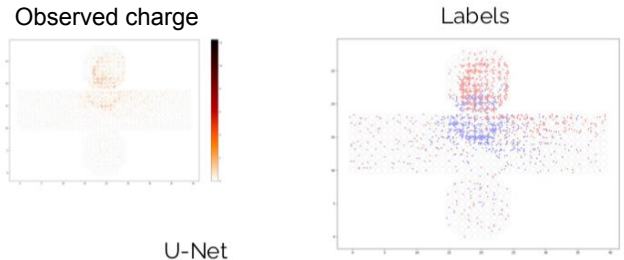
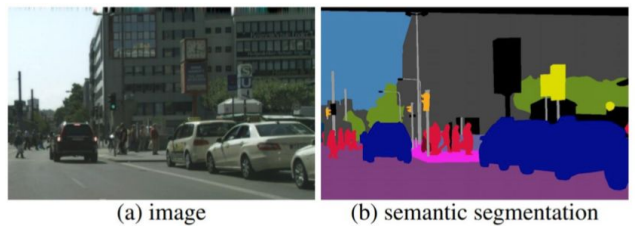
ν_{μ} CC $1\pi^{+}$



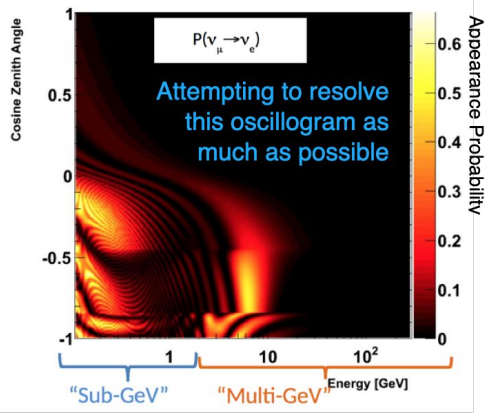
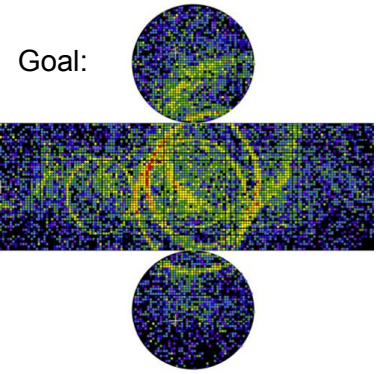
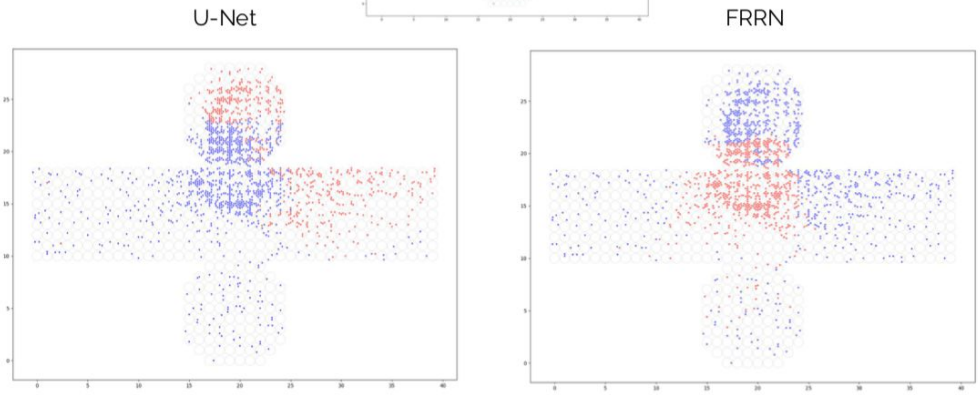


Multi-Ring Reconstruction in the Further Future

- More machine learning: panoptic segmentation
- Towards improving multi-ring & multi-GeV event classification and reconstruction
 - ν mass ordering, ν_τ appearance, δ_{CP}
- See talk by [Wojtek Fedorko \(TRIUMF\)](#)
 - Machine Learning Applications in Particle Physics: Present and Future

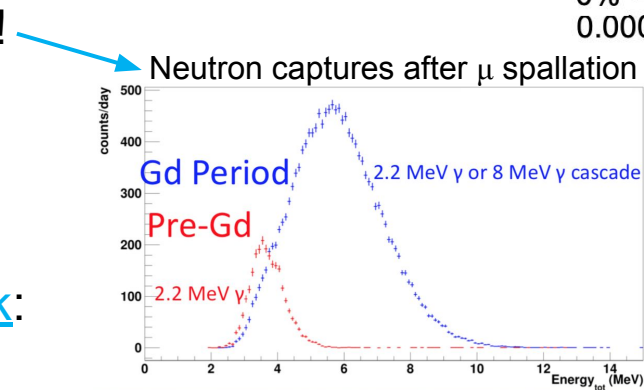
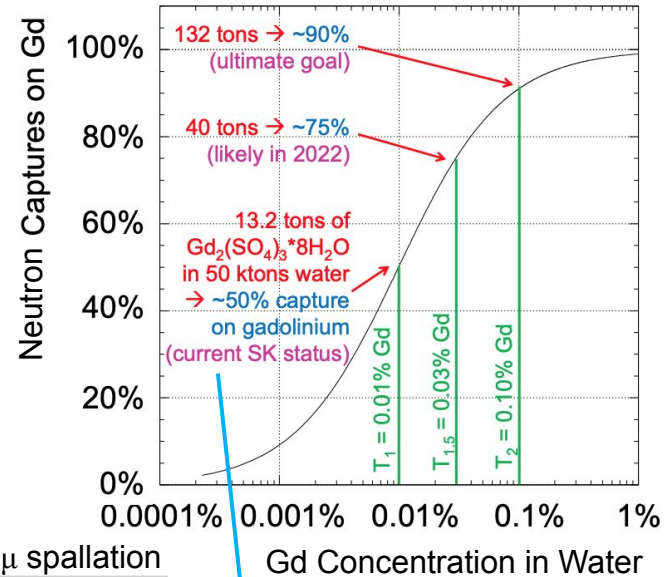
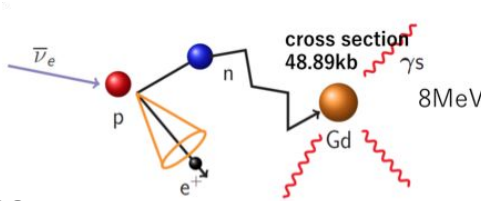


First attempt on π^0 decay events in IWCD:
~80% accuracy



Super-K Gadolinium Upgrade

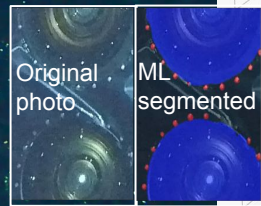
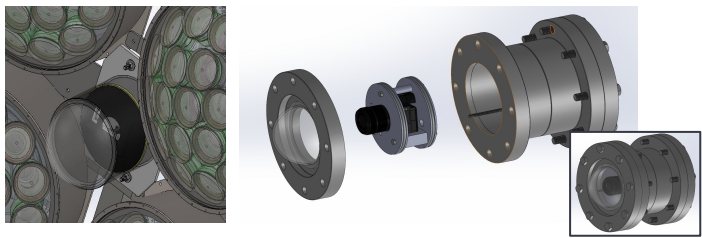
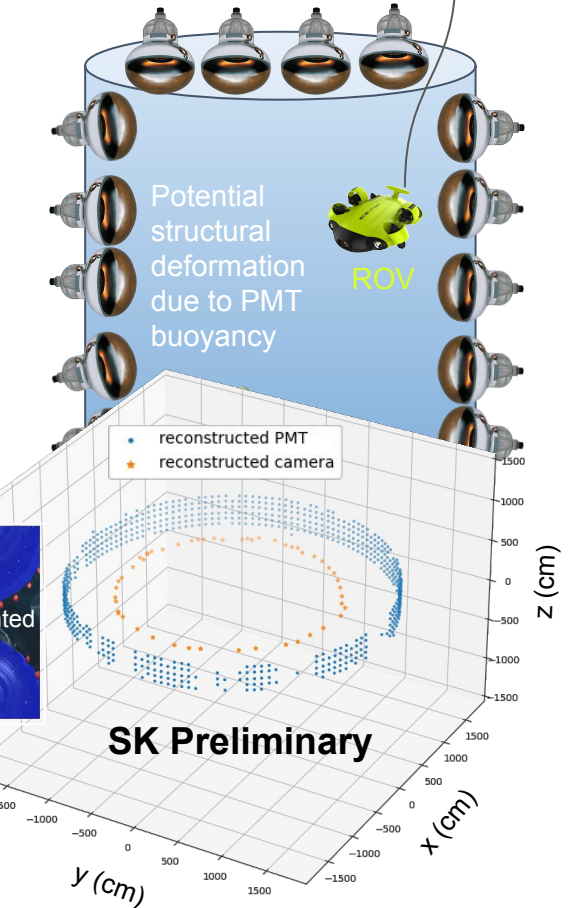
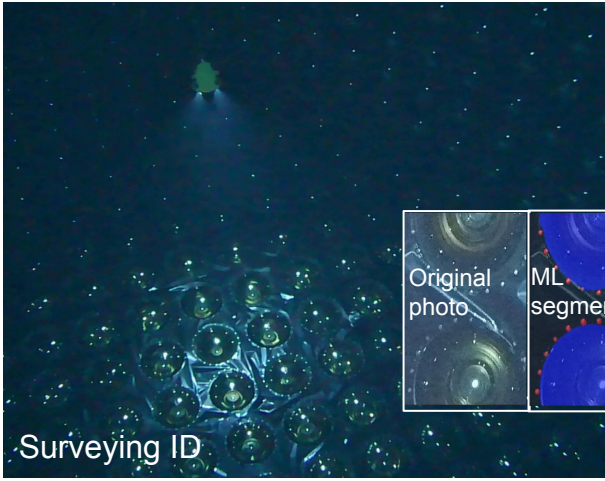
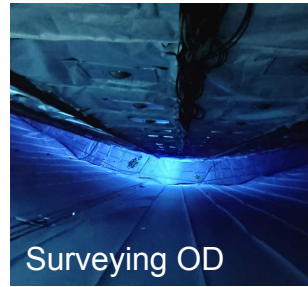
- Improved neutron detection
- More physics potentials:
 - Search for solar antineutrinos
 - Background reduction for proton decay
 - Diffuse supernova neutrino background (DSNB)
 - Improved $\nu/\bar{\nu}$ discrimination $\rightarrow \delta_{CP}$, mass ordering
- Detector refurbished in 2018 to prepare for $\text{Gd}_2(\text{SO}_4)_3$ loading in 2020 🇨🇦
- Gd neutron captures observed!
- T2K Run 11 in March brings first beam events with Gd
 - Run 12 expected after $T_{1.5}$
- See also [Ryosuke Akutsu's talk](#):
 - Study of neutrons associated with neutrino interactions in water with the IWCD





Novel Detector Geometry Calibration

- First underwater survey of Super-K detector geometry
- Challenging photogrammetry analysis ongoing
 - Demonstrated with a ring of ID barrel PMTs
- Developing new systems for Hyper-K and IWCD
 - Critical for a moving detector

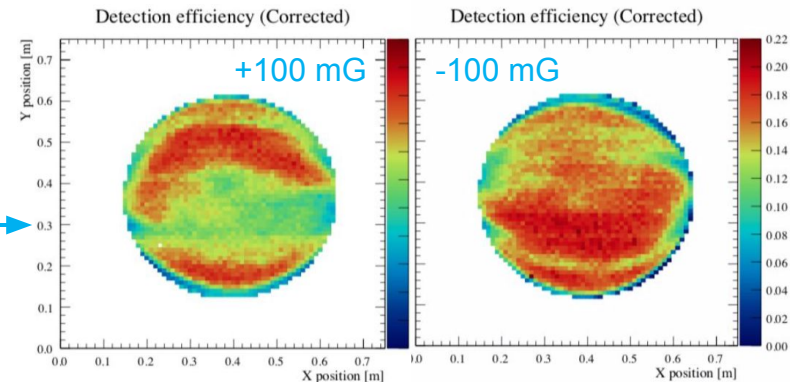
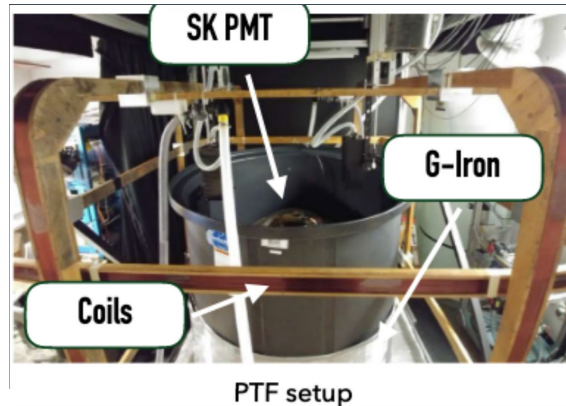
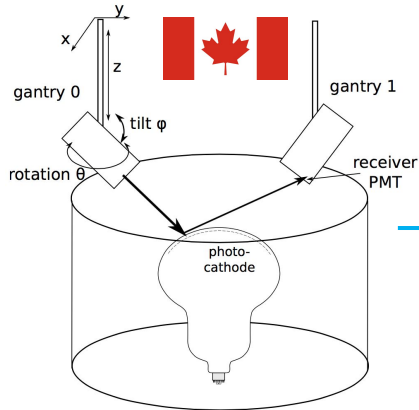
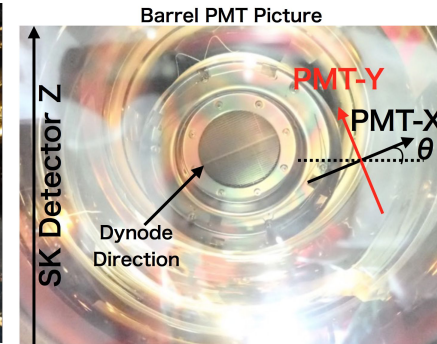


- See talk by [Blair Jamieson \(Winnipeg\)](#) and poster by [Michael Sekatchev \(TRIUMF/UBC\)](#)

Precise and Comprehensive PMT Characterization

- Uncertainties in PMT response is a major systematic in water Cherenkov detectors
- (Re)Building a photosensor test facility at TRIUMF for Super-K and Hyper-K/IWCD
- See talk by [V. Gousy-Leblanc \(UVic\)](#) and poster by [S. Wingfelder \(TRIUMF\)](#)

Magnetic field and PMT orientation survey throughout Super-K



The Water Cherenkov Test Experiment

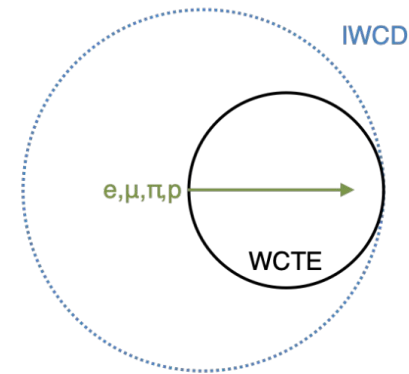
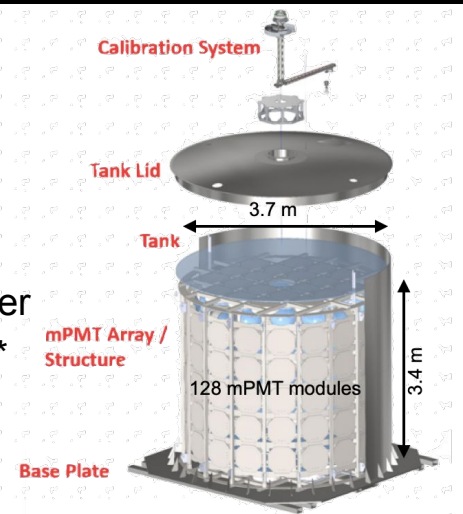
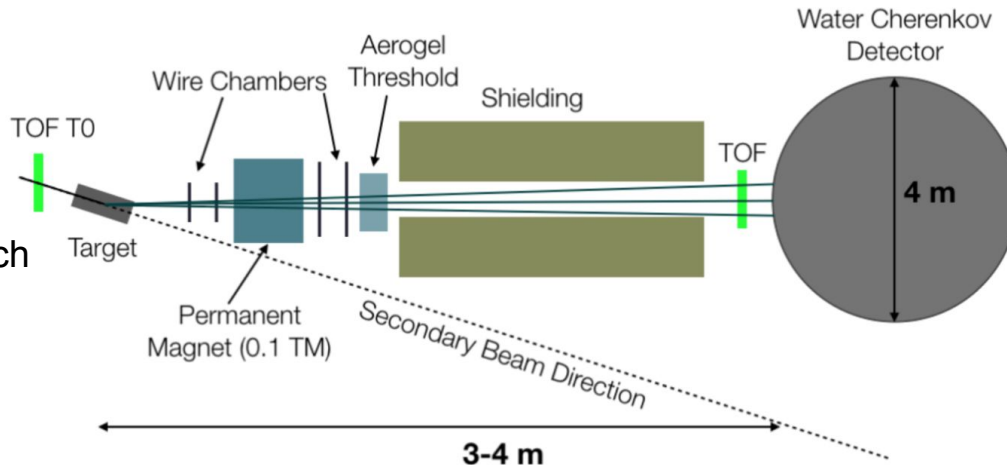


