

Status of SNO+

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On Behalf of the SNO+ Collaboration

Laurentian University

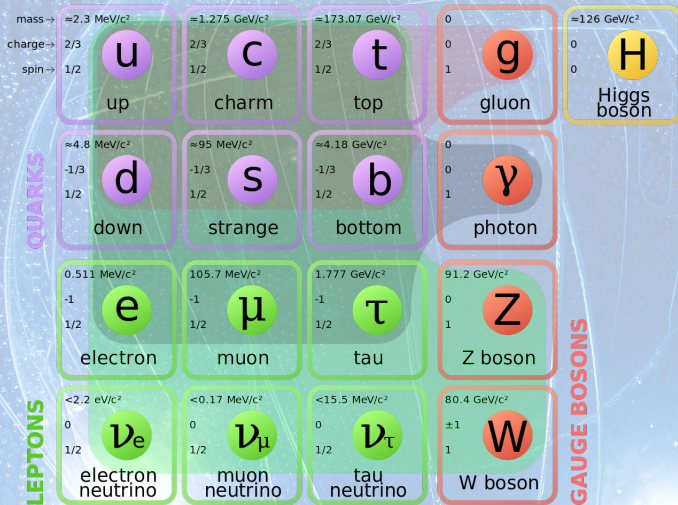
2021 CAP Congress
June 8, 2021



Laurentian University
Université Laurentienne



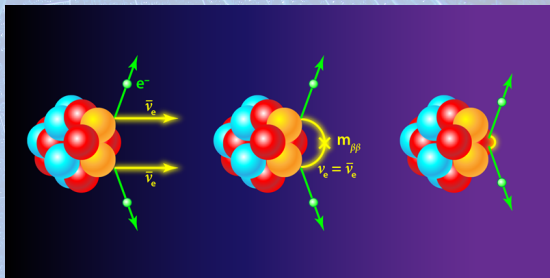
Neutrinos in the Standard Model



- Neutrinos explicitly assumed to be mass-less
- Measurements from SNO and Kamiokande show they are not

Neutrino Mass and Neutrino-less Double Beta Decay

- One path to ν mass is a Majorana mass term
- Allows for a violation of lepton number

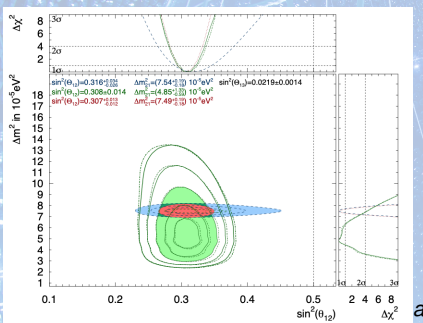


^aJ. Engel, P. Vogel, Physics 11, 30 (2018)

- 2 neutrino β decays are preferred in some isotopes
- If lepton number violation is allowed then $0\nu\beta\beta$ decays allowed
- Rate of the process is proportional to the neutrino mass
- The chance of discovery scales with $\beta\beta$ isotope mass in an active volume

Solar Oscillation Tension

- Neutrinos oscillations appear in solar and reactor experiments



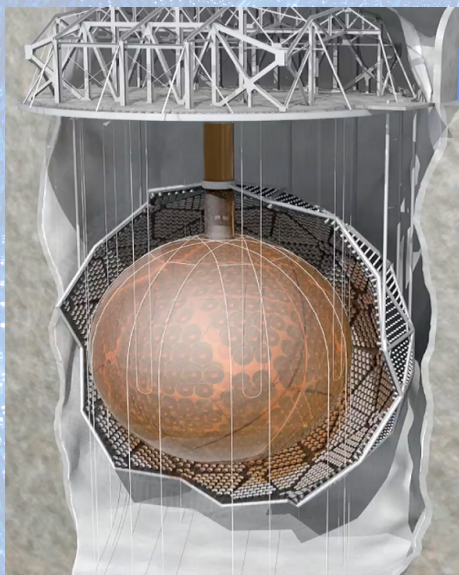
$$P_{ee} = 1 - \sin^2 2\theta_{12} \sin^2 \frac{1.267 \Delta m^2 L}{E}$$

- Solar neutrino oscillations measured in SNO and Super Kamiokande
 - $\Delta m^2 = (4.8^{+1.3}_{-0.6}) \times 10^{-5} \text{eV}^2$
- Adding Reactor measurements from KAMLAND
 - $\Delta m^2 = (7.53 \pm 0.18) \times 10^{-5} \text{eV}^2$

^aPhys. Rev. D. 94, 052010 (2016)

- Simultaneous solar and reactor neutrino experiment can address this tension

The SNO+ Detector

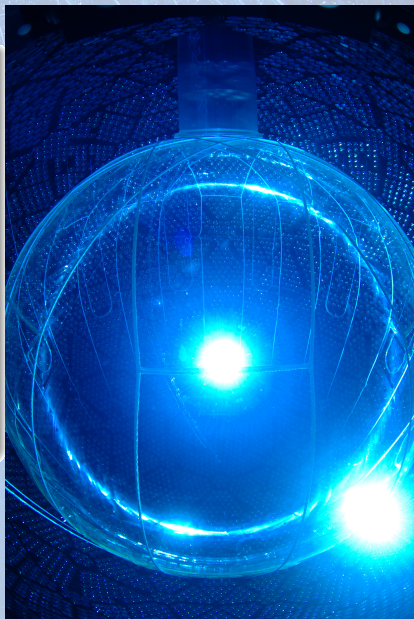


- Located at SNOLAB 2 km (≈ 6000 m.w.e.) underground
- Urylon lined cavern 30 m height, 24 m radius
 - filled with water shielding
- Consists of 12 m diameter acrylic vessel
 - Filled with 780 T linear alkylbenzene
 - Doped with PPO wavelength shifter
- Surrounded by ≈ 9400 PMTs
 - Supported by ≈ 17 m geodesic structure

^aarXiv:2104.11687

LAB Filling History

- **Nov 2018 - June 2019**
Commissioning/Early fill
- June 2019 - Nov. 2019 Pause to check backgrounds (43.4 T)
- Dec 2019 - March 2020 Fill Resumed
- March 2020 - November 2020 Pause for Covid (364 T)
- Dec 2020 - March 2021 Fill Resumed
- April 2021 Last water removed, AV prepared for recirculation
- Now undergoing a program of recirculation and PPO addition
- Scintillator Paper —
JINST16 P05009 (2021)



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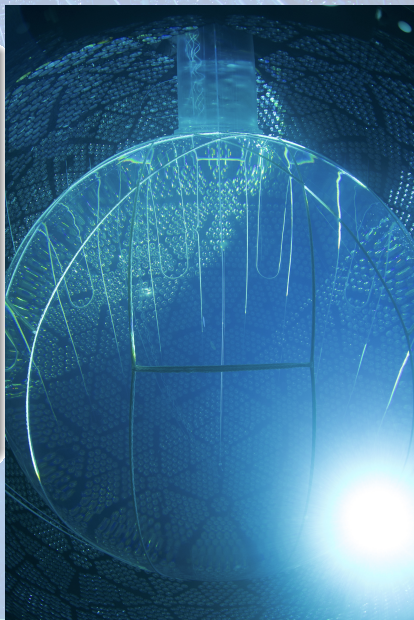
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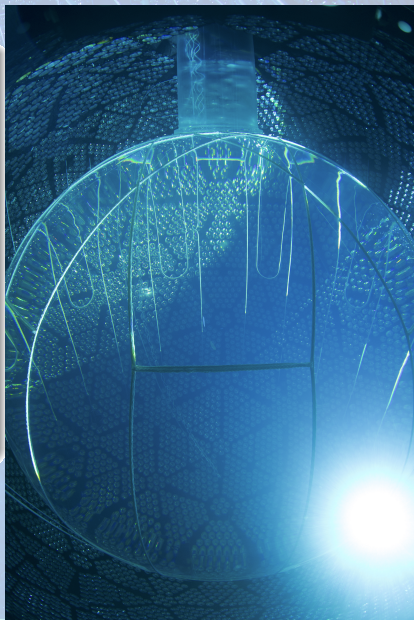
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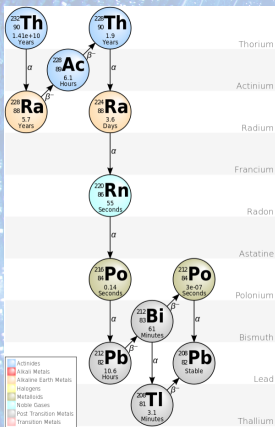