- Prototype detector for beam test at CERN in 2023
- mPMT pilot run and test-bed for precision calibration and ML
- Well-understood p , e , π^\pm , μ^\pm particle beam from 140-1200 MeV/c
 - Control samples to constrain neutrino experiment modeling:
 - Detector response: Cherenkov light emission; π^\pm re-interactions in water
 - Neutrino flux predictions: hadron production (including K not at FTBF)*

● See talk by [Matej Pavin \(TRIUMF\)](#)

● Proposal:

- [CERN-SPSC-2019-042](#) : [SPSC-I-254](#)
- Will be recommended to CERN Research Board by SPSC
- *EOI submitted
- [CERN-SPSC-2021-021](#) : [SPSC-EOI-021](#)



Generations of Kamiokande

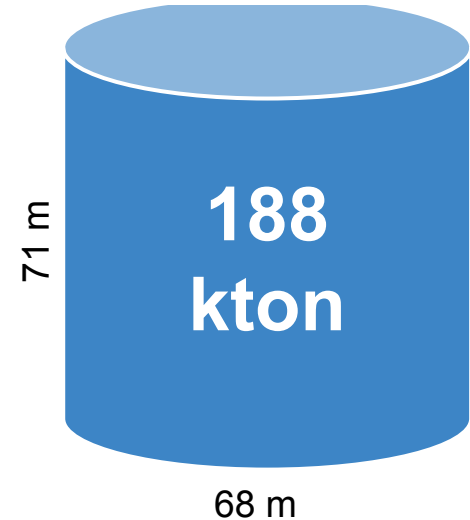
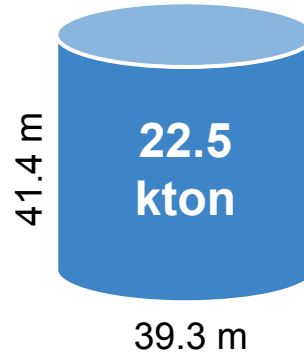
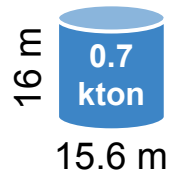
Kamiokande



Super-Kamiokande



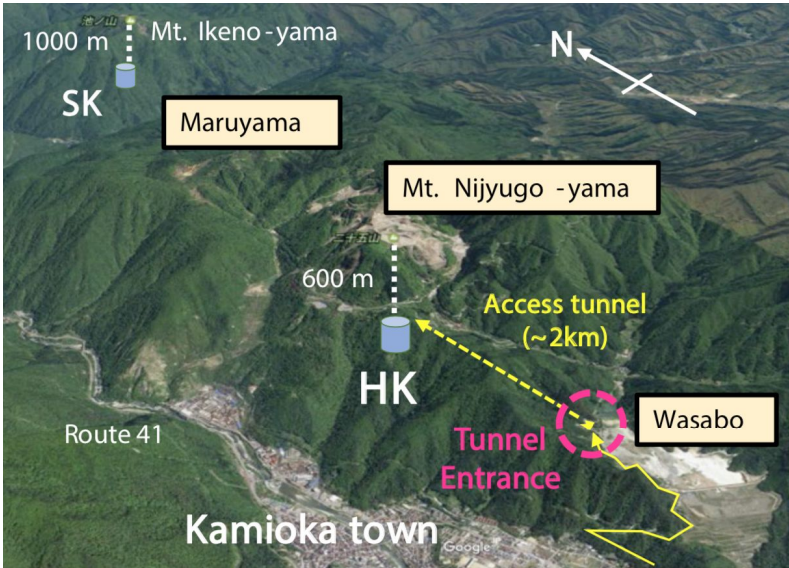
Hyper-Kamiokande

Fiducial
Volumes



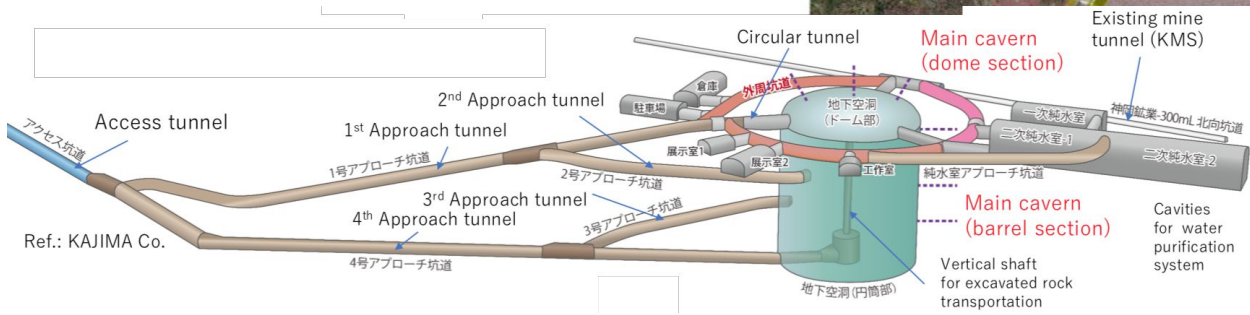
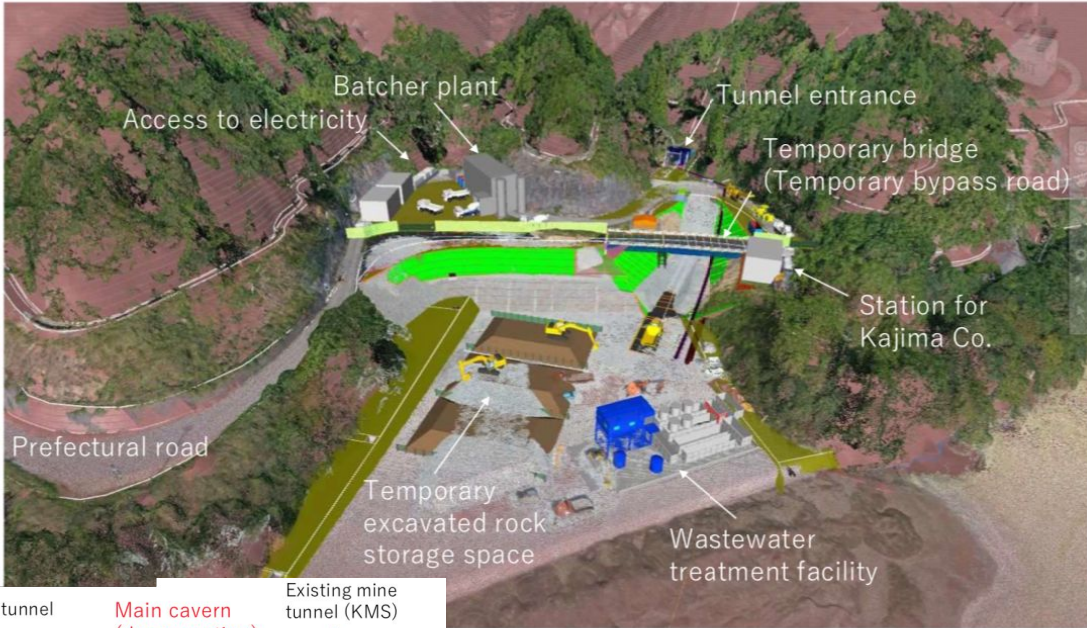
1983 - 1996	1996 - today (and beyond)	2027 - (and beyond)
Supernova 1987A	Atmospheric ν Oscillation	δ_{cp} , proton decay, indirect DM search, more SNs, ...
 2002 Physics	 2015 Physics	

Hyper-K Far Site Overview



Overall plan of the entrance yard

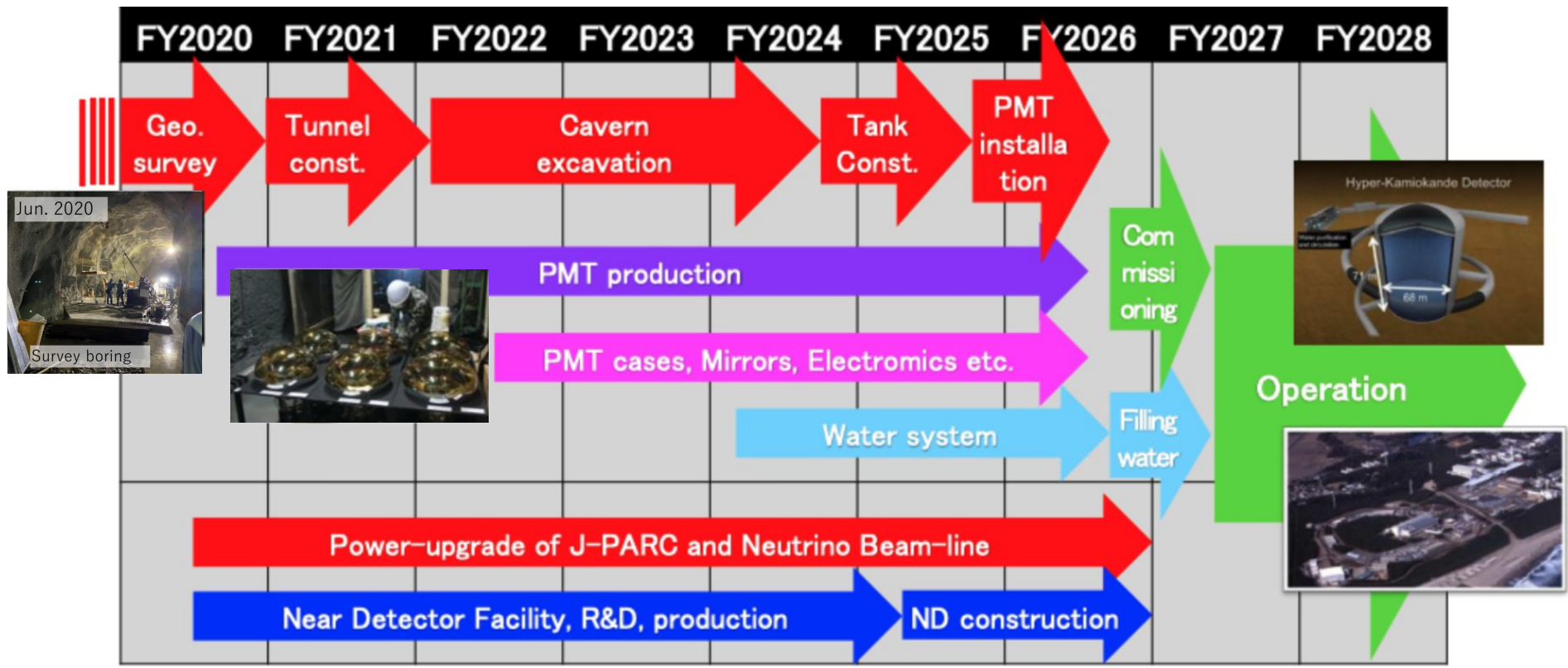
Jan 28, 2021, Kajima Corporation.



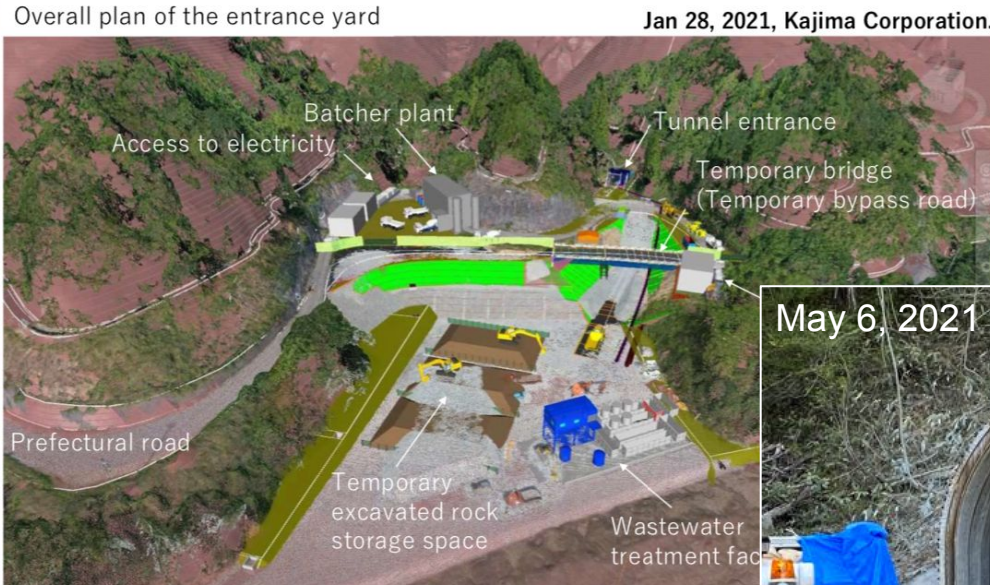
- Entrance yard, tunnel, and cavern conceptual designs

Hyper-K Schedule

Finish all preparations within ~4.5 years from now for detector installation



Hyper-K Concept Becoming Reality



- Entrance yard and tunnel/cavern excavation progressing on schedule

May 28, 2021: Groundbreaking Ceremony



東京大学

ハイパーカミオカンデ 着工記念式典

Hyper-Kamiokande Groundbreaking Ceremony

ICRR

宇宙線研究所

Rich Science with Hyper-Kamiokande

Design report: [arXiv:1805.04163](https://arxiv.org/abs/1805.04163)

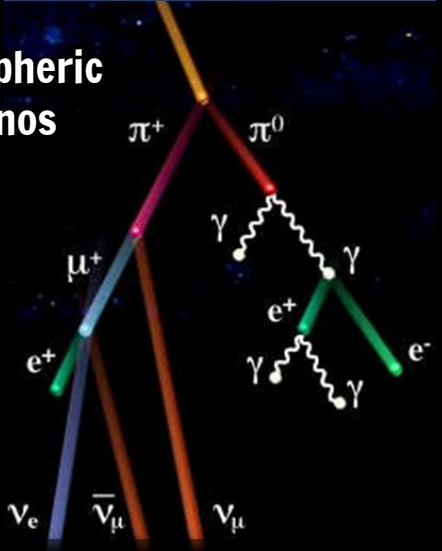
**Multi-Messenger:
Supernova, GW, ...**



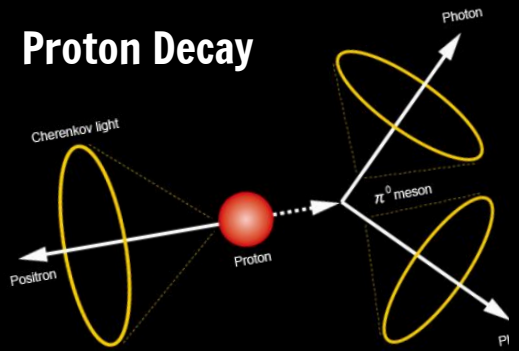
Dark Matter



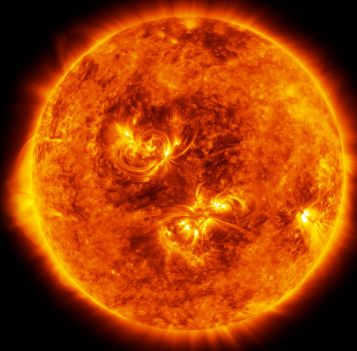
**Atmospheric
Neutrinos**



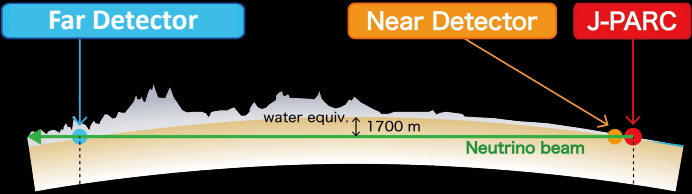
Proton Decay



Solar Neutrinos



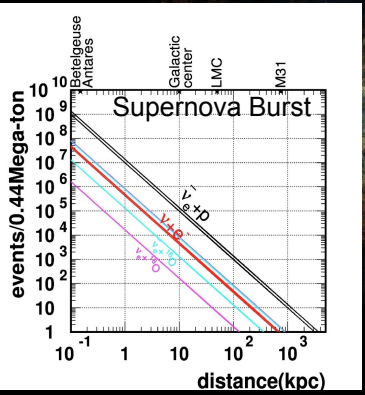
Accelerator Neutrinos



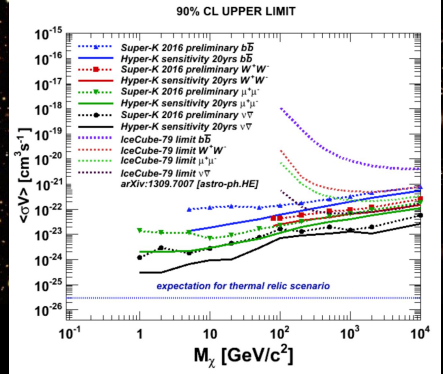
Rich Science with Hyper-Kamiokande

Multi-Messenger

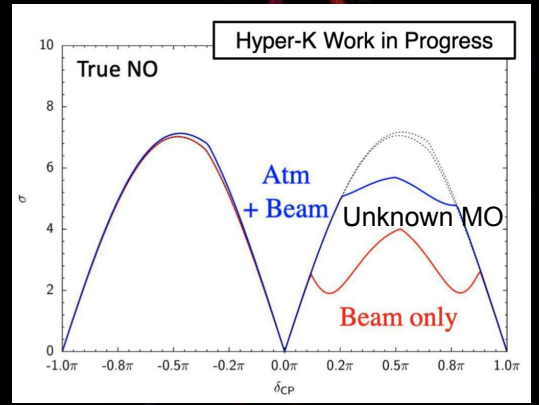
[[arXiv:2101.05269](https://arxiv.org/abs/2101.05269)]



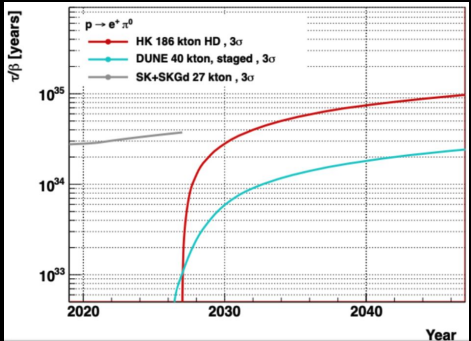
Dark Matter



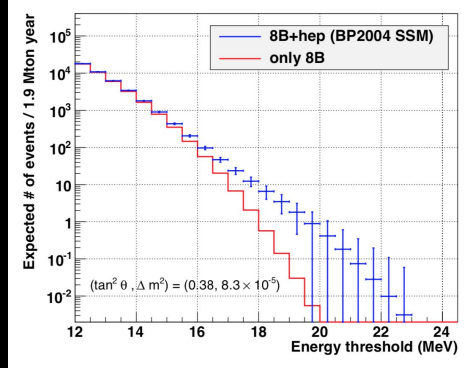
Atmospheric Neutrinos



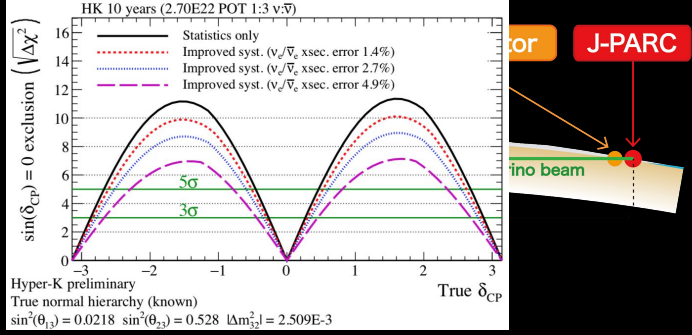
Proton Decay



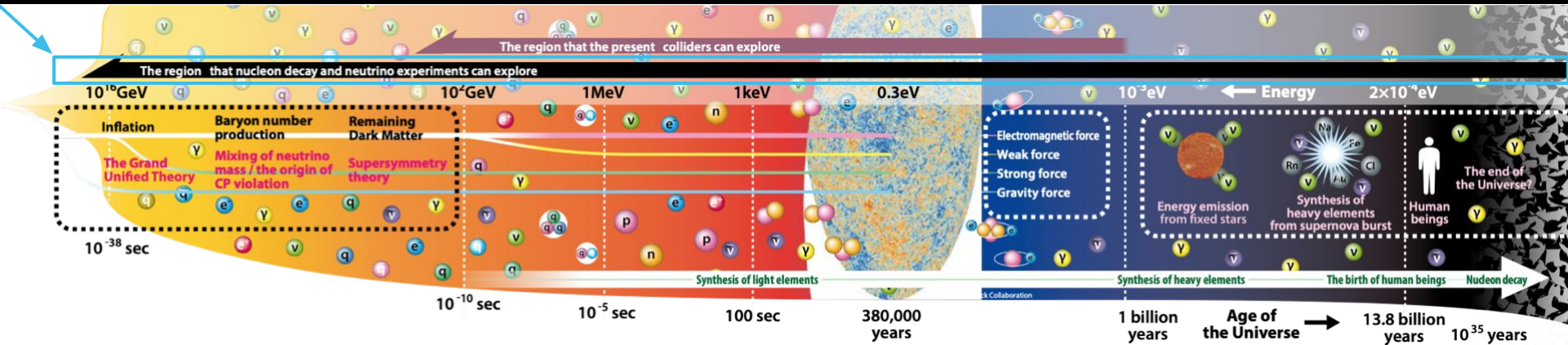
Solar Neutrinos



Accelerator Neutrinos



Summary



- Ongoing breadth of rich, world-leading physics: T2K(-II), SK-Gd, Hyper-K
- Near term auxiliary efforts to enhance all of the above:
 - WatChMaL, WCTE, EMPHATIC
- Canadian involvement on all fronts
 - *New collaborators welcome and encouraged!*



BRITISH COLUMBIA INSTITUTE OF TECHNOLOGY

+ ... ?

Overview of T2K / SK / HK Contributions at CAP Congress

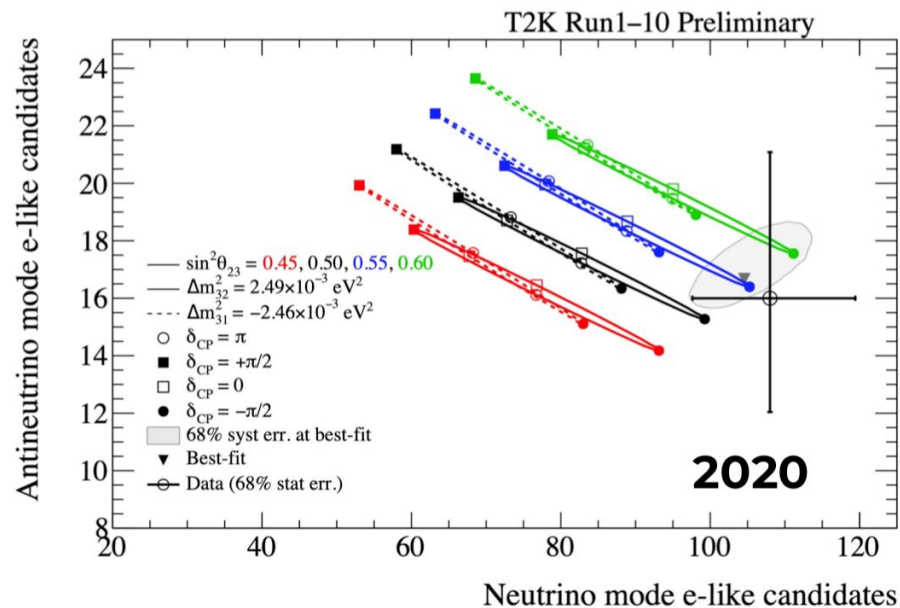
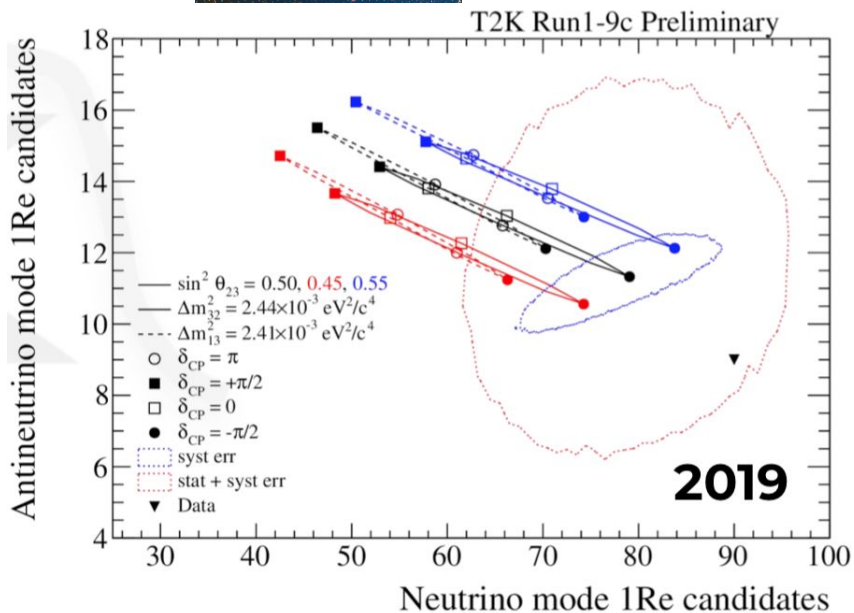
Name	Title	Date	Time (PDT)
Nick Prouse	Machine learning techniques for improving water Cherenkov event reconstruction	June 7	9:45
Luan Koerich	A Multi-Photomultiplier Photosensor Module for IWCD/Hyper-K	June 9	10:05
Michael Sekatchev	Automated Feature Detection and Camera R&D for Photogrammetry in Super-K and Future Water Cherenkov Neutrino Detectors	June 9	10:45
Skylar Wingfelder	Hyper-Kamiokande Photosensor Test Facility Recommissioning and System Upgrades.	June 9	10:55
Bruno Ferrazzi	Optical reflectors in an ARICH detector for a hadron production experiment	June 9	11:19
Vincent Gousy-Leblanc	Super-Kamiokande PMT characterizations using artificial magnetic field and robotic laser-equipped arms	June 10	8:45
Ryosuke Akutsu	Study of neutrons associated with neutrino interactions in water with the IWCD detector	June 10	9:55
Mitchell Yu	Measuring the muon (anti-)neutrino induced charged-current coherent pion production cross sections on carbon using the off-axis T2K near detector	June 10	10:05
Wojtek Fedorko	Machine Learning Applications in Particle Physics: Present and Future	June 10	12:45
Matej Pavin	Water Cherenkov Test Experiment	June 10	12:45
Blair Jamieson	Photogrammetry in Water Cherenkov Neutrino Detectors	June 10	13:15

Appendix

T2K CP Violation Constraints

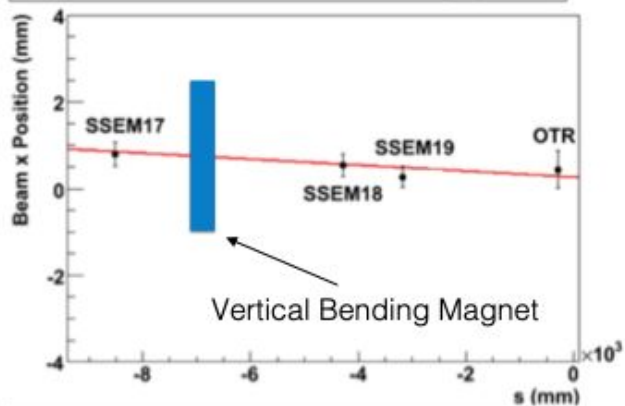


- 2019 analysis:
 - Disfavored $\delta_{CP} = 0$ at 3σ
 - Disfavored IO at 1σ
- 2020 analysis slightly looser constraints

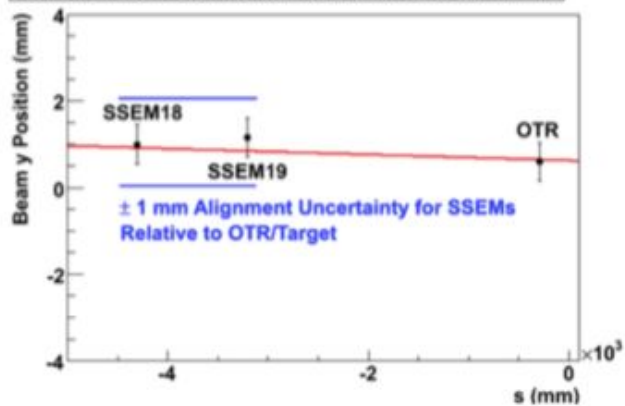


OTR Impact

Example Fit to Proton Monitors for x at the Target (s=0 mm)



Example Fit to Proton Monitors for y at the Target (s=0 mm)



Measurements are used to extrapolate the beam position

to reduce uncertainties

σ_x -position uncertainty:

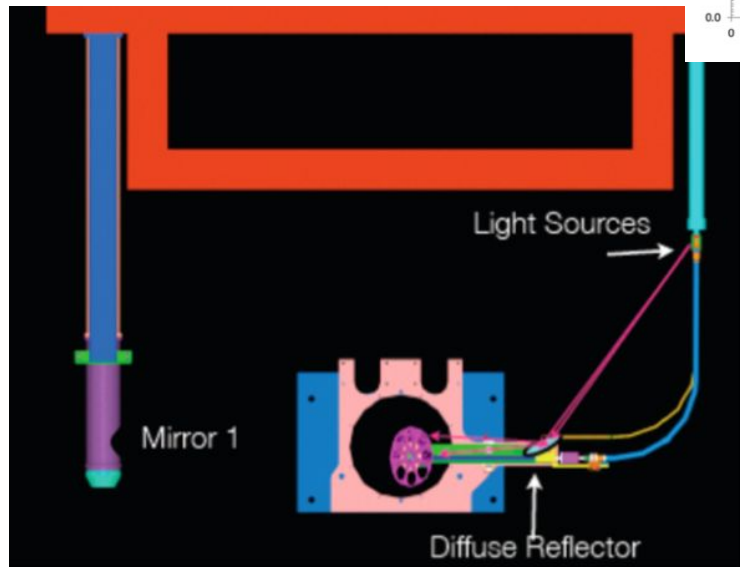
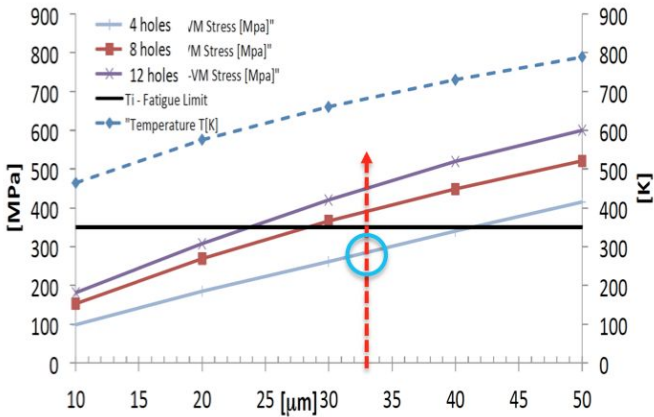
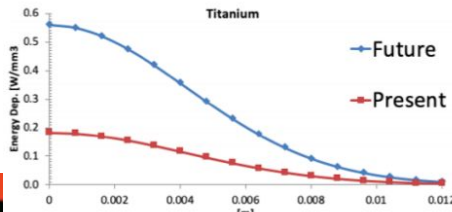
	Extrapolation uncertainty	
	Without OTR	With OTR
Pos. X (mm)	0.5	0.5
Pos. Y (mm)	2.3	0.5
Angle X (mrad)	0.08	0.08
Angle Y (mrad)	0.5	0.3



New OTR System for T2K Upgrade

- New horn for T2K upgrade era
 - New OTR installation in spring 2022
- Improving calibration systems
 - Including new simulations
- Stress test of new foils

Beam Conditions	Present	Upgrade
Nb. Protons 30 GeV	2.5e14	3.2e14
Repetition Time [s]	2.2	1.16
Beam Energy [MW]	0.545	1.32



J-PARC and Neutrino Beamline Upgrades

Horn PS upgrade
(250kA → 320kA)



New horn2 for improved cooling



- Major upgrade will be planned in FY2021-2022 long shutdown
- Budget allocation by MEXT as requested in FY2021
- Many beamline components already produced or under production