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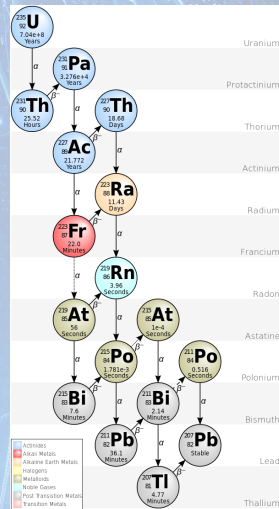
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Sources of Backgrounds

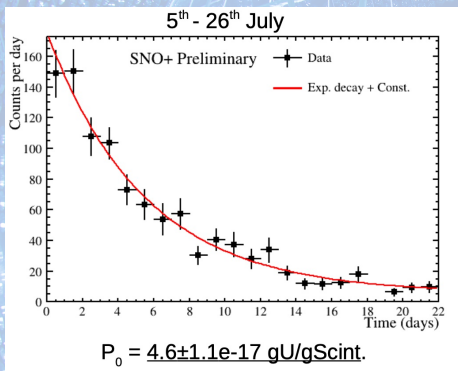


- Backgrounds from natural U and Th decay chains
- Monitor Radon ingress
 - Reduced by maintaining N_2 cover over detector volume
 - Rn content of cover gas monitored at all times
- Presence of these isotopes limit measurement sensitivities



Measured Background Rates

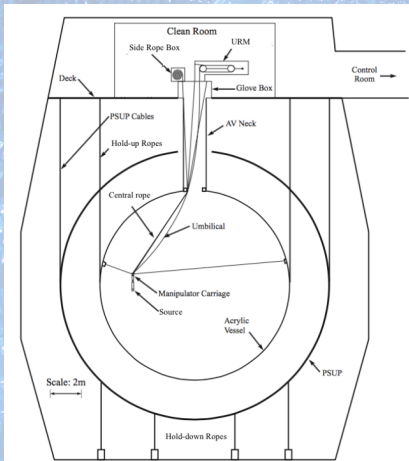
- 2020 data are being analyzed for background rates



- Focus on ^{214}Bi - ^{214}Po rates
- Rates fluctuated during filling process
- Decreased through steady state period

- Analysis continues to track and mitigate new sources of radon
- See posters by [Syed M Adil Hussain \(J85\)](#) and [Parmesh Ravi \(J86\)](#)

SNO+ Calibration Systems



- Umbilical/central rope stored in URM
- Four rope boxes control side ropes
- URM and rope boxes connected to AV through Universal Interface (UI)

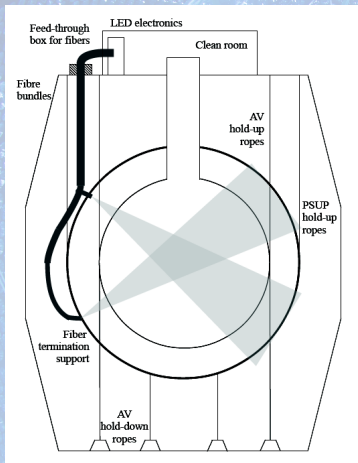
- Sources deployed via system of ropes
- Source steered in two dimensions
- Services carried via polyurethane (tygothane) "umbilical"
- Complemented with fixed fibre light injection system