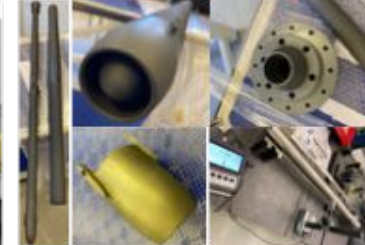
New bending magnet



Additional water disposal tank

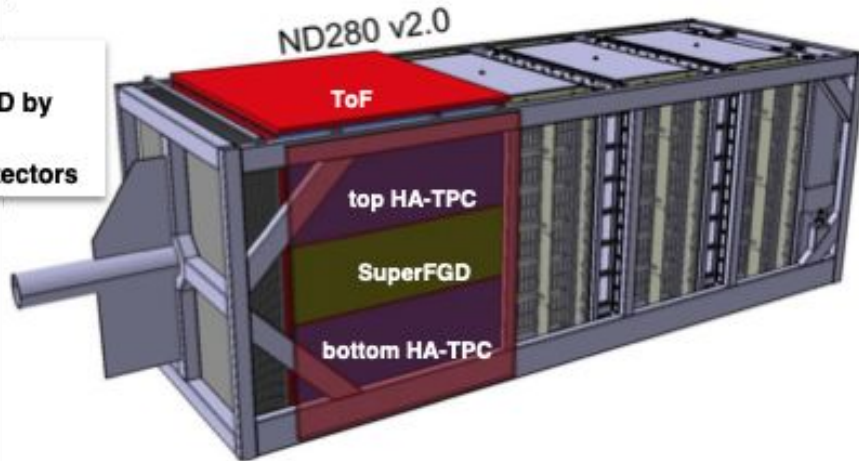


1.3MW-target development



T2K ND280 Upgrade

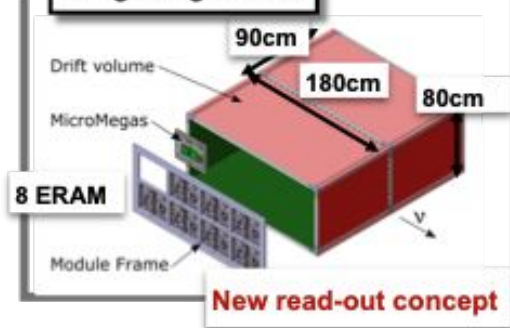
Replace P0D by
new subdetectors



Milestones

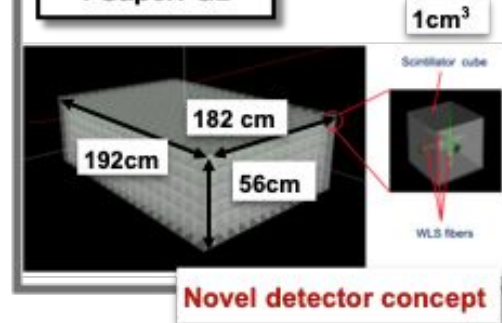
- ✦ 2018 → TDR [arXiv:1901.03750](https://arxiv.org/abs/1901.03750)
- ✦ 2021/22 final modules
- ✦ 2022 installation

2 High Angle TPCs



NIM A 957 163286 (2020)

1 SuperFGD



JINST 13, P02006 (2018)

6 ToF panels

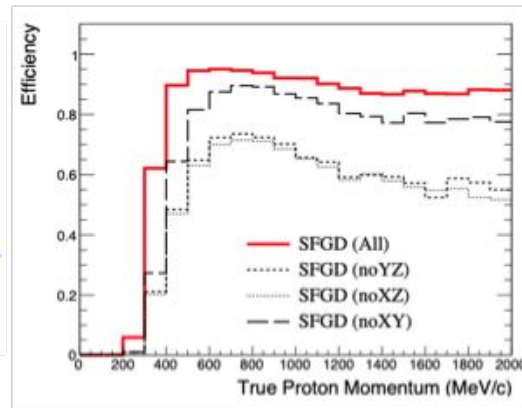
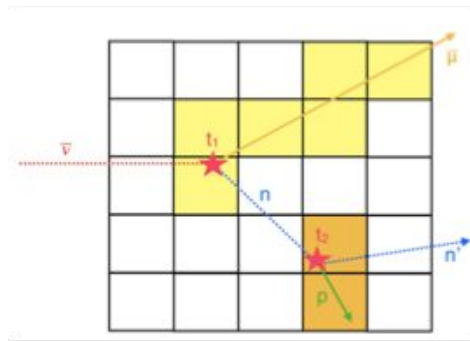


JPS Conf. Proc. 27, 011005 (2019)

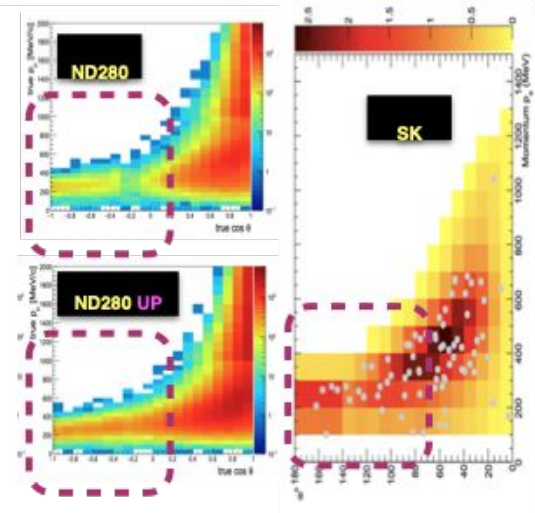
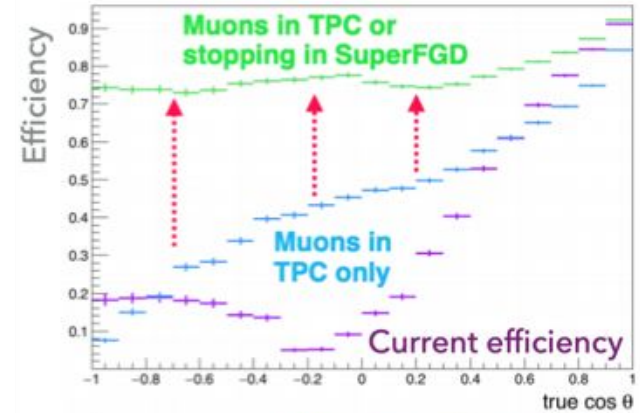
T2K ND280 Upgrade

- Upgraded ND280 covering similar phase space coverage as Super-Kamiokande
- Significant lower energy threshold
- Neutron detection capability

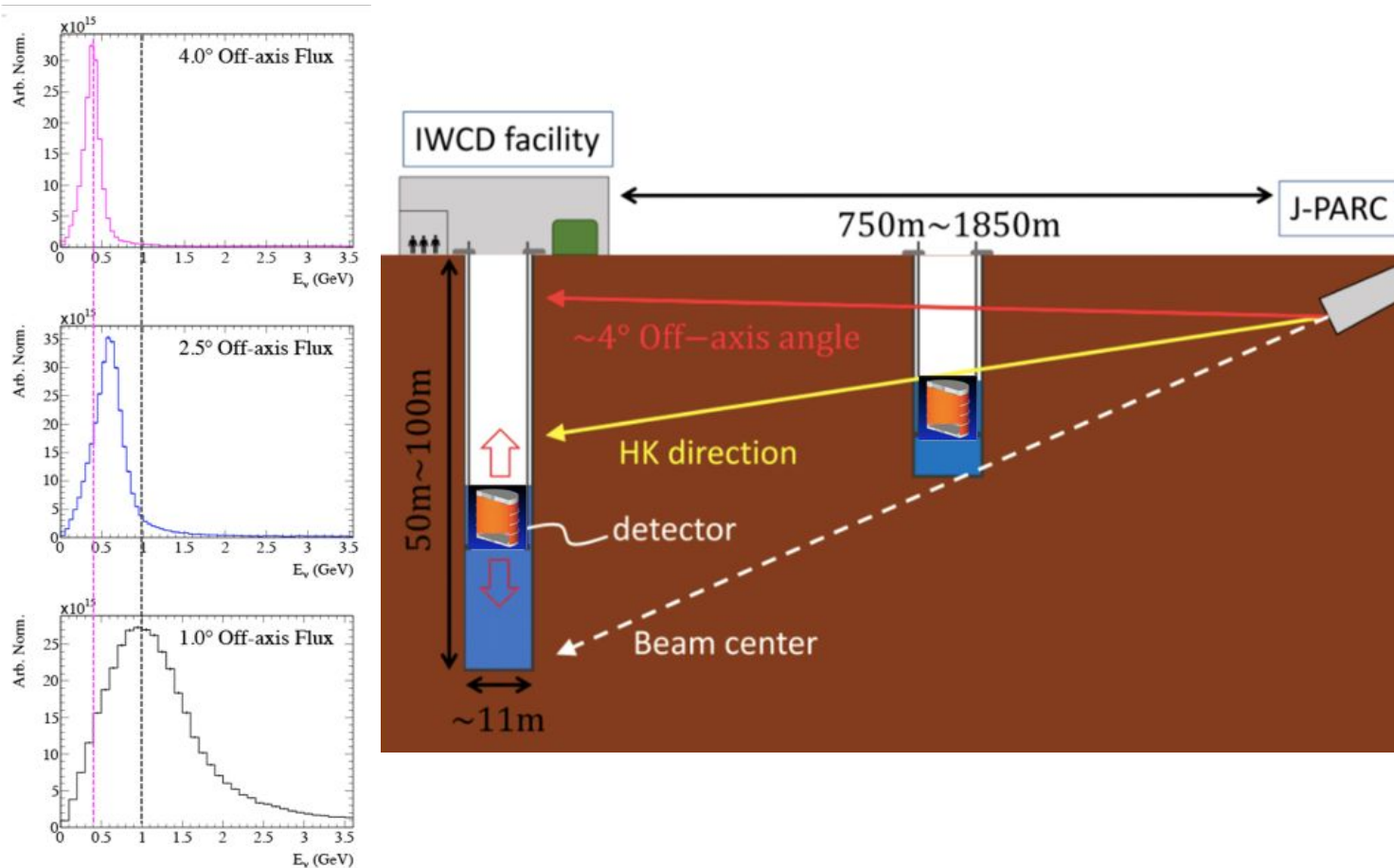
Much better constraint on beam and better cross section measurements!



Current ND280 threshold: $\sim 500\text{MeV}/c$

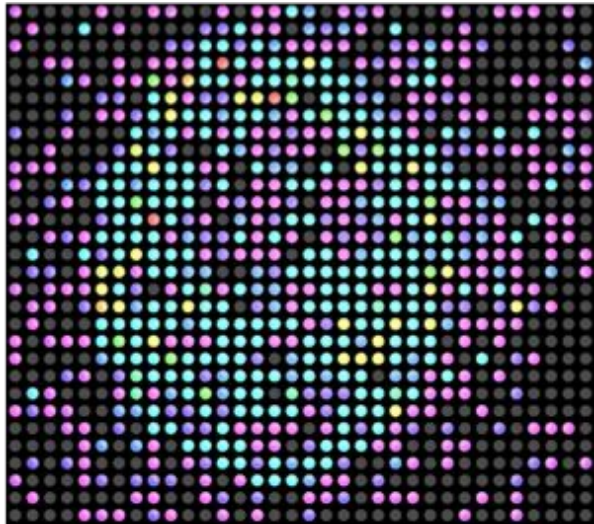


NuPRISM Off-axis Spanning Concept

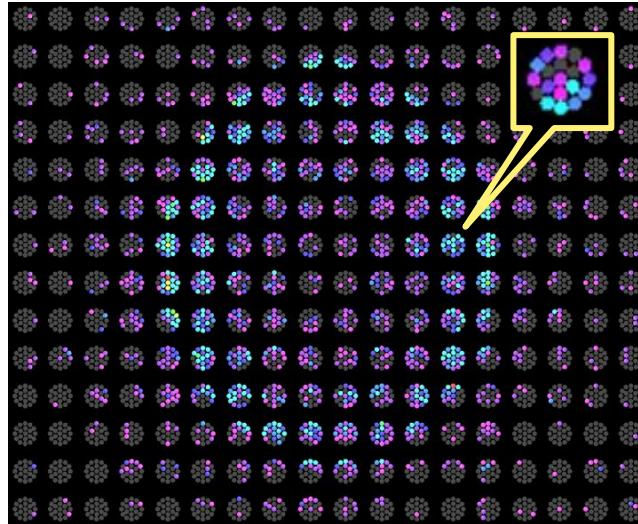


Multi-PMTs

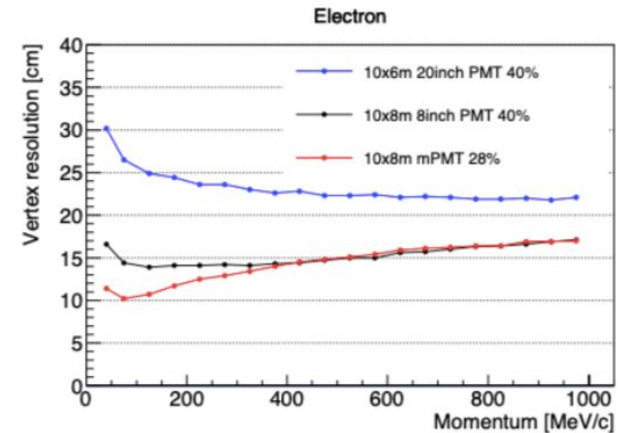
Improved spatial granularity and timing resolution (1.7 ns FWHM)



8" PMT



mPMT (3")



T2K-SK Multi-Ring Datasets for Future Analyses



ν_μ

	1.00×10^{21} POT	ν_μ CCQE	ν_μ CC1 π^+	ν_μ CC other	$\bar{\nu}_\mu$ CC	$\nu_e + \bar{\nu}_e$ CC	NC	Signal Purity
1R μ like		143.34 (Signal)	18.41	7.45	11.43	0.05	5.49	76.99%
Signal Efficiency		83.84%						
MR 1 decay e		8.93	11.15 (signal)	9.26	2.38	0.088	1.83	33.14%
Signal efficiency			52.02%					
MR 2 decay e		2.85	23.97 (signal)	11.86	0.732	0.011	1.76	58.20%
Signal efficiency			86.28%					

T2K work in progress

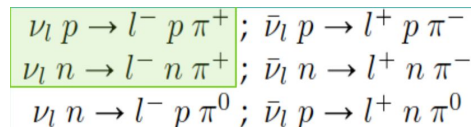
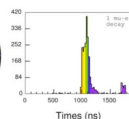
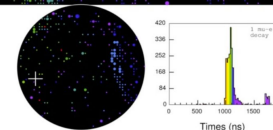
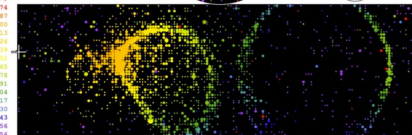
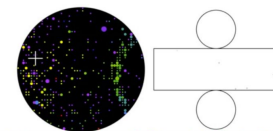
ν_e

	1.00×10^{21} POT	ν_e CCQE	ν_e CC1 π^+	ν_e CC other	$\bar{\nu}_e$ CC	$\nu_\mu + \bar{\nu}_\mu$ CC	NC	Signal purity
Existing 1 ring elike 0 decay sample		39.892 (signal)	3.347	0.884	0.482	0.201	2.771	83.85%
Signal efficiency		76.65%						
MR e like BDT sample		0.027	1.647	0.134	0.007	0.328	0.388	65.07%
Signal efficiency			9.64%					

T2K work in progress

Super-Kamiokande IV
Run: 999999 Sub: 0 Event: 897
In: 00:14:01.11100
Zener: 279 kVdc, 5176 pm
Outer: 0 Vdc, 0 pm
Trigger: 0007
E_{MC}: 111.433-0 cm
E_{MC}: 512.4-000
Z: 0-1104 (steps) Mass = 341.8 MW/10

Time (ns)
• 974
• 974-987
• 987-1000
• 1000-1013
• 1013-1026
• 1026-1039
• 1039-1052
• 1052-1065
• 1065-1078
• 1078-1091
• 1091-1104
• 1104-1117
• 1117-1130
• 1130-1143
• 1143-1156
• 1156

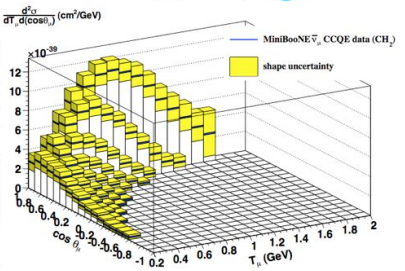
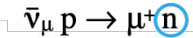
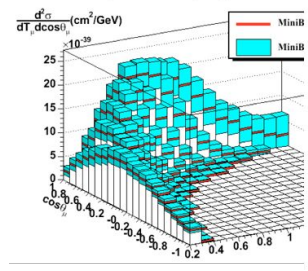
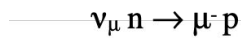
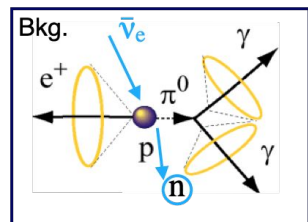
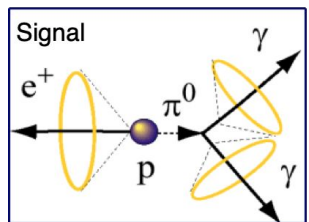
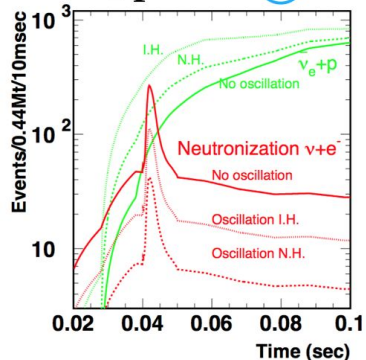
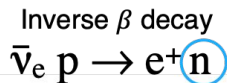
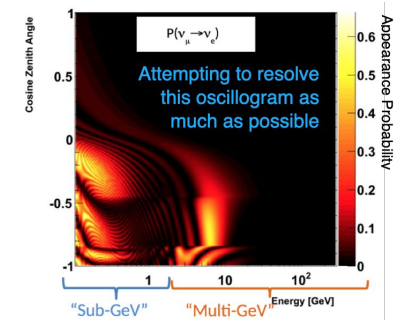


Super-K Gadolinium Upgrade Physics Potentials

- Supernova burst and relic ν
 - Separating IBD from elastic scattering for better direction resolution
 - Reduction of invisible μ decay-e bkg

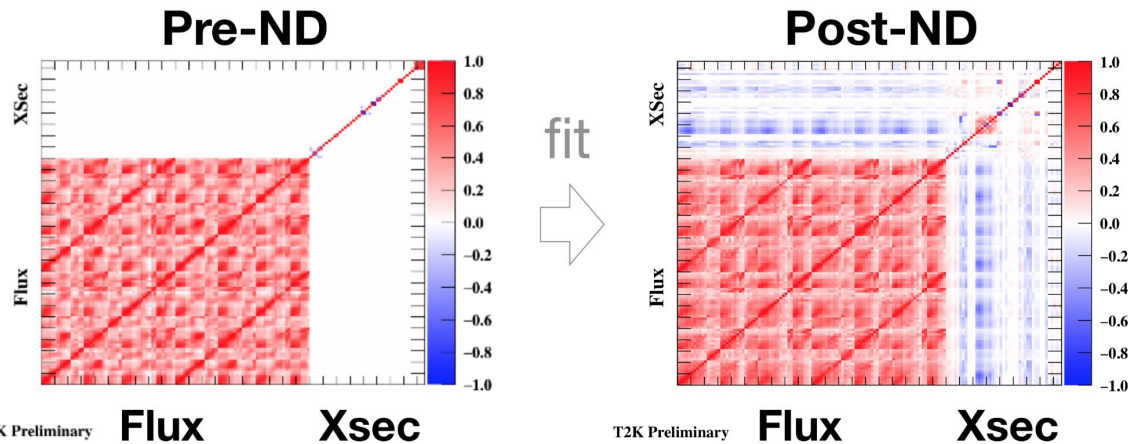
- Proton decay
 - Removal of atmospheric ν backgrounds
 - Towards background-free measurement

- Neutrino oscillation
 - Improved separation of $\nu/\bar{\nu}$
 - Reduction of beam wrong-sign bkg.
 - New probe of sub-GeV oscillogram
 - Directionality: $\bar{\nu}$ are forward Peaked
 - Full 2-body ν energy and direction reconstruction

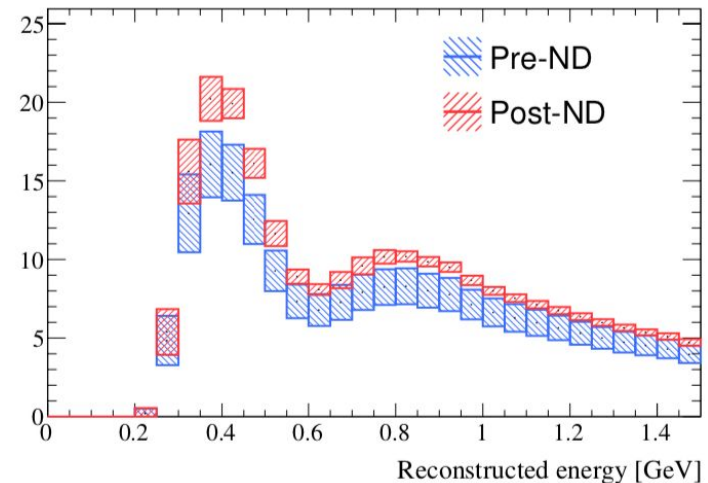


Near Detector Constraint of Systematic Uncertainties

- Neutrino flux and cross section models simultaneously constrained by near detector fit
- Priors from external hadron production and neutrino interaction data, beam monitors, and theory



Total syst uncertainty on neutrino mode $1R\mu$ events



Systematic Uncertainties

- Extrapolation of constraint from near detector isn't perfect - neutrino spectrum is different because no oscillation
- Additional errors from modeling non-quasi-elastic scattering (pion production, multi-nucleon knockout)
- Electron (anti)neutrino cross section is not constrained at near detector with 99% muon (anti)neutrino beam
- Neutral current backgrounds can fake electron (anti)neutrino candidates

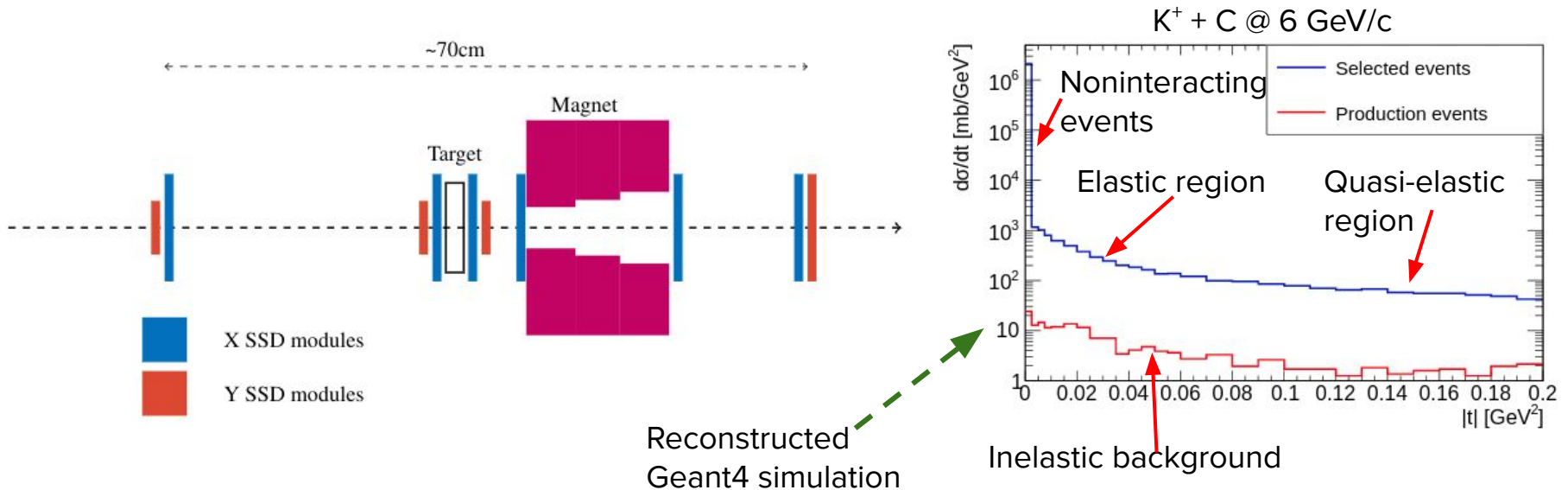
Systematic Error Source	Uncertainty on $\nu_e/\bar{\nu}_e$ Candidates (%)
Super-K Detector Model	1.47
Pion Reinteractions	1.58
Near Detector Constrained Parameters	2.31
Nuclear Binding Energy	3.74
$\sigma(\nu_e)/\sigma(\bar{\nu}_e)$	3.03
NC1 γ Production	1.49
Other NC Interactions	0.18
Total	5.87

Error Source	% Error for CP Violation search
Error from near detector constraint	1.7
Modeling of events that aren't quasi-elastic scattering	2.1
Electron (anti)neutrino cross section error	3.0
Neutral current background error	1.0
Total cross section model error	4.1

Aim to reduce total error to **<3% for Hyper-K**

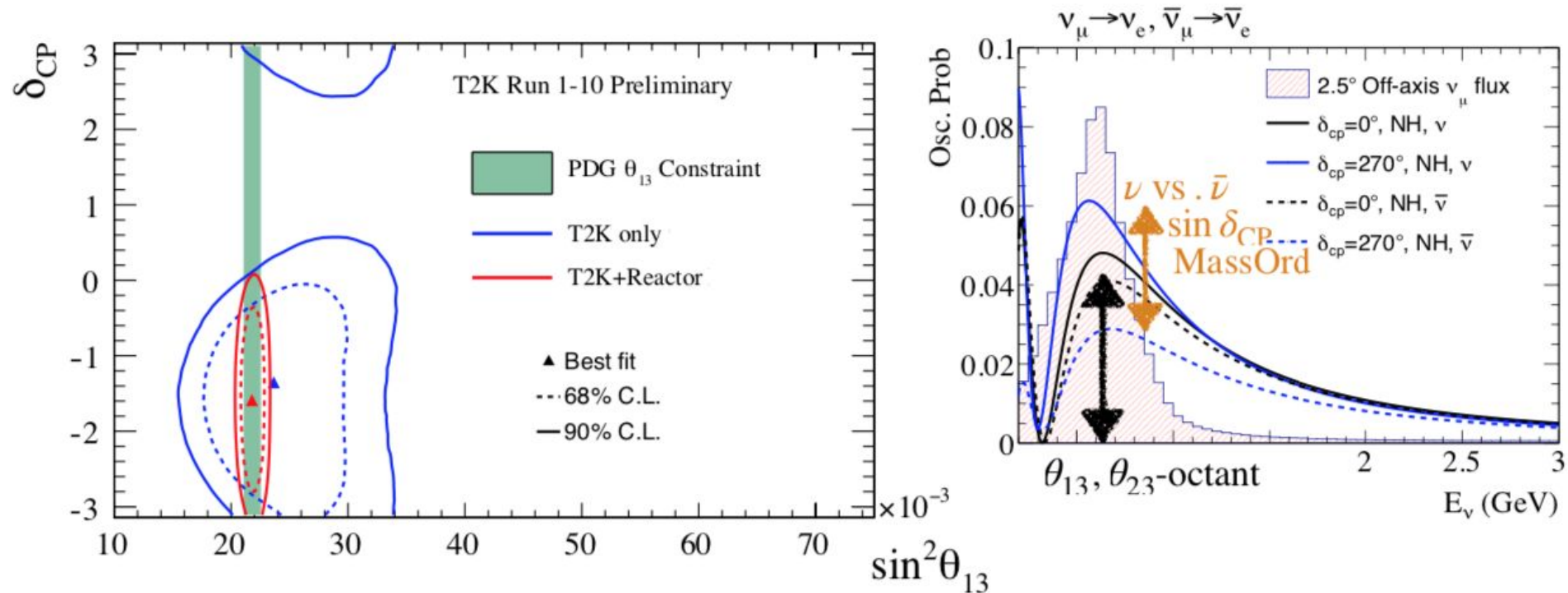
Hadron Interaction Measurements in the WCTE Facility

- Pion and kaon elastic and quasi-elastic interactions (< 10 GeV/c) are important source of systematic uncertainty in the HyperK neutrino flux prediction → **no existing data**
- Reconfiguring WCTE tertiary beam spectrometer for hadron interaction measurements → similar to EMPHATIC, but kaon beam < 10 GeV/c not possible at FTBF
- Water tank can be used to identify pions and muons in the secondary beam
- Expression of interest: <http://cds.cern.ch/record/2771386>



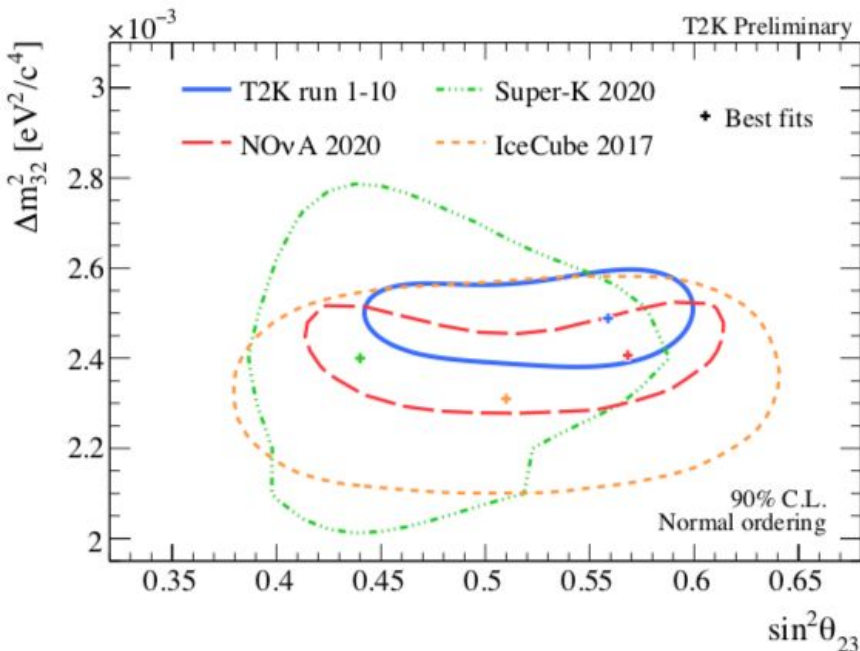
Constraints on θ_{13}

T2K's θ_{13} constraint ($\nu_{\mu} \rightarrow \nu_e$ appearance) consistent with the much stronger constraint from reactor experiments ($\bar{\nu}_e \rightarrow \bar{\nu}_e$ disappearance)

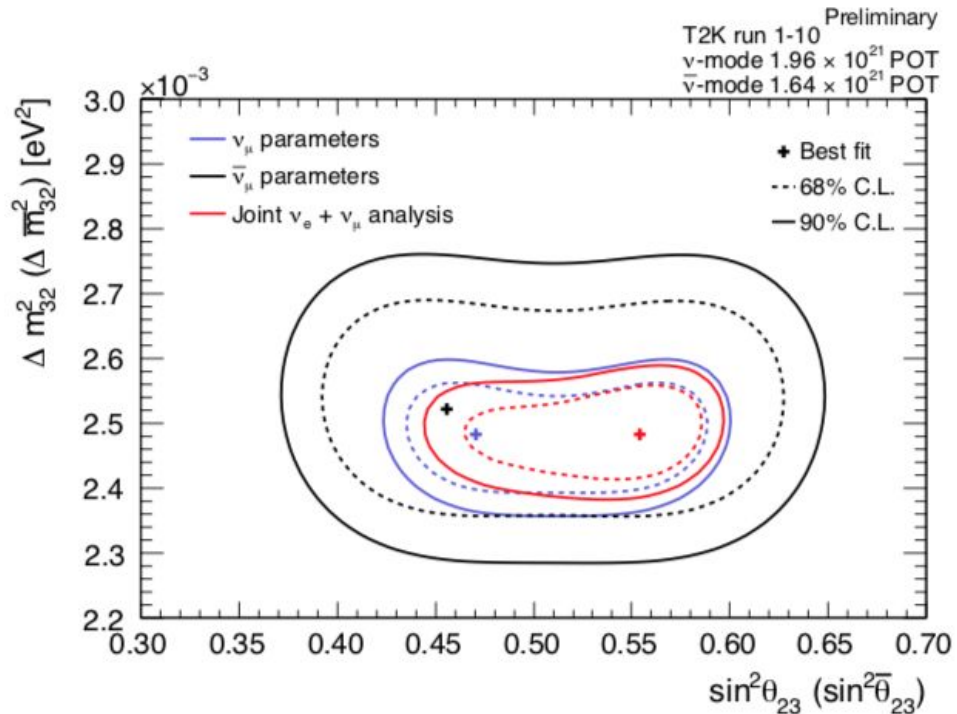


Atmospheric Mixing Parameters

- World-leading measurement of atmospheric parameters
 - Slight preference for $\sin^2\theta_{23}$ in upper octant