External Deployment

- Guide tubes pass outside AV
- Complementary source deployment



SNO+ Calibration Systems

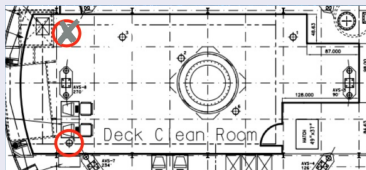


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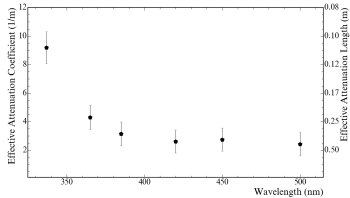
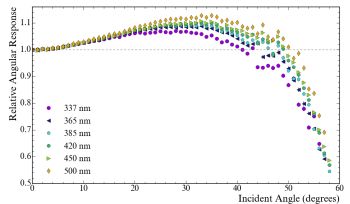
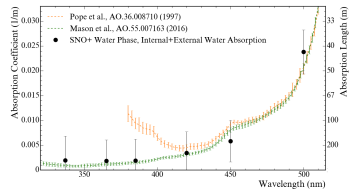
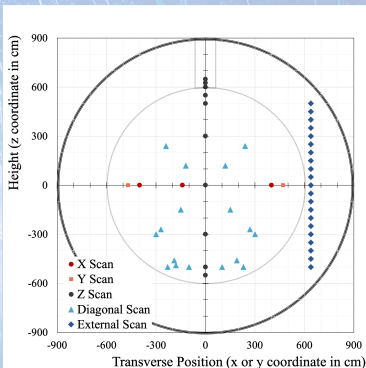
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Detector Calibrations in Water

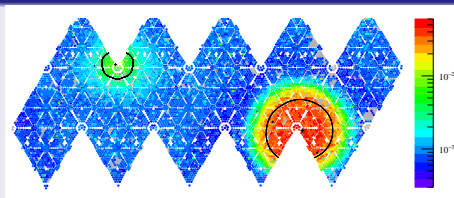
- Extensive campaigns conducted in water phase
- Determined the response of PMTs and optical properties of detector
- optical and radioactive sources
- deployed inside AV and outside AV
- Detailed paper in preparation



Calibrations in Scintillator

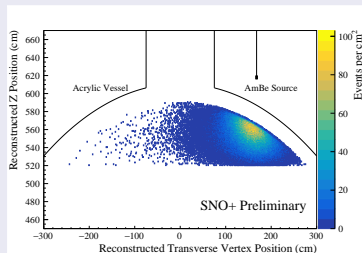
Fixed Optical Fibre System

- Fibres embedded in PSUP
- Shine LED or laser light across AV
- Measures PMT timing and light scattering, absorption



External Source Deployments

- Deployed γ and N sources in Cavity
- Translate the results from the edge to the centre of AV
- Compare neutron capture in water^a to scintillator
- See talk by [Jamie Grove](#) (R2-9)



^aPhys. Rev. C 102, 014002 (2020)

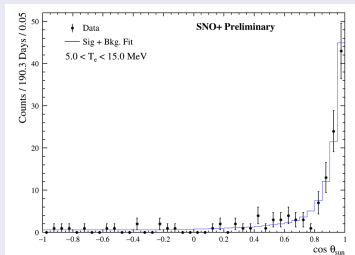
Highlights of Water Phase

Solar

- Verification of solar neutrino flux conducted with water filled detector

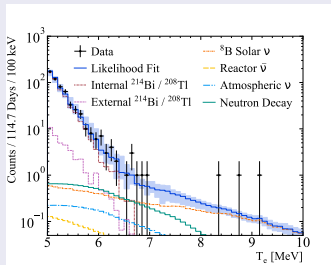
- Published ^a

$$\phi_{ES}({}^8B) = 2.53^{+0.31}_{-0.28} {}^{+0.13}_{-0.10}$$



^aPhys. Rev. D 99 012012 (2018).