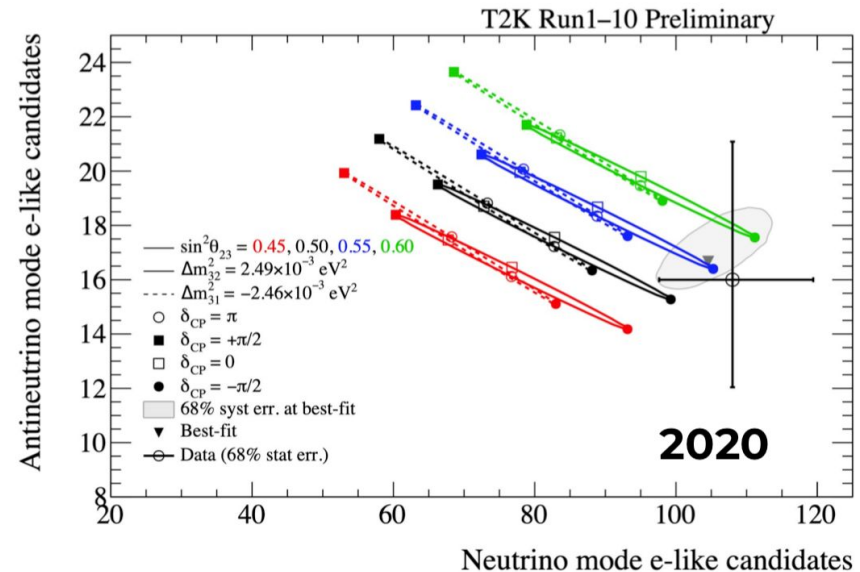


- Fit with separate parameters for $\nu_\mu/\bar{\nu}_\mu$ shows consistent results
 - Expected from CPT symmetry



ν_e vs $\bar{\nu}_e$ Appearance

- Bi-event plot illustrates origin of data constraints
- Best-fit δ_{CP} around maximal CP-violation $-\pi/2$
- Weak preference for Normal ordering with Bayes factor 4.2
- Weak preference for upper octant with Bayes factor 3.4



Mass ordering	Octant		Sum
	$\sin^2 \theta_{23} < 0.5$	$\sin^2 \theta_{23} > 0.5$	
NO ($\Delta m_{32}^2 > 0$)	0.195	0.613	0.808
IO ($\Delta m_{32}^2 < 0$)	0.034	0.158	0.192
Sum	0.229	0.771	1.000



NOvA + T2K

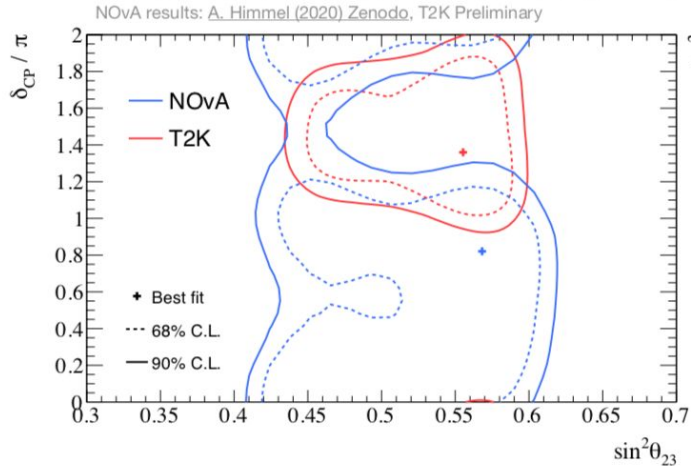
Ongoing joint fits

SK + T2K

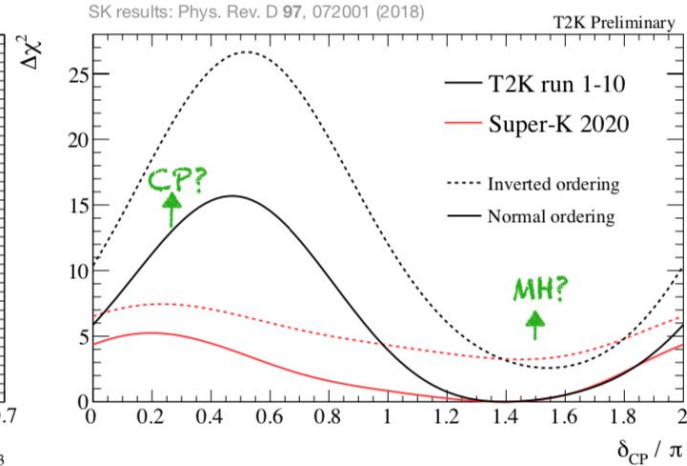
atmospheric + accelerator



Comparison of released contours (not joint fit)



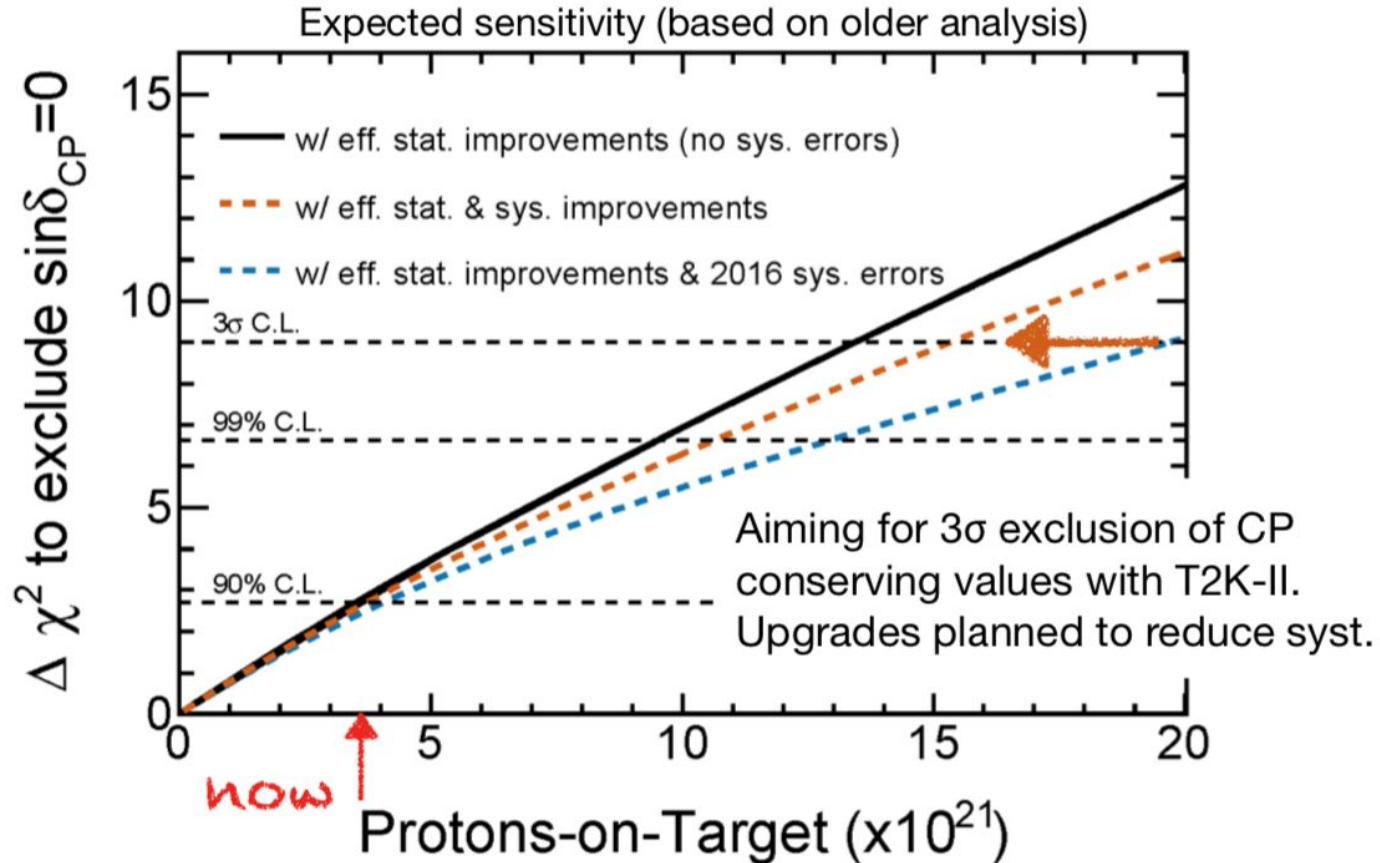
Comparison of released contours (not joint fit)



- Joint fits between experiments with different oscillation **baselines/energies** and **detector technologies**
- expect increased **sensitivity** in δ_{CP} , mass ordering, θ_{23} octant **beyond stats increase** from resolved degeneracies and syst constraints
- important to understand potentially non-trivial **syst. correlations** between experiments

Agreements are signed between experiments and joint work ongoing. Stay tuned!

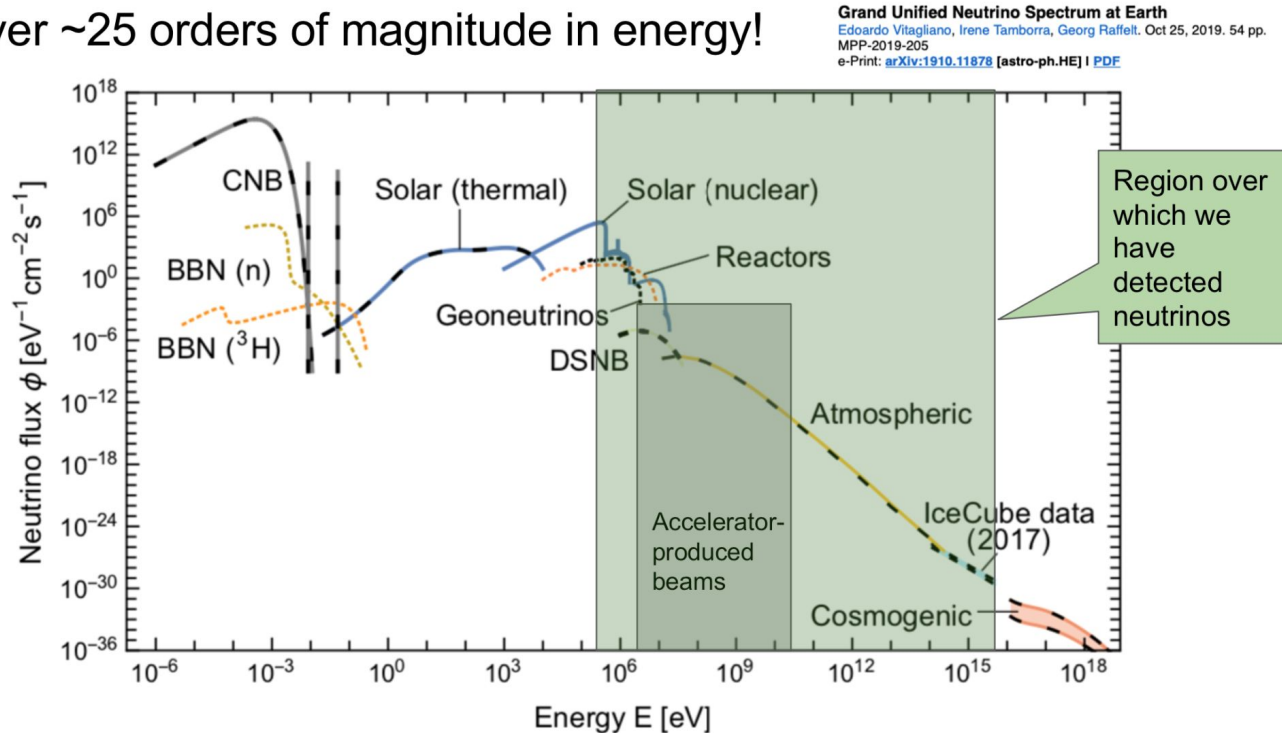
T2K Toward 3σ Evidence of CP Violation



Neutrino Flux Spectra

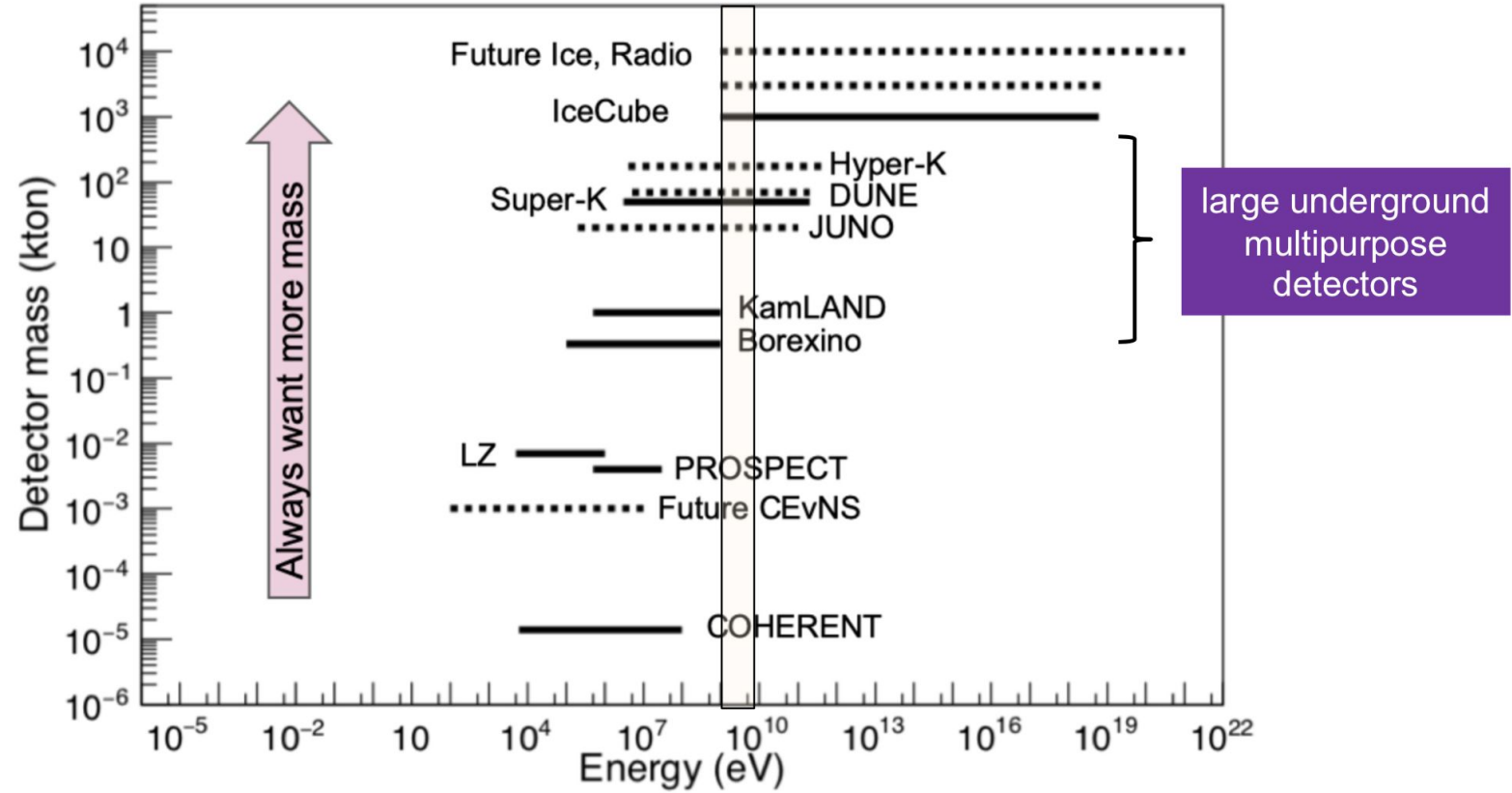
[Kate Scholberg \(Duke\), TIPP 2021](#)

Information comes from neutrinos
over ~ 25 orders of magnitude in energy!

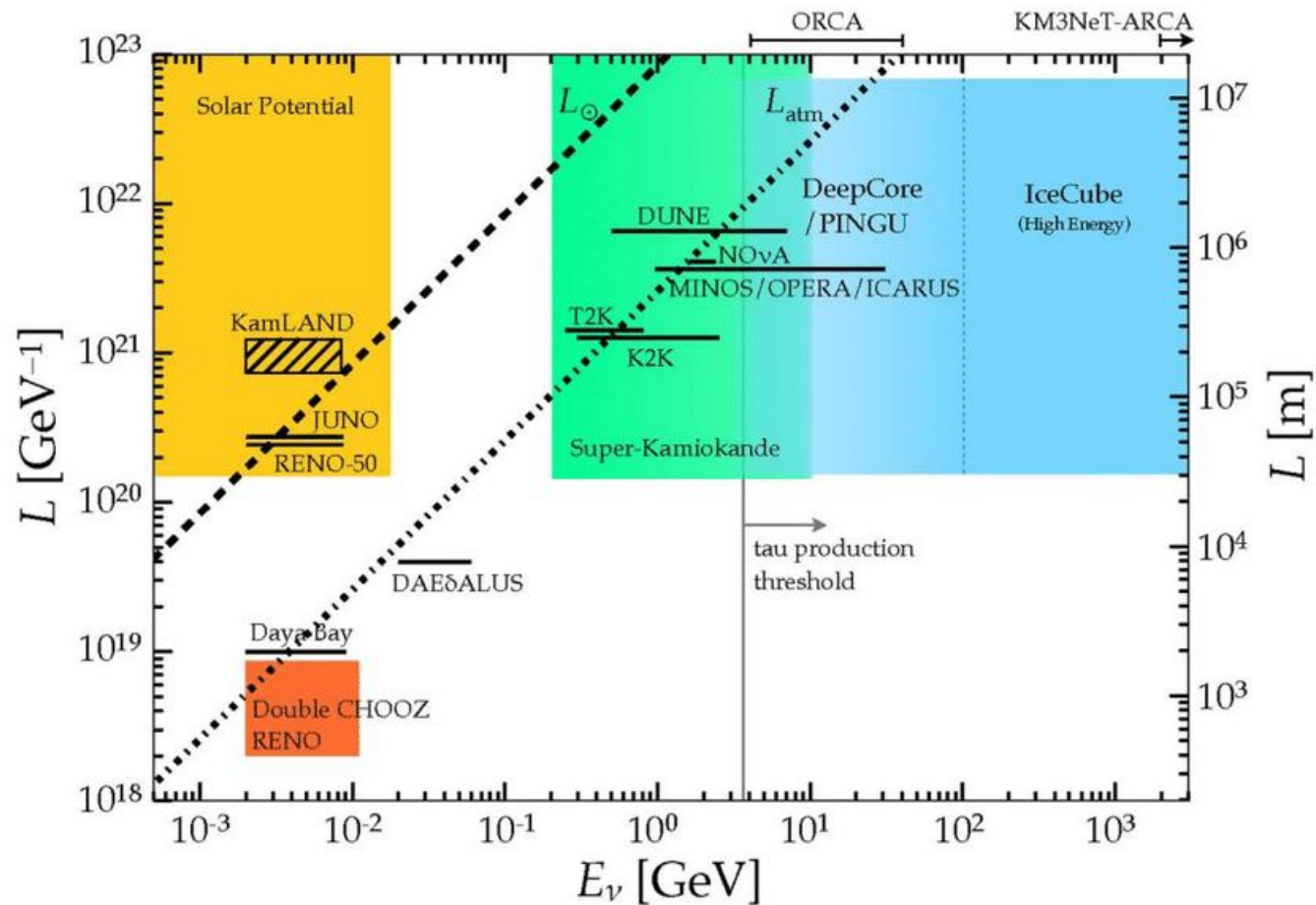


Neutrino detector masses and sensitive energy ranges

[Kate Scholberg \(Duke\), TIPP 2021](#)



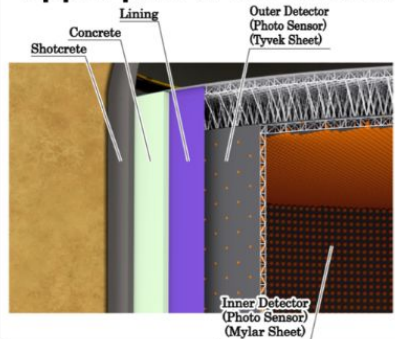
Neutrino Oscillation L/E Scales



Hyper-K Far Detector Concept

Enlarged view

Upper part of the detector



Lower part of the detector

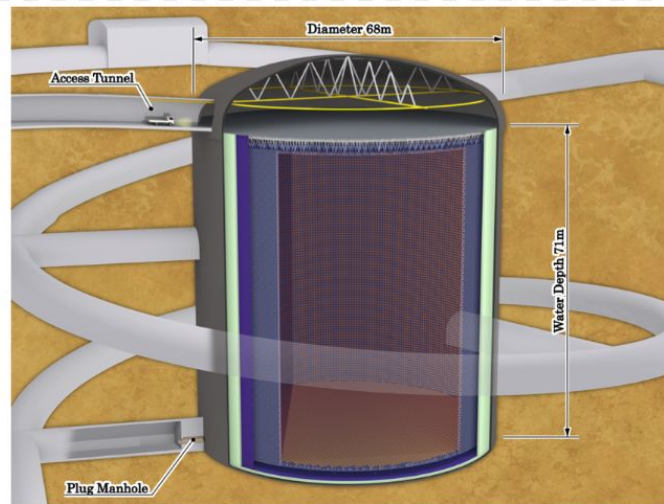
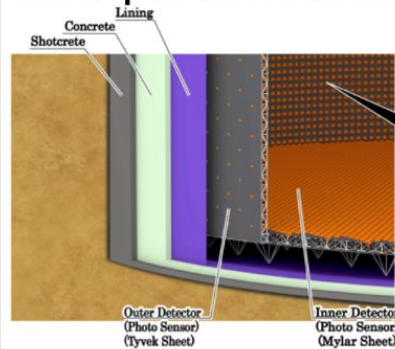
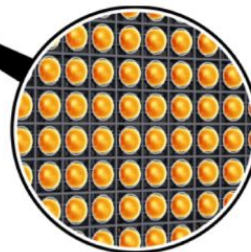
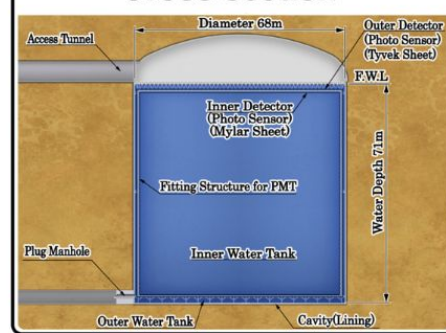


Photo-sensors



Cross section



Hyper-K PMT Progress



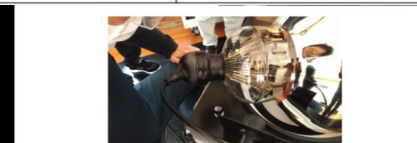
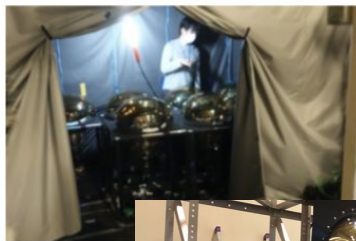
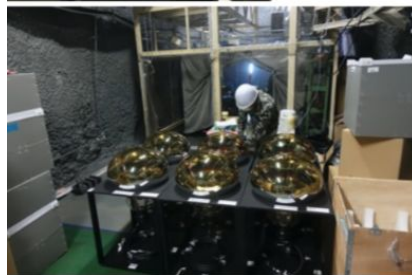
HK PMTs were checked according to the pre-calibration manual to install HK PMTs in 2018 for SK.

- Initial delivery of 20,000, 50cm PMTs in Dec. 2020; expect ~300 PMTs per month
- Inspection and long-term calibrations starting in Kamioka

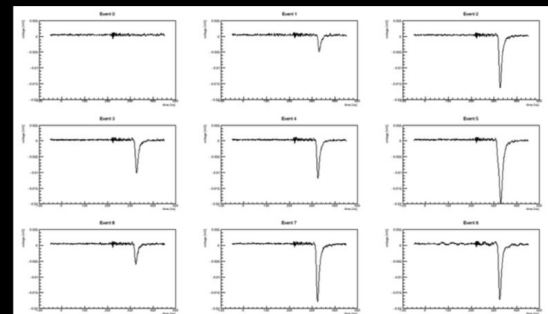
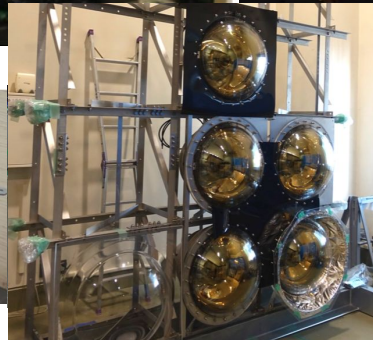
PMTs for the Inner Detector

	Super-K	Hyper-K
Number of PMTs	11,129 50cm PMTs	20,000 50cm PMTs (JPN) (+ additional PDs (Oversea))
Photo-sensitive Coverage	40 %	20 %
Single photon efficiency /PMT	~12%	~24%
Dark Rate /PMT	~4 kHz (Typical)	4 kHz (Average)
Timing resolution of 1 photon	~3 nsec	~1.5 nsec

Visual inspection

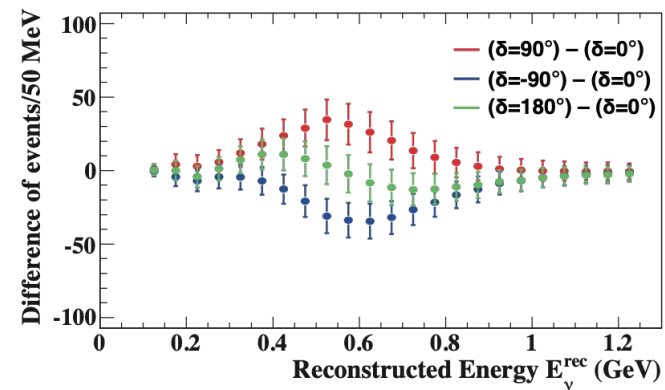
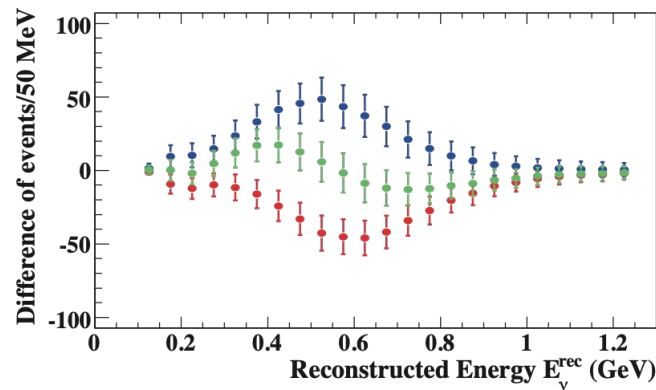
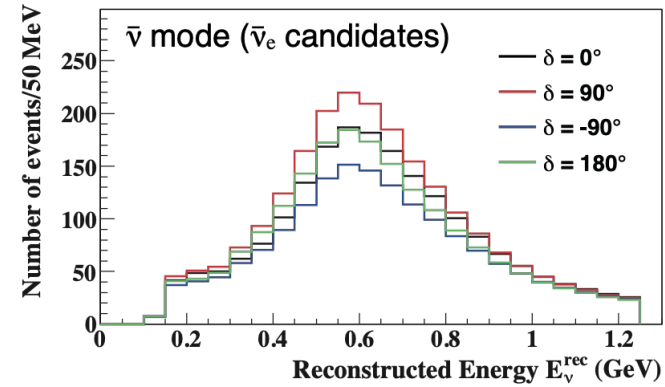
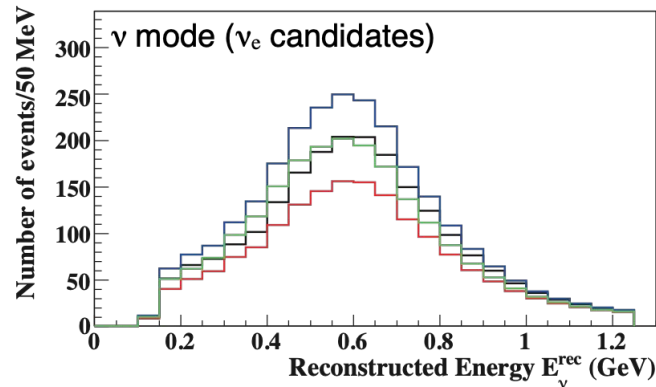


- Cover production by Spain ~2022

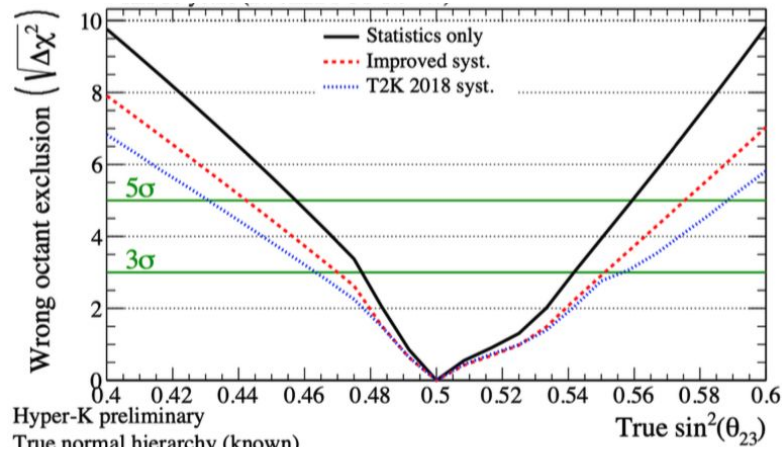
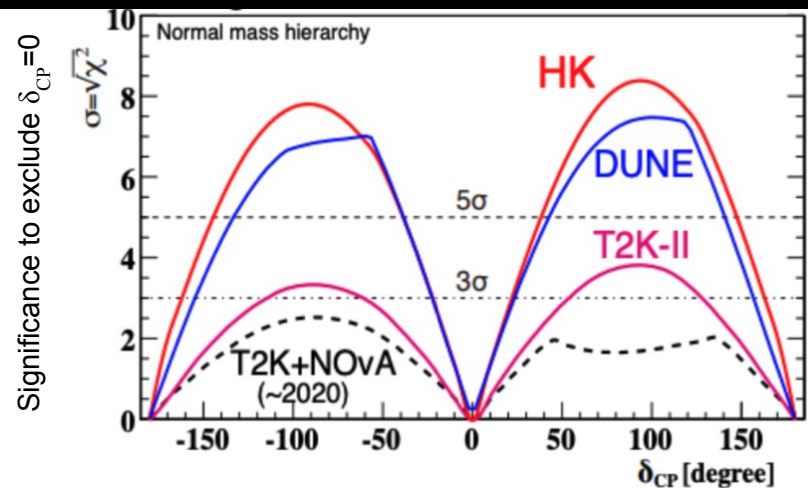
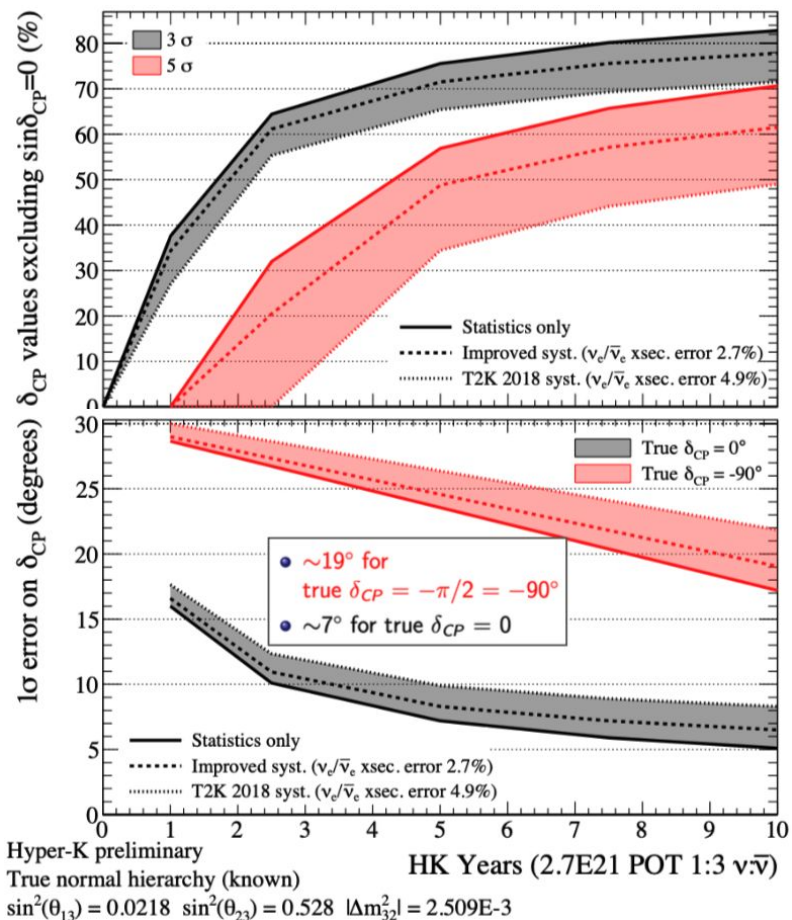