Nucleon Decay Measurements



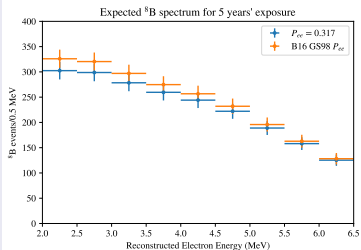
- Sought excitations from p, n decays
- Set 90% C.L. limits^a
 - $\tau(p \rightarrow \text{invis}) > 3 \times 10^{29}$ y
 - $\tau(n \rightarrow \text{invis}) > 2.5 \times 10^{29}$ y

^aPhys. Rev. D 99, 032008 (2019)

- Improved measurements with additional existing data in progress

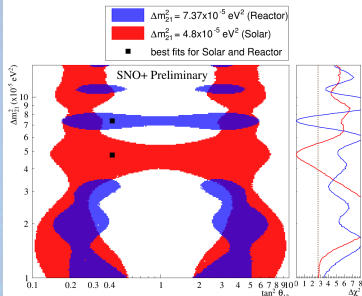
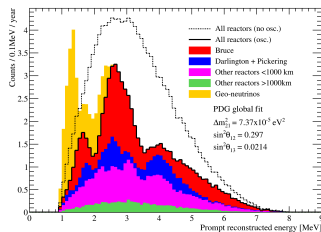
Solar and Reactor Predictions in SNO+ Scintillator

Solar Rates



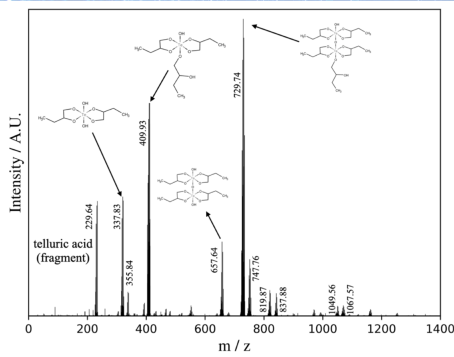
- Prediction of rates from solar and reactor events ongoing
- An analysis of reactor events from last stable period is in progress
- Can project limits for 12 months of data

Reactor Rates



Adding Double Beta Decays to SNO+ (^{130}Te)

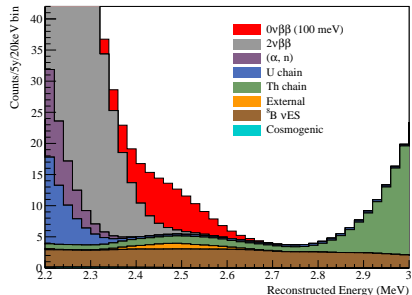
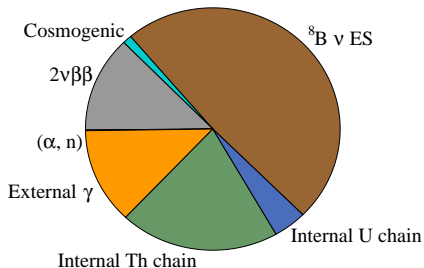
- ^{130}Te occurs with 34% natural abundance
 - No enrichment required
- Chemistry for suspending ^{130}Te in LAB has been developed



- Plants for processing and adding ^{130}Te are being commissioned underground

Projected $0\nu\beta\beta$ Measurements

ROI: 2.42 - 2.56 MeV $[-0.5\sigma - 1.5\sigma]$
Counts/Year: 9.47



- Example simulations used a cocktail LAB, PPO, 0.5%Te, Diol, DDA

- Projected limits
 - half life $T_{1/2} > 2.1 \times 10^{26}$ years
 - $m_{\beta\beta}$ range for 5 years: 37 - 89 meV
- Larger Te loading proposed for later stages in the experiment

Summary

- SNO+ has completed the fill of the detector acrylic vessel
 - Recirculation and further addition of PPO ongoing
 - Quantifying and assessing backgrounds source
- Have calibrated the detector based on data from water phase
 - Calibration of scintillator properties now underway
- Physics program is being pursued
 - $0\nu\beta\beta$ measurement planned
 - Solar and Reactor neutrino measurements underway
 - Analysis of Partial fill data in preparation
 - Other programs also ongoing (Supernova, Geoneutrinos)