Hyper-K Expected Event Rates

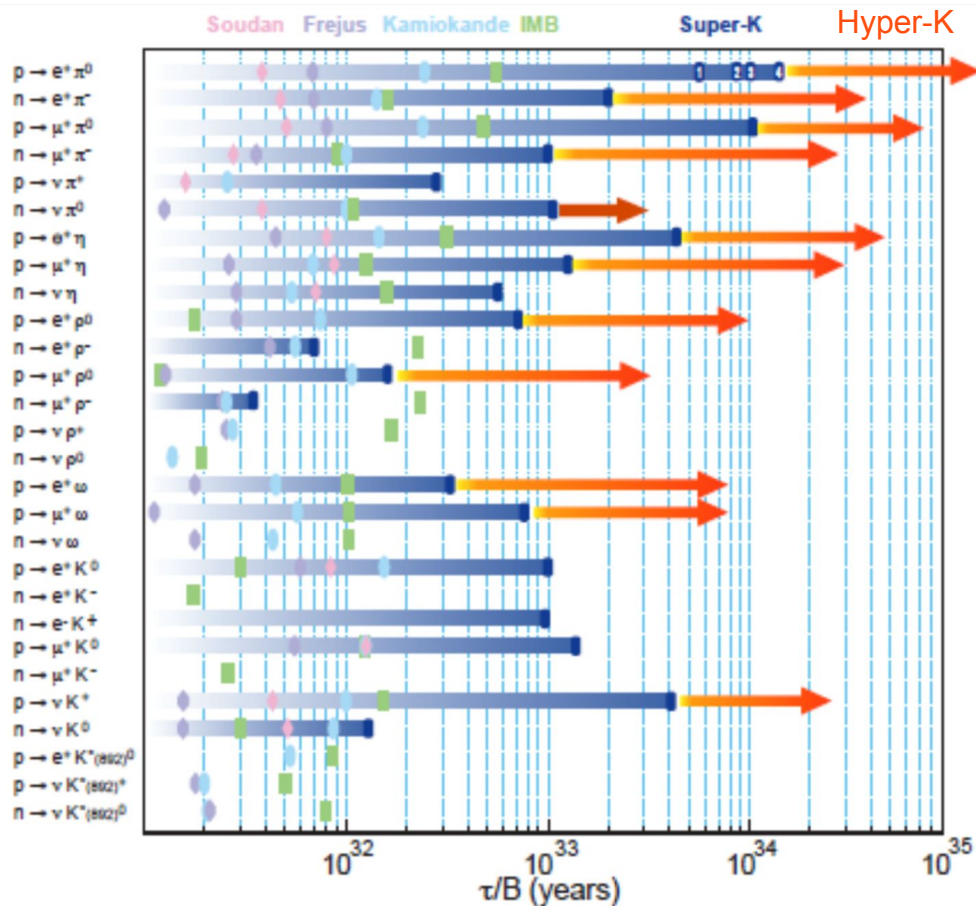
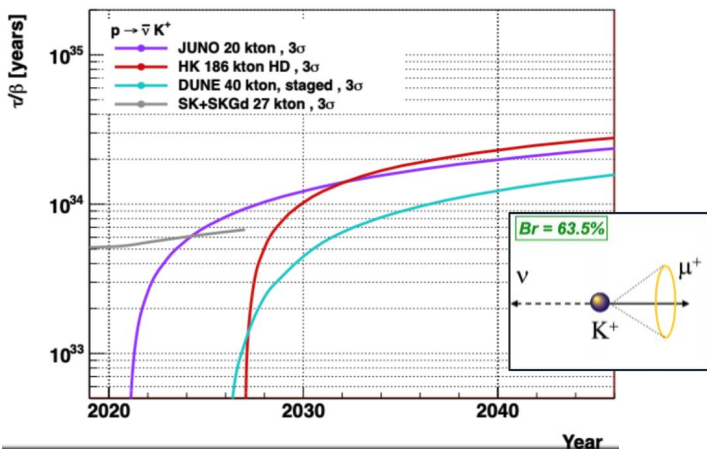
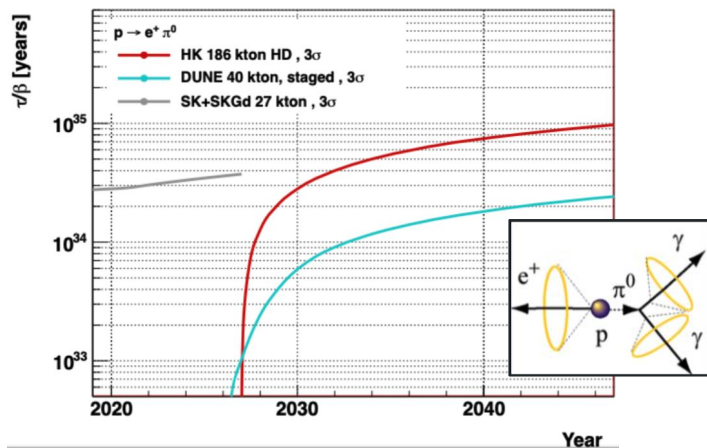
- Aim to collect ~ 2000 ν_e and $\bar{\nu}_e$ appearance events in 10 years
 - Will measure CPV with 3% statistical uncertainty!
- Controlling systematics becomes critical!



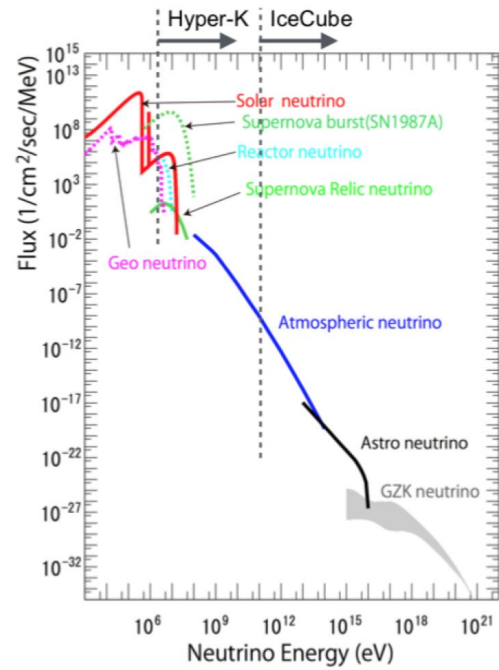
Hyper-K Long-Baseline Physics



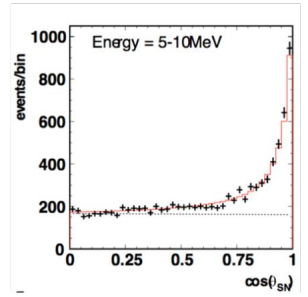
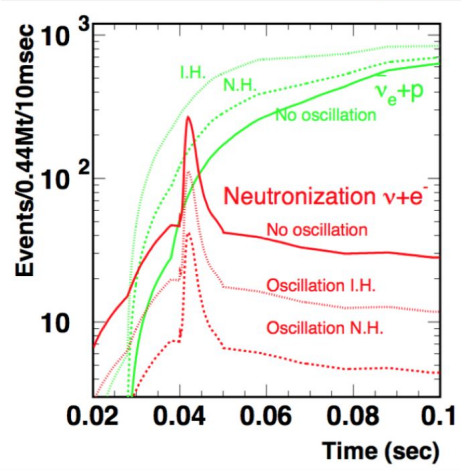
Hyper-K Proton Decay



Supernova Burst in Hyper-K

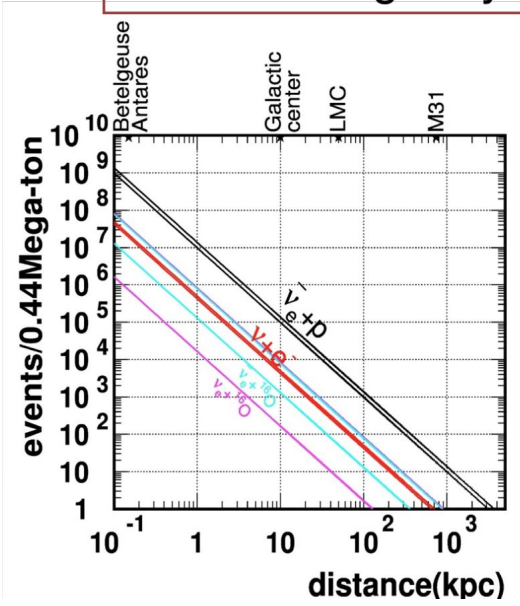


Neutrino carries information of explosion



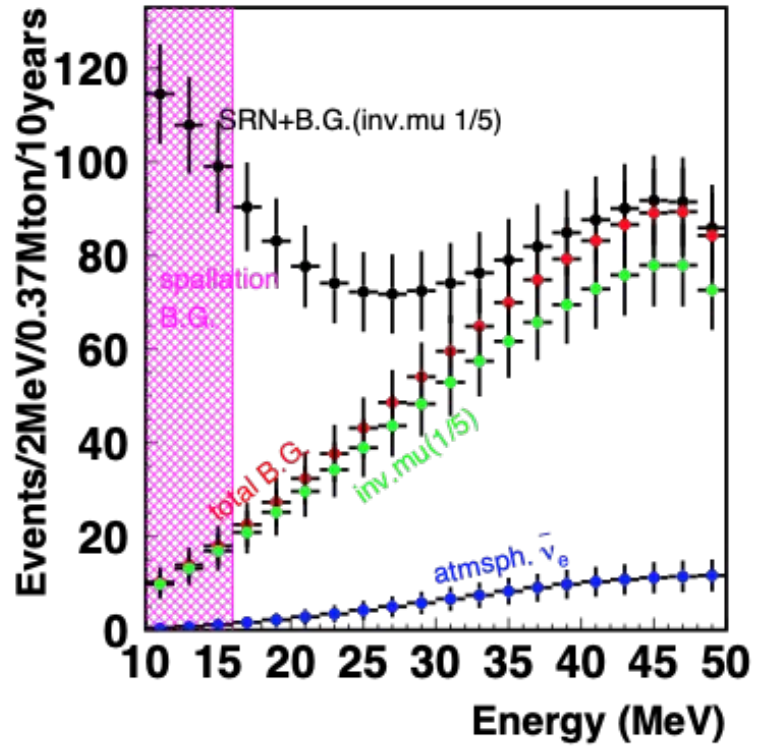
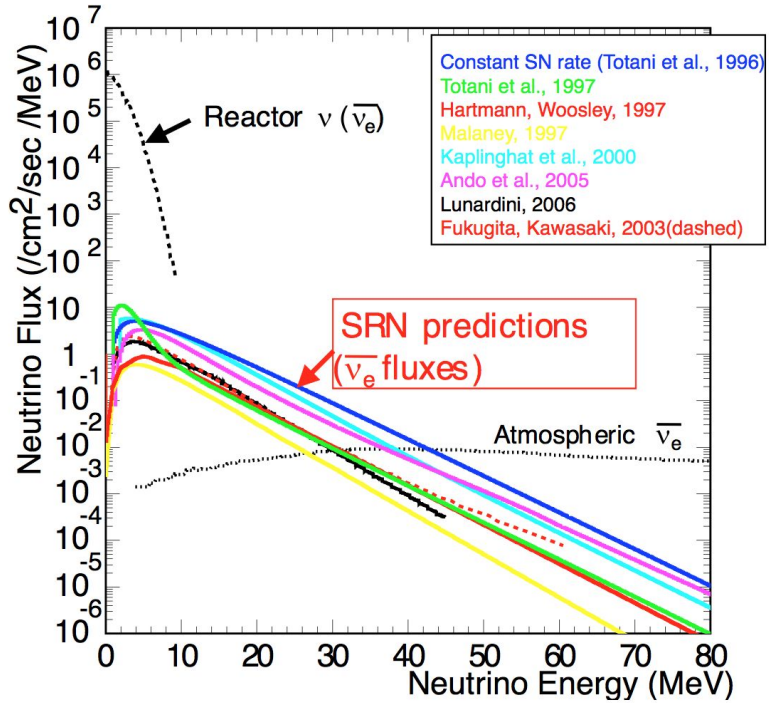
Direction alert for optical measurements (ν_e elastic)

Reaching Andromeda galaxy

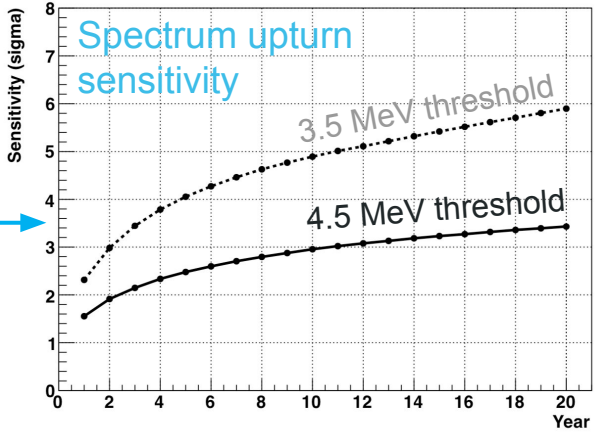
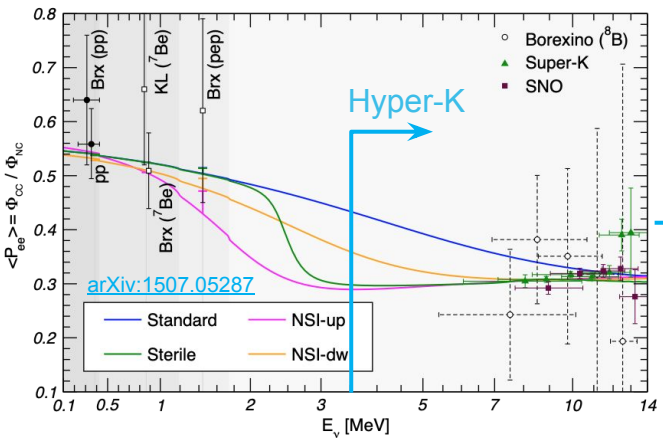
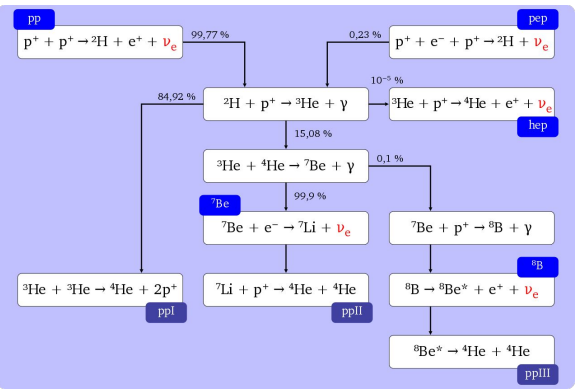
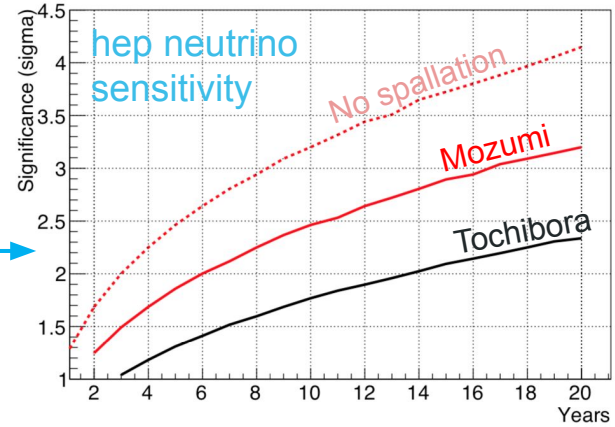
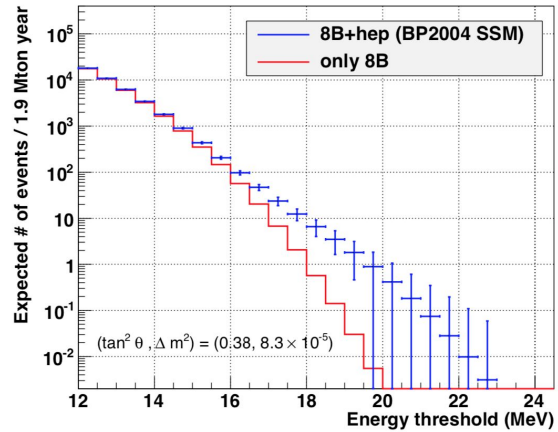
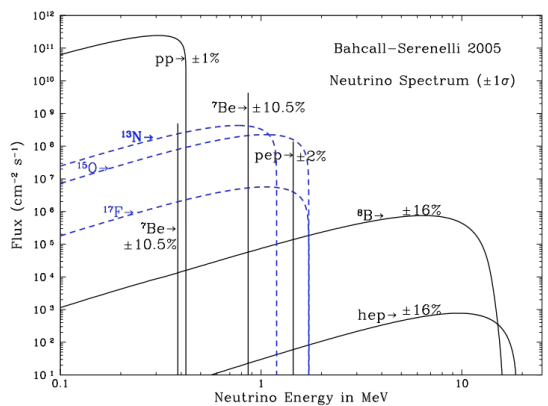


Hyper-K Supernova Relic Neutrinos

SRN can be observed by HK in 10y with $\sim 70 \pm 17$ events. It is $> 4\sigma$ for SRN signal.



Hyper-K Solar Neutrinos



Hyper-K Collaboration Membership

- 19 countries, 93 institutes, ~450 people as of May 2021, growing

